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#### **Conference and Euromech Colloquium #480**

on

**High Rayleigh Number Convection** 

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#### Power-Law scaling of Laminar natural convection with horizontal temperature gradient

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These are preliminary lecture notes, intended only for distribution to participants

### Power-law Scaling of Laminar Natural Convection with Horizontal Temperature Gradient

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### Outline

• Problem description

○ Numerical technique -- lattice Boltzmann method

O Power-law scaling and physical interpretation

○ Summary





## Dimensionless Boussinesq Equations

$$\nabla \cdot \vec{u} = 0$$
  
$$\partial \vec{u} / \partial t + \vec{u} \cdot \nabla \vec{u} = -\nabla p + v_{d} \nabla^{2} \vec{u} + Ra Pr \Theta \hat{z}$$
  
$$\partial \Theta / \partial t + \vec{u} \cdot \nabla \Theta = \kappa_{d} \nabla^{2} \Theta$$

#### **Control parameters:**

Pr (= $\nu/\kappa$ ) : Prandtl number

Ra (=GrPr=g $\beta$ (T<sub>h</sub>-T<sub>l</sub>)L<sup>3</sup>/v $\kappa$ ): Rayleigh number

AR (=H/L): Aspect ratio





#### RBC vs LHC

**Rayleigh-Bernard convection** (vertical temperature gradient)

• Important in oceanography, geophysics, meteorology, astrophysics, energy process

- Paradigm system to study nonlinear physics, turbulence properties
- Much effort expended and significant insight gained in addressing fundamental aspects of heat transfer and flow characteristics
- Heat-transfer (Nu) /turbulence (Re) scaling with Ra, Pr, AR

Laterally heated convection (horizontal temperature gradient)

- Important in industrial and engineering applications
- Less attention from physics community
- Validation example to compare numerical algorithms
- Flow structure, temperature distribution, AR effect

— stop well short of physical fundamentals







# **Mesoscopic Computation Technique** — Lattice Boltzmann Method $g_{\alpha}(\vec{x} + c\vec{e}_{\alpha}\delta t, t + \delta t) = g_{\alpha}(\vec{x}, t) - [g_{\alpha}(\vec{x}, t) - g_{\alpha}^{(eq)}(\vec{x}, t)]/\tau_{f} + \delta tF_{\alpha}$ $g_{\alpha}^{(eq)} = \lambda_{\alpha} p + s_{\alpha}(\vec{u}) \qquad \tau_{f} = 0.5 + \frac{v_{d}}{3c\delta x} \qquad Discrete \text{ buoyant force} \\ p = (\sum g_{\alpha} + s_{0}(\vec{u}))/\lambda_{0} \qquad \vec{u} = \sum c\vec{e}_{\alpha}g_{\alpha} \qquad \vec{F} = Ra \Pr \Theta \hat{z}$ $\Theta_{\alpha}(\vec{x} + c\vec{e}_{\alpha}\delta t, t + \delta t) = \Theta_{\alpha}(\vec{x}, t) - [\Theta_{\alpha}(\vec{x}, t) - \Theta_{\alpha}^{(eq)}(\vec{x}, t)] / \tau_{T}$ $\Theta_{\alpha}^{(eq)} = \Theta[\omega_{\alpha} + s_{\alpha}(\vec{u})] \qquad \tau_{T} = 0.5 + \frac{\kappa_{d}}{3c\delta x}$ $\Theta = \sum \Theta_{\alpha}$

 $\lambda_a$ ,  $s_a$ ,  $\omega_a$  are lattice related



Shi, et al, Prog. Comp. Fluid Dyn., 5, 50 (2005)



#### Validation: LBM vs. NS-based

	Values	Davis (bench-mark) (1983)	LBM (current)	Ismail (2000)	) I (	D & J (1983)	
Ra=10 <sup>3</sup>	u <sub>max</sub> (0.5,z)	3.65	3.64	3.65		3.59	
	Z	0.81 0.81		0.81		0.81	
	$w_{max}(x, 0.5)$	3.70	3.70	3.70		3.63	
	х	0.18	0.18	0.18		0.18	
	Nu	1.12	1.12	1.12		1.11	
	Values	Davis (bench-mark)	LBM	Ismail	Ding	D&J	
	Waluas				8		
	Values	(1983)	(current)	(2000)	(2004)	(1983)	
	Values u <sub>max</sub> (0.5,z)	(1983) 16.18	(current) 16.15	(2000) 16.16	(2004) 16.06	(1983) 16.19	
$\mathbf{D} = 104$	Values u <sub>max</sub> (0.5,z) Z	(1983) 16.18 0.82	(current) 16.15 0.82	(2000) 16.16 0.82	(2004) 16.06	(1983) 16.19 0.82	
Ra=10 <sup>4</sup>	Values $u_{max}(0.5,z)$ z $w_{max}(x,0.5)$	(1983) 16.18 0.82 19.62	(current)16.150.8219.70	(2000) 16.16 0.82 19.83	(2004) 16.06 — 19.27	<ul> <li>(1983)</li> <li>16.19</li> <li>0.82</li> <li>19.20</li> </ul>	
Ra=10 <sup>4</sup>	Values $u_{max}(0.5,z)$ z $w_{max}(x,0.5)$ x	(1983) 16.18 0.82 19.62 0.12	(current)16.150.8219.700.12	(2000) 16.16 0.82 19.83 0.12	(2004) 16.06 — 19.27 —	<ul> <li>(1983)</li> <li>16.19</li> <li>0.82</li> <li>19.20</li> <li>0.13</li> </ul>	
Ra=10 <sup>4</sup>	Values $u_{max}(0.5,z)$ Z $w_{max}(x,0.5)$ X Nu	(1983) 16.18 0.82 19.62 0.12 2.24	(current)16.150.8219.700.122.24	(2000) 16.16 0.82 19.83 0.12 2.26	(2004) 16.06  19.27  2.17	<ul> <li>(1983)</li> <li>16.19</li> <li>0.82</li> <li>19.20</li> <li>0.13</li> <li>2.21</li> </ul>	

NNS

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#### Validation: LBM vs. NS-based(cont'd)

Ra=10 <sup>5</sup>	Values	Davis (bench-mark)	LBM	Ismail	Ding	D&J
	u <sub>max</sub> (0.5,z)	34.73	35.24	33.42	34.60	36.46
	z	0.86	0.85	0.85		0.85
	w <sub>max</sub> (x,0.5)	68.59	69.10	70.44	69.18	62.79
	x	0.066	0.069	0.067		0.075
	Nu	4.52	4.50	4.65	4.58	4.45

	Values	Davis	Quéré (spectral)	LBM	Ismail	D&J
Ra=10 <sup>6</sup>	u <sub>max</sub> (0.5,z)	64.63	64.83	73.00	57.22	79.27
	z	0.85	0.85	0.85	0.87	0.86
	$w_{max}(x, 0.5)$	219	220	229	220	195
	х	0.038	0.038	0.042	0.045	0.045
4	Nu	8.8	8.8	9.0	8.9	9.0







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#### Steady Flow Structures (magnitude not scaled)



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#### Steady Temperature Contours



Thermally stratified, conduction dominates in thin boundary layers

#### Mean Inverse Boundary Layer Thickness



- TBL precedes VBL, true for Pr<1
- BL meaningless till fluid starts to stratify
- Nu grows faster than  $(\lambda_{\theta})^{-1}$ , conduction dominates, convection still contributes

### Threshold Ra<sub>t</sub>





### Interpretation of 1/3 Power Law

 $\bigcirc$  Ra < Ra. conduction

amos

 $\bigcirc$  Ra<sub>t</sub> = 440, convection takes place, stratification starts

- Mechanism, buoyancy? shear ?
- Relation with thermal boundary layer thickness?

 $\bigcirc$  Higher enough Ra (stratified), boundary layer/core region apparent, significant horizontal temperature gradients are in two boundary layers with thickness  $\lambda$ 

$$j_{heat} = (T_h - T_l) k / \lambda, \ j_{cond} = (T_h - T_l) k / L$$
  
Ra= g $\beta$ (T<sub>h</sub>-T<sub>l</sub>)L<sup>3</sup>/v $\kappa$ , Ra<sub>t</sub> = g $\beta$ (T<sub>h</sub>-T<sub>l</sub>)  $\lambda$ <sup>3</sup>/v $\kappa$   
Nu=j<sub>heat</sub>/j<sub>cond</sub> =L/  $\lambda$  = (Ra/Ra<sub>t</sub>)<sup>1/3</sup> = 0.13Ra<sup>1/3</sup>

•This work: Nu=0.134Ra<sup>0.305</sup> (Ra 10<sup>3.5</sup>-10<sup>7</sup>) •Turbulent RBC: Nu=0.134Ra<sup>0.309</sup> (Ra 10<sup>6</sup>-10<sup>17</sup>) Niemela et al. *Nature* (2000)

Mixing length theory, Spiegel, Annu. Rew. Astron. Astrophys. (1971)



#### Summary: Factors Driven Near 1/3 Power-law

### Nu ~ $aRa^{\beta}$

O Large scale circulation present

 $\bigcirc$  Interior thermal stratification  $\rightarrow \beta$ 

 $\bigcirc$  Threshold Rayleigh number Ra<sub>t</sub> identified  $\rightarrow$  a







○ More sophisticated or rigorous analysis

○ Nu dependency on AR

○ Nu dependency on Pr

O Power-law scaling in turbulent convection

• Experiments?



