BEYOND THE STANDARD MODEL @ THE TEV SCALE

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THE HIERARCHY PROBLEM



ADDING A SYMMETRY

... AND BREAKING IT SOFTLY

WE ASSUMED THE SYMMETRY PROTECTING THE WEAK SCALE WAS CONTINUOUS. ARE THEIR OTHER OPTIONS?

 \Rightarrow "NEUTRAL NATURALNESS"

DISCRETE SYMMETRIES

DISCRETE SYMMETRY

≲4**π/**G

HIGGS MH

SYMMETRY-BASED APPROACHES TO HIERARCHY PROBLEM EMPLOY CONTINUOUS SYMMETRIES.

DISCRETE SYMMETRY NEUTRAL PARTNERS **က** LEADS TO PARTNER STATES W/ SM QUANTUM NUMBERS.

DISCRETE SYMMETRIES CAN ALSO SERVE TO PROTECT THE HIGGS.

LEADS TO PARTNER STATES W/ NON-SM QUANTUM NUMBERS.

"NEUTRAL NATURALNESS"

THE TWIN HIGGS

CONSIDER A SCALAR H TRANSFORMING AS A FUNDAMENTAL UNDER A GLOBAL SU(4) SYMMETRY:

$$V(H) = -m^2 |H|^2 + \lambda |H|^4$$

POTENTIAL LEADS TO SPONTANEOUS SYMMETRY BREAKING,

$$|\langle H \rangle|^2 = \frac{m^2}{2\lambda} \equiv f^2$$

YIELDS SEVEN GOLDSTONE BOSONS.

 $\simeq SU(4) \rightarrow SU(3)$

THE TWIN HIGGS

NOW GAUGE SU(2)_A X SU(2)_B \subset SU(4), W/ $H = \begin{pmatrix} H_A \\ H_B \end{pmatrix}$ \uparrow \uparrow \uparrow US TWINS

THEN 6 GOLDSTONES ARE EATEN, LEAVING ONE BEHIND.

EXPLICITLY BREAKS THE SU(4); EXPECT RADIATIVE CORRECTIONS.

$$V(\mathcal{H}) \xrightarrow{9}_{64\pi64\pi^2} \frac{9}{64\pi64\pi^2} \mathcal{H}(\mathcal{H}^2_A \#^2_B \mathcal{H}^2_B \|\mathcal{H}^2_B\|^2_B |\mathcal{H}^2_B\|^2_B \|\mathcal{H}^2_B\|^2_B)^2$$

BUT THESE BECOME SU(4) SYMMETRIC IF $g_A=g_B$ FROM A Z_2 QUADRATIC POTENTIAL HAS ACCIDENTAL SU(4) SYMMETRY.

THE TWIN HIGGS

FULL THEORY: EXTEND Z_2 TO ALL SM MATTER AND COUPLINGS.

 $SM_A\,X\,\,SM_B\,X\,\,Z_2$

??

 SM_B

 $(h_B, t_B, W_B, Z_B...)$

 $(h_A, t_A, W_A, Z_A...)$

 $V(H) \supset \frac{\Lambda^2}{16\pi^2} \left(-6y_t^2 + \frac{9}{4}g^2 + \dots \right) \left(|H_A|^2 + |H_B|^2 \right)$

$$\langle H_A \rangle |^2 + |\langle H_B \rangle|^2 = f^2$$

BREAKS "QUADRATIC" SU(4), HIGGSES EWKA & EWKB

GIVES A RADIAL MODE, A GOLDSTONE MODE, AND EATEN GOLDSTONES.

V « F FOR SM-LIKE HIGGS TO BE THE GOLDSTONE

PRIMARY COUPLING BETWEEN SMA AND SMB IS VIA HIGGS PORTAL

TWIN HIGGS & THE HIERARCHY PROBLEM



THE TOP PARTNER ACTS AS EXPECTED FROM GLOBAL SYMMETRY PROTECTION, BUT IS NOT CHARGED UNDER QCD.

NO DIRECT LIMIT ON TOP PARTNER.



"NEUTRAL" NATURALNESS



SIMPLEST THEORY: EXACT MIRROR COPY OF SM

[CHACKO, GOH, HARNIK '05]

BUT THIS IS MORE THAN YOU NEED, AND MIRROR 1ST, 2ND GENS LEAD TO COSMOLOGICAL PROBLEMS

MANY MORE OPTIONS WHERE SYMMETRY IS APPROXIMATE, E.G. A GOOD SYMMETRY FOR HEAVIEST SM PARTICLES.

[NC, KNAPEN, LONGHI '14; GELLER, TELEM '14; NC, KATZ, STRASSLER, SUNDRUM '15; BARBIERI, GRECO, RATTAZZI, WULZER '15; LOW, TESI, WANG '15, NC, KNAPEN, LONGHI, STRASSLER '16]



FINDING A MIRROR



HIGGS STILL A PNGB, TUNING AS IN OTHER GLOBAL SYMMETRIES

LIMIT $v^2/f^2 < 0.1$ $\rightarrow \Delta^{\sim}10 (10\% \text{ TUNING})$ UNLIKELY TO IMPROVE MUCH IN RUN 2

- PARTNER STATES ARE SM NEUTRAL, COUPLE ONLY TO THE HIGGS. LIGHTER THAN MH/2: MODEST INVISIBLE HIGGS DECAYS.
- HEAVIER THAN M_H/2: PRODUCE THROUGH AN OFF-SHELL HIGGS.

HARD BUT VERY INTERESTING; DIRECTLY PROBE NATURALNESS



[NC, LOU, MCCULLOUGH, THALAPILLIL '14]

SM

SM

h*

h*

0++

EXOTIC HIGGS DECAYS

h

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- TWIN SECTOR MUST HAVE TWIN QCD, CONFINES AROUND QCD SCALE
- HIGGS BOSON COUPLES TO BOUND STATES OF TWIN QCD
- VARIOUS POSSIBILITIES. GLUEBALLS MOST INTERESTING; HAVE SAME QUANTUM # AS HIGGS



$$\mathcal{L} \supset -\frac{\alpha'_3}{6\pi} \frac{v}{f} \frac{h}{f} G_{\mu\nu}^{'a} G_a^{'\mu\nu}$$

PRODUCE IN RARE HIGGS DECAYS (BR~10⁻³-10⁻⁴)

$$gg \to h \to 0^{++} + 0^{++} + \dots$$

DECAY BACK TO SM VIA HIGGS

$$0^{++} \to h^* \to f\bar{f}$$

LONG-LIVED, DECAY LENGTH IS MACROSCOPIC; LENGTH SCALE ~ LHC DETECTORS

SEARCHING FOR MIRRORS



- SIGNAL: DISPLACED DECAYS OF SM HIGGS WITH BR >10⁻³ (σ .BR~20FB @ RUN 1).
- ATLAS: HCAL/ECAL & MUON CHAMBER SEARCHES POWERFUL, SENSITIVE TO DISPLACED HIGGS DECAY.
- CMS: USE INNER TRACKER, SEE VERTEX ON SHORT DECAY LENGTHS. TRIGGER THRESHOLDS TOO HIGH.
- MORE ROOM FOR INNOVATION IN THE
 DISPLACED DECAY SEARCH PROGRAM...



SELECTING A VACUUM

WE ASSUMED THAT WE ENDED UP IN THE VACUUM WITH THE OBSERVED WEAK SCALE DUE TO SOME ANTHROPIC PRESSURE. CAN WE INSTEAD DO SO DYNAMICALLY?

 \Rightarrow "RELAXION"

()

DYNAMICAL SELECTION

WHAT IF THE WEAK SCALE IS SELECTED BY DYNAMICS, NOT SYMMETRIES?

OLD IDEA: COUPLE HIGGS TO FIELD WHOSE MINIMUM SETS $M_H=0$ OLD PROBLEM: HOW TO MAKE $M_H=0$ A SPECIAL POINT OF POTENTIAL?



BUT: IMMENSE ENERGY STORED IN EVOLVING FIELD, NEED DISSIPATION.

[GRAHAM, KAPLAN, RAJENDRAN '15]

 $H_i < \Lambda_{QCD}$

EX.1: QCD/QCD' RELAXION

FIRST THOUGHT: USE AN AXION COUPLED TO QCD.



 $V(\phi)$

$$(-M^{2} + g\phi)|H|^{2} + V(g\phi) + \frac{1}{32\pi^{2}}\frac{\phi}{f}\tilde{G}^{\mu\nu}G_{\mu\nu}$$
$$\Rightarrow (-M^{2} + g\phi)|H|^{2} + V(g\phi) + \Lambda^{4}\cos(\phi/f)$$

DISSIPATION: INFLATION!
$$\ddot{\phi} + 3H\dot{\phi} + V'(\phi) = 0$$

REQUIREMENTS: (1) ϕ SCAN OVER ENTIRETY OF ITS RANGE

$$\Delta \phi = (gM^2/H_i^2)N \gtrsim M^2/g \Rightarrow N \gtrsim H_i^2/g^2$$

(2) VACUUM ENERGY DURING INFLATION EXCEEDS CHANGE IN VACUUM ENERGY DUE TO SCANNING

$$H_i > \frac{M^2}{M_{Pl}}$$

(3) BARRIERS FORM THAT ARE SUFFICIENT TO STOP SCANNING

QCD RELAXION

(4) CLASSICAL ROLLING BEATS QUANTUM FLUCTUATIONS

$$H_i < V'_{\phi}/H_i^2 \to H_i < (gM^2)^{1/3}$$

ADDITIONAL SUBSTANTIAL CONCERNS:

NON-COMPACT SHIFT SYMMETRY?
 COSMOLOGICAL CONSTANT?

JUST NEED HIGGS + NON-COMPACT AXION + INFLATION W/

• VERY LOW HUBBLE SCALE ($\ll \Lambda_{QCD}$)

10 GIGA-YEARS OF INFLATION

CARE REQUIRED TO AVOID TRANSFERRING FINE-TUNING TO INFLATIONARY SECTOR.



IN VACUUM, ϕ IS THE AXION, STOPS WELL AWAY FROM $\theta = 0 \rightarrow \text{GIVES}$ O(1) CONTRIBUTION TO θ_{QCD}



[GRAHAM, KAPLAN, RAJENDRAN '15]

QCD' RELAXION

FIX: MAKE IT SOMEONE ELSE'S QCD + AXION



1. NEW QUARKS MUST GET MOST OF MASS FROM HIGGS:

$$\mathcal{L} \supset m_L L L^c + m_N N N^c + y H L N^c + y' H^{\dagger} L^c N$$

2. MUST CONFINE, BUT WITH LIGHT FLAVOR $\Lambda^4 \simeq 4\pi f_{\pi'}^3 m_N$

DECOUPLE FROM TEV SCALE?

QCD' RELAXION

NOW $m_N \geq y y' v^2/m_L$ (SMALLEST SEE-SAW MASS FROM EWSB IF L HEAVY)

$$\text{BUT ALSO} \quad \left\{ \begin{array}{ll} m_N \geq \frac{yy'}{16\pi^2} m_L \log(M/m_L) & \text{(RADIATIVE DIRAC MASS)} \\ m_N \geq yy' f_{\pi'}^2/m_L & \text{(HIGGS WIGGLES BIGGEST)} \end{array} \right.$$

THESE BOUNDS IMPLY
$$f_{\pi'} < v$$
 and $m_L < \frac{4\pi v}{\sqrt{\log(M/m_L)}}$

NEW CONFINING PHYSICS NEAR WEAK SCALE!

COUPLES TO HIGGS, ELECTROWEAK BOSONS; HIDDEN VALLEY SIGNATURES. VARIOUS POSSIBILITIES ($N_F=1$, PIONS NOT LIGHT)

TO MY KNOWLEDGE, NO SYSTEMATIC STUDY TO DATE.

[HOOK, MARQUES-TAVARES '16]

EX. 2: INTERACTIVE RELAXION

ALTERNATIVE POSSIBILITY: KEEP BUMPS ACROSS ENTIRE POTENTIAL, TURN ON DISSIPATION AT A SPECIAL POINT OF POTENTIAL.

ANOTHER SOURCE OF DISSIPATION: PARTICLE PRODUCTION

 $\mathcal{L} \supset -\frac{\phi}{4f} F\tilde{F}$ CONSIDER AXION-LIKE COUPLINGS TO MASSIVE GAUGE FIELD: $\ddot{A}_{\pm} + \left(k^2 + m_A^2 \pm \frac{k\dot{\phi}}{f}\right)A_{\pm} = 0$ E.O.M. FOR TRANSVERSE **POLARIZATIONS:** $\omega_{\pm}^2 = k^2 + m_A^2 \pm \frac{k\phi}{f}$ FOR $\dot{\phi} \approx \text{constant}$ $A_{\pm}(k) \propto e^{i\omega_{\pm}t}$ $\omega_{+}^{2} < 0 \Rightarrow |\dot{\phi}| \gtrsim 2 f m_{A}$ EXPONENTIALLY GROWING SOLUTION FOR GROWING MODE DRAINS ENERGY FROM ϕ

EX. 2: INTERACTIVE RELAXION

APPLY TO RELAXION: USE ELECTROWEAK GAUGE FIELDS



IMPORTANT SUBTLETY: CAN'T COUPLE TO PAIRS OF PHOTONS! FOR DISSIPATION TO BECOME EFFICIENT AT H[~]V, CAN ONLY COUPLE TO BOSONS ACQUIRING MASS FROM EWSB.

(NOT A TUNING, CAN BE MADE NATURAL WITH SYMMETRIES, E.G., SU(2)_L X SU(2)_R) $\mathcal{L} \supset -\frac{1}{4g_L^2}W_L^2 - \frac{1}{4g_R^2}W_R^2 + \frac{\phi}{f}(W_L\tilde{W}_L - W_R\tilde{W}_R)$ $\phi \rightarrow \phi + \alpha \quad \theta_L \rightarrow \theta_L - \alpha \quad \theta_R \rightarrow \theta_R + \alpha \qquad \Rightarrow \mathcal{L} \propto (\theta_L + \theta_R)F\tilde{F}$

LOWERING THE CUTOFF

USUALLY ASSUME LOW CUTOFF IS DUE TO E.G. GEOMETRY OF AN EXTRA DIMENSION, GIVING UNIFORM PREDICTION FOR NEW RESONANCES & STRONG LIMITS. CAN WE DO THE SAME THING WITH ORDER INSTEAD OF DISORDER?



 \Rightarrow "GRAVITATIONAL ANDERSON LOCALIZATION"

A RANDOM DETOUR

ANDERSON LOCALIZATION



SIMPLIFY:

$$t_{ij} = t \left(\delta_{i+1}^{j} + \delta_{i-1}^{j} \right)$$

$$\textit{NEAREST-NEIGHBOR HOPPING}$$

$$\epsilon_i \in \left[-W/2, W/2 \right]$$

$$\textit{RANDOM IMPURITIES}$$

$$H = \sum_{i} \epsilon_{i} |i\rangle \langle i| - \sum_{ij} t_{ij} |i\rangle \langle j| + \text{h.c.}$$

$$\text{TUNNELING}$$

$$\text{BOUND STATE ENERGIES}$$

$$H = \begin{pmatrix} \epsilon_{1} & -t & 0 & \dots & 0 & 0 \\ -t & \epsilon_{2} & -t & \dots & 0 & 0 \\ 0 & -t & \epsilon_{3} & & & \\ \vdots & \vdots & \ddots & & \\ 0 & 0 & & & \epsilon_{N-1} & -t \\ 0 & 0 & & & -t & \epsilon_{N} \end{pmatrix}$$

$$H = \begin{pmatrix} \epsilon_{1} & -t & 0 & \dots & 0 & 0 \\ -t & \epsilon_{2} & -t & \dots & 0 & 0 \\ 0 & -t & \epsilon_{3} & & & \\ \vdots & \vdots & \ddots & & \\ 0 & 0 & & & \epsilon_{N-1} & -t \\ 0 & 0 & & & -t & \epsilon_{N} \end{pmatrix}$$

ANDERSON LOCALIZATION

S.E. FOR ENERGY EIGENSTATES

$$\psi_E = \sum_i \psi_i |i\rangle \quad \underset{(\ddagger-1)}{\text{GIVES}} \quad \psi_{i+1} + \psi_{i-1} = (E - \epsilon_i)\psi_i$$

ALL EIGENSTATES ARE **LOCALIZED** IN PRESENCE OF DISORDER, $\psi(r) \propto EXP[-r/L_{loc}]$ BUT LOCALIZATION LENGTHS NOT IDENTICAL

ANALYTIC RESULTS FOR WEAK LOCALIZATION, $\sigma \ll 1$ (FOR $\epsilon_{l} \in [-W/2, W/2]$, $\sigma^{2} = W^{2}/12$)



ANOMALOUS SCALING NEAR BAND EDGES E = ±2

$$L_{loc}^{-1} = \frac{6^{1/3}\sqrt{\pi}}{2\Gamma(1/6)}\sigma^{2/3} \approx 0.13W^{2/3}$$

STATES AT BAND EDGES MORE SHARPLY LOCALIZED THAN GENERIC EIGENSTATES AT WEAK DISORDER

THE HIERARCHY PROBLEM?

WHAT YOU'VE BEEN ASKING YOURSELF FOR THE LAST FEW MINUTES ...

HOW DOES RS SOLVE HIERARCHY PROBLEM? CURVATURE LOCALIZES THE GRAVITON ZERO MODE.

→ FIELDS LOCALIZED AT DIFFERENT POINTS IN 5TH DIMENSION SEE DIFFERENT FUNDAMENTAL SCALES



EXAMPLE 12: CAN ACHIEVE THE SAME OUTCOME IN A FLAT FIFTH DIMENSION BY LOCALIZING GRAVITON W/ DISORDER

$$S = -\int d^5x \sqrt{G} (M^3_{\star} \mathcal{R}) + \sum_{\langle ij \rangle} M^4_{\star} V(|X_i - X_j|) - \sum_i \int d^4x \sqrt{g} f_i$$

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IN THIS CASE DISORDER = RANDOMLY SPACED & TENSIONED BRANES

THE CHALLENGE: NAIVE TIGHT-BINDING MODEL DOES NOT REFLECT DIFFEOMORPHISM INVARIANCE



ON ONE HAND, SPECULATIVE INDICATIONS OF BSM ARE ON, WELL, SPECULATIVE FOOTING!

ON THE OTHER HAND, THEY POINT TO DEEP & PROFOUND (RATHER THAN PIECEMEAL) CHANGES TO THE STRUCTURE OF THE SM, WHICH PERHAPS EXPLAINS THEIR APPEAL.

SUCCESSFUL ANSWERS TO THESE SPECULATIVE PROBLEMS OFTEN ALSO FULFILL OTHER INDICATIONS OF BSM PHYSICS (E.G. DARK MATTER, UNIFICATION, & BARYOGENESIS)

CURRENT ERA IS A TIME OF OPPORTUNITY - POPULAR PARADIGMS UNDER STRESS, ROOM FOR INNOVATION.

PART 3: EVERYTHING ELSE

BEYOND THE STANDARD MODEL

TEV SCALE





$\Delta_{O} = 4$ natural ~ $\mathcal{O}(1)$ **STRONG CP PROBLEM**

NO SYMMETRY WHEN O. BUT RADIATIVE CORRECTIONS SMALL/PROPORTIONAL TO VALUE

> IN ADDITION TO GAUGE KINETIC TERMS + MATTER COUPLINGS, QCD ADMITS GENERICALLY O(1) PARITY-ODD COUPLING*

> > $\theta_{QCD} \epsilon^{\mu\nu\alpha\beta} G^a_{\mu\nu} G^a_{\alpha\beta}$

FOLLOWING IT THROUGH THE CHIRAL LAGRANGIAN, LEADS TO COUPLING BETWEEN NEUTRONS AND PHOTONS OF FORM

 $\mathcal{L} = -\frac{id_n}{8} \epsilon_{\mu\nu\alpha\beta} F^{\mu\nu} \bar{N}[\gamma^{\alpha}, \gamma^{\beta}] N \quad \text{WHERE} \quad d_n \sim \frac{em_u m_d}{(m_u + m_d) \Lambda_{QCD}^2} \theta_{QCD}$

THIS IS JUST A CLASSICAL ELECTRIC DIPOLE MOMENT.

$$H_d = -d_n(\bar{N}\sigma N) \cdot \mathbf{E}$$

BUT EXPERIMENTAL BOUND ON NEUTRON EDM GIVES

$$|d_n| \lesssim 3 \times 10^{-26} e\,{\rm cm} \Rightarrow \theta_{QCD} \lesssim 10^{-10}$$
 APPARENT NUMERICAL TUNING OF 10 ORDERS OF MAGNITUDE!

*CAN MOVE IT INTO QUARK MASSES BY REPHASINGS, BUT IT ALWAYS SHOWS UP SOMEWHERE



AXIONS?

DYNAMICALLY ADJUST θ TO ZERO?

CONSIDER PSEUDOSCALAR a COUPLING TO $G\tilde{\rm G}$

$$\mathcal{L} \supset \frac{1}{2} (\partial_{\mu} a)^2 + \frac{\theta}{32\pi^2} G\tilde{G} + \frac{a}{f_a} \frac{1}{32\pi^2} G\tilde{G} + \dots$$

REST OF THEORY HAS SHIFT SYMMETRY a
ightarrow a + lpha freedom to arbitrarily shift heta

IN FACT, QCD VACUUM ENERGY DEPENDS ON θ , $E(\theta) = (m_u + m_d)e^{i\theta}\langle \bar{q}q \rangle$ AXION VEV MINIMIZES QCD VACUUM ENERGY, WITH $\langle a \rangle = \theta f_a \Rightarrow \bar{\theta} = 0$

> SEEMS ARBITRARY, BUT COUPLING & SHIFT SYMMETRY FOLLOW DIRECTLY IF AXION IS PNGB OF SPONTANEOUSLY BROKEN U(1)

AXION LIGHT (MASS $\sim \Lambda_{QCD}^2/F$) COSMOLOGICALLY RELEVANT: COSMOLOGICAL LIMITS; DARK MATTER?



SN 1987A (a-N-coupling)

SPONTANEOUS CPV?

WHAT IF CP IS A GOOD SYMMETRY OF THE STANDARD MODEL, SPONTANEOUSLY BROKEN IN A CONTROLLED WAY (*BECAUSE CKM*)?

ONE PHYSICAL STRONG CP ANGLE: $\bar{\theta} = \theta_{QYC} - \theta_{QCD}$ WHERE FORMALLY THE QUARK MASS TERM PHASE IS $\theta_{QYC} = \operatorname{ArgDet}[Y_u Y_d]$

THE CHALLENGE:

WHY IS $\theta_{QYC} = \operatorname{ArgDet}[Y_u Y_d]$ SMALL, BUT

THE OBSERVED CKM PHASE

 $\theta_{weak} = \operatorname{ArgDet}[Y_u Y_d - Y_d Y_u]$

IS BIG?

SOUNDS LIKE IT'S TIME TO BUILD A MODEL ...

SPONTANEOUS CPV?

POPULAR CLASS OF SPONTANEOUS CPV SOLUTIONS: NELSON-BARR

AN ALTERNATIVE ALONG SIMILAR LINES: SPONTANEOUS PV. [BARR, CHANG, SENJANOVIC '91]

EXTEND SM W/ PARITY

 $SU(3)_c \times SU(2)_L \times U(1)_Y \Rightarrow SU(3)_c \times SU(2)_L \times SU(2)'_L \times U(1)_Y$

+ EXTRA "MIRROR" COPY OF SM MATTER CHARGED UNDER SU(2)_'

NOW A GENERALIZED PARITY SYMMETRY UNDER WHICH $P: SU(2)_L \leftrightarrow SU(2)'_L$ (SENDS SM MATTER INTO MIRROR MATTER)

 $\theta \rightarrow -\theta$ under this parity, so zero in UV if P is a good symmetry.

PARITY ALSO REQUIRES $Y_u HQu + Y'_u H'Q'u' = Y_u HQu + Y'_u H'Q'u'$

SO THAT $\operatorname{ArgDet}[Y_u Y_d] + \operatorname{ArgDet}[Y'_u Y'_d] = 0$ BUT CKM PHASE ALLOWED.

SPONTANEOUS CPV?

BUT: WE DON'T SEE THE MIRROR QUARKS CHARGED UNDER $SU(2)_{L}'$, SO MUST SPONTANEOUSLY BREAK $SU(2)_{L} \leftrightarrow SU(2)_{L}'$ PARITY

VIA E.G. A PARITY-ODD FIELD ϕ THAT GETS A VEV AND MAKES (H) \neq (H') $\mathcal{L} \supset g\phi(|H'|^2 - |H|^2) \Rightarrow \langle H' \rangle \sim \langle \phi \rangle \gg \langle H \rangle$

BUT **Φ** *VEV CAN'T BE TOO BIG*, BECAUSE NOW WE EXPECT OPERATORS LIKE $\frac{1}{32\pi^2}\frac{\phi}{M_{Pl}}G\tilde{G}$

NOT REINTRODUCING STRONG CP PROBLEM BOUNDS $\langle \phi \rangle \sim \langle H' \rangle \lesssim 10^{-10} M_{Pl}$

SO FIRST-GENERATION MIRROR U, D, E FERMIONS SHOULD BE BENEATH 10 TEV!

THESE FERMIONS CARRY BOTH CHARGE AND COLOR. SYMMETRIES ALLOW MIXING W/ SM FERMIONS:

$$\mathcal{L} \supset -\mu_u u u' - \mu_d dd' - \mu_e e e'$$

MIXING LEADS TO DECAYS SUCH AS E.G.

$$u' \to h + u$$
 $u' \to Z + u$ $u' \to W + d$

[D'AGNOLO, HOOK '15]

SPONTANEOUS CPV @ LHC

PARITY SOLUTION PREDICTS NEW CHARGED/COLORED FERMIONS <10 TEV W/ SM DECAY MODES



UNIFICATION

GIVEN MEASURED SM GAUGE COUPLINGS AT WEAK SCALE, CAN STUDY EVOLUTION TO HIGHER SCALES WITH RGES.

 $b_1 = 41/10$ $b_2 = -19/6$

$$\frac{\partial \alpha_i}{\partial \ln \mu} = \beta_i = b_i \frac{\alpha_i^2}{2\pi} + \dots \Rightarrow \frac{1}{\alpha_i(\mu)} - \frac{1}{\alpha_i(m_Z)} = -\frac{b_i}{2\pi} \ln\left(\frac{\mu}{m_Z}\right) + \dots \quad \left(\alpha_i \equiv \frac{g_i^2}{4\pi}\right)$$

$$\begin{array}{c}
70 \\
60 \\
50 \\
60 \\
50 \\
40 \\
30 \\
20 \\
100 \\
10^5 \\
10^8 \\
10^{11} \\
10^{14} \\
10^{17} \\
\mu [GeV]
\end{array}$$

SUGGESTIVELY, THE THREE APPEAR TO CROSS (MISSING TRIPLE INTERSECTION BY 0(10%)) AROUND 10¹⁵ GEV.

 $b_3 = -7$

CONSISTENT WITH UNIFICATION OF SU(3)XSU(2)XU(1) INTO COMMON GAUGE GROUP.

CONVENIENTLY $SO(10) \supset SU(5) \supset SU(3) \times SU(2) \times U(1)$

UNIFICATION

HOW DO THE PIECES FIT TOGETHER?

 $SU(5) \text{ rep} \to (SU(3), SU(2))_{U(1)_Y} \text{ rep} = \text{SM field}$ $\mathbf{5} \to (\mathbf{3}, \mathbf{1})_{-1/3} \oplus (\mathbf{1}, \mathbf{2})_{1/2} = T + H$ $\mathbf{\overline{5}} \to (\mathbf{\overline{3}}, \mathbf{1})_{1/3} \oplus (\mathbf{1}, \mathbf{2})_{-1/2} = \bar{d} + L$ $\mathbf{10} \to (\mathbf{3}, \mathbf{2})_{1/6} \oplus (\mathbf{\overline{3}}, \mathbf{1})_{-2/3} \oplus (\mathbf{1}, \mathbf{1})_1 = Q + \bar{u} + \bar{e}$ $\mathbf{24} \to (\mathbf{8}, \mathbf{1})_0 + (\mathbf{1}, \mathbf{3})_0 + \mathbf{1} + (\mathbf{3}, \mathbf{2})_{-5/6} + (\mathbf{\overline{3}}, \mathbf{2})_{5/6} = G + W + B + X + \bar{X}$

SM MATTER FITS TIDILY, BUT DEMANDS TRIPLET HIGGS & NEW GAUGE BOSONS.

- BEAUTIFUL IDEA, SIMPLER THEORY IN FAR UV (ORIGINAL "NATURALNESS")
- BUT UNIFICATION OF COUPLINGS IMPERFECT @ 10% LEVEL.

- PREDICTS YUKAWA UNIFICATION, NOT IN GOOD AGREEMENT.
- PREDICTS PROTON DECAY VIA
 EXCHANGE OF T & X

UNIFICATION



 $\begin{array}{l} \mathsf{X} \; \mathsf{EXCHANGE} \\ \mathsf{GENERATES} \\ \mathsf{DIM-6} \; \mathsf{OPS} \\ \frac{1}{\Lambda^2} Q L \bar{u}^\dagger \bar{d}^\dagger \\ \frac{1}{\Lambda^2} Q Q \bar{u}^\dagger \bar{e}^\dagger \end{array}$

T EXCHANGE GENERATES DIM-6 OPS $\frac{1}{\Lambda^2} \bar{u} \bar{u} \bar{d} \bar{e}$ $\frac{1}{\Lambda^2} QQQL$



WITH $\Lambda \sim M_{GUT} \sim 10^{15} \,\mathrm{GeV}$



FOR M_{GUT} =10¹⁵ GEV, PREDICT LIFETIME

 $\Gamma \sim \frac{m_p^5}{M_{GUT}^4} \sim 10^{29} \, \text{years}$



EXPERIMENTAL LIMIT (E.G. SUPER-KAMIOKANDE): τ >8*10³³ YEARS



VANILLA UNIFICATION EXCLUDED BY DATA.

IMPROVING UNIFICATION

CONSIDER THE EFFECTS OF ADDING NEW FERMIONS* AT SCALE $\ensuremath{\mathsf{M}}\xspace{\psi}$

 $\frac{1}{\alpha_{GUT}} = \frac{1}{\alpha_i(m_Z)} - \frac{b_i^{SM}}{2\pi} \ln\left(\frac{M_{GUT}}{m_Z}\right) - \frac{\Delta b_i}{2\pi} \ln\left(\frac{M_{GUT}}{M_\Psi}\right) + \dots$

UNIVERSAL Δb_i ONLY SHIFTS VALUE OF $\pmb{\alpha}_{\text{GUT}}$

DIFFERENCES $\Delta b_i - \Delta b_j$ CHANGE PRECISION OF UNIFICATION & VALUE OF M_{GUT}

SU(5)	${ m SU}(3)\otimes$	SU(2)	\otimes U(1)	n_3	\bar{n}_3	n_2	z	name	Δb_3	Δb_2	Δb_1
$5\oplus ar{5}$	$\overline{3}$	1	1/3	0	1	0	0	D	2/3	0	4/15
$5\oplus ar{5}$	1	2	$^{1}/_{2}$	0	0	1	0	L	0	2/3	2/5
$10\oplus\overline{10}$	$\overline{3}$	1	$-2/_{3}$	0	1	0	1	U	2/3	0	16/15
$10 \oplus \overline{10}$	1	1	-1	0	0	0	1	E	0	0	4/5
$10 \oplus \overline{10}$	3	2	$1/_{6}$	1	0	1	0	Q	4/3	2	2/15
$15\oplus\overline{15}$	3	2	$^{1}/_{6}$	=	=	=	=	Q	=	=	=
$15\oplus\overline{15}$	1	3	1	0	0	2	0	T	0	8/3	12/5
$15\oplus\overline{15}$	6	1	$-2/_{3}$	2	0	0	0	S	10/3	0	32/15
24	1	3	0	0	0	2	1	V	0	4/3	0
24	8	1	0	1	1	0	0	G	2	0	0
24	$\overline{3}$	2	5/6	0	1	1	0	X	4/3	2	10/3

SOME REPRESENTATIONS AND THEIR SHIFTS:

[GIUDICE, RATTAZZI, STRUMIA]

*COULD ADD SCALARS TOO, BUT MAKES MUCH SMALLER CHANGE IN RUNNING.

IMPROVING UNIFICATION



ADDING REPRESENTATIONS IMPROVES UNIFICATION PREDICTION AND RAISES GUT SCALE. IF REPRESENTATIONS NOT TOO LARGE, NEED SCALE TO BE NEAR WEAK SCALE.

FOR MGUT=10¹⁶ GEV, PROTON LIFETIME AT EDGE OF CURRENT LIMITS.



UNIFICATION @ TEV

REPS THAT 'HELP' (IMPROVE PRECISION, RAISE SCALE)

SU(5)	${ m SU}(3)\otimes$	SU(2)	\otimes U(1)	n_3	\bar{n}_3	n_2	z	name	Δb_3	Δb_2	Δb_1	SANAE OLIANITI INA #S AS
$5\oplus ar{5}$	$\overline{3}$	1	$1/_{3}$	0	1	0	0	D	2/3	0	4/15	
$5\oplus ar{5}$	1	2	$1/_{2}$	0	0	1	0	L	0	2/3	2/5	HIGGSINOS IN SUSY
$10 \oplus \overline{10}$	$\overline{3}$	1	$-2/_{3}$	0	1	0	1	U	2/3	0	16/15	
$10 \oplus \overline{10}$	1	1	-1	0	0	0	1	E	0	0	4/5	
$10 \oplus \overline{10}$	3	2	$1/_{6}$	1	0	1	0	Q	4/3	2	2/15	
$15 \oplus \overline{15}$	3	2	$1/_{6}$	=	=	=	=	Q	=	=	=	
$15 \oplus \overline{15}$	1	3	1	0	0	2	0	T	0	8/3	12/5	SAMF QUANTUM #S AS
$15 \oplus \overline{15}$	6	1	$-2/_{3}$	2	0	0	0	S	10/3	0	32/15	
24	1	3	0	0	0	2	1	V	0	4/3	0	VECTOR-LIKE QUARKS IN
24	8	1	0	1	1	0	0	G	2	0	0	COMPOSITE HIGGS
24	$\overline{3}$	2	$\frac{5}{6}$	0	1	1	0	X	4/3	2	10/3	

TAKEAWAY: SEARCHES FOR HIGGSINOS, VECTOR-LIKE QUARKS CAN BE MOTIVATED BY IMPROVED GAUGE COUPLING UNIFICATION, WHERE THE PRESSURE FOR ACCESSIBLE SCALES COMES NOT FROM NATURALNESS, BUT FROM LOGARITHMIC RUNNING OF COUPLINGS.

BARYOGENESIS

OBSERVE UNIVERSE IS PRIMARILY MADE OF BARYONS, NOT ANTI-BARYONS,

QUANTITATIVELY,
$$\eta \equiv \frac{n_B - n_{\bar{B}}}{n_{\gamma}} \sim 6 \times 10^{-10}$$

IF UNIVERSE STARTED WITH η = 0 and Baryons decoupled Like WIMPS,

$$\frac{n_B}{n_{\gamma}} \simeq \frac{n_{\bar{B}}}{n_{\gamma}} \simeq \left(\frac{m_p}{T}\right)^{3/2} e^{-m_p/T} \to 10^{-18} \ (T_f \sim 20 \,\mathrm{MeV})$$

IN BAD DISAGREEMENT! MORE OR LESS THREE OPTIONS:

 INITIAL CONDITIONS ARE TUNED. → DEEPLY UNSATISFYING, ESSENTIALLY IMPOSSIBLE W/ INFLATION
 B AND B SPATIALLY SEPARATED. → DISFAVORED BY DATA
 ASYMMETRY IS DYNAMICAL.

BARYOGENESIS

SAKHAROV CONDITIONS FOR DYNAMICAL BARYON ASYMMETRY:

1. BARYON # VIOLATION (NEED TO GET NET BARYON # FROM B=0)

2. C & CP VIOLATION (OTHERWISE RELATE B, \overline{B} -CREATING PROCESSES)

3. DEPARTURE FROM THERMAL EQUILIBRIUM

IN PRINCIPLE POSSIBLE WITHIN SM DURING ELECTROWEAK PHASE TRANSITION:

1. NONPERTURBATIVE ELECTROWEAK CONFIGURATIONS (SPHALERONS)

2. CP VIOLATION FROM CKM + DOMAIN WALL BREAKS C

3. IF PHASE TRANSITION IS STRONGLY FIRST-ORDER

IN PRACTICE, NOT ENOUGH OF ANYTHING: CPV FROM CKM PHASE TOO SMALL, EWPT NOT FIRST ORDER FOR MH=125 GEV,

BARYOGENESIS

SOME OPTIONS (NOT EXHAUSTIVE)



BARYOGENESIS@TEV1: ELECTROWEAK BARYOGENESIS

ADD MATTER TO SM TO ALTER HIGGS POTENTIAL, MAKE EWPT STRONGLY 1ST-ORDER

E.G. $\kappa |\Phi|^2 |H|^2$ WITH Φ LIGHT AND κ LARGE

IF Φ CHARGED/COLORED, AN EASY GAME: SEARCH VIA DIRECT PRODUCTION AT HADRON COLLIDERS OR LOOK FOR HIGGS COUPLING DEVIATIONS.



BARYOGENESIS@TEV 2: WIMPY BARYOGENESIS

NEW PARTICLE GETS THERMAL ABUNDANCE FROM FREEZE-OUT, LIKE DARK MATTER (WIMP MIRACLE → WEAK SCALE COUPLINGS & MASS). OUT-OF-EQUILIBRIUM DECAYS VIOLATE CP, BARYON/LEPTON



[CUI, RANDALL, SHUVE '11, CUI & SHUVE '14]

TAKEAWAY: NO GUARANTEE OF ACCESSIBLE NEW PHYSICS, BUT MANY BARYOGESIS MECHANISMS MOTIVATE SIGNALS AT THE WEAK SCALE.



BEYOND THE STANDARD MODEL

TEV SCALE



EPILOGUE: LOOKING TO THE FUTURE

LIFE AFTER THE LHC

CEPC (250 GEV), SPPC (50-100 TEV)



FCC-EE (250 GEV), FCC-HH (80-100 TEV)



<complex-block>

- THERE IS A SUPERABUNDANCE OF MOTIVATION FOR NEW PHYSICS AT THE TEV SCALE, BOTH FROM CONVENTIONAL BSM DRIVERS (THE HIERARCHY PROBLEM) AND FROM LESS CONVENTIONAL ONES (STRONG CP PROBLEM, UNIFICATION, BARYOGENESIS, DARK MATTER, NEUTRINOS,...).
- POPULAR SOLUTIONS BEING TESTED, BUT NEW SOLUTIONS TO THESE PROBLEMS ABOUND, WITH NEW SIGNALS AT THE TEV SCALE.
- MANY OF THESE SIGNALS ARE ONLY NOW COMING INTO THE REACH OF TEV-SCALE PROBES, AND MAKE INTERESTING GOALS FOR FUTURE COLLIDERS.



