

https://classic.sdss.org/dr7/coverage/index.html



https://arxiv.org/abs/1912.03302

Towards projections for the impact of LSST's detection of new Milky Way satellite dwarf galaxies on the indirect detection of dark matter

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https://www.lsst.org/scientists/understanding-simulations-and-data

Outline

- Why Dwarf Galaxies for the indirect detection search for dark matter?
- Dark Matter only N-Body Simulations used
- "Calibrating" the simulations
 - Sloan Digital Sky Survey (SDSS)
 - Dark Energy Survey (DES)
 - Panoramic Survey Telescope and Rapid Response System (Pan-STARRS1 or PS1)
- Conclusions and Future Outlook

Why Dwarf Galaxies?

- Large amount of dark matter for their size (high mass to light ratio)
- Low Astrophysical Backgrounds
- Many are outside of galactic plane (again reducing backgrounds)





Dark Matter only N-Body Simulations

- We follow the work of Nadler et al (arXiv:1912.03303) and use the L-GADGET, N-Body, cosmological dark matter only simulation (<u>https://wwwmpa.mpa-garching.mpg.de/gadget/</u>)
- 46 Milky-Way like halos were identified in the larger simulation, and then "zoom in" simulations were performed on these
- 6 of these were identified as having the most similar formation histories to the Milky Way
 - These are the halos used for the results in this talk



Image: https://wwwmpa.mpa-garching.mpg.de/galform/millennium/seqD_063a_half.jpg

Calibration: SDSS

- The Sloan Digital Sky Survey (SDSS) is a 2.5-m wide-angle optical telescope at Apache Point Observatory in New Mexico, United States
- Data collection began in 2000 and continues today



https://sloan.org/programs/research/sloan-digital-sky-survey

https://classic.sdss.org/dr7/coverage/index.html

SDSS+Classical Dwarfs: Visual Magnitude



Brighter

SDSS+Classical Dwarfs: J-Factor



Calibration: DES

- The Dark Energy Camera (DECam) was mounted on the the 4-meter Víctor M. Blanco Telescope, located at the Cerro Tololo Inter-American Observatory (CTIO) in the Chilean Andes
- The camera operated from 2013-2019.





https://www.darkenergysurvey.org/multimedia/photo-gallery/

DES: J-Factor



Calibration: Pan-STARRS1 (PS1)

- Pan-STARRS1 (PS1) is a 1.8-meter diameter telescope on the Hawaiian Island of Maui.
- One of the world's largest digital cameras, with almost 1.4 Gigapixels.



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http://www2.ifa.hawaii.edu/research/Pan-STARRS.shtml

PS1: J-Factor



10,000 realizations of random halos from 6 used in 2018 paper, subhalos random

Next Step: LSST

- The Legacy Survey of Space and Time (LSST) is under construction at the Vera Rubin Observatory in the Chilean Andes
- LSST will have 3.2 Gigapixel camera and is expected to see first light in 2024 and run for 10 years



https://gallery.lsst.org/bp/#/folder/2358042/



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Conclusions and Future Outlook

- We have used a *N*-body, dark matter only simulations of Milky Way-like halos to generate a Monte Carlo set of dwarf galaxies that seem to be representative of the actual properties of the true distribution of dwarf galaxies of the Milky Way
- We will use this set of Monte Carlo models to forecast the properties of new dwarf galaxies likely to be discovered by LSST
- We will determine the impact these new dwarf galaxies will have on the indirect detection search for dark matter