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## **Preparatory School to the Winter College on Optics: Fundamentals of Photonics – Theory, Devices and**

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### **Introduction to nonlinear optics**

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# Linear Systems

$$S(x) \longrightarrow \boxed{\text{system}} \longrightarrow r(x)$$

$$S(x) \longrightarrow \boxplus \longrightarrow r(x)$$

Linear

$$S_1(x) + S_2(x) \longrightarrow \boxplus \longrightarrow r_1(x) + r_2(x)$$

Shift invariant

$$S(x-x_0) \longrightarrow \boxplus \longrightarrow r(x-x_0)$$

consider LSI system

if  $S(x) \boxplus r(x)$ , then  $S'(x) \boxplus ?$

$$S'(x) = \lim_{\Delta x \rightarrow 0} \frac{S(x+\Delta x) - S(x)}{\Delta x} \boxplus \lim_{\Delta x \rightarrow 0} \frac{r(x+\Delta x) - r(x)}{\Delta x} = r'(x)$$

$$\underline{S'(x) \boxplus r'(x)}$$

$$\underbrace{S'(x) - i\nu 2\pi S(x)}_{\text{if this } = 0} \boxplus \underbrace{r'(x) - i 2\pi \nu r(x)}_{\text{then this } = 0}$$

$$S(x) = S_0 e^{i 2\pi \nu x} \boxplus r(x) = r_0 e^{i 2\pi \nu x}$$

$$e^{i 2\pi \nu x} \boxplus H(\nu) e^{i 2\pi \nu x}$$

In general

$$s(x) = \int_{-\infty}^{\infty} \tilde{s}(v) e^{i2\pi v x} dv \rightarrow \int_{-\infty}^{\infty} \underbrace{\tilde{s}(v) H(v)}_{\tilde{r}(v)} e^{i2\pi v x} dv = r(x)$$

$$\underline{\tilde{r}(v) = \tilde{s}(v) H(v)} \quad H(v) = \text{transfer function.}$$

$$\begin{array}{ccc}
 s(x) & \xrightarrow{\text{box}} & r(x) \\
 \downarrow \hat{\mathcal{F}} & & \uparrow \hat{\mathcal{F}}^{-1} \\
 \tilde{s}(v) & \rightarrow & \tilde{r}(v) = \tilde{s}(v) H(v)
 \end{array}$$

On the other hand:


$$s(x) = \int s(x') \delta(x-x') dx' \xrightarrow{\text{box}} r(x) = \int s(x') G(x-x') dx = s * G(x)$$

where  $\delta(x) \xrightarrow{\text{box}} G(x)$  impulse response

so the effect of a LSI system is always a convolution, which in Fourier space is a product. Therefore  $\hat{\mathcal{F}}_{x \rightarrow v} G(x) = H(v)$

The system acts as a blur, or equivalently, as a Fourier filter.

Examples: { Low pass  
High pass  
DC removal  
derivatives (including "fractional")  
combination of images

Noise  $n(x)$   random.

Fourier transform  $\hat{\int}_{x \rightarrow \nu} n(x) = \tilde{n}(\nu)$

often, we do not know  $\tilde{n}$ , but we do know roughly how it behaves

$$\langle |\tilde{n}(\nu)|^2 \rangle = S_n(\nu) \text{ power spectrum of the noise.}$$

average over many "realizations"

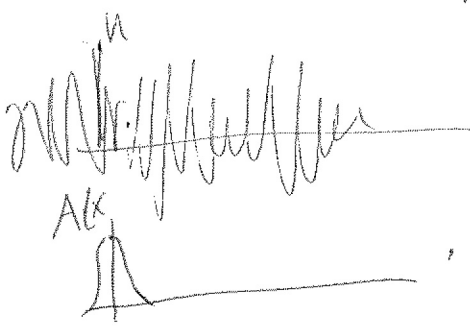
<sup>note</sup>

$$\hat{\int}_{\nu \rightarrow x} S_n(\nu) = \int S_n(\nu) e^{i2\pi\nu x} d\nu = \left\langle \int \tilde{n}^*(\nu) \tilde{n}(\nu) e^{i2\pi\nu x} d\nu \right\rangle$$

$$\left[ \int n(x') e^{-i2\pi\nu x'} dx' \right]^*$$

$$= \left\langle \int \int n^*(x') \tilde{n}(\nu) e^{i2\pi\nu(x+x')} d\nu dx' \right\rangle$$

$$= \left\langle \int n^*(x') \underbrace{\int \tilde{n}(\nu) e^{i2\pi\nu(x+x')} d\nu}_{n(x+x')} dx' \right\rangle = \left\langle \underbrace{\int n^*(x') n(x'+x) dx'}_{\text{autocorrelation of noise } A_n(x)} \right\rangle$$



$$\hat{\int}_{x \rightarrow \nu} A_n(x) = S_n(\nu)$$

Wiener-Khinchin Theorem.

Examples if  $A_n(x) \approx \delta(x)$ ,  $S_n(\nu) \approx \text{constant}$   
white noise



$$\mathcal{E} = \int |WH-1|^2 |\tilde{S}|^2 d\nu + \int |W|^2 |\tilde{n}|^2 d\nu + \int [W^* \tilde{n}^* (WH-1) \tilde{S} + \text{c.c.}] d\nu$$

Take average over possible realizations of  $\tilde{n}$ , and use  $\langle \tilde{n} \rangle = 0$ ,  $\langle |\tilde{n}|^2 \rangle = S_n(\nu)$  (noise power spectrum).

$$\langle \mathcal{E} \rangle = \int (W^* H^* - 1)(WH-1) |\tilde{S}|^2 d\nu + \int W^* W S_n(\nu) d\nu + 0$$

Now find  $W(\nu)$  that minimizes error. Use variational calculus:

$$\frac{\delta}{\delta W^*(\nu)} \int W^* (WFA) d\nu = W^* F^* A$$

$$\text{Force } \frac{\delta \langle \mathcal{E} \rangle}{\delta W^*(\nu)} = 0$$

$$H^* (WH-1) |\tilde{S}|^2 + W S_n = 0$$

$$\text{So } W [ |H|^2 |\tilde{S}|^2 + S_n ] = H^* |\tilde{S}|^2$$

$$W = \frac{H^*}{|H|^2 + S_n / |\tilde{S}|^2}$$

Wiener filter