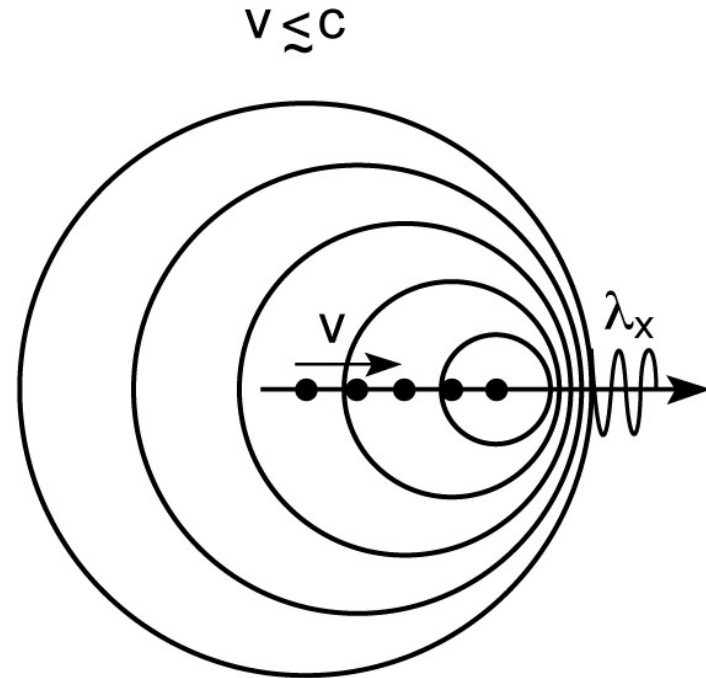
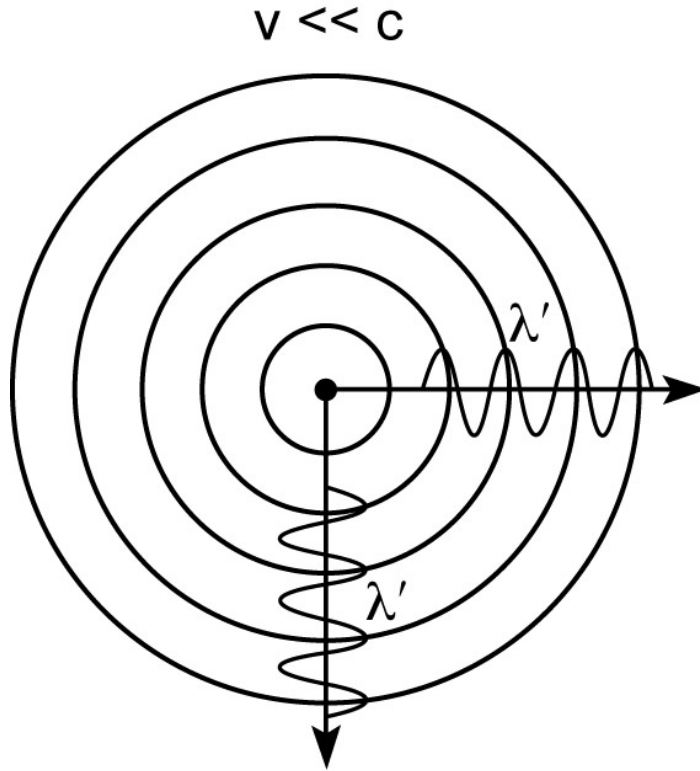




# Coherence, Brightness and Time Structure: from Synchrotrons to FELs

David Attwood  
University of California, Berkeley

# Synchrotron radiation from relativistic electrons



Note: Angle-dependent doppler shift

$$\lambda = \lambda' \left(1 - \frac{v}{c} \cos\theta\right)$$

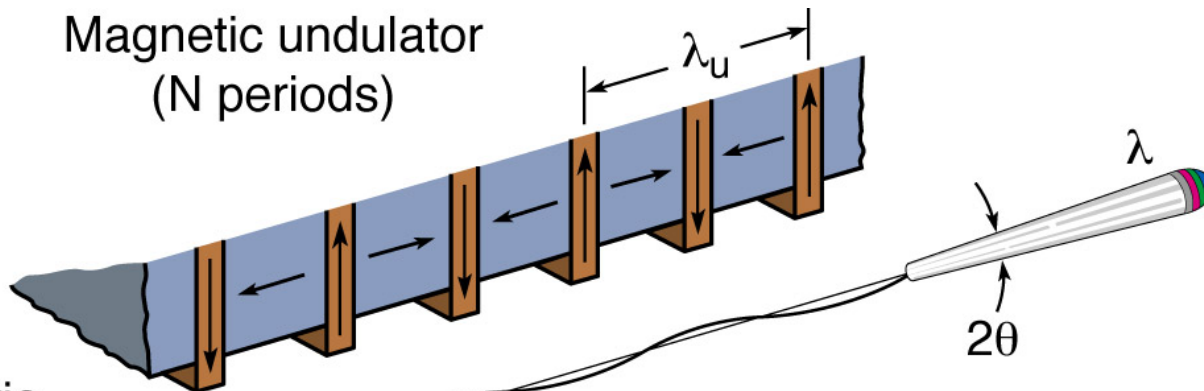
$$\lambda = \lambda' \gamma \left(1 - \frac{v}{c} \cos\theta\right)$$

$$\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}$$

# Undulator radiation from a small electron beam radiating into a narrow forward cone, is very bright



Magnetic undulator  
(N periods)

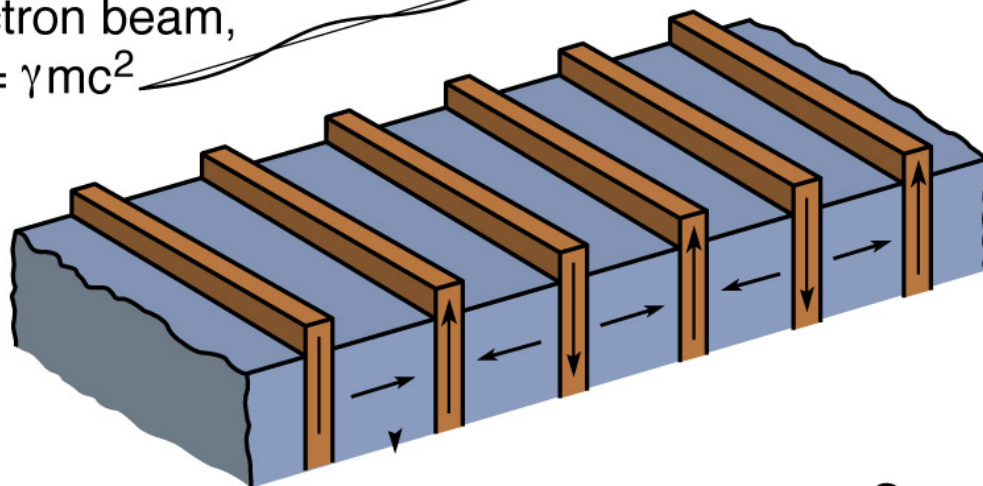


$$\lambda \approx \frac{\lambda_u}{2\gamma^2}$$

$$\theta_{\text{cen}} \approx \frac{1}{\gamma \sqrt{N}}$$

Relativistic  
electron beam,  
 $E_e = \gamma mc^2$

$$\left[ \frac{\Delta\lambda}{\lambda} \right]_{\text{cen}} = \frac{1}{N}$$



$$\text{Brightness} = \frac{\text{photon flux}}{(\Delta A) (\Delta\Omega)}$$

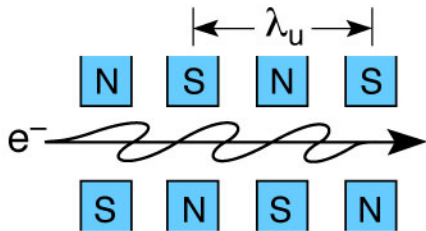
$$\text{Spectral Brightness} = \frac{\text{photon flux}}{(\Delta A) (\Delta\Omega) (\Delta\lambda/\lambda)}$$

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# Undulator radiation



## Laboratory Frame of Reference

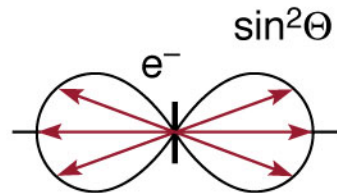


$$E = \gamma mc^2$$

$$\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}$$

$N = \#$  periods

## Frame of Moving $e^-$



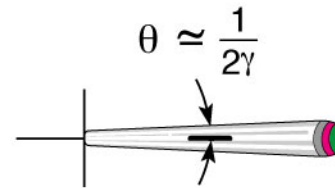
$e^-$  radiates at the Lorentz contracted wavelength:

$$\lambda' = \frac{\lambda_u}{\gamma}$$

Bandwidth:

$$\frac{\lambda'}{\Delta\lambda'} \approx N$$

## Frame of Observer



Doppler shortened wavelength on axis:

$$\lambda = \lambda' \gamma (1 - \beta \cos \theta)$$

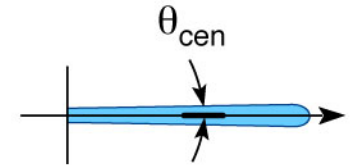
$$\lambda = \frac{\lambda_u}{2\gamma^2} (1 + \gamma^2 \theta^2)$$

Accounting for transverse motion due to the periodic magnetic field:

$$\lambda = \frac{\lambda_u}{2\gamma^2} \left( 1 + \frac{K^2}{2} + \gamma^2 \theta^2 \right)$$

where  $K = eB_0\lambda_u / 2\pi mc$

## Following Monochromator



$$\text{For } \frac{\Delta\lambda}{\lambda} \approx \frac{1}{N}$$

$$\theta_{\text{cen}} \approx \frac{1}{\gamma \sqrt{N}}$$

typically

$$\theta_{\text{cen}} \approx 40 \text{ rad}$$

# Determining the power radiated: the equation of motion of an electron in an undulator

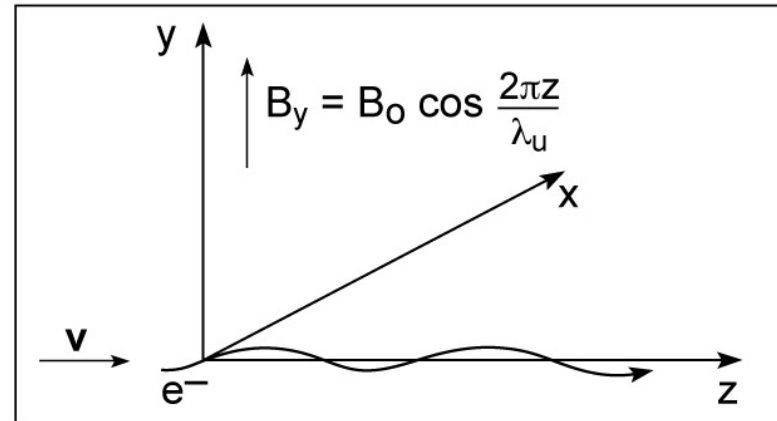


Magnetic fields in the periodic undulator cause the electrons to oscillate and thus radiate. These magnetic fields also slow the electrons axial (z) velocity somewhat, reducing both the Lorentz contraction and the Doppler shift, so that the observed radiation wavelength is not quite so short. The force equation for an electron is

$$\frac{d\mathbf{p}}{dt} = -e(\mathbf{E} + \mathbf{v} \times \mathbf{B}) \quad (5.16)$$

where  $\mathbf{p} = \gamma m \mathbf{v}$  is the momentum. The radiated fields are relatively weak so that

$$\frac{d\mathbf{p}}{dt} \simeq -e(\mathbf{v} \times \mathbf{B})$$



Taking to first order  $v \simeq v_z$ , motion in the x-direction is

$$m\gamma \frac{dv_x}{dt} = +ev_z B_y$$

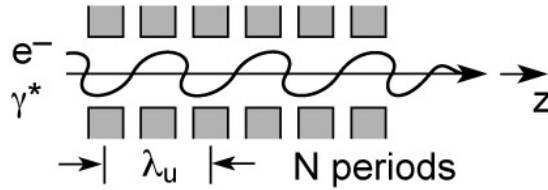
$$v_x = \frac{Kc}{\gamma} \sin\left(\frac{2\pi z}{\lambda_u}\right) \quad (5.19)$$

$$K \equiv \frac{eB_0\lambda_u}{2\pi mc} = 0.9337 B_0(\text{T})\lambda_u(\text{cm}) \quad (5.18)$$

# Calculating power in the central radiation cone: using the well known “dipole radiation” formula by transforming to the frame of reference moving with the electrons

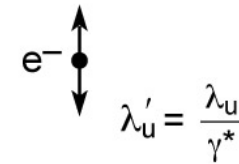


$x, z, t$  laboratory frame of reference



$$\frac{d\mathbf{p}}{dt} = -e(\mathbf{E} + \mathbf{v} \times \mathbf{B})$$

$x', z', t'$  frame of reference moving with the average velocity of the electron



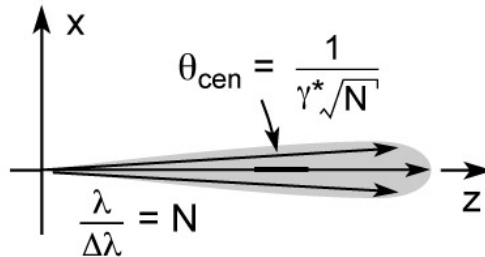
$x', z', t'$  motion  
 $a'(t')$  acceleration

Dipole radiation:

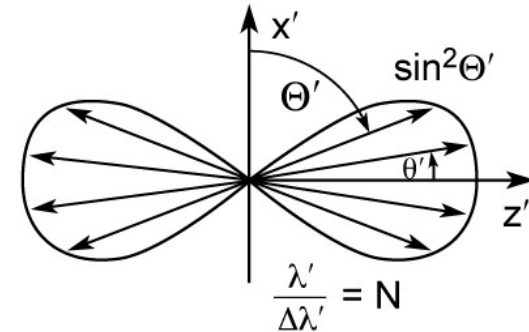
$$\frac{dP'}{d\Omega'} = \frac{e^2 a'^2 \sin^2 \Theta'}{16\pi^2 \epsilon_0 c^3}$$

Lorentz transformation

Lorentz transformation



$$\bar{P}_{\text{cen}} = \frac{\pi e \gamma^2 I}{\epsilon_0 \lambda_u} \frac{K^2}{(1 + K^2/2)^2}$$



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# Undulator radiated power in the central cone



$$\lambda_x = \frac{\lambda_u}{2\gamma^2} \left( 1 + \frac{K^2}{2} + \gamma^2 \theta^2 \right)$$

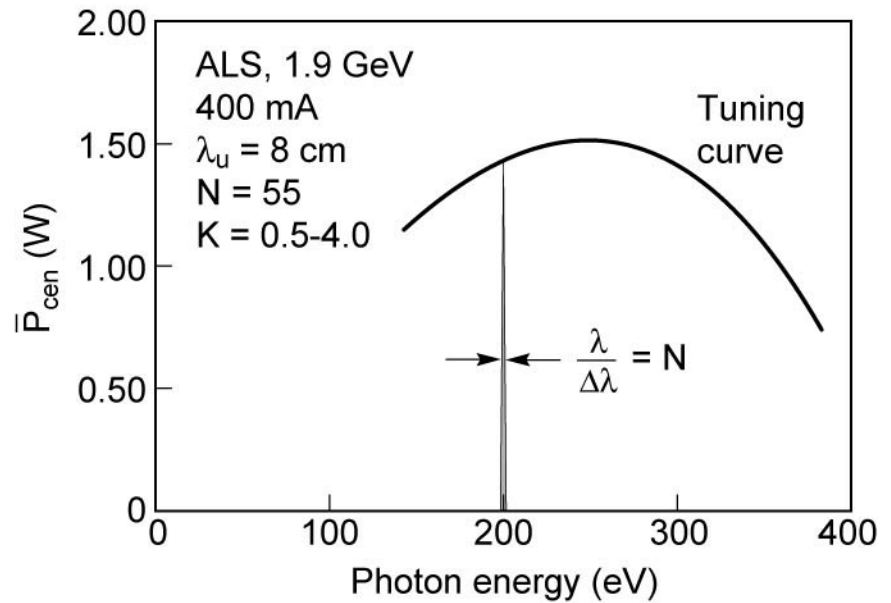
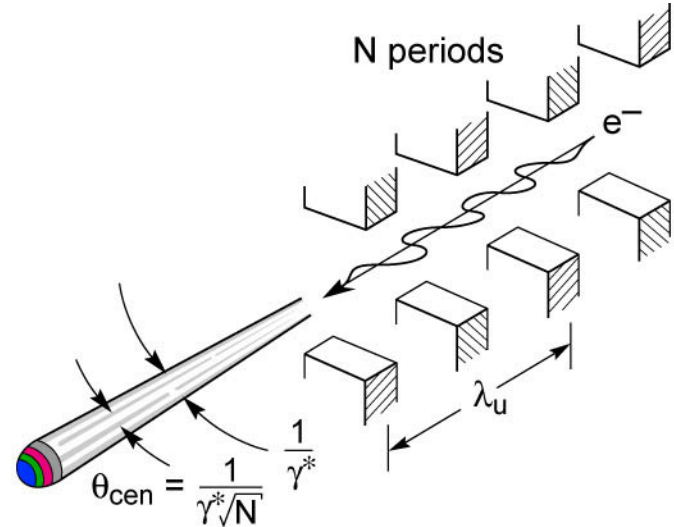
$$\bar{P}_{\text{cen}} = \frac{\pi e \gamma^2 I}{\epsilon_0 \lambda_u} \frac{K^2}{\left( 1 + \frac{K^2}{2} \right)^2} f(K)$$

$$\theta_{\text{cen}} = \frac{1}{\gamma^* \sqrt{N}}$$

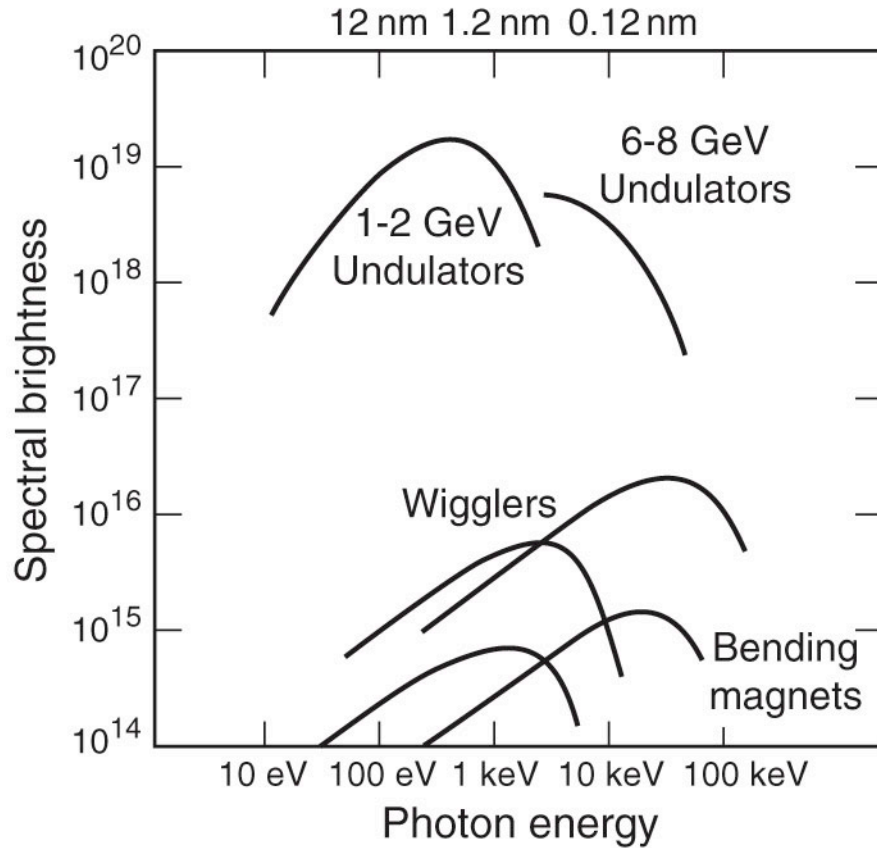
$$\left( \frac{\Delta \lambda}{\lambda} \right)_{\text{cen}} = \frac{1}{N}$$

$$K = \frac{e B_0 \lambda_u}{2 \pi m_0 c}$$

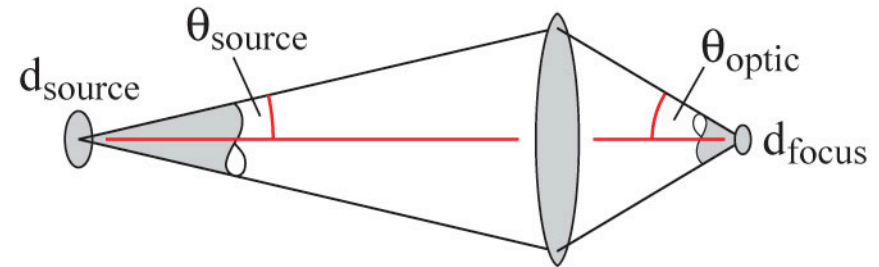
$$\gamma^* = \gamma / \sqrt{1 + \frac{K^2}{2}}$$



# Spectral brightness is useful for experiments that involve spatially resolved studies



- Brightness is conserved (in lossless optical systems)



$$d_{\text{source}} \cdot \theta_{\text{source}} = d_{\text{focus}} \cdot \theta_{\text{optic}}$$

↑ Smaller after focus      ↑ Large in a focusing optic

- Starting with many photons in a small source area and solid angle, permits high photon flux in an even smaller area

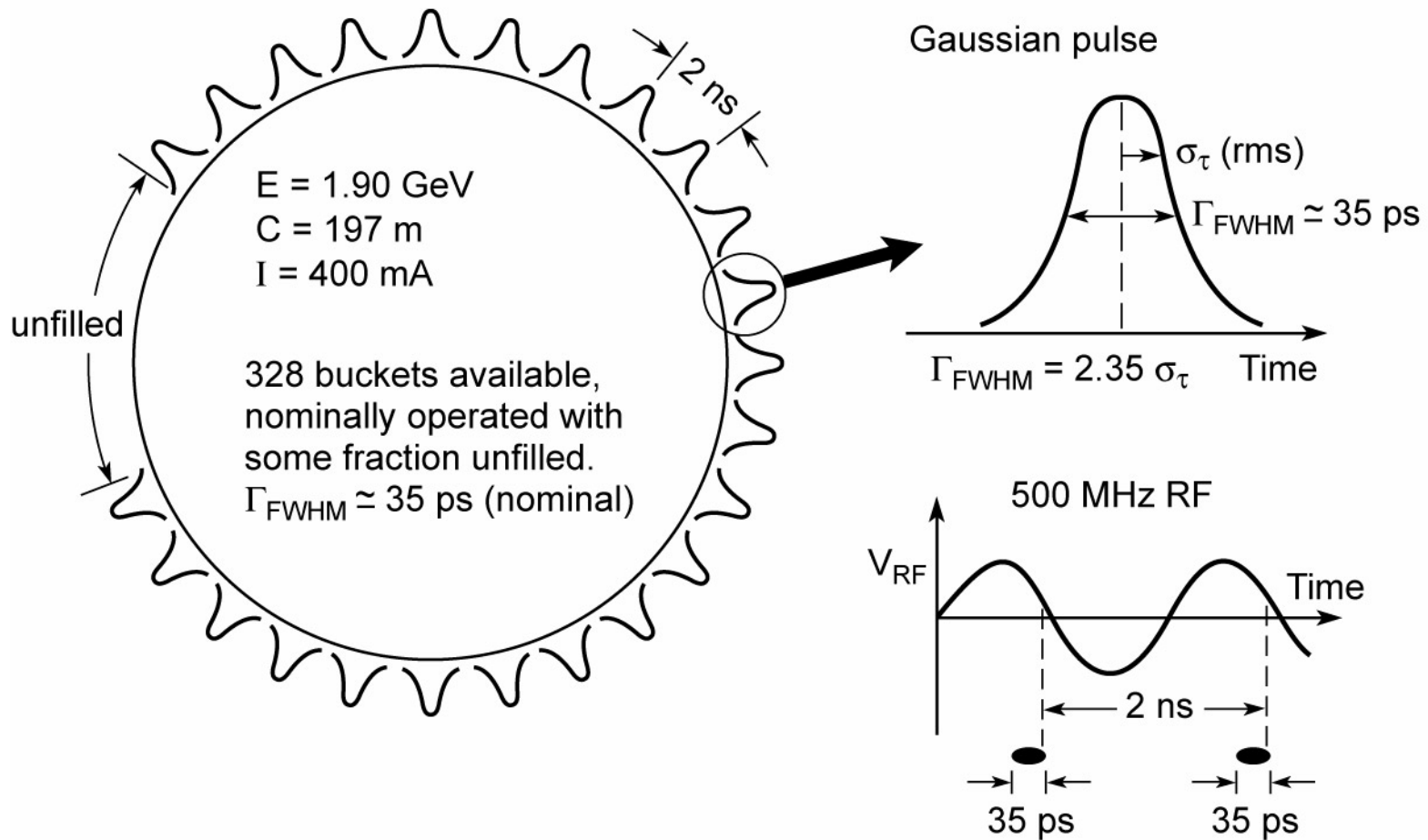
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# Time structure of synchrotron radiation



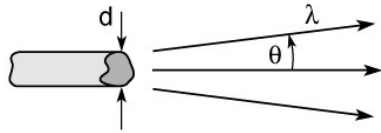
The axial electric field within the RF cavity, used to replenish lost (radiated) energy, forms a potential well “bucket” system that forces electrons into axial electron “bunches”. This leads to a time structure in the emitted radiation.



# Coherence at short wavelengths



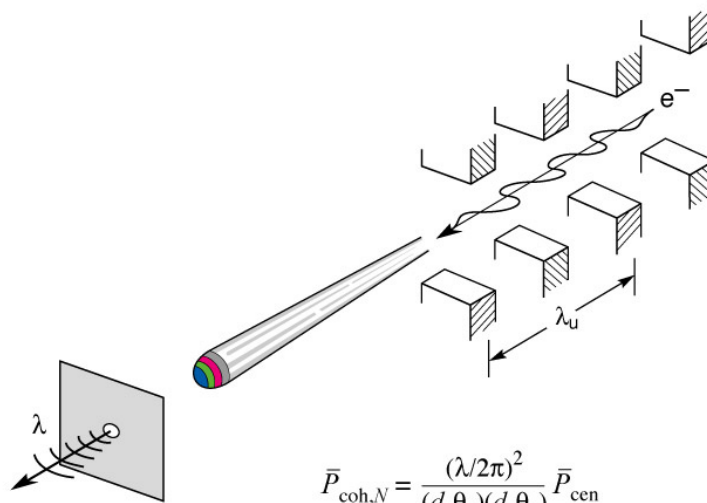
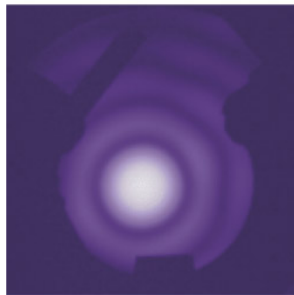
## Chapter 8



$$l_{\text{coh}} = \lambda^2 / 2\Delta\lambda \quad \{\text{temporal (longitudinal) coherence}\} \quad (8.3)$$

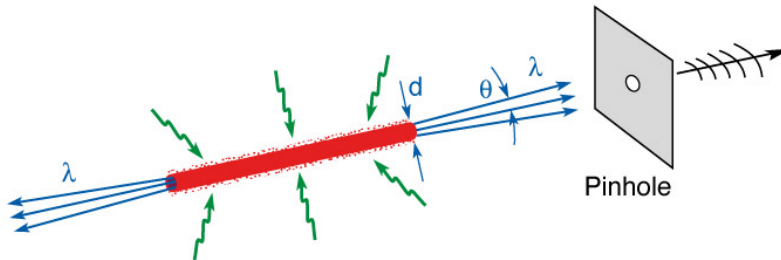
$$d \cdot \theta = \lambda / 2\pi \quad \{\text{spatial (transverse) coherence}\} \quad (8.5)$$

$$\text{or } d \cdot 2\theta|_{\text{FWHM}} = 0.44 \lambda \quad (8.5^*)$$



$$\bar{P}_{\text{coh},N} = \frac{(\lambda/2\pi)^2}{(d_x\theta_x)(d_y\theta_y)} \bar{P}_{\text{cen}} \quad (8.6)$$

$$\bar{P}_{\text{coh},\lambda\Delta\lambda} = \frac{e\lambda_u I \eta (\Delta\lambda/\lambda) N^2}{8\pi\epsilon_0 d_x d_y} \cdot \left[ 1 - \frac{\hbar\omega}{\hbar\omega_0} \right] f(K) \quad (8.9)$$



$$P_{\text{coh}} = \frac{(\lambda/2\pi)^2}{(d_x\theta_x)(d_y\theta_y)} P_{\text{laser}} \quad (8.11)$$

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**co·here** \kō-'hi(ə)r\ *vi* [L *cohaerēre*, fr. *co-* + *haerēre* to stick] **1 a** : to hold together firmly as parts of the same mass **b** : ADHERE **c** : to display cohesion **2** : to consist of parts that cohere **3 a** : to become united in principles, relationships, or interests **b** : to be logically or aesthetically consistent **syn** see STICK  
**co·her·ence** \kō-'hīr-ən(t)s, -'hēr-\ *n* : the quality or state of cohering; *esp* : systematic connection esp. in logical discourse  
**co·her·en·cy** \-ən-sē\ *n* : COHERENCE  
**co·her·ent** \kō-'hīr-ənt, -'hēr-\ *adj* [MF or L; MF *cohérent*, fr. L *cohaerent-*, *cohaerens*, prp. of *cohaerēre*] **1** : having the quality of cohering **2** : logically consistent **3** : having waves in phase and of one wavelength <~ light> — **co·her·ent·ly** *adv*

*Webster's 7th Collegiate Dictionary  
(1971)*

**co·her·ence** (kō-hīr'əns, -hēr'-) also **co·her·en·cy** (-ən-sē) *n*. The quality or state of cohering, esp. logical or orderly relationship of parts.  
**co·her·ent** (kō-hīr'ənt, -hēr'-) *adj*. **1**. Sticking together; cohering. **2**. Marked by an orderly or logical relation of parts that affords comprehension or recognition: coherent speech. **3**. *Physics*. Of or pertaining to waves with a continuous relationship among phases. **4**. Of or pertaining to a system of units of measurement in which a small number of basic

*American Heritage 2nd College Edition  
Dictionary (1985)*

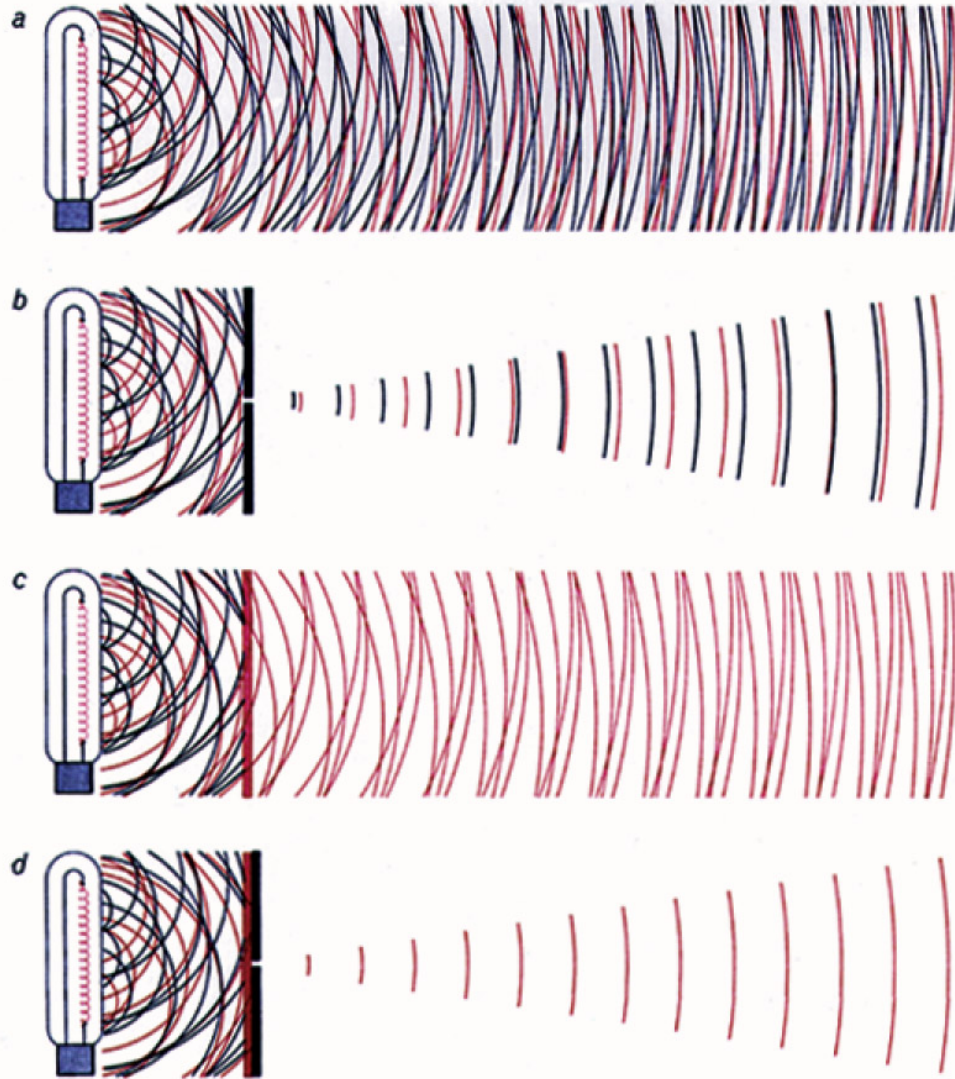
- Coherence (physics), an ideal property of waves that enables stationary (i.e. temporally and spatially constant) interference
- Coherence time, the time over which a propagating wave (especially a laser or maser beam) may be considered coherent; the time interval within which its phase is, on average, predictable

*Wikipedia,  
the free encyclopedia*

# Marching band and coherence lengths



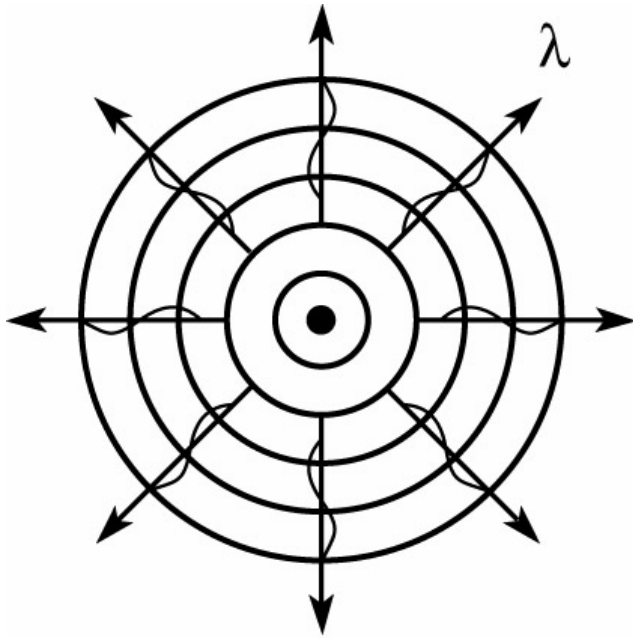
# Spatial and spectral filtering to produce coherent radiation



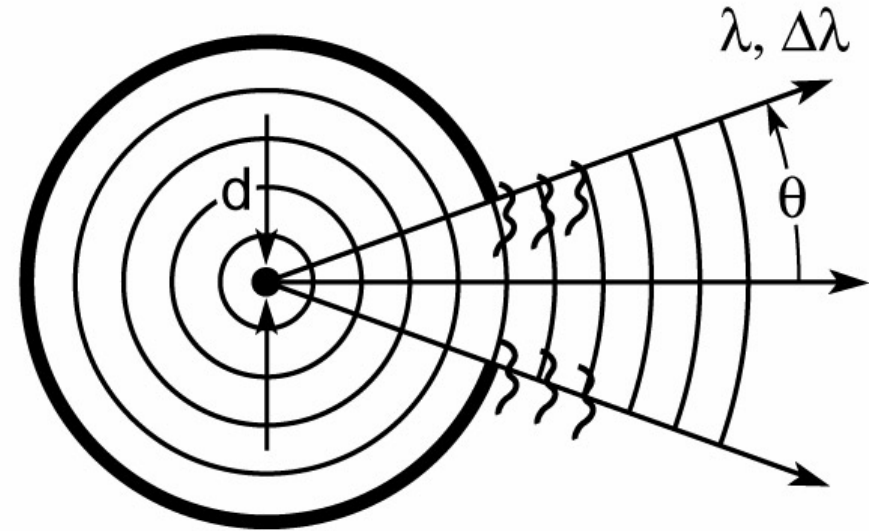
Courtesy of A. Schawlow, Stanford.

Ch08\_F08.ai

# Coherence, partial coherence and incoherence



Point source oscillator  
 $-\infty < t < \infty$



Source of finite size,  
divergence, and duration

Ch08\_F01.ai

# Spatial and temporal coherence



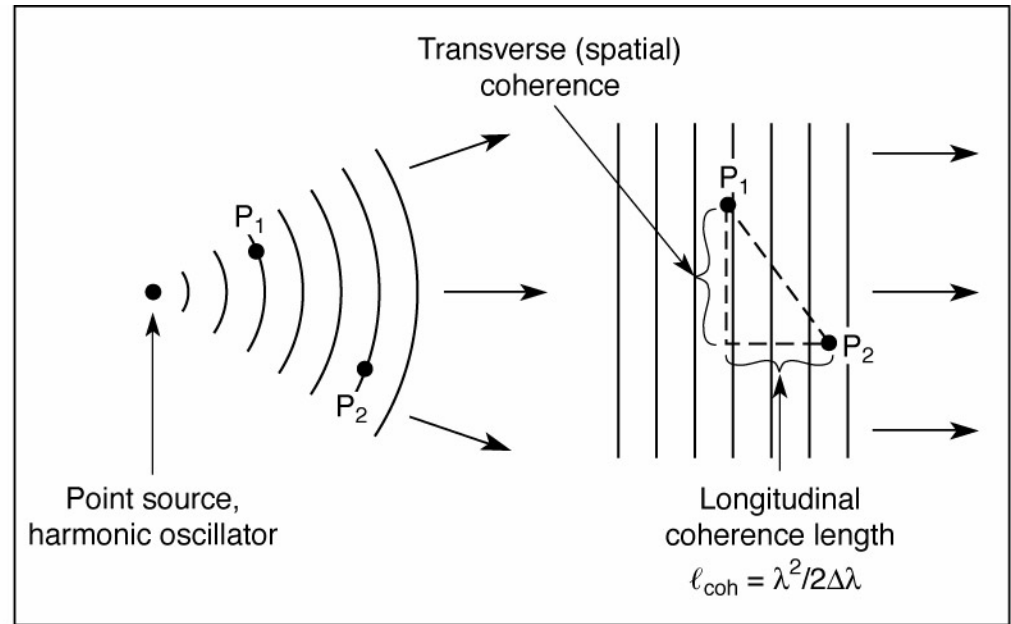
Mutual coherence factor

$$\Gamma_{12}(\tau) \equiv \langle E_1(t + \tau)E_2^*(t) \rangle \quad (8.1)$$

Normalize degree of spatial coherence  
(complex coherence factor)

$$\mu_{12} = \frac{\langle E_1(t)E_2^*(t) \rangle}{\sqrt{\langle |E_1|^2 \rangle} \sqrt{\langle |E_2|^2 \rangle}} \quad (8.12)$$

A high degree of coherence ( $\mu \rightarrow 1$ ) implies an ability to form a high contrast interference (fringe) pattern. A low degree of coherence ( $\mu \rightarrow 0$ ) implies an absence of interference, except with great care. In general radiation is partially coherent.



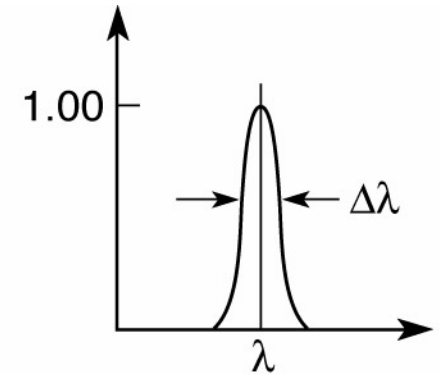
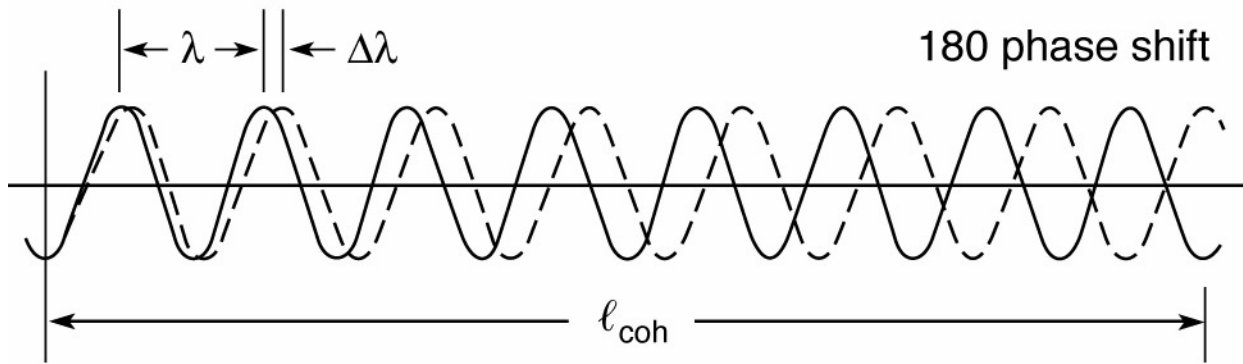
Longitudinal (temporal) coherence length

$$\ell_{\text{coh}} = \frac{\lambda^2}{2 \Delta\lambda} \quad (8.3)$$

Full spatial (transverse) coherence

$$d \cdot \theta = \lambda / 2\pi \quad (8.5)$$

# Spectral bandwidth and longitudinal coherence length



Define a coherence length  $\ell_{\text{coh}}$  as the distance of propagation over which radiation of spectral width  $\Delta\lambda$  becomes  $180^\circ$  out of phase. For a wavelength  $\lambda$  propagating through  $N$  cycles

$$\ell_{\text{coh}} = N\lambda$$

and for a wavelength  $\lambda + \Delta\lambda$ , a half cycle less ( $N - \frac{1}{2}$ )

$$\ell_{\text{coh}} = (N - \frac{1}{2})(\lambda + \Delta\lambda)$$

Equating the two

$$N = \lambda / 2\Delta\lambda$$

so that

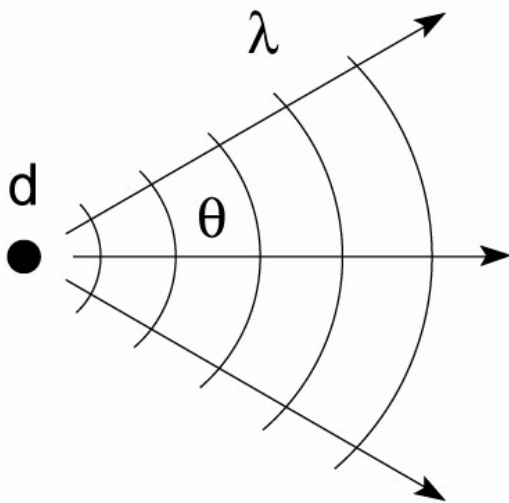
$$\boxed{\ell_{\text{coh}} = \frac{\lambda^2}{2 \Delta\lambda}} \quad (8.3)$$



# A practical interpretation of spatial coherence



- Associate spatial coherence with a spherical wavefront.
- A spherical wavefront implies a point source.
- How small is a “point source”?



From Heisenberg’s Uncertainty Principle ( $\Delta x \cdot \Delta p \geq \frac{\hbar}{2}$ ), the smallest source size “d” you can resolve, with wavelength  $\lambda$  and half angle  $\theta$ , is

$$d \cdot \theta = \frac{\lambda}{2\pi}$$

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# Partially coherent radiation approaches uncertainty principle limits



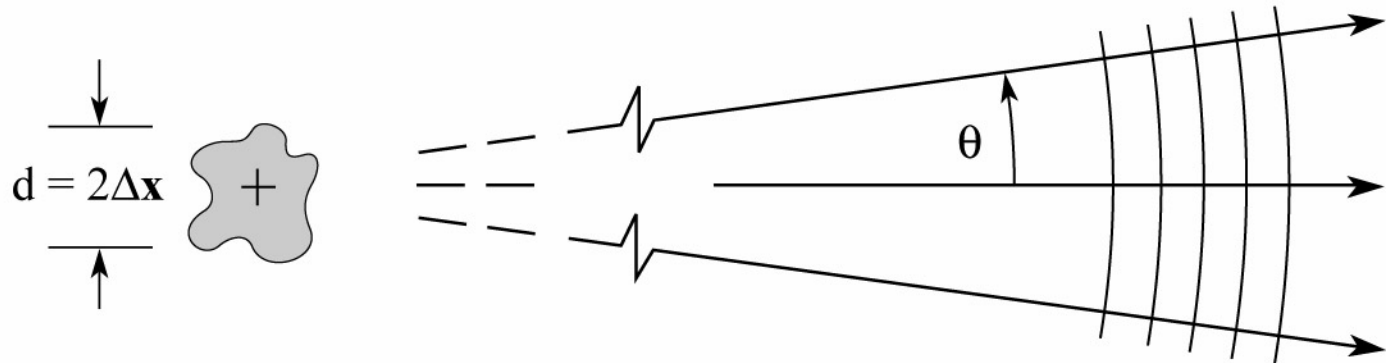
$$\Delta \mathbf{x} \cdot \Delta \mathbf{p} \geq \hbar/2 \quad (8.4)$$

$$\Delta \mathbf{x} \cdot \hbar \Delta \mathbf{k} \geq \hbar/2$$

$$\Delta \mathbf{x} \cdot \mathbf{k} \Delta \theta \geq 1/2$$

$$2\Delta \mathbf{x} \cdot \Delta \theta \geq \lambda/2\pi$$

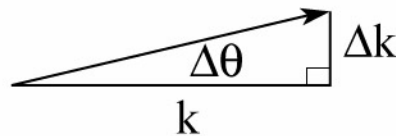
Standard deviations of Gaussian distributed functions  
(Tipler, 1978, pp. 174-189)



Note:

$$\Delta \mathbf{p} = \hbar \Delta \mathbf{k}$$

$$\Delta k = k \Delta \theta$$



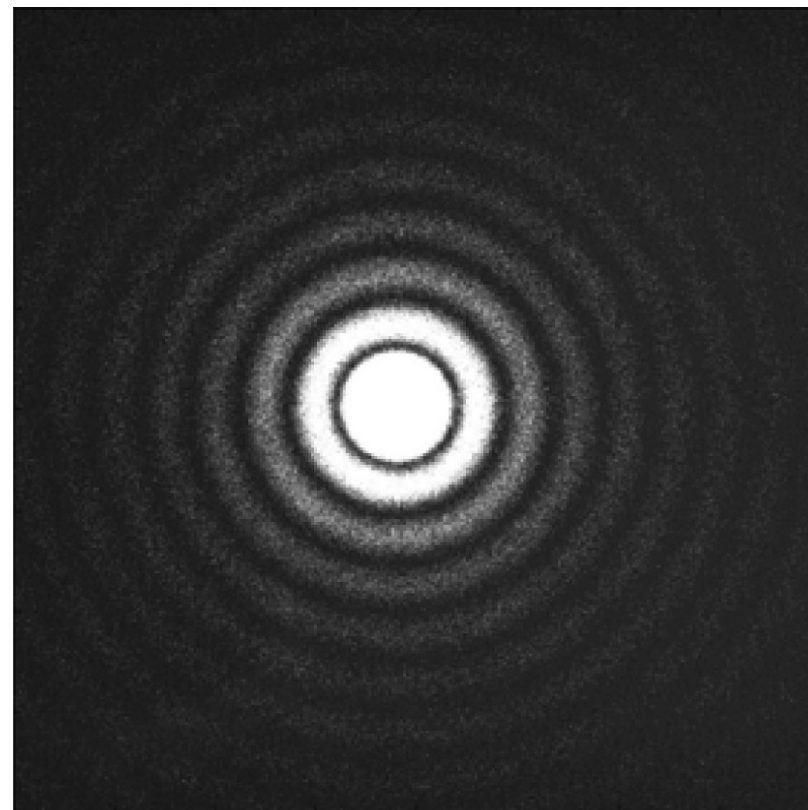
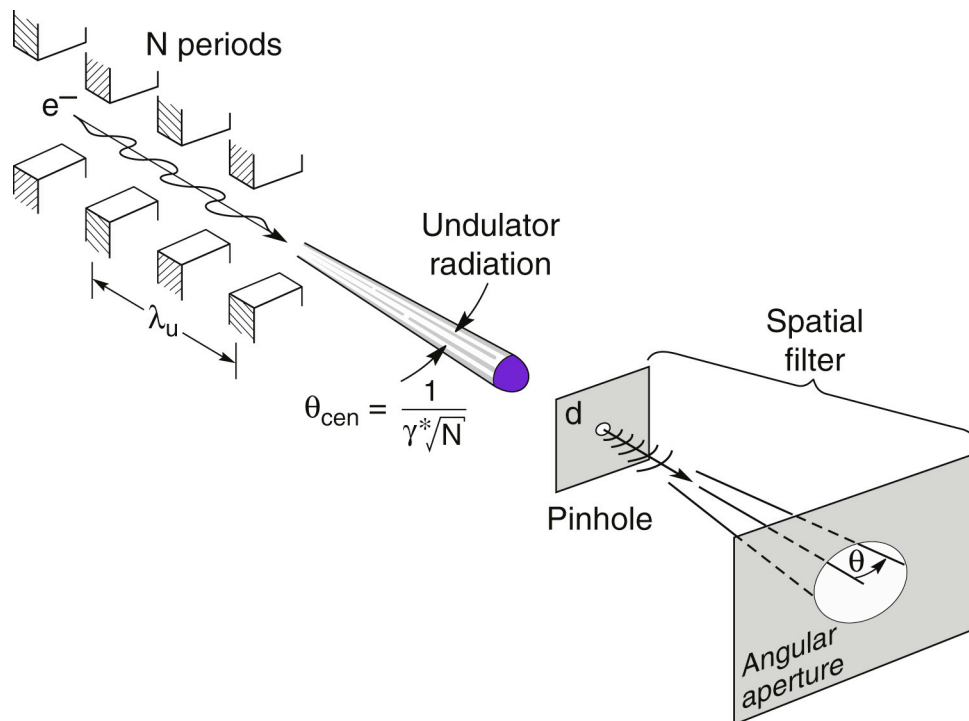
Spherical wavefronts occur  
in the limiting case

$$\left. \begin{array}{l} \boxed{d \cdot \theta = \lambda/2\pi} \\ \text{(spatially coherent)} \end{array} \right\} \frac{1}{\sqrt{e}} \text{ quantities}$$

or

$$(d \cdot 2\theta)_{\text{FWHM}} \approx \lambda/2 \left. \vphantom{(d \cdot 2\theta)_{\text{FWHM}}} \right\} \text{FWHM quantities}$$

$\lambda = 2.48 \text{ nm}$  (600 eV)  
 $d = 2.5 \mu\text{m}$   
 $t = 200 \text{ msec}$   
 ALS beamline 12.0.2  
 $\lambda_u = 80 \text{ mm}$ ,  $N = 55$ ,  $n = 3$   
 25 mm wide CCD at 410 mm



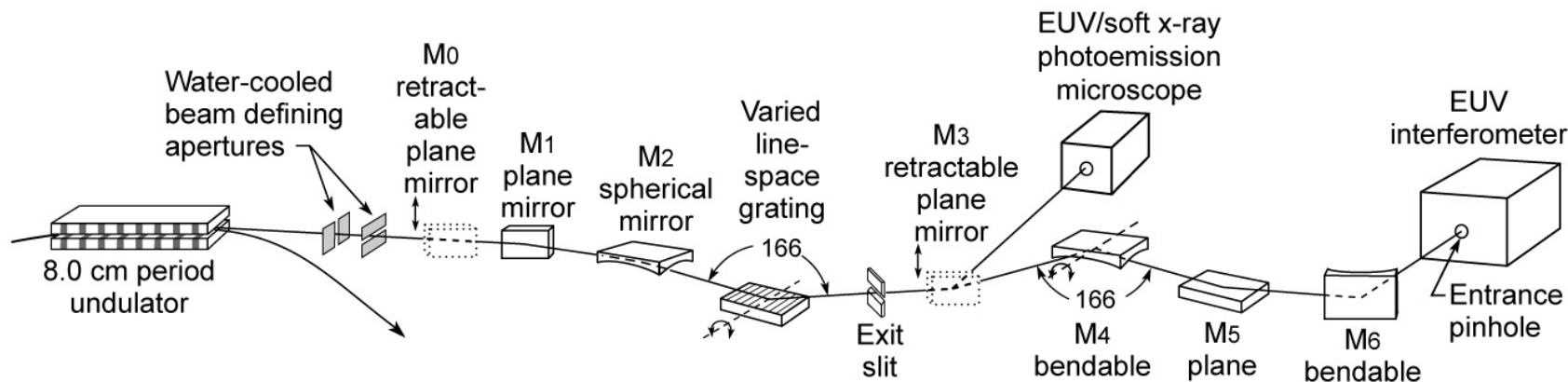
AiryPattern600eV\_Feb2013.ai

$$d \cdot \theta = \frac{\lambda}{2\pi}$$

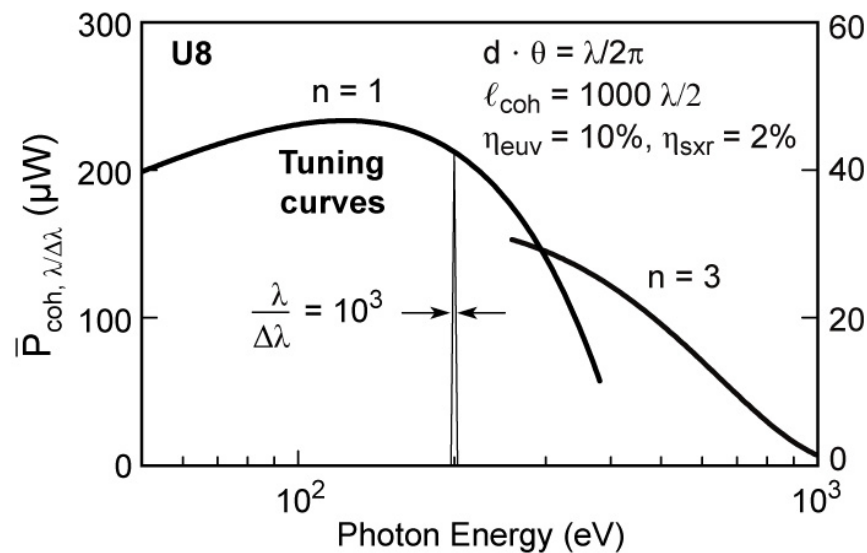
Courtesy of Patrick Naulleau, LBNL / Kris Rosfjord, UCB and LBNL

# Spatial and spectral filtering of undulator radiation

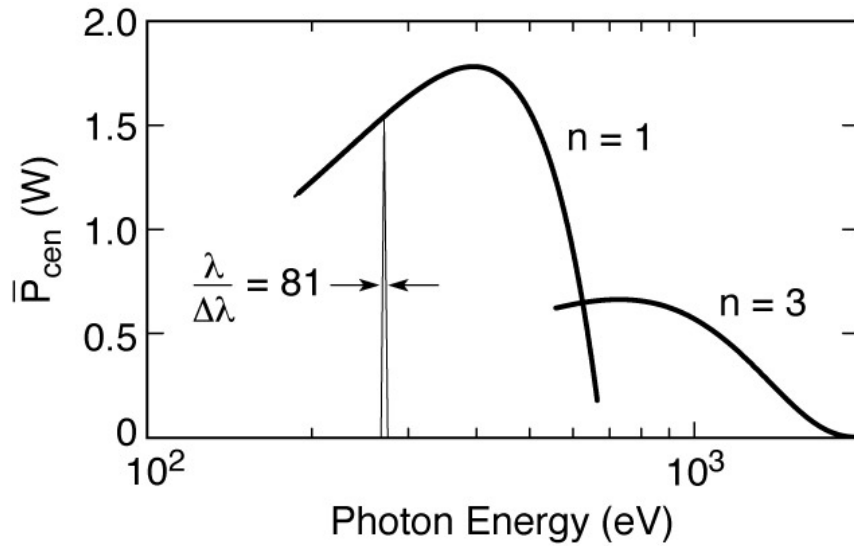
- Pinhole filtering for full spatial coherence
- Monochromator for spectral filtering to  $\lambda/\Delta\lambda > N$



$$\bar{P}_{\text{coh}, \lambda/\Delta\lambda} = \underbrace{\eta}_{\text{beamline efficiency}} \underbrace{\frac{(\lambda/2\pi)^2}{(d_x\theta_x)(d_y\theta_y)}}_{\text{spatial filtering}} \cdot \underbrace{N \frac{\Delta\lambda}{\lambda}}_{\text{spectral filtering}} \cdot \bar{P}_{\text{cen}} \quad (8.10a)$$



# Coherent power at Elettra



2.0 GeV, 300 mA

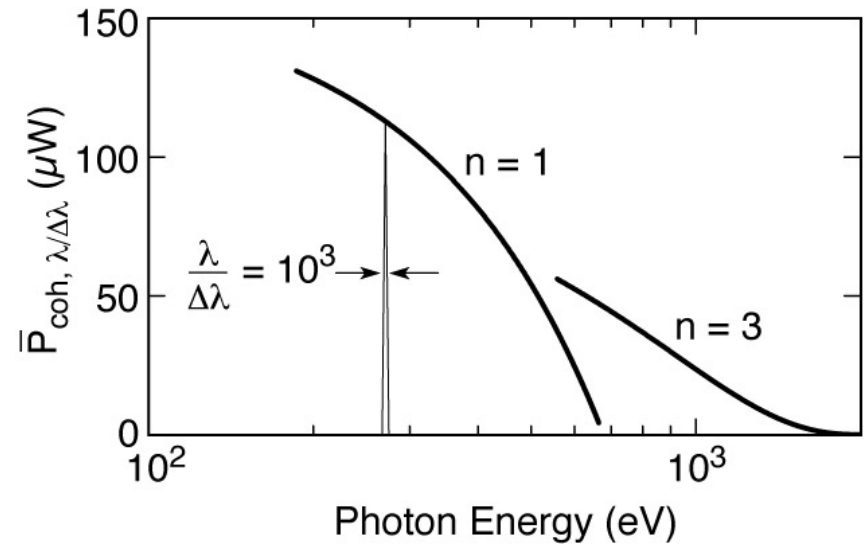
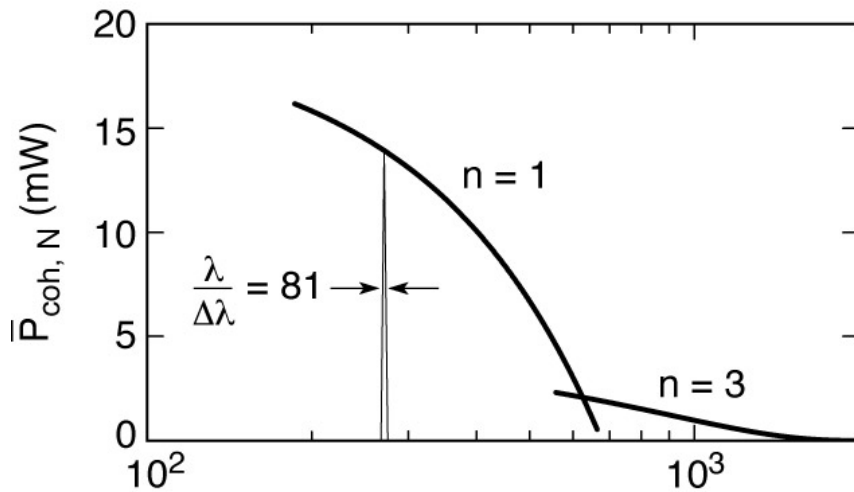
$\lambda_u = 56$  mm,  $N = 81$

$0.5 \leq K \leq 2.3$

$\sigma_x = 255 \mu\text{m}$ ,  $\sigma_x' = 23 \mu\text{r}$

$\sigma_y = 31 \mu\text{m}$ ,  $\sigma_y' = 9 \mu\text{r}$

$\eta = 10\%$



16 December 2004

International weekly journal of science

# nature

\$10.00

www.nature.com/nature

Inside this week



## X-ray holography

Lensless imaging at the nanoscale

### The 'Halloween storm'

How the Sun plays its tricks

### Protein transport

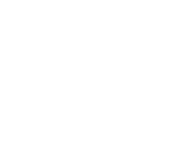
Escape from the nucleus

### Duck-billed platypus

Curiouser and curiouser

### Locusts over Africa

Time for biological control?





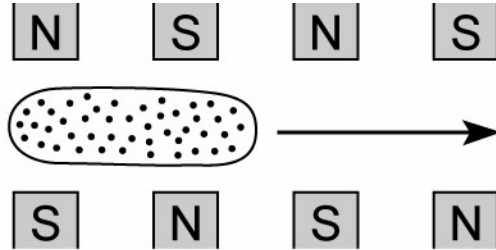
- Spatial coherence
- Temporal coherence
- Partial coherence
- Full coherence
- Spatial filtering
- Uncorrelated emitters
- Correlated emitters
- True phase coherence and mode control
- Lasers, amplified spontaneous emission (ASE) and mode control
- Undulator radiation
- SASE FEL **fsec and asec x-rays**
- Seeded FEL **true phase coherent x-rays**

# Marching band and coherence lengths



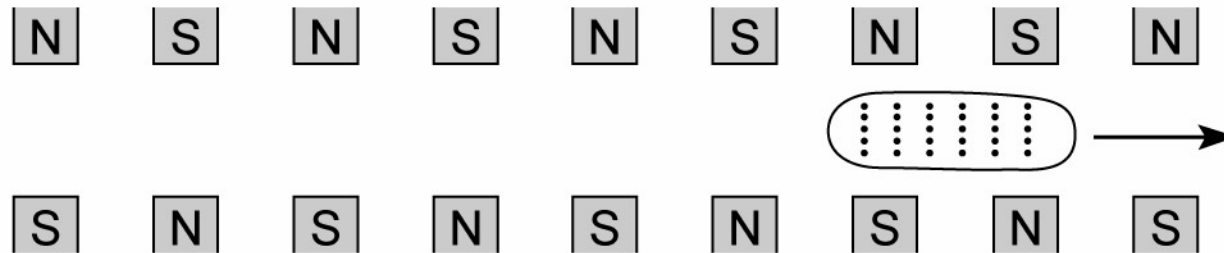


# Undulators and FELs



$$\frac{d\mathbf{p}}{dt} = -e(\mathbf{E} + \mathbf{v} \times \mathbf{B})$$

Undulator – uncorrelated electron positions, radiated fields uncorrelated, intensities add, limited coherence, power  $\sim N$ .

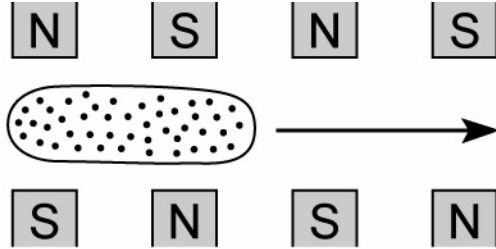


$$\frac{d\mathbf{p}}{dt} = -e(\mathbf{E} + \mathbf{v} \times \mathbf{B})$$

Free Electron Laser (FEL) – very long undulator, electrons are “microbunched” by their own radiated fields into strongly correlated waves of electrons, all radiated electric fields now add, spatially coherent, power  $\sim N^2$


“SASE” FEL – no seed (several separate “waves” of electrons possible with uncorrelated phase.)  
Less peak power, broader spectrum.

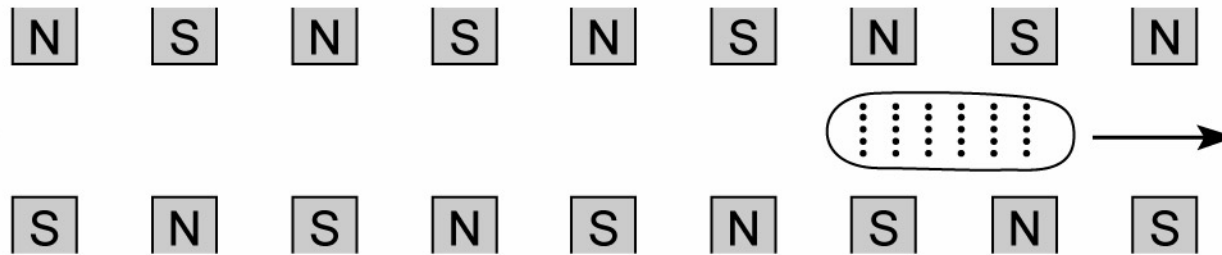
# Seeded FEL



$$\frac{d\mathbf{p}}{dt} = -e(\mathbf{E} + \mathbf{v} \times \mathbf{B})$$

Undulator – uncorrelated electron positions, radiated fields uncorrelated, intensities add, limited coherence, power  $\sim N$ .

Coherent seed pulse 



Free Electron Laser (FEL) – very long undulator, electrons are “microbunched” by their own radiated fields into strongly correlated waves of electrons, all radiated electric fields now add, spatially coherent, power  $\sim N^2$

$$\frac{d\mathbf{p}}{dt} = -e(\mathbf{E} + \mathbf{v} \times \mathbf{B})$$

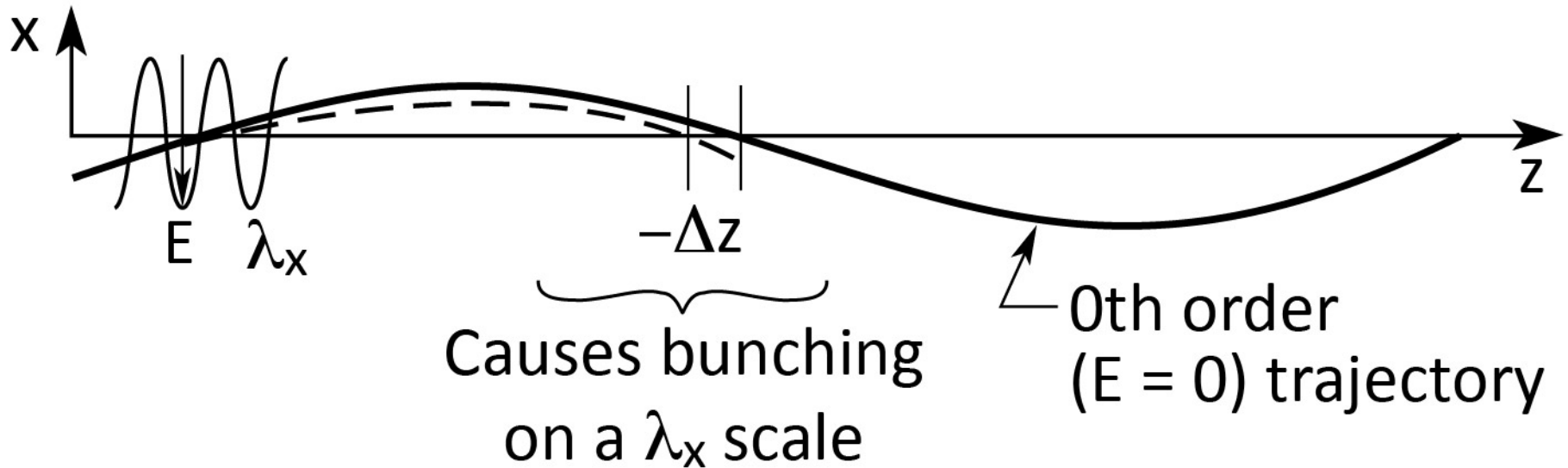
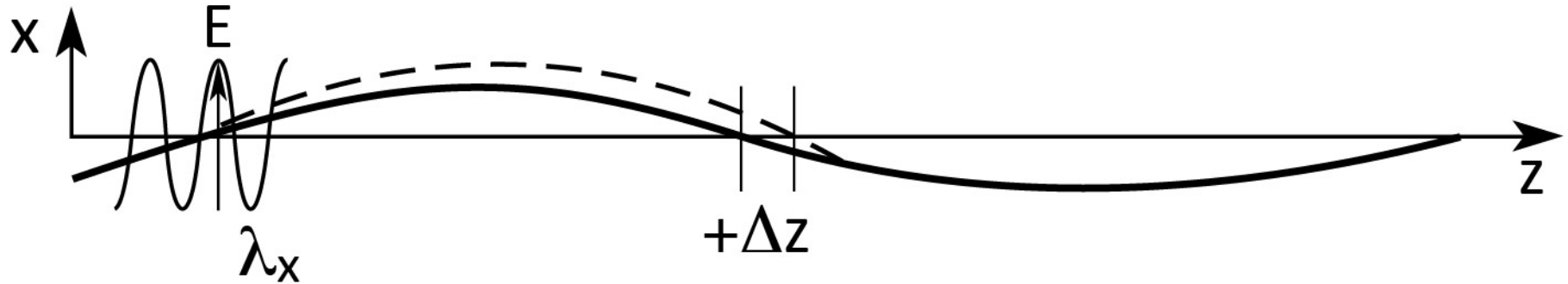
Second generation x-ray FELs.  
(Fermi in Trieste)

# The evolution of incoherent clapping (applauding) to coherent clapping



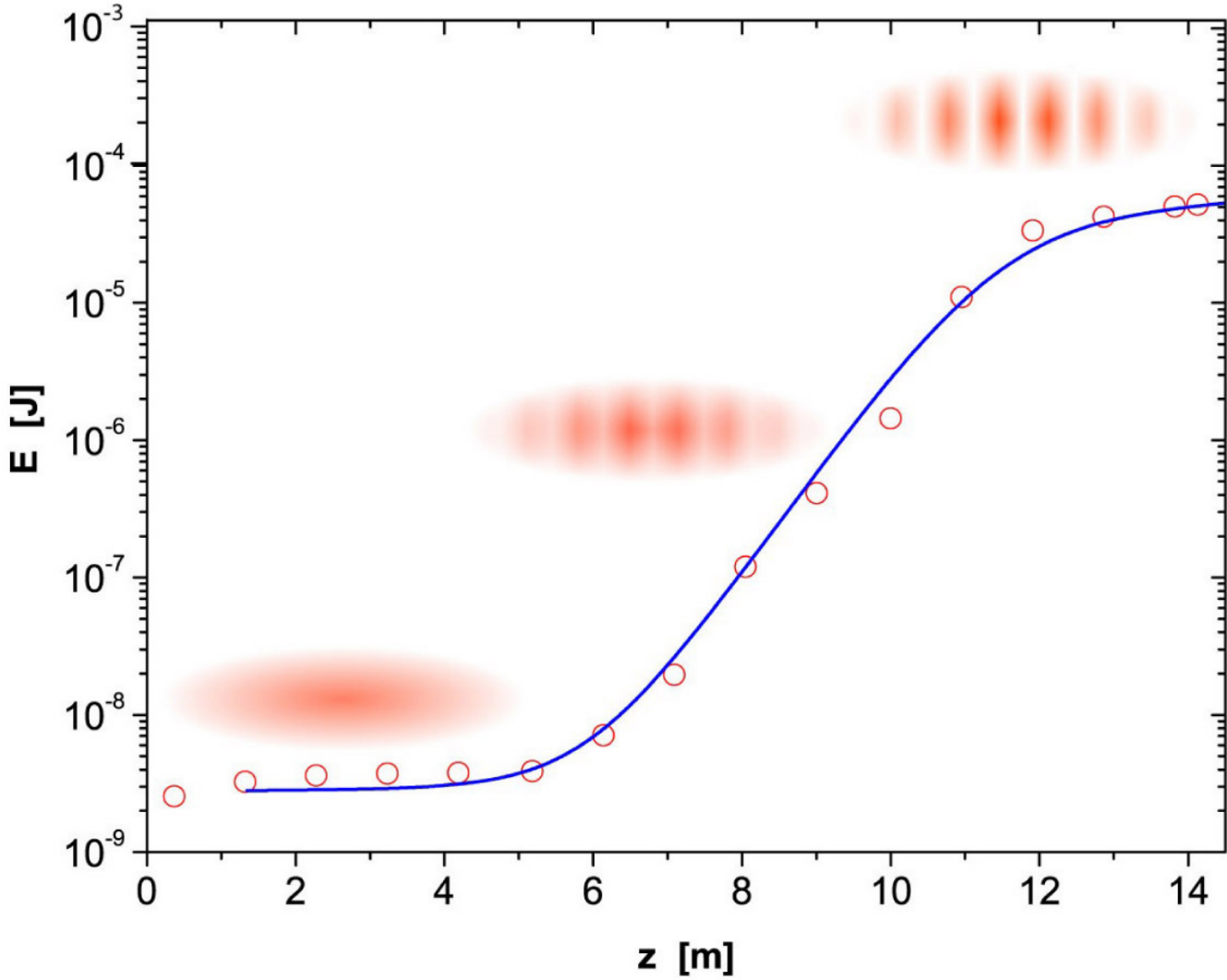
Suggested by Hideo Kitamura,  
(RIKEN)

# Electron energies and subsequent axis crossings are affected by the amplitude and relative phase of the co-propagating field



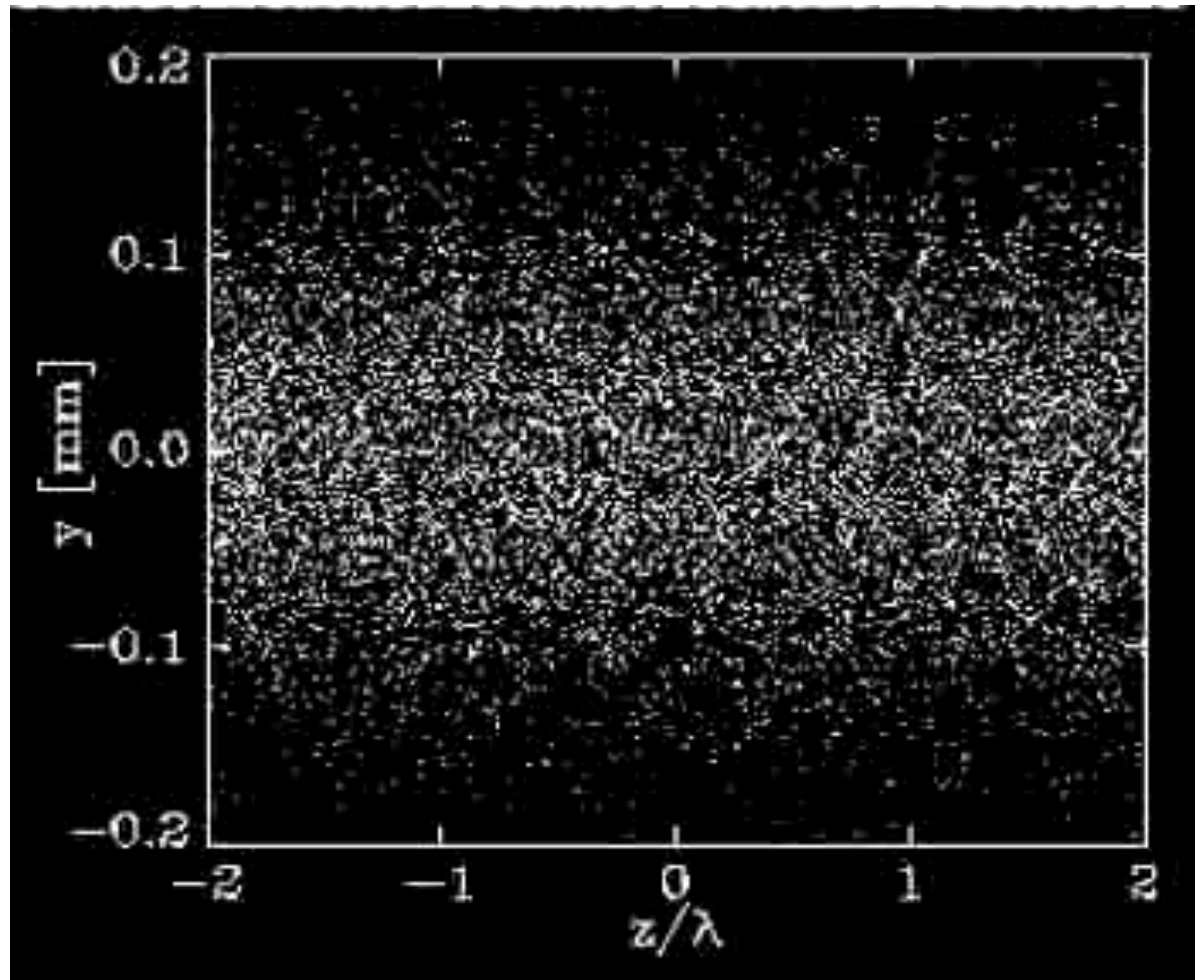
FEL\_Prebunching.ai

# Gain and saturation in an FEL



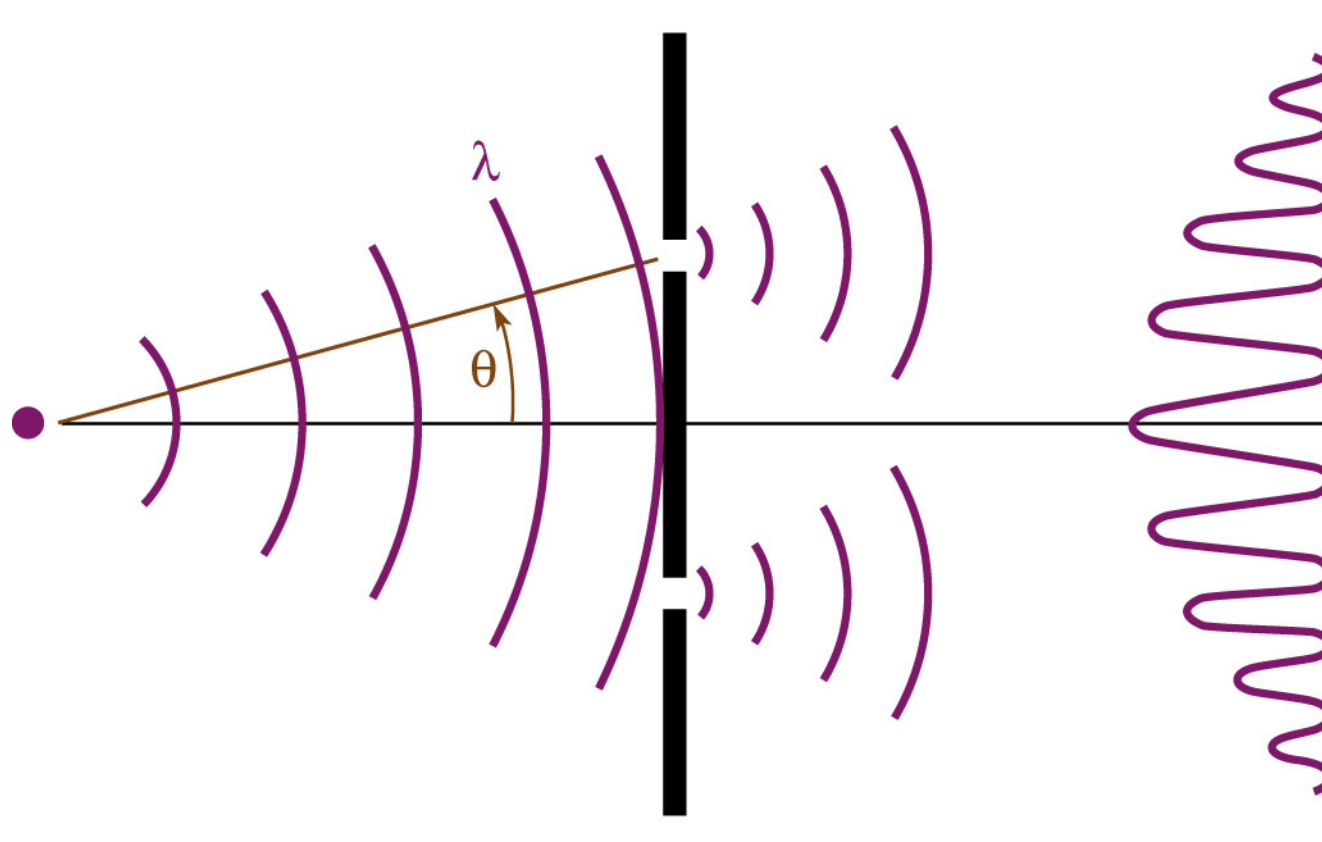
Courtesy of K-J. Kim

Gain\_Saturation\_FEL\_graph.ai



Courtesy of Sven Reiche, UCLA, now SLS

# Young's double slit experiment: spatial coherence and the persistence of fringes

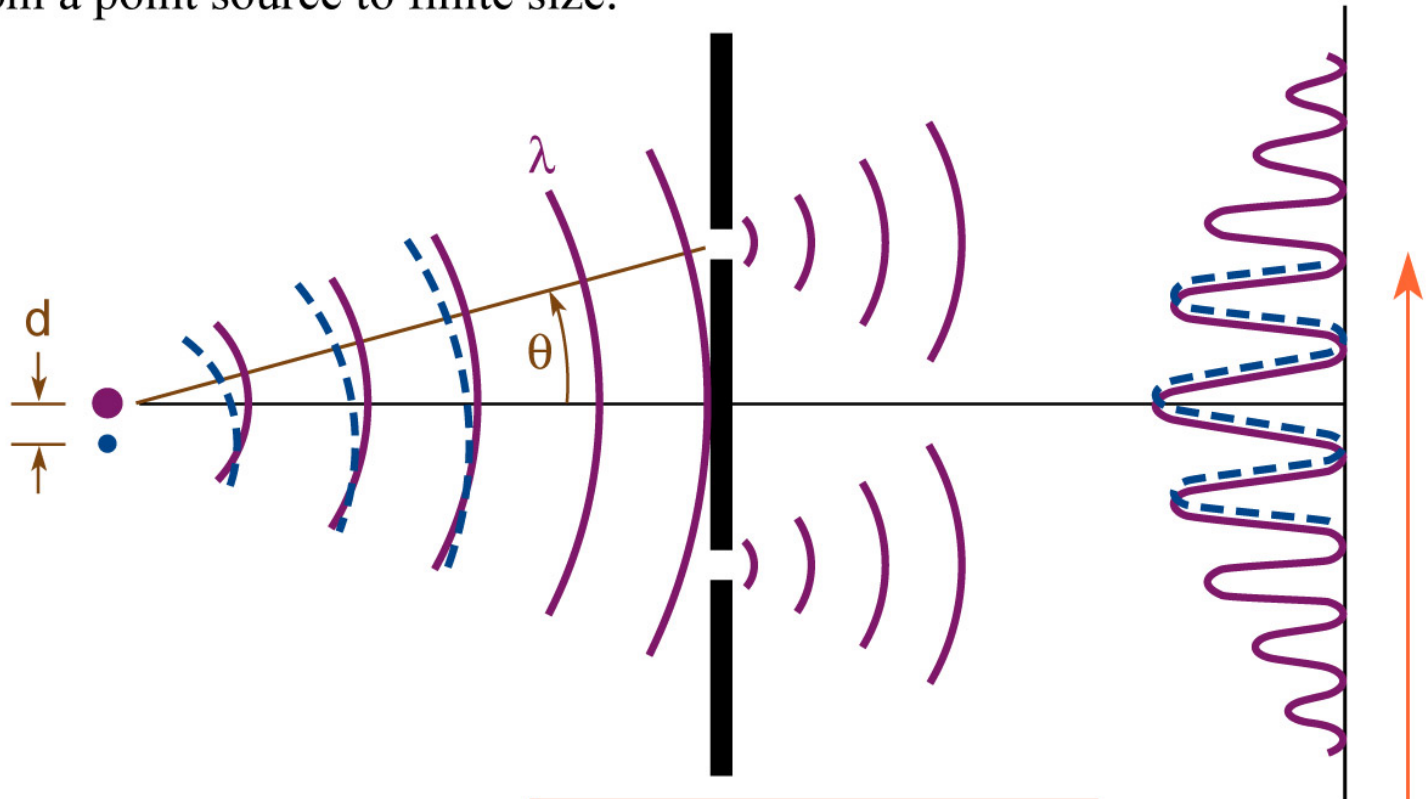


YoungsExprmt.ai

# Young's double slit experiment: spatial coherence and the persistence of fringes



Persistence of fringes as the source grows from a point source to finite size.



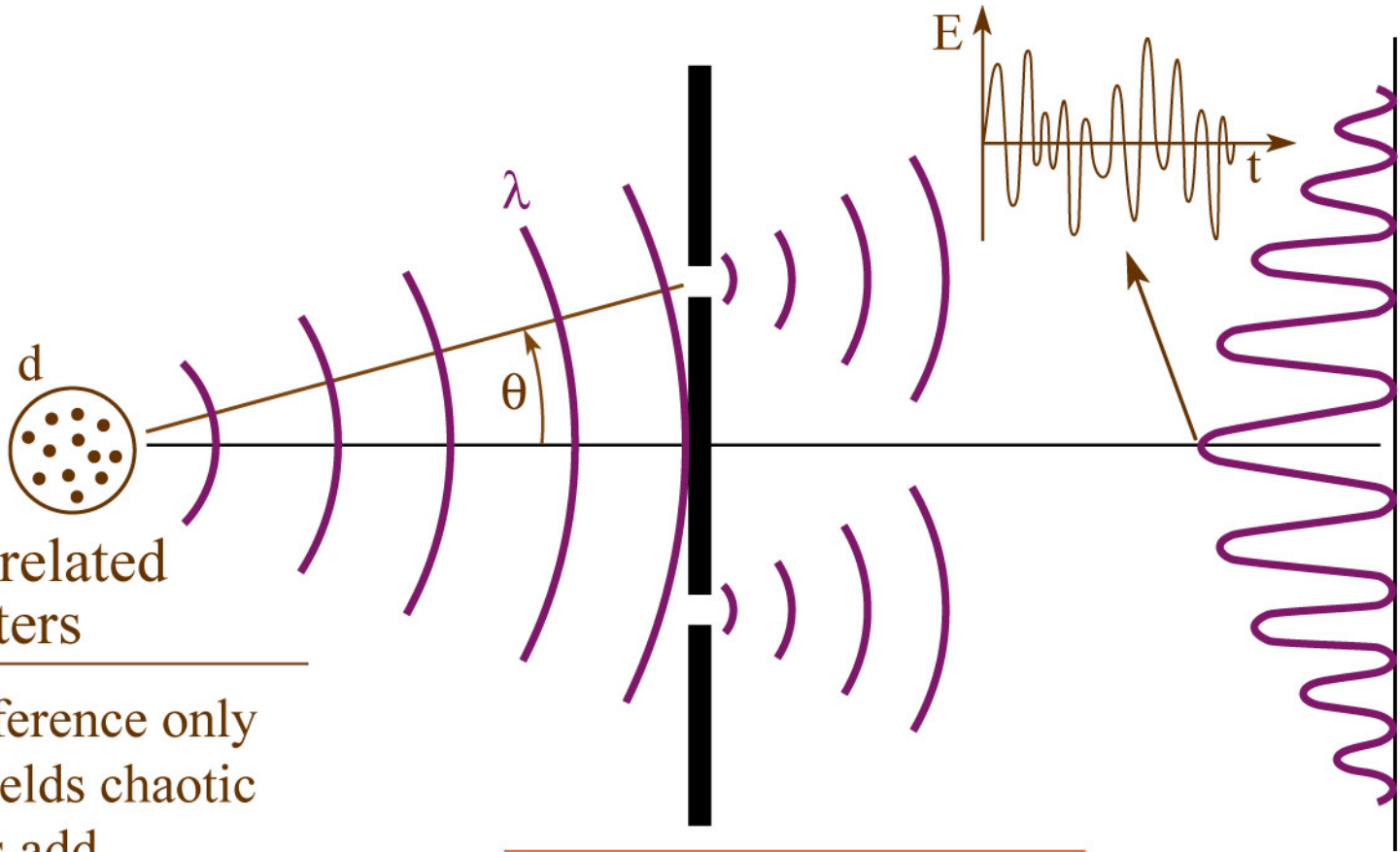
$$d \cdot 2\theta|_{FWHM} \approx \lambda/2$$

$$\lambda_{coh} = \lambda^2 / 2\Delta\lambda = \frac{1}{2} N_{coh} \lambda$$

CH08\_YoungsExprmt\_v3.ai



# Young's double slit experiment with random emitters: Young did not have a laser



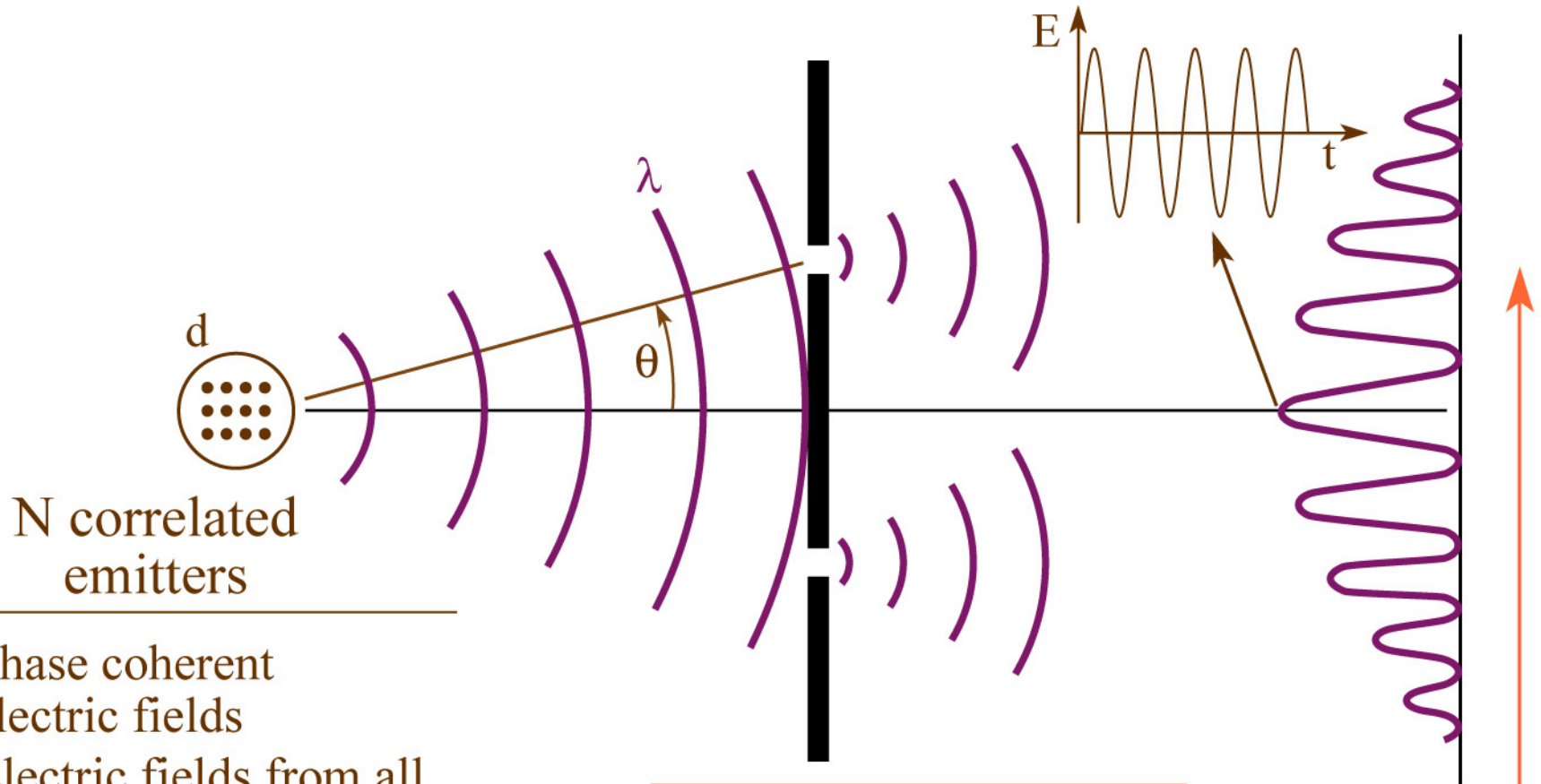
N uncorrelated emitters

- Self-interference only
- Electric fields chaotic
- Intensities add
- Radiated power  $\sim N$

$$d \cdot 2\theta|_{FWHM} \approx \lambda/2$$

$$\lambda_{coh} = \lambda^2 / 2\Delta\lambda = \frac{1}{2} N_{coh} \lambda$$

# Young's double slit experiment with phase coherent emitters (some lasers, or properly seeded FELs)

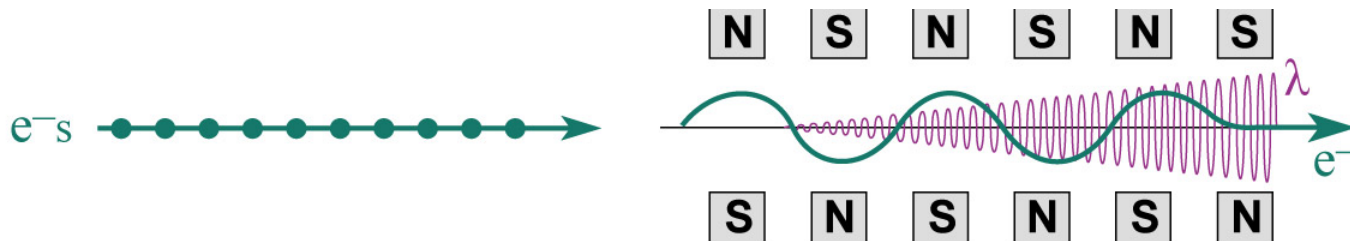


$N$  correlated emitters

- Phase coherent electric fields
- Electric fields from all particles interfere constructively
- Radiated power  $\sim N^2$
- New phase sensitive probing of matter possible

$$d \cdot 2\theta|_{\text{FWHM}} \approx \lambda/2$$

$$\lambda_{\text{coh}} = \lambda^2 / 2\Delta\lambda = \frac{1}{2} N_{\text{coh}} \lambda$$



- Uniformly distributed particles (beam) into undulator.
- Emission of radiation (“spontaneous” emission).
- Wave grows enough (undulator radiation) to begin affecting particle dynamics through  $m\mathbf{a} = -e\mathbf{E}$  radiation.
- Transverse coupling between  $\mathbf{E}_{\text{rad}}$  and transverse velocity  $\mathbf{v}_x$  (in undulator) leads to energy exchange between fields and particle (zero net at first)  $\frac{dE_e}{dt} = mc^2 \frac{d\gamma}{dt} = \mathbf{F} \cdot \mathbf{v} = -e \mathbf{E} \cdot \mathbf{v}_x$ .
- Modulated velocities with increments in  $\mathbf{v}_x$  lead to bunching on axis.
- Electron density modulation leads to stronger radiation,  

$$P_{\text{Tot}} \propto \frac{Q^4}{M^2} \sim N^2 \frac{e^4}{m^2}.$$
- Stronger fields (wave) drive stronger transverse velocity.
- Stronger  $\mathbf{v}_x$  drives stronger bunching, . . . stronger fields, . . . FEL action.

# Free Electron Lasers



Parameters	FLASH FEL (Hamburg 2005)	Fermi (Trieste, 2010)	LCLS (Stanford, 2009)	SACLA (Hyogo, 2011)	Swiss FEL (PSI, 2016)	PAL (Pohang, 2016)	EU XFEL (Hamburg, 2016)
$E_e$	1.25 GeV	1.5 GeV	13.6 GeV	8 GeV	5.8 GeV	10 GeV	17.5 GeV
$\gamma$	2,450	2,900	26,600	15,700	11,300	19,600	35,000
$\hat{I}$	1.3 kA	300 A	3.4 kA	9 kA	3 kA	3 kA	5 kA
$\lambda_u$	27.3 mm	55 mm	30 mm	18 mm	15 mm	26 mm	40 mm
N	989	216	3733	4986	3200	3456	4375
$L_u$	27 m	14 m	112 m	90 m	48 m	90 m	175 m
$\hbar\omega$	30-300 eV (4-40 nm)	20-300 eV (4-60 nm)	250 eV-10 keV (1.2-50 Å)	4.5-15 keV (0.8-2.8 Å)	2-12 keV (1-6 Å)	2-12 keV (1-6Å)	3-25 keV (0.05-4Å)
$\lambda/\Delta\lambda_{FWHM}$	100	1000	200-500	200-400	200-500	200-500	1000
$\Delta\tau_{FWHM}$	50 fsec	85 fsec	<10-70 fsec	<10 fsec	2-20 fsec	5-60 fsec	100 fsec
$\dot{J}$ (ph/pulse)	$3 \times 10^{12}$	$5 \times 10^{12}$	$2 \times 10^{12}$	$4 \times 10^{11}$	$10^{11}$	$5 \times 10^{11}$	$10^{12}$
rep rate	1 MHz @ 10 Hz	10 Hz	120 Hz	60 Hz	100 Hz	60 Hz	2.7 kHz @ 5 Hz
$\hat{P}$	1 GW	1 GW	30-60 GW	60 GW	3 GW	15 GW	20 GW
L	260 m	200 m	2 km	710 m	730 m	1.1 km	3.4 km
Polarization	linear	variable	linear	linear	Linear	Linear	Linear
Mode	SASE (3 $\omega$ Ti: sapphire)	Seeded	SASE & self-seeded	SASE	SASE & self-seeded	SASE & self-seeded	SASE

FreeElectronLasersChart\_July2015.ai



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# 2<sup>nd</sup> Edition in progress: new FEL, HHG, Coherence, and X-ray Imaging chapters



**Soft X-Rays and Extreme Ultraviolet Radiation**  
Principles and Applications  
David Attwood

Wavelength: 100 nm, 10 nm, 1 nm, 0.1 nm = 1 Å

Photon energy: 10 eV, 100 eV, 1 keV, 10 keV

Regions: UV, VUV, Extreme Ultraviolet, Soft X-rays, Hard X-rays

Key transitions:  $iS_{1L}$ ,  $C_K$ ,  $O_K$ ,  $iS_{1K}$ ,  $CuK_{\alpha_1}$ ,  $2\alpha_0$ ,  $CuK_{\alpha_2}$

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