

2458-3

Workshop on GNSS Data Application to Low Latitude Ionospheric Research

6 - 17 May 2013

Ionospheric Structure and Scintillation

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Ionospheric Structure and Scintillation

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**Workshop on GNSS Data
Applications to Low Altitude
Ionospheric Research**

International Center for
Theoretical Physics

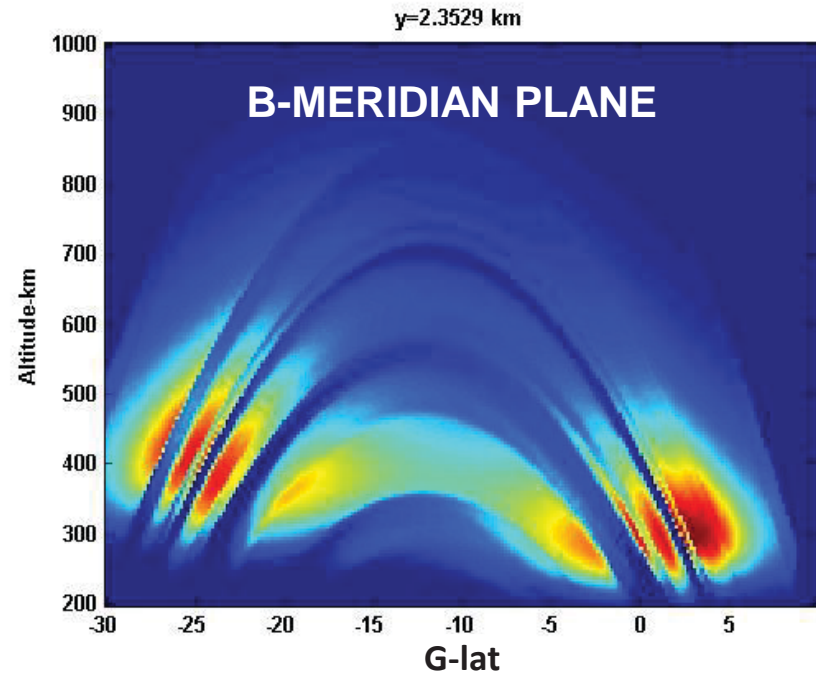
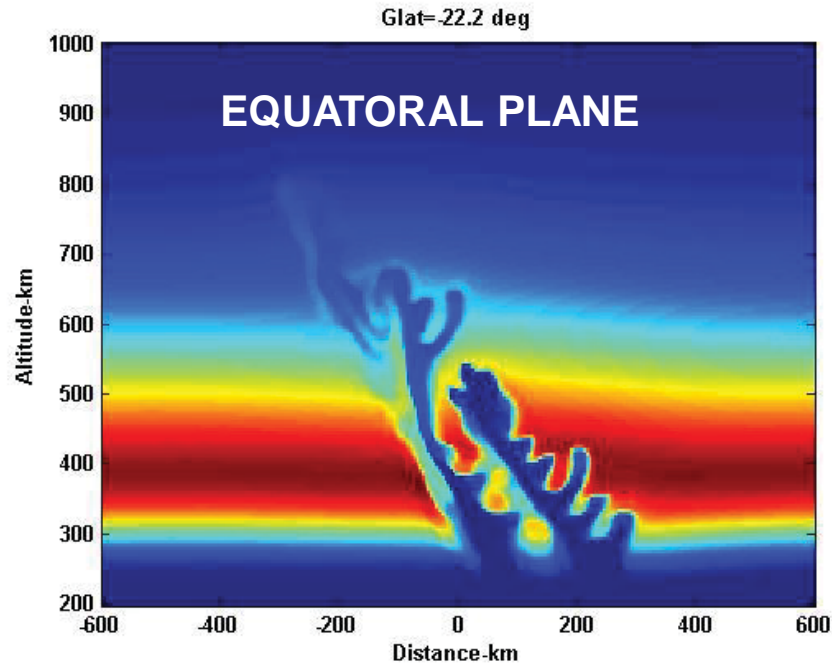
Trieste, Italy

May 2013

OUTLINE

1. Ionospheric Structure
 - Structure Characterization
2. Scintillation
 - Overview
 - Recent Applications

A systematic characterization of **Ionospheric Structure**



STRUCTURE CHARACTERISTICS

- Inhomogeneous
- Anisotropic
- **Current resolution limit of ~5 km excludes intermediate scale structure that causes scintillation**

Simulations provided by John Retterer, Boston College

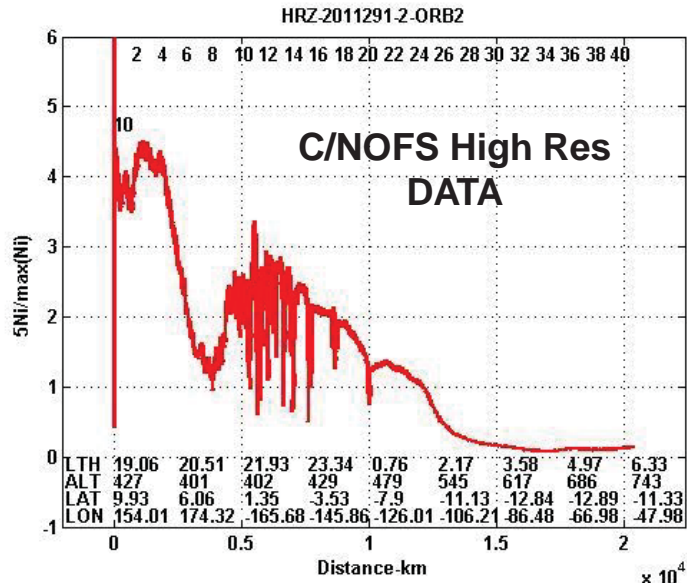
$$N(\mathbf{r}, t) = \bar{N}(\bar{\mathbf{r}}, \bar{t}) (1 + \delta N(\Delta \mathbf{r} - \mathbf{v} \Delta t) / N_0)$$

Slowly varying average
 centered at $\bar{\mathbf{r}}$ & \bar{t}

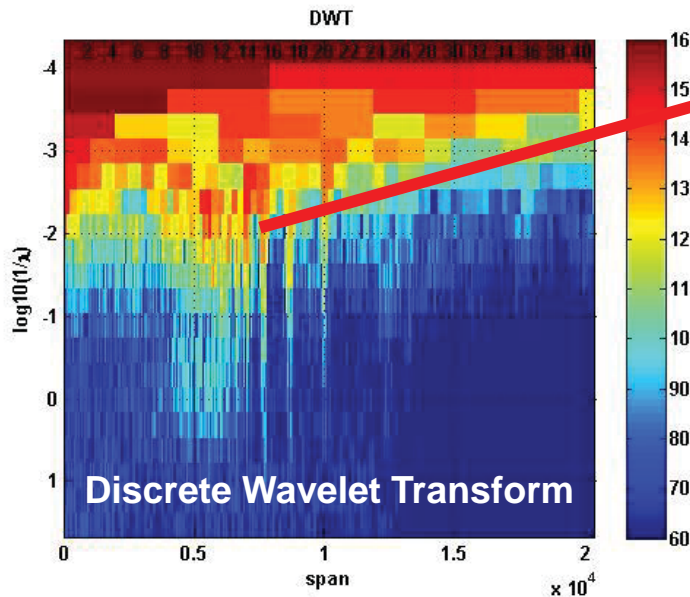
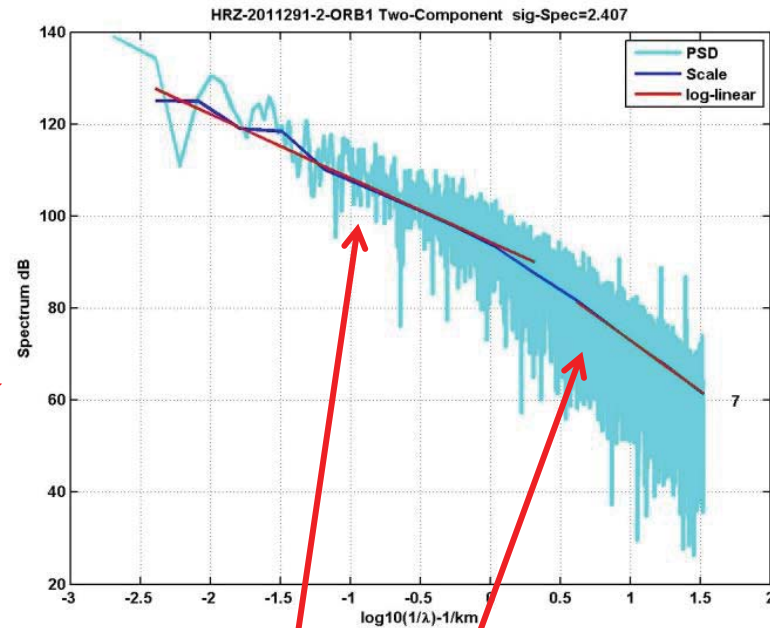
Structure frozen within
 volume defining $\bar{N}(\bar{\mathbf{r}}, \bar{t})$

Random component with
 mean zero over volume
 defining $\bar{N}(\bar{\mathbf{r}}, \bar{t})$

$\delta N(\mathbf{r})/N_0$ is typically **imposed** as a stochastic overlay with a well-defined spectral density function

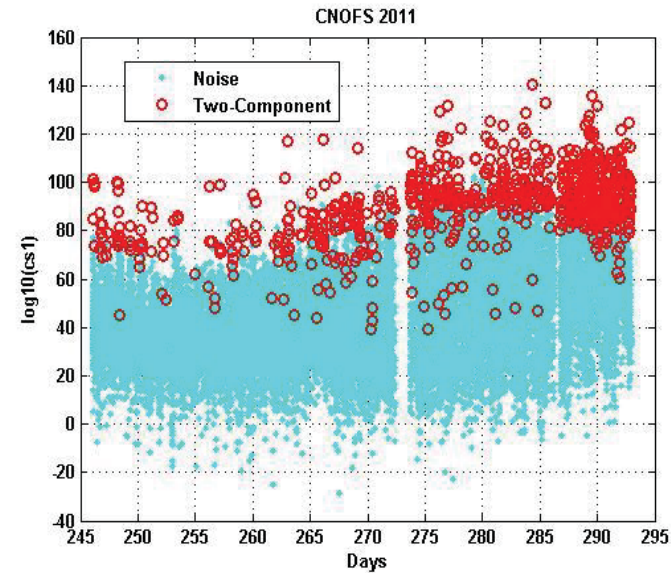
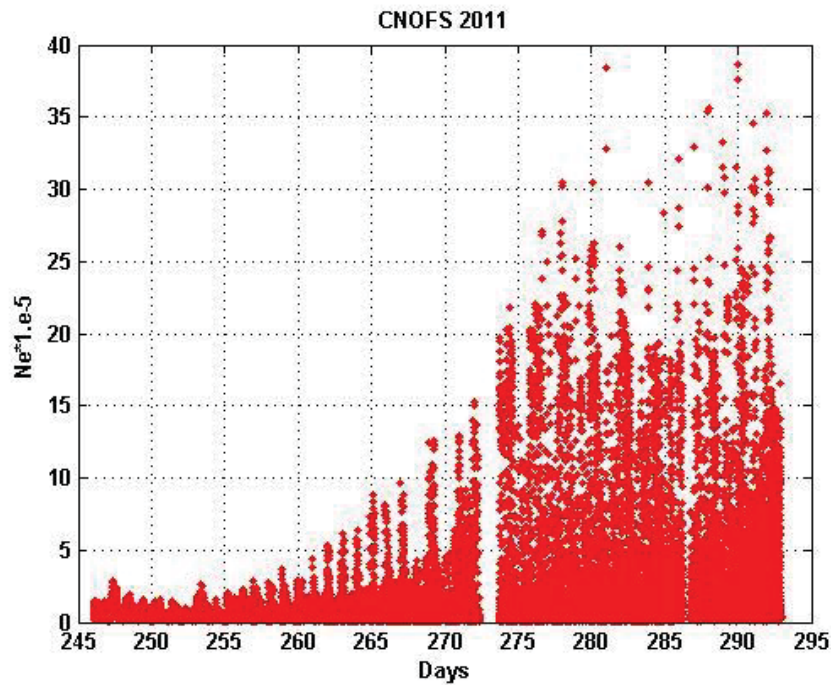


Spectral Density Function for well-developed segments

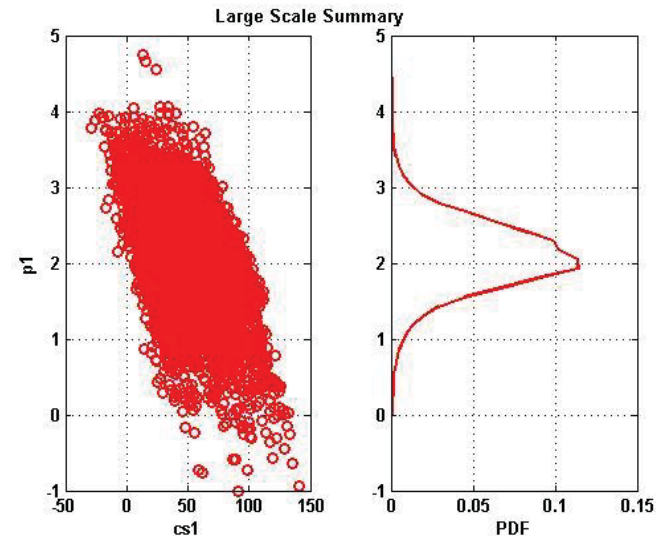


$$\varphi(q) = \begin{cases} c_s q^{-p_1} & \Delta q \leq q < q_0 \\ \underbrace{(c_s q_0^{(p_2-p_1)})}_{C_s^{(2)}} q^{-p_2} & q_0 \leq q < q_{\max} \end{cases}$$

285 consecutive C/NOFS data sets recorded 2011 day 246 through day 292



79,449 segments analyzed, 62,512 (79%) achieved overall least-square errors less than 10, 59,849 segments noise-limited ($p_2 < p_1$). Only 2663 noise free ($p_1 < p_2$).



- Data interpretation and extrapolation require a 3D structure model
- The model must accommodate inhomogeneous distributions with field-aligned structure that subtend scale sizes from hundreds of kilometers to hundreds of meters.
- Transition from quasi-deterministic variation to stochastic structure in the scale range from 100 to 10 is ill-defined
 - Fractional Brownian motion provides a theoretical framework that captures the transition
 - Configuration space models have been introduced as a 2.5D alternative
- An inverse power law subtending the entire intermediate scale range (>100 km to <100 m) with 3D index 4 (Kolmogorov index = $11/3$) will be used as a canonical reference

PRE PUBLICATION REFERENCES:

<http://chuckrino.com/wordpress/wp-content/uploads/2013/02/PLPHighResDataPaperRev15.pdf>

<http://chuckrino.com/wordpress/wp-content/uploads/2013/02/PLPHighResDataPaperRev21.pdf>

Scintillation

with modern computational resources

Integral form of Helmholtz Equation

$$\psi(\mathbf{r}) = \psi_i(\mathbf{r}) + 2k^2 \iiint \frac{\exp\{ik|\mathbf{r} - \mathbf{r}'|\}}{2\pi|\mathbf{r} - \mathbf{r}'|} \delta n(\mathbf{r}') \psi(\mathbf{r}') dV$$

Propagation

Media Interaction

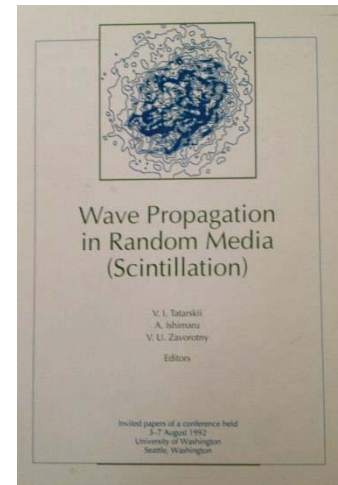
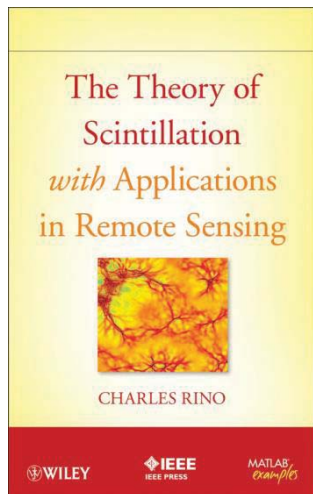
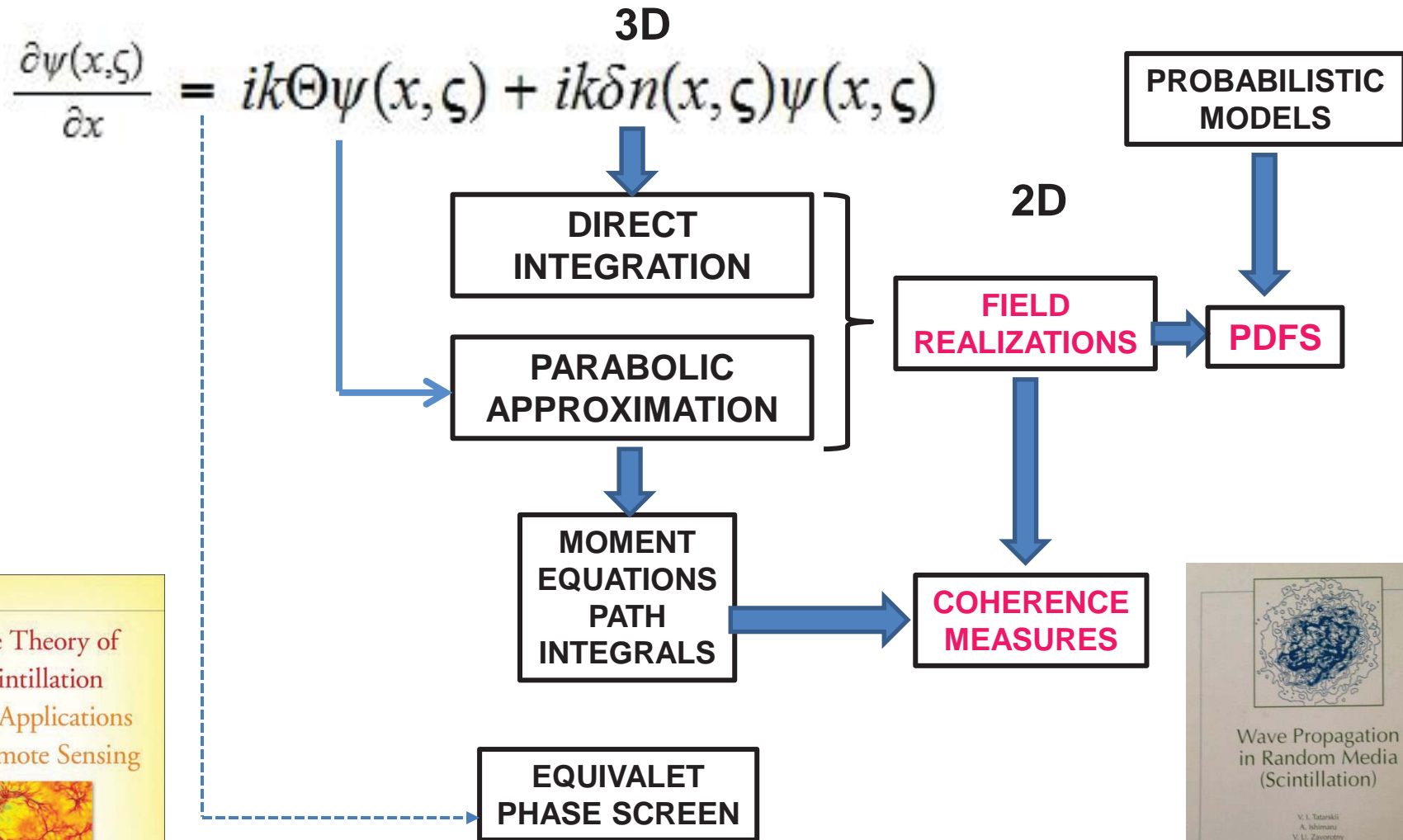
$$\frac{\partial \psi(x, \zeta)}{\partial x} = ik\Theta\psi(x, \zeta) + ik\delta n(x, \zeta)\psi(x, \zeta)$$

Forward Propagation Equation (FPE)

$$\hat{\psi}(\boldsymbol{\kappa}; x_0) = \iint \psi(x_0, \zeta) \exp\{-i\boldsymbol{\kappa} \cdot \boldsymbol{\zeta}\} d\boldsymbol{\zeta}$$

$$\Theta\psi(x, \zeta) = \iint \hat{\psi}(\boldsymbol{\kappa}; x_0) \exp\left\{ik\sqrt{1 - (\boldsymbol{\kappa}/k)^2} x\right\} \exp\{i\boldsymbol{\kappa} \cdot \boldsymbol{\zeta}\} \frac{d\boldsymbol{\kappa}}{(2\pi)}$$

$$\sqrt{1 - (\boldsymbol{\kappa}/k)^2} \simeq 1 - (\boldsymbol{\kappa}/k)^2/2 \text{ Parabolic Approximation}$$



- **The forward propagation equation is valid for any refractive index configuration that has no gradients steep enough to require explicit boundary treatment**
- **Split-step integration is suggested by the FPE and has proven to be very effective**
 - **The forward integration step can subtend hundreds of wavelengths, as long the associated amplitude change is small**
 - **Transverse sampling must capture the smallest significant structure**
- **Phase-screen realizations are generated by filtering uncorrelated noise**
 - **Fractional Brownian motion realizations are generated by extending the inverse power law to the largest resolved scale**
 - **The effects of constrained power-law distributions will be demonstrated**

$$\Phi_I(\boldsymbol{\kappa}; U, \rho_F) = \iint (\exp\{-U^{p_1-2} \gamma(\boldsymbol{\eta}, \boldsymbol{\kappa})\} - 1) \exp\{-i\boldsymbol{\kappa} \rho_F \cdot \boldsymbol{\eta}\} d\boldsymbol{\eta}$$

$$\rho_F = \sqrt{x/k}$$

$$U = \left(C_p^{\frac{1}{p_1-2}} \right) \rho_F$$

$$C_p = r_e^2 \lambda^2 l_p C_s$$

p_1 is the 2D phase power-law index, which is equal to the 3D in-situ power-law index

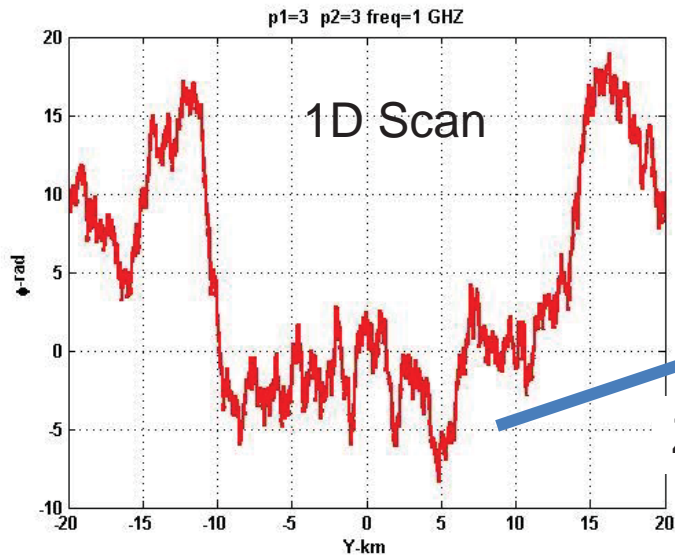
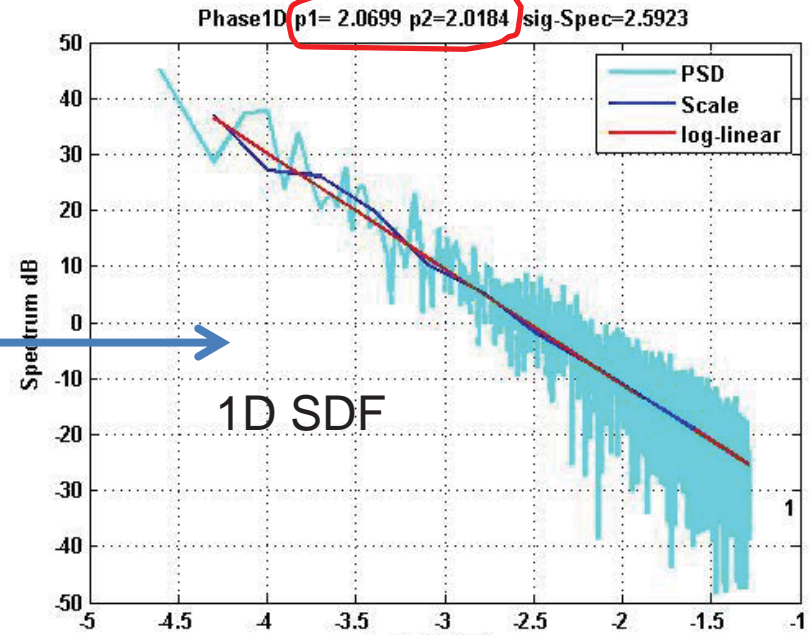
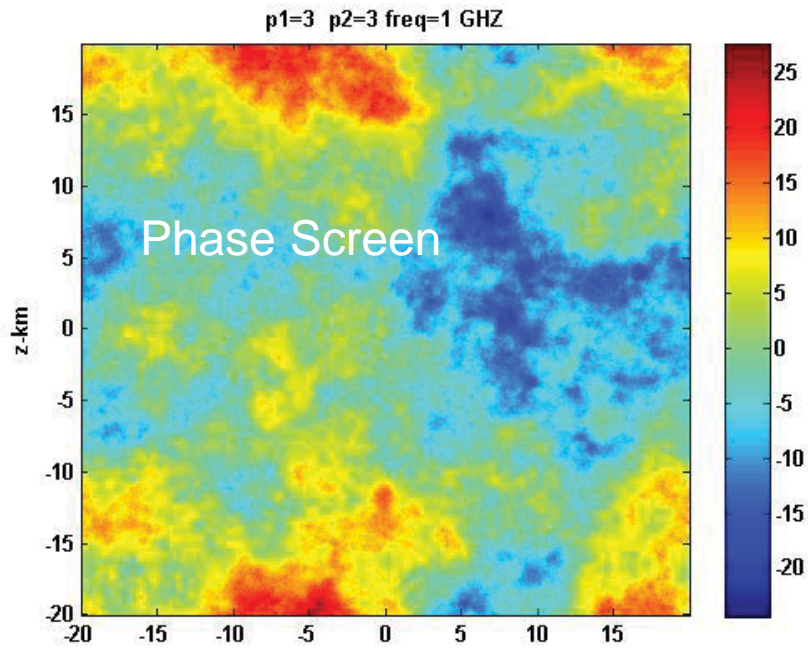
$$\gamma(\boldsymbol{\eta}, \boldsymbol{\mu}) = 8 \iint \begin{cases} \chi^{-p_1} & \chi < \chi_0 \\ \chi_0^{p_2-p_1} \chi^{-p_2} & \chi > \chi_0 \end{cases} \sin^2(\boldsymbol{\chi} \cdot \boldsymbol{\eta}/2) \sin^2(\boldsymbol{\chi} \cdot \boldsymbol{\mu}/2) d\boldsymbol{\chi} / (2\pi)^2$$

$$\chi_0 = q_0 \rho_F$$

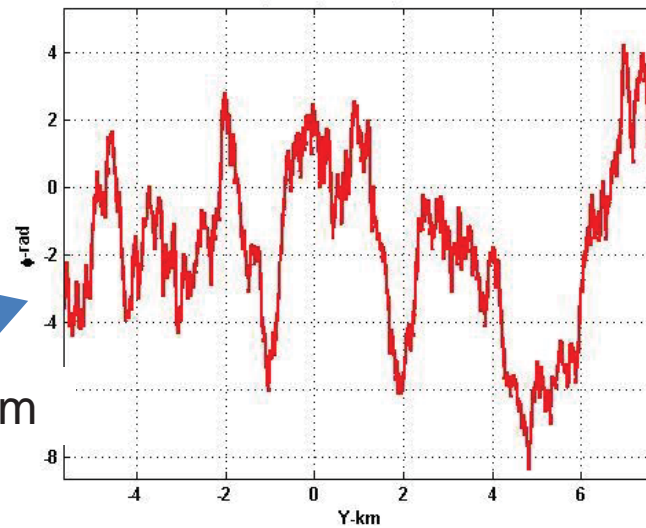
$$\lim_{U \rightarrow 0} \Phi_I(\boldsymbol{\kappa}; U, \rho_F) = 4C_p \varphi(\boldsymbol{\kappa}) \sin^2((\boldsymbol{\kappa} \rho_F)^2/2)$$

Power-Law Index Dependence

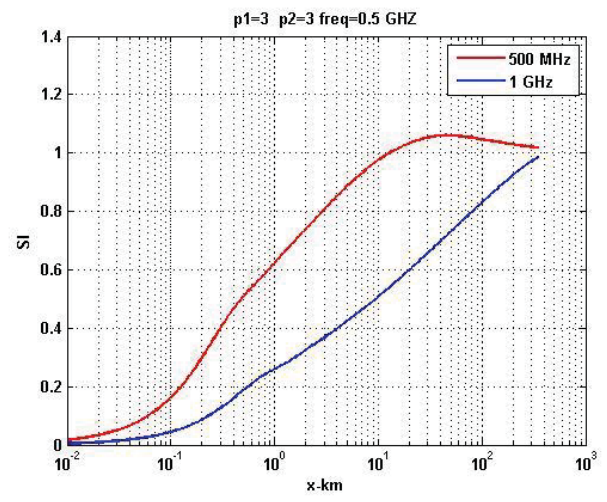
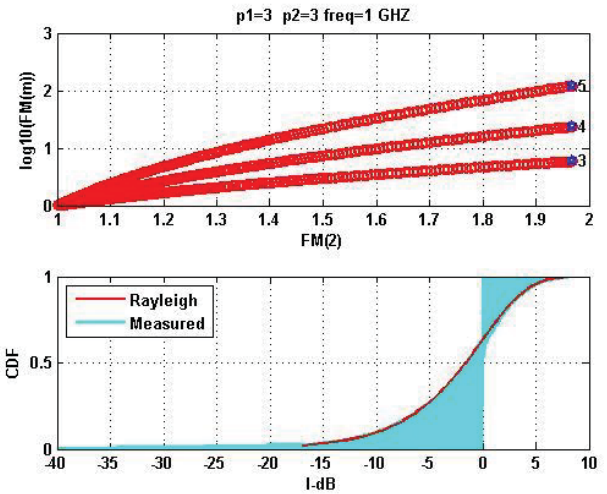
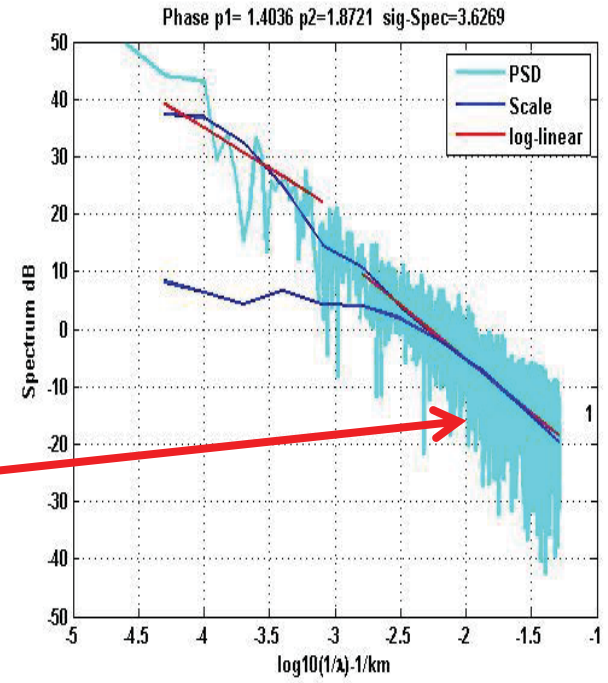
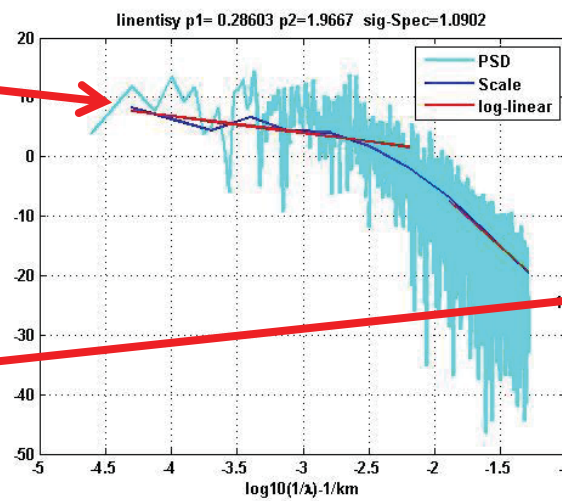
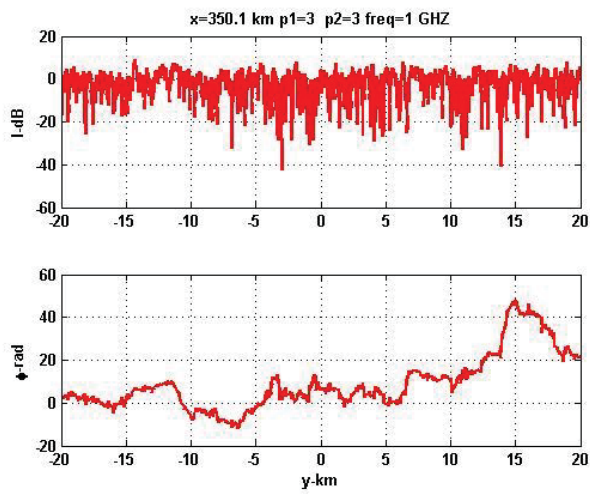
P(3D)=3 Phase Screen Realization Infinite Outer Scale



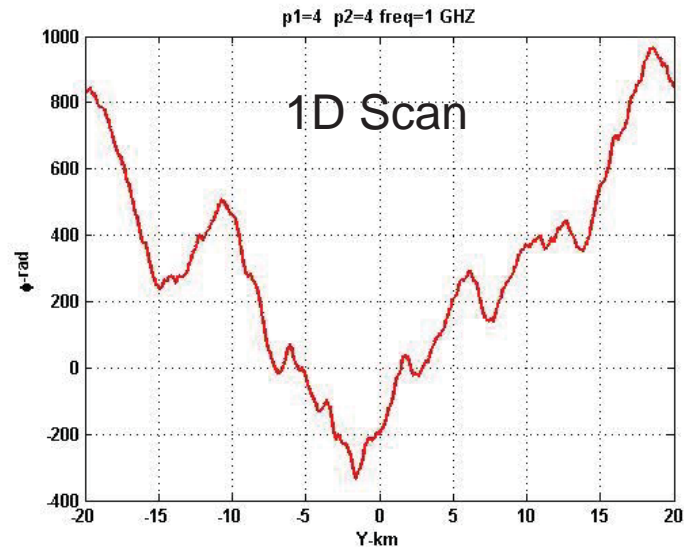
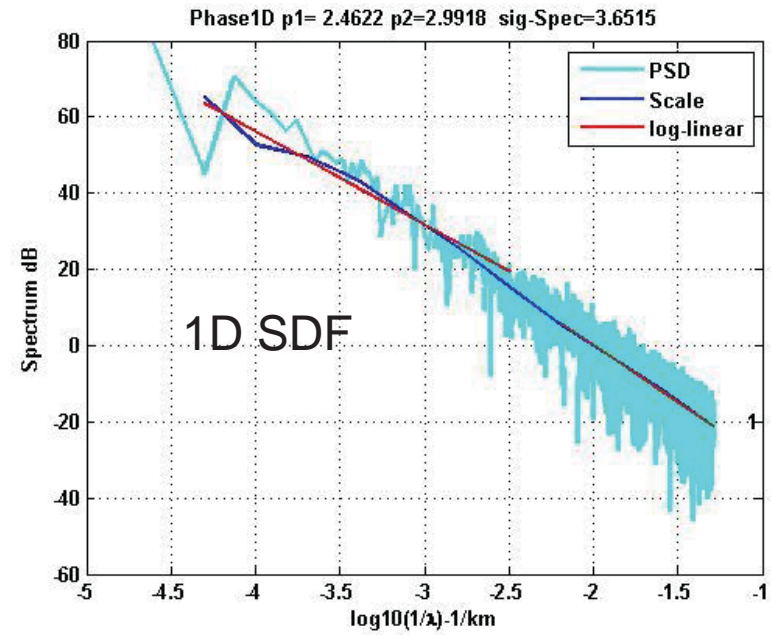
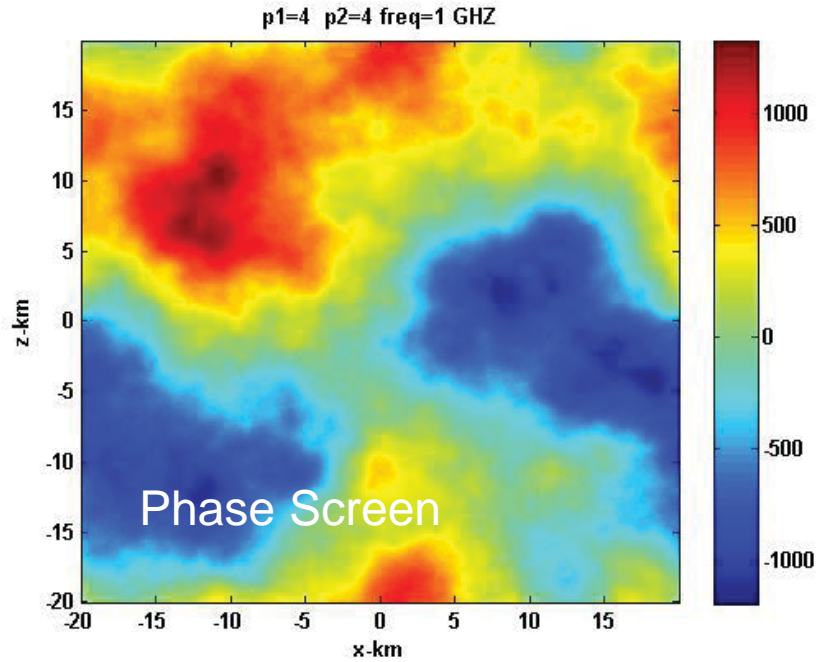
Zoom



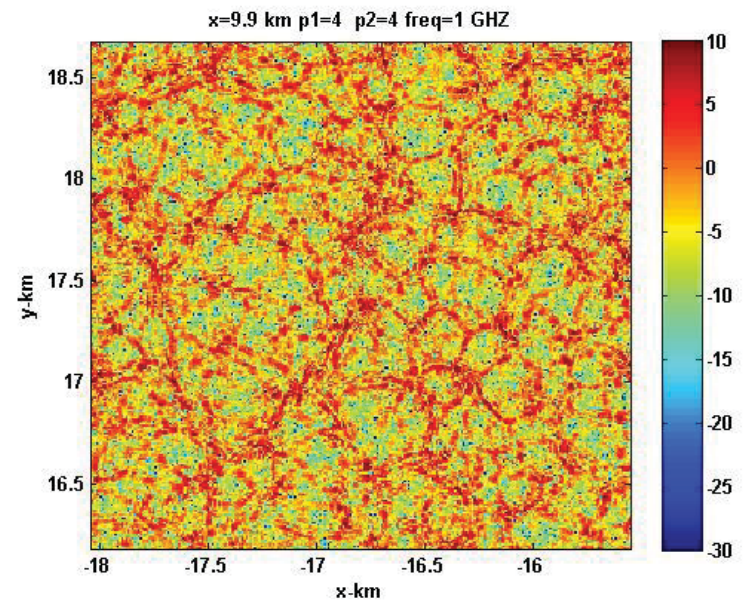
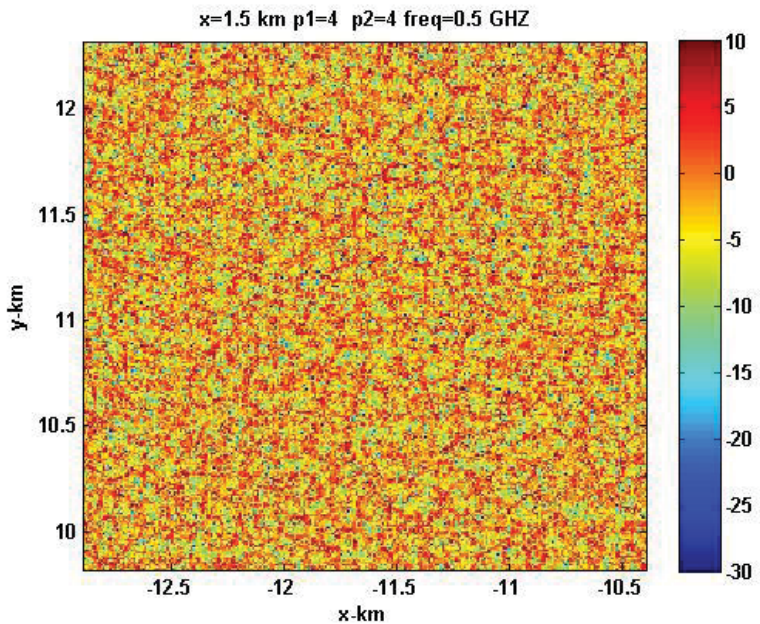
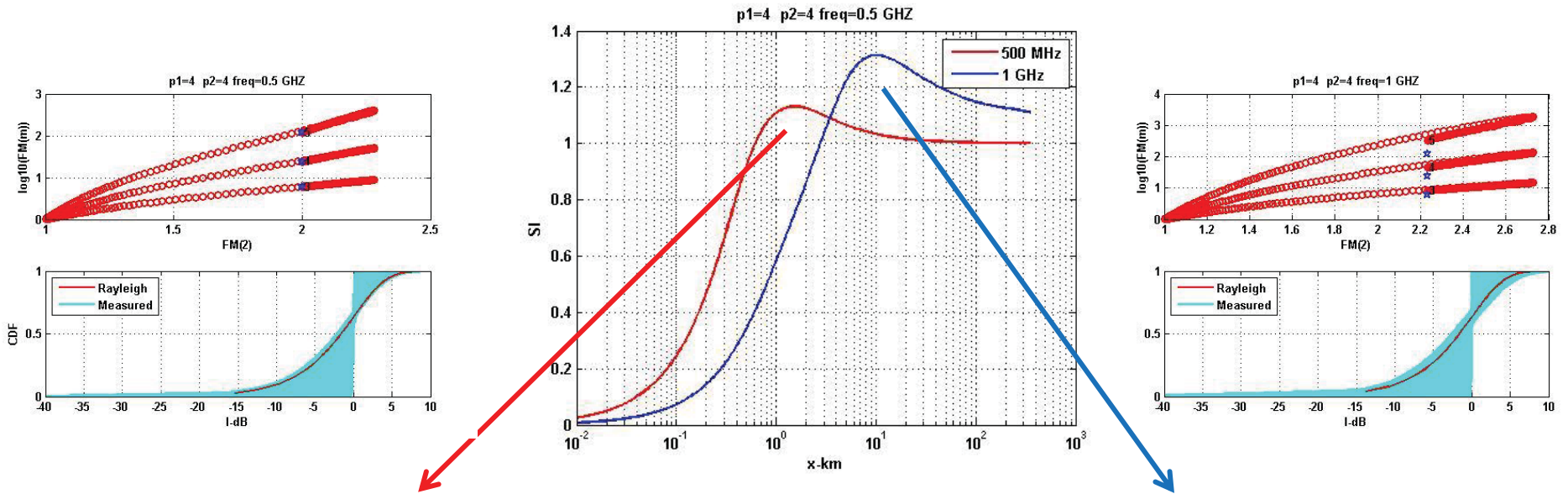
P(3D)=3 Phase Screen Field Evolution



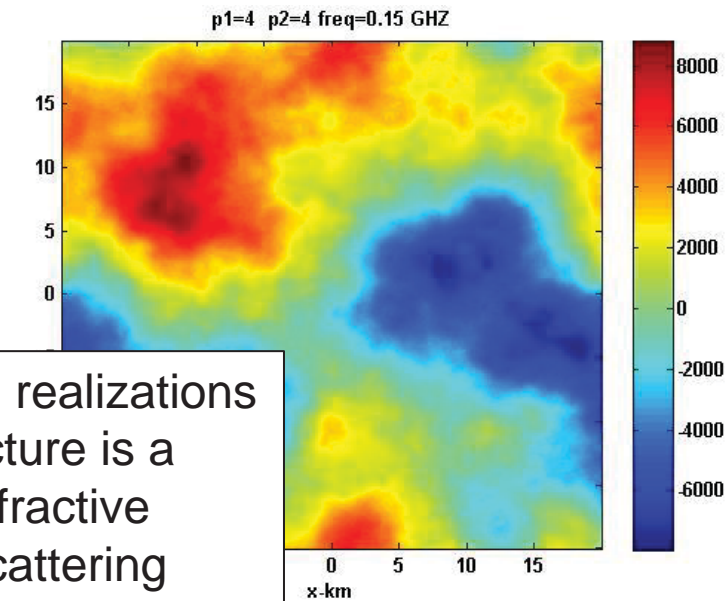
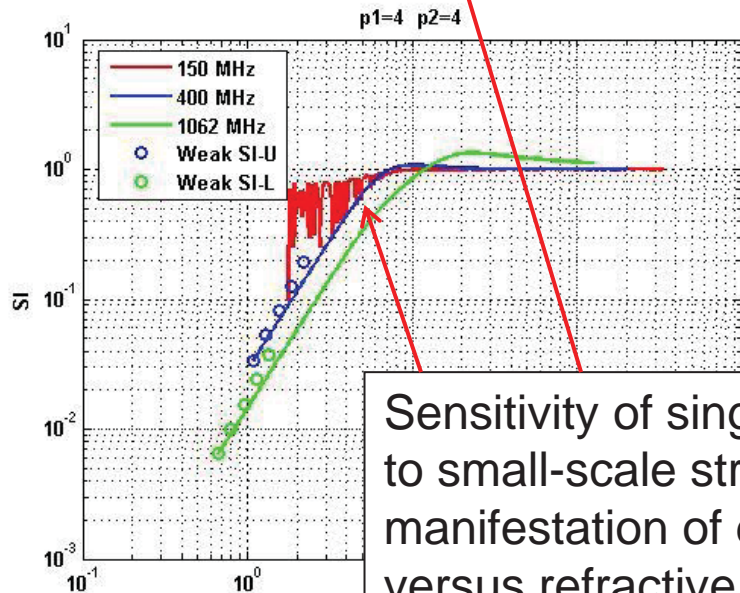
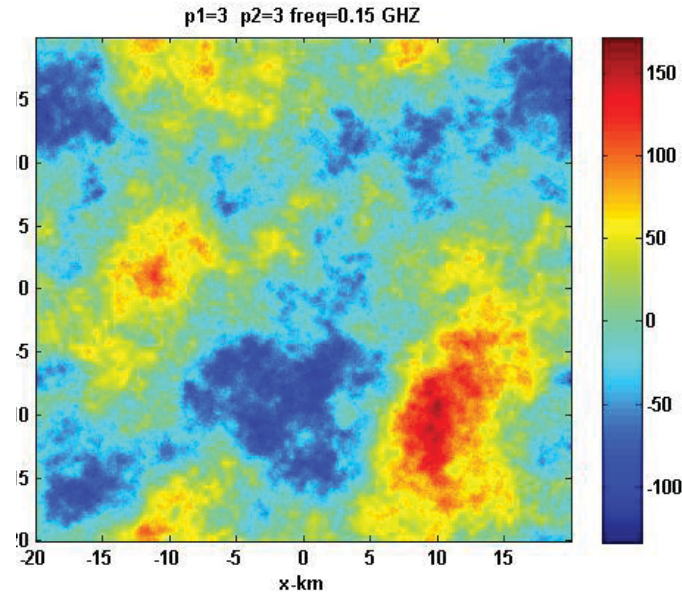
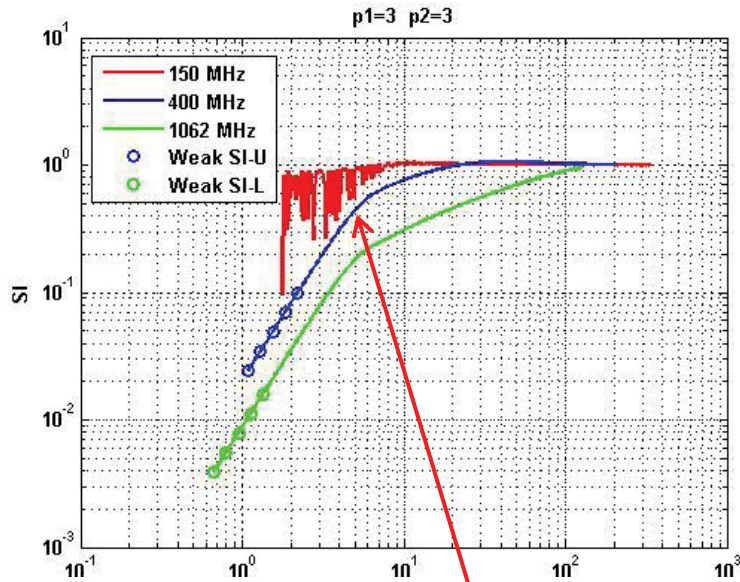
P(3D)=4 Phase Screen Realization Infinite Outer Scale



P(3D)=4 Phase Screen Field Evolution

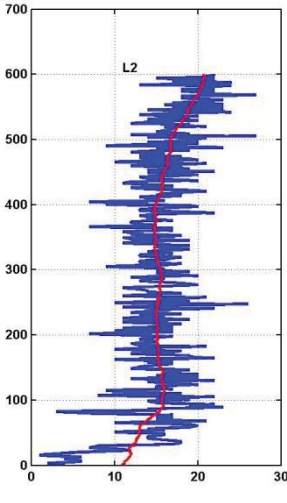
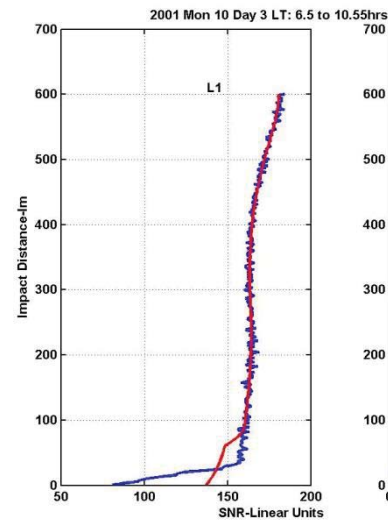
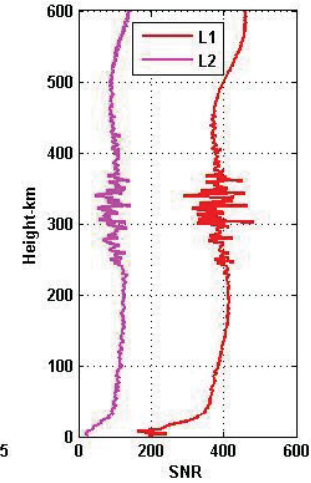
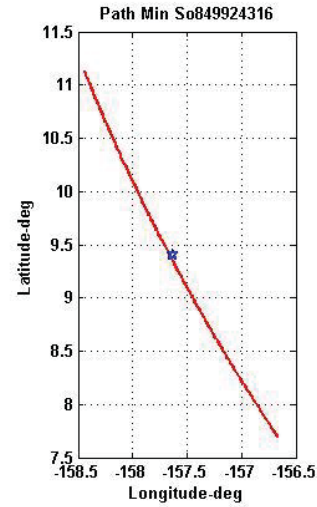
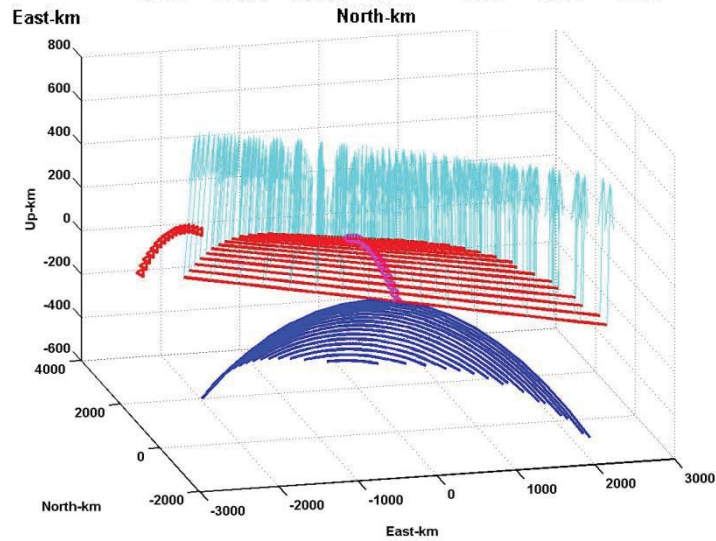
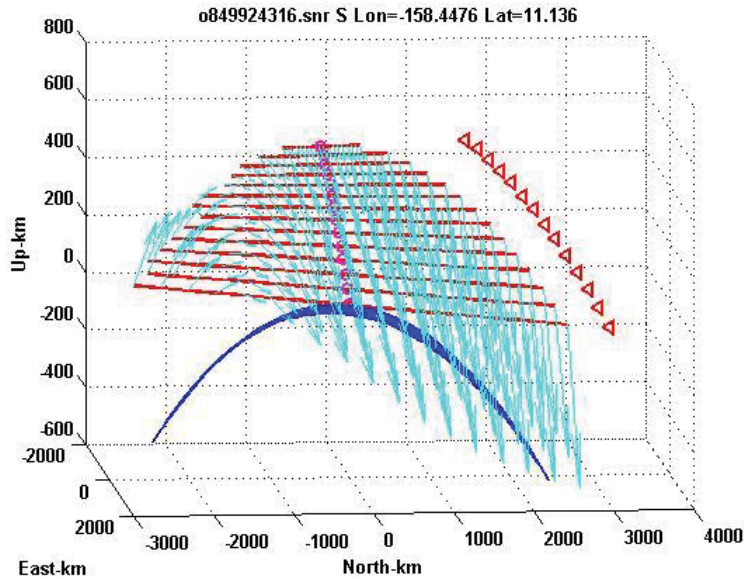


Frequency Dependence



Sensitivity of single realizations to small-scale structure is a manifestation of diffractive versus refractive scattering

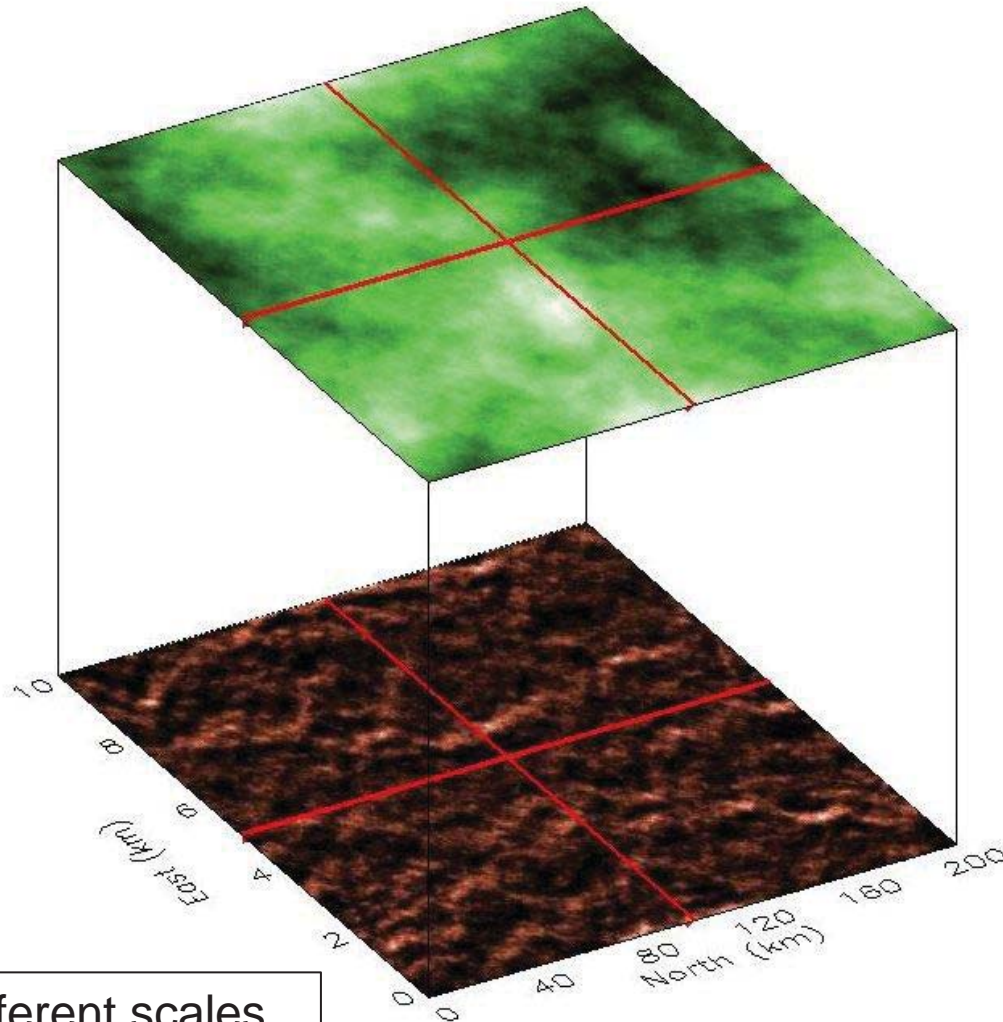
- **3D Propagation simulations using 2D phase-screen realizations illustrate critical dependencies on power-law index, phase turbulent strength, and Fresnel radius**
 - **Ongoing research is pursuing the critical parameter dependencies and extensions to the two-component model**
 - **The phase-screen model can accommodate oblique propagation geometries and anisotropy, but the amount of computation is prohibitive for data interpretation and predictive modeling**
- ***Equivalent* 2D propagation models provide a viable alternative with very encouraging results, which will occupy the remainder of this presentation**



RED => GPS to COSMIC links <800 km
BLUE => Earth surface projection of links
CYAN => Magnetic field direction along

Simplifying the theory for real-world **Applications**

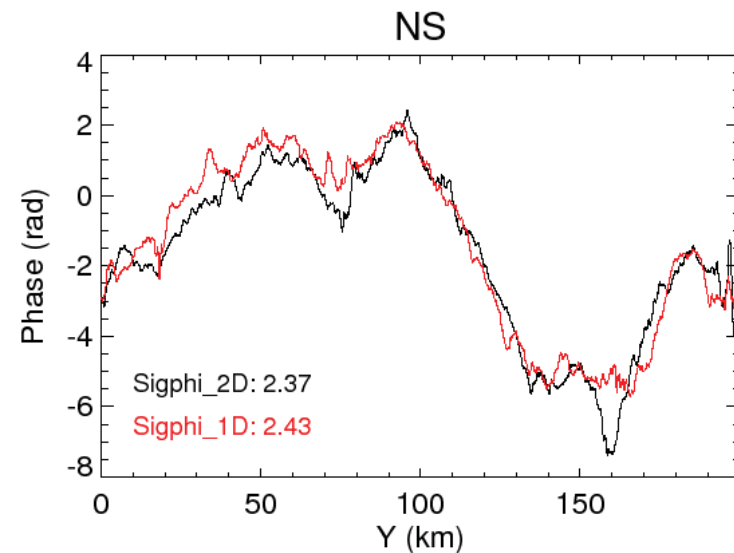
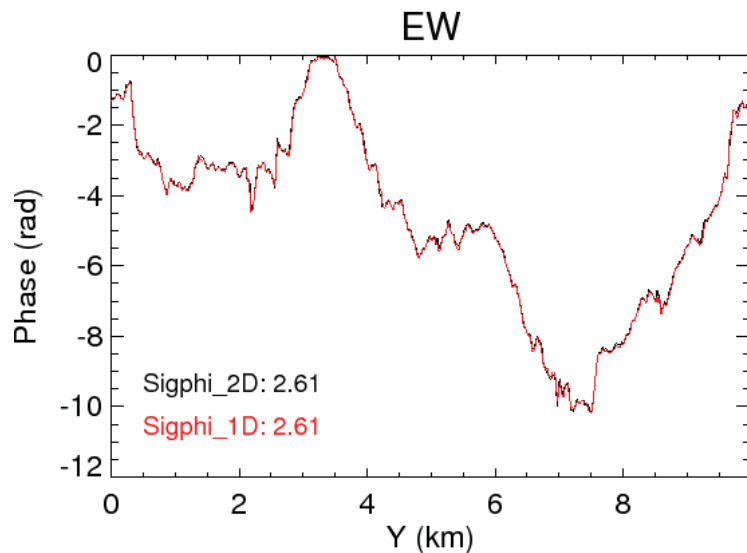
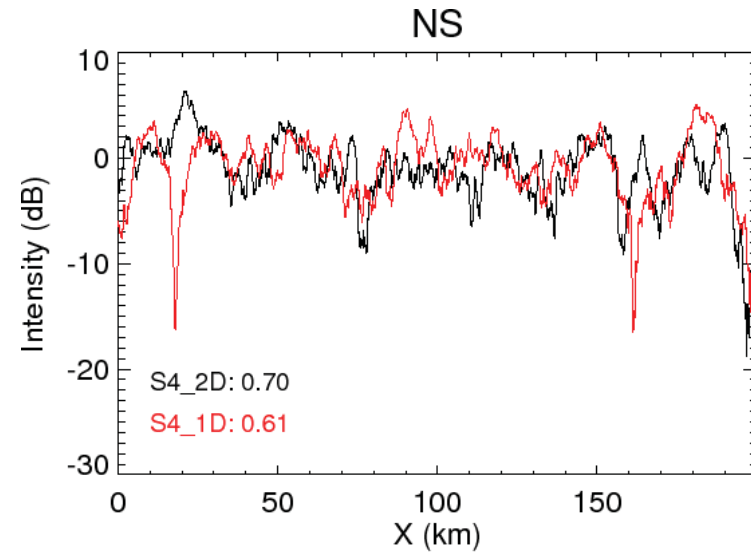
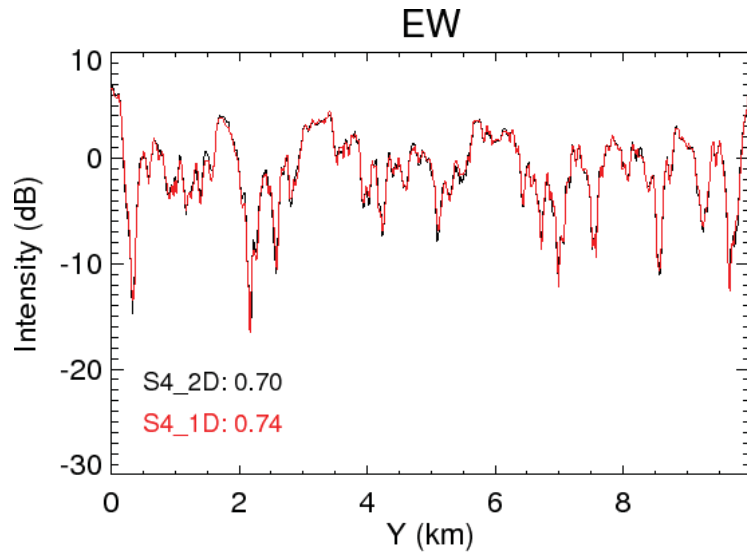
All the material in this section was generously provide by Charles Carrano
A complete list of his publications and preprints can be found on his BC web page:
<https://www2.bc.edu/~carranoc/>



Different scales
accommodate
10:1 elongation

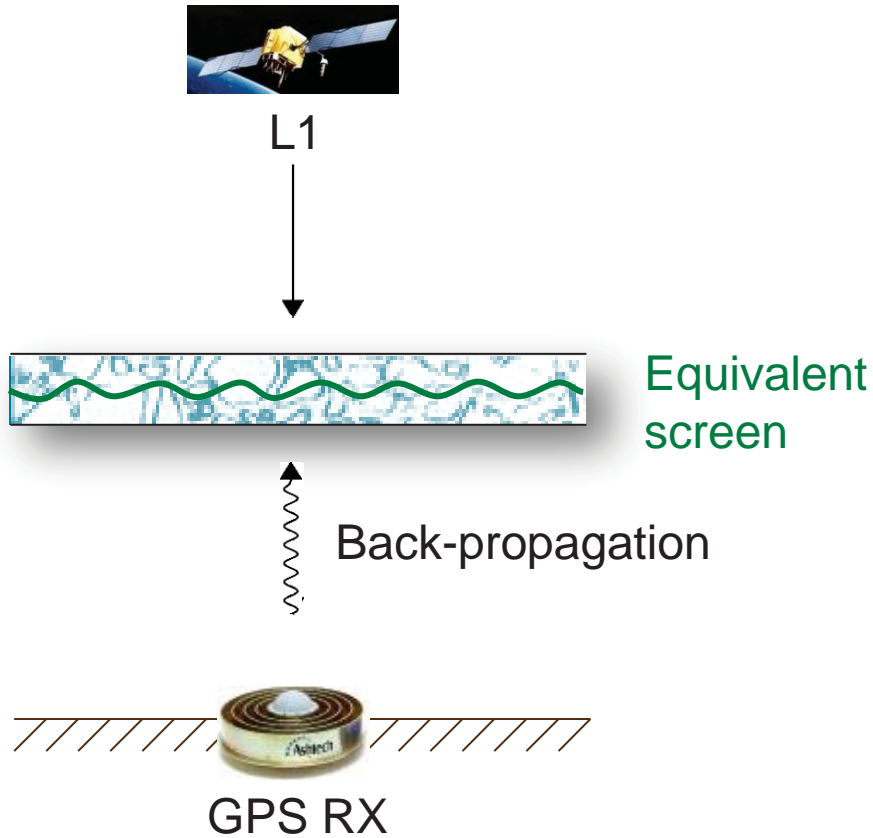
- Path integration at oblique incidence through elongated structure projects the elongation onto the phase screen
- Path motion relative structure presents a 1D scan to an observer
 - Space to time conversion imposes an effective velocity
- **The projection of the scan at the phase screen could be used to initiate an in-plane two-dimensional forward propagation calculation**
- Charles Carrano has investigated and exploited this concept

Comparing Forward-Propagation through 2D and 1D Screens



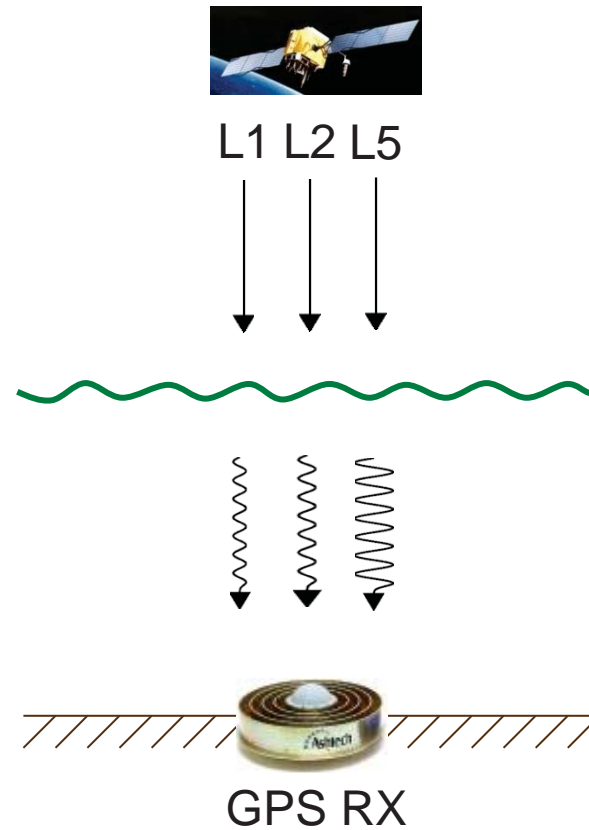
1D propagation results agree with 2D results only when scan is east-west

Field Measurements

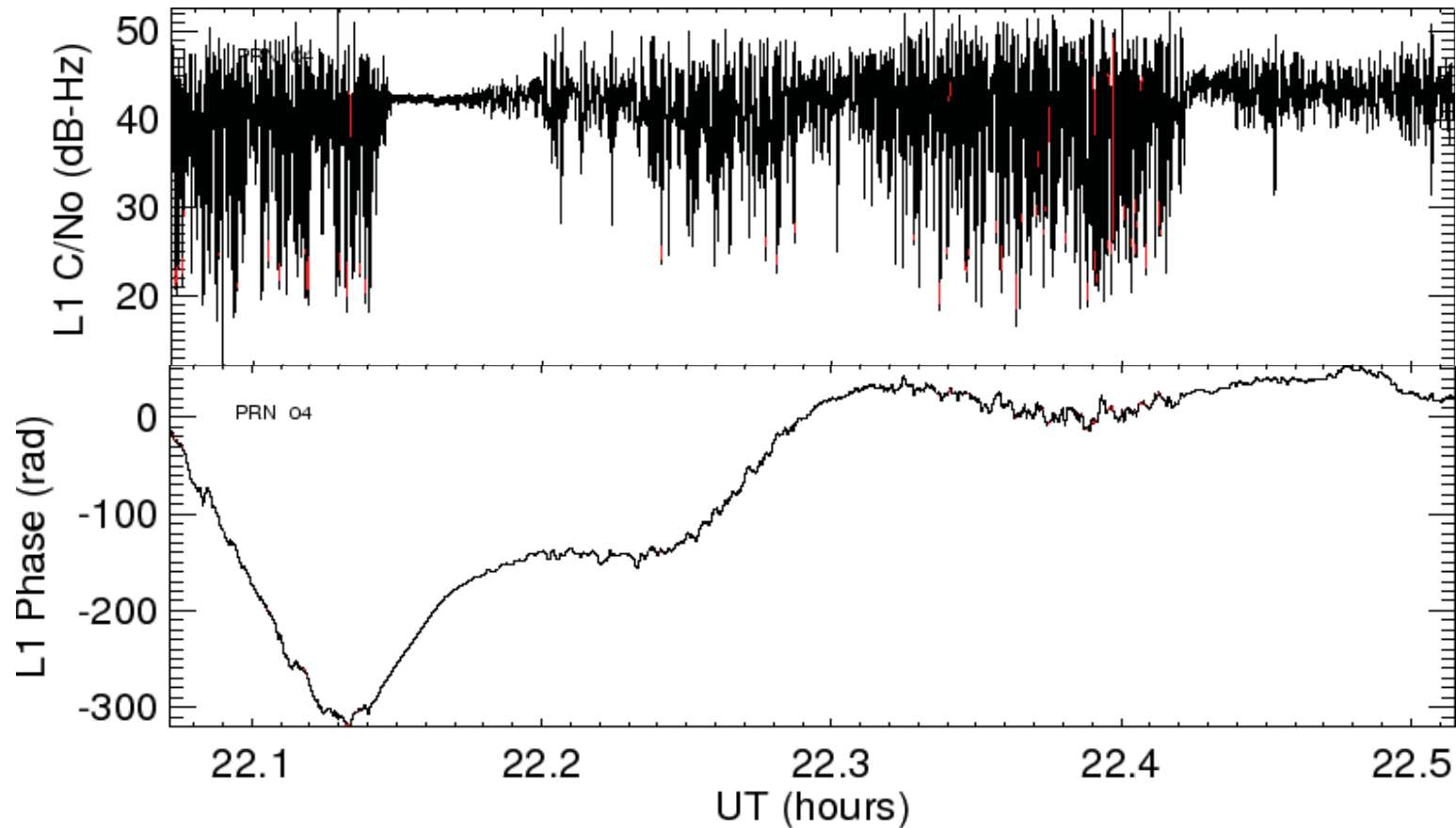


Amplitude and phase
on L1 carrier

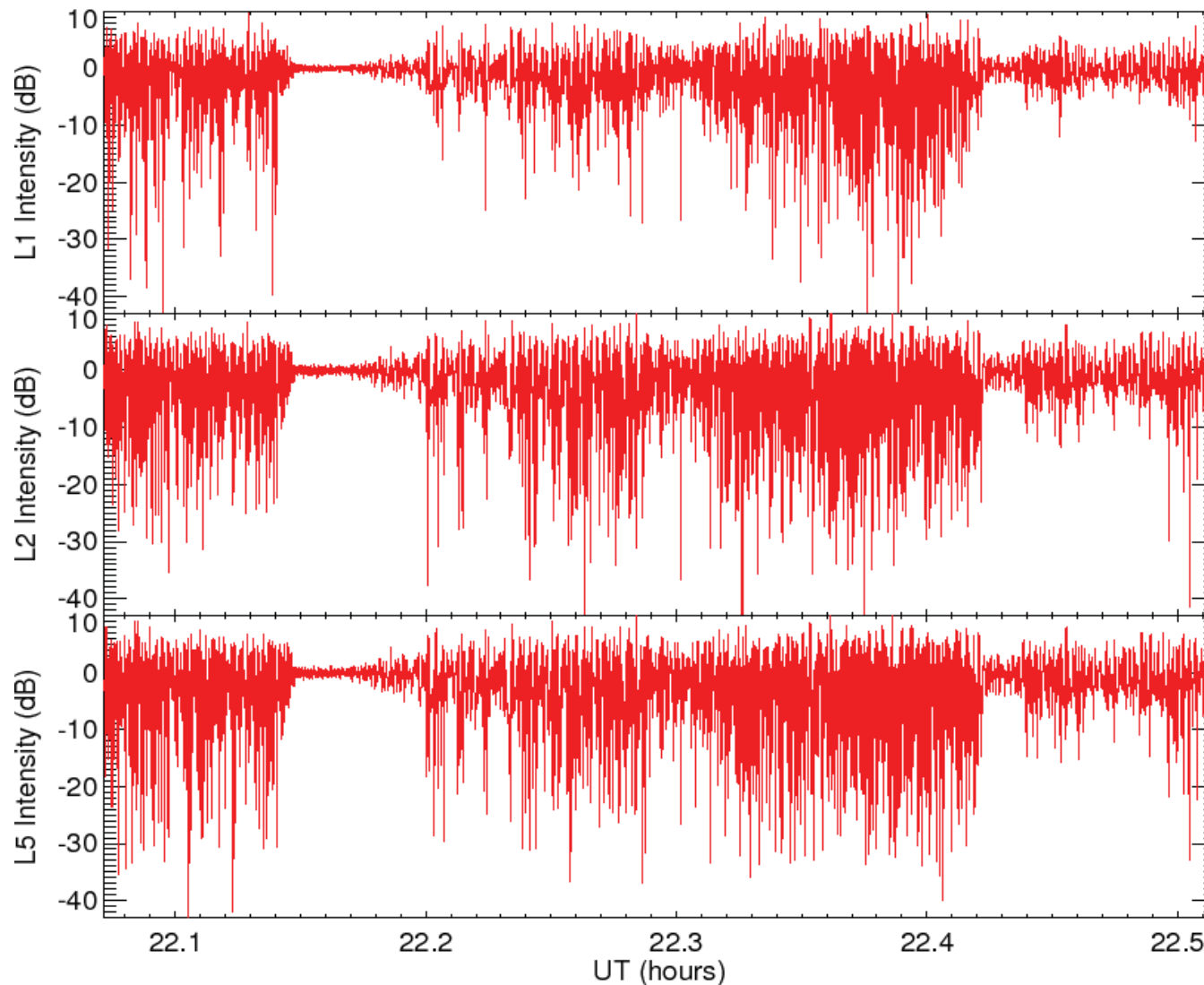
Phase Screen Simulation



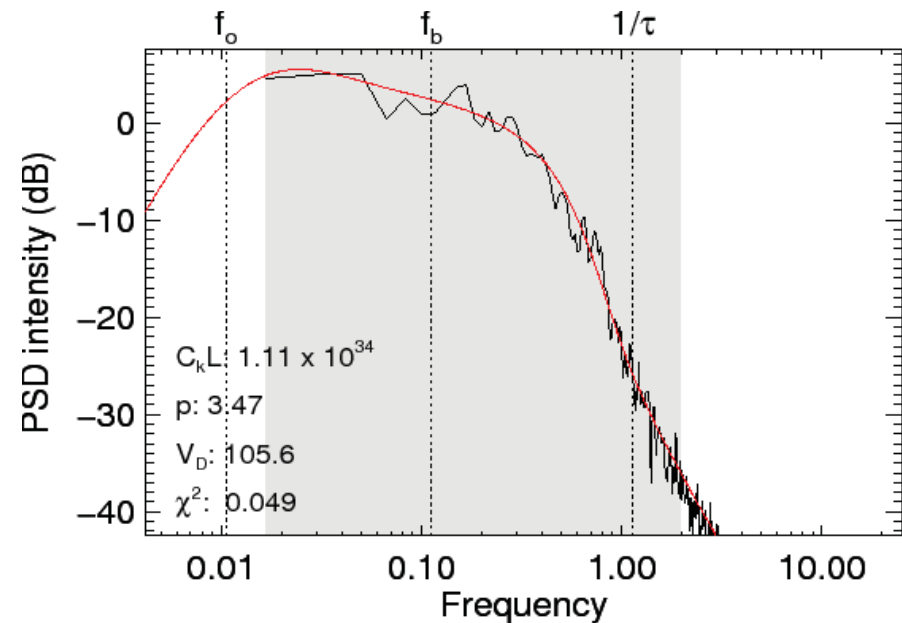
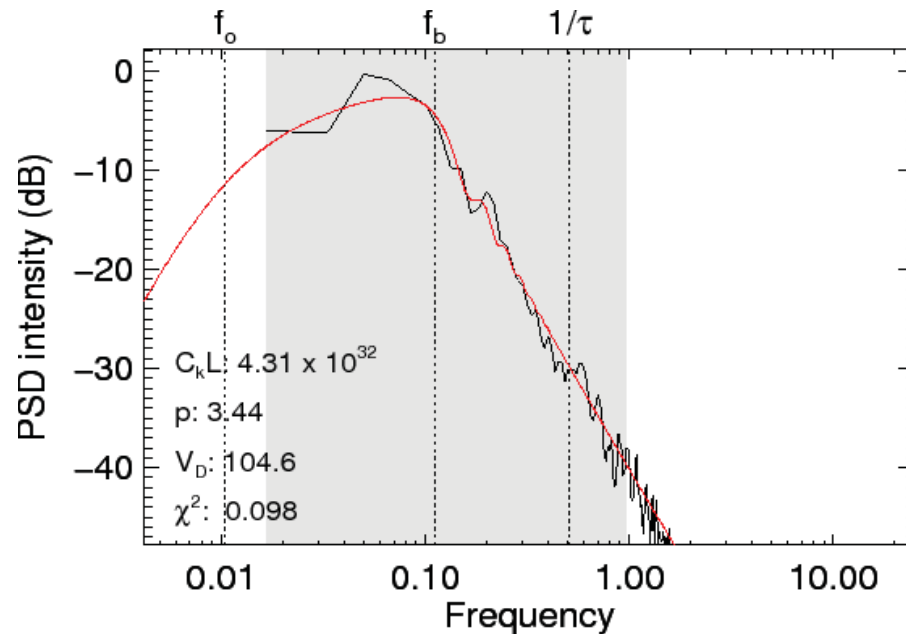
Amplitude and phase
on L1, L2, L5 carriers



Data collected by AFRL on March 13, 2002 (solar max)
 at Ascension Island (7.98°S, 14.4°W, 15°S dip latitude)
 using NovAtel GSV4004 receiver



Intensity correlations between carrier pairs are calculated from these realizations



Theoretical predictions derived from two-dimensional phase-screen theory

Accuracy of recovered parameters:

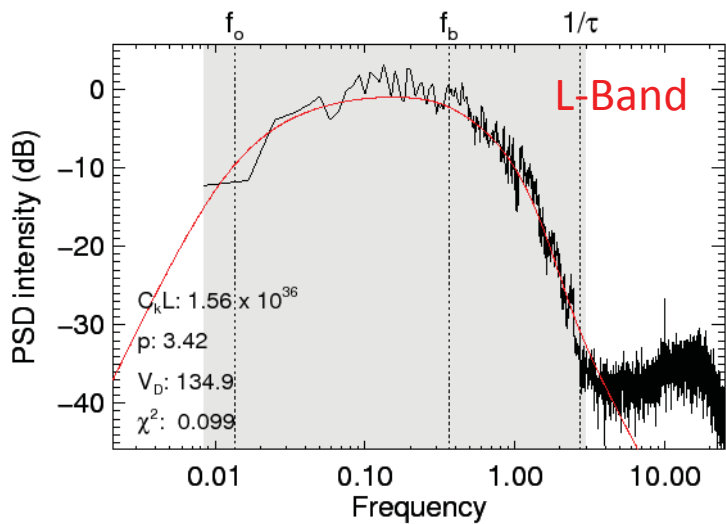
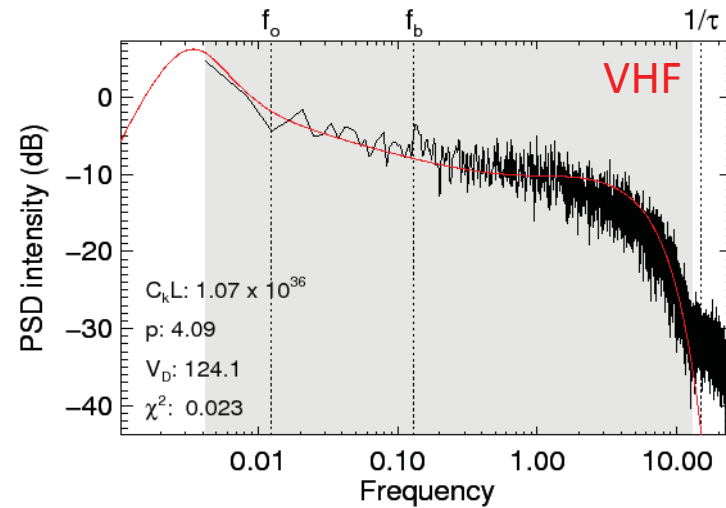
Weak Scatter

$C_k L$ (14%), p (2%), and V_D (5%) accuracy.

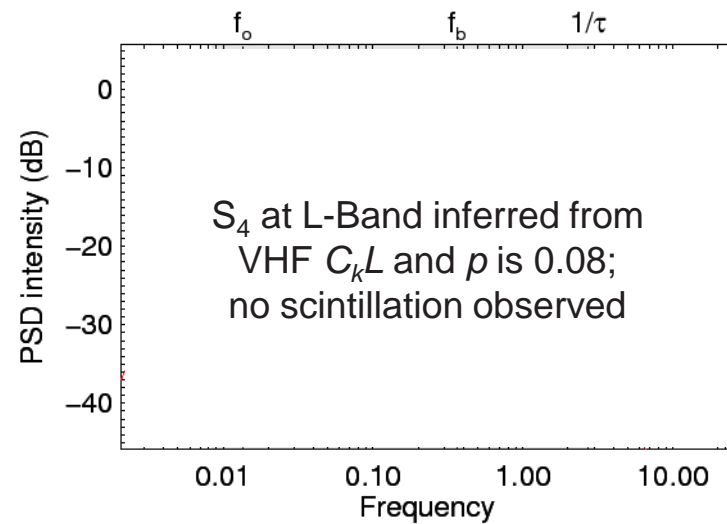
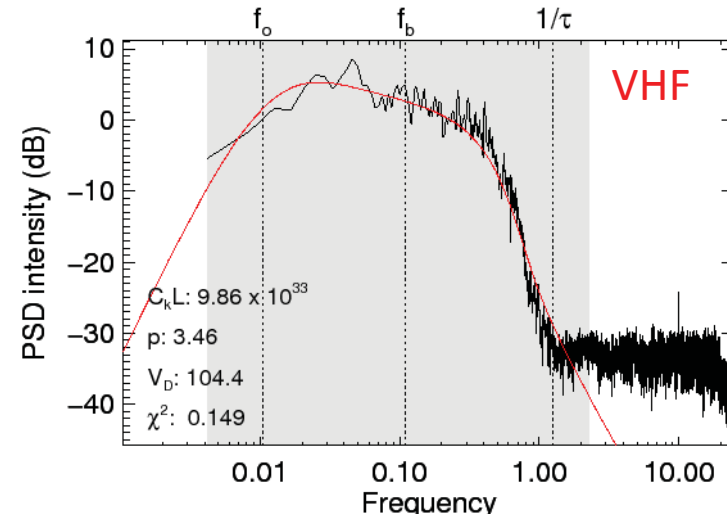
Strong Scatter

$C_k L$ (11%), p (1%), and V_D (6%)

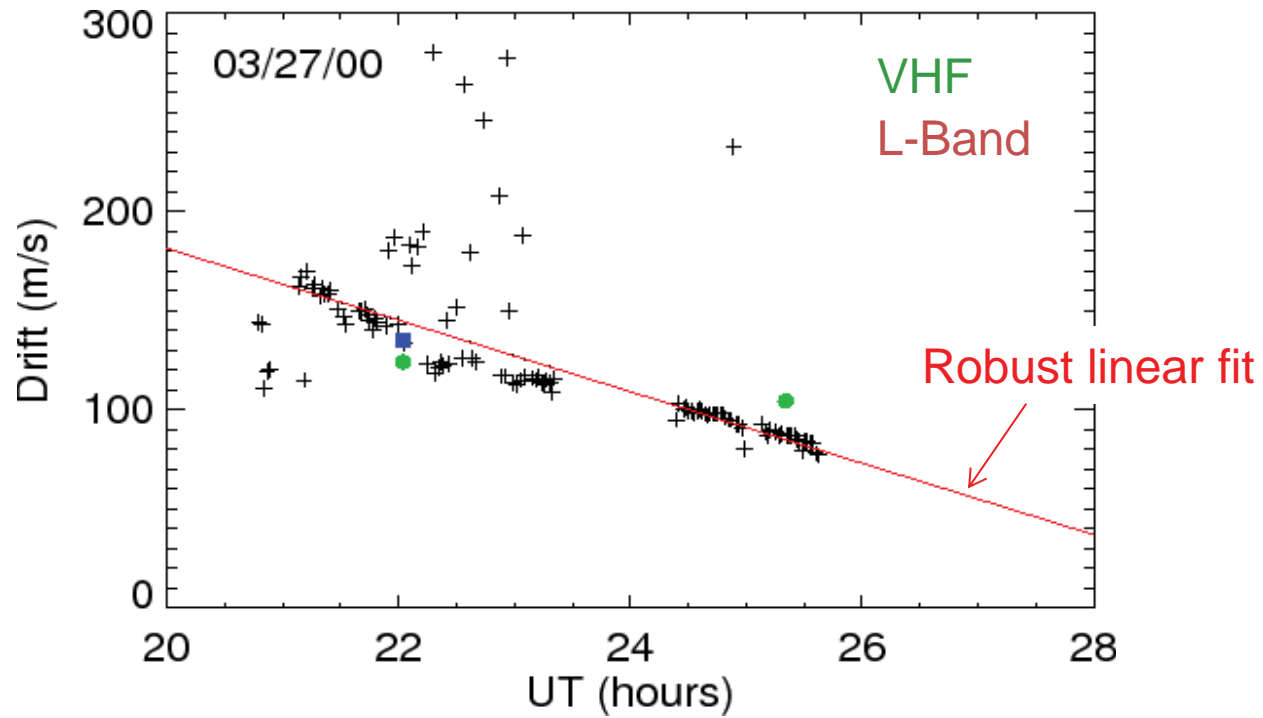
Early Evening (22:02 UT)



Post Midnight (25:20 UT)



Excellent agreement between screen parameters inferred from VHF and L-Band



IPE analysis provides good estimates of the zonal drift using only a single receiver

- **Our objective was to present a consistent structure and scintillation model framework with analysis procedures that can be used to validate and refine models**
- **The strong-scatter theory was emphasized because these conditions are the most stressing for satellite navigation and communication**
 - **With modern computational resources efficient modeling and data analysis procedures are readily realized**
- **The challenge for planned experiments is to make sure data quality supports the refined analysis procedures**
- **Good research topics abound!**

**It has been my pleasure
to present this material
Thank You**

