



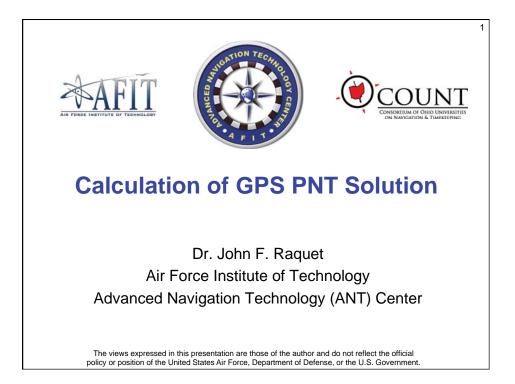
2458-1

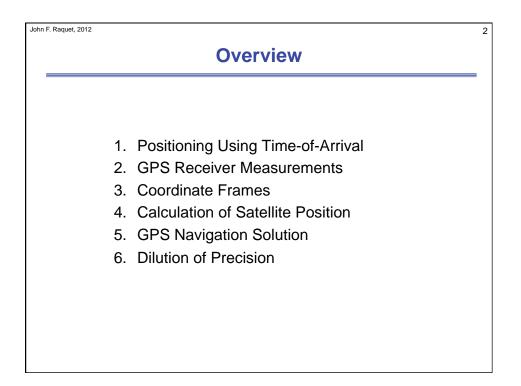
## Workshop on GNSS Data Application to Low Latitude Ionospheric Research

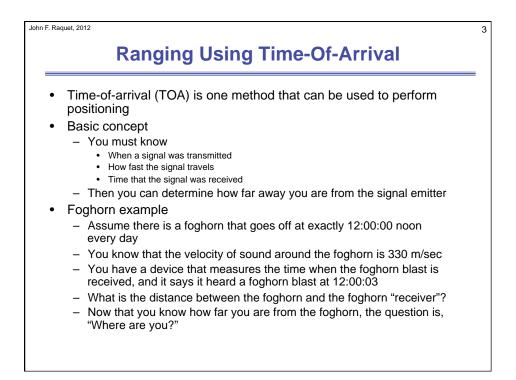
6 - 17 May 2013

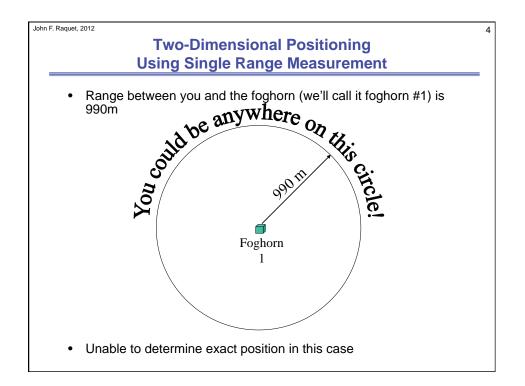
## **Calculation of GPS PNT Solution**

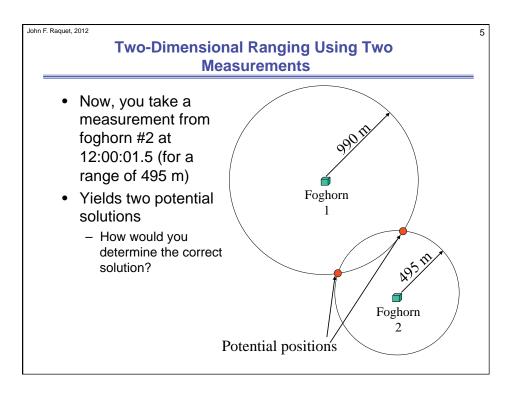
John F. Raquet Air Force Institute of Technology USA

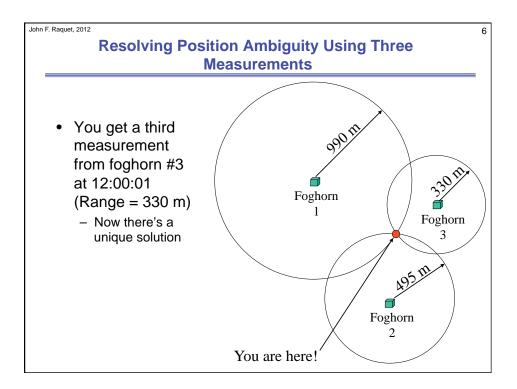


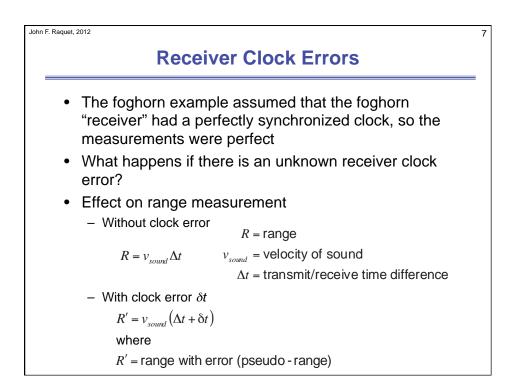


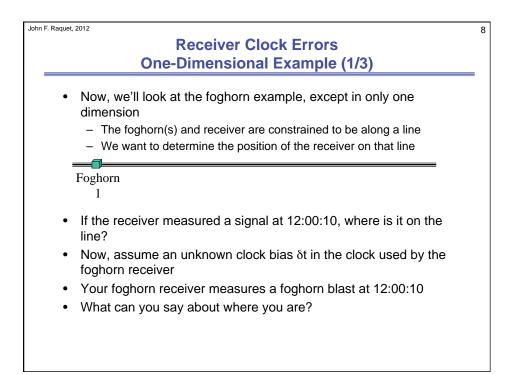


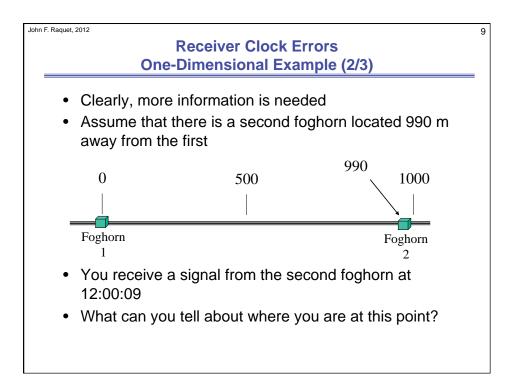




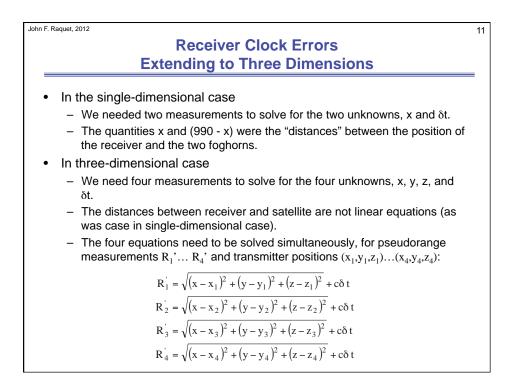


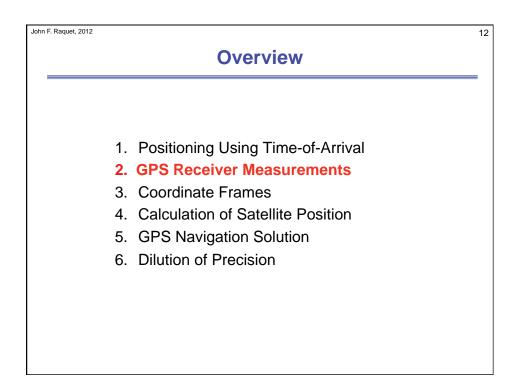


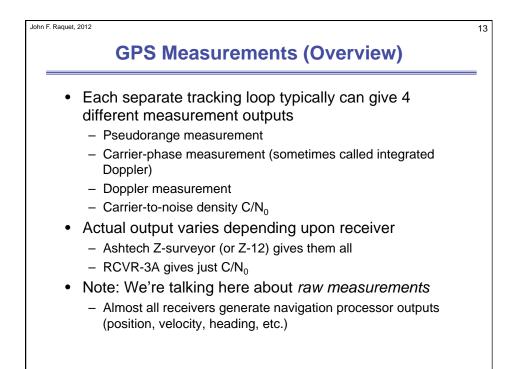


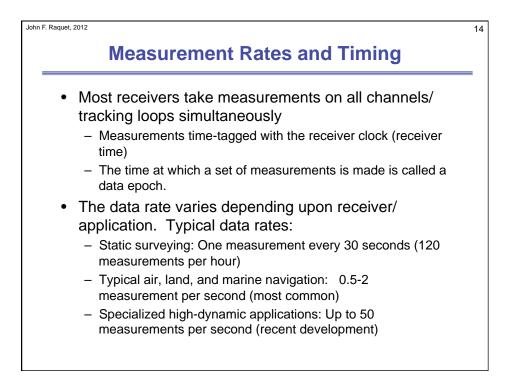


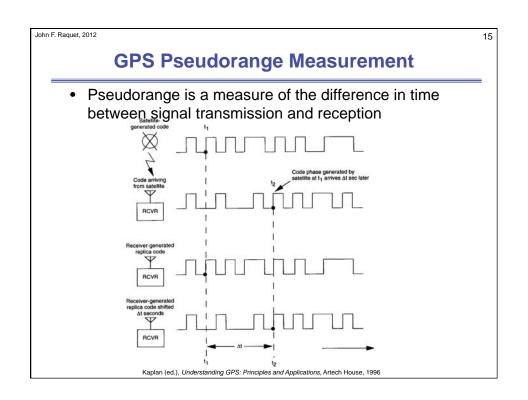
Raquet, 2012	
Receive	er Clock Errors
One-Dimens	ional Example (3/3)
<ul> <li>Here are the measure</li> </ul>	ments we have:
Pseudorange?	$1 = 330 \times 10 = 3300 = R_1'$
Pseudorange	$2 = 330 \times 9 = 2970 = R_2'$
<ul> <li>From the pseudorange</li> </ul>	equation:
$R_1' = v_{sound} \left( \Delta t_1 + \delta t \right)$	$= x + v_{sound} \delta t = 3300$
$R_2' = v_{sound} \left( \Delta t_2 + \delta t \right)$	$) = 990 - x + v_{sound} \delta t = 2970$
<ul> <li>Rearranging terms we</li> </ul>	get
$x + v_{sour}$	$d_{d}\delta t = 3300$
$x - v_{sour}$	$d_{d}\delta t = -1980$
<ul> <li>We can then solve for</li> </ul>	the two unknowns
$\delta t =$	8 seconds Does this work?
x = 6	560 m

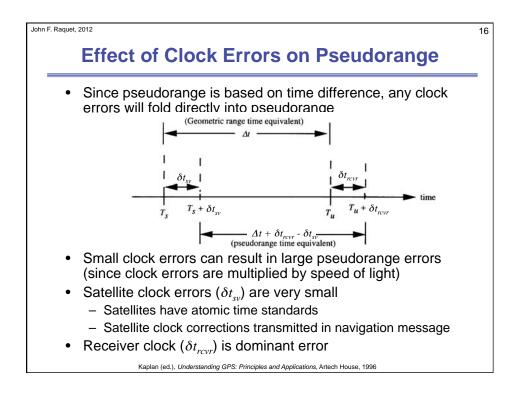


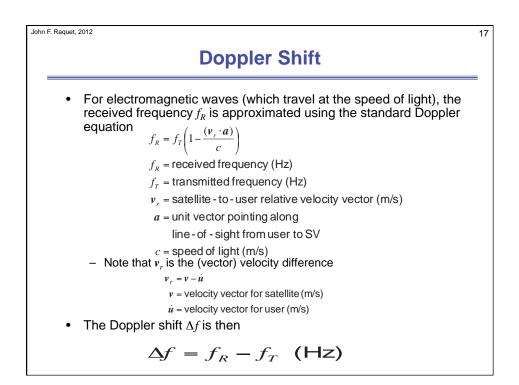


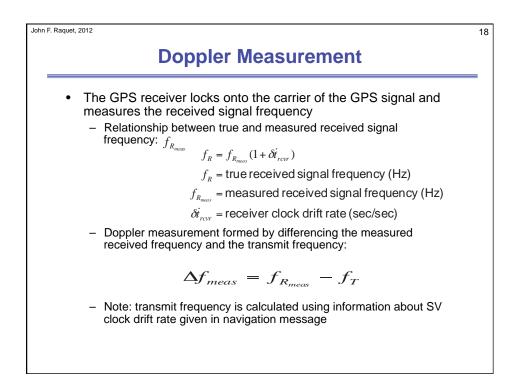


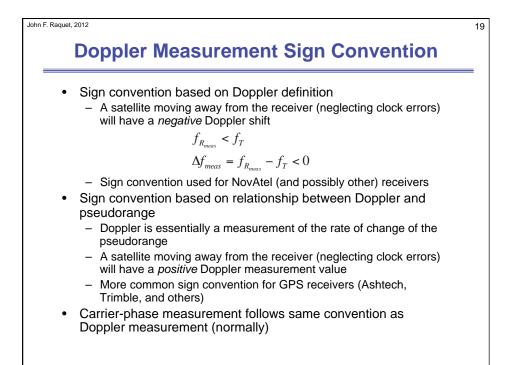


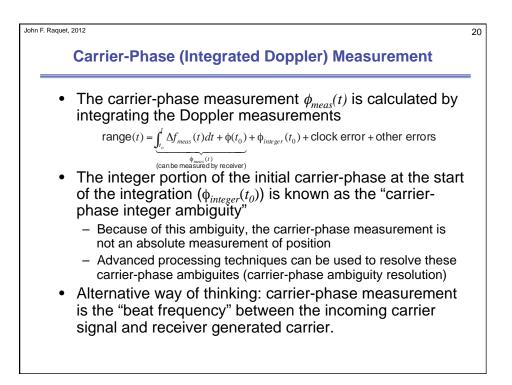


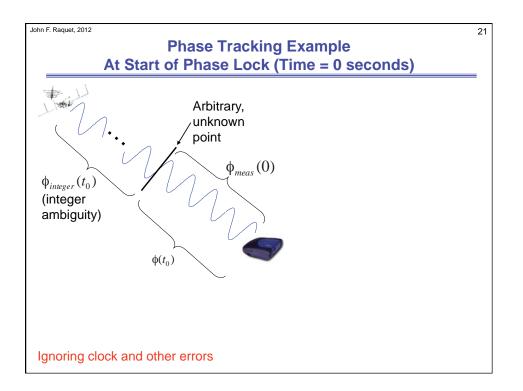


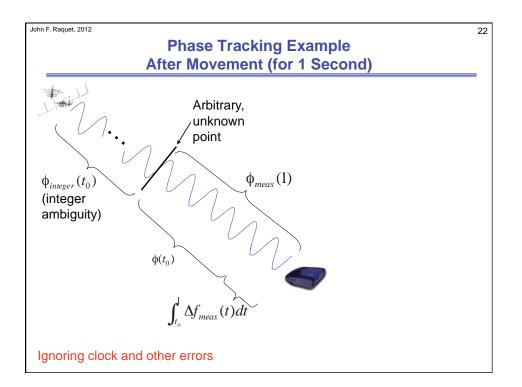


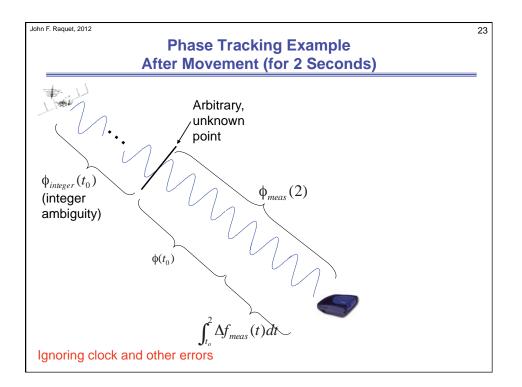




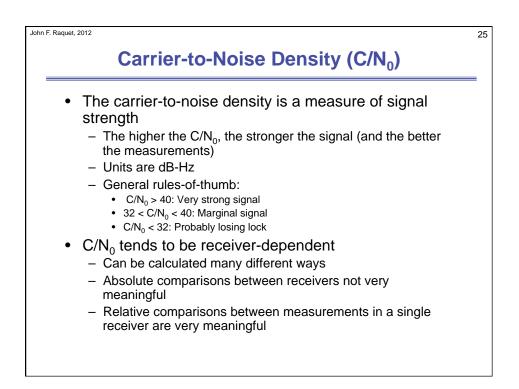


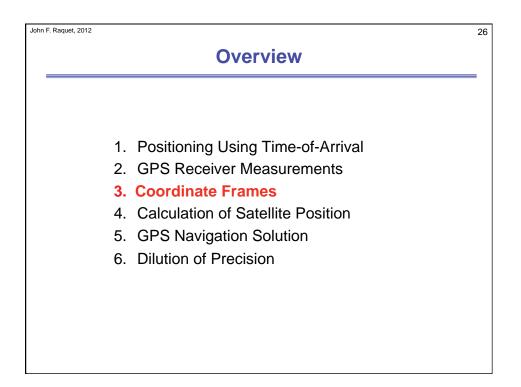


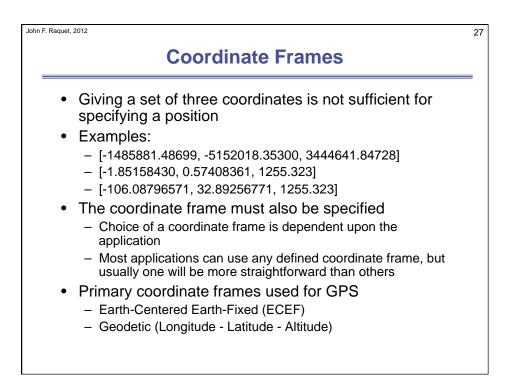


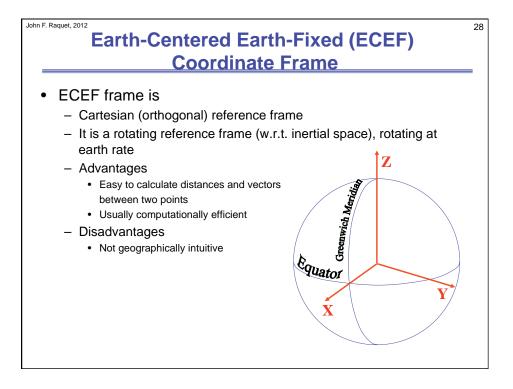


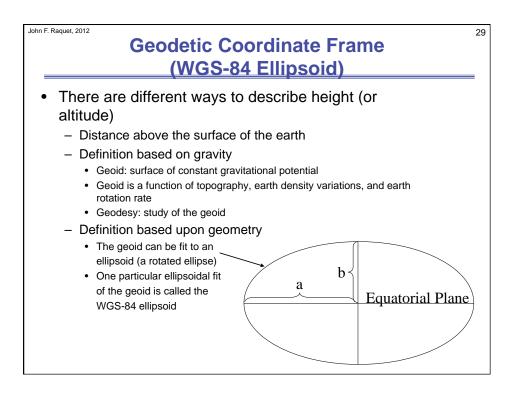
	Pseudorange	Carrier-Phase
Type of measurement	Range (absolute)	Range (ambiguous)
Measurement precision	~1 m	~0.01 m
Robustness	More robust	Less robust (cycle slips possible)
		Necessary for high precision GPS

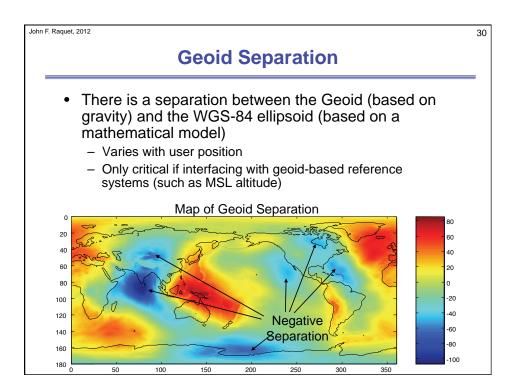


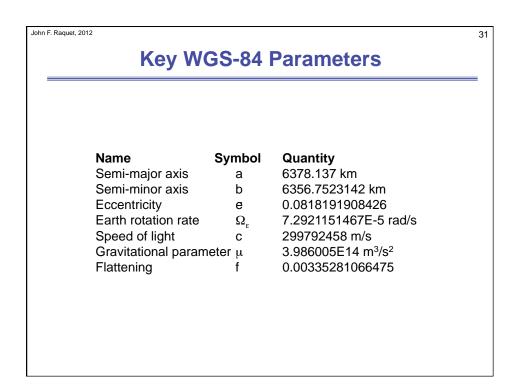


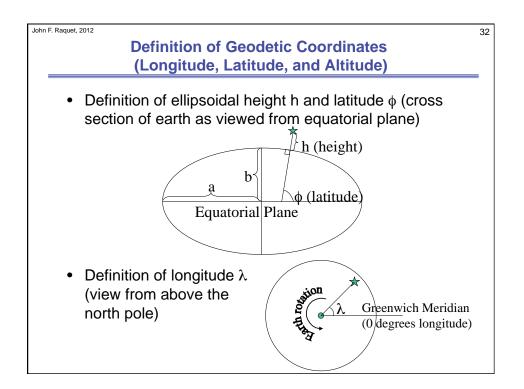


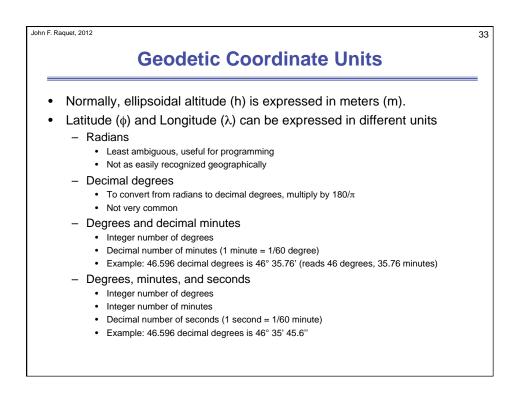


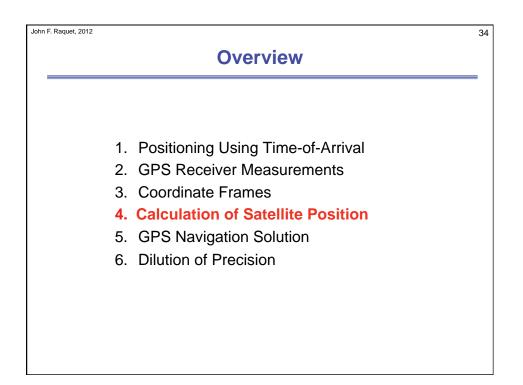


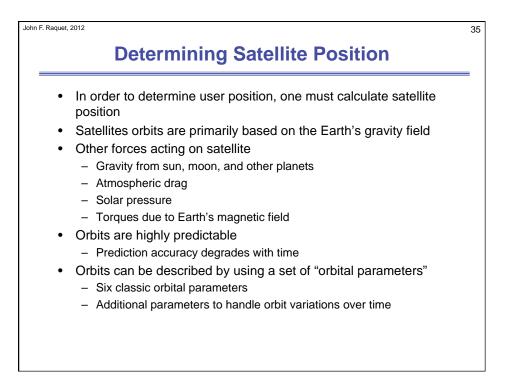


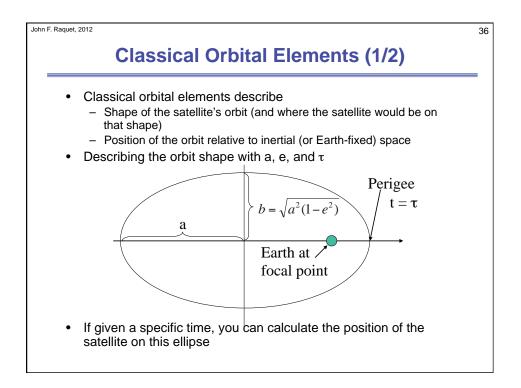


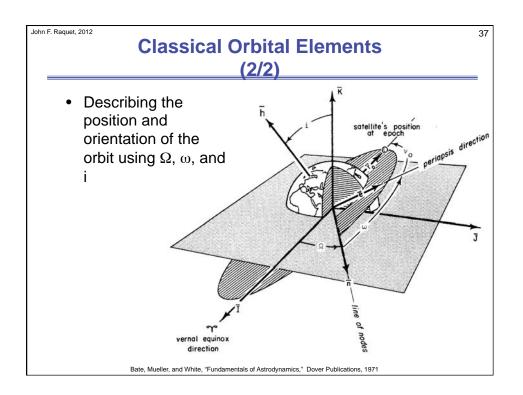


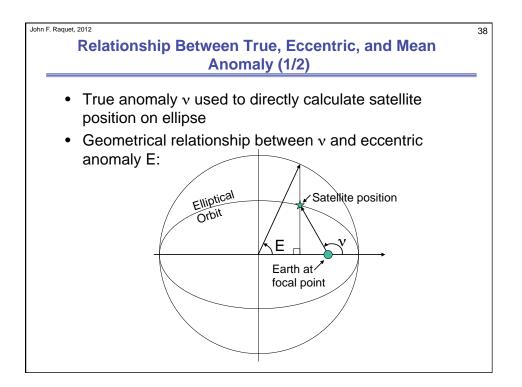














39

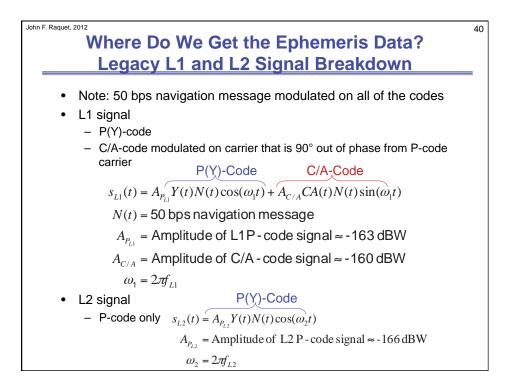
Mean anomaly M varies linearly with time (unlike E or v), so it can be easily calculated

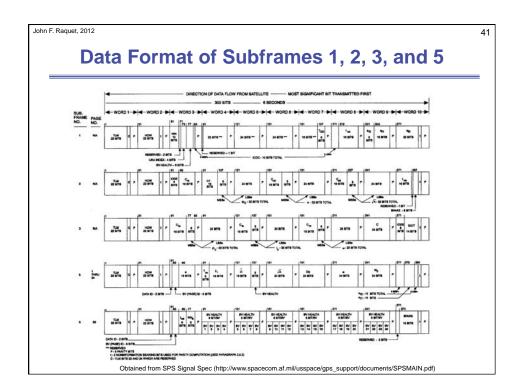
$$M(t) = M_0 + n(t - t_0)$$
$$M_0 = M(t_0)$$
$$\mu$$

 $n = \sqrt{\frac{\mu}{a^3}}$  = mean motion Eccentric anomaly and mean anomaly related through Kepler's equation

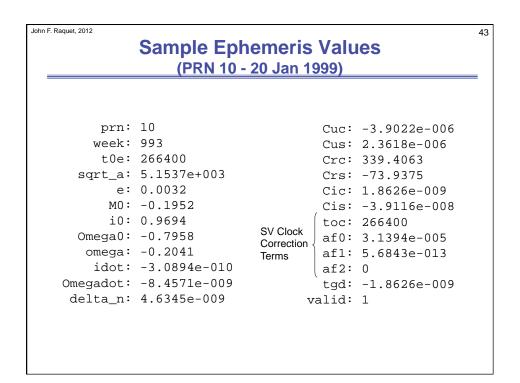
$$M = E - e\sin E$$

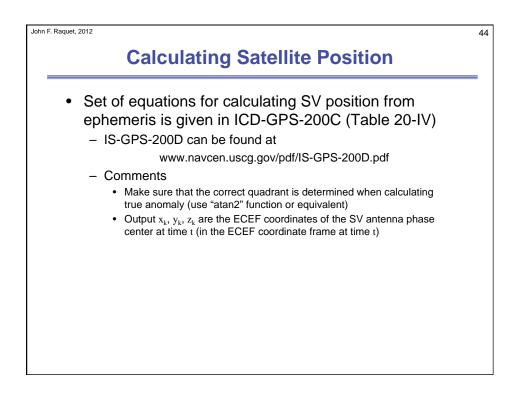
- Finally, true anomaly calculated from arctangent\* function, using  $\sin v = \frac{\sqrt{1 - e^2} \sin E}{1 - e \cos E}$   $\cos v = \frac{\cos E - e}{1 - e \cos E}$ 
  - \*Be sure to use the 4-quadrant arctangent function (atan2 in MATLAB).

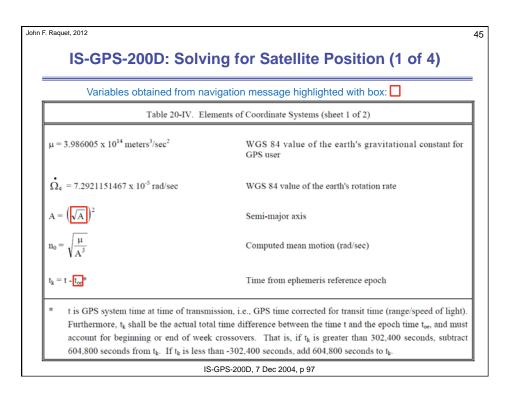


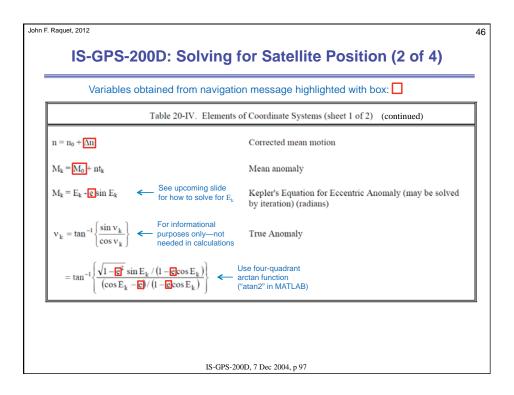


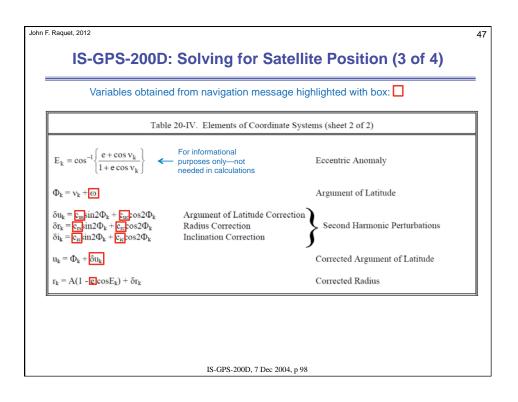
. Raquet, 2012	<b>GPS Ephemeris Data</b>	
(From Navigation Message)		
• For d	lefining orbit shape and timing	
	$t_{0}$ = Reference time of ephemeris (sec)	
	$\sqrt{a}$ = Square root of semi - major axis (m <sup>1/2</sup> )	
	<i>e</i> = Eccentricity	
	$M_0$ = Mean anomaly at time $t_{0_c}$ (rad)	
<ul> <li>For d</li> </ul>	lefining orientation/position of orbit <i>i</i> <sub>0</sub> = inclination at time <i>t</i> <sub>0</sub> (rad)	
	$\Omega_0$ = Longitude of ascending node at $t_0$ (rad)	
	$\omega$ = Argument of perigee at $t_{0_c}$ (rad)	
Corre	ection Terms	
	$\dot{i}$ = Rate of change of inclination (rad/sec)	
	$\dot{\Omega}$ = Rate of change of $\Omega$ (rad/sec)	
	$\Delta n$ = Mean motion correction (rad/sec)	
	$C_{uc}, C_{us}$ = Argument of latitude correction coefficients	
	$C_{rc}, C_{rs}$ = Orbital radius correction coefficients	
	$C_{ic}, C_{is}$ = Inclination correction coefficients	

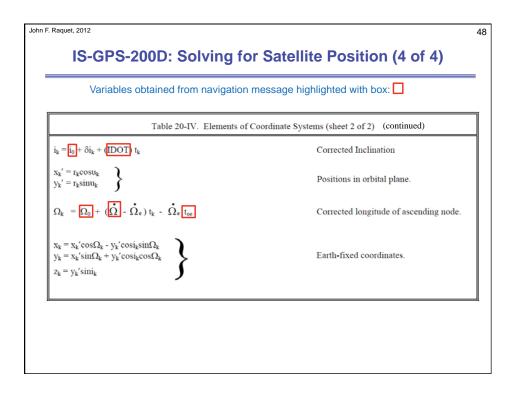


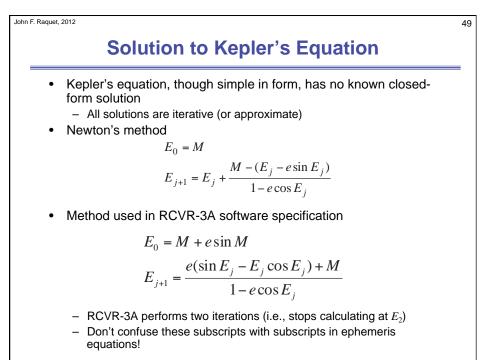


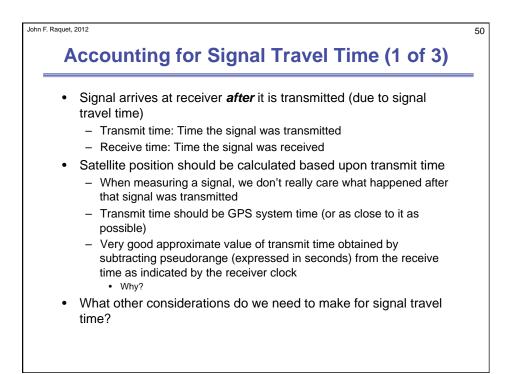


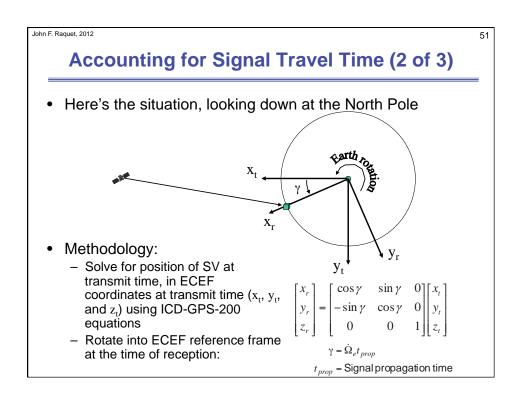


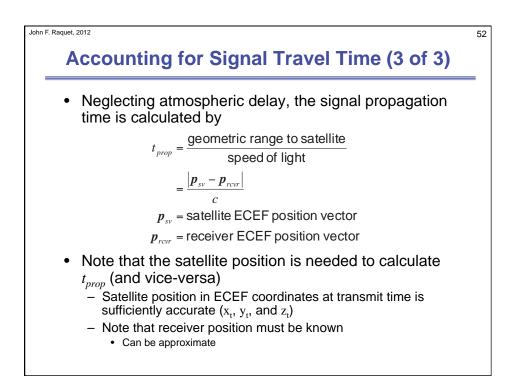


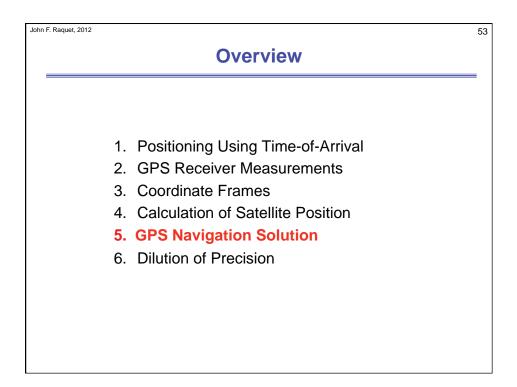




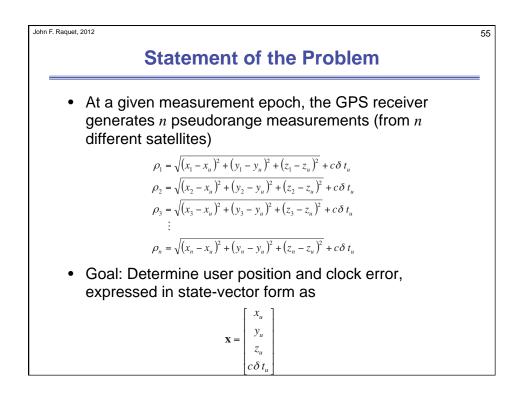


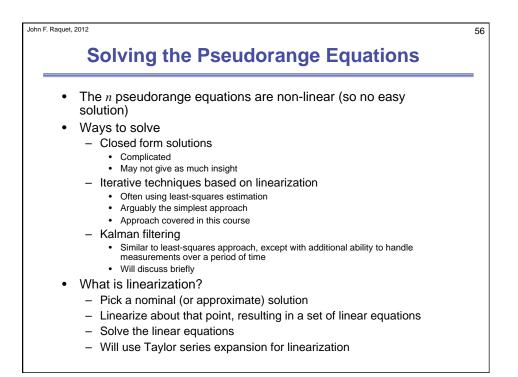


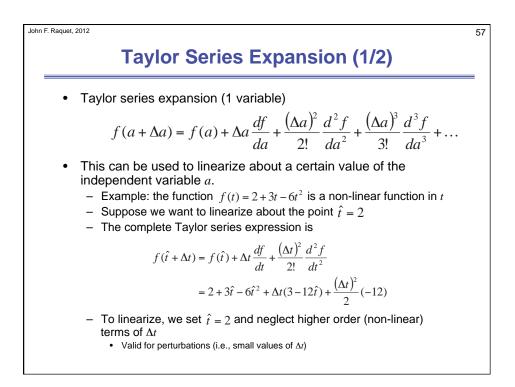


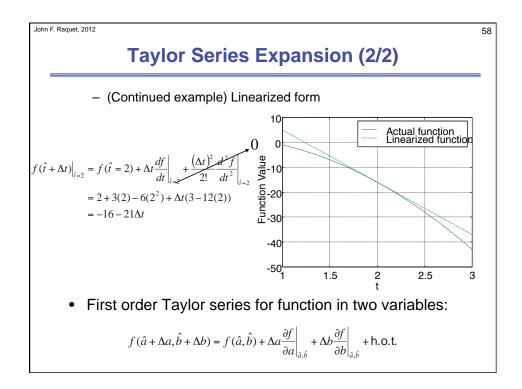


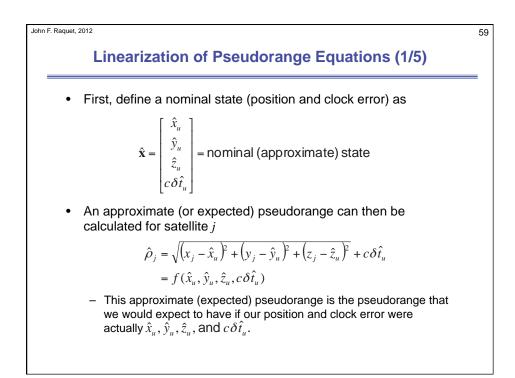
<b>Pseudorange Equation</b>	
•	The pseudorange is the sum of the true range plus the receiver
	clock error
	<ul> <li>We're assuming (for now) that the receiver clock error is the only remaining error</li> </ul>
	SV clock error has been corrected for
	<ul> <li>All other errors are deemed negligible (or have been corrected)</li> </ul>
	$\rho_{j} = \sqrt{(x_{j} - x_{u})^{2} + (y_{j} - y_{u})^{2} + (z_{j} - z_{u})^{2}} + c\delta t_{u}$
	$=f(x_u, y_u, z_u, \delta t_u)$
	$\rho_{i}$ = pseudorange measurement from satellite $j$ (m)
	$x_i, y_i, z_i = \text{ECEF}$ position of satellite j (m)
	$x_u, y_u, z_u = \text{ECEF position of user (m)}$
	$\delta t_u$ = receiver clock error (sec)
•	For now, only use one type of pseudorange (L1 C/A, L1 P, or L2 P
	······································

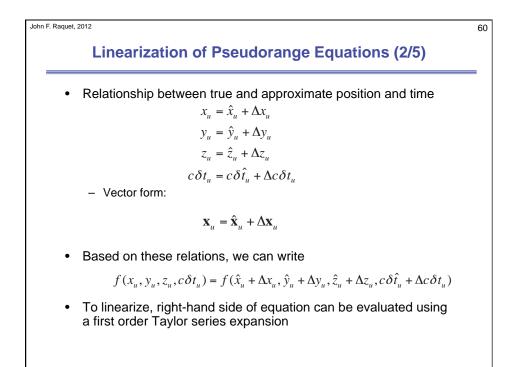


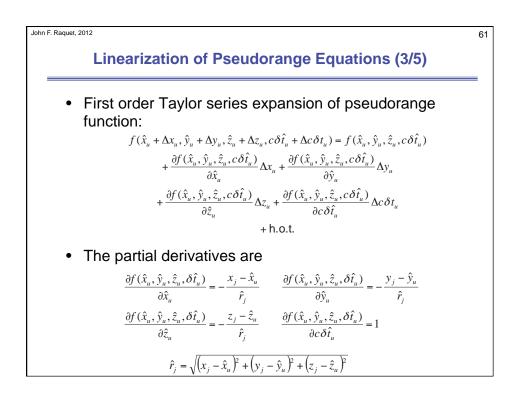


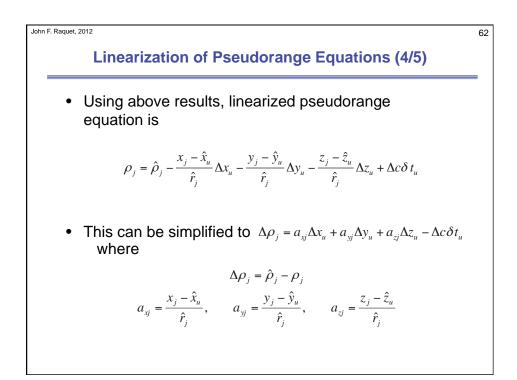


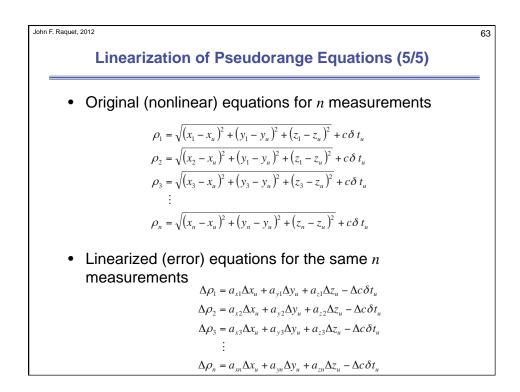


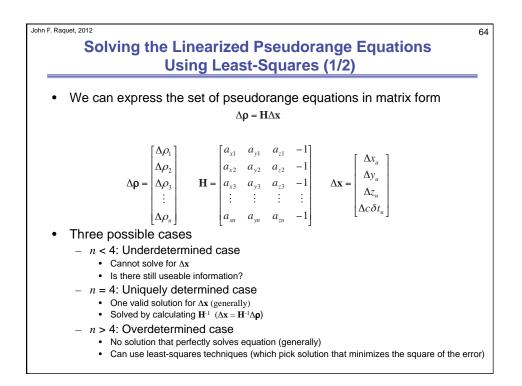


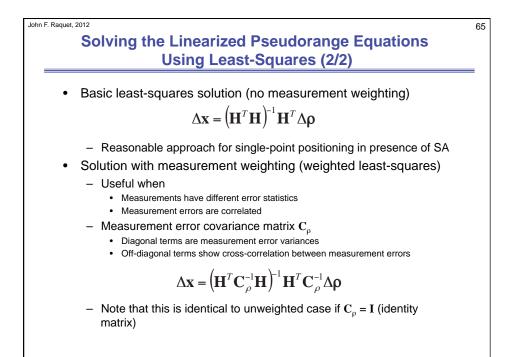


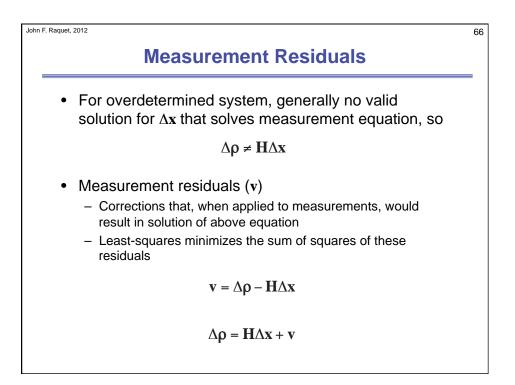


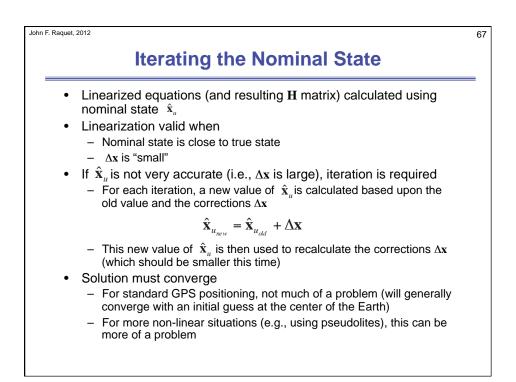


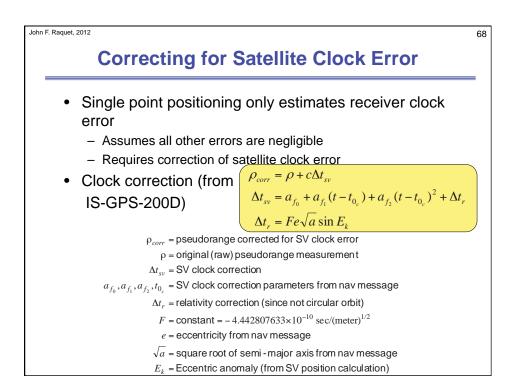


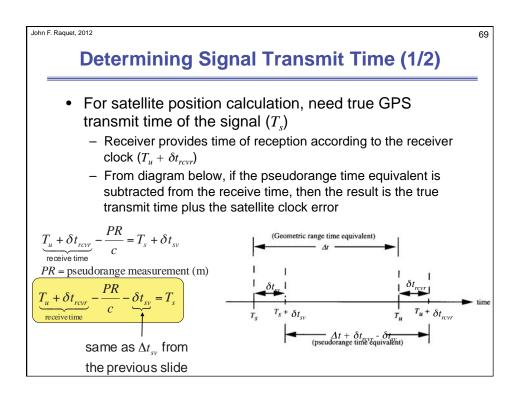


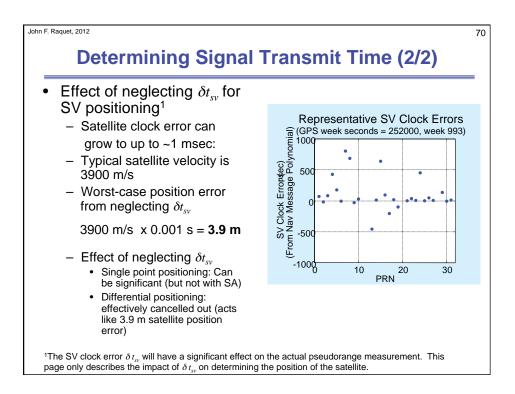














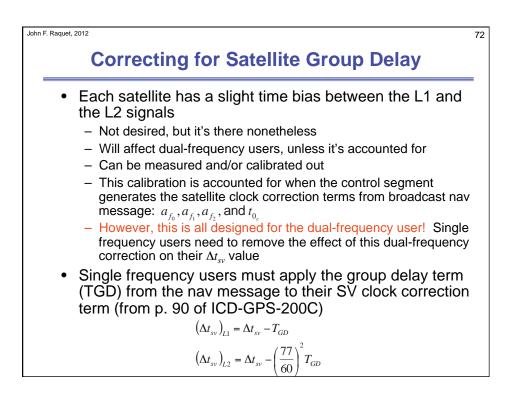
71

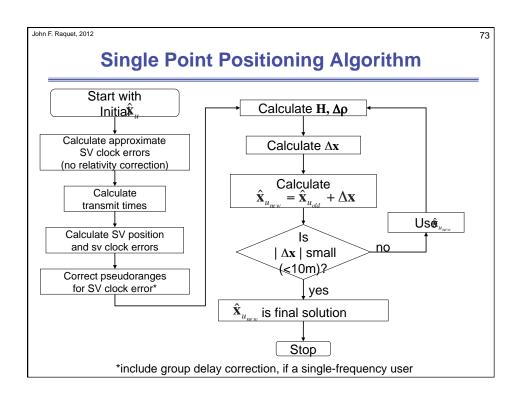
• L1 ionospheric delay calculated by

$$\begin{split} \Delta S_{iono,corr_{L1}} &= \left(\frac{f_2^2}{f_2^2 - f_1^2}\right) \left(\rho_{L1} - \rho_{L2}\right) \\ \Delta S_{iono,corr_{L1}} &= \text{L1ionospheric delay (m)} \\ f_1, f_2 &= \text{L1and L2 carrier frequencies} \\ \rho_{L1}, \rho_{L2} &= \text{L1and L2 pseudorange measurements} \end{split}$$

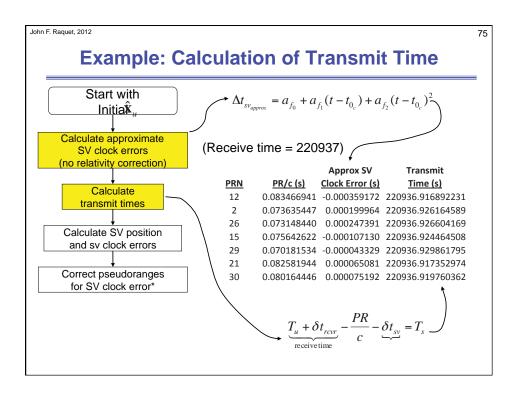
$$\text{L2 ionospheric delay can be calculated by} \\ \Delta S_{iono,corr_{L2}} &= \left(\frac{f_1}{f_2}\right)^2 \Delta S_{iono,corr_{L1}} \\ \text{e lonospheric-free pseudorange:} \\ \rho_{IF} &= \frac{\rho_{L2} - \gamma \rho_{L1}}{1 - \gamma}, \qquad \gamma = \left(\frac{f_{L1}}{f_{L2}}\right)^2 = \left(\frac{77}{60}\right)^2 \end{split}$$

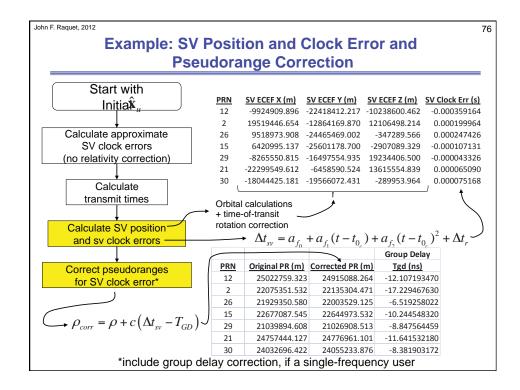
• Multipath and measurement noise will corrupt this measurement of ionosphere

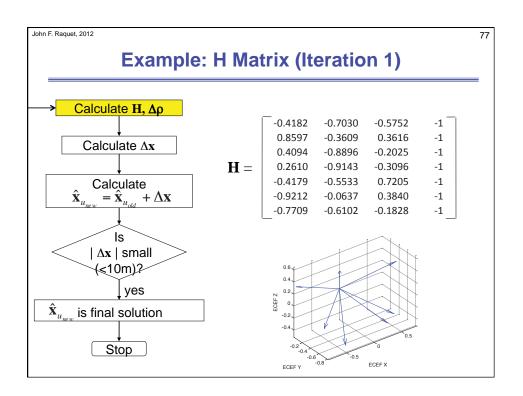


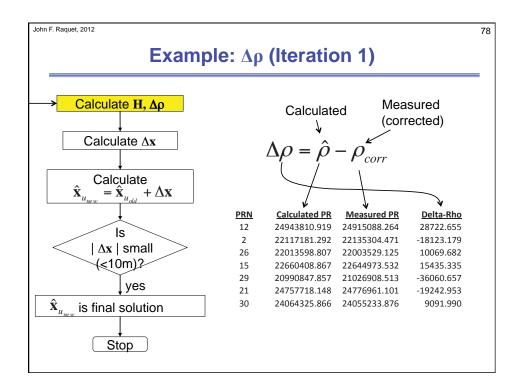


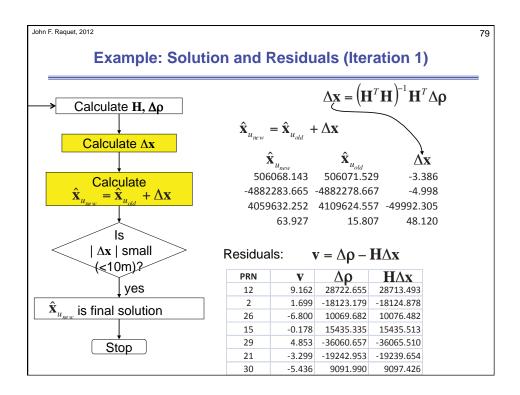
F. Raquet, 2012	ositi	oning Exampl	e
_	rement 529 -/ Initia	Ase to give an exar time (GPS week second 4882278.667 4109624.5 al guess of position error by ~50 km) Pseudorange 25022759.323 22075351.532 21929350.580 22677087.545 21039894.608 24757444.127 24032696.422	nds): 220937

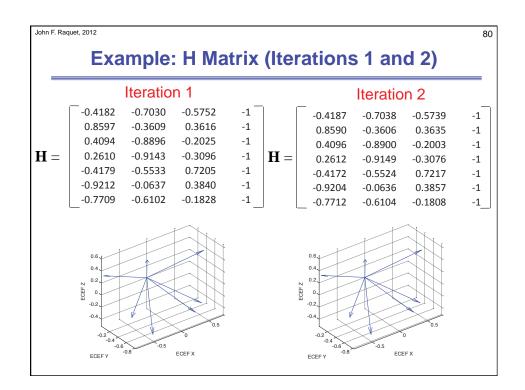


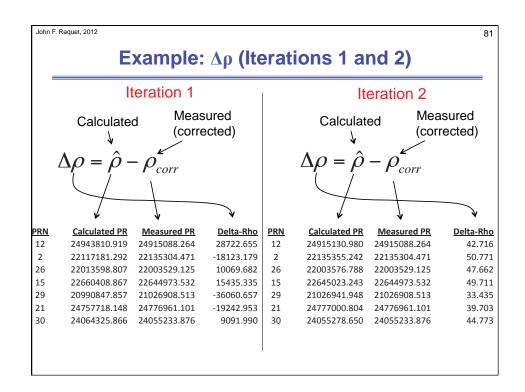




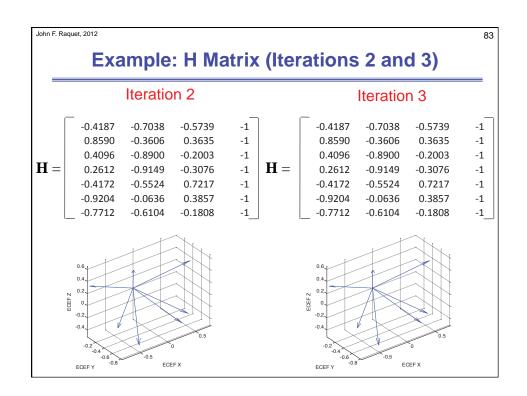


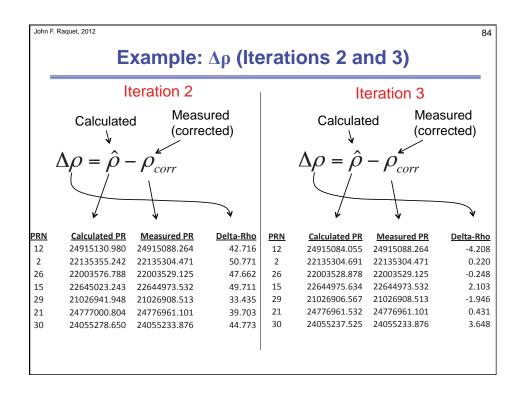


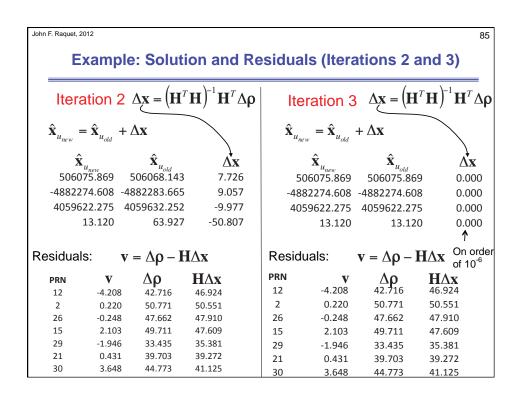




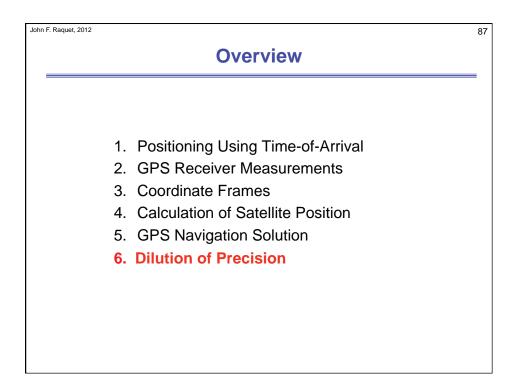
John F. Raquet,	2012							82	
E	Example: Solution and Residuals (Iterations 1 and 2)								
lter	Iteration 1 $\Delta \mathbf{x} = (\mathbf{H}^T \mathbf{H})^{-1} \mathbf{H}^T \Delta \boldsymbol{\rho}$					ation 2	$\Delta \mathbf{x} = \left( \mathbf{H}^{T} \right)$	$(\mathbf{H})^{-1}\mathbf{H}^T\Delta\mathbf{\rho}$	
$\mathbf{\hat{x}}_{u_{new}}$	$\hat{\mathbf{x}}_{u_{new}} = \hat{\mathbf{x}}_{u_{old}} + \Delta \mathbf{x}$				$\mathbf{\hat{x}}_{u_{new}}$	$= \hat{\mathbf{X}}_{u_{old}} +$	$-\Delta \mathbf{x}$		
ź	ŝ,	Â,	$\Delta$	x	ź	ŝ,	Â,	$\Delta \mathbf{x}$	
506	6068.143	506071.52	-3.3	386	5060	075.869	506068.143	7.726	
-4882	283.665	-4882278.66	57 -4.9	998	-48822	274.608 -	4882283.665	9.057	
4059	632.252	4109624.55	57 -49992.3	305	40590	622.275	4059632.252	-9.977	
	63.927	15.80	07 48.3	120		13.120	63.927	-50.807	
Residua	Residuals: $\mathbf{v} = \Delta \boldsymbol{\rho} - \mathbf{H} \Delta \mathbf{x}$				Residua	ls: v	$v = \Delta \rho - \mathbf{E}$	I∆x	
PRN	V	Δρ	HΔx		PRN	V	Δρ	HΔx	
12	9.162	28722.655	28713.493		12	-4.208	42.716	46.924	
2	1.699	-18123.179	-18124.878		2	0.220	50.771	50.551	
26	-6.800	10069.682	10076.482		26	-0.248	47.662	47.910	
15	-0.178	15435.335	15435.513		15	2.103	49.711	47.609	
29	4.853	-36060.657	-36065.510		29	-1.946	33.435	35.381	
21	-3.299	-19242.953	-19239.654		21	0.431	39.703	39.272	
30	-5.436	9091.990	9097.426		30	3.648	44.773	41.125	



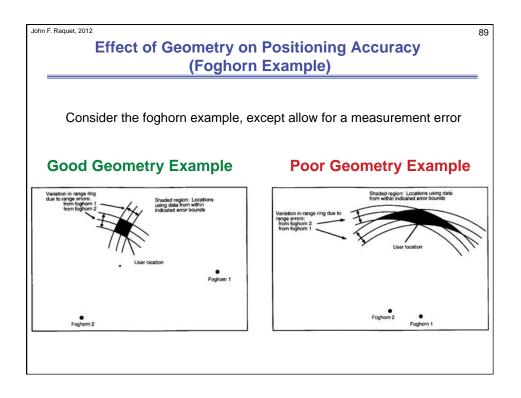




Convergence	
Practically speaking, getting the system to conv with GNSS is easy	/erge
<ul> <li>Example showed case where initial guess was 50 km error</li> </ul>	ı in
<ul> <li>Can start with the center of the Earth as a guess, and would only add an iteration or two</li> </ul>	l it
<ul> <li>Normally, a receiver will use its last solution as a star point, so only a single iteration is necessary</li> </ul>	ting
Nonlinearities (which drive the need for iteration more severe when dealing with pseudolites – Much closer to receiver than satellite	ו) are
<ul> <li>H matrix varies more quickly as a function of position</li> </ul>	



<ul> <li>Measurement Domain vs. Position Domain</li> <li>Pseudorange errors are errors in "measurement domain" <ul> <li>Errors in the measurements themselves</li> <li>UERE is one example</li> </ul> </li> <li>Ultimately, we'd like to know errors in "position domain" <ul> <li>The position errors that result when using the measurements</li> <li>Errors in position domain are different than measurement errors! <ul> <li>Can be larger</li> </ul> </li> </ul></li></ul>	_
<ul> <li>Errors in the measurements themselves</li> <li>UERE is one example</li> <li>Ultimately, we'd like to know errors in "position domain"</li> <li>The position errors that result when using the measurements</li> <li>Errors in position domain are different than measurement errors!</li> <li>Can be larger</li> </ul>	
<ul> <li>Ultimately, we'd like to know errors in "position domain"</li> <li>The position errors that result when using the measurements</li> <li>Errors in position domain are different than measurement errors!</li> <li>Can be larger</li> </ul>	
<ul> <li>Can be smaller</li> <li>Dependent on measurement geometry</li> </ul>	
<ul> <li>Mathematical representation         <ul> <li>We have covariance matrix of measurements (C<sub>ρ</sub>).</li> <li>We want covariance matrix of calculated position and clock error (C<sub>x</sub>)</li> </ul> </li> <li>In GPS applications, this problem is approached using concept called Dilution of Precision (DOP)</li> </ul>	



John F. Raquet, 2012 Obtaining $C_x$ from Least-Squares Analysis (1	90 /2)
• Definition of $\mathbf{C}_{\mathbf{x}}$ $C_{\mathbf{x}} = \begin{bmatrix} \sigma_{x_u}^2 & \sigma_{x_u y_u} & \sigma_{x_u z_u} & \sigma_{x_u \delta t_u} \\ \sigma_{x_u y_u} & \sigma_{y_u}^2 & \sigma_{y_u z_u} & \sigma_{y_u \delta t_u} \\ \sigma_{x_u z_u} & \sigma_{y_u z_u} & \sigma_{z_u}^2 & \sigma_{z_u \delta t_u} \\ \sigma_{x_u \delta t_u} & \sigma_{y_u \delta t_u} & \sigma_{z_u \delta t_u} & \sigma_{\delta t_u}^2 \end{bmatrix}$	
where, for example,	
$\sigma_{x_u}^2 = E\left[\left(x_u - E[x_u]\right)^2\right]$	
= variance of $x_u$	
$\sigma_{x_u y_u} = E\left[\left(x_u - E[x_u]\right)\left(y_u - E[y_u]\right)\right]$	
= covariance of $x_u$ and $y_u$	
• Definition of $\mathbf{C}_{\mathbf{p}} = \begin{bmatrix} \sigma_{\rho_1}^2 & \sigma_{\rho_1\rho_2} & \cdots & \sigma_{\rho_1\rho_n} \\ \sigma_{\rho_1\rho_2} & \sigma_{\rho_2}^2 & \cdots & \sigma_{\rho_2\rho_n} \\ \vdots & \vdots & \ddots & \sigma_{\rho_3\rho_n} \\ \sigma_{\rho_1\rho_n} & \sigma_{\rho_2\rho_n} & \sigma_{\rho_3\rho_n} & \sigma_{\rho_n}^2 \end{bmatrix}$	

