

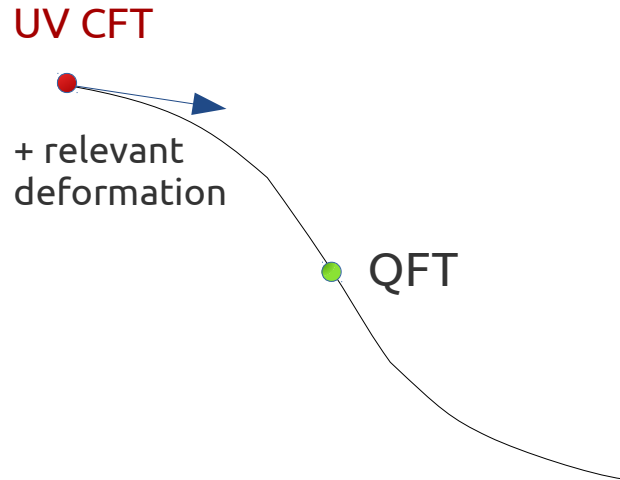
\overline{TT} and the mirage of a bulk cutoff

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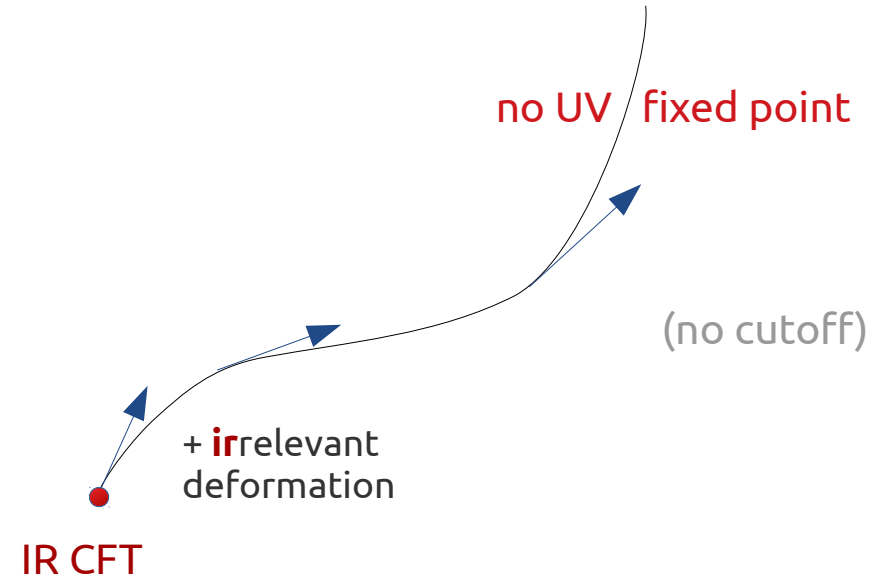
based on 1906.11251: with Ruben Monten

Motivation

Usual framework: **local, UV complete** QFTs



Examples of **non-local, UV complete** QFTs



- Quantum gravity ?
- Holography in **non-asymptotically AdS** spacetimes

$T\bar{T}$ - deformed CFTs

- **universal** deformation of 2d CFTs/QFTs

$$\frac{\partial S}{\partial \mu} = \int d^2 z \underbrace{(T_{zz}T_{\bar{z}\bar{z}} - T_{z\bar{z}}^2)}_{\text{“}T\bar{T}\text{”}} \mu$$

- deformation **irrelevant** (dim = (2,2)) but **integrable**

finite size spectrum, partition function, thermodynamics

Smirnov & Zamolodchikov, Cavaglia et al, Cardy

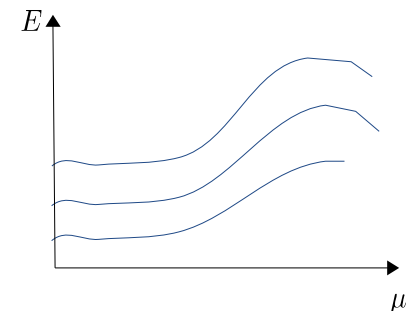
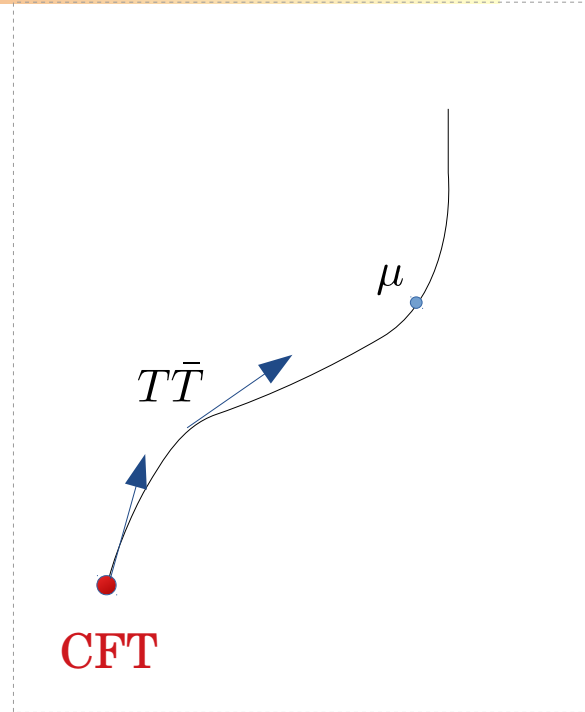
- energy levels **smoothly deformed**

$$E(\mu) = -\frac{R}{2\mu} \left(1 - \sqrt{1 + \frac{4\mu E_0}{R} + \frac{4\mu^2 P_0^2}{R^2}} \right)$$

- deformed theory **non-local** (scale μ) but argued **UV complete**

S-matrix (2 \rightarrow 2): $\mathcal{S}_\mu = e^{\frac{i\mu s}{4}} \mathcal{S}_0$

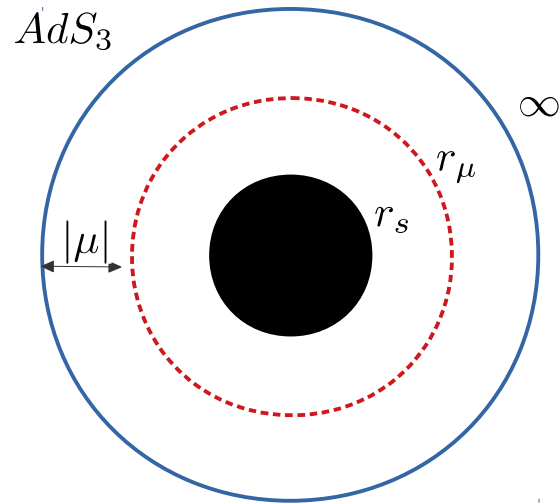
Dubovsky et al.



$\overline{\text{T}\overline{\text{T}}}$ and the finite bulk cutoff

- energy spectrum of $\overline{\text{T}\overline{\text{T}}}$ -deformed CFTs with $\mu < 0$ **exactly matches** energy of a “black hole in a box”

McGough, Mezei, Verlinde '16



$$E(\mu) = \frac{R}{2|\mu|} \left(1 - \sqrt{1 - \frac{4|\mu|M}{R} + \frac{4\mu^2 J^2}{R^2}} \right)$$

= energy measured by an observer on a fixed radial slice r_μ

- imaginary energies** for large M $\iff r_s > r_\mu$
at μ, R **fixed**
- matter fields? $\mu > 0$?



Finite bulk cutoff usually associated with **integrating out degrees of freedom** in bulk/boundary

(holographic Wilsonian RG)



Integrability & UV completeness of $\overline{\text{T}\overline{\text{T}}}$?

This talk

- first principles derivation of the holographic dictionary for $T\bar{T}$ - deformed CFTs for both signs of μ
- as expected for double trace: AdS_3 with $\left\{ \begin{array}{l} \text{mixed boundary conditions at } \infty \text{ for the metric} \\ \text{unchanged (Dirichlet) for the matter fields} \end{array} \right.$
- for $\mu < 0$ and pure gravity and on-shell \approx Dirichlet at finite radius r_μ independent of the mass
? pure coincidence
- when matter field profiles (vevs) are present, no special reinterpretation in terms of Dirichlet at finite radius

Double-trace deformations in AdS/CFT

- $T\bar{T}$ is a **double-trace** deformation \rightarrow **mixed** boundary conditions for dual bulk fields

- e.g. **scalar**

$$\Phi = \mathcal{J} z^{d-\Delta} + \dots + \langle \mathcal{O} \rangle z^{\Delta} + \dots$$

source (fixed) **vev (fluctuates)**

- $S_{\mu} = S_{CFT} + \mu \int \mathcal{O}^2$

- **1) variational principle** (equivalent to Hubbard-Stratonovich, only uses **large N field theory**)

$$\delta S_{\mu} = \delta S_{CFT} - \delta \left(\mu \int \mathcal{O}^2 \right) = \int \mathcal{O} \delta \mathcal{J} - \mu \int \delta \mathcal{O}^2 = \int \underbrace{\mathcal{O}}_{\substack{\text{new vev} \\ \uparrow}} \underbrace{\delta(\mathcal{J} - 2\mu\mathcal{O})}_{\text{new source } \tilde{\mathcal{J}}}$$

- **2)** translate into boundary conditions on the bulk field

Sources and vevs in $T\bar{T}$ - deformed CFTs

- variational principle approach:

$$\delta S_{CFT} - \Delta\mu \delta S_{T\bar{T}} = \int d^2x \underbrace{\sqrt{\gamma} T_{\alpha\beta} \delta\gamma^{\alpha\beta}}_{\text{CFT}} - \Delta\mu \int d^2x \delta(\underbrace{\sqrt{\gamma} \epsilon^{\alpha\gamma} \epsilon^{\beta\delta} T_{\alpha\beta} T_{\gamma\delta}}_{\text{deformation}}) = \int d^2x \underbrace{\sqrt{\tilde{\gamma}} \tilde{T}^{\alpha\beta} \delta\tilde{\gamma}^{\alpha\beta}}_{\text{new sources \& vevs}}$$

- flow equations

$$\partial_\mu \gamma_{\alpha\beta} = -2(T_{\alpha\beta} - \gamma_{\alpha\beta} T) \equiv -2\hat{T}_{\alpha\beta} \quad \partial_\mu \hat{T}_{\alpha\beta} = -\hat{T}_\alpha{}^\gamma \hat{T}_{\gamma\beta} \quad \partial_\mu (\hat{T}_\alpha{}^\gamma \hat{T}_{\gamma\beta}) = 0$$

- exact solution

$$\gamma_{\alpha\beta}(\mu) = \gamma_{\alpha\beta}(0) - 2\mu \hat{T}_{\alpha\beta}(0) + \mu^2 \hat{T}_\alpha{}^\gamma \hat{T}_{\gamma\beta}(0)$$

$$\hat{T}_{\alpha\beta}(\mu) = \hat{T}_{\alpha\beta}(0) - \mu \hat{T}_\alpha{}^\gamma \hat{T}_{\gamma\beta}(0)$$

- both signs of μ
- other (matter) vevs can be on
- large N field theory

- sources for matter operators unaffected at linear level

$$\sqrt{\gamma} \mathcal{O} \delta\mathcal{J} = \sqrt{\tilde{\gamma}} \tilde{\mathcal{O}} \delta\tilde{\mathcal{J}}$$

- flow equations

$$\longrightarrow \partial_\mu (R\sqrt{\gamma}) = 0, \quad \partial_\mu (\sqrt{\gamma} \mathcal{O}_{T\bar{T}}) = 0, \quad T(\mu) = \frac{c}{24\pi} R(\mu) - \mu \mathcal{O}_{T\bar{T}}(\mu)$$

The $T\bar{T}$ holographic dictionary

- **new sources** $\gamma_{\alpha\beta}(\mu) = \gamma_{\alpha\beta}(0) - 2\mu\hat{T}_{\alpha\beta}(0) + \mu^2\hat{T}_{\alpha}^{\gamma}\hat{T}_{\gamma\beta}(0)$

- **new vevs** $\hat{T}_{\alpha\beta}(\mu) = \hat{T}_{\alpha\beta}(0) - \mu\hat{T}_{\alpha}^{\gamma}\hat{T}_{\gamma\beta}(0)$

} large N
field theory

Holography → Fefferman Graham expansion

in original CFT

$$ds^2 = \frac{\ell^2 d\rho^2}{4\rho^2} + \left(\frac{g_{\alpha\beta}^{(0)}}{\rho} + g_{\alpha\beta}^{(2)} + \dots \right) dx^{\alpha} dx^{\beta}$$

$$g_{\alpha\beta}^{(0)} \leftrightarrow \gamma_{\alpha\beta}(0), \quad g_{\alpha\beta}^{(2)} \leftrightarrow 8\pi G\ell \hat{T}_{\alpha\beta}(0)$$

- $\gamma_{\alpha\beta}(\mu)$ **fixed** ↔ **mixed non-linear** boundary conditions for the AdS_3 metric

$$g_{\alpha\beta}^{(0)} - \frac{\mu}{4\pi G\ell} g_{\alpha\beta}^{(2)} + \frac{\mu^2}{(8\pi G\ell)^2} g_{\alpha\gamma}^{(2)} g^{(0)\gamma\delta} g_{\delta\beta}^{(2)} = \text{fixed}$$

- stress tensor expectation value **non-linearly related** to $g_{\alpha\beta}^{(2)}$
- **matter** field boundary conditions **unchanged**, since $\mathcal{J}(\mu) = \mathcal{J}(0)$

Pure gravity

- pure 3d gravity → Fefferman-Graham expansion truncates

$$ds^2 = \frac{\ell^2 d\rho^2}{4\rho^2} + \frac{g_{\alpha\beta}^{(0)} + \rho g_{\alpha\beta}^{(2)} + \rho^2 g_{\alpha\beta}^{(4)}}{\rho} dx^\alpha dx^\beta \qquad g_{\alpha\beta}^{(4)} = \frac{1}{4} g_{\alpha\gamma}^{(2)} g^{(0)\gamma\delta} g_{\delta\beta}^{(2)}$$

- mixed boundary conditions at ∞ → coincide precisely with Dirichlet at $\rho_c = -\frac{\mu}{4\pi G\ell}$

$\mu < 0$ coincides with McGough, Mezei, Verlinde

- deformed stress tensor → coincides precisely with Brown-York + counterterm at ρ_c

$$T_{\alpha\beta}(\mu) = \hat{T}_{\alpha\beta}(\mu) - \gamma_{\alpha\beta}(\mu) \hat{T}(\mu) = -\frac{1}{8\pi G} \overbrace{(K_{\alpha\beta} - K g_{\alpha\beta} - g_{\alpha\beta}/\ell)} (\rho_c)$$

- fixed by variational principle → no ambiguity!

The “asymptotically mixed” phase space

- most general pure gravity solution with $\gamma_{\alpha\beta}(\mu) = \eta_{\alpha\beta}$ (TT on flat space with coordinates U, V)
- $R(\mu) = 0 \Rightarrow R(0) = 0 \Rightarrow \gamma_{\alpha\beta}(0)dX^\alpha dX^\beta = dudv$ for some auxiliary coordinates u, v
- in these coordinates, the most general bulk solution is

$$ds^2 = \frac{\ell^2 d\rho^2}{4\rho^2} + \frac{dudv}{\rho} + \underline{\mathcal{L}(u)}du^2 + \underline{\bar{\mathcal{L}}(v)}dv^2 + \rho\mathcal{L}(u)\bar{\mathcal{L}}(v)$$

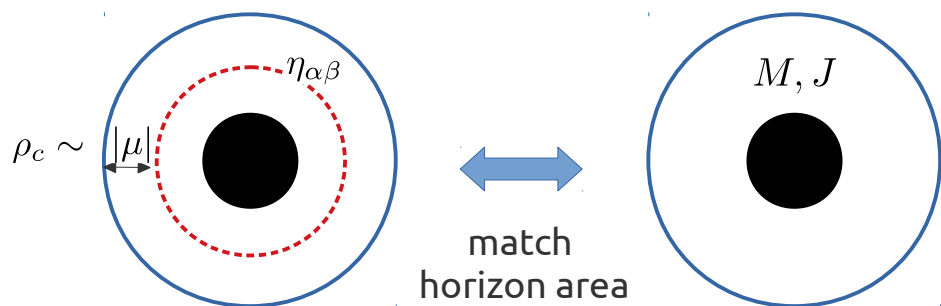
- boundary condition: $ds_c^2 = \rho_c ds^2(\rho = \rho_c) = dUdV \rightarrow$ relation between u, v and TT coordinates U, V

$$U = u + \rho_c \int^v dv' \bar{\mathcal{L}}(v'), \quad V = v + \rho_c \int^u du' \mathcal{L}(u')$$

- $g^{(0)}, g^{(2)}$ metric above in the U, V coordinate system (asymptotically mixed)
- most general solution parametrized by **two arbitrary functions** of the state-dependent coordinates

Energy match

- high energy eigenstates \rightarrow black holes : can we reproduce $E(\mu) = -\frac{R}{2\mu} \left(1 - \sqrt{1 + \frac{4\mu M}{R} + \frac{4\mu^2 J^2}{R^2}} \right) ?$
- deformed black hole: **constant** $\mathcal{L}_\mu, \bar{\mathcal{L}}_\mu$; **energy** $E(\mu) = \int d\phi T_{TT}(\mu)$
- relation to undeformed $M \sim \mathcal{L}_0 + \bar{\mathcal{L}}_0, J \sim \mathcal{L}_0 - \bar{\mathcal{L}}_0 ?$



deformed state

undeformed state

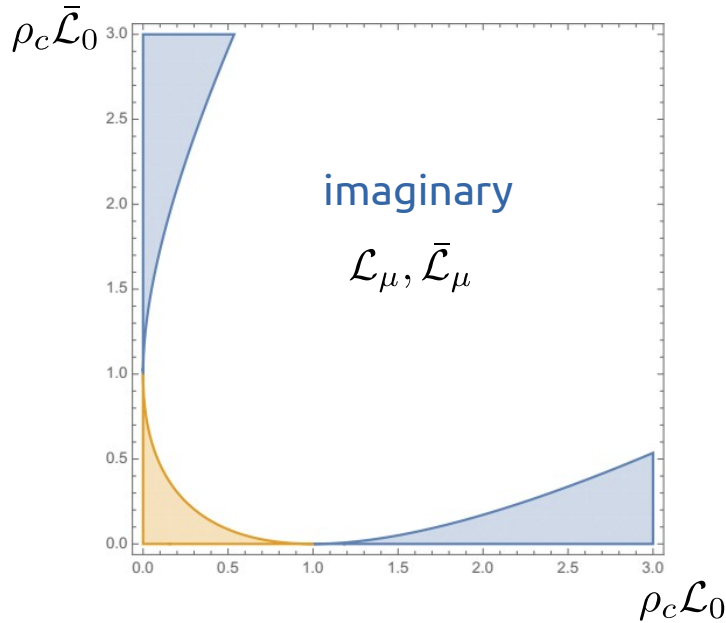
- energy eigenstates smoothly deformed \rightarrow **unchanged degeneracy**
- angular momentum quantized \rightarrow **unchanged**
- perfect match** for **both signs** of μ ✓

- McGough et al computed energy on **undeformed BTZ** at **Schwarzschild coordinate** $r_c \sim |\mu|^{-\frac{1}{2}}$

- map: $r^2(\rho) = \frac{(1 + \mathcal{L}_\mu(\rho - \rho_c) - \mathcal{L}_\mu \bar{\mathcal{L}}_\mu \rho \rho_c)(1 + \bar{\mathcal{L}}_\mu(\rho - \rho_c) - \mathcal{L}_\mu \bar{\mathcal{L}}_\mu \rho \rho_c)}{\rho(1 - \rho_c^2 \mathcal{L}_\mu \bar{\mathcal{L}}_\mu)^2}$ $r^2(\rho_c) = \frac{1}{\rho_c}$

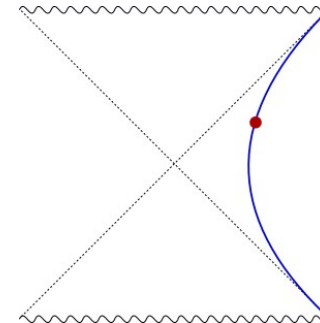
Imaginary energies

- for $\mu < 0$ the energy $E(\mu) = -\frac{R}{2\mu} \left(1 - \sqrt{1 + \frac{4\mu M}{R} + \frac{4\mu^2 J^2}{R^2}} \right)$ can become **imaginary**



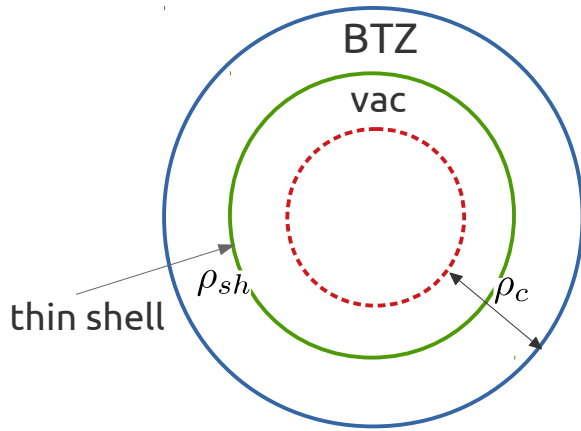
- orange** region ~ energies measured by observer outside outer horizon
- blue** region ~ energies measured by observer inside inner horizon ($g^{(0)}$ has CTCs)

- McGough et al picture still valid in **typical states**



Adding matter

- difference between **mixed at infinity** and **Dirichlet at finite radial distance** for $\mu < 0$



- shell **outside** ρ_c

→ **mixed** b.c. picture only depends on the asymptotic behaviour of the metric = BTZ → **energy matches** field theory ✓

→ **Dirichlet** b.c. yield vacuum answer ✗

- configurations **outside** this surface ✓ → 2d $\bar{T}\bar{T}$ describes **entire spacetime** : UV completeness & integrability

- imaginary energies? → breakdown of coordinate transformation used to make $\gamma_{\alpha\beta}(\mu) = \eta_{\alpha\beta}$ which **only depends on the asymptotic value** of the metric (no details of the interior matter)

Take-home:

universal formula for energy ↔ universal **asymptotic behaviour**

Conclusions

Summary and future directions

- large N holographic dictionary for $\overline{T\overline{T}}$ – deformed CFTs
 - derivation from variational principle: precision holography
 - both signs of μ and in presence of matter
 - mixed boundary conditions at infinity for the metric (no finite bulk cutoff)
 - ASG: non-local & state-dependent generalization of Virasoro

Future directions:

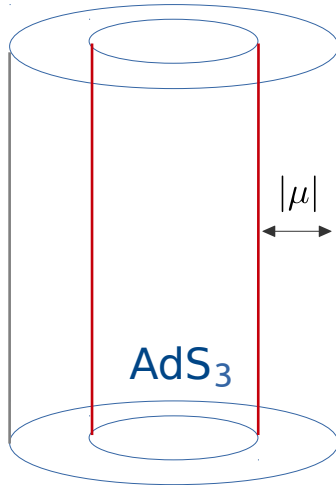
- precision match between all observables (e.g. correlation functions)? can holography help?
- $1/N$ corrections?
- field theory interpretation of the Virasoro symmetries → constraints on the theory/ non-locality?
- generic single trace generalisations of these UV-complete irrelevant deformations?
non- aAdS spacetimes

Thank you !

Holography: why interesting

Double-trace $T\bar{T}$ deformation

- **universal**, \forall large c CFT



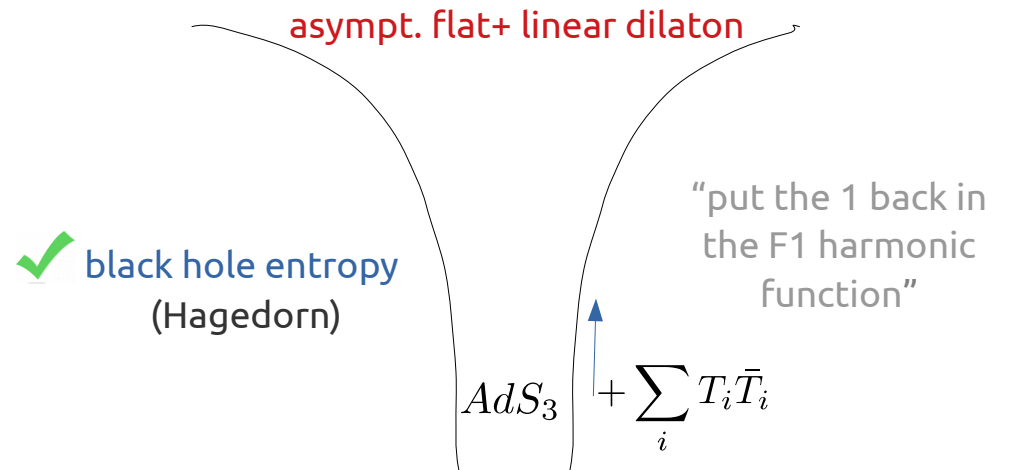
AdS_3 with **mixed** bnd. conditions at ∞

$\mu < 0 \approx$ **Dirichlet** at finite radius

McGough, Mezei, Verlinde

Single-trace $T\bar{T}$ deformation $\sum_{i=1}^p T_i \bar{T}_i$

- near horizon **NS5-F1** $\rightarrow \mathcal{M}^p/S_p$



Giveon, Itzhaki, Kutasov

Generalisations?

- **tractable single-trace** irrelevant flows with no UV fixed point?