Optical lattice atomic clocks

Michał Zawada KL FAMO, Nicolaus Copernicus University

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Optical lattice atomic clocks

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Time standards

Optical clocks 1

Lamb-Dicke regime Wannier-Stark regime

Laser cooling and trapping

Optical clocks 2

Magic wavelenght Ultra-stable laser Clock cycle

La seconde

est la durée de 9 192 631 770 périodes de la radiation correspondant à la transition entre les deux niveaux hyperfins de l'état fondamental de l'atome de césium 133.

The second is the duration of 9 192 631 770 periods of the radiation corresponding to the transition between the two hyperfine levels of the ground state of the caesium 133 atom.

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Deep space missions



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stock exchange



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Fundamental research

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Transport



E-commerce, banking,

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Fundamental research



smart grid



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Fundamental research

Electric power

smart grid





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Global GNSS market size: 250 billion € in 2016

source: European GNSS Agency (GSA), 4rd Market Report (2015)

Telecommunications: one trillion \$ market by 2019

source: Global Wireless Infrastructure Market 2013 Forecast to Industry Size, Shares, Strategies, Trends, and Growth 2019

Smart electrical grid: 400 billion € by 2020

source: GTM Research, Global Smart Grid Technologies and Growth Markets, 2013-2020

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Direct applications

JRP-s16 OC18 "Optical clocks with 1E-18 uncertainty" Letters of support

Applications of optical clocks

Affiliation	Country	Area of expertise
CCTF (Consultative Committee on Time and	International	SI second
Frequency)		
Space Optical Clocks	International	Science in space
ESA (European Space Agency)	International	Science in space
ACES (Atomic Clock Ensemble in Space)	International	Science in space
University of Liverpool	UK	Oceanography
Institut für Erdmessung, Hannover	Germany	Geodesy
British Geological Survey	UK	Geodesy
International Association of Geodesy	International	Geodesy
JIVE (Joint Institute for VLBI ERIC)	International	Radio astronomy
British Telecommunications	UK	Telecommunications
Orange Polska	Poland	Telecommunications

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Atomic clock

$$\Delta E = \text{const.}$$

Atomic clock diagram



Ideal clock: a signal with stable and universal frequency.

 Energy levels of unperturbed atoms are stable and universal.

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Atomic clock

$$\Delta E = \text{const.}$$

Atomic clock diagram



$$\omega(t) = \omega_{ef} * (1 + \varepsilon + y(t))$$

 ε -fractional offset of frequency y(t) - fractional fluctuations of frequency

Accuracy

- uncertainty of ε

Stability

- statistical properties of y(t), interrogation characterized by the Allan deviation $\sigma_{\gamma}(\tau)$

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Accuracy and stability



source: nist.gov

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Allan deviation



 f_n is a set of frequency offset measurements that consists of individual measurements, f_1 , f_2 , f_3 , and so on and the data are equally spaced in segments τ seconds long.

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Time standards

- 1. Atomic fountains, accuracy of $\sim 10^{-16}$
- 2. Commercial caesium clocks, with good long term stability $\sim 10^{-15}$ over few months and accuracy of $\sim 10^{-13}$
- 3. Hydrogen masers: 1.4 GHz hyperfine structure transition in atomic hydrogen. Much better short-time stability than any commercial caesium clock: $\sim 10^{-15}$ over few hours

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Caesium clocks and hydrogen masers



HP5071A caesium clock and VCH-1005

hydrogen maser in the Central Office of

Measures in Poland



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Caesium clocks and hydrogen masers



Feng-Lei Hong et al. Opt. Express 13, 5253-5262 (2005)

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Atomic fountain



Best accuracy: 2.6×10⁻¹⁶ Real-time control of collision shift with adiabatic

passage: Phys. Rev. Lett. 89, 233004 (2002)

Resolution 6x10⁻¹⁷ at 50 days (assuming white noise)

v(FO2-Rb) (2007) =6 834 682 610.904 309 (8) Hz

Total uncertainty 1.1x10-15

Total uncertainty	4.9	2.6	4.5	6.6
Others (quantum motion, Background gas collisions, Ramsey & Rabi pulling,)	< 2.0	< 2.0	< 2.5	< 1.7
Microwave leakage and phase perturbations	< 0.5	< 0.5	< 0.1	< 6
First order Doppler effect	< 3.0	-0.96 ± 0.84	< 2.0	
Cold collisions and cavity pulling	245.3 +/- 3	191.6 +/- 0.8	0 +/- 2.5	16.7 +/- 1.7
Blackbody radiation	164.9 +/- 0.6	167.2 +/- 0.6	120.6 +/- 1.6	165.3 +/- 0.6
Quadratic Zeeman effect	-1276.2 +/- 0.2	-1919.5 +/- 0.2	-3472 +/- 0.2	-305.7 +/- 1.1
Corr. +/-Uncert. (10 16)	FO1	F02-Cs	FO2-Rb	FOM

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Source: Systèmes de Références Temps-Espace



Helen Margolis, Nature Physics 10, 82-83 (2014)

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Quality of the clock: $Q = \frac{\nu}{\Delta \nu} \times \frac{S}{N}$

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Quality of the clock: $Q = \frac{\nu}{\Delta \nu} \times \frac{S}{N} \sim \nu T \times \frac{S}{N}$

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Quality of the clock: $Q = \frac{\nu}{\Delta \nu} \times \frac{S}{N} \sim \nu T \times \frac{S}{N}$ Stability of an atomic clock: $\sigma_y(\tau) \sim \frac{\sigma_{spect}}{Q} \sqrt{\frac{T_c}{\tau}}$ Optical lattice atomic clocks

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 $\Delta \nu$

Quality of the clock: $Q = \frac{\nu}{\Delta \nu} \times \frac{S}{N} \sim \nu T \times \frac{S}{N}$ Stability of an atomic clock: $\sigma_y(\tau) \sim \frac{\sigma_{spect}}{Q} \sqrt{\frac{T_c}{\tau}}$ Quantum Shot Noise limitation: $\sigma_y(\tau) = \frac{1}{\pi Q} \times \frac{1}{N_{at}} \times \sqrt{\frac{T_c}{\tau}}$ Optical lattice atomic clocks

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Optical clocks: trapped ions and neutral atoms

lon traps: lons trapped in the Paul trap by the RF field

► Trap is perturbed only slightly \Rightarrow excellent accuracy 3×10^{-18}

N. Huntemann et al. Phys. Rev. Lett.

116, 063001 (2016)

► Good stability $(3 \times 10^{-15}/\sqrt{\tau})$, but restricted by Quantum Shot Nose - only 1 ion. Optical lattice atomic clocks

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> N. Huntemann et al. Phys. Rev. Lett. 116. 063001 (2016)

- ► Good stability $(3 \times 10^{-15}/\sqrt{\tau})$, but restricted by Quantum Shot Nose only 1 ion
 - only 1 ion.

Neutral atoms: Optical lattice

• A trap with a hight-intensity light \Rightarrow high perturbation, though well under control. $2 * 10^{-18}$

T.L. Nicholson et al. Nature Communications, 6, 6896 (2015)

► High number of atoms (10⁴)⇒ high stability possible. $1.8 \times 10^{-16}/\sqrt{\tau}$ down to 2×10^{-18}

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Optical clocks: trapped ions and neutral atoms



A. G. Smart, Phys. Today 67, 3, 12 (2014)

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A transition good for optical clock

Natural linewidth of an electric dipole (E1) transition is typically bigger than 1 MHz \Rightarrow Q below 10⁸

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A transition good for optical clock

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Clock transition should be:

- Narrow (forbidden)
- Mostly insensitive to external fields.

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A transition good for optical clock

Natural linewidth of an electric dipole (E1) transition is typically bigger than 1 MHz \Rightarrow Q below 10^8

Clock transition should be:

- Narrow (forbidden)
- Mostly insensitive to external fields.

Possible candidates:

- two-photon transitions and higher order electric transitions (quadrupole, octupole ...)
- a low energy nuclear transition (still sought at ²²⁹Th)
- an intercombination transition

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- Forbidden ${}^{1}S_{0} {}^{3}P_{1}$ transition:
 - Fine structure interaction

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- Forbidden ${}^{1}S_{0} {}^{3}P_{1}$ transition:
 - Fine structure interaction
 - Double forbidden ${}^{1}S_{0} {}^{3}P_{0}$ transition:
 - Fermions: hyperfine interaction
 - Bosons: quenching by a static B field

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4He 20Ne

⁴⁰Ar ¹³²Xe

$$f(v) \sim \frac{v^2}{T^{3/2}} e^{-\frac{mv^2}{2kT}}$$

Maxwell-Boltzmann Molecular Speed Distribution for Noble Gases

Speed (m/s)



500 1000 1500 2000 2500

0.004

0.003

0.002

0.001

Probability density (s/m)

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source: Pdbailey at Wikipedia

$f(v) \sim \frac{v^2}{T^{3/2}} e^{-\frac{mv^2}{2kT}}$

• Doppler shift:
$$\nu = \nu_0 \left(1 + \frac{v}{c}\right)$$

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source: Pdbailey at Wikipedia

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$$\nu = \nu_0 \left(1 + \frac{v}{c}\right)$$

Distribution of the observed frequencies:

$$P_{\nu}(\nu)d\nu = P_{\nu}(\nu)\frac{d\nu}{d\nu}d\nu$$

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 Distribution of the observed frequencies:

$$P_{\nu}(\nu)d\nu = P_{\nu}(\nu_{\nu})\frac{d\nu}{d\nu}d\nu$$

$$\int P_{\nu}(\nu) d\nu \sim \sqrt{\frac{1}{T^2 \nu_0^2}} \exp\left(-\frac{mc^2(\nu-\nu_0)^2}{2kT\nu_0^2}\right) d\nu$$

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• Doppler shift:
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 Distribution of the observed frequencies:
 D (x) dx - D (x) dy dy

$$P_{\nu}(\nu)d\nu = P_{\nu}(\nu_{\nu})\frac{d\nu}{d\nu}d\nu$$

$$\sqrt{\frac{1}{T^2\nu_0^2}} \exp\left(-\frac{mc^2(\nu-\nu_0)^2}{2kT\nu_0^2}\right) d\nu$$
$$\blacktriangleright \Delta\nu_{FWHM} = \sqrt{\frac{8kT\ln 2}{mc^2}}\nu_0$$

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source: Pdbailey at Wikipedia

•
$$f(v) \sim \frac{v^2}{T^{3/2}} e^{-\frac{mv^2}{2kT}}$$

• Doppler shift:
$$\nu = \nu_0 \left(1 + \frac{v}{c}\right)$$

Distribution of the observed frequencies: $P_{\nu}(\nu)d\nu = P_{\nu}(v_{\nu})\frac{d\nu}{d}d\nu$

$$P_{\nu}(\nu)d\nu \sim \sqrt{\frac{1}{T^{2}\nu_{0}^{2}}} \exp\left(-\frac{mc^{2}(\nu-\nu_{0})^{2}}{2kT\nu_{0}^{2}}\right)d\nu$$

• $\Delta \nu_{FWHM} = \sqrt{\frac{8kT \ln 2}{mc^2}} \nu_0 \sim GHz$ at room temperature

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Maxwell-Boltzmann Molecular Speed Distribution for Noble Gases

source: Pdbailey at Wikipedia

Sub-Doppler spectroscopy, or cool down atoms...

$$f(v) \sim \frac{v^2}{T^{3/2}} e^{-\frac{mv^2}{2kT}}$$

• Doppler shift:
$$\nu = \nu_0 \left(1 + \frac{v}{c}\right)$$

 Distribution of the observed frequencies:

$$P_{\nu}(\nu)d\nu = P_{\nu}(\nu_{\nu})\frac{d\nu}{d\nu}d\nu$$

$$\int \mathcal{P}_{\nu}(\nu) d\nu \sim \sqrt{\frac{1}{T^2 \nu_0^2}} \exp\left(-\frac{mc^2(\nu-\nu_0)^2}{2kT\nu_0^2}\right) d\nu$$

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Unperturbed Hamiltonian of the atom: $H_0 = \hbar \omega_g \ket{g} \bra{g} + \hbar \omega_e \ket{e} \bra{e}$

Zero point energy can be chosen at will: $H_0 = \hbar \omega_0 |e\rangle \langle e|$

Optical lattice atomic clocks

Michał Zawada KL FAMO, Nicolaus Copernicus University

Time standards

Optical clocks 1

Lamb-Dicke regime Wannier-Stark regime

Laser cooling and trapping

Optical clocks 2

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Monochromatic electromagnetic wave: $\vec{E}(t, \vec{r}) = \vec{E_0}e^{-\hat{\imath}(\omega t + \phi(\vec{r}))} + c.c.$

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 $\implies \mathcal{H}_{int} = \frac{\hbar}{2} \Omega(\vec{r}) e^{-\hat{\imath}\omega t} \left(|e\rangle \langle g| + |g\rangle \langle e| \right) + \frac{\hbar}{2} \Omega^*(\vec{r}) e^{\hat{\imath}\omega t} \left(|g\rangle \langle e| + |e\rangle \langle g| \right)$ where $\Omega(\vec{r}) = -2d \cdot \vec{E_0} e^{-\phi(\vec{r})} / \hbar$ is the so-called Rabi frequency.

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After rotating wave approximation ($|\Delta| = |\omega - \omega_0| \ll \omega_0$):

$$H_{int}^{RWA} = rac{\hbar}{2} \Omega(\vec{r}) e^{-\hat{\imath}\omega t} \ket{e} ra{g} + rac{\hbar}{2} \Omega^*(\vec{r}) e^{\hat{\imath}\omega t} \ket{g} ra{e}$$

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Using the abbreviation for the generalized Rabi frequency: $\tilde{\Omega}=\sqrt{\Delta^2+\Omega^2}$

the probability of finding the atom in the upper level state is:

$$egin{aligned} P_e(t) &= rac{\Omega}{ ilde{\Omega}}\sin^2(ilde{\Omega}t/2) \ P_e(\omega) &\sim rac{\sin^2(\omega t/2)}{(\omega t/2)^2}. \end{aligned}$$

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$$H_{int}^{RW\!A}=rac{\hbar}{2}\Omega(ec{r})e^{-\widehat{\imath}\omega\,t}\ket{e}ig\langle g
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the probability of finding the atom in the upper level state is:

 $P_e(t) = rac{\Omega}{ar{\Omega}} \sin^2(ar{\Omega}t/2).$ $P_e(\omega) \sim rac{\sin^2(\omega t/2)}{(\omega t/2)^2}.$

In the presence of sufficient decoherence, $P_e(\omega)$ becomes a Lorentzian line shape whose width is the decoherence rate Γ divided by 2π .

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In the presence of sufficient decoherence, $P_e(\omega)$ becomes a Lorentzian line shape whose width is the decoherence rate Γ divided by 2π . Doppler shift across the atomic velocity distribution broadens the line shape into a Gaussian or Voigt line shape.

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In a harmonic potential the atomic motion is not a continuous variable, but is restricted to the quantized motional states $|n\rangle$ of the system.

 \implies New term in Hamiltonian: $H_{osc} = \hbar \omega_h \left(a^{\dagger} a + \frac{1}{2} \right)$.

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Modified Rabi rate of transitions in two-level atom from $|g\rangle |n\rangle$ state to $|e\rangle |m\rangle$ state:

$$\begin{split} \Omega_{mn} &= \tilde{\Omega} \langle m | e^{\hat{i} \eta \left(a + a^{\dagger} \right)} | n \rangle = \tilde{\Omega} e^{-\eta^2 / 2} \sqrt{\frac{n < !}{n > !}} \eta^{|m-n|} L_{n <}^{|m-n|} (\eta^2) \\ \text{where } \eta &= \sqrt{\frac{\hbar k^2}{2m\omega_h}} = \frac{k x_0}{\sqrt{2}} \text{ is the so-called Lamb-Dicke parameter} \end{split}$$

 $\eta \ll 1$ required!

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frequency (units of trap frequency, ω)

source: Ludlow et al. Rev. Mod. Phys. 87, 637 (2015)

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Back to two-level atom ($\hbar\omega_0$) in the EM field (ω).

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Back to two-level atom $(\hbar\omega_0)$ in the EM field (ω) .

Electric dipole moment: $\vec{p}(\vec{r}, t) = \alpha \vec{E}(\vec{r}, t)$, where α is the complex polarizability of the atom.

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Potential of interaction between the light and the dipole moment: $U_{dip} = -\frac{1}{2} \left\langle \vec{p} \cdot \vec{E} \right\rangle = -\frac{1}{2\varepsilon_0 c} Re(\alpha) I(\vec{r})$

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In the classical Lorentz Oscillator model of an atom $(\ddot{x} + \Gamma_\omega \dot{x} + \omega_0^2 x = -eE(t)/m_e)$

$$\alpha = 6\pi\varepsilon_0 c^3 \frac{\Gamma/\omega_0^2}{\omega_0^2 - \omega^2 - \hat{\imath}(\omega^3/\omega_0^2)\Gamma}, \text{ where } \Gamma = (\omega_0/\omega)^2 \Gamma_\omega.$$

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After rotating wave approximation ($|\Delta| = |\omega - \omega_0| \ll \omega_0$):

$$U_{dip} = \frac{3\pi c^2}{2\omega_0^3} \frac{\Gamma}{\Delta} I(\vec{r})$$

$$\Gamma_{sc} = \frac{3\pi c^2}{2\hbar\omega_0^3} \left(\frac{\Gamma}{\Delta}\right)^2 I(\vec{r})$$

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 $\Gamma_{sc} = rac{3\pi c^2}{2\hbar\omega_0^3} \left(rac{\Gamma}{\Delta}
ight)^2 I(\vec{r})$

- The interaction potential U_{dip} is proportional to the light intensity $I(\vec{r})$ and its sign depends on the detuning.
 - $\blacktriangleright \Delta < 0$ red detuned force directed to potential maxima
 - ► Δ > 0 blue detuned force directed to potential minima

•
$$\Gamma_{sc}/U_{dip} \propto 1/\Delta \Longrightarrow$$
 large detuning are better for us

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The most basic optical lattice trap consists of two focused, counter-propagating red-detuned laser beams $(\vec{k} \text{ and } -\vec{k})$

Intensity in the direction of propagation (z) becomes a standing wave:

$$I(z) = \left| \vec{E_1}(z,t) + \vec{E_2}(z,t) \right|^2 = 2 \left| E_0 \cos(\omega t) \cos(kz) \right|^2 \approx E_0^2 \cos^2(kz)$$

and the potential $U(z) \sim \cos^2(kz)$

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One-dimensional optical lattice: $U(r, z) = U_o \left(1 - e^{-2r^2/w_o^2} \cos^2(kz)\right)$ Two-level atoms in the lattice probed by the clock laser (ω_p, k_p) can be described by:

$$\begin{aligned} \hat{H} &= \hbar \omega_{eg} |e\rangle \langle e| + (\hbar \Omega \cos(\omega_p t - k_p \hat{z}) |e\rangle \langle g| + H.C.) + \hat{H}_{ext}, \\ \hat{H}_{ext} &= \frac{\hbar^2 \hat{\kappa}^2}{2m} + U_0 (1 - e^{-2(\hat{\kappa}^2 + \hat{y}^2)/w_0^2} \cos^2(k\hat{z})) \end{aligned}$$

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source: PhD thesis Rodolphe Le Targat

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\Rightarrow Bloch states \Rightarrow tunneling between sites \Rightarrow no longer in L-D regime.



source: PhD thesis Rodolphe Le Targat

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Solution: vertical lattice

$$\hat{H}_{ext} = rac{\hbar^2 \hat{\kappa}^2}{2m} + U_0(1 - e^{-2(\hat{x}^2 + \hat{y}^2)/w_0^2} \cos^2(k\hat{z})) + mg\hat{z}$$



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Resolved sidebands in ⁸⁸Sr optical lattice clock



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Lattice depth

Typical depth of the lattice: few tens of μ K.

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Lattice depth

Typical depth of the lattice: few tens of μ K.

Atoms have to be cooled down to tens of $\mu {\rm K}$ before loading into the lattice \ldots

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The basic idea - absorption and spontaneous emission

$$\Delta \vec{p} = \sum \hbar \vec{k}_{abs} - \sum \hbar \vec{k}_{em} = N \hbar \vec{k}_{abs} - 0$$

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The basic idea - absorption and spontaneous emission

$$\Delta ec{p} = \sum \hbar ec{k}_{abs} - \sum \hbar ec{k}_{em} = N \hbar ec{k}_{abs} - 0$$

Example:

Na atoms, $\lambda = 590$ nm, m = 23, v = 600 m/s at T = 400 K. Absorption of 1 photon $\implies \Delta v = \hbar k/m = 3$ cm/s. \implies it takes $\sim 20~000$ photons to stop. If I = 6 mW/cm², then the atomic beam will stop in 1 ms at 0.5 m. Deceleration: 10^6 m/s².

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A complication: atoms will go off resonance due to the Doppler shift.



source: WD. Phillips, Rev. Mod. Phys., 70, 721, (1998)

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Stability and accuracy Stability Accuracy

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A complication: atoms will go off resonance due to the Doppler shift.

Solutions:

- Tune the frequency of the laser.
- Tune the transition frequency in atoms.

Optical lattice atomic clocks

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Time standards

Optical clocks 1

Lamb-Dicke regime Wannier-Stark regime

Laser cooling and trapping

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Magic wavelenght Ultra-stable laser Clock cycle

Stability and accuracy Stability Accuracy

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A complication: atoms will go off resonance due to the Doppler shift.

"Chirping" the laser frequency:



source: WD. Phillips et al., j. Opt. Soc. Am. B, 2, 1751 (1985)

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A complication: atoms will go off resonance due to the Doppler shift.

Zeeman slower



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Laser cooling — Zeeman slower



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Laser cooling — Optical molasses

Two counter propagating, red-detuned laser beams with frequency $\omega < \omega_0$



Atoms are more in resonance with counter-propagating beam. Atoms loose energy when emitting

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Laser cooling — Optical molasses



For low velocities $F \propto v \Longrightarrow$ cooling.

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Laser cooling — Magneto-Optical trap



Force which depends on position $F \propto x$ \implies Magneto-Optical Trap Optical lattice atomic clocks

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Laser cooling — Magneto-Optical trap



source: TM Brzozowski PhD Thesis Cooling limit — In emissions of photons $\langle v^2 \rangle \neq 0$ \implies Doppler limit: $K_B T = \hbar \Gamma/2$

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Vacuum System

And now back to Optical Lattice Clocks...



source: P. Morzynski PhD Thesis

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Alkaline earth-like



Two stages of magneto-optical trapping needed.

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Alkaline earth-like



Real atoms in vast majority are not two-level atoms...

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Alkaline earth-like



Solution: Repumping lasers.

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Lamb-Dicke regime



Lamb-Dicke regime: $\sqrt{\frac{\omega_{rec}}{\omega_V}}\sqrt{(n+1)} \ll 1$ Potential with depth $10E_{rec}$ or more is needed to neglect motions of atoms.

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Lamb-Dicke regime



Lamb-Dicke regime: $\sqrt{\frac{\omega_{rec}}{\omega_V}}\sqrt{(n+1)} \ll 1$ Potential with depth $10E_{rec}$ or more is needed to neglect motions of atoms. Strong trapping light

- huge light shift (AC Stark shift), at least several tens of kHz,
- polarisation dependent effects



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Magic

 $U = -\frac{1}{4} \alpha(\epsilon, \omega_L) E^2$ Solution: "Magic" wavelength



Source: Katori, Nature Photonics 5, 203, (2011) where polarizabilites, i.e. AC-Stark shifts, are equal for both clock states

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Lasers for Sr optical lattice clock



source: PG. Westergaard PhD Thesis

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Clock laser



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Ultra-stable laser Clock cycle

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Ultra-stable cavity



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Timing scheme of one cycle in ⁸⁸Sr clock



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Bosons vs fermions



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Source: Katori, Nature Photonics 5, 203, (2011)

Timing scheme of one cycle in ⁸⁷Sr clock

Optical Blue MOT Red MOT pumping Interrogation Detection Blue MOT Zeeman Slower Red MOT Quadrupole magnetic field Homogeneois magnetic field Optical lattice Repumpers Blue probe Clock light

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Interrogation and detection



Fluorescence imaging with electron shelving technique probes directly the probability of the transition $P_c = \frac{N_e}{N_e + N_r}$

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Detection



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Remember Rabi oscillations?

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Rabi interrogation



Single clock laser π pulse.

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Ramsey interrogation



source: PG. Westergaard PhD Thesis

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Ramsey interrogation



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source: MN Bishof PhD Thesis

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Lock to the clock line



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Stability

- noises of an oscillator (laser)
- noises of detection
- quantum projection noise

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Dick effect

Dead time in each clock cycle leads to degrading the long term clock stability due to fast fluctuation of clock laser



source: PG Westergaard et al., IEEE Trans. Ultrason., Ferroelect., Freq. Control, 57, 623

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Dick effect

Dead time in each clock cycle leads to degrading the long term clock stability due to fast fluctuation of clock laser



Depends on:

- S(f): noise of the laser
- d: Duty cycle
- g(t): type of interrogation used



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Dick effect

Dead time in each clock cycle leads to degrading the long term clock stability due to fast fluctuation of clock laser



Depends on:

- S(f): noise of the laser
- d: Duty cycle
- g(t): type of interrogation used

Limitation of the fractional Allan variance:

$$\sigma_L^2(\tau) = \frac{1}{\tau} \sum_{m=1}^{\infty} \left| \frac{g_m}{g_0} \right|^2 S(m/T_c)$$

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Dick effect - synchronous vs asynchronous comparison





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Detection noise

How the fluctuations of the clock transition frequency transfer into fluctuation of the locked clock laser

- An example Ramsey interrogation, π/2 impulses τ_d long, separated by T
- Probability of transition at frequency ν is $p(\nu) = \frac{1}{2} (1 + \cos(2\pi T ((\nu \nu_0))))$
- Fluctuations of the probability measurements δp and fluctuations of transition frequency $\delta \nu_0$ are connected by: $\delta p = \pi T \delta \nu_0$
- ► For any measurement $\delta p = \pi \int_{cycle} g(t) \delta \nu_0(t) dt$, where g(t) is an atomic sensivity function
- ► δp is applied as an error signal to the PI lock of the laser. Therefore any noise in probability measurement induce noise of the frequency of the locked laser $\delta P/\pi T_{cycle} < g(t) >$

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Quantum projection noise

- A two-level system prepared as a linear superposition $|\psi\rangle = c_A |A\rangle + c_B |B\rangle$
- We detect if the system is in $|A\rangle$ or $|B\rangle$
- Probability of the system in state $|A\rangle$ is equal to $p_A = |c_A|^2$
- Measurements can be predicted in certainty only when $c_A = 0$ or $c_B = 0$

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Quantum projection noise

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- We detect if the system is in $|A\rangle$ or $|B\rangle$
- Probability of the system in state $|A\rangle$ is equal to $p_A = |c_A|^2$
- Measurements can be predicted in certainty only when $c_A = 0$ or $c_B = 0$
- ▶ N independent atoms, N_A i N_B are in |A⟩ and |B⟩ states, respectively
- ► $P(N_B, N, p_B) = \frac{N!}{N_B!(N!-N_B!)} (p_B)^{N_B} (1-p_B)^{N-N_B}$
- variance of the binomial distribution: $\sigma^2 = Np_B(1 p_B)$

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State-of-the art - above QPN



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source: BJ Bloom et al. Nature 506, 71, (2014)

Accuracy

How well we can control shifts due to environment

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Zeeman effect: 1 and 2 order

linear



quadratic



Optical pumping + interleaved $^{(2015)}$ measurements $m_F = \pm 1/2$

Measurement with different value of total magnetic field and interpolation to zero

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Lamb-Dicke regime Wannier-Stark regime

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Linear AC Stark shift

Lattice light



Probe light



source: M. Takamoto et al Nature 435, 321



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(2005)



Room temperature $\Rightarrow \lambda \sim 10 \mu m$

⁸⁷Sr:
$$\Delta \nu_{|{}^{1}S_{0}>\rightarrow|{}^{3}P_{0}>,E1} = -2.354 \pm 0.032 Hz$$

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How to fight?

Fractional correction from BBR shift at 300 K

- Sr: -5.5×10^{-15}
- ▶ Yb: -2.6 × 10⁻¹⁵
- ► Hg: -1.6 × 10⁻¹⁶

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How to fight?

- Fractional correction from BBR shift at 300 K
 - Sr: -5.5×10^{-15}
 - ▶ Yb: −2.6 × 10⁻¹⁵
 - ► Hg: -1.6 × 10⁻¹⁶

Measure surroundings as good as possible and do the ray-tracing



source: P. Morzynski et al. Sci Rep, 5, 17495 (2015)

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source: TL Nicholson et al Nat. Commun.6, 6896 (2015)

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How to fight?

Fractional correction from BBR shift at 300 K

- Sr: -5.5×10^{-15}
- ▶ Yb: -2.6 × 10⁻¹⁵
- ► Hg: -1.6 × 10⁻¹⁶

Measure surroundings as good as possible and do the ray-tracing

Cryogenic environment



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source: I. Ushijima Nat. Photon. 9, 185, (2015)

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Accuracy of $10^{-18} \leftrightarrow$ light-shift control better than 10^{-8}

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Accuracy of $10^{-18} \leftrightarrow$ light-shift control better than 10^{-8}

• hyperpolarisability $\Delta \nu = -\frac{1}{4h} \Delta \alpha(\omega, \boldsymbol{e}) E^2 - \frac{1}{16h} \Delta \gamma(\omega, \boldsymbol{e}) E^4$

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Accuracy of $10^{-18} \leftrightarrow$ light-shift control better than 10^{-8}

• hyperpolarisability $\Delta \nu = -\frac{1}{4h} \Delta \alpha(\omega, \boldsymbol{e}) E^2 - \frac{1}{16h} \Delta \gamma(\omega, \boldsymbol{e}) E^4$

• three contributions of electrical dipol polarisability $\alpha(\omega, \boldsymbol{e})$: scalar, vector and tensor $\Delta \nu_{\alpha}^{E1} = (\Delta \kappa^s + \Delta \kappa^v \xi m_F \sin \psi + \Delta \kappa^t f(\theta, \psi, \xi) [3m_F^2 - F(F-1)]) U_0$

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Accuracy of $10^{-18} \leftrightarrow$ light-shift control better than 10^{-8}

• hyperpolarisability $\Delta \nu = -\frac{1}{4h} \Delta \alpha(\omega, \boldsymbol{e}) E^2 - \frac{1}{16h} \Delta \gamma(\omega, \boldsymbol{e}) E^4$

• three contributions of electrical dipol polarisability $\alpha(\omega, \boldsymbol{e})$: scalar, vector and tensor $\Delta \nu_{\alpha}^{E1} = (\Delta \kappa^s + \Delta \kappa^v \xi m_F \sin \psi + \Delta \kappa^t f(\theta, \psi, \xi) [3m_F^2 - F(F-1)]) U_0$

• higher multipoles besides E1 $\Delta \nu^{M1/E2} = \zeta (n+1/2) \sqrt{U_0}$

Optical lattice atomic clocks

Michał Zawada KL FAMO, Nicolaus Copernicus University

Time standards

Optical clocks 1

Lamb-Dicke regime Wannier-Stark regime

Laser cooling and trapping

Optical clocks 2

Magic wavelenght Ultra-stable laser Clock cycle

Stability and accuracy Stability Accuracy

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Other effects

Line pulling



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Other effects

- Line pulling
- Collisions

In fermions cold-cold collisions are limited, since *s*-wave scattering is forbidden (possible *p*-wave scattering of ${}^{3}P_{0}$ states)

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aser cooling

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Magic wavelenght Ultra-stable laser Clock cycle

Other effects

- Line pulling
- Collisions
- DC Stark shift

Static charges on dielectrics inside the vaccum



Can be high (up to \sim 40 Hz), but easy to remove. (IEEE TUFFC 59, 411, 2012)

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Magic wavelenght Ultra-stable laser Clock cycle

Gravitational redshift



Clock raised by 33 cm.

Optical lattice atomic clocks

Michał Zawada KL FAMO, Nicolaus Copernicus University

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Stability and accuracy Stability Accuracy

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