Transport Barriers -The GK charge Flux Constraint

G. Merlo Max Planck Institute for Plasma Physics

ICTP Trieste 2024



Transport Barriers in Magnetized Plasmas- General Theory with Dynamical Constraints

M. Kotschenreuther,^{1,*} X. Liu,¹ S. M. Mahajan,¹ D. R. Hatch,¹ and G. Merlo¹ ¹University of Texas, Austin, Texas, USA (Dated: October 27, 2023)

A fundamental dynamical constraint - that fluctuation induced charge-weighted particle flux must vanish- can prevent instabilities from accessing the free energy in the strong gradients characteristic of Transport Barriers (TBs). Density gradients, when larger than a certain threshold, lead to a violation of the constraint and emerge as a stabilizing force. This mechanism, then, broadens the class of configurations (in magnetized plasmas) where these high confinement states can be formed and sustained. The need for velocity shear, the conventional agent for TB formation, is obviated. The most important ramifications of the constraint is to permit a charting out of the domains conducive to TB formation and hence to optimally confined fusion worthy states; the detailed investigation is conducted through new analytic methods and extensive gyrokinetic simulations.

Particle motion in an B field with an external force

$$m\frac{d\vec{v}}{dt} = q\,\vec{v}\times\vec{B} + \vec{F},$$

Is the superposition of a gyromotion and a drift due to F with velocity

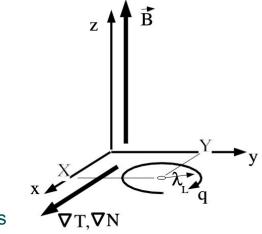
$$\vec{v}_F = \frac{\vec{F} \times \vec{B}}{qB^2},$$

The equilibrium solution is a Maxwellian distribution whose mean velocity is

$$\vec{V}_d(x=0) = \frac{T}{qB} \left[\frac{d\ln\mathcal{N}}{dx} \Big|_{x=0} + \frac{d\ln T}{dx} \Big|_{x=0} \right] \vec{e}_y = \frac{T}{qB} \left[\frac{d\ln N}{dx} \Big|_{x=0} + \frac{d\ln T}{dx} \Big|_{x=0} - \frac{F}{T} \right] \vec{e}_y$$
$$= \frac{1}{qB^2} \left(-\frac{\nabla p}{N} + \vec{F} \right) \times \vec{B},$$

i.e. superposition of diamagnetic drifts and F induced drift







Let's consider a kinetic description of a collisionless plasma, satisfies Vlasov

$$\left[\frac{\partial}{\partial t} + \vec{v} \cdot \frac{\partial}{\partial \vec{r}} + \frac{1}{m} \left(q\vec{v} \times \vec{B} + \vec{F}\right) \cdot \frac{\partial}{\partial \vec{v}}\right] f = 0.$$

And add a perturbation

$$\phi = \hat{\phi}(x) \exp i(k_y y + k_z z - \omega t)$$



Let's consider a kinetic description of a collisionless plasma, satisfies Vlasov

$$\left[\frac{\partial}{\partial t} + \vec{v} \cdot \frac{\partial}{\partial \vec{r}} + \frac{1}{m} \left(q\vec{v} \times \vec{B} + \vec{F}\right) \cdot \frac{\partial}{\partial \vec{v}}\right] f = 0.$$

And add a perturbation

$$\phi = \hat{\phi}(x) \exp i(k_y y + k_z z - \omega t)$$

Each distribution function will respond with a similar perturbation

$$\delta f = f_0 + \delta f$$
 $\delta f = \delta f \exp i(k_y y + k_z z - \omega t).$

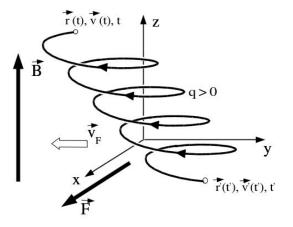
Satisfying the linearized Vlasov equation

$$\frac{D}{Dt}\Big|_{u.t.p.}\delta f = \left[\frac{\partial}{\partial t} + \vec{v}\cdot\frac{\partial}{\partial \vec{r}} + \frac{1}{m}\left(q\vec{v}\times\vec{B}+\vec{F}\right)\cdot\frac{\partial}{\partial \vec{v}}\right]\delta f = \frac{q}{m}\nabla\phi\cdot\frac{\partial f_0}{\partial \vec{v}}.$$

We can solve for the perturbation integrating along unperturbed trajectories

$$\delta f(\vec{r},\vec{v},t) = \frac{q}{m} \int_{-\infty}^{t} dt' \left. \nabla \phi \cdot \frac{\partial f_0}{\partial \vec{v}} \right|_{\vec{r}'(t'),\vec{v}'(t'),t'}$$

$$\begin{array}{lll} \displaystyle \frac{d\vec{r}^{\,\prime}}{dt^{\prime}} & = & \vec{v}^{\,\prime}, \\ \displaystyle \frac{d\vec{v}^{\,\prime}}{dt^{\prime}} & = & \displaystyle \frac{1}{m} \left(q\vec{v}^{\,\prime} \times \vec{B} + \vec{F} \right), \end{array}$$





We can solve for the perturbation integrating along unperturbed trajectories

$$\delta f(\vec{r}, \vec{v}, t) = \frac{q}{m} \int_{-\infty}^{t} dt' \left. \nabla \phi \cdot \frac{\partial f_0}{\partial \vec{v}} \right|_{\vec{r}'(t'), \vec{v}'(t'), t'}$$

$$\begin{aligned} \frac{d\vec{r}'}{dt'} &= \vec{v}', \\ \frac{d\vec{v}'}{dt'} &= \frac{1}{m} \left(q\vec{v}' \times \vec{B} + \vec{F} \right), \end{aligned}$$

$$\hat{\delta f} = -\frac{q\hat{\phi}}{T} \left[1 - (\omega_d' - \omega) \sum_{n,n'=-\infty}^{+\infty} \frac{J_n \left(\frac{k_y v_\perp}{\Omega}\right) J_{n'} \left(\frac{k_y v_\perp}{\Omega}\right) e^{i(n-n')\theta}}{k_z v_z + n\Omega + \omega_F - \omega} \right] f_0,$$

$$\mathbf{y}$$



We can solve for the perturbation integrating along unperturbed trajectories

$$\delta f(\vec{r}, \vec{v}, t) = \frac{q}{m} \int_{-\infty}^{t} dt' \left. \nabla \phi \cdot \frac{\partial f_0}{\partial \vec{v}} \right|_{\vec{r}'(t'), \vec{v}'(t'), t'}$$

$$\begin{aligned} \frac{d\vec{r}'}{dt'} &= \vec{v}', \\ \frac{d\vec{v}'}{dt'} &= \frac{1}{m} \left(q\vec{v}' \times \vec{B} + \vec{F} \right), \end{aligned}$$

Which can be solved and used to derive a dispersion relation from the dielectric function

$$\hat{\delta f} = -\frac{q\hat{\phi}}{T} \left[1 - \left(\omega_d' - \omega\right) \sum_{n,n'=-\infty}^{+\infty} \frac{J_n\left(\frac{k_y v_\perp}{\Omega}\right) J_{n'}\left(\frac{k_y v_\perp}{\Omega}\right) e^{i(n-n')\theta}}{k_z v_z + n\Omega + \omega_F - \omega} \right] f_0,$$

$$\begin{split} - \bigtriangleup \phi &= k^2 \phi = \frac{1}{\epsilon_0} \sum_{\text{species}} q \delta N \\ \epsilon(\vec{k}, \omega) &\doteq 1 - \sum_{\text{species}} \frac{q}{\epsilon_0 k^2} \frac{\delta \hat{N}}{\hat{\phi}} = 0, \end{split}$$



Explicitly performing integral and n=0 limit (long wavelength low frequency)

$$1 + \sum_{\text{species}} \frac{1}{(k\lambda_D)^2} \left\{ 1 + \frac{\omega - \omega'_d}{\omega - \omega_F} \left[W\left(\frac{\omega - \omega_F}{|k_z|v_{th}}\right) - 1 \right] \Lambda_0(\xi) \right\} = 0.$$

One can consider various limiting cases:

- only density gradient and F not dependent on q -> flute instability (like Rayleigh-Taylor)
- only B gradients
- only density gradient -> drift waves, modification of sound waves
- only temperature gradients -> slab ITG

Explicitly performing integral and n=0 limit (long wavelength low frequency)

$$1 + \sum_{\text{species}} \frac{1}{(k\lambda_D)^2} \left\{ 1 + \frac{\omega - \omega'_d}{\omega - \omega_F} \left[W\left(\frac{\omega - \omega_F}{|k_z|v_{th}}\right) - 1 \right] \Lambda_0(\xi) \right\} = 0.$$

One can consider various limiting cases:

- only density gradient and F not dependent on q -> flute instability (like Rayleigh-Taylor)
- only B gradients
- only density gradient -> drift waves, modification of sound waves
- only temperature gradients -> slab ITG

Neglect FLR, expand W, neglect resonances

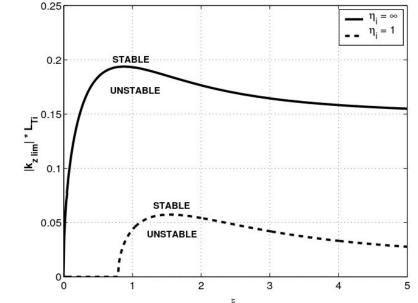
$$1 - \left(\frac{k_z c_s}{\omega}\right)^2 \left(1 - \frac{\omega_{T_i}}{\omega}\right) = 0$$



In practice we usually have T and n gradients

$$\eta_i \doteq \frac{d\ln T_i}{d\ln N} = \frac{L_N}{L_{T_i}}$$

Proceed same way, but solve numerically now



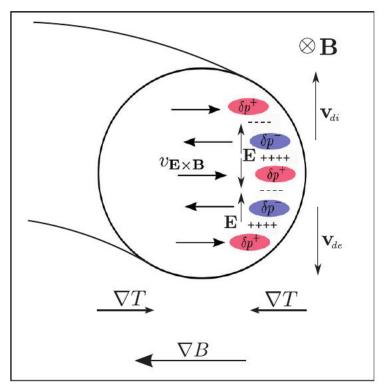


Why it is so relevant? Understood as one of the most important turbulence drive mechanisms

$$\omega = \pm \sqrt{-2\frac{T_e}{T_i}(k_\theta \rho_i)^2 \nabla \log T \cdot \nabla \log B}$$

Unstable at the outboard midplane.

Electron contribution is not necessarily negligible. They are responsible for the TEM, the other major instability in a Tokamak.

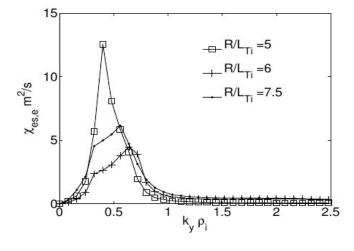


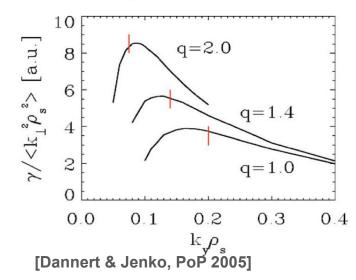
Mixing length estimates

Mixing length argument, balance linear growth and nonlinear convection $D_{mix}=\chi\sim\gamma/k_\perp^2$

Variations incorporating the eigenfuction work better

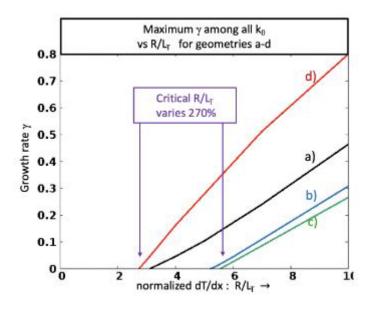
$$\langle k_{\perp}^2 \rangle(k_y) = \sum_{k_x} \int k_{\perp}^2(k_x, k_y, z) \left| \hat{\phi}(k_x, k_y, z) \right|^2 \\ \times J(z) \, dz \Big/ \sum_{k_x} \int \left| \hat{\phi}(k_x, k_y, z) \right|^2 J(z) \, dz$$







Let's consider a series of variations around nominal parameter sets representative of transport barrier conditions

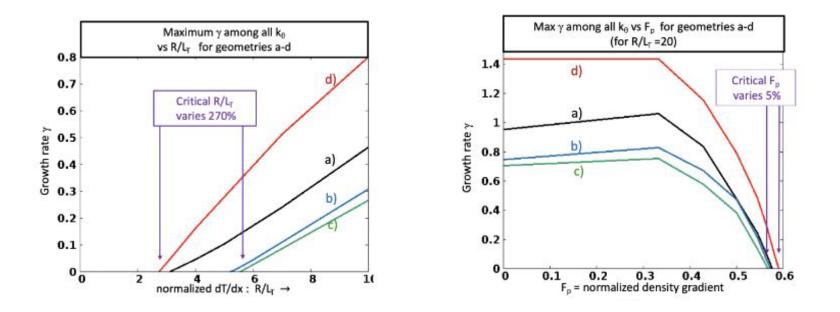


Very strong variation of critical gradients with plasma parameters. Best is to use F_p the fraction of pressure gradient in density:

$$F_P = (Tdn/dx)/[d(nT)/dx] = \frac{1/L_n}{(1/L_n + 1/L_T)}$$

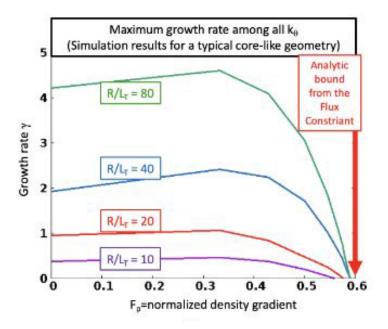


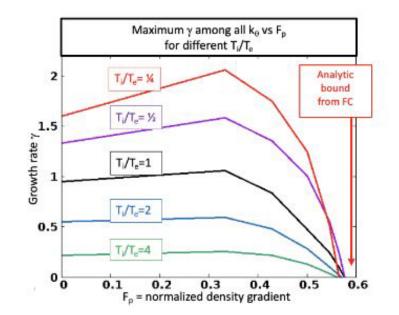
Let's consider a series of variations around a nominal parameter set repernsetative of transport barrier parameters



 ${\rm F}_{\rm p}$ is much more robust indicator of the intrinsic system behavior

For a lot of variations

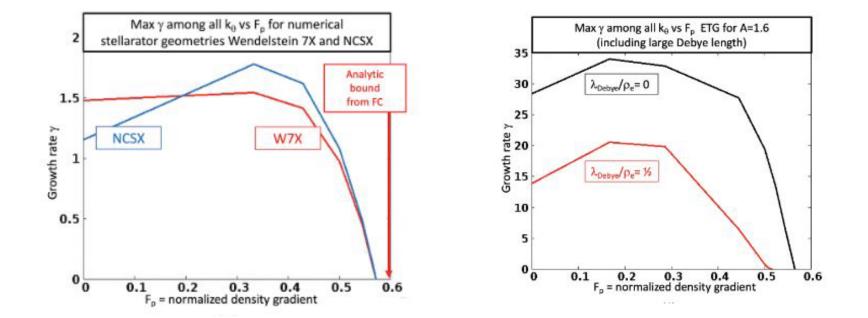






For a lot of variations





Regardless of the completely different parameters, the system follows the same behavior, why?

The charge Flux Constraint - FC

The reason of such a robust behavior stems from a very simple constraint

$$\sum_{s} q_s \Gamma_{rs} = 0$$

The charge weighted flux must be zero.



The charge Flux Constraint - FC

The reason of such a robust behavior stems from a very simple constraint

$$\sum_{s} q_s \Gamma_{rs} = 0$$

The charge weighted flux must be zero.

This acts together with the energetics dynamics of the system

$$\begin{aligned} &\frac{\partial}{\partial t} \sum_{s} \left[\frac{T_s}{2} \int d\mathbf{x} d\mathbf{v} \frac{\delta f_s^2}{F_M} + \frac{m_s n_s \delta V_{E \times B}^2}{2} + \frac{\delta E^2}{8\pi} + \frac{\delta B^2}{8\pi} \right] = \\ &\sum_{s} \left[\left(Q_s \frac{1}{T_s} \frac{dT_s}{dx} + T_s \Gamma_s \frac{1}{n_s} \frac{dn_s}{dx} \right) \right] - \sum_{s} \int d\mathbf{x} d\mathbf{v} \frac{\delta f_s}{F_M} \mathbf{C}(\delta f_s) \end{aligned}$$

Free energy in fluctuations grows because of fluxes (minus entropy)



A mean field theory - SKiM

Gyrokinetics is extremely complicated, many degrees of freedom. Simplify it with a mean field theory (Simplified Kinetic Model - SKiM), constructed by taking an average along the field line:

$$< k_{\parallel} >^{2} = \frac{\int \mathrm{d}l |\frac{\partial \phi}{\partial l}|^{2}}{\int \mathrm{d}l |\phi|^{2}|}$$
$$< k_{\perp} >^{2} = \frac{\int \mathrm{d}l |\phi|^{2} k_{\perp}^{2}}{\int \mathrm{d}l |\phi|^{2}},$$
$$< \omega_{di} > = \frac{\int \mathrm{d}l |\phi|^{2} \omega_{di}}{\int \mathrm{d}l |\phi|^{2}}$$

Eigenfunction averaged wave vectors and drift frequencies



A mean field theory - SKiM

 $< k_{\parallel} >^2 = \frac{\int \mathrm{d}l |\frac{\partial \phi}{\partial l}|^2}{\int \mathrm{d}l |\phi|^2}$

 $\langle k_{\perp} \rangle^2 = \frac{\int \mathrm{d}l |\phi|^2 k_{\perp}^2}{\int \mathrm{d}l |\phi|^2},$

 $<\omega_{di}>=rac{\int \mathrm{d}l|\phi|^2\omega_{di}}{\int \mathrm{d}l|\phi|^2}$

Gyrokinetics is extremely complicated, many degrees of freedom. Simplify it with a mean field theory (Simplified Kinetic Model - SKiM), constructed by taking an average along the field line:

Non-adiabatic part of distribution function h

$$h_s = \frac{\frac{q_s \phi}{T_s} J_0(\langle k_\perp > \rho_s)(\omega - \omega_s^\star) f_{Ms}}{(\omega - \langle \omega_{ds} \rangle) - v_{\parallel} \langle k_{\parallel} \rangle + iC}$$

Yielding with QN the dispersion relation

$$\sum_{ion \ species} \left(\int \mathrm{d}\vec{v} \frac{F_{Ms} J_0(\langle k_\perp > \rho_i)^2 (\omega - \omega_s^\star)}{\omega - \langle k_\parallel > v_\parallel - \langle \omega_{ds} >} - 1 \right) \frac{q_s^2}{T_s} - \frac{n_e e^2}{T_e} = 0$$

Eigenfunction averaged wave vectors and drift frequencies

One can solve SKiM and compare to "standard" dispersion relation, but there is some basic information one can extract



A mean field theory - SKiM

Compute the particle flux resulting from h gives

$$\int d\boldsymbol{v} f_M J_0(\langle k_\perp \rangle \rho_i)^2 \Big[\frac{\gamma}{\gamma^2 + (\omega_r - \langle k_\parallel \rangle v_\parallel - \langle \omega_d \rangle)^2} \Big] \\ \Big[1/L_n + (m_s v^2/2T_s - 3/2)/L_T \Big] = 0$$

Cannot be solved if

$$1/L_n > 3/(2L_T)$$

or

 $F_P > 0.6$

Simple prediction, no need to perform any complex GK simulation, valid for any fluctuation!

The violation of the charge flux constraint is extremely robust and fundamental.



FC vs energetics - a graphical picture

The crucial role of the FC can be more easily understood visually

$$D(\omega) = \sum_{ion \ species} \left(\int \mathrm{d}\vec{v} \frac{F_{Ms} J_0(\langle k_\perp > \rho_i)^2 (\omega - \omega_s^\star)}{\omega - \langle k_\parallel > v_\parallel - \langle \omega_{ds} >} - 1 \right) \frac{q_s^2}{T_s} - \frac{n_e e^2}{T_e} = 0$$
 Negative flux

Solved only if

 $Im(D(\omega)) = 0$ Flux constraint

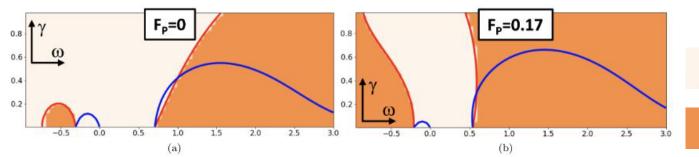
 $Re(i\omega D(\omega)){=}0$ Free energy

Solution of the system is the intersection of the two curves in the complex plane. Both are important and provide two different dynamics to the system evolution



Positive flux

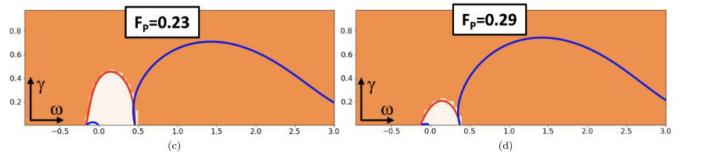
FC vs energetics - a graphical picture

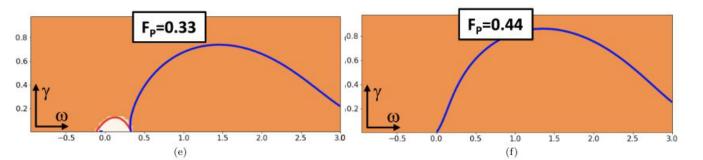




Positive flux

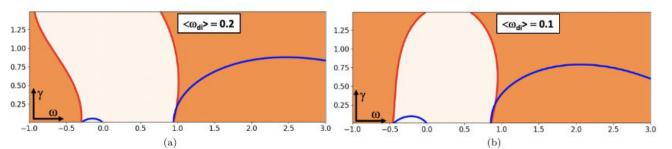
Negative flux

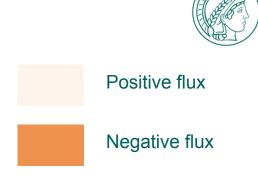


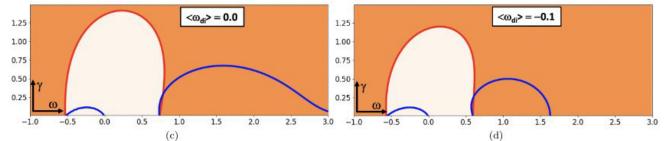


System is stabilized by the FC. Increasing F_p shrinks and finally removes the zero flux region, even if there is a lot of free energy available.

FC vs energetics - a graphical picture

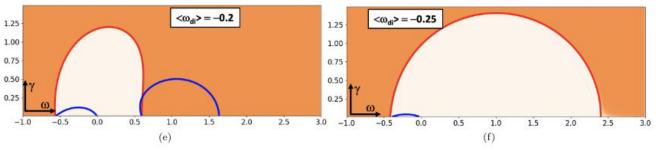




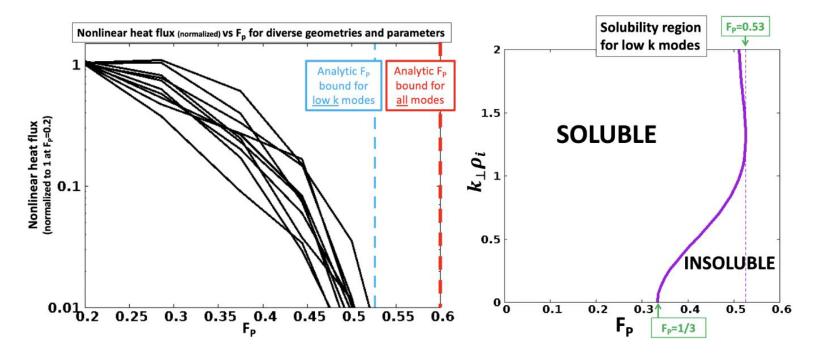


Scan in curvature, from positive to very negative

In this case the stabilization is via energetics, the FC can always be satisfied

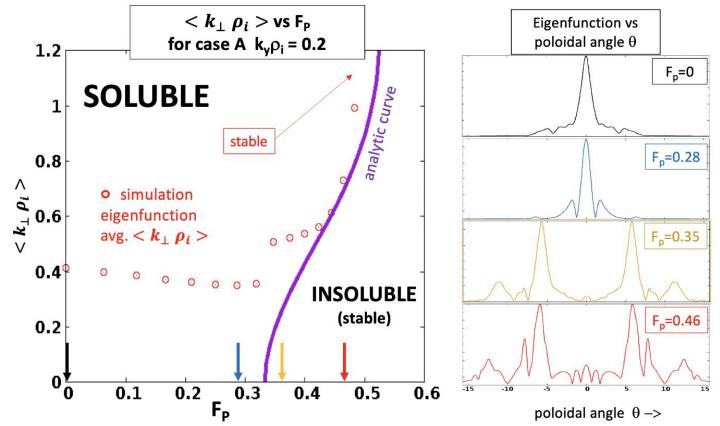


Nonlinear transport is dominated by low k_v modes & follows similar trend



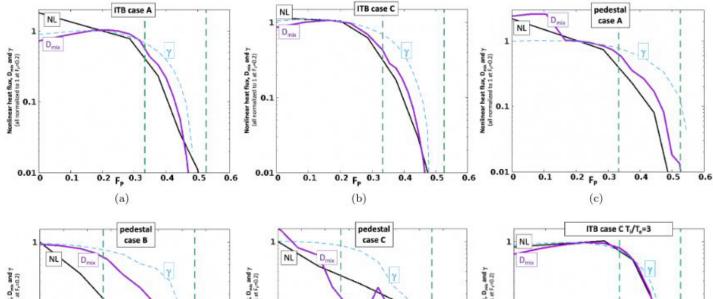
Dependence of F_p on details of the eigenfunction, max $F_p \sim 0.53$ & k~1, but for low k_y dramatic changes once near 1/3

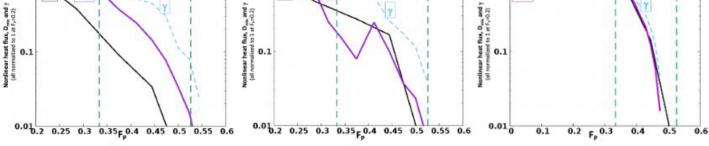




As F_n is increased and the solubility is approached, the system adapts with major changes in the eigenfunction

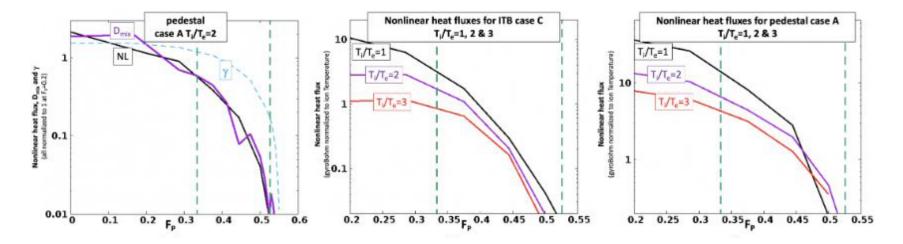






 k_{\perp} matters, adaptivity





Q_{NL} doesn't



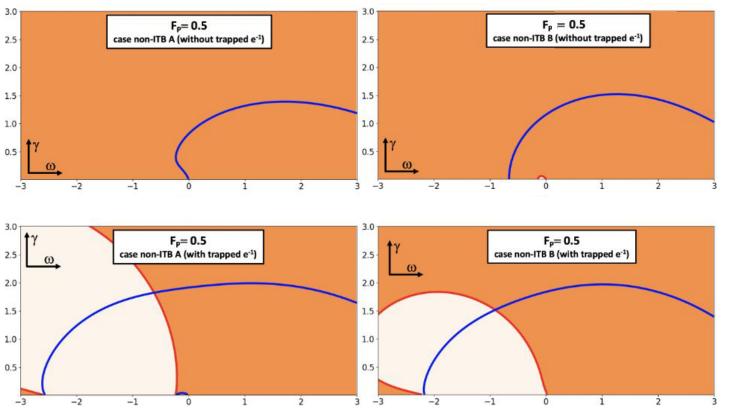
Non-adiabatic electron response often plays a crucial role in gyrokinetic instabilities. For trapped electrons, it can lead to the TEM.

Since non-adiabatic electrons contribute charge flux in the opposite direction violation of FC becomes more difficult.

It is possible that when the non-adiabatic electron charge flux is large enough, density gradients do not, necessarily, lead to insolubility. This situation occurs for many typical tokamak geometries, in fact.

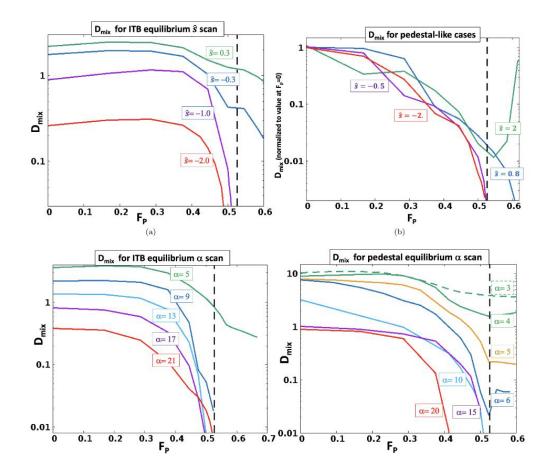
In these cases, the stabilization is greatly reduced because the FC remains soluble for much larger density gradients. But the flux constraint may still be potent enough to suppress instabilities and allow TB formation. Negative shear and/or large Shafranov shifts are experimental cases where TB for without velocity shear

The FC still applies, and can still be used to understand how to drive instabilities!



There is more energetics from trapped particles (not really necessary) and FC is soluble -> instability appear



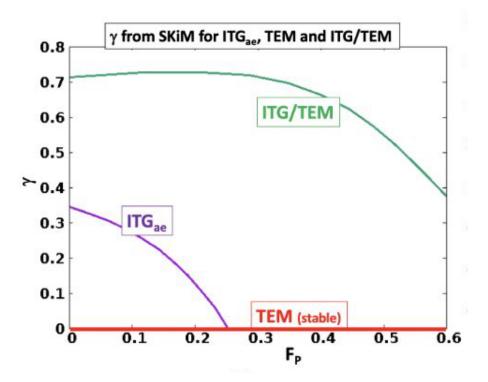


The behavior is however quite similar to what happened with adiabatic electrons.

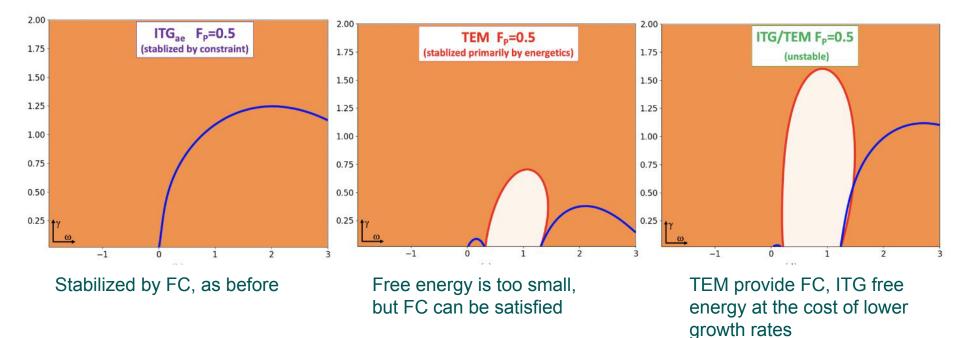
The system adapts to stabilizing effect (e.g. curvature by getting closer to adiabatic e)



Consider a single case where we vary the gradients to (de)-stabilized a mode



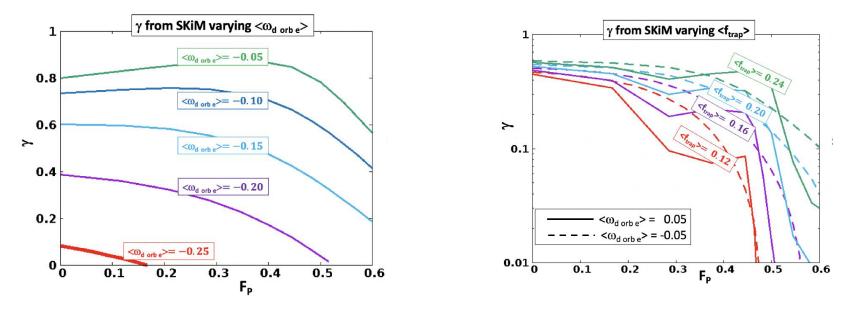




ITG and TEM are strongly coupled, their fate goes together



SKiM can be generalized to include kinetic electrons and used to understand how the system behaves.

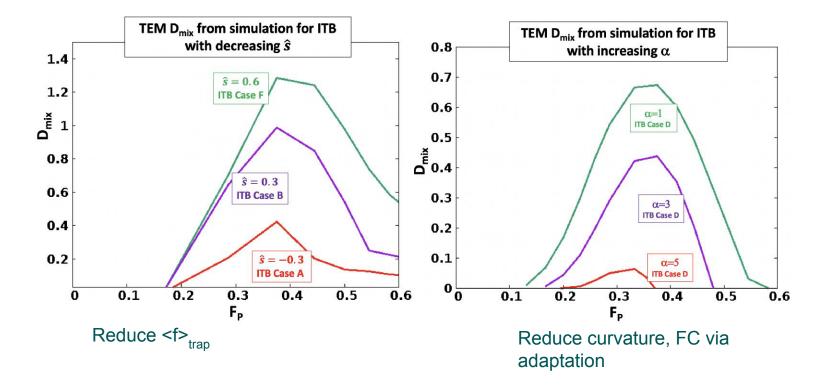


Stabilization is in fact the effect of a reduced effective trapped fraction.

The eigenfunction adapts to avoid curvature stabilization, and in doing so decouples from trapped electrons -> flux decrease

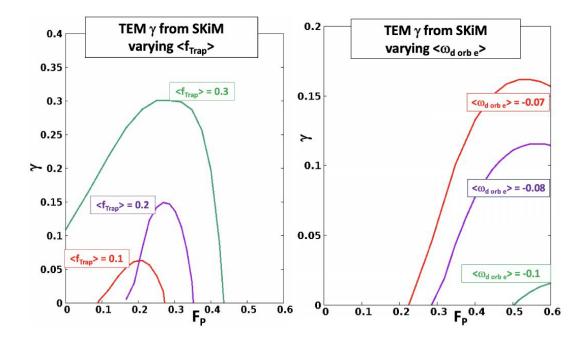


This is the regime we want, adaptation makes non-adiabatic response small, hence the FC will be dominating stability



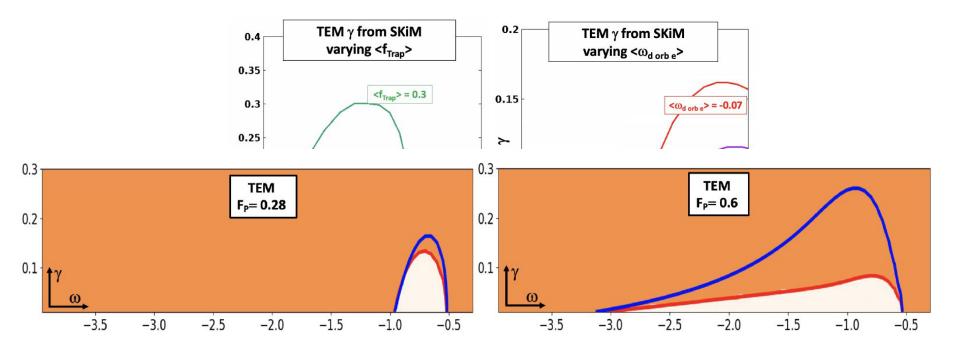


SKiM confirms that the stabilization is due to $< f_{trap} >$ and FC actively working

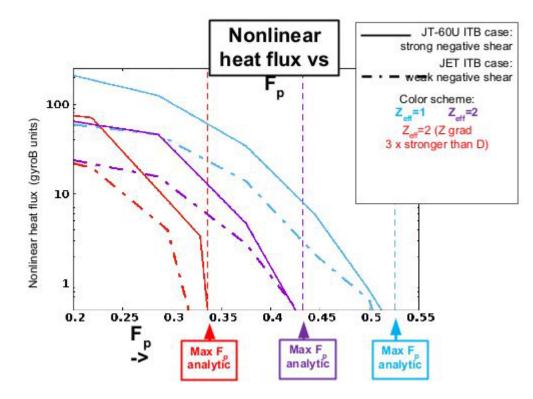




SKiM confirms that the stabilization is due to $< f_{trap} >$ and FC actively working



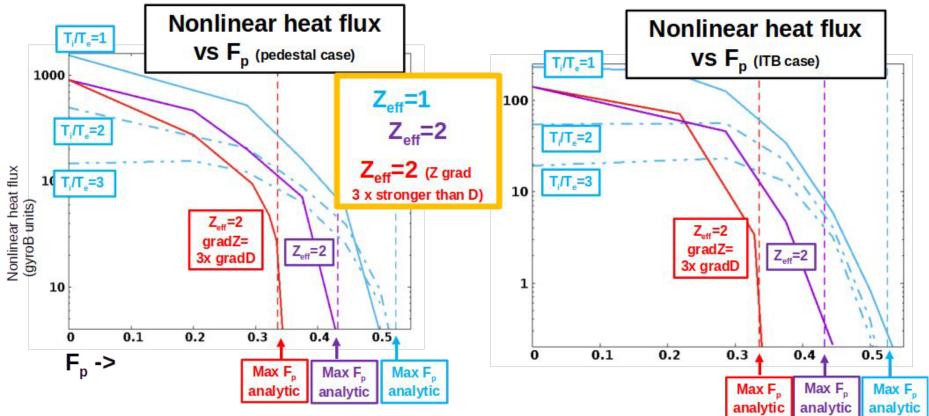
Impurities can amplify the FC effect



As we approach the limit F heat flux goes down by $\sim 2^{p}$ orders of magnitude, enough for a TB

The FC overpowers energetics





Summary

Progress towards understanding the TB formation

There are two basic dynamics that regulate the system: free energy is available for microinstability to grow if they satisfy basic laws: the radial charge flux must be zero.

More energy means more transport but only if the FC is not violated, otherwise the system cannot sustain instabilities.

A way to produce a TB is to drive the system towards conditions where violation of the FC is approached. Not an easy or granted, but there are knobs that one can leverage, i.e. tuning density gradients and impurity content.





