Introduction to FLYCHK

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FLYCHK COLLISIONAL-RADIATIVE MODEL

Population Kinetics Modeling

Rate equations are solved for level population distributions for given plasma conditions

$$\frac{dn_i}{dt} = -n_i \sum_{j \neq i}^{N \max} W_{ij} + \sum_{j \neq i}^{N \max} n_j W_{ji}$$

$$W_{ij} = B_{ij}\overline{J_{ij}} + n_e C_{ij} + \beta_{ij} + n_e \gamma_{ij} \quad W_{ji} = A_{ij} + B_{ji}\overline{J_{ji}} + n_e D_{ji} + n_e (\alpha_{ji}^{RR} + \alpha_{ji}^{DR}) + n_e^2 \delta_{ij}$$

- B_{ij} Stimulated absorption
- C_{ij} Collisional excitation
- γ_{ij} Collisional ionization
- B_{ij} Photoionization (+st. recom)

- A_{ii} Spontaneous emission
- B_{ii} Stimulated emission
- D_{ii} Collisional deexcitation
- a_{ij}^{DR} Dielectronic recombination
- a_{ij}^{RR} Radiative recombination
- δ_{ii} Collisional recombination

FLYCHK uses screened hydrogenic levels (super configurations)



Hydrogenic

Level energy obtained with ionization potential from its 1st continuum level



Atomic processes included in FLYCHK



electron capture

FLYCHK (nl) (nl) (nlj) HULLAC / FAC / MCDF (n) (nl) (nlj) (detailed-term)

- Screened hydrogenic energy levels with relativistic corrections
- Relativistic Hartree-Slater oscillator strengths (M. Chen) and photoionization cross-sections (J. Scofield,+ Kramer)
- Fitted collisional cross-section to PWB approximation (M. Chen)
- Semi-empirical cross-sections for collisional ionization (A. Burgess)
- Detailed counting of autoionization and electron capture (M. Chen)
- Continuum lowering (Stewart-Pyatt, Ecker-Kroll)

Application to a wide range of Z & experiments:

Excitation autoionization (EA) / Dielectronic recombinationa (DR) processes are modeled with extensive inner-shell (IS) states



FLYSPEC SPECTROSCOPIC MODULE

FLYSPEC uses detailed (H, He, Li-like) and Super Transition Array for spectra



Data Types for Spectroscopic Model

Z < 27 H, He and Li	FLY model
Z > 27 H, He and Li	HULLAC data (term levels up to n=4)
Be-like and lower charge states	Super Transition Array (STA) made with Configurations (jj) 1s, 2s, 2p ⁻ , 2p ⁺ , 3s, 3p ⁻ , 3p ⁺ , 3d ⁻ , 3d ⁺ , Up to n=6

Energy-dependent spectral intensity in the STA formalism

Spectra for specific E/ ranges: STA formalism

Spectra using configuration-average atomic data generated by the DHS (Dirac-Hartree-Slater) code (M.Chen)

$$\eta(\mathbf{v}) = n_A A_{AB} E_{AB} \phi(\mathbf{v}) = \frac{n_A \sum_{i \in A: j \in B} g_i \exp(-E_i / kT_e) A_{ij} E_{ij} \phi(\mathbf{v})}{\sum_{i \in A: j \in B} g_i \exp(-E_i / kT_e)} \quad \text{[ergs/s/Hz/cm³/ster]}$$

$$A_{AB} = \frac{\sum_{i \in A, j \in B} \exp(-E_i / kT_e) A_{ij}}{\sum_{i \in A, j \in B} g_i \exp(-E_i / kT_e) A_{ij}} \qquad E_{AB} = \frac{\sum_{i \in A, j \in B} g_i \exp(-E_i / kT_e) A_{ij}}{\sum_{i \in A, j \in B} g_i \exp(-E_i / kT_e) A_{ij}} \qquad \mu_{AB}^2 = \left[\frac{\sum_{i \in A, j \in B} g_i \exp(-E_i / kT_e) A_{ij} E_{ij}}{\sum_{i \in A, j \in B} g_i \exp(-E_i / kT_e) A_{ij}}\right]^2 - E_{AB}^2$$

$$\frac{\text{Run time: Thu Mar 24 12:00:45}}{\text{Input and output files @ vertex is vertex vertex vertex is vertex v$$

Total line emissivity in the STA formalism

Approximate total line emissivity:

A plot show approximate line emission spectra and provides information on energy range of dominant emission

$$S = n_u A_{ul} E_{ul} / N_e$$



Screened-Hydrogenic



FLYCHK APPLICATIONS

FLYCHK Help Pages

- <u>http://nlte.nist.gov/FLY/Doc/</u> <u>Manual_FLYCHK_Nov08.pdf</u>
- <u>http://nlte.nist.gov/FLY/README.html</u>
- <u>http://nlte.nist.gov/FLY/EXAMPLE.html</u>
- Click on the Question Marks
 - -<u>http://nlte.nist.gov/FLY/Help/runfile.html</u>
 - <u>http://nlte.nist.gov/FLY/Help/opacity.html</u>

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Available to the community at passwordprotected NIST website: <u>http://nlte.nist.gov/FLY</u>

Advantages: simplicity and versatility→ applicability

- <Z> for fixed any densities: electron, ion or mass
- Mixture-supplied electrons (eg: Argon-doped hydrogen plasmas)
- External ionizing sources : a radiation field or an electron beam.
- Multiple electron temperatures or arbitrary electron energy distributions
- Optical depth effects

Outputs: population kinetics code and spectral synthesis

- <Z> and charge state distribution
- Radiative Power Loss rates under optically thin assumption
- Energy-dependent spectral intensity of uniform plasma with a size

Caveats: simple atomic structures and uniform plasma approximation

- Less accurate spectral intensities for non-K-shell lines
- Less accurate for low electron densities and for LTE plasmas
- When spatial gradients and the radiation transport affect population significantly

Example: Radiative loss rates are important as an energy loss mechanism of high-Z plasmas

Calculated Kr radiative cooling rates per N_e [eV/s/atom/cm⁻³]



of radiative transitions using HULK code

lon	HULLAC+DHS
1	3049
2	27095
3	30078
4	404328
5	3058002
6	5882192
7	7808014
8	6202123
9	5544814
10	1050919
11	841094
Sum	30,851,708

Data for Radiation Hydrodynamics: Kr Radiative loss rates over (Ne, Te)



List of Selected Cases		
Dens	Data	
1.e+15	file	
1.e+16	file	
1.e+17	file	
1.e+18	file	
1.e+19	file	
1.e+20	file	
1.e+21	file	
1.e+22	file	
1.e+23	file	
1.e+24	file	

The radiative loss rates show the similar coronal behavior up to $N_e = 10^{17}$ and the rate/ N_e stays constant. As N_e increases, the rate/ N_e decreases from the coronal \Rightarrow value

Example: Gold ionization balance in high temperature hohlraum (HTH) experiments

- High-T hohlraum reach temperatures: ~ 10 keV
- Spectrum from $n_e \sim 4x10^{21}$ cm⁻³, $T_e \sim 7-10$ keV measured for first time



FLYCHK gives an estimate of Gold L-shell spectra

Relative Intensity

Long pulse laser plasmas: Gold L-shell spectroscopy

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		2 nd Te [eV] 🚱	2nd Te:		Fraction:		Or history file: 🔲	
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		Radiation Field 🤪				Browse		
		EEDF 设				Browse		
							Run FLYCHK	Clear

FLYCHK at NIST is developed and managed by H.-K. Chung, M. Chen and R. W. Lee at LLNL and Yu. Ralchenko at NIST. This work was performed under the auspices of the U.S. Department of Energy by University of California Lawrence Livermore National Laboratory under Contract No. W-7406-Eng-48

STA spectra compared with configurationaverage spectra



Example: Cu K α radiation measured by single hit CCD spectrometer and 2-D imager for T_e diagnostics



Single Hit CCD K α yield is higher than that of 2-D imager for smaller target volumes : An experimental evidence of shifting and broadening of K α emission lines in small targets with high temperatures

Shifts and Broadening of $K\alpha$ emission as a function of electron thermal temperature



Short pulse laser plasmas: Cu Kα Spectroscopy



Example: Photoionized plasmas produced by Z-Machines - Astrophysical model benchmark



 ξ =20-25 ergs-cm/s

Charge state distribution is a function of N_e and Radiation field strength

 $N_e = 1.95E19cm^{-3}$ Radiation field of 165 eV and 0.01 dilution



Photoionization equilibrium plasmas: Fe Z-Pinch Plasma



Example: XFEL provides an opportunity for HEDS plasma spectroscopy



In Warm Dense Matter regime the hollow ions provide timeresolved diagnostic information

- XFEL forms unique states and provides in situ diagnostics with ~100 fs res.
 - $5x10^{10}$ 1.85 keV photons in 30 µm spot into a $n_e = 10^{23}$ cm⁻² plasma
 - Strong coupling parameter, Γ_{ii} = Potential/Kinetic Energy ~ 10



XFEL ionized plasma: Mg time-dependent K-shell spectroscopy



Postprocessing electron kinetic simulation

- SiO₂ aerogel targets doped with Ge or Ti for X-ray backlighter development
- 1-D e- kinetic code FPI shows Non-maxwellian energy functions due to strong laser heating and nonlocal electron heat flow -- J-P. Matte & K.B. Fournier

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lon T _i [eV] 🤪 📑	T ₁ /T _e :	Fixed T ₄ :	Or history file:				3906E+02
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linear or log grid only	number: 41				114570016F+07**	E7374E+E052P;55F41E+051;41	
linear followed by log	Final time for linear:	Linear step number:	Log step number:		.time 0. 1.e	2-9 41	
log followed by linear	Final time for log:	Log step number:	Linear step number:		end		
			Run FLYCHK)	Clear)			

Output: <Z> is quite similar with/without thermal e-

Using fe(E) with thermal e-

Using fe(E) only without thermal e-



Aluminum Opacity (NIST data)



Electron temperature (eV)

Useful Examples http://nlte.nist.gov/FLY/EXAMPLE.html

Please check the Screen shot of each case

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	Radiation T _r [eV] T _{rad} : Dilution :	Or history file:
	Radiation Field 💿	Browse
	EEDF 💿	Browse

Theory and Modeling:

R. W. Lee, M. H. Chen, H. A. Scott, M. Adams, M. E. Foord, S. J. Moon, S. B. Libby, S. B Hansen, K. B. Fournier, B. Wilson, C. Iglesias, M. May, S. C. Wilks, A. Kemp, R. Town, M. F. Gu, M. Tabak (LLNL), Y. Ralchenko (NIST), A. Bar-Shalom (HULLAC, Israel), J. Oreg (HULLAC, Israel), M. Klapisch (HULLAC), M. S. Wei, R. B. Stephens (GA), B. Ziaja, S. Son (CFEL, Germany), M. Bussman, T. Kluge, L.Huang (HZDR, Germany), E. Stambulchik (Weizmann, Israel). M.S. Cho (GIST, Korea)

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Thank you!