



Kinetic simulations of the NTM polarisation current

Emanuele Poli
Andreas Bergmann, Arthur G. Peeters

Max-Planck-Institut für Plasmaphysik, Garching bei München, Germany
EURATOM Association

Motivations

- Reliable description of NTMs necessary in order to determine onset conditions and stabilisation requirements (→ ITER)
- Problem at the meeting point of MHD and kinetic theory (→ required for accurate predictions, e.g. NTM polarisation current)

Outline

- Polarisation current in the presence of a magnetic island
- Solving the drift kinetic equation
- Single-particle motion and full 3D simulations: new conditions for island stability

The Neoclassical Tearing Mode

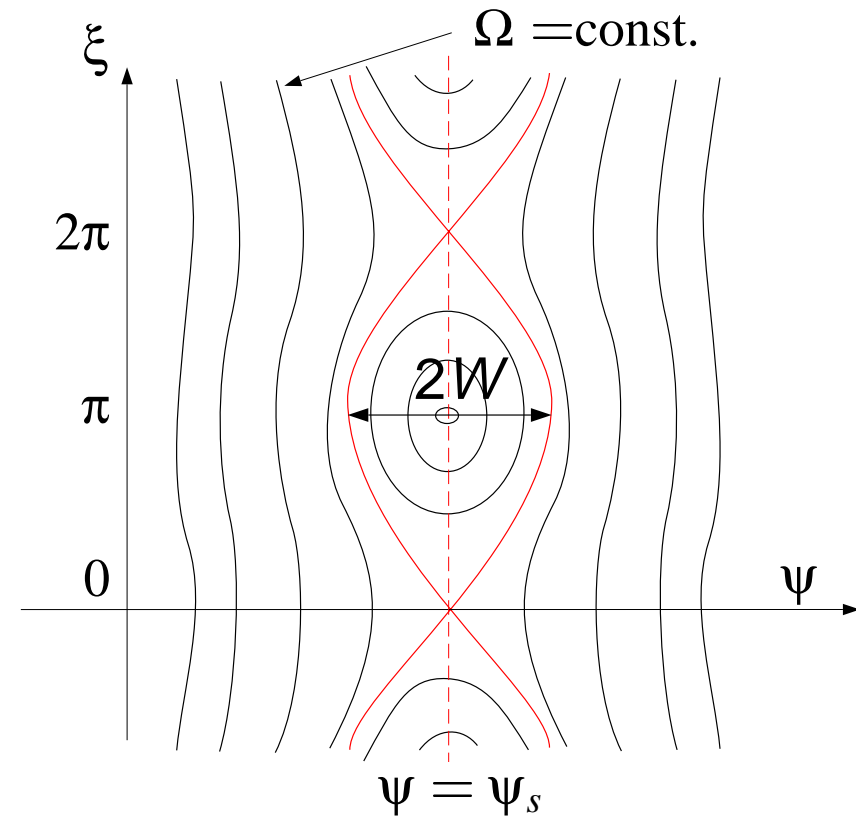
- Island evolution connected with the parallel currents flowing near the resonant surface

$$\frac{dW}{dt} = c_1 \Delta' + \frac{c_2}{W} \int_{-1}^{\infty} d\Omega \oint \frac{d\xi \cos \xi}{\sqrt{\cos \xi + \Omega}} j_{\parallel}^{n.i.}$$

New flux coordinates:

helical flux $\Omega \equiv 2(\psi - \psi_s)^2 / W_{\psi}^2 - \cos \xi$

helical angle $\xi \equiv m\theta - n\zeta$



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- **Destabilising term:**

Bootstrap current loss

[Qu and Callen, UWPR1985;
Carrera et al., PoF 1986]

- **Stabilising terms:**

(large W)

→ Δ' (current profile, $m \geq 2$)

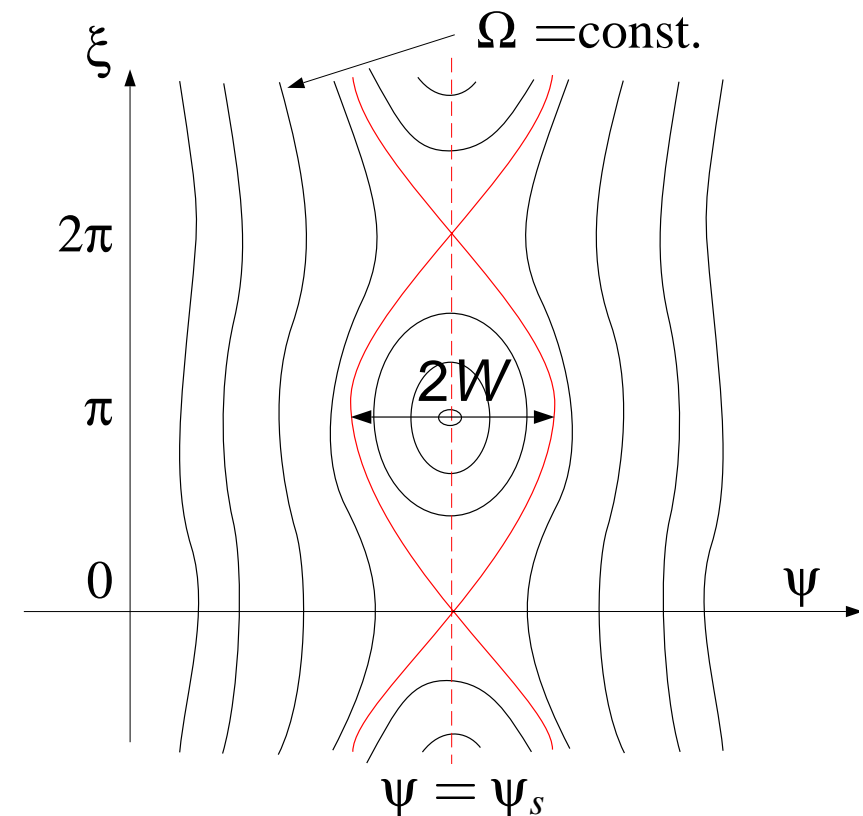
[Rutherford, PoF 1973]

(small W)

→ ...

→ Polarisation current (?)

[Smolyakov et al., PoP 1995;
Wilson et al., PoP 1996]



The island polarisation current

- Island motion with respect to the plasma
⇒ electric field induced (Faraday)
- $E \times B$ motion in the island rest frame: plasma acceleration and deceleration around the O -point
- Variation of plasma inertia balanced by a Lorentz force provided by

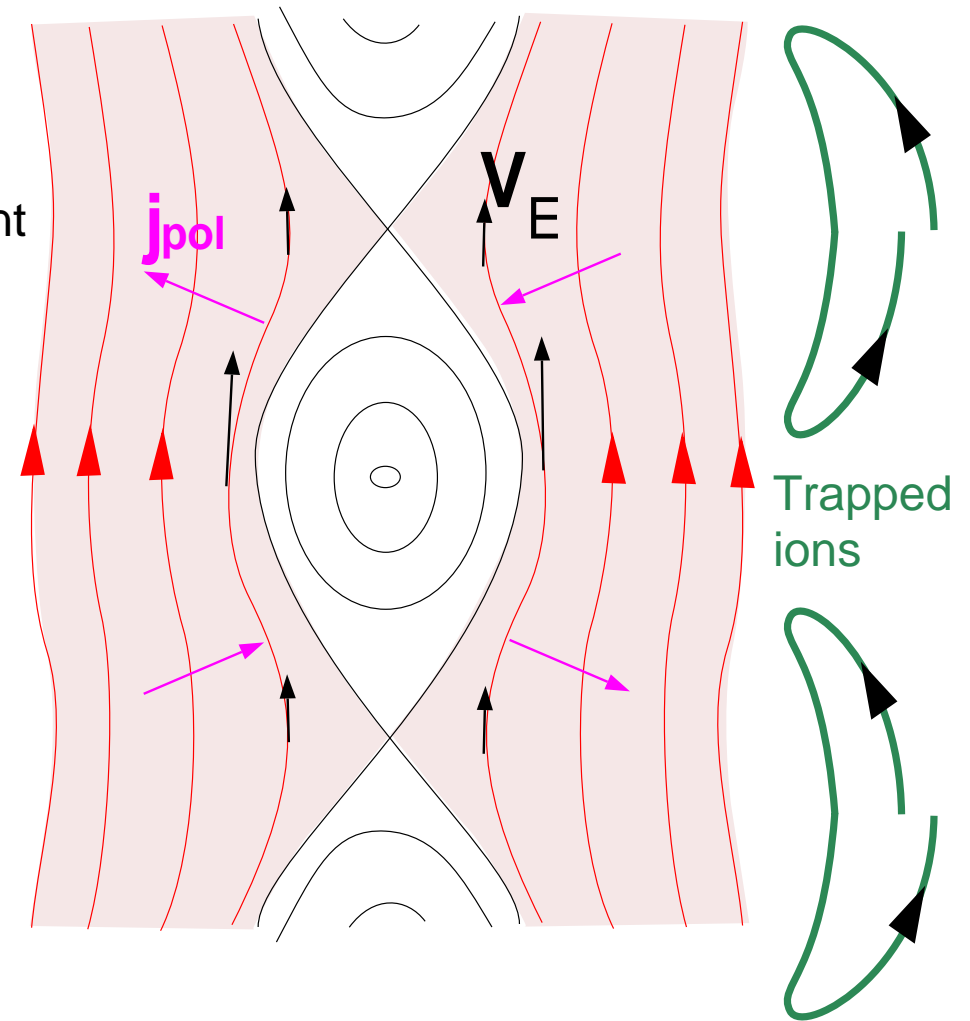
$$j_{\text{pol}}^{\text{class}} = \frac{en}{\omega_c} \frac{dv_E}{dt}$$

Polarisation current

(⇒ mainly carried by the ions)

[Smolyakov, PPCF 1993]

- Current continuity ($\nabla \cdot \mathbf{J} = 0$) ensured by an electron **parallel** current contributing to the Rutherford equation



Solving the drift kinetic equation



- **Analytical** determination of $j_{\parallel}^{n.i.}$ from the drift kinetic equation possible employing the expansion parameters $W/r, w_b/W \ll 1$ (and further simplifications...)

- Drift kinetic equation in toroidal geometry with an island structure to be solved

$$\frac{df}{dt} = \frac{\partial f}{\partial t} + \left(v_{\parallel} \hat{\mathbf{b}} + \mathbf{v}_d + \mathbf{v}_E \right) \cdot \frac{\partial f}{\partial \mathbf{r}} - \frac{e}{m} \frac{\mathbf{v}_d \cdot \nabla \Phi}{v} \frac{\partial f}{\partial v} = C(f)$$

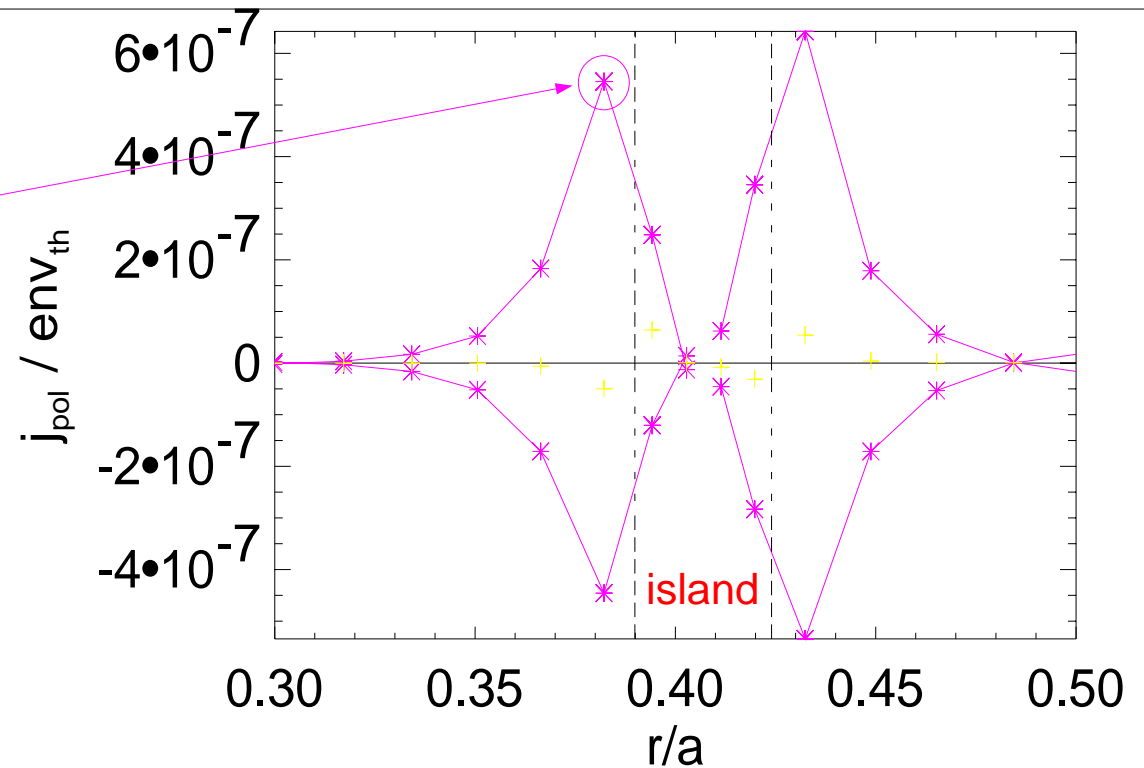
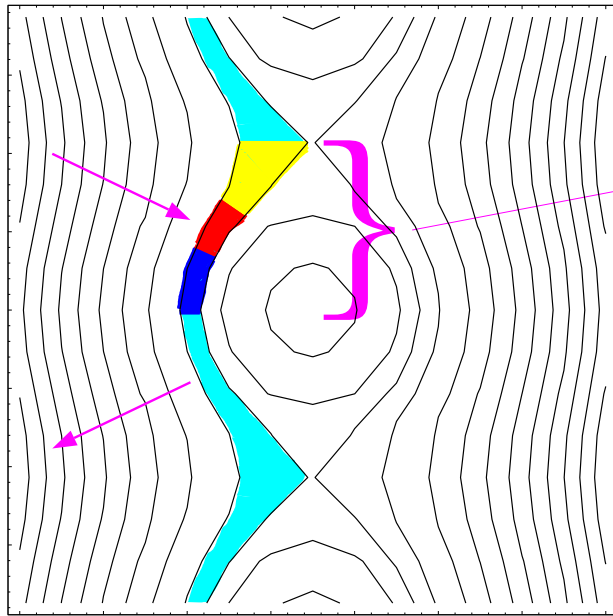
parallel motion
magnetic & electric drift
electric field
collisions

- Representation of the distribution function: $f = f_0 + \delta f = f_M(\psi, \mathcal{E}) + \delta f$
if $\delta f \ll f_0$: reduction of the numerical noise

- The equation for δf is $\frac{d(\delta f)}{dt} = C(\delta f) - \mathbf{v}_d \cdot \nabla f_M - \frac{e f_M}{T} \mathbf{v}_d \cdot \nabla \Phi$

- Solution:
- $\delta f \rightarrow$ markers (ions) \rightarrow Hamiltonian equations of motion in Boozer coordinates (\rightarrow **HAGIS**) [Pinches et al., CPC 1998]
 - Collisions: **Monte Carlo procedure** [Bergmann et al., PoP 2001]

Current profiles from the drift kinetic equation



- Macroscopic quantities as moments of the distribution function
- Flux surface averages → **cells**

- “Radial” profiles of the polarisation current available!
- Binning in velocity space also possible

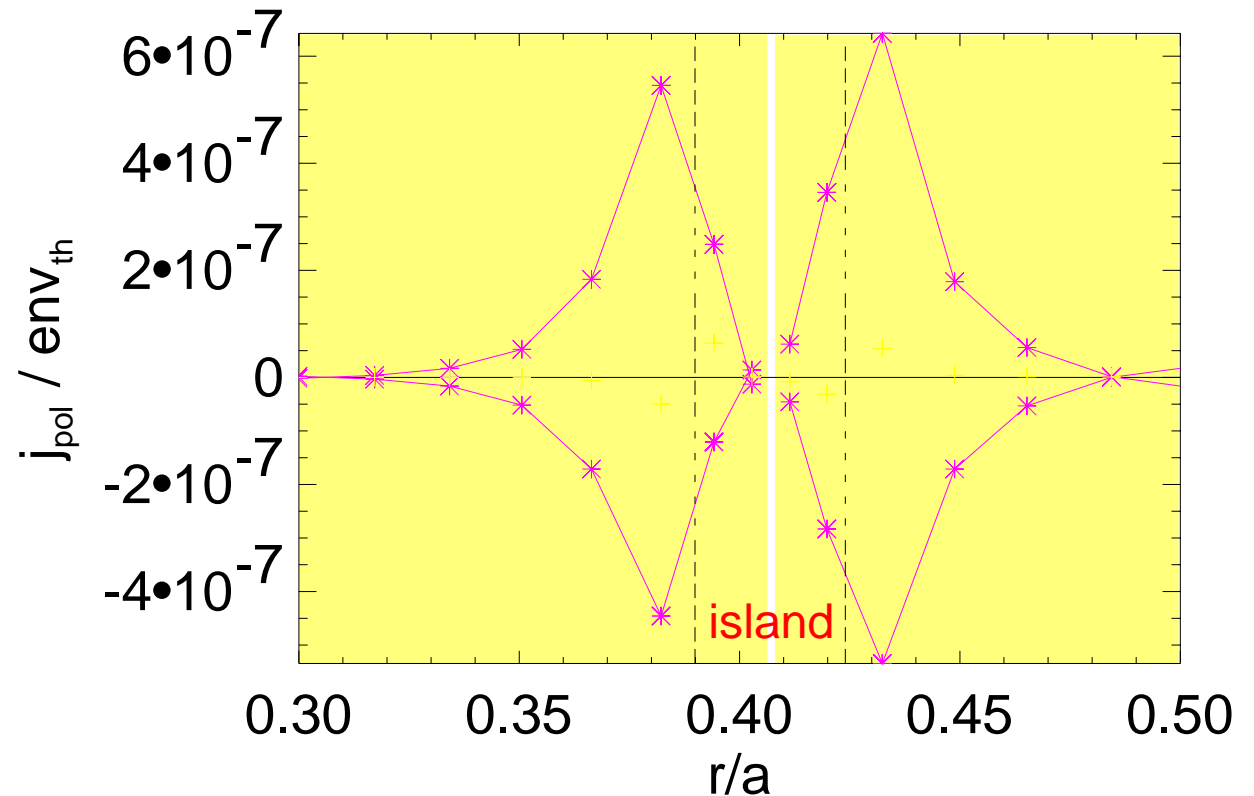
$$\langle A \rangle = \lim_{\delta\Omega \rightarrow 0} \frac{\int A d^3\mathbf{r}}{\int d^3\mathbf{r}} \Rightarrow \frac{1}{n} \left\langle \int A \delta f d^3\mathbf{v} \right\rangle \simeq \frac{\int_{\Omega-\delta\Omega}^{\Omega+\delta\Omega} A \delta f d^3\mathbf{r} d^3\mathbf{v}}{\int_{\Omega-\delta\Omega}^{\Omega+\delta\Omega} f_0 d^3\mathbf{r} d^3\mathbf{v}}$$

Stabilising or destabilising?

- Contribution of the polarisation current to Rutherford equation

$$\Delta'_{\text{pol}} \propto \int_{-1}^{\infty} d\Omega \oint \frac{d\xi \cos \xi}{\sqrt{\cos \xi + \Omega}} j_{\parallel}$$

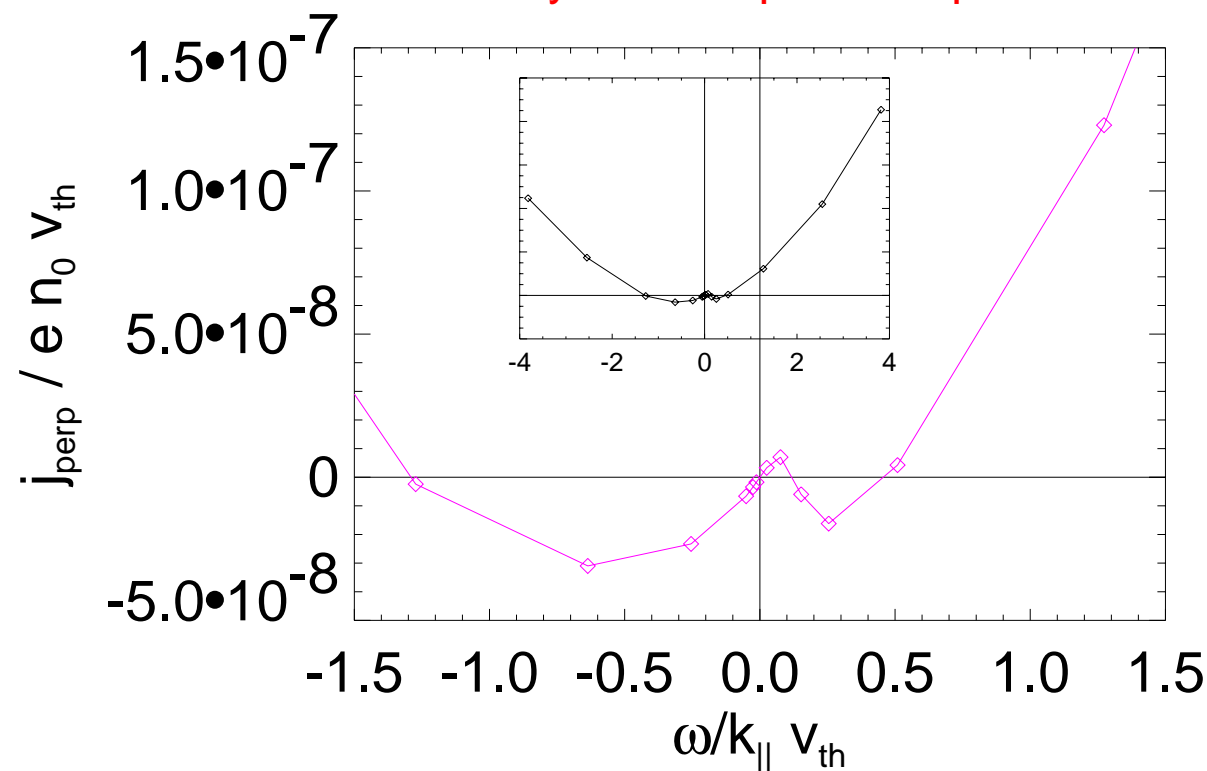
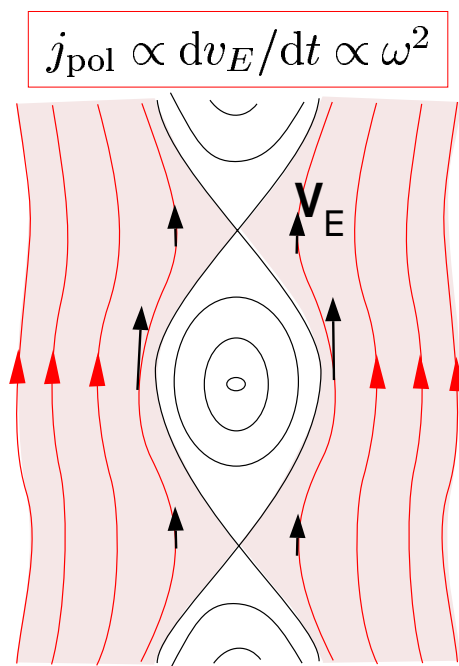
- Parallel current obtained from integration of $\nabla_{\parallel} j_{\parallel} = -\nabla_{\perp} \cdot j_{\perp}$ (j_{\perp} -profile numerically available)



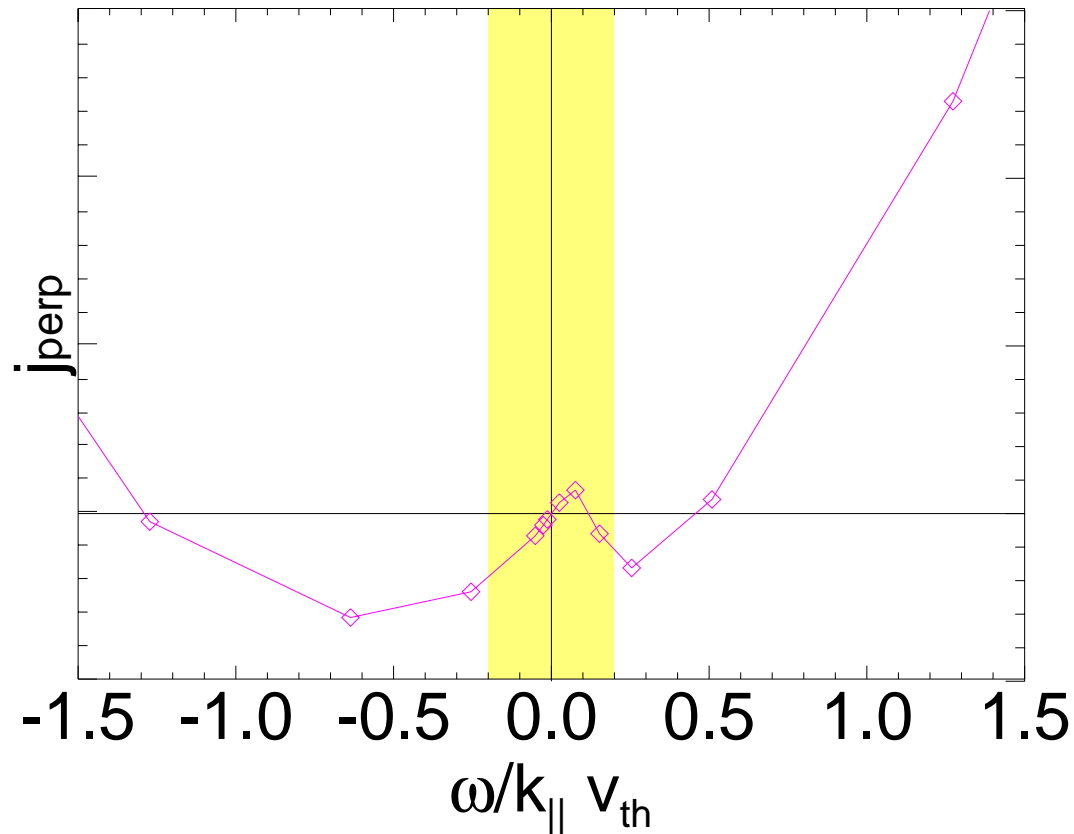
- Numerical results (“standard” polarisation current) in agreement with the current understanding (**flat pressure**): polarisation current
 - ⇒ **stabilising** if the separatrix is excluded from the radial integration
 - ⇒ **destabilising** if it is included in the radial integration
- [Waelbroeck and Fitzpatrick, PRL 1997]

Perpendicular current vs. island rotation frequency

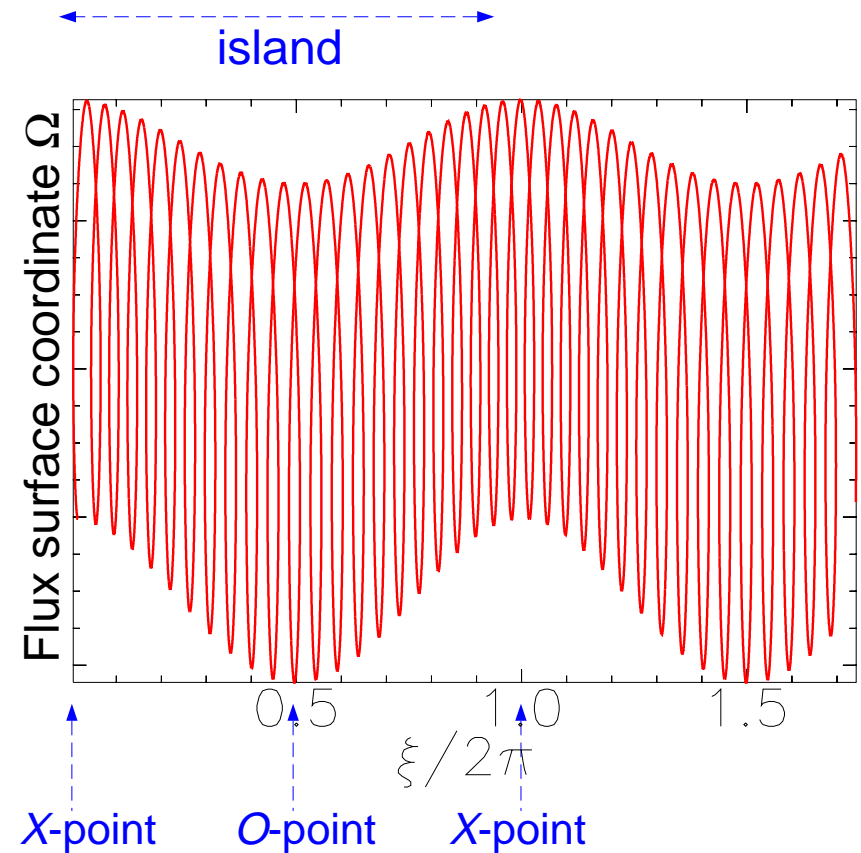
- Scan over ω important because of theoretical and experimental uncertainties about its actual value
- Behaviour of j_{\perp} vs. ω puzzling (quadratic scaling with ω expected from fluid picture)
- Simulation parameters: (3,2) mode, $R = 8$ m, $B_0 = 8$ T, $n_i = 10^{20}$ m⁻³, $T_i = 5$ keV
flat density and temperature profiles



Perpendicular current vs. ω : low frequencies

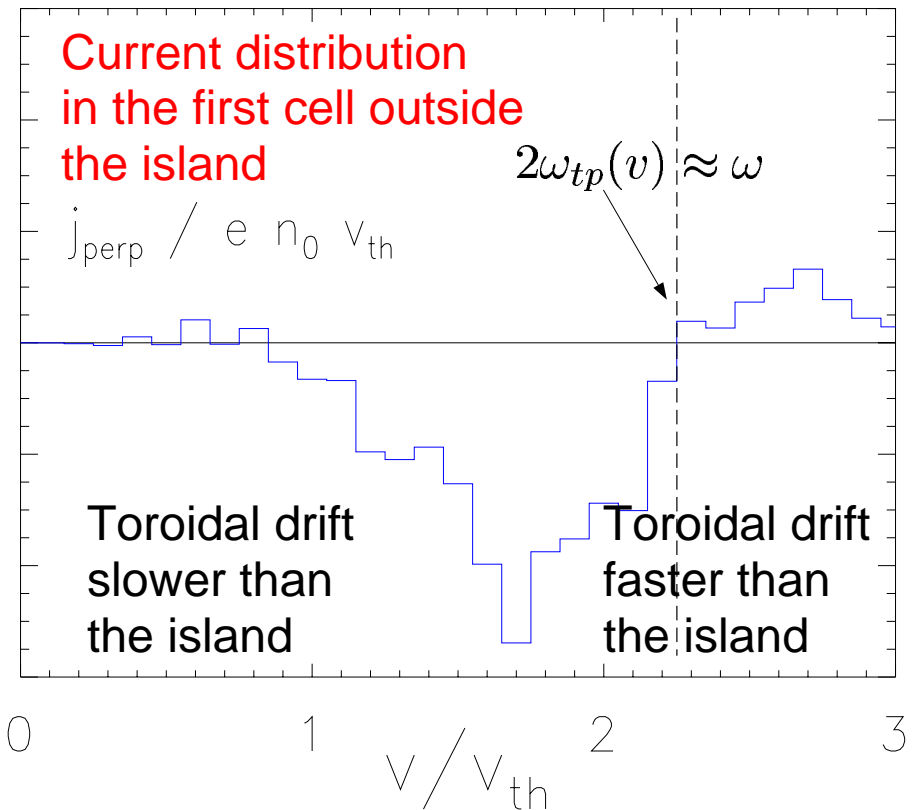
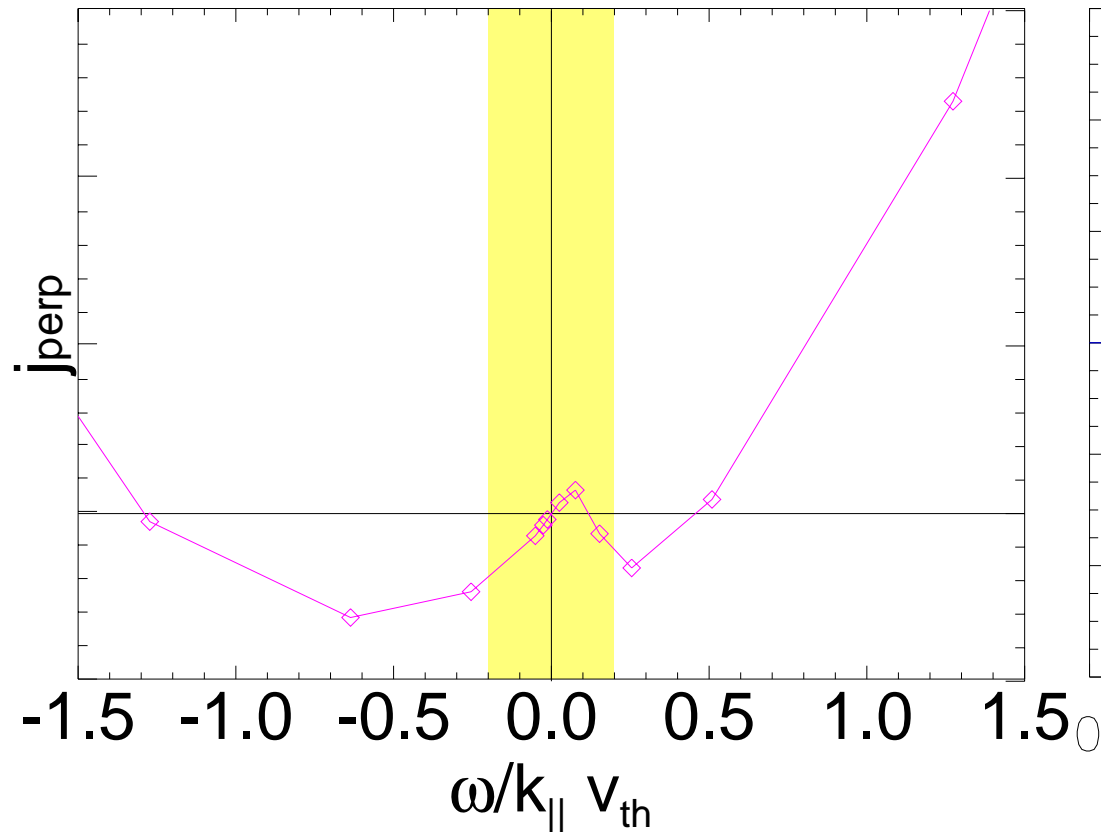


- Toroidal precession compensated by the $E \times B$ drift (island frame) when $2\omega_{tp} \approx \omega$



- Deviation from the perturbed magnetic surfaces due to a combination of magnetic and electric drift (dominates over the polarisation drift)

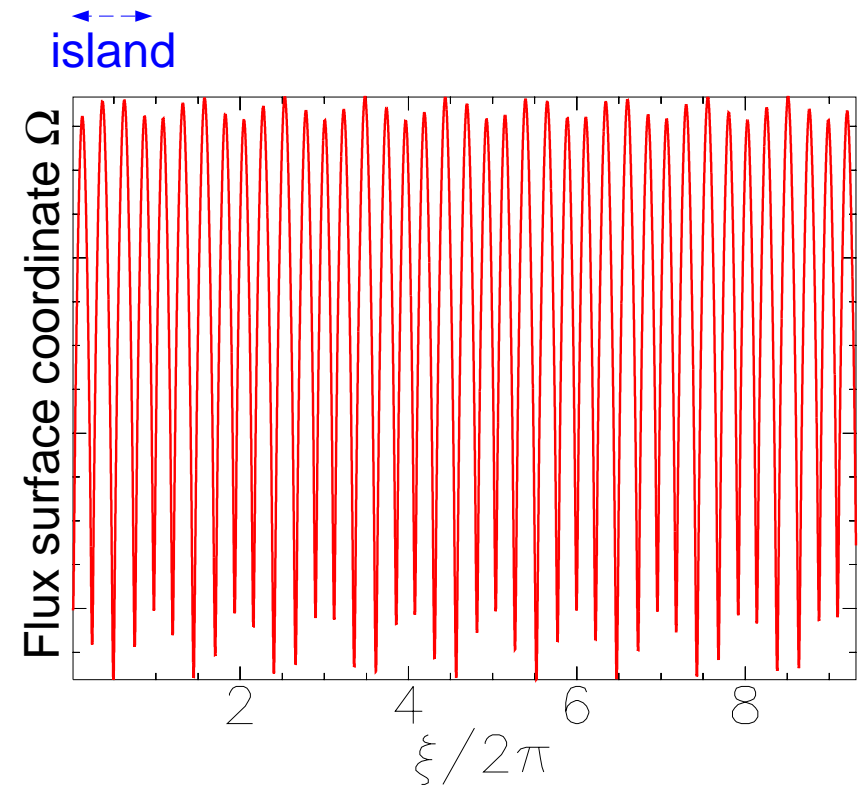
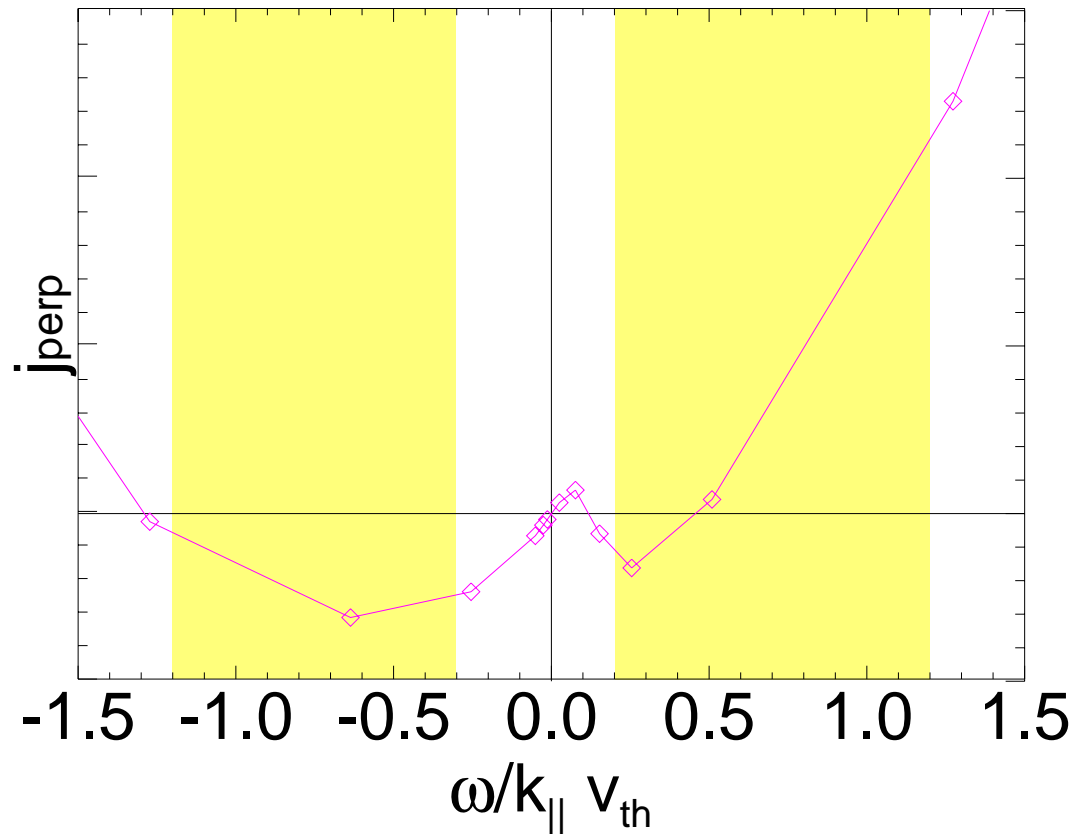
Perpendicular current vs. ω : low frequencies



- Toroidal precession compensated by the $E \times B$ drift (island frame) when $2\omega_{tp} \approx \omega$
- Motion dominated by the radial component of the $E \times B$ velocity

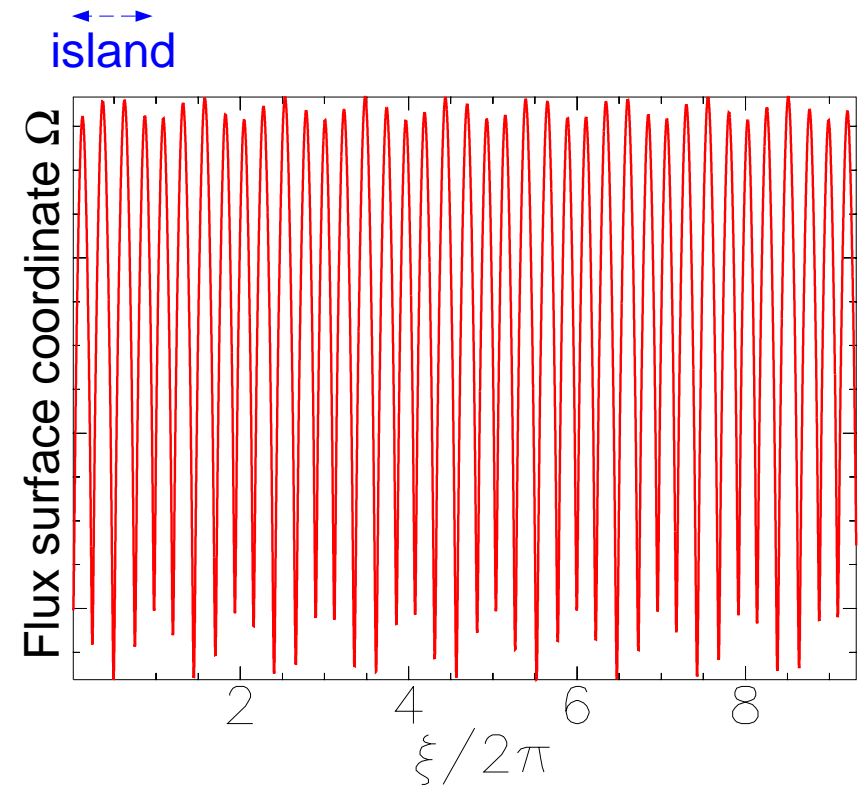
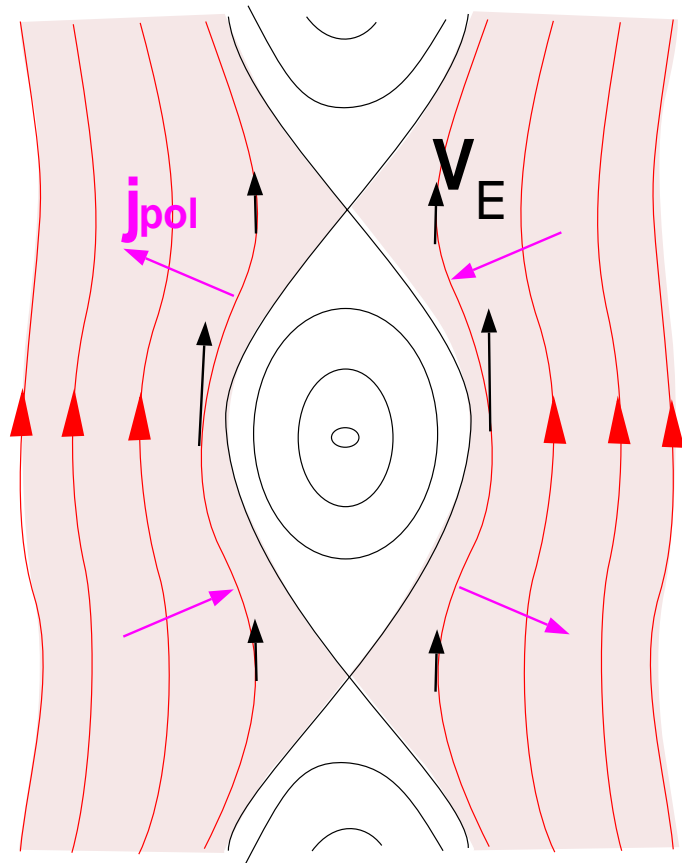
- Different sign of the current for particles drifting slower or faster than the island

Perpendicular current vs. ω : transition to higher frequencies



- Transition to higher frequencies
→ toroidal precession less and less important
- Bounce motion along the perturbed surfaces → polarisation current sets on

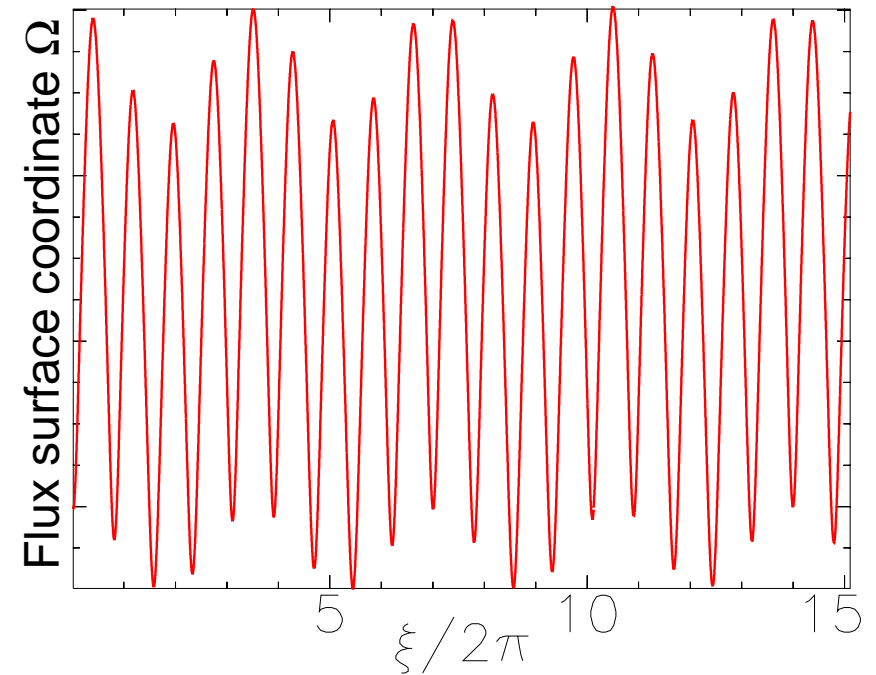
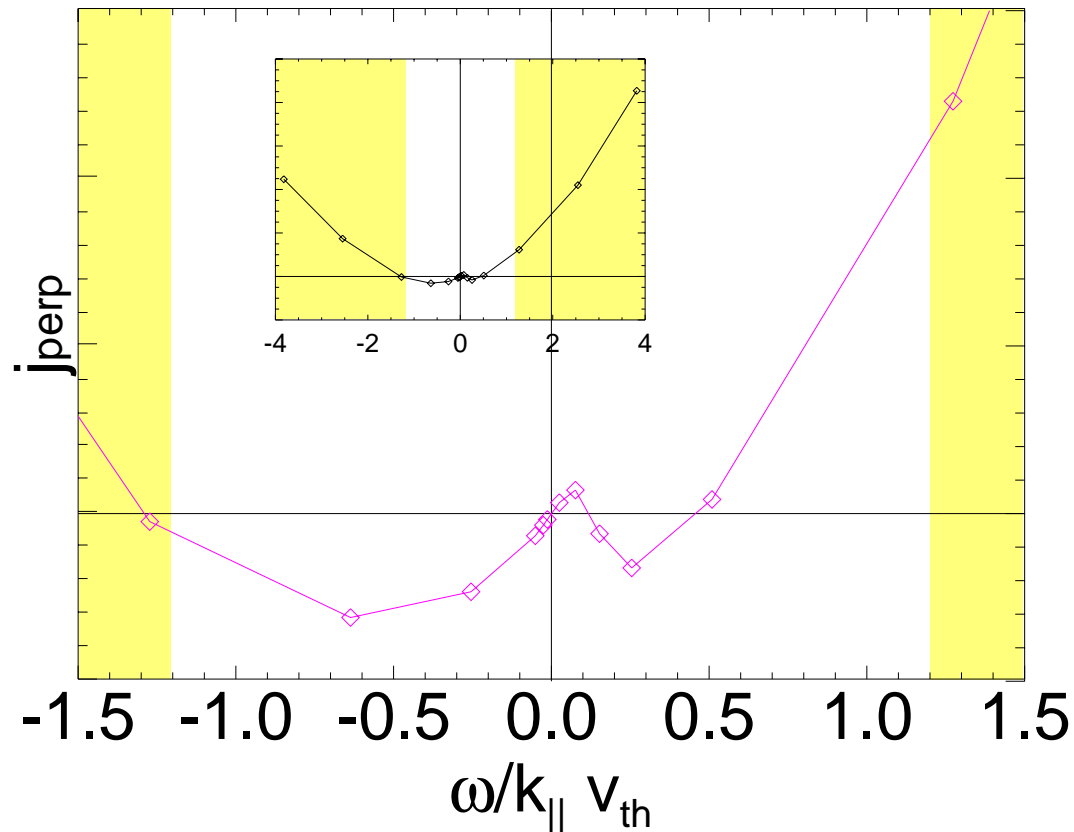
Perpendicular current vs. ω : transition to higher frequencies



- Transition to higher frequencies
→ toroidal precession less and less important

- Bounce motion along the perturbed surfaces → polarisation current sets on (cf. fluid picture)

Perpendicular current vs. ω : the “standard” polarisation current



- High frequencies: polarisation current close to “fluid” behaviour \rightarrow quadratic dependence on ω found

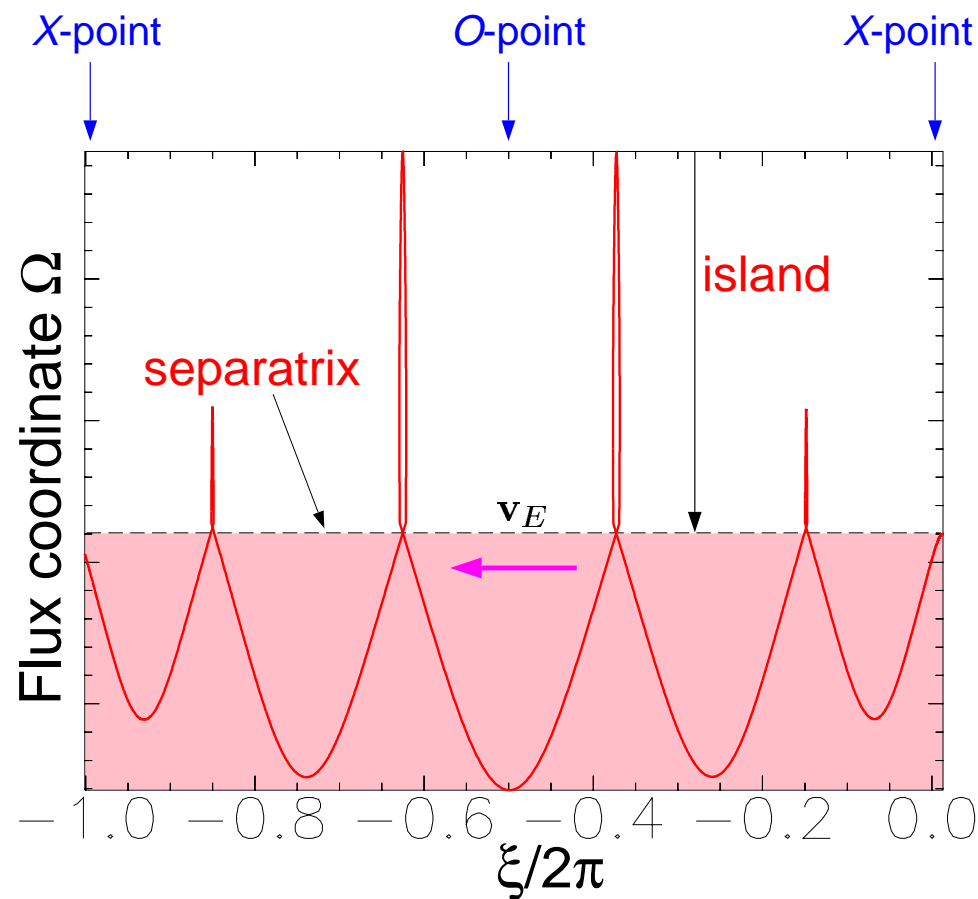
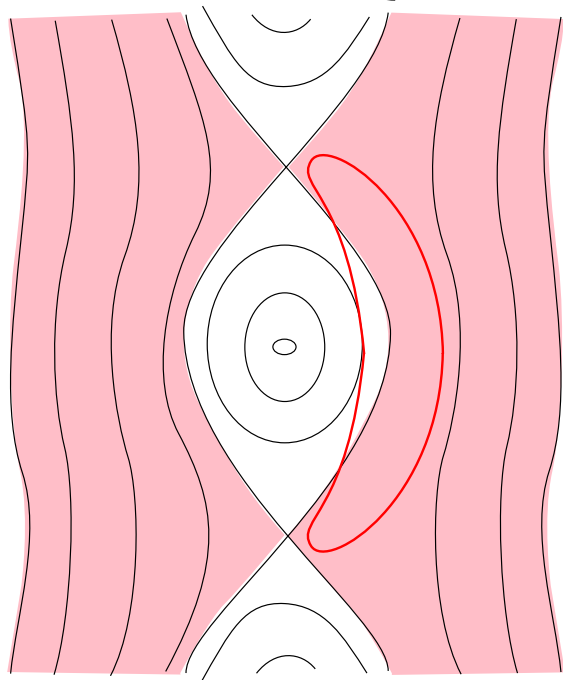
- Superposition of island motion and bounce motion \rightarrow current reduction due to slower particles

Polarisation current vs. island width

- Simulation parameters: (3,2) mode, $R = 8$ m, $B_0 = 2$ T, $n_i = 10^{20}$ m⁻³, $T_i = 5$ keV, flat temperature and density profiles

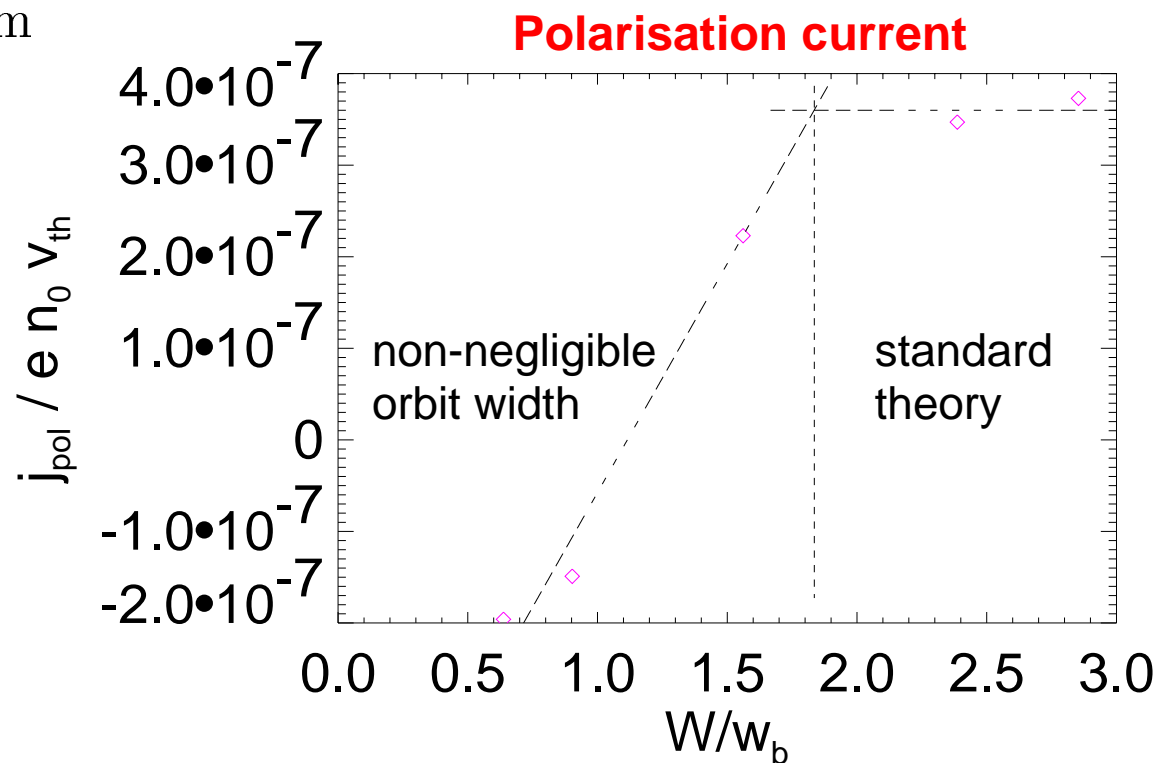
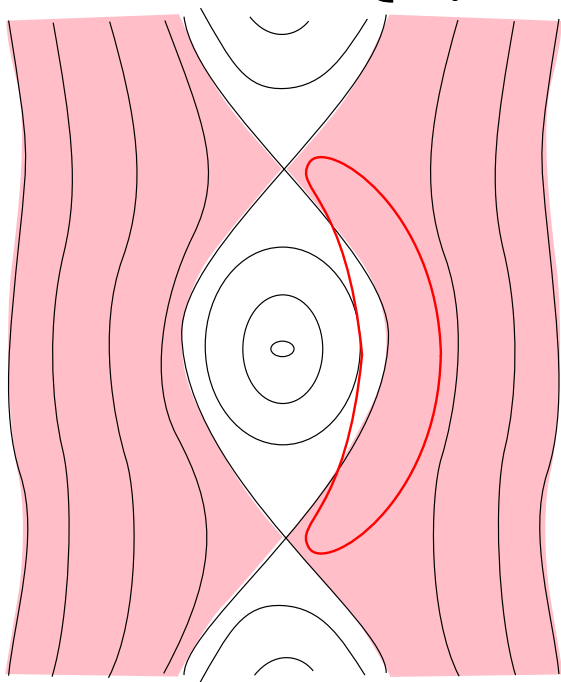
- Local effects “smeared out” by trapped particles overlapping the island

- ASDEX Upgrade: $\begin{cases} W_{\text{seed}} \approx 1 \div 5$ cm \\ $w_b \approx 0.7 \div 3$ cm \end{cases}



Polarisation current vs. island width

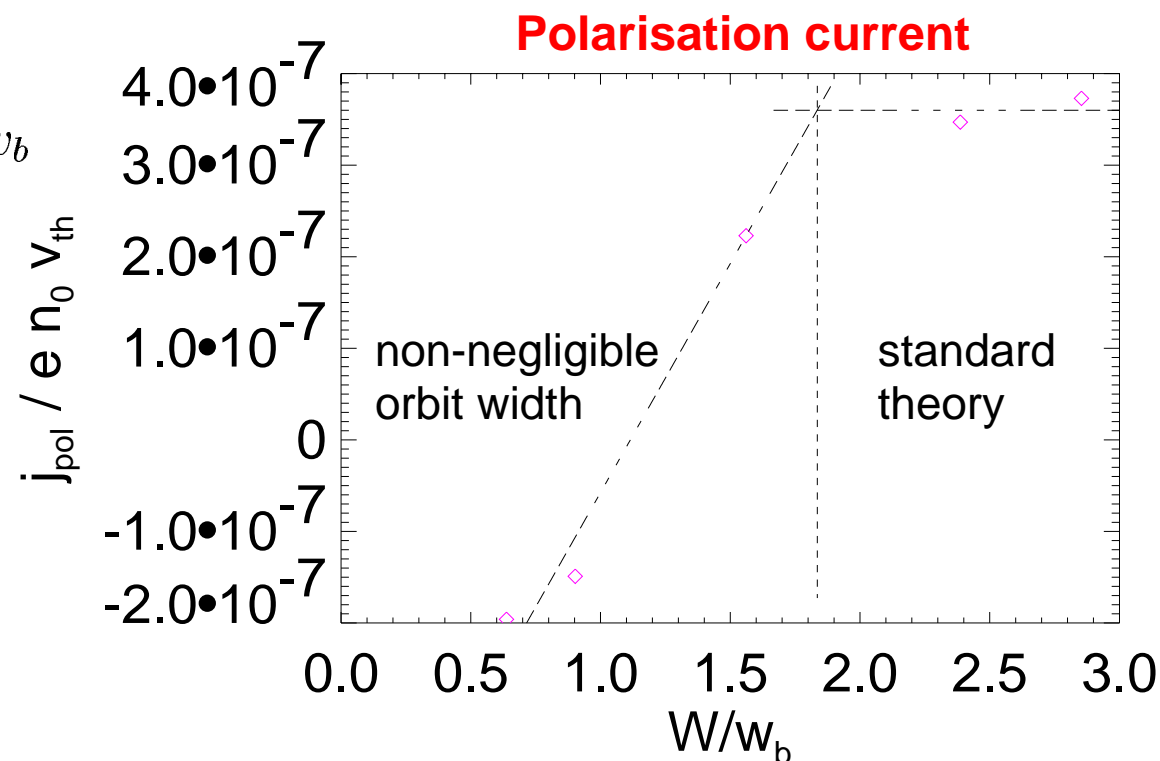
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- Local effects “smeared out” by trapped particles overlapping the island

- Same results if ω/ω_B , $\omega/k_{||}v_{th}$, W/w_b are kept constant
- Toroidal precession effects less important for small islands



- Sign of the polarisation current influenced by competition between electric and magnetic drift
- Polarisation current strongly reduced for small island widths (comparable to banana width)

Conclusions

- Kinetic effects are essential to capture the whole dynamics of the NTM
- Our present model needs to be extended to include self-consistent electron response