



**The Abdus Salam  
International Centre for Theoretical Physics**



**2024-3**

## **Spring School on Superstring Theory and Related Topics**

***23 - 31 March 2009***

**Holography and strongly coupled model building**

**Lecture 3**

S. Kachru  
*Stanford University*  
*U.S.A.*

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Tneste '09, Lecture #3

Kachru

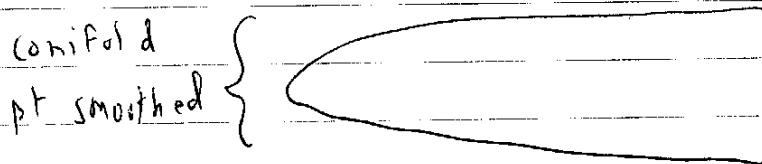
### III. Solutions with exp small SUSY

At the end of the last lecture, we had constructed gravity solutions dual to the gauge theory ( $N=1$ )

	$SU(M+N)$	$SU(N)$
$A_{1,2}$	$\overline{M+N}$	$\overline{N}$
$B_{1,2}$	$M+N$	$N$

$$W_{\text{tree}} \sim \text{tr} ABAB$$

For  $N = kM$   $k \in \mathbb{Z}^+ \Rightarrow$



$$\sum z_i^2 = 0 \rightarrow \sum z_i^2 = \epsilon^2 \quad \epsilon \approx e^{-\frac{2\pi k}{gM}}$$

Tip geometry  $\sim$  a large  $S^3$  pierced by  $M$  units of  $F_3^{\text{RR}}$  flux

$$F_3 \sim f \epsilon_{ijk}$$

$$a_0^2 = e^{-\frac{4\pi k}{3gM}}$$

$$f \approx \frac{2}{\sqrt{g_5^3 M} b_0} \quad U(1)$$

$$ds^2 = a_0^2 dx^\mu dx_\mu + g_5 M b_0^2 \left( \frac{1}{2} dr^2 + r^2 d\tilde{\Omega}_2^2 + d\Omega_3^2 \right)$$

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a) Susy solutions via D-brane probes (hep-th/0112157)

Natural idea: Given the tiny warped scale at the IR tip, can we locally do Susy there to generate  $F_{\text{Susy}} \sim a_0^2 \sim e^{-4\pi k z / 3g_m}$ ?

Sure.

Consider a  $\overline{D3}$  probe in the KS solution with  $N = kM$ .

- No "extra  $\overline{D3}$ s" in smooth SUGRA backgrd  $\rightarrow$

$D3 / \overline{D3}$  annihilation is impossible, despite the presence of  $N$  units of  $D3$  charge.

- Dynamics? Determined by the DBI action on the probes.

$$S_{\overline{D3}} = - \frac{T_3}{g_s} \int d^4x \text{Tr} \sqrt{\det(G_{\mu\nu}) \det(Q)}$$

$$- T_3 \int \text{Tr} [2\pi i \tilde{L}_\mu \tilde{L}_\nu B_6 + C_4]$$

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where:

$$Q^i_j = \delta^i_j + \frac{2\pi i}{g_s} [\Phi^i, \Phi^k] (G_{kj} + g_s C_{kj})$$

and

$$\tilde{L}_\Phi \tilde{L}_\Phi B_6 = \Phi^n \Phi^m B_{mnpqrs} \frac{dy^p dy^q \dots dy^s}{4!}$$

Note:  $dB_6 = \frac{1}{g_s^2} \star_{10} H_3 = -\frac{1}{g_s} dV_4 \wedge F_3$

with  $dV_4 = d\phi^4 d^4x$ , & used ISD nature of the KS solution which relates  $H_3$  &  $F_3$  ( $\star_6 G_3 = i G_3$  w/  $G = F - \tau H$ ),

• First, imagine a single  $\overline{D3}$  probe  $\rightarrow$  drop all commutator terms in  $S_{\overline{D3}}$ :

$$S = \frac{T_3}{g_s} \int d^4x \sqrt{g_4} \text{Tr} e^{\phi_A} \left[ 2 + \frac{1}{2} e^{-2A} \partial^\mu \Phi^i \partial_\mu \Phi^j g_{ij} \right]$$

So } a potential  $\sim e^{\phi_A} \rightarrow$

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$\overline{D3}$  [or COM of  $p > 1$   $\overline{D3}$ s]  
attracted to the TIP.

What is the dynamics AT the tip? Expanding  
out the matrices in  $S_{\overline{D3}} \rightarrow$

$$V_{\text{eff}} = \frac{T_3}{g_s} \left\{ p - i \frac{4\pi^2}{3} f \epsilon_{ijk} \text{Tr} [\Phi^i, \Phi^j] \Phi^k - \frac{\pi^2}{g_s^2} \text{Tr} ([\Phi^i, \Phi^j]^2) + \dots \right\}$$

Where are the extrema?

Not hard to see that  $\exists$  extrema where

$$[\Phi^i, \Phi^j] = -i g_s^2 f \epsilon_{ijk} \Phi^k$$

-- which up to a rescaling, is just the  
commutation relations satisfied by  $SU(2)$   
generators!

•  $p$   $\overline{D3}$ s  $\rightarrow$  need a  $p$  dim'l  $SU(2)$  representation

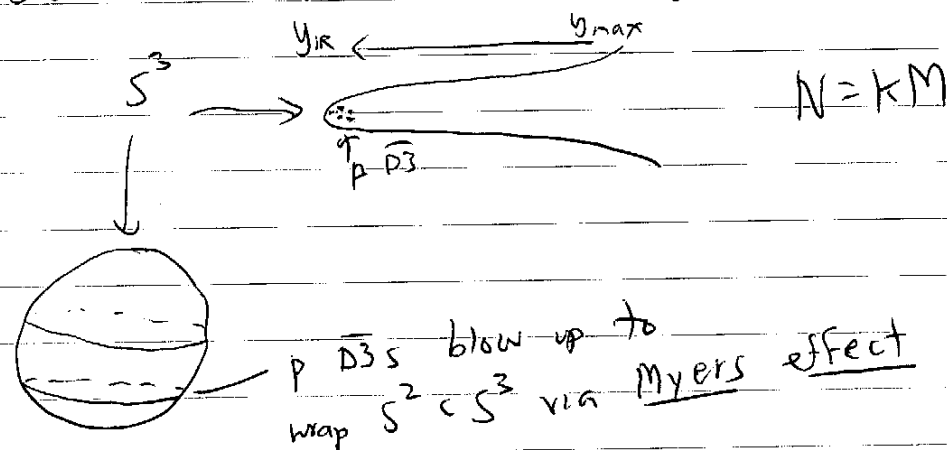
(5)

- $\exists$  critical pts for each one, ie any partition of  $p$
- Minimal  $V \rightarrow p$  dim'l irrep
- Can check that the radius of the blob of expanded  $\overline{D3}$  is

$$R^2 \approx \frac{4\pi^2}{M^2} p^2 \times (R_{S^3})^2$$

$\rightarrow$  only reliable if  $p \ll M$ ; otherwise, blob  $\approx$  size of space @ the tip!

So in the gravity regime,  $N$   $D3_s$  +  $M$   $D5_s$  +  $p$   $\overline{D3}_s$  has one branch of  $\{\text{vacua}\}$  where:

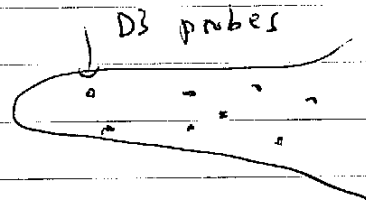


$$\text{SUSY energy} \sim p T_{\overline{D3}} \times e^{-\frac{8\pi k}{3gM}} \left. \vphantom{\text{SUSY energy}} \right\} \text{exp small.}$$

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Note that  $\exists$  another state, visible in same gravity theory, w/ same charges:

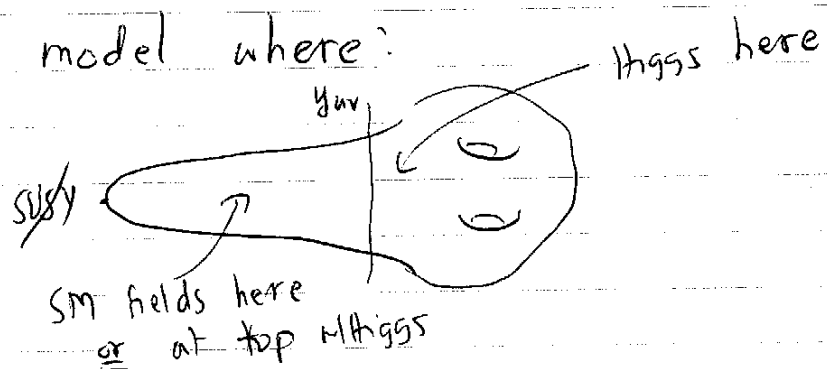
- KS solution with  $\tilde{N} = (k-1)M$ ,  $\tilde{M} = M$ ,  
and  $M-p$  D3s:



In fact,  $\exists$  a brane/flux annihilation process that allows ~~SUSY~~ states to decay to these SUSY states. But, metastable if  $p \ll M$ .

- b) SUSY solutions as normalizable perturbations of IIB SUGRA modes in conifold throat  
(cf ~~hep~~ arXiv: 0801.1520)

Our eventual goal is to use these SUSY states in a model where:



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As long as no SM modes have support @ the tip, the SUSY effects on SM should be capturable if we know the SUSY gravity solution away from the tip at  $r \gg \epsilon^{2/3}$ .

In the SUSY theory with  $M$  D5s and  $N$  D3s, this sol'n was given by Klebanov & Tseytlin.

$$ds^2 = e^{2A(r)} \eta_{\mu\nu} dX^\mu dX^\nu + e^{-2A(r)} \left[ dr^2 + r^2 e_\psi^2 + r^2 \left( \sum_{i=1}^2 e_{\theta_i}^2 + e_{\phi_i}^2 \right) \right]$$

where :

$$e_\psi = \frac{1}{3} \left( d\psi + \sum_{i=1}^2 \cos \theta_i d\phi_i \right)$$

$$e_{\theta_i} = \frac{1}{\sqrt{6}} d\theta_i$$

$$e_{\phi_i} = \frac{1}{\sqrt{6}} \sin \theta_i d\phi_i$$

The warp factor is given by :



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$$e^{-4A} = \frac{27\pi g_s}{4r^4} \left[ N + \frac{3g_s M^2}{2\pi} \left( \log \left( \frac{r}{r_{uv}} \right) + \frac{1}{4} \right) \right]$$

$$T_{DB} = \text{constant}$$

$$F_5 = (1 + *) \cdot -27\pi \left[ N + \frac{3g_s M^2}{2\pi} \log \left( \frac{r}{r_{uv}} \right) \right]$$

$$\times \text{Vol}(T^{1,1}) \leftarrow = e_\psi \wedge e_{\theta_1} \wedge e_{\phi_1} \wedge e_{\theta_2} \wedge e_{\phi_2}$$

+ 3-form fluxes:

$$F_3 = \frac{g_s M}{2} e_\psi \wedge (e_{\theta_1} \wedge e_{\phi_1} - e_{\theta_2} \wedge e_{\phi_2})$$

$$B_2 = \frac{g_s M}{2} [e_{\theta_1} \wedge e_{\phi_1} - e_{\theta_2} \wedge e_{\phi_2}] \log \left[ \frac{r}{r_{uv}} \right]$$

This kT solution preserves the full  $SU(2) \times SU(2)$  global symmetry of the conifold.

The SUSY state:

- Open string probe analysis  $\rightarrow$  embiggened  $\overline{D3}$ s @ the IR end of the geometry yield SUSY.
- At long distance from tip, should be able to find a gravity sol'n encapsulating SUSY

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(a "KT" version of the  $SU(2)$  state).

The smeared  $SU(2)$  solution is:

$$ds^2 = r^2 e^{2a(r)} \eta_{\mu\nu} dx^\mu dx^\nu +$$

$$e^{-2a(r)} \left[ \frac{dr^2}{r^2} + e^{2b(r)} e_\psi^2 + e^{2c(r)} \sum_i (e_{\phi_i}^2 + e_{\tilde{\phi}_i}^2) \right]$$

where - at leading order in  $M/N$

- at leading order in  $S = p e^{\frac{-8\pi k}{3gM}}$

$$e^{-2a} = \left( \frac{1}{2} + \frac{S}{32r^4} \right) \sqrt{27\pi gN}$$

$$e^{2b} = 1 + \frac{S}{r^4}$$

Dilaton  $\phi = \log g_s + \frac{1}{r^4} [-3S \log(r)]$

+ ... ( $\Delta$  3-form fluxes at  $O(S)$ )

The UPSHOT of all of the changes, is that:

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- The  $T^{11}$  is "squashed"
- Non- $ISD$   $(1,2)$  flux sourced
- $\phi$  now runs

All are  $O(S')$ , normalizable perturbations;  
this is a SUSY state in KS field theory.

#### IV. Coupling SUSY to SM

##### a) Adding a SM / GUT gauge group

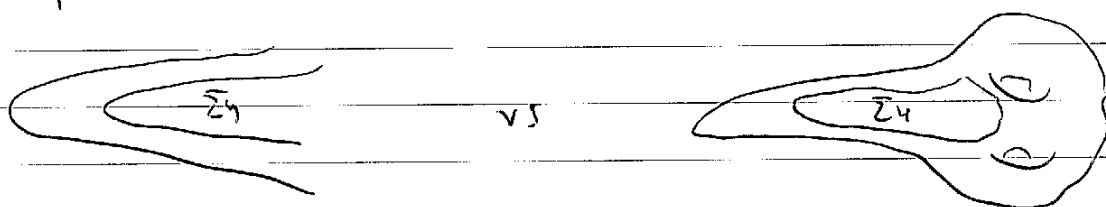
Compared to the strongly coupled KS theory, the SM group is weakly ganged.

So we should try to embed it in the throat by adding eg an  $SU(5)$  global symmetry.

Intuitively, we can get new gauge sector by adding D7s in throat. The D7 gauge fields are at  $g_{YM}^{D7} \rightarrow 0$  in non-compact

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limit, but have finite  $g_{\text{YM}}^{\text{D7}}$  after compactification:



$$\frac{1}{(g_{\text{YM}}^{\text{D7}})^2} \sim \frac{\text{Vol}(\Sigma_4)}{g_s} \rightarrow \begin{cases} \infty & \Sigma_4 \text{ non-compact} \\ \text{finite} & \Sigma_4 \text{ compact} \end{cases}$$

Want SUSY, MSSM or GUT, so choose a SUSY D7 embedding  $\rightarrow$  specified by a holomorphic embedding (we'll avoid gauge bundle subtleties for now).

Easiest choice:

Conifold  $\sum z_i^2 = \epsilon^2$

$\Sigma_4$

$\Sigma_4 = \mathcal{M}$

"Kuperstein embedding"

$k$  D7s on  $\Sigma_4 \Rightarrow$  get additional  $SU(k)$

"gauge" group!  $k \geq 5 \rightarrow$  this is good enough.