

# Small Scale experiments for fundamental physics

ICTP Summer School on Particle  
Physics, June 12-15

# Fundamental Physics

## Energy Frontier

### Large Hardon Collider, CERN

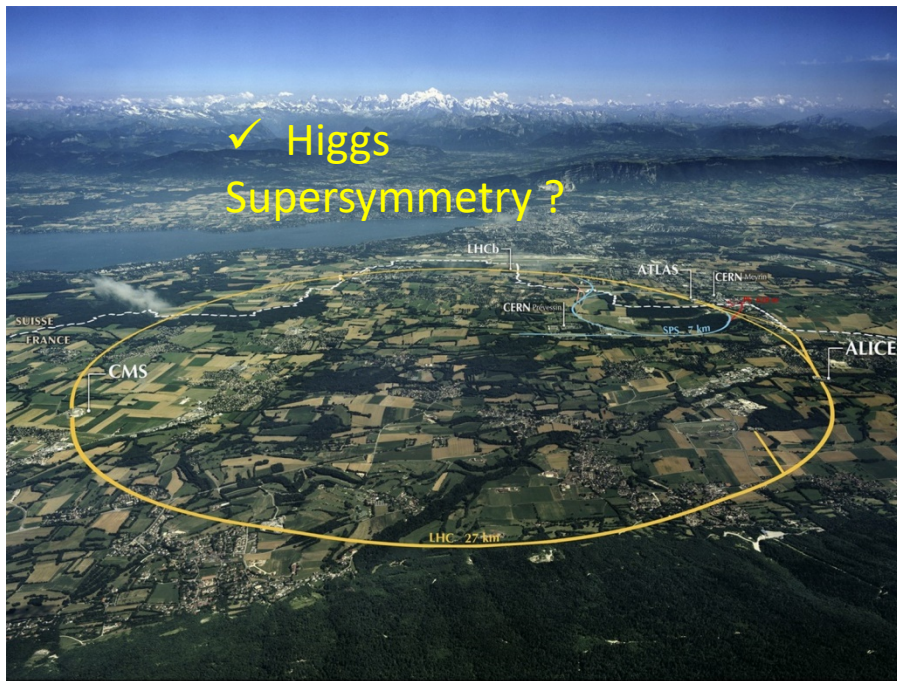
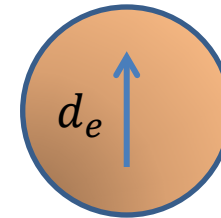


Photo by CERN

## Precision Frontier

- EDM experiments: search for supersymmetry



n, e, nuclei

- Dark Matter: Axion Dark Matter Experiment (ADMX) uses  $\mu$ -wave cavity



# EDM experiments

## Searching for extra CP Violation

- Neutrons ~200
  - @ILL
  - @ILL, @PNPI
  - @PSI
  - @FRM-2
  - @RCNP, @TRIUMF
  - @SNS
  - @J-PARC
- Ions-Muons ~200
  - @BNL
  - @FZJ
  - @FNAL
  - @JPARC

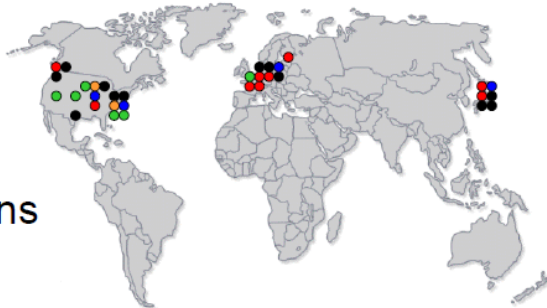
- Molecules ~50
  - YbF@Imperial
  - PbO@Yale
  - ThO@Harvard
  - HfF+@JILA
  - WC@UMich
  - PbF@Oklahoma
- Solids ~10
  - GGG@Indiana
  - ferroelectrics@Yale

Rough estimate of numbers of researchers, in total  
 ~500 (with some overlap)

- Atoms ~100
  - Hg@UWash
  - Xe@Princeton
  - Xe@TokyoTech
  - Xe@TUM
  - Xe@Mainz
  - Cs@Penn
  - Cs@Texas
  - Fr@RCNP/CYRIC
  - Rn@TRIUMF
  - Ra@ANL
  - Ra@KVI
  - Yb@Kyoto

- $|de| < 8.7 \times 10^{-29}$  e cm (ThO Molecule, ACME)
- (199Hg)  $< 7.4 \times 10^{-30}$  e cm, Seattle  
 $\rightarrow < 1.2 \times 10^{-26}$  e cm for nedm (95%)
- $n_{\text{edm}}$  (direct, trapped Ultracold Neutrons)  $< 3.6 \times 10^{-26}$  e cm (95%)

EDM measurements probing physics beyond energy scale of LHC



Adapted from "EDM Searches", K. Kurch, PSI

ACME Collaboration: Science 1248213 (2013)  
 Mercury experiment: Phys. Rev. Lett. 116, 161601 (2016)  
 Neutron EDM: Phys. Rev. D 92, 092003 (2015)

# Tabletop fundamental physics

The  
Economist

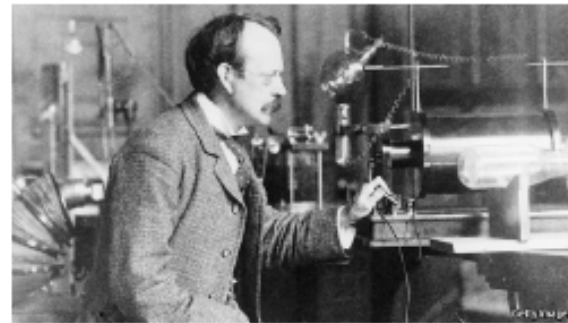
**Fundamental physics**

## Searching for particles on a benchtop

**Making precise measurements of tiny forces**

Jan 28th 2017

THE beams of protons that circulate around the 27km-circumference ring of the Large Hadron Collider (LHC), the world's biggest particle accelerator, carry as much kinetic energy as an American aircraft-carrier sailing at just under six knots. Andrew Geraci's equipment, on the other hand, comprises a glass bead 300 billionths of a metre across, held in a lattice of laser light inside an airless chamber. The power it consumes would run a few old-fashioned light bulbs. Like researchers at the LHC, Dr Geraci and his team at the University of Nevada, in Reno, hope to find things unexplained by established theories such as the Standard Model of particle physics and Newton's law of gravity. Whereas the LHC cost around SFr4.6bn (\$5bn) to build, however, Dr Geraci's set-up cost a mere \$300,000 and fits on a table about a metre wide and three long.



# Goal of these lectures

- Focus on small-scale (i.e. non-accelerator, non-large scale DM detectors) experiments
- Give particle physics students sense of experimental techniques available, newly developing techniques, what they may be useful for
- Generate new ideas for experiments in the future?!

# Syllabus

- Introduction
- New (scalar) forces
- Gravitational Waves and Ultralight Dark Matter
- New (spin-dependent) forces  
(relation to axions, EDMS, Cosmic DM experiments)

# Outline for Lectures

- Lecture 1 –

## New (scalar) Forces

Background/Motivation

Gravitational experiments

Principles of Force sensing

New Techniques:

→ Optical levitation

→ Atomic-based sensors

→ Matter-wave interferometry

# Outline for Lectures

- Lecture 2 –
  - Gravitational waves
    - New Techniques:
      - Levitated sensors
      - Atom interferometry
    - Ultralight scalar field dark matter



# Outline for Lectures

- Lecture 3 –
  - New (spin-dependent) forces
    - Background
    - Torsion balance tests
    - Magnetometry
    - New techniques: (ARIADNE)
    - Relation to axions, EDM experiments,
    - Cosmic DM experiments

# Syllabus for Lectures

- Lecture 1 –

New (scalar) Forces

Background/Motivation

Gravitational experiments

Principles of Force sensing

New Techniques:

→ Optical levitation

→ Matter-wave interferometry

# The Standard Model

Provides an adequate description of the electromagnetic, weak, and strong interactions.

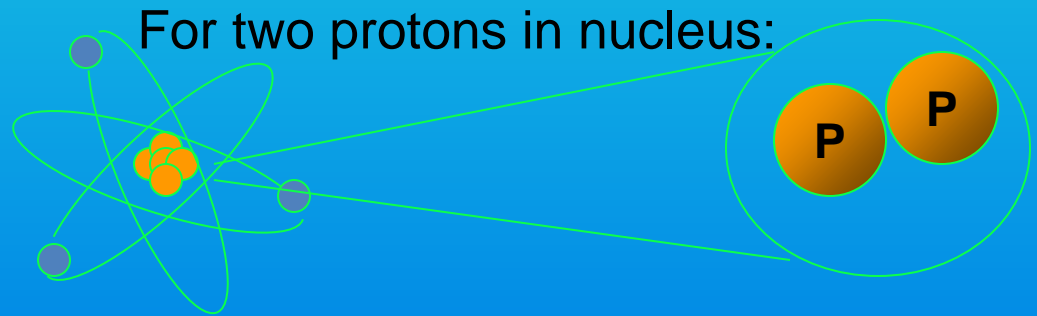
## The Interactions:

**Strong:** Holds nucleons together

**Electromagnetic:** Acts between charged particles

**Weak:** Causes certain decays

**Gravity:** Attraction between masses



Strong : Electromagnetic : Weak : Gravity = 20 : 1 :  $10^{-7}$  :  $10^{-36}$

The Hierarchy Problem: Why is Gravity so small?

# Solving the Hierarchy Problem

Quantum Gravity  $\sim 10^{19}$  GeV



Electro-weak  $\sim 10^3$  GeV

Supersymmetry?

Searching for it now at LHC!

Large Extra Dimensions (sub-mm)?

Gravity's mass (i.e. Planck) scale of  $\sim 10^{19}$  GeV

Is not a fundamental scale. Its magnitude comes from the size of the extra dimensions.

These effects may cause gravity to change below a characteristic scale  $\lambda < 1$  mm :

- Change its power law from  $1/r$
- Acquire a new exponential form

Can we measure this?

# 1) Light Moduli

Moduli: Massless scalar particles generic in string theory

- describe the shape and size of (small OR large) extra dimensions

Moduli must become massive to avoid conflict with experiment

Mass may come from Supersymmetry breaking

Gravity-Mediated  
(high scale  $10^{11}$  GeV)



Moduli at EW scale

Gauge Mediated  
(scale can be low  $\sim 30$  TeV)



Moduli are light ( $\text{mm}^{-1}$ ) →

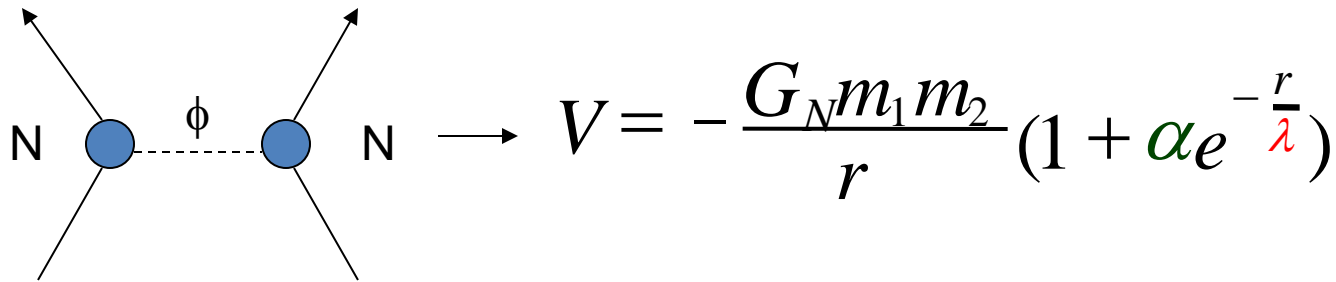
Observable signatures!

# Standard Model Couplings

May depend on moduli

- Yukawa couplings  $\rightarrow$  masses of quarks, leptons  $\sim \lambda_{ij}(\phi) \bar{u}_i Q_j H_u$
- Gauge couplings  $\rightarrow$  W,Z bosons, gluons  $\sim \lambda_g \frac{\phi}{M} G_{\mu\nu}^a G^{\mu\nu a}$

e.g. 2 nucleons can exchange gluon modulus:



$$V = -\frac{G_N m_1 m_2}{r} \left(1 + \alpha e^{-\frac{r}{\lambda}}\right)$$

range

$$\lambda(\text{mm}) \approx 0.8 \frac{(100 \text{ TeV})^2}{F} \left( \frac{M \lambda_g^{-1}}{5 \times 10^{17} \text{ GeV}} \right)$$

Strength relative to gravity

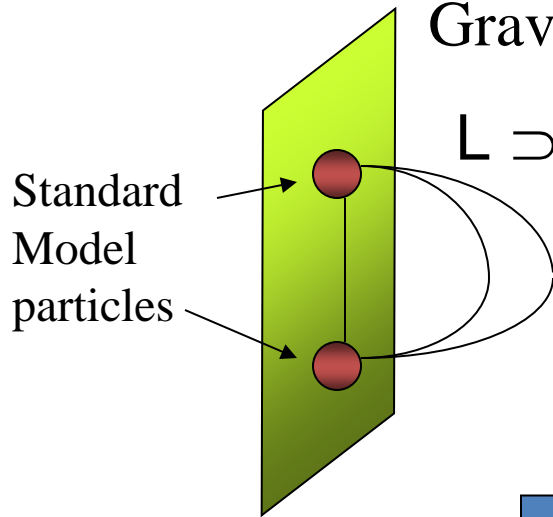
$$\alpha \propto \left( \frac{5 \times 10^{17} \text{ GeV}}{M} \right)^2 \lambda_g^2$$

# 2) Large Extra Dimensions

N. Arkani-Hamed, S. Dimopoulos, G.Dvali, Phys.Lett. B429 (1998)

Standard Model fields confined to a “brane”, except gravity

Gravity is “diluted” by the volume of extra space



$$L \supset \int d^4x d^n y M_*^{2+n} \sqrt{GR} \longrightarrow \int d^4x M_{\text{Pl}}^2 \sqrt{GR}$$

$$V_n M_*^{2+n} = M_{\text{Pl}}^2$$

Now  $M^*$  can be closer to weak scale at  $\sim \text{TeV}$



No more fine-tuning problem

For  $V_n \sim L^n$ , taking  $n=2$ ,  $L$  can be **1mm** for  $M_* \sim \text{TeV}$

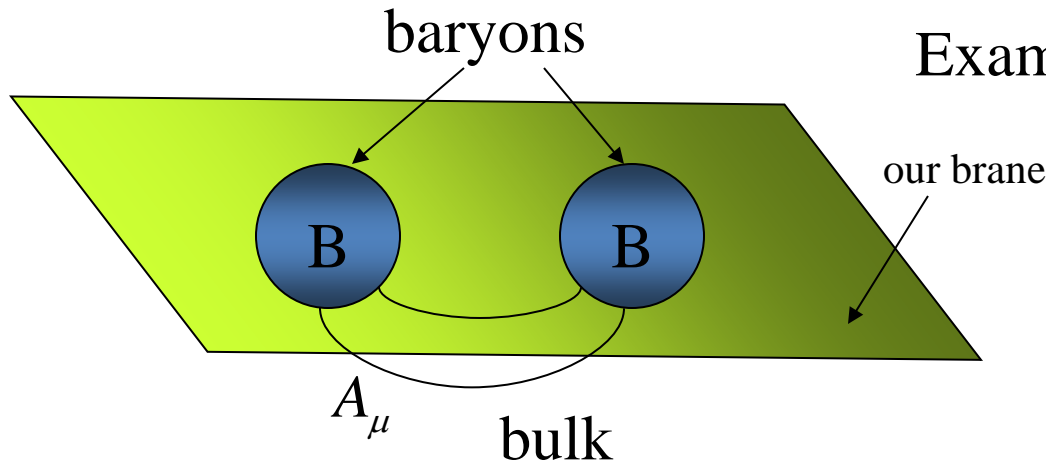
--Newton's law changed below radius of compactification

$$F(r) = \frac{G^{(4)} m_1 m_2}{r^2} \longrightarrow \frac{G^{(4+n)} m_1 m_2}{r^{2+n}}$$

# Gauge particles in the bulk

N. Arkani-Hamed, S. Dimopoulos, G. Dvali, Phys.Rev. **D59** (1999)

- 1) can mediate forces  $10^6$  x gravity
- 2) range of zero mode not limited by the radius of compact dimensions (so  $n > 2$  still accessible to tabletop)



Example – gauged baryon #

$$\frac{F_{\text{gauge}}}{F_{\text{gravity}}} = \frac{g_4^2}{G_N m_{\text{nucleon}}^2} \approx \frac{M_*^2}{m_{\text{nucleon}}^2} \approx 10^6 - 10^8$$

$$m_{A_\mu} \approx g_4 M_* \approx (\text{mm}^{-1})$$

other brane --  
gauge symmetry broken spontaneously by some "Higgs" field with vev  $\sim M_*$



# Scalars in Bulk

“Yukawa Messengers”

N. Arkani-Hamed, S. Dimopoulos, PRD 65, 052003 (2002)

- Also can get sub-mm range from SUSY breaking
- Coupling much stronger than moduli, and independent fundamental scale  $M^*$

effective coupling to the zero mode  $\rho \text{ ff } \chi_0$   $\rho = \frac{v}{M_{\text{Pl}}} \approx 10^{-16}$

but compare to gravity:

$$\frac{\rho^2}{G_N m_{\text{nucleon}}^2} \approx 10^6$$

- quite strong
- independent of “n”
- Independent of  $M^*$

Both vectors and scalars in the bulk can also produce a Yukawa-potential:

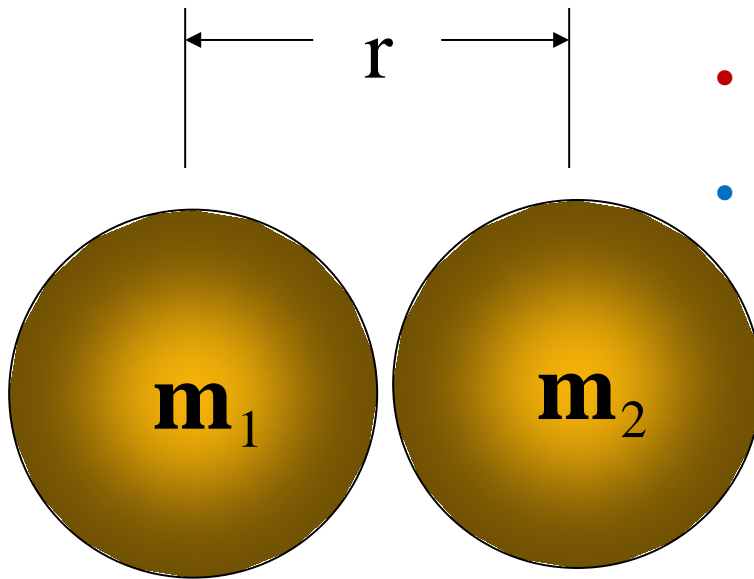
$$V = -\frac{G_N m_1 m_2}{r} \left(1 + \alpha e^{-\frac{r}{\lambda}}\right)$$

$\alpha > 0$  attractive (scalar)  
 $\alpha < 0$  repulsive (vector)

# Testing gravity at short range

$$V_N = -G \frac{m_1 m_2}{r} \left( 1 + \alpha e^{-r/\lambda} \right)$$

Exotic particles (new physics)

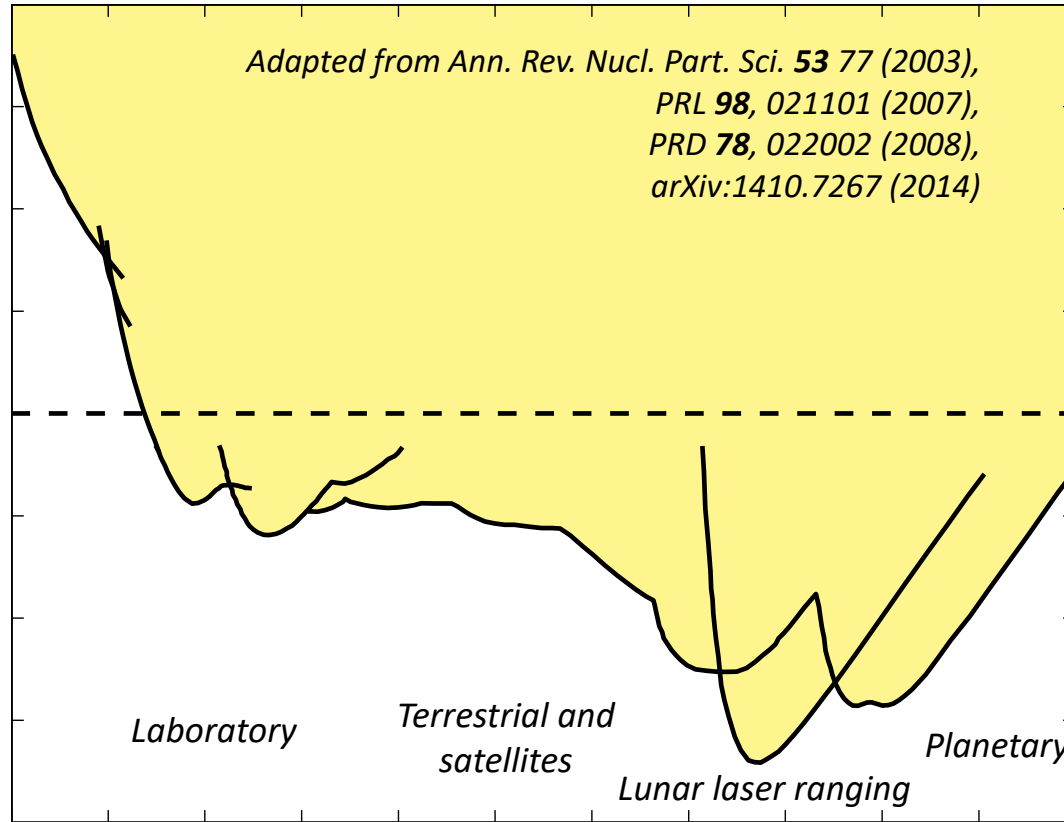


$$\lambda < 1 \text{ mm}$$

- Supersymmetry/string theory (moduli, radion, dilaton)
- Particles in large extra dimensions (Gravitons, scalars, vectors?)

# Landscape for ISL corrections

$$V_N = -G \frac{m_1 m_2}{r} \left( 1 + \alpha e^{-r/\lambda} \right)$$

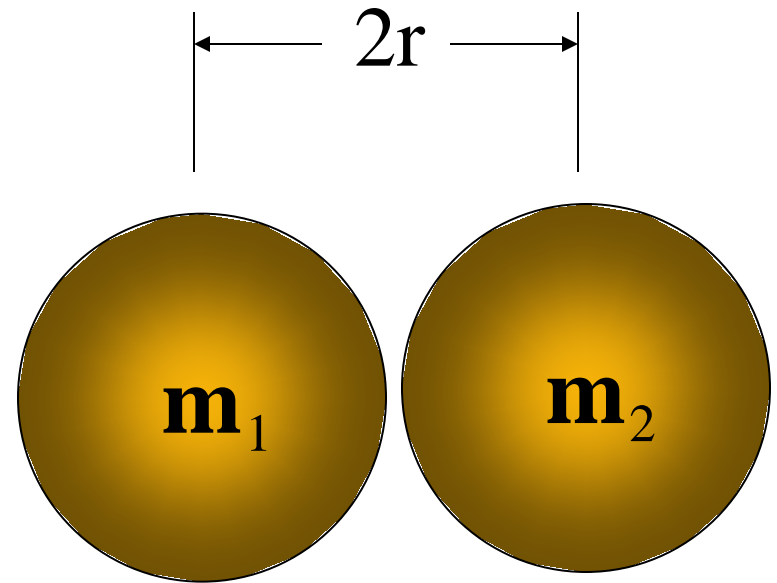


# Experimental challenge: scaling of gravitational force

$$V_N = -G \frac{m_1 m_2}{r}$$

$$F_N = G_N \frac{\rho^2 (4\pi r^3 / 3)^2}{4r^2} \sim G_N \rho^2 r^4$$

$$F_N \cong 0.1 r^4 \quad \text{for } \rho \sim 20 \text{ gr} / \text{cm}^3$$



In the range of experimental interest:

$$r \sim 10 \mu\text{m} ;$$

$$F_N \sim 10^{-21} \text{ N}$$

# Small forces



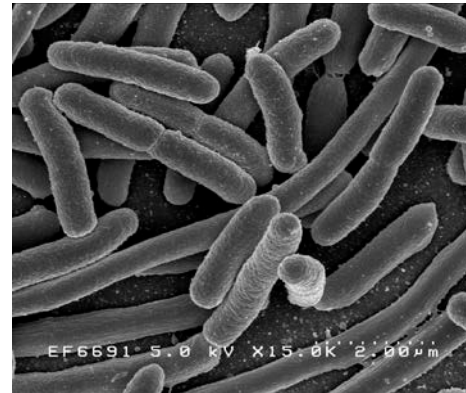
70 kg ~ 700 N

- Bathroom scales measure  $10^{-1}$  N

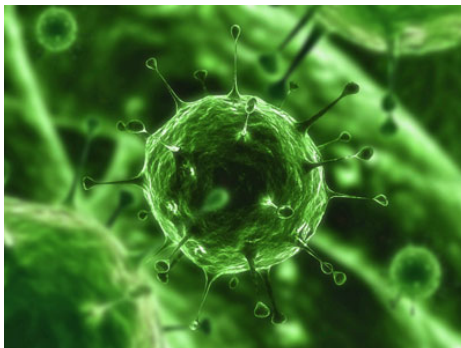
Dust mite  $10^{-7}$  N



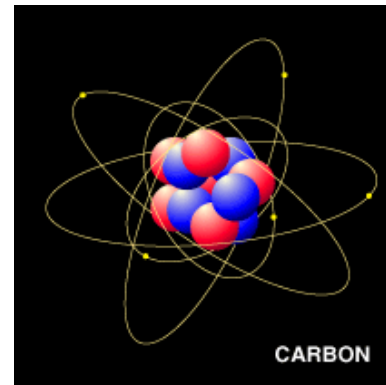
E. coli  $10^{-15}$  N



Virus  $10^{-19}$  N



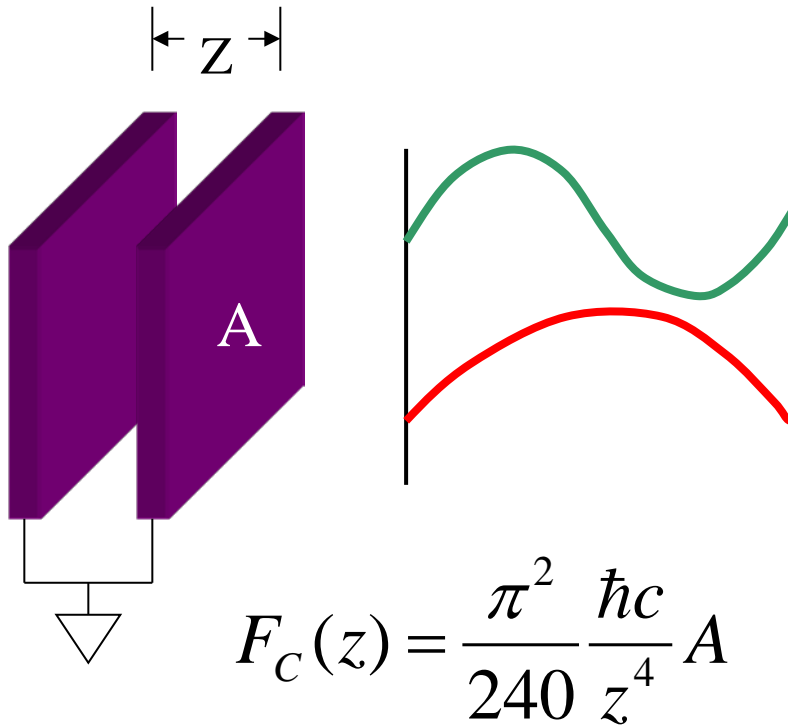
Carbon atom  $10^{-25}$  N



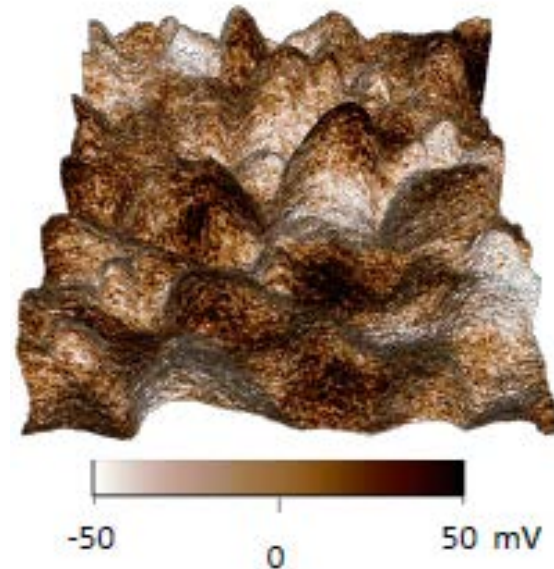
- AFM measures  $10^{-11}$  N

# Experimental challenge: Electromagnetic Background forces

Casimir effect (1948):

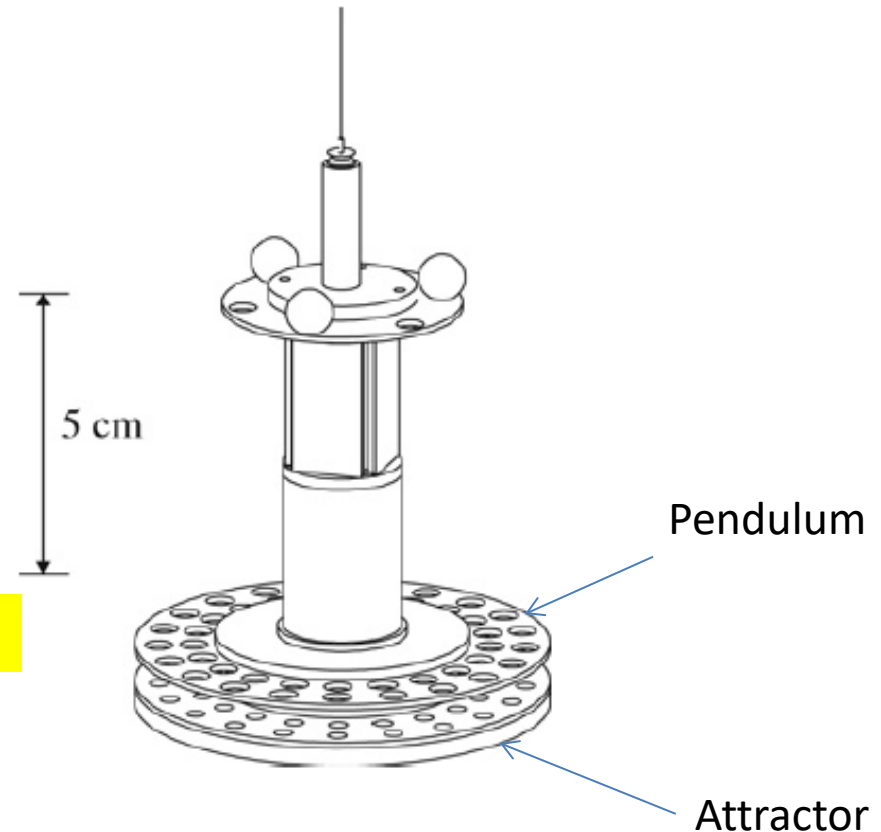
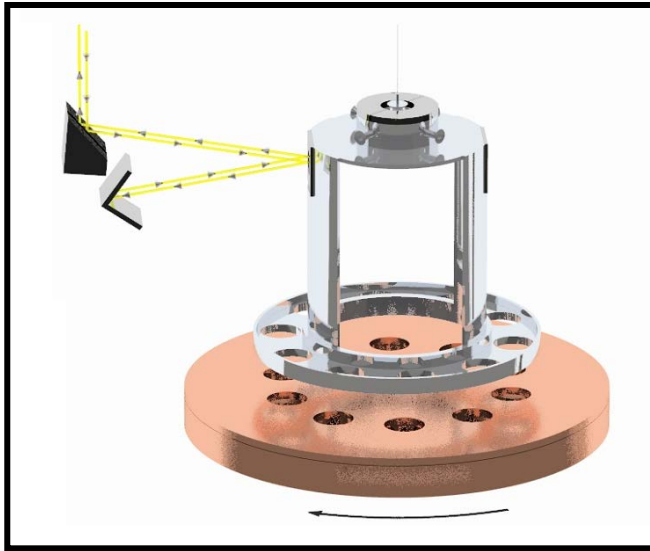


Electrostatic Patch Potentials:



J. L. Garrett, D. Somers, J. N. Munday  
J. Phys.: Condens. Matter 27 (2015) 214012

# UW Torsion Balance Experiments



Best yukawa constraints at  $\sim 10 \mu\text{m} - 5 \text{mm}$  range

1x Gravity tested at  $\sim 50 \mu\text{m}$

Kapner et. al., PRL 98, 021101 (2007)

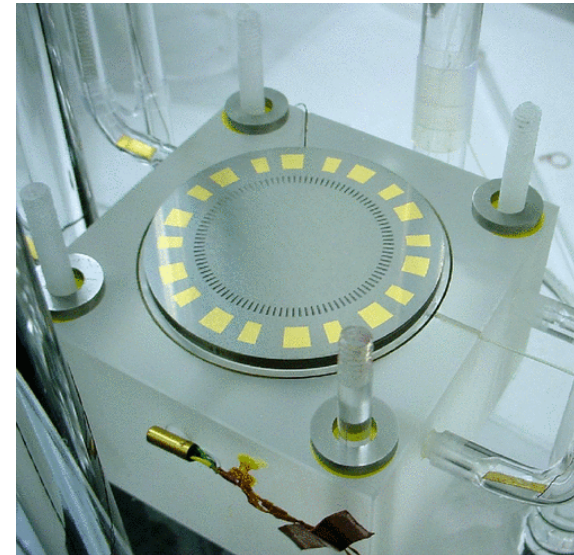
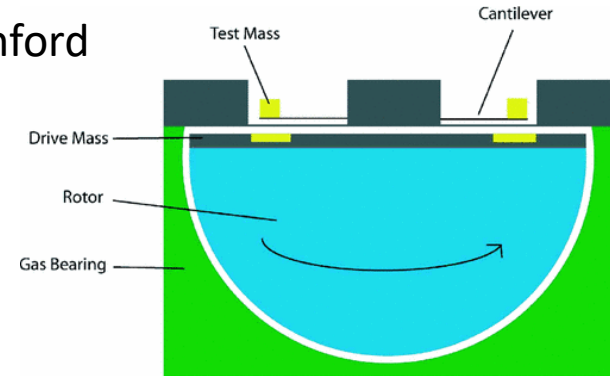
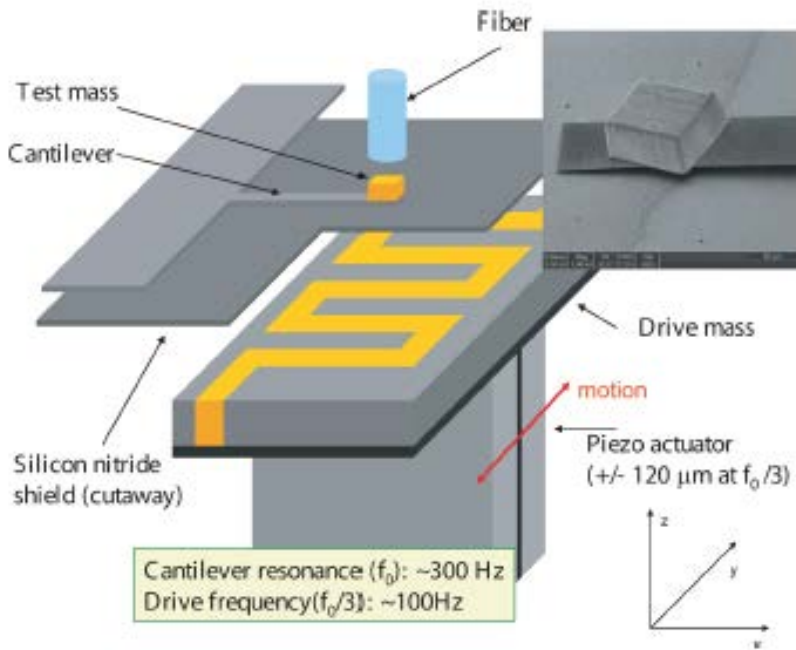
C.D. Hoyle et. al., PRL 86, 1418 (2001)

Review: E.G. Adelberger et. al. Progress in Particle and Nuclear Physics 62, 102 (2009)

Future: Cryogenic torsion balance experiments with improved sensitivity!

# Micro- and nano- resonators

Kapitulnik group, Stanford



A.A. Geraci, S.J. Smullin, D. M. Weld, J. Chiaverini, and A. Kapitulnik, *Phys. Rev. D* **78**, 022002 (2008).

D. M. Weld, J. Xia, B.Cabrera, and A. Kapitulnik *Phys. Rev. D* **77**, 062006 (2008)

Best yukawa constraints at  $\sim 10 \mu\text{m}$  range:

Next generation apparatus  
 $\sim 10\text{-}100 \times$  sensitivity

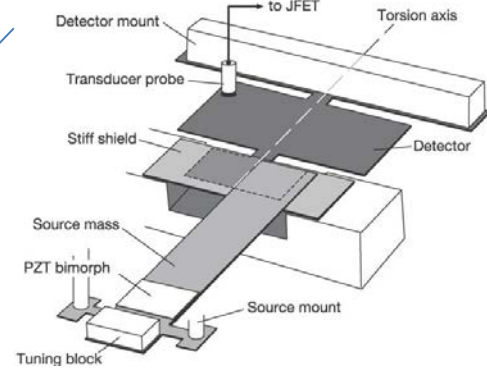


# Yukawa phase space

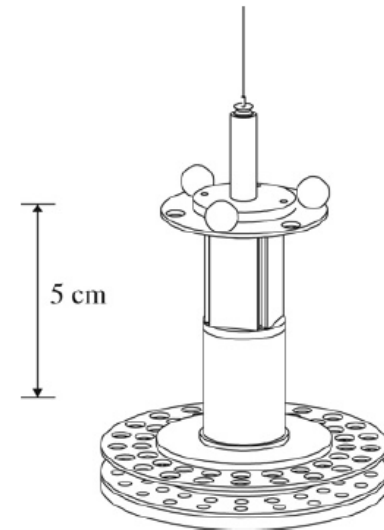
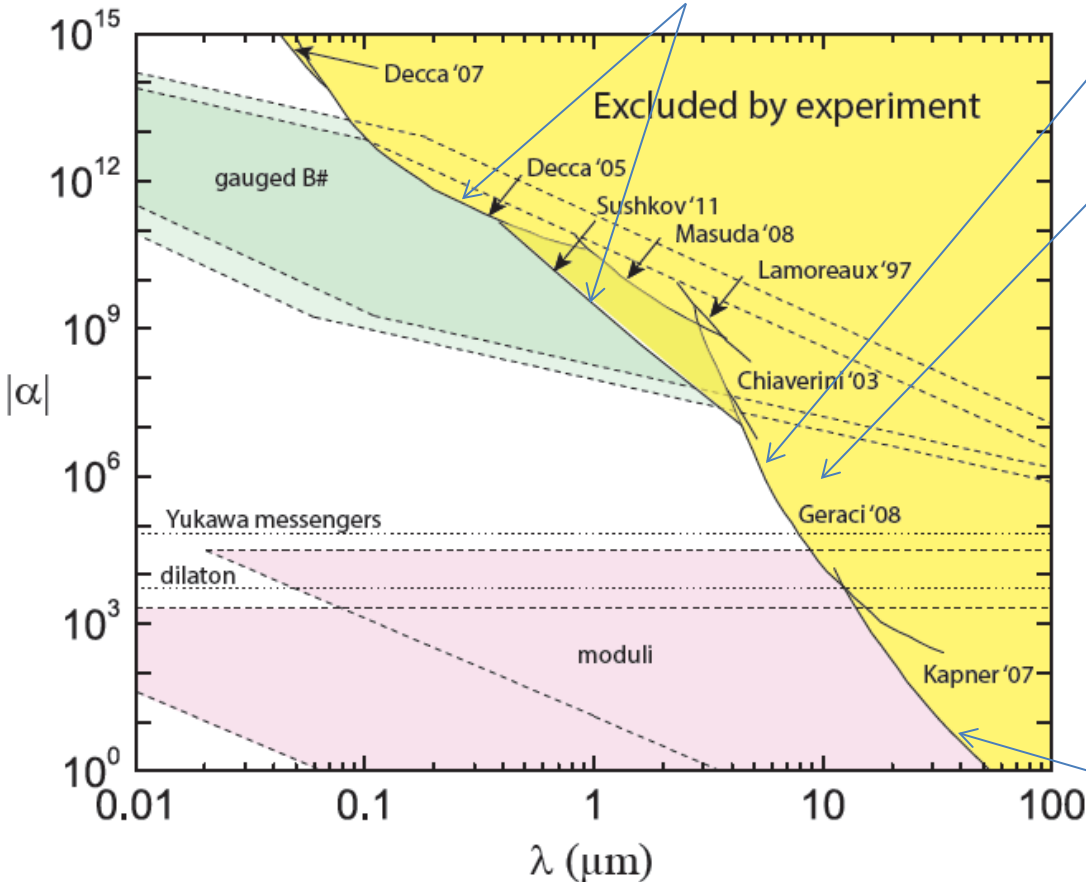
$$V_N = -G \frac{m_1 m_2}{r} \left( 1 + \alpha e^{-r/\lambda} \right)$$

Cantilevers  
(Stanford)

Torsion oscillators  
(Colorado, IU)



Casimir Measurements



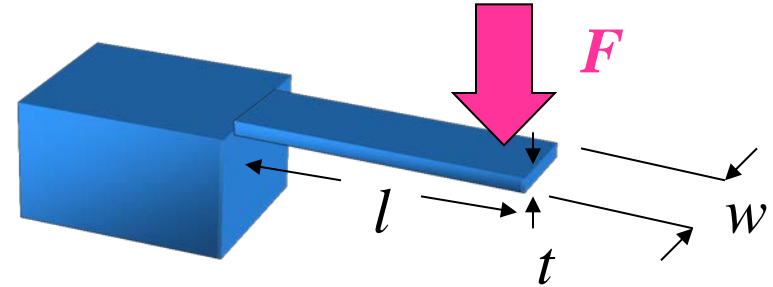
Torsion balance experiments  
(U Washington)

# Resonant force detection

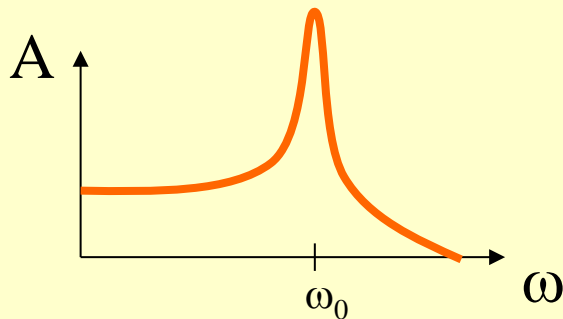
- Cantilever is like a spring:

$$F = -Kx$$

$$\omega_0 = \sqrt{\frac{K}{m}}$$



Amplitude:



$$A_{(\omega=0)} = \frac{F}{k}$$

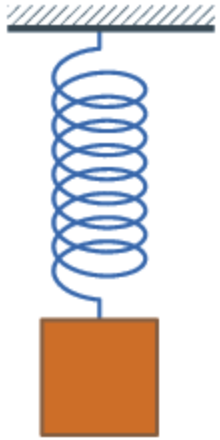
Constant force

$$A_{(\omega=\omega_0)} = \frac{F}{k} Q$$

Driving force on resonance  
of cantilever  $\omega_0$

Q can be very large >100,000

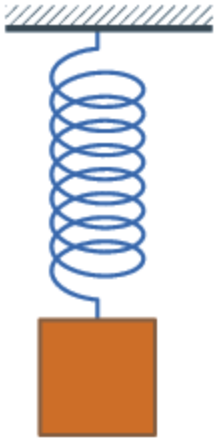
# Dissipation



$$E = \frac{1}{2}m\dot{x}^2 + \frac{1}{2}kx^2$$

Energy stored in oscillation

# Dissipation



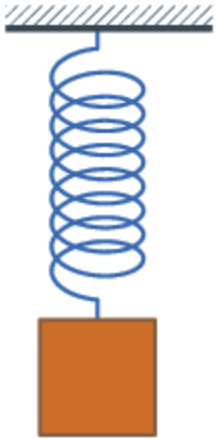
$$E = \frac{1}{2} m \dot{x}^2 + \frac{1}{2} k x^2$$

Energy stored in oscillation

$$Q \equiv 2\pi \frac{E}{\delta E}$$

Energy stored in oscillation / energy dissipated in 1 cycle

# Dissipation



$$E = \frac{1}{2} m \dot{x}^2 + \frac{1}{2} k x^2$$

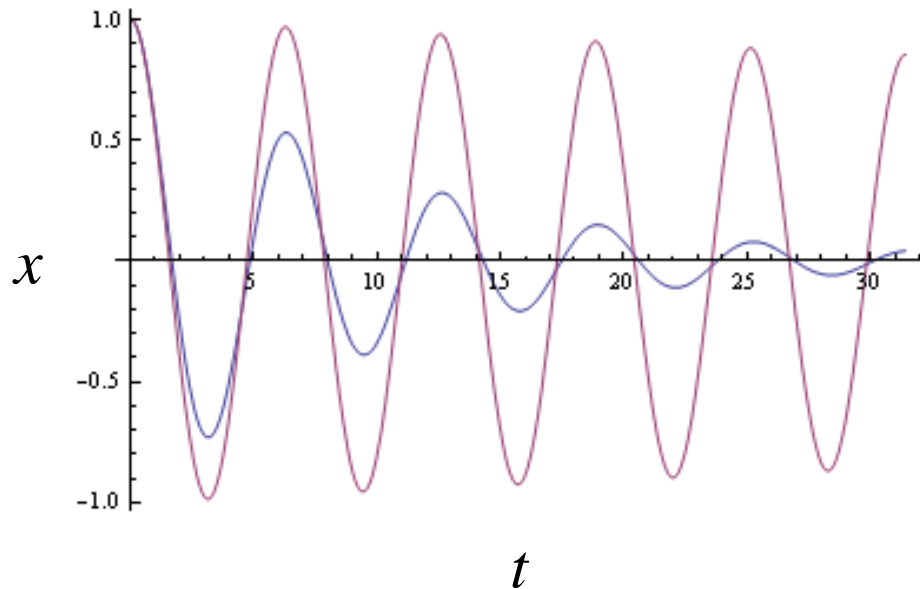
Energy stored in oscillation

$$Q \equiv 2\pi \frac{E}{\delta E}$$

Energy stored in oscillation / energy dissipated in 1 cycle

$$m \ddot{x} = -kx - \gamma \dot{x}$$

$$Q \approx \frac{\sqrt{mk}}{\gamma}$$



Plays crucial role in force detection

# Cantilever response

$$\ddot{x} + \omega_0^2 x + \gamma \dot{x} = \frac{F(t)}{m}$$

For a harmonic driving force:  $F(t) = f_0 e^{-i\omega t}$

$$-\omega^2 \ddot{x} + \omega_0^2 x - i\gamma\omega x = \frac{f_0}{m}$$

# Cantilever response

$$\ddot{x} + \omega_0^2 x + \gamma \dot{x} = \frac{F(t)}{m}$$

For a harmonic driving force:  $F(t) = f_0 e^{-i\omega t}$

$$-\omega^2 \ddot{x} + \omega_0^2 x - i\gamma\omega \dot{x} = \frac{f_0}{m}$$

Susceptibility:

$$\chi(\omega) = \frac{x(\omega)}{f(\omega)} = \frac{1}{m} \left( \frac{1}{-\omega^2 + \omega_0^2 - i\omega\gamma} \right)$$

$$\text{DC: } \omega \rightarrow 0, x = \frac{f_0}{k}$$

$$\lim_{\omega \rightarrow \infty} \chi(\omega) = -\frac{1}{m\omega^2}$$

# Cantilever response

$$\chi(\omega) = \frac{x(\omega)}{f(\omega)} = \frac{1}{m} \left( \frac{1}{-\omega^2 + \omega_0^2 - i\omega\gamma} \right)$$

Steady state  $x(t)$  in presence of driving force

$$f_0 \cos(\omega t)$$

$$\begin{aligned} x(t) &= \text{Re}[\chi(\omega) f_0 e^{-i\omega t}] \\ &= \text{Re}[|\chi(\omega)| f_0 e^{-i\omega t} e^{i\phi}] \\ &= f_0 |\chi(\omega)| \cos(\omega t - \phi(\omega)) \end{aligned}$$



# Cantilever response

$$\chi(\omega) = \frac{x(\omega)}{f(\omega)} = \frac{1}{m} \left( \frac{1}{-\omega^2 + \omega_0^2 - i\omega\gamma} \right)$$

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$$|\chi(\omega)| = \frac{1}{m} \frac{1}{\sqrt{(\omega^2 - \omega_0^2)^2 + \omega^2 \gamma^2}}$$

$$\tan \phi(\omega) = \frac{\omega\gamma}{\omega_0^2 - \omega^2}$$

# Cantilever response

$$\chi(\omega) = \frac{x(\omega)}{f(\omega)} = \frac{1}{m} \left( \frac{1}{-\omega^2 + \omega_0^2 - i\omega\gamma} \right)$$

Steady state  $x(t)$  in presence of driving force

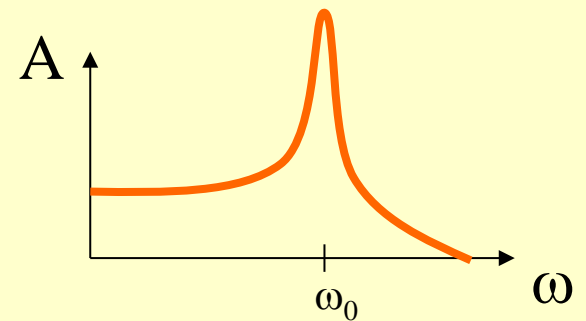
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$$\tan \phi(\omega) = \frac{\omega\gamma}{\omega_0^2 - \omega^2}$$

Amplitude:



$$A(f) = A_{dc} \frac{f_0^2}{\sqrt{(f_0^2 - f^2)^2 + \frac{f_0^2 f^2}{Q^2}}}$$

# Dissipation

$$\frac{dW}{dt} = F(t)\dot{x}(t)$$

Steady-state dissipated power:

$$P = \frac{1}{2} \omega f_0^2 |\chi(\omega)| \sin \phi(\omega) = \frac{1}{2} f_0^2 \omega \operatorname{Im}[\chi(\omega)]$$

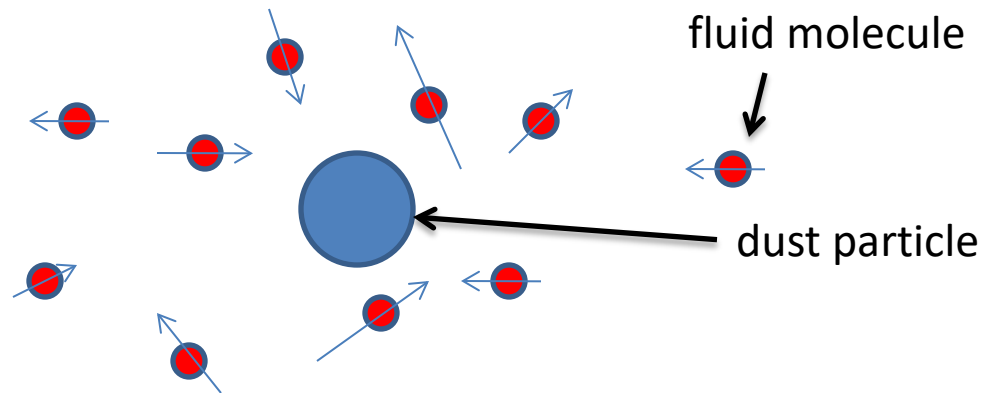
$$\operatorname{Im}[\chi(\omega)] = \frac{1}{m} \frac{\omega \gamma}{(\omega^2 - \omega_0^2)^2 + (\omega \gamma)^2}$$

On resonance:

$$P = \frac{1}{2} f_0^2 \frac{1}{m \gamma}$$

# Fundamental limitation: thermal noise

Brownian motion – random “kicks” given to particle due to thermal bath



- Random “kicks” are given to cantilever due to finite  $T$  of oscillator

$$\frac{1}{2}k\langle x^2 \rangle = \frac{1}{2}k_B T$$



$$F_{\min} = \left( \frac{4kk_B T b}{Q\omega_0} \right)^{1/2}$$

# Fundamental limitation: thermal noise

- White noise background due to finite T of oscillator

$$S_X(\omega) = |\chi(\omega)|^2 S_F$$

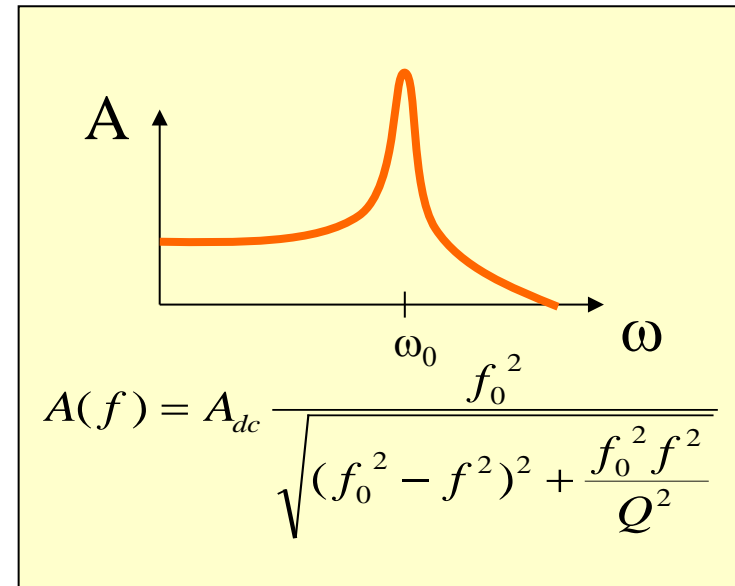
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- White noise background due to finite T of oscillator

$$S_X(\omega) = |\chi(\omega)|^2 S_F$$

$$\langle x^2 \rangle = \frac{1}{k^2} \int_0^\infty S_F \left( \frac{A(f)}{A_{dc}} \right)^2 df$$

$$S_F^{1/2} = \left( \frac{2}{\pi Q f_0} \right)^{1/2} k x_{rms}$$



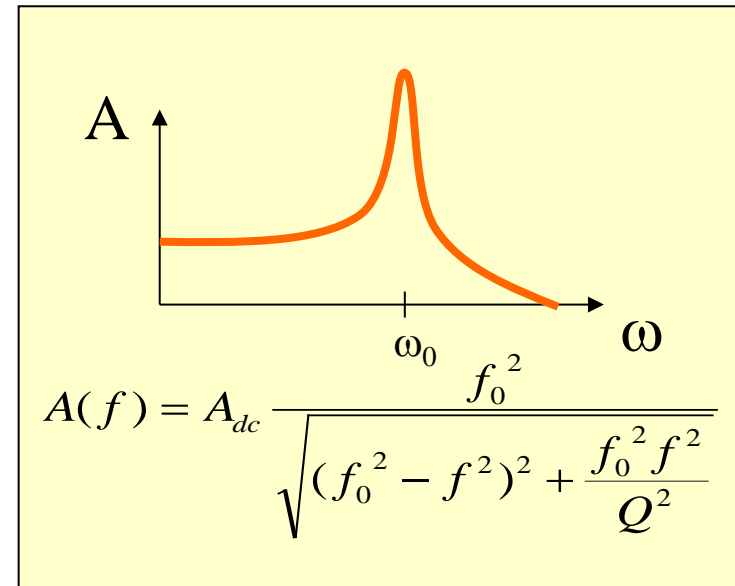
# Fundamental limitation: thermal noise

- White noise background due to finite T of oscillator

$$S_X(\omega) = |\chi(\omega)|^2 S_F$$

$$\langle x^2 \rangle = \frac{1}{k^2} \int_0^\infty S_F \left( \frac{A(f)}{A_{dc}} \right)^2 df$$

$$S_F^{1/2} = \left( \frac{2}{\pi Q f_0} \right)^{1/2} k x_{rms}$$



$$\frac{1}{2} k \langle x^2 \rangle = \frac{1}{2} k_B T \longrightarrow S_F^{1/2} = \left( \frac{4 k k_B T}{Q \omega_0} \right)^{1/2}$$

$$F_{min} = S_F^{1/2} B^{1/2} \longrightarrow$$

$$F_{min} = \left( \frac{4 k k_B T b}{Q \omega_0} \right)^{1/2}$$

# Fluctuation-Dissipation theorem

- In equilibrium, thermal fluctuations are related to dissipation:

$$S_F^{1/2} = \left( \frac{4k k_B T}{Q \omega_0} \right)^{1/2}$$

$$S_F = 4 k_B T m \Gamma$$

$$\Gamma = \omega_0 / Q$$

e.g. Johnson noise in a resistor:

$$S_V = 4 k_B T R$$



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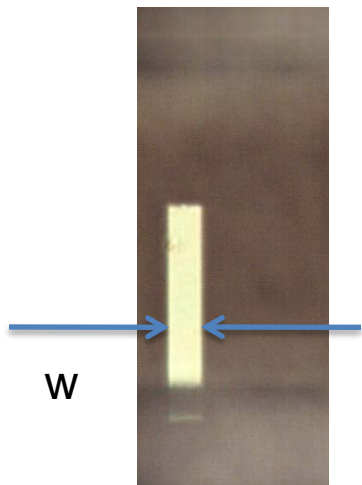
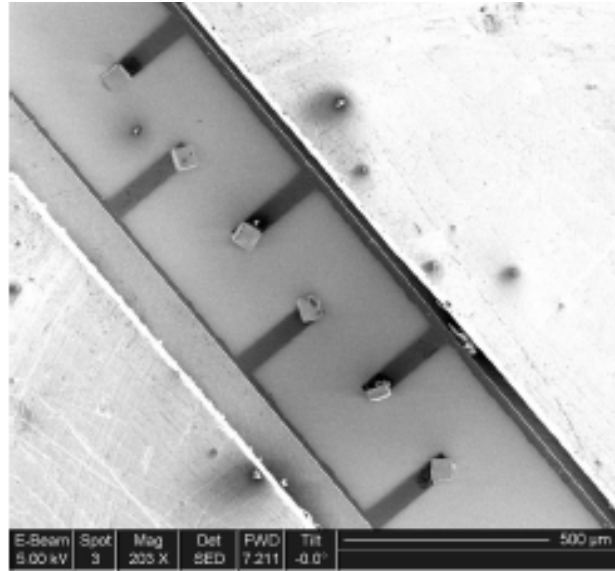
e.g. Johnson noise in a resistor:

$$S_V = 4k_B T R$$

$$\text{FDT: } S_X = 2k_B T \frac{\text{Im}[\chi(\omega)]}{\omega} \quad \text{Im}[\chi(\omega)] = \frac{1}{m} \frac{\omega\gamma}{(\omega^2 - \omega_0^2)^2 + (\omega\gamma)^2}$$

$$S_X(\omega) = |\chi(\omega)|^2 S_F$$

# Fundamental limitation: thermal noise



w= 50 μm  
l= 250 μm  
t=0.3 μm

$$F_{\min} = \left( \frac{4kk_B T b}{Q\omega_0} \right)^{1/2}$$

$$F_{\min} = (k_B T B)^{1/2} (E\rho)^{1/4} \left( \frac{wt^2}{lQ} \right)^{1/2}$$

*Silicon Cantilevers:*

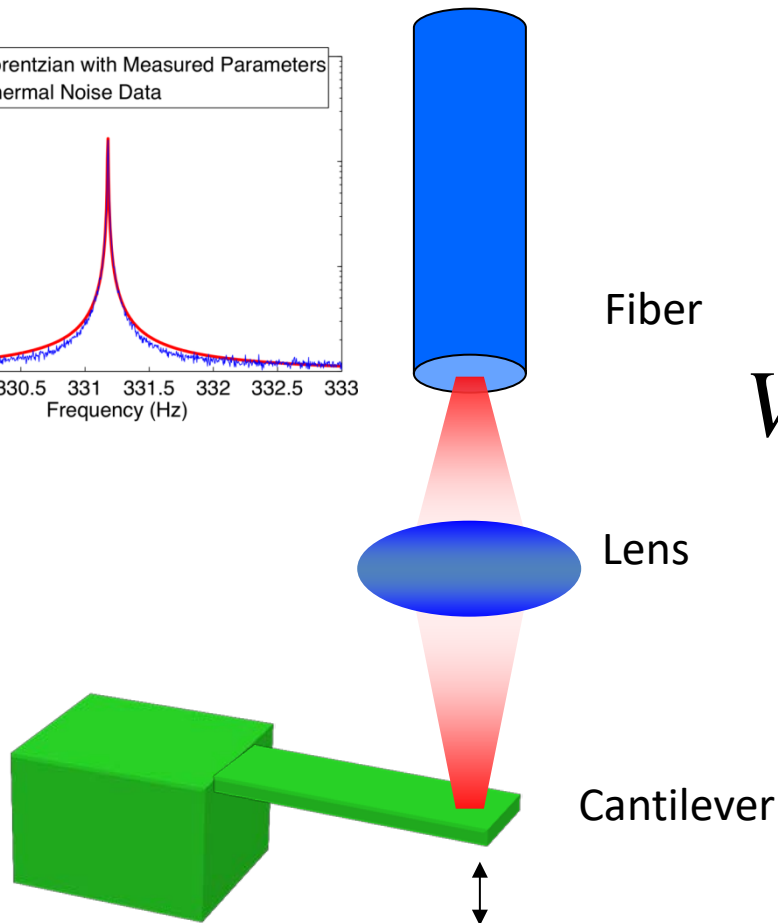
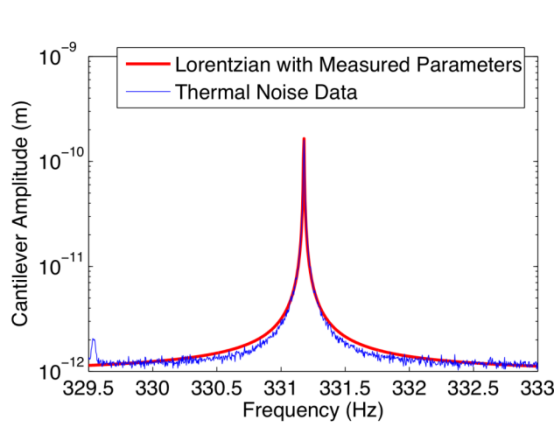
$F_{\min} \sim 10 \times 10^{-18} \text{ N}/\sqrt{\text{Hz}}$  at 4 K at  $Q=10^5$

To improve sensitivity:

- Make cantilever small
- Lower temperature
- Raise the quality factor

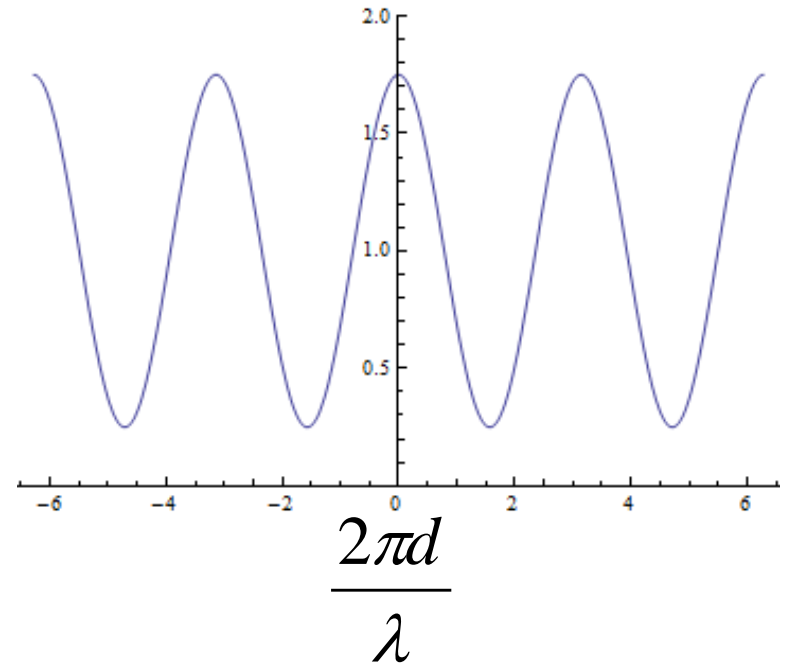
# Displacement detection

Cantilever is only moving few angstroms



$V$

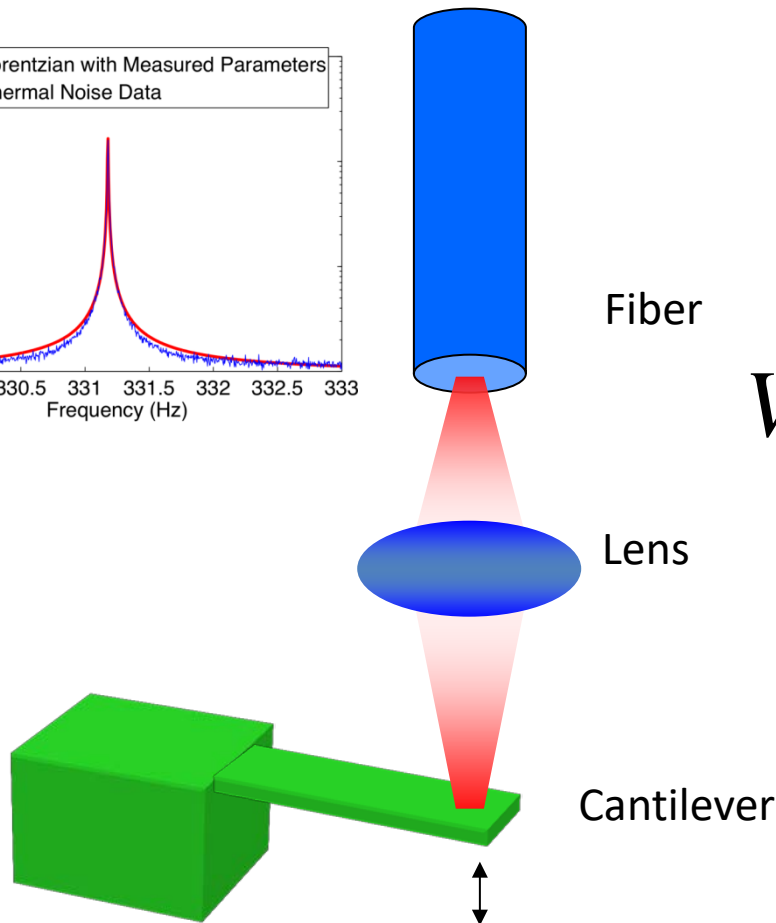
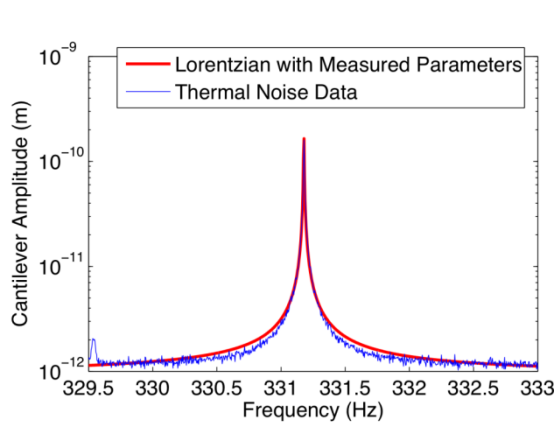
Interferometer signal



Can detect  $10^{-12}$  m/Hz<sup>1/2</sup> ( $\mu$ W)

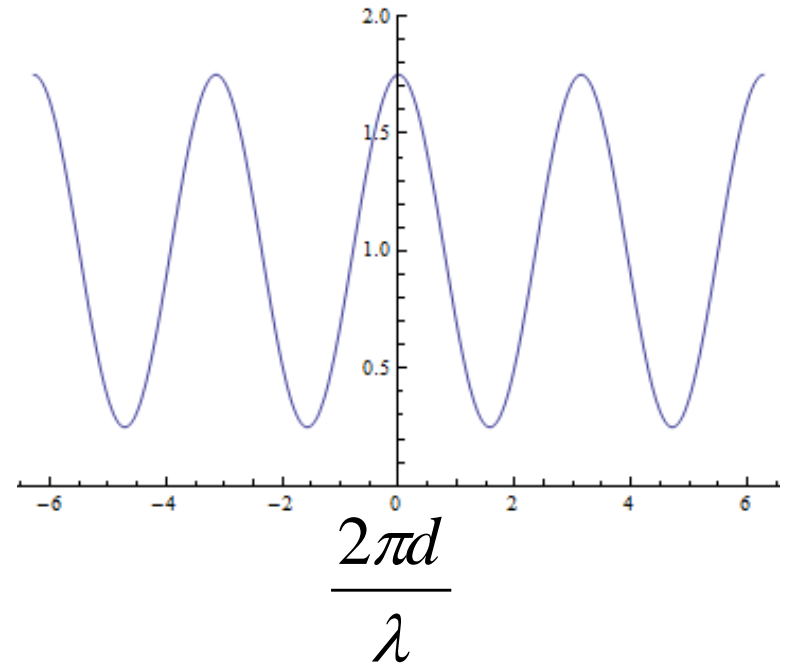
# Displacement detection

Cantilever is only moving few angstroms



$V$

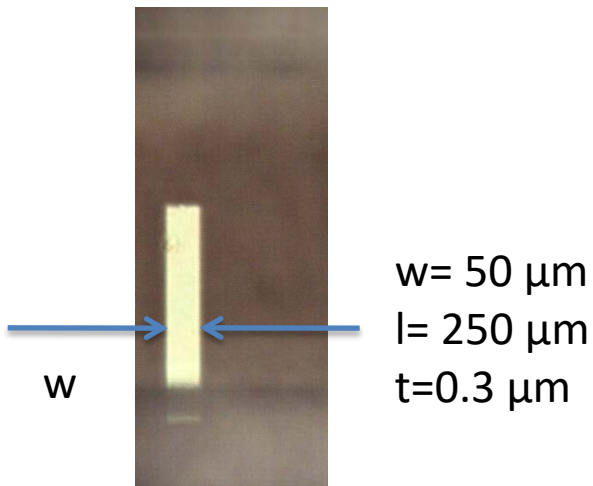
Interferometer signal



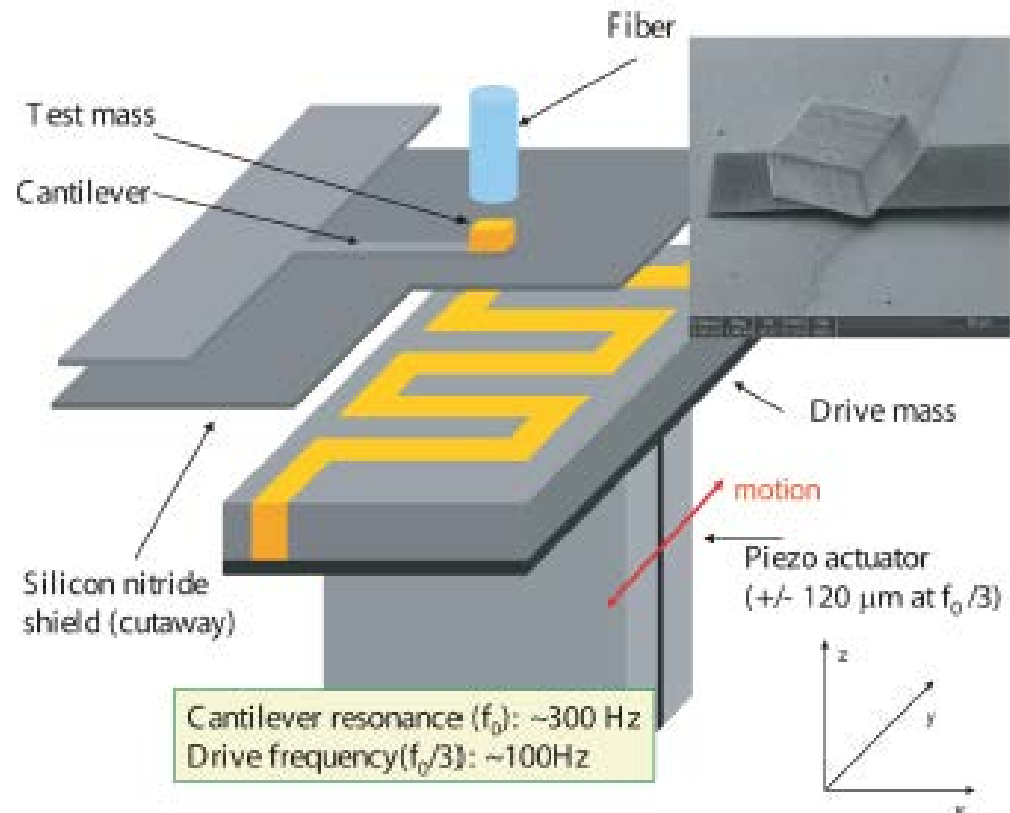
Can detect  $10^{-12} \text{ m/Hz}^{1/2}$  ( $\mu\text{W}$ )

Other methods: capacitance, QPD, high finesse optical cavity

# Stanford cantilever experiment



*Silicon Cantilevers:*  
 $F_{min} \sim 10 \times 10^{-18} \text{ N}/\sqrt{\text{Hz}}$   
at 4 K at  $Q = 10^5$

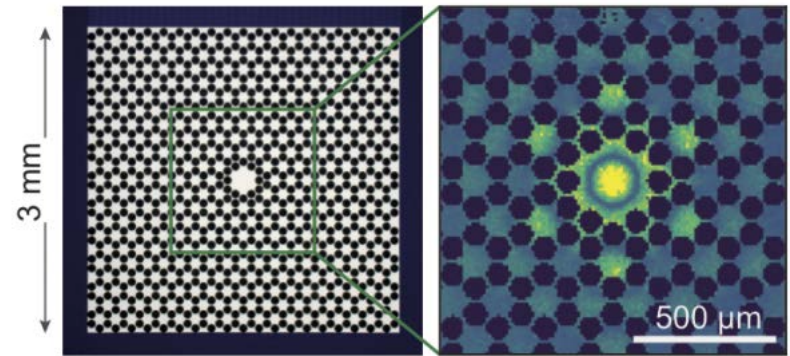
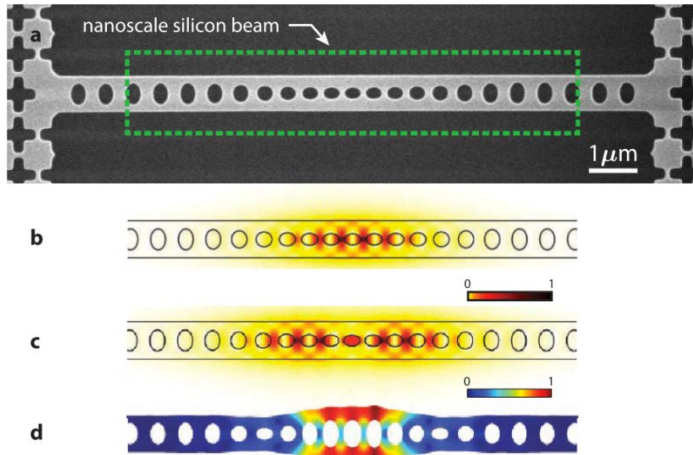


Best Yukawa constraints at  $\sim 10 \mu\text{m}$  range:

A.A. Geraci, S.J. Smullin, D. M. Weld, J. Chiaverini, and A. Kapitulnik,  
*Phys. Rev. D* 78, 022002 (2008).

# Advances in cryogenic nano-oscillators

Significantly improved sensitivities (higher frequencies)



Schliesser group, Copenhagen

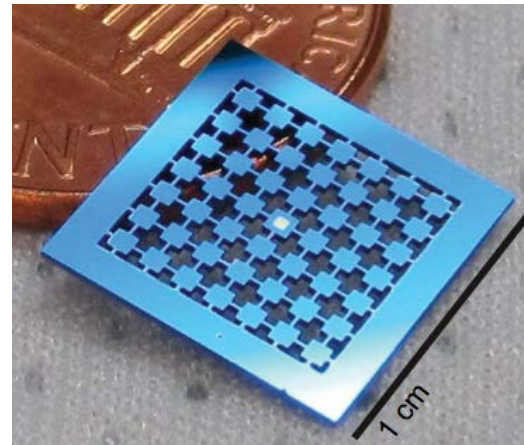
Painter group, Caltech

Si:  
freq=5 GHz  
 $Q_m=5 \times 10^{10}$   
mass =136 fg  
T=60 mK

Also nanotubes:

J. Moser, J. Guttinger, A. Eichler, M. J. Esplandiu, D. E. Liu, M. I. Dykman, and A. Bachtold, Nat. Nanotechnol. **8**, 493 (2013).

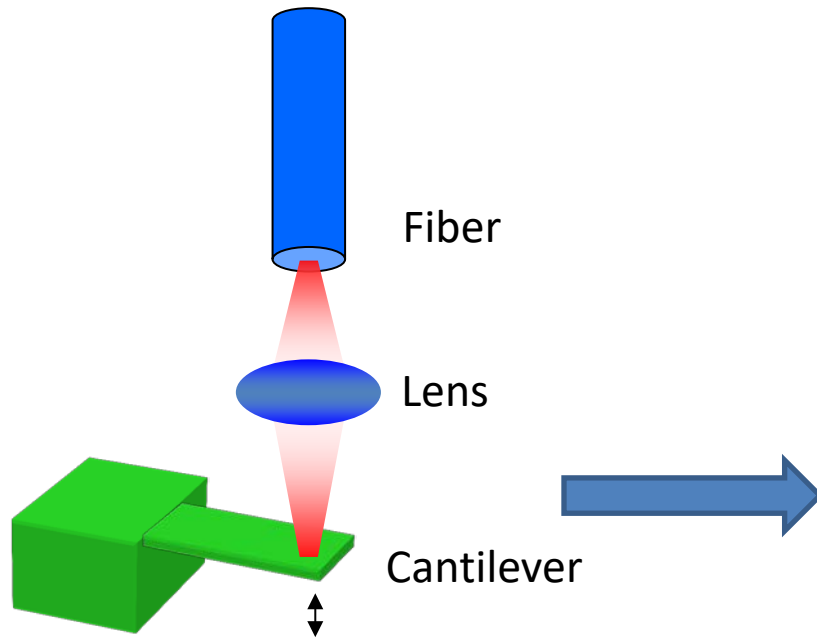
$$\sim 10 \text{ zN}/\sqrt{\text{Hz}}$$



Regal group, JILA

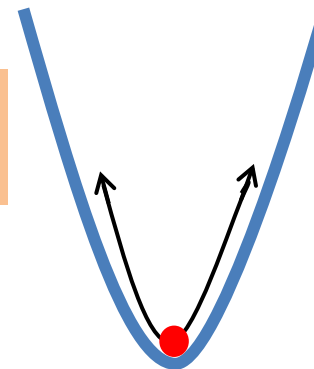
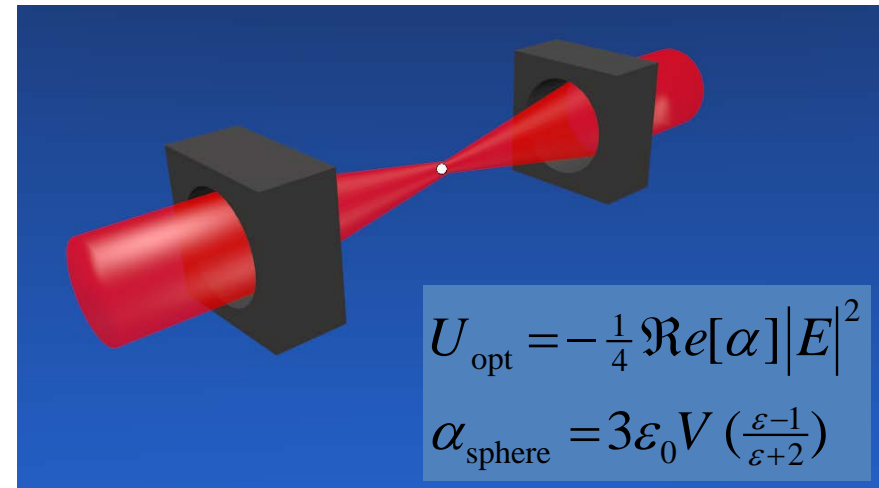
SiN:  
freq=1.5 MHz  
 $Q_m=2 \times 10^8$   
mass=10 ng  
T=30 mK

# Improving the sensitivity



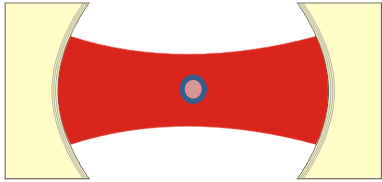
Limitations on Q: Clamping, surface imperfections, internal materials losses

Levitate the force sensor!



CM motion decoupled from environment – no clamping, materials losses

# Optically-levitated sensors



Decoupled from environment – no clamping, materials losses!

Pressure-limited damping

$$\frac{dp}{dt} = -\gamma_g \frac{p}{2} \qquad \frac{\gamma_g}{2} = \left( \frac{8}{\pi} \right) \frac{P}{\bar{v} r \rho}$$

$$P=10^{-10} \text{ Torr}, r=0.2 \text{ } \mu\text{m}, \omega/2\pi=100 \text{ kHz}, Q=10^{12} !$$

$$F_{\min} = \left( \frac{4kk_B T b}{Q\omega_0} \right)^{1/2}$$

$$Q \sim 10^{12}$$

$$T \sim 300$$

$$\omega_0/2\pi \sim 10^5$$

$$m \sim 10^{-(14-17)} \text{ kg}$$

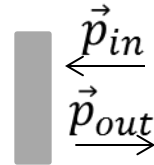
$$\rightarrow F \sim 10^{-21} \text{ N/Hz}^{1/2}$$



# Optical Trapping

Optical forces on objects due to radiation pressure

Light incident on a totally reflecting mirror:



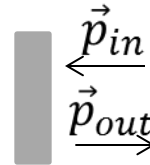
$$F_{rad} = 2 \frac{P}{c}$$

$$F_{rad} = \frac{2 \cdot 1W}{3 \cdot 10^8 m/s} = 6.67 nN$$

# Optical Trapping

Optical forces on objects due to radiation pressure

Light incident on a totally reflecting mirror:



$$F_{rad} = 2 \frac{IA}{c}$$

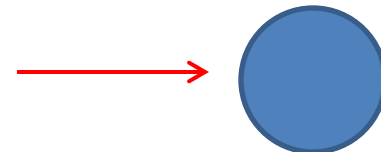
$$F_{rad} = \frac{2 \cdot 1W}{3 \cdot 10^8 m/s} = 6.67 nN$$

Scattering Force

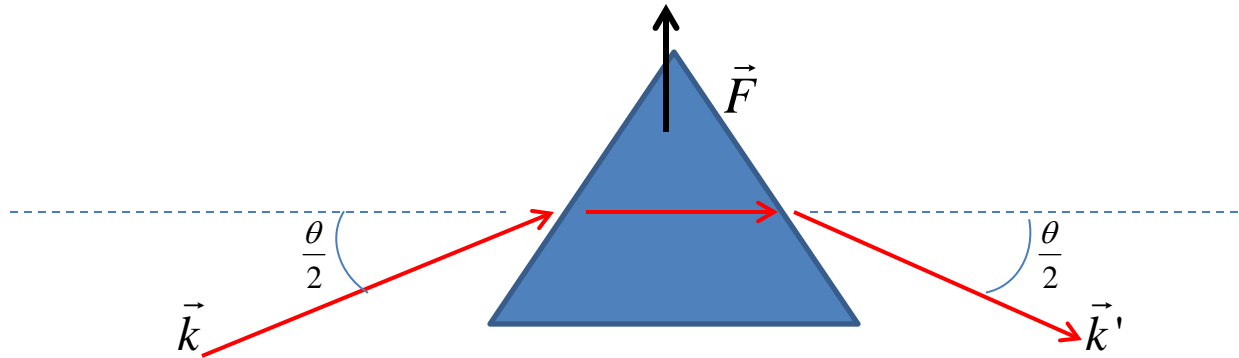
Imparted by traveling wave

$$F_{scat} = \hbar k R_{scat}$$

Dielectric sphere



# Optical Trapping



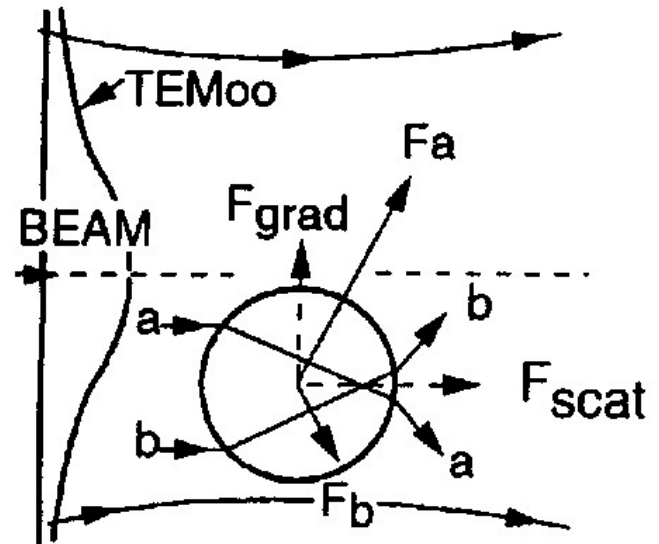
$$F = \frac{IA}{c} 2 \sin\left(\frac{\theta}{2}\right)$$

- Force arising from refraction of light

# Optical Trapping

## Ray Optics Approximation ( $R \gg \lambda$ )

- When photons strike bead, they are refracted and momentum changes
  - Forces
    - Scattering force occurs along the axial direction
      - Axial component of  $F_a + F_b$
    - Gradient force is along direction of increasing intensity



# Rayleigh Approximation ( $R \ll \lambda$ )

- Glass bead treated as induced dipole
- Scattering force
  - Electric field oscillates in time
  - Sphere acts as oscillating electric dipole
  - Radiates secondary (scattering) waves in all directions
  - Changes magnitude and direction of energy flux of electromagnetic field
  - Field due to momentum changes in electromagnetic field due to scattering by dipole

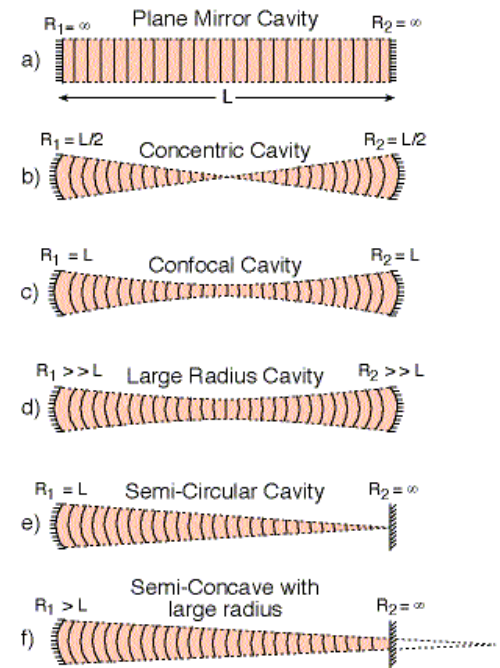
$$\vec{F}_{Scat}(\vec{r}) = \left(\frac{n}{c}\right) C_{scat} I(\vec{r}) \hat{z} = \frac{128\pi^5 R^6}{3c\lambda_0^4} \left(\frac{m^2 - 1}{m^2 + 2}\right)^2 n_{md}^5 I(\vec{r}) \hat{z}$$

- Gradient Force

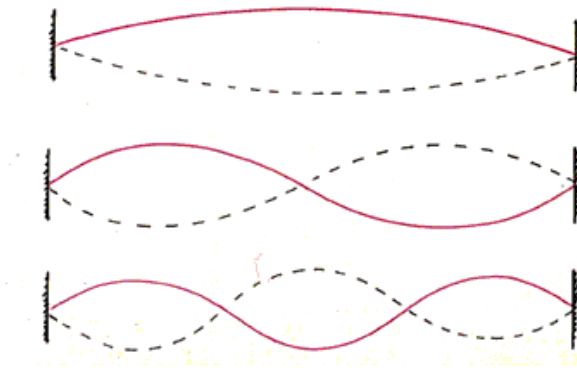
$$\vec{F}_{grad}(\vec{r}, t) = \vec{\nabla}[\vec{p}(\vec{r}, t) \cdot \vec{E}(\vec{r}, t)] = \frac{2\pi n_{md} R^3}{c} \left(\frac{m^2 - 1}{m^2 + 2}\right)^2 \vec{\nabla} I(\vec{r})$$

# Optical Cavity

- An arrangement of mirrors that forms a standing wave of light waves.
- Cavity Resonance Condition
  - Length between mirrors must be an integer multiple of one half the laser wavelength.
  - $L = m \left(\frac{\lambda}{2}\right)$
  - Standing wave

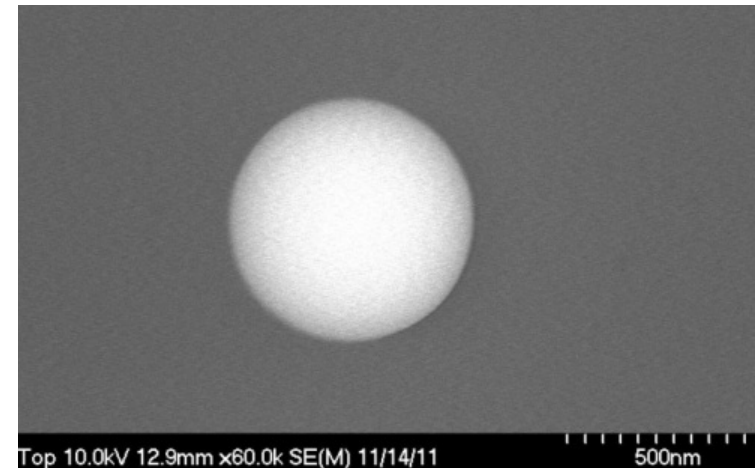
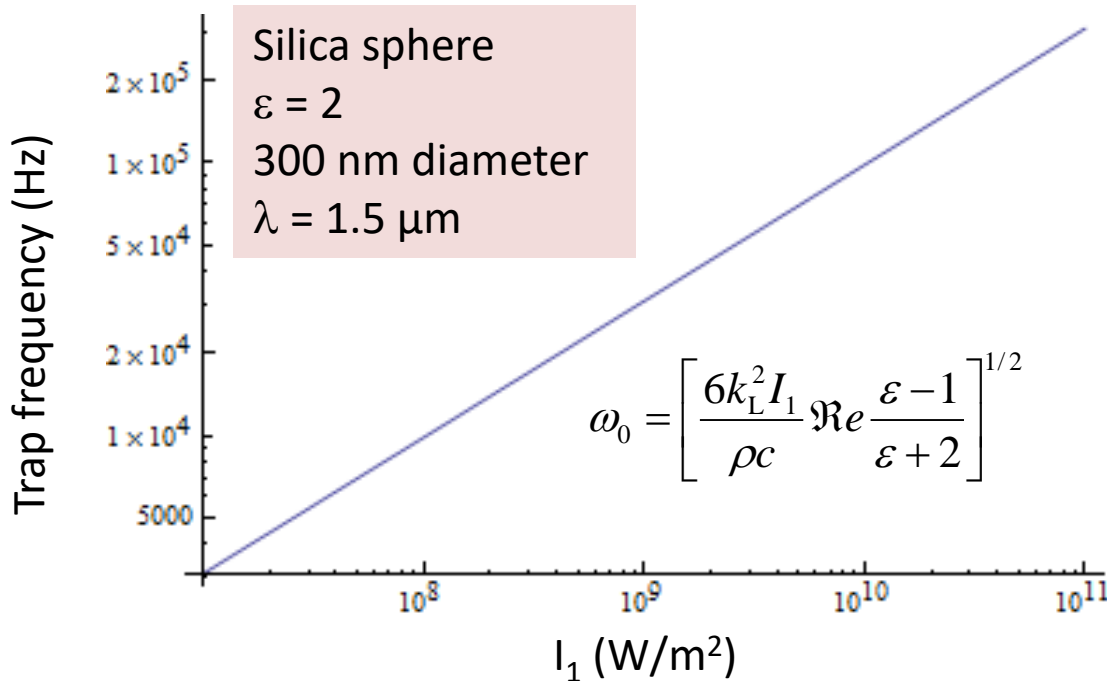
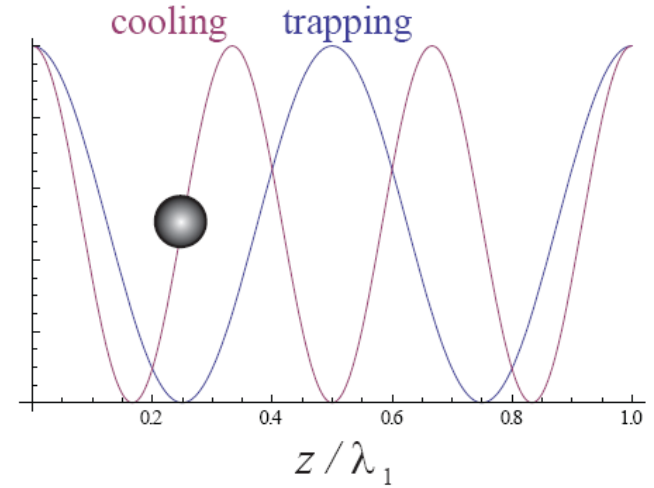
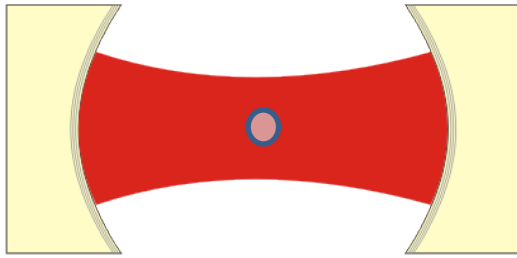


<http://perg.phys.ksu.edu/vqm/laserweb/ch-8/f8s1t1p1.htm>



<http://www.watervalley.net/users/bolen/page5.html>

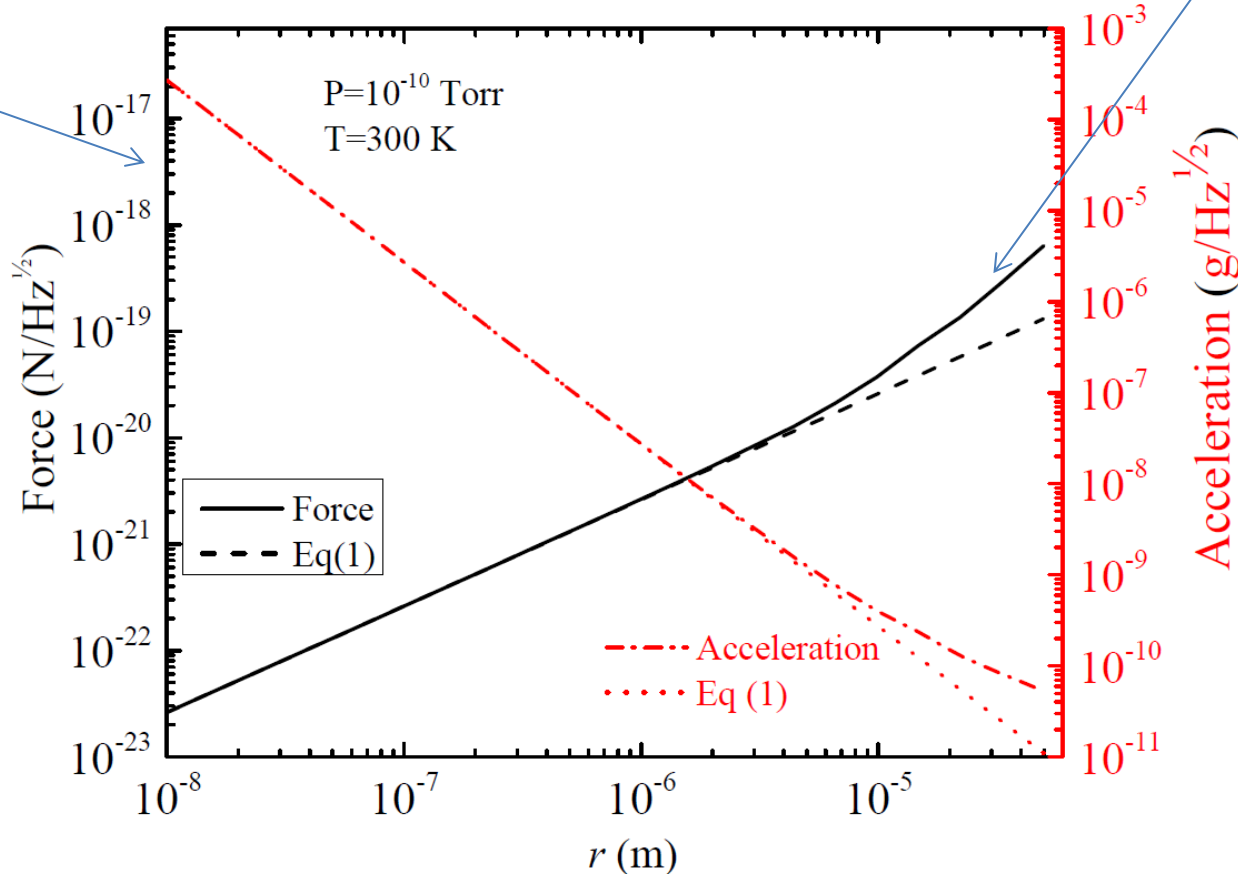
# Standing-wave trap



# Projected sensitivity

$$F_{\min} = (4k_B T \gamma m)^{1/2} \quad (1)$$

Cantilevers



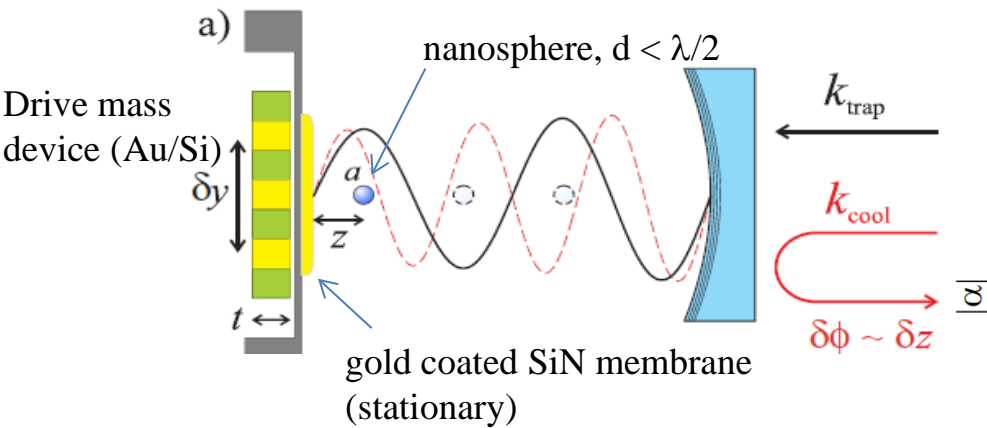
Photon recoil heating  
Seen recently by  
Novotny group  
V. Jain et. al.,  
PRL 116, 243601  
(2016)

20 zN/Hz<sup>1/2</sup> Gieseler, Novotny, Quidant (Nature Phys. 2013)

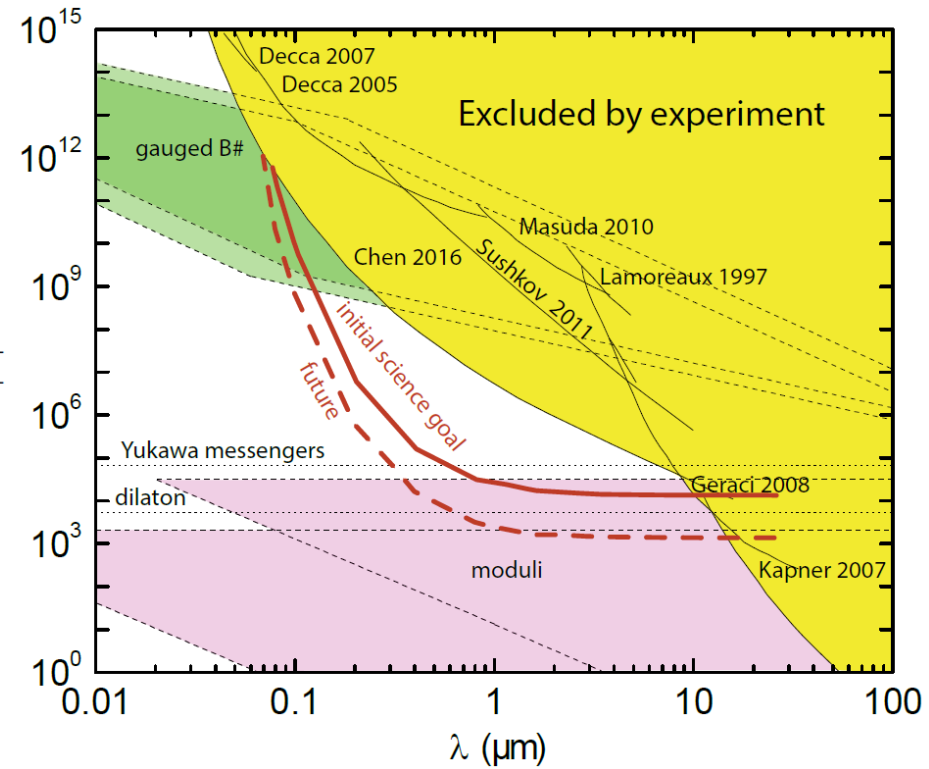
Z. Yin, A. Geraci, T. Li, Int. J. Mod. Phys. B 27,1330018 (2013).



# Projected sensitivity

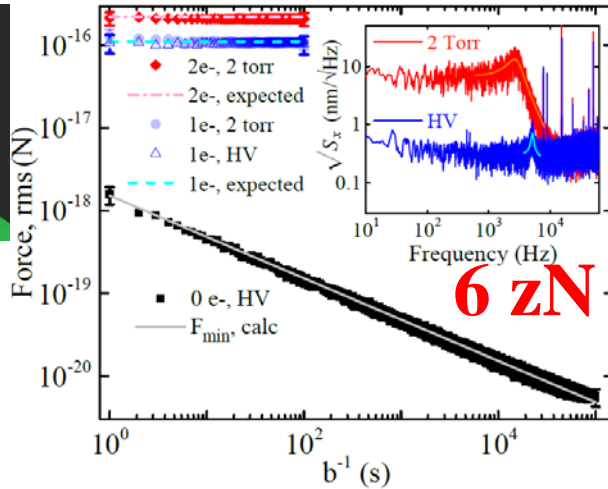
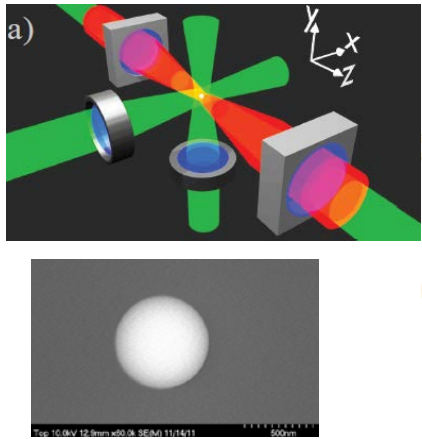


## Projected Reach



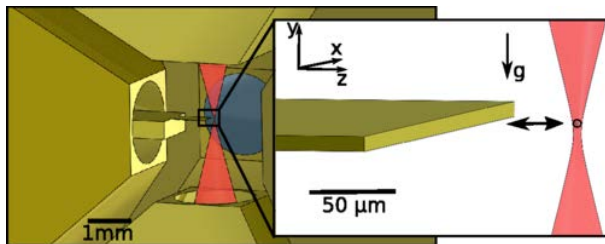
# Optically-levitated micro-particles

High Q Mechanical Resonance:  
Optically levitated microspheres



G. Ranjit et.al., *Phys. Rev. A* **91**, 051805(R) (2015).

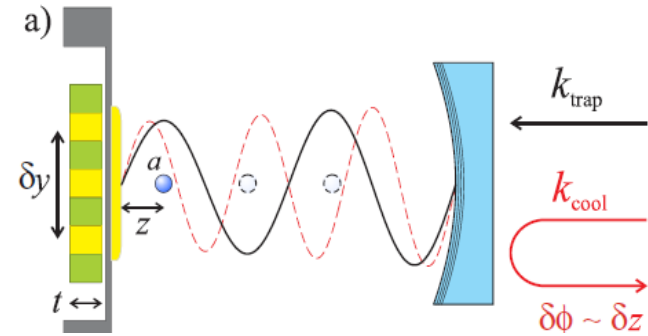
G. Ranjit, M. Cunningham, K. Casey, and AG, *Phys. Rev. A*, **93**, 053801 (2016).



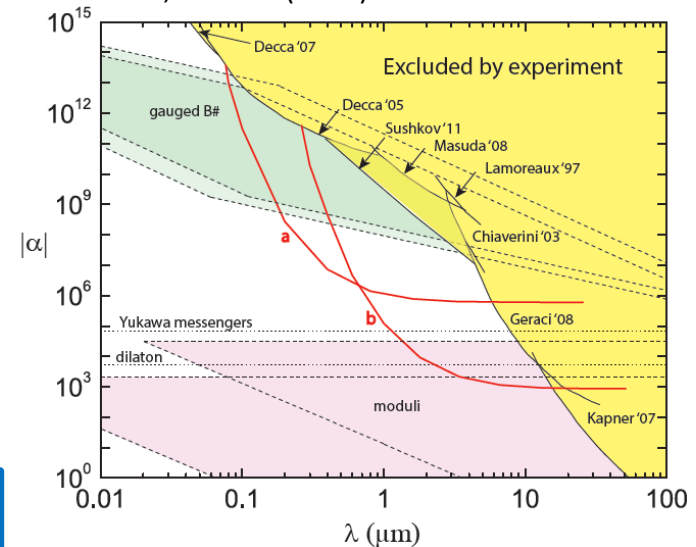
A. D. Rider, D. C. Moore, C. P. Blakemore, M. Louis, M. Lu, and G. Gratta  
*Phys. Rev. Lett.* **117**, 101101 (2016)

## New Physics

Gravity at micron scales



AG., S. Papp, and J. Kitching, *Phys. Rev. Lett.* **105**, 101101 (2010).

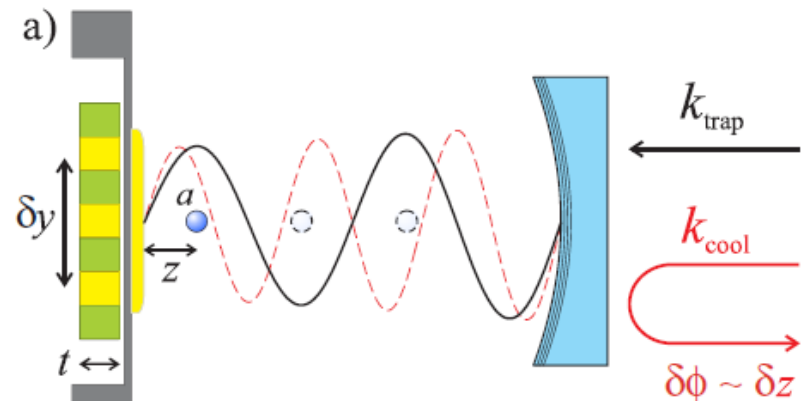
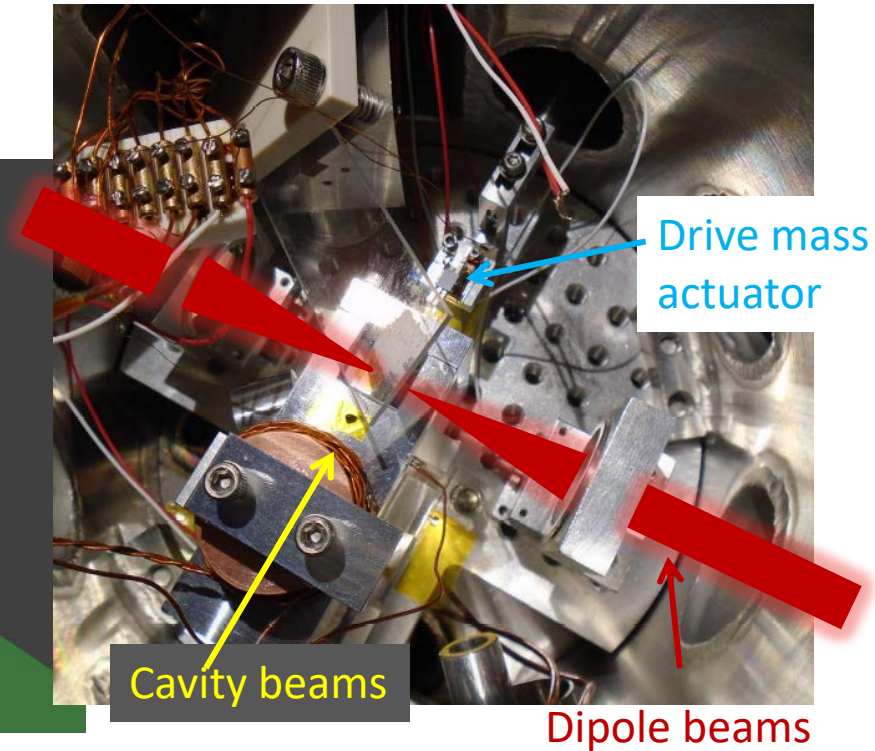
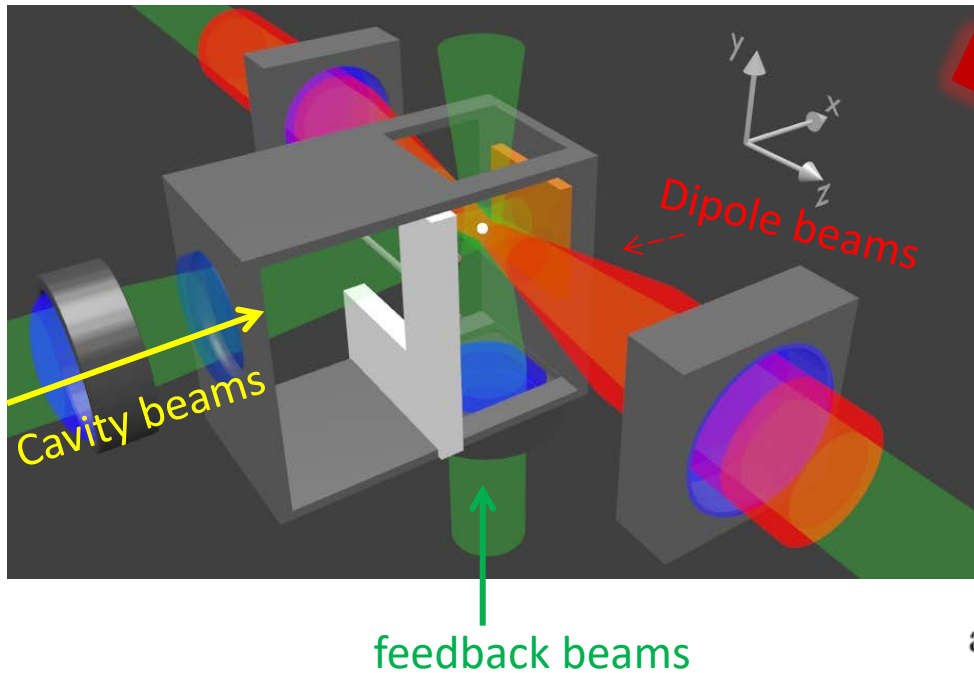


Gravitational Waves

A. Arvanitaki and AG., *Phys. Rev. Lett.* **110**, 071105 (2013).

# Experimental Setup

Combined cavity/dipole trap



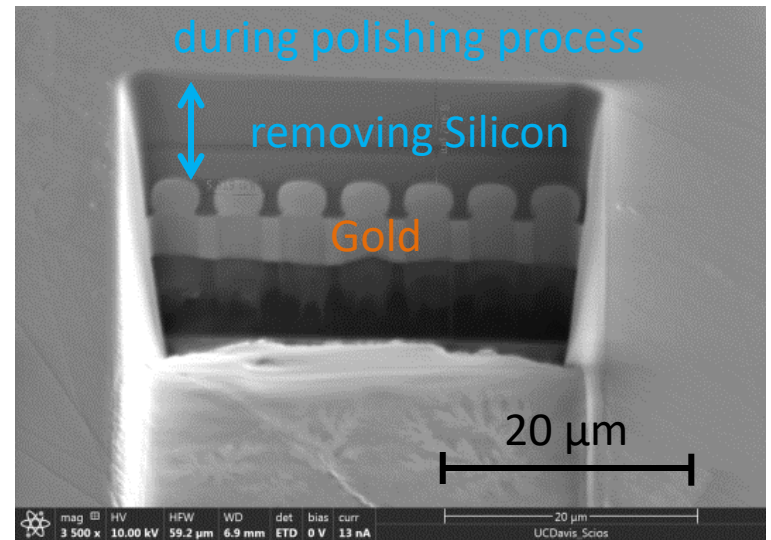
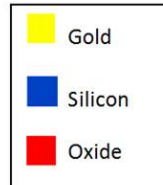
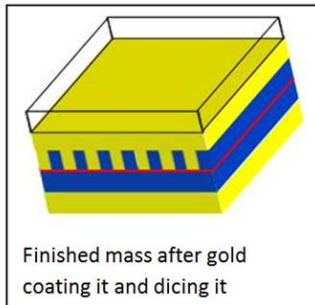
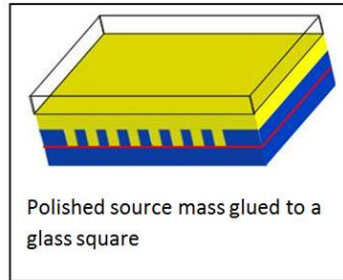
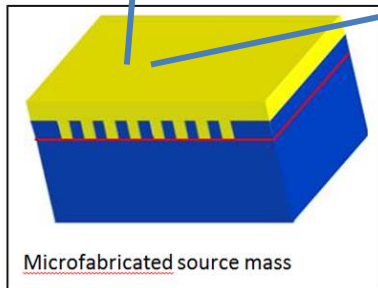
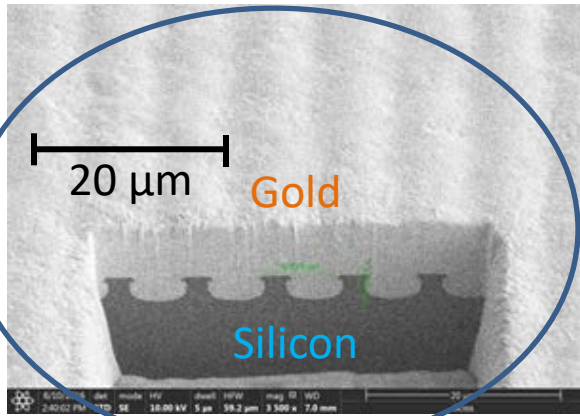
AG, S.B. Papp, and J. Kitching, *Phys. Rev. Lett.* **105**, 101101 (2010)

G. Ranjit et.al., PRA 91, 051805(R) (2015).

G. Ranjit et.al. , *Phys. Rev. A*, 93, 053801 (2016).

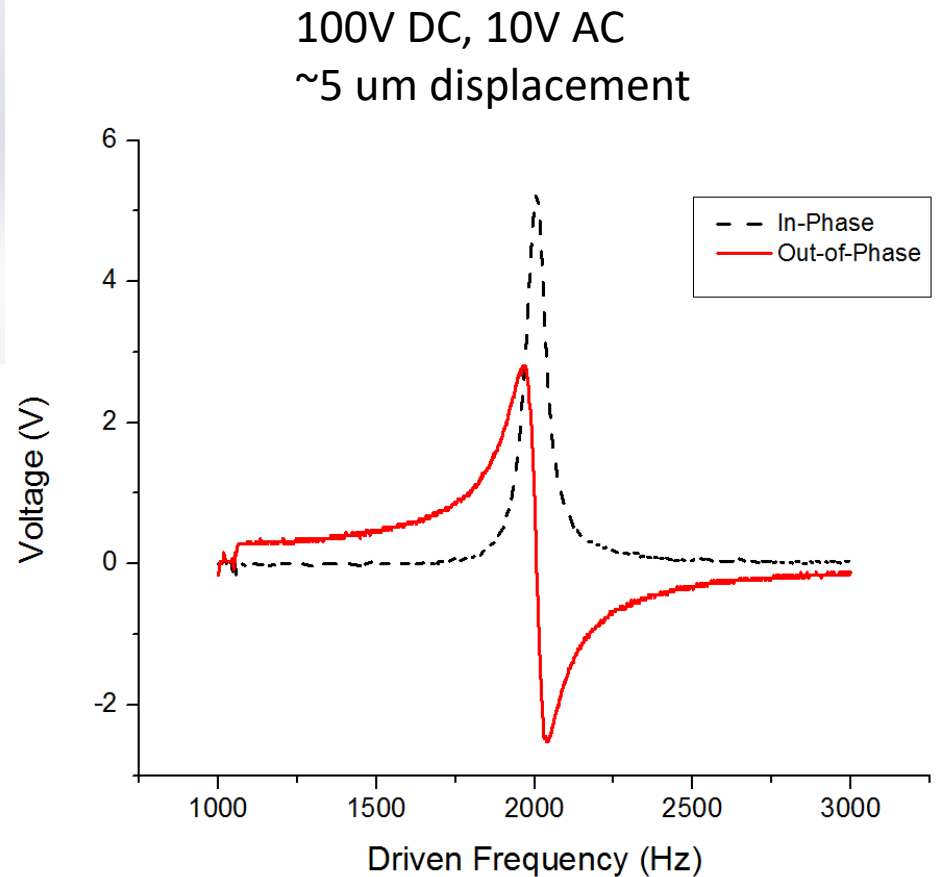
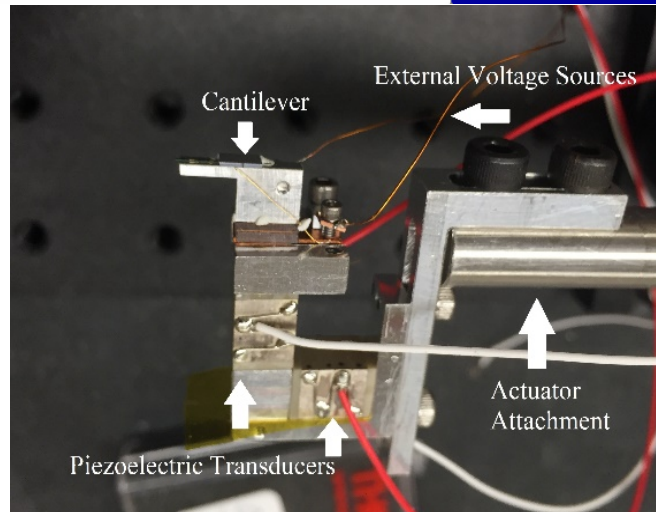
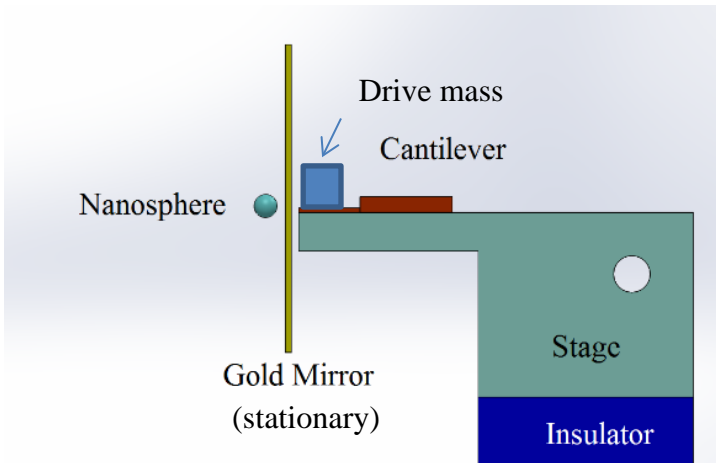
# Drive Mass fabrication

Buried drive mass technique – eliminates corrugation

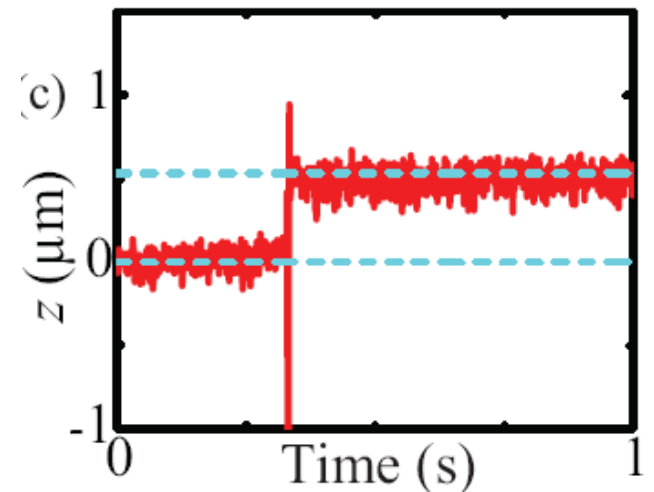
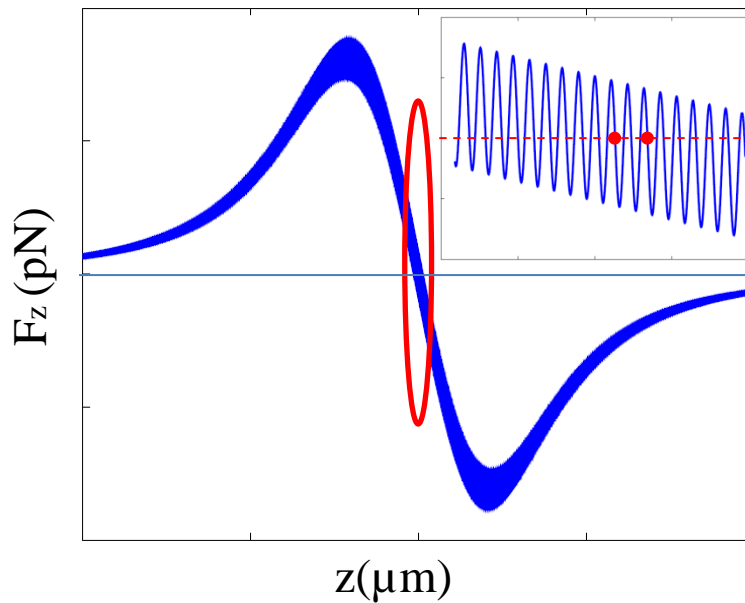
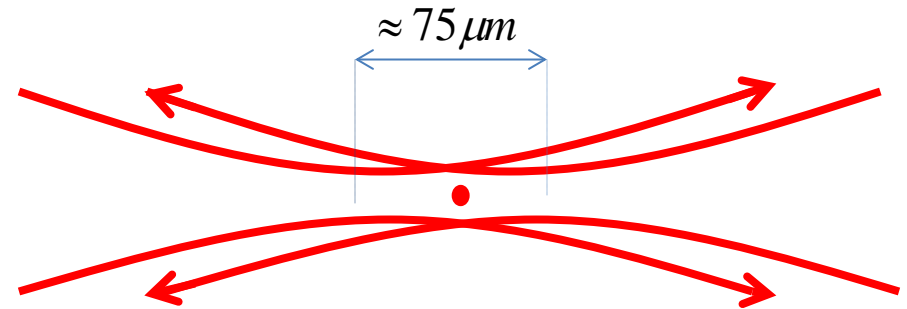
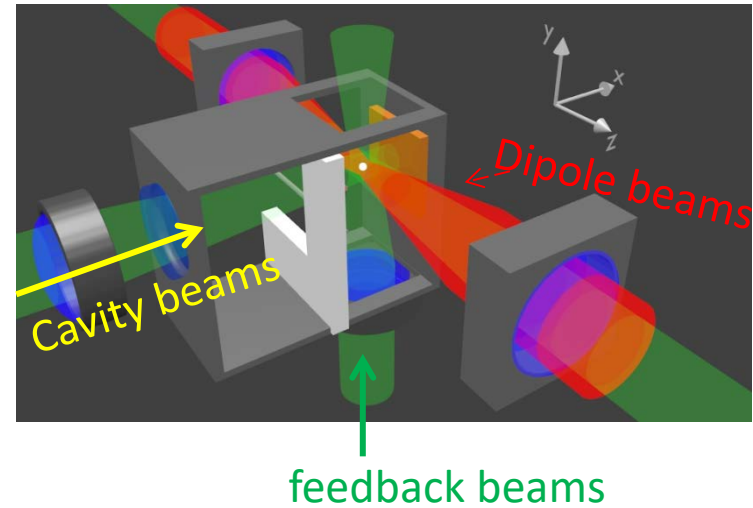


# MEMS actuator

- Device for positioning drive mass

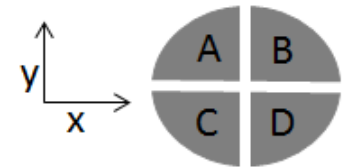
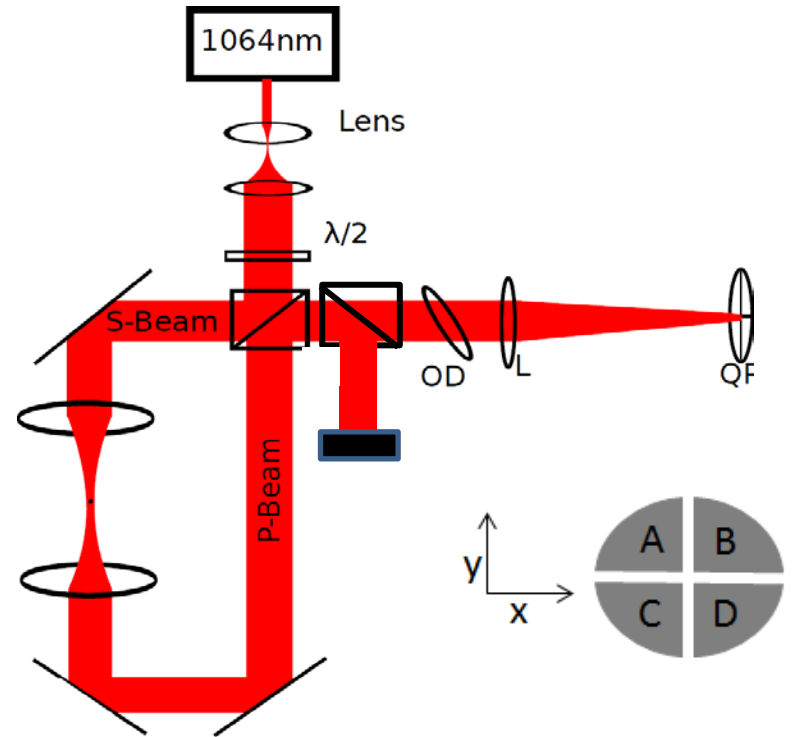
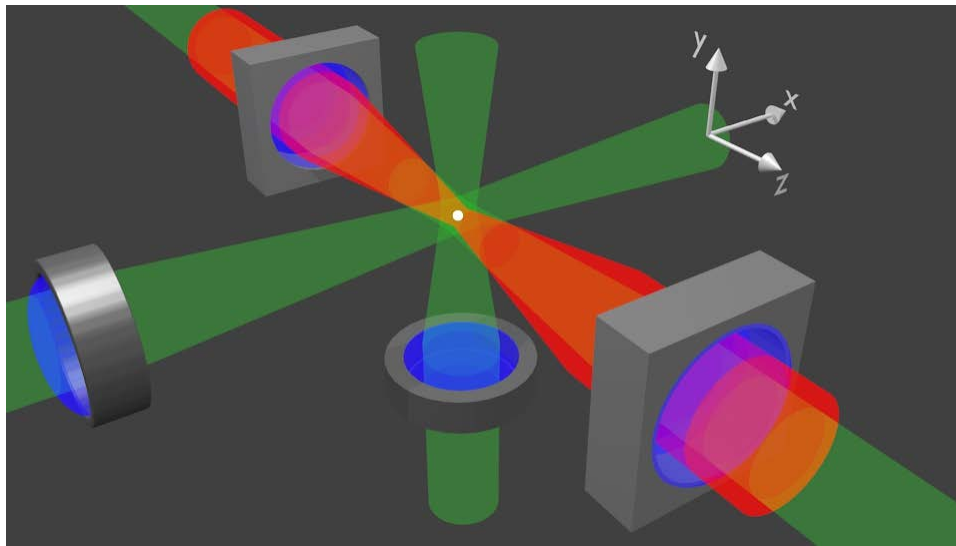


# Standing wave optical trap



# 3D feedback cooling of a nanosphere

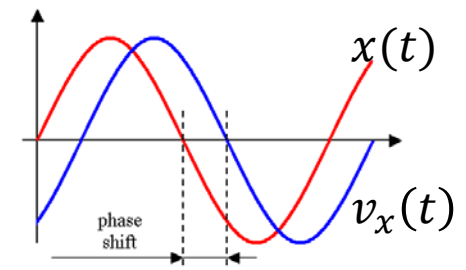
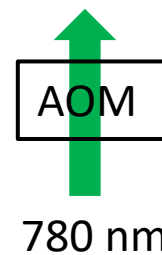
Needed to stabilize the particle, damp and cool it  
Mitigate photon recoil heating



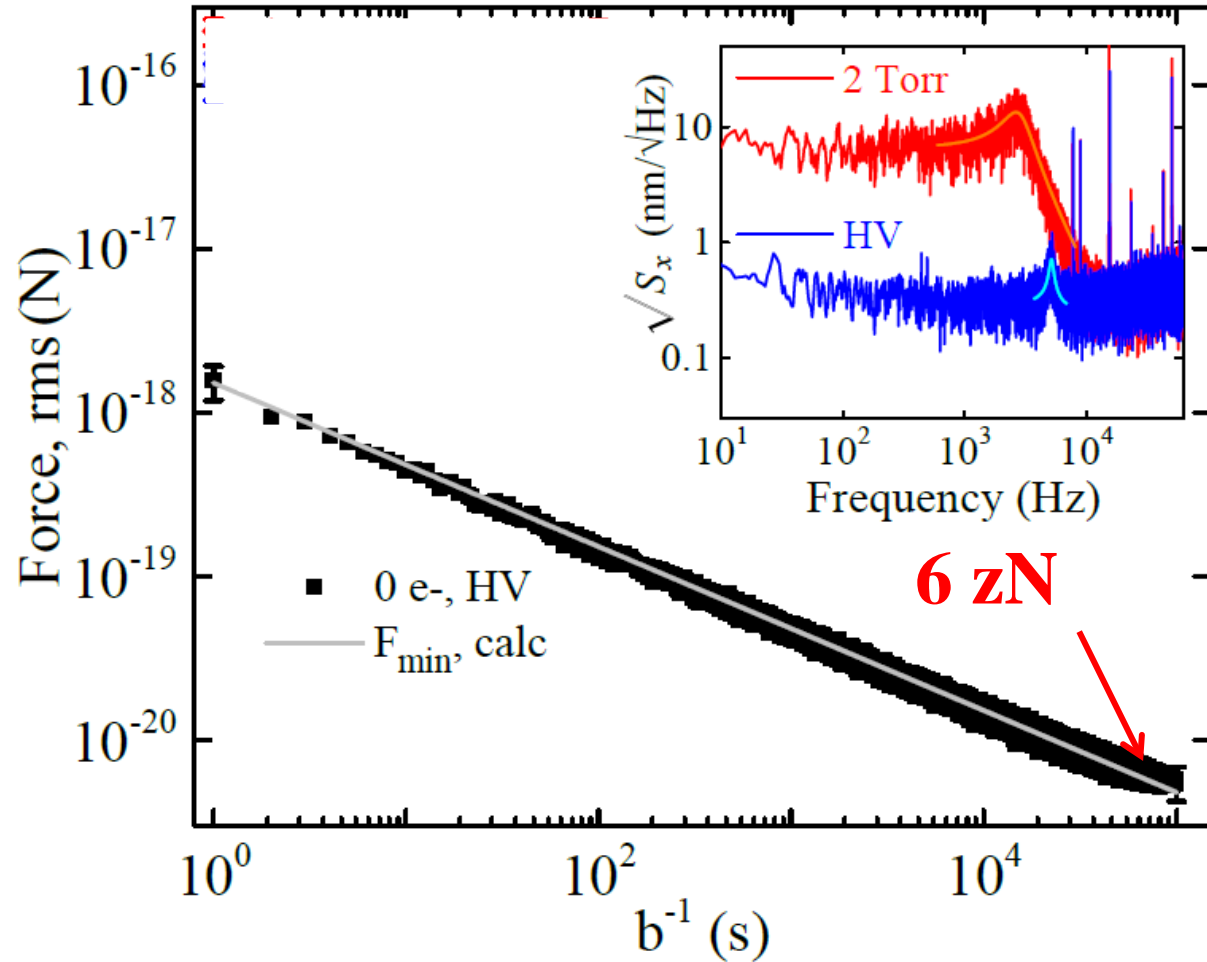
$$F_{\min} = \sqrt{\frac{4kK_B T B}{\omega_0 Q}}$$

$$Q_{\text{eff}} = \frac{Q_0 \Gamma_0}{\Gamma_0 + \Gamma_{\text{cool}}}$$

$$T_{\text{eff}} = \frac{T_0 \Gamma_0}{\Gamma_0 + \Gamma_{\text{cool}}}$$



# Zeptonewton force sensing



**Sensitivity**

$$S_{F,x} = 1.63 \pm .37 \text{ aN} / \sqrt{\text{Hz}}$$



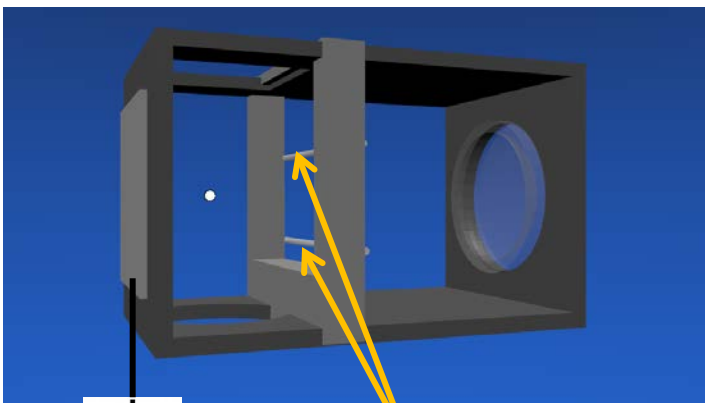
# Zeptonewton force sensing

## Electrostatic Calibration

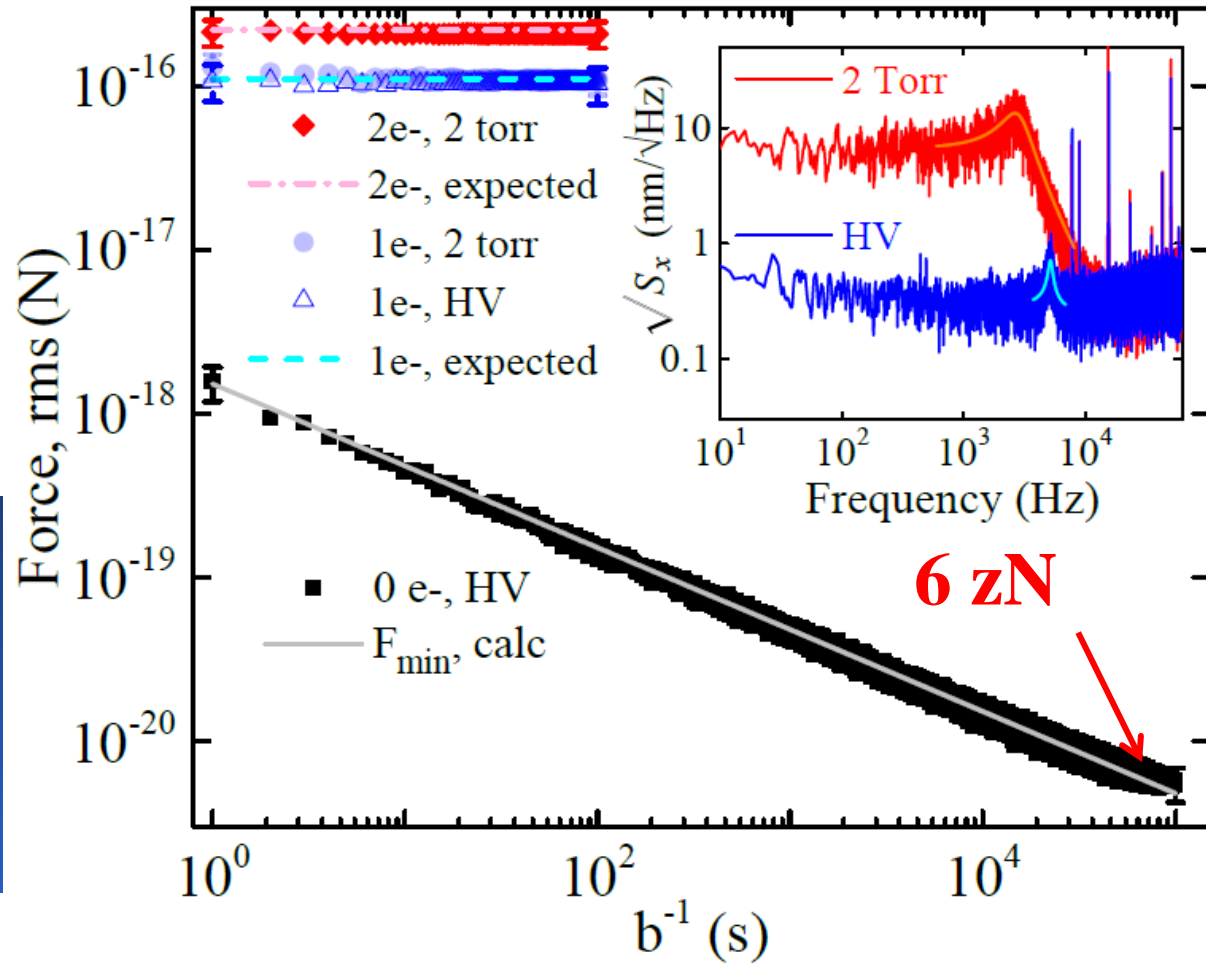
90% of beads are neutral

Neutral beads stay neutral

Charge stays constant over days



Calibration electrodes

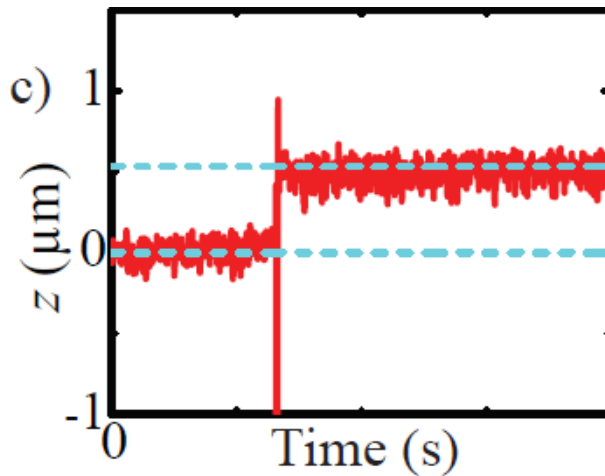


**Sensitivity**

$$S_{F,x} = 1.63 \pm .37 \text{ aN} / \sqrt{\text{Hz}}$$

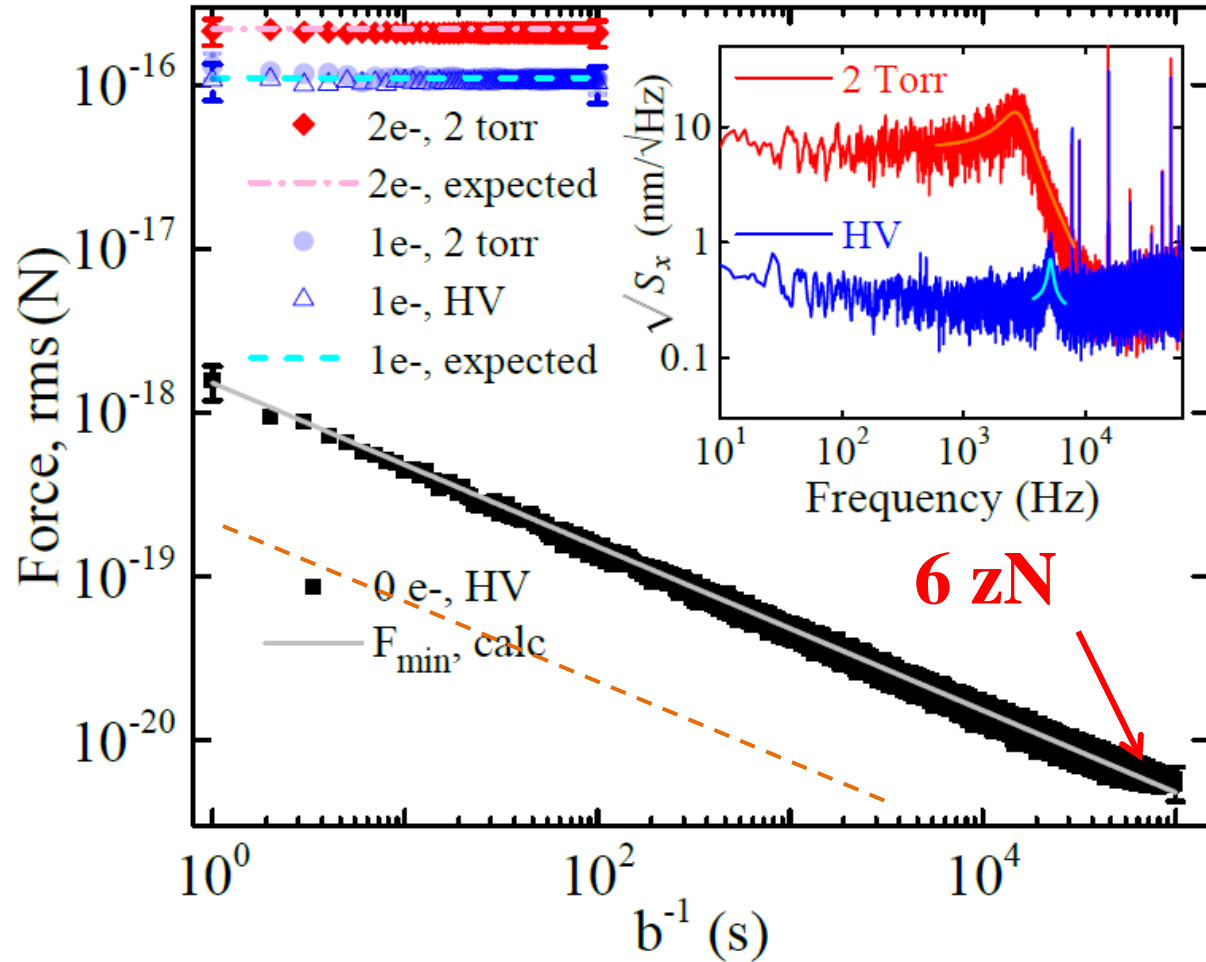
# Zeptonewton force sensing

## Optical lattice calibration



Useful for neutral objects

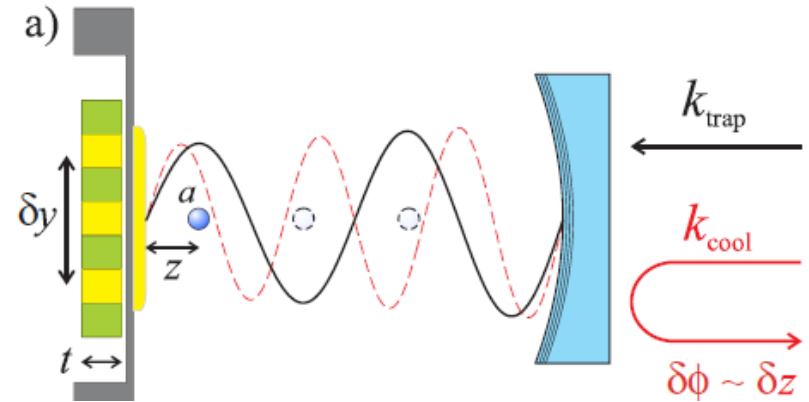
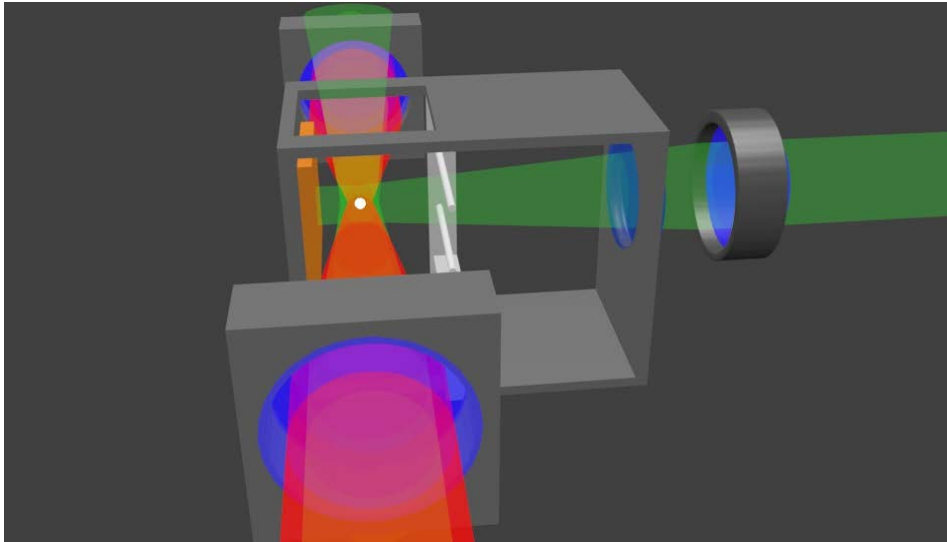
Method consistent with electric field approach



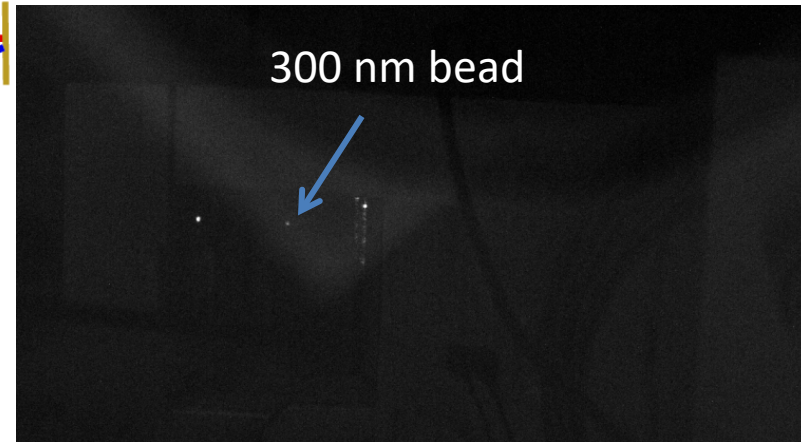
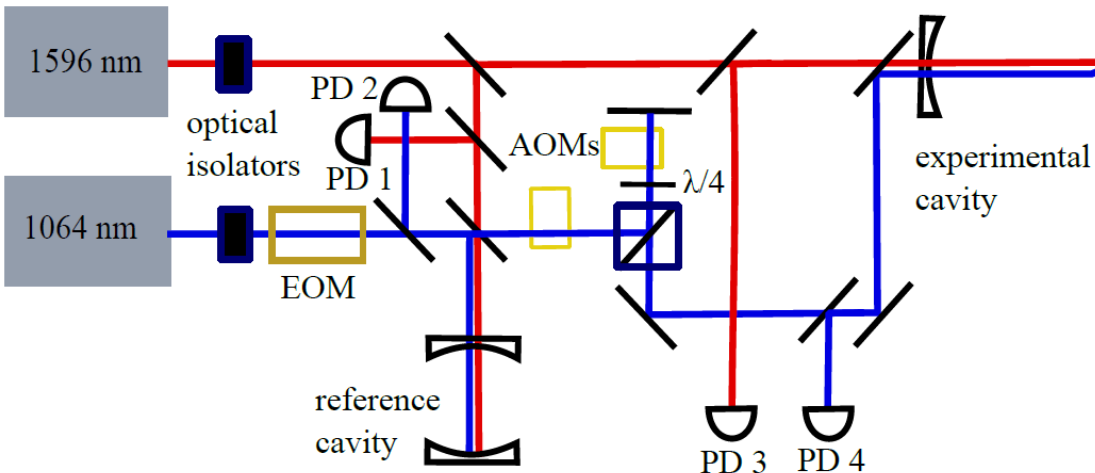
Sensitivity

$$S_{F,x} = 1.63 \pm .37 \text{ aN} / \sqrt{\text{Hz}}$$

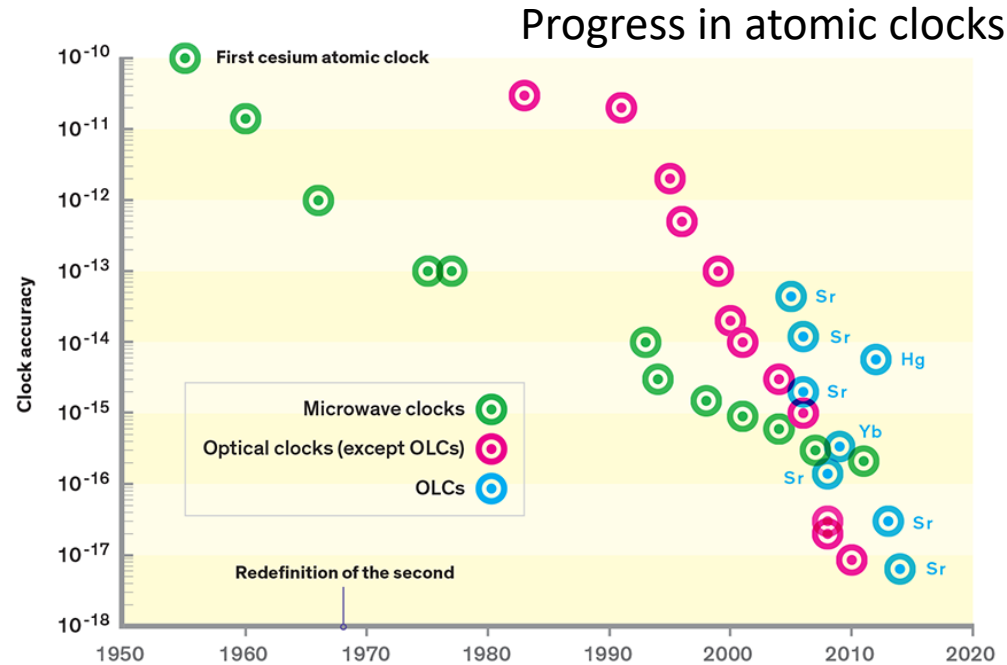
# Next: Cavity Trapping and cooling



1596nm beam to trap a bead at its antinode  $\rightarrow$  localization  
 1064nm beam to cavity cool the CM of bead  $\rightarrow$  position readout



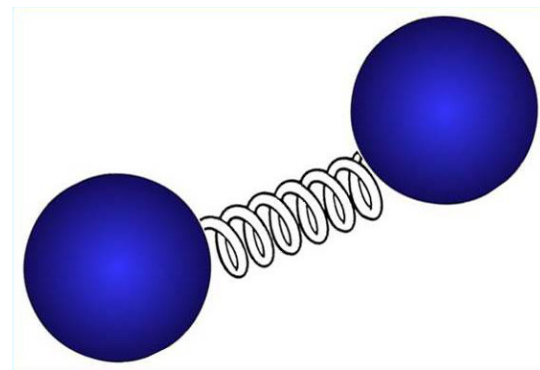
# AMO- based techniques



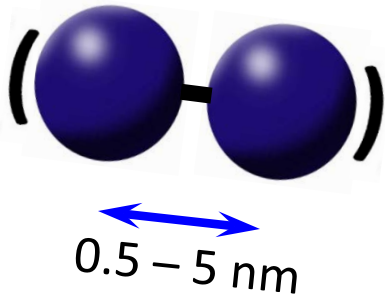
A molecular quantum clock, a device that precisely measures the vibrations of this spring,

- Time variation of electron-to-proton mass ratio
- Test if unknown forces - besides electromagnetism and Newtonian gravity - modify the vibrations of the spring at the nanometer length scale.

(Sr Molecular clock)

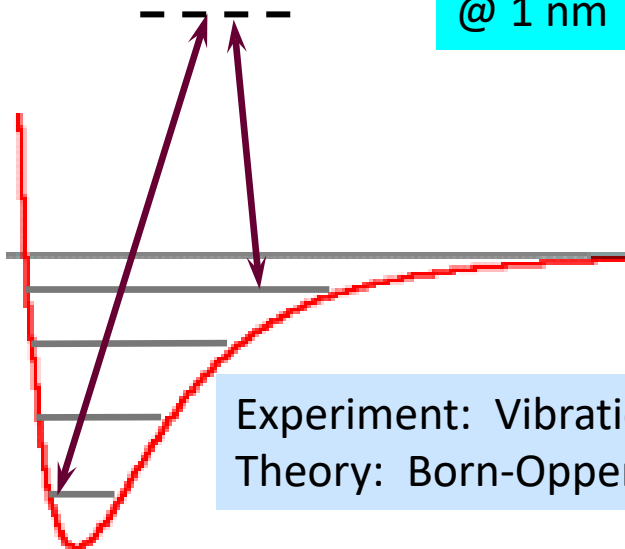


# Nanometer-Scale Mass-Dependent Forces with Ultracold Molecules

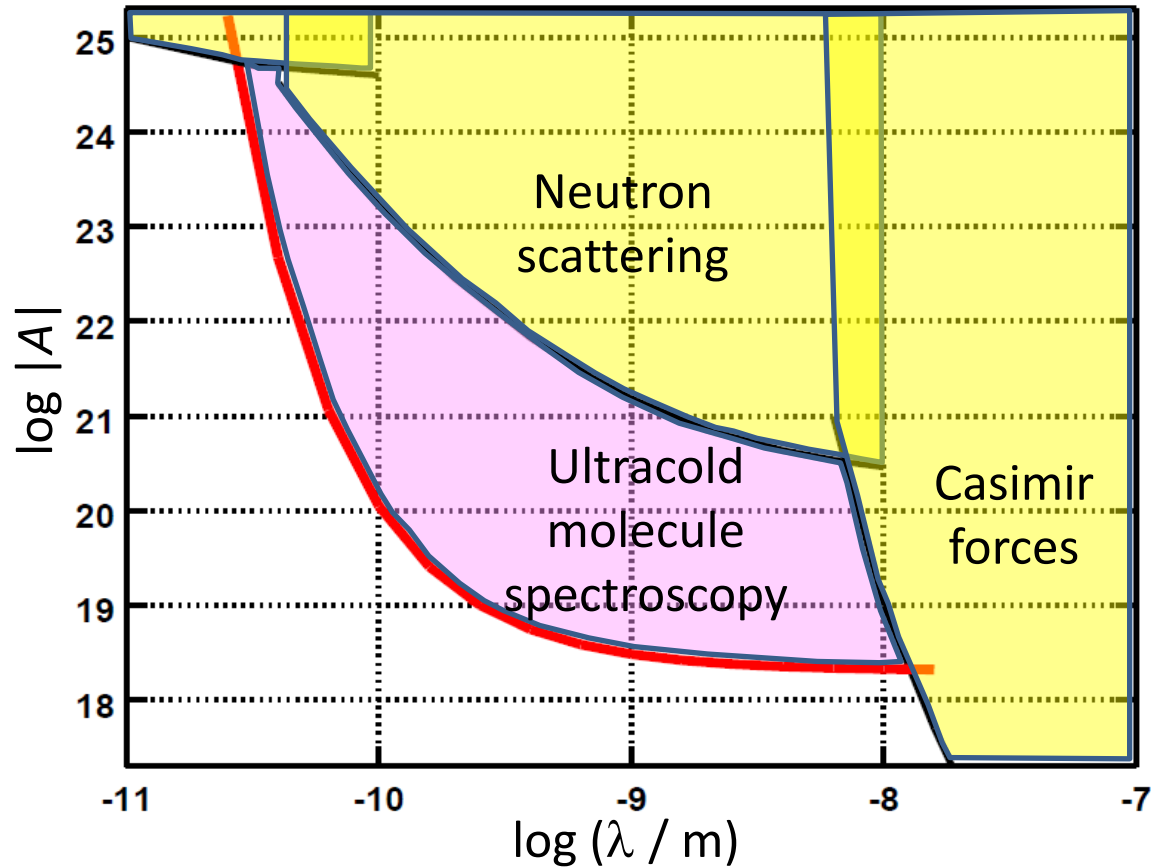


$$V = -\frac{GM^2}{r} (1 + Ae^{-r/\lambda})$$

$A < 10^{21}$   
@ 1 nm !



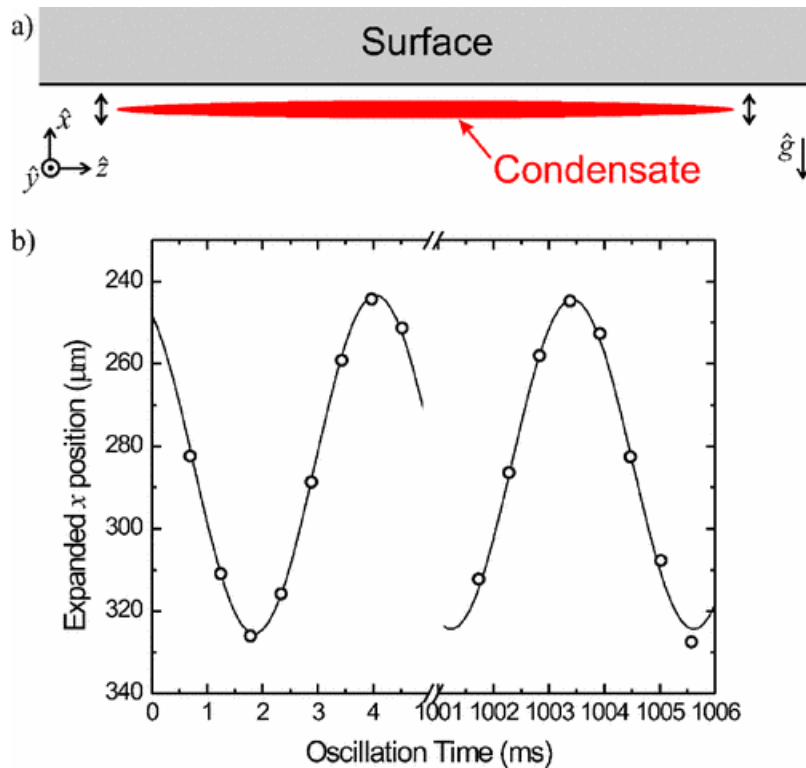
Experiment: Vibrational spectroscopy of molecular isotopes  
Theory: Born-Oppenheimer and mass-dependent corrections



Slide: T. Zelevinsky

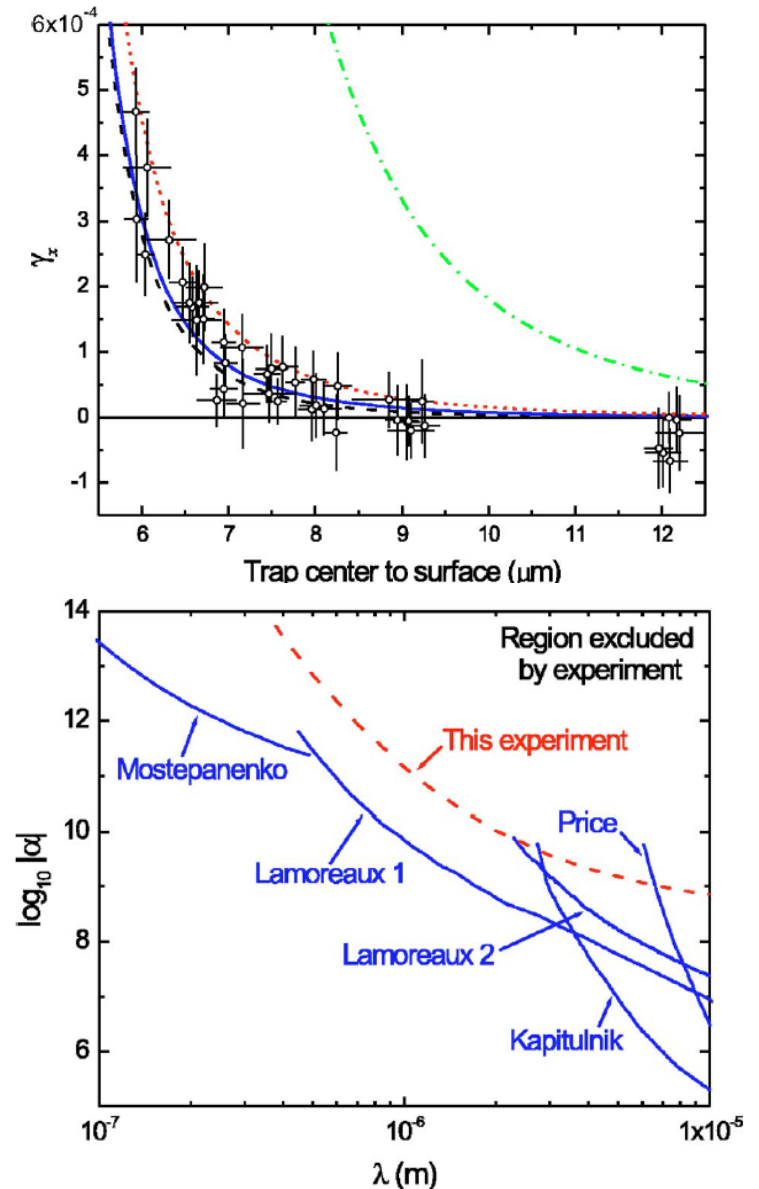
# Atomic Bose-Einstein condensates

## Casimir-Polder force measurements

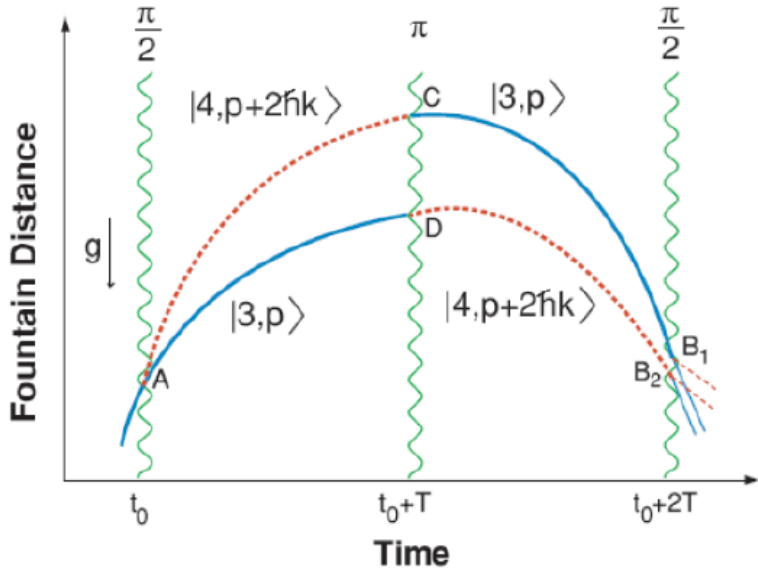


D. M. Harber, J. M. Obrecht, J. M. McGuirk, and E. A. Cornell, PRA **72**, 033610 2005

Cornell group, JILA

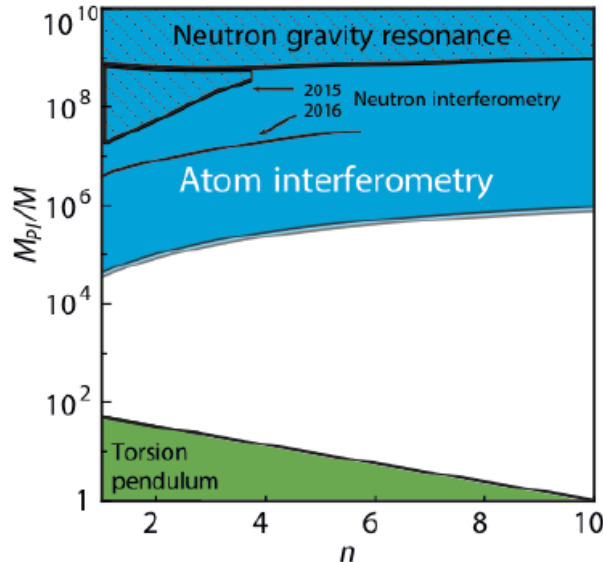
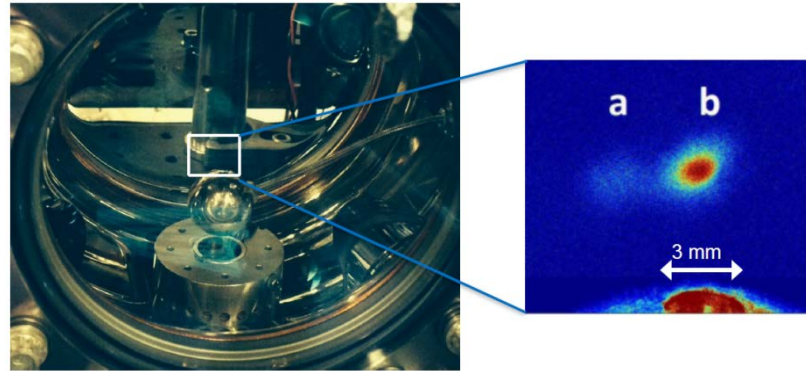


# Atom interferometers



*Probing Dark Energy with Atom Interferometry*  
 C. Burrage, E. J. Copeland, E. A. Hinds  
 JCAP 1503 (2015) 03, 042

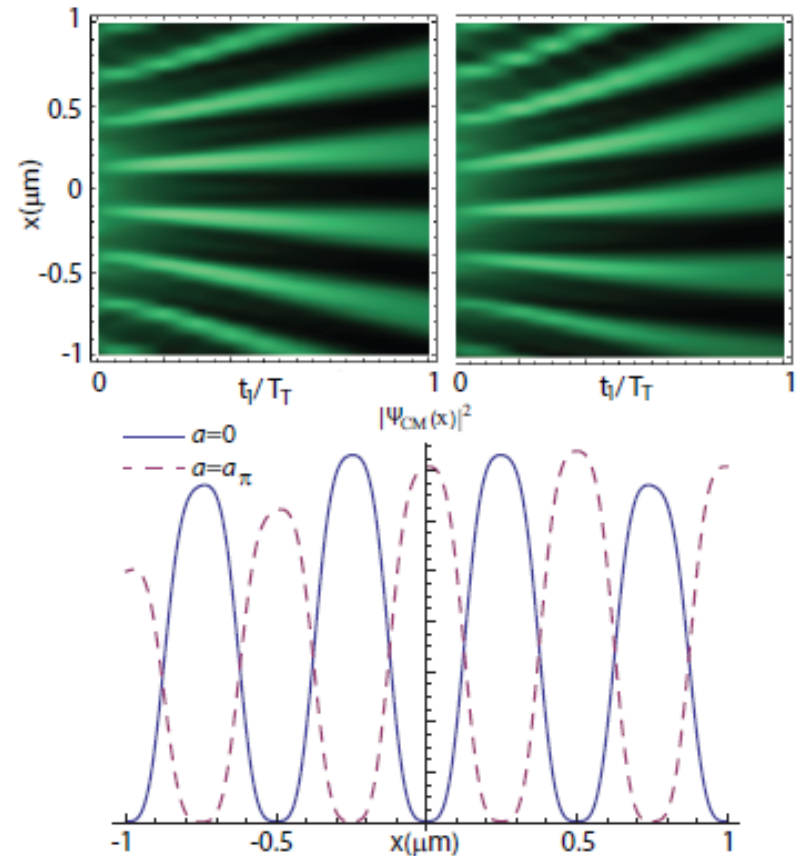
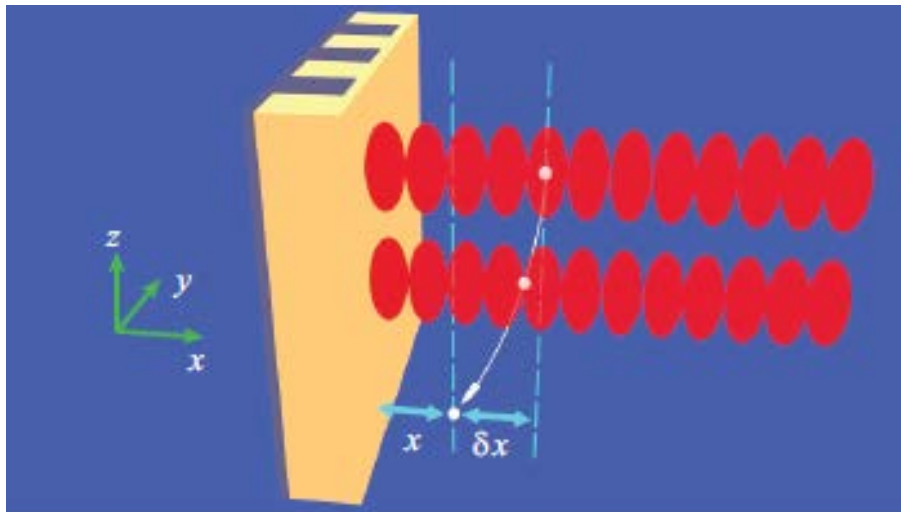
**Search for Chameleons (screened within matter, only shell contributes, except for atoms!)**



P. Hamilton, M. Jaffe, P. Haslinger,  
 Q. Simmons, H. Müller, J. Khoury  
 Science 349, 849 (2015).

Benjamin Elder, Justin Khoury, Philipp Haslinger, Matt Jaffe,  
 Holger Müller, Paul Hamilton, [PRD 94, 044051](#) (2016).

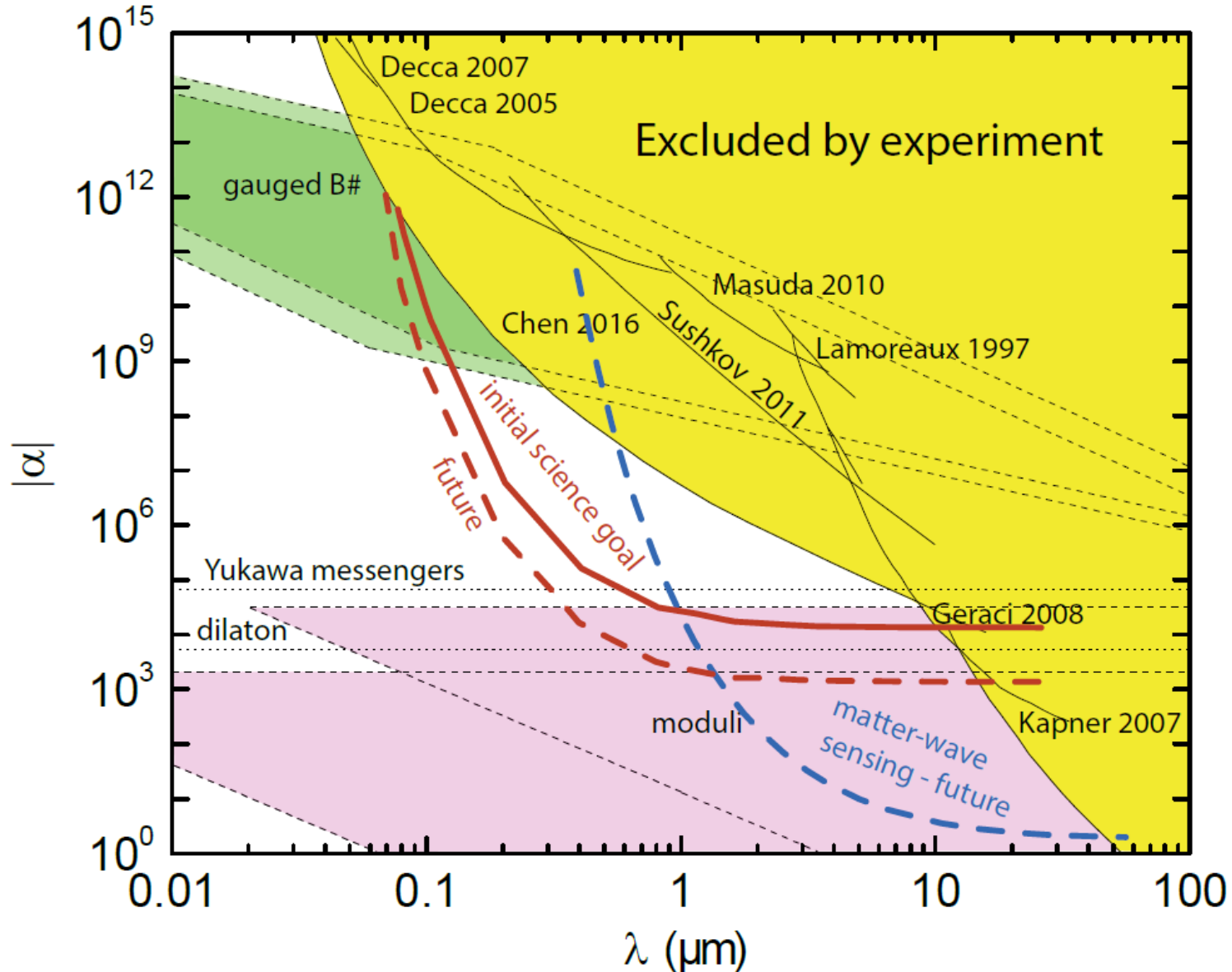
# Matter-wave interferometry



ng acceleration  
sensing



# Projected reach- nanosphere matter-wave interferometer



# Summary

- Rich possibilities for new science
  - ✓ Tests of gravity
  - ✓ Dark Sector Physics
  - ✓ Beyond the Standard Model
- Diverse set of techniques from researchers of varying backgrounds and expertise
  - Atomic, Molecular, Optical Physics
  - Condensed matter and low-temperature physics
  - Torsion balances/precision mechanical measurements