

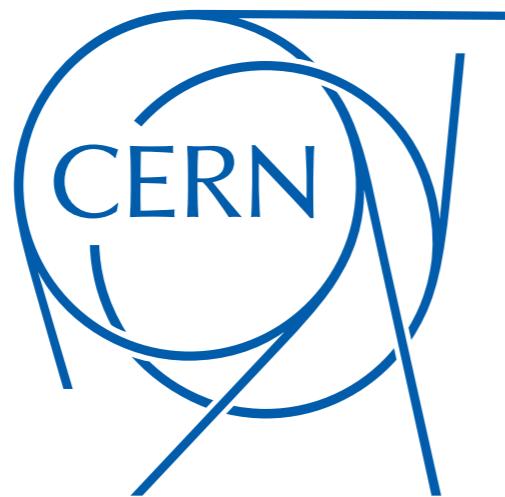
Gravitational Imprints of Flavour Hierarchies

Toby Opferkuch

CERN Theory Department

Based on 1910.02014 with Admir Greljo and Ben A. Stefanek

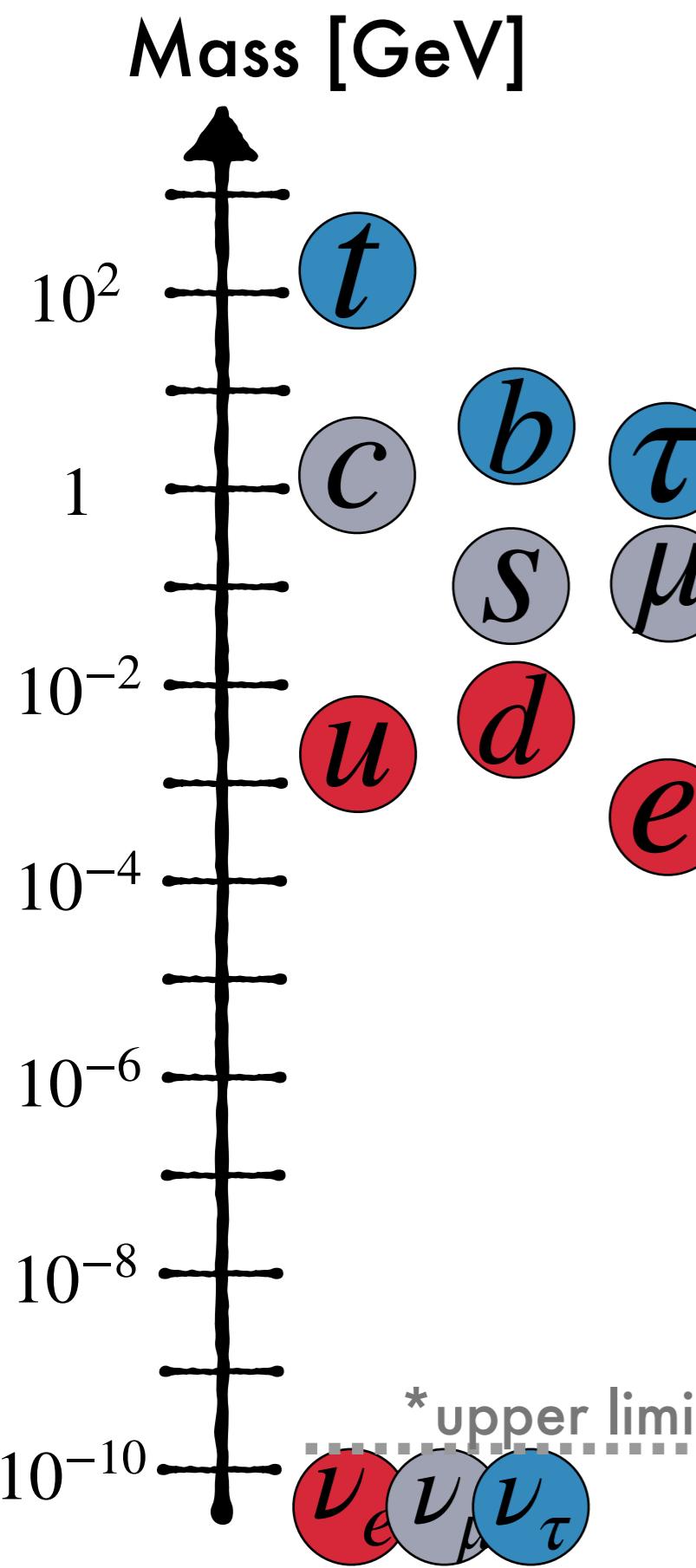
Challenges and Opportunities of High Frequency
Gravitational Wave Detection – 14th October 2019



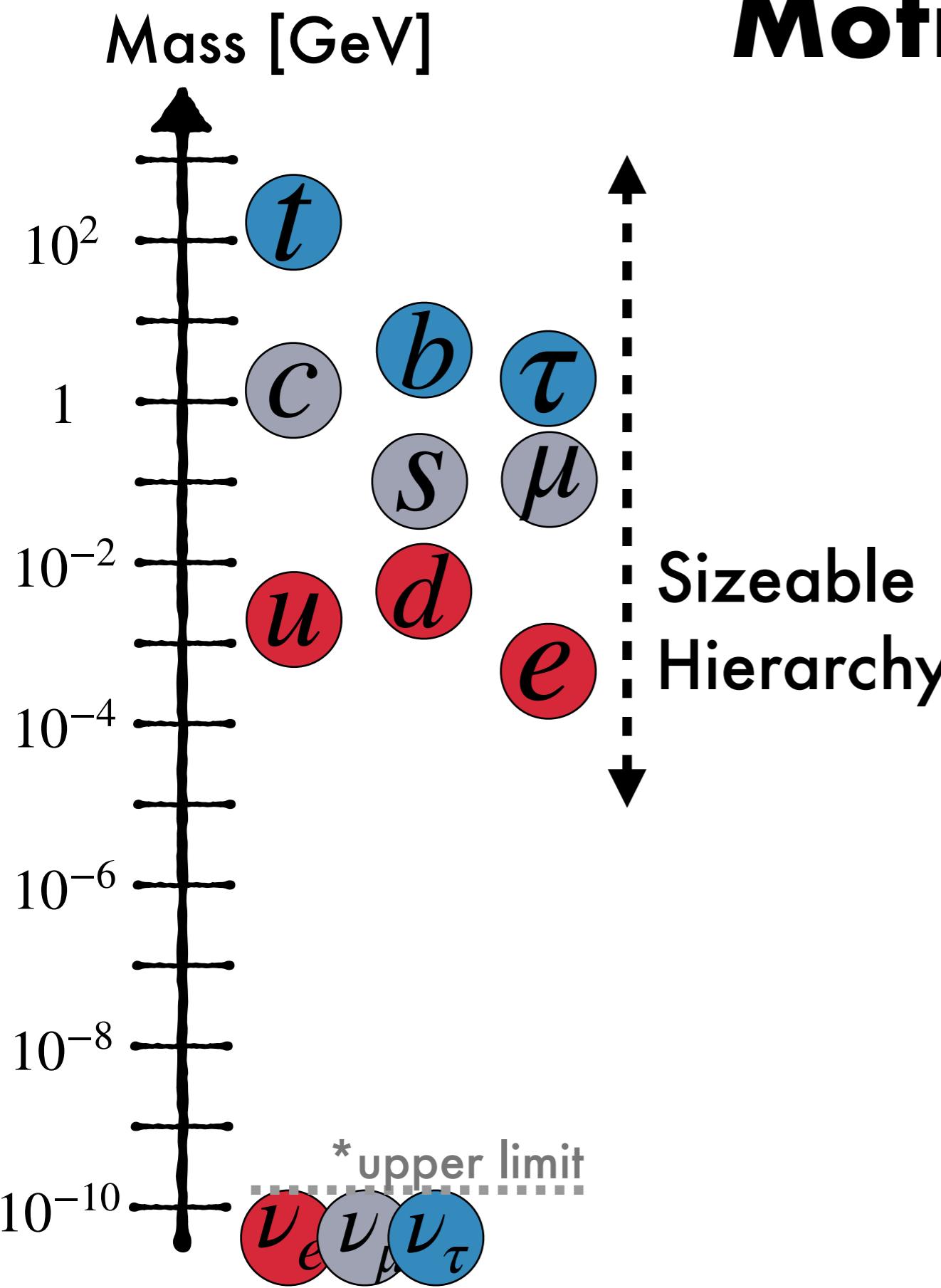
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Motivation



Motivation



In the Standard Model:

[Hierarchical complex numbers "Yukawas"]

$$Y = \frac{\text{Mass}}{\text{Scale}}$$

[Higgs vacuum expectation value (vev)]

Motivation

Flavour hierarchies



Additional (hierarchical)
symmetry breakings

Motivation

Flavour hierarchies



Additional (hierarchical)
symmetry breakings

New flavoured gauge symmetries:

[Additional scalar vev(s)]

$$Y_i = y \left(\frac{\langle \Phi \rangle}{\Lambda} \right)^{n_i}$$

Annotations:

- A black arrow points from the left towards the y term.
- A black arrow points from the bottom left towards the Λ term.
- A red arrow points from the bottom right towards the n_i term.

*[Current experimental B-anomalies
provide strong motivation for TeV-scale]*

[order-one
complex
numbers]

Motivation

- Stochastic gravitational wave background from 1st order phase transitions *[Peak frequency proportional to vev]*
- Plethora of ground- and space-based GW observatories in 2030s *[Experiments can probe far above the TeV scale]*
- Hierarchical range of optimised frequencies: $f_{\text{opt}} \sim L^{-1}$
- Interestingly: $\sqrt{m_t m_b} : \sqrt{m_s m_c} : \sqrt{m_u m_d}$
 $1 : 10^{-2} : 10^{-4}$
 $f_{\text{LISA}}^{-1} : \dots : f_{\text{ET}}^{-1}$

Model Example

- Expand Standard Model to $P \cong SU(4) \times SU(2) \times SU(2)$

**Quark-Lepton
Unification**

$$\Psi_L^i \equiv \begin{pmatrix} Q_L^i \\ L_L^i \end{pmatrix}$$

Quarks
Leptons

Model Example

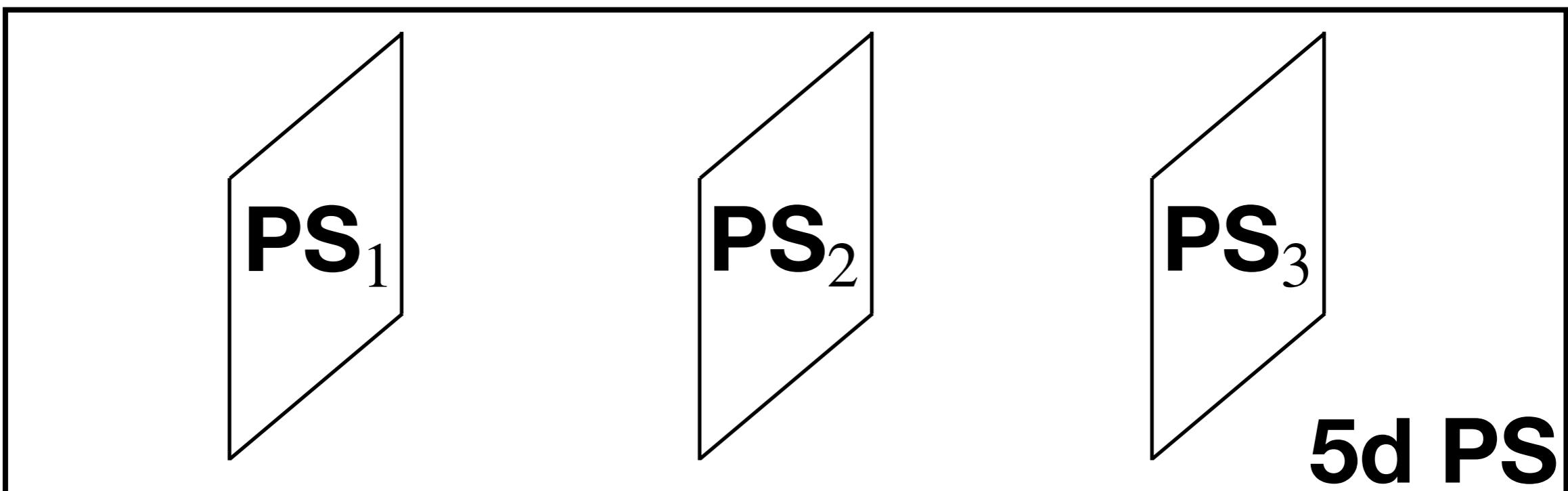
- Expand Standard Model to $PS \equiv SU(4) \times SU(2) \times SU(2)$

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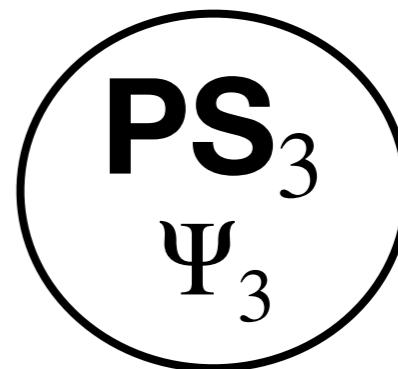
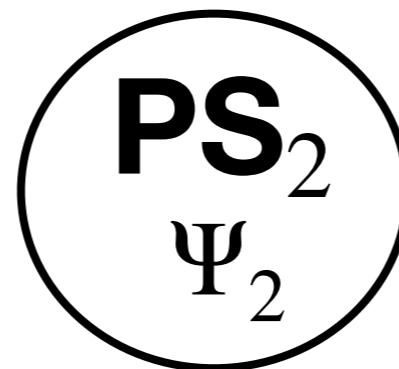
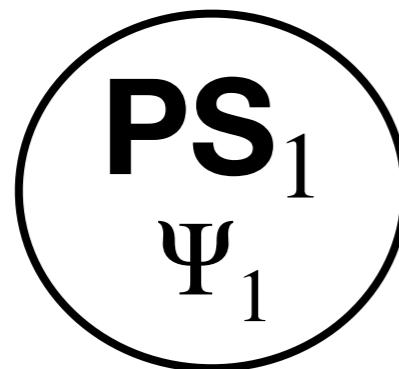
- Consider generation dependent higher-dimensional model



Model Example

- 4d Picture – “Pati-Salam Cubed”

[Bordone, Cornelle, Fuentes-Martin,
Isidori 1712.01368]



Fermions:

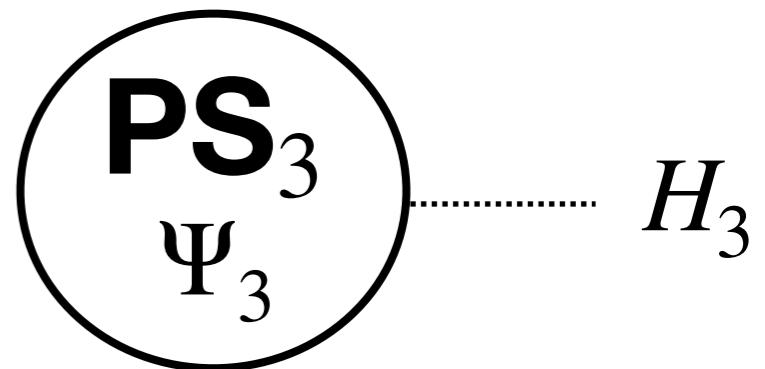
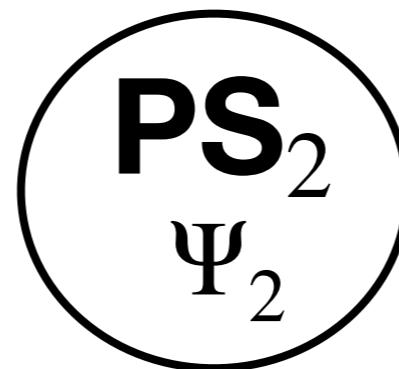
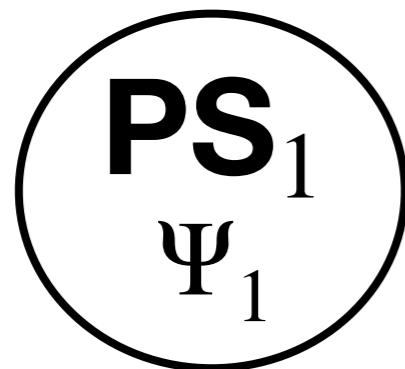
$$\Psi_L^i \equiv (4, 2, 1)_i$$

$$\Psi_R^i \equiv (4, 1, 2)_i$$

Model Example

- 4d Picture – “Pati-Salam Cubed”

[Bordone, Cornelle, Fuentes-Martin,
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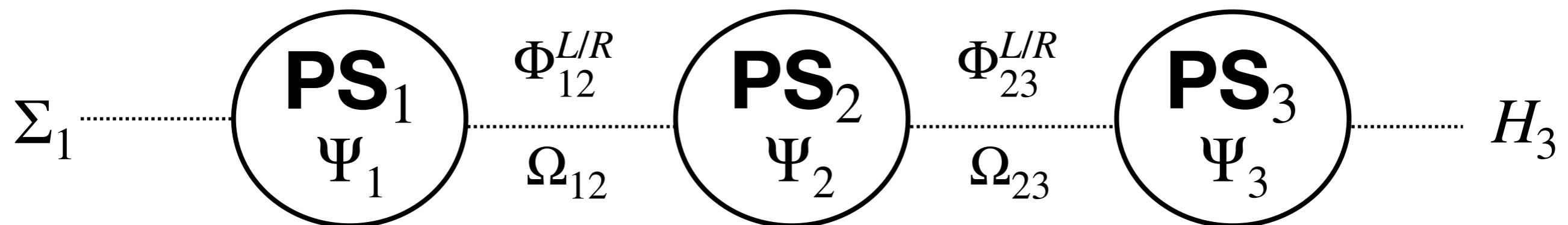
Scalars:

$$H_3 \equiv (1, 2, \bar{2})_3$$

Model Example

- 4d Picture – “Pati-Salam Cubed”

[Bordone, Cornelle, Fuentes-Martin,
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Fermions:

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$$\Psi_R^i \equiv (\mathbf{4}, \mathbf{1}, \mathbf{2})_i$$

Scalars:

$$H_3 \equiv (\mathbf{1}, \mathbf{2}, \bar{\mathbf{2}})_3$$

$$\Sigma_1 \equiv (\mathbf{4}, \mathbf{1}, \mathbf{2})_1$$

$$\Phi_{ij}^L \equiv (\mathbf{1}, \mathbf{2}, \mathbf{1})_i \times (\mathbf{1}, \bar{\mathbf{2}}, \mathbf{1})_j$$

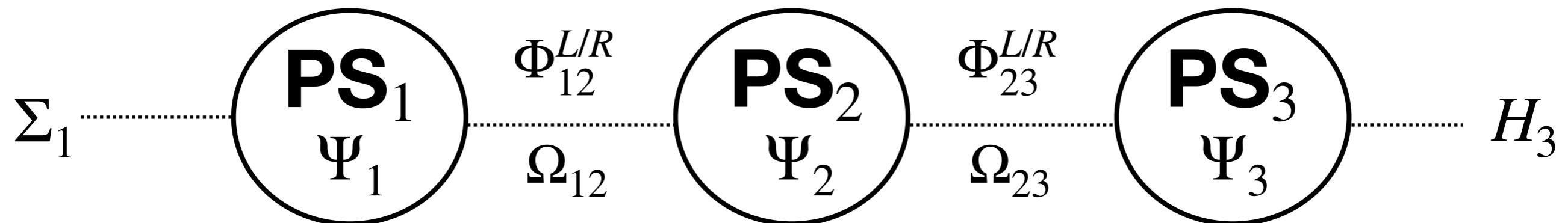
$$\Phi_{ij}^R \equiv (\mathbf{1}, \mathbf{1}, \mathbf{2})_i \times (\mathbf{1}, \mathbf{1}, \bar{\mathbf{2}})_j$$

$$\Omega_{ij} \equiv (\mathbf{4}, \mathbf{2}, \mathbf{1})_i \times (\bar{\mathbf{4}}, \bar{\mathbf{2}}, \mathbf{1})_j$$

Model Example

- 4d Picture – “Pati-Salam Cubed”

[Bordone, Cornelle, Fuentes-Martin, Isidori 1712.01368]



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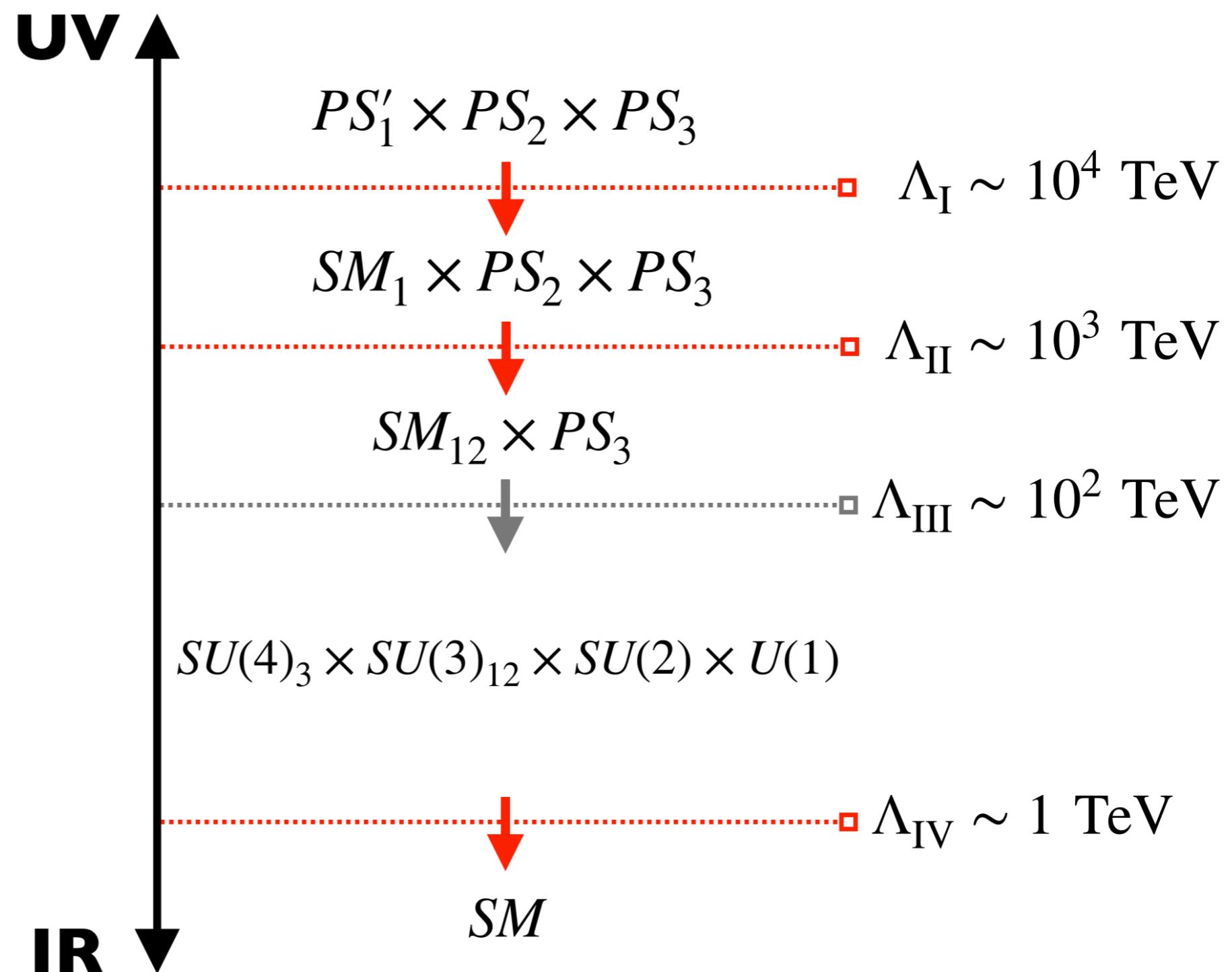
$$\Phi_{ij}^R \equiv (1, 1, 2)_i \times (1, 1, \bar{2})_j$$

$$\Omega_{ij} \equiv (4, 2, 1)_i \times (\bar{4}, \bar{2}, 1)_j$$

*Extra-dimensional
picture implies
suppressed quartics*

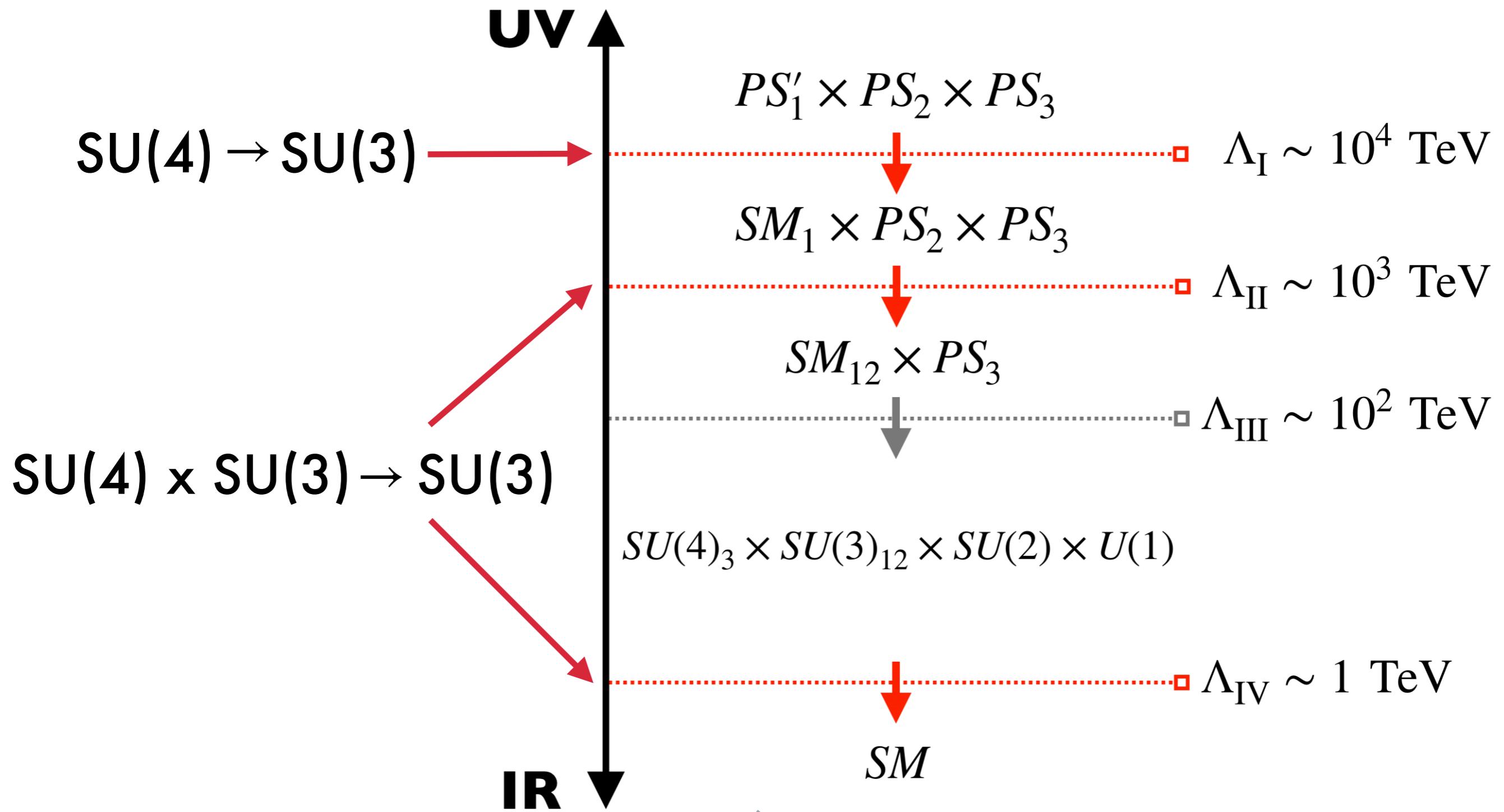
Model Example

- Relevant phase transitions:



Model Example

- Relevant phase transitions:



Cosmological Phase Transitions

Phase transition dynamics
controlled by effective potential:

$$V_{\text{eff}}(\underline{g}, \underline{\lambda}, \underline{v}, \phi, T) = V_0 + V_{\text{CW}} + V_{T \neq 0}$$

[fundamental parameters] *[temperature]*

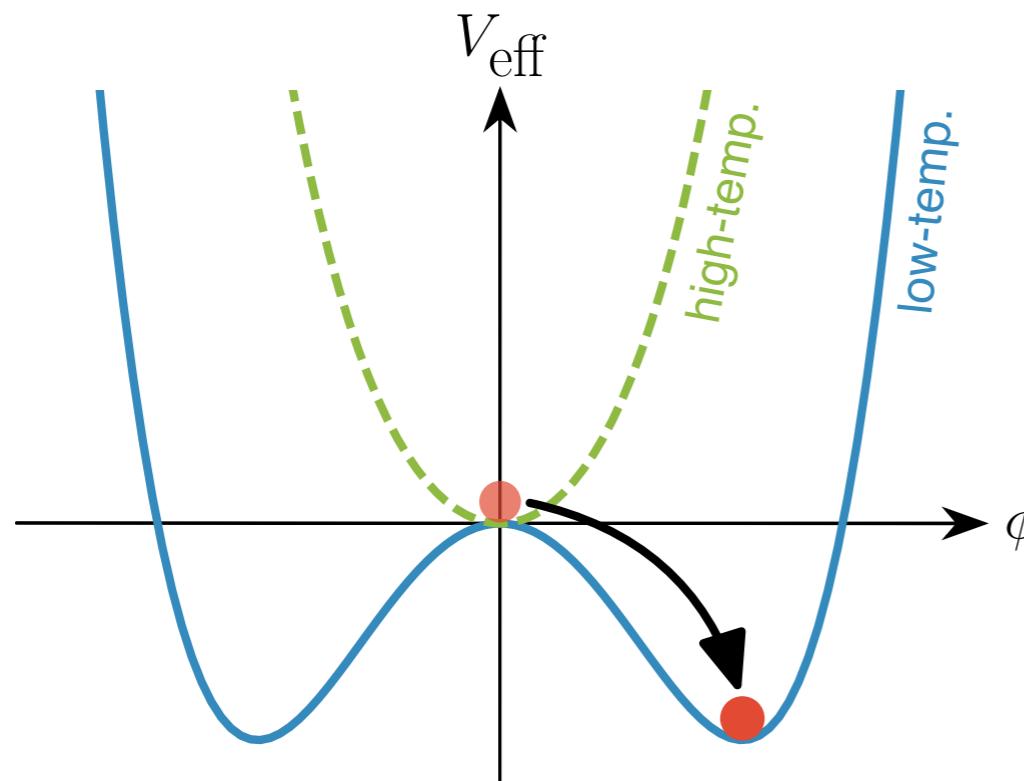
Cosmological Phase Transitions

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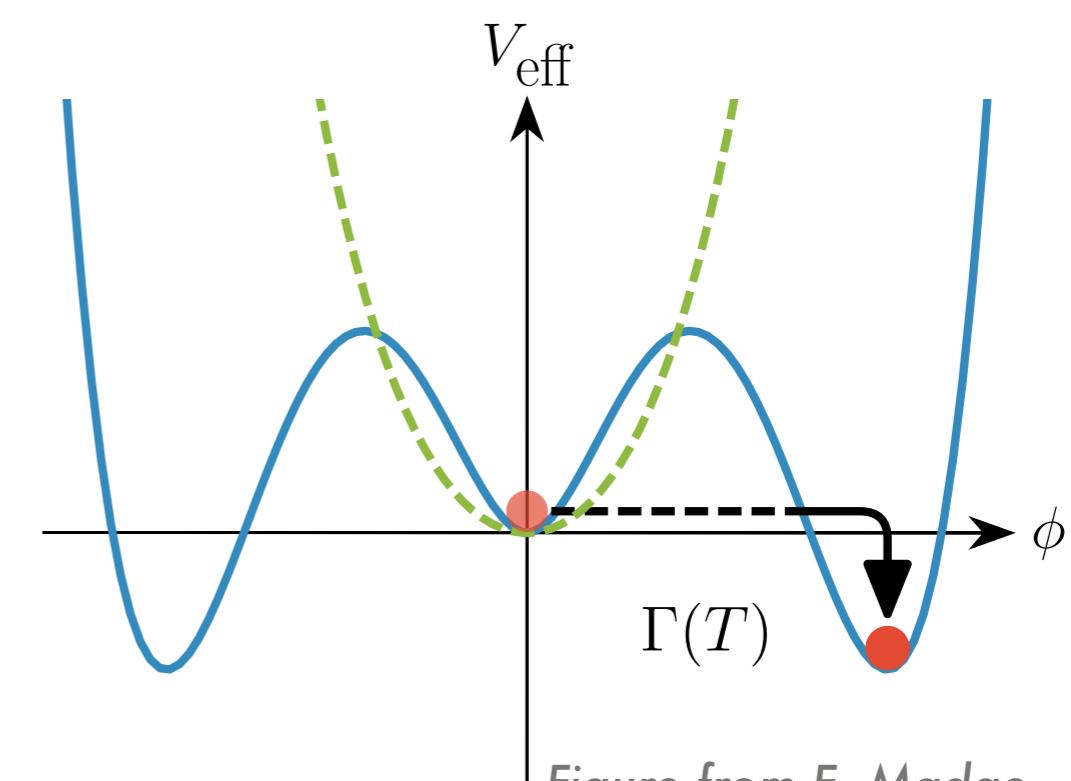
$$V_{\text{eff}}(g, \lambda, v, \phi, T) = V_0 + V_{\text{CW}} + V_T \neq 0$$

[fundamental parameters] [temperature]

Cross-over



1st-order



Cosmological Phase Transitions

- Nucleation temperature:

$$\Gamma(T_n) \sim H^4(T_n)$$

- Contributions to the GW spectrum:

Ω_ϕ : Collision of bubble walls

Ω_{sw} : Sound waves in plasma

Ω_{turb} : Turbulence in plasma

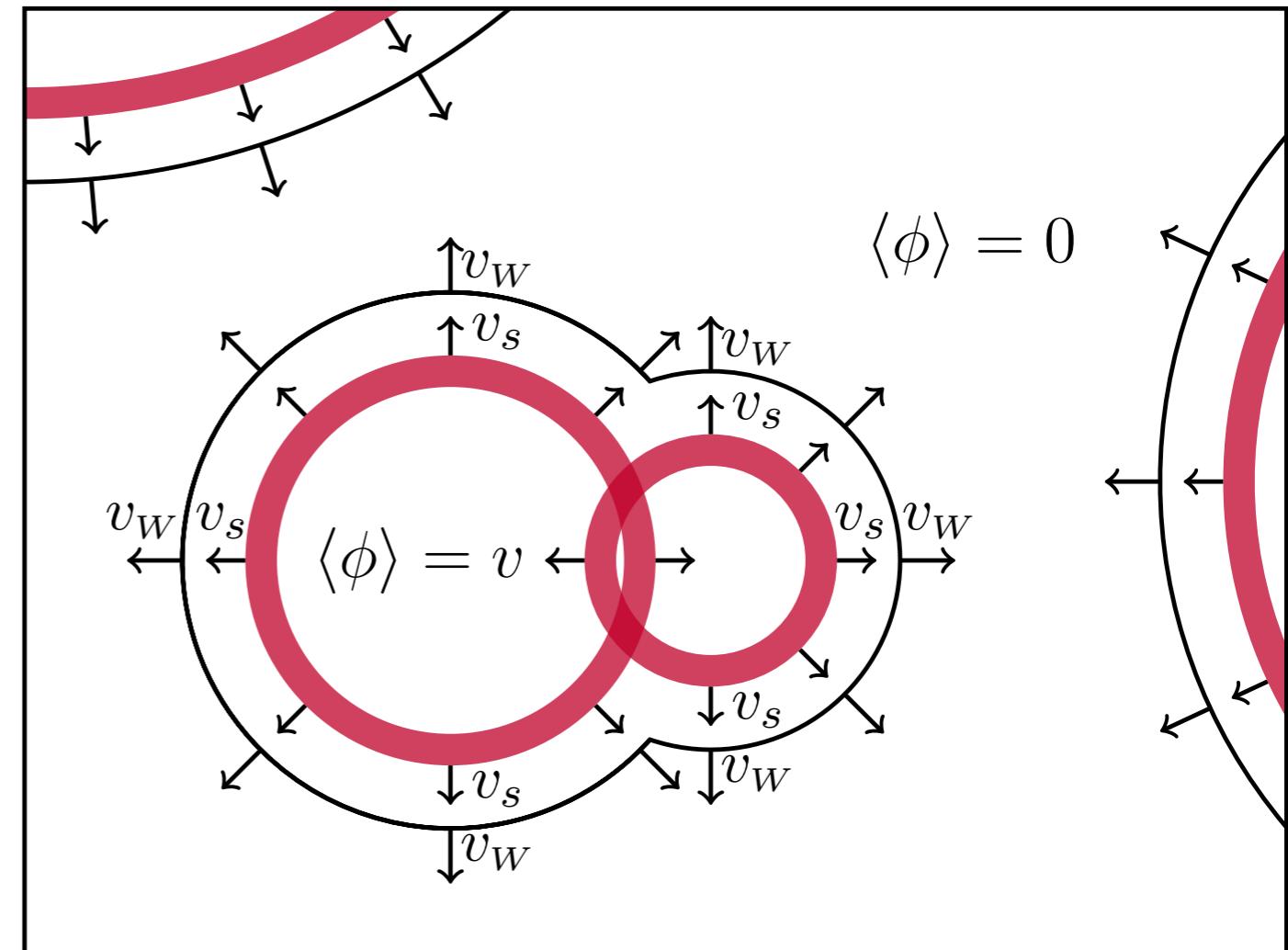


Figure from E. Madge

See for example:

[Huber, Konstandin 0806.1828]

[Hindmarsh, Huber, Rummukainen, Weir 1504.03291]

[Hindmarsh, Huber, Rummukainen, Weir 1704.05871]

[Cutting, Hindmarsh, Weir 1802.05712]

[Pol et. al. 1903.08585]

[Cutting, Hindmarsh, Weir 1906.00480]

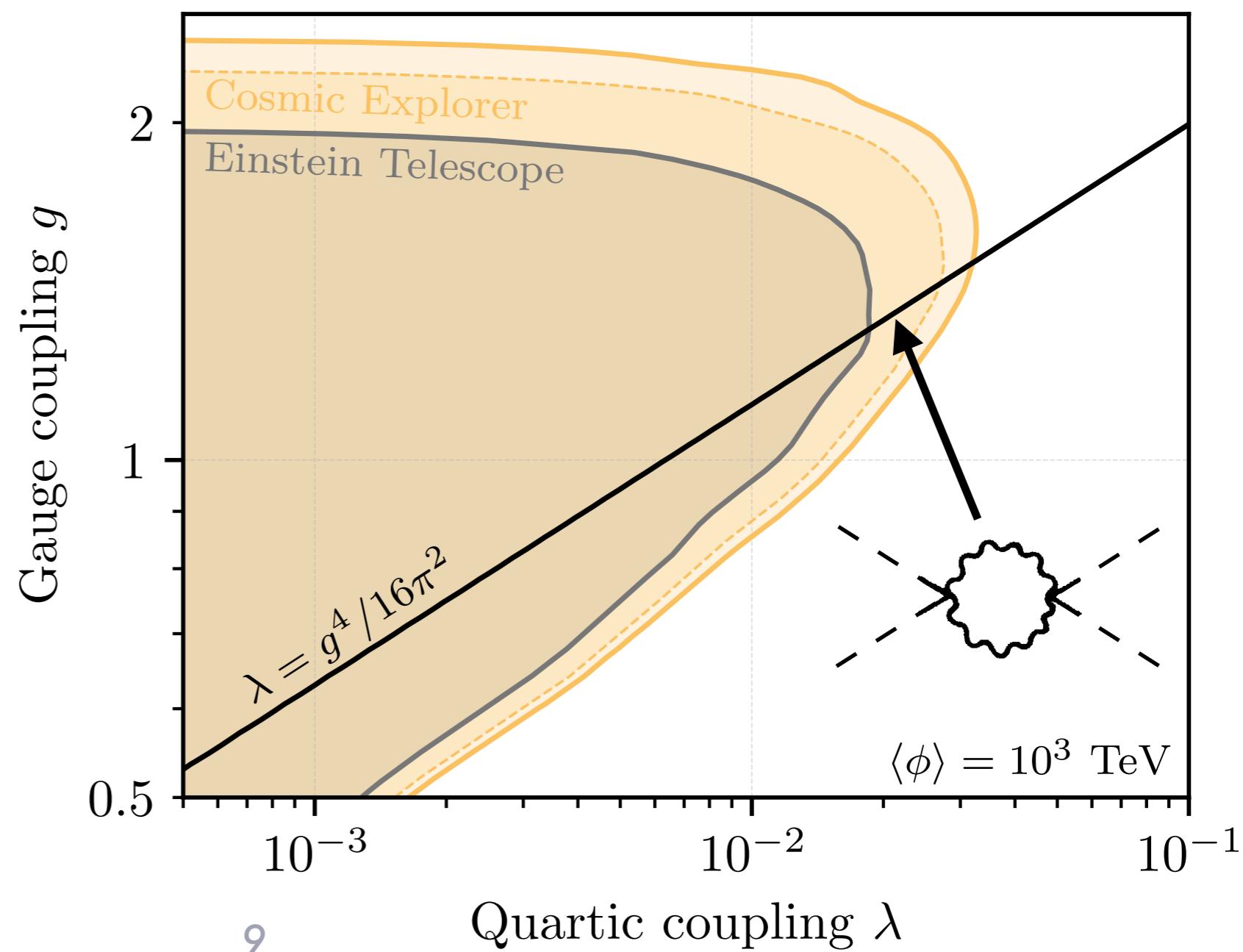
Example Transition

- $SU(4) \rightarrow SU(3)$ transition [relevant for Λ_I]

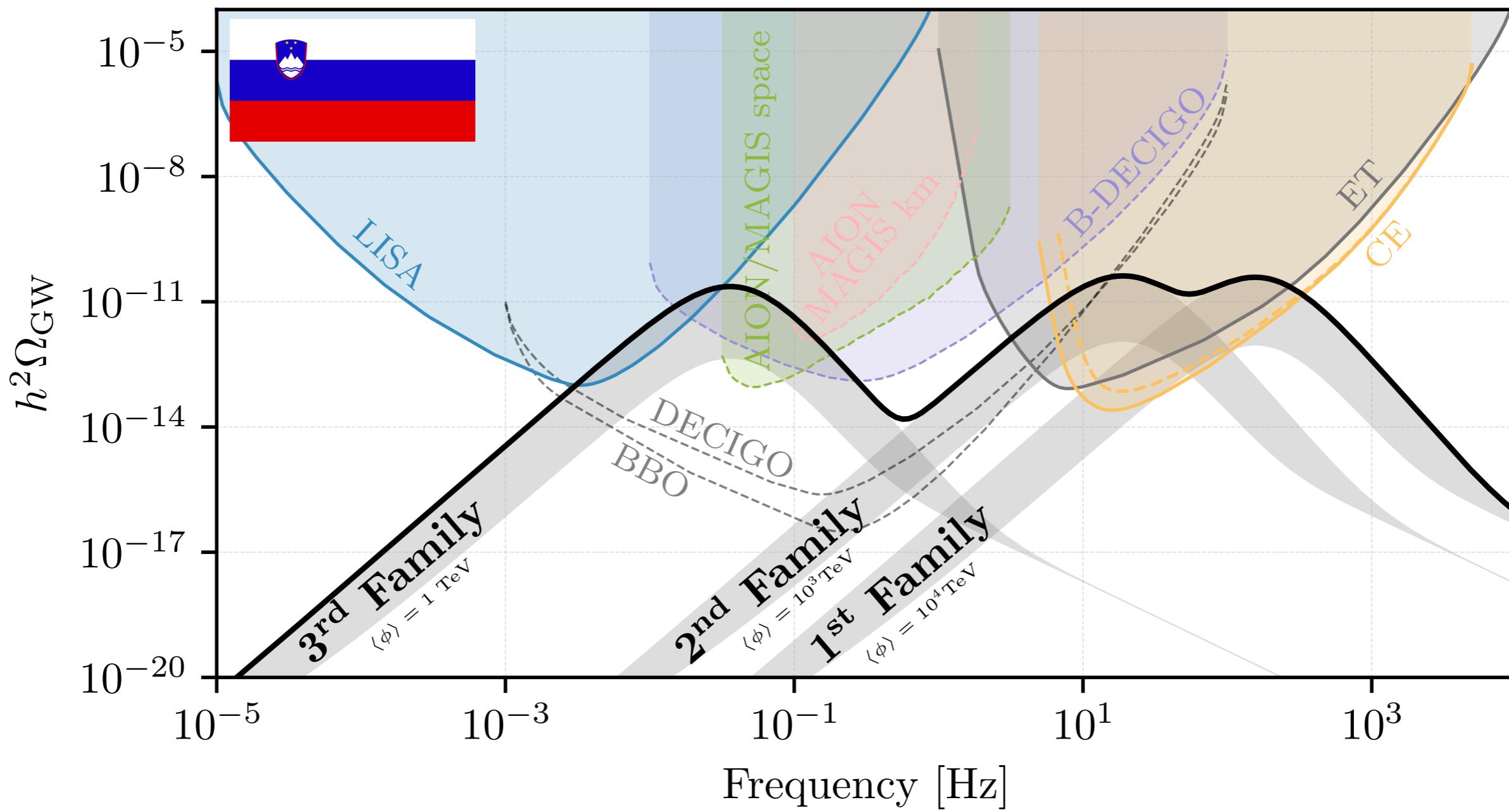
[similar transition for breaking with a 15 of $SU(4)$ considered in Croon et. al. 1812.02747]

- SM QCD coupling requires $g \sim 1$

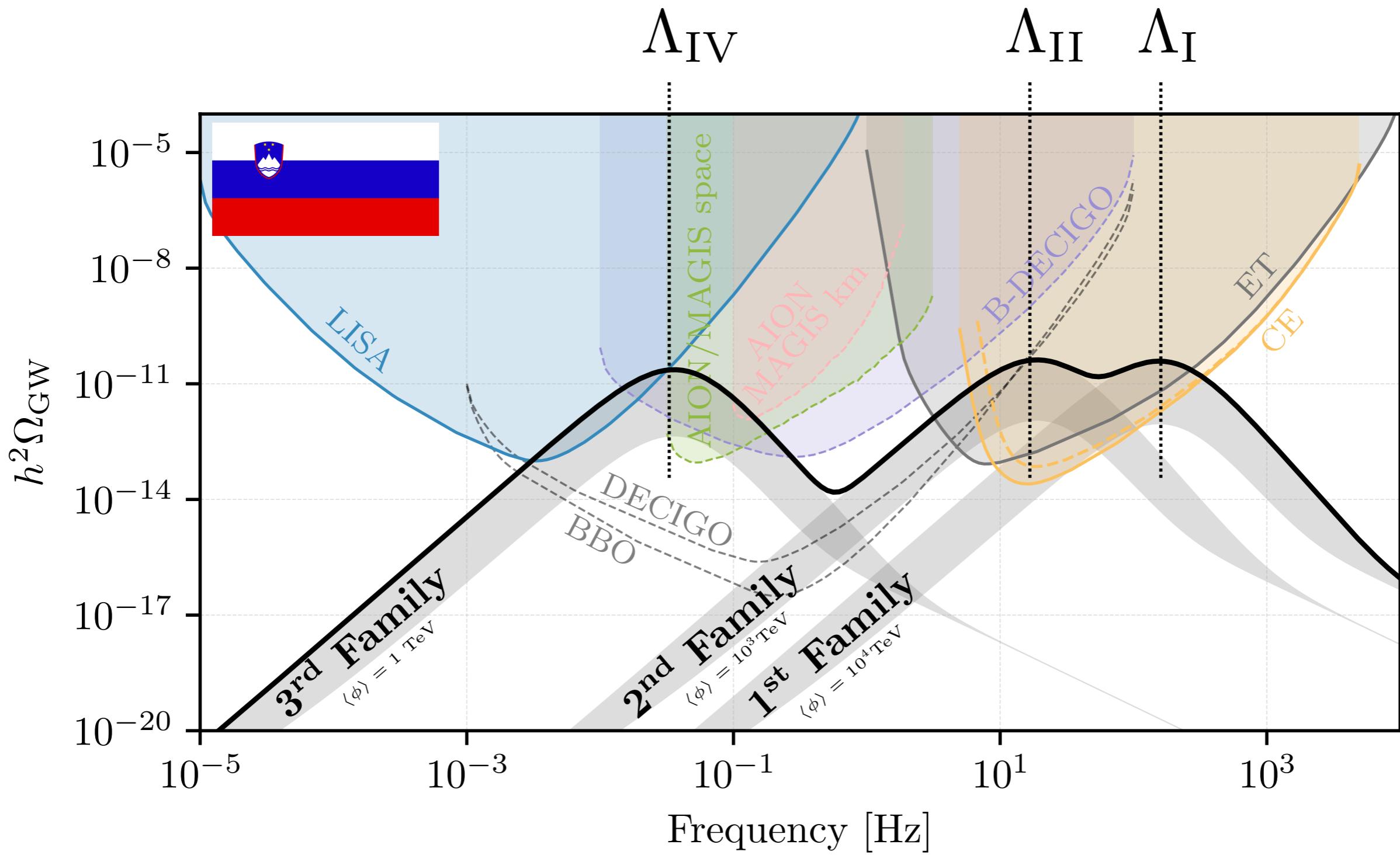
Detectability over natural range of expected parameters



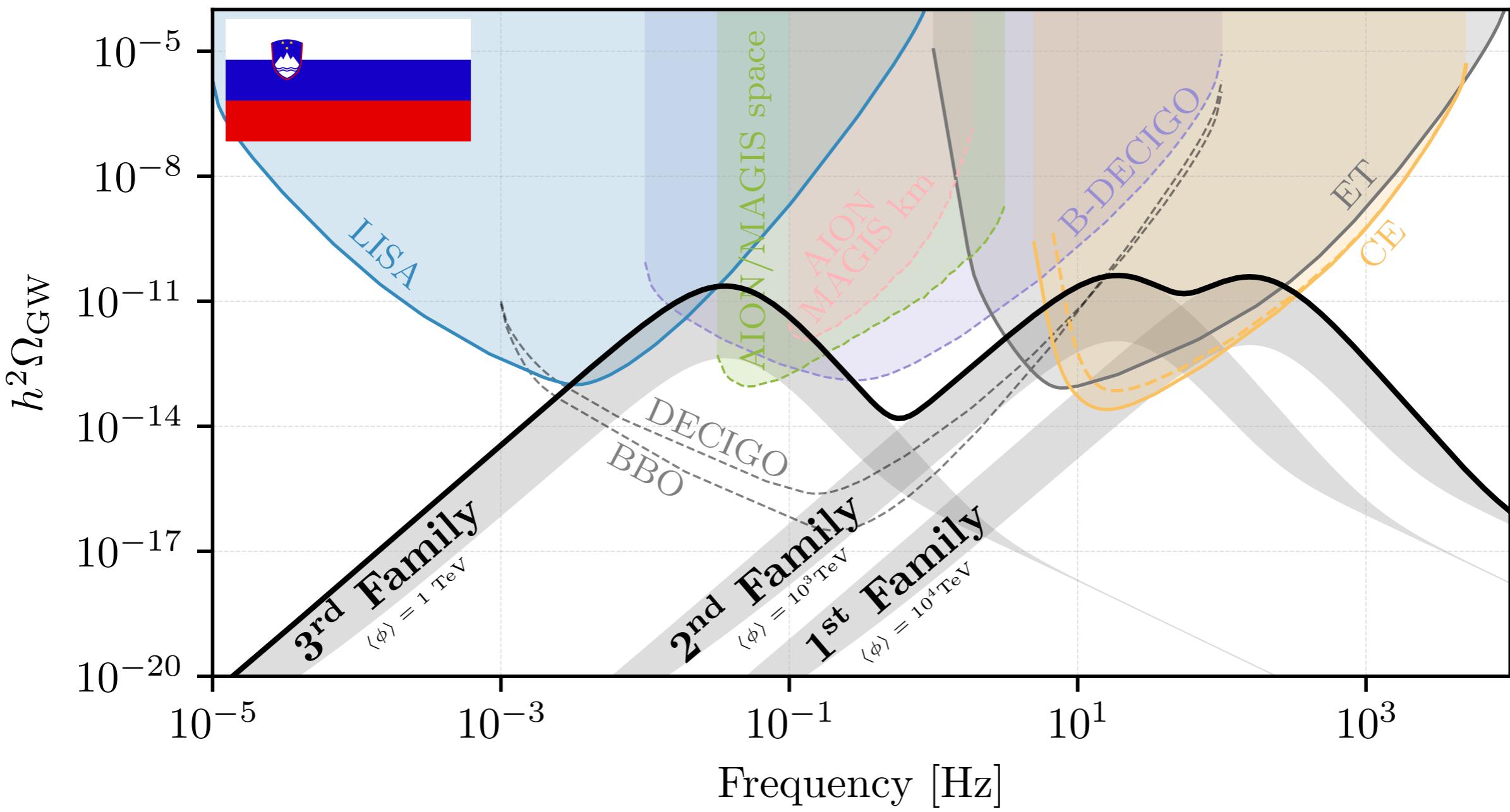
“Triglav Signature”



“Triglav Signature”

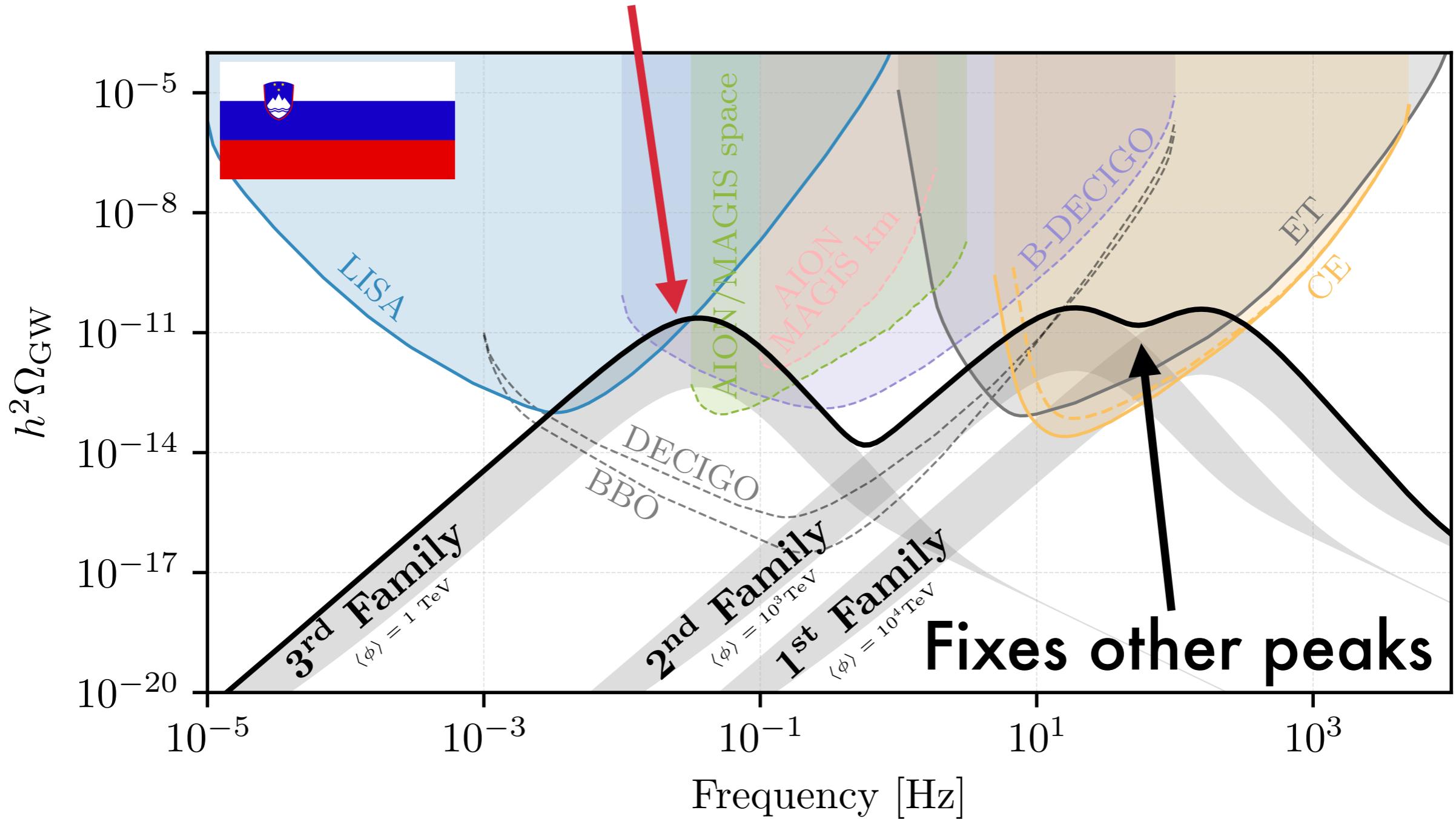


“Triglav Signature”

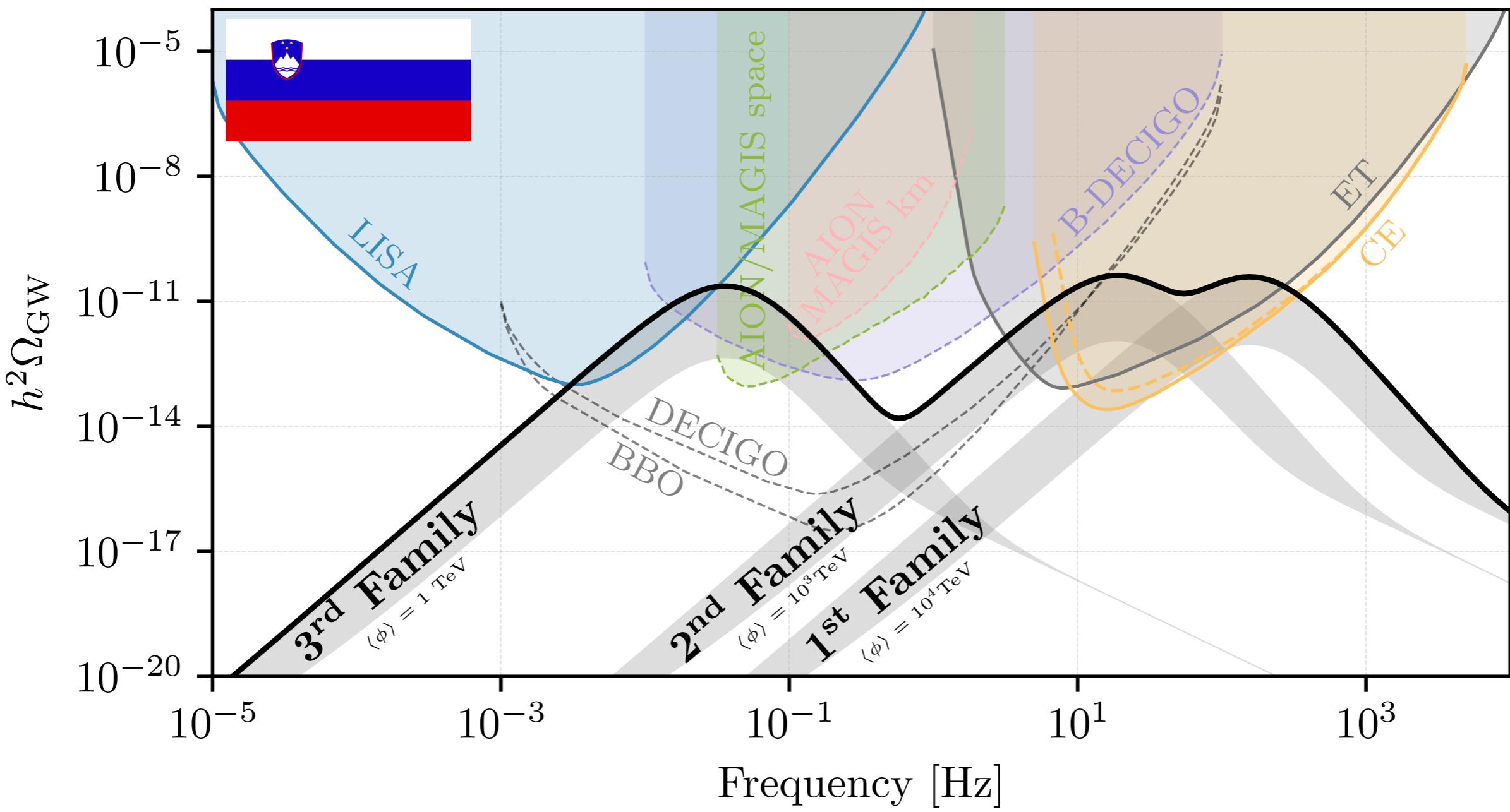


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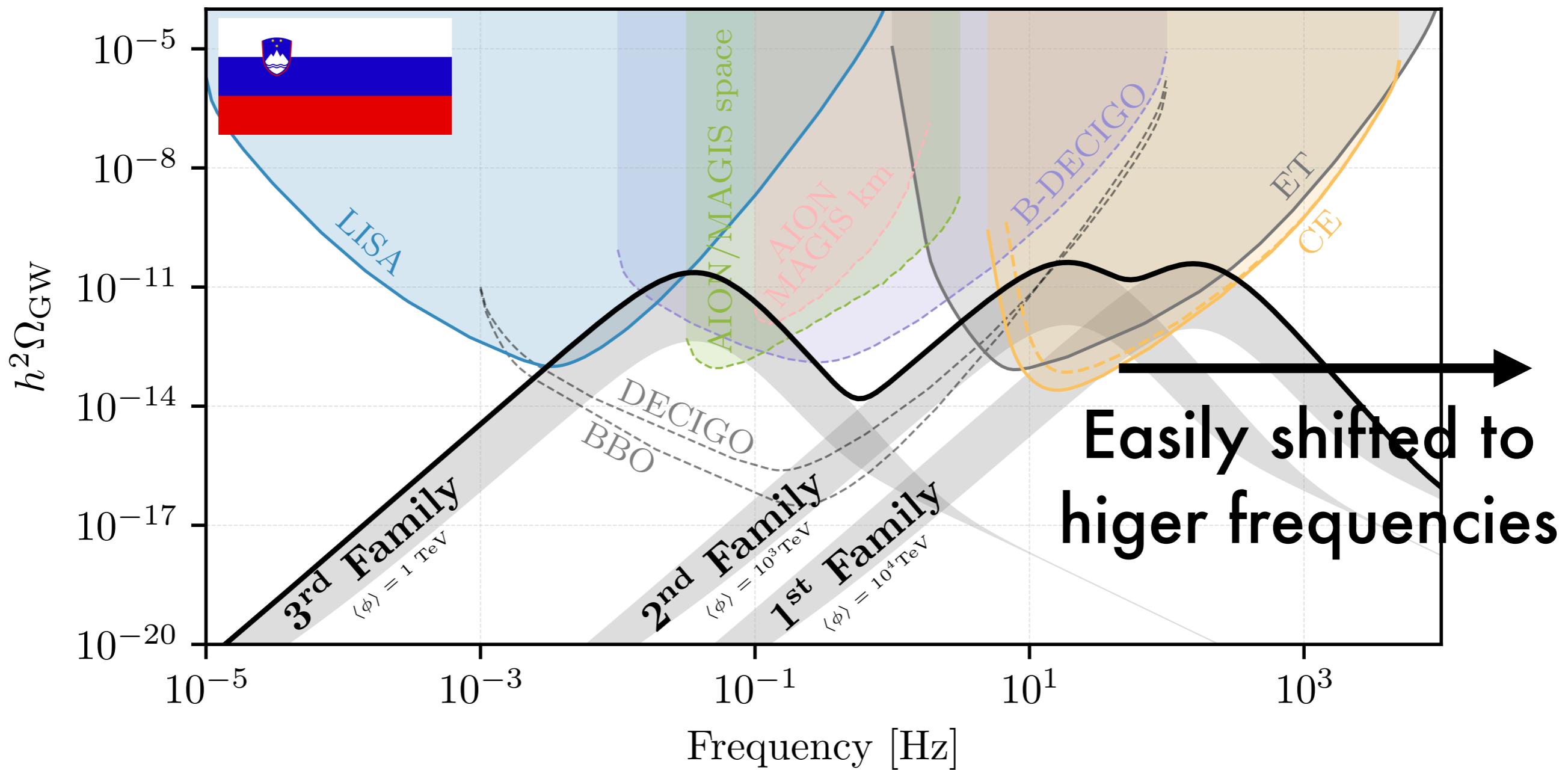
$\Lambda_{\text{IV}} \sim 1 \text{ TeV}$ B-anomalies motived



“Triglav Signature”



“Triglav Signature”



Conclusions

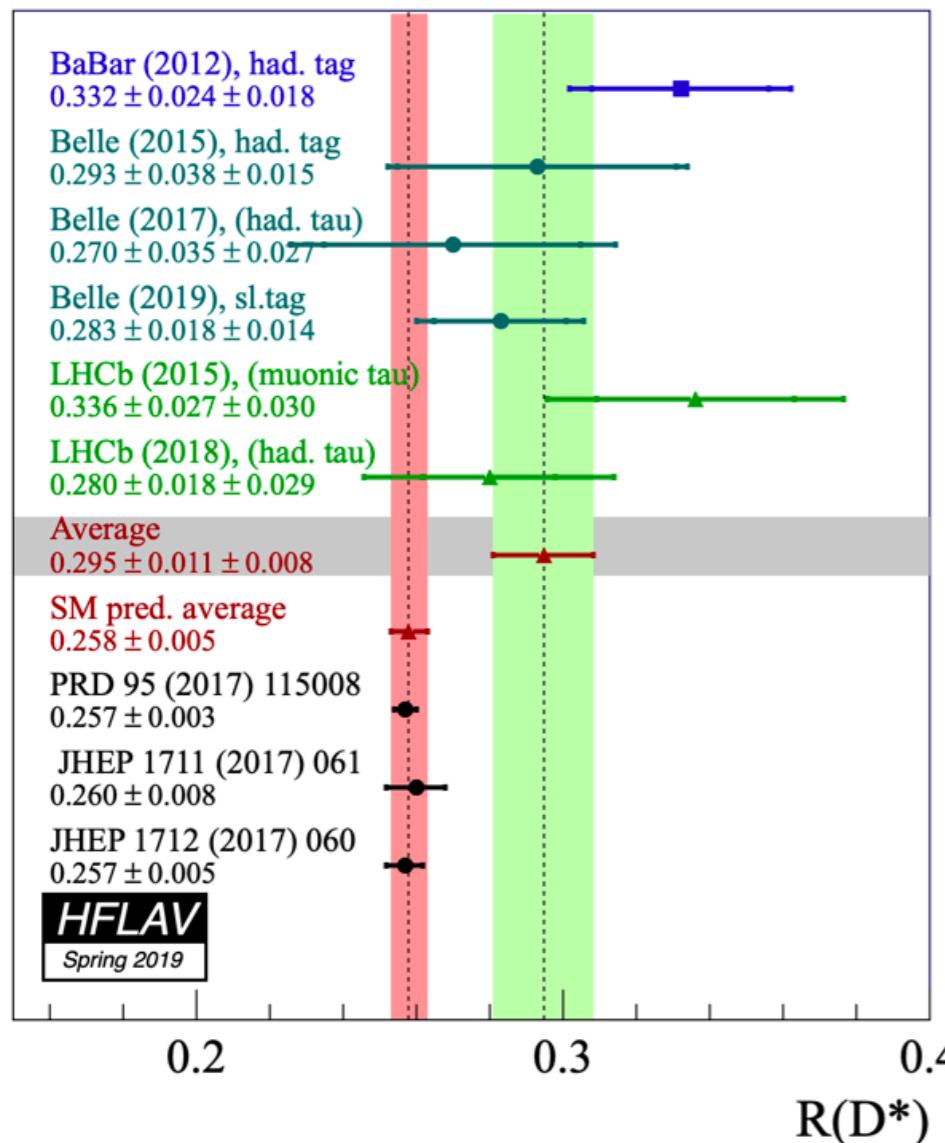
- Rich stochastic GW background from phase transitions
[Sensitive to phase transitions far beyond collider reach]
- Multi-peaked spectrum arising from the flavour hierarchy
[Mass hierarchy encoded in relative peak frequencies]
- These peaks may lie above planned future experiments
[Clear appeal for high frequency GW detectors!]

Backup Slides

B-Anomalies Status

$b \rightarrow c\ell\bar{\nu}$ transitions:

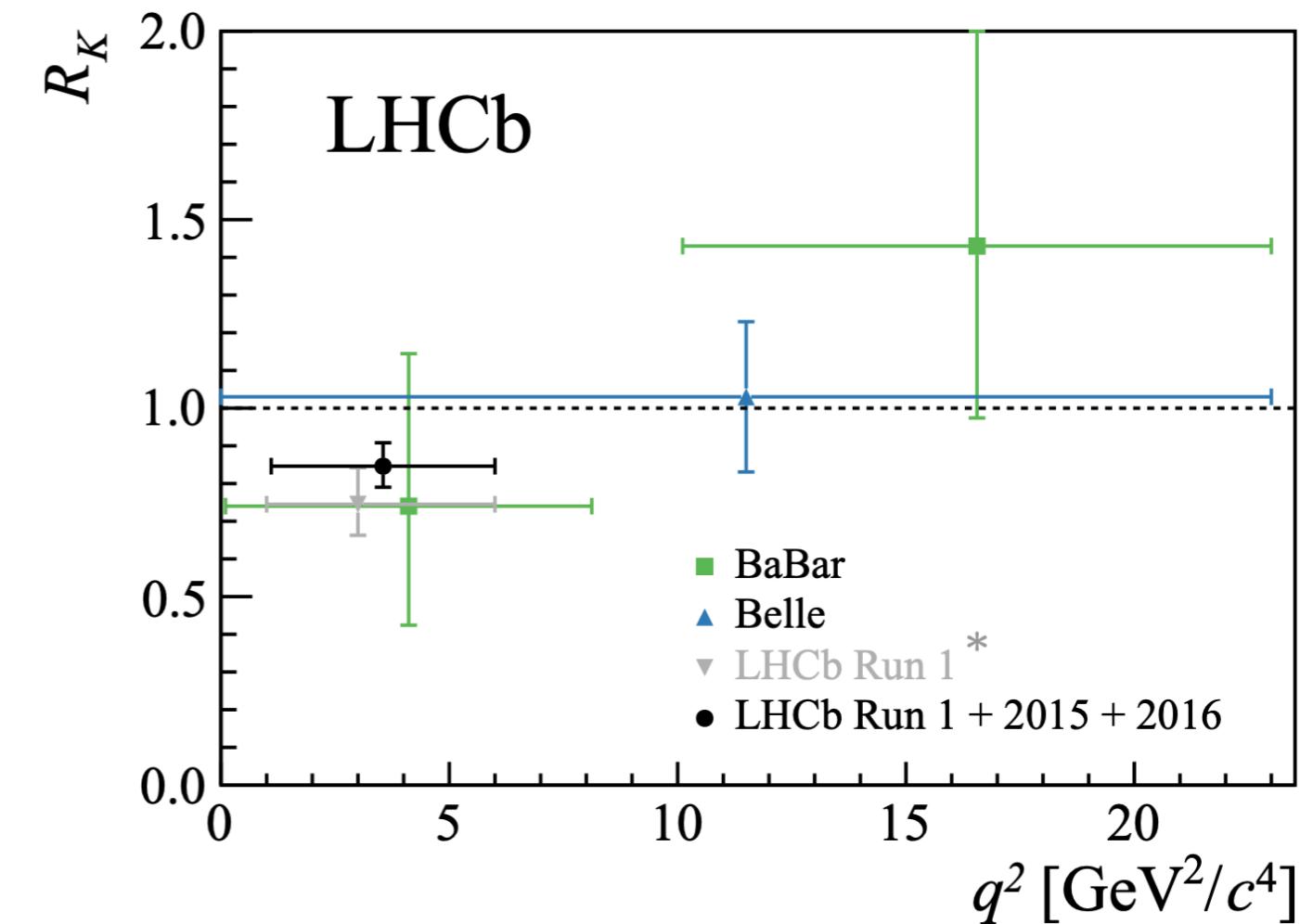
$$R(D^{(*)}) = \frac{\text{Br}(B \rightarrow D^{(*)}\tau\nu_\tau)}{\text{Br}(B \rightarrow D^{(*)}\ell\nu_\ell)}$$



3.1σ discrepancy

$b \rightarrow s\ell\ell$ transitions:

$$R(K^{(*)}) = \frac{\text{Br}(B \rightarrow K^{(*)}\mu^+\mu^-)}{\text{Br}(B \rightarrow K^{(*)}e^+e^-)}$$

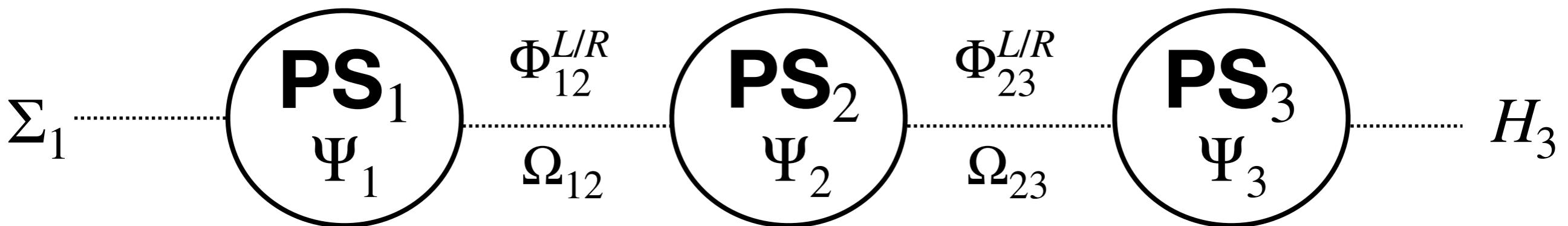


2.5σ discrepancy

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[Bordone, Cornelle, Fuentes-Martin, Isidori 1712.01368]



- Light Fermion interactions through mixing:

$$\mathcal{L}_{23} = \frac{1}{\Lambda_{\text{III}}} \bar{\Psi}_L^{(2)} \Omega_{23} H_3 \Psi_R^{(3)} + \text{h.c.} \quad \rightarrow \quad |V_{ts}| \sim \frac{\langle \Omega_{23} \rangle}{\Lambda_{\text{III}}} \sim \frac{\Lambda_{\text{IV}}}{\Lambda_{\text{III}}}$$

$$\mathcal{L}_{12} = \frac{1}{\Lambda_{\text{II}}^2} \bar{\Psi}_L^{(k)} \Phi_{k3}^L H_3 \Phi_{3l}^R \Psi_R^{(l)} + \text{h.c.} \quad \rightarrow \quad Y_c \sim \frac{\langle \Phi_{23}^L \rangle \langle \Phi_{32}^R \rangle}{\Lambda_{\text{II}}^2} \sim \frac{\Lambda_{\text{III}}^2}{\Lambda_{\text{II}}^2}$$

See for example:
[Greljo, Stefanek 1802.04274]

Strong coupling running

