Nucleon Decay Searches in Atmospheric Neutrino Detectors

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 — lightest baryon (proton) stable

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- Grand Unified Theories (GUTs) unify SM gauge groups
 - -- explain charge quantization, unification of couplings, anomaly cancellation
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- Predicted lifetime $\tau \sim 10^{29-36} {
 m yrs}$
- Minimal model ruled out (IMB-3, Kamiokande, Super-K)

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- Minimal (TeV-)SUSY model ruled out (Super-K) [Kobayashi+ (SK), 2005]

Proton decay: models

Model	Ref.	Modes	τ_N (years)
Minimal $SU(5)$	Georgi, Glashow [2]	$p \rightarrow e^+ \pi^0$	$10^{30}-10^{31}$
Minimal SUSY $SU(5)$	Dimopoulos, Georgi [11], Sakai [12]	$p \rightarrow \bar{\nu}K^+$	
	Lifetime Calculations: Hisano,	$n \rightarrow \bar{\nu} K^0$	$10^{28} - 10^{32}$
	Murayama, Yanagida [13]		
SUGRA $SU(5)$	Nath, Arnowitt [14, 15]	$p \rightarrow \bar{\nu}K^+$	$10^{32} - 10^{34}$
SUSY $SO(10)$	Shafi, Tavartkiladze [16]	$p \rightarrow \bar{\nu}K^+$	
with anomalous		$n \rightarrow \bar{\nu} K^0$	$10^{32} - 10^{35}$
flavor $U(1)$		$p \rightarrow \mu^+ K^0$	
SUSY $SO(10)$	Lucas, Raby [17], Pati [18]	$p \rightarrow \bar{\nu}K^+$	$10^{33} - 10^{34}$
MSSM (std. $d = 5$)		$n \rightarrow \bar{\nu} K^0$	$10^{32} - 10^{33}$
SUSY $SO(10)$	Pati [18]	$p \rightarrow \bar{\nu}K^+$	$10^{33} - 10^{34}$
ESSM (std. $d = 5$)			$\lesssim 10^{35}$
SUSY $SO(10)/G(224)$	Babu, Pati, Wilczek 19, 20, 21],	$p \rightarrow \bar{\nu}K^+$	$\lesssim 2 \cdot 10^{34}$
MSSM or ESSM	Pati [18]	$p \rightarrow \mu^+ K^0$	
$(new \ d = 5)$	0`	Br	$\sim (1 - 50)\%$
SUSY $SU(5)$ or $SO(10)$	Pati [18]	$p \rightarrow e^+ \pi^0$	$\sim 10^{34.9\pm1}$
MSSM $(d = 6)$			
Flipped $SU(5)$ in CMSSM	Ellis Danopoulos and Wlaker[22]	$p \rightarrow e/\mu^+ \pi^0$	$10^{35}-10^{36}$
Split SU(5) SUSY	Arkani-Hamed, et. al. [23]	$p \rightarrow e^+ \pi^0$	$10^{35}-10^{37}$
SU(5) in 5 dimensions	Hebecker, March-Russell[24]	$p \rightarrow \mu^+ K^0$	$10^{34} - 10^{35}$
		$p \rightarrow e^+ \pi^0$	
SU(5) in 5 dimensions	Alciati et.al.[25]	$p \rightarrow \bar{\nu}K^+$	$10^{36} - 10^{39}$
option II			
GUT-like models from	Klebanov, Witten[26]	$p \rightarrow e^+ \pi^0$	$\sim 10^{36}$
Type IIA string with D6-branes			

TABLE I: Summary of the expected nucleon lifetime in different theoretical models.

A. Bueno et al. hep-ph/0701101

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Proton decay: general



Proton decay: general



- Many related processes ($n \overline{n}$, di-/tri-nucleon decays, exotic n-decays)
- Specific predictions for main channel and lifetimes strongly model-dependent

Proton decay: searches

• Proton lives $> 10^{30}$ years how to test if decays?

A) look at 1 proton for VERY long time

B) look at many protons simultaneously for few years

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large underground water Cherenkov experiments



cheap, proven technology

State of the Art

Super-Kamiokande

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Kamioka Nucleon Decay Experiment



The Super-Kamiokande Experiment

- <u>Experimental setup</u> (Kamioka, Japan)
 - -- 22.5 kton fiducial volume
 - -- inner (11k PMTs, 40% coverage) & outer (2k PMTs) detectors

-- 4 run periods: SK-I (1996 - 2001)

- **SK-II** (2003 2005): accident, $\frac{1}{2}$ PMT coverage
- **SK-III** (2006 2008): restore PMT coverage
- **SK-IV** (2008 now): upgraded electronics





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• **<u>AMAZING</u>** multipurpose physics detector ($\sim 10 - 10^4$ MeV)

-- v-oscillations, Lorentz invariance, sterile-v

- -- nucleon decay
- -- solar-v, day/night effect, supernovae relic v's
- -- indirect dark matter detection
- -- exotics (monopoles, Q-balls, etc.)



Super-Kamiokande (SK)



Technique: Cherenkov ring imaging



Cherenkov radiation

charged particle travels in medium (ref. index n) faster than light, builds up wave-front

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Technique: Cherenkov ring imaging

SK event display



(recorded on 1998.04.04 08:35:22)

- -- energy 603 MeV
- -- time width 162 ns



(recorded on 1998.04.04 21:26:08)

-- energy 492 MeV -- time width 130 ns

Detecting nucleon decay

(1) Signal – proton decay

"Golden channel":
$$p \rightarrow \pi^0 e^+$$
 $p_{\pi} = p_e = 459 \text{ MeV}$
 $\pi^0 \rightarrow 2\gamma$ $p_{Y/\pi R} = 68 \text{ MeV}$



Detecting nucleon decay

(2) Background – atmospheric neutrinos







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Detecting nucleon decay

(3) Setting Limit



• Motivation: "golden" dominant GUT channel

- [Abe+ (SK), 2016]
- Signature: $e^+, \pi^0 (\rightarrow 2\gamma)$ visible --> fully reconstruct parent proton
- Analysis:

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exposure: 306 kt-yrs

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SK I-IV	R1+R2
Efficiency	~ 40 %
BKG	0.62
Data	0

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Search: $\mathbf{p} \rightarrow \mu^+ \pi^0$

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• Results:

exposure: 306 kt-yrs

SK 1-1V	KI+KZ
Efficiency	~ 35 %
BKG	0.88
Data	2

[Abe+ (SK), 2016]

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• Results: 2 candidates, consistent with background



exposure:306 kt-yrsSK I-IVR1+R2Efficiency~ 35 %BKG0.88Data2

[Abe+ (SK), 2016]

• Motivation: "silver" dominant SUSY-GUT channel

[Abe+ (SK), 2014]

- Signature: $\nu, K^+(\rightarrow \mu^+ \nu, \pi^0 \pi^+)$ both invisible --> reconstruct Kaon decays
- Analysis:

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Search (A) $(K^+ \to \mu^+ \nu)$ spectral fit to μ^+ momentum Search (B) $(K^+ \to \mu^+ \nu)$ tag γ from nuclear de-excitation Search (C) $(K^+ \to \pi^0 \pi^+)$ reconstruct pions







SK I-IV

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	SK I-IV	Α	В	С
SK I-IV	Efficiency	~ 34%	~ 8%	~ 8%
	BKG	579.4	0.39	0.56
	Data	566	0	0

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SK 1-1V	BKG	579.4	0.39	0.56
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exposure: 260 kt-yrs

• Results: NO excess
$$\tau_p > 6.6 \times 10^{33} \text{ yrs.}$$

Search: $n - \overline{n}$ oscillation

- Motivation: $\Delta B = 2$, parametrizes $U(1)_{B-L}$ (as Majorana v's), test intermediate scales
- Signature: \overline{n} captured by nucleon (n or p) produces pions ---> reconstruct initial state
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(improved machine learning-based analysis on SK I-IV yields similar results)

Search: spectral 1-ring modes

- Motivation: broad search, Pati-Salam & extended Higgs models
- Signature: visible $e^+/\mu^+/\gamma$ --> spectral fit to ring momenta [Takh

[Takhistov+ (SK), *PRL* 2014] [Takhistov+ (SK), *PRL* 2015]

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from S. Sussman (PRELIMINARY)

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Probing lifetime $\tau \gtrsim 10^{32} \text{ yrs}$ across channels !

Future Prospects

Near term (~2019): SK-Gd

Super-K Gadolinium upgrade

--> efficient neutron tagging (~80%) will help reduce background

--> allows to claim discovery if candidates observed



Medium term (~2020+): JUNO

Large liquid scintillator detector

- 20 kt FV liquid scintillator (China)
- Clean timing signature --> specialize in charged Kaon (+ invisible mode)



Long term (~2025): Hyper-Kamiokande

Next generation water Cherenkov detector

- Well established technology,
 builds on success of Super-K
- 650m underground in Kamioka
- 10x Super-K FV
- Photo-sensors with

2x single-photon sensitivity



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Long term (~2025): Hyper-Kamiokande

• Shear Hyper-K size allows up to an order gain in lifetime sensitivity across the board, can finally probe $au \sim 10^{35}~{
m yrs}$

$\begin{array}{ll} p \to e^+ \pi^0 & 6.3 \ (8.0) \times 10^{34} \\ p \to \overline{\nu} K^+ & 2.0 \ (2.5) \times 10^{34} \\ \hline p \to \mu^+ \pi^0 & 6.9 \ (8.7) \times 10^{34} \\ p \to e^+ \eta^0 & 3.0 \ (3.9) \times 10^{34} \\ p \to \mu^+ \eta^0 & 3.4 \ (4.7) \times 10^{34} \\ p \to e^+ \rho^0 & 3.4 \ (5.0) \times 10^{33} \\ p \to \mu^+ \rho^0 & 1.3 \ (1.6) \ \times 10^{33} \\ p \to e^+ \omega & 5.4 \ (6.9) \times 10^{33} \\ p \to \mu^+ \omega & 0.78 \ (1.0) \times 10^{34} \end{array}$	Mode	$\tau_{disc} \ 3\sigma$ [years]
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$n \rightarrow e^+ \rho^-$	$1.1 (1.5) \times 10^{33}$
$n ightarrow \mu^+ ho^-$	$6.2 (8.1) \times 10^{32}$

from HK design report

Deep Underground Neutrino Experiment (DUNE)

--> flagship experiment for US neutrino program

- 1 km underground in Sanford Lab (Homestake, South Dakota)
- 40 kton FV liquid Argon detector



LArTPC technique



Final state Kaon visible in LArTPC (below threshold in w. Cherenkov)
 --> well suited to probe SUSY modes

 $p \to K^+ \overline{\nu}$



Multi-prong events without neutrinos clearly visible

--> potential for strong background descrimination

$n - \overline{n}$ oscillations



Where do we stand?



from E. Kearns

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- Nucleon decay and related processes offer unique window into energies not accessible to accelerators as well as plethora of BSM physics
- Uncertainty in theory predictions requires a broad search program, well suited for large underground experiments, with the field led by Super-K --> many modes left to search (e.g. $n \rightarrow \nu\nu\nu$, $p \rightarrow e^+e^-e^+$)
- Approaching interesting limit range of $\tau \sim 10^{35} {
 m yrs.}$

→ great complimentary physics program to neutrinos (Hyper-K/DUNE)