

WIMPy Baryogenesis/Leptogenesis Miracle

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Arnab, Chandan, **SP**, Utpal, [[arXiv:1605.01292](https://arxiv.org/abs/1605.01292)] **Appeared Today !!**

2-6 May, ICTP, Trieste, Italy



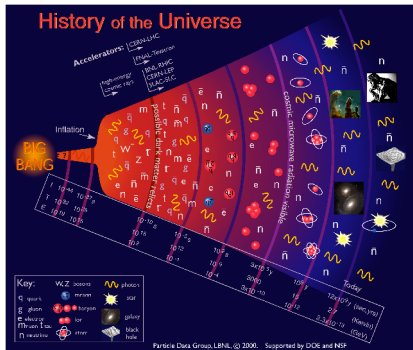
Workshop on perspectives on the Extragalactic
origin: Astrophysics to Fundamental Physics



Plan

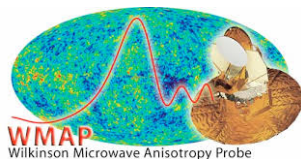
- ★ A cosmic puzzle: Link between Baryon asymmetry and Dark Matter of the Universe
- ★ WIMPY Baryogenesis
- ★ WIMPY Leptogenesis: a TeV scale Miracle.
 - Basic Picture
 - CP-asymmetry
 - Boltzmann Equation
- ★ Summary

Unsolved Problems in Astro-Particle Physics



- ★ **Baryon Asymmetry:** Why our present Universe is matter dominated ?
- ★ **Dark Matter:** Constituting 25% energy density of Universe ?
- ★ **Theoretical origin of non-zero Neutrino masses?**

Observational Facts



- ★ **Cosmic microwave background (CMB) anisotropy observations by the Wilkinson Microwave Anisotropy Probe (WMAP):**

Baryon to Photon ratio as

$$\eta_B^{\text{CMB}} \equiv \frac{n_B - n_{\bar{B}}}{n_\gamma} = (6.3 \pm 0.3) \times 10^{-10} \quad \text{[From Planck 2015 also]}$$

- ★ **Dark Matter (DM) relic abundance:** $\Omega_{DM} h^2 \simeq 0.1123$
- ★ **Baryonic Matter abundance:** $\Omega_B h^2 \simeq 0.02$

Can DM and Baryon asymmetry connected ...?

- ★ **Coincidences:** $\Omega_{DM}/\Omega_B \simeq 5$ related to Asymmetric Dark Matter
- ★ **WIMP Miracle:** $\Omega_{DM} \simeq 0.1123$ Weak scale mass and coupling, thermal freeze-out
- ★ **Generally, baryogenesis and the establishment of the dark matter number density are treated as independent processes.**
- ★ **Can weakly interacting massive particle (WIMP) dark matter connected dark matter physics with baryogenesis ?**

★ **WIMPy Baryogenesis:-** Cui, Randall, Shuve, JHEP 1204 (2012) 075

Different from Asymmetric Dark Matter: **Not discussed here !**

Keep the WIMP miracle and ask the question

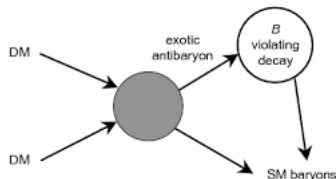
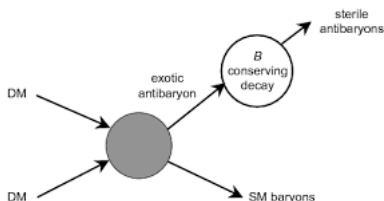
can we have a framework where the baryon asymmetry can be related to the thermal WIMP relic abundance?

WIMPy Baryogenesis: Miracle and Miracle !!

Cui, Randall, Shuve, JHEP 1204 (2012) 075

★ Two miracle happen in one framework

- **WIMP Miracle:**
weak-scale DM, thermal relic abundance
- **WIMPy baryogenesis miracle:**
Weakly Interacting Massive Particle (WIMP) Dark Matter annihilation is directly responsible for Baryon asymmetry



Sakharov conditions for WIMPy Baryogenesis

★ Baryon number violation:

- WIMP annihilations violate baryon or lepton number.
We propose a minimal $U(1)_{B-L}$ gauge model to realize.

★ CP violation:

- WIMP couplings to Standard Model fields violate CP

★ Departure from thermal equilibrium:

- the cooling of the universe provides the necessary departure from thermal equilibrium.

A TeV scale model for WIMPy Leptogenesis

Arnab, Chandan, **SP**, Utpal [arXiv:1605.01292]

- ★ propose a novel framework to explain dark matter abundance and matter-antimatter asymmetry simultaneously via WIMPy leptogenesis.
- ★ SM + DM fermions (N 's)+ Extra scalars (η, ζ)
- ★ The role of N 's are two fold:-
 - To cancel the gauge triangle anomalies induced by $U(1)_{B-L}$ gauge model:

$$\begin{array}{cc} \mathcal{A} \left[(U(1)_{B-L})^3 \right], & \mathcal{A} \left[U(1)_{B-L} (U(1)_Y)^2 \right] \\ \mathcal{A} \left[U(1)_{B-L} (SU(2)_L)^2 \right] & \mathcal{A} \left[\text{gravity}^2 \times U(1)_{B-L} \right] \end{array}$$

- excellent WIMP Dark Matter while their annihilation can provide DM abundances as well as correct Baryon asymmetry of the Universe.

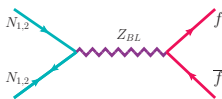
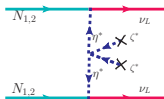
Basic Ingredients of WIMPy Leptogenesis

$$\mathcal{L} \supset y_u \bar{q}_L \tilde{H} u_R + y_d \bar{q}_L H d_R + y_e \bar{\ell}_L H e_R + y_\nu \bar{\ell}_L \tilde{\eta} N_{1,2} + \sum_{\alpha,\beta=1,2} h_{\alpha\beta} \bar{N}_\alpha^c \chi N_\beta + \sum_{\alpha,\beta=1,2} h_{3\alpha} \xi \bar{N}_\alpha^c N_3. \quad (1)$$

★ Sakharov Conditions Revisited:

- **Baryon Number Violation**

WIMP annihilations violate L number via $\lambda''(\zeta^\dagger \eta)^2$



- **CP violation**

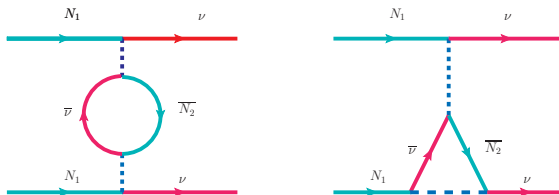
WIMP couplings to SM fields violate CP $y_\nu \bar{\ell}_L \tilde{\eta} N_{1,2}$

- **Departure from thermal Equilibrium**

The coolong/Expansion of the Universe provides the necessary departure from thermal equilibrium

CP-asymmetry

- ★ CP-violation comes through the interference with the tree and loop diagrams:



- ★ CP asymmetry parameter for Lepton asymmetry

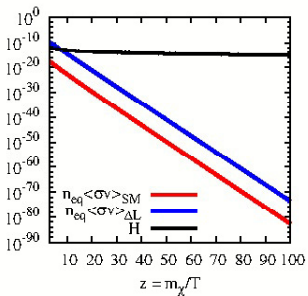
$$\epsilon = \frac{\sigma_{N_1 N_1 \rightarrow \nu \nu} + \sigma_{\bar{N}_1 \bar{N}_1 \rightarrow \nu \nu} - \sigma_{\bar{N}_1 \bar{N}_1 \rightarrow \bar{\nu} \bar{\nu}} - \sigma_{N_1 N_1 \rightarrow \bar{\nu} \bar{\nu}}}{\sigma_{N_1 N_1 \rightarrow \nu \nu} + \sigma_{\bar{N}_1 \bar{N}_1 \rightarrow \nu \nu} + \sigma_{\bar{N}_1 \bar{N}_1 \rightarrow \bar{\nu} \bar{\nu}} + \sigma_{N_1 N_1 \rightarrow \bar{\nu} \bar{\nu}}}. \quad (2)$$

Cross-check for WIMPy Leptogenesis !!

[arXiv:1605.01292]

★ Successful generation of lepton asymmetry through WIMPy leptogenesis requires:

- **Strength of CP-violating DM annihilation cross-section should be larger than CP-conserving ones** $\frac{\langle \sigma_{N_{1,2} N_{1,2} \rightarrow \nu \nu} \rangle}{\langle \sigma_{N_{1,2} N_{1,2} \rightarrow \bar{f} \bar{f}} \rangle} > 1$



FREEZE OUT: QUALITATIVE

(1) Assume a new heavy particle X is initially in thermal equilibrium:



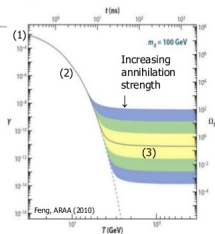
(2) Universe cools:



(3) Universe expands:



Zeldovich et al. (1960s)

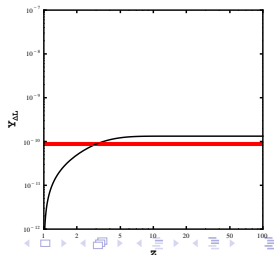
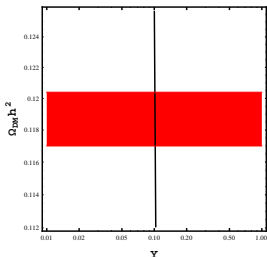


- **Out-of-equilibrium condition:** $\frac{n_{eq}^{N_{1,2}} \langle \sigma_{N_{1,2} N_{1,2} \rightarrow \nu \nu} | \nu \rangle}{H} < 1$

Effect of Wash-out processes: **Central Result**

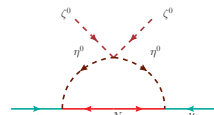
[arXiv:1605.01292]

- ★ WIMP annihilation can generate a baryon asymmetry, there are other processes that have the potential to wash out the asymmetry, and their freeze-out is crucial to create the observed baryon asymmetry.
 - If washout processes freeze out before WIMP freeze-out, then a large baryon asymmetry may accumulate
 - Its final value is proportional to the WIMP abundance at the time that washout becomes inefficient.

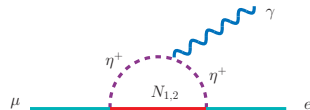


Other Phenomenology

★ Neutrino Mass via Radiative Mechanism



★ Lepton Flavour Violation $\mu \rightarrow e \gamma$



Potential LHC prospects with TeV scale spectrum and hence, can easily be probed by upcoming search experiments

★ WIMPy Baryogenesis/Leptogenesis Miracle New !!

- **A miracle by generating the observed baryon asymmetry through annihilations of weak-scale dark matter**
- **provides a dynamical connection between the dark matter and baryon abundances**
- **Viable Window for new Physics searches at TeV scale: LHC !**