

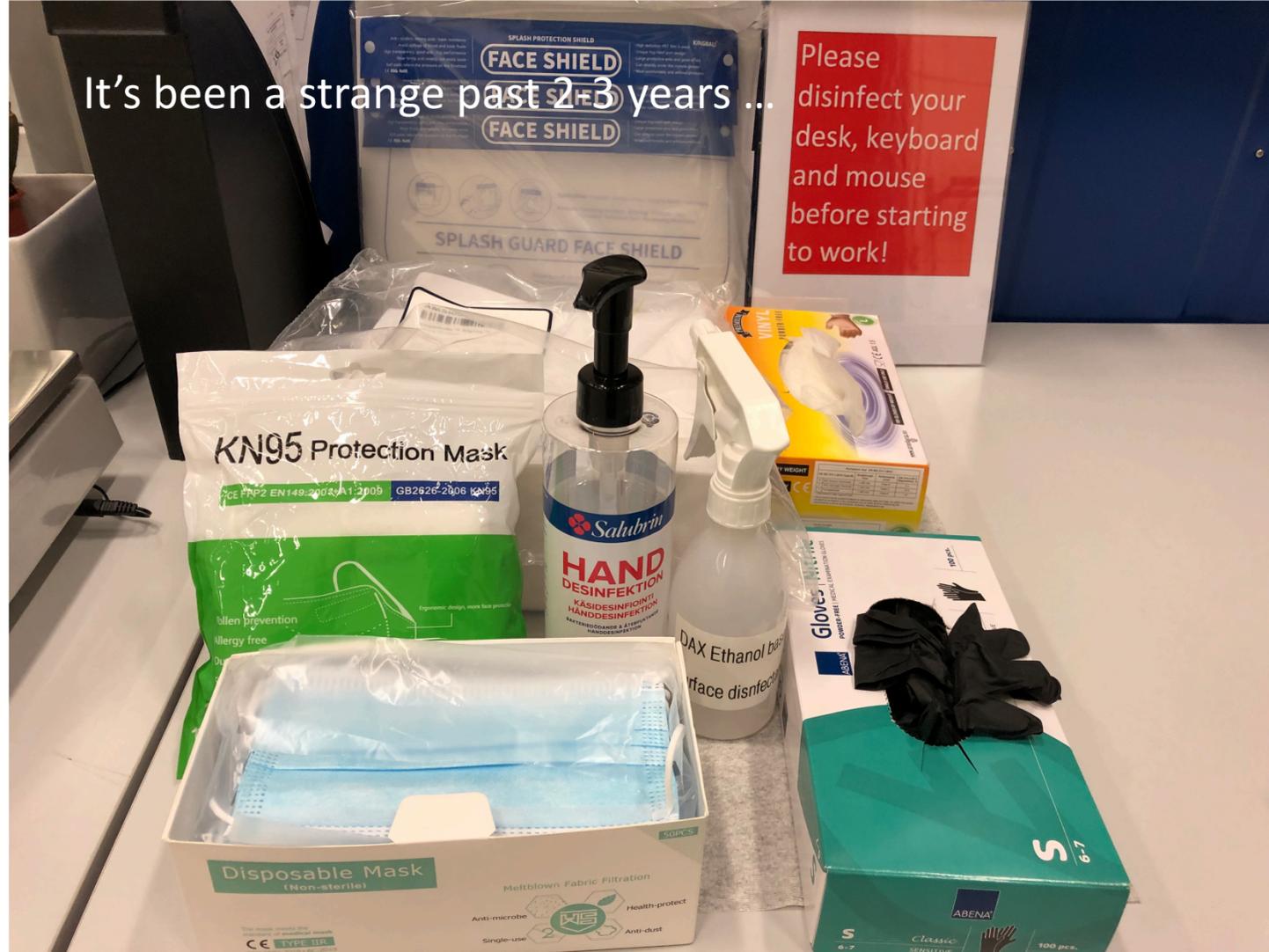


SENSORS
AND DEVICES

The Detector Data Acquisition System of the European Spallation Source

Prof. Richard Hall-Wilton

FBK-SD director

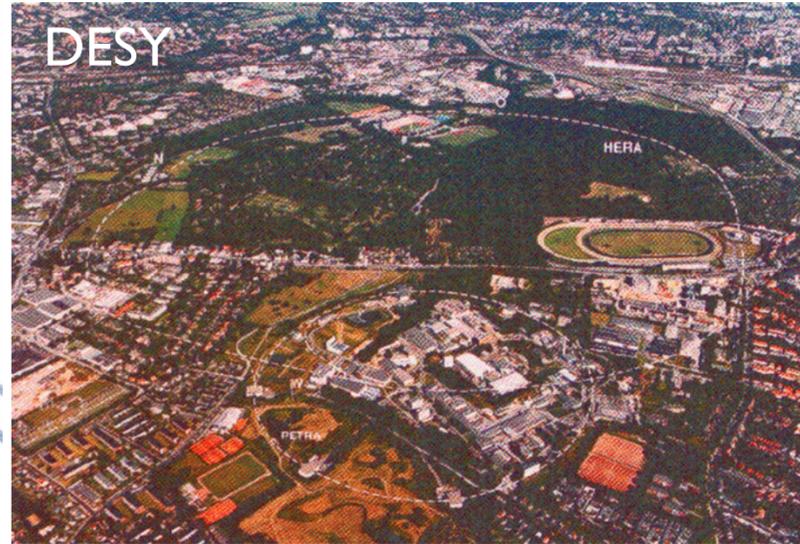
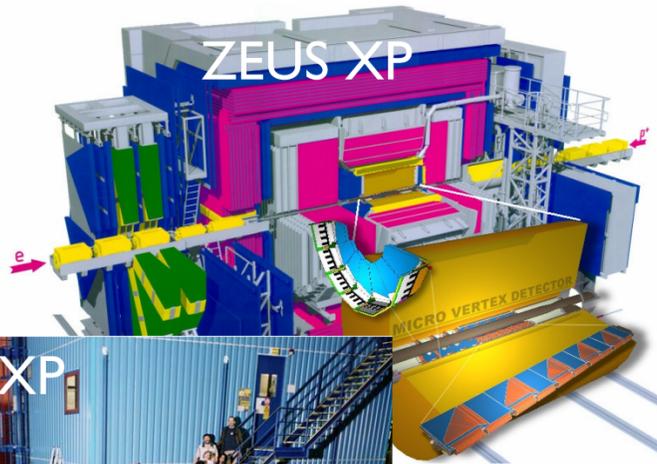


It's been a strange past 2-3 years ...

It's really exciting that we can (re-)start to meet in person again and meet each other ...

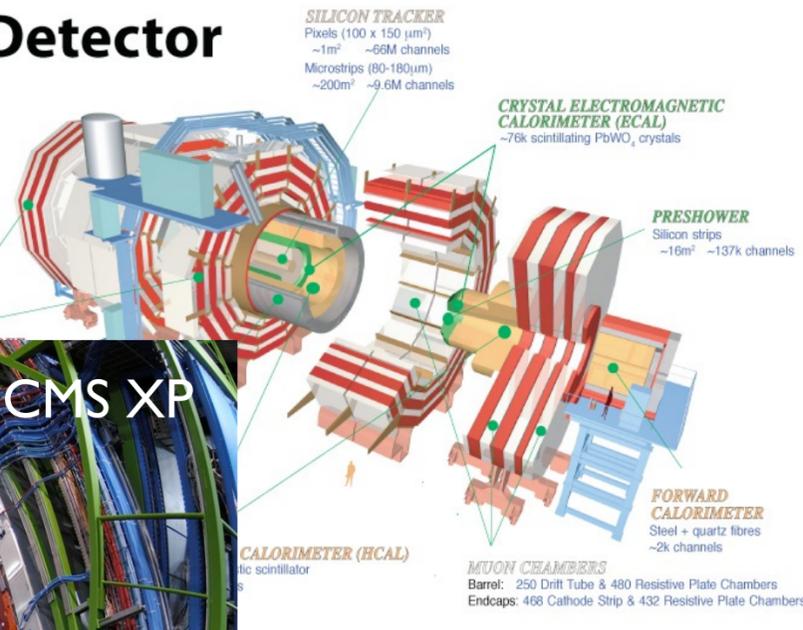


Projects that I have worked on ...

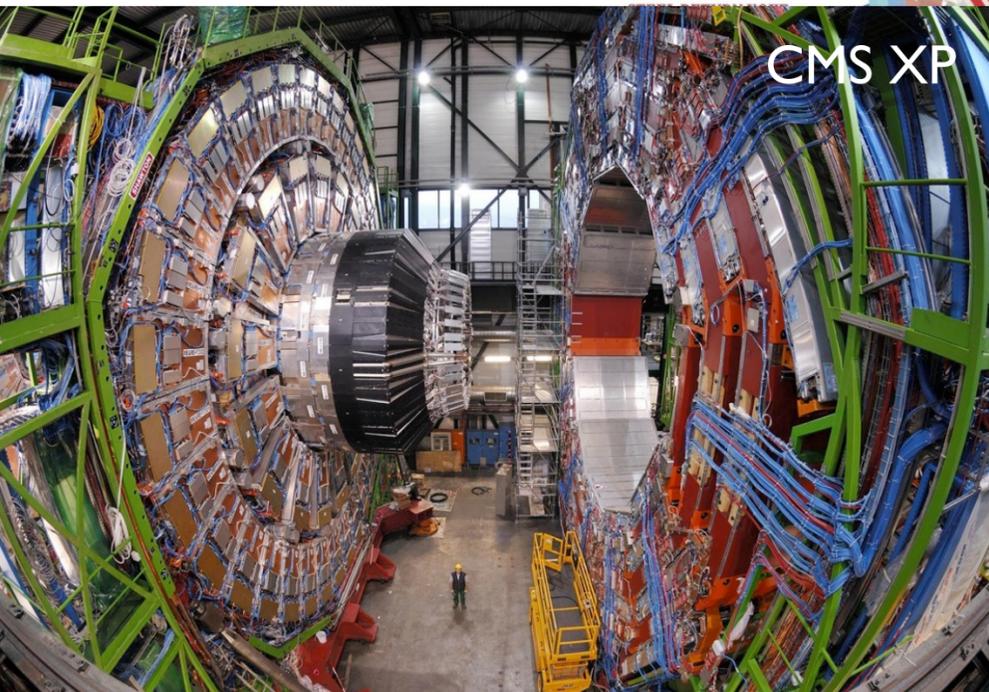
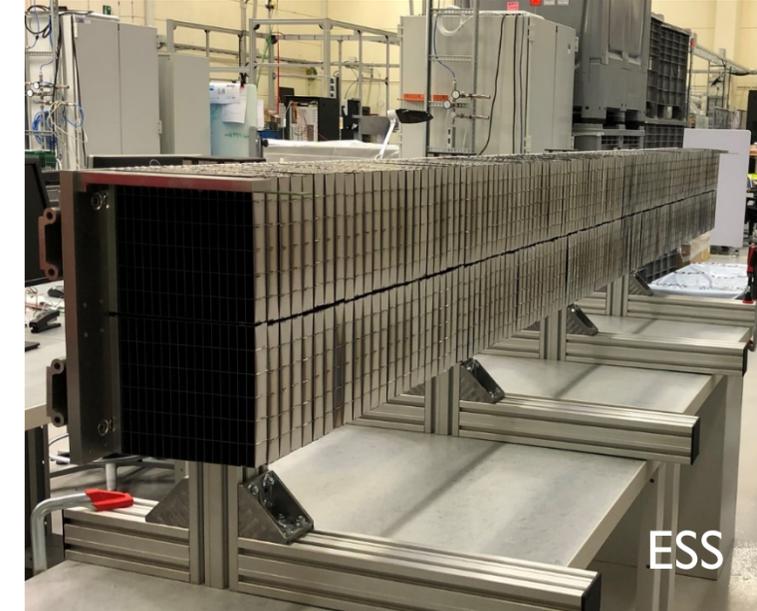


CMS Detector

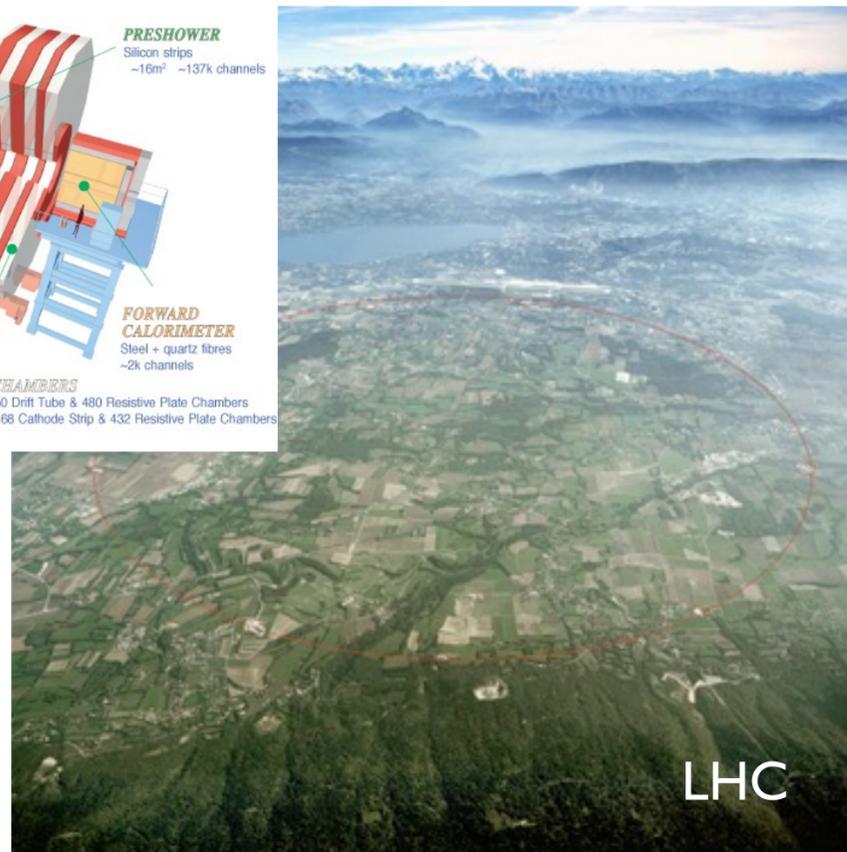
Pixels
 Tracker
 ECAL
 HCAL
 Solenoid
 Steel Yoke
 Muons



- RD42 Collaboration (Diamond Detector Development) (1995-6, 2006-11)
- ZEUS Collaboration for HERA at DESY (1995-2006)
- LHC Machine-Experiment Interface group (2005-2008)
- CMS Collaboration for LHC at CERN (2006-11)
- RD51 Collaboration (Micro Pattern Gaseous Detectors) (2013—)
- European Spallation Source (2011-2022)



CMS XP



LHC



ESS

WARNING:

The Disneyland Resort
contains chemicals
known to the state of
California to cause
cancer and birth
defects or other
reproductive harm.

Proposition 65,
California Health &
Safety Code Section
25249.6 et seq.

Caveat Emptor

... the contents of this talk
are reflections from
personal experience ...



EUROPEAN
SPALLATION
SOURCE



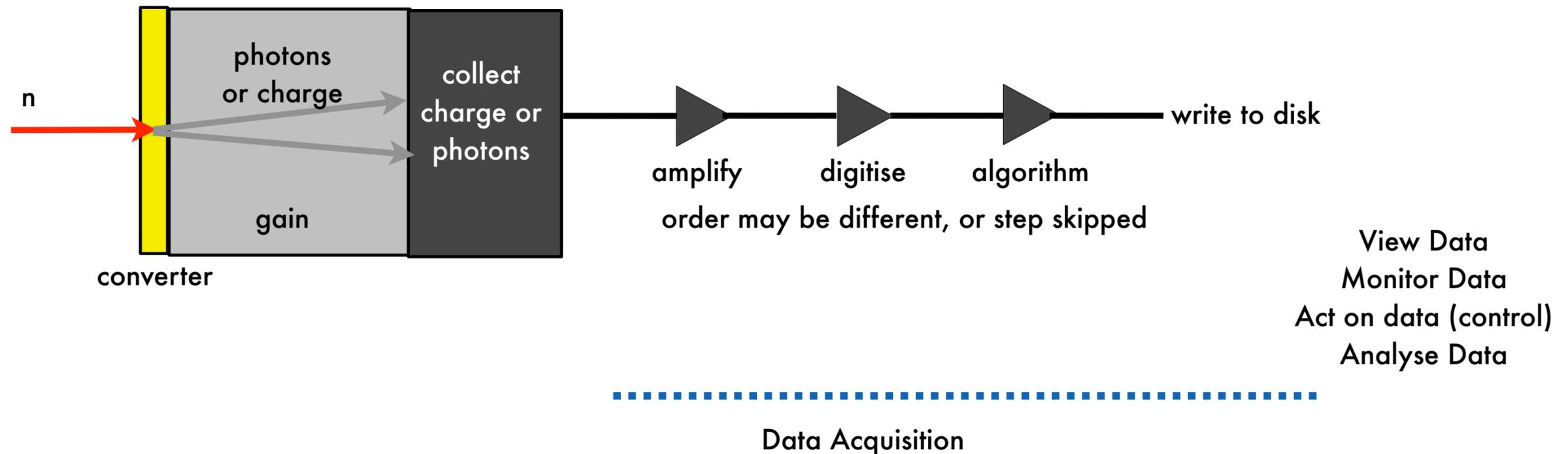
What am I going to talk about ...



- You have been learning about: Systems-on-Chip Based on FPGA for Scientific Instrumentation and Reconfigurable Computing
 - Detailed courses on how to perform a very particular and valuable technical skill
- Here, I am going to give a bird's eye view of how a large data acquisition system might look, utilising these skills you have been looking at ...
 - Am going to go through - with very broad strokes and very little detail - what needs to be taken into account in the design of the data acquisition system
- Like everyone, I will show this from my own experience - of the detector data acquisition system for the European Spallation Source ...

Basic Principles of Neutron Detectors

- Need to produce a measurable electric signal
- Not possible to directly detect slow neutrons - energy is too low
- Need to use nuclear reactions to convert neutrons into charged particles
- Then indirectly detect the charged particles in a charged particle detector
- Amplify, digitise, process as needed.
- Store data on disk



What is your end goal?

“horses for courses”

- Data acquisition is about being able to extract the information from the sensors to be able to carry out the measurements of interest as simply and **best** as possible
- Optimisation can be done for performance, cost, simplicity, off-the-shelf, size, energy usage, ...



Instrumentation



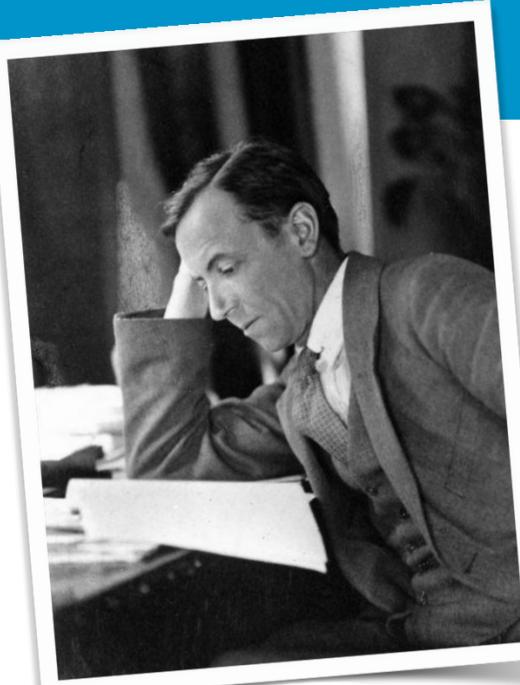
What camera you use has a big impact on the quality of photos that you get out of it ...



Bleeding edge Instrumentation enables novel and future science

What is the aim of the European Spallation Source?

Neutrons



1932: Chadwick discovers "a radiation with the more peculiar properties", the neutron.

Cliff Shull: where are the atoms?



1994 Nobel Prize in Physics

Bert Brockhouse: what do they do?





Lighting

New materials

Solar energy

Food

Medicine

Tailor made materials

Mobile phones

Cosmetics

Pacemakers

Bio fuels

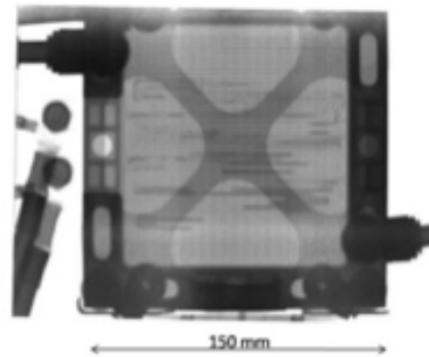
Implants

Transportation

Geo science

Charge neutral

Deeply penetrating



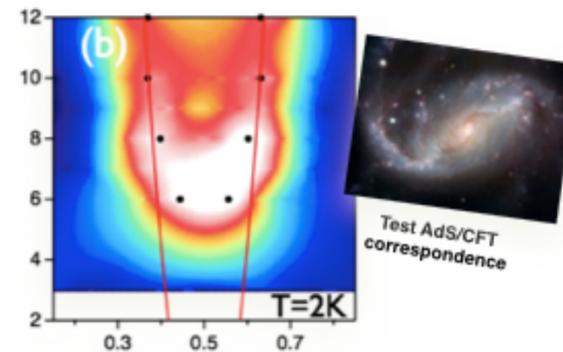
Li motion in fuel cells



**Help build
electric cars**

S=1/2 spin

Directly probe magnetism



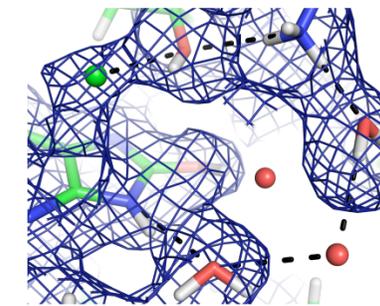
**Solve the puzzle of High-Tc
superconductivity**



Efficient high speed trains

Nuclear scattering

**Sensitive to light elements
and isotopes**

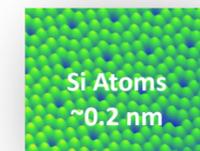
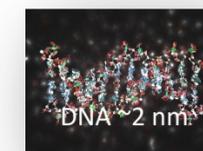
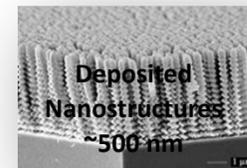


Active sites in proteins

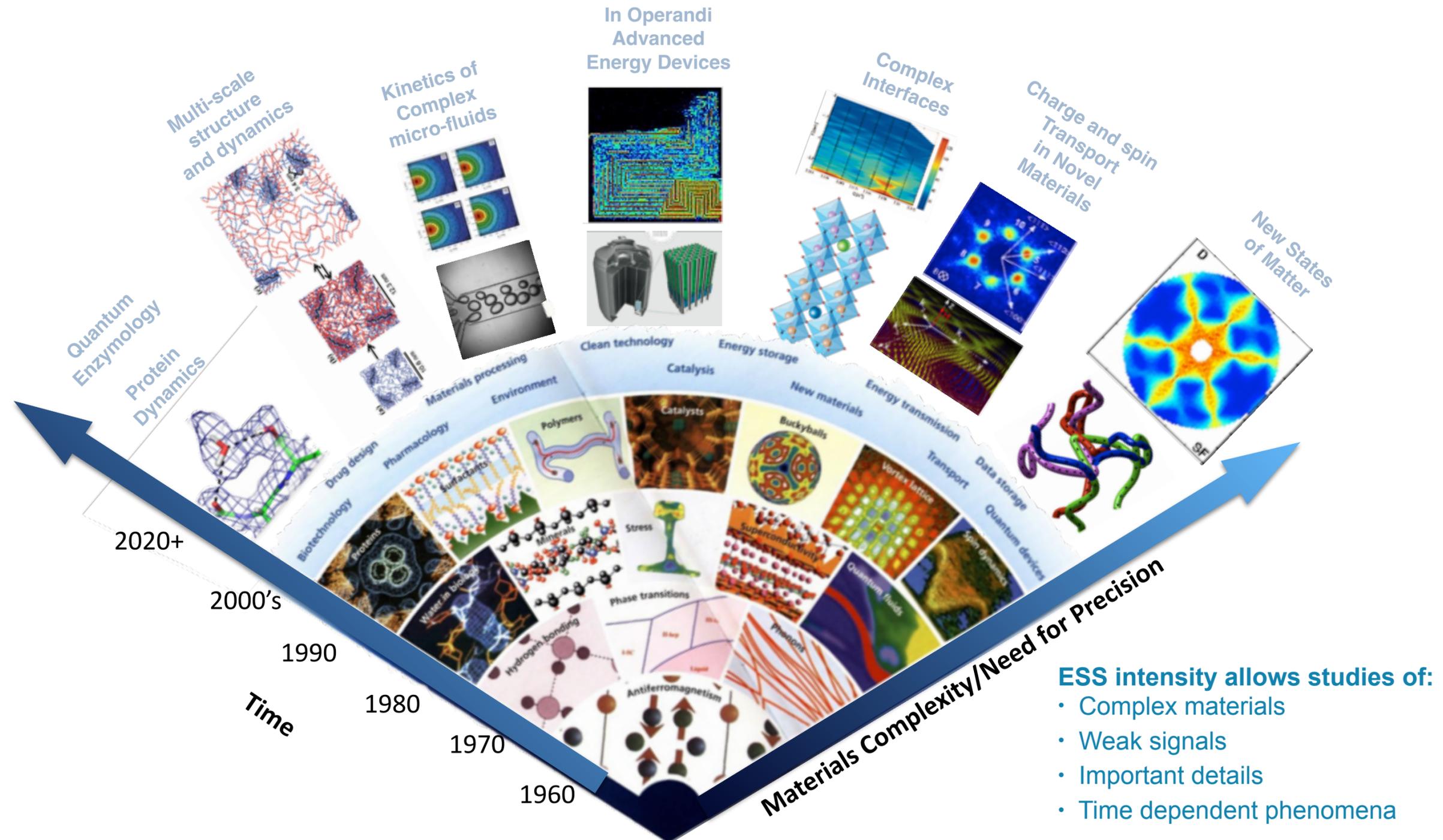


Better drugs

Probing length scales and dynamics



Neutron Science Pushes the Boundaries



The European Spallation Source: view to the Southwest in 2025



The ESS Site



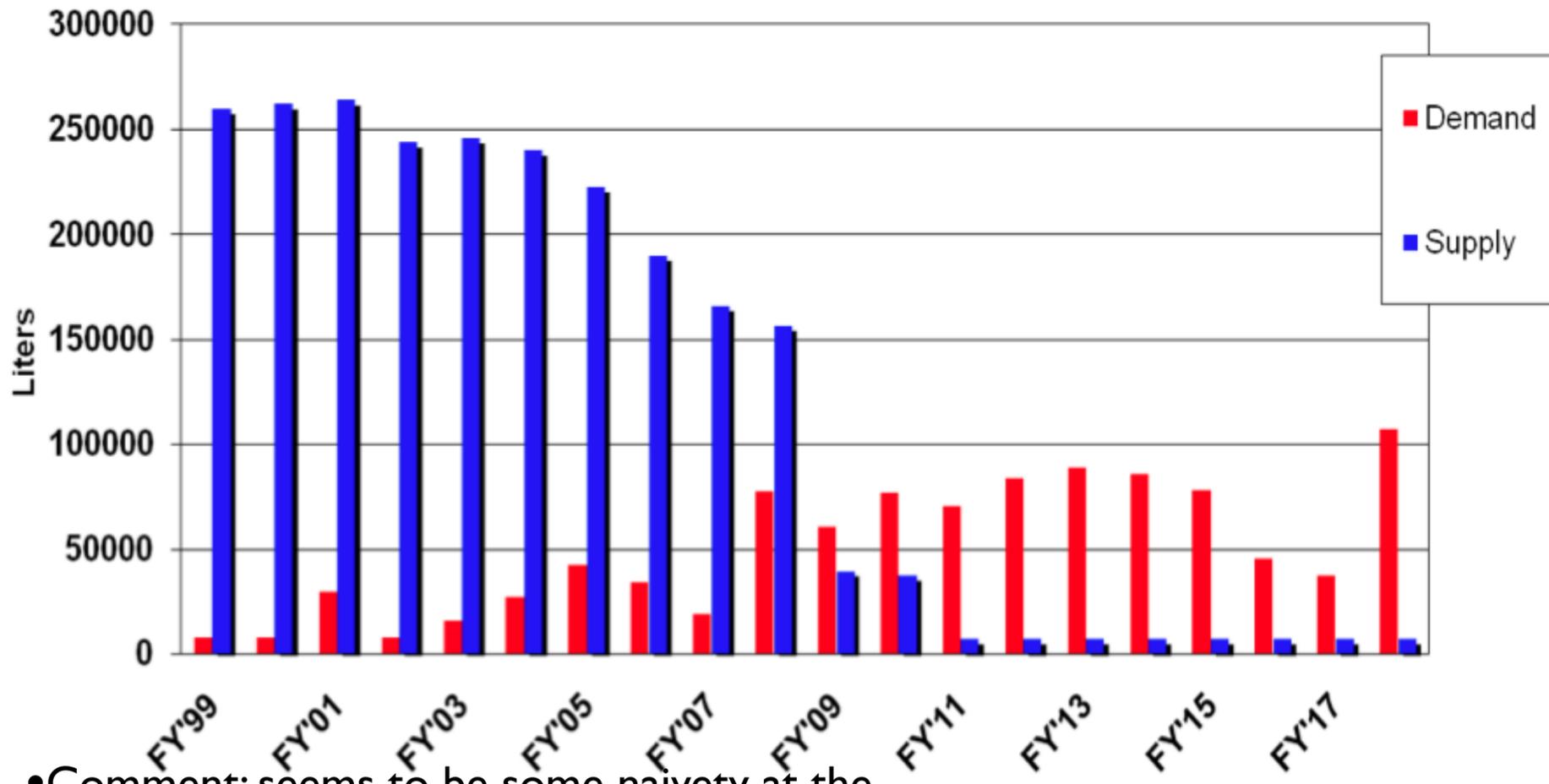
2011

- He-3 Crisis
- Schedule: 10 years until initial operations start

ESS Construction - October 2020



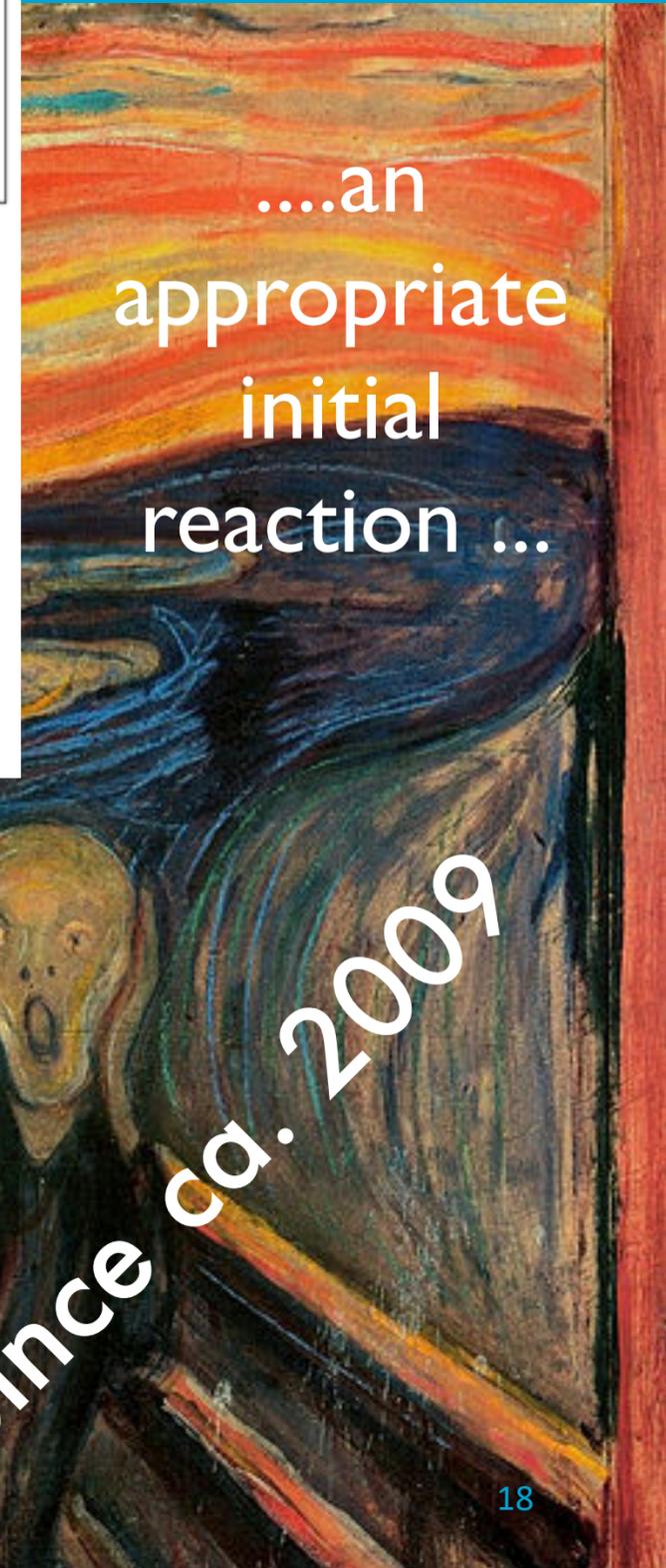
Helium-3 Crisis



- Comment: seems to be some naivety at the moment as stocks are being emptied rapidly
- Aside ... maybe He-3 detectors are anyway not what is needed for ESS?
eg rate, resolution reaching the limit ...

Crisis or opportunity ... ?

For almost all instruments, detectors are a limitation on performance
Opportunity to implement modern DAQ



...an appropriate initial reaction ...

Since ca. 2009

Schedule ...



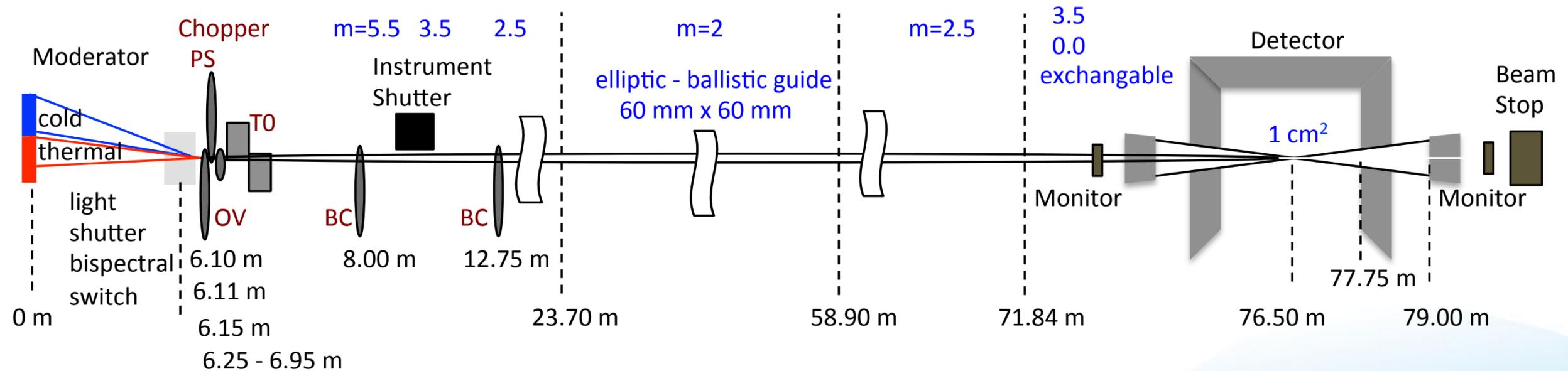
- 10 years until first neutrons ...

Instruments and their Requirements



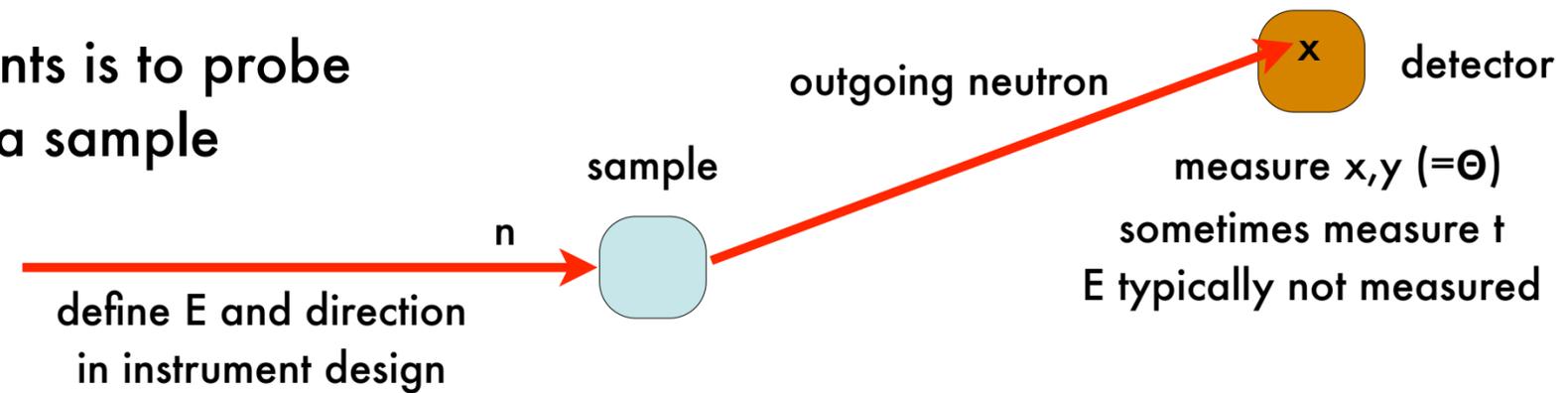
Instrument Design

- Instrument Design is about selecting the phase space of interest and maximising that
- Phase space here primarily means flux (6D: position, divergence) and neutron energy/wavelength
- Remember that as the neutron energy is not measurable, need to use time-of-flight or diffractive scattering to determine neutron energy
- Remember Liouville's theorem:
 - Phase space density is constant for conservative force fields
- It implies that high resolution measurements are low flux and vice-versa



Why?

- The purpose of the instruments is to probe with neutrons some aspect of a sample



- Very generically, this can be divided into elastic and inelastic categories
 - elastic: gives information on where atoms are
 - inelastic: gives information on what atoms do (ie move)
- This is measuring the cross sections:

elastic

$$\frac{d\sigma}{d\Omega}(\lambda, 2\theta, \psi)$$

- cross section / scattering probability into a solid angle, as a function of wavelength, scattering angle and aximuthal angle

inelastic

$$\frac{d^2\sigma}{d\Omega dE}(\lambda_{in}, \lambda_{sc}, 2\theta, \psi)$$

- double differential cross section / scattering probability into a solid angle, as a function of wavelength, scattered wavelength scattering angle and aximuthal angle

Science Drivers for the Reference Instrument Suite from the Technical Design Report



Multi-Purpose Imaging



General-Purpose SANS



Broadband SANS



Surface Scattering



Horizontal Reflectometer



Vertical Reflectometer



Thermal Powder Diffractometer



Bispectral Power Diffractometer



Pulsed Monochromatic Powder Diffractometer



Materials Science Diffractometer



Extreme Conditions Instrument



Single-Crystal Magnetism Diffractometer



Macromolecular Diffractometer



Cold Chopper Spectrometer



Bispectral Chopper Spectrometer



Thermal Chopper Spectrometer



Cold Crystal-Analyzer Spectrometer



Vibrational Spectroscopy



Backscattering Spectrometer



High-Resolution Spin-Echo



Wide-Angle Spin-Echo



Fundamental & Particle Physics

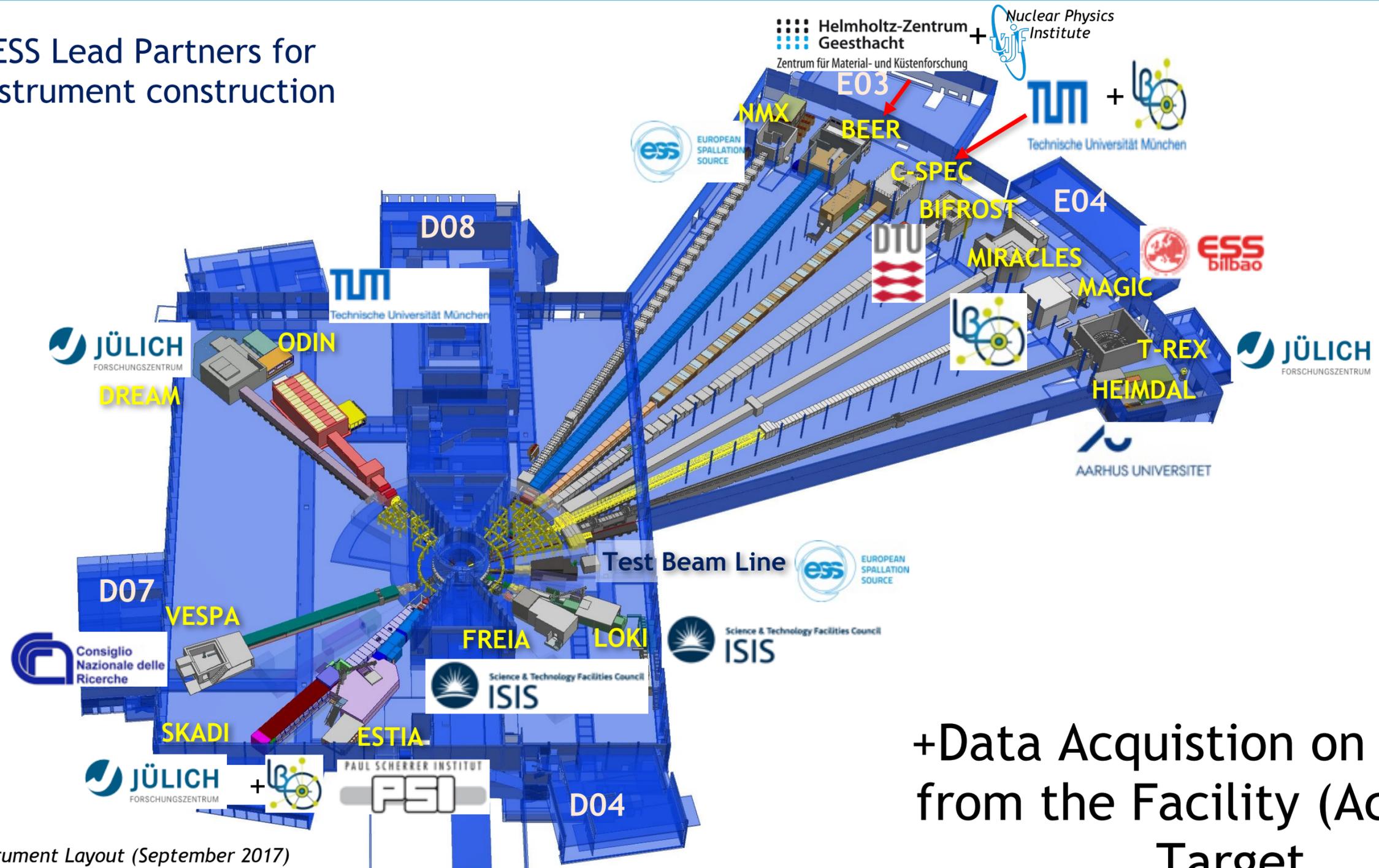


	life sciences		magnetism & superconductivity
	soft condensed matter		engineering & geo-sciences
	chemistry of materials		archeology & heritage conservation
	energy research		fundamental & particle physics

NSS Project scope: 15 neutron instruments + test beamline + support labs

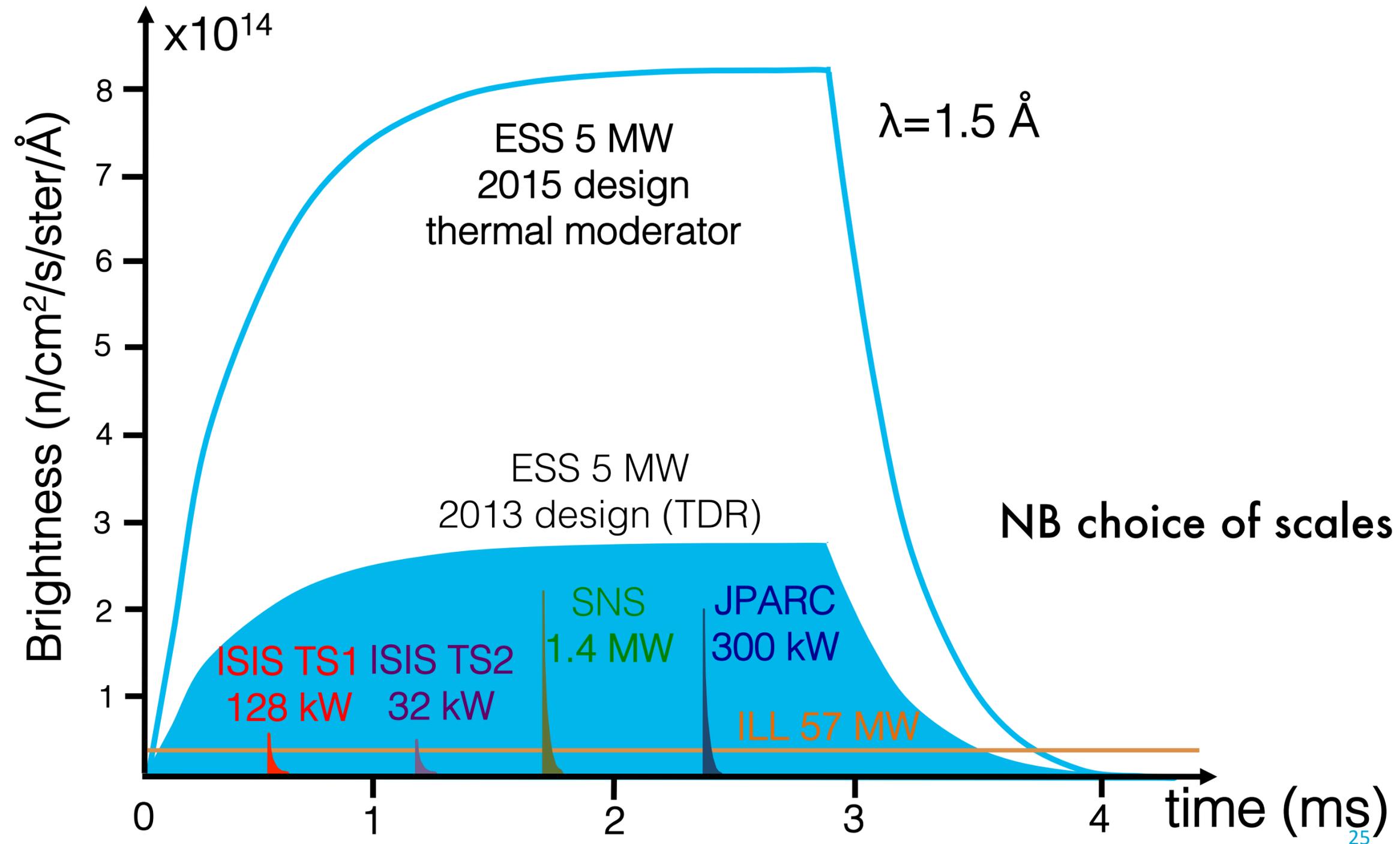


ESS Lead Partners for instrument construction

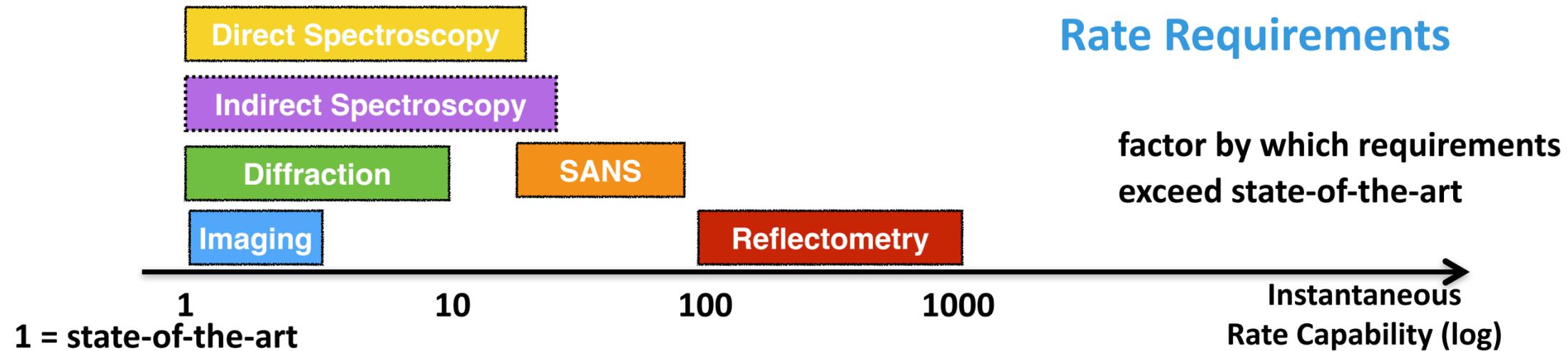


+Data Acquisition on Monitoring from the Facility (Accelerator, Target, ...)

Source BrightnESS

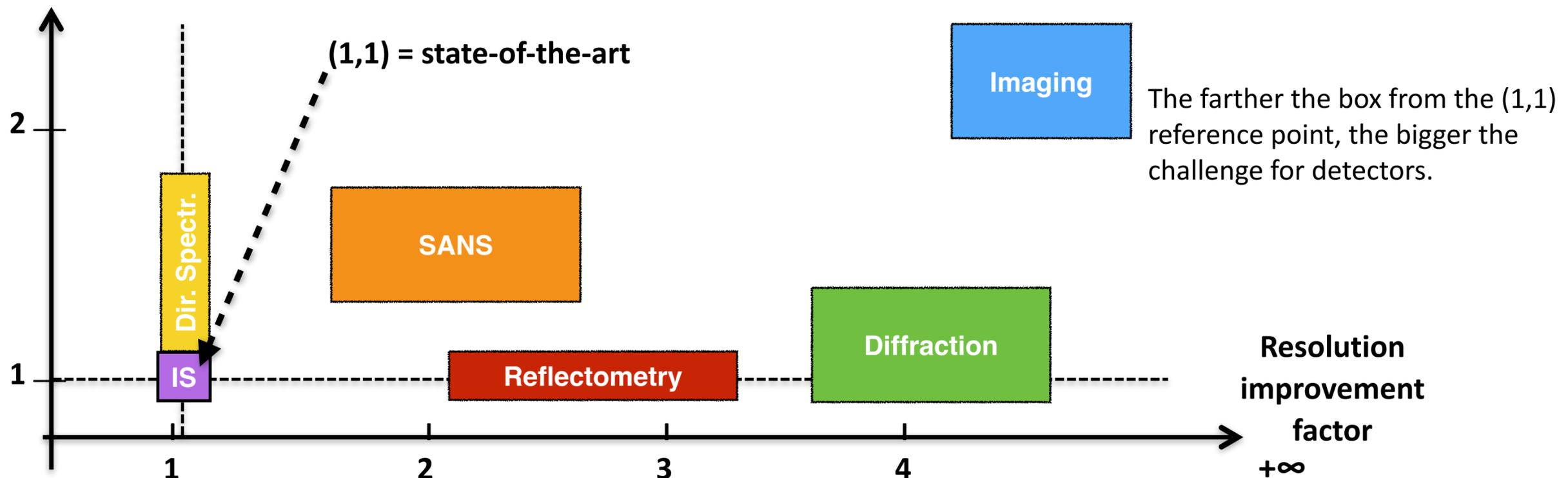


Requirements Challenge for Detectors for ESS: *beyond detector present state-of-the-art*



Increase factor
detector area

Resolution and Area Requirements



Typical Detector Requirements



- Size: from 0.25m² up to 30m²
- Position resolution: 100um - 10mm
- Time Resolution: <1us

- Rate and DAQ requirements very much defined by the instrument and data topology
- Can be >MHz/channel instantaneous in some cases
- Average rates much lower

- Every instrument is very much a bespoke individual design ...
- For a good user experience ...
- ... It's important to design the detector to the individual detector
- ... it's important to make sure that the DAQ can cope with the, and has a homogeneous look and feel

Instrument Control System Design

- Need a modular system to cope with diversity in design
- Use EPICS as the control system for the DAQ
- Use accelerator timing system: only 1 timing system needed for the facility
- Can access monitoring & diagnostic data across the facility data
- High rate & high number of channels = high data volume
- Use a dedicated data interface
- Unify the data where you actually want to utilise it
- Where possible use network equipment & standard PCs: lower the expert knowledge needed
- Use FPGAs only where performance needed ...

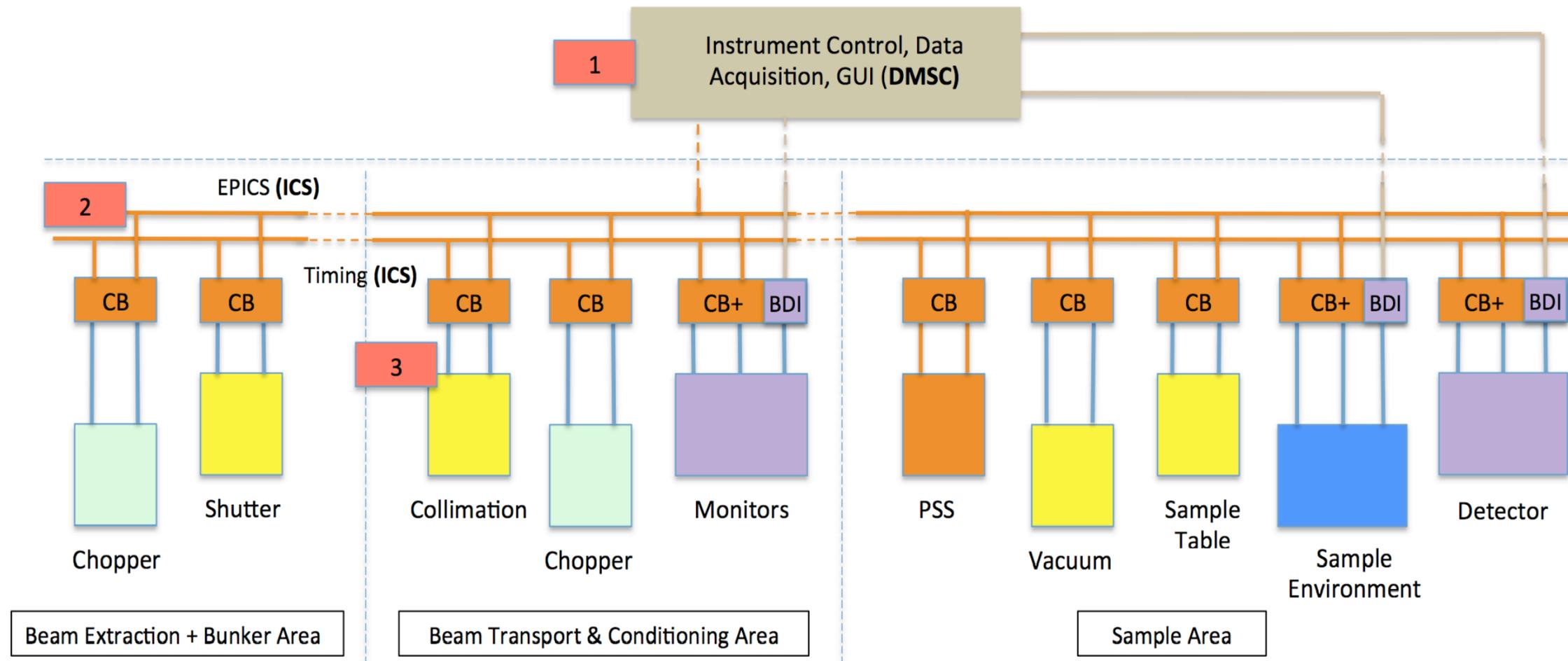
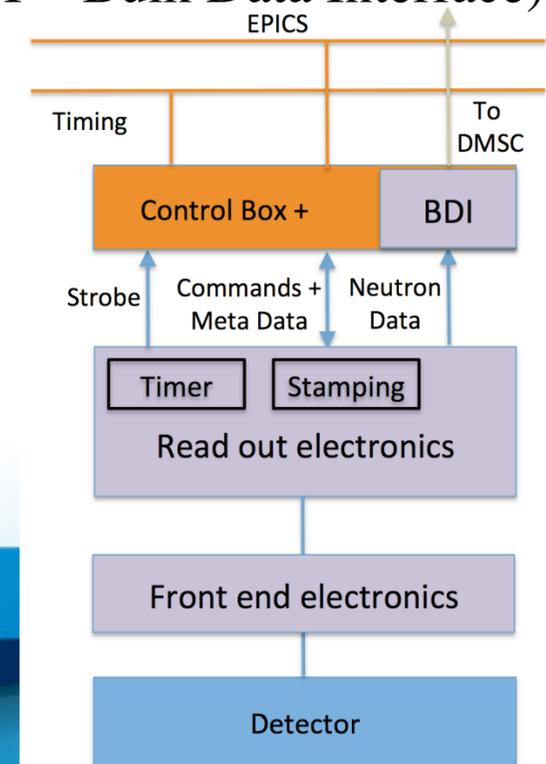


Figure 3. Topology of the modular instrument control concept for a 160m long Neutron Scattering Instrument at ESS. (CB = Control Box, BDI = Bulk Data Interface).





TEAMWORK

Share Victory. Share Defeat.

Everyone should play to their strengths

ESS Partners on Detectors



INTERNATIONAL COLLABORATION FOR THE DEVELOPMENT OF NEUTRON DETECTORS

Detectors for ESS Instruments



Instrument	Neutron Converter	Detector Type	Gas Gain	Number of Channels	Front-End Type/ASIC
CSPEC	10B4C	MWPC	ca. 10	ca.20k	VMM3A
TREX	10B4C	MWPC	ca. 10	ca. 15k	VMM3A
ESTIA	10B4C	MWPC	ca. 10	ca. 6k	VMM3A
FREIA	10B4C	MWPC	ca. 10	ca. 4k	VMM3A
NMX	Gd	GEM	ca. 1000	ca. 15k	VMM3A
DREAM	10B4C	MWPC	<100	400k	CDT/CIPIX
MAGIC	10B4C	MWPC	<100	165k	CDT/CIPIX
HEIMDAL	10B4C	MWPC	<100	250k	CDT/CIPIX
LOKI	10B4C	Straws	>1000	5k	Discrete Preamp/ CAEN R5560
BIFROST	3He	Tube	>1000	ca. 100	Discrete Preamp/ CAEN R5560
VESPA	3He	Tube	>1000	ca. 100	Discrete Preamp/ CAEN R5560
MIRACLES	3He	Tube	>1000	ca. 100	Discrete Preamp/ CAEN R5560
SKADI	6Li	Scintillator	N/A	25000	IDEAS/IDE3465
ODIN	6Li	Various	N/A	ca. 1M	TIMEPIX4
BEER	10B4C	MWPC	>100	40	Delay Line + Custom FPGA
Beam Monitors	Various	MWPC/GEM/IC	1-100	50	Discrete Preamp / ADC OHWR FMC- ADC-100m14b4Cha
TestBeam Line	10B4C	MWPC	ca. 10	ca. 1k	VMM3A

- Step change # channels cf. current instruments
- From 100's to 10k's
- Need for using ASICs to handle large channel count at moderate cost
- Different detector partners means a variety of choices for front-end
- Key requirement for DAQ system is to be able to integrate a multiplicity of detector types and approaches
- Unify the “look and feel” within the electronics DAQ

Instrument	Neutron Converter	Detector Type	Gas Gain	Number of Channels	Front-End Type/ASIC
CSPEC	10B4C	MWPC	ca. 10	ca.20k	VMM3A
TREX	10B4C	MWPC	ca. 10	ca. 15k	VMM3A
ESTIA	10B4C	MWPC	ca. 10	ca. 6k	VMM3A
FREIA	10B4C	MWPC	ca. 10	ca. 4k	VMM3A
NMX	Gd	GEM	ca. 1000	ca. 15k	VMM3A
DREAM	10B4C	MWPC	<100	400k	CDT/CIPIX
MAGIC	10B4C	MWPC	<100	165k	CDT/CIPIX
HEIMDAL	10B4C	MWPC	<100	250k	CDT/CIPIX
LOKI	10B4C	Straws	>1000	5k	Discrete Preamp/ CAEN R5560
BIFROST	3He	Tube	>1000	ca. 100	Discrete Preamp/ CAEN R5560
VESPA	3He	Tube	>1000	ca. 100	Discrete Preamp/ CAEN R5560
MIRACLES	3He	Tube	>1000	ca. 100	Discrete Preamp/ CAEN R5560
SKADI	6Li	Scintillator	N/A	25000	IDEAS/IDE3465
ODIN	6Li	Various	N/A	ca. 1M	TIMEPIX4
BEER	10B4C	MWPC	>100	40	Delay Line + Custom FPGA
Beam Monitors	Various	MWPC/GEM/IC	1-100	50	Discrete Preamp / ADC OHWR FMC-ADC-100m14b4Cha
TestBeam Line	10B4C	MWPC	ca. 10	ca. 1k	VMM3A

Detectors for ESS Instruments



Have managed to settle on 3 main front ends for 13/16 instruments:

- CAEN R5560 for detectors using charge division (He-3 PSDs and Boron Straws). 4 instruments.
- CIPIX for diffraction detectors using CDT Jalousie detectors. 3 instruments.
- VMM3A for Boron wire detectors (MultiGrid, MultiBlade and GEM detectors). 6 instruments.

Detector Electronics Integration Models

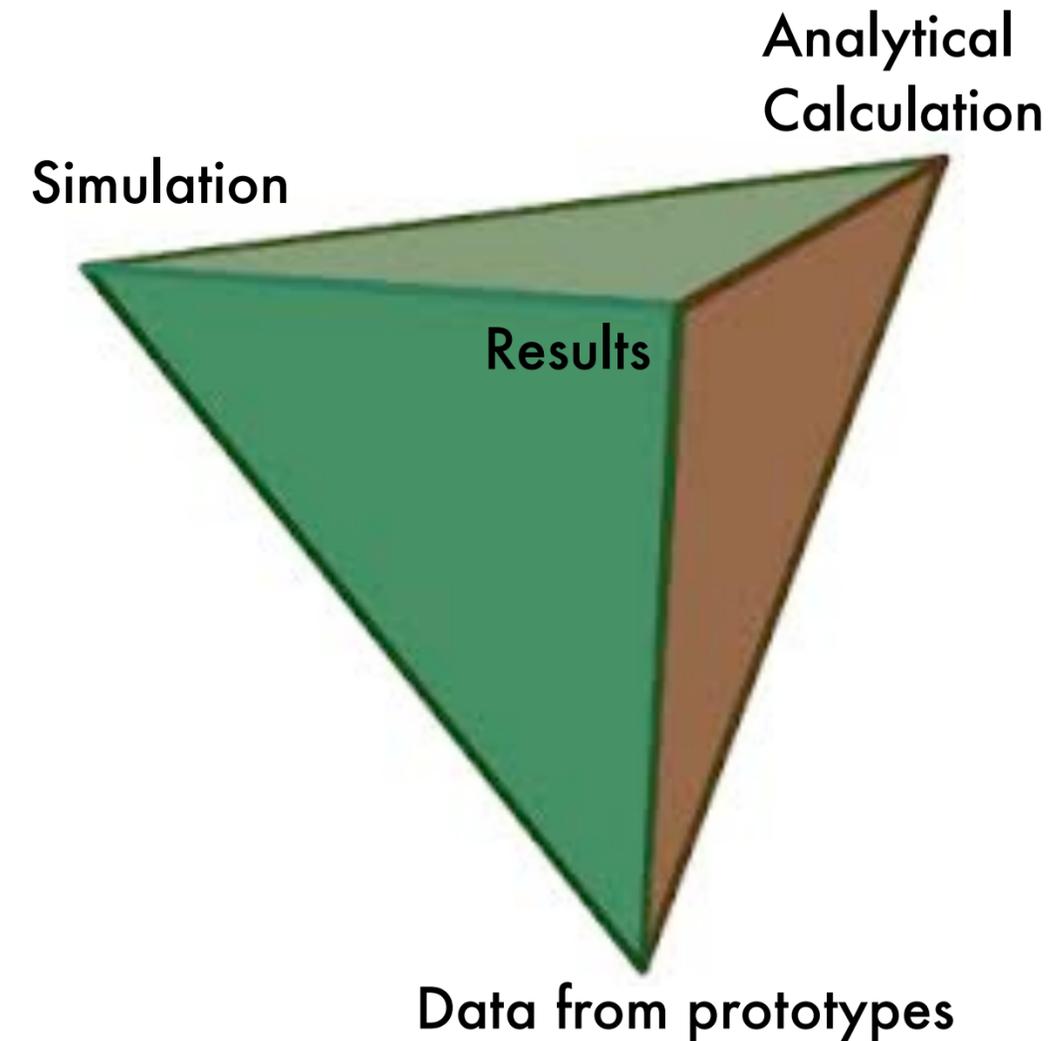


Data



Data: Calculation, Simulation, Prototype Data

- As you have heard, simulation is a very powerful tool
- ... but the computer will always lie to you ...
- Data from prototype tests is golden
- Lack of ability to trigger independently on the neutron means some degree of arbitrariness in defining the measurement
 - Checking that your measured data is correct is complicated
- Additionally, always try and calculate analytically or “back of envelope” what your expectation is
 - (Or at least upper and lower limits)
- Use all 3 of these **together** to understand the performance of your prototypes
 - Expect “features” and non-agreement and investigate them
 - Iterative



Definitions and Standards



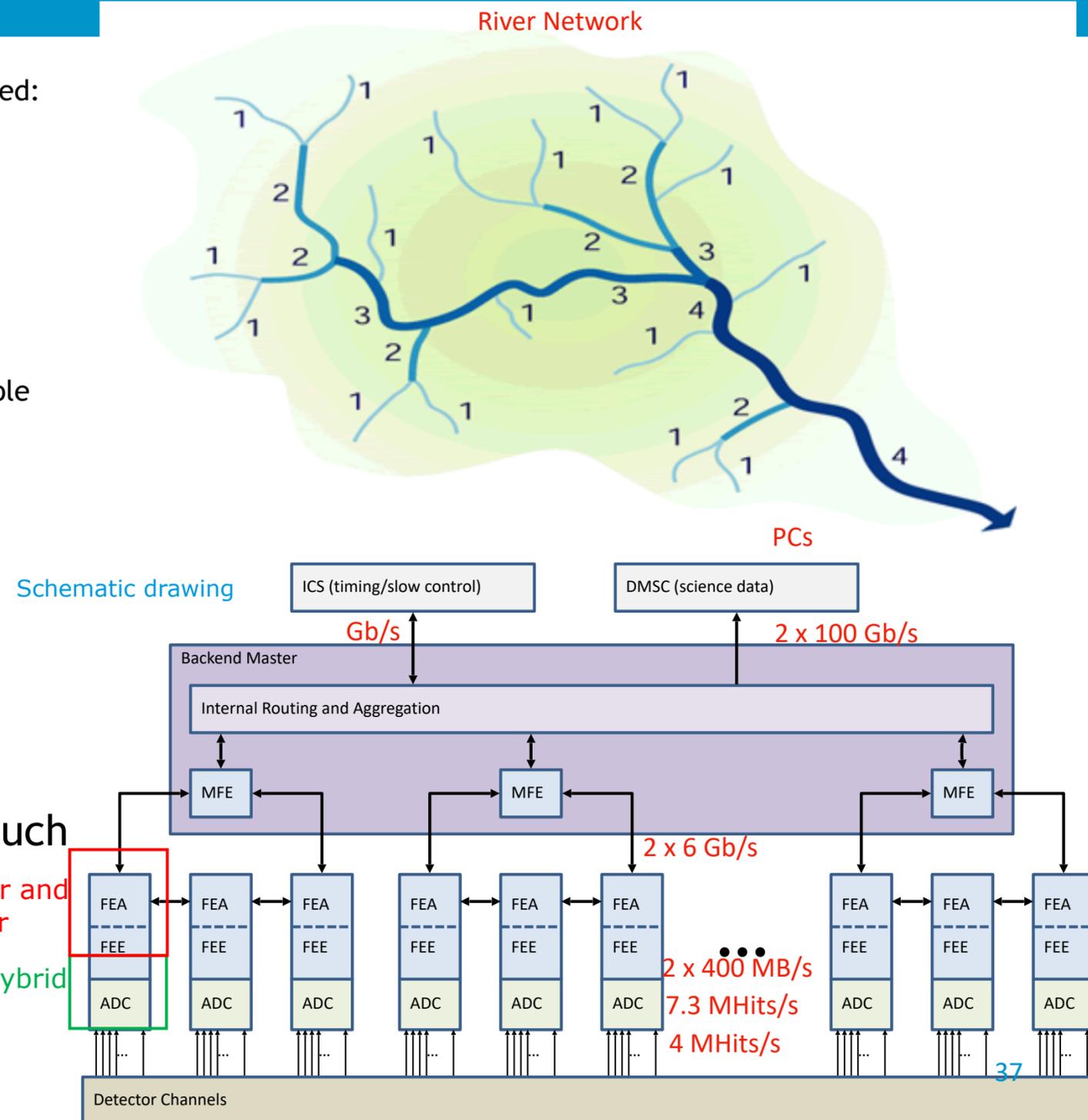
Detector Rate and Data Rate

Important to define what is meant by rate when it is quoted:

- **Global time-averaged incident/detection rate:** the total number of neutrons per second entering/recorded by the entire detector
- **Local time-averaged incident/detection rate:** the total number of neutrons per second entering/recorded in a detector pixel, channel or unit
- **Global peak incident/detection rate:** the highest instantaneous neutron incident/detection rate on the whole detector
- **Local peak incident/detection rate:** the highest instantaneous neutron incident/detection rate on the brightest detector pixel, channel or unit

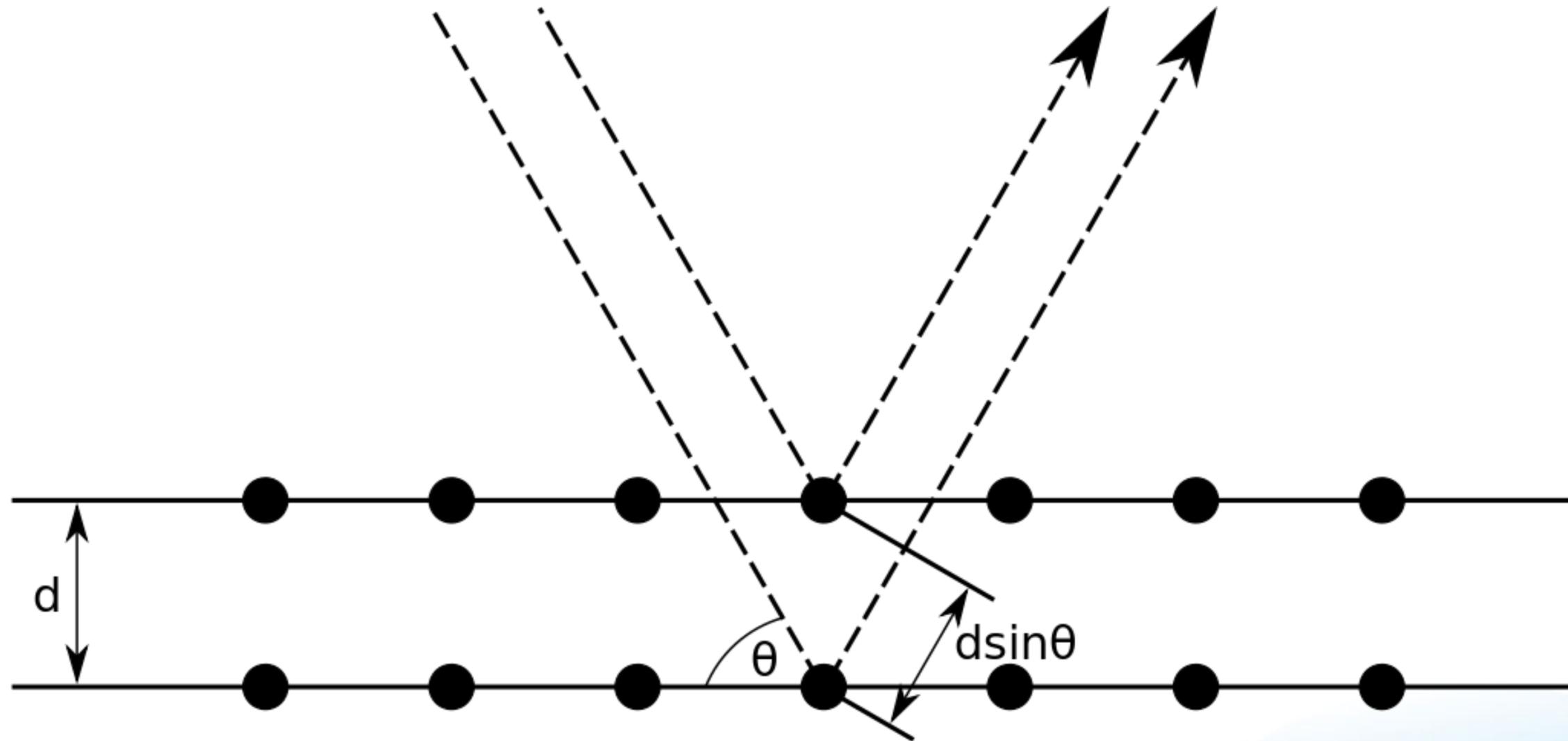
*I. Stefanescu et al., JINST12 (2016) P01019

- Any DAQ system is vulnerable to "flooding" at any point in the system
- Important to with realistic data as much as possible
- Pattern & quantity



Diffraction

Sizes probed = "atomic structures" = 0.1 nm - 10 nm



Position and intensity of diffraction peaks gives atomic positions

Detectors are tools

Basically, in some form,
you want to measure
Bragg's equation

$$n\lambda = 2d \sin \theta$$

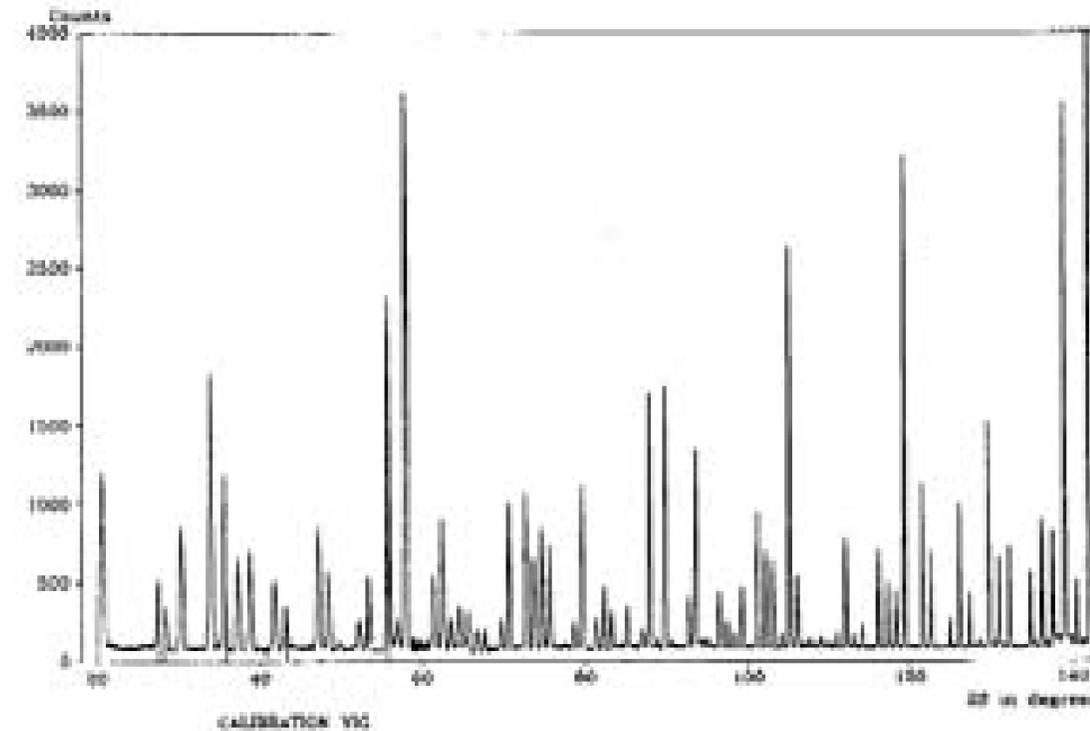
Define the neutron wavelength with your instrument design

Detectors allow you to measure theta

It means that you can calculate "d"

Therefore the detector should be designed to give you the most appropriate measurement of scattering angle for a instrument class

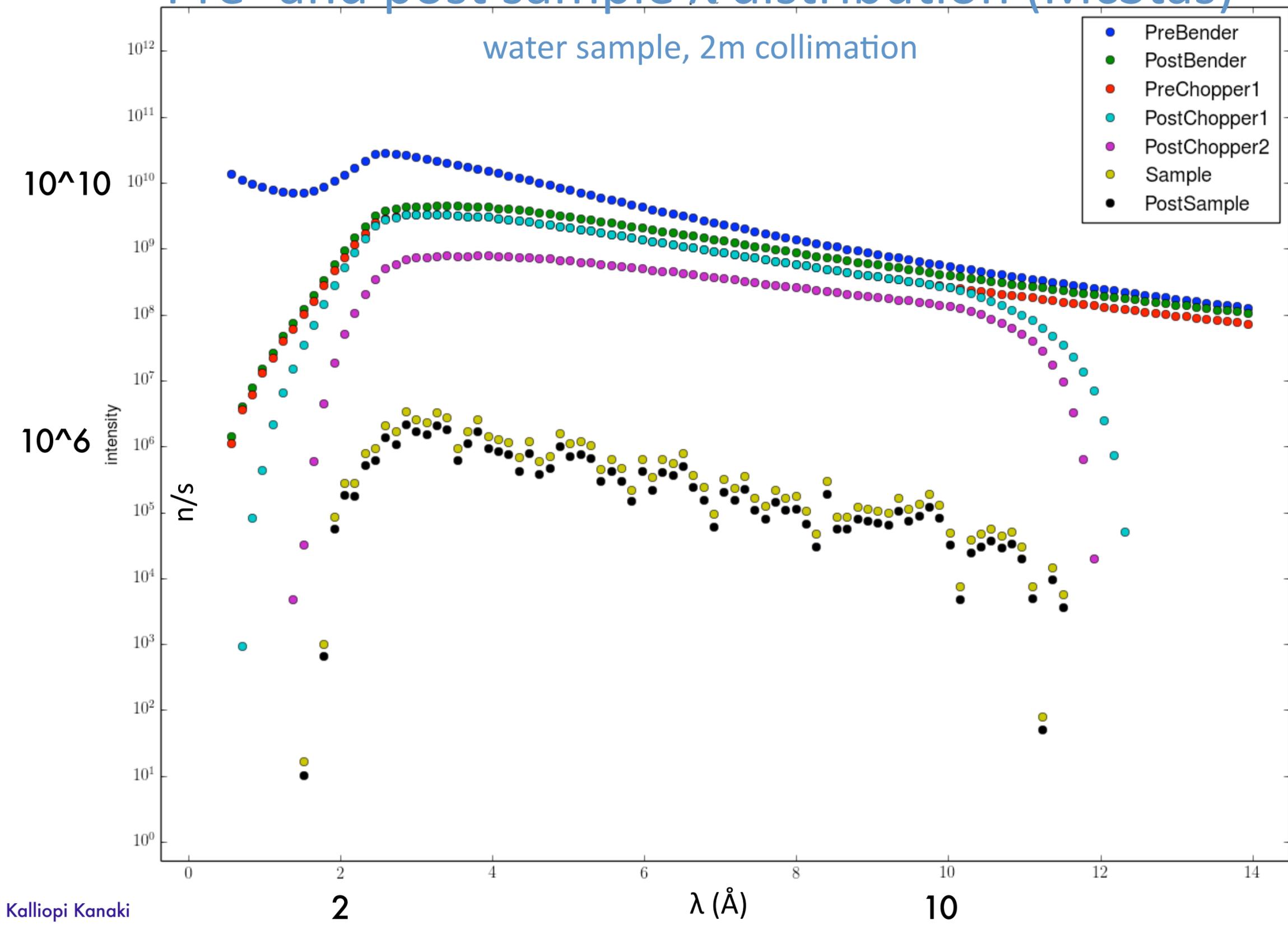
"horses for courses"





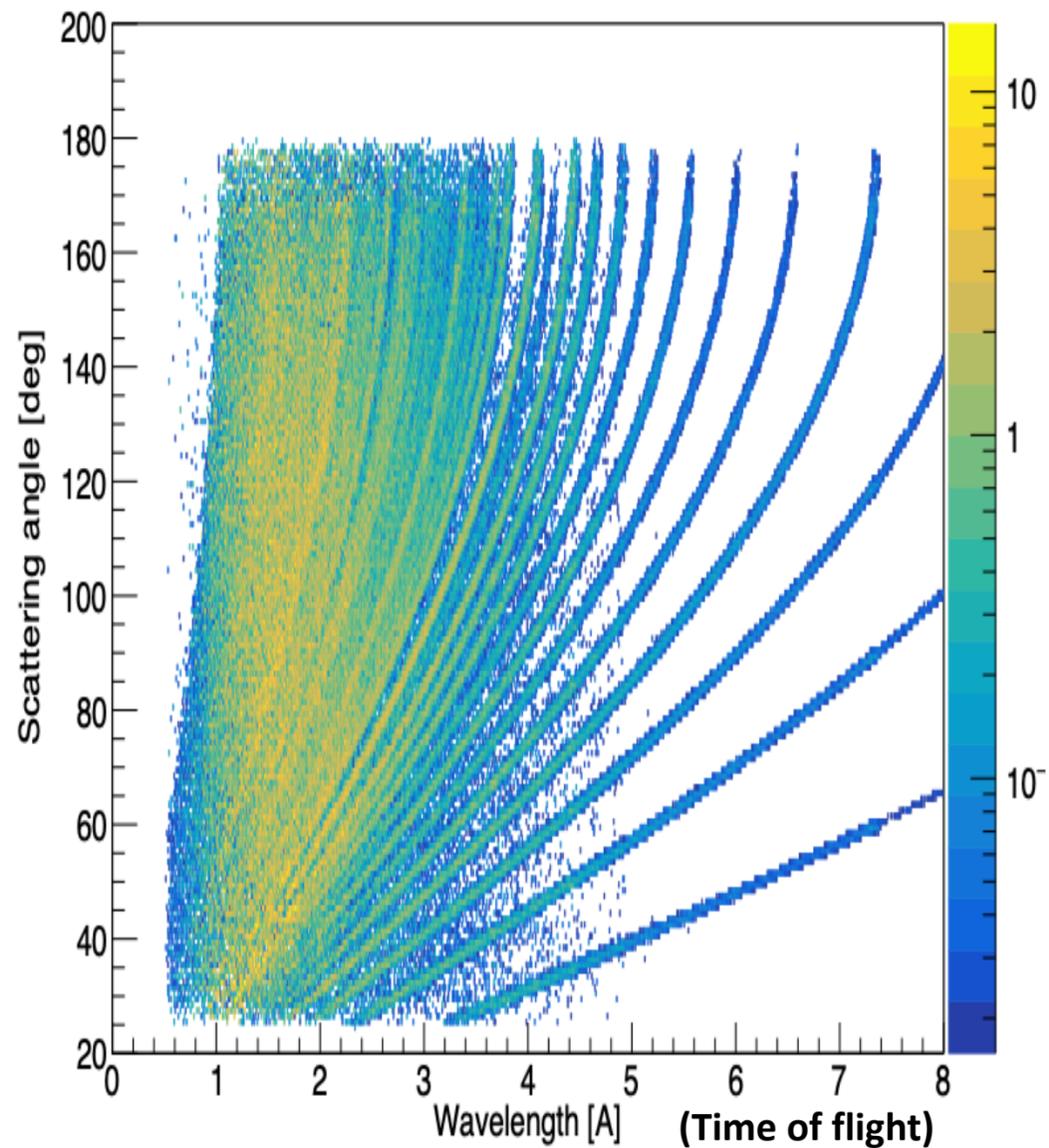
Selecting phase space ...

Pre- and post sample λ distribution (McStas)

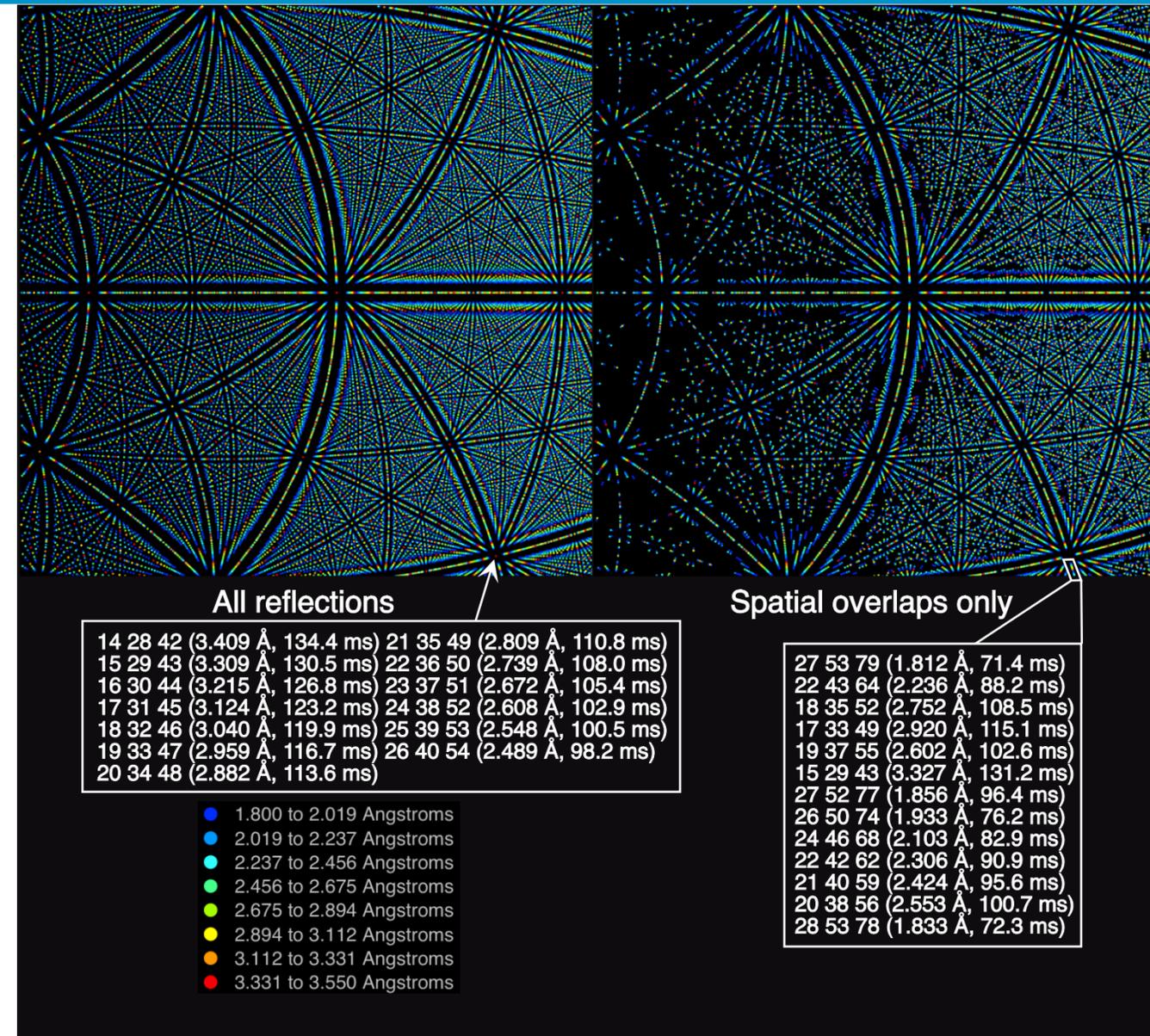


What does data look like?

Neutron Diffraction



Powder Diffraction

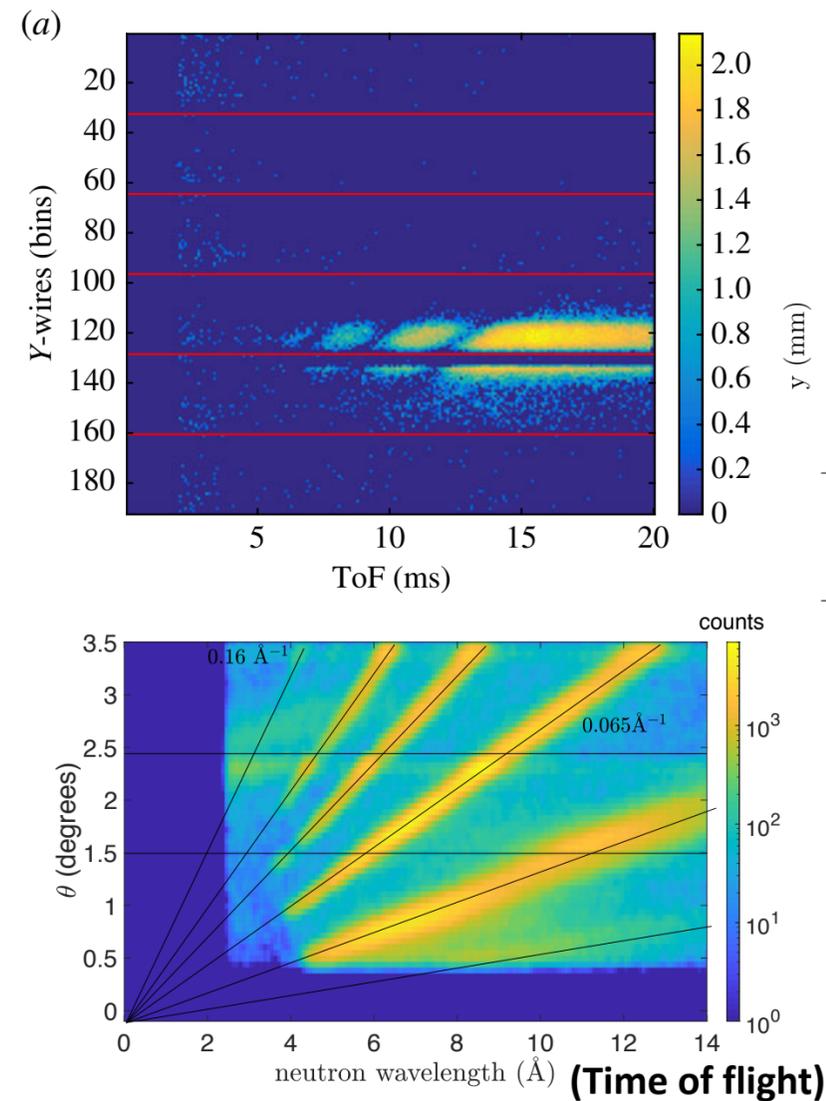


Protein Diffraction

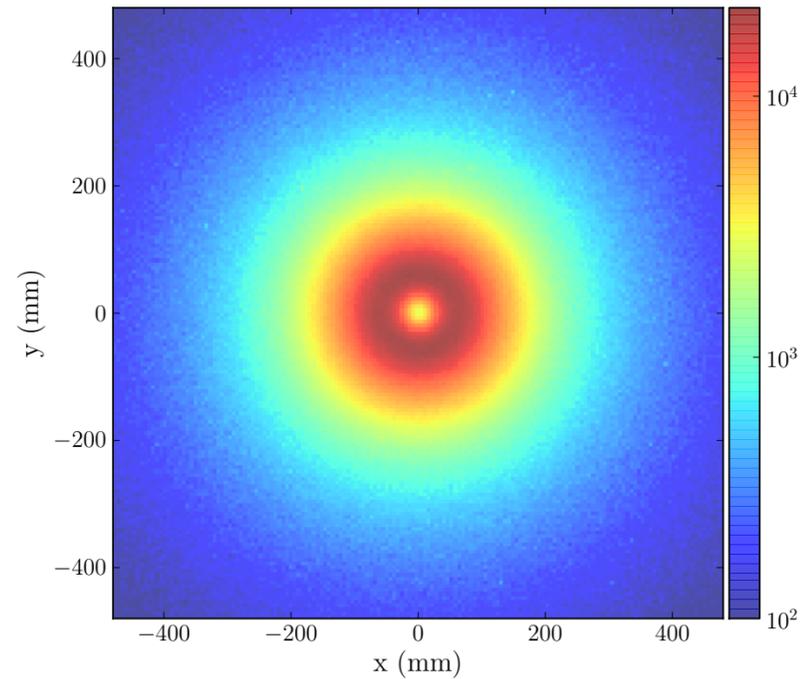
Data is sparse and peaky

What does data look like?

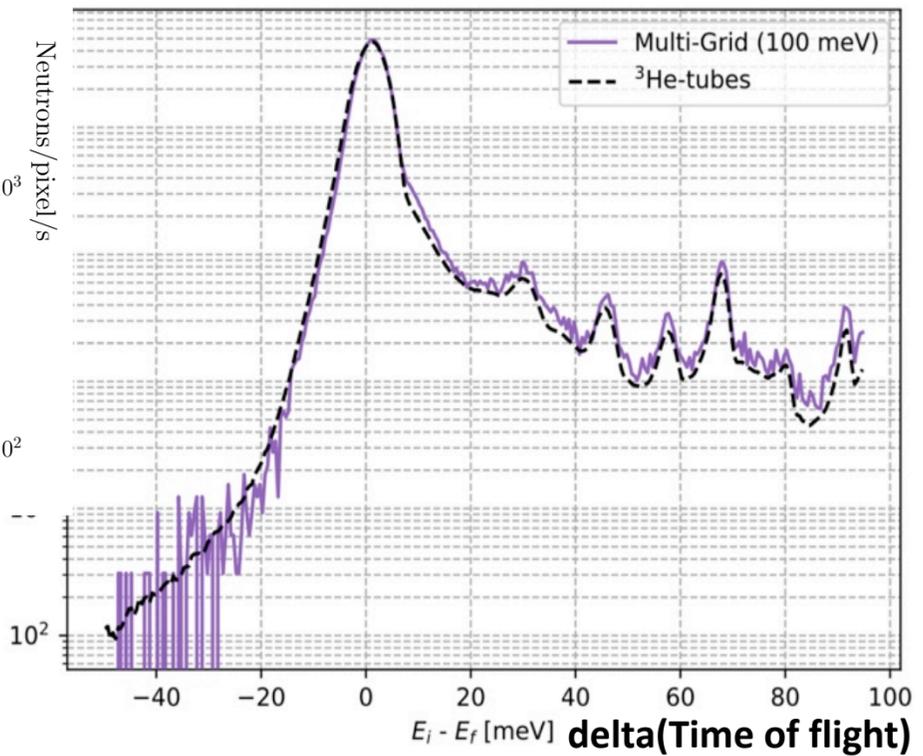
Reflectometry, Small Angle Scattering, Spectroscopy



Reflectometry



Small Angle Neutron Scattering



Spectroscopy

Different types of instruments have very different data characteristics

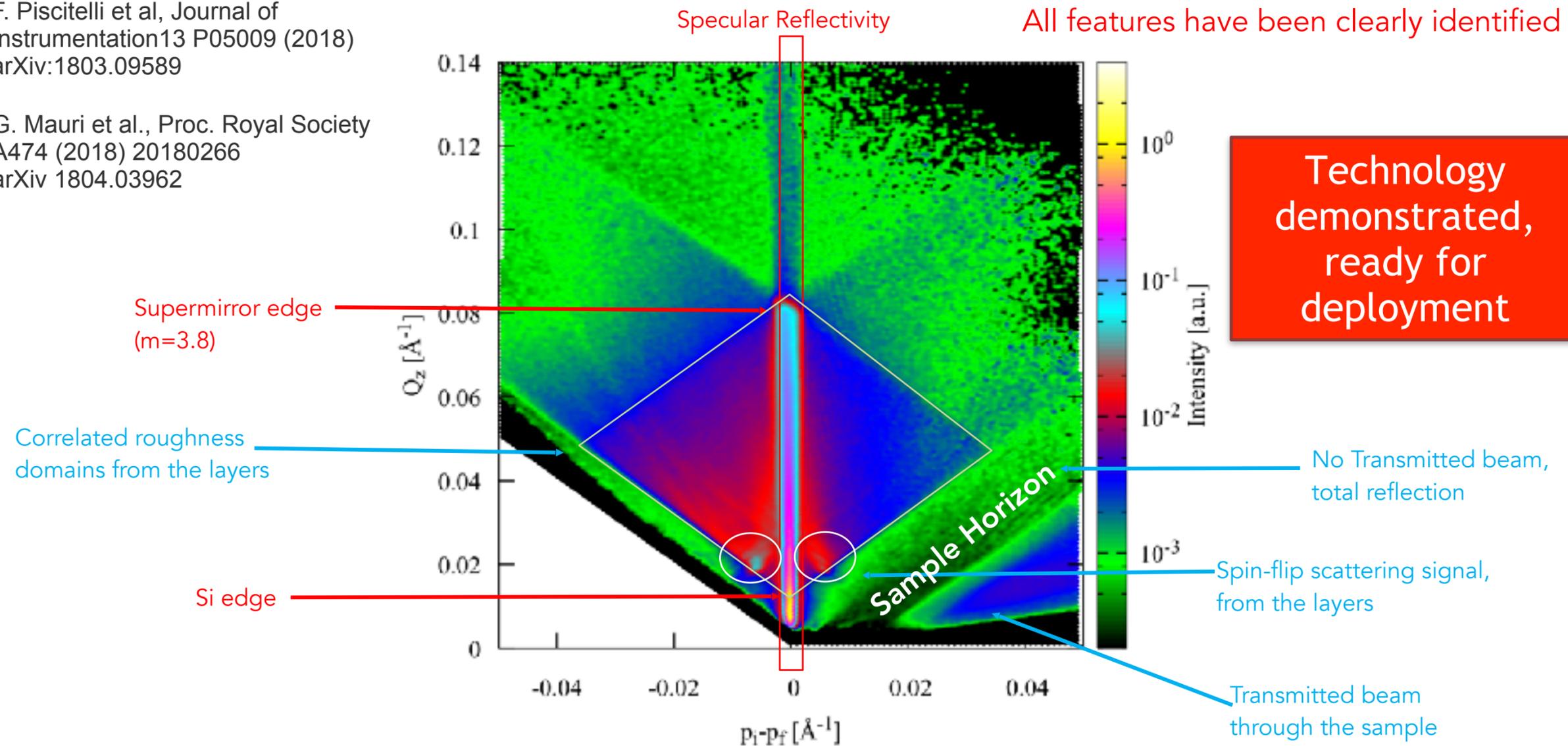
Scientific Results from CRISP: Scattering from Fe/Si Supermirror

Results

F. Piscitelli et al, Journal of Instrumentation 13 P05009 (2018)
arXiv:1803.09589

G. Mauri et al., Proc. Royal Society A474 (2018) 20180266
arXiv 1804.03962

Off-specular scattering from Fe/Si supermirror



Detector Data Acquisition System

Data Acquisition Chain for ESS Instruments



Most detectors for instruments are provided by in-kind partners

The integration of the DAQ for the instruments is done at the backend readout electronics

- Upstream the DAQ looks different for different instruments.
- In general ESS instruments have a very high number of channels
- However there was a desire to reduce the number of integrations: 3 main types.

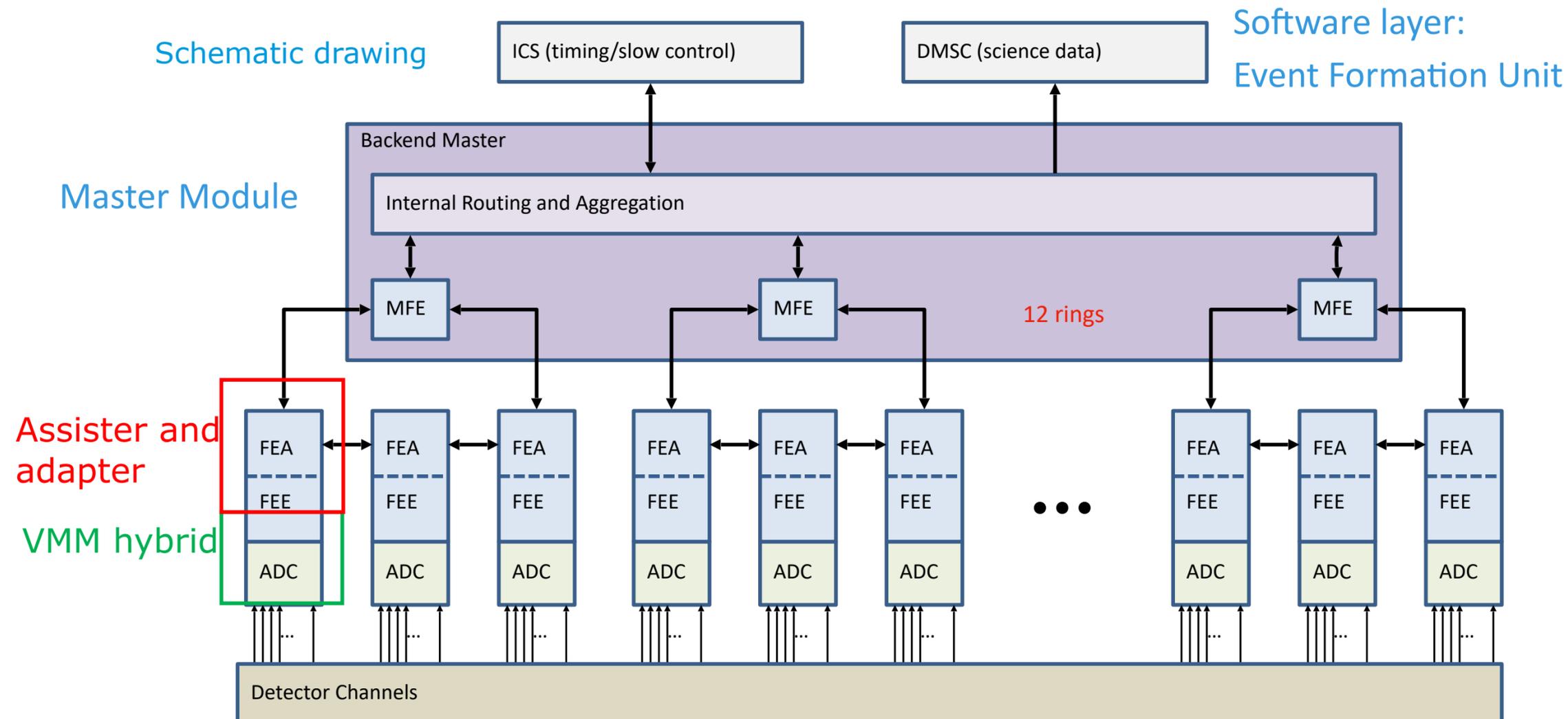
- Downstream, there is a compute layer (the EFU: event formation unit) to form the events
- Aim is to do whatever can be done in standard PCs is done (i.e. reduce development effort in the electronics).

- Integration includes facility time (ICS timing) and slow control (EPICS)

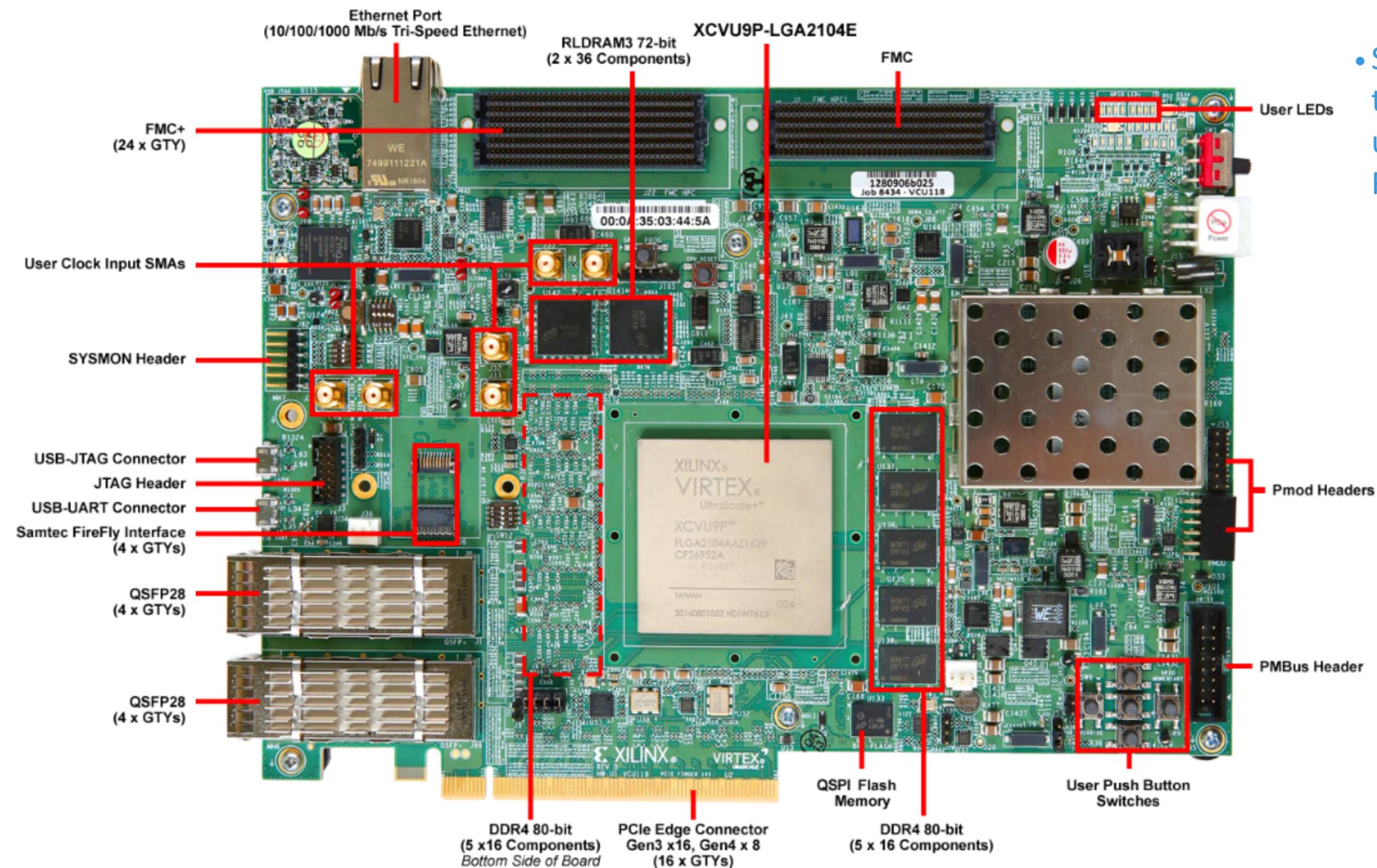
Maintenance:

- Integration needs to be as simple as possible to reduce level of effort needed.

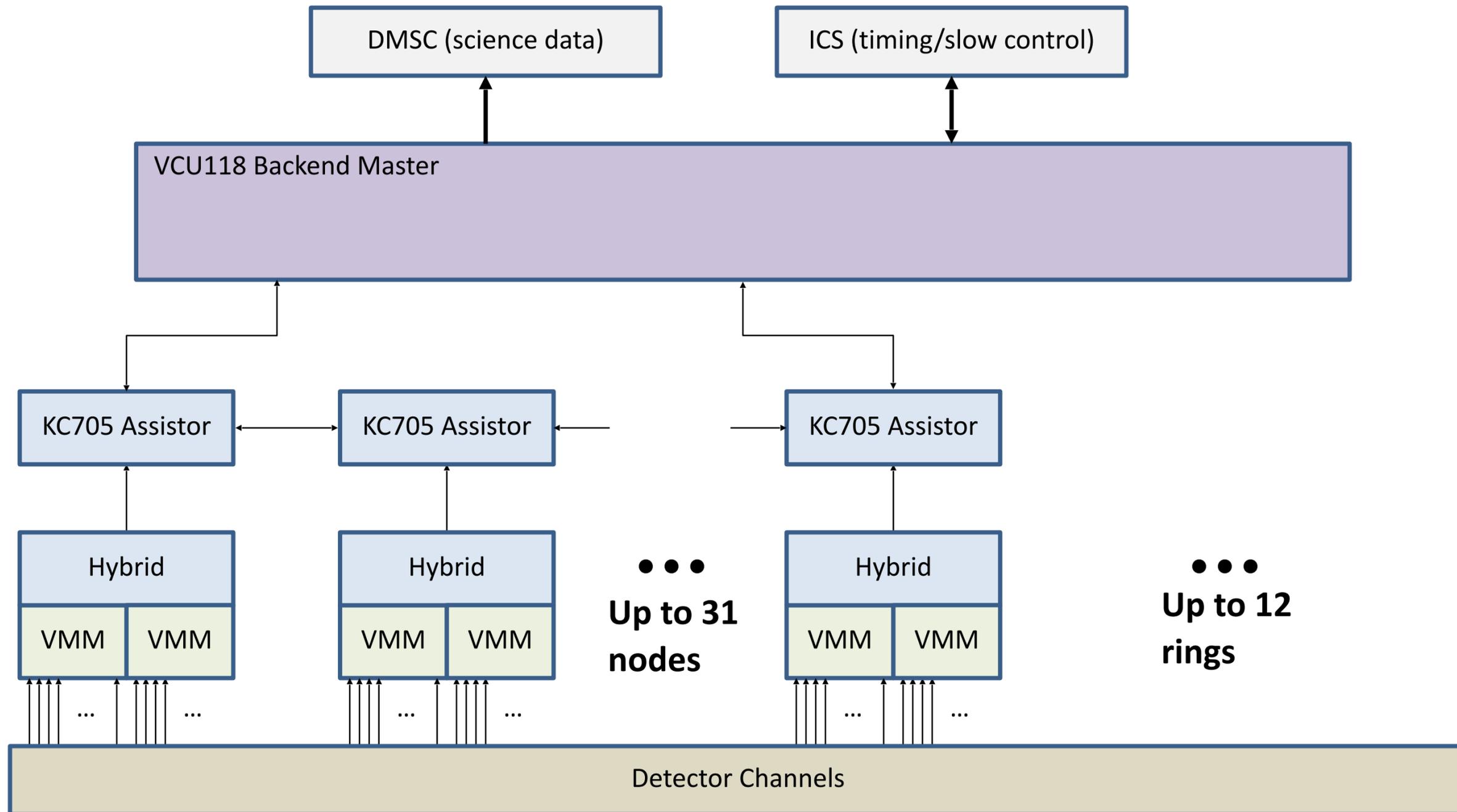
- Standardised Instrument Data Acquisition at the Electronics Backend: All instruments will use the "Master Module" using a commercial FPGA dev board (VCU118)
- Front-ends handled using 12 data rings of "assistor" boards
- Facility ("accelerator") timing distributed to the front-end via the rings



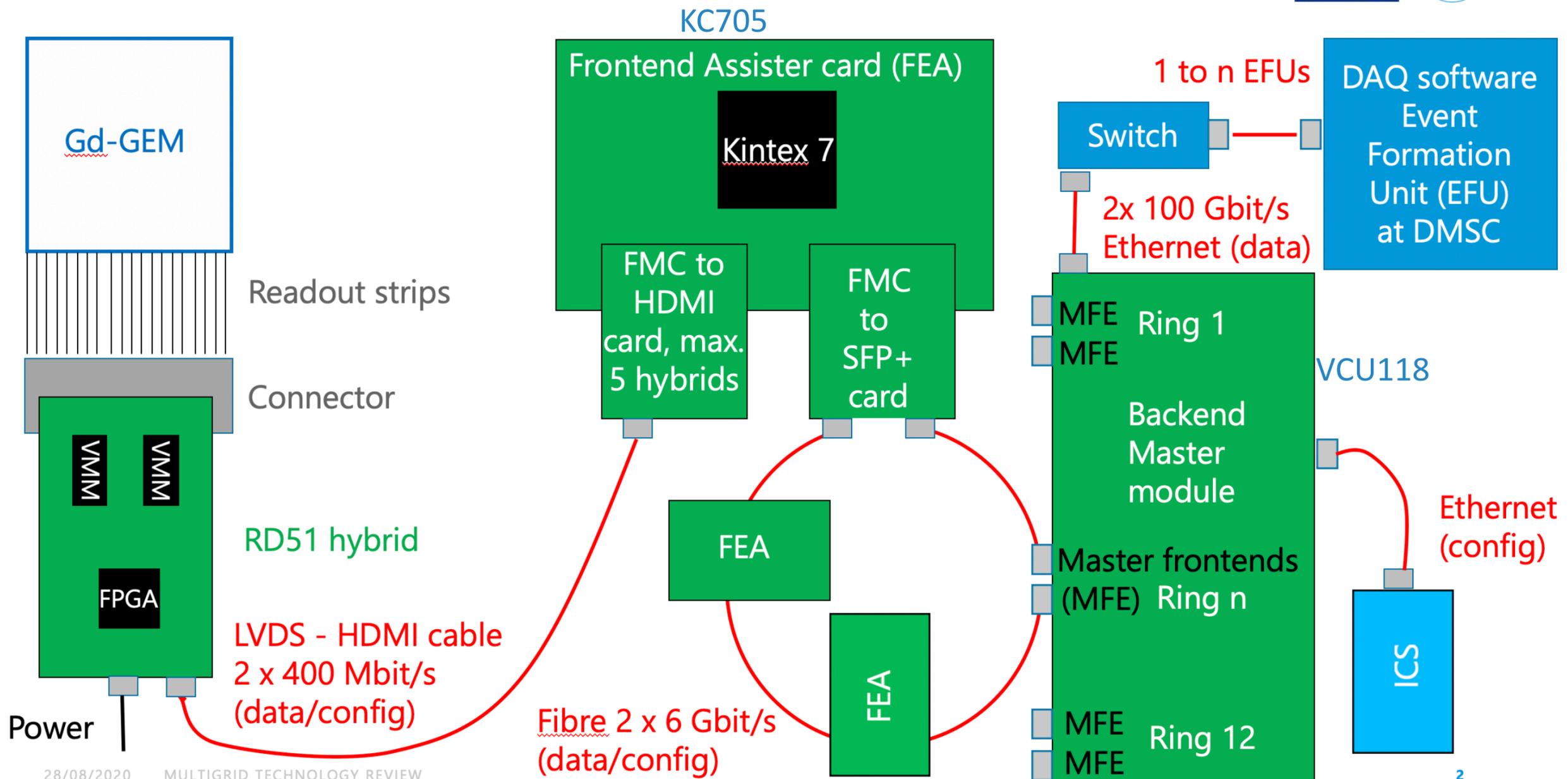
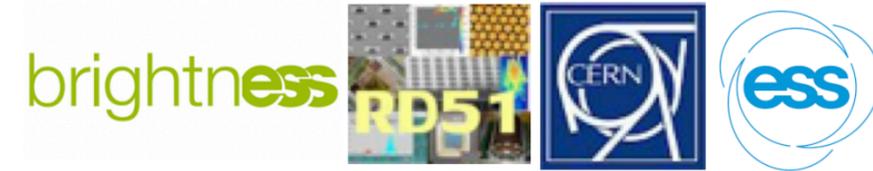
- Standardised Instrument Data Acquisition at the Electronics Backend: All instruments will use the "Master Module" using a commercial FPGA dev board (Xilinx VCU118)



ESS Readout Architecture (VMM Implementation)

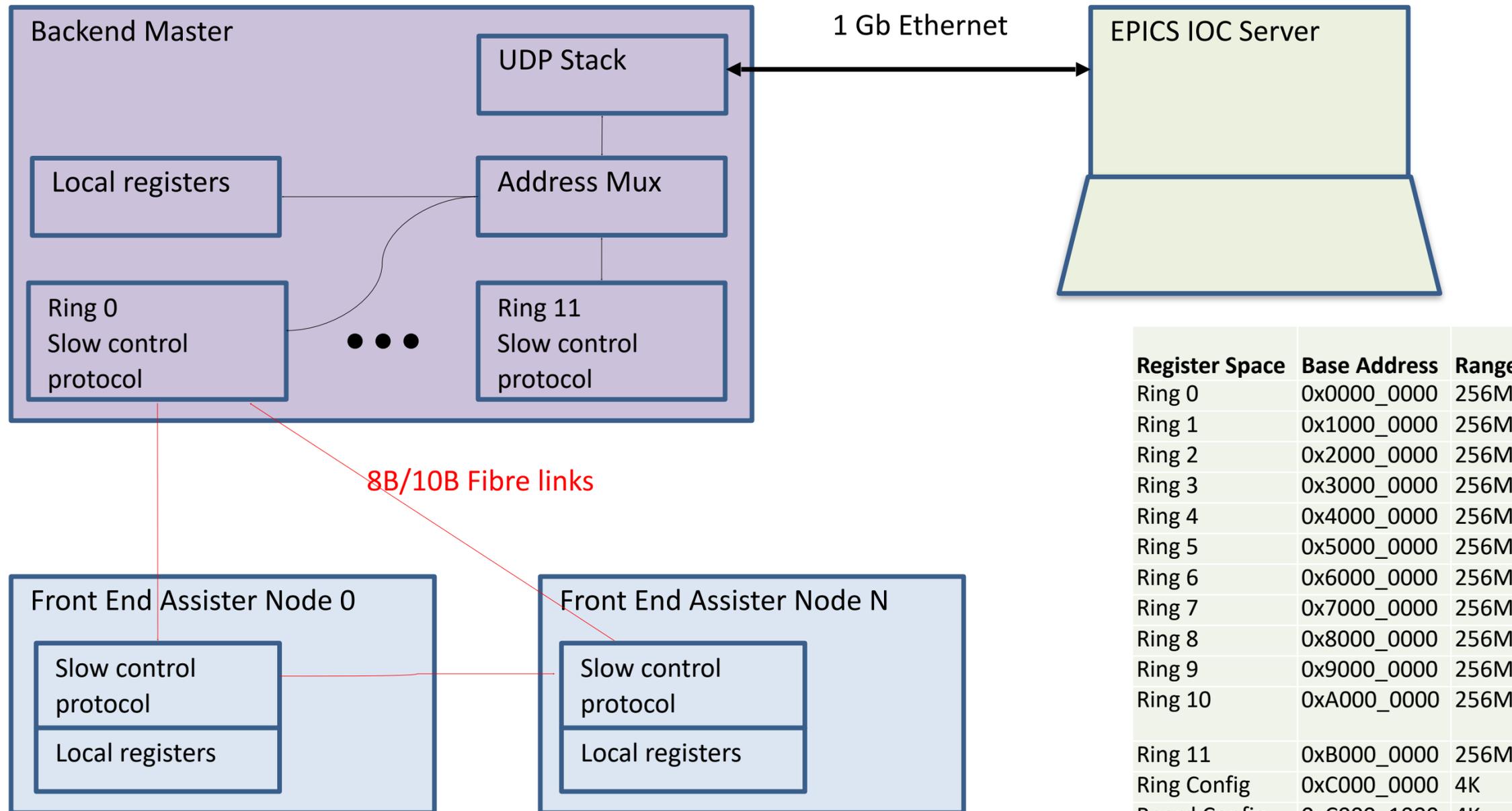


Readout chain NMX



- A very simple UDP-based Ethernet protocol has been implemented on the backend Master for configuration and monitoring of the entire readout system.
- UDP was selected instead of TCP for ease of firmware implementation. A Master-Slave configuration is adopted whereby the EPICS IOC is the Master, and every command must be acknowledged by the backend Master (which is actually the slave in this context). In other words, link reliability is achieved in the application layer of the stack.
- Each Instrument will have a dedicated IOC and a dedicated backend Master, provisioning a 32-bit address space for each readout system.
- The backend Master forwards read/write requests to the relevant register based on the address.
- To reduce cabling and simplify grounding, read/write requests to registers on front end nodes are sent over the same 8B/10B fibre used for timing distribution.
- ICS have implemented a baseline IOC which handles generic read/write requests, and are actively working on Instrument-specific functionality.

ICS EPICS Slow Control Integration 2/2



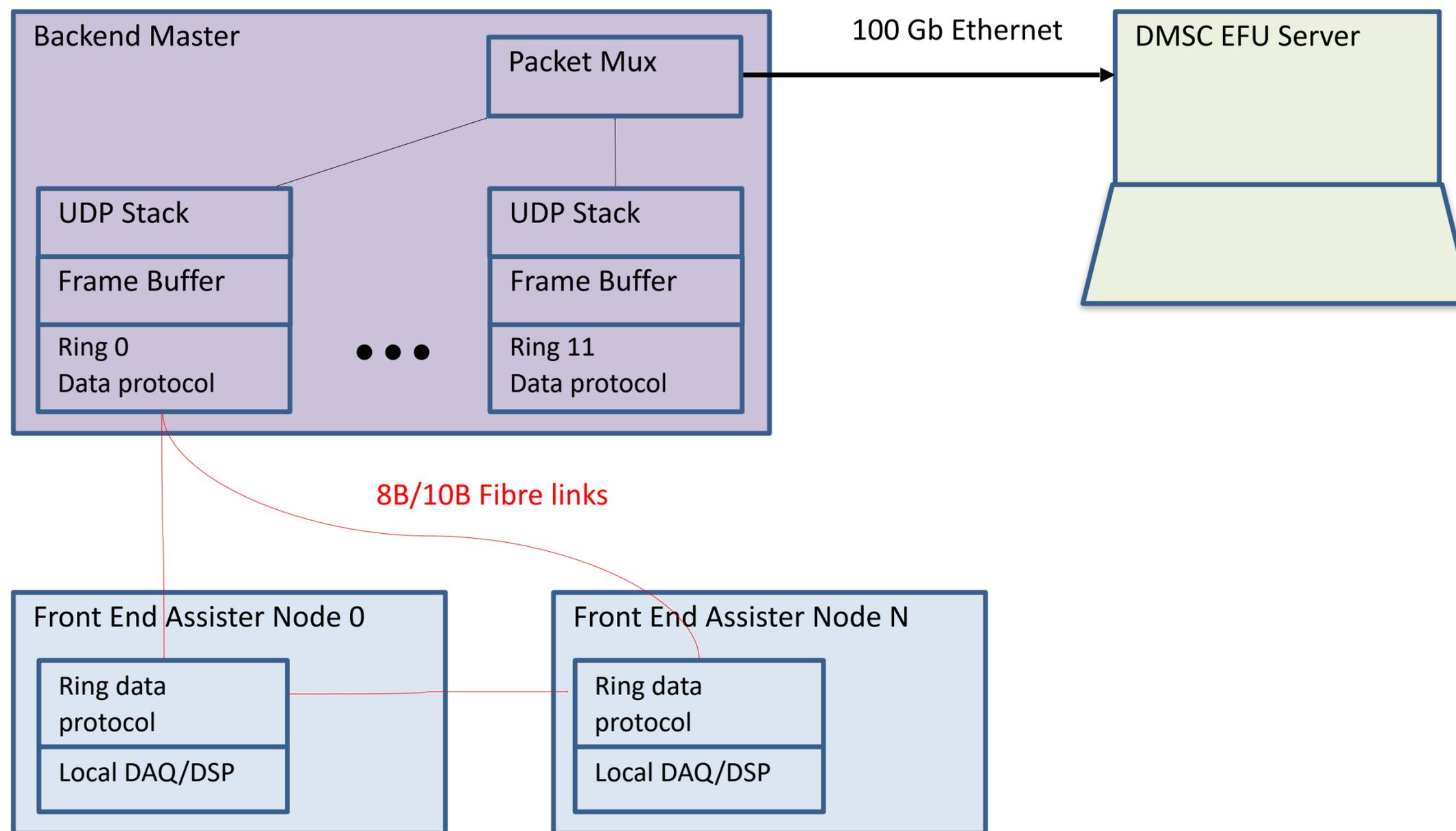
Register Space	Base Address	Range
Ring 0	0x0000_0000	256M
Ring 1	0x1000_0000	256M
Ring 2	0x2000_0000	256M
Ring 3	0x3000_0000	256M
Ring 4	0x4000_0000	256M
Ring 5	0x5000_0000	256M
Ring 6	0x6000_0000	256M
Ring 7	0x7000_0000	256M
Ring 8	0x8000_0000	256M
Ring 9	0x9000_0000	256M
Ring 10	0xA000_0000	256M
Ring 11	0xB000_0000	256M
Ring Config	0xC000_0000	4K
Board Config	0xC000_1000	4K
DMSC Config	0xC000_2000	4K
Timing Config	0xC000_3000	4K
100 G Config	0xC000_4000	8K

DMSC Neutron Event Integration 1/2



- The standardised “front end assistor” (FEA) firmware makes the ESS clock and timestamp available to the rest of the Instrument-specific “front end electronics” (FEE), which is responsible for detector data acquisition and signal processing.
- Timestamped neutron event data is transmitted over the 8B/10B fibre links to the backend Master.
- Each fibre ring can support approximately 10 Gbps: a 12-ring system can theoretically egress approximately 120 Gbps.
- The backend Master aggregates this data into jumbo frames associated with the current 14 Hz ESS accelerator pulse time, and these jumbo frames are then sent to the DMSC Event Formation Unit via UDP packets over one or two 100 Gb Ethernet links.

DMSC Neutron Event Integration 2/2



- The VMM3A is the primary choice of ASIC for the MultiBlade, MultiGrid and NMX detectors in order to satisfy high channel count requirements.
- Significant development (the VMM Hybrid) already undertaken by CERN RD51 to read out and control this ASIC, but targeting the SRS readout system.
- Best solution in terms of design-reuse is to implement the FEA logic on a separate FPGA board – a so-called “Assistor Board”:
- The current assistor board is a commercial off-the-shelf Xilinx development board (the KC705).
- The SRS FEC firmware has been ported to the KC705, allowing the VMM hybrid to be controlled and read out using existing firmware. A new interface has been written to the ESS assistor logic.
- This work does not preclude the use of custom solutions in the future with smaller size/weight/power/cost.

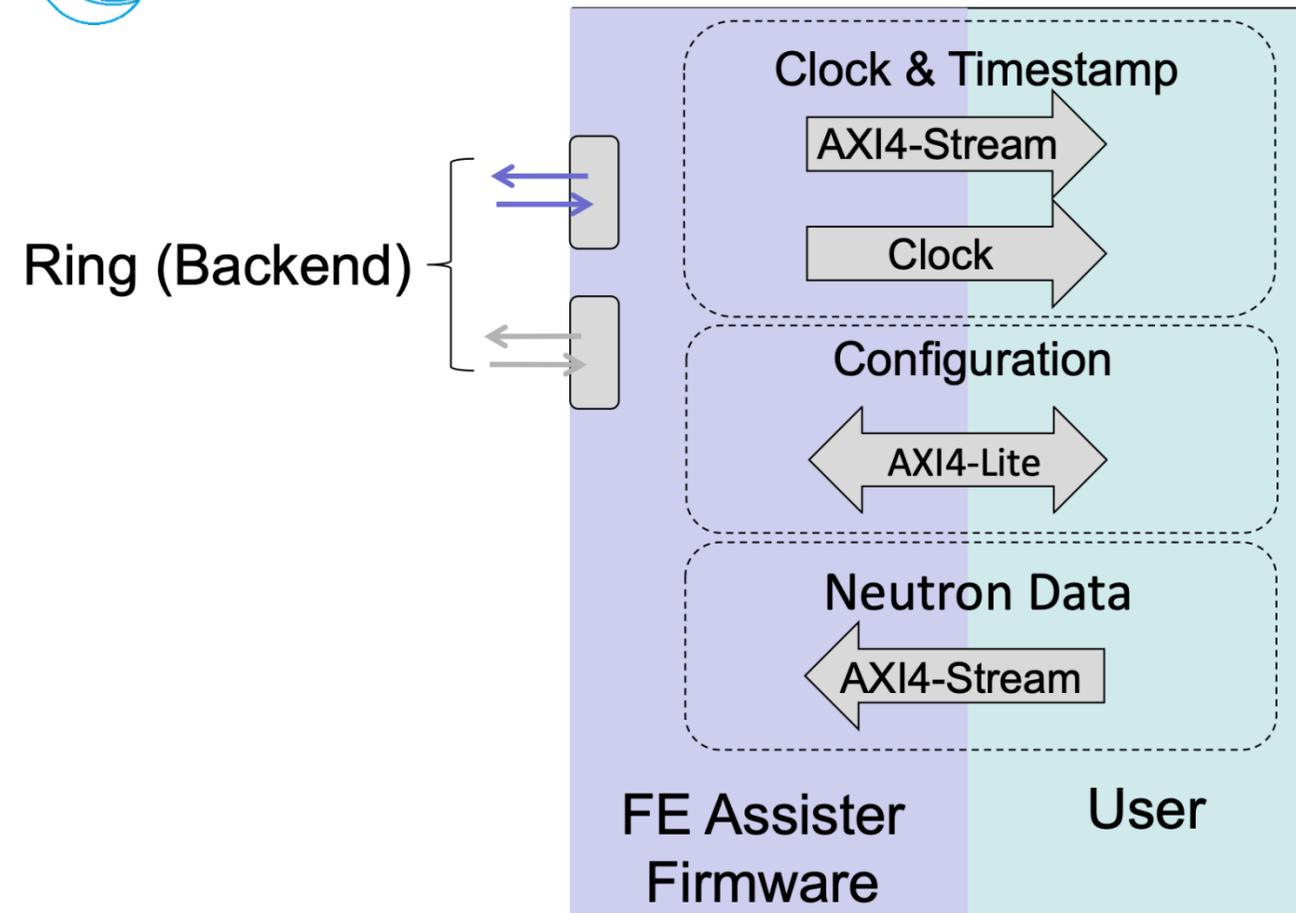
Assister Specification Document

Chess document: ESS-2055809

- FEA assister is an FPGA which glues the integration together
- FEA (front end assister) firmware communicates with the ring/backend
- FEE (front end user firmware) communicates with frontend ASIC like VMM3a or ADC
- FEA and FEE part of firmware communicate via AXI4 streams

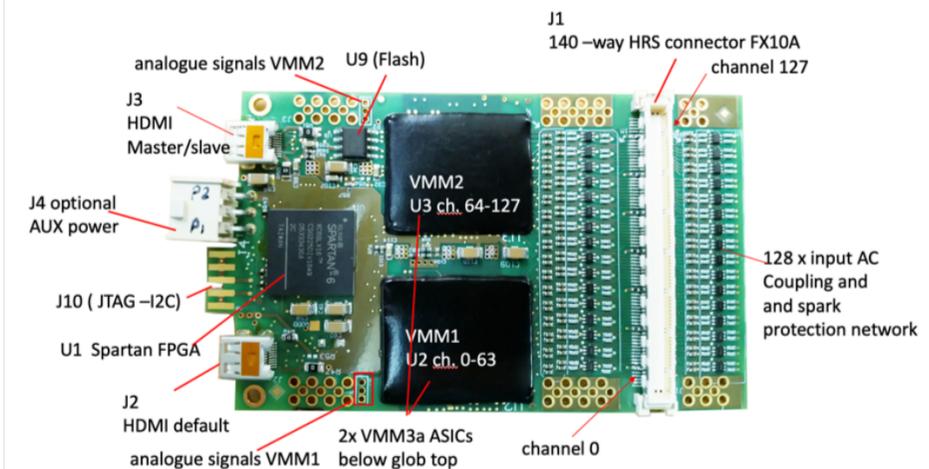


Front End User Interface

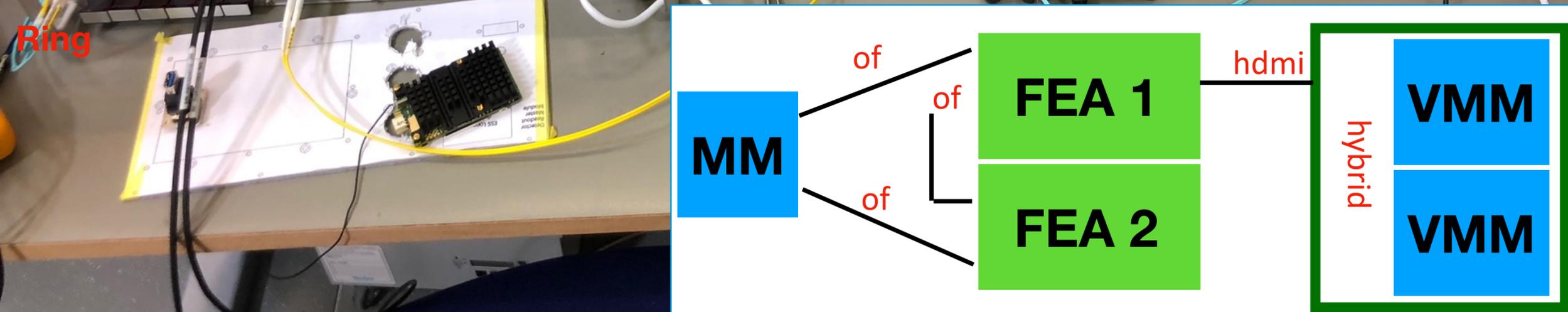
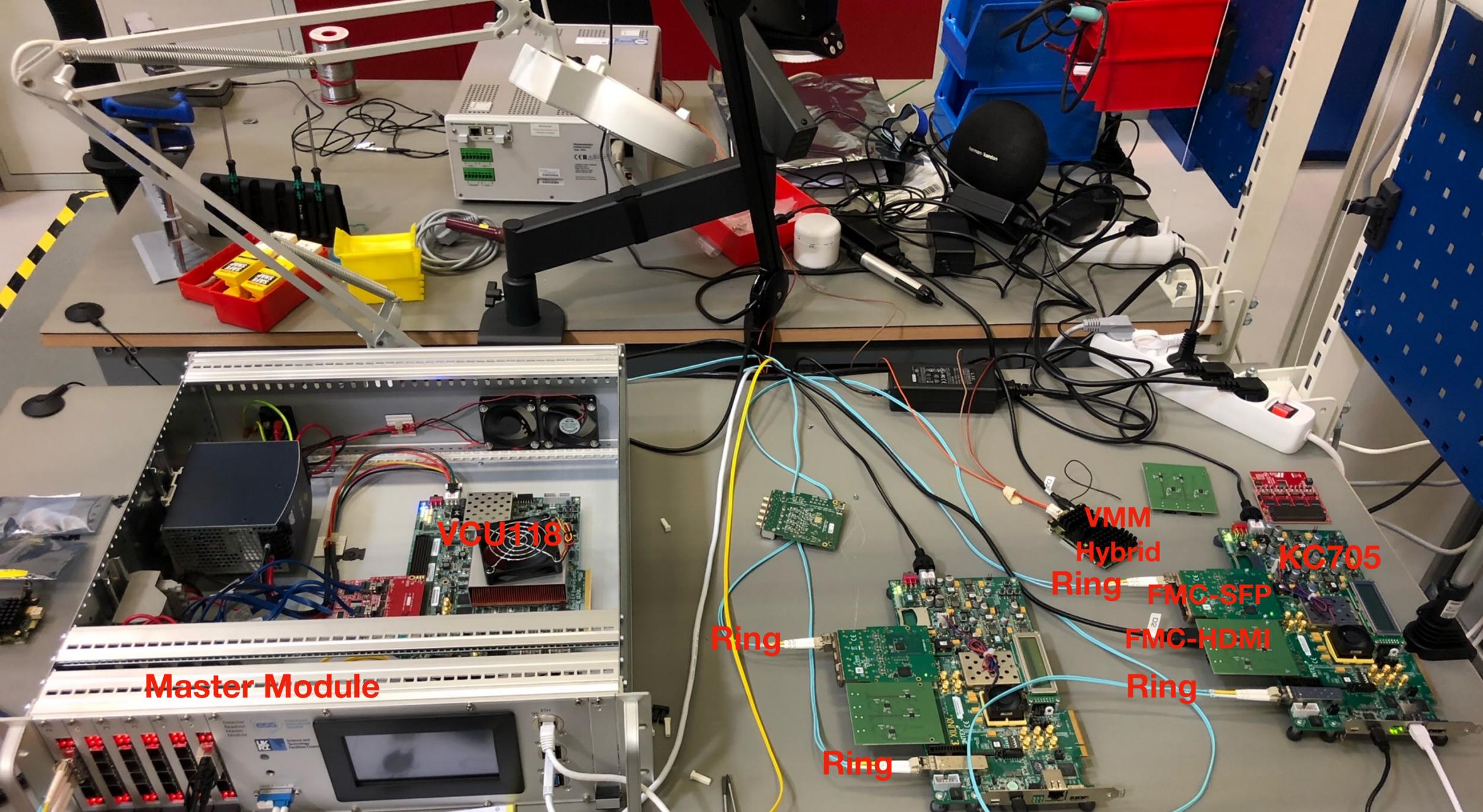


For FE integration models see
BrightnESS Deliverable Report: D4.1 – *Integration Plan
for Detector Readout*. Scott Kolva et.al. 2017

- VMM3a is the 4th version of an ASIC developed by Brookhaven National lab for the ATLAS New Small Wheel upgrade at CERN
- ASIC developed to read out Micro Pattern Gaseous detectors (MPGD)
- ASIC is high rate, sub-ns time resolution
- RD51 VMM3A hybrid common ESS-CERN project: successful integration of the VMM3a ASIC into the CERN Scalable Readout System (SRS) during BrightnESS
- 7.3 Mhits/s per VMM3a ASIC
- Per single VMM3a channel 4 Mhits/s
- Works well also for wire-based gaseous detectors

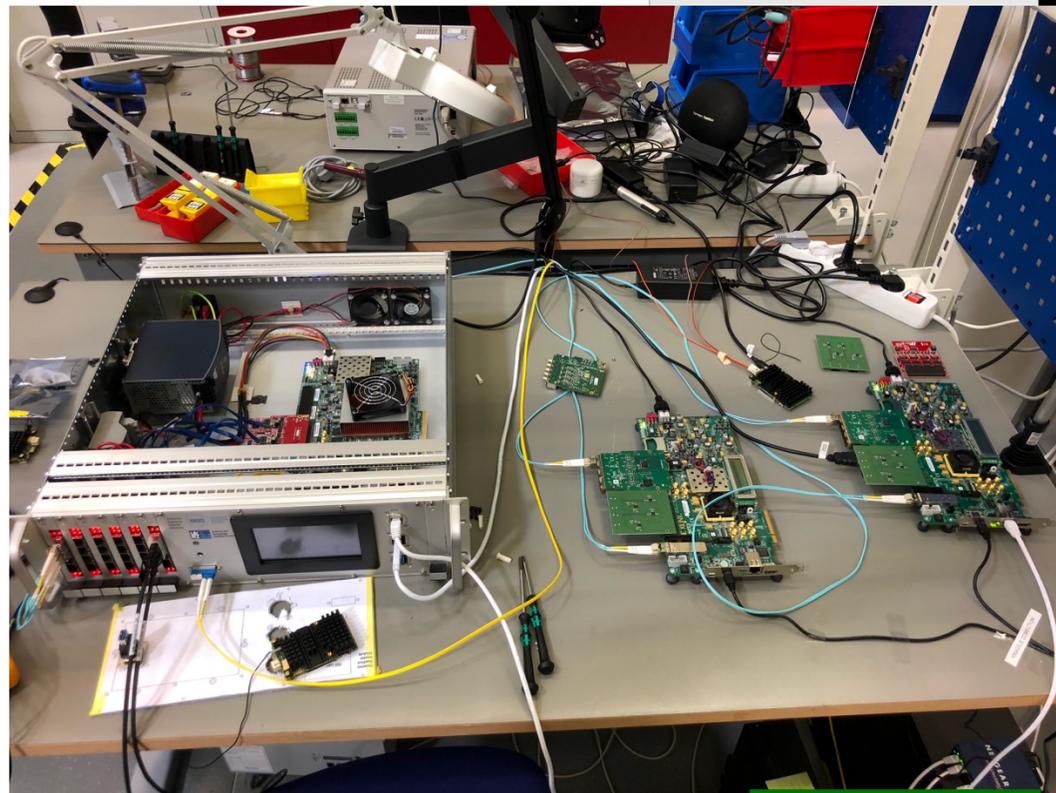


VMM Integration



VMM Integration

- Data seen from VMM input channel through to data egress from Master Module
- Through data rings
- End-end data in electronics



Remote View Bookmarks Help

Connect [Icons] Send Ctrl-Alt-Del

Applications Places Wireshark Network Analyzer

*p4p1 [Wireshark 1.10.14 (Git Rev Unknown from unknown)]

File Edit View Go Capture Analyze Statistics Telephony Tools Internals Help

Filter: frame Expression... Clear Apply Save

No.	Time	Source	Destination	Protocol	Length	Info
1	0.000000000	192.168.1.5	192.168.1.100	ESSR/VMM3A	8992	Source port: bctp Destination port: cslistener
2	0.000006145	192.168.1.2	192.168.1.100	ESSR/VMM3A	8992	Source port: bctp Destination port: cslistener
3	0.000635090	192.168.1.2	192.168.1.100	ESSR/VMM3A	8992	Source port: bctp Destination port: cslistener
4	0.000651062	192.168.1.5	192.168.1.100	ESSR/VMM3A	8992	Source port: bctp Destination port: cslistener
5	0.001287267	192.168.1.2	192.168.1.100	ESSR/VMM3A	8992	Source port: bctp Destination port: cslistener
6	0.001311423	192.168.1.5	192.168.1.100	ESSR/VMM3A	8992	Source port: bctp Destination port: cslistener
7	0.001936822	192.168.1.2	192.168.1.100	ESSR/VMM3A	8992	Source port: bctp Destination port: cslistener
8	0.001953035	192.168.1.5	192.168.1.100	ESSR/VMM3A	8992	Source port: bctp Destination port: cslistener
9	0.002587484	192.168.1.2	192.168.1.100	ESSR/VMM3A	8992	Source port: bctp Destination port: cslistener
10	0.002607622	192.168.1.5	192.168.1.100	ESSR/VMM3A	8992	Source port: bctp Destination port: cslistener
11	0.003238650	192.168.1.2	192.168.1.100	ESSR/VMM3A	8992	Source port: bctp Destination port: cslistener
12	0.003241928	192.168.1.5	192.168.1.100	ESSR/VMM3A	8992	Source port: bctp Destination port: cslistener
13	0.003889807	192.168.1.2	192.168.1.100	ESSR/VMM3A	8992	Source port: bctp Destination port: cslistener

Frame 1: 8992 bytes on wire (71936 bits), 8992 bytes captured (71936 bits) on interface 0

Ethernet II, Src: 0e:05:05:00:00:05 (0e:05:05:00:00:05), Dst: 04:3f:72:f2:f1:32 (04:3f:72:f2:f1:32)

Internet Protocol Version 4, Src: 192.168.1.5 (192.168.1.5), Dst: 192.168.1.100 (192.168.1.100)

User Datagram Protocol, Src Port: bctp (8999), Dst Port: cslistener (9000)

ESSR Header

Readout 1, Ring 5, FEN 1, VMM: 1, CH:56, Time 1614073922 s 96146358.75 ns, Overflow 1043, BC 103, OTHR 1, ADC 145, TDC:104 GEO 0

Readout 2, Ring 5, FEN 1, VMM: 0, CH:45, Time 1614073922 s 96146358.75 ns, Overflow 1043, BC 103, OTHR 1, ADC 144, TDC:114 GEO 0

Readout 3, Ring 5, FEN 1, VMM: 1, CH:59, Time 1614073922 s 96146358.75 ns, Overflow 1043, BC 103, OTHR 1, ADC 131, TDC:104 GEO 0

Readout 4, Ring 5, FEN 1, VMM: 1, CH:60, Time 1614073922 s 96146358.75 ns, Overflow 1043, BC 103, OTHR 1, ADC 140, TDC:113 GEO 0

Readout 5, Ring 5, FEN 1, VMM: 0, CH:47, Time 1614073922 s 96146358.75 ns, Overflow 1043, BC 103, OTHR 1, ADC 142, TDC:120 GEO 0

Readout 6, Ring 5, FEN 1, VMM: 1, CH:63, Time 1614073922 s 96146358.75 ns, Overflow 1043, BC 103, OTHR 1, ADC 127, TDC:105 GEO 0

Readout 7, Ring 5, FEN 1, VMM: 0, CH:49, Time 1614073922 s 96146358.75 ns, Overflow 1043, BC 103, OTHR 1, ADC 140, TDC:111 GEO 0

Readout 8, Ring 5, FEN 1, VMM: 0, CH:51, Time 1614073922 s 96146358.75 ns, Overflow 1043, BC 103, OTHR 1, ADC 136, TDC:128 GEO 0

Readout 9, Ring 5, FEN 1, VMM: 0, CH:53, Time 1614073922 s 96146358.75 ns, Overflow 1043, BC 103, OTHR 1, ADC 137, TDC:119 GEO 0

Readout 10, Ring 5, FEN 1, VMM: 0, CH:55, Time 1614073922 s 96146358.75 ns, Overflow 1043, BC 103, OTHR 1, ADC 128, TDC:115 GEO 0

Readout 11, Ring 5, FEN 1, VMM: 0, CH:57, Time 1614073922 s 96146358.75 ns, Overflow 1043, BC 103, OTHR 1, ADC 128, TDC:133 GEO 0

Readout 12, Ring 5, FEN 1, VMM: 0, CH:59, Time 1614073922 s 96146358.75 ns, Overflow 1043, BC 103, OTHR 1, ADC 128, TDC:137 GEO 0

Readout 13, Ring 5, FEN 1, VMM: 0, CH:61, Time 1614073922 s 96146358.75 ns, Overflow 1043, BC 103, OTHR 1, ADC 132, TDC:115 GEO 0

Readout 14, Ring 5, FEN 1, VMM: 0, CH:63, Time 1614073922 s 96146336.25 ns, Overflow 1043, BC 102, OTHR 1, ADC 124, TDC: 64 GEO 0

Readout 15, Ring 5, FEN 1, VMM: 1, CH: 0, Time 1614073922 s 96238518.75 ns, Overflow 1044, BC 103, OTHR 1, ADC 144, TDC:122 GEO 0

Readout 16, Ring 5, FEN 1, VMM: 1, CH: 1, Time 1614073922 s 96238518.75 ns, Overflow 1044, BC 103, OTHR 1, ADC 138, TDC:119 GEO 0

Readout 17, Ring 5, FEN 1, VMM: 1, CH: 2, Time 1614073922 s 96238518.75 ns, Overflow 1044, BC 103, OTHR 1, ADC 134, TDC:115 GEO 0

Readout 18, Ring 5, FEN 1, VMM: 1, CH: 3, Time 1614073922 s 96238518.75 ns, Overflow 1044, BC 103, OTHR 1, ADC 152, TDC:113 GEO 0

Readout 19, Ring 5, FEN 1, VMM: 1, CH: 4, Time 1614073922 s 96238518.75 ns, Overflow 1044, BC 103, OTHR 1, ADC 146, TDC:112 GEO 0

Readout 20, Ring 5, FEN 1, VMM: 1, CH: 5, Time 1614073922 s 96238518.75 ns, Overflow 1044, BC 103, OTHR 1, ADC 151, TDC:129 GEO 0

```

0000 04 3f 72 f2 f1 32 0e 05 05 00 00 05 08 00 45 00  .?r..2.. ..E.
0010 23 12 00 00 00 00 05 11 0f 22 c0 a8 01 05 c0 a8  #.....".....
0020 01 64 23 27 23 28 22 fe 00 00 00 00 45 53 53 48  .d#("#. ...ESSH
0030 f6 22 05 02 42 d0 34 60 33 78 60 00 42 d0 34 60  .".B.4` 3x`.B.4`
0040 f9 7f 00 00 53 20 00 00 05 01 14 00 42 d0 34 60  ....S .. ...B.4`
0050 27 68 82 00 67 00 91 80 00 68 01 38 05 01 14 00  'h.g... .h.8....
0060 42 d0 34 60 27 68 82 00 67 00 90 80 00 72 00 2d  B.4`h.. g....r.-
0070 05 01 14 00 42 d0 34 60 27 68 82 00 67 00 83 80  ....B.4` 'h.g...

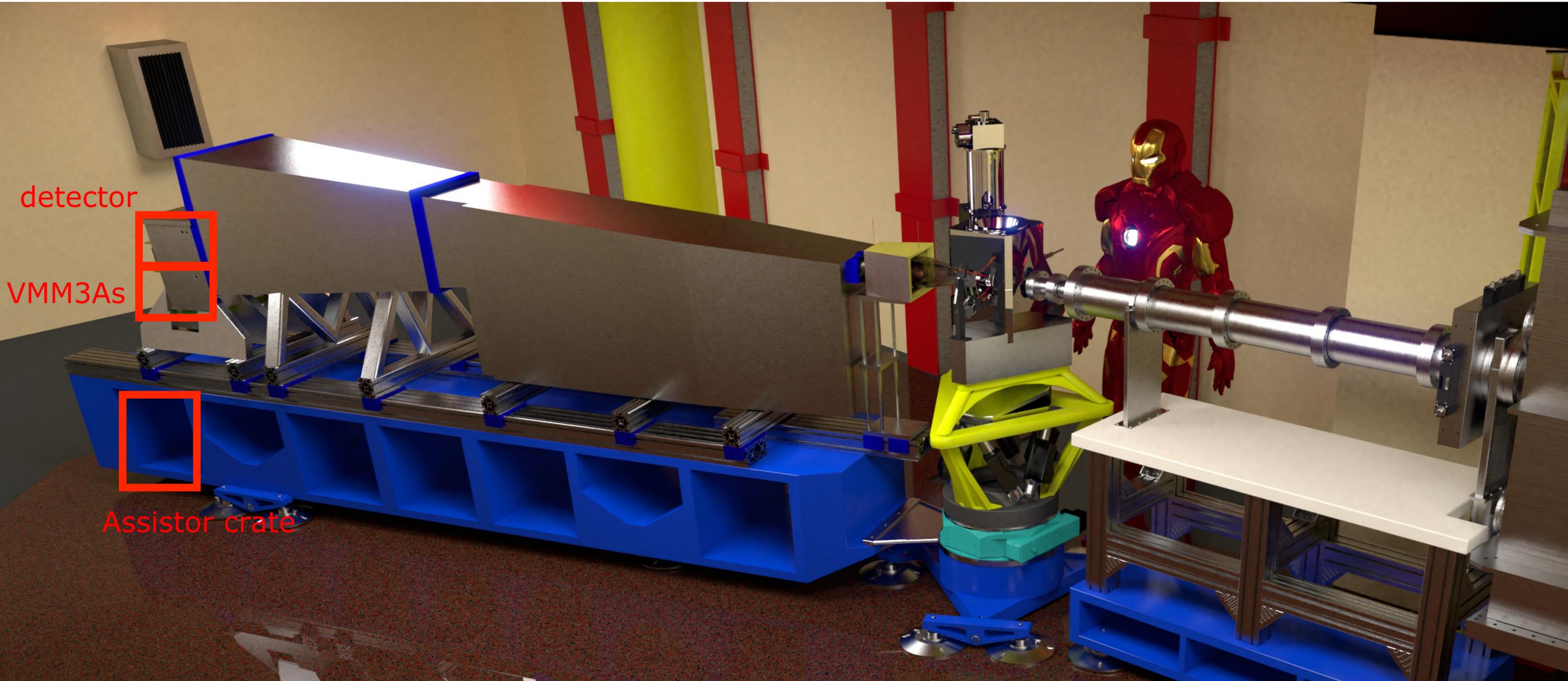
```

File: /tmp/wireshark_pcapng_p4p1... Packets: 400 · Displayed: 400 (100.0%) · Dropped: 0 (0.0%)

stevencock@dgro-vmm4:~/g... [Calculator] *p3p1 [Wireshark 1.10.14 (...)] VMM3a.pcapng [Wireshark 1....] Vivado Lab Edition 2019.1 *p4p1 [Wireshark 1.10.14 (...)]

Engineering the whole system ...
It's not just digital ...
... or only about data transport and manipulation ...

MultiBlade

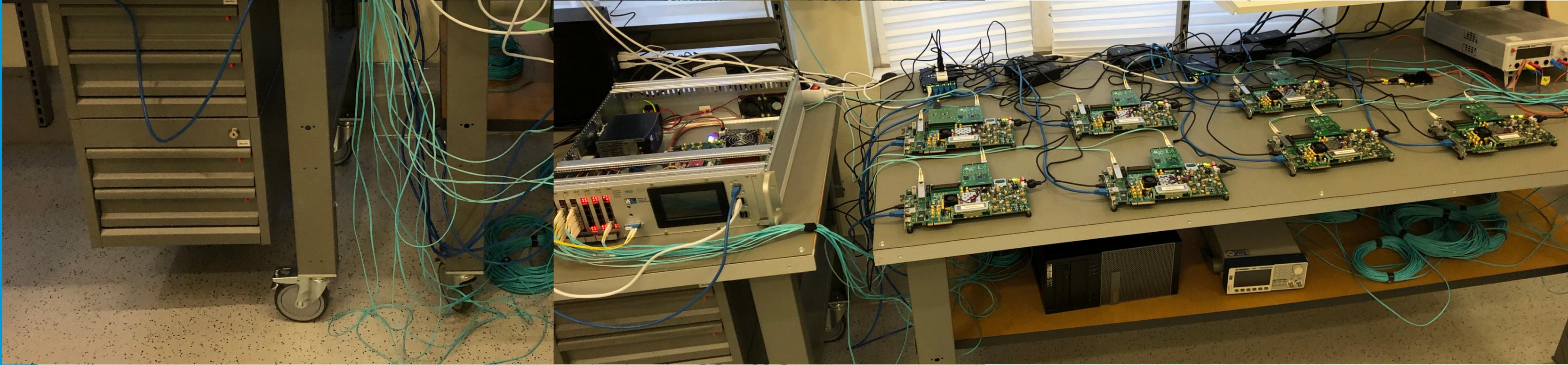


ca. 1-2m services path from VMM3As to Assistor crate
Detector rack ca. 15-20m distant

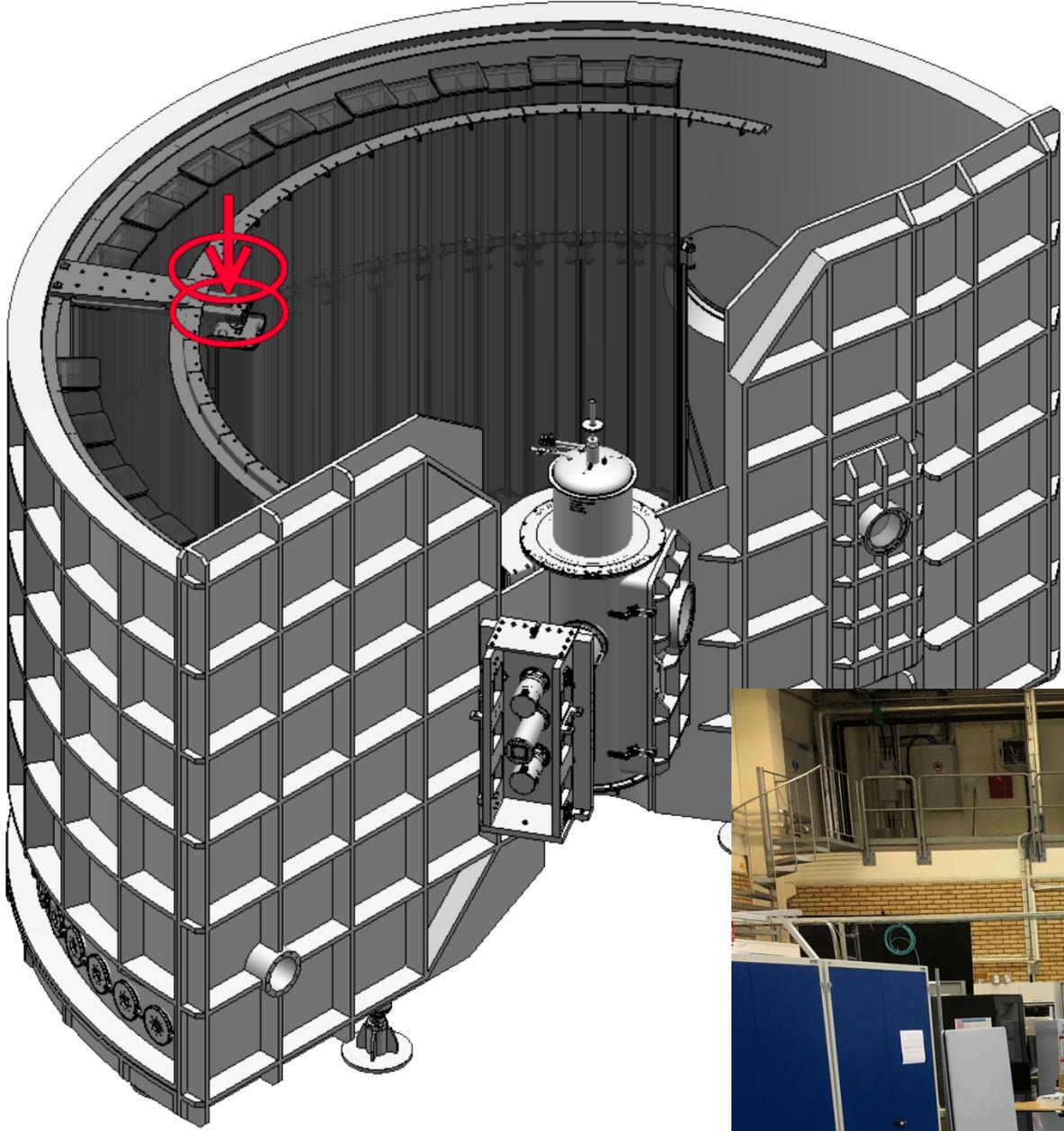
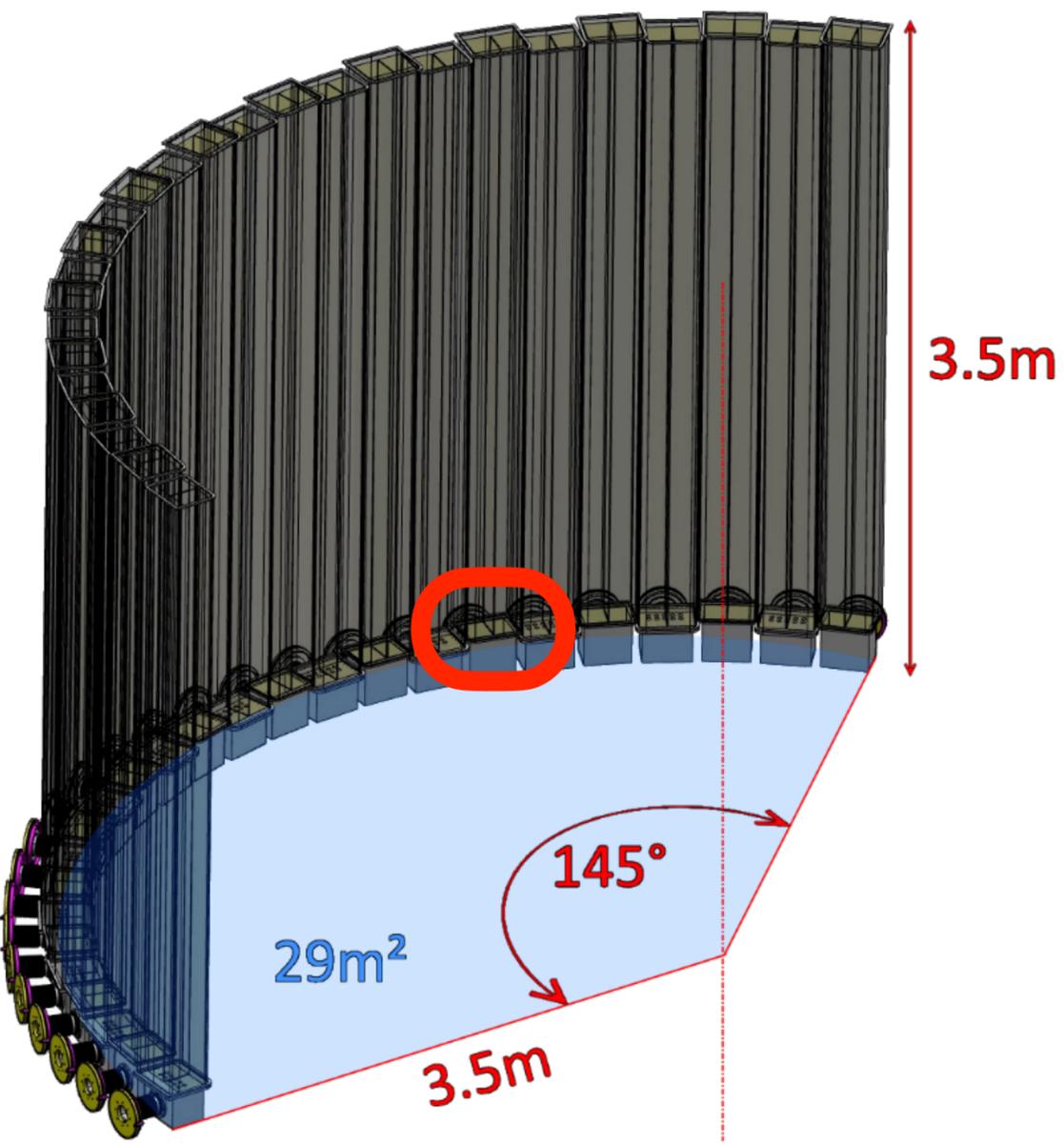
Backend Readout Electronics



Electronics DAQ system works
Electronics DAQ system scales



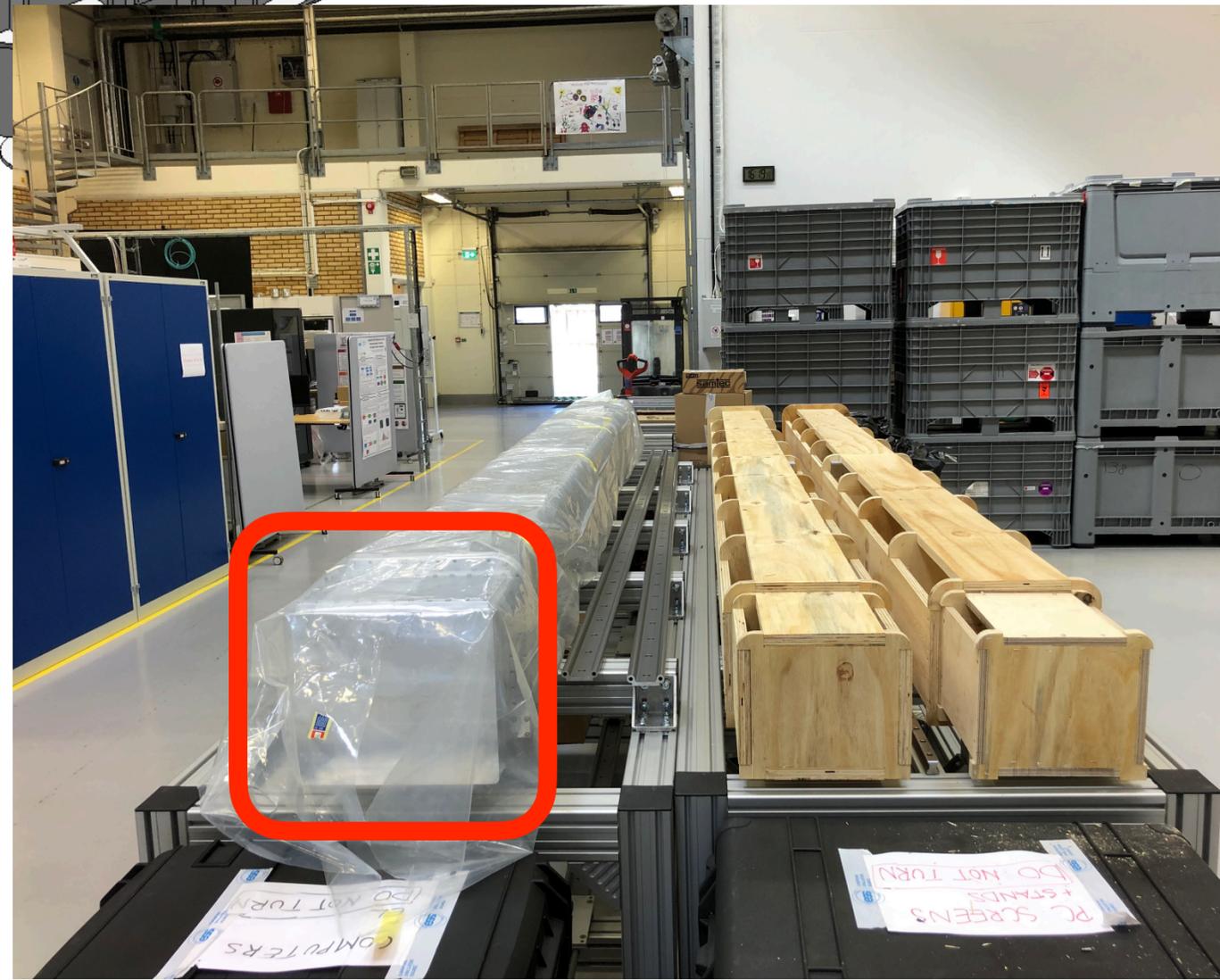
MultiGrid



VMM3A and assister inside electronics box

ca. 50cm services path from electronics box to outside world.

Detector rack ca. 10-15m distant



Eg CSPEC instrument

CSPEC D

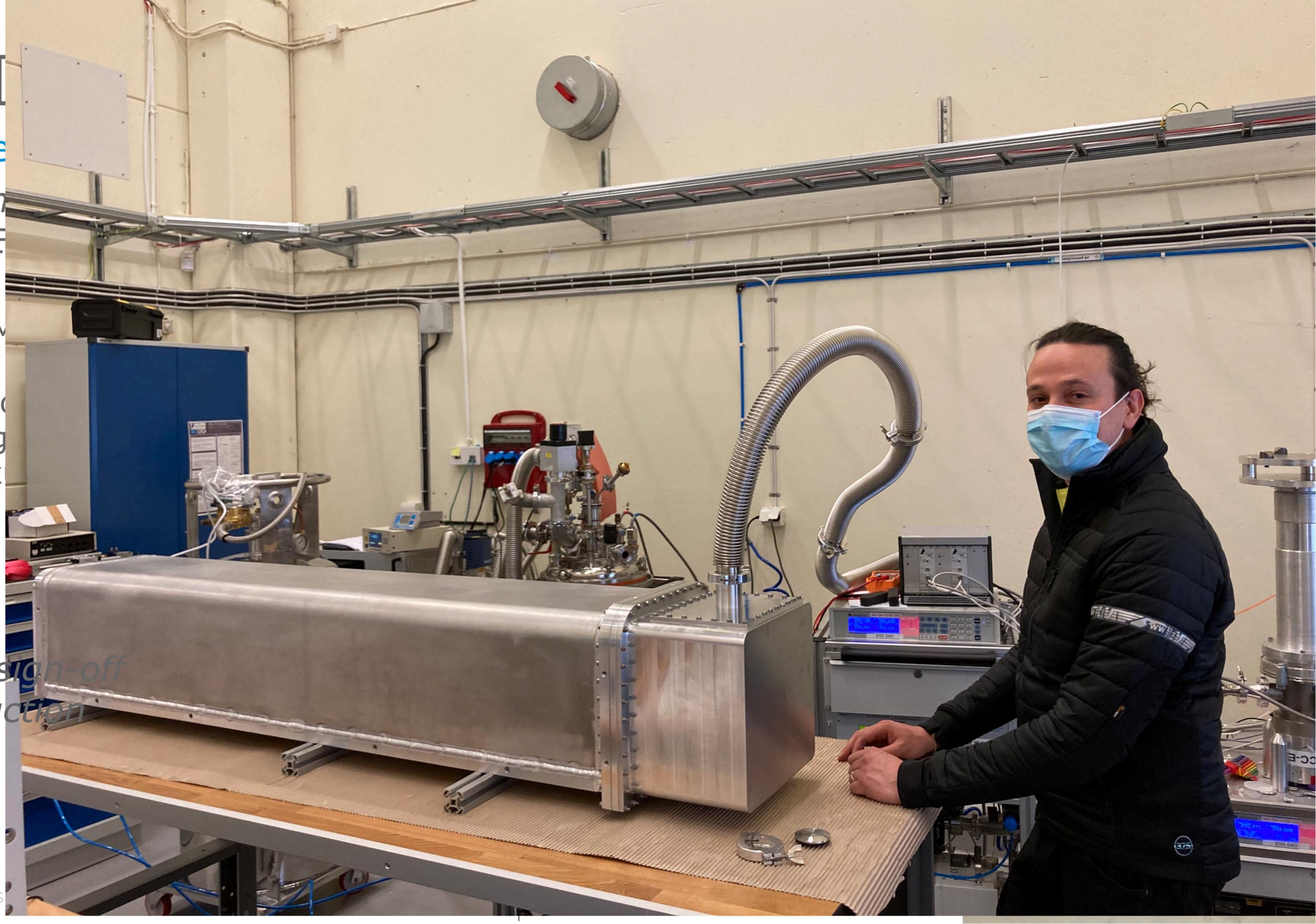
MultiGrid detector

*Aim: Testing h
detector on LE*

Status:

2x2 day test in M
Vessel ready
Coatings, Blade o
Preparation for g
Successful mock

*Aim: final sign-off
for construction*



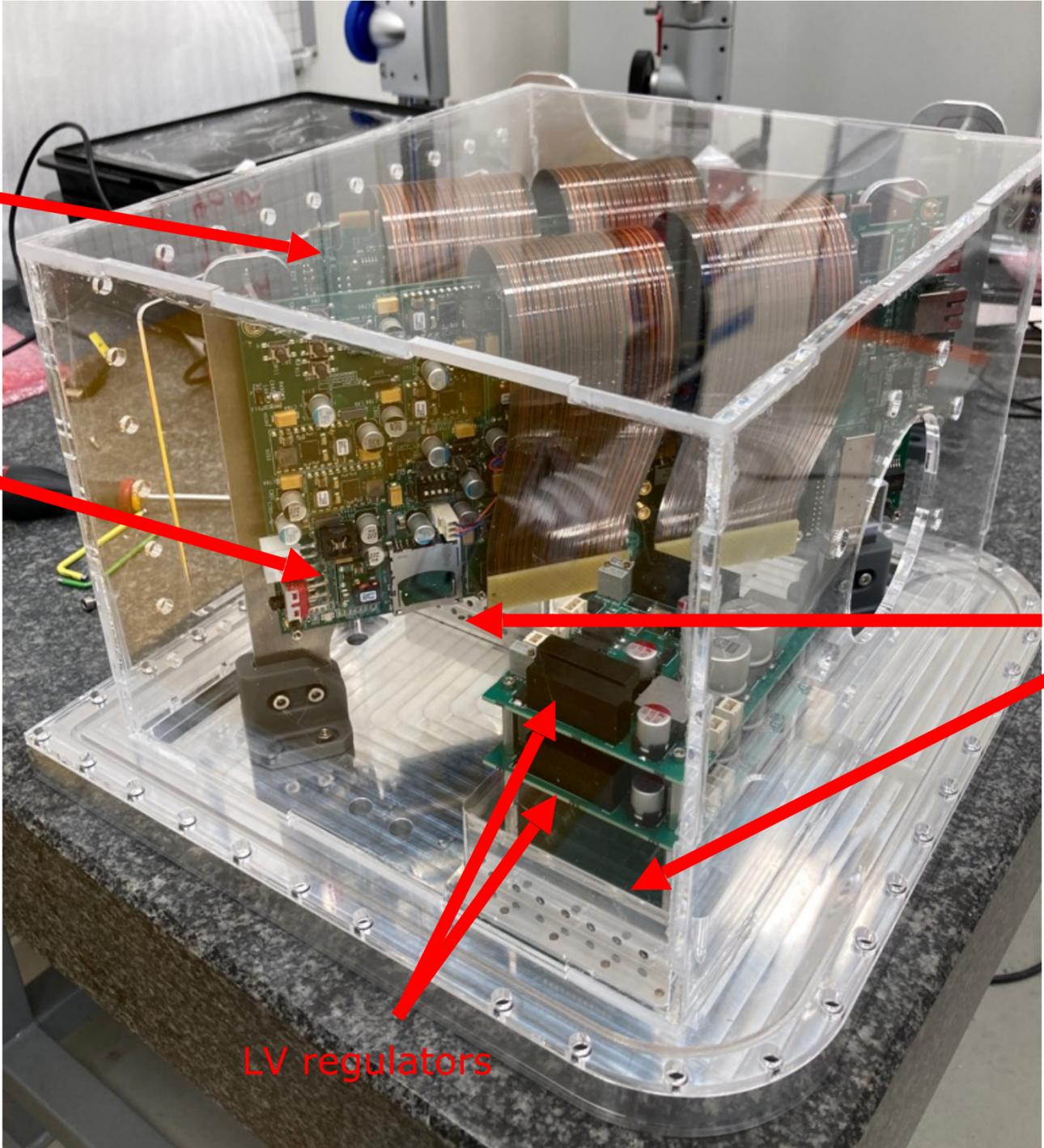
MultiGrid Electronics Box

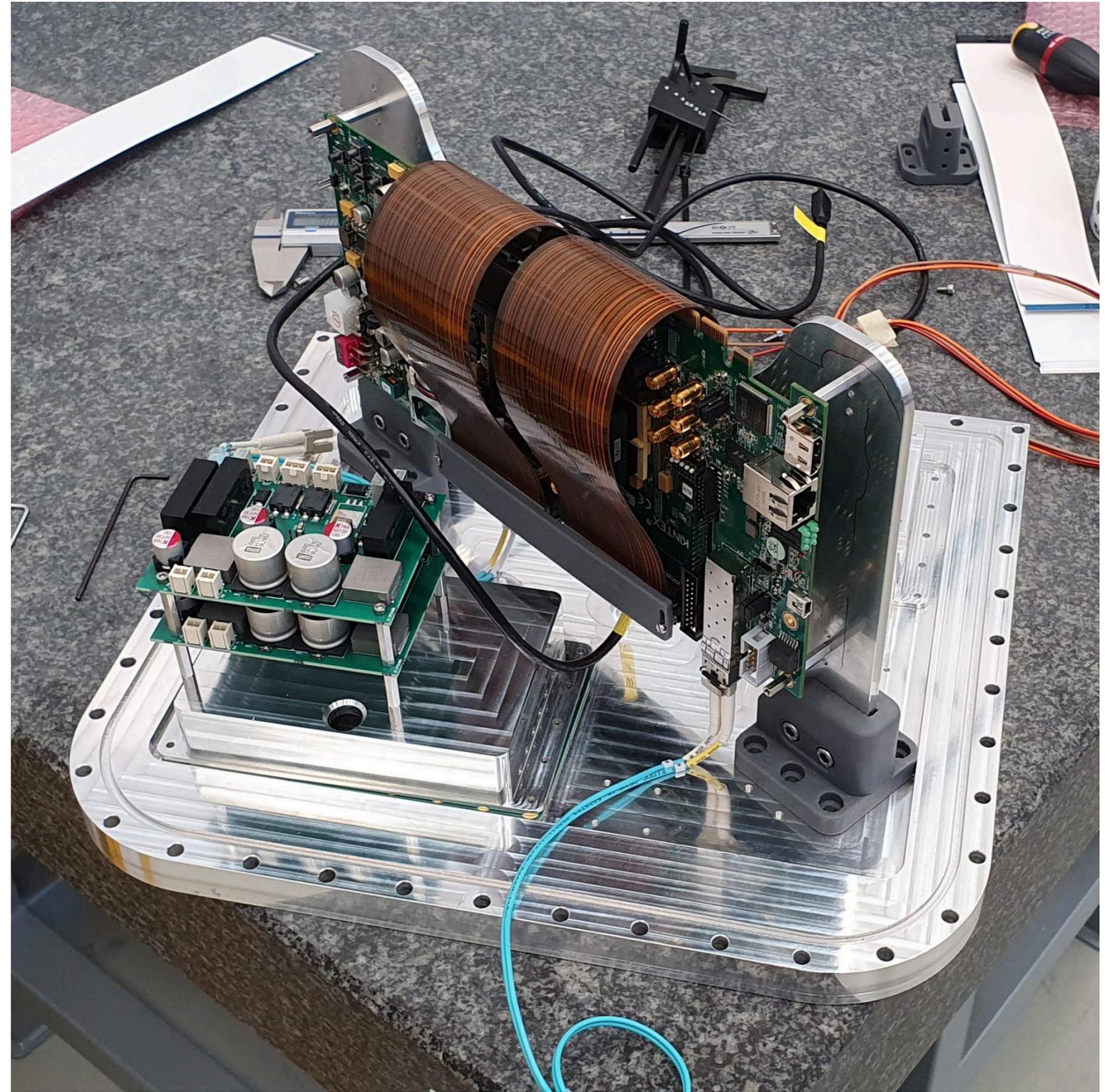
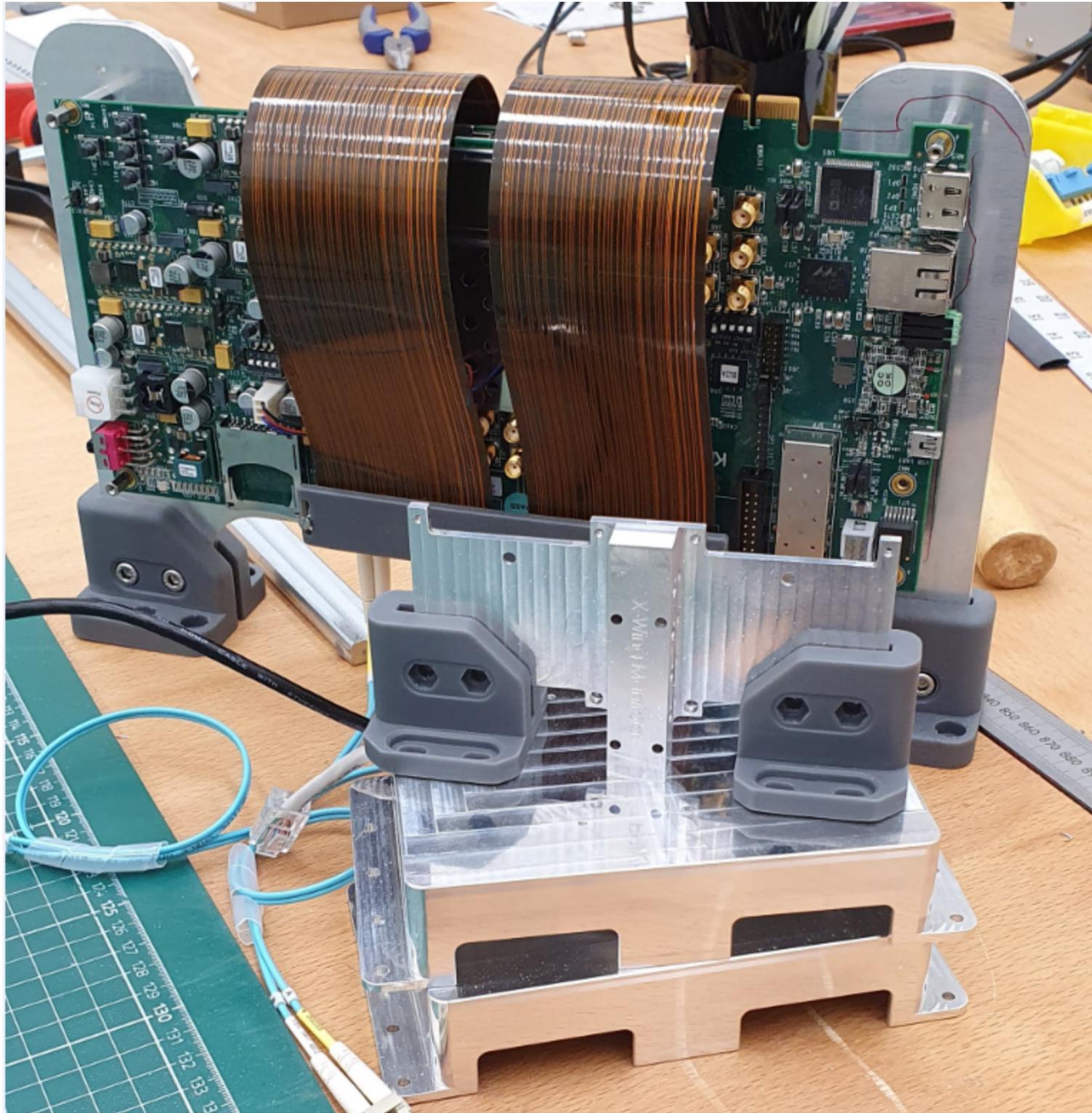
Assistor board 2

Assistor board 1

EMI shielding on top of 2x VMM hybrids

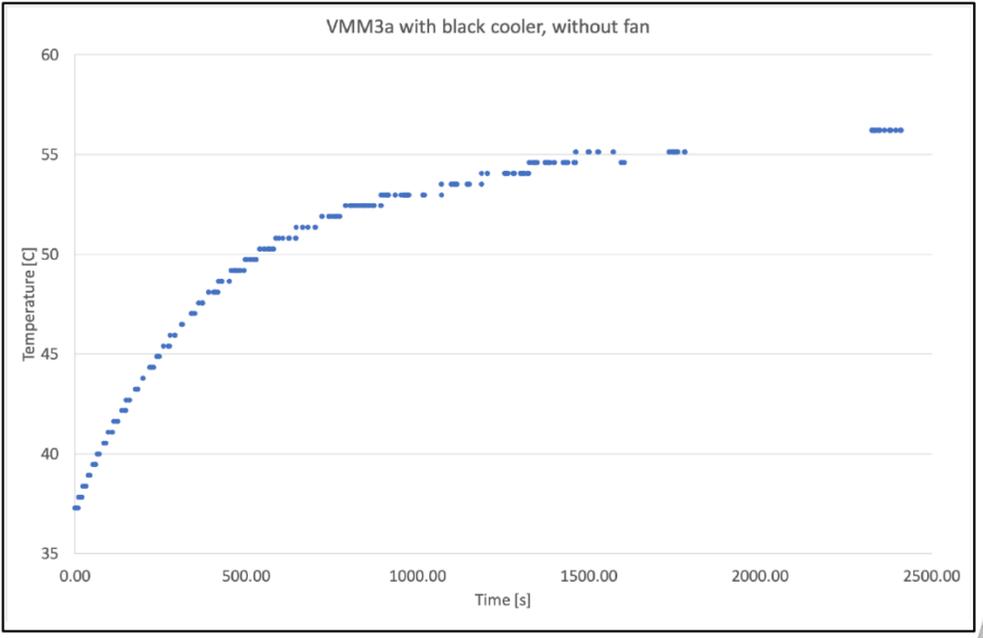
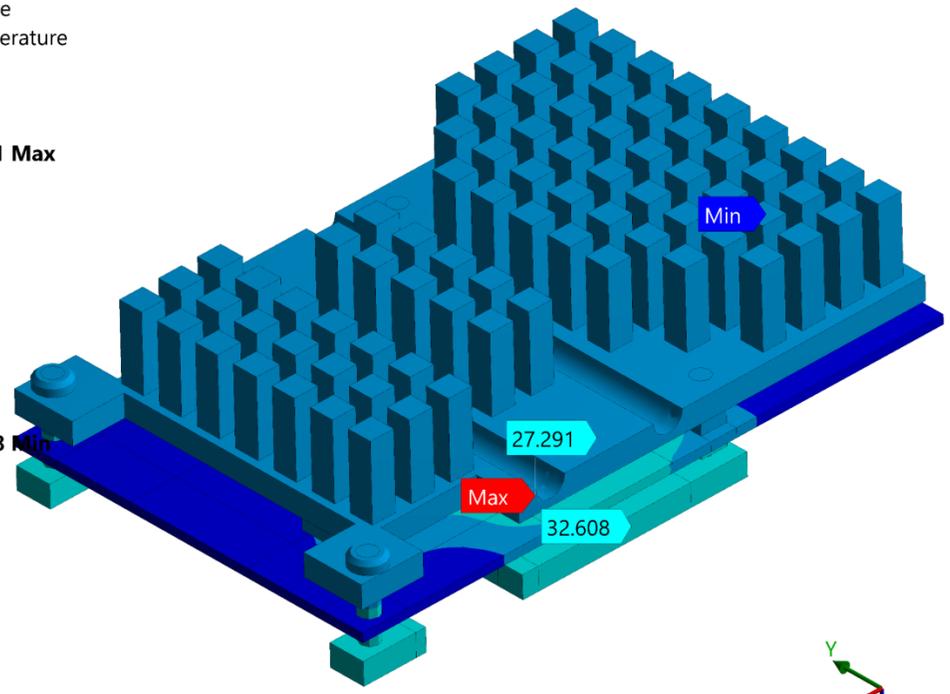
LV regulators





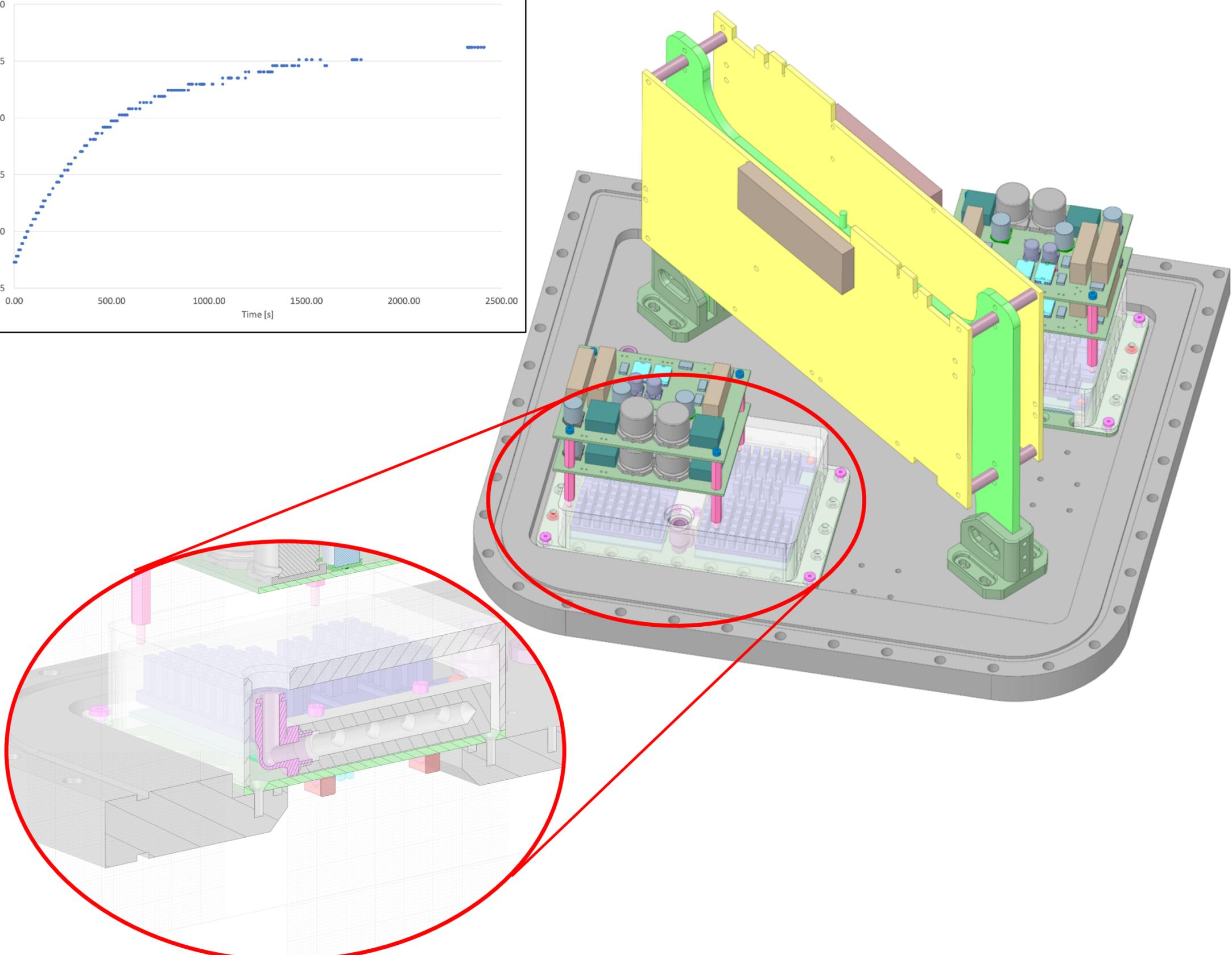
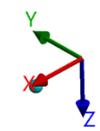
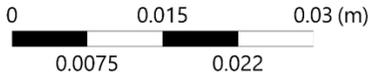
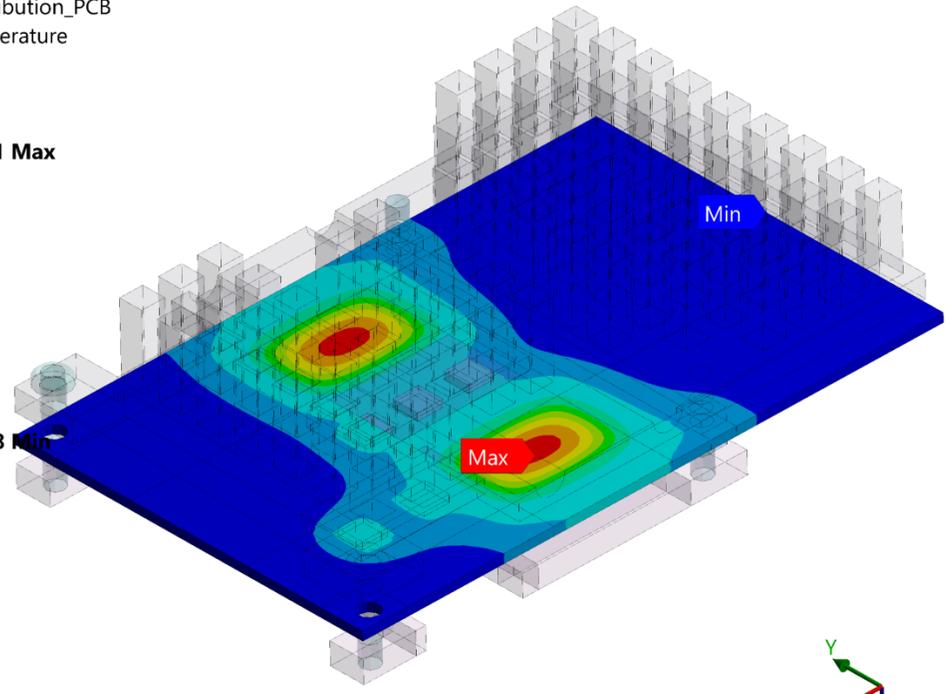
D: Steady-State Thermal_Aluminium1050
 Temperature
 Type: Temperature
 Unit: °C
 Time: 1

56.761 Max
 52.902
 49.042
 45.183
 41.324
 37.465
 33.606
 29.746
 25.887
22.028 Min



D: Steady-State Thermal_Aluminium1050
 Temp_Distribution_PCB
 Type: Temperature
 Unit: °C
 Time: 1

56.761 Max
 52.902
 49.042
 45.183
 41.324
 37.465
 33.606
 29.746
 25.887
22.028 Min



Principle

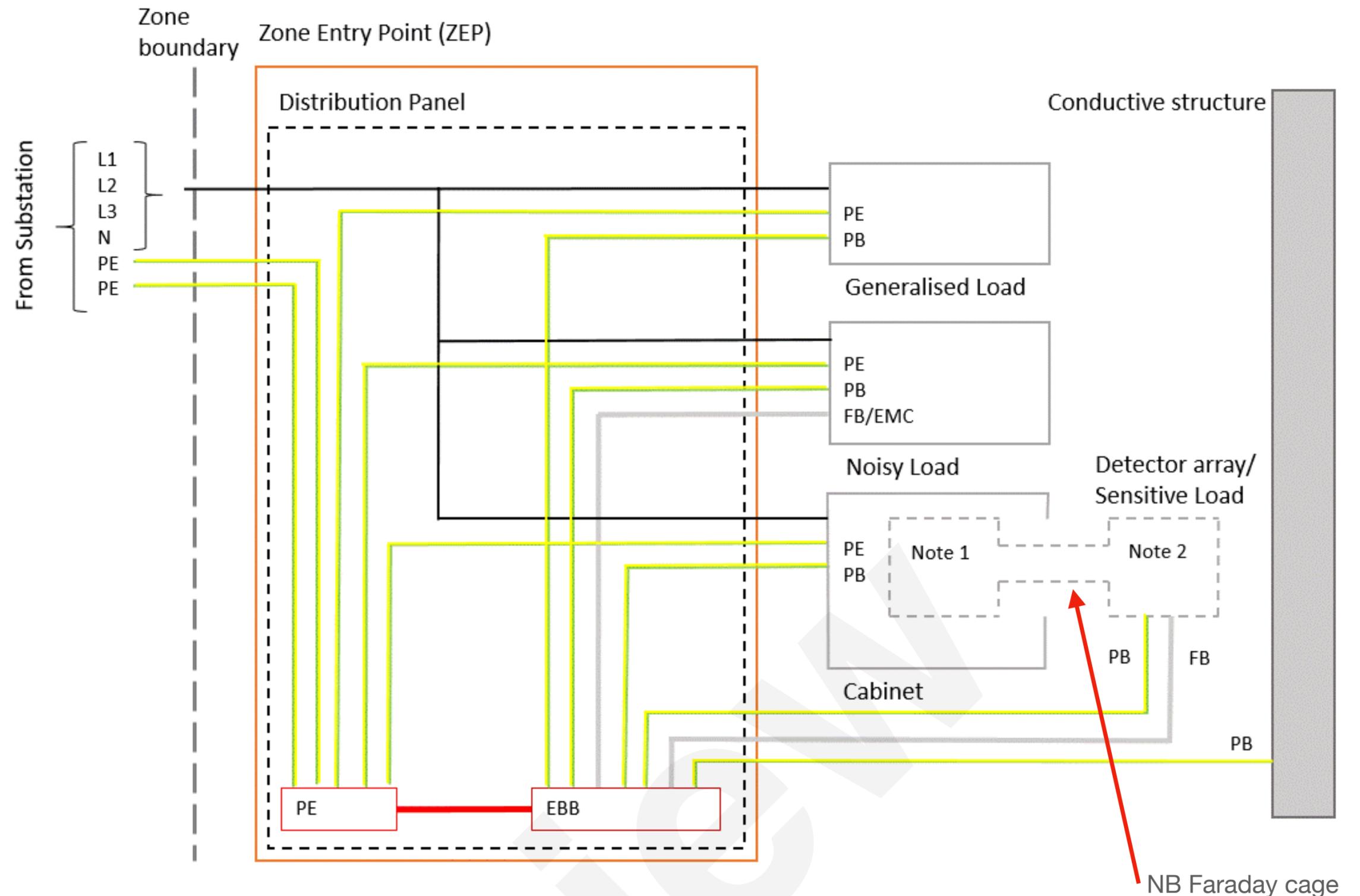
Sensitive part of detector will be surrounded by Faraday cage

i.e. Faraday cage encompasses from voltage supply (in rack) through detector sensitive element, through to Front End ASIC

For NMX this Faraday cage encompasses the detector module,
As this is the metallic object that surrounds readout plane
Therefore this should be isolated from robot/support

It is always wise to foresee bonding points at all isolated points - as it may be necessary to connect later

Details of NMX grounding will be drawn up Autumn this year



Note 1: This detector-corresponding part is supplied by its own UPS and therefore galvanically separated from the cabinet.

Note 2: The detector-complex is isolated from everything except the PB/FB connectors shown.

Figure 9 Principle drawing of earthing and bonding in a generic zone. ZEP in orange frame.

NB Caveat: indicative.
 This diagram is not
 uptodate.
 New version coming

NB all networks, & similar
 will be optically decoupled

Prototype Rack exists in Utgard and
 it is starting to be populated to
 determine configuration
 Example for LOKI



Rack electronics

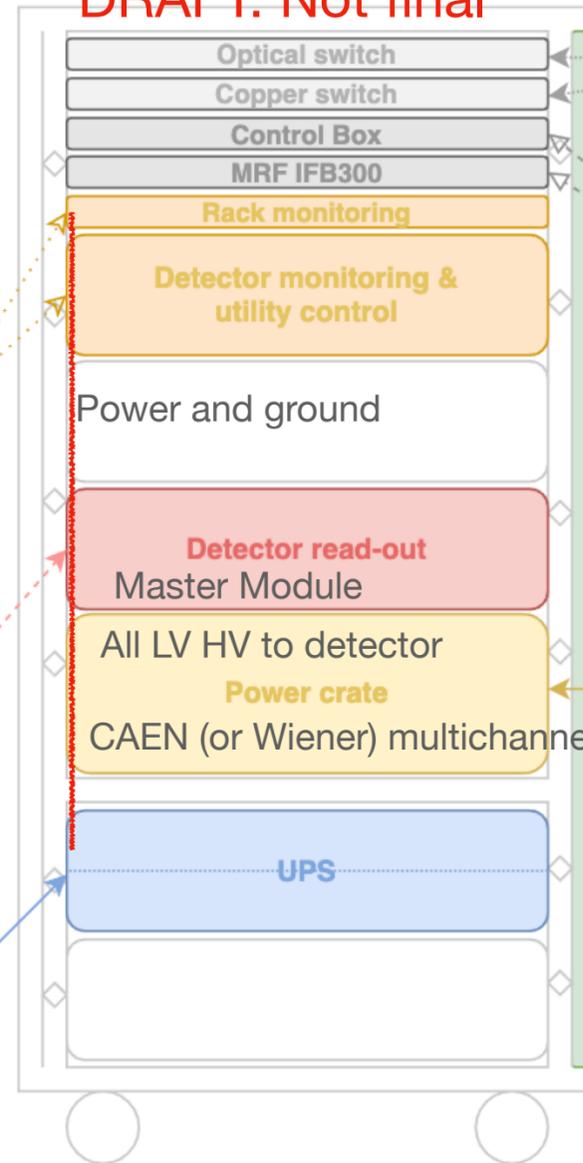
The standard detector rack is considered part of the readout back-end.

The rack environment will be monitored for e.g. humidity and temperature by use of a commercial-off-the-shelf system. Beckhoff terminals (3U-6U) are used to read and control detector parameters.

The ESS readout back-end provides a common interface for detector front-ends including controls and time distribution.

Six (6) horizontal units of space in the lower frame are reserved for the double-converting Uninterruptable Power Supply (UPS) which will be grounded through the mains input and supply floating ground for all downchain detector electronics.

DRAFT: Not final



Switches/patch panels to provide access to networks for timing distribution, event data, and controls. Only routed in via optical fibre.

Server to act as controls interface and to host the PCIe Event Receiver (EVR) and an interface board ("IFB300") to provide timing signal to ESS back-end readout system.

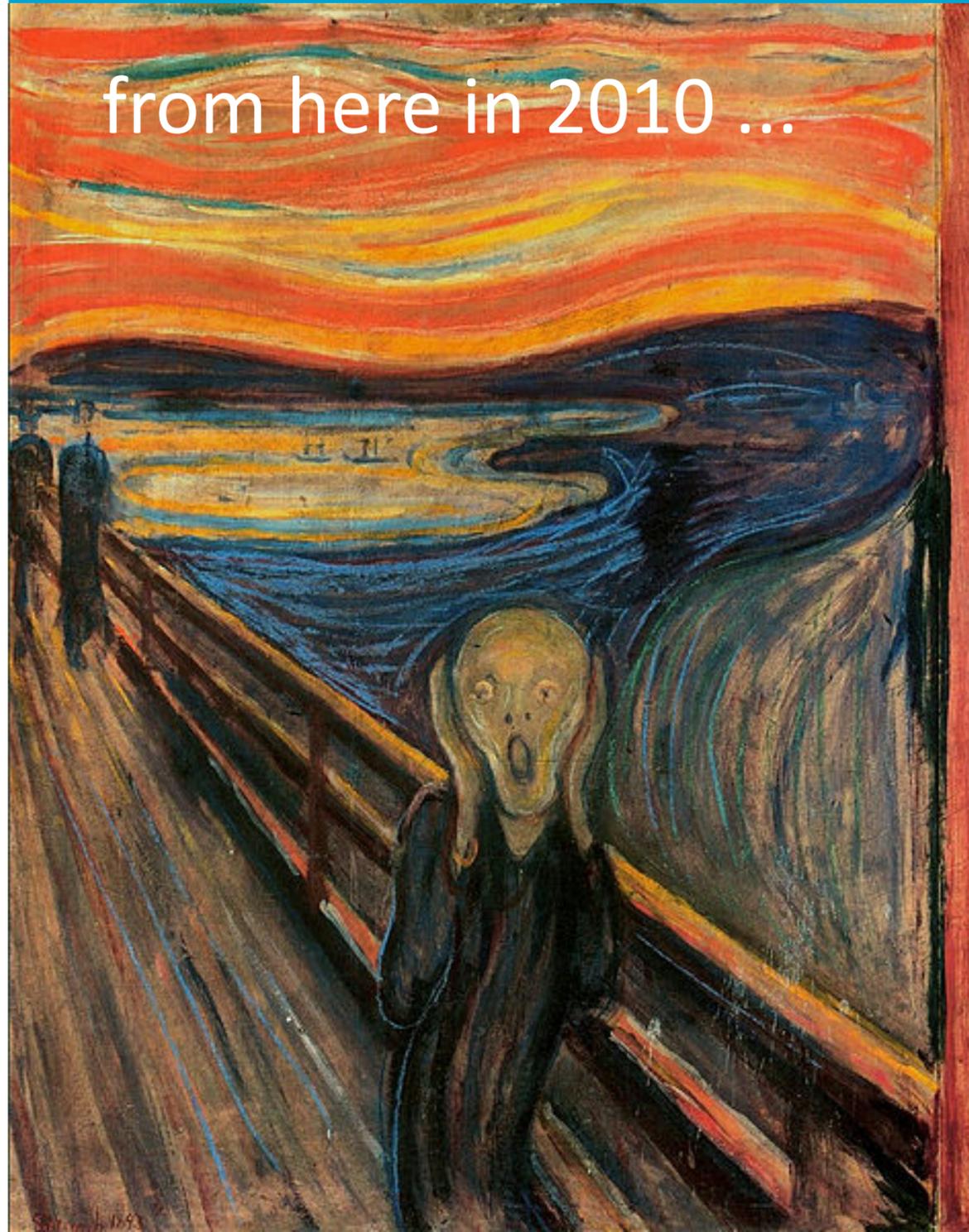
Up to eight (8) horizontal units in the upper frame will host the multichannel Low and High Voltage system, supplying e.g. bias for the detector(s).

A Power Distribution Unit (PDU) will take input from the UPS and feed power to all other detector components.

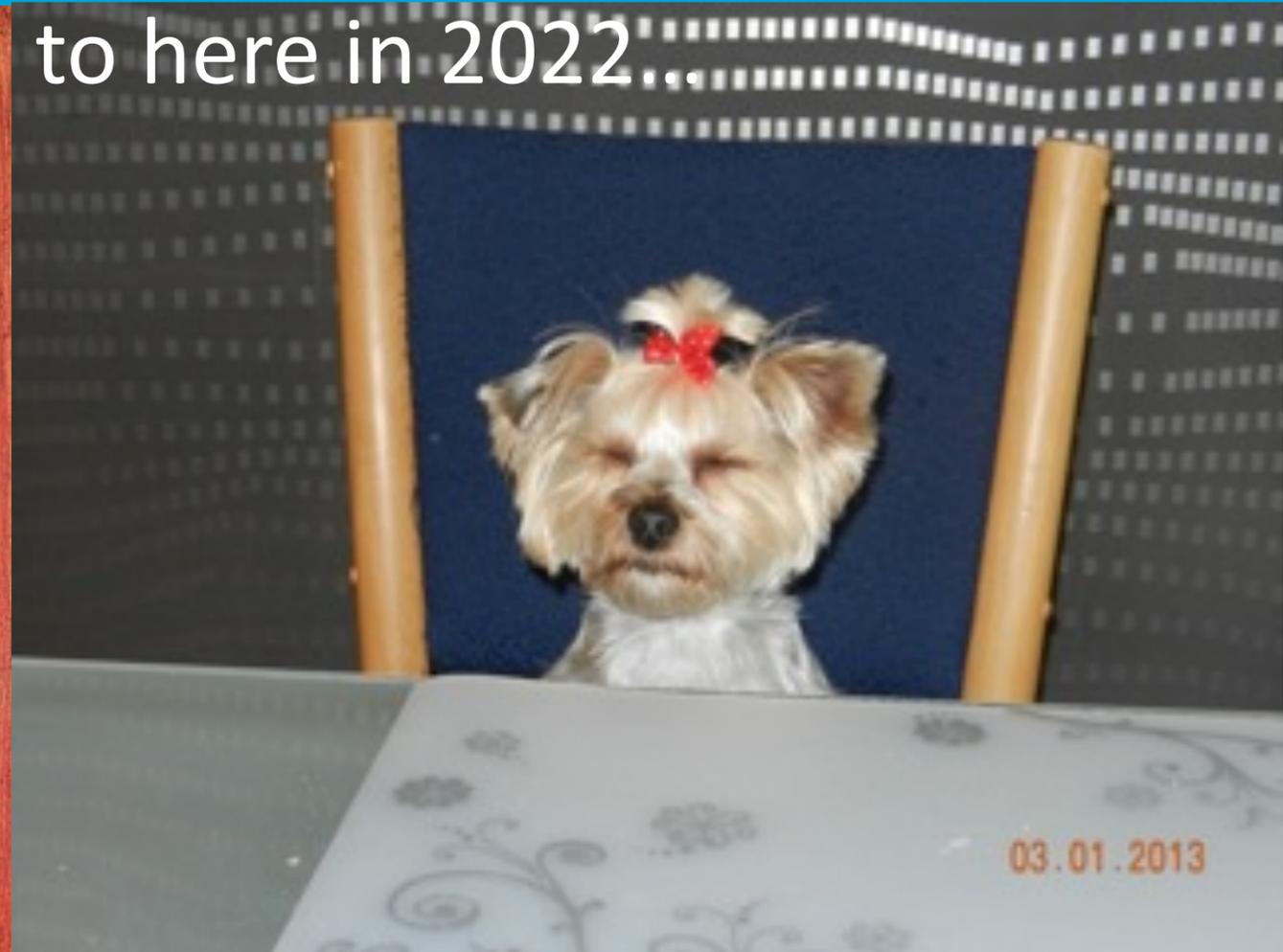
Some Final Thoughts ...

Mood Message for the Development so far ...

from here in 2010 ...

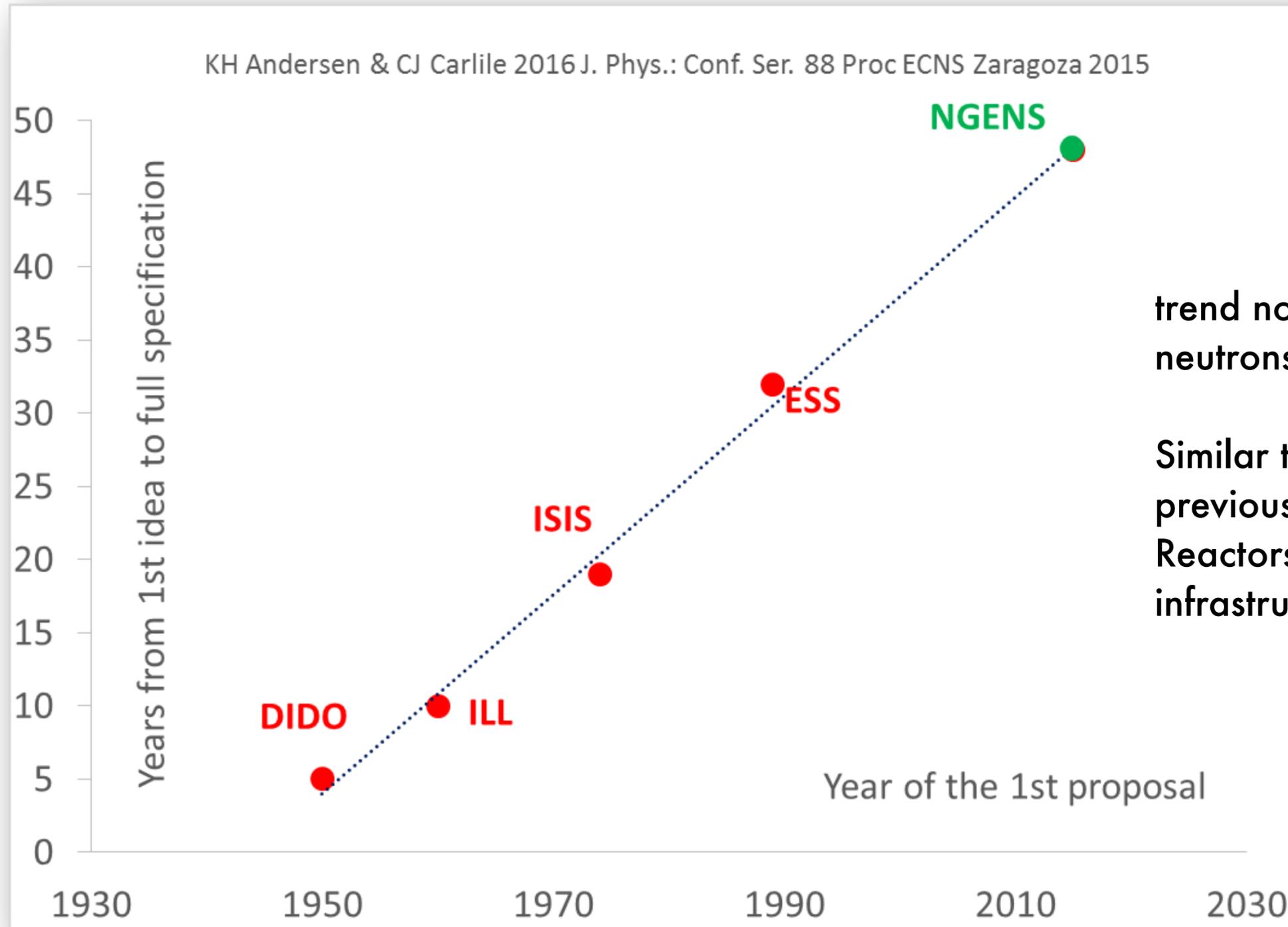


to here in 2022...



- Development time is long: typically 10 years from conception to utilisation
- Solve challenges one at a time, and remain calm

Time from first proposal of a neutron source to operation at full specification



trend not unique to neutrons

Similar trends noted previously for Power Reactors, HEP, civil infrastructure, etc ...

Collaborations for Construction Phase



Big projects are very very difficult

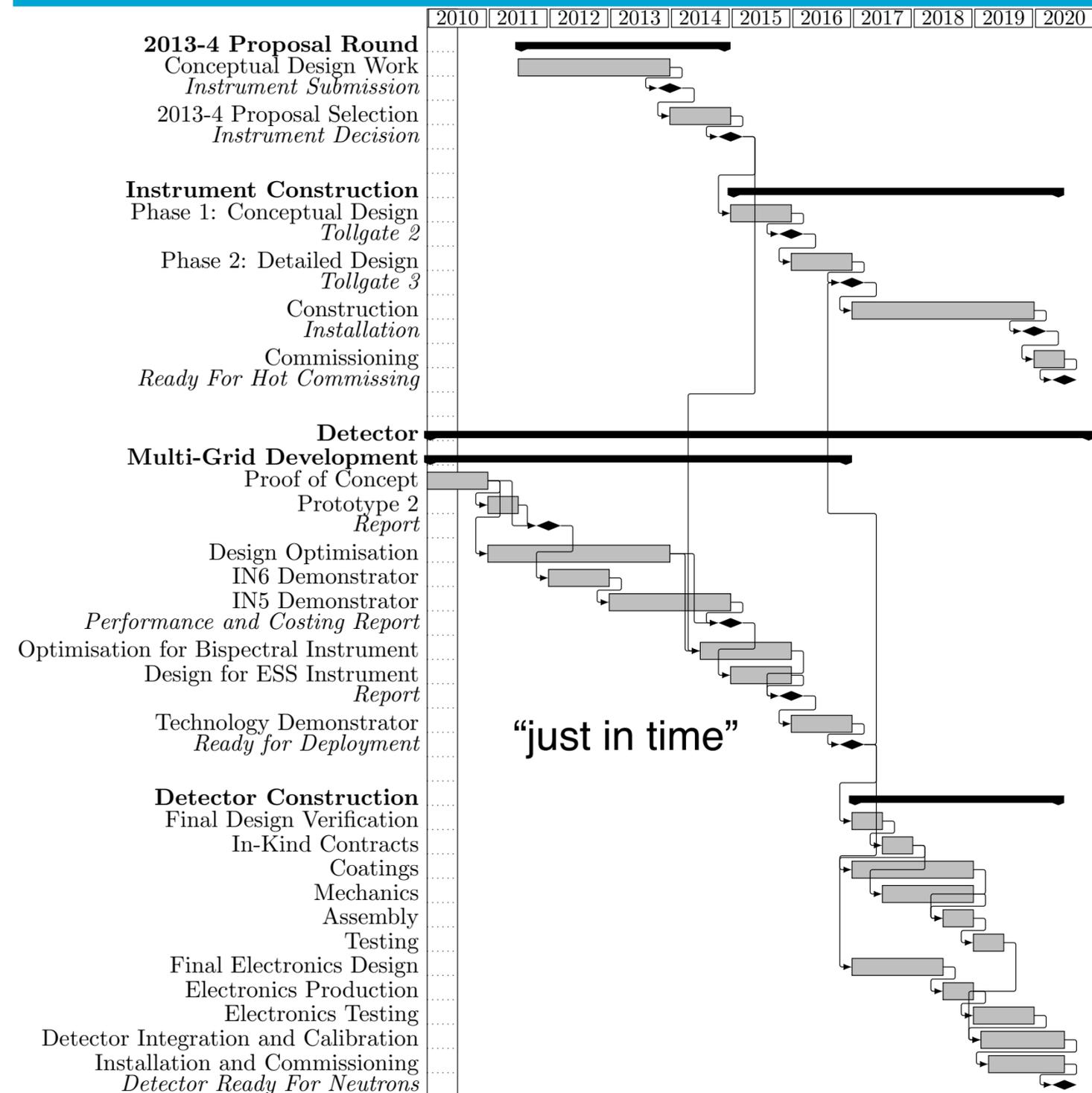
- Berlin airport ...
- Every military contract ...
- Your local hospital ...

It's not going to be smooth ...

... don't worry about this months crisis ...

Focus on the goal and the vision and work towards this

Timeline for detectors (schedule from 2012)



- Here is the timeline for a thermal chopper spectrometer with one concept for detector technology

- note: 10 years from concept to (potential) utilisation

- note: neither the proof of concept nor construction phases dominate the timeline, but rather the numerous prototyping and demonstration phases in between

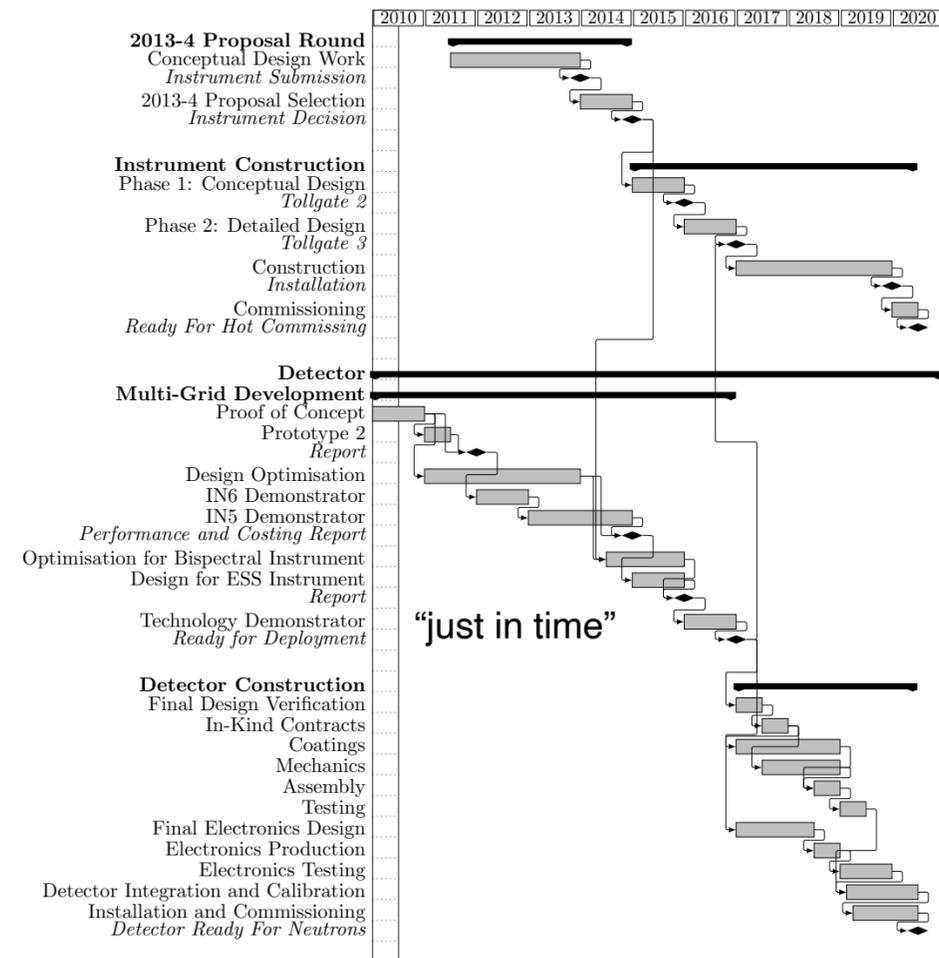
- Design friction: Only review decisions made if you really really need to

- I have seen many systems redesigned without any need to ... wasted effort.

- Get it working - keep it working -

- and then only then improve it

An aside on Project Management ...



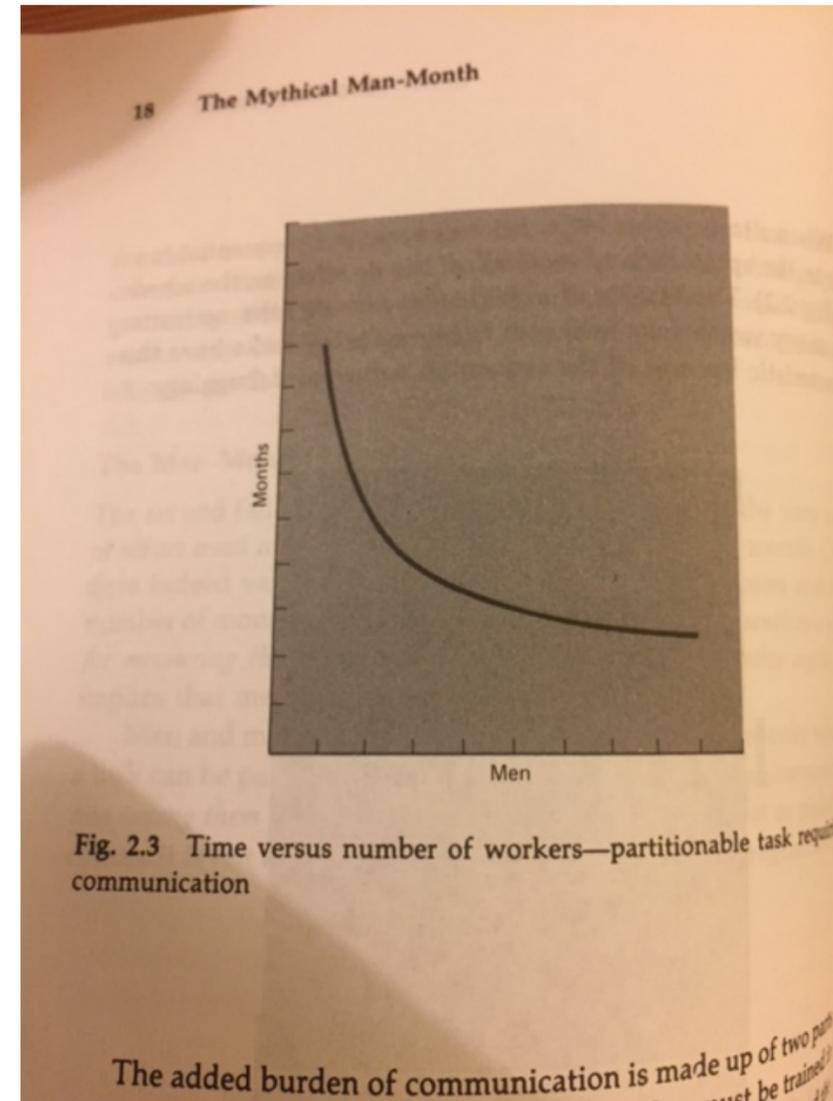
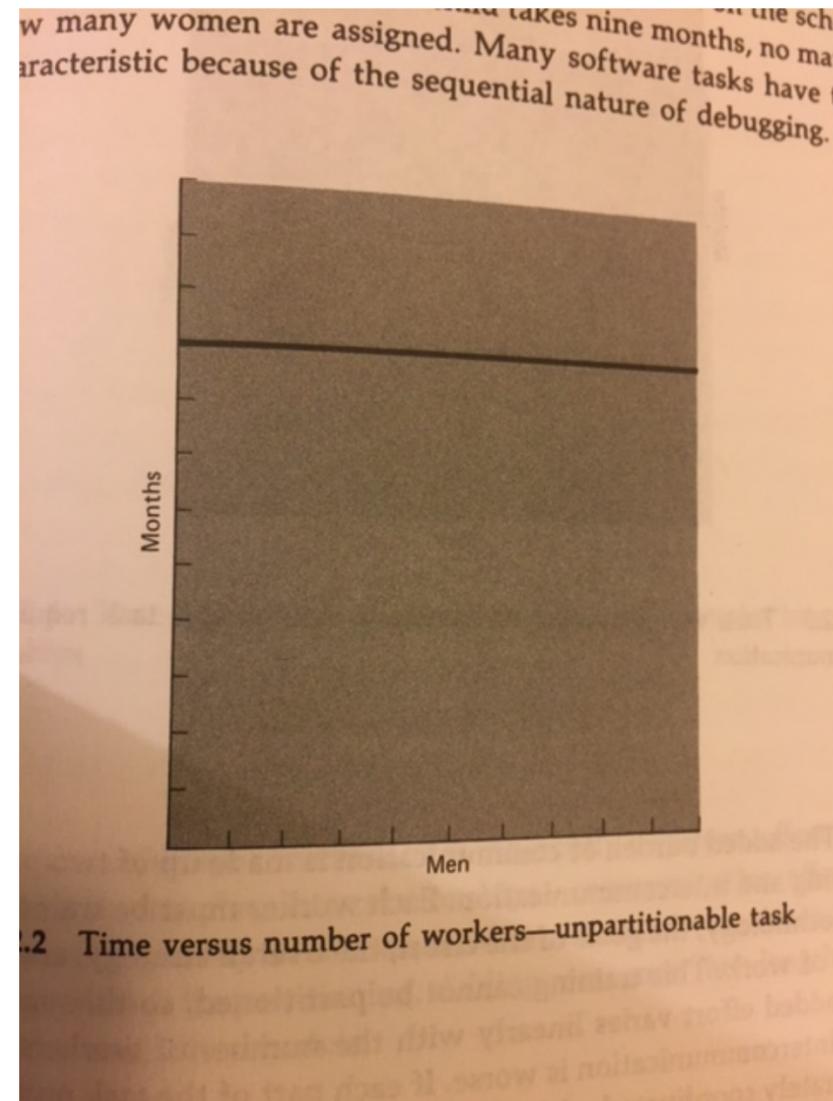
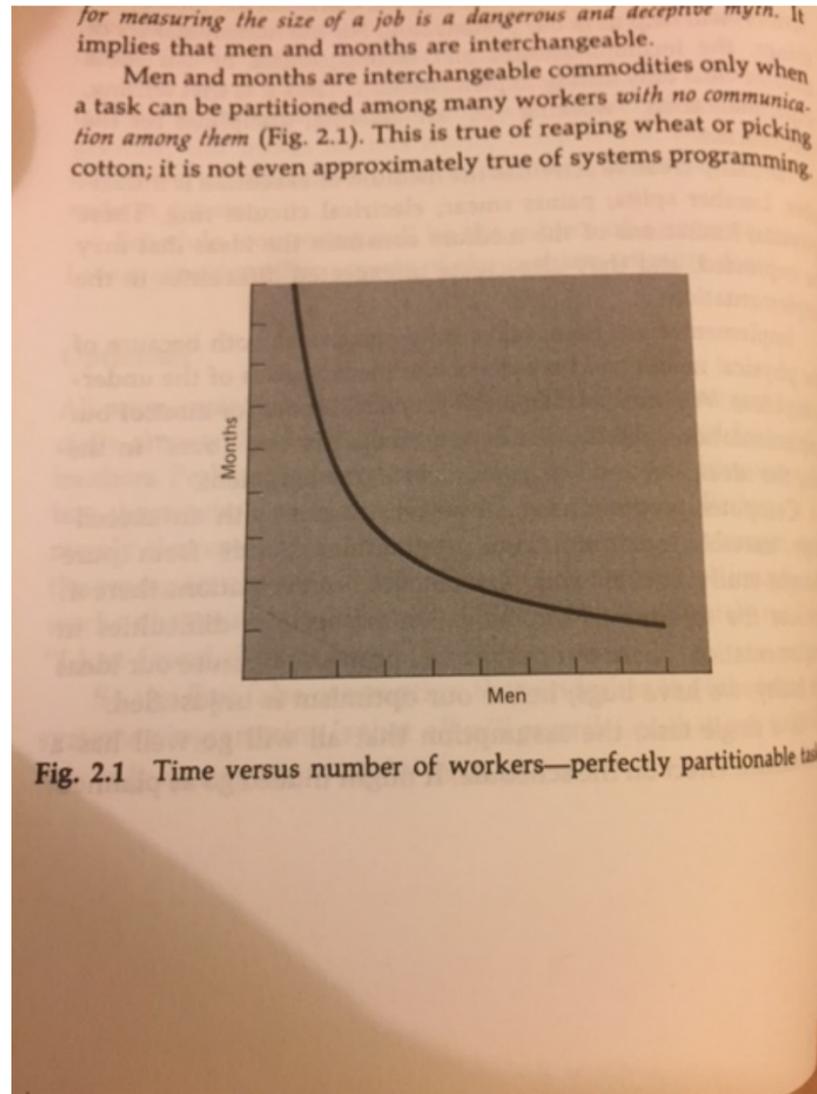
- You must to use “waterfall” project management (Gantt charts, etc).
- Most appropriate for “*known*” builds eg house...
- Focuses on immediate issues and critical path
- Use it to make sure that decisions are made, and interfaces understood and to understand progress

- However, detectors are very much a hi-tech item ...
- ... agile technique much better reflects actual work methodology
- You don’t know what exactly you will achieve during R&D
- Be aware that this might better reflect day-day work

- 3rd method: “successive principle”.
- Focus on the goal, and work back from these.
- Use non-experts to evaluate schedule given by experts ...

- Making progress and getting things done is complex in big projects ...

- From "the mythical man month"



- From "the mythical man month"

for measuring the size of a job is a dangerous and deceptive myth. It implies that men and months are interchangeable.

Men and months are interchangeable commodities only when a task can be partitioned among many workers *with no communication among them* (Fig. 2.1). This is true of reaping wheat or picking cotton; it is not even approximately true of systems programming.

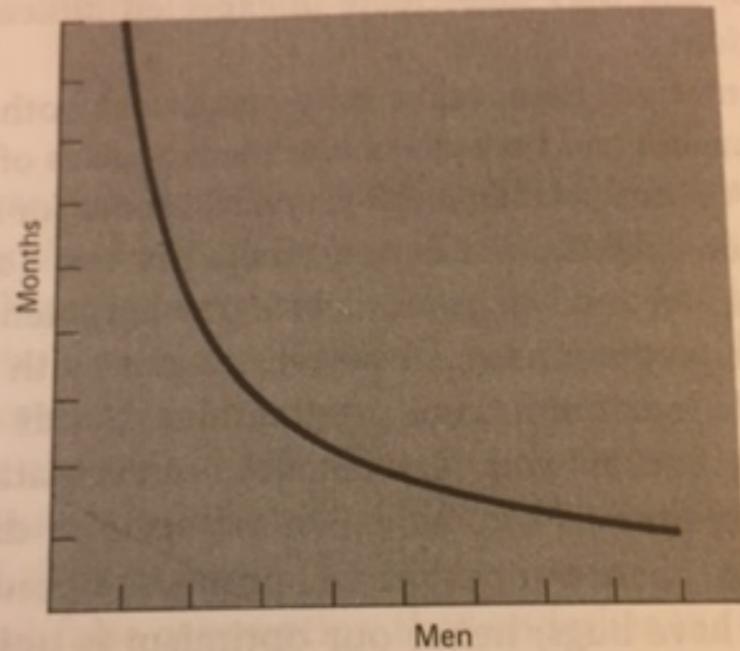
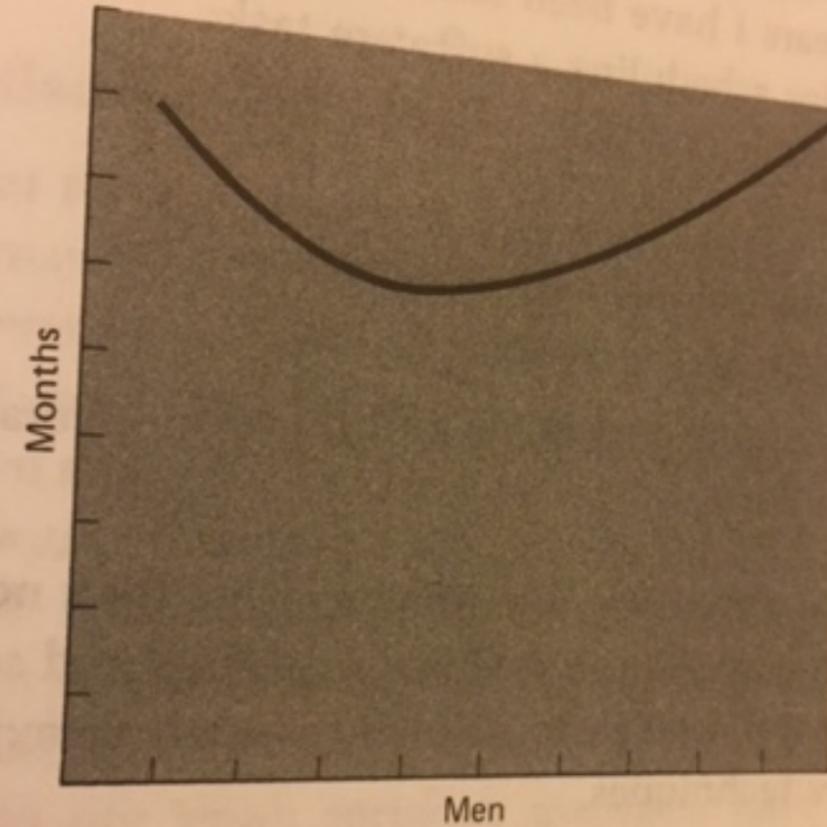


Fig. 2.1 Time versus number of workers—perfectly partitionable task



Time versus number of workers—task with complex interrelationships

tion is inherently a systems effort—an
tion effort is

Teamwork ...

- A pleasant working atmosphere ...
- A working culture that is open to diversity
- Creativity ...

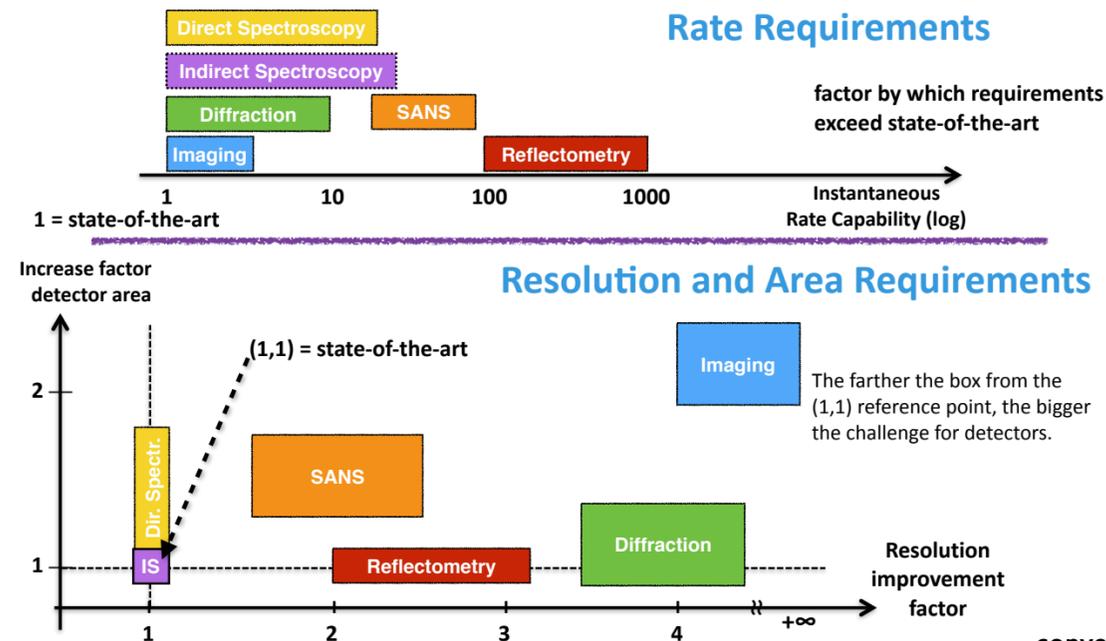


EUROPEAN
SPALLATION
SOURCE



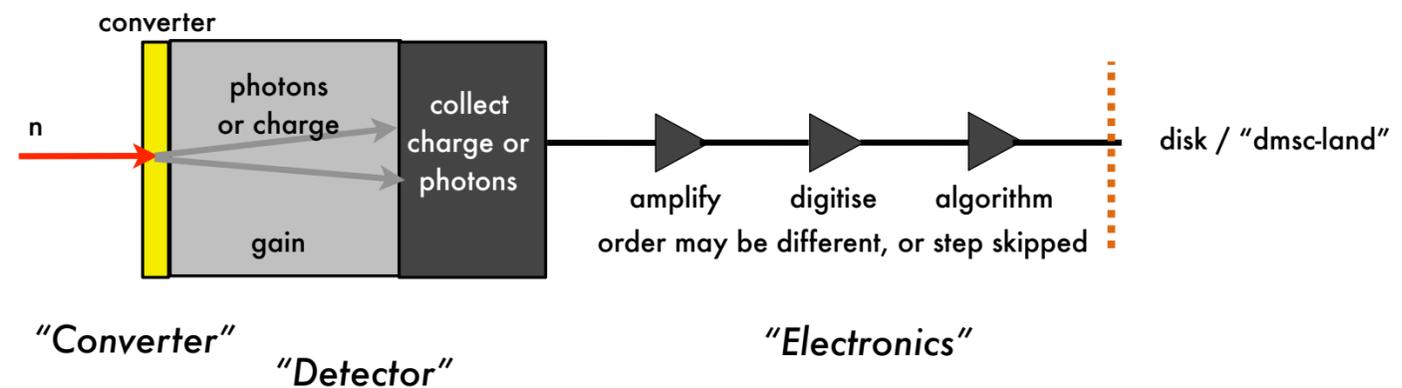
Thoughts ...

- Detector & DAQ development takes time
- Very difficult to go from concept to beam line in less than a decade
- e.g. Multi Grid started 2009/10. On ESS instrument ca. 2024/25
- Detector development time >> Instrument construction time
- This should be our aspiration level for a (every?) decade ... :



- Computing and electronics have become ubiquitous and very cheap
- Future detectors will be much more designed around electronics
- Don't divide DAQ & detector efforts
- Simulation will play a much bigger role in design
- Specialisation and collaboration

...and cost improvement of a factor of few ...



Summary

- Every DAQ system has different requirements: “horses for courses”
 - Used ESS as an example for how a data acquisition system is designed
 - What do you actually want to do with the DAQ system?
 - What is important?
-
- Optimise your time. Recycle what you can
 - You will underestimate the amount of time for changes (NRE).
 - Only change what you need to.
 - Don't be afraid to take off the shelf solutions.
-
- What can go wrong?
 - Remember: typically 10 years concept to implementation
 - Don't forget all the engineering factors ... not just data transport ...
-
- Make sure that you define what you want to measure clearly and unambiguously
 - Publish what you do: too many of the best results remain forgotten and are redone 3-10 years later
-
- It's all about people ...
 - ... and how they work with each other



Thanks!