

A flux-scaling scenario for moduli stabilization and dS vacua

Anamaría Font V.

Universidad Central de Venezuela



in collaboration with:

R. Blumenhagen, D. Herschmann, M. Fuchs, E. Plauschinn, Y. Sekiguchi, F. Wolf. Nucl.Phys. B897 (2015) 500

R. Blumenhagen, C. Damian, D. Herschmann, R. Sun. arXiv 1510.01522

The First Millenium

The origins, 1956-1979



The origins, 1956-1979

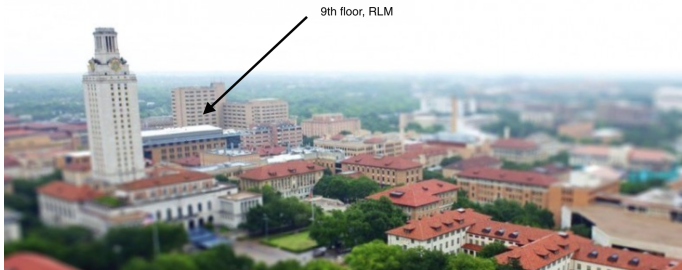
TE
UVG Quevedo Rodríguez, Fernando de la Trinidad
FIS Grupos de simetría en física de partículas. -
.Q6 Guatemala : U.V.G, 1979.
1979 173 p. - (UVG-Tesis).
 Tesis (licenciatura en física)--Universidad del
 Valle de Guatemala, Facultad de Ciencias y
 Humanidades.
 1. Partículas (Física nuclear) 2. Simetría
 (Física)



The origins, 1956-1979



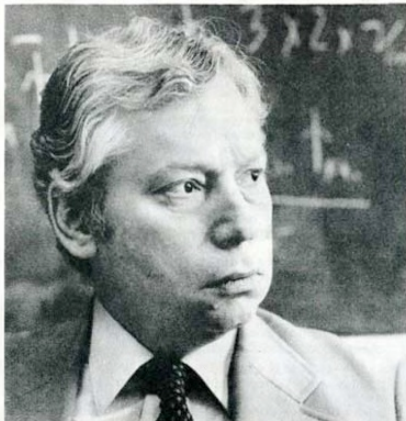
The Austin years, 1981-1986



9th floor, RLM

THE PROPAGATOR

A newsletter of the Department of Physics
The University of Texas at Austin
Austin, Texas 78712
Winter, 1982



NOBEL LAUREATE

Steven Weinberg now permanent faculty member

Nobel laureate Steven Weinberg is now a permanent member of the UT Physics faculty. Dr. Weinberg has been designated a Regener into 15 languages. His best known scholarly publication is *Gravitation and Cosmology: Principles and Applications of the*

Many good courses

Lectures on Quantum Field Theory

Physics 389M

Fall 1982

University of Texas at Austin

© Steven Weinberg 1982

Many BBQs and happy hours ...



Many BBQs and happy hours . . .



photo courtesy Elisa Quevedo

Summer schools



TASI, Ann Arbor, 1984

Summer schools



TASI, Ann Arbor, 1984

Summer schools



TASI, Yale, 1985

Summer schools



TASI, Yale, 1985

Fernando's first string phenomenology paper

Nuclear Physics B272 (1986) 661–676
North-Holland, Amsterdam

LOW-ENERGY EFFECTIVE ACTION FOR THE SUPERSTRING

C.P. BURGESS¹

Institute for Advanced Study, Princeton, NJ 08540, USA

A. FONT²

Center for Particle Theory, University of Texas at Austin, Austin, TX 78712, USA

F. QUEVEDO³

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Received 2 December 1985

We construct the low-energy $D = 4$, $N = 1$ supergravity that arises in superstring theories for an arbitrary number of generations. The couplings of all massless modes that carry low-energy gauge quantum numbers are calculated by truncating the heavy Kaluza-Klein modes of the ten-dimensional effective field theory. The resulting action is compared to the most general effective action compatible with the symmetries of the underlying ten-dimensional field (and string) theories. This comparison indicates which features of the truncation correctly approximate the exact low-energy action.

Topics In Supergravity And Superstring Phenomenology

Fernando Quevedo Rodríguez (Texas U.)

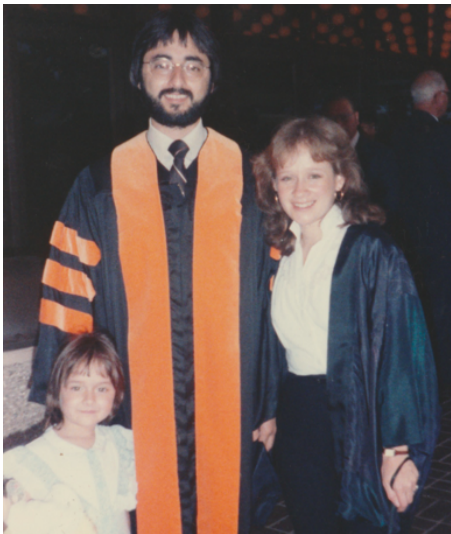
May 1986 - 84 pages

UMI-86-18568

Note: Ph.D. Thesis

Keyword(s): INSPIRE: [THESIS](#)

Graduation, 1986



Austin, May 1986

$Z_N \times Z_M$ ORBIFOLDS AND DISCRETE TORSION

A. FONT

LAPP, B.P. 909, F-74019 Annecy-le-Vieux, France

L.E. IBÁÑEZ

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and

F. QUEVEDO

Department of Physics, McGill University, Montreal, PQH3A 2T8, Canada

Received 1 November 1988

We extend previous work on Z_N -orbifolds to the general $Z_N \times Z_M$ abelian case for both $(2, 2)$ and $(0, 2)$ models. We classify the corresponding $(2, 2)$ compactifications and show that a number of models obtained by tensoring minimal $N=2$ superconformal theories can be constructed as $Z_N \times Z_M$ -orbifolds. Furthermore, $Z_N \times Z_M$ -orbifolds allow the addition of discrete torsion which leads to new $(2, 2)$ compactifications not considered previously. Some of the latter have negative Euler characteristic and Betti numbers equal to those of some complete intersection Calabi-Yau (CICY) manifolds. This suggests the existence of a previously overlooked connection between CICY models and orbifolds.

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Table 1
 $Z_N \times Z_M$ -orbifolds. Here $\alpha = \pm 2\pi/3$ and $\beta = \pm 2\pi/6$. $N=2$ superconformal models with similar massless spectrum are shown in the last column.

Model #	Orbifold	a	b	c	$\chi/2$	$N_{\overline{37}}$	N_{27}	$N=2$ cand.
0	$Z_6 \equiv Z_2 \times Z_3$	$(\frac{1}{3}, 0, -\frac{1}{3})$	$(0, \frac{1}{3}, -\frac{1}{3})$	1	24	11	35	$1^4 2^2 4$ $1^3 2 4 (-10)$
1	$Z_3 \times Z_2$	$(\frac{1}{3}, 0, -\frac{1}{3})$	$(0, \frac{1}{3}, -\frac{1}{3})$	1	48	3	51	$1^2 2^2 4^2$ $1 2 4^3 (-10)$
2	mirror pair			-1	-48	51	3	$4^2 (-10)^2$

Strong-weak coupling duality and non-perturbative effects in string theory

A. Font ^a, L.E. Ibáñez ^b, D. Lüst ^b and F. Quevedo ^c

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Received 13 July 1990

We conjecture the existence of a new discrete symmetry of the modular type relating weak and strong coupling in string theory. The existence of this symmetry would strongly constrain the non-perturbative behaviour in string partition functions and introduces the notion of a maximal (minimal) coupling constant. An effective lagrangian analysis suggests that the dilaton vacuum expectation value is dynamically fixed to be of order one. In supersymmetric heterotic strings, supersymmetry (as well as this modular symmetry itself) is generically spontaneously broken.

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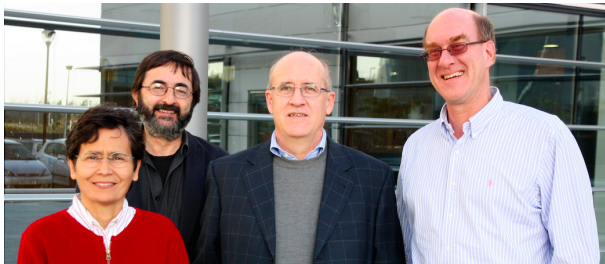
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S-duality Quartet, Madrid, December 2011

On the family front, three lovely daughters



Geneva, September 1990

ICTP Prize Winners 1998



Anamaria Font, Miguel Virasoro and Fernando Quevedo

Anamaria Font Fernando Quevedo

The ICTP Prize in the field of High Energy Physics (in honour of Professor Chen Ning Yang) has been awarded jointly to Anamaria Font from Universidad Central de Venezuela, Caracas, Venezuela, and Fernando Quevedo from Universidad Nacional Autonoma de Mexico (UNAM), Mexico D.F.,

Mexico,

for their contribution to the phenomenological studies in superstring theory based on orbifold compactifications and many works on Calabi-Yau compactifications, mirror symmetry and duality symmetries. These works have contributed to a greater understanding of the low energy string physics, as well as various stringy symmetries. In particular, the important concept of S-duality has been introduced by them and their collaborators.

The Second Millenium

Bottom-up, LVS, string inflation, dS, ...



In Cambridge, Elisa Quevedo

Bottom-up, LVS, string inflation, dS, ...



In Trieste

A comment on continuous spin representations of the Poincaré group and perturbative string theory

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Received 17 September 2014, accepted 17 September 2014

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Key words Poincaré group, perturbative string theory.

We make a simple observation that the massless continuous spin representations of the Poincaré group are not present in perturbative string theory constructions. This represents one of the very few model-independent low-energy consequences of these models.

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No continuous spin representations in perturbative string theory

Generic prediction !

A flux-scaling scenario for moduli stabilization and dS vacua

based on Nucl.Phys. B897 (2015) 500, arXiv 1510.01522

in with R. Blumenhagen, C. Damian, D. Herschmann, M. Fuchs, E. Plauschinn, R. Sun, Y. Sekiguchi, F. Wolf

Outline

- Introduction
- Flux-scaling scenario
 - Generalized superpotential
 - Criteria for moduli stabilization
 - Example
 - Results
- Minkowski and de Sitter vacua
- Summary and final comments

This talk Mostly

- **Getting de Sitter**
- **Getting inflation**

Crucial issue: Moduli stabilisation

(concentrate on IIB strings: KKLT, LVS)

This talk Mostly

Crucial issue: Moduli stabilisation

(concentrate on IIB strings: ~~KKLT, LVS~~)

with geometric and non-geometric fluxes

- **Getting de Sitter**
- **Getting inflation** (only comments)

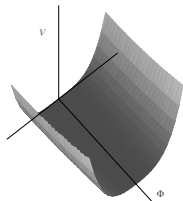
Introduction

fluxes: non-trivial backgrounds for field strengths

e.g. NS-NS: $\int_{\Pi_3} \langle H \rangle = h \neq 0$; R-R: $\int_{\Pi_3} \langle \mathfrak{F} \rangle = \mathfrak{f} \neq 0$ Π_3 : internal 3-cycle

can be used to

- ▷ fix moduli, i.e. $\langle \Phi \rangle$, in the 4d effective theory



Φ : massless scalar with flat potential

\exists in generic standard compactifications

- ▷ trigger supersymmetry breaking
- ▷ induce axion monodromy inflation
- ▷ ...

moduli Φ

typical: axiodilaton S , Kähler (size) T , complex structure (shape) U

in $\mathcal{N}=1$, $S, T, U \in$ chiral multiplets, $\text{Re } \Phi$: saxion, $\text{Im } \Phi$: axion

it is necessary to fix $\langle \Phi \rangle$

to avoid fifth forces

gauge couplings determined by vevs, $\frac{1}{g_{\text{YM}}^2} \sim \text{Re} \langle aS + bT + cU \rangle$

first attempt to fix S (heterotic): $W = h + ce^{-\gamma S}$ Dine, Rohm, Seiberg, Witten

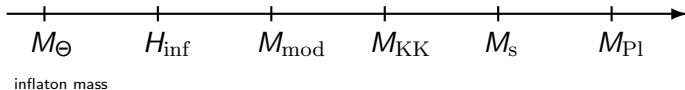
more modern (IIB): $W = W_{\text{flux}}(S, U) + W_{\text{np}}(T) = W_{\text{flux}} + ce^{-\gamma T}$

KKLT, Kachru, Kallosh, Linde, Trivedi ; Large Volume Scenario, Balasubramanian, Berglund, Conlon, Quevedo

in KKLT, LVS $M_T \ll M_{S,U}$

Idea

combine moduli stabilization and axion monodromy inflation
in string scenario with



to obtain $M_{\text{mod}} > M_\Theta$, fix all S, U, T , moduli at tree level

it can be done using non-geometric fluxes !

Flux-scaling scenario

flux induced moduli potentials in 4d

$$\mathcal{S} = \frac{1}{\ell_s^8} \int d^{10}x \sqrt{-G} \left\{ e^{-2\varphi} [\mathcal{R} - H^2] - \sum_n \mathfrak{F}_n^2 + \dots \right\}$$

$$\int_{\Pi_3} \langle H \rangle = h \quad \Rightarrow \quad V = \frac{h^2 e^{2\varphi}}{R^{12}} \quad \Rightarrow \quad W = \begin{cases} h & \text{heterotic} \\ hS & \text{orientifolds} \end{cases}$$

superpotentials in $\mathcal{N}=1$ type II orientifolds with NS-NS and R-R fluxes

$$\text{IIB} \quad \begin{matrix} \int_{\Pi_3} \langle H \rangle \\ \int_{\Pi_3} \langle \mathfrak{F} \rangle \end{matrix} \Rightarrow W(S, U) \quad ; \quad \text{IIA} \quad \begin{matrix} \int_{\Pi_3} \langle H \rangle \\ \int_{\Pi_{2m}} \langle \mathfrak{F}_{2m} \rangle \end{matrix} \Rightarrow W(S, U, T)$$

to recover T -duality introduce non-geometric fluxes Shelton, Taylor, Wecht

$$H \xleftrightarrow{T} F \xleftrightarrow{T} Q \xleftrightarrow{T} R$$

* $\mathcal{N}=1$ V_F from W and Kähler potential K : $V_F = e^K \left\{ K^{I\bar{J}} D_I W D_{\bar{J}} \bar{W} - 3 |W|^2 \right\}$

Generalized superpotential I

Aldazabal, Cámara, A.F., Ibáñez;
Graña, Louis, Waldram; Benmachiche, Grimm

IIB orientifolds on Calabi-Yau 3-fold \mathcal{M} characterized by $(J, \Omega, h_{\pm}^{2,1}, h_{\pm}^{1,1})$

W generalizes $\int_{\mathcal{M}} (\mathfrak{F} - iSH) \wedge \Omega$

Gukov, Vafa, Witten

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$$W = \int_{\mathcal{M}} (\mathfrak{F} + \mathcal{D}\Phi_c) \wedge \Omega, \quad \Phi_c = iS - iT_{\alpha}\sigma^{\alpha} - iG^a\omega_a; \quad \alpha = 1, \dots, h_+^{1,1} \\ a = 1, \dots, h_-^{1,1}$$

$$\Omega = X^{\lambda}\alpha_{\lambda} - \mathcal{F}_{\lambda}\beta^{\lambda}, \quad \mathcal{F}_{\lambda} = \partial_{\lambda}\mathcal{F}, \quad \mathcal{F} : \text{prepotential}; \quad \lambda = 0, \dots, h_-^{2,1}$$

$$U^i = -i\frac{X^i}{X^0}; \quad i = 1, \dots, h_-^{2,1}$$

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$$\mathcal{D} = d - H \wedge -F \circ -Q \bullet -R_{\perp} \quad \left\{ \begin{array}{lll} F \circ : & p\text{-form} & \rightarrow (p+1)\text{-form} \\ Q \bullet : & p\text{-form} & \rightarrow (p-1)\text{-form} \\ R_{\perp} : & p\text{-form} & \rightarrow (p-3)\text{-form} \end{array} \right.$$

e.g. on 4-form: $\mathcal{D}\sigma^{\alpha} = -\tilde{q}^{\lambda\alpha}\alpha_{\lambda} + q_{\lambda}^{\alpha}\beta^{\lambda} \quad \tilde{q}, q : \text{non-geometric fluxes}$

$$\mathcal{D}^2 = 0 \Rightarrow \text{Bianchi identities, e.g.} \quad \tilde{q}^{\lambda\alpha}q_{\lambda}^{\beta} - q_{\lambda}^{\alpha}\tilde{q}^{\lambda\beta} = 0$$

Generalized superpotential II

$$\begin{aligned}
 W &= \int_{\mathcal{M}} \left[\mathfrak{F} - iSH + iT_{\alpha} (Q \bullet \sigma^{\alpha}) + iG^a (F \circ \omega_a) \right] \wedge \Omega = W(S, U, T, G) \\
 &= -(\mathfrak{f}_{\lambda} X^{\lambda} - \tilde{\mathfrak{f}}^{\lambda} \mathcal{F}_{\lambda}) + iS (h_{\lambda} X^{\lambda} - \tilde{h}^{\lambda} \mathcal{F}_{\lambda}) \\
 &\quad + iT_{\alpha} (q_{\lambda}^{\alpha} X^{\lambda} - \tilde{q}^{\lambda \alpha} \mathcal{F}_{\lambda}) + iG^a (f_{\lambda a} X^{\lambda} - \tilde{f}_a^{\lambda} \mathcal{F}_{\lambda})
 \end{aligned}$$

Kähler potential

Grimm, Louis

$$\begin{aligned}
 K &= -\log \left(-i \int_{\mathcal{M}} \Omega \wedge \overline{\Omega} \right) - \log(S + \overline{S}) - 2 \log \mathcal{V} \\
 \mathcal{V} &= \frac{1}{6} e^{-3\phi/2} \int_{\mathcal{M}} J \wedge J \wedge J
 \end{aligned}$$

Flux induced R-R tadpoles

$$N_{D3}^{\text{flux}} = \int_{\mathcal{M}} H \wedge \mathfrak{F} \quad ; \quad [N_{D7}^{\text{flux}}]^{\alpha} = - \int_{\mathcal{M}} (Q \bullet \sigma^{\alpha}) \wedge \mathfrak{F} \quad ; \quad \text{also } [N_{D5}^{\text{flux}}]_a$$

Criteria for moduli stabilization

- ▷ supersymmetric minima with unconstrained axions have tachyons Conlon
thus, search for non-supersymmetric minima with saxions stabilized
in perturbative regime (weak string coupling, large radius)
- ▷ only axions can remain massless
- ▷ values of $\langle \Phi \rangle$ and M_{mod} are parametrically controlled by adjusting
fluxes
- ▷ mass of lightest massive axion (inflaton candidate) is parametrically
or numerically controlled
- ▷ require $M_{\text{mod}} < M_{\text{KK}}, M_s$

STU Example, $h_{-}^{2,1} = 1, h_{+}^{1,1} = 1$

$$K = -\log(S + \overline{S}) - 3\log(T + \overline{T}) - 3\log(U + \overline{U})$$

$$W = -f - 3\tilde{f}U^2 - hUS - qUT$$

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stable AdS non-supersymmetric minimum with flux scaling

axions: $\text{Im } U = 0$, $h\text{Im } S + q\text{Im } T = 0$

saxions: $\text{Re } T = -15 \frac{\tilde{f}}{q} v$, $\text{Re } S = -12 \frac{\tilde{f}}{h} v$, $(\text{Re } U)^2 = v^2 = \frac{1}{3\sqrt{10}} \frac{f}{\tilde{f}}$

$q < 0$, $h < 0$, $f > 0$, $\tilde{f} > 0$, $f \gg \tilde{f} \Rightarrow \text{Re } T, \text{Re } S > 1$

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$q < 0$, $h < 0$, $f > 0$, $\tilde{f} > 0$, $f \gg \tilde{f} \Rightarrow \text{Re } T, \text{Re } S > 1$

mass scales

$$M_{\text{mod}}^2 = \mu \frac{M_{\text{Pl}}^2}{4\pi \cdot 2^7} \frac{h|q|^3}{f^{\frac{3}{2}} \tilde{f}^{\frac{1}{2}}}, \quad \mu \approx (2.1, 0.37, 0.25; 1.3, 0.013, 0), \quad M_{3/2} \sim M_{\text{mod}}$$

$$\frac{M_s^2}{M_{\text{KK}}^2} = 62.5 \left(\frac{h}{q}\right)^{\frac{1}{2}}, \quad \frac{M_{\text{mod}}^2}{M_{\text{KK}}^2} \sim h q \left(\frac{\tilde{f}}{f}\right)^{\frac{1}{2}}, \quad M_{\text{Pl}} \gtrsim_{\tilde{\rho}} M_s \gtrsim_{\tilde{\rho}} M_{\text{KK}} \gtrsim_{\tilde{\rho}} M_{\text{mod}}$$

Some general results

- ▷ analyzed several models with non-supersymmetric flux-scaling extrema, also including more T 's and odd Kähler G 's
- ▷ not always possible to have $M_{\text{mod}} < M_{\text{KK}}$

Some general results

- ▷ analyzed several models with non-supersymmetric flux-scaling extrema, also including more T 's and odd Kähler G 's
- ▷ not always possible to have $M_{\text{mod}} < M_{\text{KK}}$
- ▷ when $h^{1,1} > 1$ new tachyons appear but can be lifted by a D-term potential due to magnetized D7-branes wrapping Σ_4

$$V_D = \frac{M_{\text{Pl}}^4}{2\text{Re}(f)} \xi^2, \quad \xi = \frac{1}{\mathcal{V}} \int_{\Sigma_4} J \wedge c_1(L) \quad \text{FI depends on } T_\alpha$$

Freed-Witten condition $\Rightarrow \xi = 0$ at the AdS supersymmetric minimum

$\xi = 0$ also at AdS non-supersymmetric extremum with same ratios of vevs

extremum is not shifted but negative mass of Kähler tachyon is uplifted

Minkowski and de Sitter vacua I

via Q -fluxes only

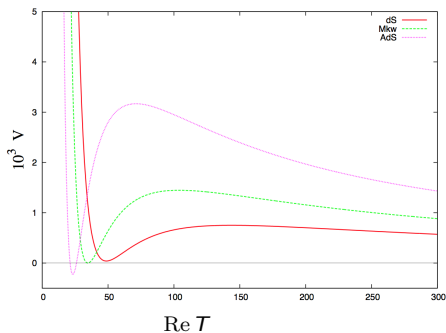
de Carlos, Guarino, Moreno; Blabäck, Danielsson, Dibitetto

in STU models need at least 2 non-geometric fluxes

$$W = -qT(U - U^3) + iS(\epsilon_2 + 3i\epsilon_1 U + 3\epsilon_2 U^2 + i\epsilon_1 U^3) - \xi_3(\epsilon_1 - 3i\epsilon_2 U + 3\epsilon_1 U^2 - i\epsilon_2 U^3) - q\xi_7(1 - U^2)$$

$\mathfrak{so}(3, 1)^{\oplus 2}$

de Carlos, Guarino, Moreno



$$q = -3, \epsilon_1 = 1, \xi_3 = 1, \xi_7 = 16$$

$$\epsilon_2 = 45 \quad \text{AdS}$$

$$\epsilon_2 = 44 \quad \text{dS}$$

no flux-scaling

lightest mode is saxion

Minkowski and de Sitter vacua II

via anti D3-branes

KKLT, KKLM MT

$$V_{\text{up}} = \frac{A}{\mathcal{V}^{\frac{4}{3}}}, \quad \text{for } \overline{\text{D3}} \text{ on warped throat}$$

STU example: $W = -ifU + ih_0S - 3ihSU^2 - iqT$

without V_{up} , non-tachyonic susy AdS with flux scaling

$$\text{Re } T = -\frac{5^{\frac{1}{2}}f}{2q} \left(\frac{h_0}{h} \right)^{\frac{1}{2}}, \dots$$

with V_{up} , **new** stable non-susy Minkowski with flux scaling

$$\text{Re } T = \frac{f}{3^{\frac{1}{4}}q} \left(\frac{h_0}{h} \right)^{\frac{1}{2}}, \dots \quad ; \quad A = A_{\text{Mink}} = \frac{3^{\frac{1}{4}}}{2}qh \left(\frac{h}{h_0} \right)^{\frac{1}{2}}$$

taking $A > A_{\text{Mink}}$ leads to dS, lightest mode is saxionic

Minkowski and de Sitter vacua III

via D-terms from $U(1)_{\hat{\lambda}}$ multiplets, $\hat{\lambda} = 1, \dots, h_+^{2,1}$

Robbins, Wrase

$$V_D = \left[(\text{Re } f)^{-1} \right]^{\hat{\lambda} \hat{\sigma}} D_{\hat{\lambda}} D_{\hat{\sigma}}, \quad D_{\hat{\lambda}} = i \partial_I K \delta_{\hat{\lambda}} \Phi^I$$

flux dependence from $\delta \Phi^I = \xi^{\hat{\lambda}} \delta_{\hat{\lambda}} \Phi^I$, e.g. for $h_-^{1,1} = 0$

$$D_{\hat{\lambda}} = \frac{1}{\mathcal{V}} \left(f_{\hat{\lambda} \alpha} t^\alpha - r_{\hat{\lambda}} e^{\varphi} \mathcal{V} \right), \quad \mathcal{V} = \frac{1}{6} \kappa_{\alpha \beta \gamma} t^\alpha t^\beta t^\gamma, \quad T_\alpha = \frac{1}{2} \kappa_{\alpha \beta \gamma} t^\beta t^\gamma + \dots$$

↗
↖
geometric F-flux
R-flux

$D_{\hat{\lambda}}$ depends on S and T , gauge kinetic function f depends on U

Minkowski and de Sitter vacua III

via D-terms from $U(1)_{\hat{\lambda}}$ multiplets, $\hat{\lambda} = 1, \dots, h_+^{2,1}$

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$D_{\hat{\lambda}}$ depends on S and T , gauge kinetic function f depends on U

toy ex. $h_+^{2,1} = 1$: $f = cU$, $W = i\tilde{f}U + i\tilde{f}U^3 - ihS + iqT$, $V_D = \frac{f_{11}^2}{c U_R T_R^2} \left(3 + \frac{q}{h} \frac{T_R}{S_R} \right)^2$

$V = V_F + V_D$ admits Mink. vacuum with massless axion \rightarrow inflaton candidate Θ

Θ becomes massive at 2nd stage, $W' = \lambda W + \Delta W$ $\lambda \gg 1$

$M_{KK} > M_{\text{mod}} > M_\Theta$ requires fractional fluxes

↓
can arise from polynomial corrections to prepotential

Summary and Final Comments

- Using non-geometric fluxes, constructed non-susy non-tachyonic models with all moduli, except some axions, stabilized in AdS. Moduli vevs and masses can be controlled by flux scaling but there is tension with $M_{\text{mod}} < M_{\text{KK}}$.
- Minkowski and dS flux-scaling vacua can be achieved via $\overline{\text{D3}}$ or D-term uplift.
- Natural set-up for F-term axion monodromy inflation.
- 10d origin of $\mathcal{N} = 1$, 4d scalar potential with non-geometric fluxes
Blumenhagen, A.F., Plauschinn
 - ★ Dimensional reduction of double field theory on a Calabi-Yau 3-fold with small fluxes gives scalar potential of $\mathcal{N} = 2$, 4d gauged supergravity of D'Auria, Ferrara, Trigiante.
 - ★ Upon orientifold projection, $\mathcal{N} = 2 \rightarrow \mathcal{N} = 1$, $V = V_{\text{F}} + V_{\text{D}} + V_{\text{tad}}$



Happy 60th Birthday !

Best wishes for the future.





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Best wishes for the future.

Especially, . . .



wish you success with the crusades



wish you success with the crusades



String Phenomenologists:

Strategic (long term) Plan:

**String theory scenario that satisfies
all particle physics and
cosmological observations and
hopefully lead to measurable
predictions**

F. Quevedo, PASCOS 2011

wish you success with the crusades



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In line with the dramatic geo-economic changes taking place, we are now entering a new phase in which ICTP should take a leading role in promoting better working conditions for scientists in the developing world. We aspire to assist science policy makers and scientists of these countries in the creation of local centres of excellence and active scientific networks. In order to confront these new challenges, ICTP will collaborate with scientists in emerging countries, assisting in this way the poorest regions in Africa, Latin America and Asia whilst, simultaneously, strengthening and broadening the research activities of our Centre.

Greetings from ICTP Director Fernando Quevedo



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Thanks !