

Lecture II: Slow-roll inflation

Recap: Horizon & Flatness problem $\left(\begin{array}{l} \frac{1}{aH} \text{ decreases,} \\ \ddot{a} < 0 \end{array} \right)$
 $\rightarrow$ solved $\omega = P/\rho < -1/3$ in the past

Slow-roll inflation:

$$S = \int d^4x \sqrt{-g} \left(\frac{1}{2} R - \frac{1}{2} g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi - V(\phi) \right), \quad \phi(x,t) = \phi(t)$$

$$\hookrightarrow \omega_\phi = \frac{P_\phi}{\rho_\phi} = \frac{\frac{1}{2} \dot{\phi}^2 - V(\phi)}{\frac{1}{2} \dot{\phi}^2 + V(\phi)} \xrightarrow{\dot{\phi}^2 \ll V} -1 < -1/3$$

"slow roll"

1) equation of motion:

$$\frac{\partial S}{\partial \phi} - \partial_\mu \frac{\partial S}{\partial \phi_{,\mu}} = \frac{1}{\sqrt{-g}} \partial_\mu (\sqrt{-g} \partial^\mu \phi) + V_{,\phi} = 0$$

↓ $\phi(t, \vec{x}) = \phi(t)$, FRW metric

$$\ddot{\phi} + 3H\dot{\phi} + V_{,\phi} = 0 \quad \& \quad H^2 = \frac{1}{3} \left(\frac{1}{2} \dot{\phi}^2 + V(\phi) \right)$$

cont. equation →

$$\left[\begin{aligned} &S = a^{-3(1+\omega)} \rightarrow \text{const for } \omega = -1 \\ &(\mp 1): H = \dot{a}/a = \sqrt{S/3} = \text{const.} \rightarrow a(t) \propto e^{Ht} \end{aligned} \right] \text{ recall}$$

→ exponentially expanding universe, dominated by potential energy of scalar field.

2) Slow roll approximation

$$\epsilon \equiv -\dot{H}/H^2 = \frac{1}{2} \frac{\dot{\phi}^2}{H^2} < 1 \quad \Leftrightarrow \text{accelerated expansion} \quad (\hat{=} \omega < -\frac{1}{3})$$

$$\eta \equiv -\ddot{\phi}/(\dot{\phi}H) < 1 \quad \Leftrightarrow \epsilon \text{ changes slowly} \rightarrow \text{sustained inflation}$$

→ slow-roll inflation: $\epsilon, |\eta| < 1$.

"Potential slow-roll parameters": $\epsilon_v = \frac{M_p^2}{2} \left(\frac{V_{,\phi}}{V} \right)^2, \quad \eta_v = M_p^2 \frac{V_{,\phi\phi}}{V}$
 $\epsilon_v, |\eta_v| \ll 1 \rightarrow \epsilon \approx \epsilon_v, \quad \eta \approx \eta_v - \epsilon_v$

$$\Rightarrow \left[3H\dot{\phi} + V_{,\phi} = 0, \quad H^2 \approx \frac{1}{3} V \approx \text{const}, \quad a(t) \sim e^{Ht} \right]$$

slow-roll eom

More convenient time coordinates (X)

• conformal time : $d\tau = dt/a \rightarrow \tau = -\frac{1}{aH}$
 $(\rightarrow ds^2 = a^2(-d\tau^2 + d\vec{r}^2))$ $\leftarrow < 0$
during inflation

• e-folds : $dN = -H dt$ (elapsed Hubble times)
 $= -\frac{d\alpha}{\alpha} \frac{dt}{dt} = -d \ln a$ linear measure of exp. expansion

$$N(\phi) = \ln(a_{\text{end}}/a) = \int_{\phi}^{\phi_{\text{end}}} \frac{H}{\dot{\phi}} d\phi = \int_{\phi_{\text{end}}}^{\phi} \frac{1}{\sqrt{2}\epsilon} d\phi$$

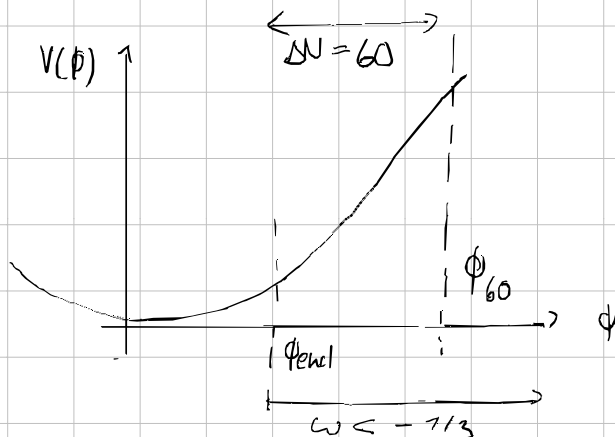
Solving horizon & flatness problem requires
 $N \gtrsim 60$

3) An example : $V(\phi) = \frac{1}{2} m^2 \phi^2$

$$\epsilon_v = \frac{M_p^2}{2} \left(\frac{V'}{V} \right)^2 = 2 \left(\frac{M_p}{\phi} \right)^2 \leq 1 \rightarrow \phi_{\text{end}} (\epsilon=1) = \sqrt{2} M_p$$

$$N(\phi) = \frac{1}{M_p} \int_{\phi_{\text{end}}}^{\phi} \frac{1}{\sqrt{2}\epsilon} d\phi = \frac{\phi^2}{4M_p^2} - \frac{1}{2} \stackrel{!}{=} 60 \rightarrow \phi_{60} = 15 M_p$$

(inflation for $\phi > \phi_{\text{end}}$)
 (start of observable inflation)



→ working inflation model

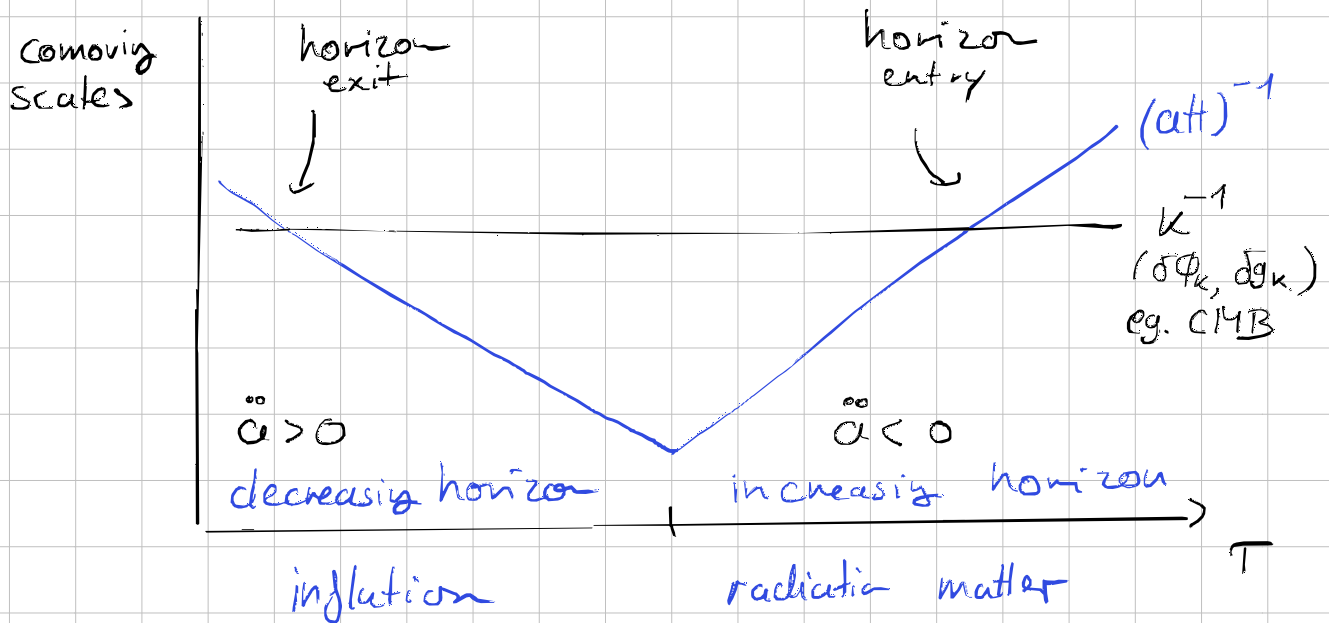
4) Scales and horizons

$$\Phi(x,t) = \bar{\Phi}(t) + \delta\Phi(x,t)$$

$$g_{\mu\nu}(x,t) = \bar{g}_{\mu\nu}(t) + \delta g_{\mu\nu}(x,t)$$

homogeneous background, evolves under slow-roll eom

small perturbations, to first order, evolve linearly
 $\rightarrow$ Fourier transform $\Phi_{\vec{k}}, h_{\vec{k}}$



this lecture

$\hookrightarrow$ see Marko & Mathew's lectures

SVT decomposition and gauge invariance (GR)

$\delta\phi(x,t) \rightarrow$ scalar clbf

$\delta g_{\mu\nu}(x,t) \rightarrow$ 4 scalars, 2 vectors, 1 tensor

↓
2 can be gauged away under coordinate transformation $x^\mu \rightarrow x^\mu + \xi^\mu$

↓
decay

↓
Lecture 4

define gauge invariant scalar

eg. $R = \psi + \frac{H}{\dot{\phi}} \delta\phi$

$g_{ij} = a^2[(1+2\psi)\delta_{ij} + \dots]$

"comoving curvature perturbation"

= spatial curvature on constant- ϕ hypersurface

choose a gauge:

eg. Newtonian gauge
(Matthews lecture)

or: $\delta\phi = 0$

$g_{ij} = a^2[(1-2R)\delta_{ij} + h_{ij}]$

"comoving gauge"

↓
one dynamical scalar

$R(\vec{x}, t)$, h_{ij} = pure tensor

↓
Lecture 3

Notes :

slow roll parameters.

$$\omega = \frac{x-1}{x+1} \simeq -1+2x = -1 + \frac{2}{3}\epsilon$$

$$x = \frac{1/2 \dot{\phi}^2}{V \ll 3H^2} = \frac{1}{3}\epsilon \quad \Rightarrow \quad \omega < -1/3 \quad \stackrel{\wedge}{=} \quad \epsilon < 1$$