

# *Some aspects of Vorticity fields in Relativistic and Quantum Plasmas*

Felipe A. Asenjo<sup>1</sup>

Universidad Adolfo Ibáñez,  
Chile

- ▶ *Part I: Non-relativistic and Special relativistic Plasmas*
- ▶ *Part II: General relativistic Plasmas*
- ▶ *Part III: Quantum and Quantum Relativistic Plasmas*

---

<sup>1</sup>felipe.asenjo@uai.cl

*Part II:*  
*VORTICITY IN GENERAL*  
*RELATIVISTIC PLASMAS*

Felipe A. Asenjo

# Today...

- ▶ We explore the concept of vorticity fields in general relativistic plasmas
- ▶ We study the generation of vorticity
- ▶ We introduce the concept of Generalized helicity in general relativistic plasmas

# Motivation

- ▶ We saw previously that the motion of a charged fluid in space-time generates a magnetic field, it stands to reason that if spacetime were distorted in the region occupied by a charged fluid, a magnetic field would emerge.
- ▶ According to this idea, we can explore if the properties of the plasmas can generate vorticities and magnetic fields in general relativity.
- ▶ The General Relativistic effects can open exciting possibility of spontaneous generation of magnetic fields near gravitating sources.

It was recently demonstrated<sup>2</sup> that the generalized vorticity

$$\Omega = \mathbf{B} + \frac{m}{q} \nabla \times (f \gamma \mathbf{v})$$

The dynamics of  $\Omega$  is given by

$$\frac{\partial \Omega}{\partial t} - \nabla \times (\mathbf{v} \times \Omega) = \chi_B + \chi_R$$

$$\chi_B = -\frac{c}{q\gamma} \nabla T \times \nabla \sigma, \quad \chi_R = \frac{cT}{q\gamma^2} \nabla \gamma \times \nabla \sigma,$$

---

<sup>2</sup>Mahajan and Yoshida, PRL **105**, 095005 (2010); PoP **18**, 055701 (2011).

# *Plasma dynamics in General Relativity*

# general relativity (I): plasma dynamics of unified fields

The dynamics of an ideal plasma is obtained through  $T^{\mu\nu}{}_{;\nu} = qnF^{\mu\nu}U_\nu$  (using the usual symbol ; for covariant derivatives) for the energy-momentum tensor  $T^{\mu\nu} = hU^\mu U^\nu + pg^{\mu\nu}$  where  $F^{\mu\nu}$  is the electromagnetic field tensor and  $U^\mu$  is the normalized four-velocity ( $U^\mu U_\mu = -1$  with  $c = 1$ ),  $n$  is the density,  $h$  is the enthalpy density and  $p$  is the pressure. The charge  $q$  and the mass  $m$  of the fluid particles are invariants.

The plasma fluid fulfill the continuity equation  $(nU^\mu)_{;\mu} = 0$ .

The equation of motion could be written in terms of unified fields<sup>3</sup>

$$q U_\nu M^{\mu\nu} = T\sigma^{,\mu}, \quad (1)$$

where  $M^{\mu\nu} = F^{\mu\nu} + (m/q)S^{\mu\nu}$  in terms of  $S^{\mu\nu} = (fU^\nu)_{;\mu} - (fU^\mu)_{;\nu}$  and  $f = h/mn$ . All kinematic and thermal aspects of the fluid are now given by  $S^{\mu\nu}$ . The function  $\sigma$  is the entropy density of the fluid (where  $T$  is the temperature)

$$\sigma^{,\mu} = \frac{p^{,\mu} - mnf^{,\mu}}{nT}. \quad (2)$$

The antisymmetry of  $M_{\mu\nu}$  guarantees that the fluid is isentropic  $U_\mu \sigma^{,\mu} = 0$ . Inclusion of the Maxwell equations completes the system description

$$F^{\mu\nu}{}_{;\nu} = 4\pi qnU^\mu. \quad (3)$$

---

<sup>3</sup>Mahajan, Phys. Rev. Lett. **90**, 035001 (2003).

## general relativity (II): 3+1 decomposition

The 3 + 1 formalism allows us to obtain a set of equations that is similar to those found in special relativity.

The interval is (the shift vector is zero)

$$ds^2 = -\alpha^2 dt^2 + \gamma_{ij} dx^i dx^j, \quad (i, j = 1, 2, 3) \quad (4)$$

$\alpha$  is the lapse function and  $\gamma_{ij}$  is the 3-metric of the spacelike hypersurfaces of metric  $g_{\mu\nu}$ .

$$\alpha = \sqrt{-g_{00}}$$

it corresponds to the gravitational potential.

The timelike vector field  $n^\mu = (-1/\alpha, 0, 0, 0)$  and  $n_\mu = (\alpha, 0, 0, 0)$  [ $n^\mu n_\mu = -1$  and  $n^\mu \gamma_{\mu\nu} = 0$ ]

$$g_{\mu\nu} = \gamma_{\mu\nu} - n_\mu n_\nu$$

Thus, the 3 + 1 decomposition is achieved by projecting every tensor onto  $n^\mu$  in timelike hypersurfaces and onto  $\gamma_{\mu\nu}$  in spacelike hypersurfaces. For example, the four-velocity  $U^\mu = (\Gamma, \Gamma v^i)$ , such that  $n_\mu U^\mu = \alpha\Gamma$ , the decomposition

$$U^\mu = -\alpha\Gamma n^\mu + \Gamma\gamma^\mu{}_\nu v^\nu, \quad (5)$$

allows us to write the Lorentz factor as

$$\Gamma = \left( \alpha^2 - \gamma_{\mu\nu} v^\mu v^\nu \right)^{-1/2}. \quad (6)$$



# *Vorticity generation and helicity in General Relativity*

# Magnetofluid unification

The generalized electric ( $\xi^\mu$ ) and magnetic ( $\Omega^\mu$ ) fields in terms of  $M^{\mu\nu}$  are

$$\xi^\mu = n_\nu M^{\nu\mu}, \quad \Omega^\mu = \frac{1}{2} n_\rho \epsilon^{\rho\mu\sigma\tau} M_{\sigma\tau}, \quad (7)$$

both spacelike ( $n_\mu \xi^\mu = 0$  and  $n_\mu \Omega^\mu = 0$ ). The magnetofluid tensor reads

$$M^{\mu\nu} = \xi^\mu n^\nu - \xi^\nu n^\mu - \epsilon^{\mu\nu\rho\sigma} \Omega_\rho n_\sigma$$

$$\xi = \mathbf{E} - \frac{m}{\alpha q} \nabla (f \alpha^2 \Gamma) - \frac{m}{\alpha q} \frac{\partial}{\partial t} (f \Gamma \mathbf{v}), \quad (8)$$

$$\Omega = \mathbf{B} + \frac{m}{q} \nabla \times (f \Gamma \mathbf{v}). \quad (9)$$

The generalized magnetic field  $\Omega$  is the generalized vorticity, GV. General relativity enters the definition of GV through  $\Gamma$  and  $\nabla$  (calculated with  $\gamma_{ij}$ ).

The plasma equations are

$$q \alpha \Gamma \mathbf{v} \cdot \xi = -T \frac{\partial \sigma}{\partial t}, \quad (10)$$

while the plasma **momentum evolution equation** is

$$\alpha \Gamma \xi + \Gamma \mathbf{v} \times \Omega = \frac{T}{q} \nabla \sigma. \quad (11)$$

# Generalized Vorticity generation <sup>4</sup>

The antisymmetry of  $M^{\mu\nu}$  implies that its dual must obey  $M^{*\mu\nu}{}_{;\nu} = 0$ . The 3 + 1 decomposition of this equation leads to  $\partial\Omega/\partial t = -\nabla \times (\alpha\xi)$ . Using this

$$\frac{\partial\Omega}{\partial t} - \nabla \times (\mathbf{v} \times \Omega) = \Xi_B + \Xi_R, \quad (12)$$

$\Xi_B$  and  $\Xi_R$  are the sources of the vorticity  $\Omega$ . These drives are nonzero only for inhomogeneous thermodynamics

$$\Xi_B = -\left(\frac{1}{q\Gamma}\right) \nabla T \times \nabla \sigma, \quad \Xi_R = \frac{T\Gamma}{2q} \left[ -\nabla \alpha^2 + \nabla \left( \gamma_{ij} v^i v^j \right) \right] \times \nabla \sigma, \quad (13)$$

$\Xi_B$  is the traditional Biermann battery corrected by curvature.  $\Xi_R$  is the general relativistic drive and it is the principal object of this search.

- ▶ The relativistic drive  $\Xi_R$  is radically transformed from its flat space antecedent. The striking result is that the gravitational potential  $\alpha$ , can produce a magnetic field in any region populated by charged particles even if their local velocities are negligible.
- ▶  $\Xi_B$  and  $\Xi_R$  are non-magnetic thermodynamic source terms that create the conditions for the linear growth of the magnetic fields from zero initial value (batteries).

---

<sup>4</sup>Asenjo, Mahajan & Qadir, PoP20, 022901 (2013);

# Generalized relativistic helicity

Again we define

$$K^\mu = \frac{1}{2\sqrt{-g}} \varepsilon^{\mu\nu\alpha\beta} P_\nu M_{\alpha\beta}$$

where

$$P_\mu = A_\mu + \frac{m}{q} f U_\mu$$

$$\begin{aligned} K^\mu{}_{;\mu} &= \frac{1}{\sqrt{-g}} \partial_\mu (\sqrt{-g} K^\mu) = \frac{1}{2} \varepsilon^{\mu\nu\alpha\beta} P_{\nu;\mu} M_{\alpha\beta} + \frac{1}{2} \varepsilon^{\mu\nu\alpha\beta} P_\nu M_{\alpha\beta;\mu} \\ &= \varepsilon^{\mu\nu\alpha\beta} P_{\nu;\mu} M_{\alpha\beta} \end{aligned}$$

And the Generalized vorticity

$$h \equiv \int \sqrt{-g} K^0 d^3x = \int \varepsilon^{0ijk} P_i M_{jk} d^3x = \int \mathcal{P} \cdot \Omega d^3x$$

$$\begin{aligned} \int \sqrt{-g} K^\mu{}_{;\mu} d^3x &= \int \partial_\mu (\sqrt{-g} K^\mu) d^3x = \int \partial_0 (\sqrt{-g} K^0) d^3x = \partial_t h \\ &= \int \partial_t \mathcal{P} \cdot \Omega d^3x + \int \mathcal{P} \cdot \partial_t \Omega d^3x = \int \frac{2}{q\Gamma} \nabla \sigma \cdot \Omega d^3x \end{aligned}$$

*Around a black hole*

# accreting plasma around a Schwarzschild black hole (I)

The plasma moves in an accretion disk in the equatorial plane ( $\theta = \pi/2$ ). Its orbital velocity is  $v^\phi = r\dot{\phi} = c\sqrt{r_0/2r}$  (being  $r_0$  the Schwarzschild radius), and it is larger than the radial velocity.

At  $5r_0$ , the usual definition for entropy is valid. If the plasma obeys a barotropic equation of state, then  $(T/c)\nabla\sigma \equiv \zeta k_B \nabla T$  where  $\zeta$  is of order unity and the Biermann battery vanishes.

For the model described above, the general relativistic drive becomes

$$\Xi_R = \frac{3\zeta c k_B r_0 \alpha}{4 e r^3} \left(1 - \frac{3r_0}{2r}\right)^{-1/2} \frac{\partial T}{\partial \phi} \hat{e}_z, \quad (14)$$

where the variations of the temperature have been taken in cylindrical geometry, and we have neglected the toroidal temperature gradients compared with the poloidal variations,  $\partial_\theta T \ll \partial_\phi T$ .

All the charged matter of the accretion disk contributes to  $\Xi_R$ , and acts as a source for  $\Omega$ . For the stable orbit at  $r = 5r_0$ , the total relativistic drive is ( $M_\odot$  is the solar mass)

$$\Xi_{R\text{total}} = \int_0^{2\pi} d\phi \Xi_R \approx 3 \times 10^{-2} \zeta \left(\frac{M_\odot}{M}\right)^{9/4} \hat{e}_z, \quad (15)$$

where the disk radiates like a blackbody<sup>5</sup>  $\int \partial_\phi T d\phi \approx 5 \times 10^7 (M_\odot/M)^{1/4} \text{K}$ .

<sup>5</sup>M. Vietri, *Foundations of high-energy astrophysics* (2008).

## accreting plasma around a Schwarzschild black hole (II)

For a short time  $\tau$  when the nonlinear terms involving  $\Omega$  are negligible,  $\Omega$  grows linearly with time  $\Omega_{\text{total}} \approx \Xi_{R\text{total}}\tau$ .

A measure of  $\tau$  is provided by  $|\Omega_{\text{total}}|\tau^{-1} \simeq |\nabla \times (\mathbf{v} \times \Omega_{\text{total}})|$  implying that  $\tau \simeq L/|\mathbf{v}|$ , where  $L$  is the length of variation of the  $|\mathbf{v} \times \Omega|$  force. Taking the length  $L$  on which  $|\mathbf{v}|$  varies to be of the order of  $5r_0/\alpha$ , the time for initial linear phase of GV seed is

$$\tau = \frac{5r_0}{|\mathbf{v}|\alpha} \approx 1.7 \times 10^{-4} \left( \frac{M}{M_{\odot}} \right), \quad (16)$$

in seconds, where the velocity is of order  $v^{\phi}$ . The total strength of the magnetic field generated (in gauss) for the “test” plasma matter accreting at a distance  $5r_0$  is

$$|\Omega_{\text{total}}| \approx 5 \times 10^{-6} \zeta \left( \frac{M_{\odot}}{M} \right)^{5/4}. \quad (17)$$

The initial seed is supposed to be small. It is created in a very short initial time in a state where there was no magnetic field to begin with. The existence of this seed is crucial to the startup of the standard processes of long-time magnetic field generation, like the dynamo process or the magneto-rotational instability. The dynamo process can operate only when it has some initial magnetic field to amplify; we have shown that the General Relativistic drive can provide the needful.

*In cosmology*



# Friedmann–Robertson–Walker

The spatially flat universe

$$ds^2 = -dt^2 + a^2 \gamma_{ij} dx^i dx^j, \quad (i, j = 1, 2, 3) \quad (18)$$

$a = a(t)$  is the time-dependent scale factor of the Universe, and  $\gamma_{ij} = (1, 1, 1)$  is the 3-metric of the spacelike hypersurfaces of the flat spacetime.

The four-velocity  $U^\mu = (\Gamma, \Gamma v^i)$ ,

$$U_\mu = -\Gamma n_\mu + a^2 \Gamma \gamma_{\mu\nu} v^\nu, \quad (19)$$

such that  $n_\mu U^\mu = \Gamma$ , and the Lorentz factor is given by

$$\Gamma = \left(1 - a^2 v^2\right)^{-1/2}, \quad (20)$$

Generalized magnetofluid fields

$$\xi^\mu = n_\nu M^{\nu\mu}, \quad \Omega^\mu = \frac{1}{2} n_\rho \epsilon^{\rho\mu\sigma\tau} M_{\sigma\tau}, \quad (21)$$

$$M^{\mu\nu} = \xi^\mu n^\nu - \xi^\nu n^\mu - \epsilon^{\mu\nu\rho\sigma} \Omega_\rho n_\sigma. \quad (22)$$

$$\xi = \mathbf{E} - \frac{m}{qa^2} \left[ \nabla (f\Gamma) + \frac{\partial}{\partial t} (fa^2 \Gamma \mathbf{v}) \right], \quad (23)$$

$$\Omega = \mathbf{B} + \frac{ma^2}{q} \nabla \times (f\Gamma \mathbf{v}) = \nabla \times \left( \mathbf{A} + \frac{a^2 m f \Gamma}{q} \mathbf{v} \right), \quad (24)$$

$$a^2 \mathbf{v} \cdot \xi = \frac{T}{q\Gamma} \frac{\partial \sigma}{\partial t}, \quad (25)$$

$$\xi + \mathbf{v} \times \Omega = -\frac{T}{q\Gamma} \nabla \sigma. \quad (26)$$

$$\frac{\partial}{\partial t} (a^3 \Omega) + a^3 \nabla \times \xi = 0, \quad (27)$$

$$\frac{1}{a^3} \frac{\partial}{\partial t} (a^3 \Omega) - \nabla \times (\mathbf{v} \times \Omega) = \Xi_B + \Xi_R. \quad (28)$$

The vector fields  $\Xi_B$  and  $\Xi_R$  are the sources of the vorticity  $\Omega$ . The cosmological Biermann battery

$$\Xi_B = \frac{1}{q\Gamma} \nabla T \times \nabla \sigma, \quad (29)$$

and the general relativistic drive for a cosmological background

$$\Xi_R = \frac{T}{q} \nabla \left( \frac{1}{\Gamma} \right) \times \nabla \sigma = -\frac{a^2 T \Gamma}{2q} \nabla v^2 \times \nabla \sigma. \quad (30)$$

# Relativistic cosmological drive<sup>6</sup>

Universe in the radiation-dominated era  $a \propto t^{1/2}$  and  $T \propto a^{-1} \propto t^{-1/2}$ . In this case  $\sigma = \sigma(T)$ , and the Biermann battery drive vanishes.

The plasma velocity goes as  $|\mathbf{v}| \propto a^{-1}$ . If the temperature  $T(t, r, \theta) = T_s(r, \theta)/a$ , and the velocity  $\mathbf{v}(t, r, \theta) = \mathbf{v}_s(r, \theta)/a$ , then

$$|\Xi_R| \propto \frac{\Gamma_s}{q} t^{1/2} |\nabla v_s^2 \times \nabla T_s|, \quad (31)$$

where  $v_s^2 = \mathbf{v}_s \cdot \mathbf{v}_s$ , and the Lorentz factor  $\Gamma_s = (1 - v_s^2)^{-1/2}$

For some short enough time (the initial seed generation phase) we can neglect the nonlinear terms involving  $\Omega$ . The vorticity seed is found to grow as

$$|\Omega| \propto \frac{\Gamma_s}{q} |\nabla v_s^2 \times \nabla T_s| t^{1/2}, \quad (32)$$

being proportional to the scale factor of the Universe expansion ( $\propto a$ ).

The vorticity generated (and therefore the magnetic field) grows as the early-Universe.

---

<sup>6</sup>Asenjo & Mahajan, submitted to PoP

## *More general spacetimes*

- ▶ Rotating black holes (Kerr metric)  
[Qadir, Asenjo & Mahajan, Phys. Scr. **89**, 084002 (2014);  
Bhattacharjee, Das & Mahajan, PRD **91**, 123005 (2015)]
- ▶ Non-minimal gravity coupling  
[Bhattacharjee, Das & Mahajan, PRD **91**, 064055 (2015)]

That's all (for now).

Thanks!