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Some peculiarities of the polar summer mesosphere

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# Some peculiarities of the polar summer mesosphere

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# OUTLINE

- Polar summer mesosphere: important parameters
- Charging of an isolated dust particle in a plasma
- Impact of polar mesospheric clouds on ionospheric plasma: the simple model
- Heating of mesospheric particles by the solar radiation
- Growth/sedimentation/charging model of polar mesospheric clouds
- Ice particles formation in the RF chamber: experimental data and simulations
- Conclusions



N.B. Source of smoke particles in mesosphere: meteoroid's ablation & soot from volcanic eruptions

#### **Polar summer mesosphere (cntnd)**



#### Charging of isolated particle in mesospheric plasma

$$\begin{split} \frac{\partial n_e}{\partial t} &= q_e - \alpha_s^{\text{rec}} n_e n_i^s - \alpha_c^{\text{rec}} n_e n_i^c + \\ &+ \pi \langle \nu_{ed}^{\text{photo}} a^2 n_d \rangle - \pi \langle \nu_{ed} a^2 n_d \rangle, \\ \frac{\partial n_i^s}{\partial t} &= q_e - \alpha_s^{\text{rec}} n_e n_i^s - \beta_c n_i^s - \pi \langle \nu_{sd} a^2 n_d \rangle, \\ \frac{\partial n_i^c}{\partial t} &= \beta_c n_i^s - \alpha_c^{\text{rec}} n_e n_i^c - \pi \langle \nu_{cd} a^2 n_d \rangle, \\ \frac{\partial Z_d^a}{\partial t} &= \nu_d^{\text{photo}} + \nu_{sd} + \nu_{cd} - \nu_{ed}, \end{split}$$
$$\nu_{ed} = \begin{cases} \pi a^2 \left(\frac{8T_e}{\pi m_e}\right)^{\frac{1}{2}} n_e \exp\left(\frac{e^2 Z_d}{a T_e}\right) & \text{if } Z_d < 0, \\ \pi a^2 \left(\frac{8T_e}{\pi m_e}\right)^{\frac{1}{2}} n_e \left(1 + \frac{e^2 Z_d}{a T_e}\right) & \text{if } Z_d > 0. \end{cases}$$
$$\nu_{id} = \begin{cases} \pi a^2 \left(\frac{8T_i}{\pi m_i}\right)^{\frac{1}{2}} n_i \left(1 - \frac{e^2 Z_d}{a T_i}\right) & \text{if } Z_d < 0, \\ \pi a^2 \left(\frac{8T_i}{\pi m_i}\right)^{\frac{1}{2}} n_i \exp\left(-\frac{e^2 Z_d}{a T_i}\right) & \text{if } Z_d > 0. \end{cases}$$
$$\nu_{ed}^{\text{photo}} = \pi a^2 \int_{-\infty}^{N_W} \sigma_{abs}(\lambda, a, m) \Phi_\lambda Y(\lambda, m, a) d\lambda, \end{cases}$$

 $J_0$ 



• Impact of collisions on the charge of isolated particle: increase of ion flux due to the trapped particles, so decrease the particle charge.

#### **Optical properties of mesospheric particles**



 Metals: low work functions (few eV), very effective photoelectric effect even for visible wavelenghts.

• Water ice: high work function (about 8 eV), photoelectric effect possible only for UV light (Lyman-alpha hydrogen)

# **Charging of dust particles (cntnd)**



Dimensionless grain charge versus ion to electron temperature ratio and ion mass Dust grain charge (OML) in PSM versus dust number density and effective solar flux

#### Impact of dust on plasma parameters



#### Heating of dust particles in the upper atmosphere

$$P^{\text{sol}} + P^{\text{rad}} + P^{\text{col}} = 0$$

$$P^{\text{sol}} = \pi a^{2} h c \int \sigma_{\text{abs}}(\lambda, a, m) (\Phi_{\lambda}^{\text{sol}}(H, \alpha_{\text{sa}}) + \Phi_{\lambda}^{\text{sca}}(H, \alpha_{\text{sa}})) d\lambda/\lambda,$$

$$P^{\text{rad}} = \pi a^{2} h c \int \sigma_{\text{abs}}(\lambda, a, m) B_{\lambda}(T_{d}) d\lambda/\lambda,$$

$$P^{\text{col}} \approx 4\pi a^{2} \alpha_{\text{T}} n_{n} v_{th} k_{\text{B}}(T_{d} - T_{n})$$

$$\int_{0}^{0} \frac{1}{10^{1}} \frac{1}{10$$



$$egin{aligned} &rac{\partial f_d}{\partial t} + rac{lpha_{
m w}}{4} m_{
m w} v_{
m w}^T 
ho_d (n_{
m w} - n_{
m w}^s) rac{\partial f_d}{\partial a} + v rac{\partial f_d}{\partial H} + \ &+ \left[ g - rac{\pi 
ho c_s a^2 F_d (v + v_{
m wind})}{m_d} 
ight] rac{\partial f_d}{\partial v} = 0. \end{aligned}$$





$$\begin{split} \frac{\partial f_d}{\partial t} &+ \frac{\alpha_{\rm w}}{4} m_{\rm w} v_{\rm w}^T \rho_d (n_{\rm w} - n_{\rm w}^s) \frac{\partial f_d}{\partial a} + v \frac{\partial f_d}{\partial H} + \\ &+ \left[ g - \frac{\pi \rho c_s a^2 F_d (v + v_{\rm wind})}{m_d} \right] \frac{\partial f_d}{\partial v} = 0. \end{split}$$



PMC's particle size distribution vs time

Particles growth/sedimentation in phase space: (r, h, t)





#### NLC: growth/sedimentation model (cntnd)





## Ice particle formation in RF chamber

#### (S. Shimizu, T. Shimizu, 2009)



Pressure = 1 mbar, oxygen/hydrogen = 0.1 Typical size of ice particle is about 1 micron; the timescale <1s

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# Conclusions

- Simple model of dust-induced plasma chemistry of polar summer mesosphere is presented.
- Presence of dust particles significantly changes the plasma parameters @ very moderate dust densities:
  - both depletion and increase of electron density are possible
  - complicated behavior of ion density
  - fast charge-induced coagulation of dust particles
  - significant charge (as positive as negative) of dust particle depending on the photoelectric properties of the particle
- Growth/sedimentation model reproduces most of the NLC data
- Possible impact on neutral chemistry
- Water ice particles production if RF chamber was studied (both experimentally and numerically)