



**2333-27**

**Workshop on Science Applications of GNSS in Developing Countries (11-27 April), followed by the: Seminar on Development and Use of the Ionospheric NeQuick Model (30 April-1 May)**

*11 April - 1 May, 2012*

**Introduction to Space Weather and its Impact on our Daily Lives**

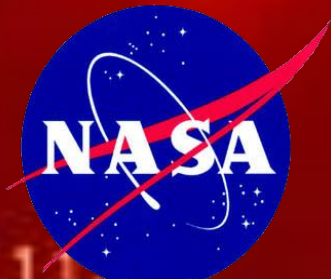
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# Introduction to Space Weather and its Impact on our Daily Lives

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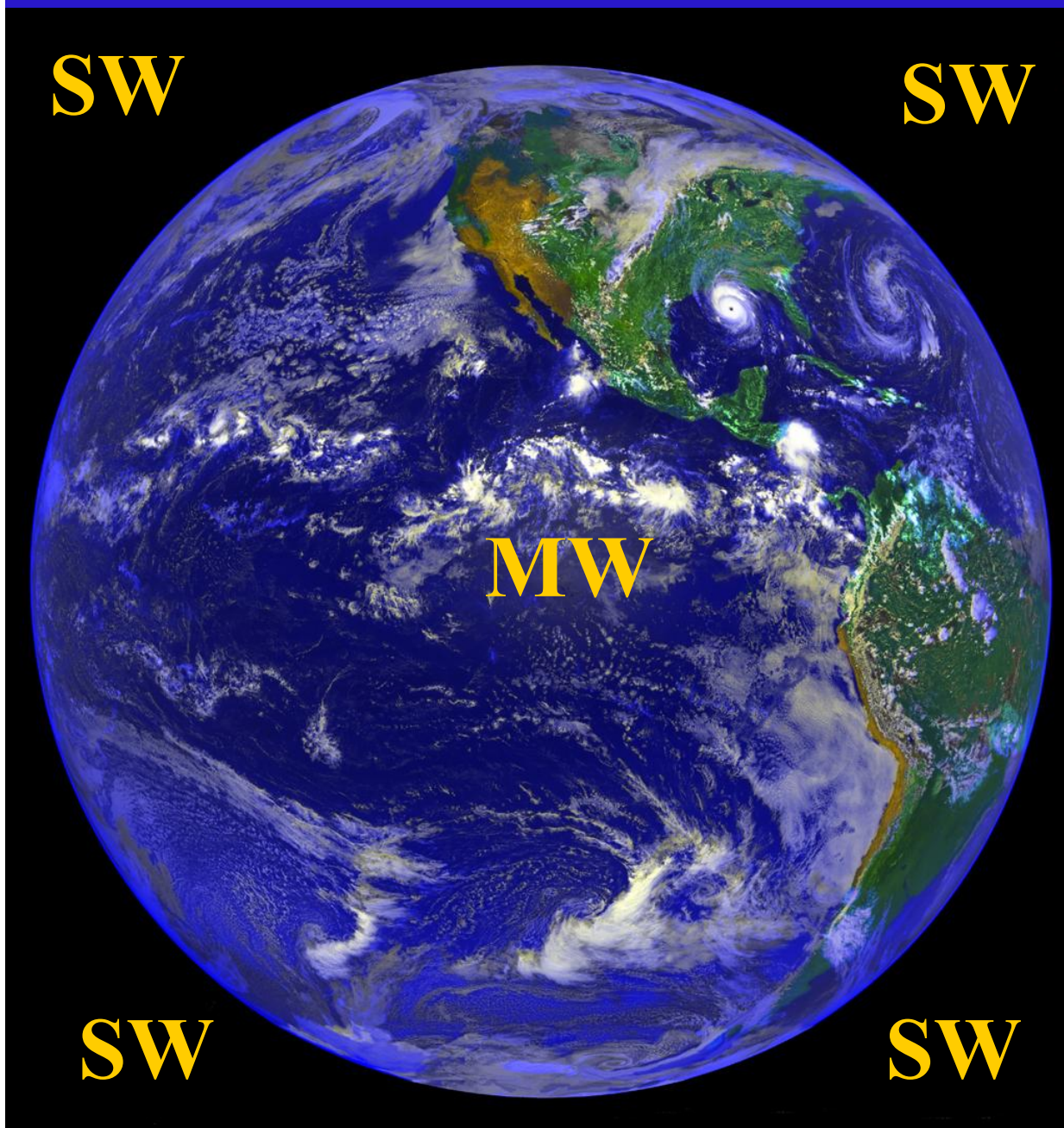




# Outline

- Introduction about space weather
- Space Weather impacts and the consequence of economic loss
- Space weather impact on Communication and navigation systems
- Space weather impacts due to lower thermospheric tides
- Space weather impacts due to tsunami and/or earthquake generated Gravity wave
- Space weather impacts due to lunar tides
- Remark and conclusion

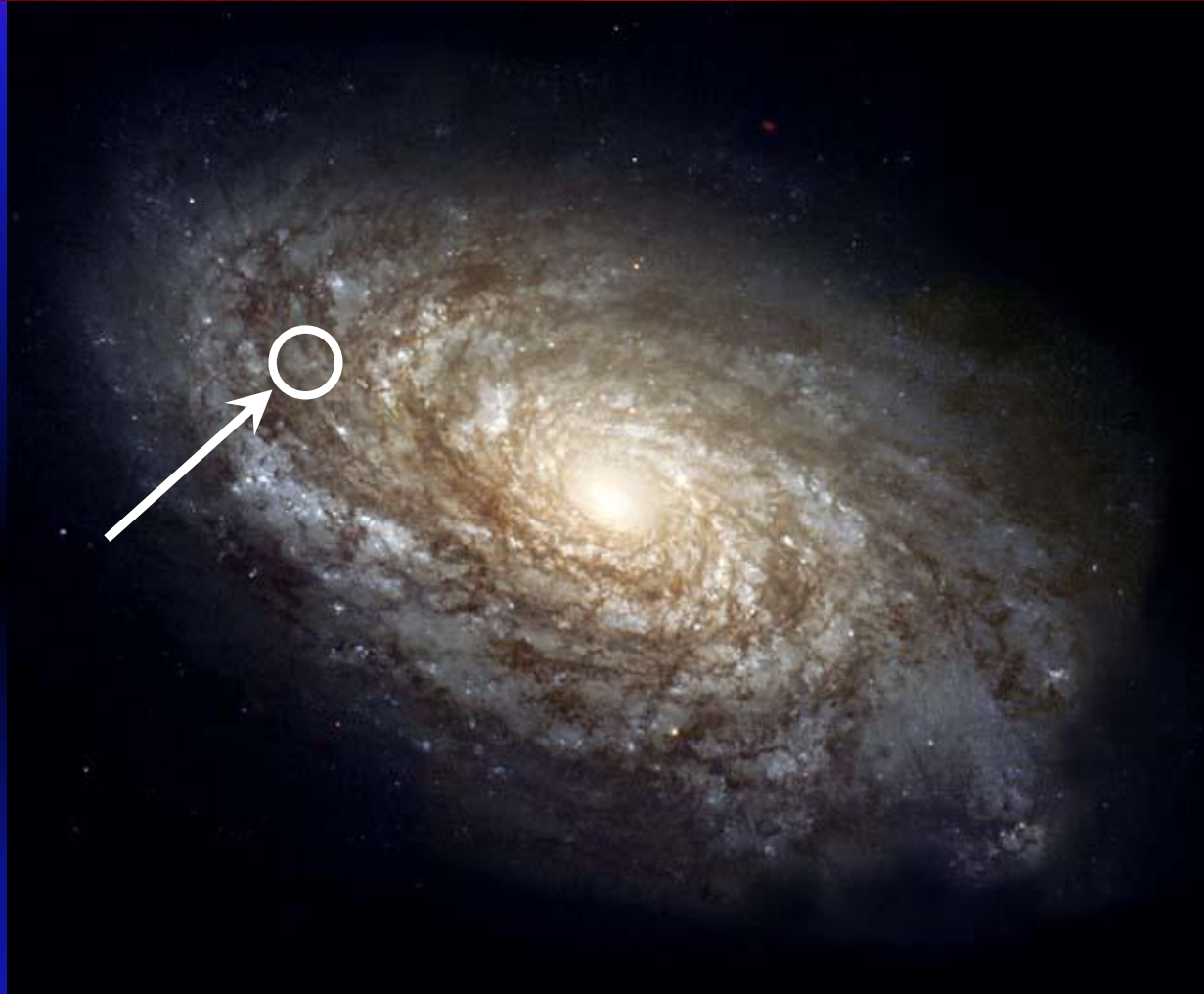
Who do we have to  
blame for the  
source of space  
weather problems?



The SUN  
of course!

It can unleash the  
energy of 100 billion  
atomic bombs

# The Sun's place in the Galaxy



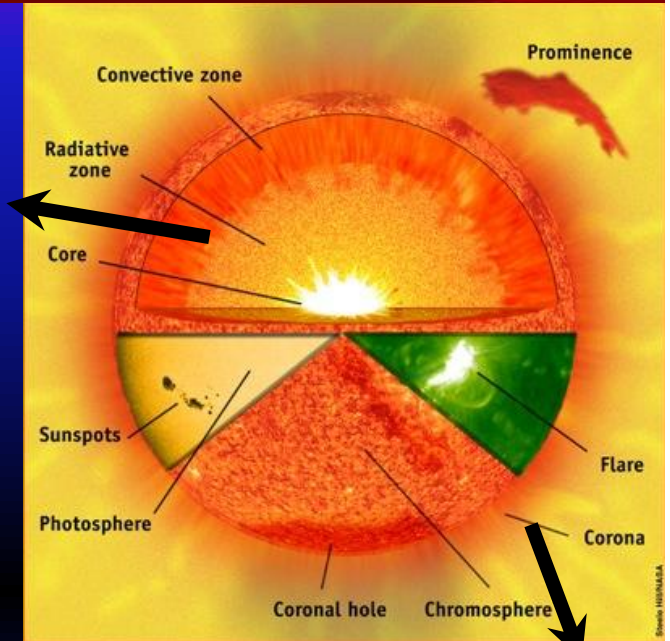
The Sun is one of about 200 billion stars in a galaxy we call Milky Way. It resides on the outskirts, about 28,000 light years from the center.



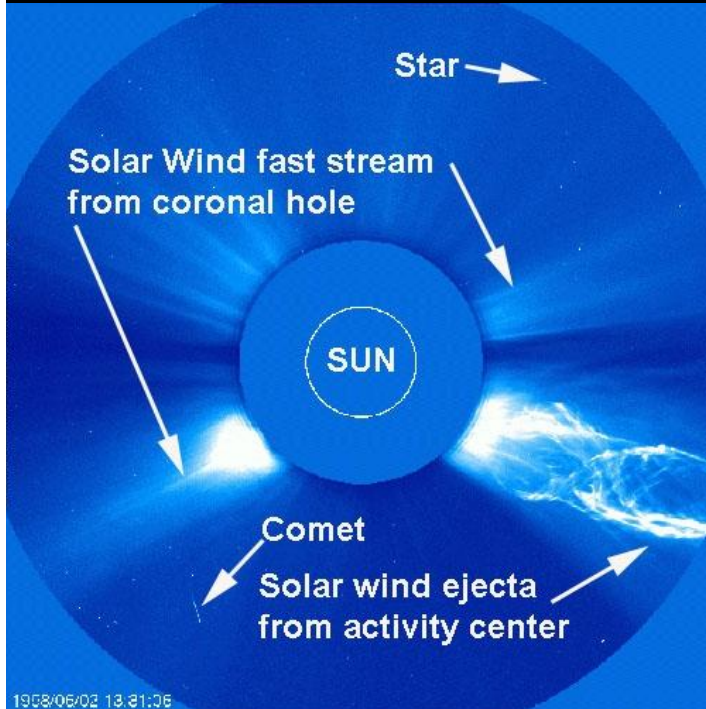
# What is the Solar mysteries?



6000 K



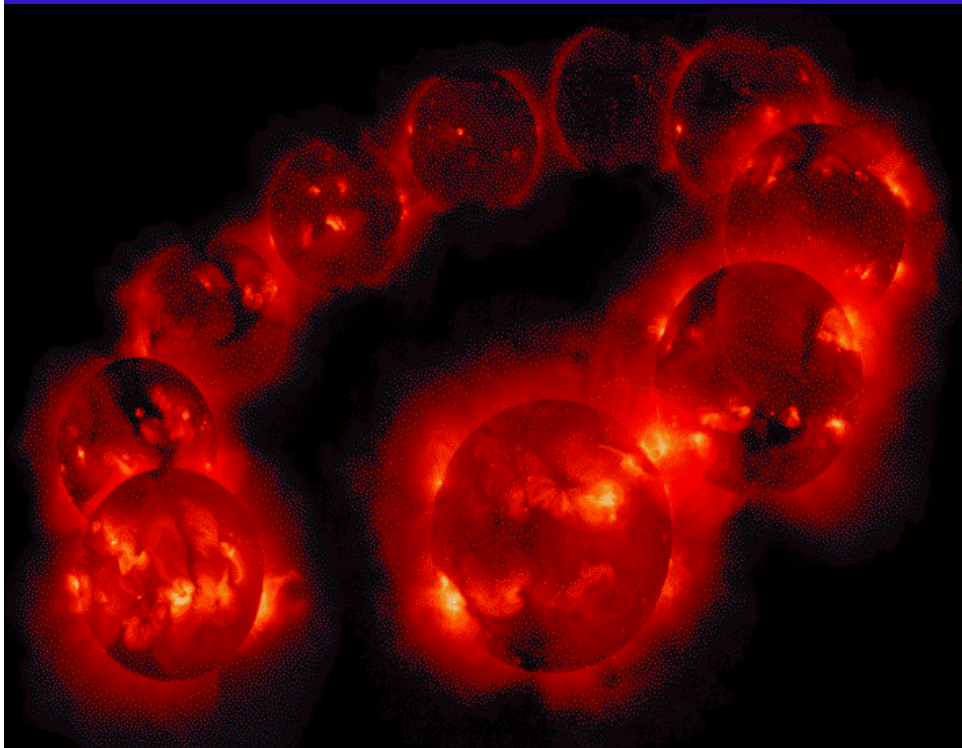
Millions of C



**Current Speculation:** powerful MHD waves (specifically Alfven waves) propagating along magnetic field-lines from the solar interior to the corona, interacting with the plasma and heating it to astonishing temperatures

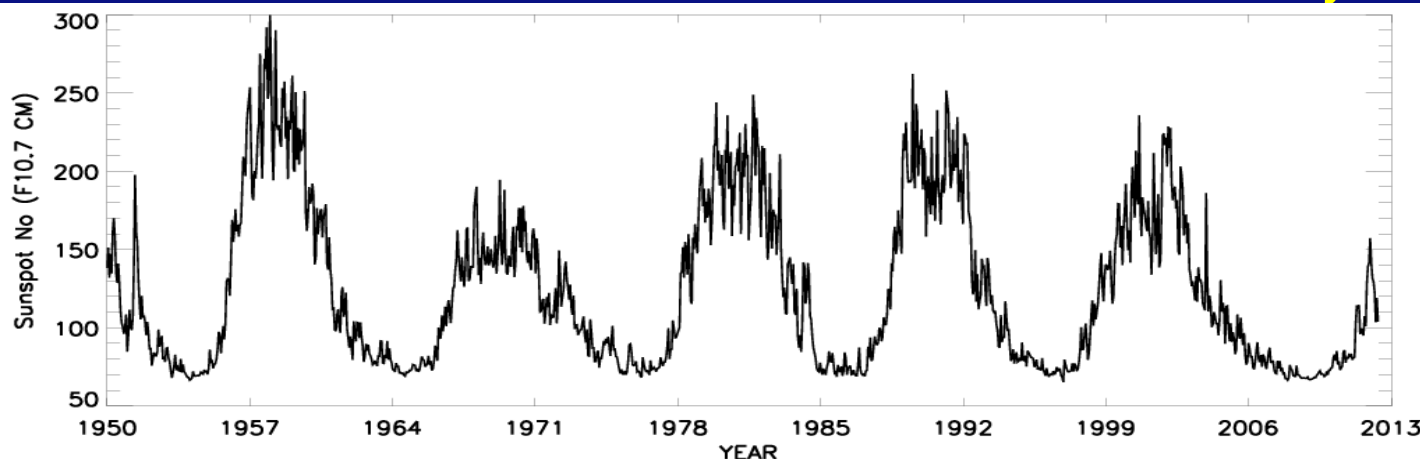
**NASA plan to solve the mysteries and decided to send Solar Probe Plus deep into corona in 2018**

# Just facts about the Sun



Taken by SOHO Satellite

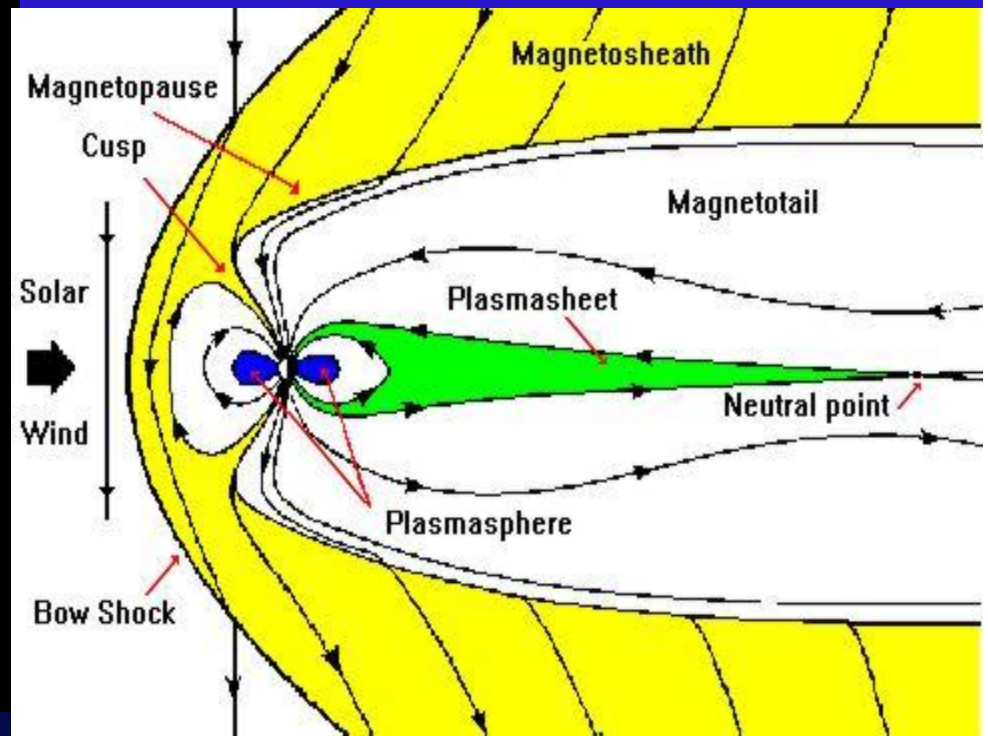
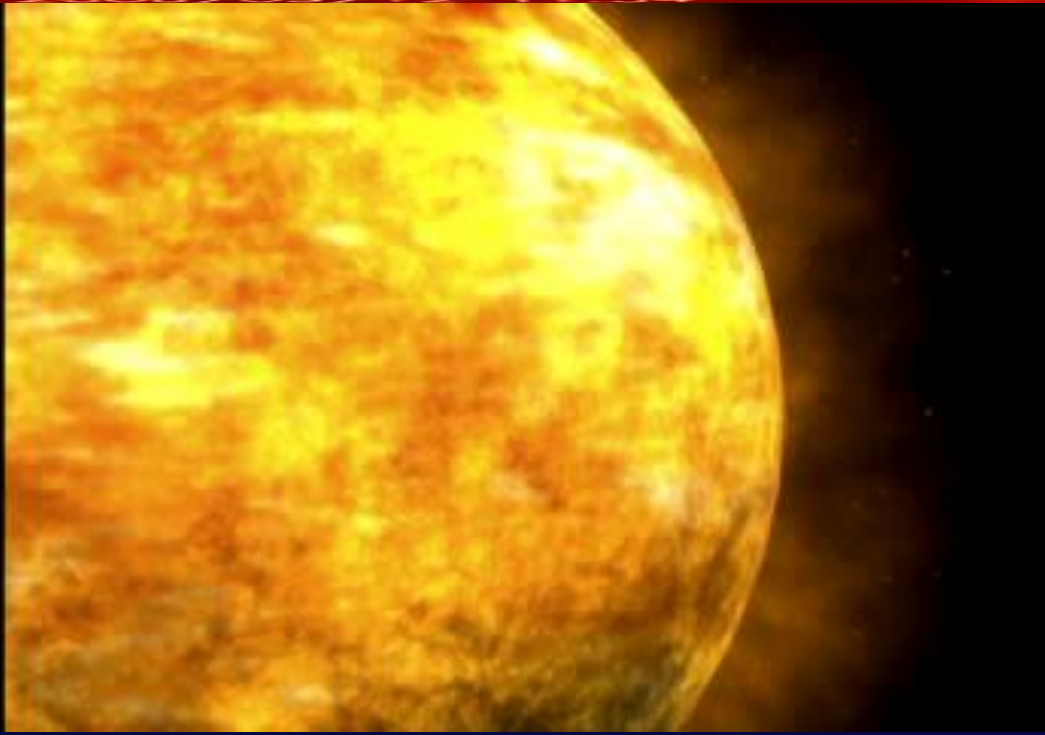
- ➔ 150 million km from Earth
- ➔ 1.4 million km across = 109 **Earth's** diameter
- ➔ 4.5 billion years old
- ➔  $2 \times 10^{30}$  kg mass. (333,000 ME)
- ➔ 75% hydrogen, 24% helium, 1% everything else
- ➔ Makes up 99.98% of the mass of the solar system



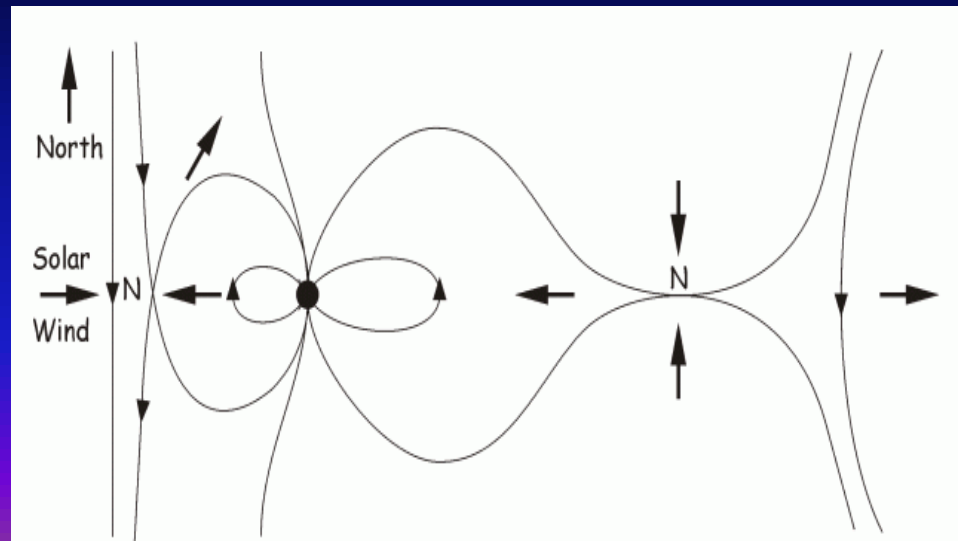
We're heading to  
solar maximum  
era



# Why do we study the Sun?

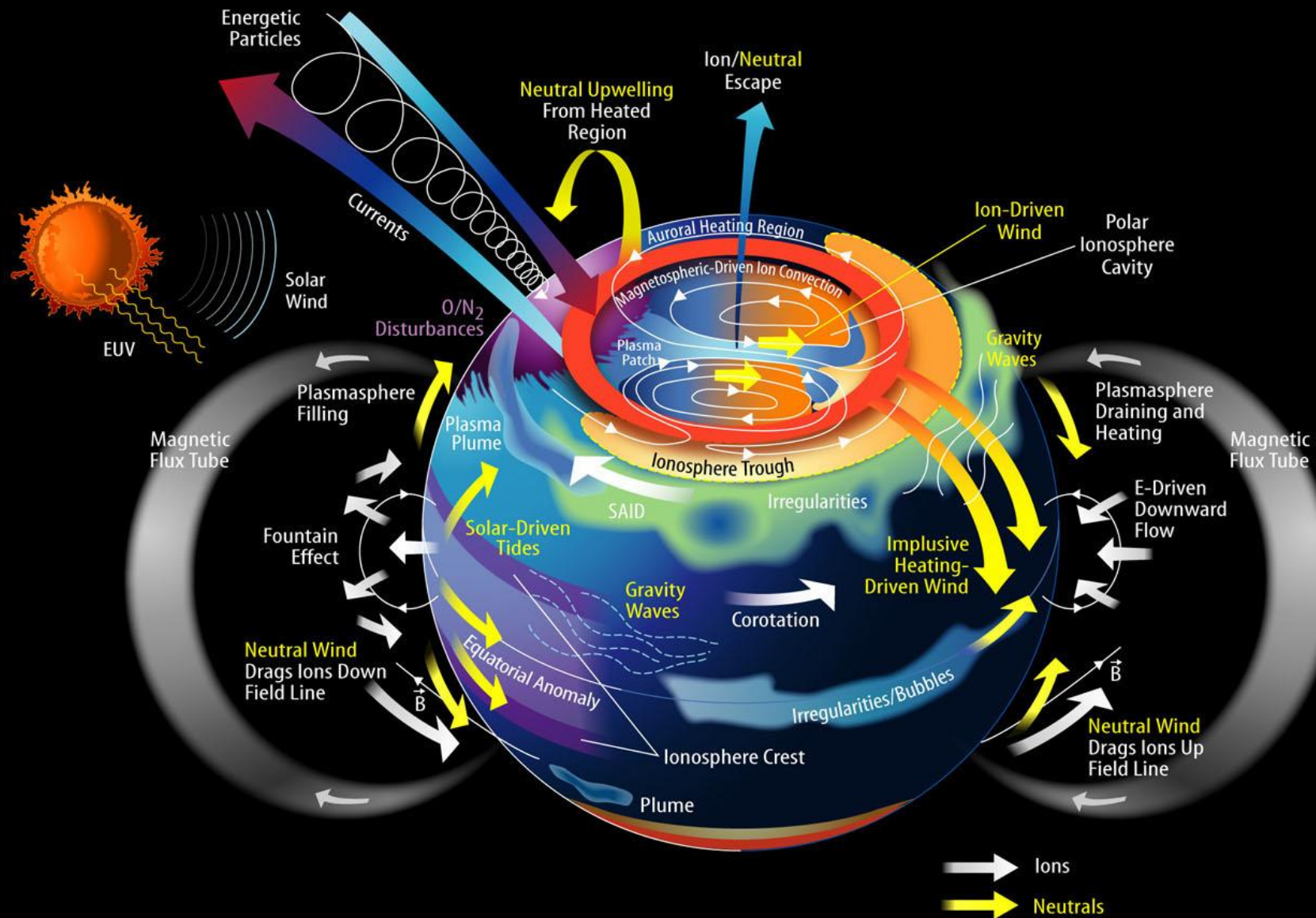


CMEs are bubbles of plasma wrapped in a magnetic field, and if  $B_z$  turns south this magnetic bubble can penetrate into our magnetosphere and inject high energy particles (mainly protons) deep into the Earth's atmosphere.



# Why do we care about Sun-Earth connection?

## Addition of Geomagnetic Storms





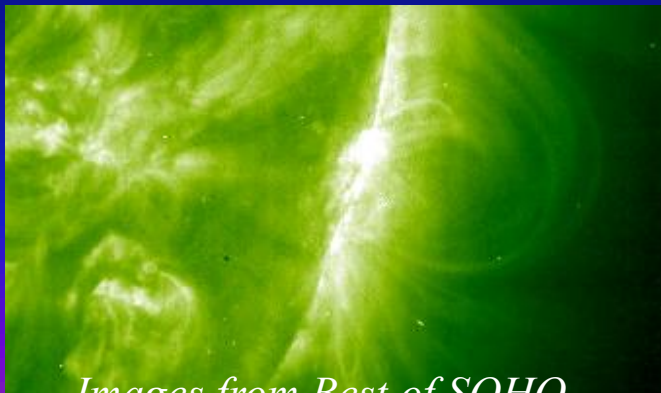
# 1. Radio Blackouts 2. Radiation Storms 3. Geomagnetic Storms

- Solar Flares send out x-rays
- Arrive at Earth in 8 minutes
- Modify the ionosphere
- Degrades HF radio communication blackouts possible

## Impacts:

- Airline communication
- HF radio operators
- DoD Communications
- Satellite Communications

*Nov 4, 2003*



*Images from Best of SOHO*

- Solar Flares and CMEs send out Energetic Particles
- Increasing radiation
- Arrive at Earth in 15 minutes to 24 hours
- Modify the high latitude ionosphere
- Disrupt HF radio communication

## Impacts:

- Airline communication
- Degradation of satellite tracking and power systems
- Radiation hazards to humans flying over the poles or at high altitudes astronauts
- GPS Navigation Errors

*Doherty, 2010*

- CMEs send out bursts of Magnetic Clouds
- Arrive at Earth in 1-4 days
- Strikes magnetosphere and can result in accelerated particles to enter the ionosphere

## Impacts:

- HF radio communication
- GPS Navigation Errors
- Electric Power Grids
- Increased Satellite Drag
- Weaken **Earth's** magnetic field



*13 April, 2010*



# How many type of storms do we know?

CME-driven Storms	CIR-driven Storms
→ Generated by CME sheaths, magnetic clouds, and ejecta;	→ Generated by the associated recurring high-speed streams.
→ Cause denser plasma sheets, strong ring currents and Dst, have solar energetic particle events, can produce great auroras and dangerous geomagnetically induced currents	→ Have longer duration, have hotter plasmas and stronger spacecraft charging, and produce high fluxes of relativistic electrons.
→ CME-driven storms pose more of a problem for Earth-based electrical systems	→ CIR-driven storms pose more of a problem for space-based assets

# How many type of storms do we know?

Phenomenon	CME	CIR
Phase of the solar cycle	Solar max.	Declining phase
Occurrence pattern	irregular	27-day repeating
Calm before the storm	sometimes	usually
Solar energetic particles (SEP)	sometimes	none
Storm sudden commencement (SSC)	common	infrequent
Formation of new radiation belts	Sometimes	no
Dipole distortion	Very Strong	Strong
ULF pulsations	Shorter duration	Longer duration
Spacecraft surface charging	Less severe	More Severe
Ring current (Dst)	Stronger	Weaker
Great aurora	Sometimes	Rare
Geomagnetically induced current (GIC)	Sometimes	No

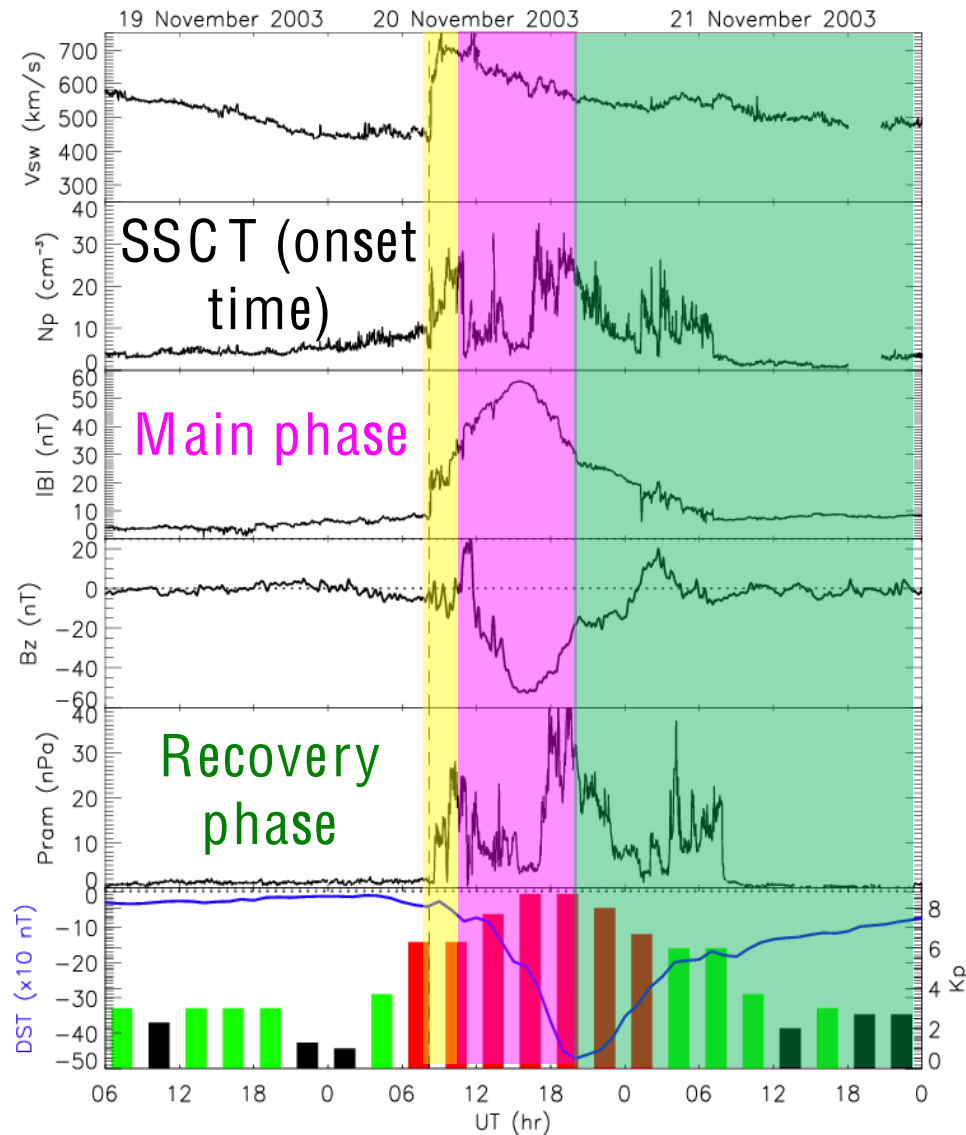
*Borovsky and Denton, JGR, 2006*

# How do we know whether there is storm or not?

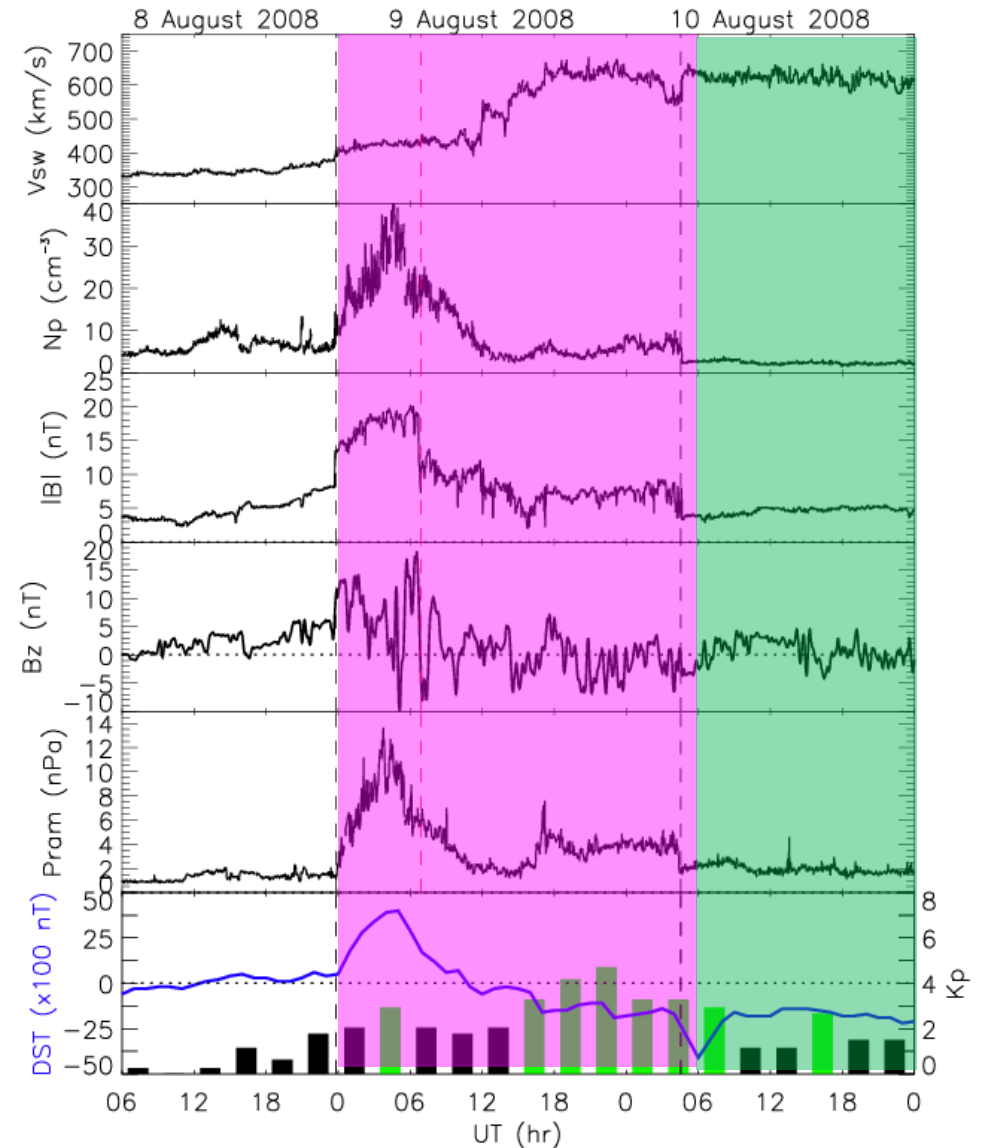
## Geomagnetic indices

- ( Kp index: is obtained as the mean value of the geomagnetic disturbance levels in the two horizontal field components, observed at 13 selected, subauroral stations. The name *Kp* originates from "*p*lanetarische *K*ennziffer" (= planetary index).
- ( Dst index: is a measure of geomagnetic activity used to assess the severity of magnetic storms based on the average value of the horizontal component of the Earth's magnetic field measured hourly at four near-equatorial geomagnetic observatories.
- ( AE index: is derived from geomagnetic variations in the horizontal component observed by magnetograms from observatories at several different longitudes around the auroral zone.
- ( IMF Bz: is much more important than the remaining IMF's component, and can be used as useful indicator for external energy injection into higher latitude region.

# How do we know whether there is storm or not?



CME-driven storm (with clear shock)



CIR-driven storm (without clear shock)

# Geomagnetic Storms effects

Kp  
value # of storm  
events (# of  
storm days)

Extreme

Power systems: widespread voltage control problems and protective system problems can occur, some grid systems may experience complete collapse or blackouts. Transformers may experience damage.

Spacecraft operations: may experience extensive surface charging, problems with orientation, uplink/downlink and tracking satellites.

Other systems: pipe line currents can reach hundreds of amps, HF radio propagation may be impossible in many areas, satellite navigation may be degraded, LF radio navigation can be out, and aurora can be seen as low as 40 geom. lat.

Kp=9

4 per cycle  
(4 days per  
cycle)

Severe

Power systems: possible widespread voltage control problems.

Spacecraft operations: may experience surface charging and tracking problems, corrections may be needed for orientation problems.

Other systems: induced pipe line currents affect preventive measures, HF radio propagation sporadic, satellite navigation degraded, LF radio navigation disrupted, and aurora can be seen as low as 45 geom. lat .

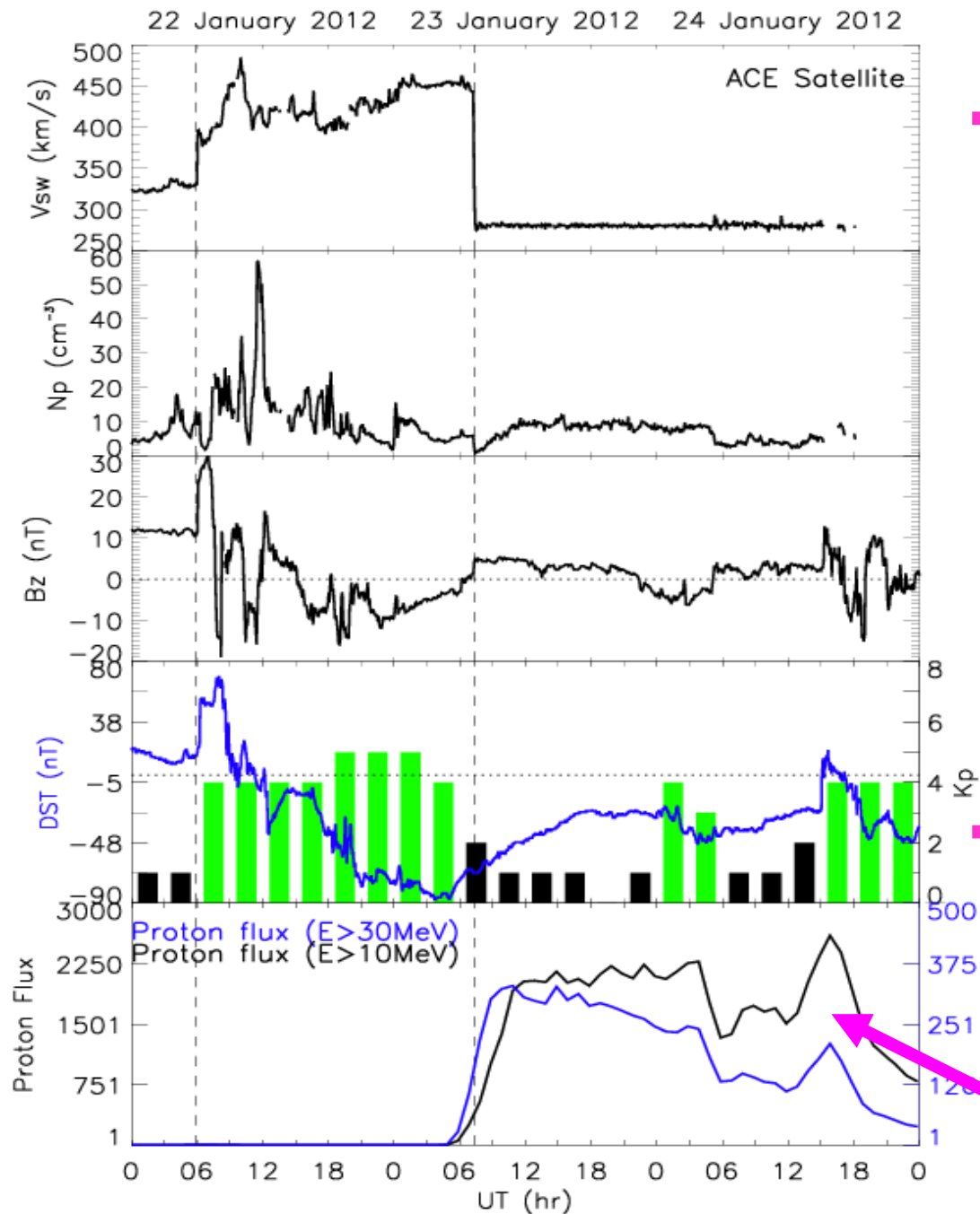
Kp=8

100 per  
cycle  
(60 days  
per cycle)



	Geomagnetic Storms effects	Kp value	# of storm events (# of storm days)
Strong	<p>Power systems: voltage corrections may be required, false alarms triggered on some protection devices.</p> <p>Spacecraft operations: surface charging may occur on satellite components, drag may increase on LEO satellites, and corrections may be needed for orientation problems.</p> <p>Other systems: intermittent satellite navigation and low-frequency radio navigation problems may occur, HF radio may be intermittent, and aurora has been seen as low as 50 geom. Lat.</p>	Kp=7	200 per cycle (130 days per cycle)
Moderate	<p>Power systems: high-latitude power systems may experience voltage alarms, long duration storms may cause transformer damage.</p> <p>Spacecraft operations: corrective actions to orientation may be required by ground control; possible changes in drag affect orbit predictions.</p> <p>Other systems: HF radio propagation can fade at higher latitudes, and aurora has been seen as low as 55 geomagnetic lat.</p>	Kp=6	600 per cycle (360 days per cycle)
Minor	<p>Power systems: weak power grid fluctuations can occur.</p> <p>Spacecraft operations: minor impact on satellite operations possible.</p> <p>Other systems: migratory animals are affected at this and higher levels; aurora is commonly visible at high latitudes.</p>	Kp=5	1700 per cycle (900 days per cycle)

# Can Kp index tell us all possible SW effects?



Not that much magnetic storm activity happened

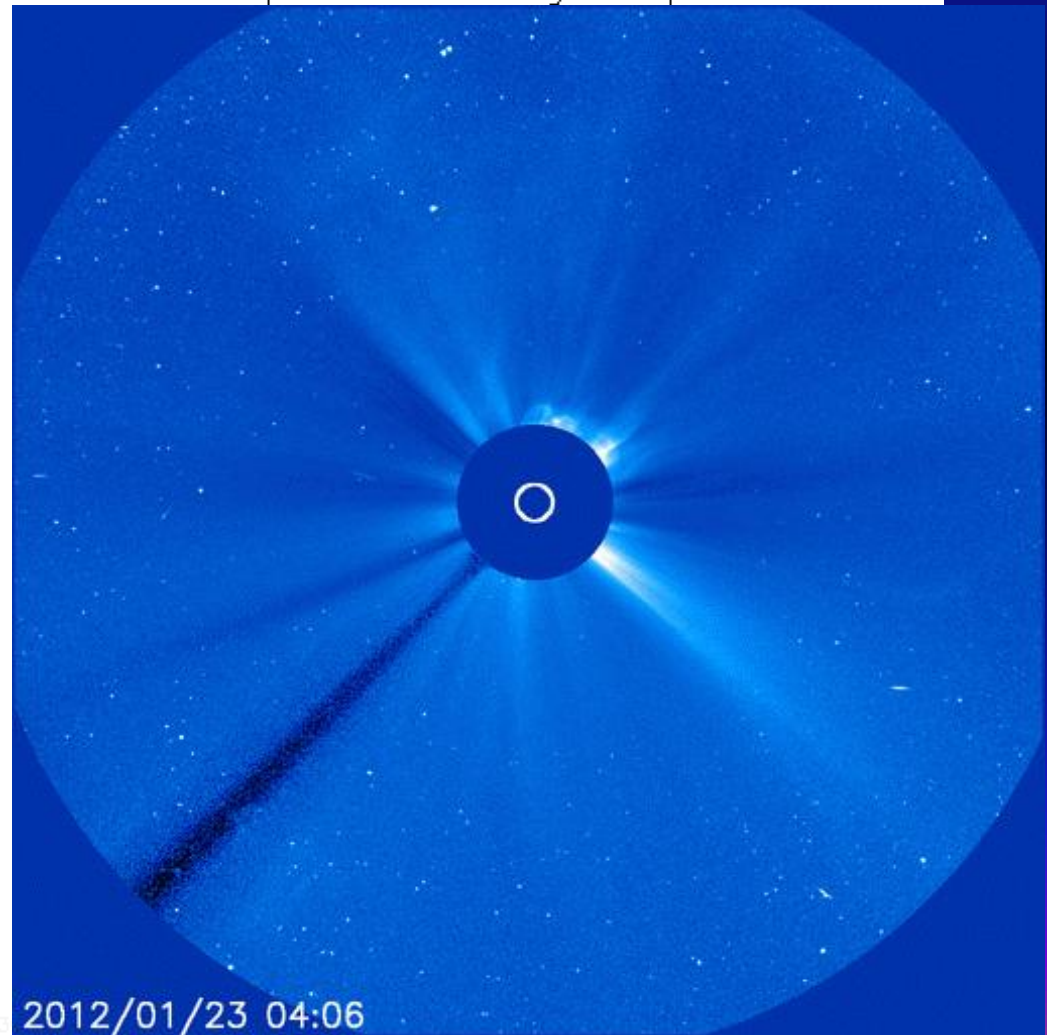
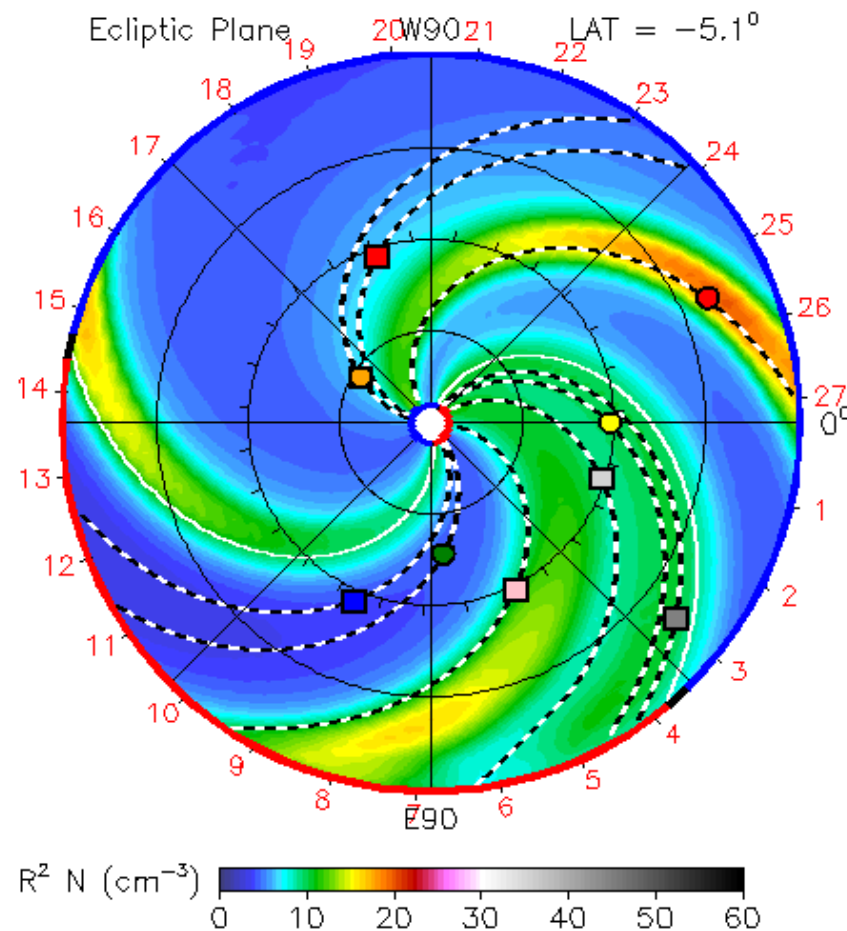
However, there were significant (S3 class) Solar Radiation Storms

# CME which was not aiming to us

2012-01-22T00:00

2012-01-22T00 +0.00 day

● Earth ● Mars ● Mercury ● Venus ■ Juno ■ Kepler ■ Messenger ■ Spitzer  
■ Stereo\_A ■ Stereo\_B





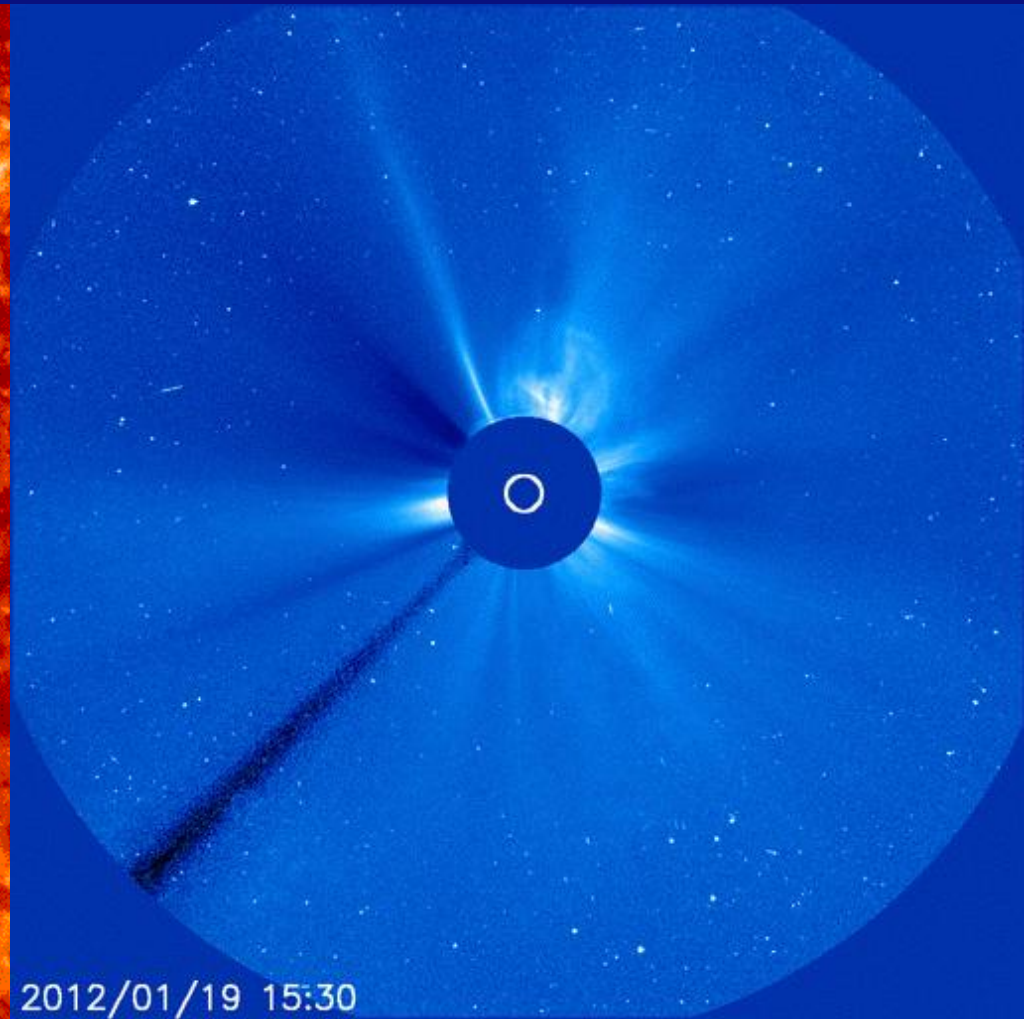
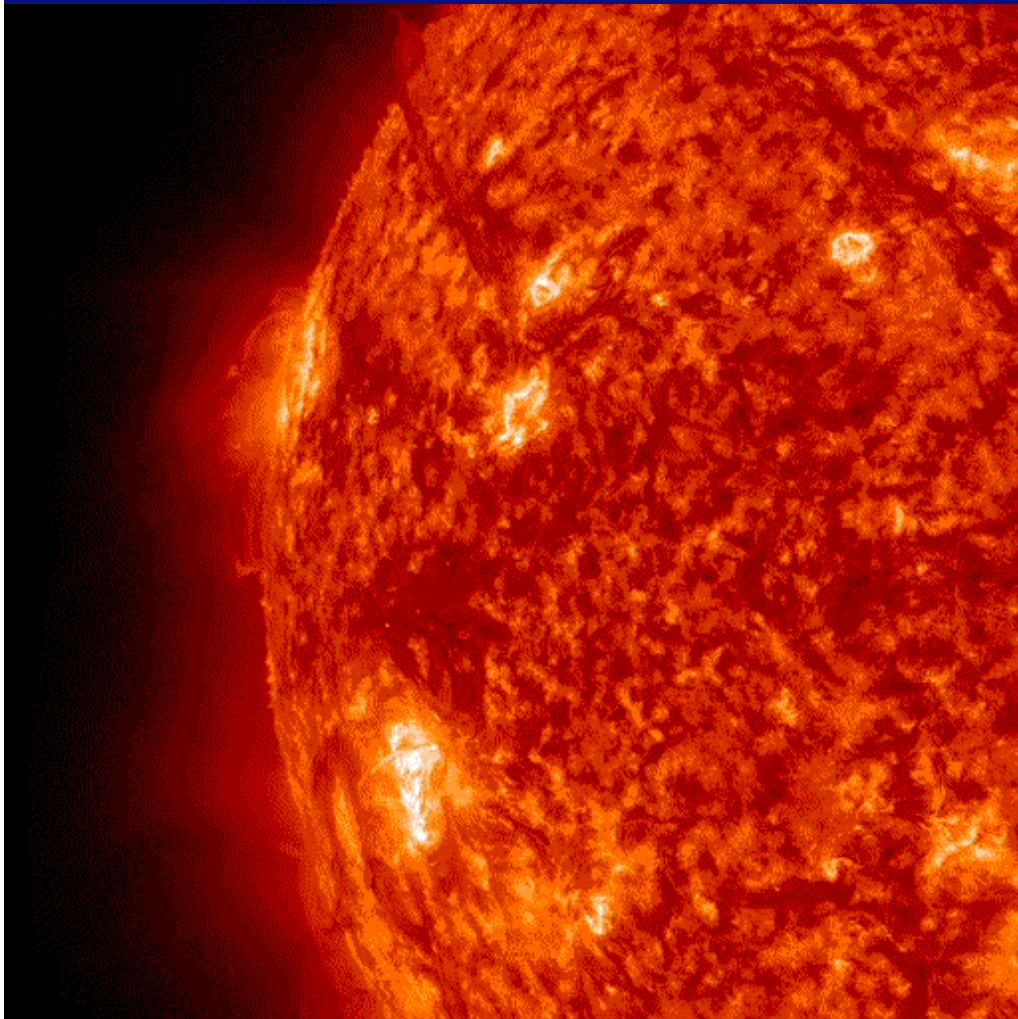
Solar Radiation Storms effects		Flux level of ≥ 10 MeV	# of events
Extreme Scale: S5	<p>Biological: unavoidable high radiation hazard to astronauts; passengers and crew in high-flying aircraft at high latitudes may be exposed to radiation risk.</p> <p>Satellite operations: satellites may be rendered useless, may cause serious noise in data imaging, permanent damage to solar panels possible.</p> <p>Other systems: complete blackout of HF communications, and position errors make navigation extremely difficult.</p>	10E5	< 1 per cycle
Severe Scale: S4	<p>Biological: unavoidable radiation hazard to astronauts; passengers and crew in high-flying aircraft at high latitudes may be exposed to radiation risk.</p> <p>Satellite operations: may experience memory device problems and noise on imaging systems; and solar panel efficiency can be degraded.</p> <p>Other systems: blackout of HF radio communications and increased navigation errors for several days</p>	10E4	3 per cycle

Solar Radiation Storms effects		Flux level of ≥ 10 MeV	# of events
Strong Scale: S3	<p>Biological: radiation hazard avoidance recommended for astronauts; passengers and crew flying at high latitudes may be exposed to radiation risk.</p> <p>Satellite operations: noise in imaging systems, and slight reduction of efficiency in solar panel.</p> <p>Other systems: degraded HF radio propagation (in the polar regions) and navigation errors.</p>	10E3	10 per cycle
Moderate Scale: S2	<p>Biological: passengers and crew flying at high latitudes may be exposed to elevated radiation risk.</p> <p>Satellite operations: infrequent single-event upsets.</p> <p>Other systems: small effects on HF propagation and navigation errors at polar cap locations.</p>	10E2	25 per cycle
Minor Scale: S1	<p>Biological: none.</p> <p>Satellite operations: none.</p> <p>Other systems: minor impacts on HF radio in the polar regions.</p>	10E1	50 per cycle

# More CM Es which were not aiming to us

Going to the other side  
just last week (16 Apr. 12)

Just missed us on 19 Jan. 12

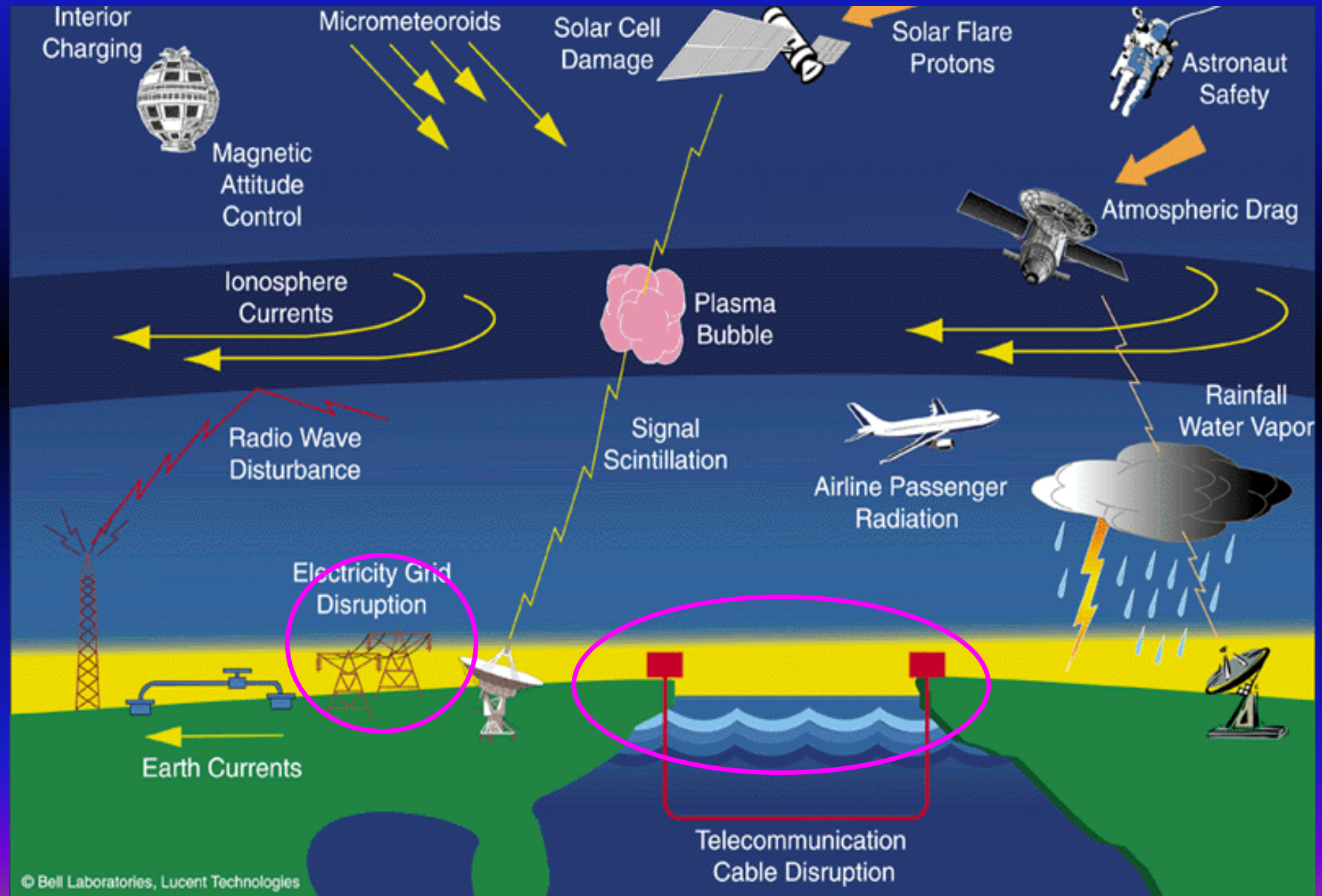


2012/01/19 15:30



# Why do we care about Space Weather?

## Space Weather Impacts: Society and Technology



# Commercial Communication satellites

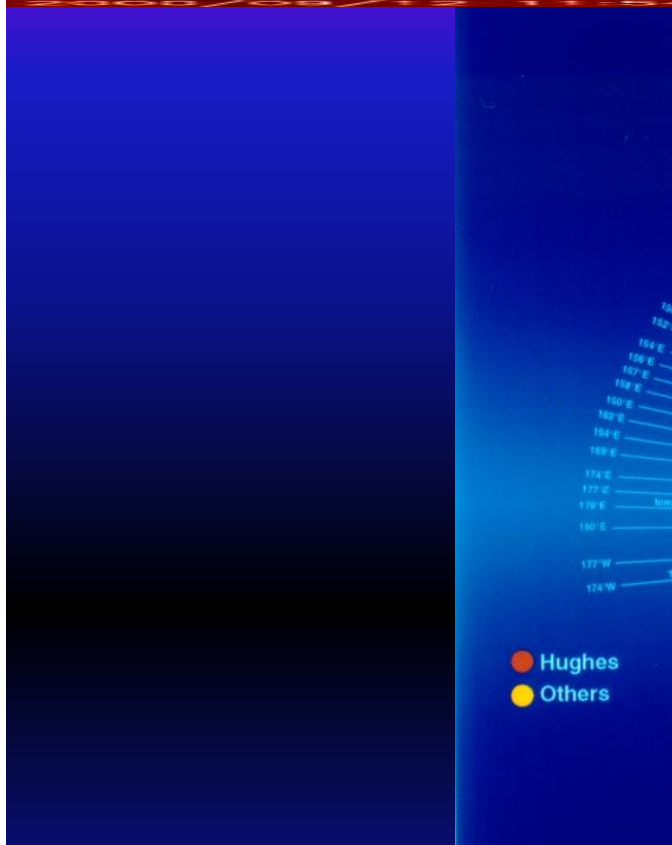


Image Credit: L. J. Lanzerotti, Bell Laboratories, Leont Technologies, Inc.

Many African Countries do not have fiber optics and their communication system is highly dependent on wireless with satellite communication system

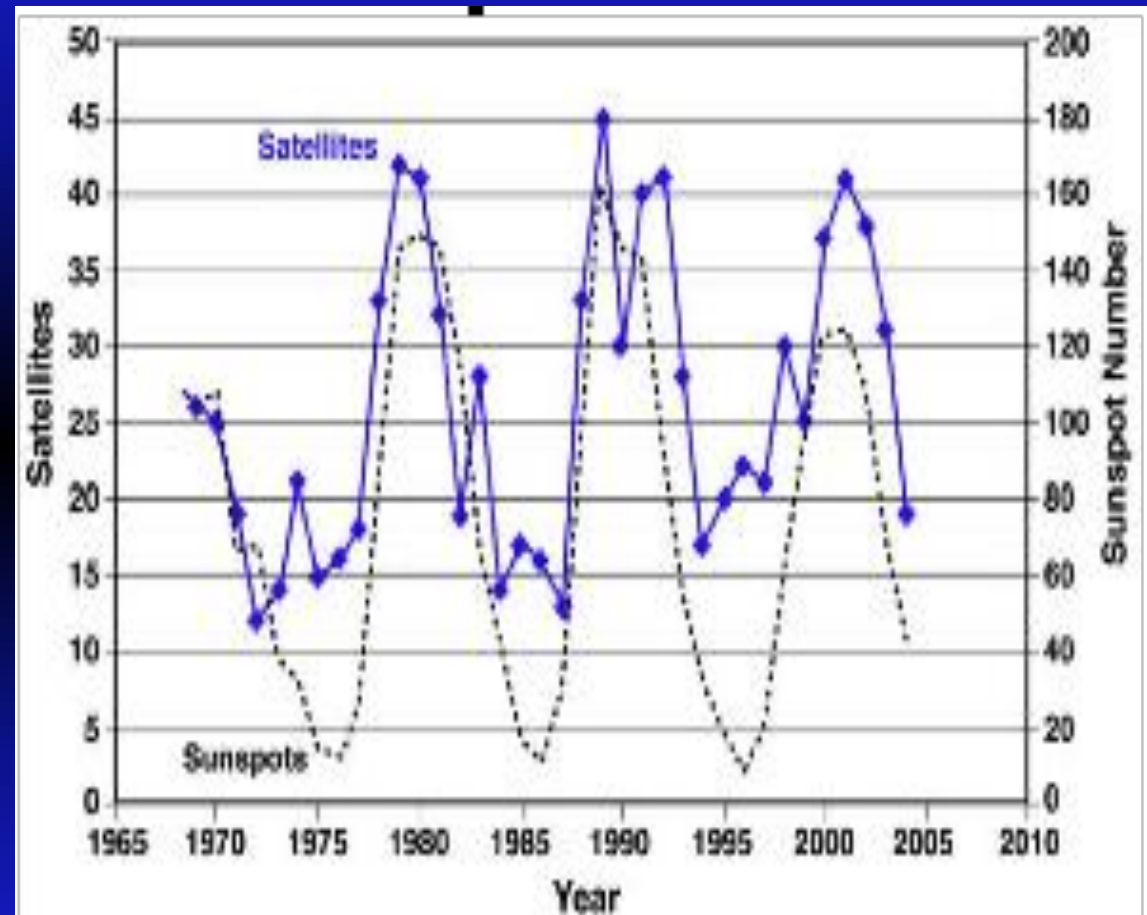


# Potential Economic impact due to Spacecraft Damage/Loss

→ Solar Panel  
degradation from  
SEP event

→ Lifetime of industry  
transponder usage

Total estimated  
cost \$ 30B



*Baker et al.*

# Space Weather impact on Power Grids

- What are the issues?
- What are the risks?
- What are the economic impacts?
- What are the societal impacts?



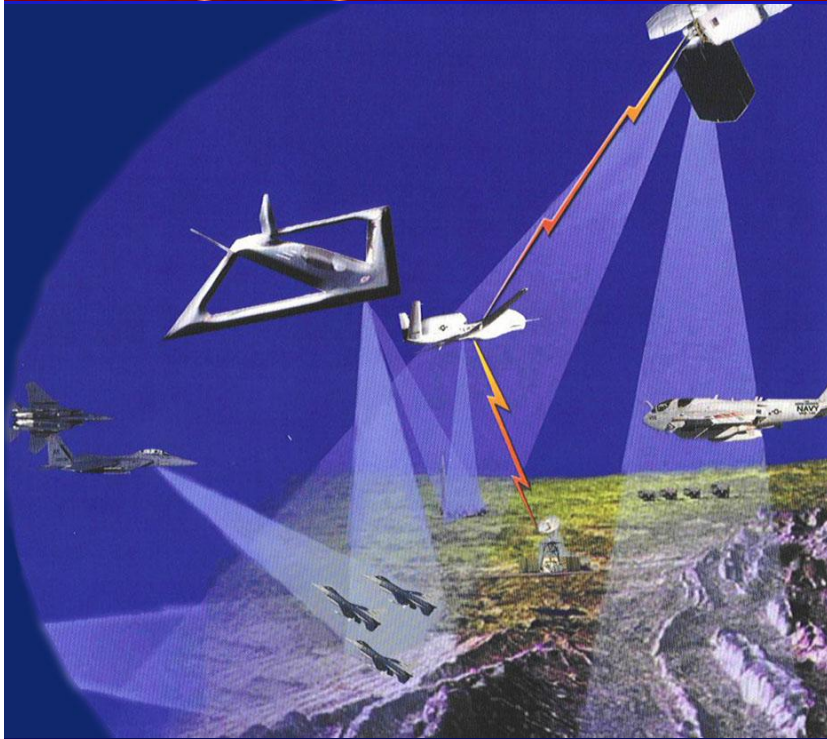
# Other large economic events

- San Francisco Earthquake ..... 1906 .....\$ 500B
- Hurricane Katrina ..... 2005 .....\$ 120B
- Annual loss from Electric interruption .....\$ 80B
- **North American Power Grid Blackout .....\$ 30B**
- GEO satellite revenue loss .....\$ >25B
- Blackout of East Coast ..... 1955 .....\$ 10B
- Mt Lassen Volcanic Eruption .. 1915 .....\$ 5B
- Quebec Blackout ..... 1906 .....\$ 2B

*Baker et al.*



# Space weather and our communication and navigation systems



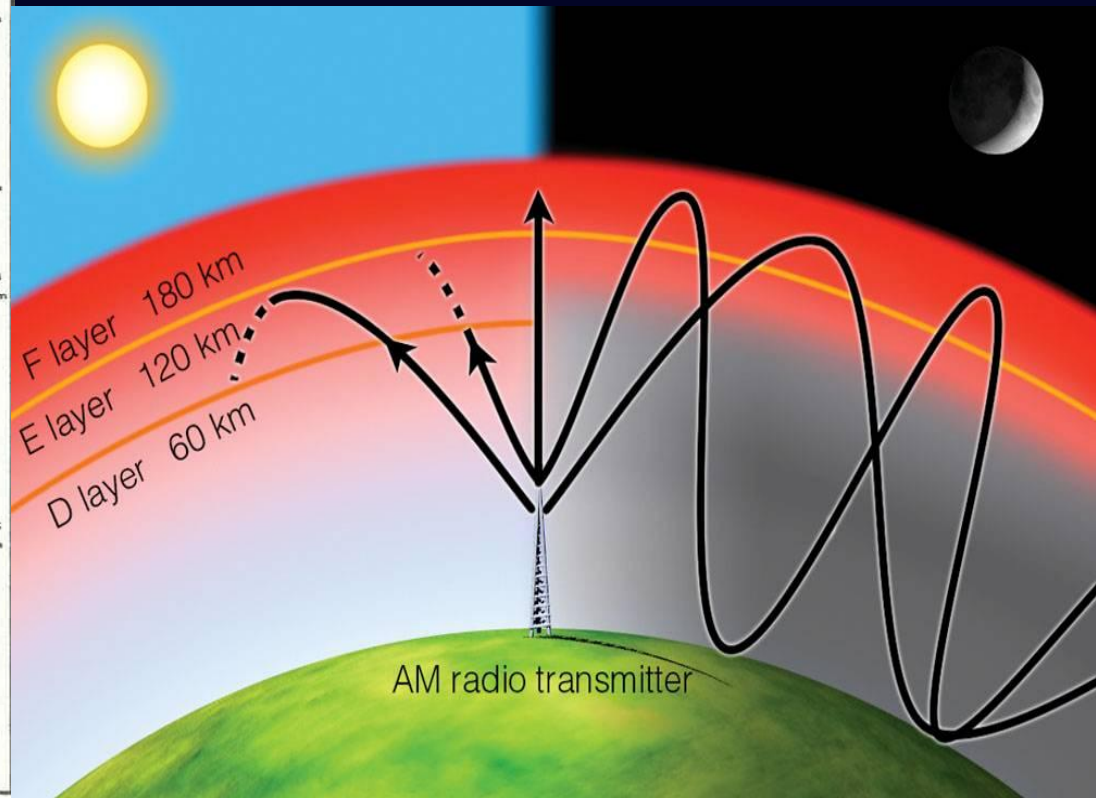
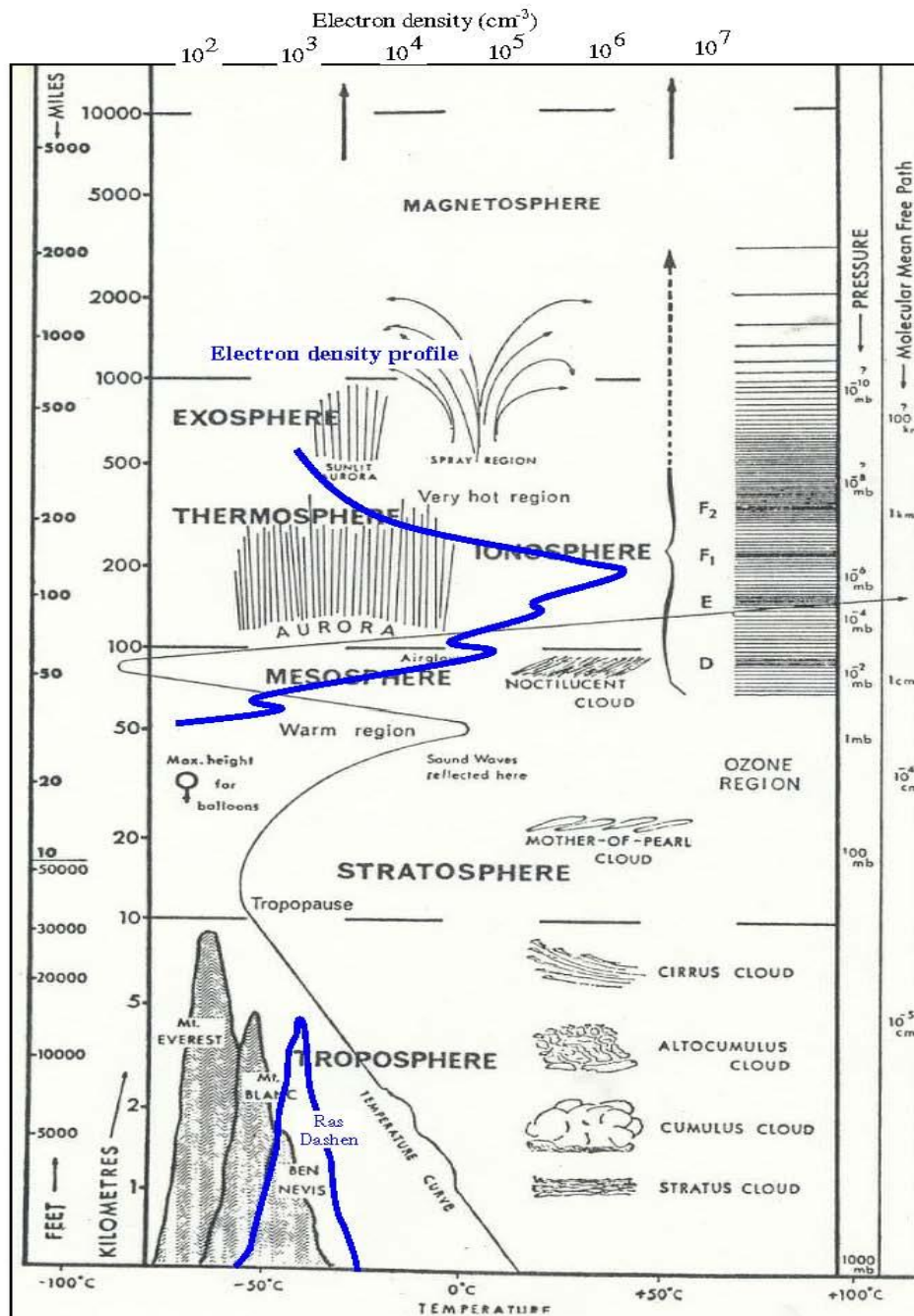
→ The main factor that affect the navigation and communication system is the variability of structure and dynamics of plasma in our atmosphere.

## How and where?

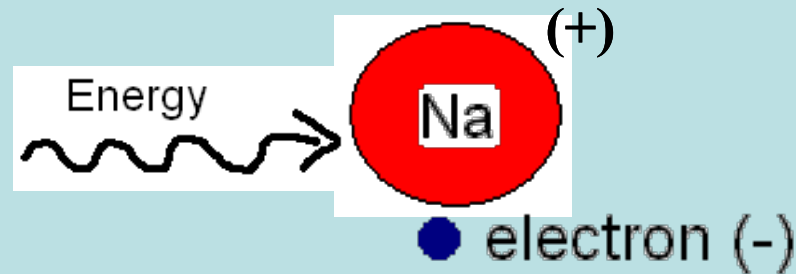
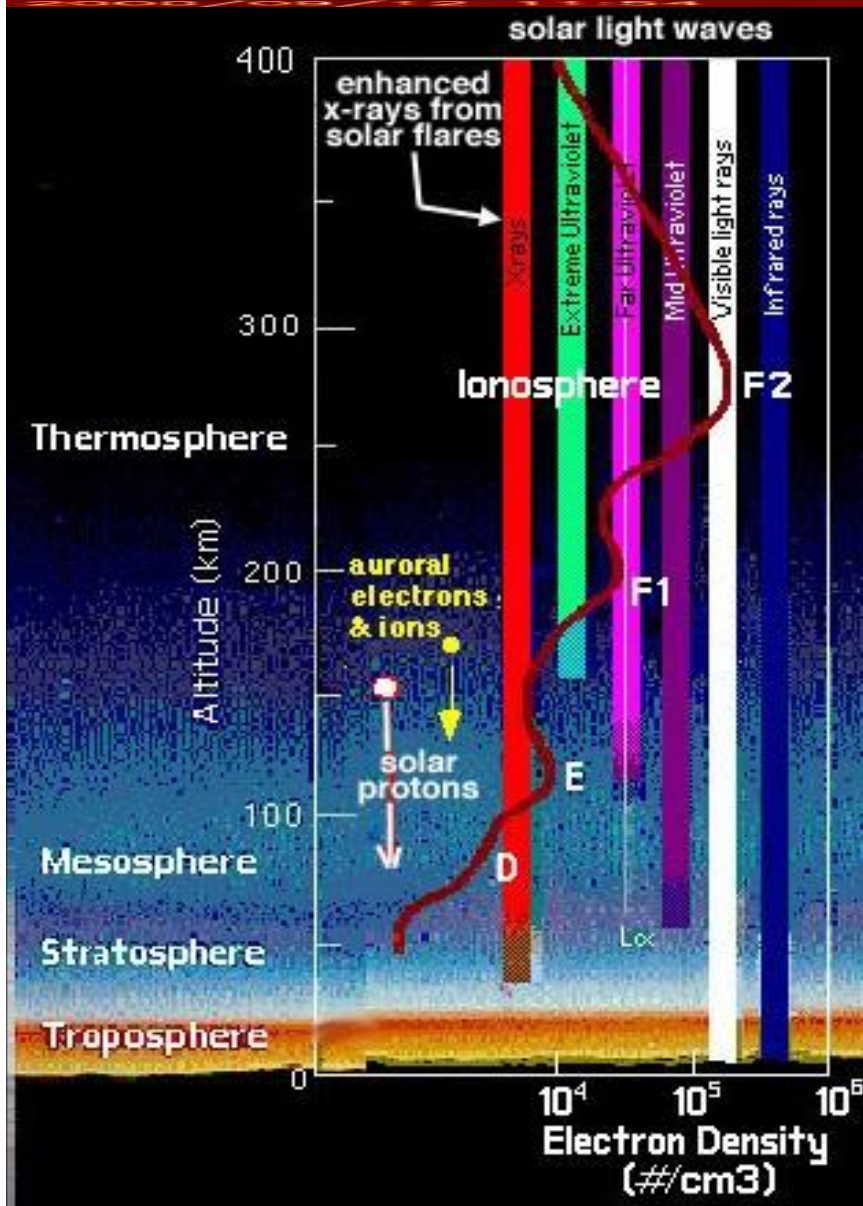


# Atmospheric region of interest

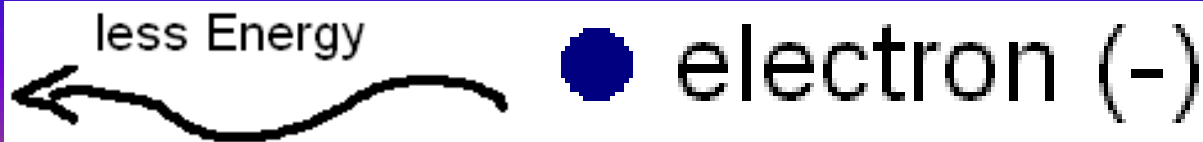
Ionized layer of our atmosphere, known as **Ionosphere**, is both important and trouble making region for our technological systems.



# How the ionosphere is formed?

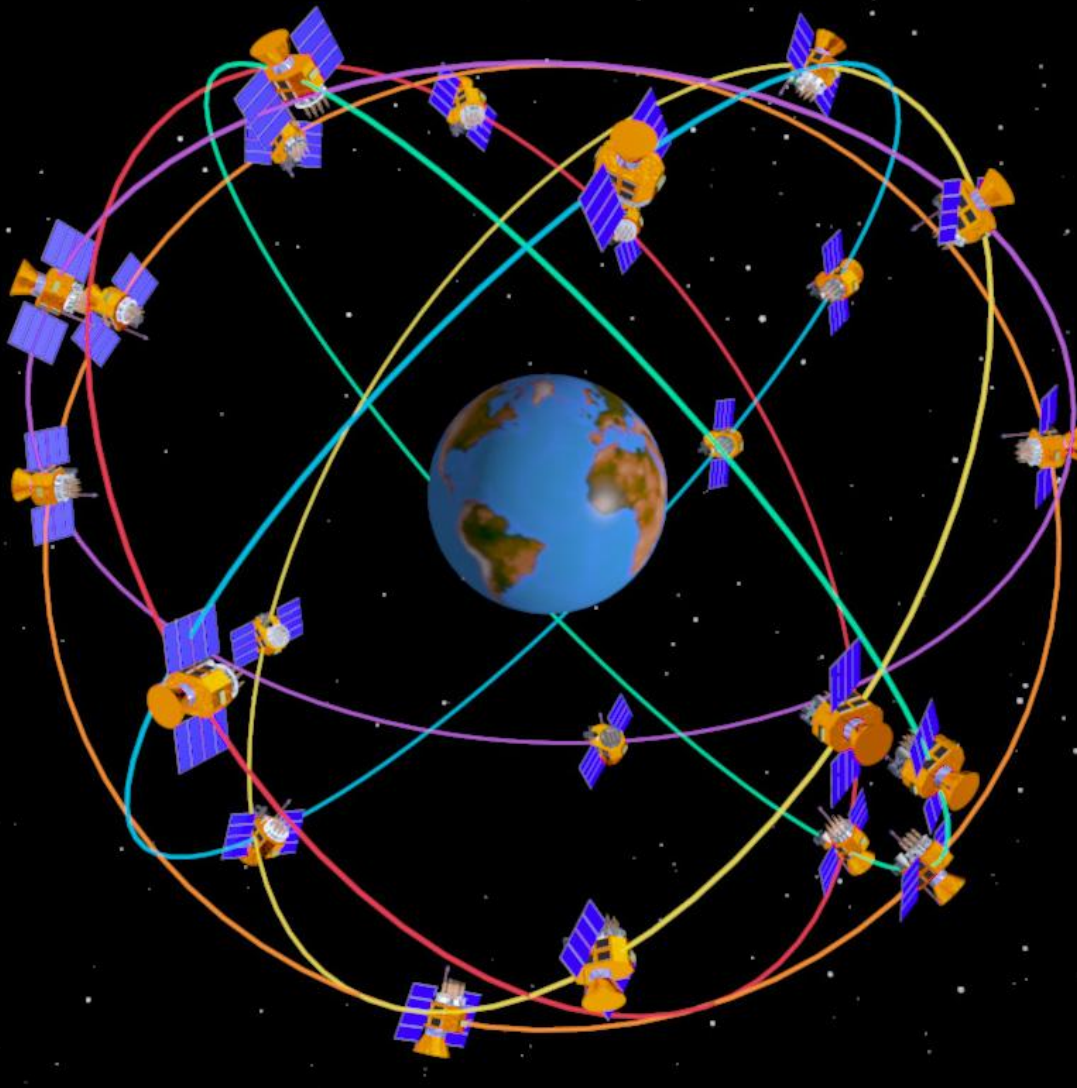


→ UV light with high energy from the sun hits atoms in Earth's upper atmosphere and knocks an electron off the atom, leaving a free electron and an ion. This type of ionized gas is called plasma, which can conduct an electric charge, reflect and deflect EM wave, and is affected by magnetic field.





# GPS Navigation problems



GPS satellite information

- ➔ Altitude: ~20,200 km
- ➔ Orbit inclination:  $55^\circ$
- ➔ Orbiting period: 12 hrs

Transmit dual frequency

$$f_1 = 1.575 \text{ GHz}$$

$$f_2 = 1.228 \text{ GHz}$$

GLONASS of Russia

GALILEO of Europe

QZSS of Japan

COMPASS of China

# Does GPSs suffer most?

Space Segment

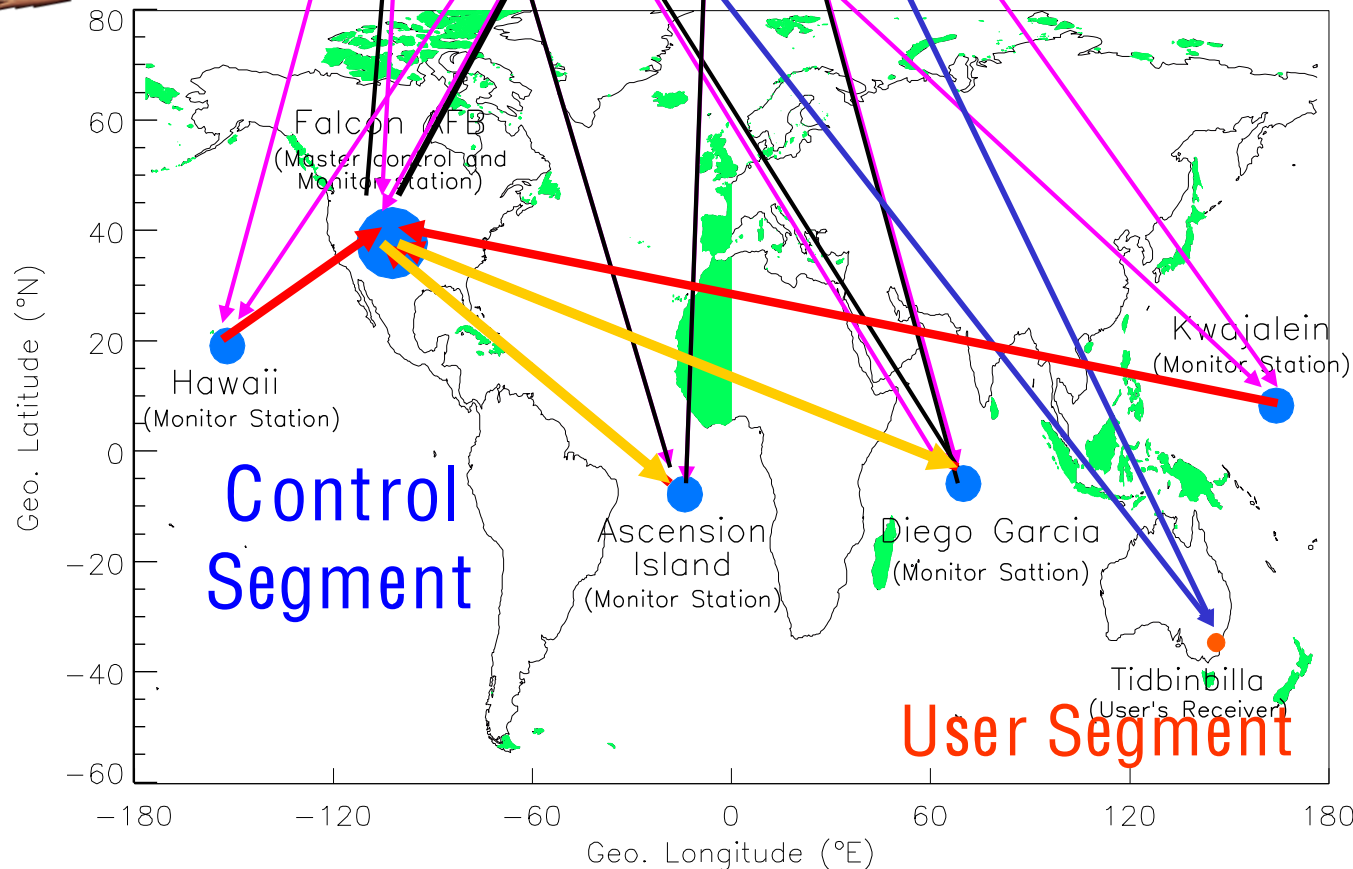
Ranging information from satellite broadcast

LEO

Send the information to master control station that precisely calculate the satellite orbit

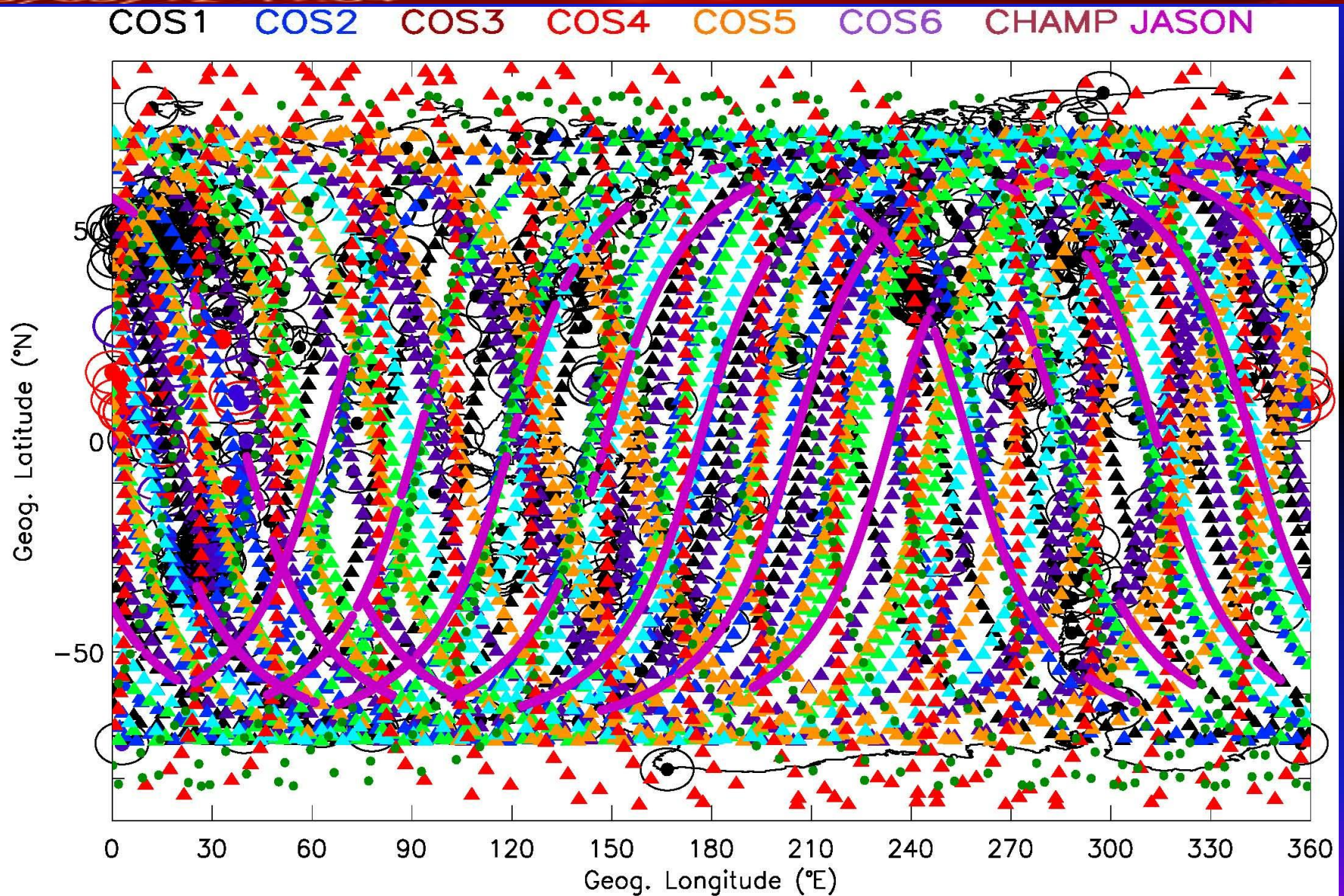
Converted to updated navigation message for each satellite and transmitted back to the satellites

The satellites then broadcast coded signals to the user segment



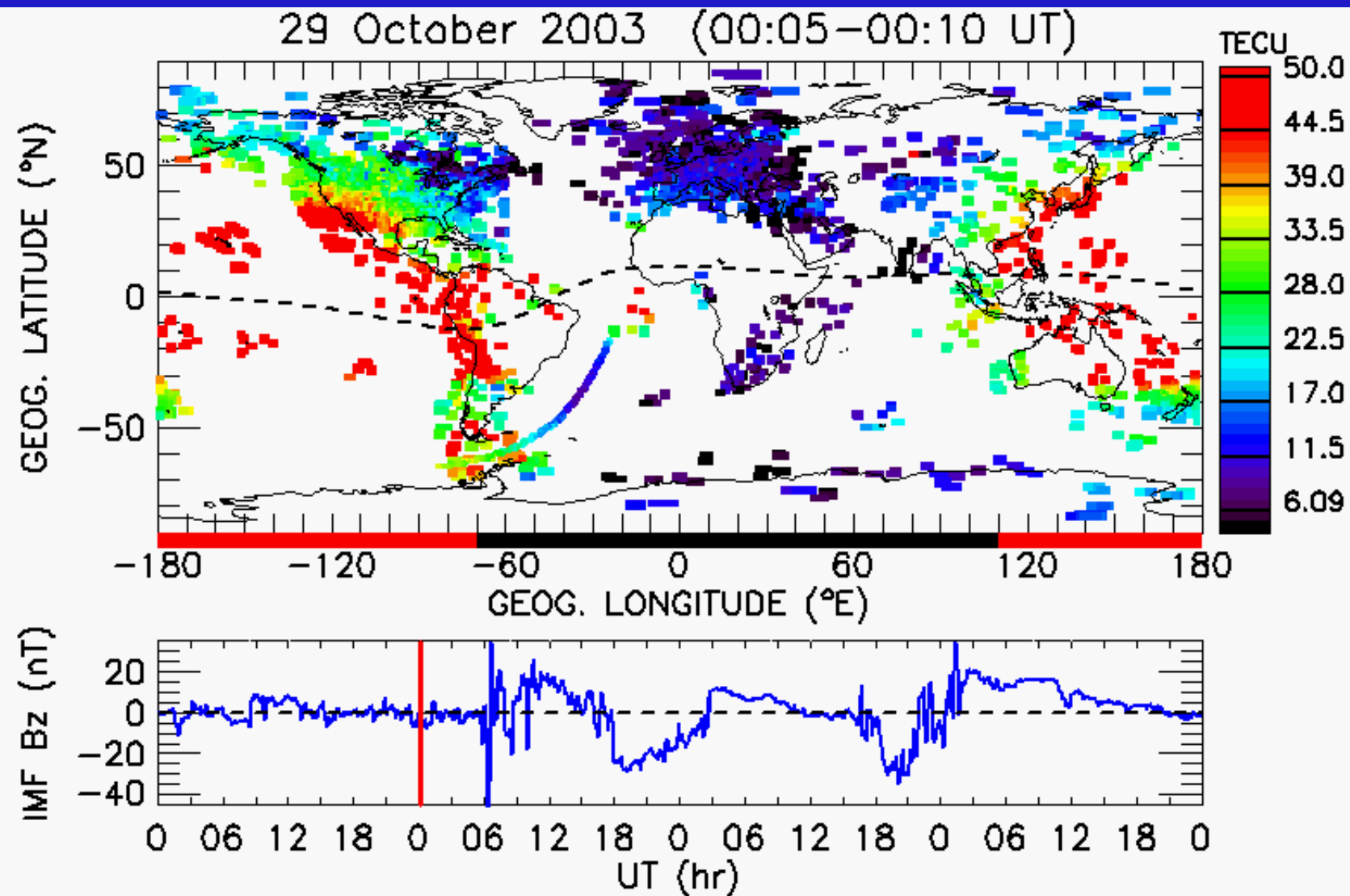


# GPS can be used monitor the space weather impact on navigation errors

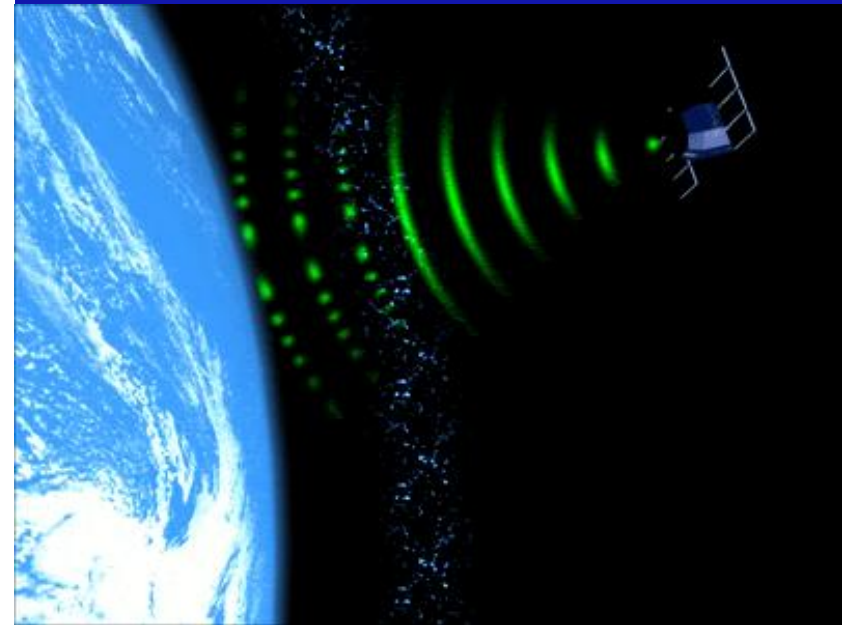
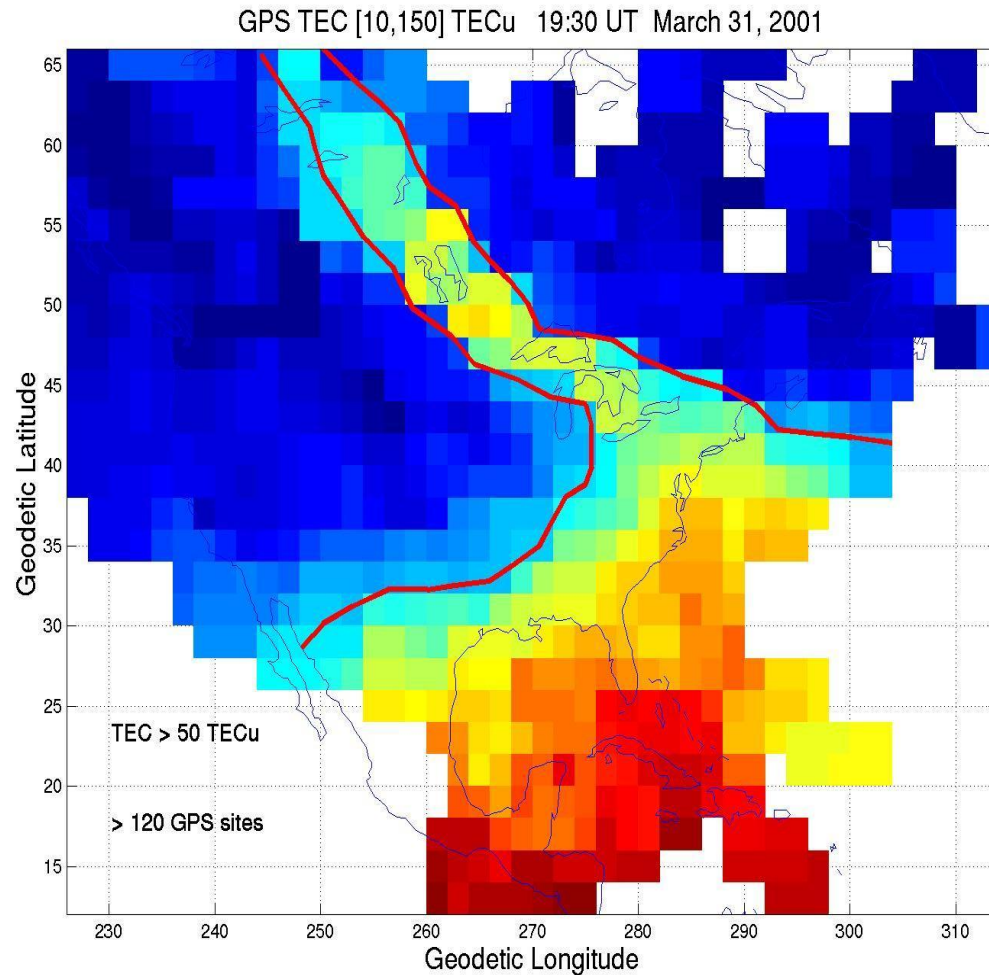




# Ionosphere is boiling like ---?



# Space weather cause strong density gradient and irregularities

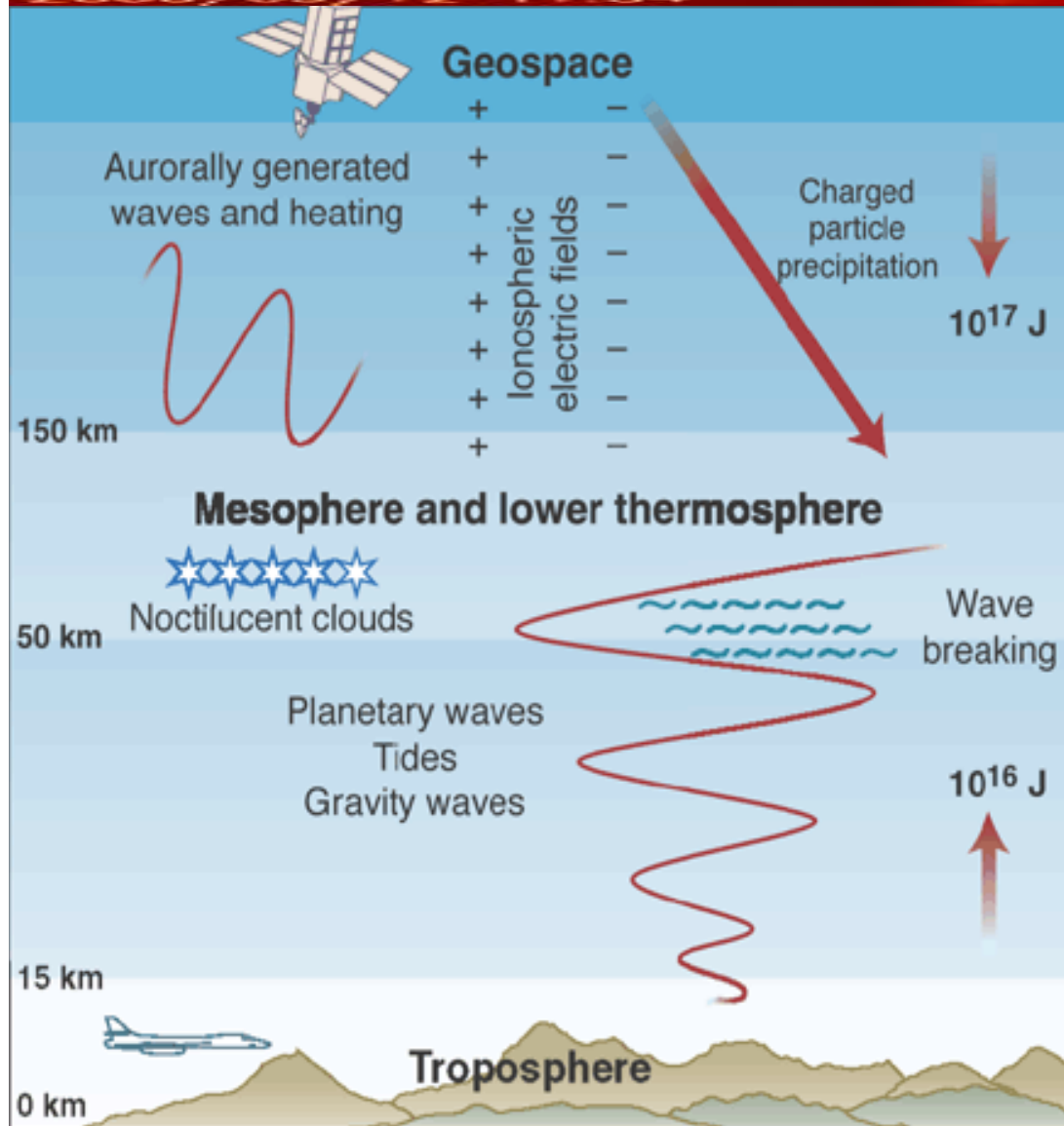


*After C. Mitchell*

*Foster et al., GRL 2002*



# Space weather impact from the lower atmosphere

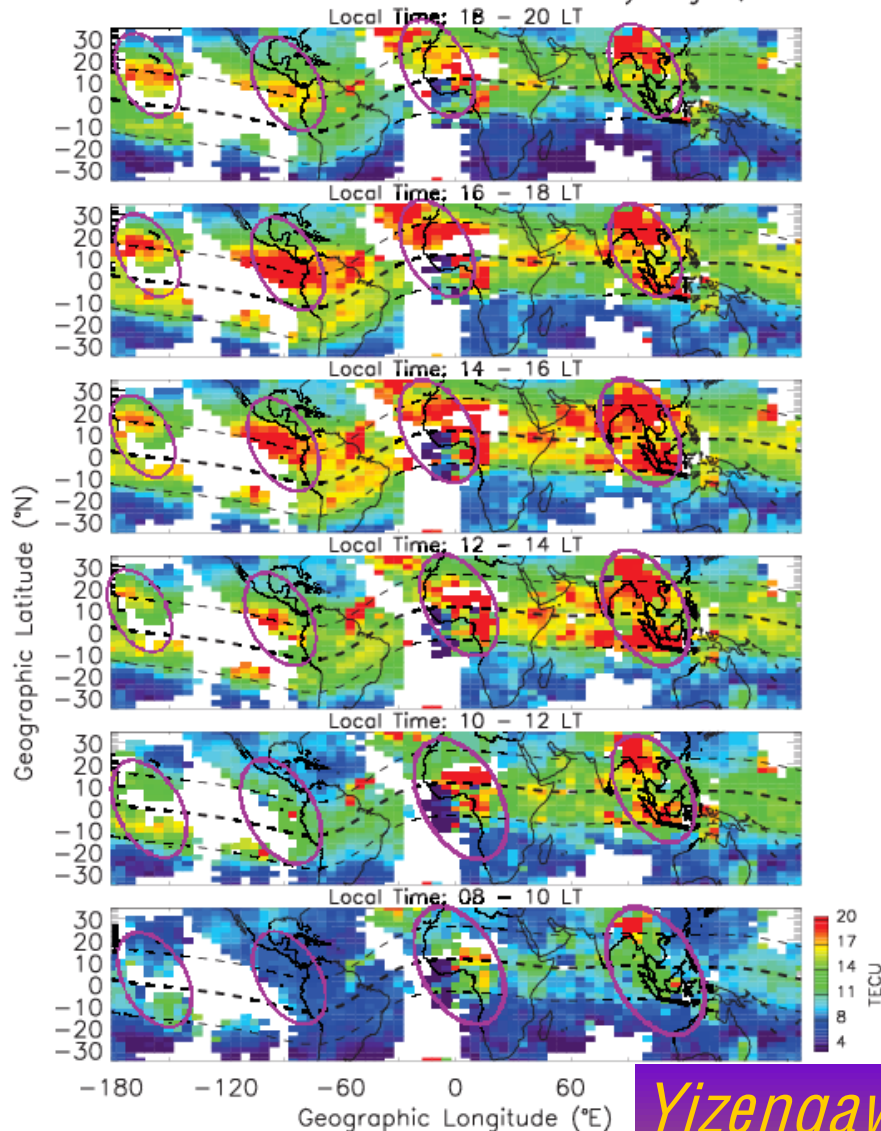


- Gravity waves can propagate upward and produce density irregularities
- Tsunami and earthquake can generate strong gravity wave
- Thermospheric tides also modulate the equatorial electrodynamics and thus the longitudinal density distribution

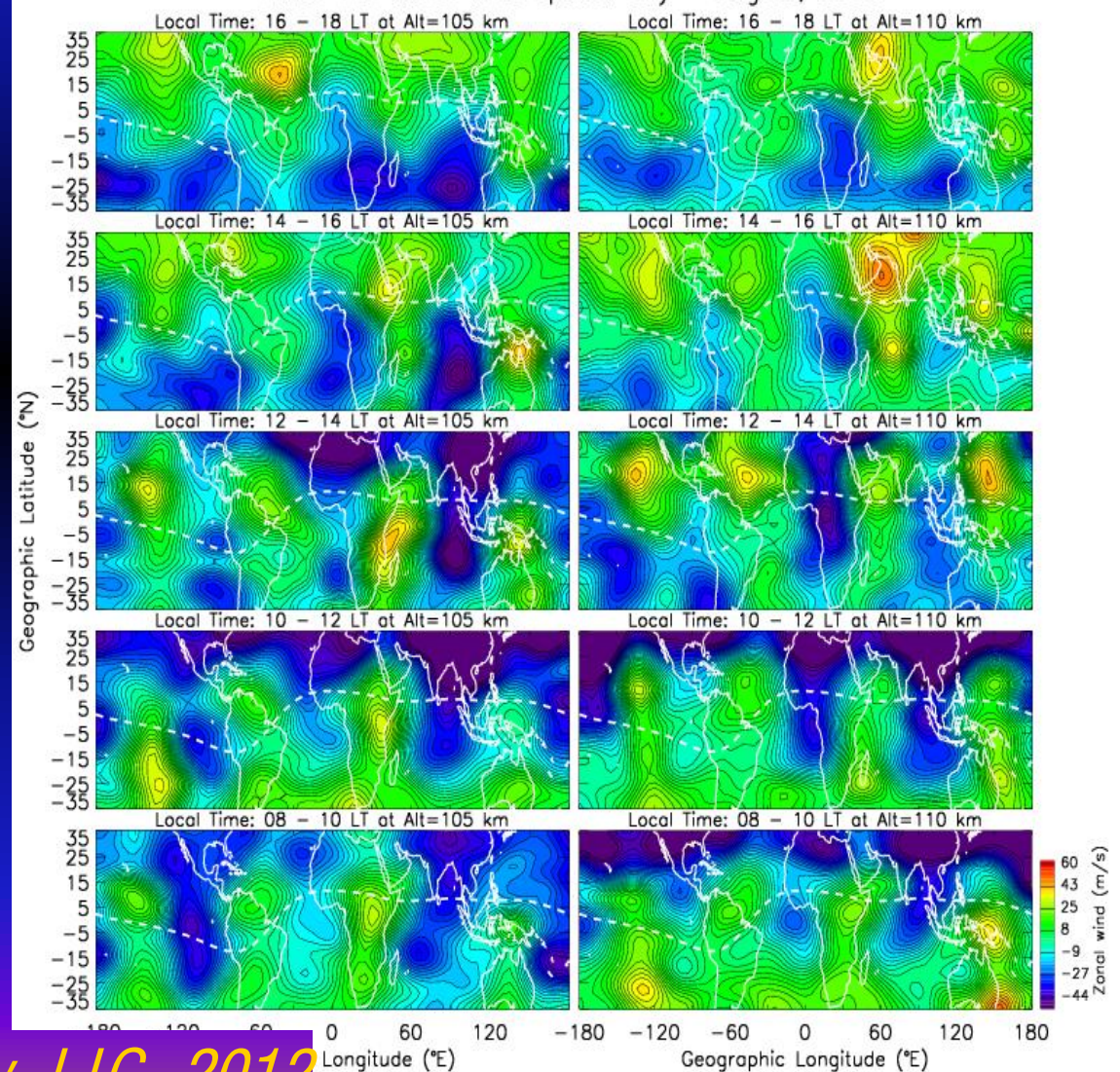


# Thermospheric tides impact on global density distribution

Global GPS TEC distribution between for July–August, 2008

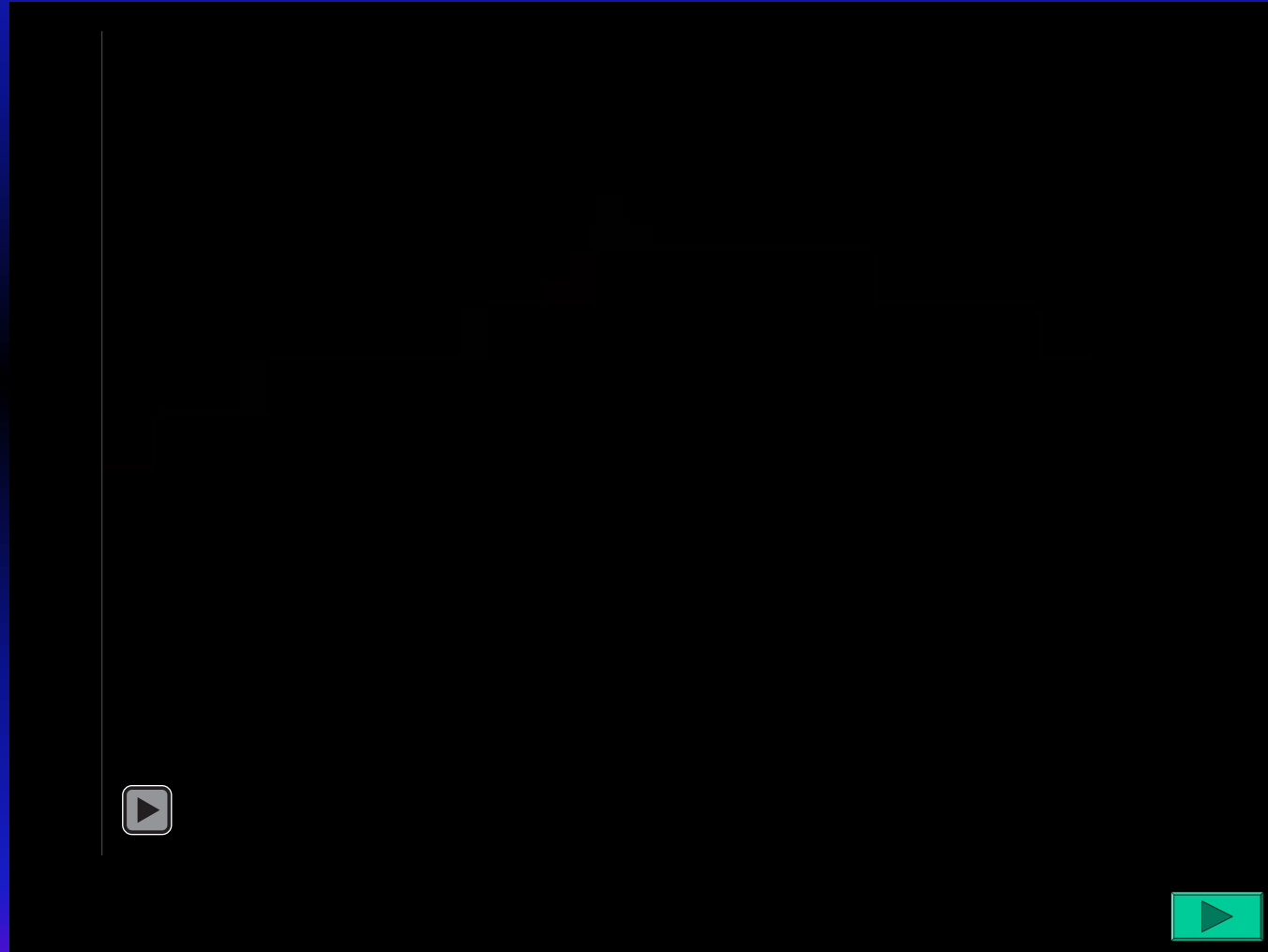


TIMED TIDI Zonal wind speed: July – August, 2008



Yizengaw, IJG, 2012

# Tsunami and Earthquake impact on ionospheric density

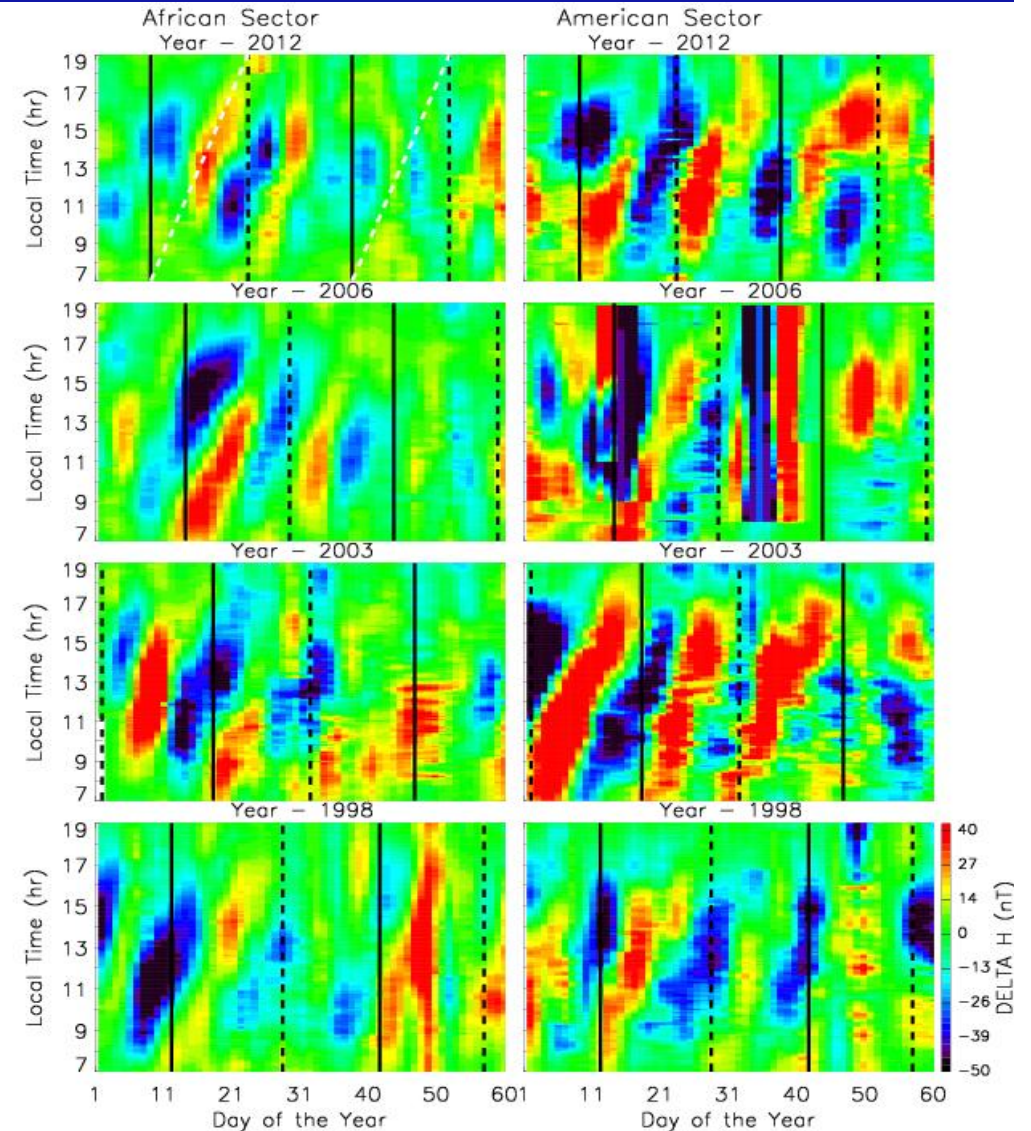
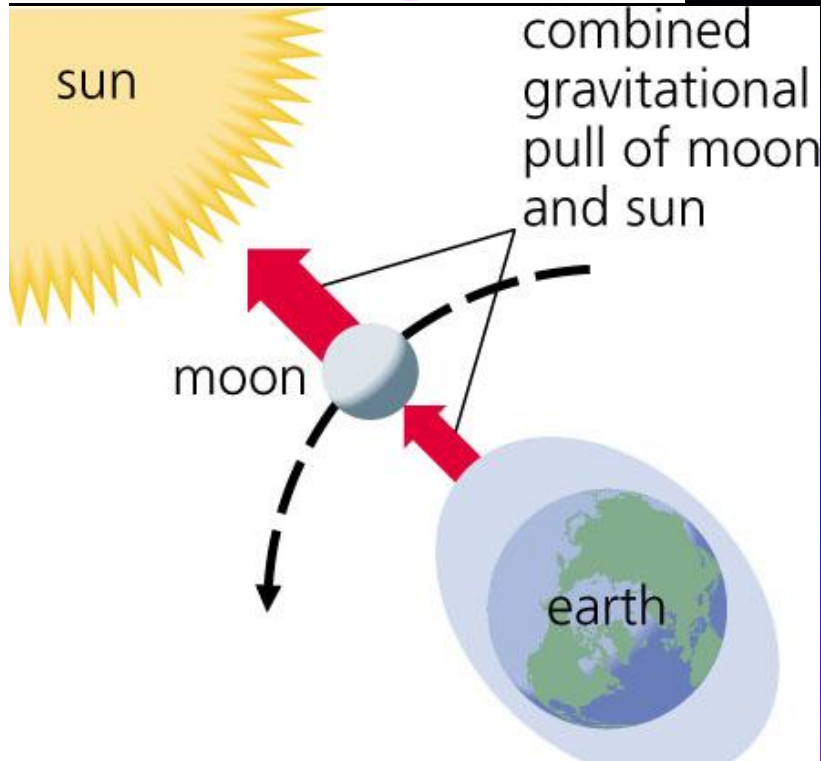
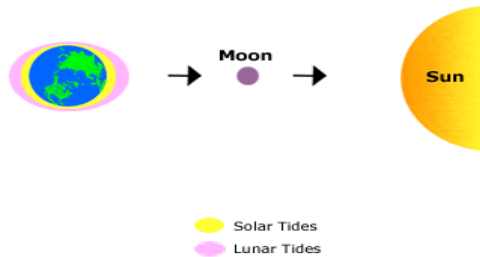


*Courtesy to David Galvan*



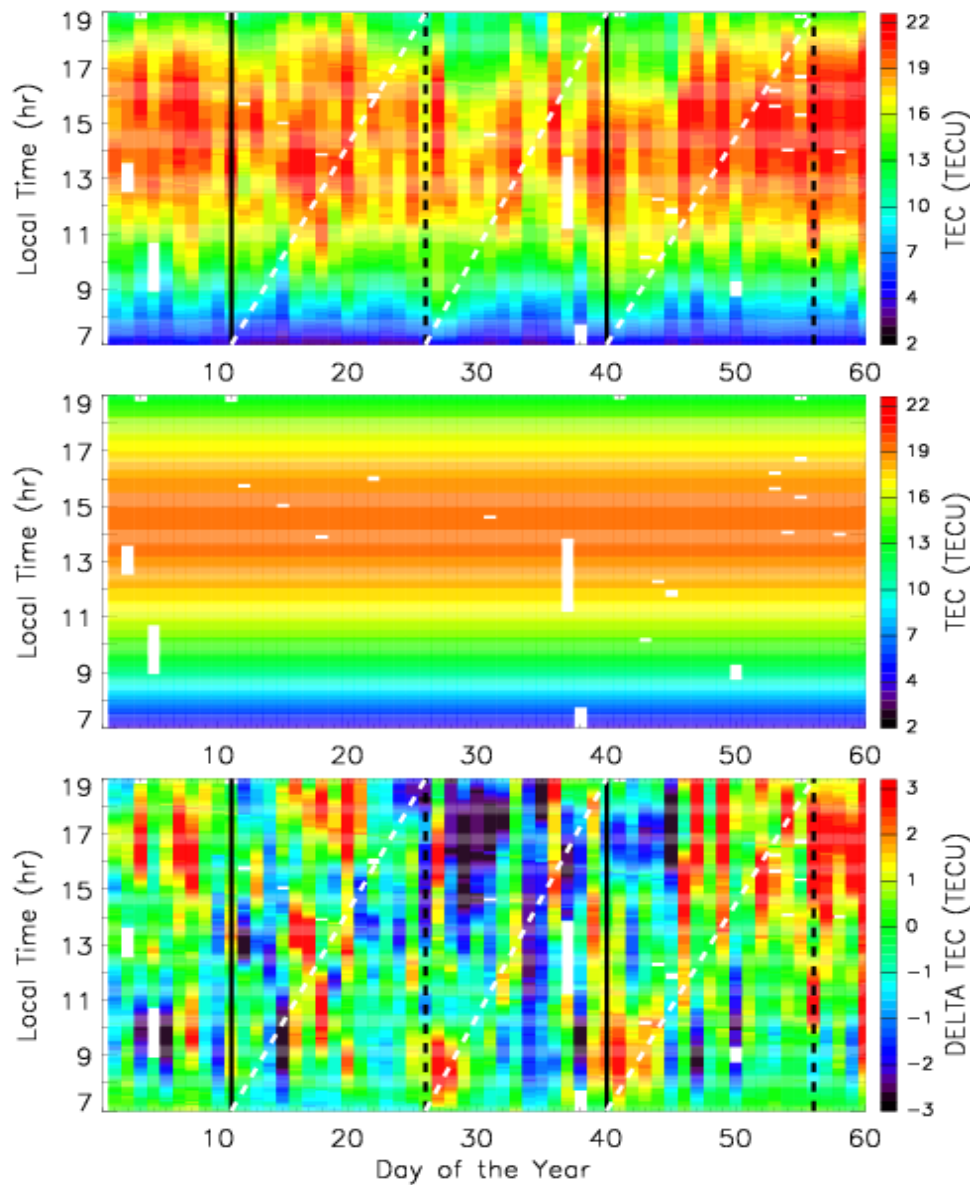
# Lunar tide impact on vertical drift ionospheric density

Spring Tides

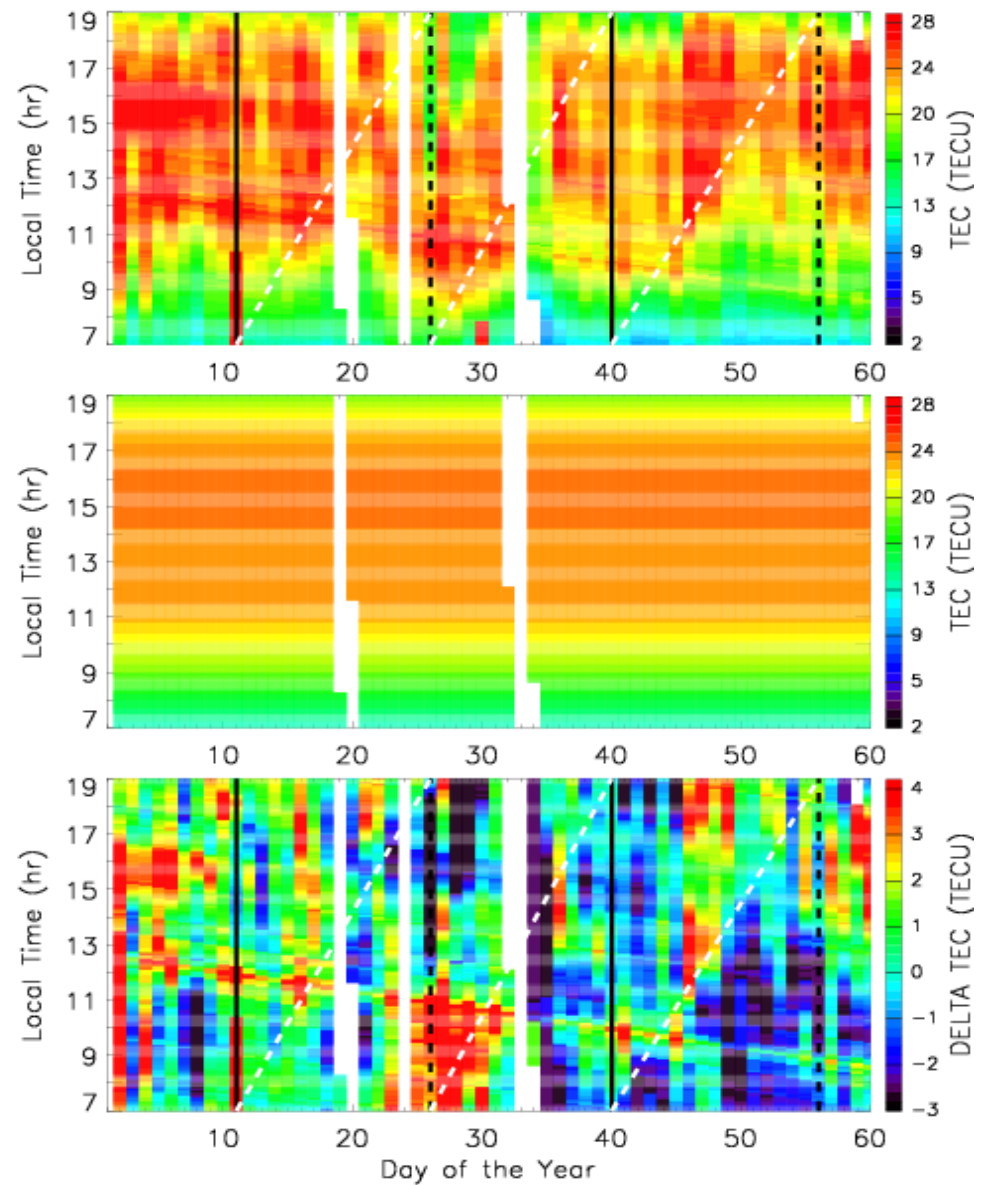


# Lunar tide impact on GPS TEC

## African Sector (ADIS)



## American Sector (ANCN)



# How can such impact be reduced?

- ➔ Provide advice to the power station crew to reduce the output current in advance by looking at the Sun
- ➔ However, experienced space weather expert is required to provide such advice.
- ➔ For Africa, encouraging the next generation of African scientists to become inspired by space science is important.
- ➔ One way to do this is creating opportunities for Africans to make interaction with international scientists.
- ➔ The instrument deployment in Africa through IHY & ISWI program significantly spark interest in space science education and research throughout Africa.
- ➔ A follow-up training workshop, like GNSS workshop, is highly essential to produce more young well trained fellows.





Last month, the British government declare that space weather is one of the greatest threats to the country

2009/09/12 11:54

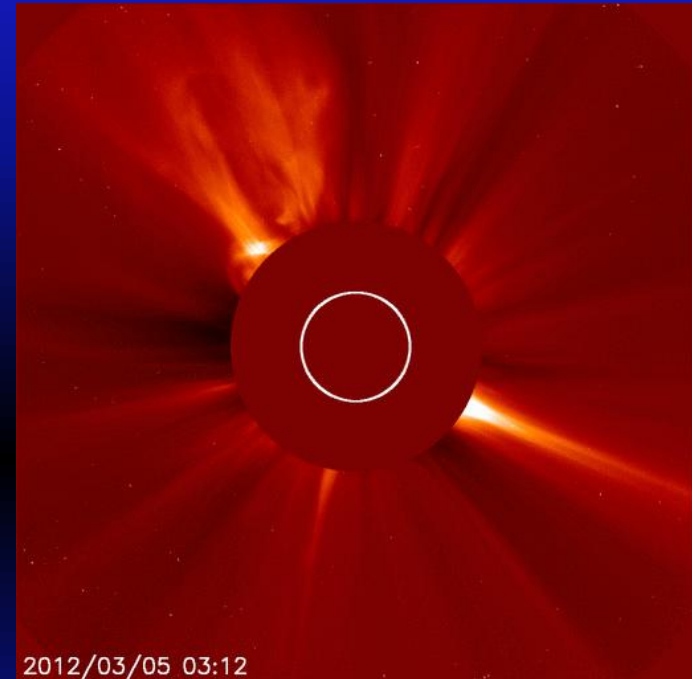
*“Severe space weather can cause disruption to a range of technologies and infrastructure, including communications systems, electronic circuits and power grids.”*

Cabinet Office's National Risk of Civil Emergencies report

# Why space weather is so important? (summary)

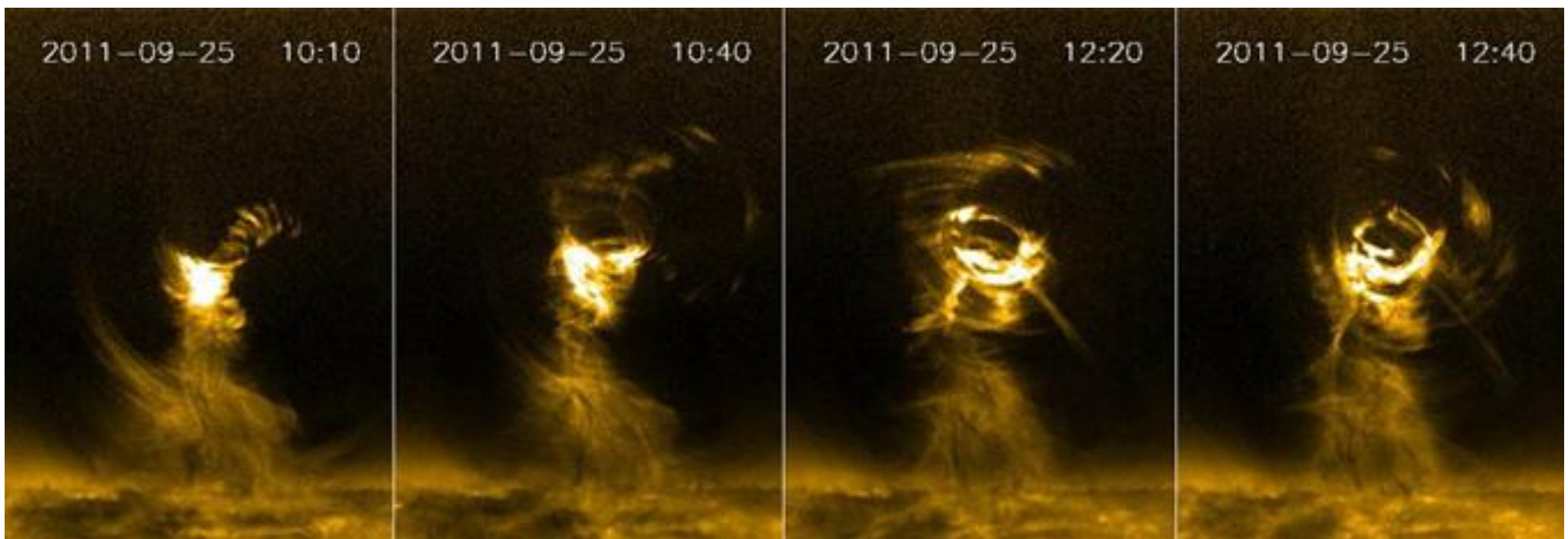
## Solar wind-Earth interaction

- Generates Electrical Currents and cause a power outage, energizes particles (radiation), moves plasma and affect our communication and navigation systems
- heat the upper atmosphere, causing it to expand, increasing drag on LEO satellites



The 2008 US National Research Council report estimated the cost if a September 1859 sized CME hit us; first it could take us 10 years to recover, and cost could be between \$1 trillion and \$2 trillion (in the first year alone) to repair the damage.

Let's just hope, with the NASA's Solar Probe Plus, we have plenty of warning before the next major event overloads the planet.



With the help of **NASA's** Solar Dynamic Observatory (SDO), it is now possible to see previously unseen small scale features of the Sun, link Solar Tornadoes

*Courtesy to NASA and Aberystwyth University*

Thank You



# What is the cost of a blackout?

Assume a model power station is hit by very large (800A) GIC Event/geomagnetic Storm

5,000 MW of generation

3,000 employees

3,000,000 population served

1,000,000 in a metropolitan area

## Consider the costs from 2 viewpoints:

- The utility
- The whole society

Demsky, NASA/NOAA, 2008

# What is societal cost when a solar storm hits?

For a country with GNP \$7,400 annual per capita

- \$29 per working day

- Assume 80% lost: \$23

Total for city of 1M people: \$23M/day, \$115M/week

- City without power: Reduced GNP \$460M
- Wide area outage: Reduced GNP \$2.3B

**Total: \$2.76B**

# What does it cost a utility when a solar storm hits?

City without power:

- Idle workers \$29M/day

- Transformers \$30M

- Sales less fuel \$500M/day?

Total: \$529M/day + 30M

*Baker et al.*