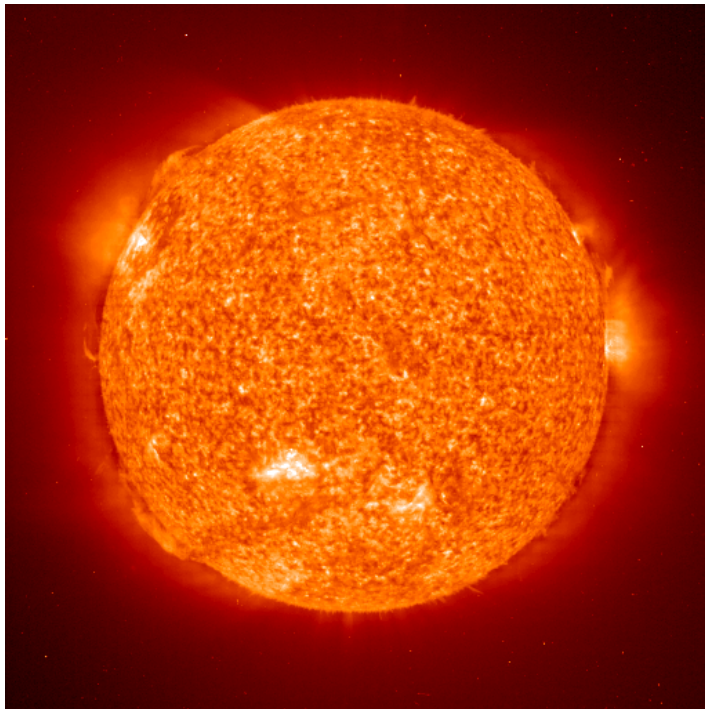


The tokamak principle – magnetic fusion



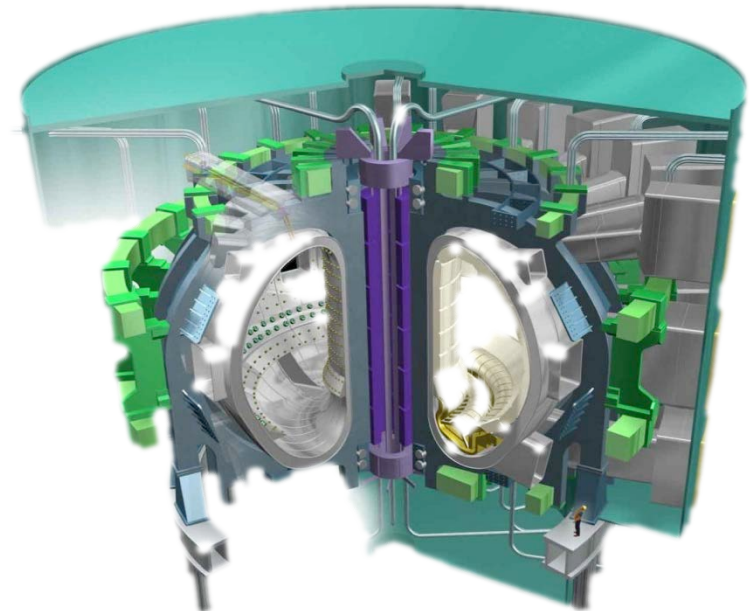
Joint ICTP-IAEA College on Advanced
Plasma Physics, Triest, Italy, 2016

Marco Wischmeier

Max-Planck-Institut für Plasmaphysik

85748 Garching

marco.wischmeier at ipp.mpg.de





Based on lectures given by Hartmut Zohm, Ursel Fantz, Sebastijan Brezinsek, Sibylle Günter and additional material



Introduction: The inspiration

Nuclear Fusion - The tokamak principle – magnetic fusion (~60 min)

Progressing Performance – Tokamaks Core physics (~60 min)

Power Exhaust – Materials for Tokamaks (~60 min)



The inspiration



Particles fuse inside the sun

IPP

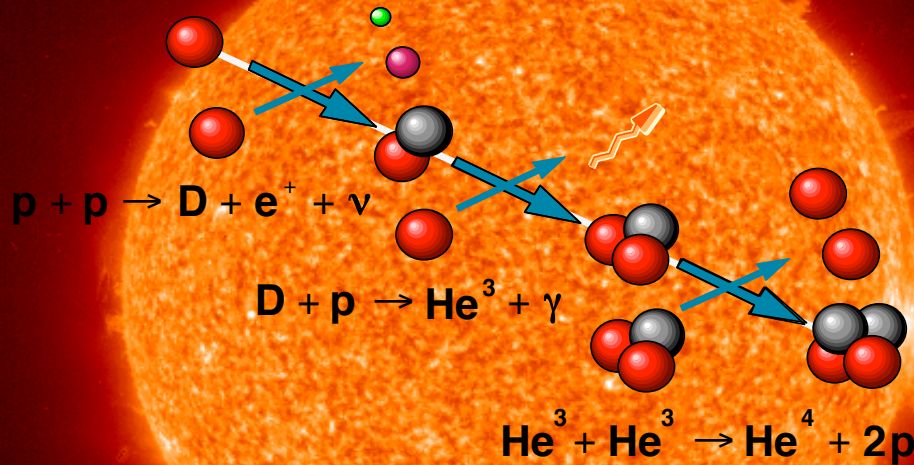




Nuclear fusion - the energy source of the sun and the stars



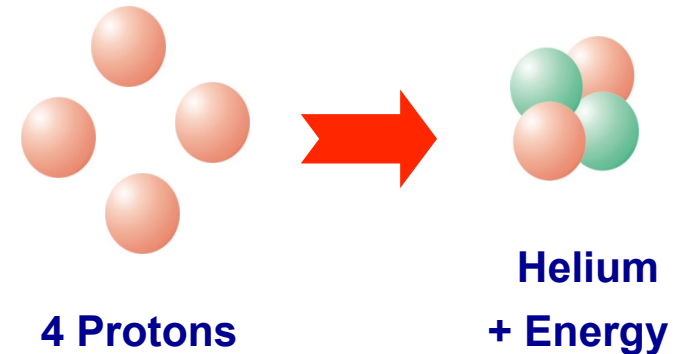
SOHO EIT, He II line, 304 Å
May 18, 1996 at 20:02



Energy set free by fusion of light nuclei, in the sun:

pp-cycle

Essentially:



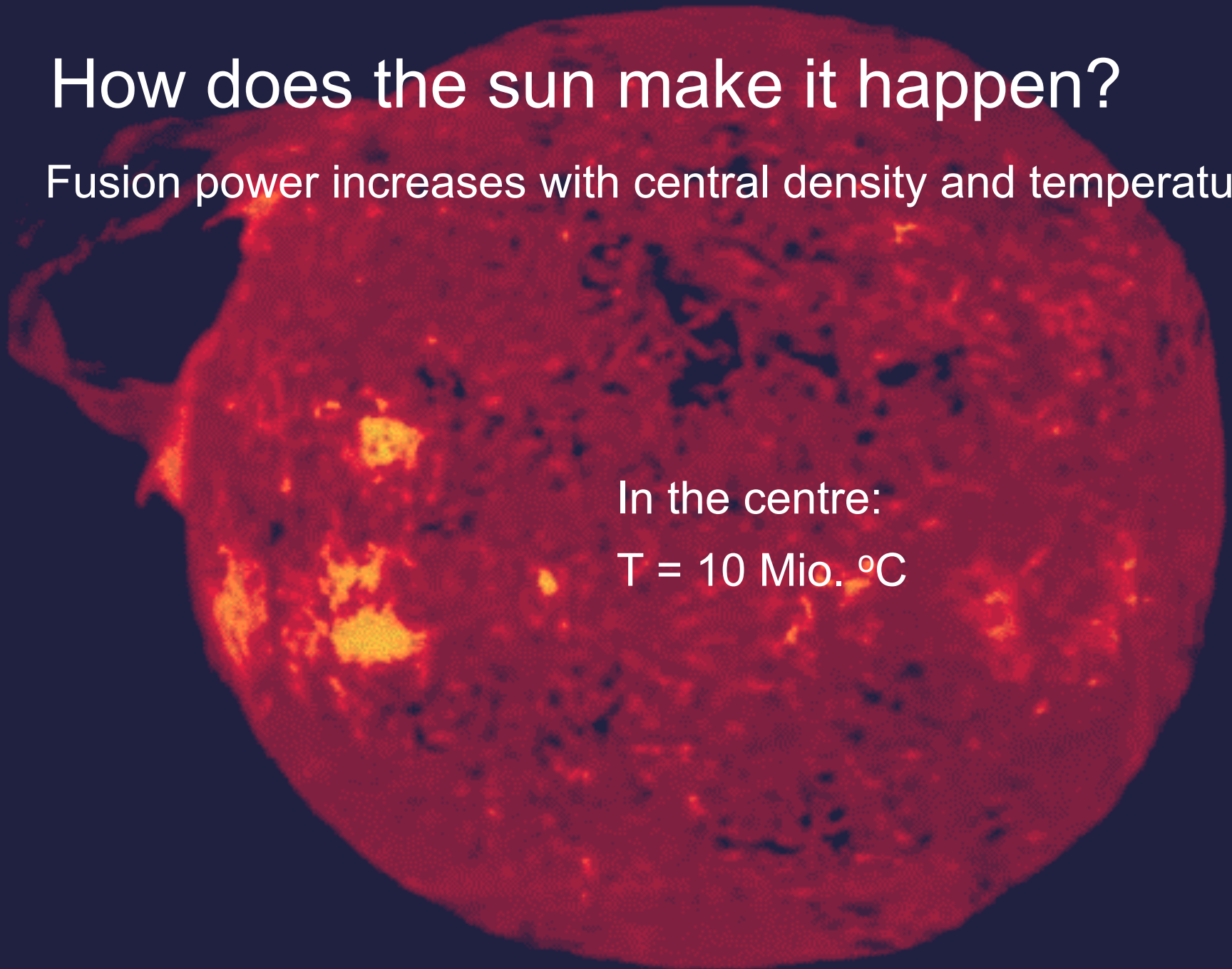
weak interaction involved,
thus small fusion rates, on
earth: 10^{27} times larger fusion
rates required

How does the sun make it happen?

Fusion power increases with central density and temperature

In the centre:

$T = 10 \text{ Mio. } ^\circ\text{C}$



How does the sun make it happen?

Fusion power increases with central density and temperature

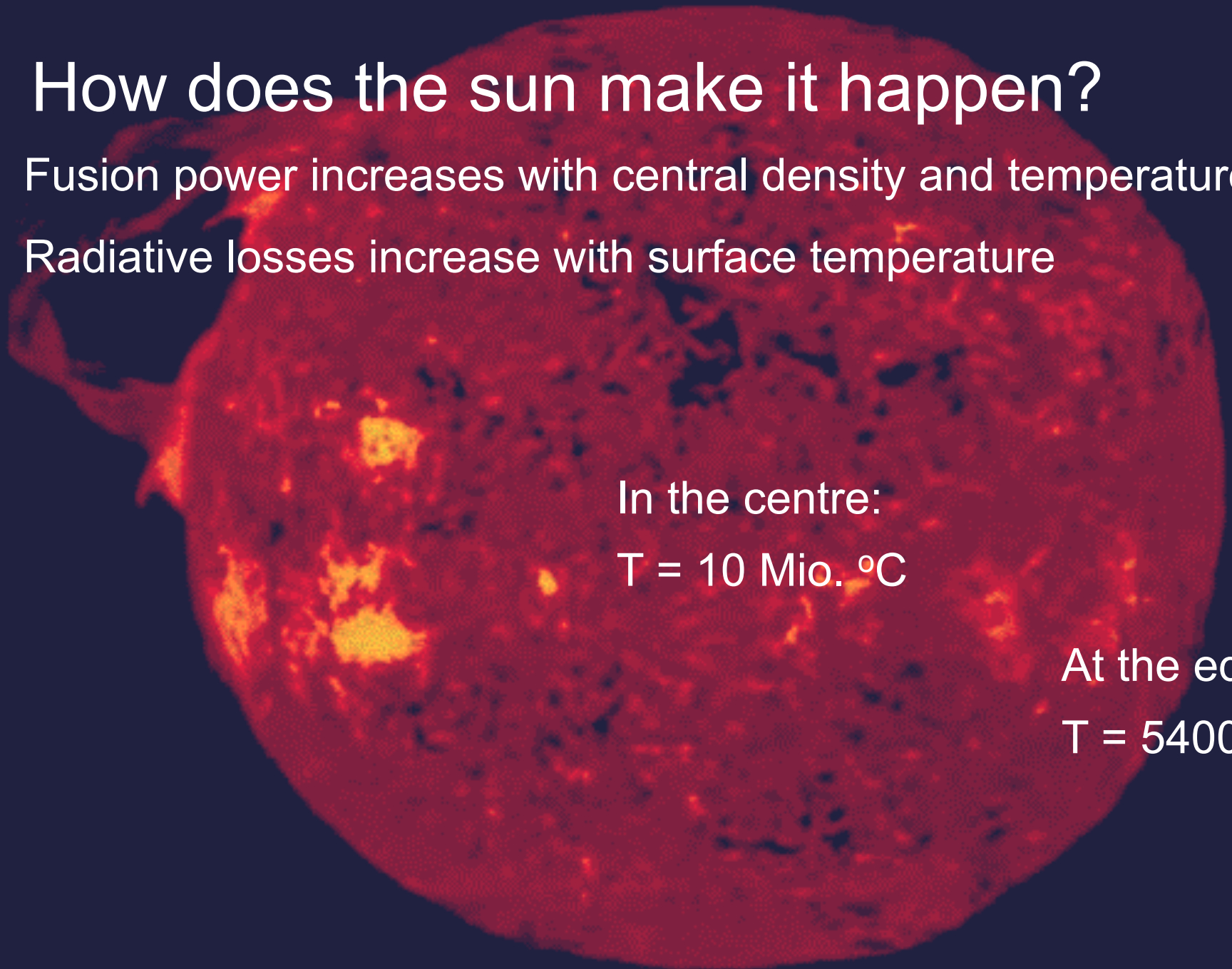
Radiative losses increase with surface temperature

In the centre:

$T = 10 \text{ Mio. } ^\circ\text{C}$

At the edge:

$T = 5400 \text{ } ^\circ\text{C}$



How does the sun make it happen?

Fusion power increases with central density and temperature

Radiative losses increase with surface temperature

In the centre:

$T = 10 \text{ Mio. } ^\circ\text{C}$

At the edge:

$T = 5400 \text{ } ^\circ\text{C}$

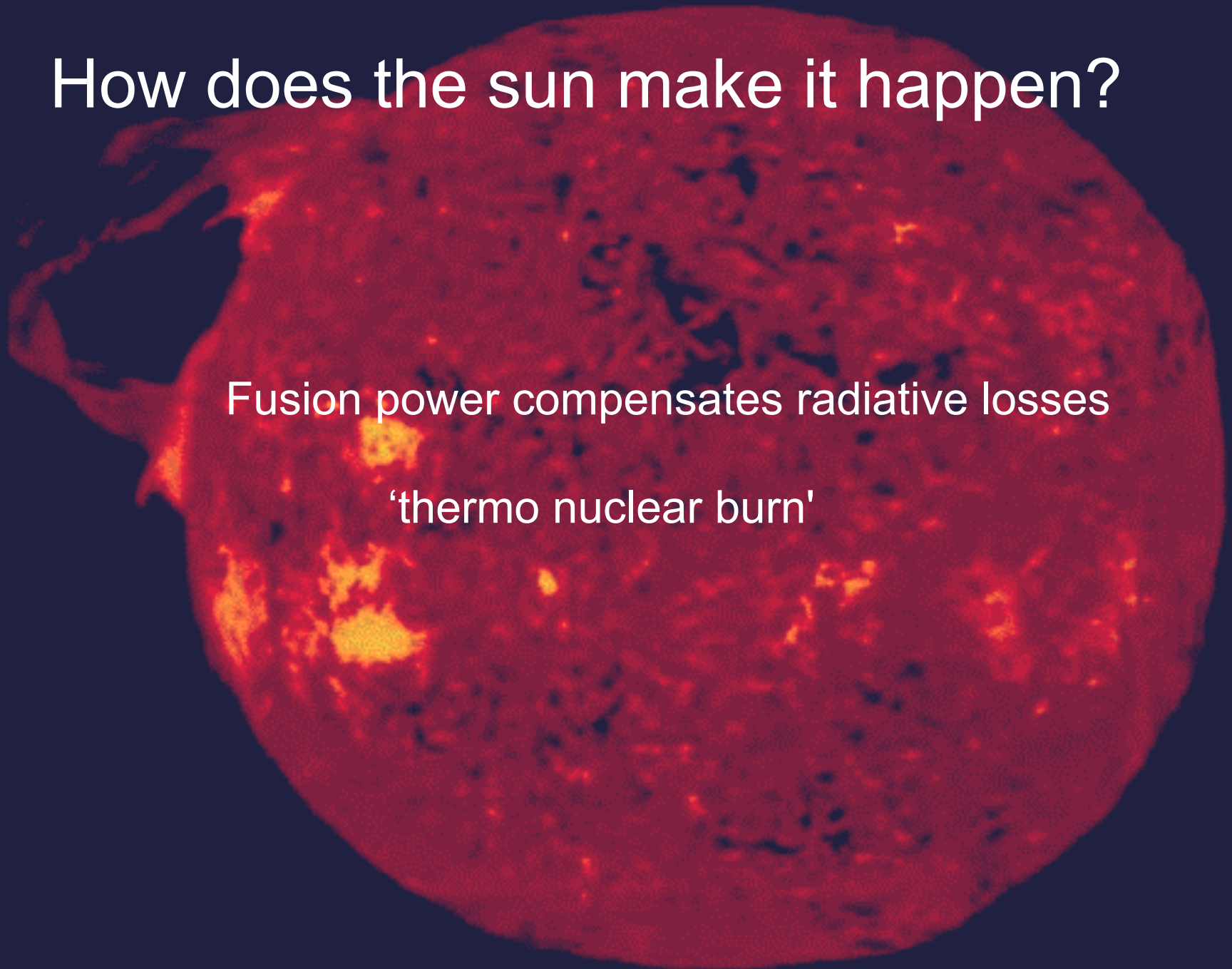
⇒ high central temperature, low edge temperature

⇒ good heat insulation necessary!

How does the sun make it happen?

Fusion power compensates radiative losses

'thermo nuclear burn'



How does the sun make it happen?

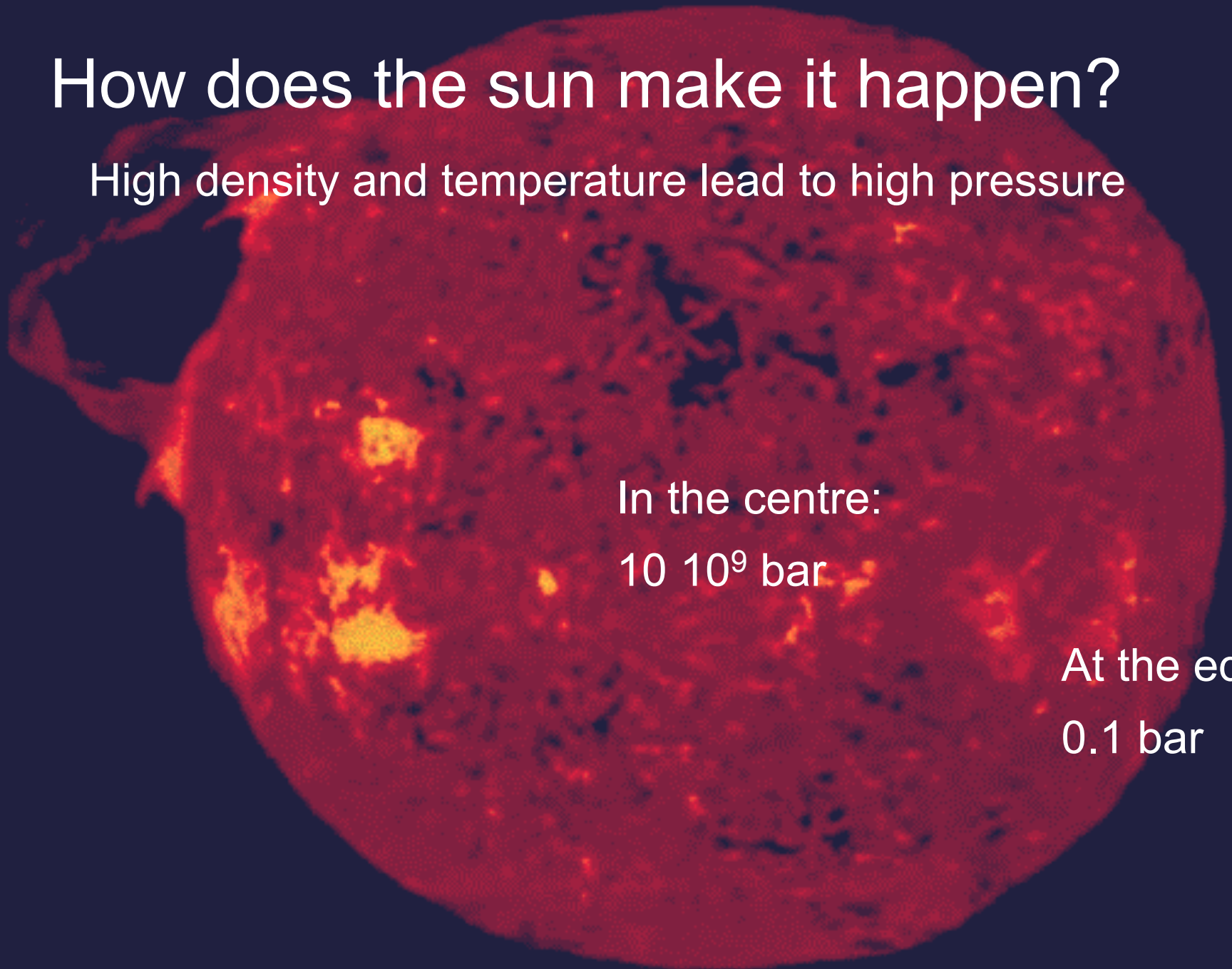
High density and temperature lead to high pressure

In the centre:

10^{10} bar

At the edge:

0.1 bar



How does the sun make it happen?

High density and temperature lead to high pressure

In the centre:

10^{10} bar

At the edge:

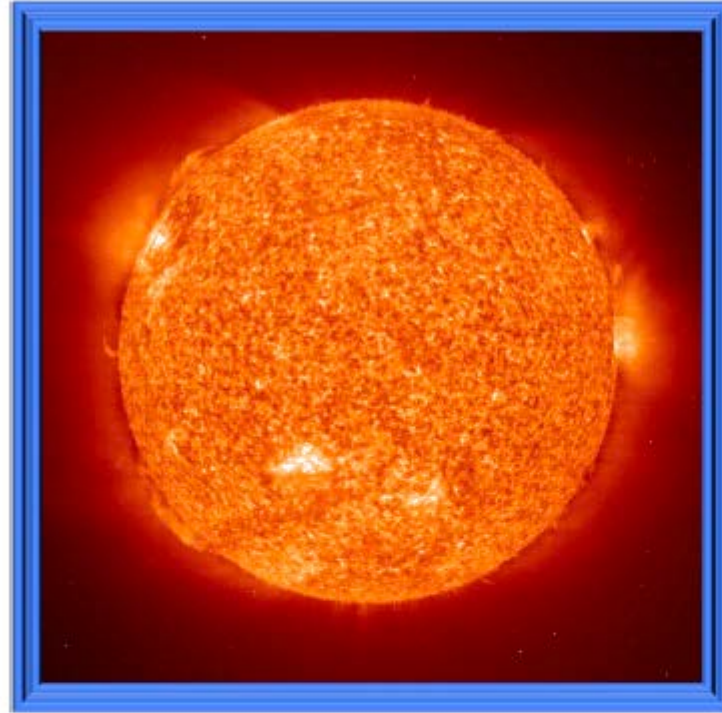
0.1 bar

Pressure difference leads to strong expansion force
⇒ confining force necessary (e.g. gravitation)



Aim of nuclear fusion research

IPP



We say we want to put the sun into a box.
The idea is great. The problem is, that we
don't know how to build the box.

Sébastien Balibar, research director at CNRS



Methods for energy production

Helium



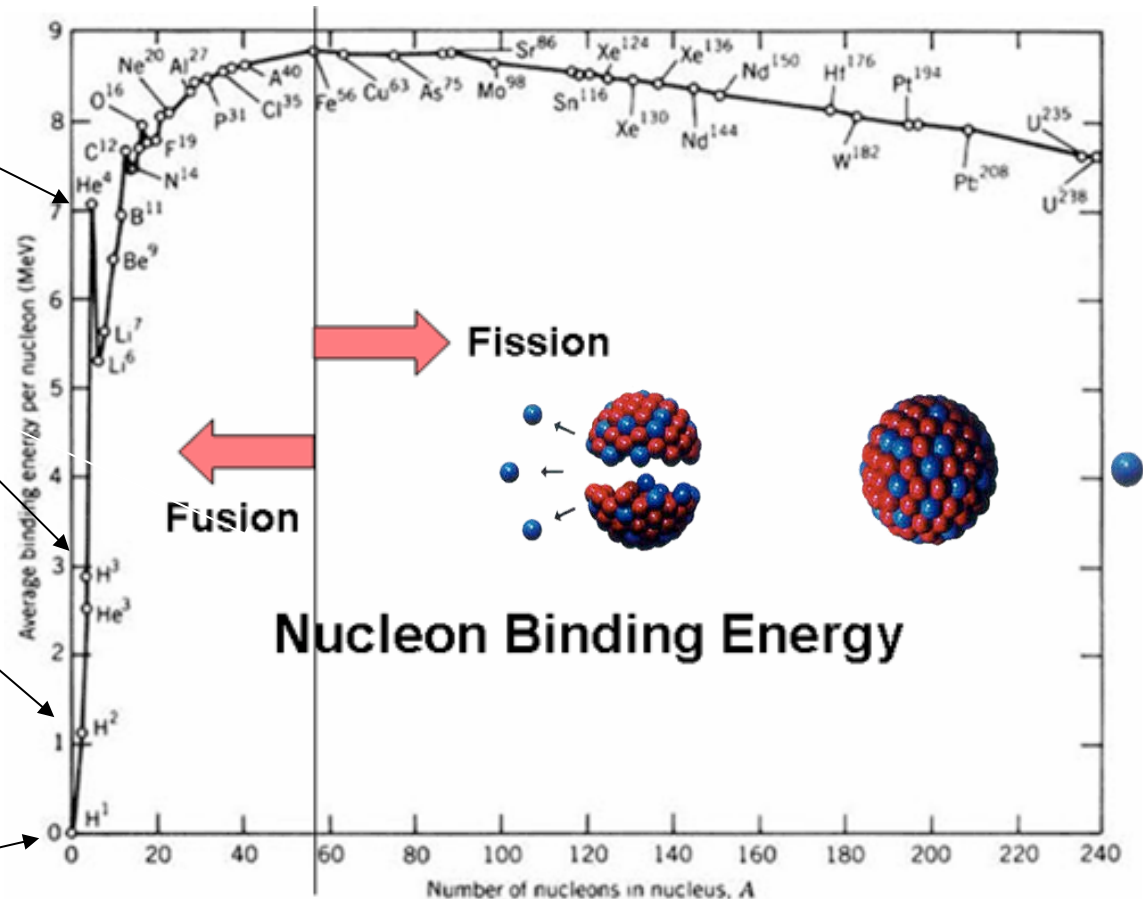
Tritium



Deuterium

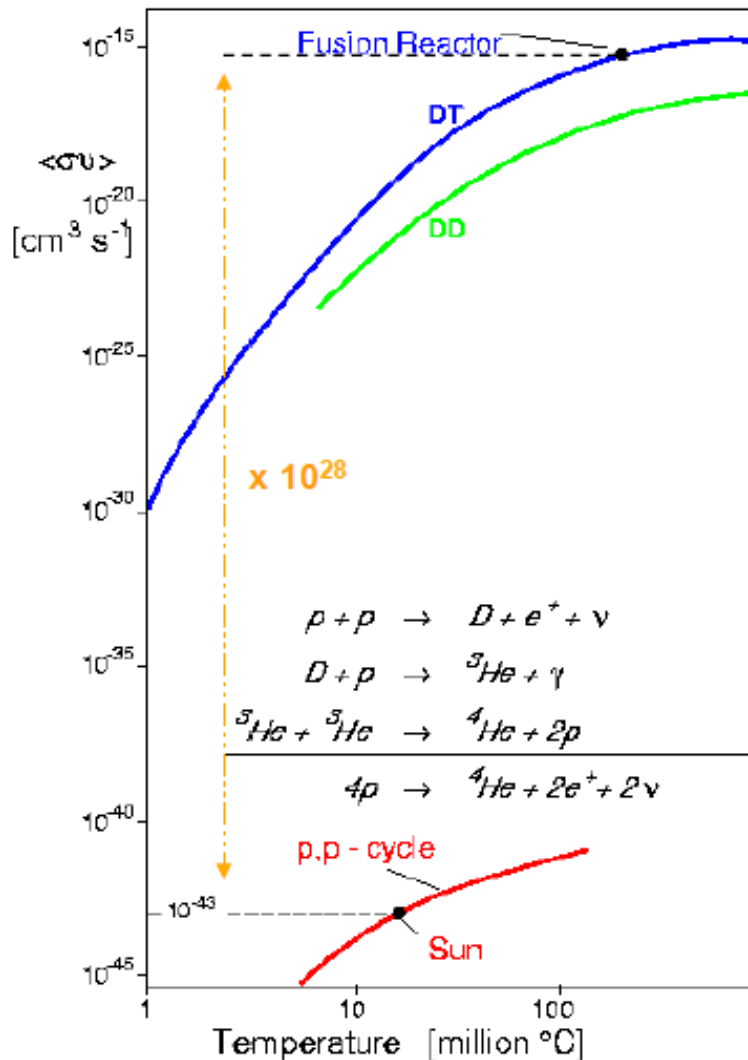


Proton





Nuclear fusion on earth – the right reaction



- Rate for DD and DT reactions is much larger than pp reaction
- DT plasma is a candidate for fusion reactor
- Problems:
 - ❖ high temperature (100Mio C)
 - ❖ T is radioactive: $t_{1/2} = 12.3 \text{ y}$
 - ❖ T is not easily accessible and must be bred from Li



In a power plant: fusion of deuterium and tritium

IPP

Deuterium



Fuel for fusion

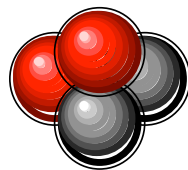
Tritium is very rare in nature, because of its short decay length (12.3y). It can be bred from Lithium in a reactor.

Tritium

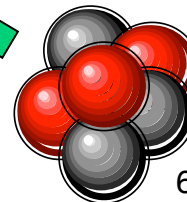


4.8 MeV

^4He

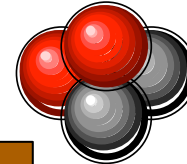


$^6\text{Lithium}$



„the ash“

$^4\text{Helium}$

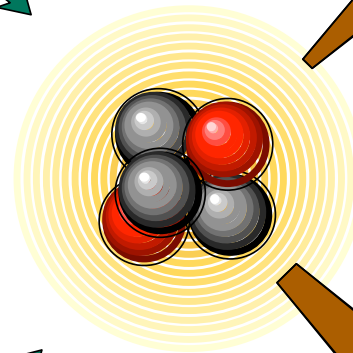


3.5 MeV

14.1 MeV

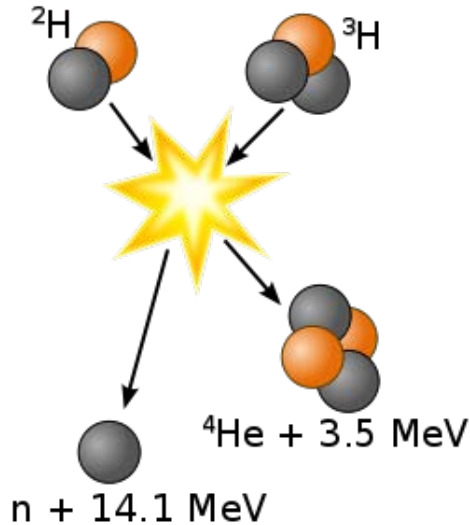
energy

Neutron





Energy gain in a D-T-fusion reaction



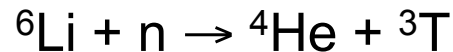
Very high energy density and availability:

10000 t coal corresponds to

800 g D-T

Deuterium is contained in water to 0.015%.

Tritium must be generated in situ:



One bathtub of water
Li in a battery



Energy for 50 years for
an average family





Easiest idea for fusion is:
collide two particle beams against each other



Energy needed for acceleration $\sim 0.1 \text{ MeV}$

Energy gain from fusion reaction: 17.6 MeV

Good deal!



Easiest idea for fusion is:
collide two particle beams against each other



Energy needed for acceleration $\sim 0.1 \text{ MeV}$
Energy gain from fusion reaction: 17.6 MeV

Good deal!

Problem:



Particles are deflected by coulomb forces (nearly always)



Nuclear fusion on earth – energy balance



Fusion reaction:

- Difficult obstacle
- Very deep hole

Solution:

- Many attempts without having to accelerate the particles for each attempt
- Confining a DT mixture at high temperature (200Mio C)
- Every gas becomes a plasma at these temperatures

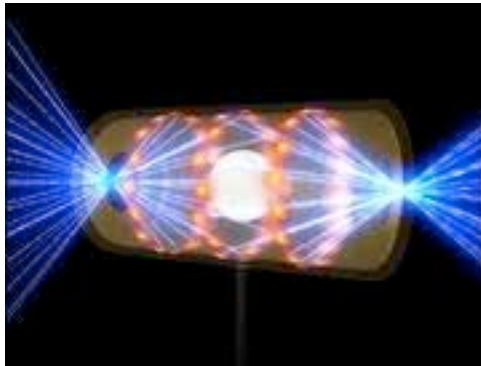


Two ways to fusion energy

**D-T fusion needs temperatures 10 times larger than in the solar interior:
~150 Mio degrees**

Inertial fusion:

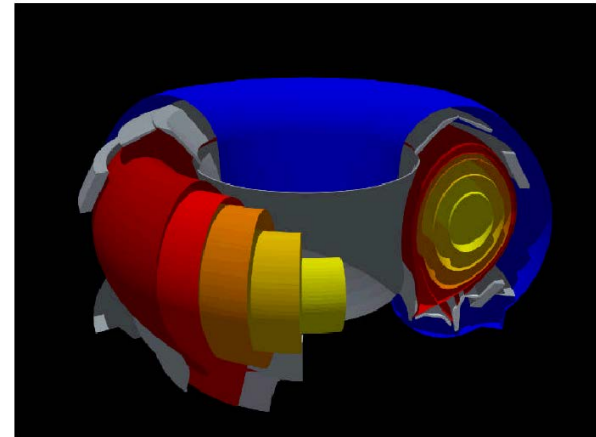
- fast heating (e.g. laser) of small pellets, minor explosions
- pressure comparable to solar interior ($n \sim 10^{31} \text{ m}^{-3}$)
- confinement time $\sim 10^{-10}$



Mainly in the US and Japan, keep-in-touch activities in Europe

Magnetic fusion:

- confinement through magnetic fields
- very low pressure
- confinement time: a few seconds





The tokamak principle – magnetic fusion (~60 min)

Progressing Performance – Tokamaks Core physics (~60 min)

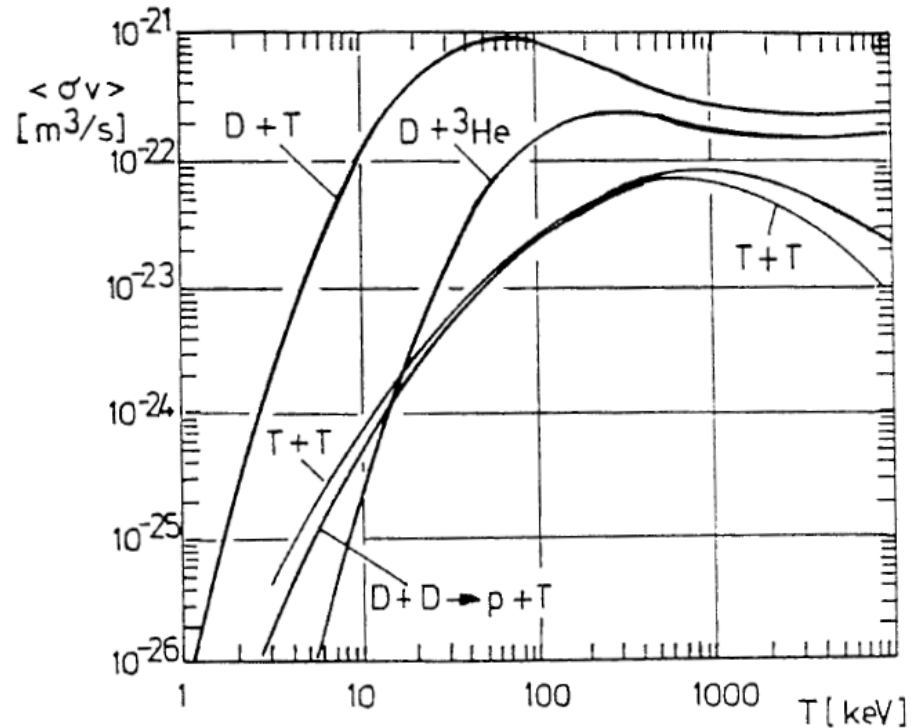
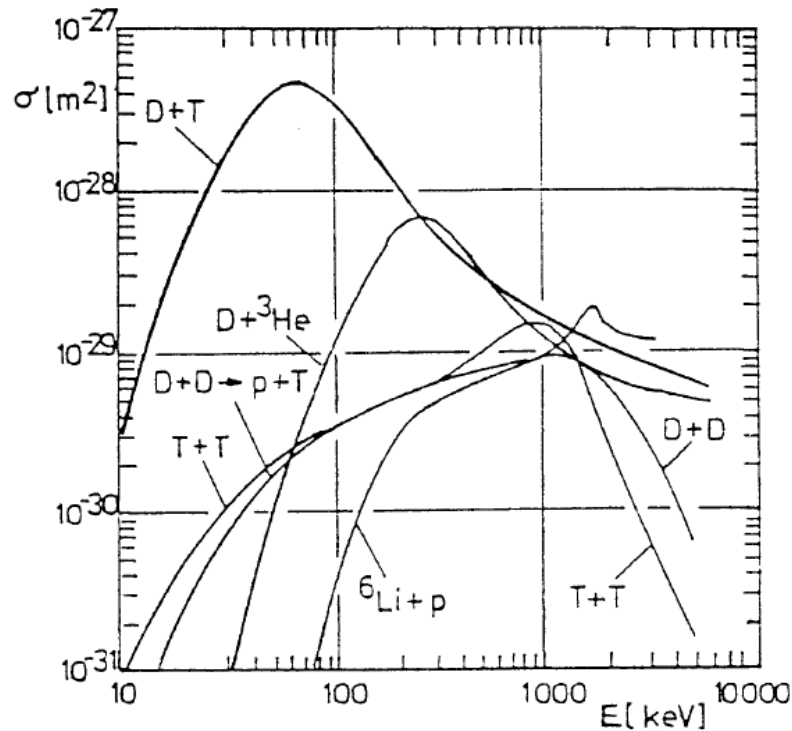
Power Exhaust – Materials for Tokamaks (~60 min)



Basic Fusion Physics



D-T reaction most favorable for energy gain



Highest cross section with maximum at lowest energy

At these energies, elastic collision still 100 x more likely than fusion

- crossed beam configurations would not be efficient enough
- have to confine the particles to allow many collisions – thermal plasma



Reactor energetics: the 'Lawson' criterion for $n\tau_E$



α -heating compensates losses:

- radiative losses (Bremsstrahlung)
- heat conduction and convection

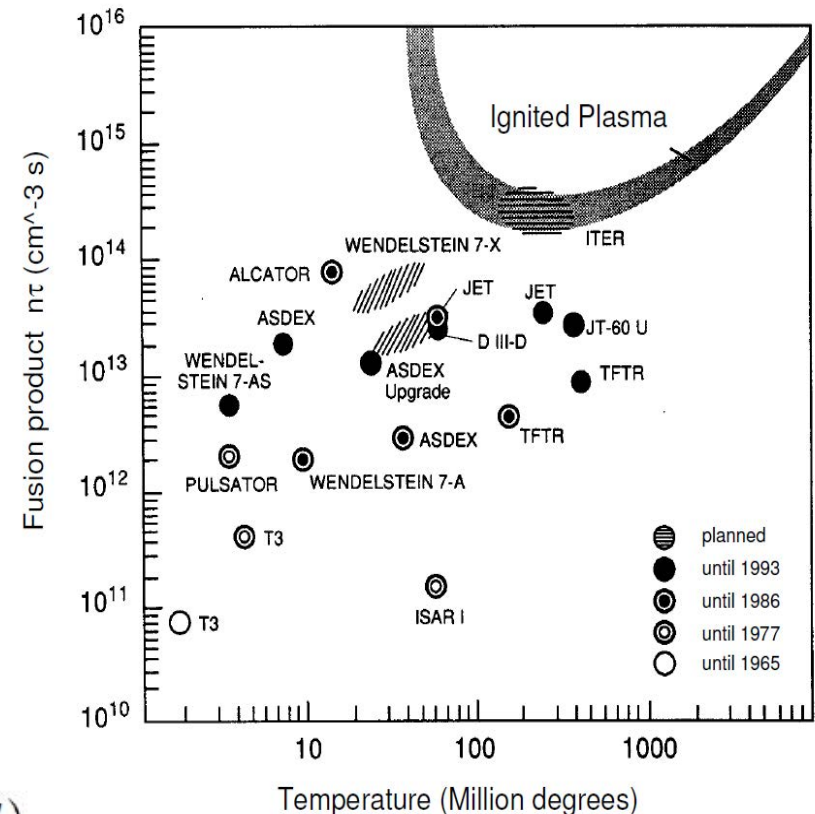
$$\frac{n_e^2}{4} \langle \sigma u \rangle E_\alpha > c_{Br} n_e^2 Z_{eff} \sqrt{k_B T} + \frac{3n_e k_B T}{\tau_E}$$

$$\tau_E = W_{plasma} / P_{loss} \quad (\text{'energy confinement time'})$$

leads to

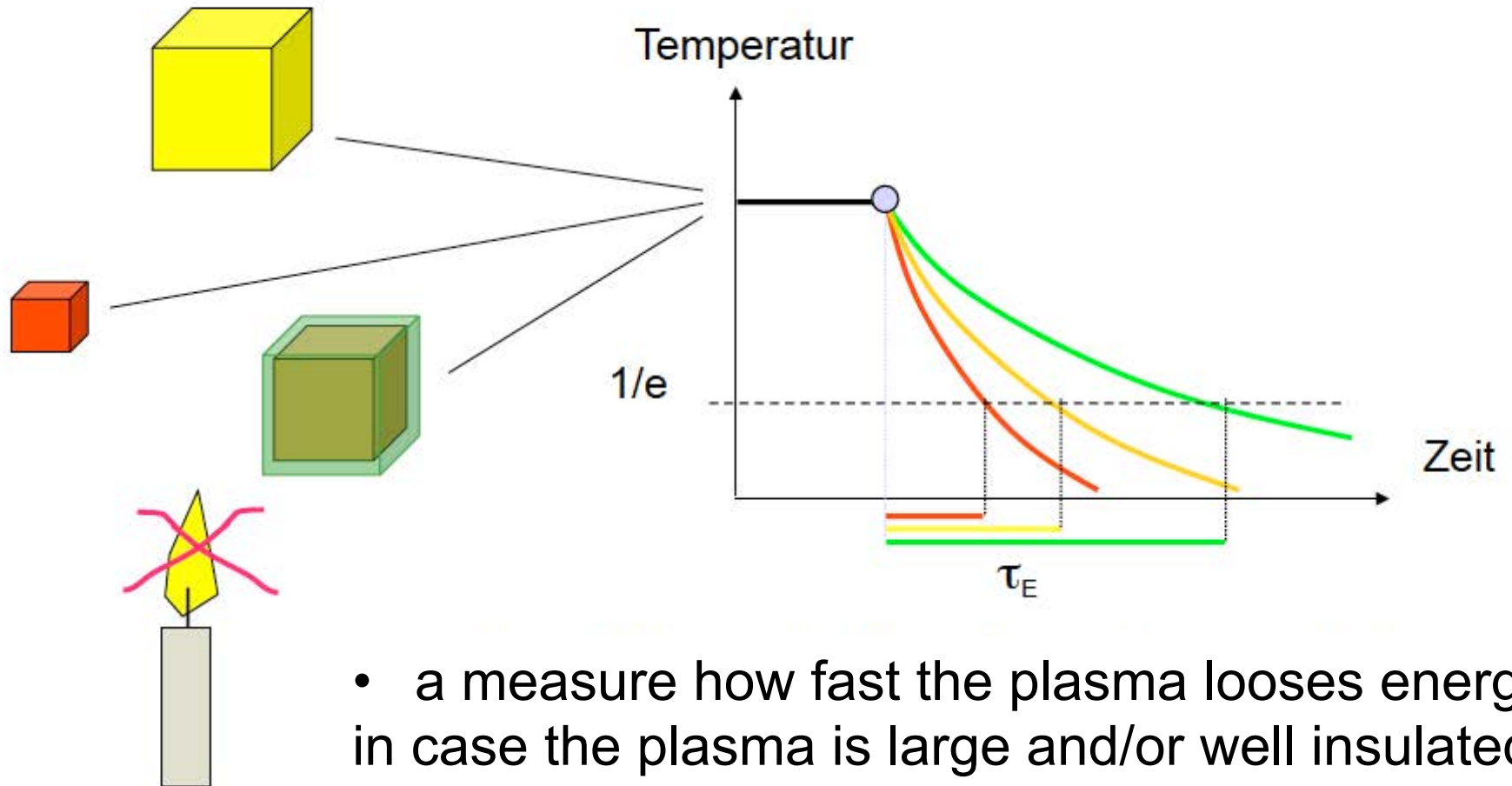
$$n_e \tau_E > \frac{3k_B T}{\langle \sigma u \rangle E_\alpha / 4 - c_{Br} Z_{eff} \sqrt{k_B T}} = f(T)$$

which has a minimum for $n\tau_E = 2 \times 10^{20} \text{ m}^{-3} \text{ s}$ at $T = 20 \text{ keV}$





Confinement time



- a measure how fast the plasma loses energy, in case the plasma is large and/or well insulated
→ small losses / large value,



Figure of merit for fusion performance $nT\tau$



Power P_{loss} needed to sustain plasma

- determined by thermal insulation:

$$\tau_E = W_{plasma}/P_{loss} \text{ ('energy confinement time')}$$

Fusion power increases with W_{plasma}

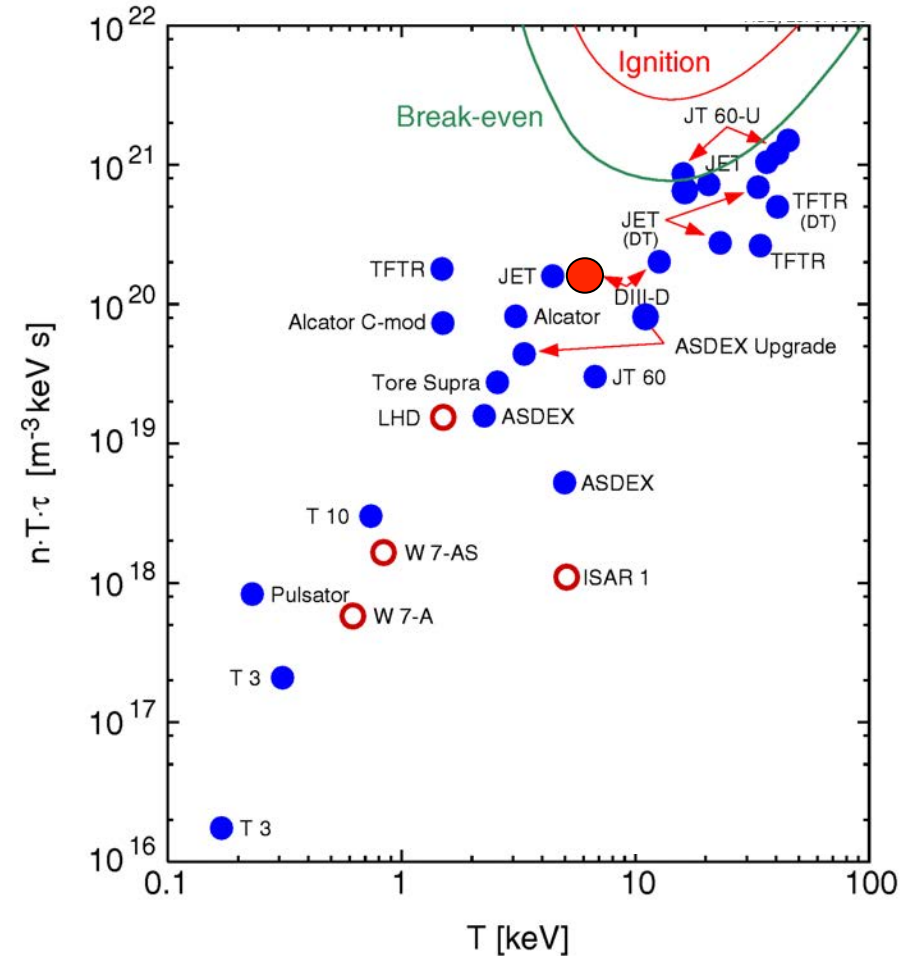
$$P_{fus} \sim n_D n_T \langle \sigma v \rangle \sim n_e^2 T^2 \sim W_{plasma}^2$$

Present day experiments: P_{loss} compensated by external heating

$$Q = P_{fus}/P_{ext} \approx P_{fus}/P_{loss} \sim nT\tau_E$$

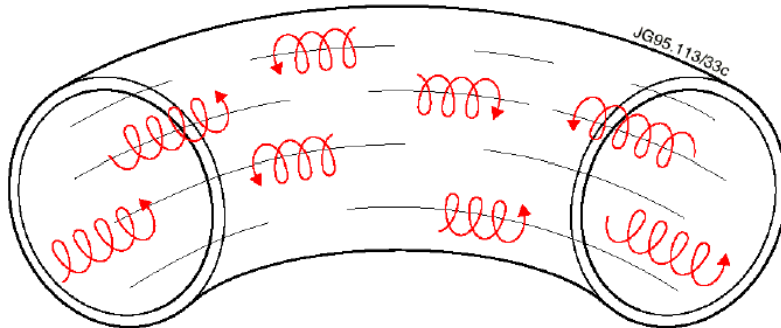
Reactor: P_{loss} compensated by α -(self)heating

$$Q = P_{fus}/P_{ext} = P_{fus}/(P_{loss} - P_{\alpha}) \rightarrow \infty \text{ (ignited plasma)}$$



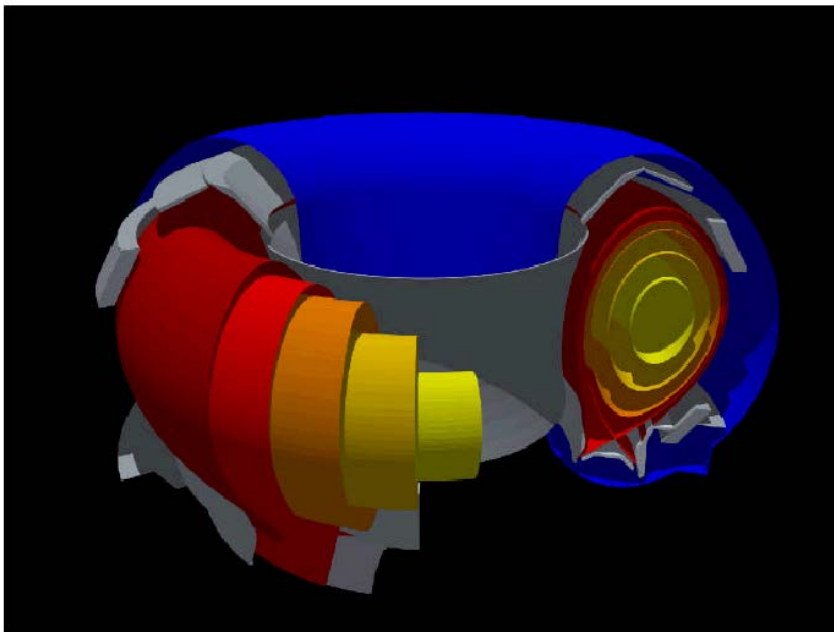


Magnetic confinement of fusion plasmas



Magnetic field:

- reduces perpendicular particle motion (particle and energy confinement)
- balances plasma pressure (~ 10 atm)



No end losses in torus
Field lines on magnetic surfaces

System with low power per volume

ITER: 0.5 kW/l

Combustion engine: 50 kW/l

Fission reactor: 100 kW/l



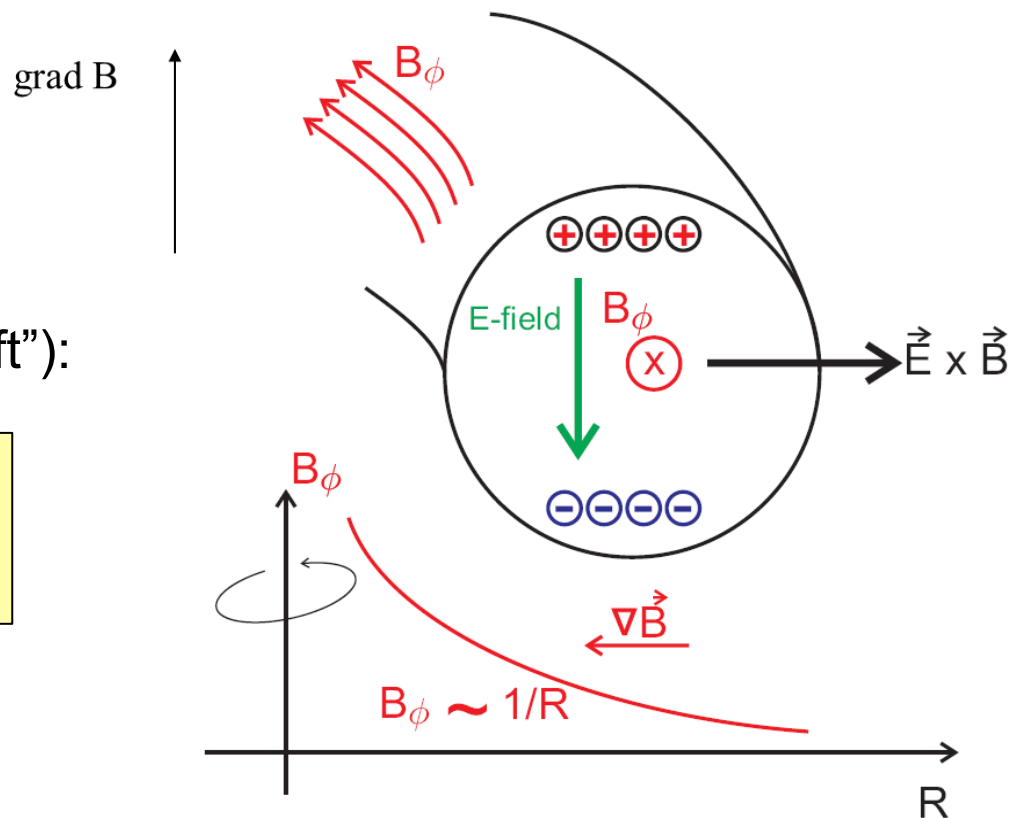
Is a simple toroidal confinement sufficient?

Drift of particles due to forces:

$$\vec{V}_D = \frac{\vec{F} \times \vec{B}}{qB^2} \Rightarrow$$

∇B and curvature drift (“torus drift”):

$$\vec{V}_D = \frac{m}{qB^3} \left(v_{\parallel}^2 + \frac{1}{2} v_{\perp}^2 \right) \vec{B} \times \nabla B$$



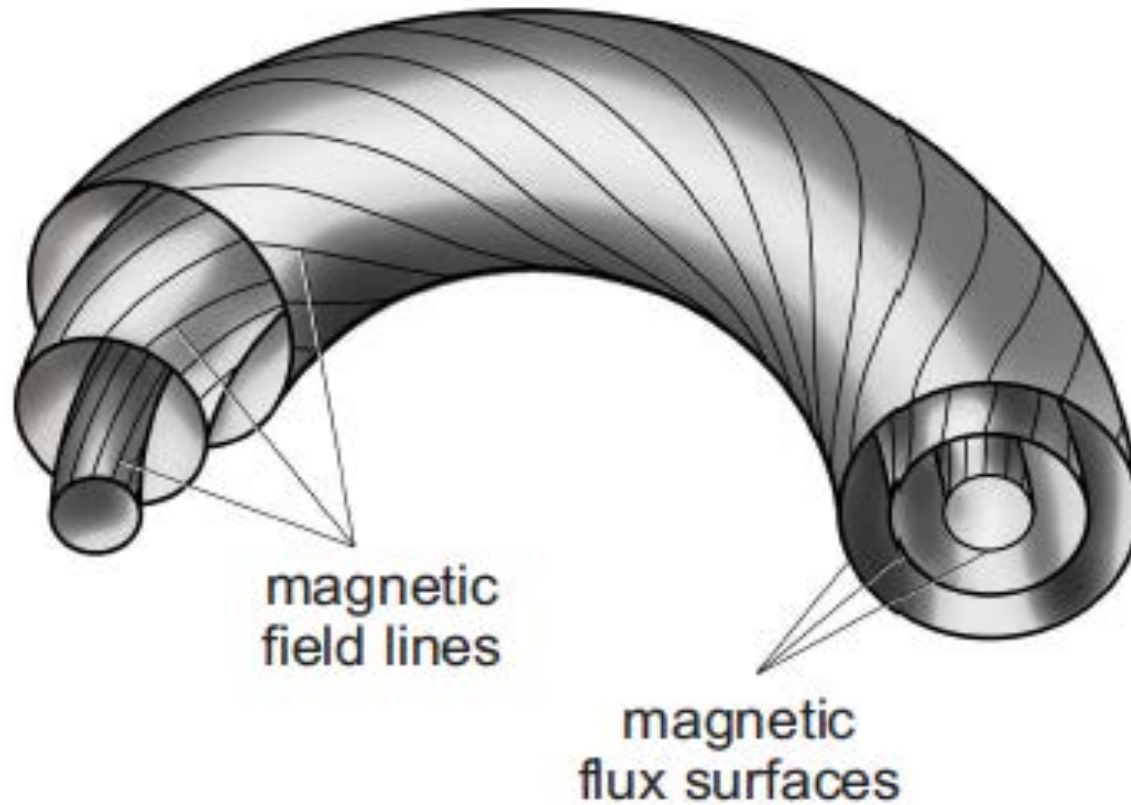
Pure toroidal magnetic field $\rightarrow$ charge separation via ∇B drift and curvature drift lead to a vertical E-field.

$E \times B$ drift would drive all charged particles towards the outboard side

$\rightarrow$ Need helical magnetic field structure



Plasma can be confined in a magnetic field

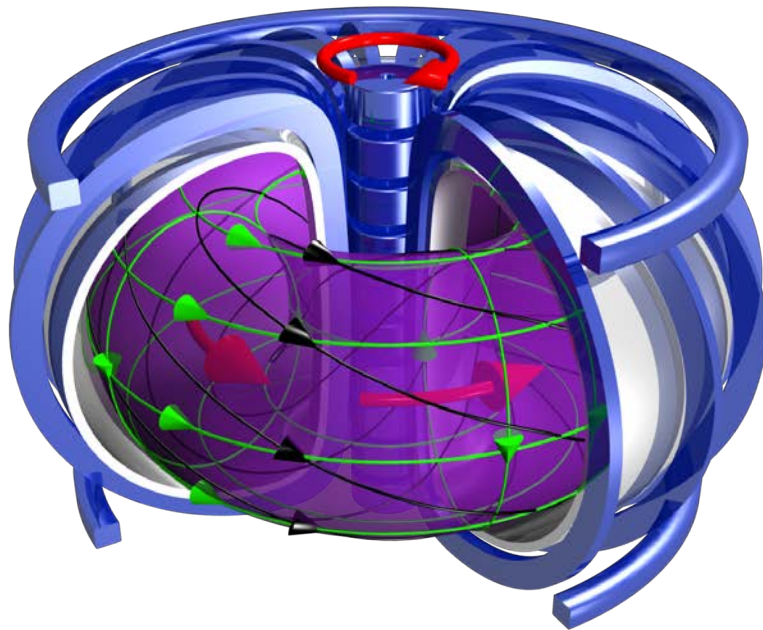


Toroidal systems avoid end losses along magnetic field
⇒ Need to twist field lines helically to compensate particle drifts



Concepts of magnetic confinement

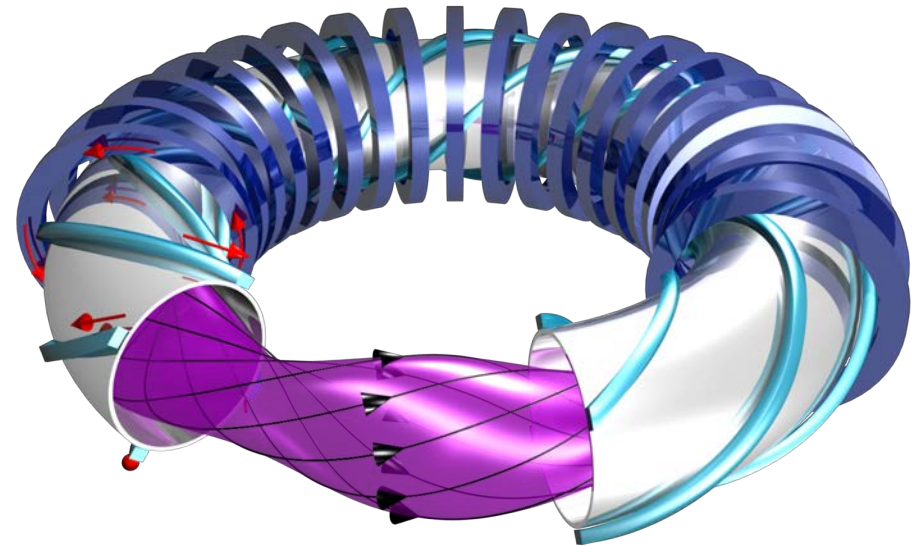
Tokamak (axis symmetric)



Magnetic field by external coils
and plasma current

pulsed

Stellarator (3D)

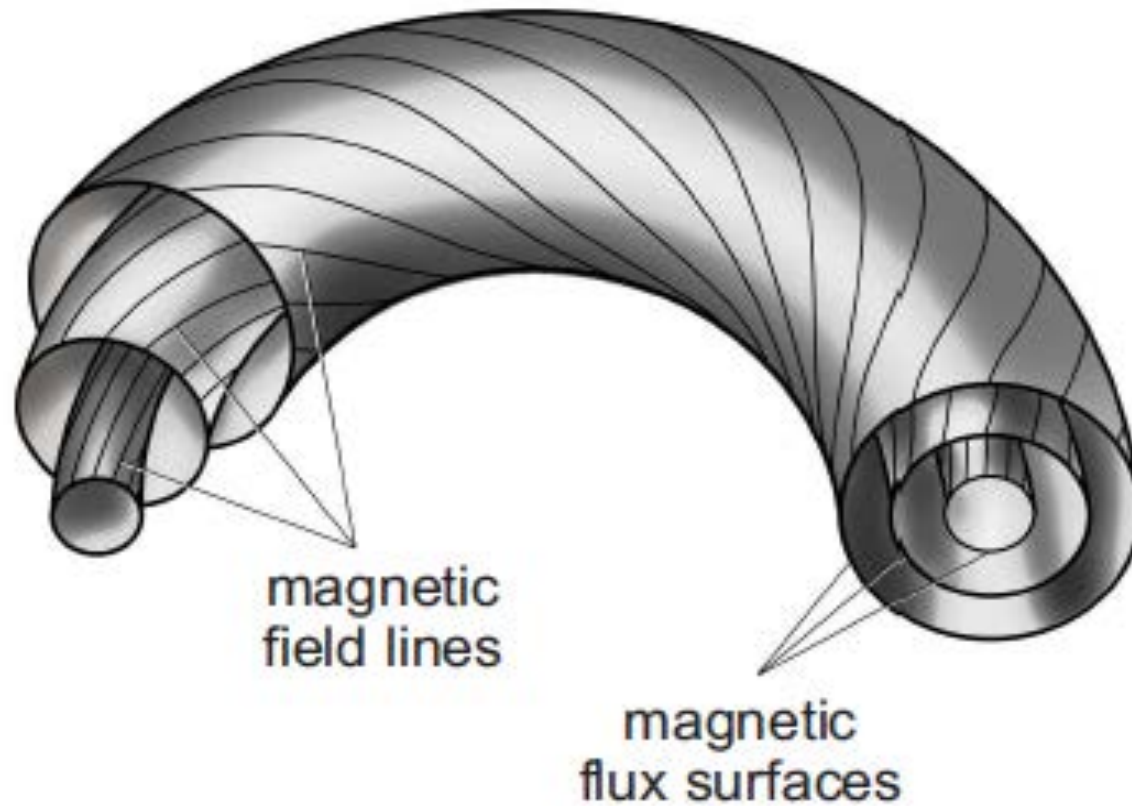


Magnetic field only by external coils

intrinsically stationary



Field line geometry described by safety factor q



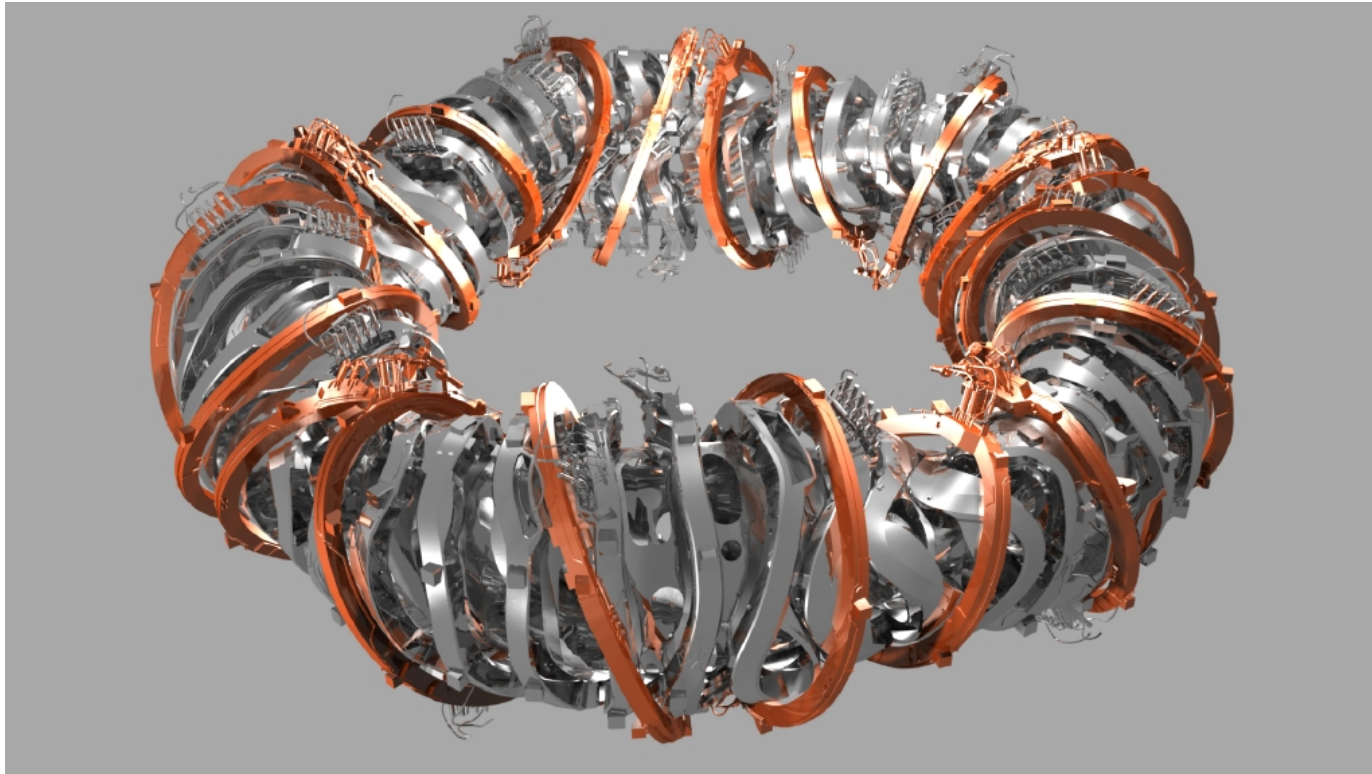
$$q = \frac{\text{number of toroidal windings}}{\text{number of poloidal windings}}$$



Plasma can be confined in a magnetic field



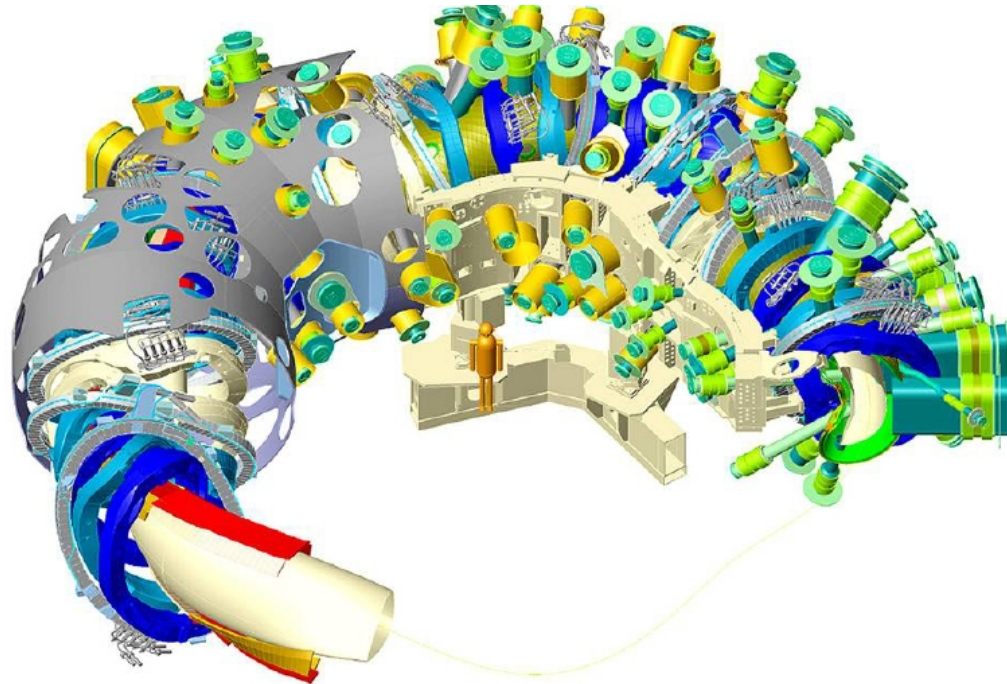
'Stellarator': magnetic field exclusively produced by coils



Example: Wendelstein 7-X (IPP Greifswald)



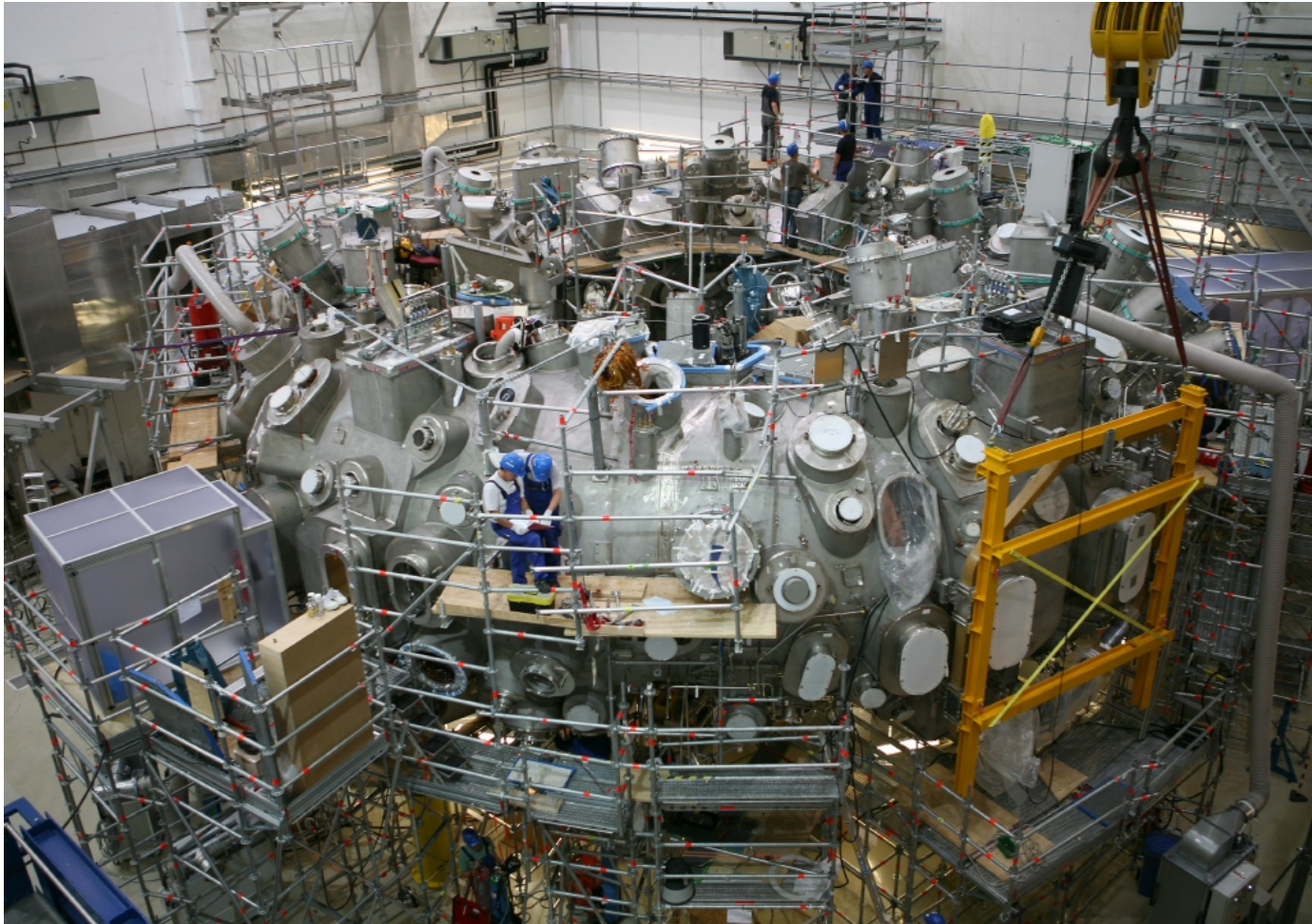
Stellarator (I)



CAD design of Wendelstein 7-X including ports
and magnetic flux surfaces



Stellarator (II)



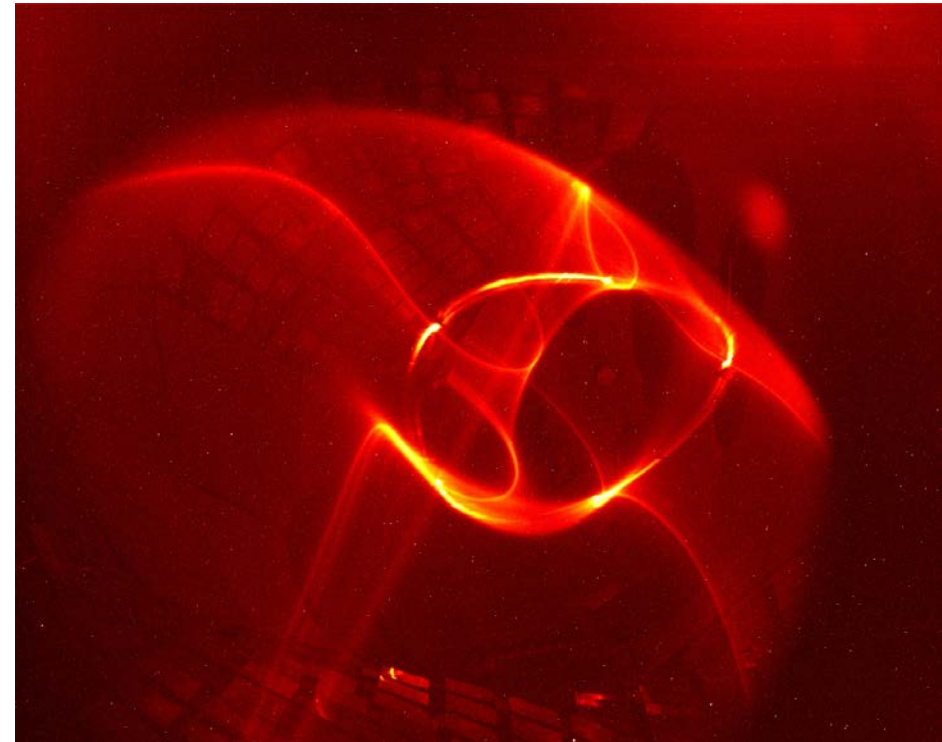
Wendelstein 7-X, May 2014



Magnetic field tests at full field successful

temperature

12.5 kA



First plasma in He in December 2015

First plasma in H in February 2016

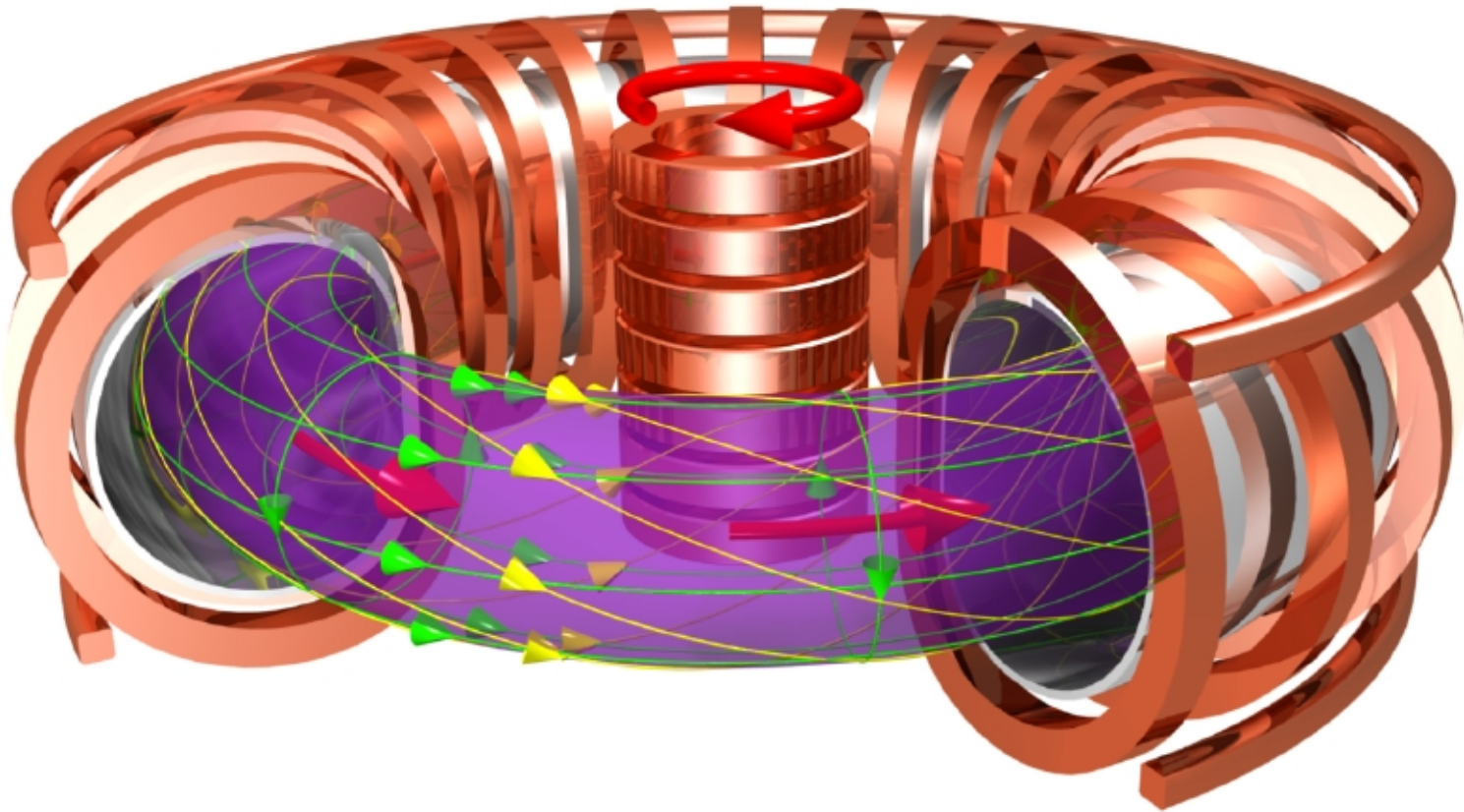
➔ Currently under maintenance



Plasma can be confined in a magnetic field



'Tokamak': poloidal field component from current in plasma

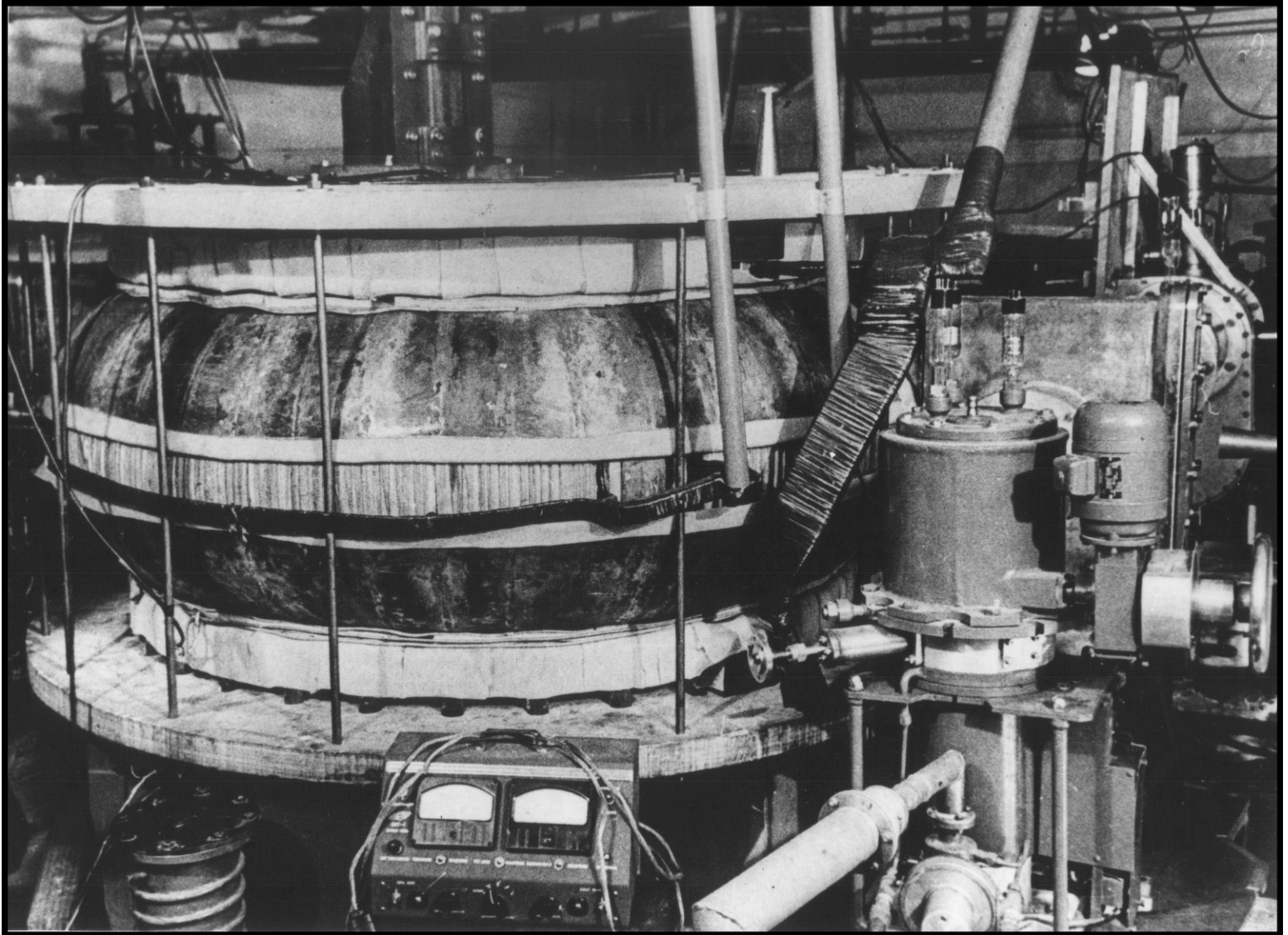


Simple concept, but not inherently stationary!

Example: ASDEX Upgrade (IPP Garching)

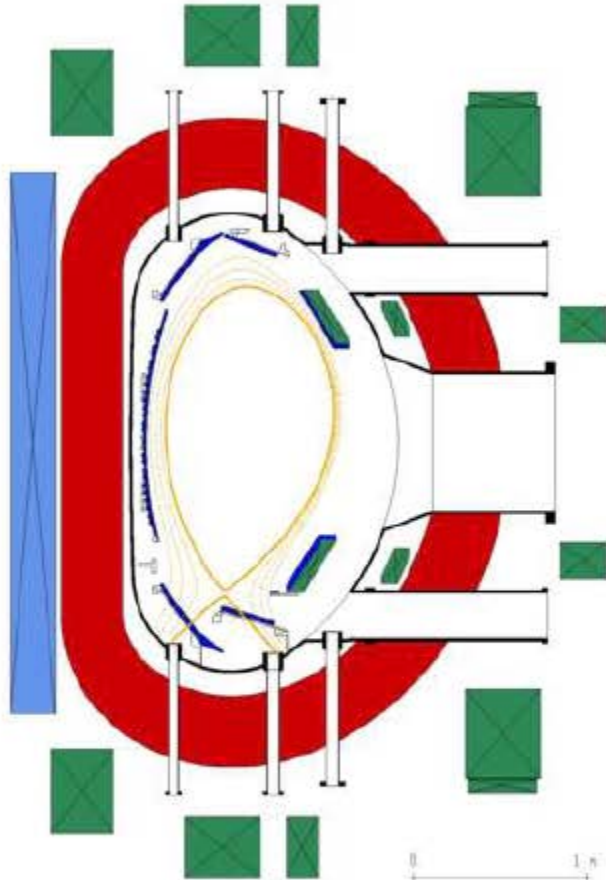


Tokamak T1

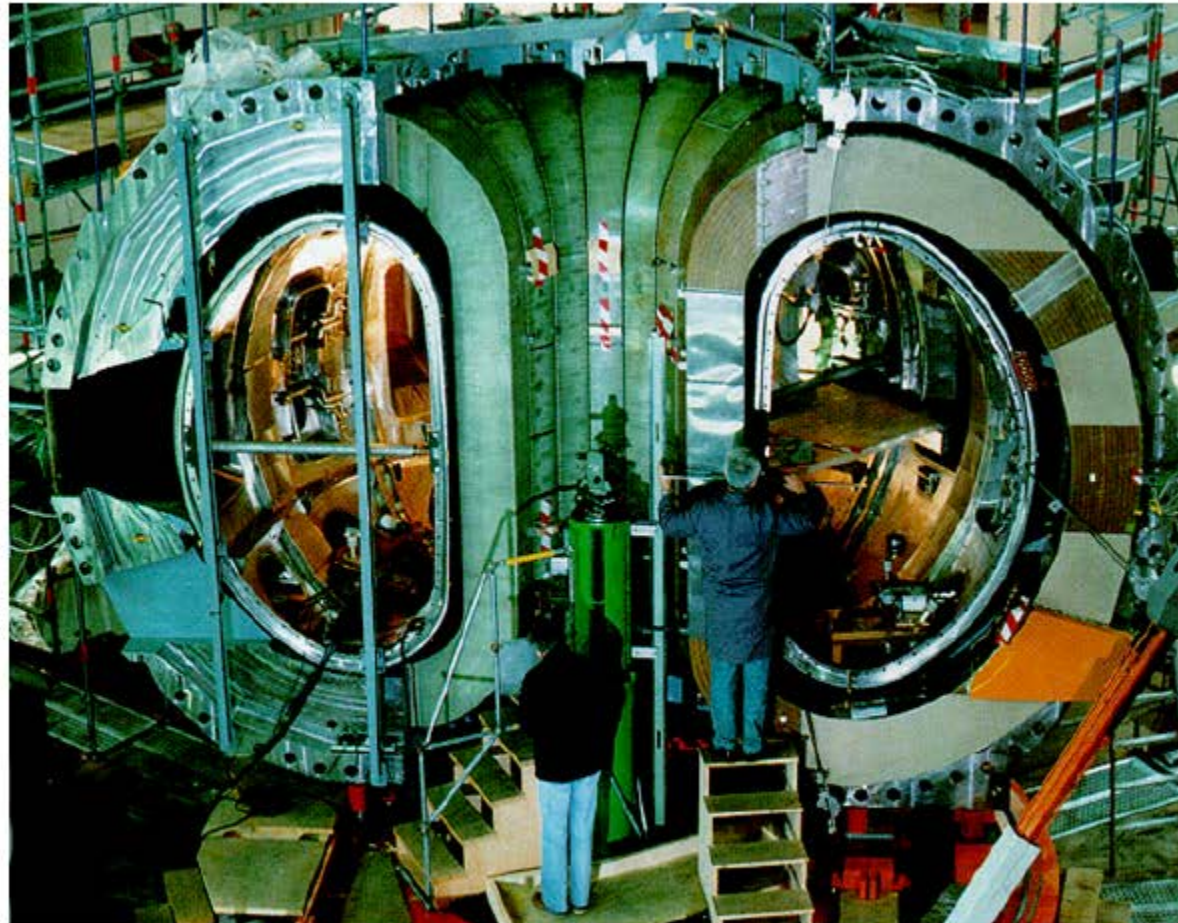




ASDEX Upgrade – Largest German tokamak



Großer Radius = 1.65 m, $B_t \leq 3,5$ T
Kleiner Radius = 0.5 m, $I_p \leq 1,4$ MA
 $P_{\text{Heiz}} \leq 28$ MW,

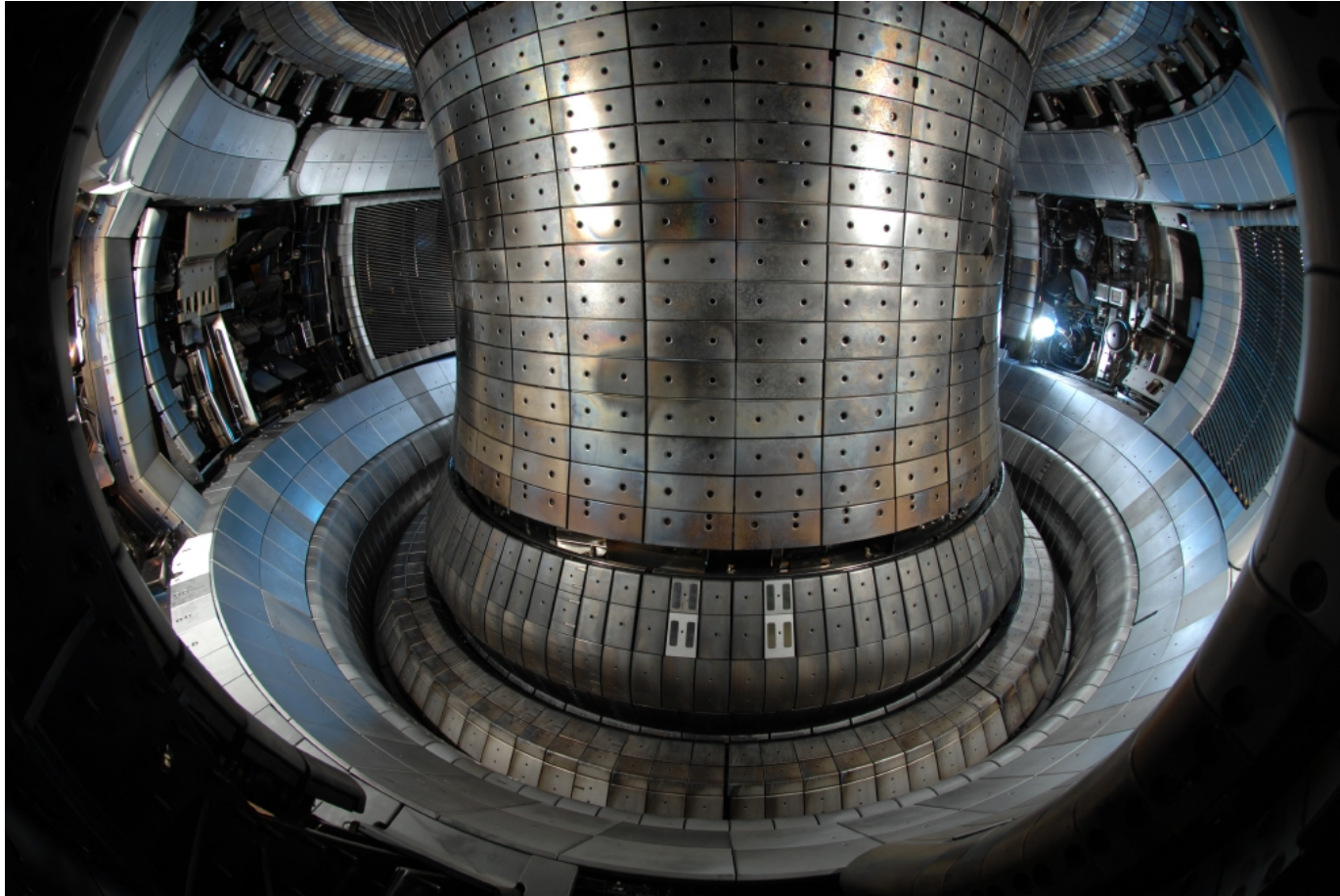


Start des Betriebs in 1991;
hier während der Konstruktion in 1989



Plasma can be confined in a magnetic field

'Tokamak': poloidal field component from current in plasma

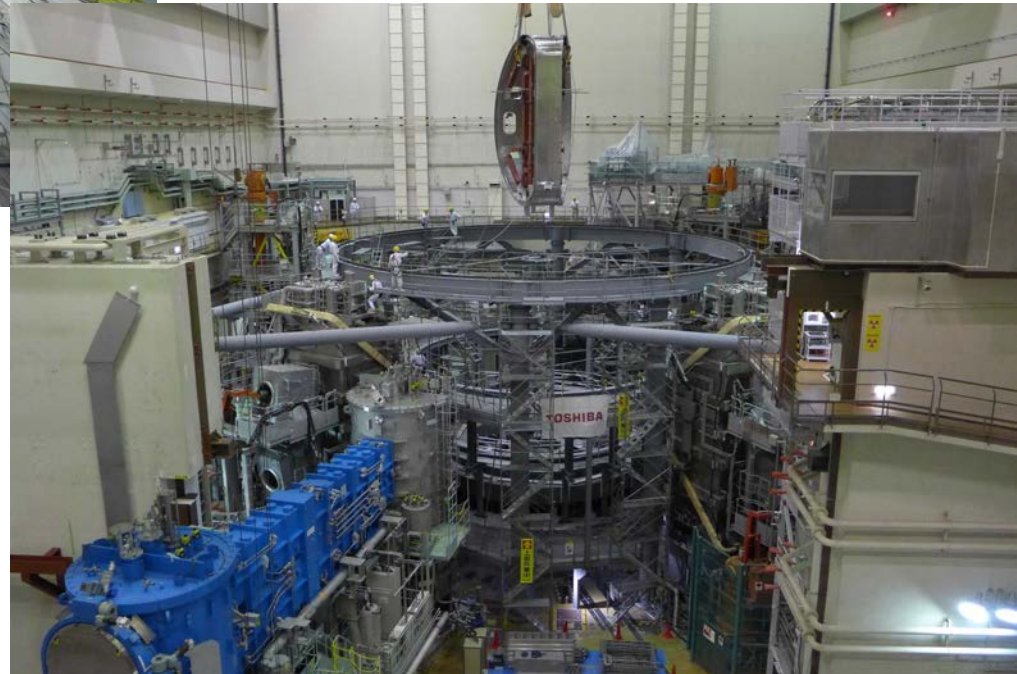


Simple concept, but not inherently stationary!

Example: ASDEX Upgrade (IPP Garching)

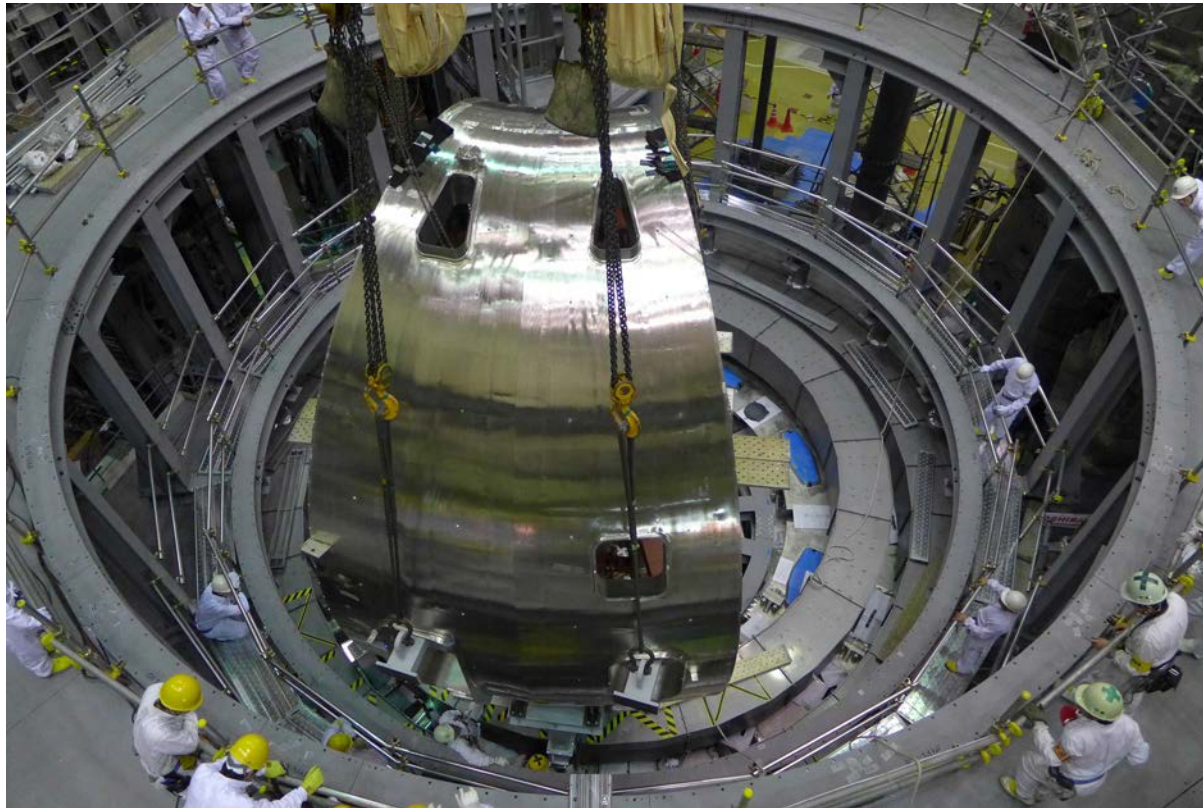


Tokamak under construction – JT60-SA Japan





Tokamak under construction (II) – JT60-SA Japan



Spring 2014



Tokamak under construction (III) – JT60-SA Japan

IPP

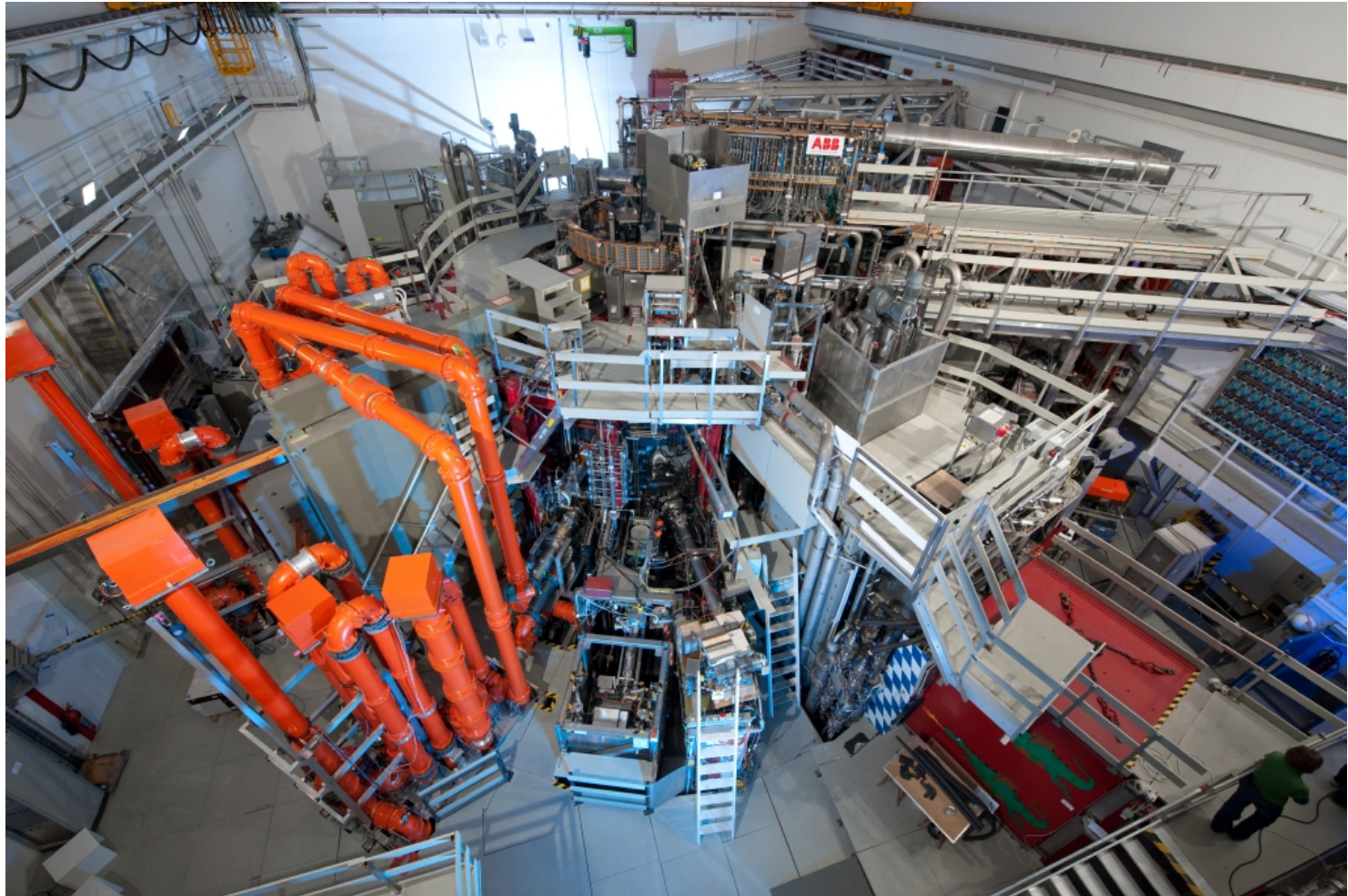


Spring 2016



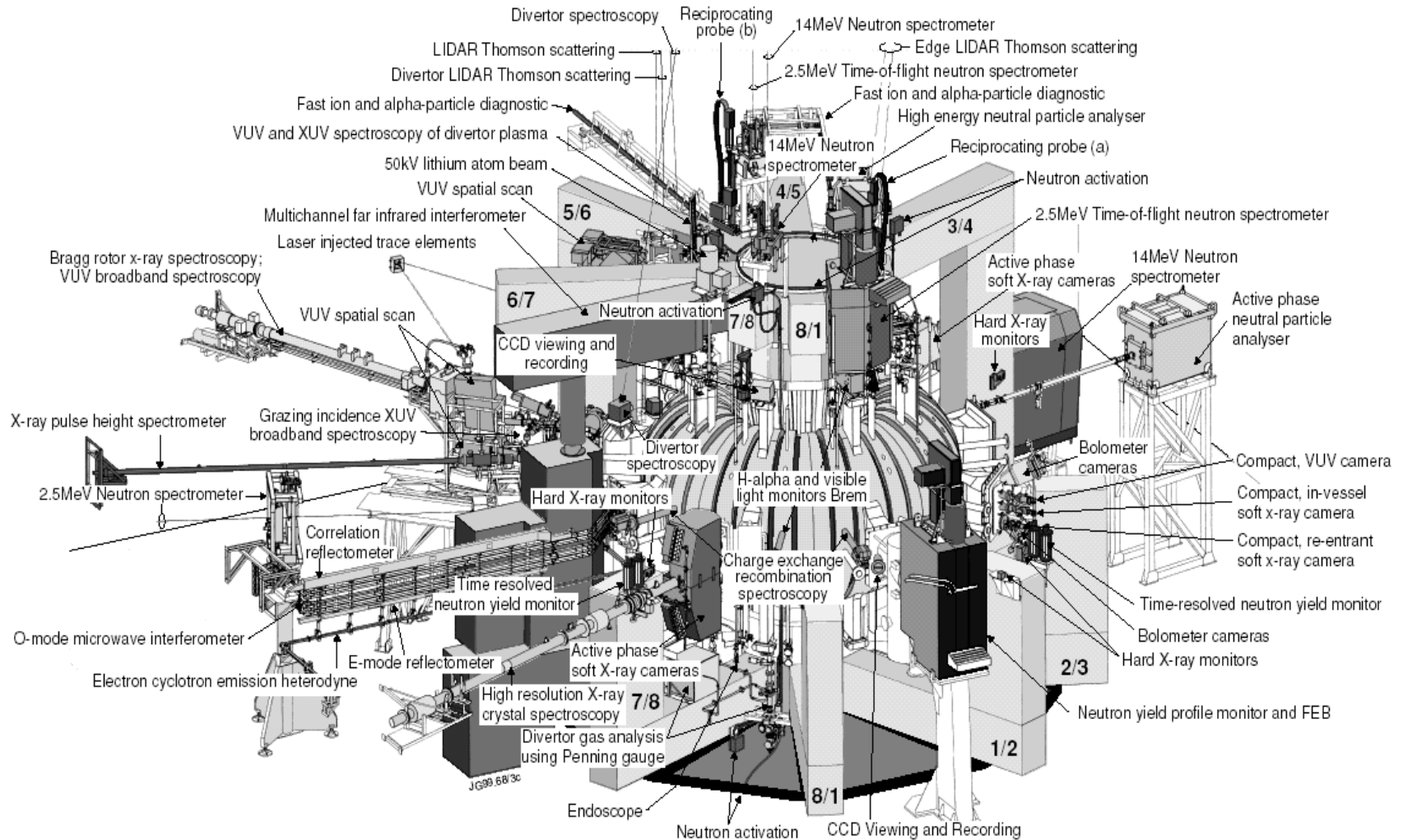


ASDEX Upgrade viewed inside torus hall



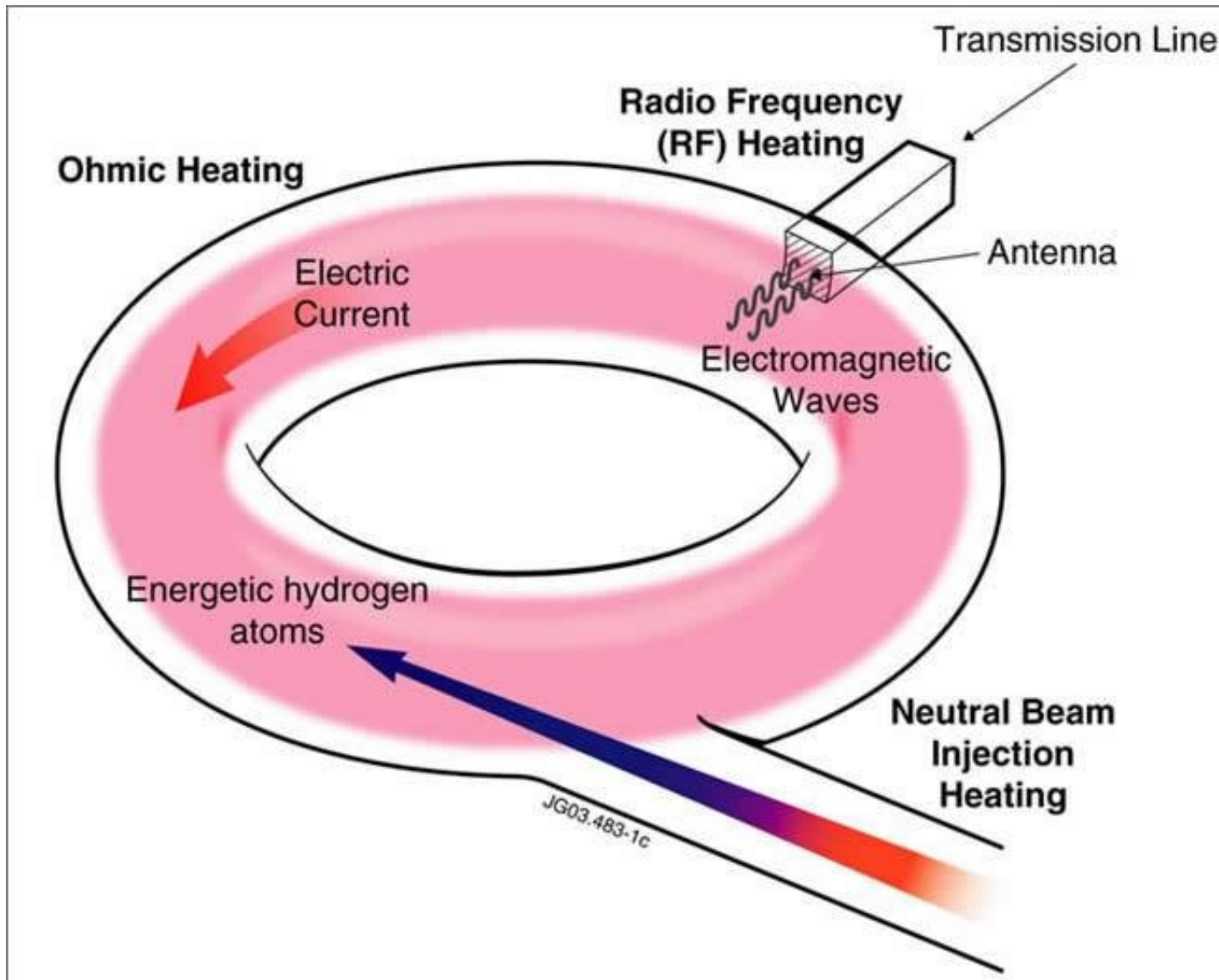


Diagnostics of Fusion Plasmas: JET-Design





Tokamak plasma heating systems





Ohmic heating (only tokamaks) :

$$P_{OH} = I_p^2 \cdot R_p \quad \text{mit} \quad R_p \sim 1/T^{3/2}$$

→ Very inefficient at high T

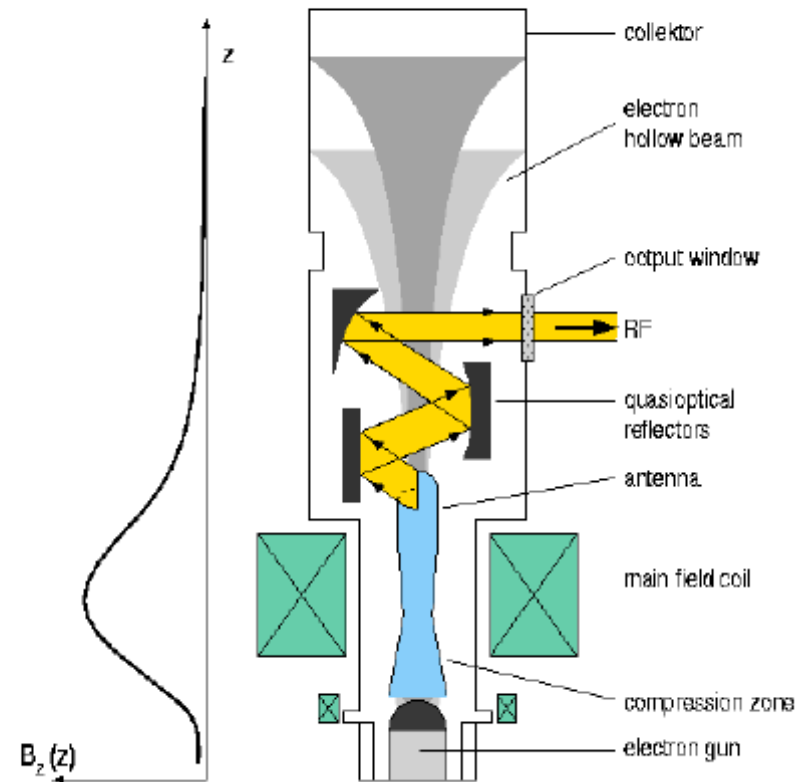
RF heating:

Injecting waves at cyclotron frequency

$$\omega_C = qB/m$$

Ionen (ICRH): < 100 MHz

Elektr. (ECRH): < 180 GHz



Schematic of a high frequency amplifier

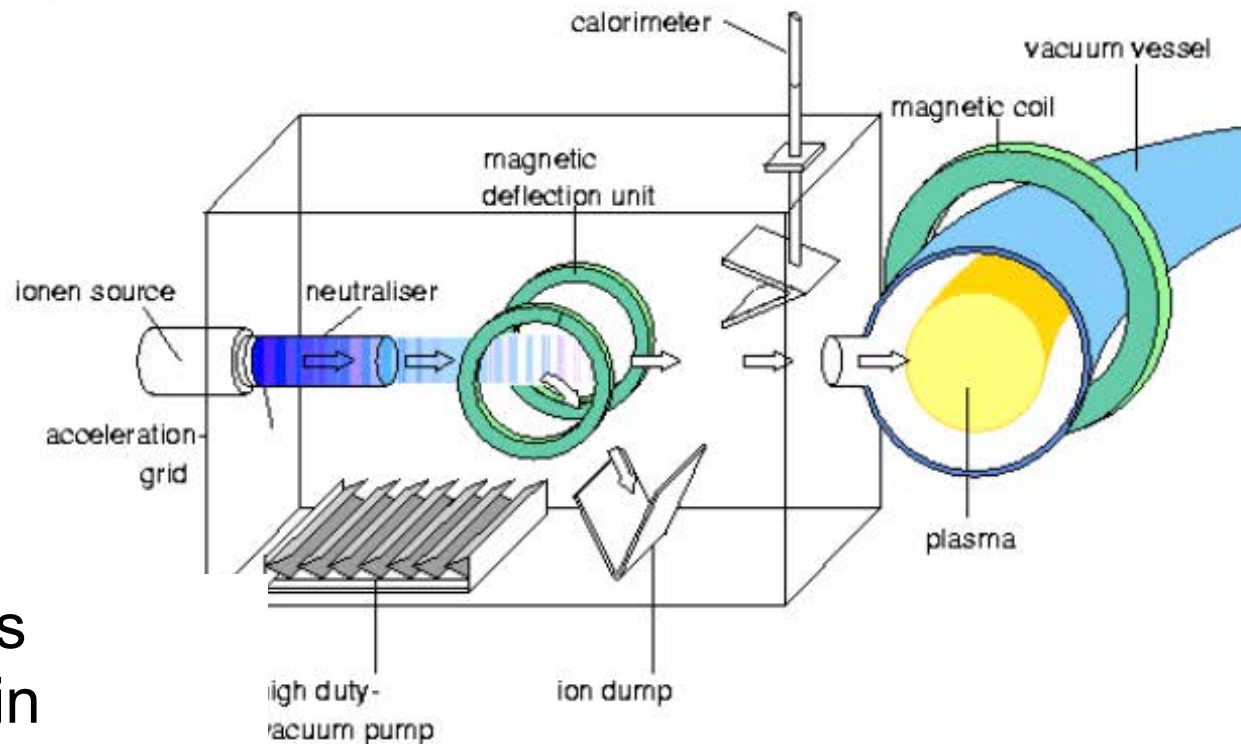
Gyrotron ($f \approx 80 - 180$ GHz)



Methods for plasma heating

Neutral beam injection

$$P_{\text{NBI}} = U_B \cdot I_B$$



$$E_{\text{beam}} \gg 100 \text{ keV}$$

- Acceleration of ions
- Charge exchange in neutral gas
- Negative ions necessary for



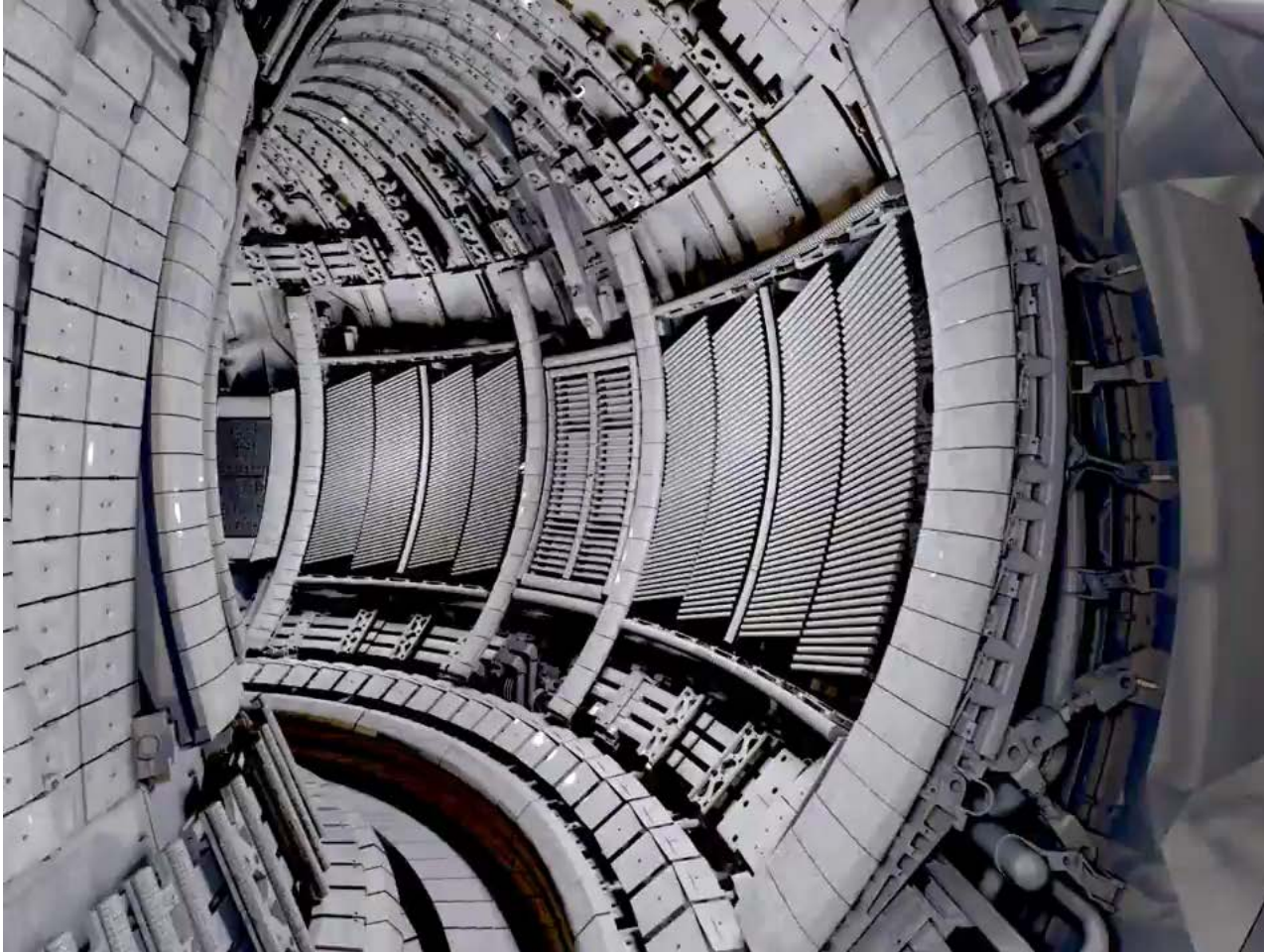
Periphery of non super conducting tokamak: Power supply

IPP





Inside the torus of the JET tokamak





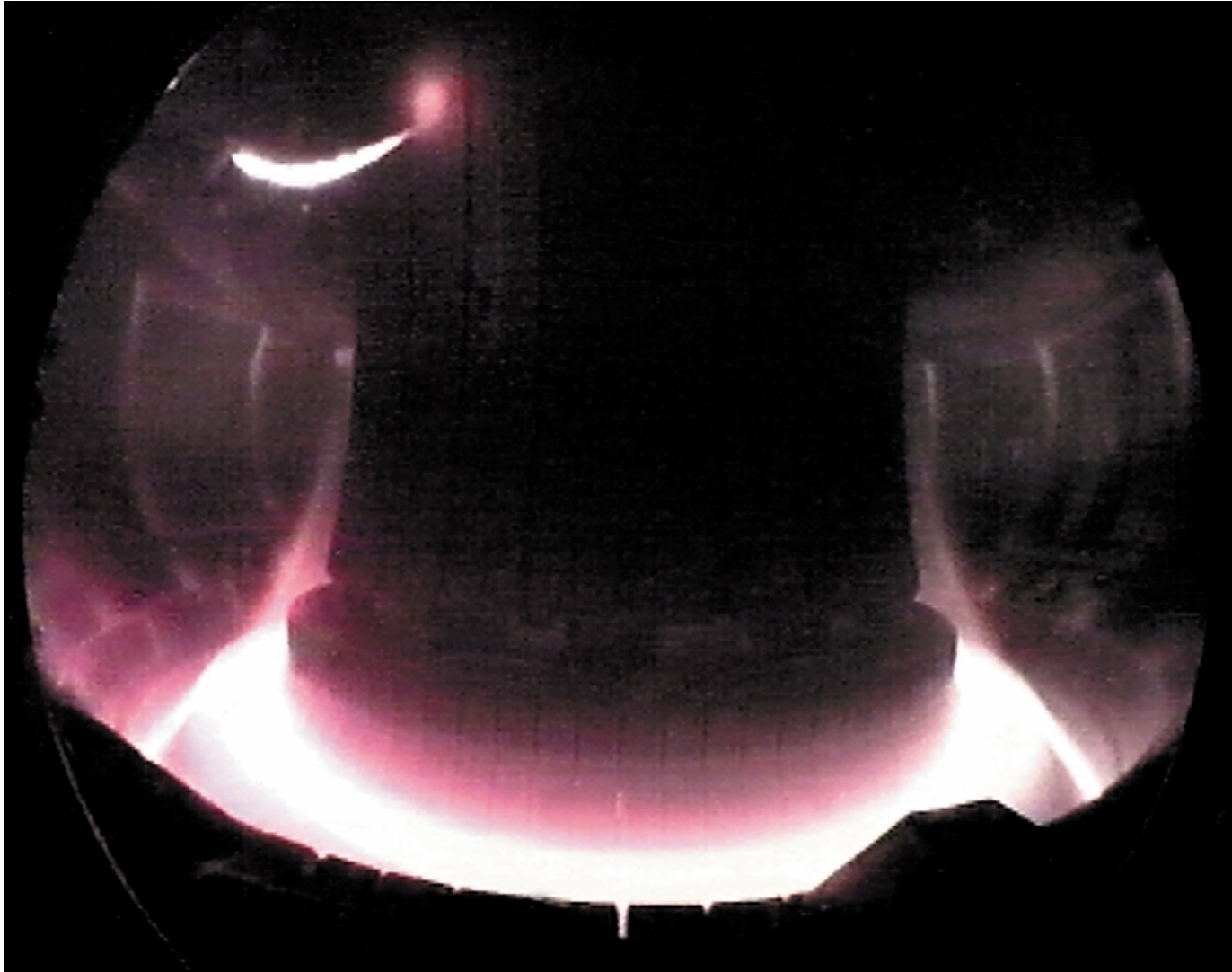
A discharge at JET (movie)



A discharge at JET with the ILW – can find it on youtube



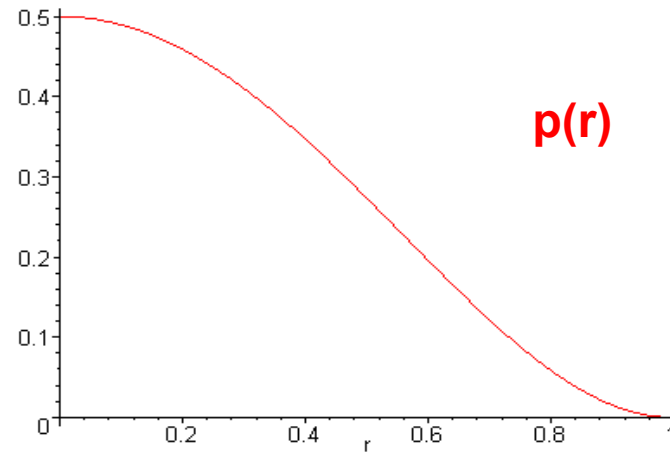
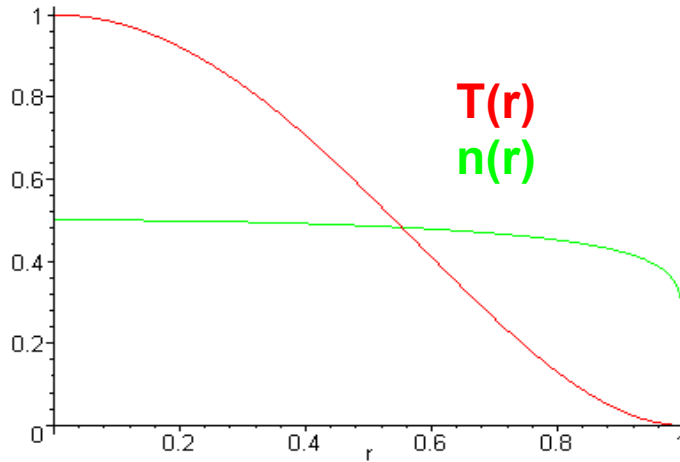
Plasma can be confined in a magnetic field



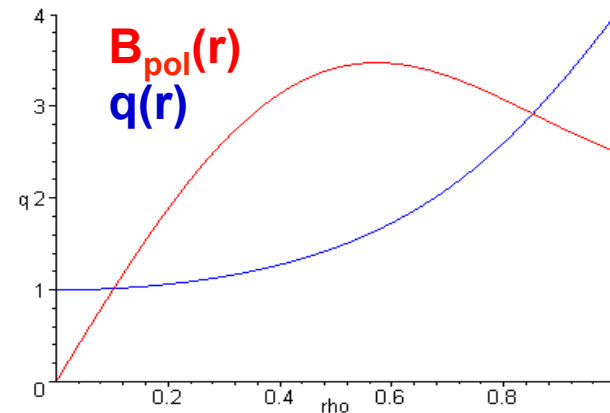
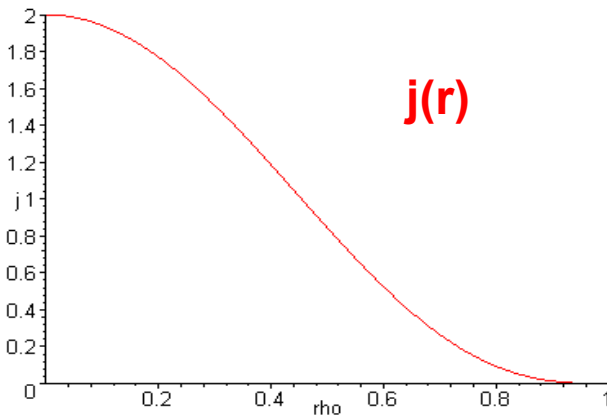
ASDEX Upgrade (with trace of a pellet)



Typical radial profiles in fusion experiments (schematic)



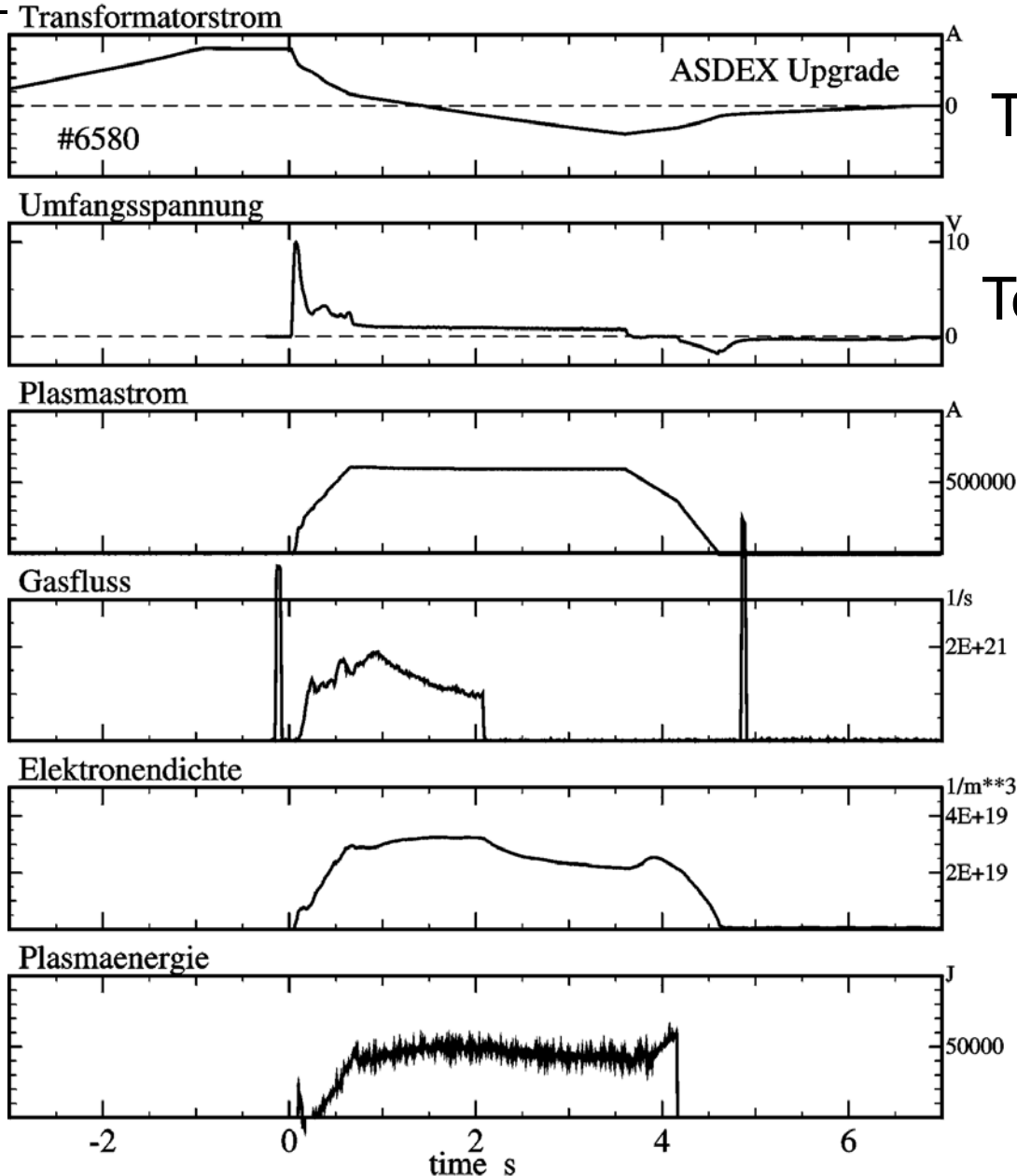
Temperature profiles peaked on axis, density usually flatter, $p = n k_B T$



Tokamak current profiles peaked on axis ($\sigma \sim T^{3/2}$ is highest) – B_{pol} and safety factor q increase from centre to the edge



Typical evolution of a tokamak discharge



Transformer Current

Toroidal loop voltage

Plasma current

Gas inlet

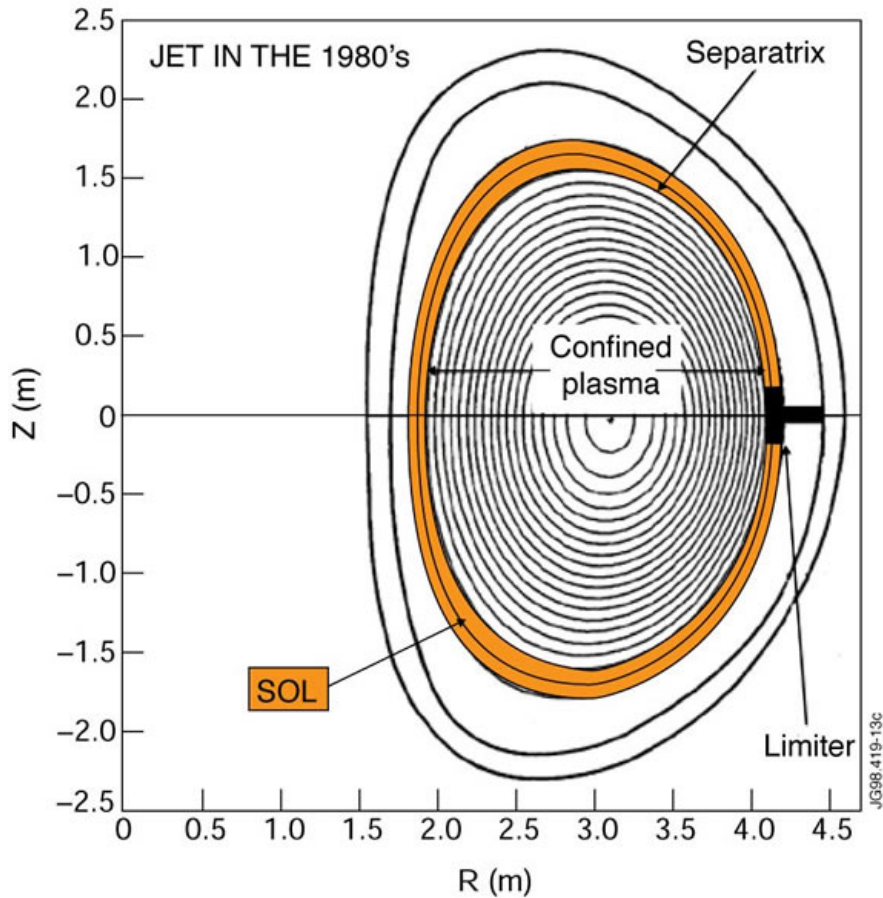
Electron density

Plasma stored energy

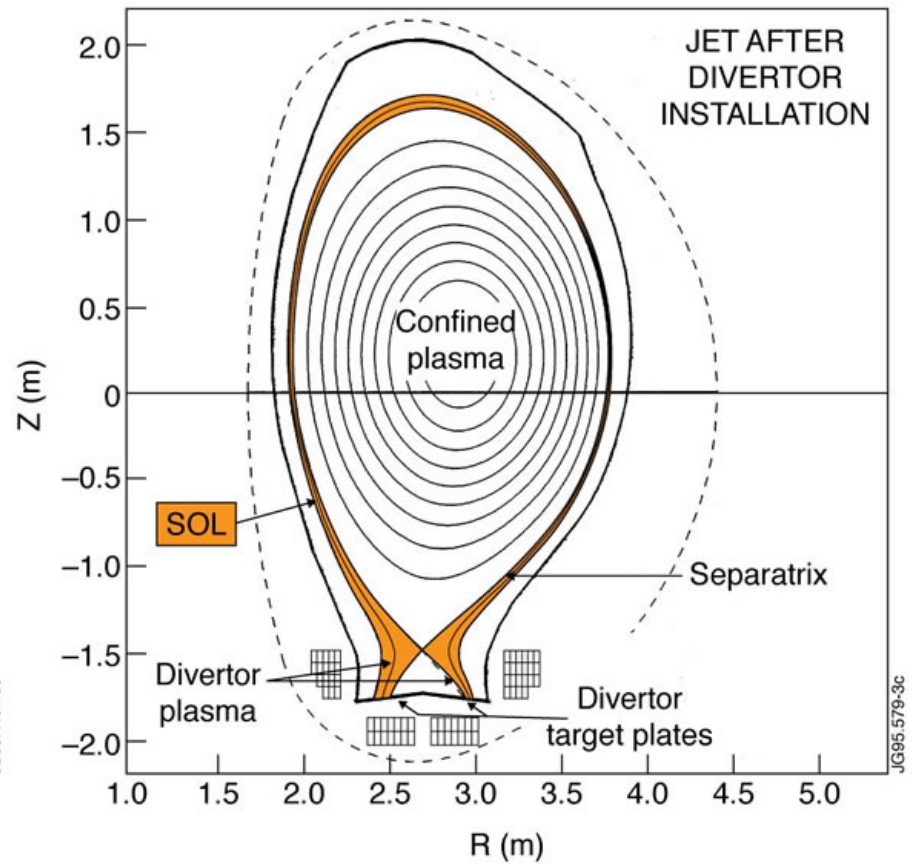


Tokamak configurations

Limiter



Divertor

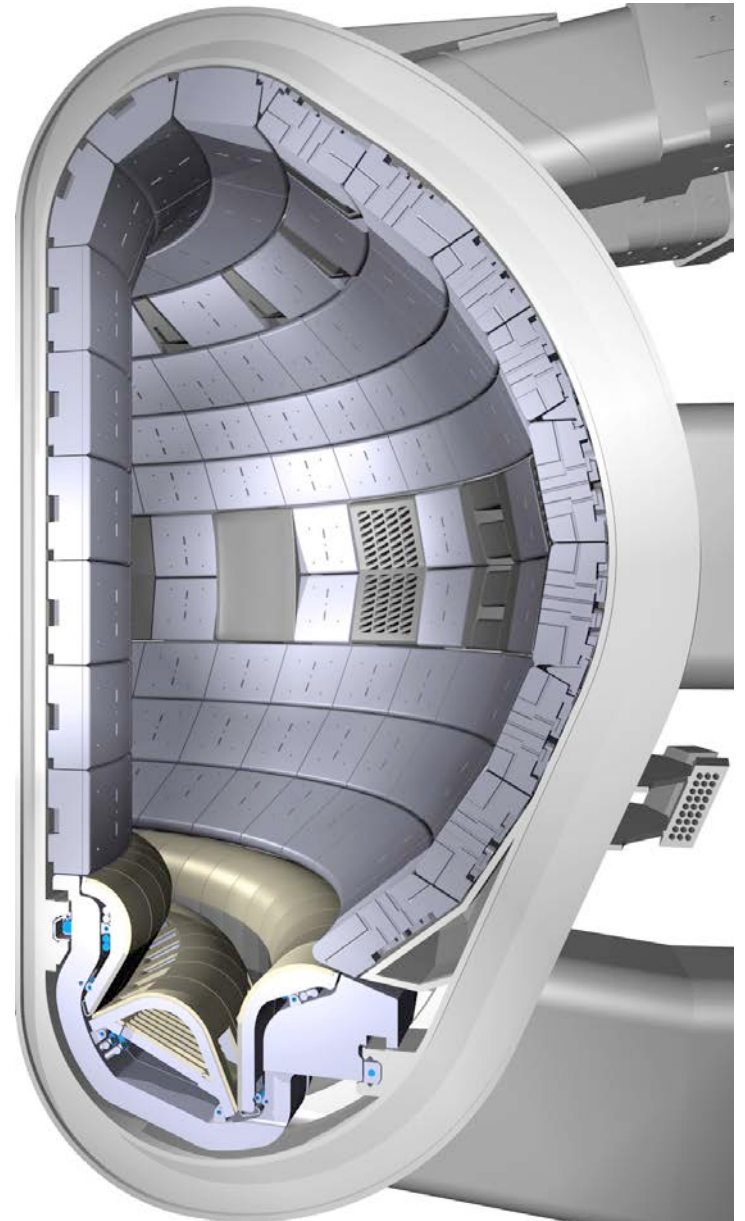




ITER divertor



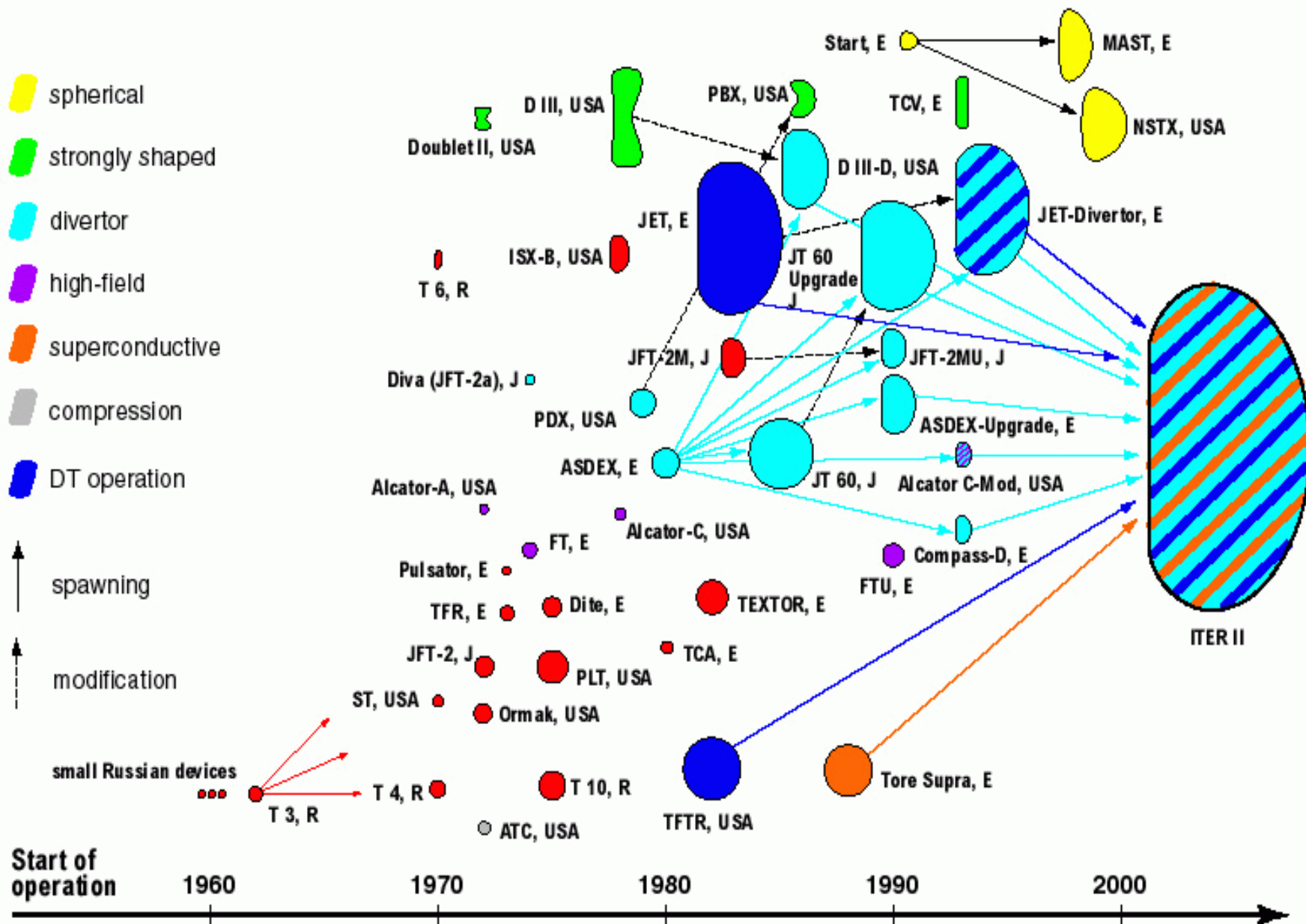
www.iter.org





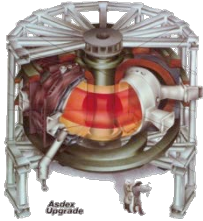
Tokamak Evolution

IPP

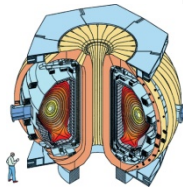




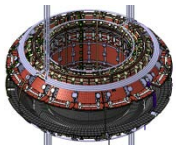
Tokamak experiments worldwide



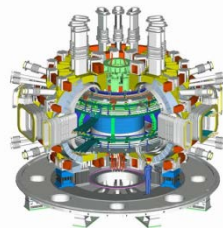
ASDEX
Upgrade
Garching (D)



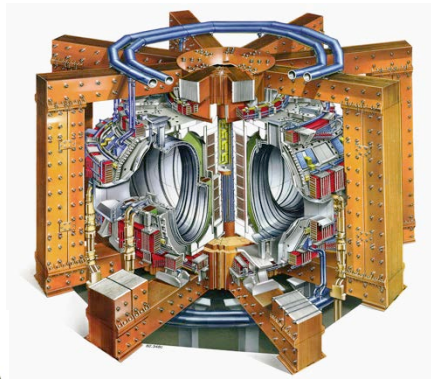
DIII-D
San Diego
(USA)



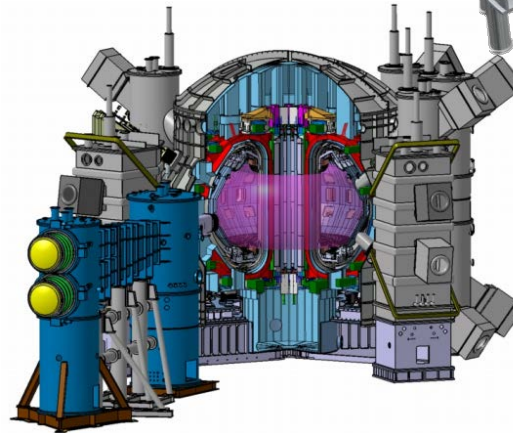
EAST
Chengdu (C)



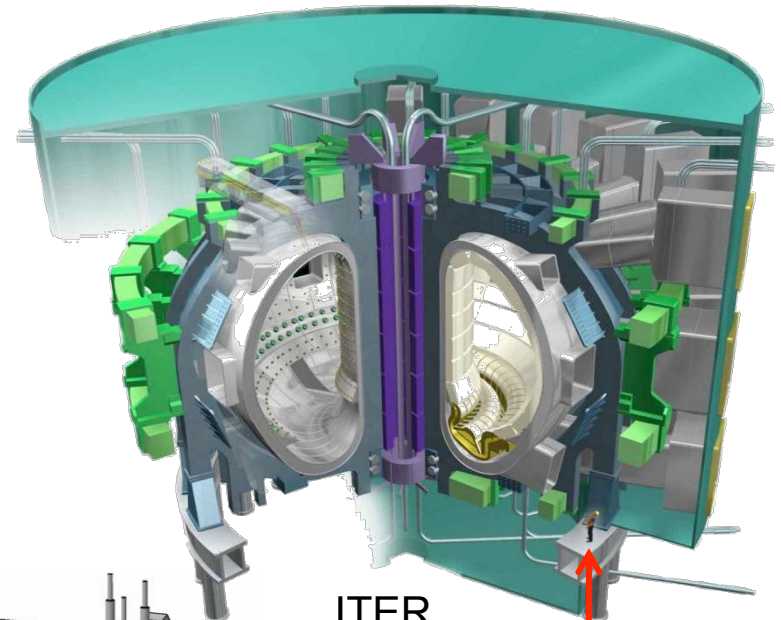
KSTAR
Daejeon (KR)



JET
Culham (GB)



JT-60SA
Naka (JA)



ITER
Cadarache (F)

person



Planing ITER based on existing devices

IPP

Extrapolation towards ITER like in
wind tunnel experiments

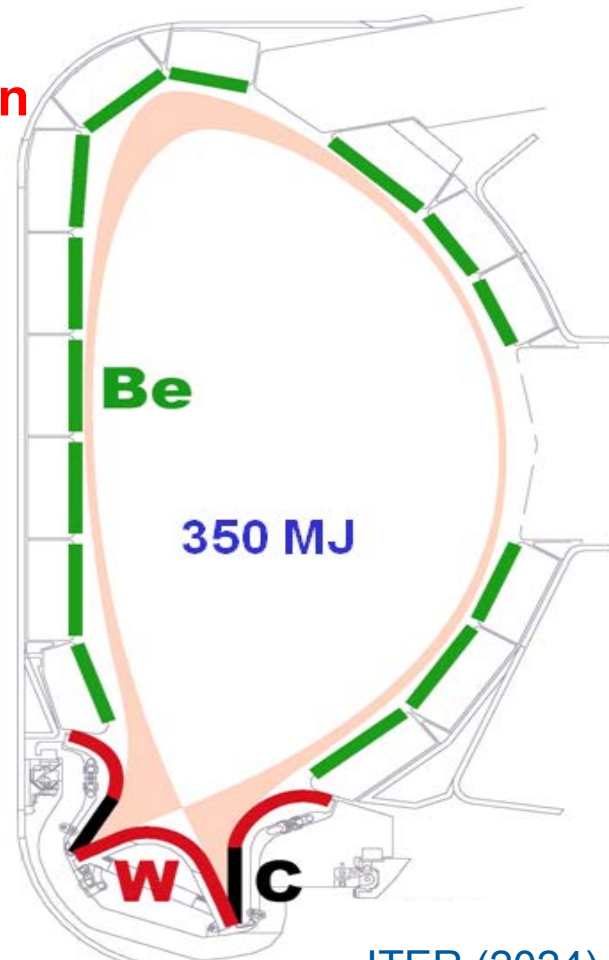
1 m



ASDEX Upgrade
(Germany)



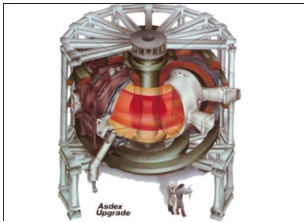
JET
(EU)



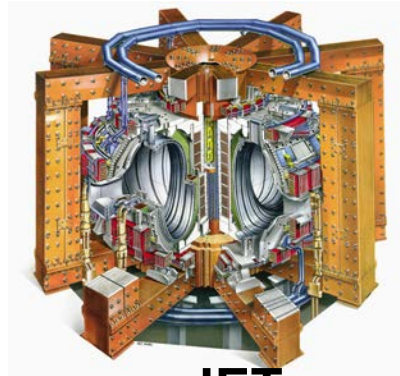
ITER (2024)
(international)



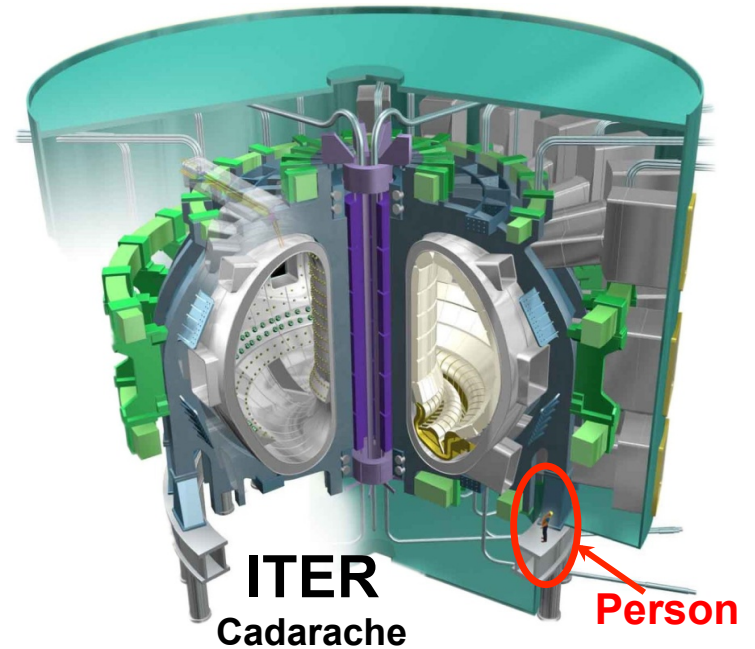
Tokamaks



ASDEX Upgrade
Garching



JET
Culham



ITER
Cadarache

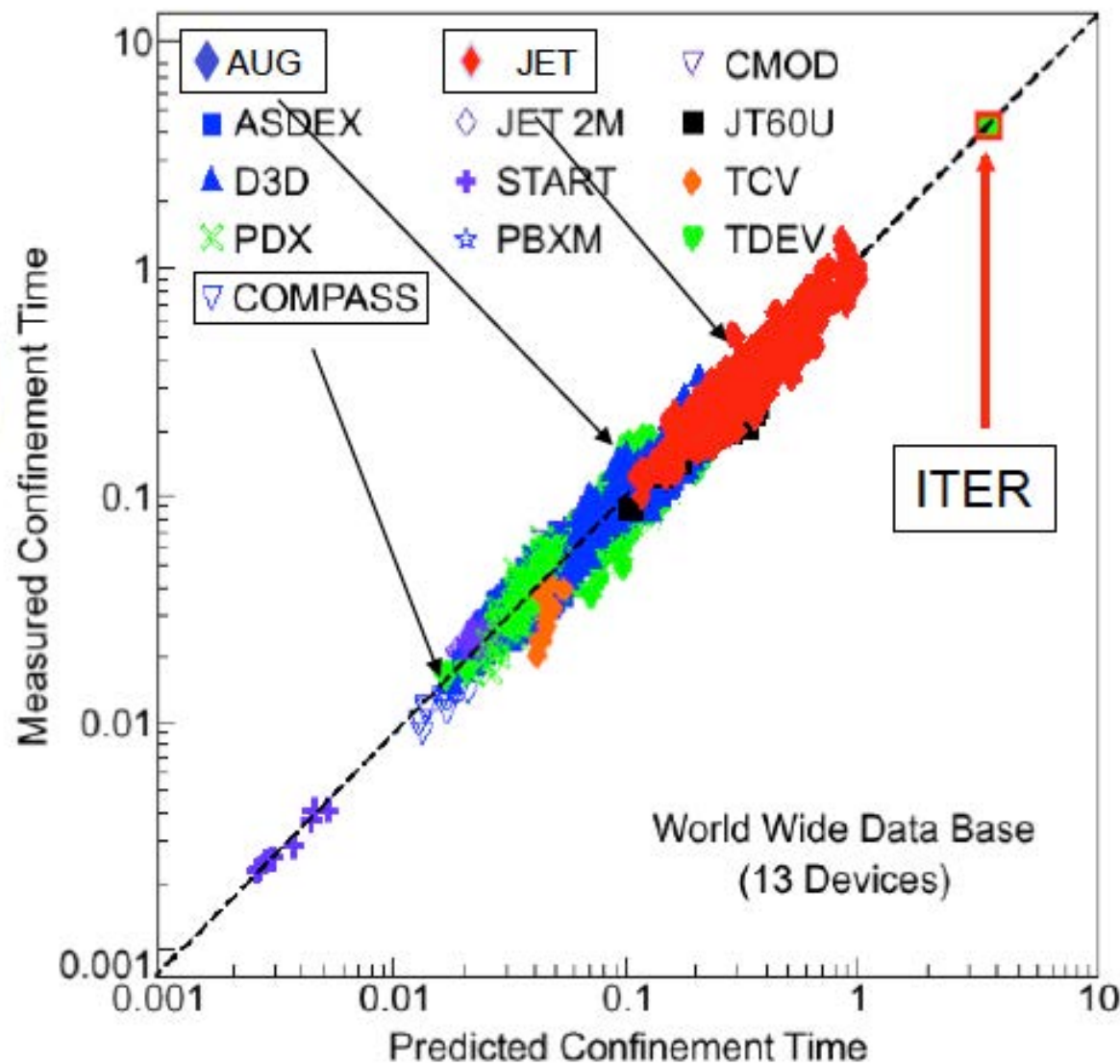


Extrapolation der Einschlusszeit zu ITER

Erreichte Kennzahlen:

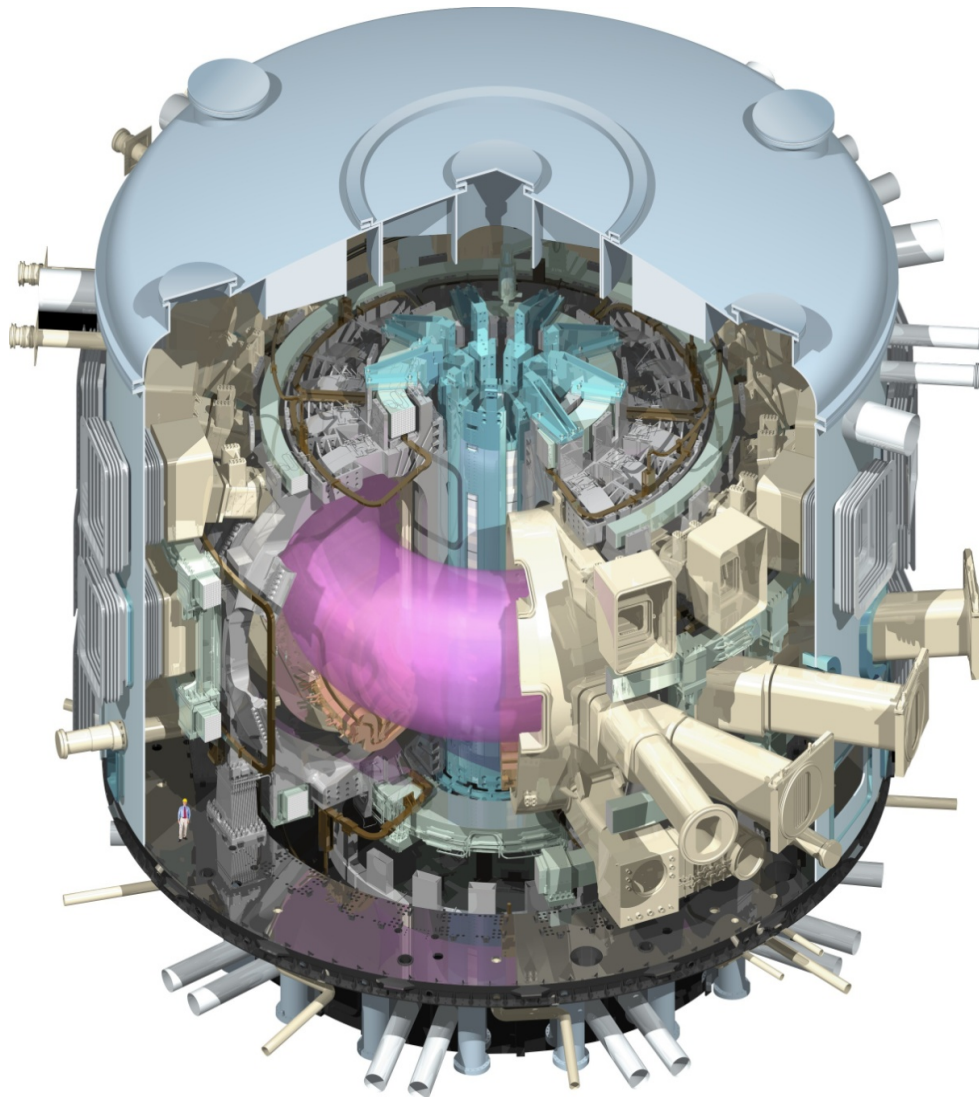
- Temperatur T 400 Mio. °C ✓
- Dichte n 10^{20} m^{-3} ✓
- Energie Einschlusszeit $\tau_E \sim 1 \text{ s}$,
ist noch zu klein
(etwa 4s benötigt)

ITER:
Due to larger volumes
and higher confinement
time $Q=10$ is expected





The ITER Design



	ITER
Major Radius	6.2 m
Minor Radius	2.0 m
Plasma current	15 MA
Magnetic field	5.3 T (Supercond.)
Power amplification Q	≥ 10
Fusion power	400 (800)MW
Duration of burn	400 (3000) s
External heating	73 (110) MW

Cost: ~ 15 Billion €

Requires world-wide effort

ITER will be built in Cadarache (F) as joint effort – Cn, EU, In, Jp, Ko, RF, US



ITER site – artists vision





ITER site – artists vision

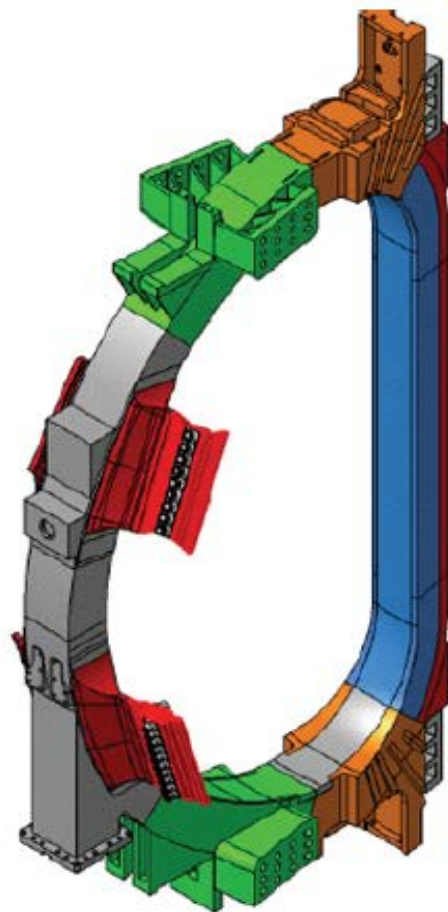


**First Plasma slipped from 2019 – 2025 and
2027/28**

Energy gain $Q = 10$ (~2037 – 2039)

**Demonstrate the physical feasibility of
nuclear fusion by magnetic confinement**

Toroidal field coils



16 x 9 m, **~360 t**
(EU, JP – 18 coils)



Boeing 747-300
(maximum take-off weight **~377 t**)



Example: toroidal field coils

TF Coils – A Worldwide Collaboration





ITER site – aerial view 2014





ITER site 2015 – tokamak pit





ITER site – aerial view 2016





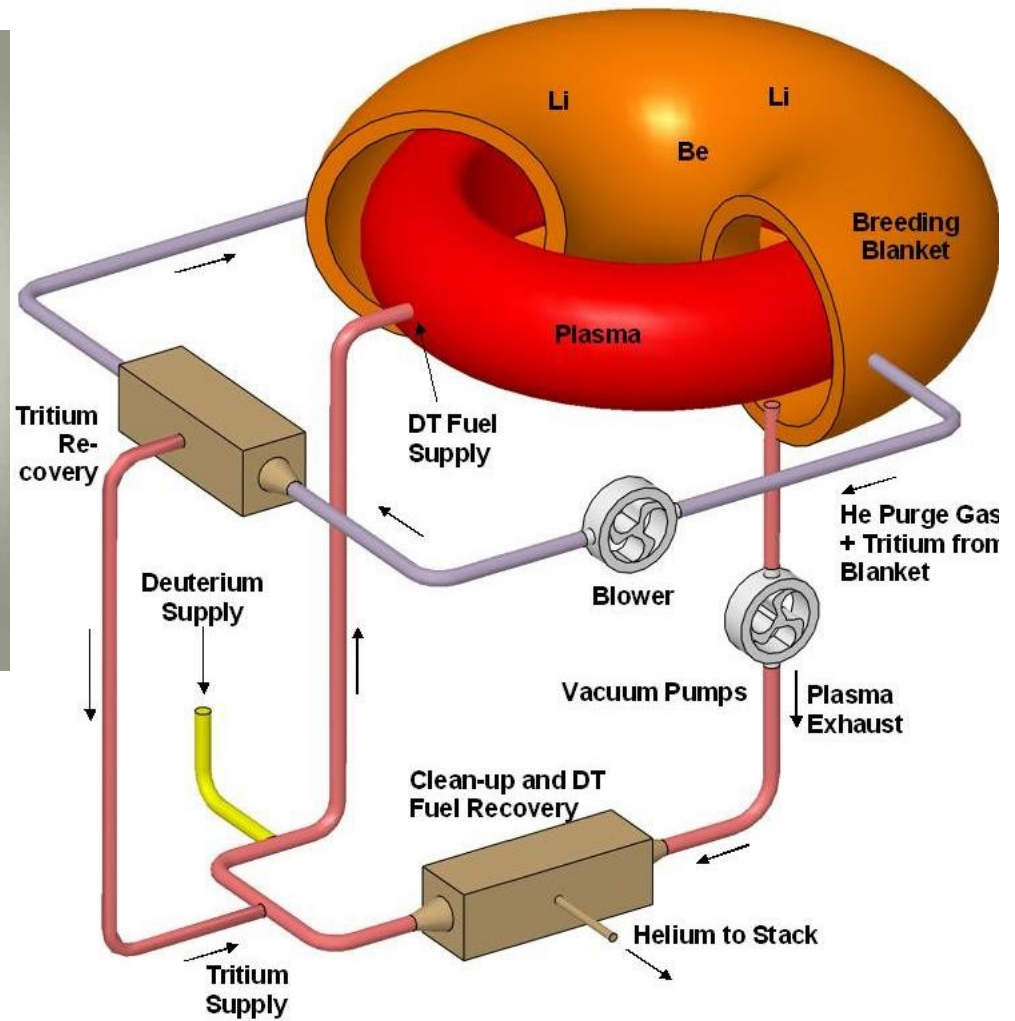
Tritium breeding



Lithium

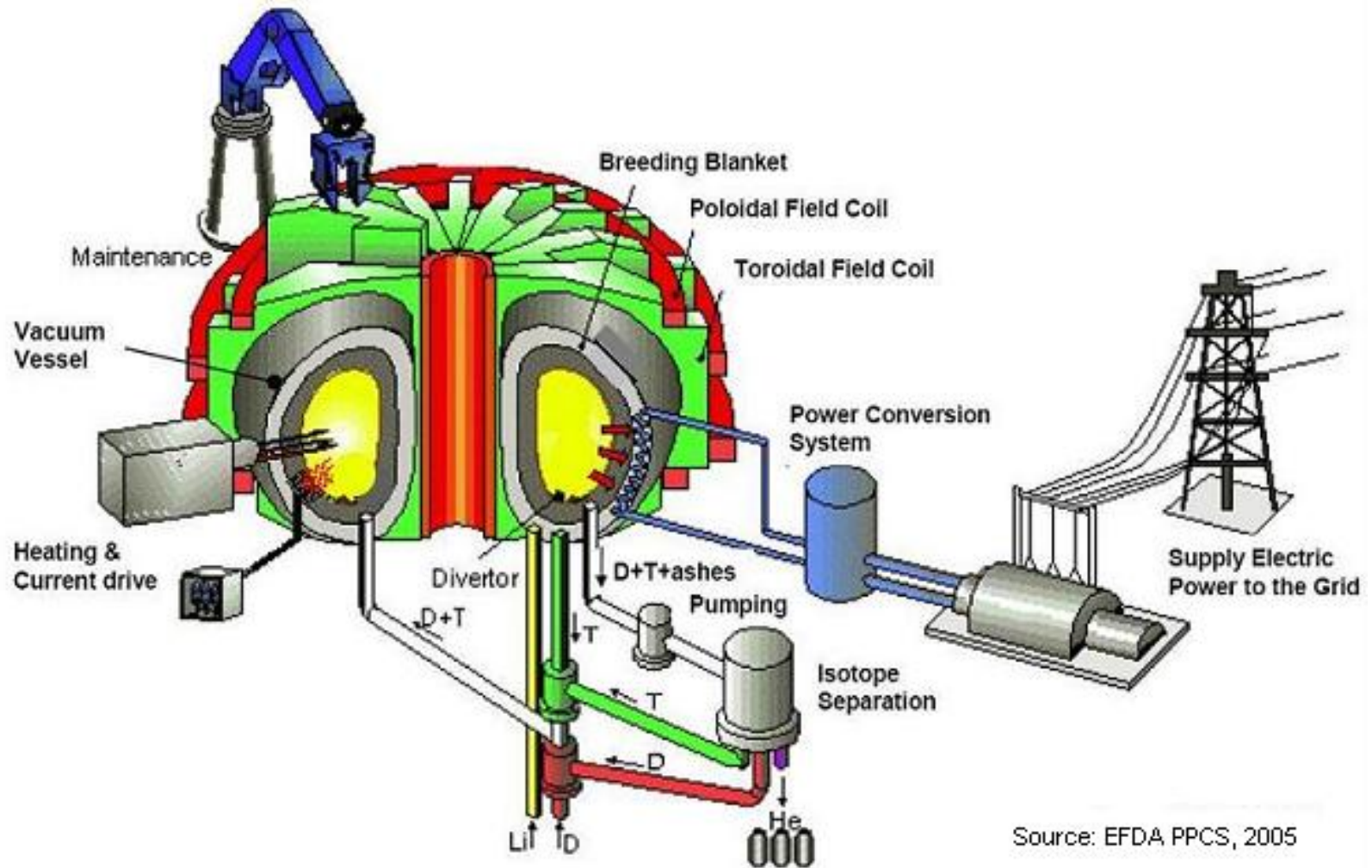


Neutron





Fusion power plant

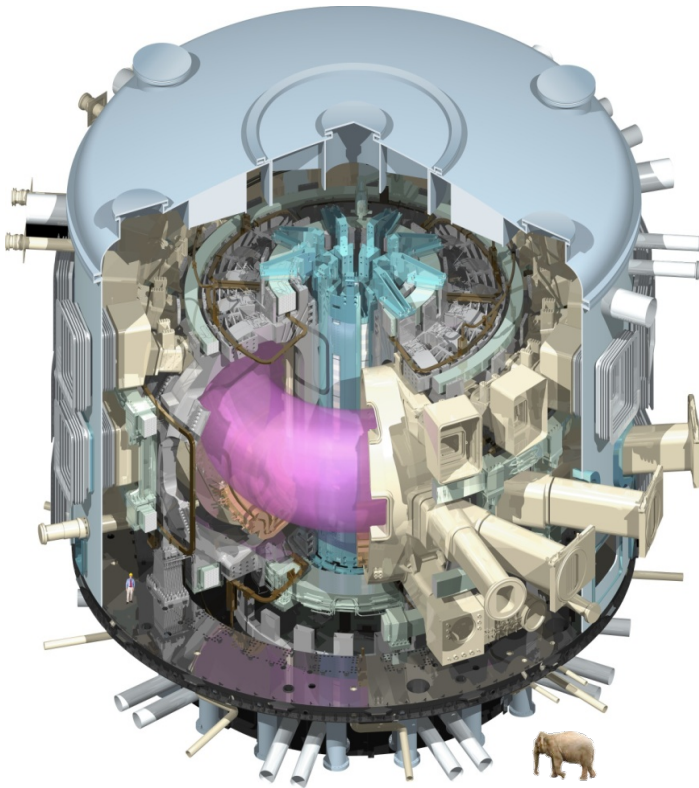




Path to a fusion power plant

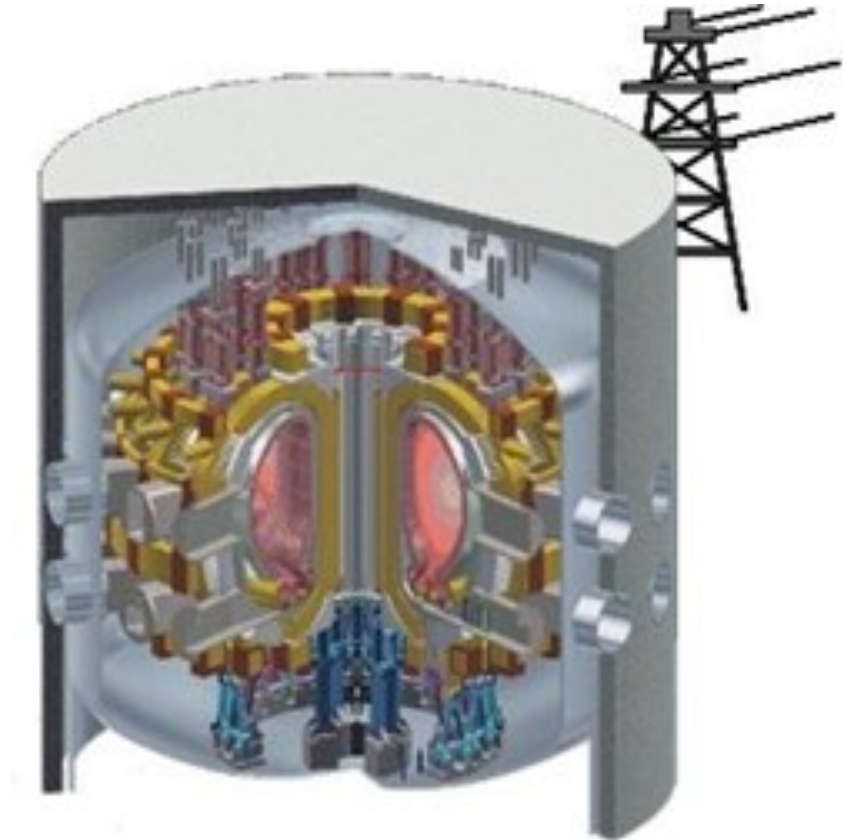
IPP

ITER



7min discharge
Tritium supplied externally
Experiment

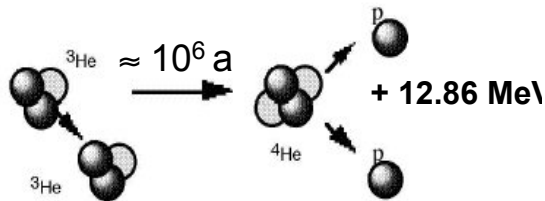
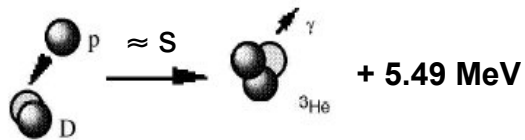
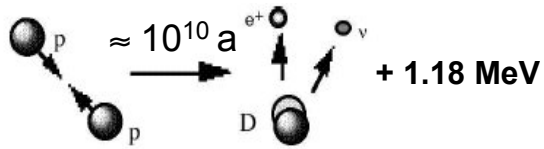
DEMO



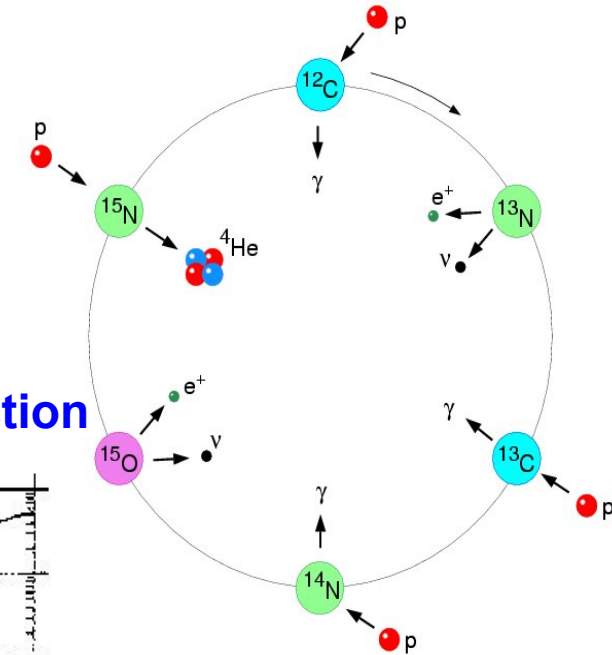
Continuous operation
Tritium bred from Lithium
Net electricity output to the network

Fusion process in stars

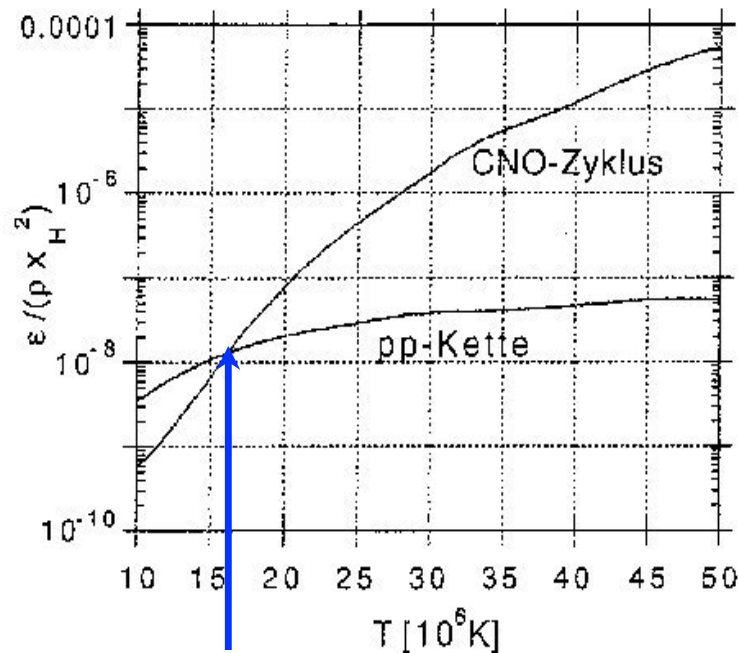
Protonen chain



CNO-Zyklus



Normalized energy production



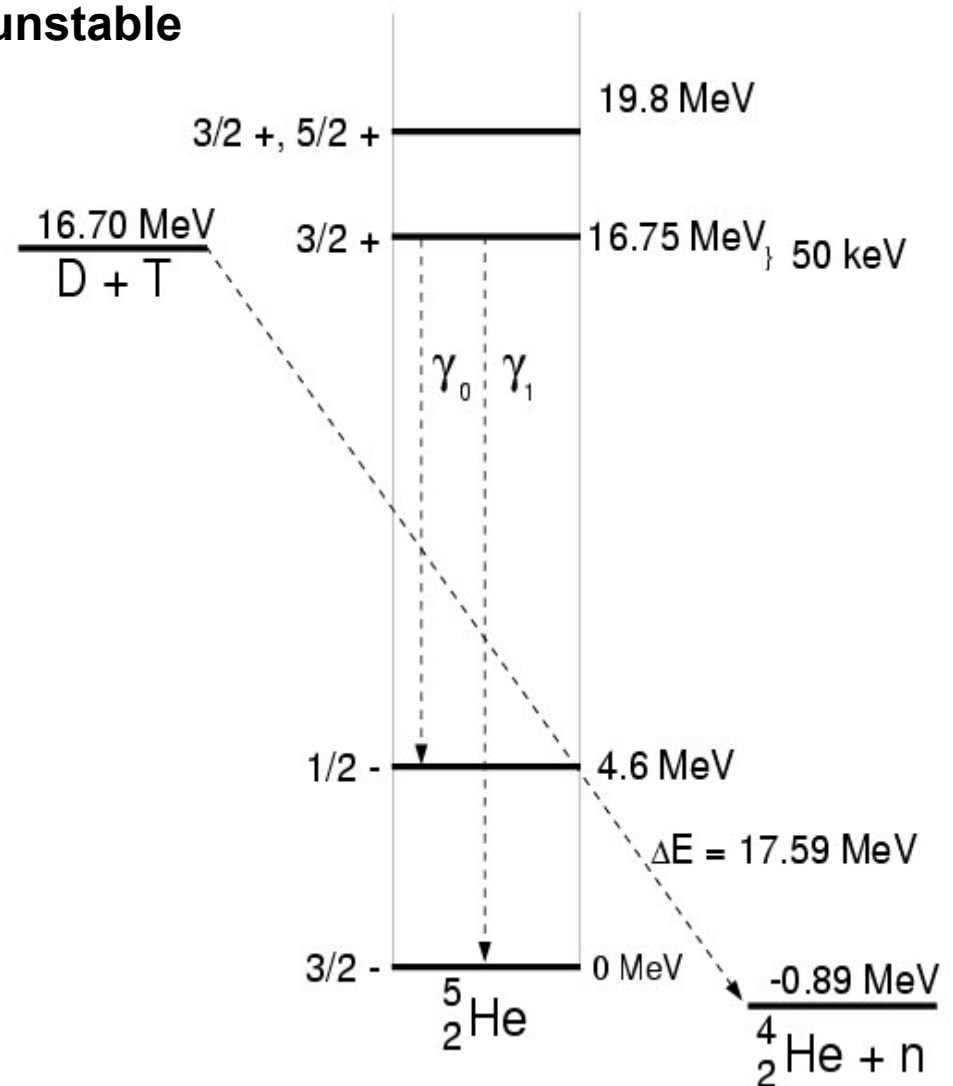
$T \approx 1.5 \text{ keV}$

Fusion process in laboratory



Energy level diagram of the unstable He_2^5

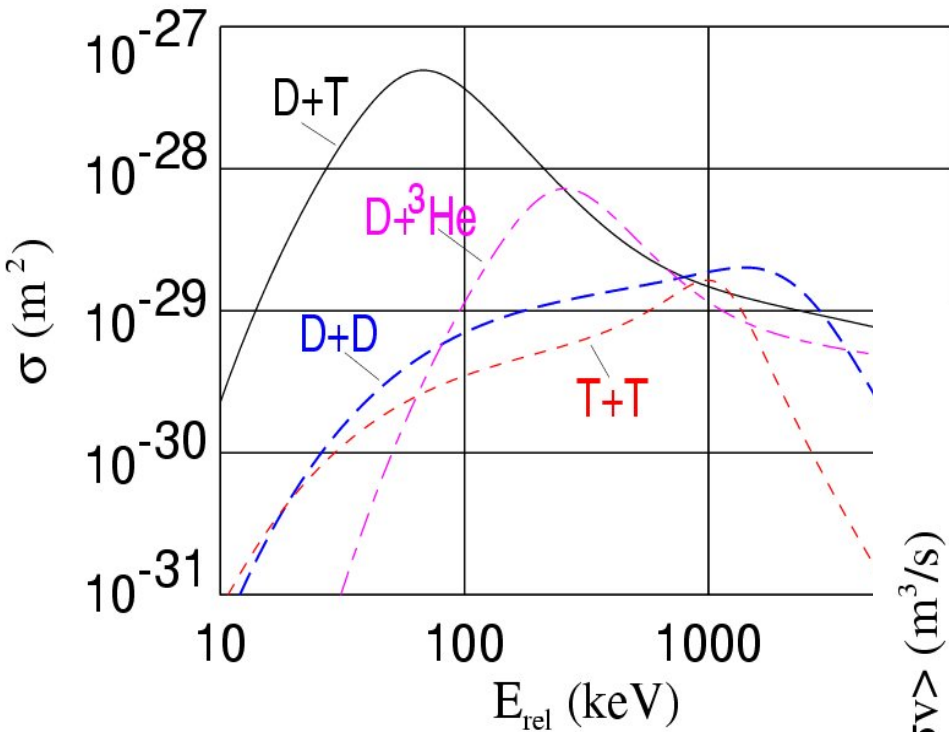
Resonance for $E = 64 \text{ keV}$



Fusion processes in a laboratory



Cross section



Reaction parameter – rate coefficient Assuming a Maxwell distribution

