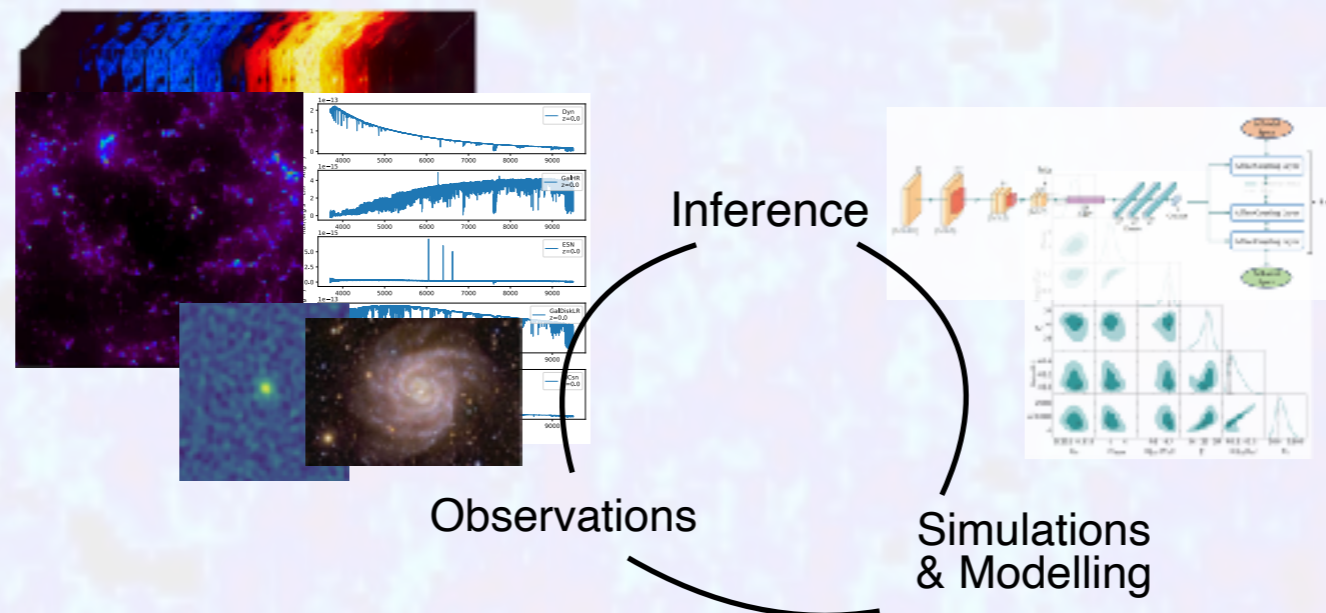


Machine Learning for Astrophysics

-

Tutorial

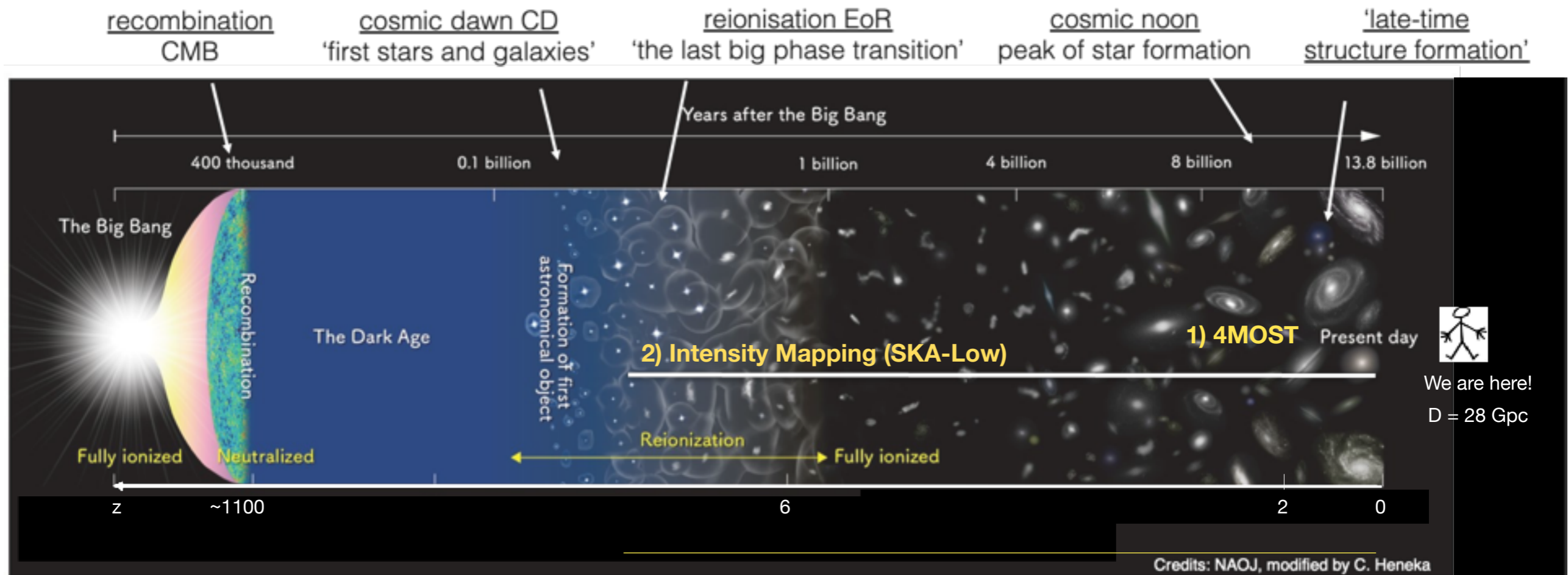


Caroline Heneka, ITP Heidelberg

Group 'Computer Vision Astrophysics and Cosmology'

Advanced School on Applied Machine Learning, ICTP Trieste, May 28th 2024

Machine Learning for Astrophysics - Tutorial



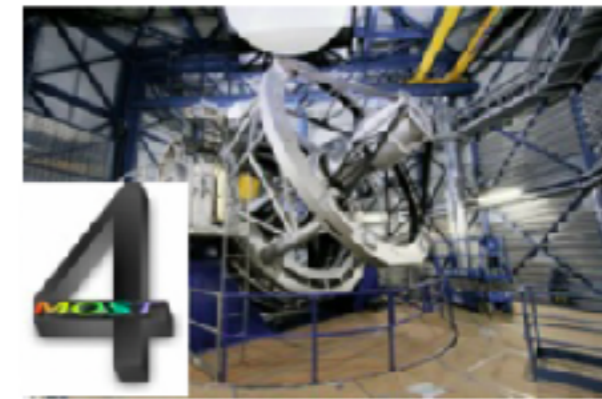
Two routes for this tutorial:

- 1) Low-z: Classification (4MOST)
- 2) High-z: Inference (SKA)

1) Low-redshift (late time): Classification of 4MOST spectra

4MOST: On-the-fly classification of spectra (1D)

- 5-year survey
- wide-field, fibre-fed, optical spectroscopy
- on ESO's 4-m-class telescope VISTA
- 2.5-degree diameter field-of-view, 2436 fibres
- HRS $R \approx 18000 - 21000$, LRS $R \approx 4000 - 7500$
- 20mio. (LRS), 3mio. (HRS) sources



<https://www.4most.eu> Credit: ESO

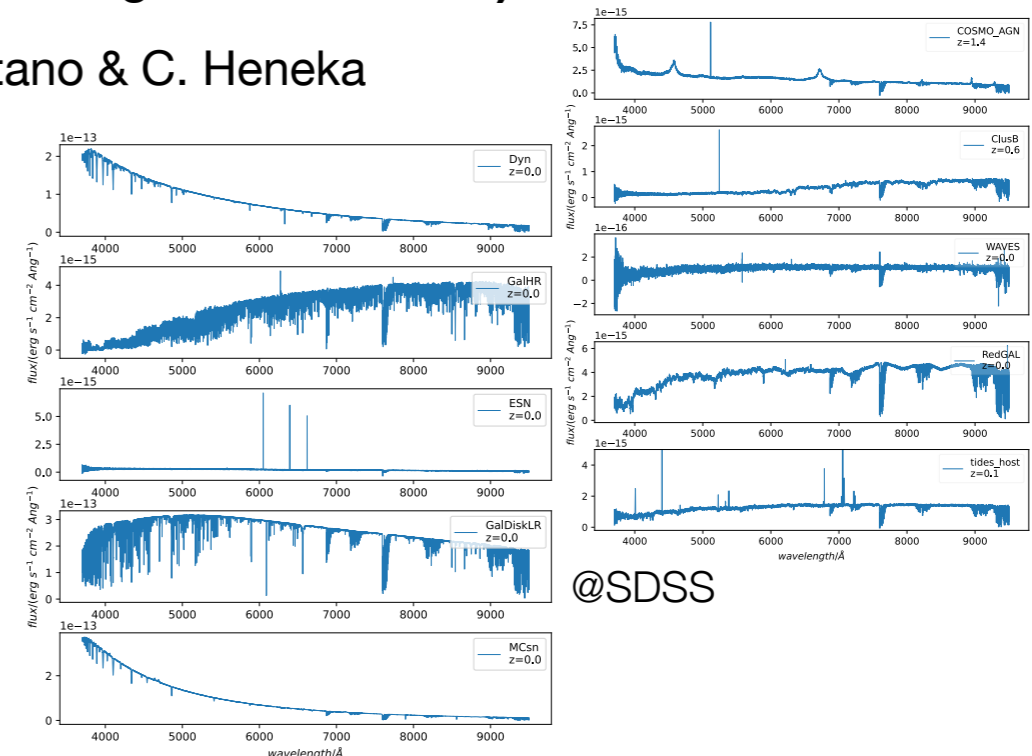
Goal: Data-driven classification pipeline layer (galactic & extragalactic sources)

Classification infrastructure working group, led by: N. Napolitano & C. Heneka



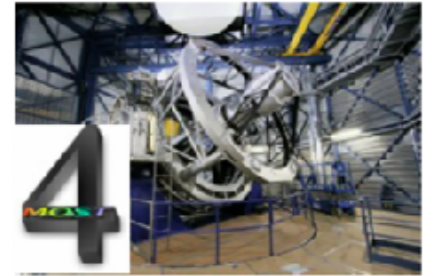
*Benchmark with
SDSS archival spectra:*

See now our tutorial!



@SDSS

1) Low-redshift (late time): Classification of 4MOST spectra



<https://www.4most.eu> Credit: ESO

4MOST: On-the-fly classification of spectra (1D)

Goal: Data-driven classification pipeline layer (galactic & extragalactic sources)

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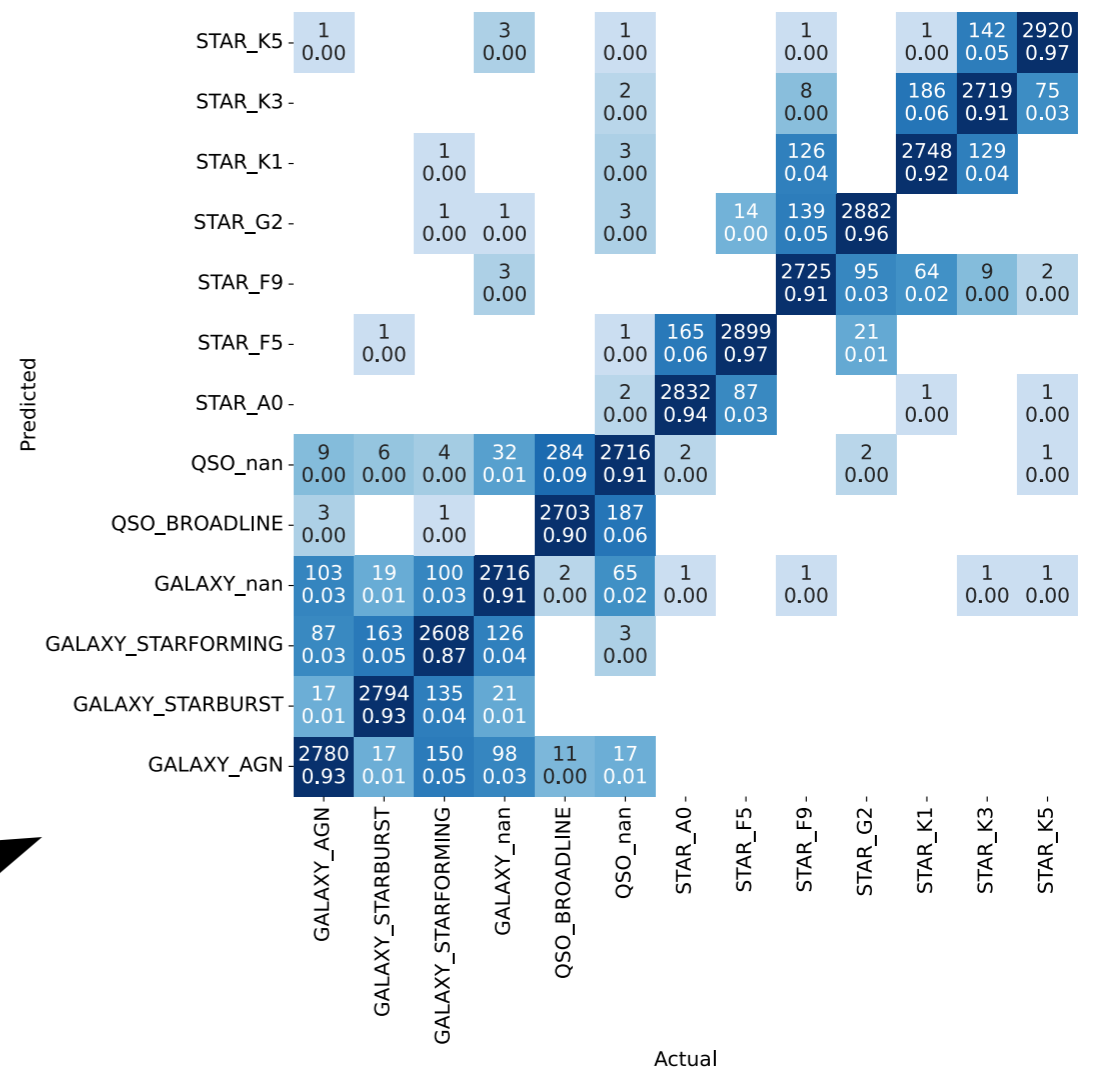
→ Probabilistic multi-classifier

*For class:
Convolutional
network variants*

*For class uncertainties:
Bayesian neural networks
and contrastive learning*

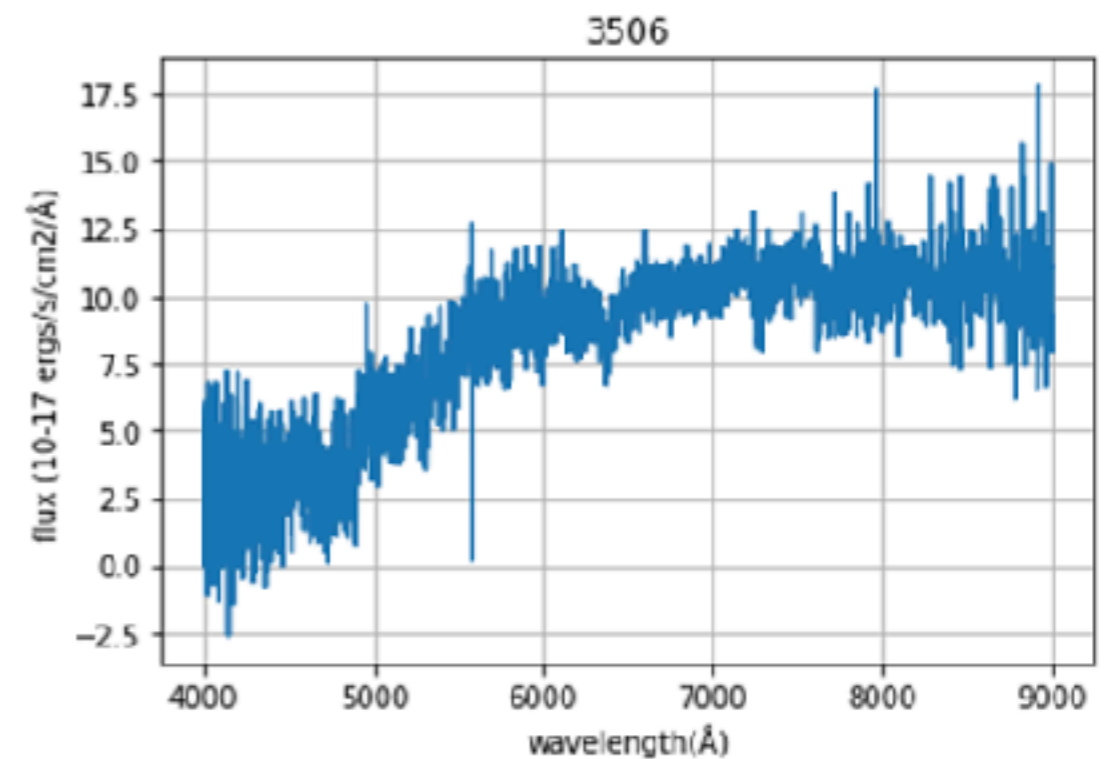
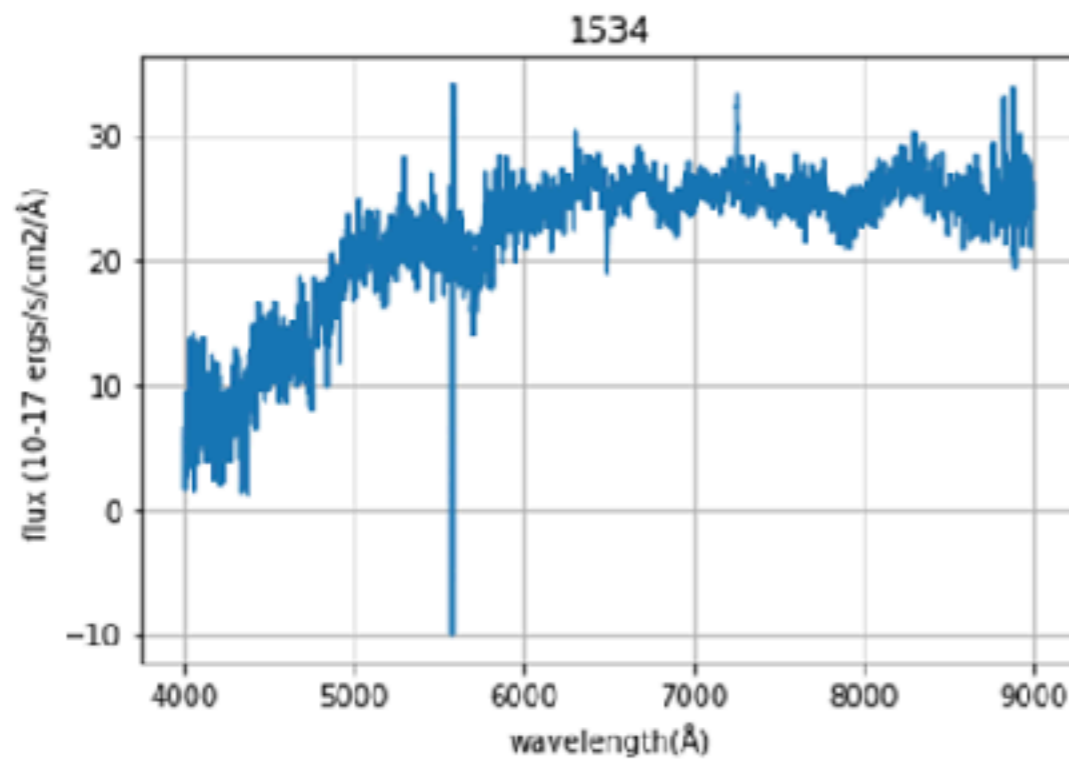
++ competitive with template fitting

Zhong, Napolitano, Heneka+ arXiv:2311.04146



This is what one wants to beat!

1) Tutorial classification challenge: 4MOST spectra



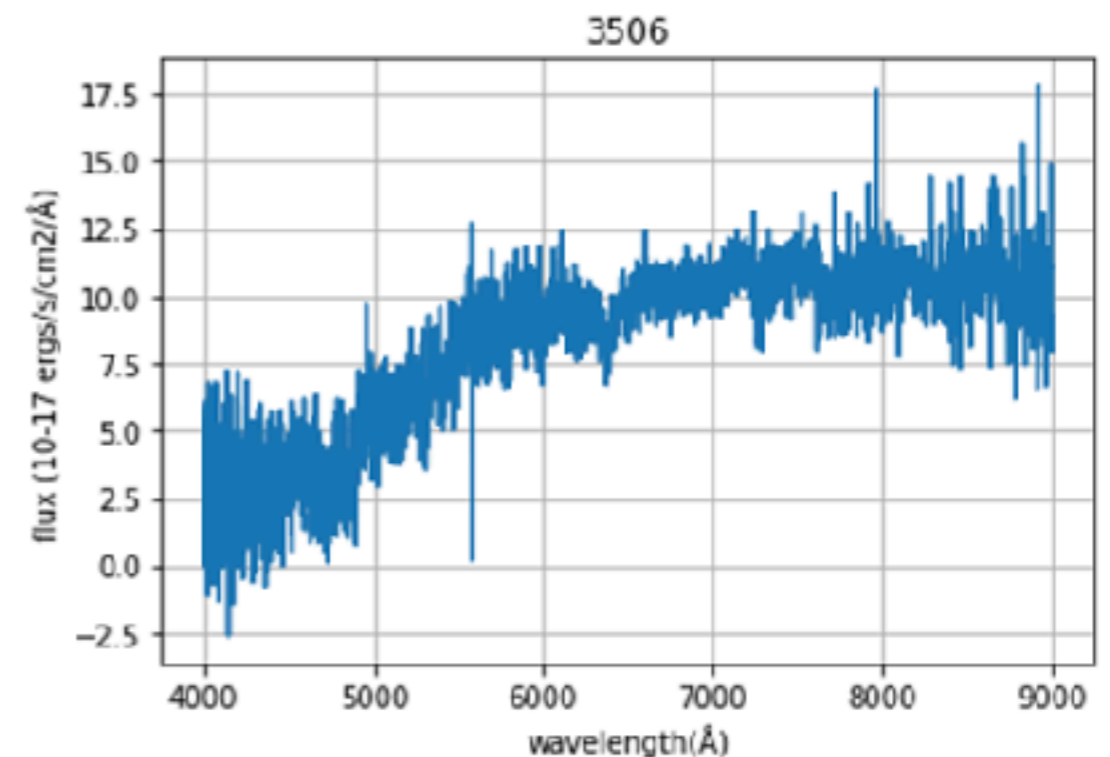
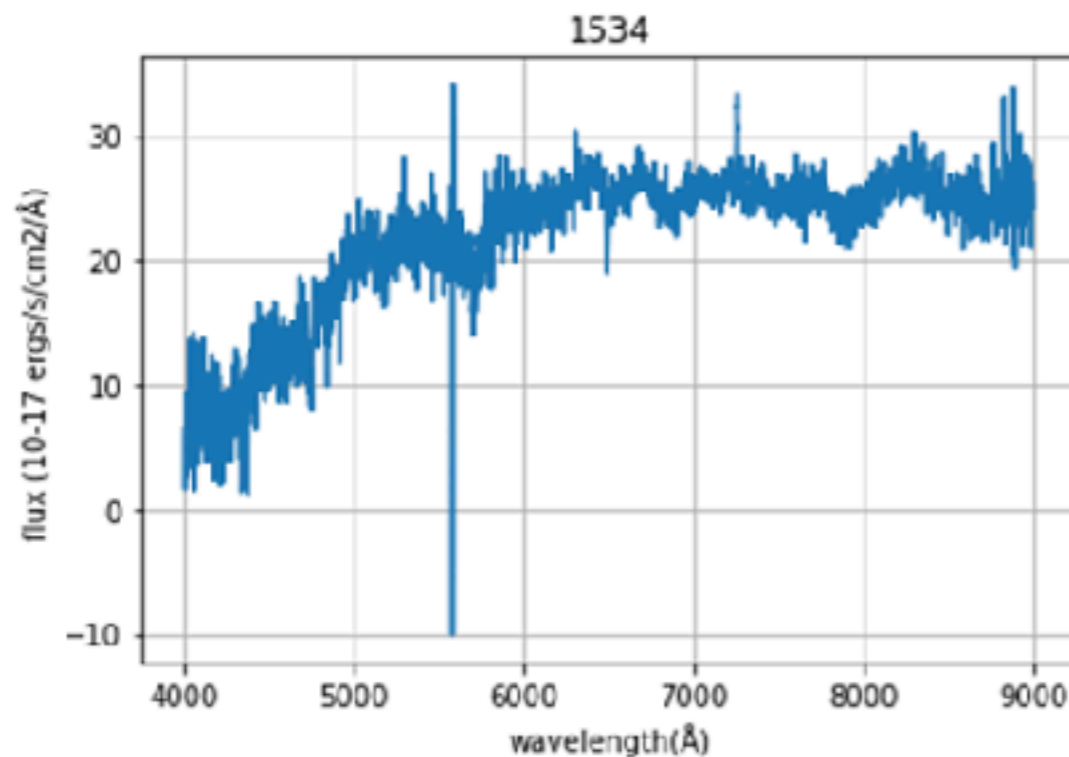
Class: Galaxy or Star? Other? A subtype?

See tutorial part:

<https://github.com/csheneka/ML-for-Astro-tutorial>

spectral_classifier.ipynb

1) Tutorial classification challenge: 4MOST spectra

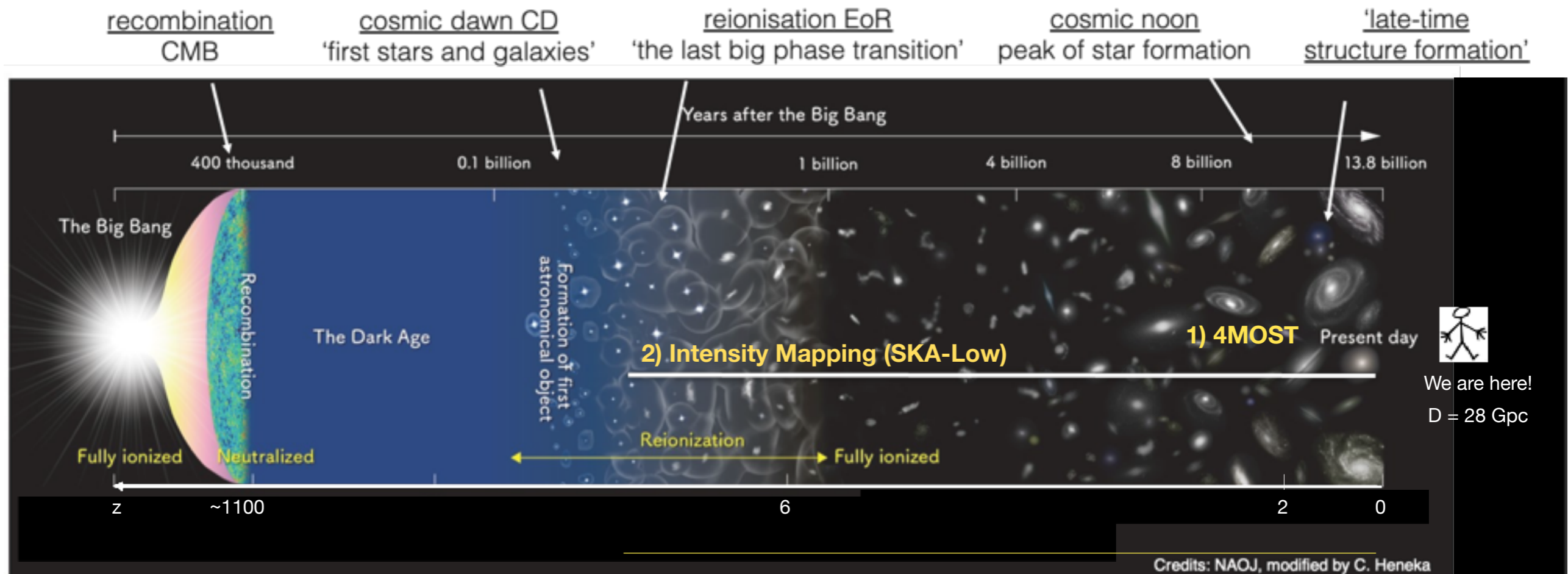


Class: Galaxy or Star? Other? A subtype?

More data (Credit: Fucheng Zhong):

https://huggingface.co/datasets/Fucheng/GaSNet-II-SDSS-dataset/tree/main/train_data

2) High-redshift: Regression for the Square Kilometre Array (SKA)



Two routes for this tutorial:

- 1) Low- z : Classification (4MOST)
- 2) High- z : Regression (SKA)

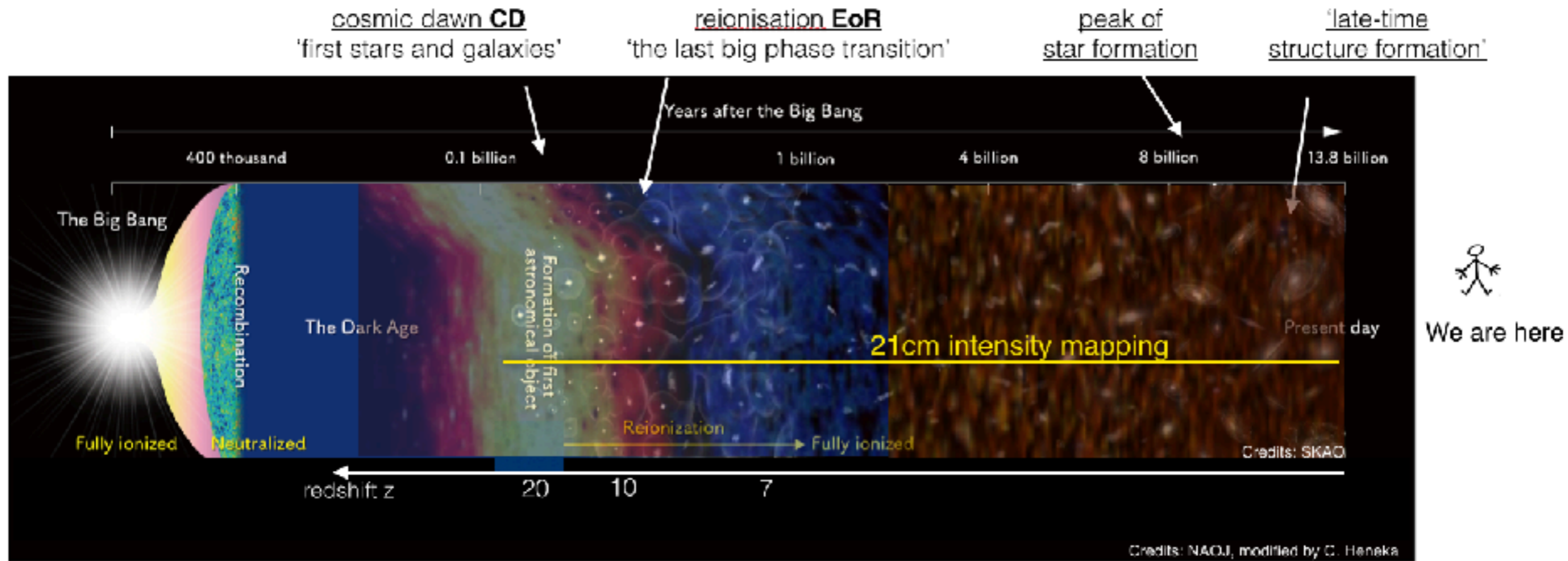
2) High-redshift: Regression for the Square Kilometre Array (SKA)



Why care?
Tomography of >80% of the Universe
Square Kilometre Array - true 'Big Data'
non-linear, non-Gaussian signal

Expected data SKA rate:
TB/s, few EB/day
Archive: ~700 PB/yr

→ Tutorial: Introspection on regression with networks



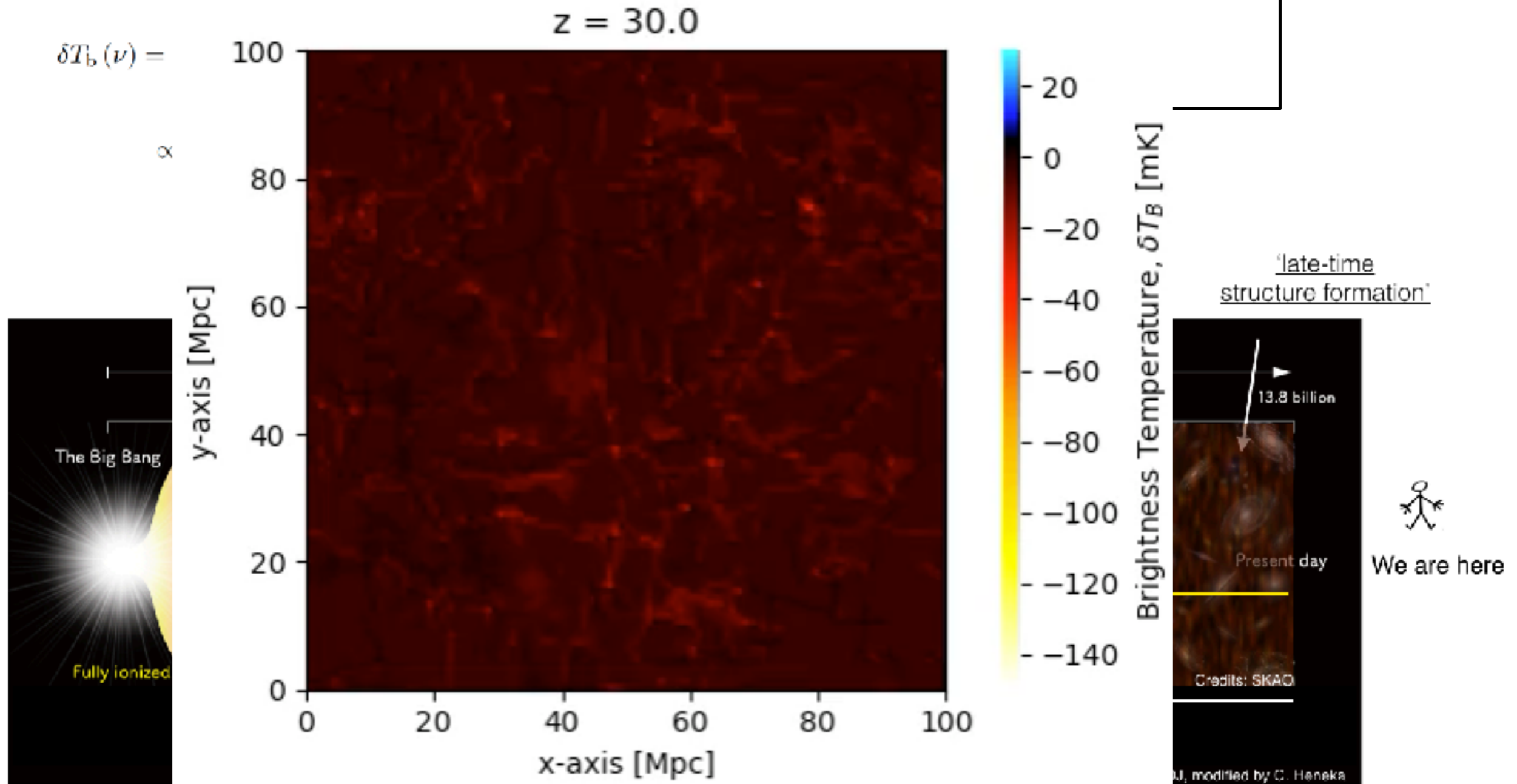
2) High-redshift: Regression for the Square Kilometre Array (SKA)

21cm signal

a tracer of neutral hydrogen:

Why care?

Tomography of >80% of the Universe



2) Brief Introspection basics for our tutorial

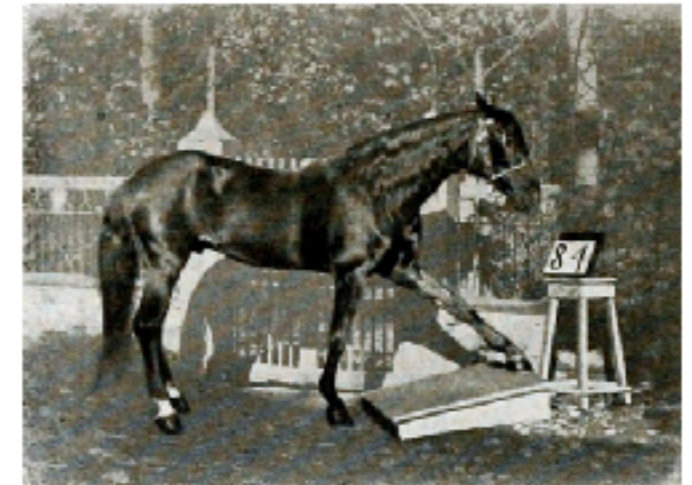
Network models are data driven.

Need introspection to:

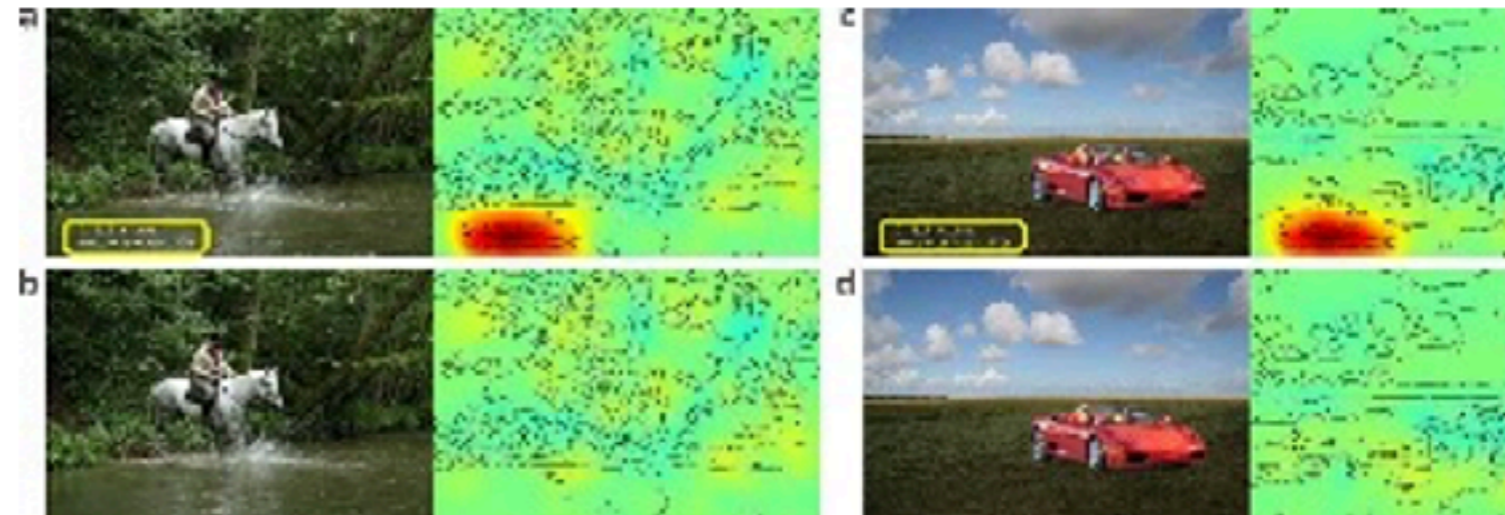
- Evaluate training **data**
- Debug, detect ‘anomalies’
- Understand and verify your **model**
- Understand **predictions**

Famous examples:

‘Clever Hans’ effect
(wrong cue, ‘overfitting’)
Over-confidence
(e.g. miss-classifying)



Hans am Tretbrett (1909)



Lapuschkin et al. 2019, arXiv: 1902.10178

2) Brief Introspection basics for our tutorial

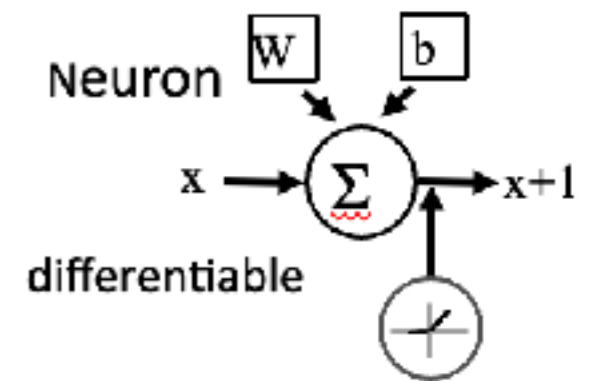
Input
Data



prediction
'the answer'



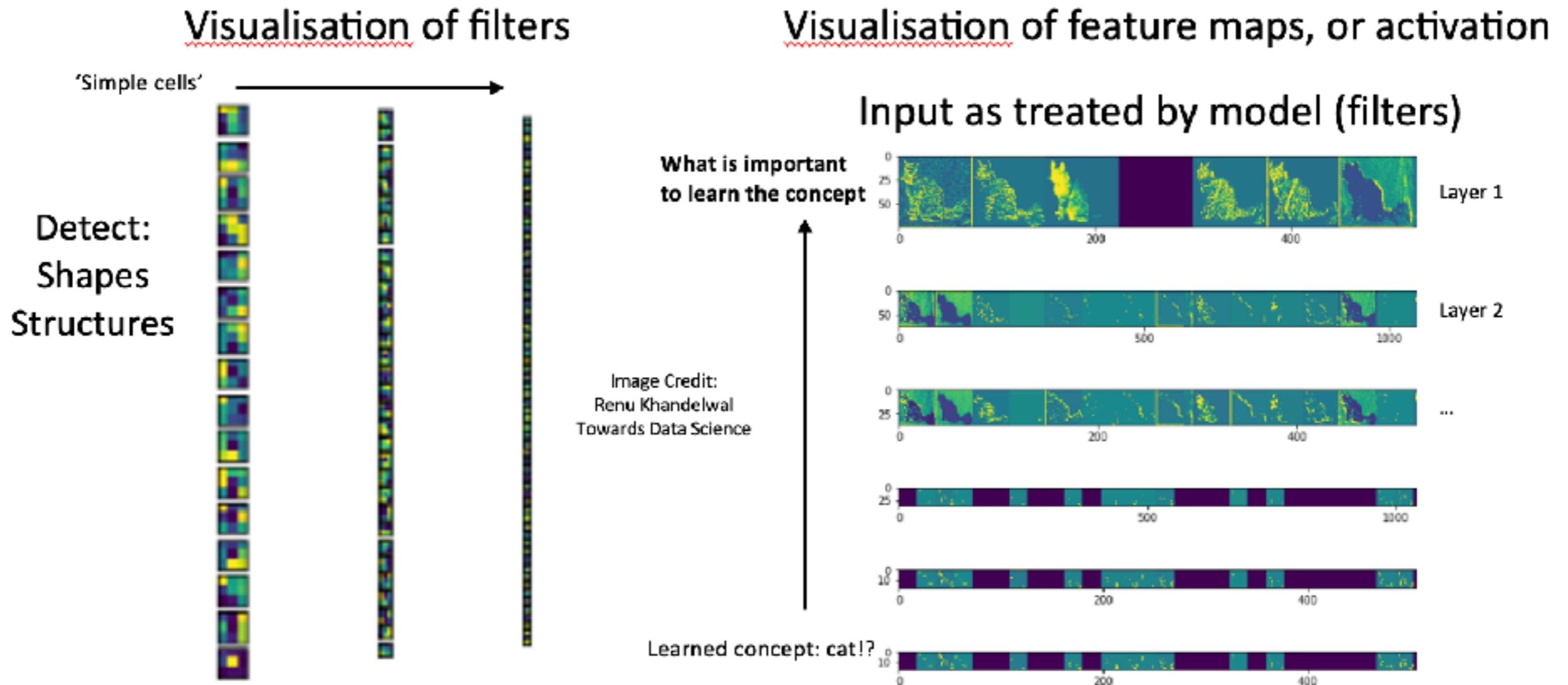
Easily order(million) parameters
~~'black box'~~



$$x_j^{(l+1)} = g \left(\sum_j w_{j,k}^{(l)} x_k^{(l)} + b_j^{(l)} \right)$$

We do know: architecture (model), best-fitting parameters (weights/bias/filter)
+ how these change during training + reaction to data → analyse

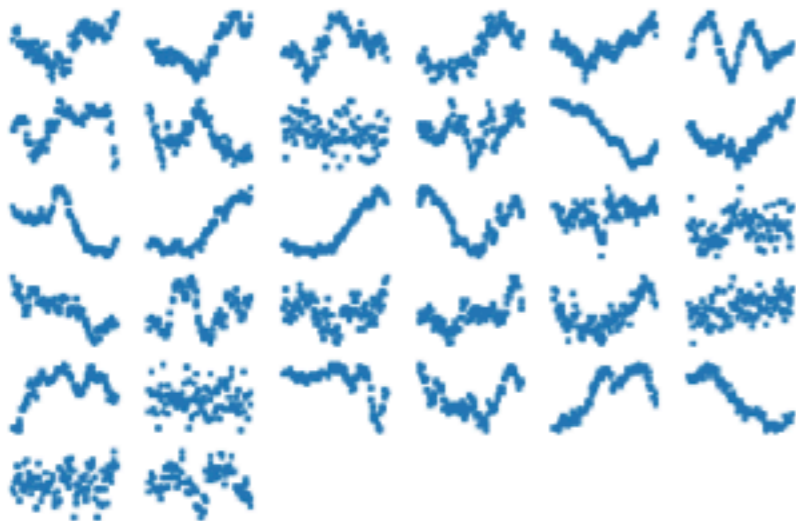
2) Tutorial part 1: Network visualisation



See tutorial part: https://github.com/csheneka/introspection-tutorial/blob/main/visualisation_3D-21cmPIE-Net.ipynb

2) Tutorial part 1: Network visualisation

```
# Task 2: Spatially average and plot filters
nfilter = 32
nplots = nfilter
n_axis = 6
for i in range(nplots):
    weights = model.layers[1].weights[0].numpy()
    weights = weights.reshape(3,3,102,nfilter)[:,:,:,:i]
    weights = np.mean(weights,axis=(0,1)) # spatial dims
    x = np.linspace(0, len(weights), len(weights))
    plt.subplot( n_axis, n_axis, 1 + i )
    plt.scatter(x,weights,s=4)
    plt.axis('off')
```



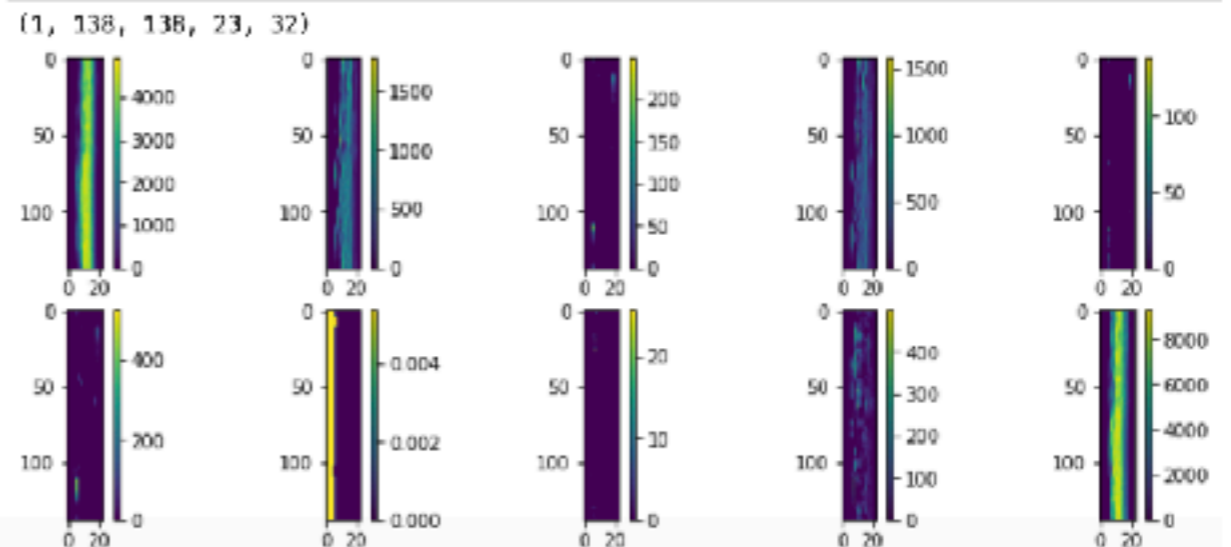
```
# plot the activation, or features
nslice = 1
nfilter = 32

lc_predict = data[0,0]
img_tensor = np.expand_dims(lc_predict, axis=0)

# get the activation for example after the first layer
activation_model = keras.Model(inputs=model.inputs, outputs=model.layers[1].output)
activation = activation_model(img_tensor)
print(activation.shape)

plt.figure(figsize=(15,15))
for i in range(32):
    plt.subplot(6,6,i+1)
    feature_map = activation[0,nslice,:,:i]
    xy = plt.imshow(feature_map)
    plt.colorbar(xy,orientation="vertical",aspect=40)

# alternate approach:
# directly multiply and scan the data with the filters to see the effect of convoluti
```



See tutorial part: <https://github.com/csheneka/introspection-tutorial>

visualisation_3D-21cmPIE-Net.ipynb

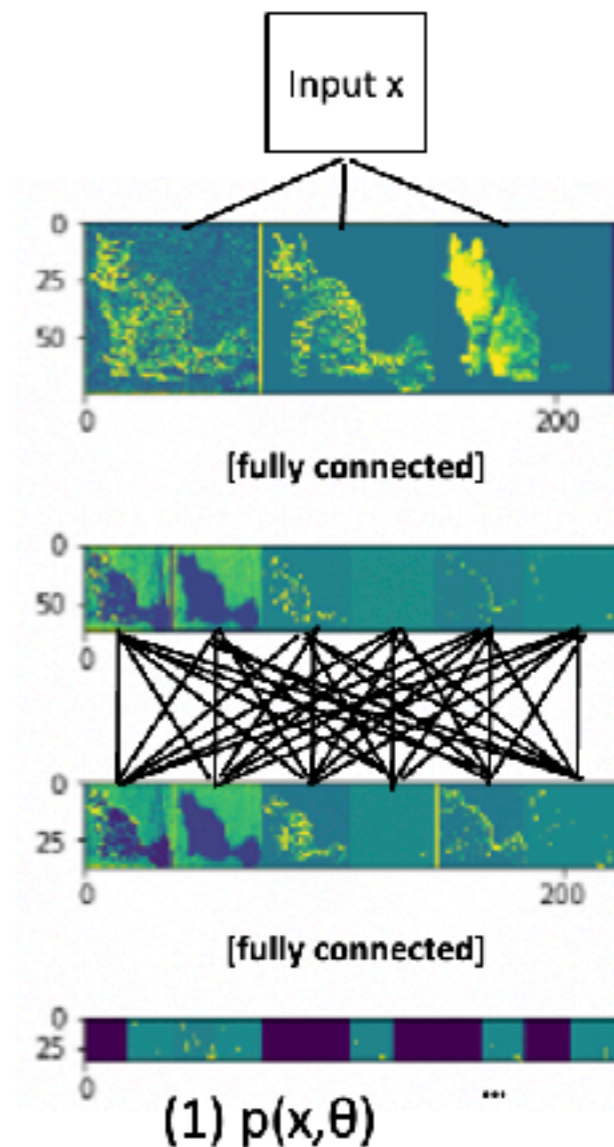
2) Tutorial part 2: Saliency maps

(1) For fixed model parameters θ , obtain for input space x the NN prediction $p(x, \theta)$

(2) Calculate the gradient map (same shape as input):

$$S = \left. \frac{\partial p(x, \theta)}{\partial x} \right|_{x=x_0}$$

(3) Potentially: Same procedure for other class of p



See tutorial part: <https://github.com/csheneka/introspection-tutorial>

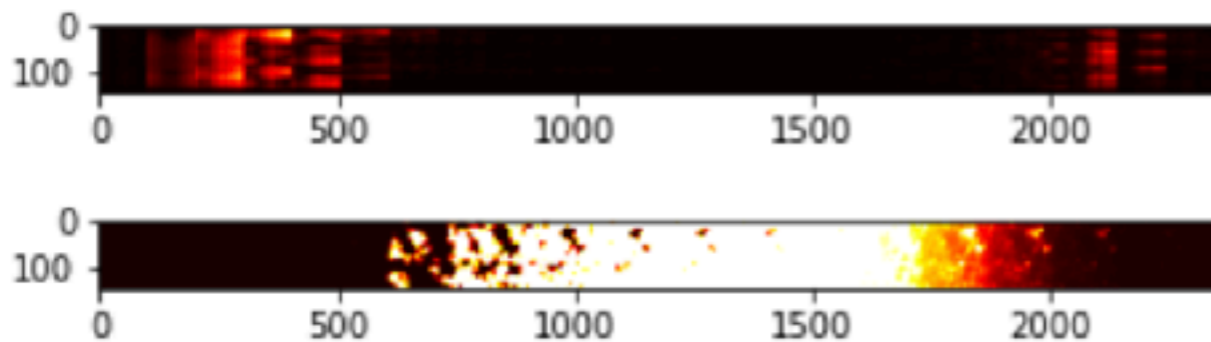
attention_exercise.ipynb

2) Tutorial part 2: Saliency maps

```
# generate saliency maps for parameters i.e. classes
parameters=[0,1,2,3,4,5]
lc = data[0,0]
for para in parameters:
    def loss(output):
        return output[0][para] # shape (samples,classes)
    map_saliency = obj_saliency(loss,lc.reshape(140,140,2350,1))

print(map_saliency.shape)
nslice = 130
plt.figure()
plot_saliency = plt.imshow(map_saliency[0,nslice,:,:], cmap=cm.hot)
plt.figure()
plot_lc = plt.imshow(lc[nslice,:,:], cmap="EoR", vmin=-150, vmax=30)
```

(1, 140, 140, 2350)



Question: What do you notice, which areas spatially and temporally are important? How does saliency shift with the inclusion of noise?

See tutorial part: <https://github.com/csheneka/introspection-tutorial>

attention_exercise.ipynb