

# The type-II seesaw mechanism in alternative gauged $U(1)_X$ and dark sector

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# § Introduction

# Nonvanishing neutrino mass

- Mass of right-handed neutrinos  $\nu_R$

$$\mathcal{L} = \frac{1}{2} \overline{\nu_R^c} M_{\nu_R} \nu_R + \text{H. c.}$$

*B-L violating*

- Majorana mass  $M_{\nu_R}$  might come from the breaking of the gauged  $U(1)_{B-L}$  symmetry.  
[Davidson (1979), Mohapatra and Marshak (1980), ...]

# Anomaly cancellation for gauged B-L

- Once the charges are set to be anomaly free under  $U(1)_Y$  and  $U(1)_{B-L}$ , it is so under a linear combination  $X = x_H Y + (B - L)$  too.
  - Convenient parametrization
  - $x_H = 0, U(1)_{B-L}$
  - $x_H = -2, U(1)_R$  [Jung et al (2010)]
- Two choices of RH neutrino charges

Field	standard	alternative
$\nu_R^1, \nu_R^2, \nu_R^3$	-1,-1,-1	-4,-4,+5

# Anomaly cancellation for gauged B-L

- Two choices

Field	standard	alternative
$\nu_R^1, \nu_R^2, \nu_R^3$	-1,-1,-1	-4,-4,+5

- The standard is standard.
- For the alternative, Type-I seesaw and Majorana mass is not easy [Ma and Srivastava, Sanchez-Vega and Schmitz (2015), ..., Asai, Nakayama and Tseng (2021)]
- How about Type-II?
- Type-II by triplet Higgs  $\Delta$  [Schechter and Valle, Magg and Wetterich, Cheng and Li (1980)]

$$\mathcal{L} \supset \frac{1}{2} \bar{L}^c \Delta L + \text{H. c.}$$

## § Model

- Alternative charge of  $U(1)_X$
- Type-II

# § § Particle content

- The anomaly cancellation fixes fermions
- Triplet  $\Delta$  for neutrino mass
- $\Phi_1$  for the tadpole of  $\Delta$ 
  - No Yukawa
- $\Phi_X$ 
  - $U(1)$  breaking
  - Mass of  $\nu_R$
  - Mass of NG (CP odd Higgs)
- $\nu_R^i \nu_R^3$  : stable, dark matter

	$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	$U(1)_X$
$q_L^i$	<b>3</b>	<b>2</b>	$\frac{1}{6}$	$\frac{1}{6}x_H + \frac{1}{3}$
$u_R^i$	<b>3</b>	<b>1</b>	$\frac{2}{3}$	$\frac{2}{3}x_H + \frac{1}{3}$
$d_R^i$	<b>3</b>	<b>1</b>	$-\frac{1}{3}$	$-\frac{1}{3}x_H + \frac{1}{3}$
$l_L^i$	<b>1</b>	<b>2</b>	$-\frac{1}{2}$	$-\frac{1}{2}x_H - 1$
$e_R^i$	<b>1</b>	<b>1</b>	$-1$	$-x_H - 1$
$\nu_R^1$	<b>1</b>	<b>1</b>	$0$	$-4$
$\nu_R^2$	<b>1</b>	<b>1</b>	$0$	$-4$
$\nu_R^3$	<b>1</b>	<b>1</b>	$0$	$5$
$\Phi_1$	<b>1</b>	<b>2</b>	$\frac{1}{2}$	$\frac{1}{2}x_H + 1$
$\Phi_2$	<b>1</b>	<b>2</b>	$\frac{1}{2}$	$\frac{1}{2}x_H$
$\Delta_3$	<b>1</b>	<b>3</b>	$1$	$x_H + 2$
$\Phi_X$	<b>1</b>	<b>1</b>	$0$	$1$

# § § Interactions and masses

- $\mathcal{L} \supset -\frac{1}{\sqrt{2}} Y_{\Delta}^{ij} \overline{l_L^i}^C \cdot \Delta l_L^j - \sum_{i=1,2} Y_{\nu_R^i} \Phi_X^\dagger \overline{\nu_R^3}^C \nu_R^i + \text{H. c.}$
- LH Neutrino mass  
$$(m_\nu)_{ij} = Y_{\Delta}^{ij} \nu_\Delta$$

# § § Interactions and masses

- RH Neutrino mass

$$\mathcal{L} \supset -(\nu_R^{1\dagger}, \nu_R^{2\dagger}, \nu_R^{3T}) \begin{pmatrix} 0 & 0 & Y_{\nu_R^1} \\ 0 & 0 & Y_{\nu_R^2} \\ Y_{\nu_R^1} & Y_{\nu_R^2} & 0 \end{pmatrix} \frac{\nu_X}{\sqrt{2}} \begin{pmatrix} \nu_R^1 \\ \nu_R^2 \\ \nu_R^{3*} \end{pmatrix}$$

- One massless  $\nu_R$  : **Dark radiation**
- Two compose one Dirac  $\chi$  : **Dark matter candidate**

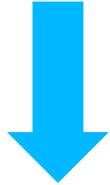
$$m_\chi = \sqrt{\frac{Y_{\nu_R^1}^2 + Y_{\nu_R^2}^2}{2}} \nu_X$$

# § § Interactions and masses

- One singlet, two doublet, one triplet

$$\Phi_1^\dagger \Phi_2, (\Phi_1^\dagger \Phi_2)^2 \notin V$$

- $V \supset \lambda_{12} \Phi_X (\Phi_1^\dagger \Phi_2) - \frac{\Lambda_6}{\sqrt{2}} (\Phi_1^T \cdot \Delta_3 \Phi_1) + \text{H. c.}$



pseudo-scalar mass



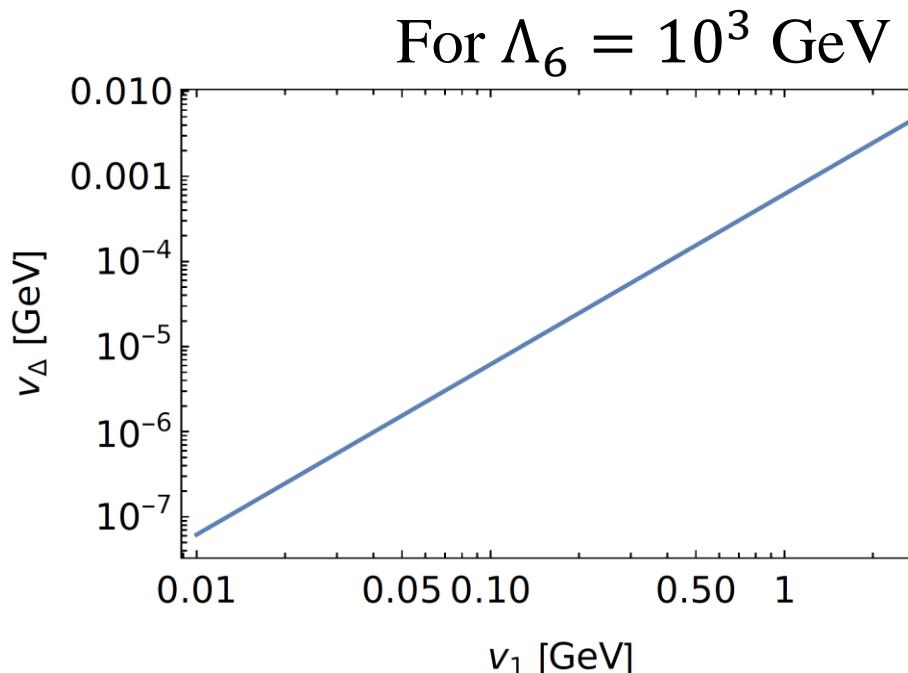
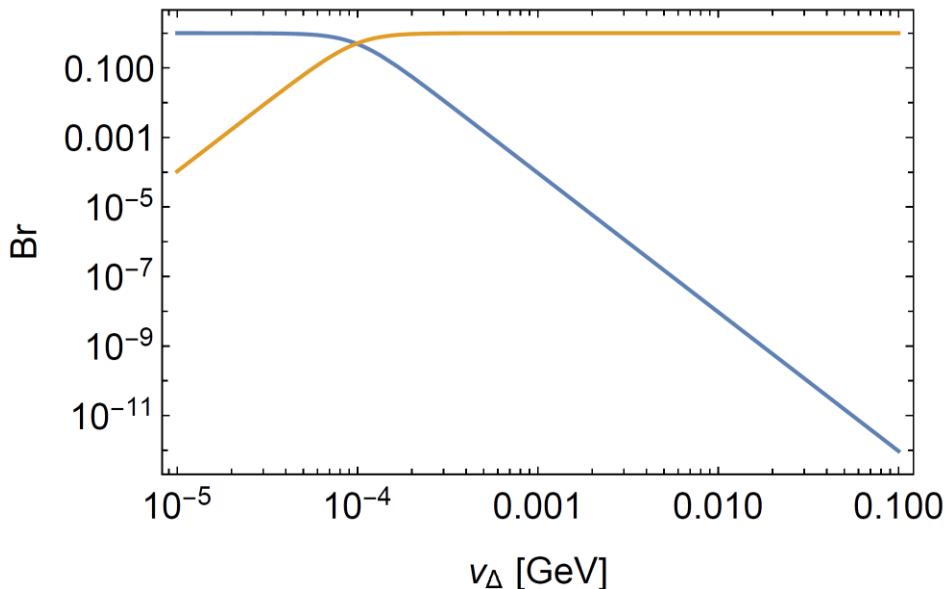
$$v_\Delta \cong \frac{v_1^2 \Lambda_6}{2 \mu_3^2} \propto v_1^2$$

Small neutrino mass due to the small Higgs VEV

Philosophy of neutrophilic Higgs model

# § § Implication

- $H^{\pm\pm}$  decay
  - $H^{\pm\pm} \rightarrow W^\pm W^\pm$
  - $H^{\pm\pm} \rightarrow l^\pm l^\pm$
- Very small  $\nu_\Delta$  from slightly small  $\nu_1$



For  $m_{H^{\pm\pm}} = 900$  GeV

# § § Experimental constraints

- $\rho$  parameter
  - $v_\Delta \lesssim 1$  GeV
- The Z boson decay width
- $H^{\pm\pm}$  search at the LHC
  - $\gtrsim 880$  GeV for  $ll$  [ATLAS (2018)]
  - $\gtrsim 350$  GeV for  $WW$  [ATLAS (2021)]
- LFV [Chun et al, Kakizaki et al (2003), Akeroyd et al (2009)]
  - SINDRUM for  $\mu \rightarrow eee$
  - MEG for  $\mu \rightarrow e\gamma$
  - $v_\Delta \gtrsim 1$  eV



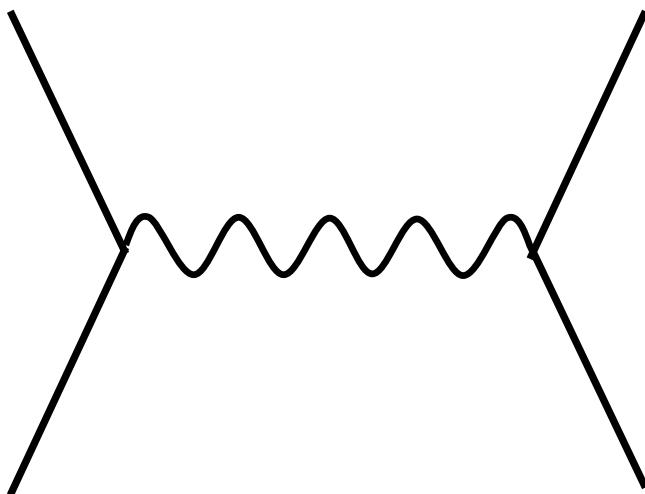
# § § Experimental constraints

- **X boson search at the LHC**
  - $pp \rightarrow X \rightarrow l\bar{l}$ ,  $pp \rightarrow X \rightarrow jj$
  - The most severe bound from dilepton  $l\bar{l}$  ( $l = e, \mu$ )
- Under the narrow width approximation, we calculate
  - $\sigma(pp \rightarrow X)\text{Br}(X \rightarrow l\bar{l})$
  - $\text{Br}(X \rightarrow l\bar{l}) = \frac{8 + 12x_H + 5x_H^2}{F(x_H)}$
  - $F(x_H) = 13 + 16x_H + 10x_H^2$



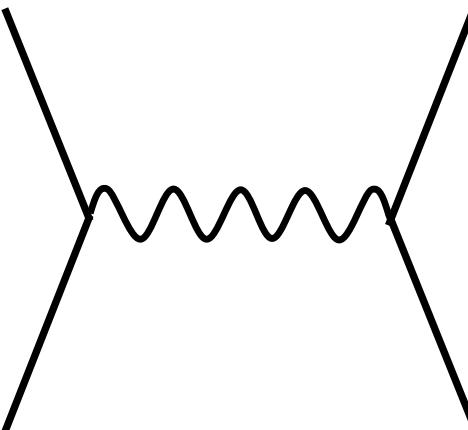
# § § Cosmological constraints

- Dark radiation
  - One massless: behaves as one more neutrino
  - Interaction  $f\bar{f} \leftrightarrow \nu_R \bar{\nu}_R$  through  $U(1)_X$   
[Heeck (2014), Fileviez Perez, Murgui and Plascencia (2019)]



# § § Cosmological constraints

- Dark matter
  - Stable due to charge mismatch, not a parity
  - $U(1)_X$  interacting Dirac fermion
  - Abundance :  $Z'$  resonance annihilation
  - Direct search bounds : SI from the vector interaction via  $Z'$



# § § Results

- LHC

$N_{\text{eff}}$

- 3.5 solid

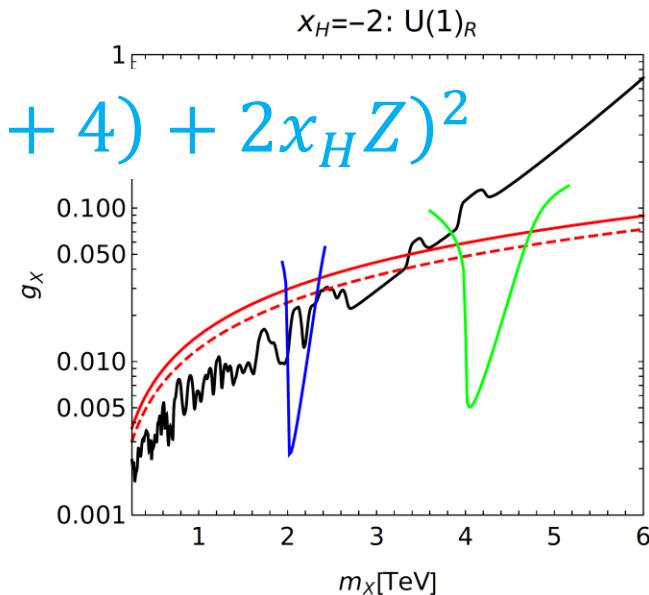
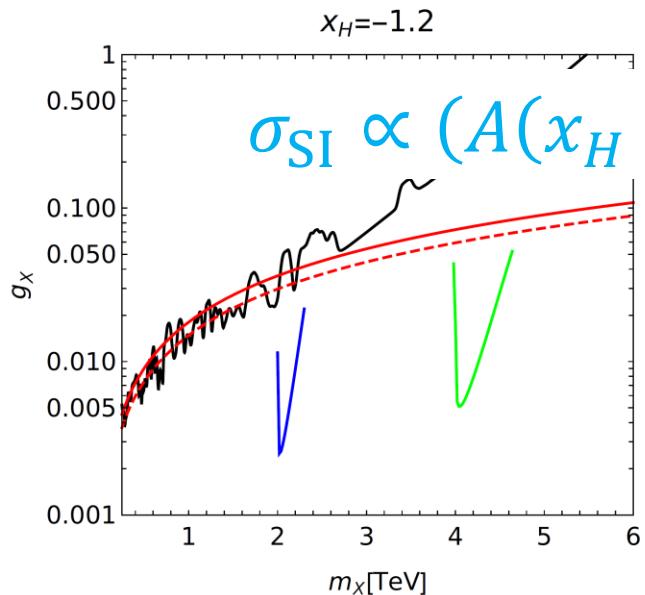
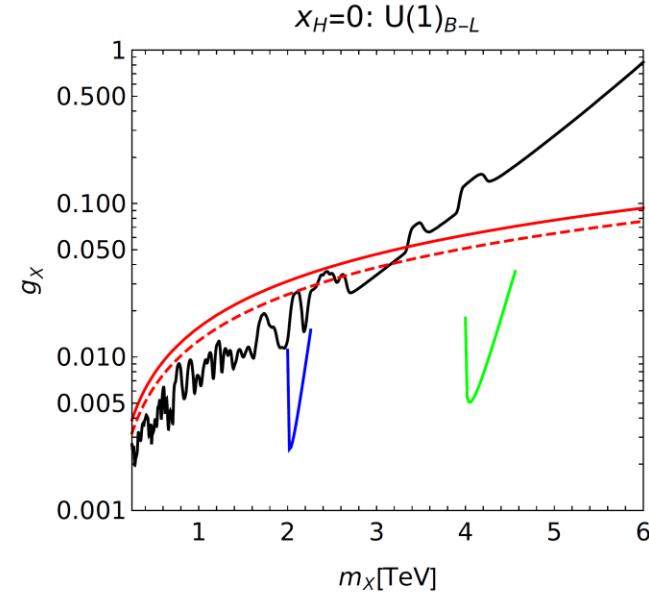
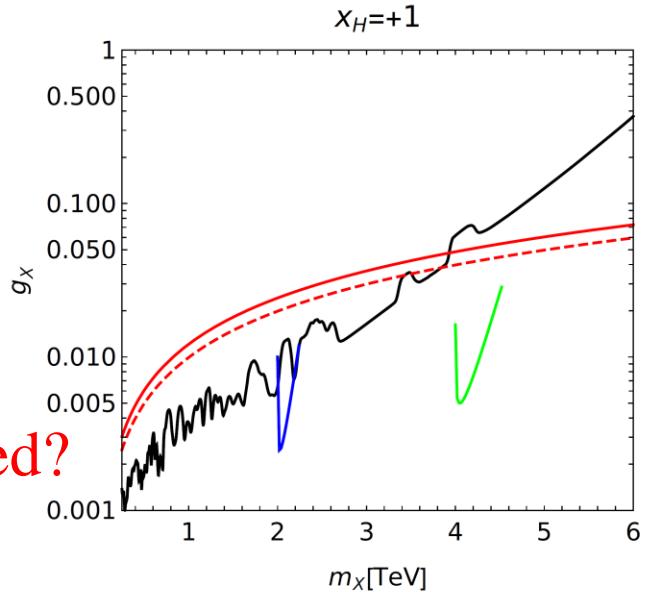
Hubble tension favoured?

- 3.1 dashed

- DM( $\Omega_{\chi} h^2$ ,  
XENON1T)

$m_{\chi} = 1 \text{ TeV}$

$m_{\chi} = 2 \text{ TeV}$



# § Summary

- Alternative  $B-L$  charge: possibility
  - Compatible with Type-II seesaw
  - Minimal Higgs sector
    - Same sign di-lepton is likely
    - Anomaly cancellation requires DR and Dirac DM
    - WIMP DM without a parity by hand
- Constraints are derived for LHC, DR and DM