

Effect of scintillation on GNSS-based TEC metrology

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"There is Physics behind every measurement technique"

Anonymous





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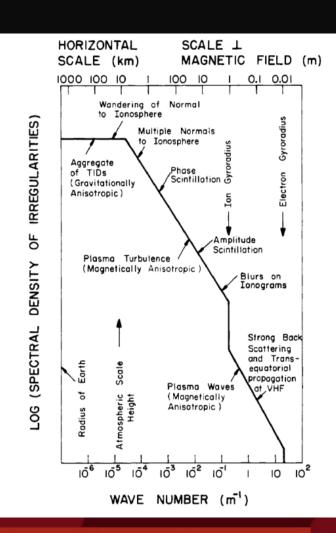
Outline

- Introduction and background
- GNSS metrology pertaining to TEC estimation
- The issue
- Concluding remarks



The ionosphere

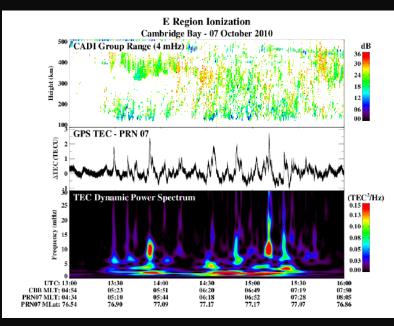
- Region of ionized plasma above 60 km from the Earth's surface
- Primarily produced by the solar EUV and particle precipitation
- Temporal variation scales from "milliseconds" to solar cycle
- Spatial variations from cm to few thousand km
- Affects the propagation of radio waves
- Significant impact on navigation and communication systems

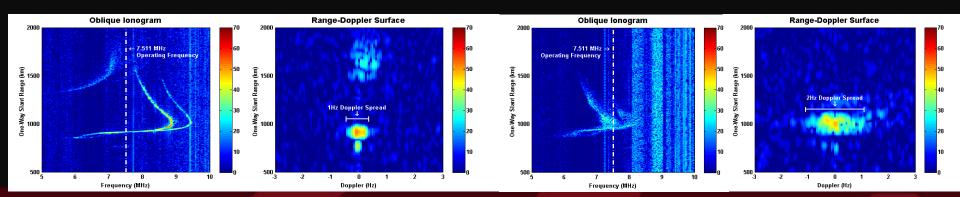




The ionosphere

- Refraction and dispersion-Deterministic
 - TEC variations
 - HF Propagation
- Diffraction and scattering-Stochastic
 - Scintillation
 - Coherent backscatter







Sources or errors in GNSS-Correctable/Manageable

- satellite:
 - orbit
 - clock
 - instrumental delays

~5 m

- signal path
 - ionosphere 5-40 m
 - troposphere 50 cm 2 m
 - multipath 1-2 m

- receiver
 - clock

- 1-2 m
- instrumental delays
- other
 - spoofing

- 1-2 m
- interference

Total: 14 - 55 m

	Range Delay					
TEC	S-Band	L-Band	<u>UHF</u>	VHF	Elev	Mapping Function
50	2.4 m	12 m	104 m	787 m	90 °	x 1
110	5.1 m	26 m	223 m	1.7 km	20 °	x 2.12



A world of applications

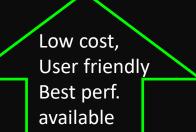
Safety of life

- Transportation (air, rail, and water)
- Public safety (Police, fire, ambulance....)
- Traffic
- ADAS
- SatCom
- Military
-

Integrity (error-free)
Standards, regulation,
continuity, availability,
accuracy

Mass market

- Personal communication and navigation
- Recreation
- LCVs
- Trucks and buses
- IoT and LBS
-



Professional

High precision High accuracy High reliability

- Oil and natural Gas
- Mining
- Timing
- Surveying
- Geodesy
- Agriculture
- SatCom

- IoT and LBS
- Vehicle control and robotics
- Construction
- Meteorology and weather forecast
- Fisheries
- Environmental monitoring
- •



Scintillation

- Rapid random fluctuations of the amplitude and phase of a trans-ionospheric radio signal
- Caused by Fresnel scale irregularities in the ionosphere
- Earlier observation using signal from radio stars
- Early studies using geo-stationary satellite signals
- Morphological characteristics are well known
- Two regions of <u>intense</u> scintillations Polar and equatorial regions (just stating this for historical reasons). This is outdated as 'scintillation' is not intense in the high-latitudes
- Produce stochastic errors in PNT applications



Disclaimer-Scintillation

Fluctuations in the amplitude of the signal

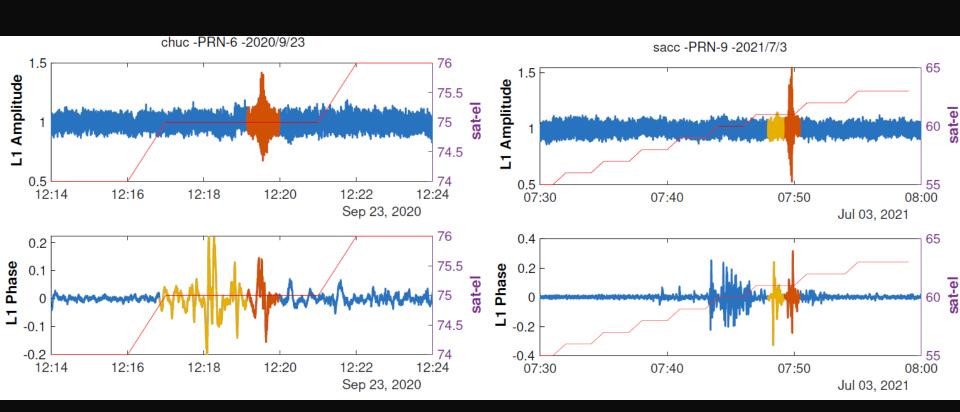
Variations in phase that does not follow 1/f² dependence (if it does, it is called TEC)

Fresnel scale irregularities





Scintillation

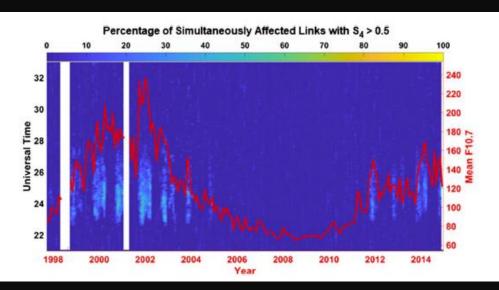


Song, et al., 2022



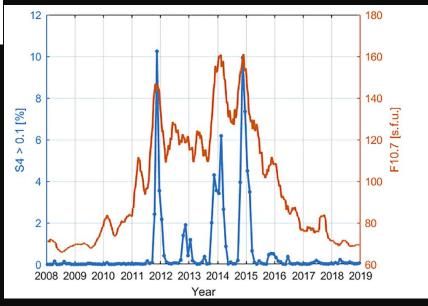


Scintillation occurrence – Tale of two regions



Equatorial region – Salles, et al., 2021

In term of probability of occurrence, Scintillation is not a major concern for the polar region



Polar region – Meziane, et al., 2019



Probability Distribution Function (PDF) -**Moments**

Zero moment – Normalization

$$1 = \int_{-\infty}^{+\infty} f(u) \, du$$

First moment – Mean μ

$$\mu = \int_{-\infty}^{+\infty} u f(u) \, du$$

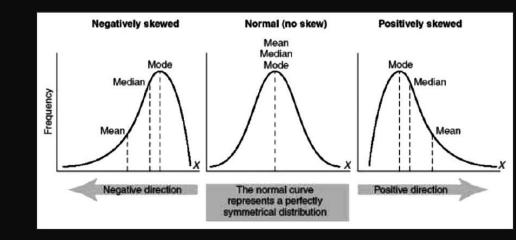
Second moment – Variance σ^2

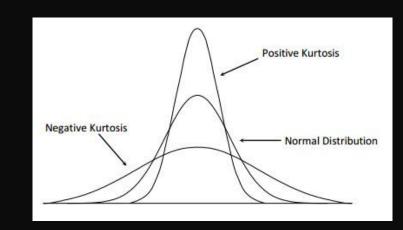
$$\sigma^2 = \int_{-\infty}^{+\infty} (u - \mu)^2 f(u) \, du$$

[Standard] Third moment – Skewness S

$$S = \frac{1}{\sigma^3} \int_{-\infty}^{+\infty} (u - \mu)^3 f(u) \, du \quad [S_G = 0]$$

[Standard] Fourth moment – Kurtosis K







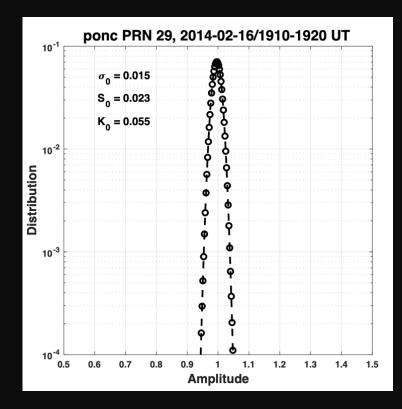
UNB



Which distribution?

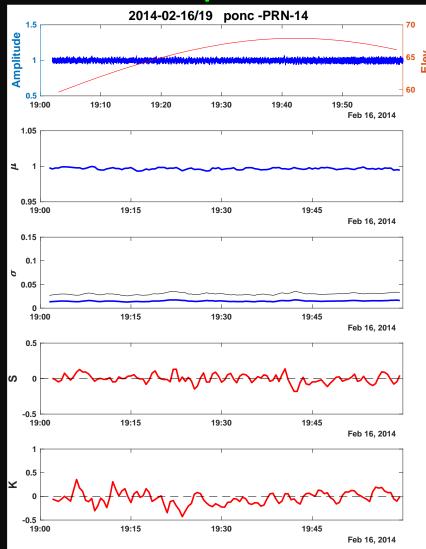
- Normal [Rino et al., 1976; ...]
- Log-normal [Wheelon, 2003; ...]
- Nakagami [Fremouv et al., 1980]
- Rician [Mercer, 1962]
- α-μ distribution [Moraes *et al.*, 2014]
- Generalized Gaussian [?]





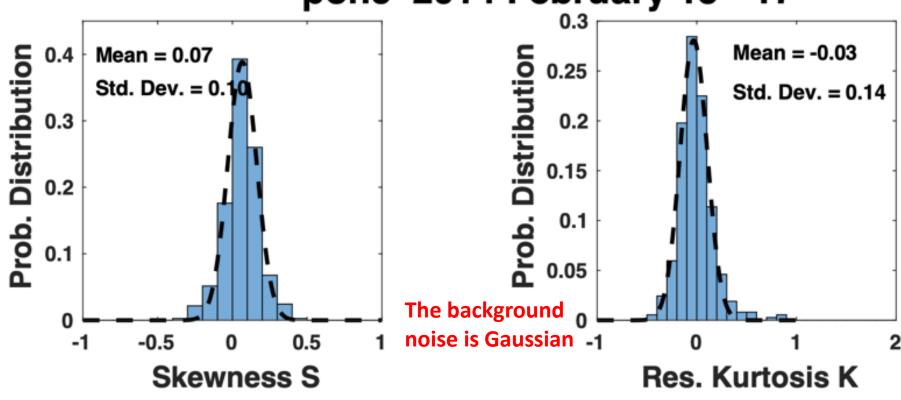
Best Gaussian fit

An example - Noise





ponc 2014 February 15 - 17



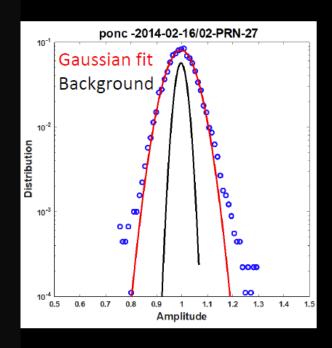
$$S_{Rayleigh} = 0.63$$

$$K_{Rayleigh} = 0.25$$

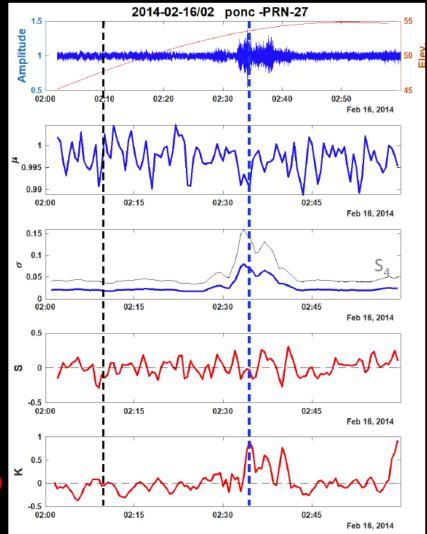




An Example - Scintillation

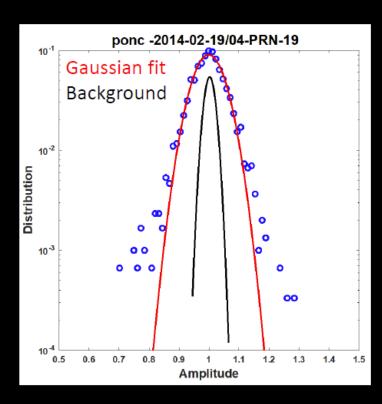


Skewness S ≈ 0 Kurtosis K > 0

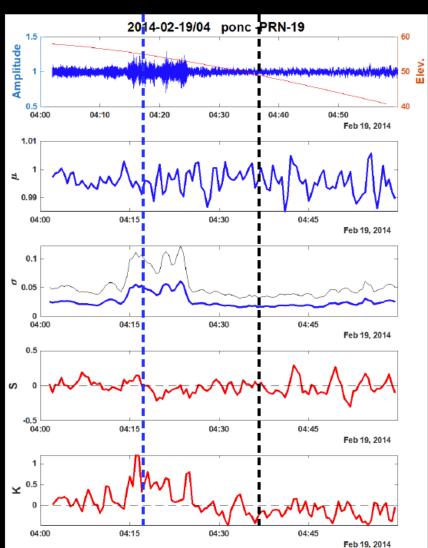


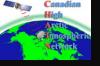


Another example

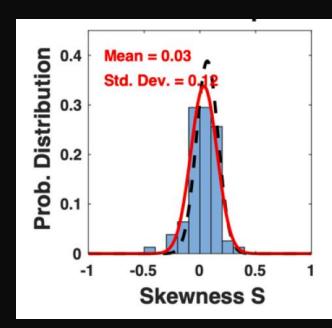


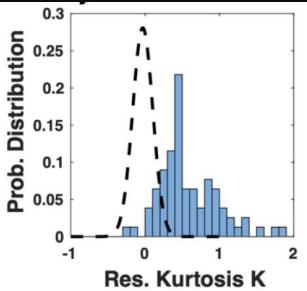
Skewness S ≈ 0 Kurtosis K > 0





Statistics



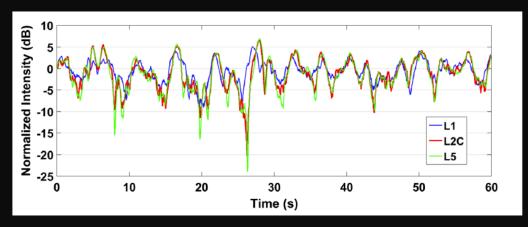


- Absence of scintillation [background]
 - Gaussian distribution.
 - Weakly scattered signal.
- Presence of scintillation
 - Symmetric (S = 0), Leptokurtic (K > 0) distribution.
 - Best fit: Generalized Gaussian distribution with β < 2.
 - Familiar distributions (Log-normal, Rician, Nakagami) not satisfactory.

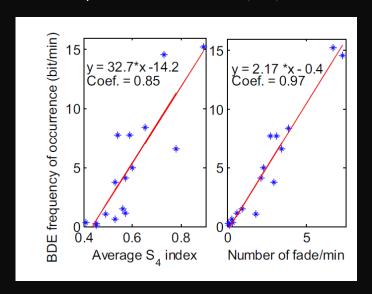


Effects on systems and measurements

- Decrease positioning accuracy and precision
 - Range errors
 - Cycle slips
 - Loss of lock
- Communications
 - Information or bit loss
 - Navigation message BDE
- Estimation of TEC using GNSS observables
- In the worst case scenario, GNSS based applications will not function



An example scintillation at Eureka, NU, Canada



Xu and Morton, 2018





Indices derived using GNSS observables

- Scintillation indices (amplitude and phase)
- Rate of change of TEC (ROT) Pi et al., 1997

$$ROT = \frac{TEC_i - TEC_{i-1}}{\tau_t - \tau_{t-1}}$$

ROT (ROTI) index – Pi et al., 1997

$$ROTI(i) = \sqrt{\frac{1}{N} \sum_{j=i-N}^{i} (ROT(j) - \overline{ROT})^2}$$

- Disturbance ionosphere index (DIX) Jakowski et al., 2012
- Gradient ionospheric index (GIX) Jakowski and Hoque, 2018
- Delta phase rate (DPR) Ghoddousi-Fard and Lahaye,

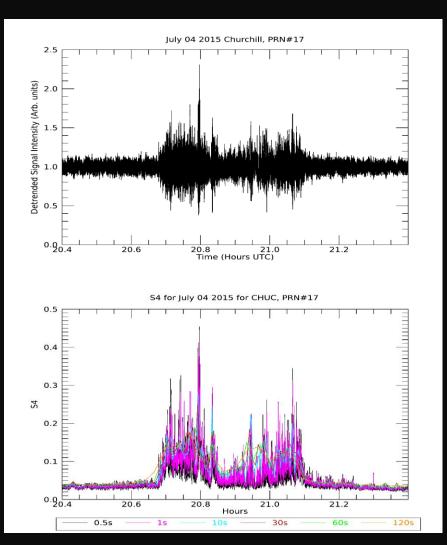
$$DPR = \frac{\sum_{i=0}^{n} \psi_i}{n}$$

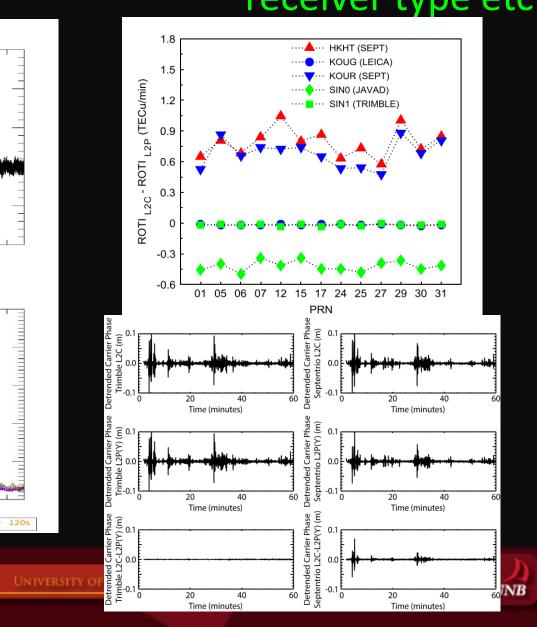
$$\psi_i = \frac{\varphi_{\mathsf{g}}^{t_{i+1}} - \varphi_{\mathsf{g}}^{t_i}}{t_{i+1} - t_i}$$

$$\varphi_{\mathsf{g}} = \lambda_1 L_1 - \lambda_2 L_2$$



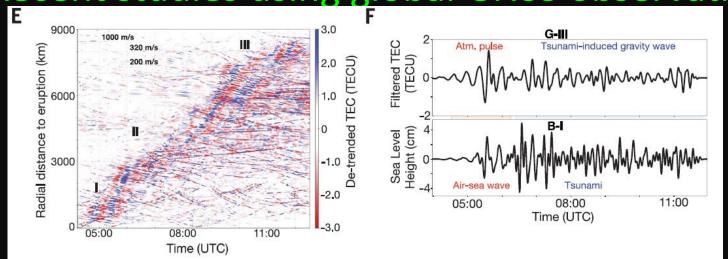
These indices depend on the sampling rate, receiver type etc.







Recent studies using global GNSS observations



Matoza et al., 2022

Zhang et al., 2020





GNSS Metrology-TEC

$$\Phi_L^{s,r}(t) = \rho^{s,r}(t) + c \Big(dt^s(t) - dt^r(t) \Big) + T^{s,r}(t) - I_L^{s,r}(t) + \lambda_L N_L^{s,r} + \epsilon_L^{s,r}(t)$$

Ionosphere is dispersive

$$n(h) = 1 - \frac{40.3N_e(h)}{f^2}, \quad n_g(h) = \frac{c}{v_g} = 1 + \frac{40.3N_e(h)}{f^2}$$

Change in Phase or Code

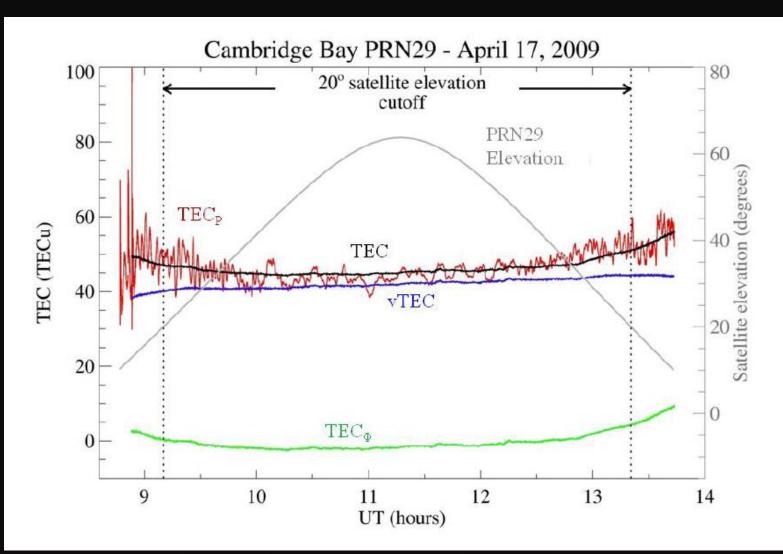
$$\Delta \Phi = -\Delta P = -\frac{40.3}{f^2} \int N_e dl$$

Total electron content (TEC)

$$TEC = \frac{f_{L1}^2 f_{L2}^2}{40.3(f_{L1}^2 - f_{L2}^2)} \Delta \Phi$$

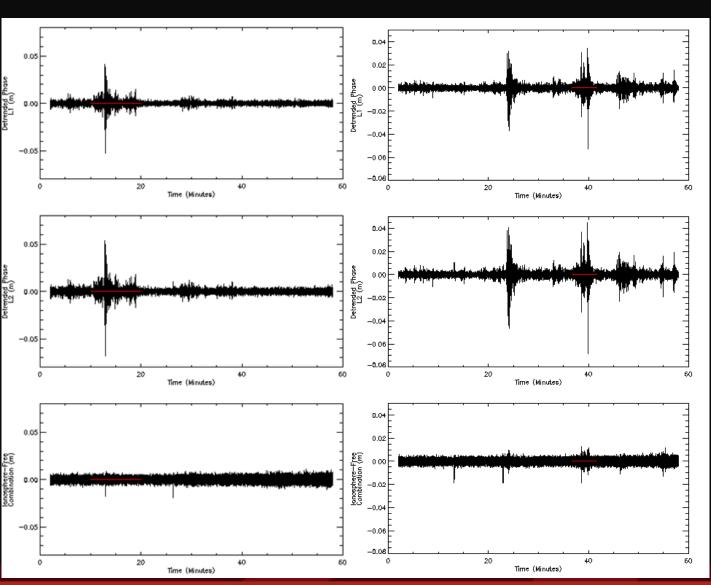


TEC





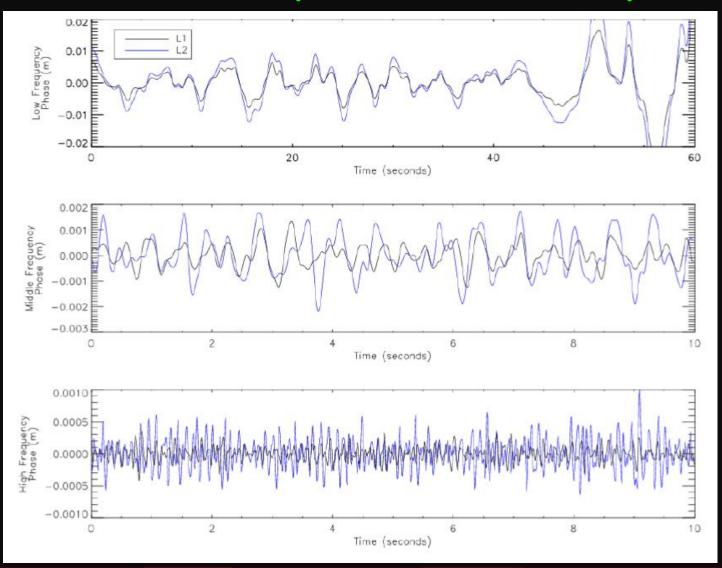
Ionosphere free linear combination (IFLC)



$$\Phi_{IFLC} = \frac{\Phi_1 f_1^2 - \Phi_2 f_2^2}{f_1^2 - f_2^2}$$

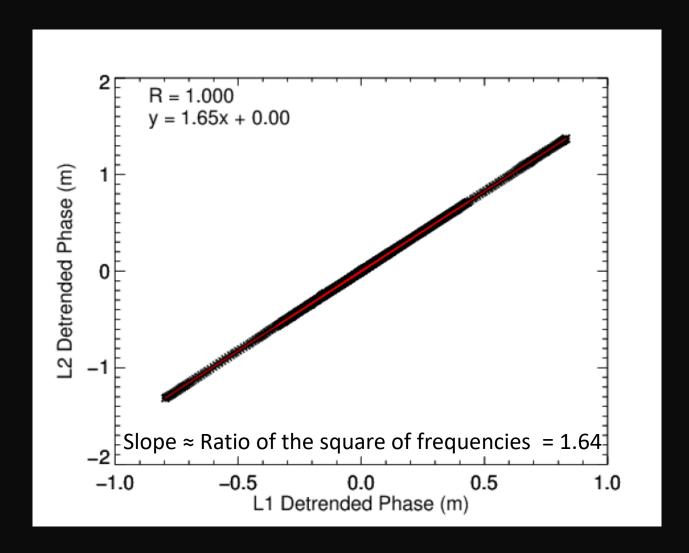


Variation of phase at two frequencies



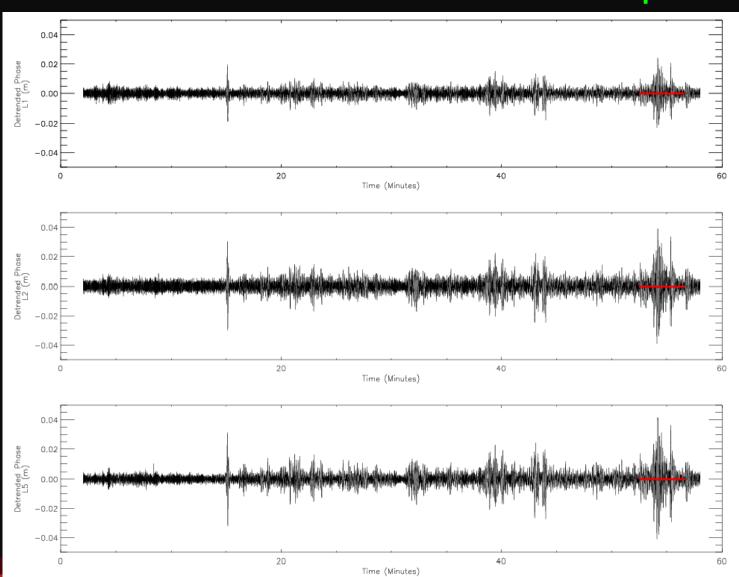


It works most of the time



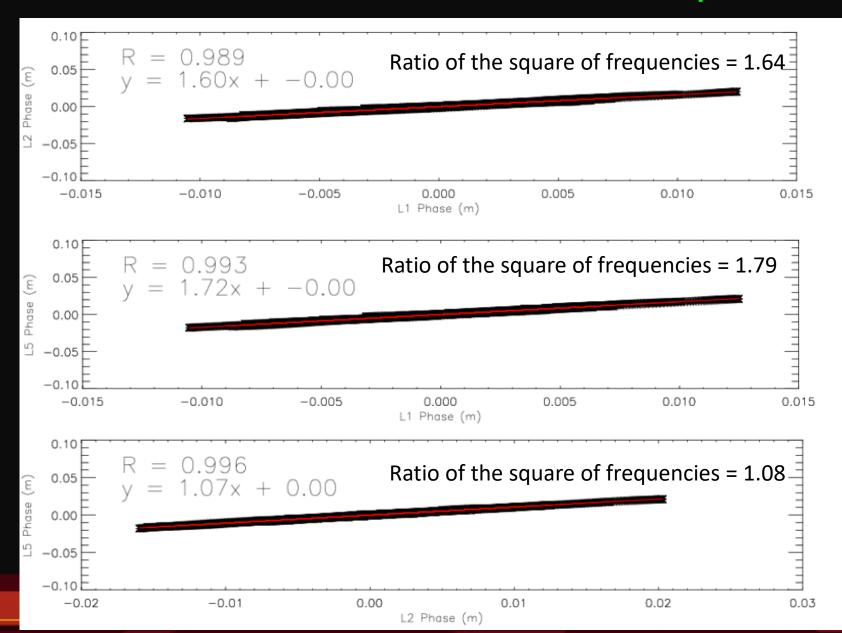


Three frequencies



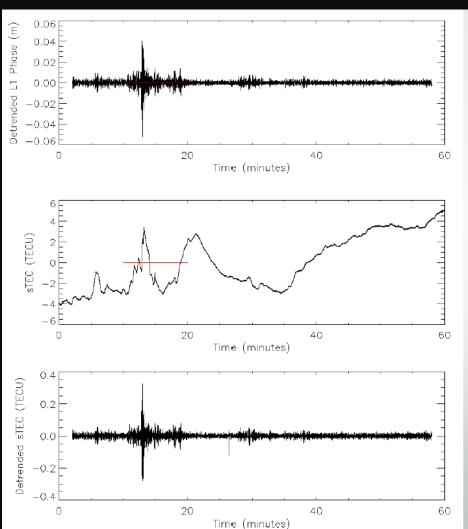


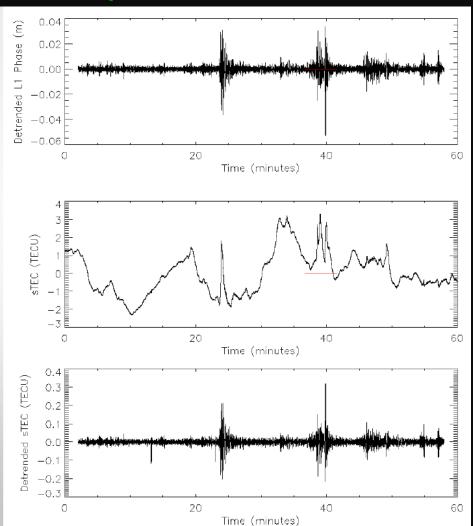
Three frequencies





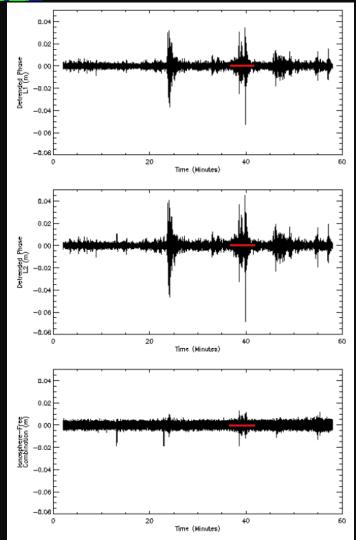
TEC and phase variations



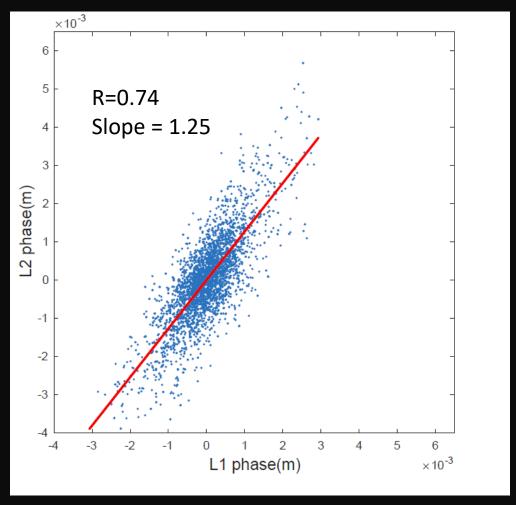








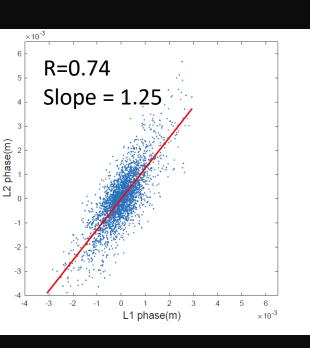
Under real scintillation

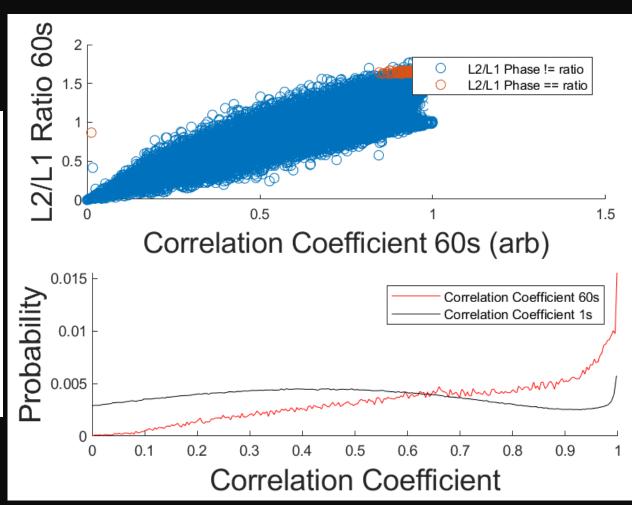


TEC estimate under scintillation is not accurate because 1/f² dependence is violated



Under scintillation

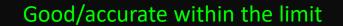


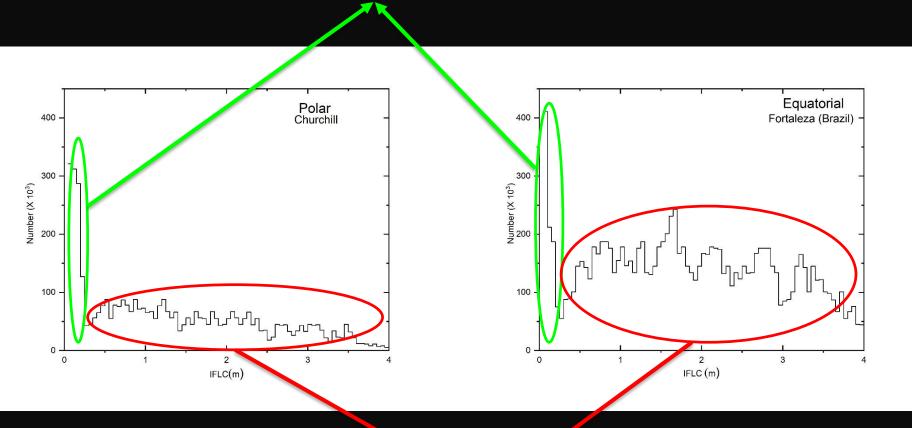


TEC estimate under scintillation is not accurate because 1/f² dependence is violated



In a nutshell





Bad/significant error in TEC estimate or not TEC at all





Conclusions

- Scintillation is a major issue in estimating accurate TEC based on GNSS observables
- Under "actual scintillation", this 1/f² dependence is violated. Ergo the TEC estimated will not be accurate and will be nonphysical
- Any Space weather indices (e.g. ROTI) is derived using the estimated TEC will not represent the actual situation
- We need to develop a method to flag or provide an error estimate for the TEC derived using GNSS observables
- This can be based on the correlation value of phase at multiple frequencies.