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"The Empirical Optimal Estimation Method Applied to the Inversion of
Atmospheric FTIR Remote Sensing Data"

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Please note: These are preliminary notes intended for internal distribution only.

The empirical optimal estimation method applied to the inversion of atmospheric FTIR remote sensing data.

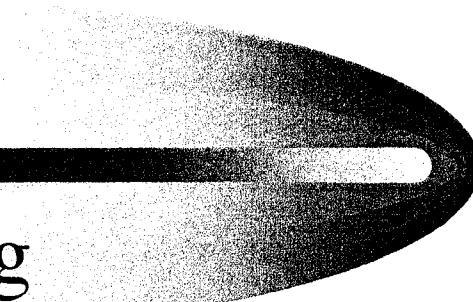
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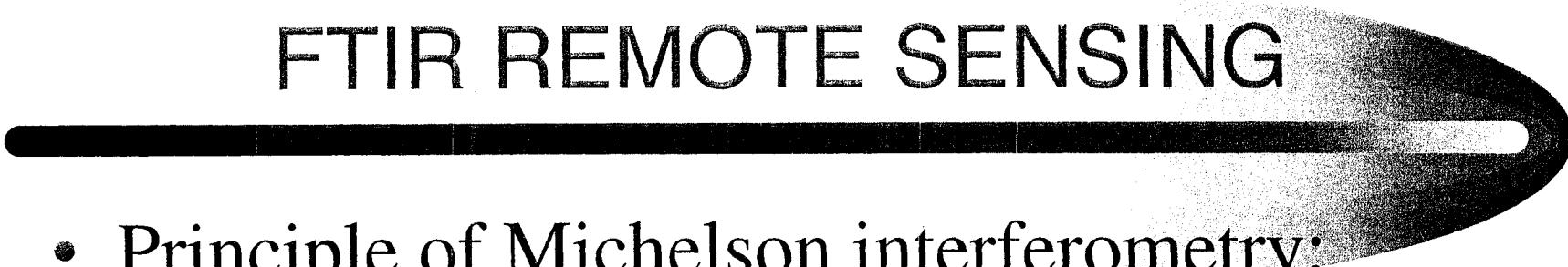
martine@oma.be

OUTLINE

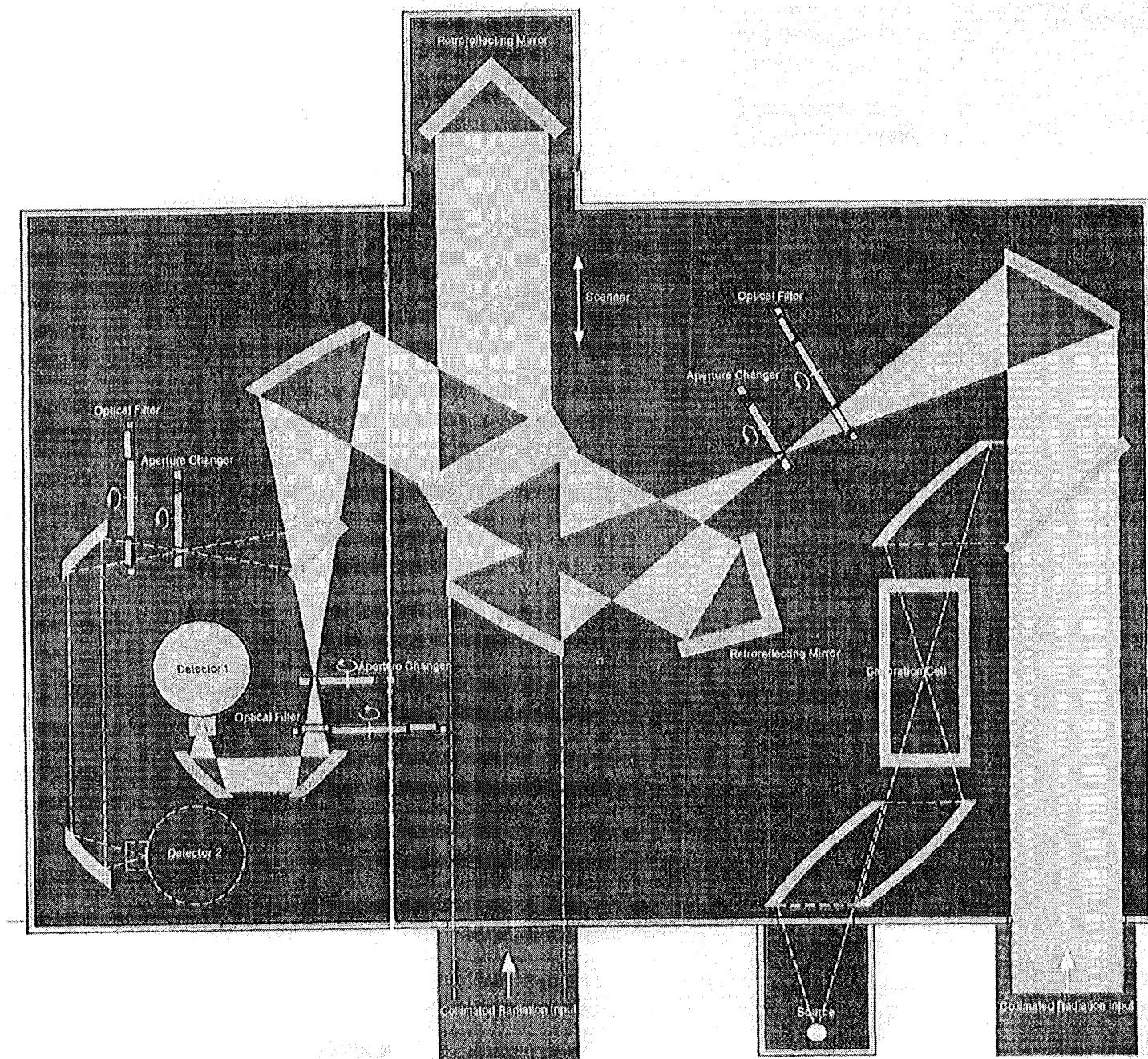


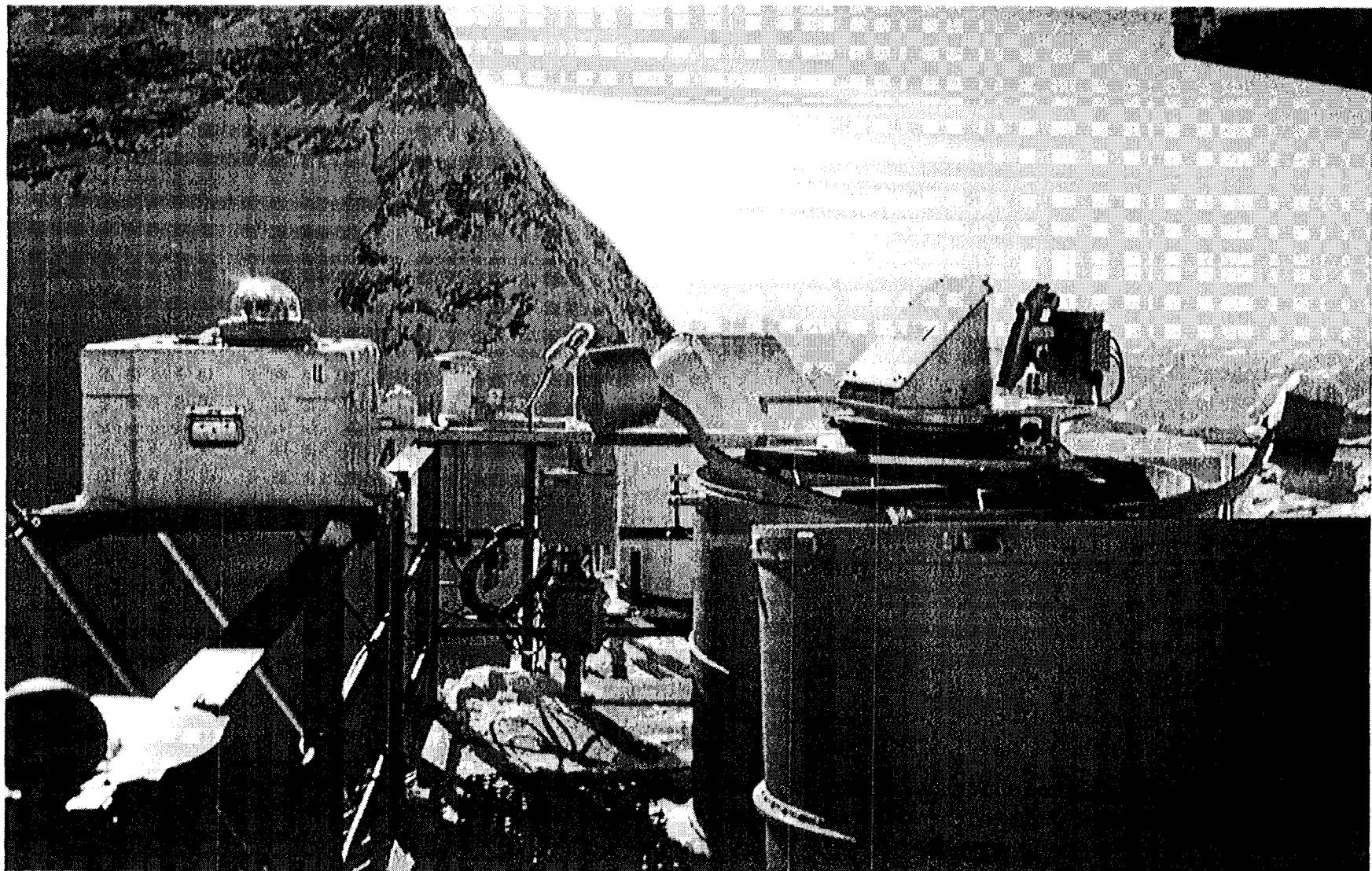
- Principle of FTIR remote sensing
- Inversion
 - Total columns
 - Improved total columns
 - vertical profiles
- Characterisation of the retrieval
- Application:
1996-2001 time series of O₃ at the Jungfraujoch

FTIR REMOTE SENSING



- Principle of Michelson interferometry:
 - ⇒ recording of interferogram of incident light (sun, thermal emission,)
 - ⇒ FFT
 - ⇒ spectrum
 - high spectral resolution
 - multiplex advantage
 - ⇒ distinction of many absorption/emission signatures simultaneously

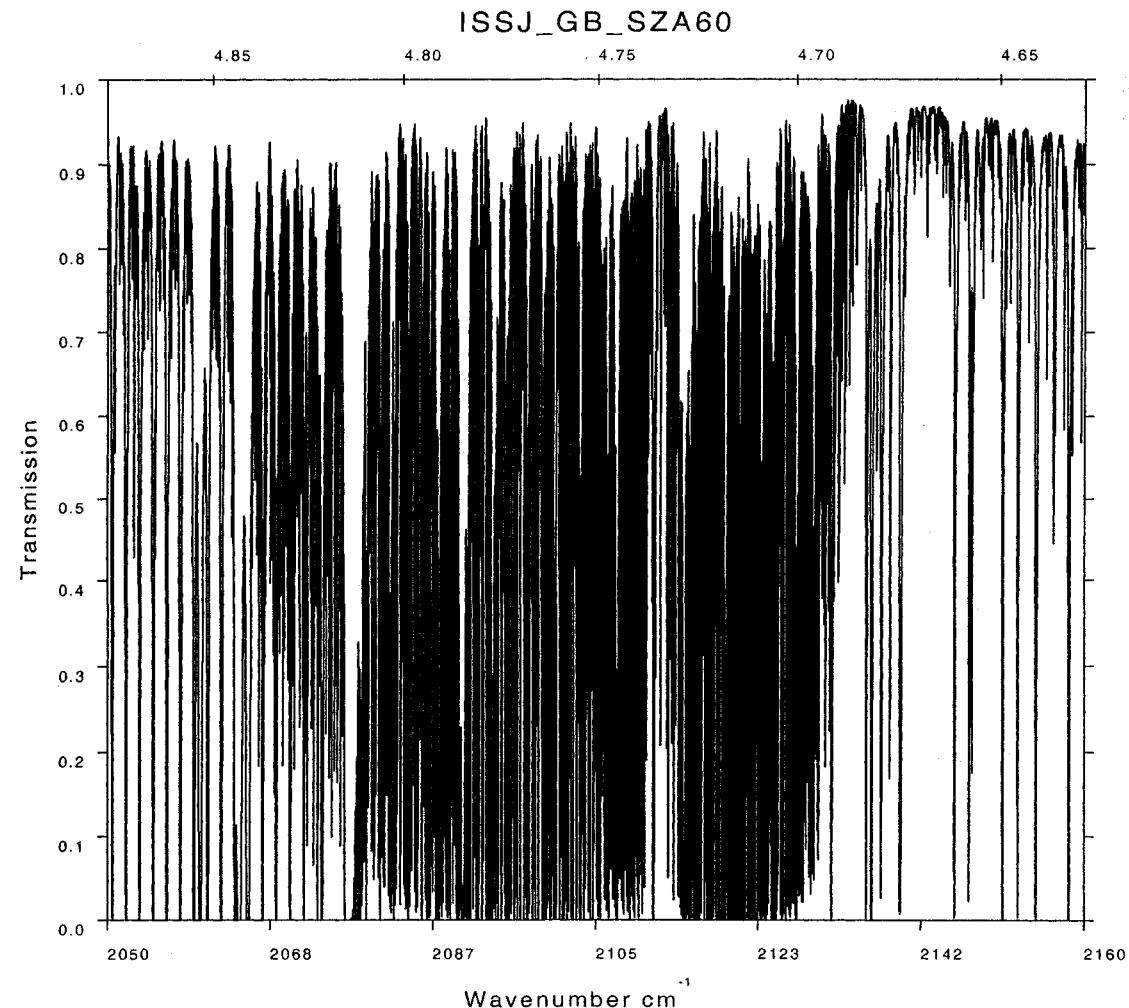




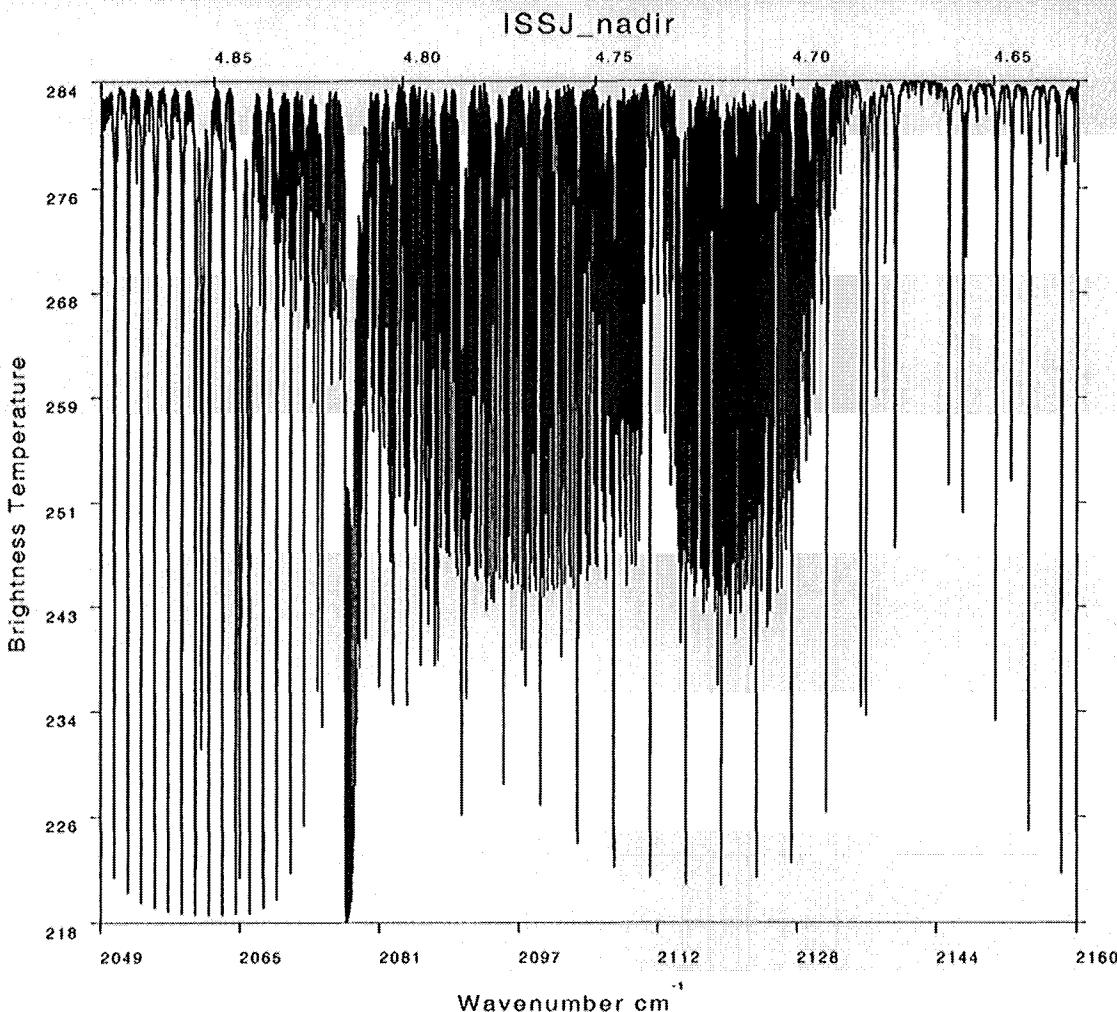
ICTP 2001

OEM applied to FTIR
M. De Mazière, BIRA-IASB

Absorption ground-based spectrum



Emission nadir spectrum



FTIR atmospheric 'PRODUCTS'

- Reference gas: $\text{N}_2(<1\%)$
- Minor constituents: $\text{CO}_2(<2\%), \text{N}_2\text{O}(\leq 2\%), \text{CH}_4(\approx 2\%), \text{CO}(\leq 4\%), \text{O}_3(\leq 5\%), [\text{H}_2\text{O}]$
- Halogenated trace species:
 $\text{HCl}(\leq 4\%), \text{ClONO}_2(\approx 20\%), \text{CCl}_2\text{F}_2(\leq 4\%), \text{CHClF}_2(\leq 6\%), \text{HF}(\leq 3\%), \text{COF}_2(\leq 20\%), \text{SF}_6(\approx 30\%)$
- Nitrogenated trace species:
 $\text{NO}(\leq 6\%), \text{NO}_2(\leq 10\%), \text{HNO}_3(\approx 5\%), \text{HNO}_4(\text{monthly avg.})$
- Other trace species: $\text{C}_2\text{H}_2(\approx 20\%), \text{C}_2\text{H}_6(\leq 6\%), \text{HCN}(\leq 8\%), \text{OCS}(\leq 8\%), \text{H}_2\text{CO}(\text{monthly avg.}), [\text{H}_2\text{CO}_2, \text{CH}_3\text{Cl}, \text{CCl}_4, \text{CCl}_3\text{F}, \dots]; \text{various isotopes}$

(Precisions estimated for Jungfraujoch conditions)

FTIR retrieval: forward model

$$y = f(x, b, b') + \varepsilon$$

$$F(x, b) \approx f(x, b, b'); \quad \Delta f = f(x, b, b') - F(x, b)$$

$$y \approx F(x, b) + \varepsilon$$

y is the experimental measurement, x is the unknown state

ε is the experimental error term

f, F are the forward function, model, respectively

b are forward function parameters, that are in the model,

w / best estimates \hat{b} ; some may be retrieved \rightarrow included in x

b' are forward function parameters, ignored in the model

FTIR Forward Line By Line model

Solar absorption mode

RT model: Transmission $T(\tilde{v}) = I(\tilde{v}) / I_0(\tilde{v}) = \exp(-\alpha(\tilde{v}))$

with $\alpha(\tilde{v})$ absorption coefficient:

$$\alpha(\tilde{v}) = \int_{z_0}^{\infty} dz \ l(z) \sum_i n_i(z) \sum_j S_{ij}(\tilde{v}, z) f_{ij}(\tilde{v}, z)$$

$i \rightarrow$ molecule ; $j \rightarrow$ absorption line

S_{ij} line strength / molecule ; f_{ij} lineshape

dependence on $z \Leftrightarrow$ dependence on P, T

Instrument model: $\tilde{T}(\tilde{v}) = T(\tilde{v}) \otimes ILS(\tilde{v})$

FTIR LBL model parameters

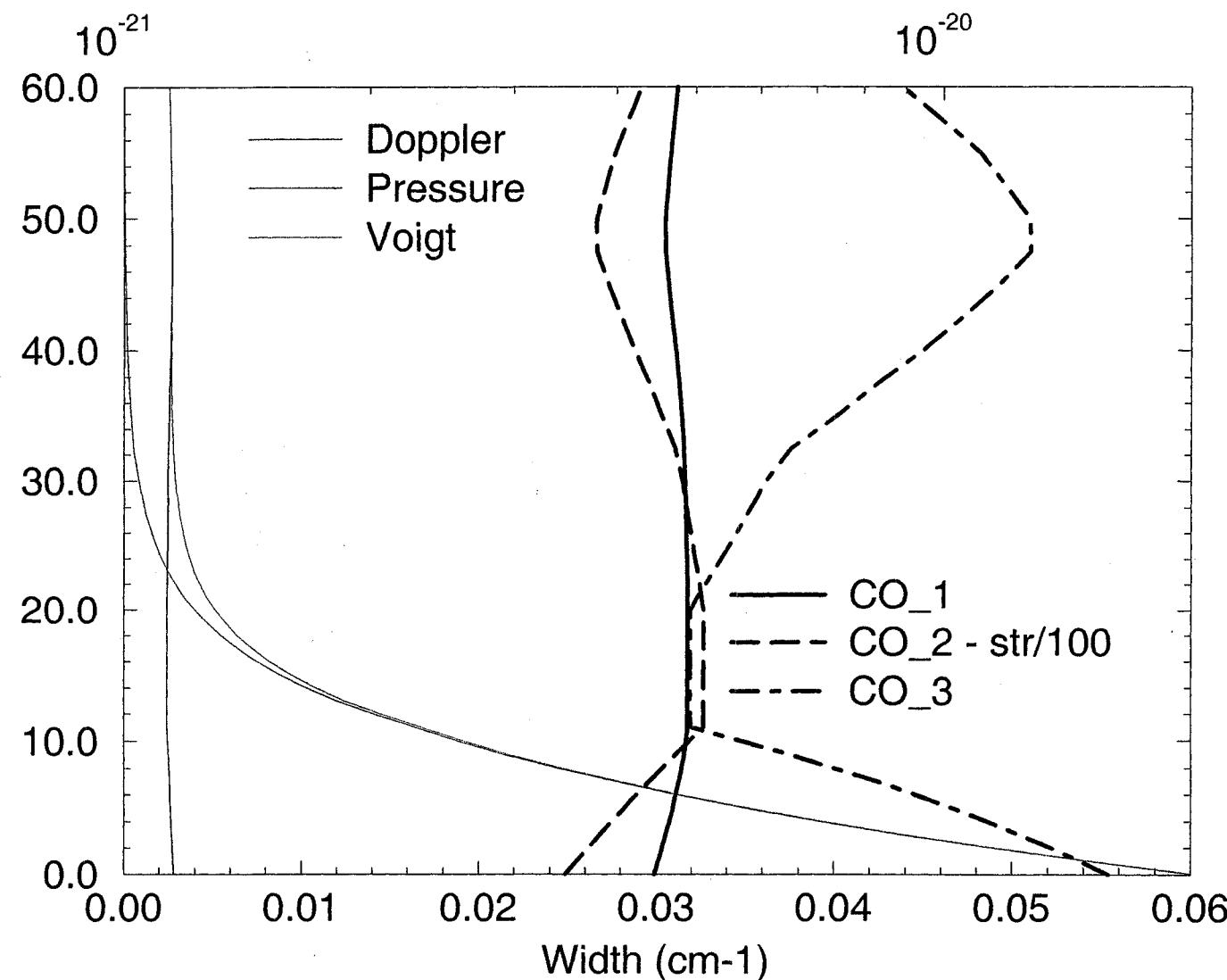
- Linestrength and -shape $\Rightarrow b$
 - Strength $S(T)$ - depending on E''
 - Voigt lineshape is adopted,
close to Lorentz at high pressure,
close to Doppler at low pressure
 - *neglects* Dicke narrowing, line mixing, ... $\Rightarrow b'$

$$\text{Doppler HWHM } \alpha_D = v_0 \sqrt{\frac{2kT \ln 2}{M}}$$

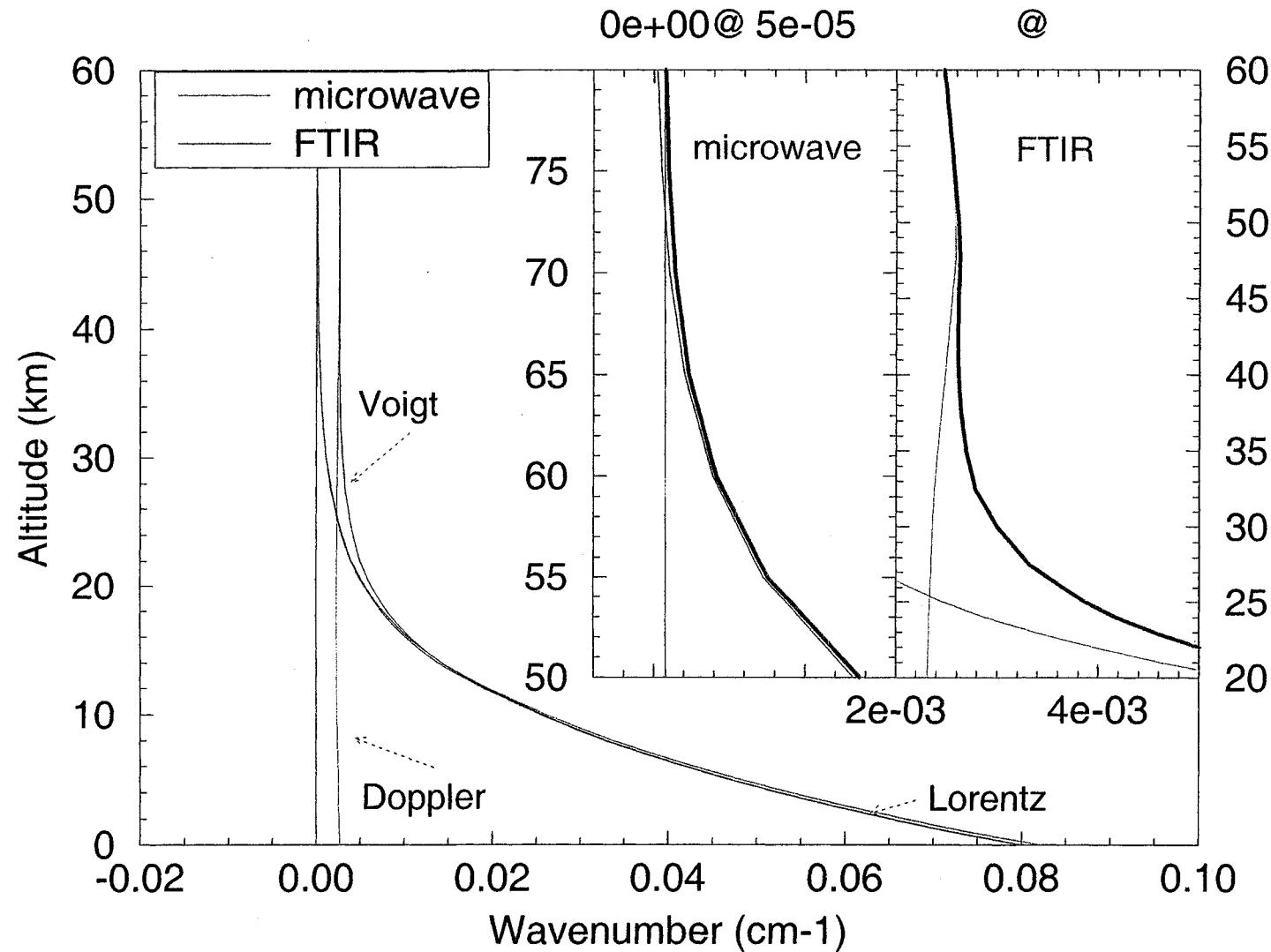
$$\text{Lorentz HWHM } \alpha_L = \alpha_{L,0} (P/P_0)^m (T_0/T)^n$$

$$m \leq 1 \quad n: 0.5 \rightarrow 1 \quad (\text{databases for } \alpha_{L,0} \text{ and } n)$$

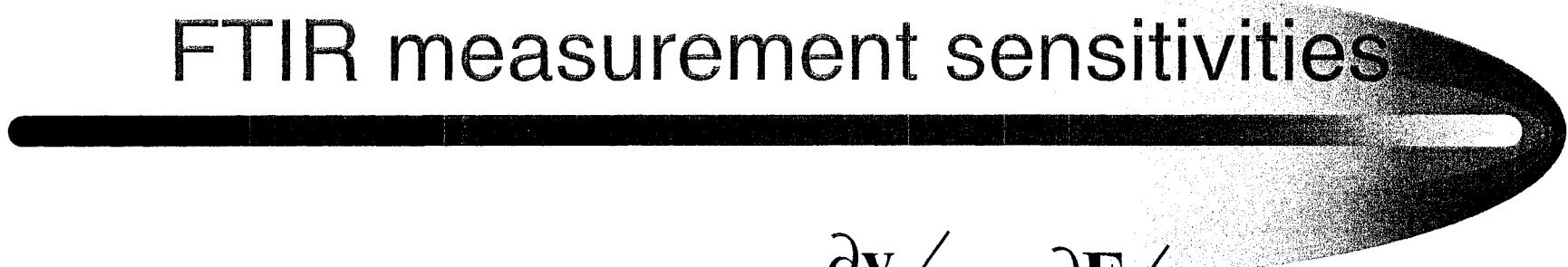
Linestrength



Linewidth contributions - comparison

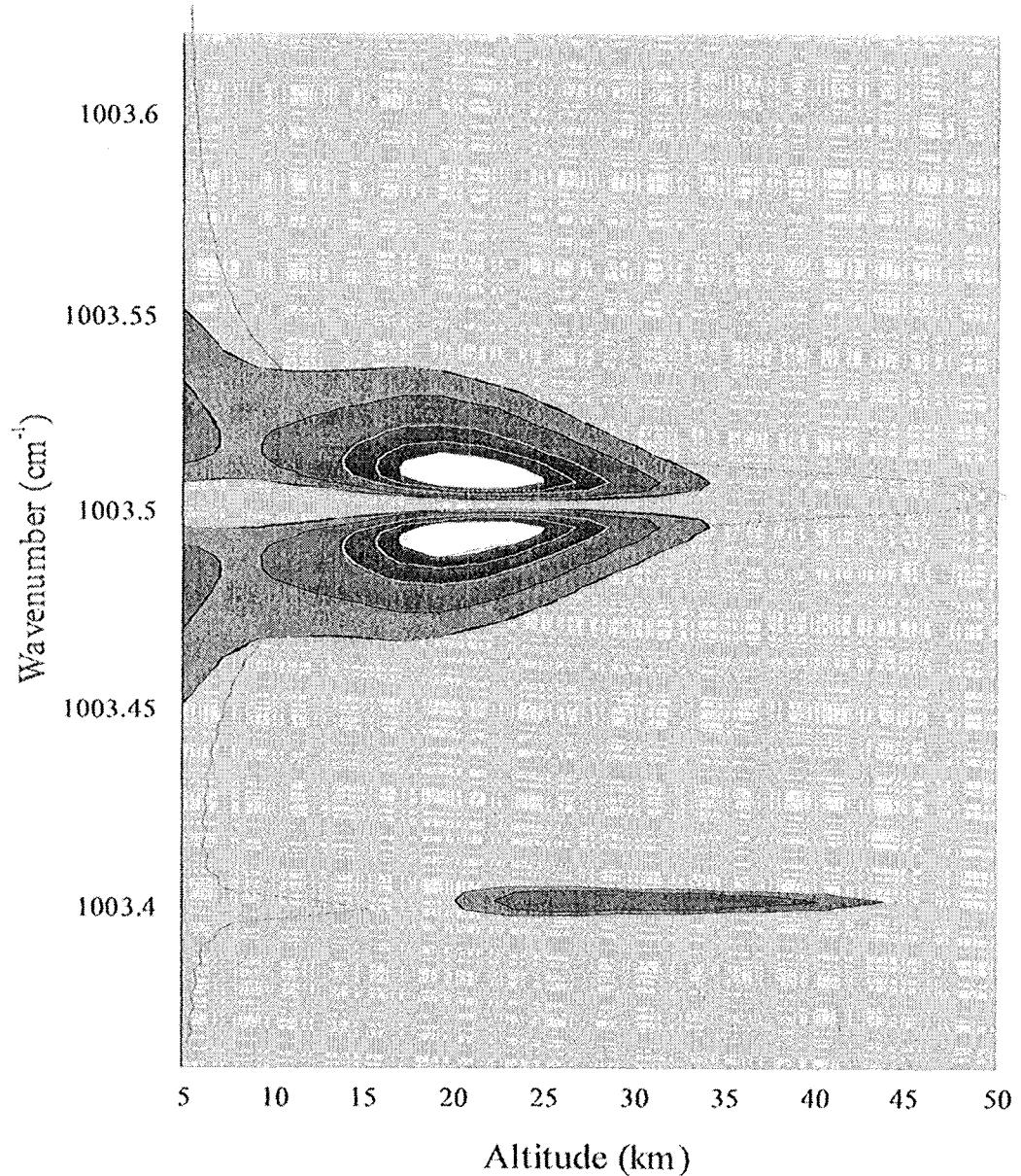


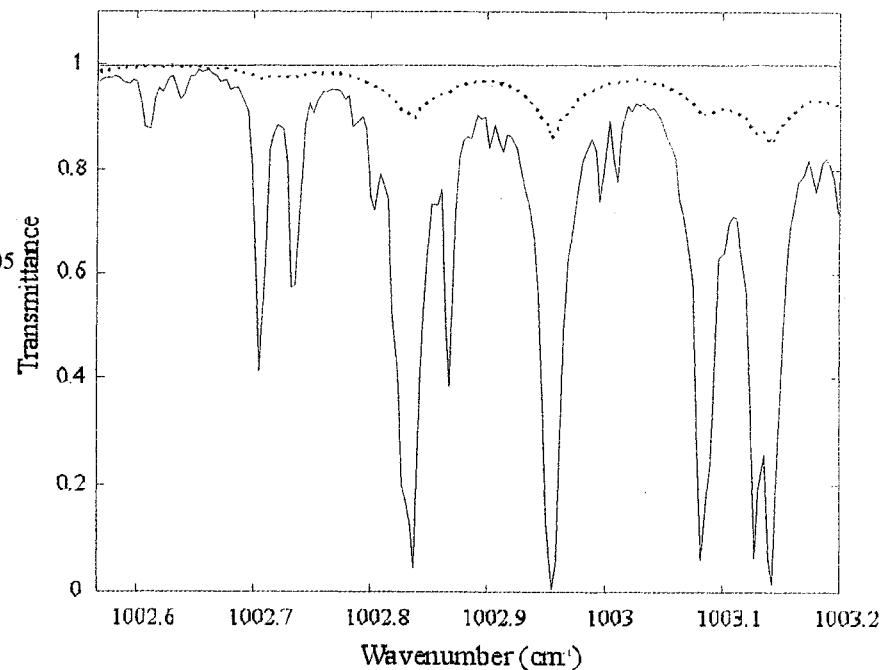
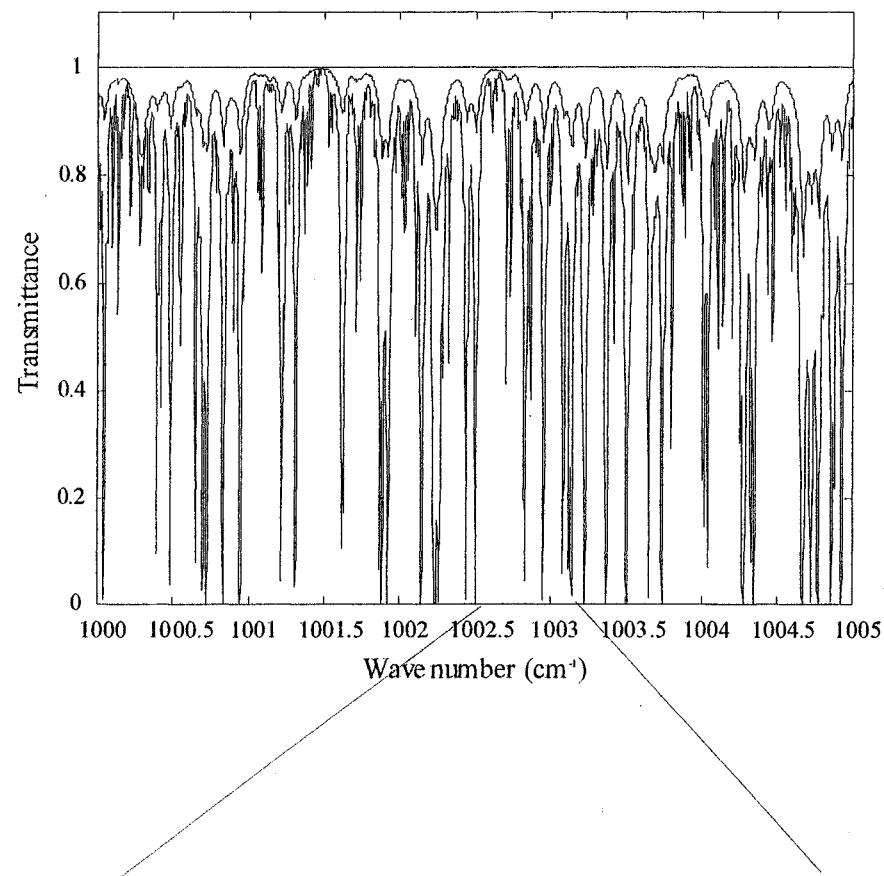
FTIR measurement sensitivities



- Weighting f^{ion} : $K_x = \frac{\partial y}{\partial x} \approx \frac{\partial F}{\partial x}$
 - saturated lines:
 - more sensitive to tropospheric contributions,
 - reduced information content as to stratosphere
 - weak lines:
 - more information about the stratosphere

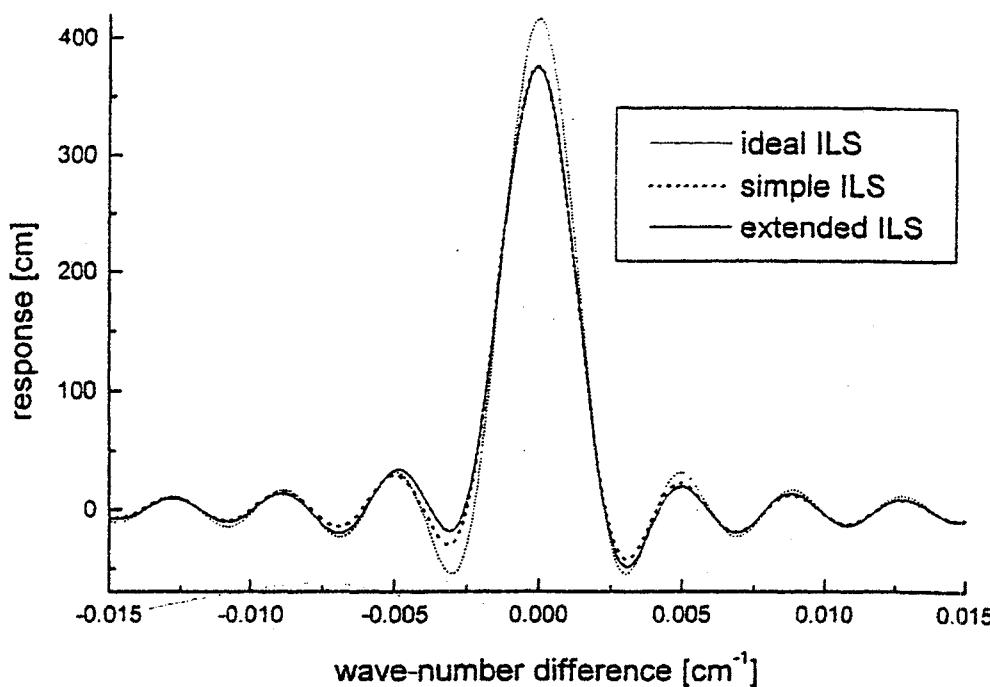
$$\mathbf{K}_x = \frac{\partial \mathbf{y}}{\partial \mathbf{x}} \approx \frac{\partial \mathbf{F}}{\partial \mathbf{x}}$$





FTIR instrument function

- ILS $\Rightarrow b$
 - FTS-ILS (v): $\text{sinc}(\propto \text{OPD}) \otimes \text{Boxcar}(\propto \Omega)$
 $\otimes EAP\text{-function}$



FTIR retrieval: Inverse Model -1

- General formulation

$$\hat{x} = R(F(x, b) + \Delta f(x, b, b') + \varepsilon, x_a, \hat{b}, c)$$

\hat{x} is the retrieval result for x , through the inverse method R

x_a and c : retrieval method parameters

x_a the a priori estimate of x

- FTIR:

– $x \Leftrightarrow$ target concentrations (column w/ fixed profile $\rightarrow \dots \rightarrow$ profile), selected model pars;

– $c \Leftrightarrow x_a, S_a, S_\varepsilon \Leftrightarrow SNR_{retr}$, convergence criterium

FTIR retrieval: Inverse Model - 2

- OEM - Non-linear case- Newtonian iteration:

$$\hat{x} = (S_a^{-1} + K^T S_\varepsilon^{-1} K)^{-1} (K^T S_\varepsilon^{-1} y + S_a^{-1} x_a)$$

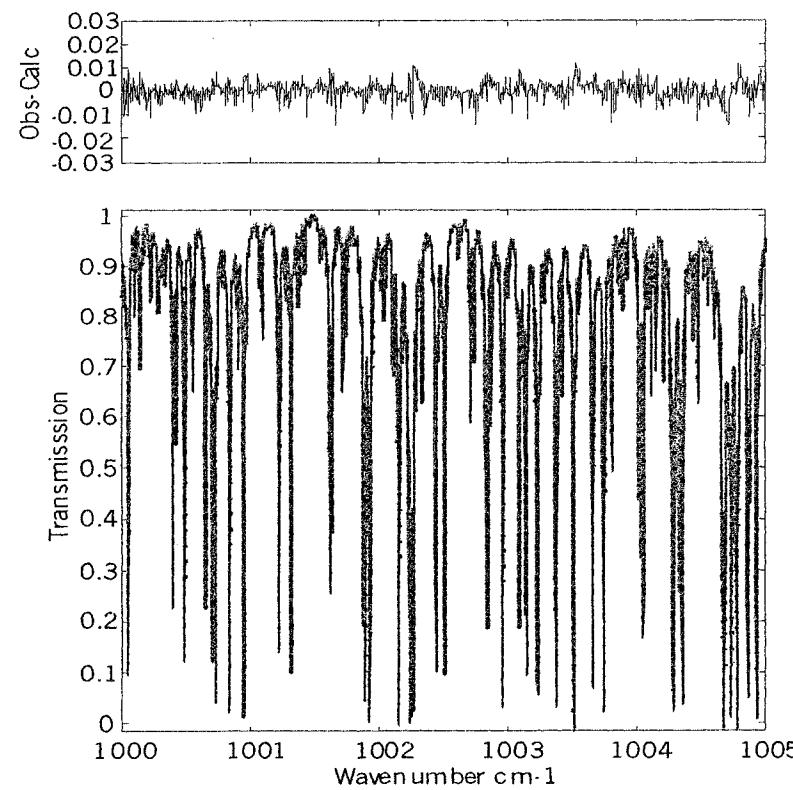
$$x_{n+1} = (S_a^{-1} + K_n^T S_\varepsilon^{-1} K_n)^{-1} \left[K_n^T S_\varepsilon^{-1} (y_m - y_n + K_n x_n) + S_a^{-1} x_a \right]$$

$$x_{n+1} = x_a + S_a K_n^T (K_n S_a K_n^T + S_\varepsilon)^{-1} \left[(y_m - y_n) - K_n (x_a - x_n) \right]$$

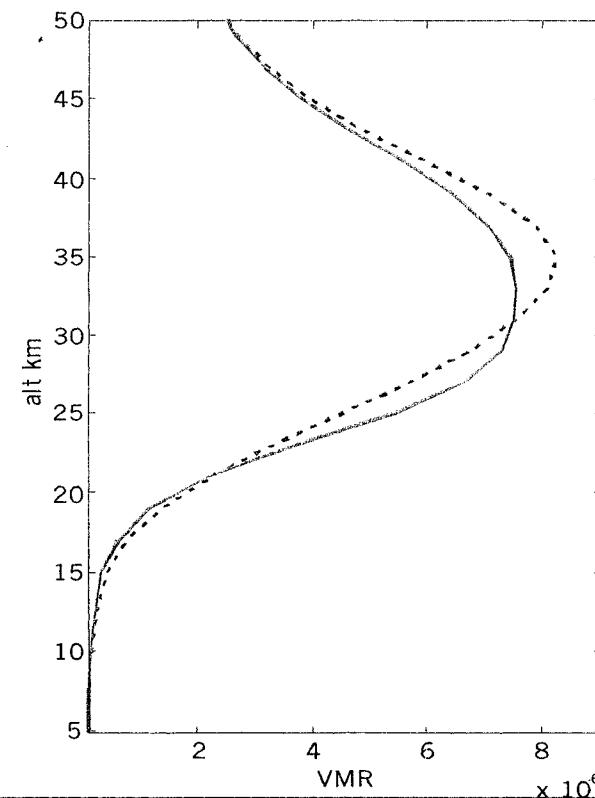
- For vertical inversion of ground-based solar absorption
FTIR: *SFIT2* code, empirical implementation of OEM
 - empirical estimations of best S_ε , S_{x_a} and of covariances of all retrieved parameters (S_b)
 - best ? best fit, w/o instabilities in retrieved state = best compromise between a priori and measurement info

FTIR inversion: examples for O₃

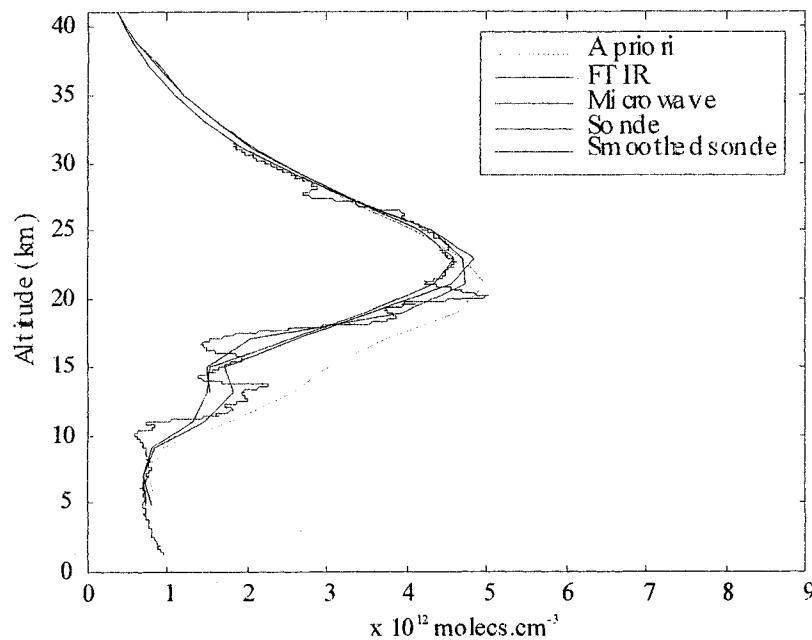
Obs R00814sw.moy
SZA=36.78
OPD=125 cm⁻¹
VCA=8.145e+018 cm⁻³
RMS=0.37



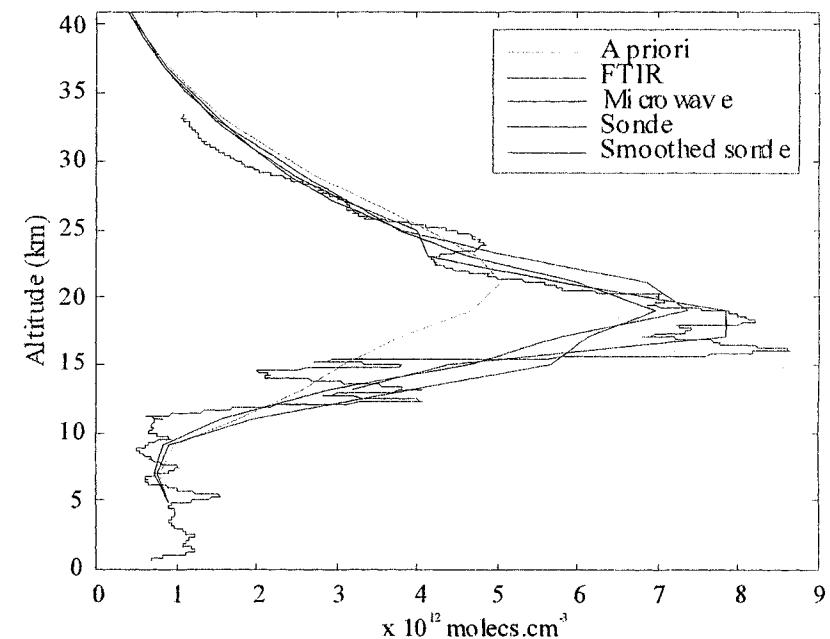
Retrieved
A priori



FTIR inversion: examples for O₃



1999-03-12



1999-03-17

Retrieval characterisation - 1

After linearisation and re-arrangement

$$\hat{x} - x_a = R(F(x_a, \hat{b}), \hat{b}, x_a, c) - x_a \quad \text{bias} > 0$$

$$+ A(x - x_a) \quad \text{smoothing}$$

$$+ G_y \varepsilon_y \quad \text{retrieval error}$$

$$G_y = \frac{\partial R}{\partial y} \quad \text{sensitivity of retrieval to measurement}$$

$$K_x = \frac{\partial F}{\partial x} \quad \begin{aligned} &\text{sensitivity of forward model to unknown state} \\ &\text{(weighting function)} \end{aligned}$$

$$A = G_y K_x \quad \text{sensitivity of retrieved to real state (averaging kernel)}$$

Retrieval characterisation - 2

Or else

$$\hat{x} - x = (A - I)(x - x_a) + G_y K_b (b - \hat{b}) + G_y \Delta f(x, b, b') + G_y \varepsilon$$

smoothing error
model parameter error
forward model error
'retrieval' (measurement) error

$$K_b = \frac{\partial F}{\partial b}$$

sensitivity of forward model to model parameters

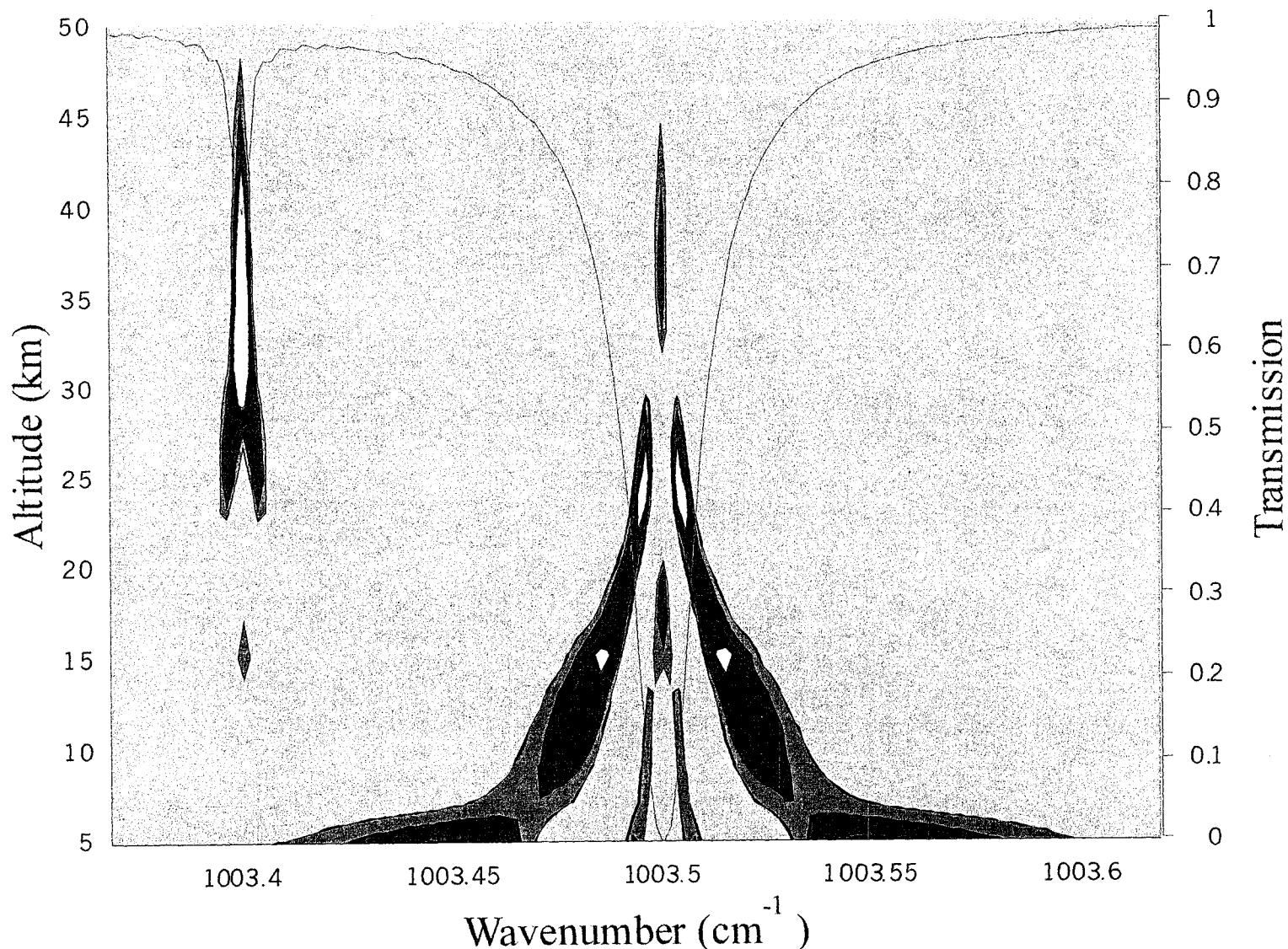
$$S_A = (A - I)S_x(A - I)^T$$

covariance of smoothing error

$$S_y = G_y S_\varepsilon G_y^T$$

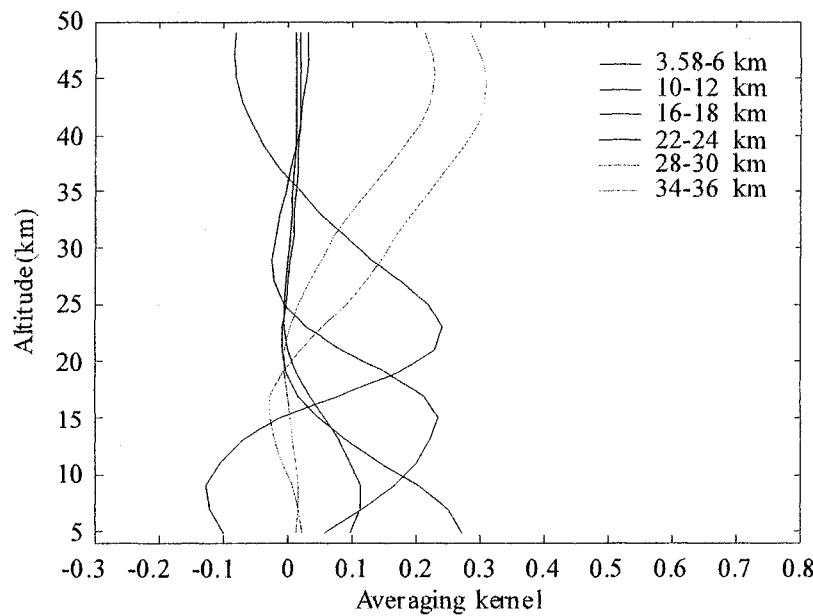
covariance of measurement error

$$\text{Contribution } f^{\text{ion}} \ G_y = \frac{\partial \hat{\mathbf{x}}}{\partial y} \equiv \frac{\partial \mathbf{R}}{\partial y}$$

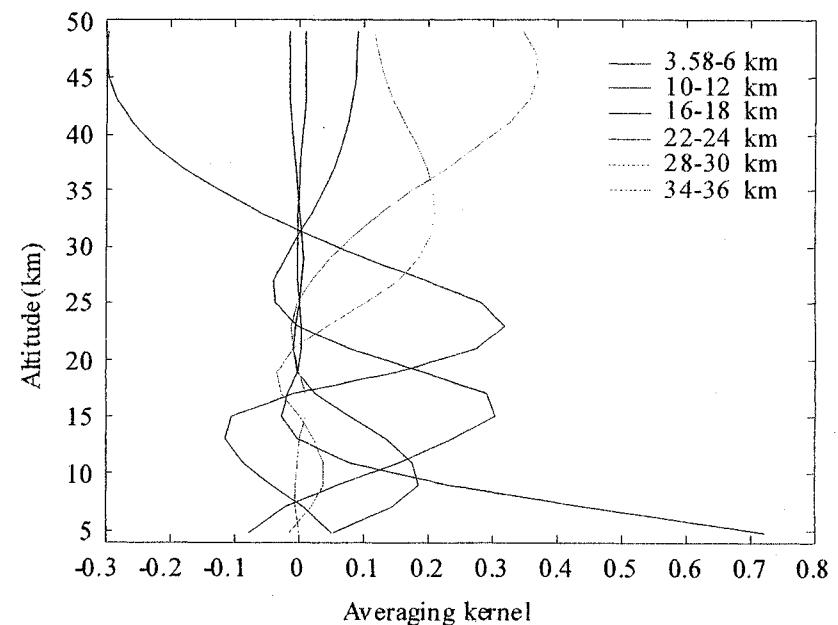


FTIR inversion: Averaging kernel

O_3



1002.6-1003.2 cm^{-1}

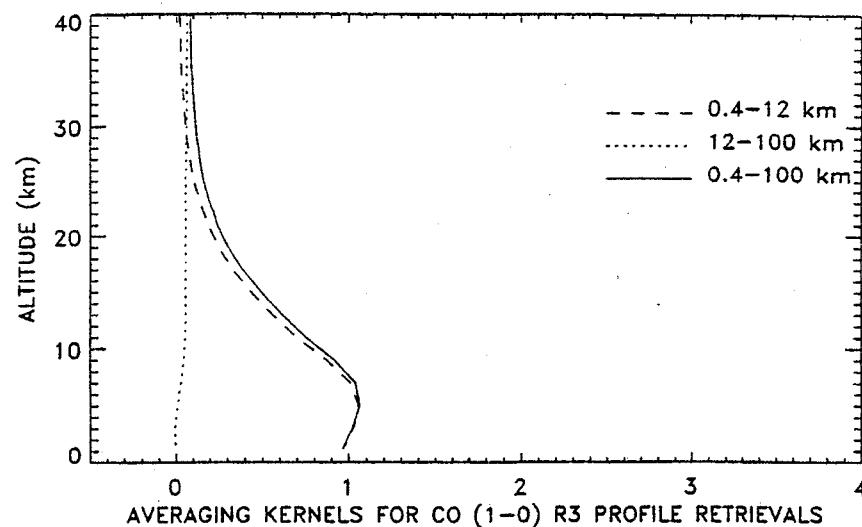


1000.0-1005.0 cm^{-1}

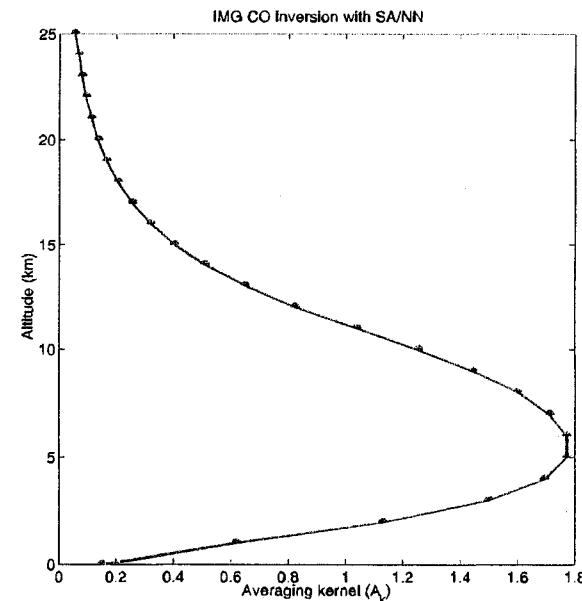
FTIR inversion: Averaging kernel

CO

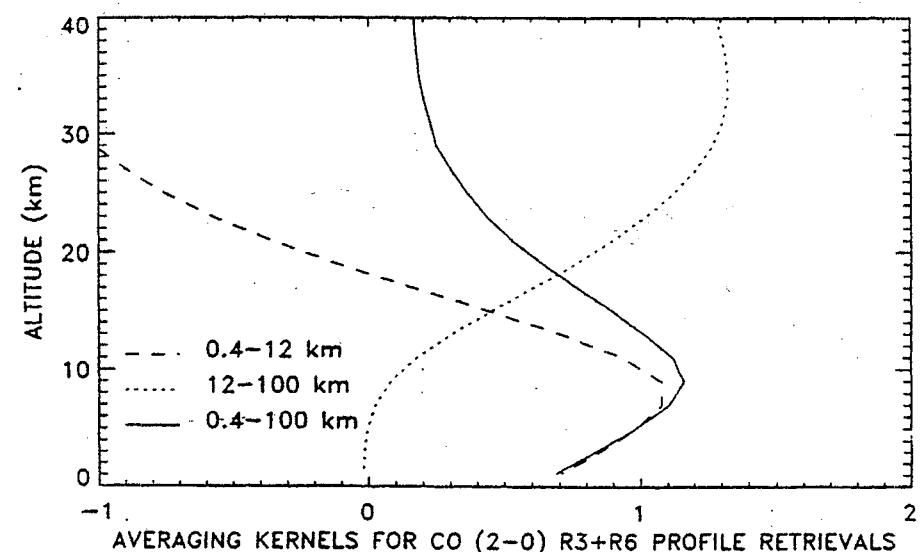
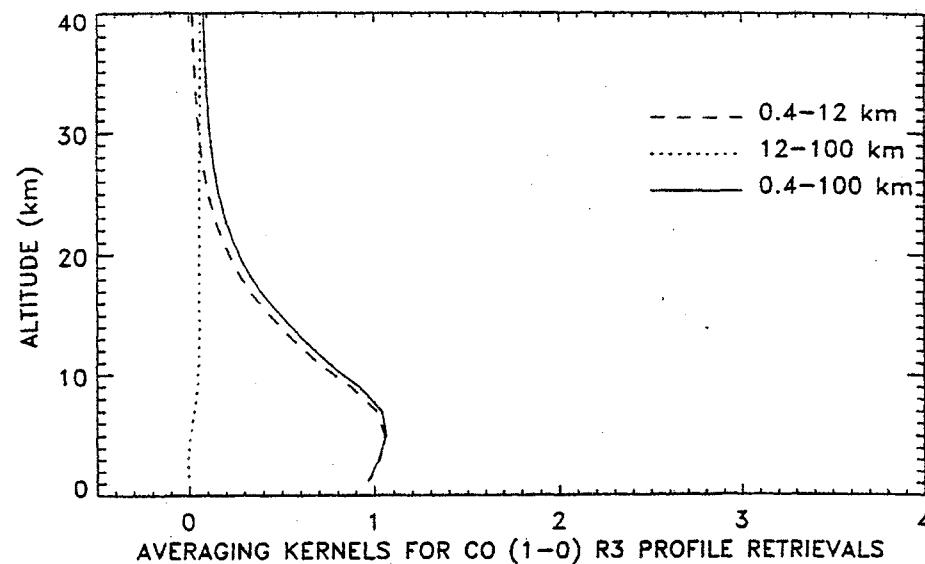
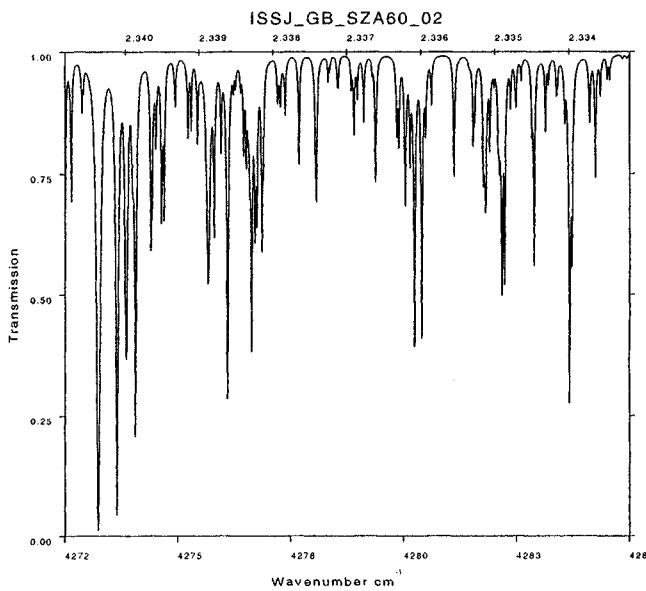
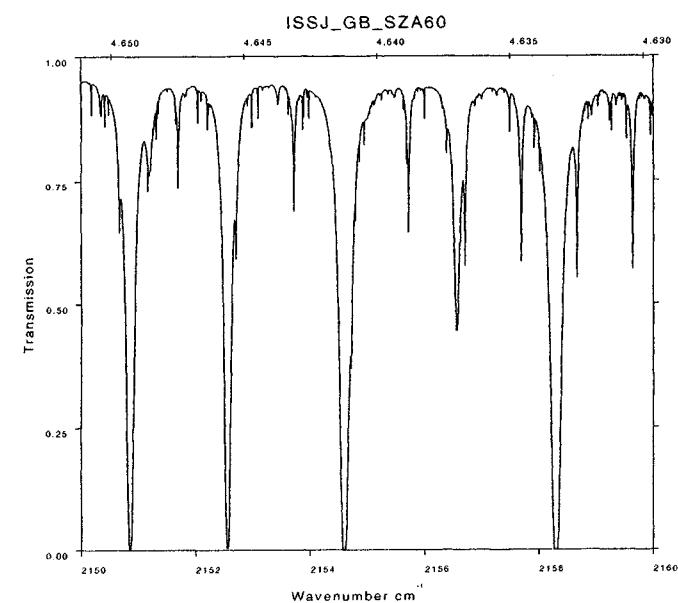
Solar absorption-GB



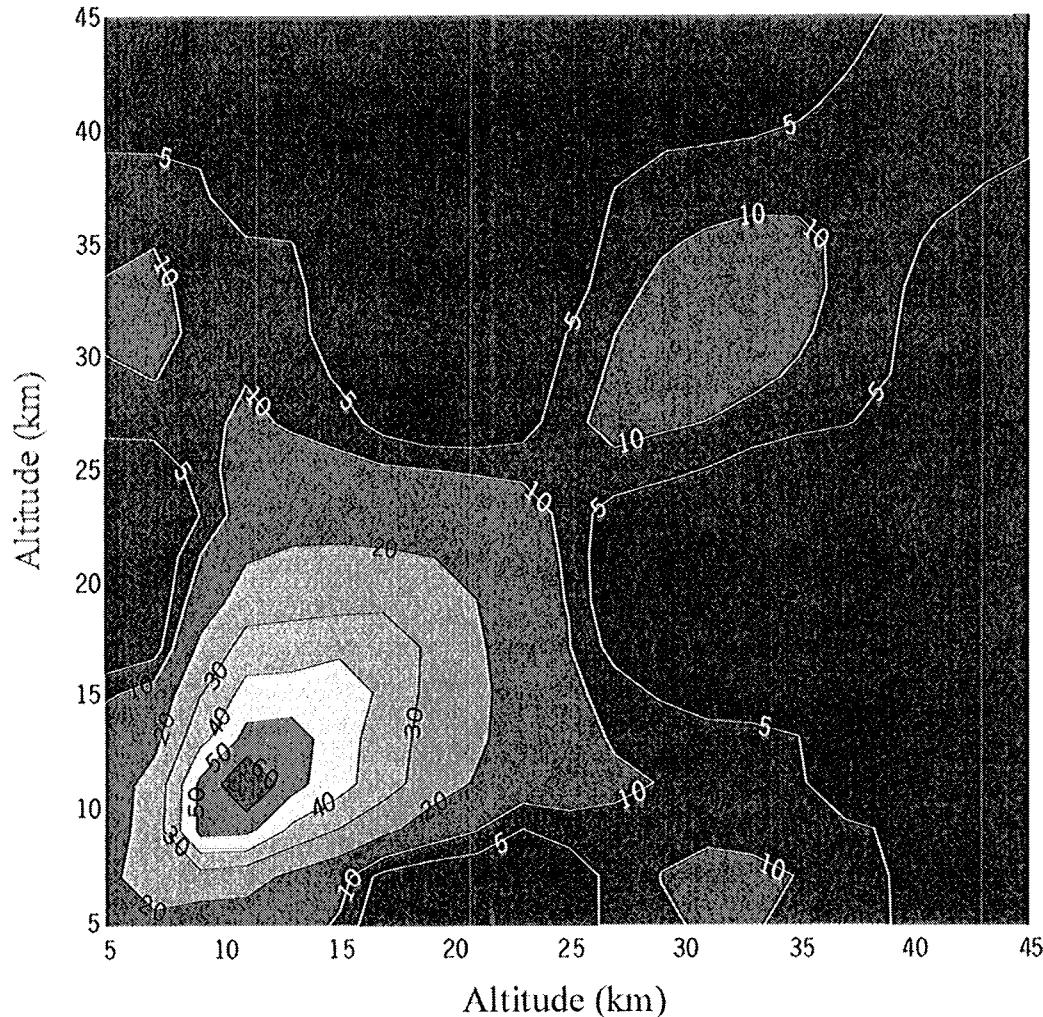
Emission - nadir space-based



CO



FTIR inversion: smoothing and measurement errors - 1



$$S_A = (A - I)S_x(A - I)^T$$
$$S_y = G_y S_\varepsilon G_y^T$$

$$S_x(i, j) = \sqrt{\frac{\text{cov}(x_i, x_j)}{\bar{x}_i \bar{x}_j}}$$

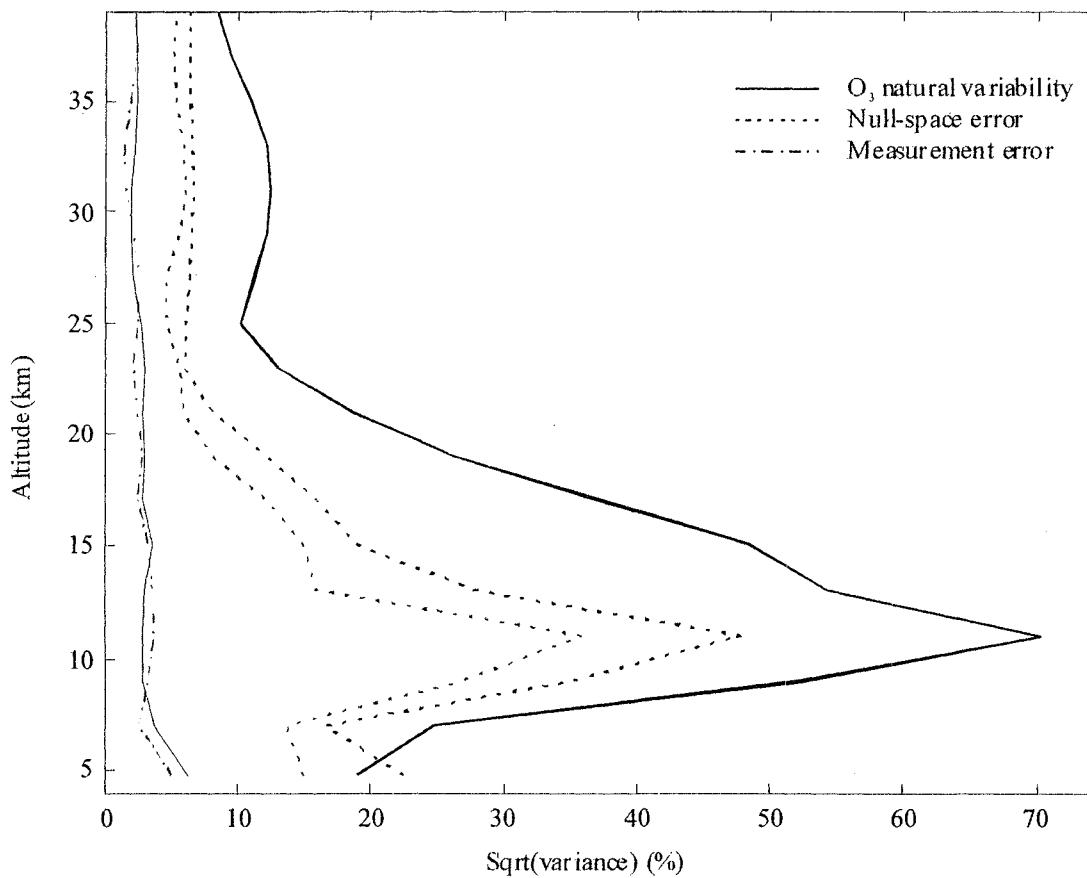
(in %);

$$x \equiv x_a$$

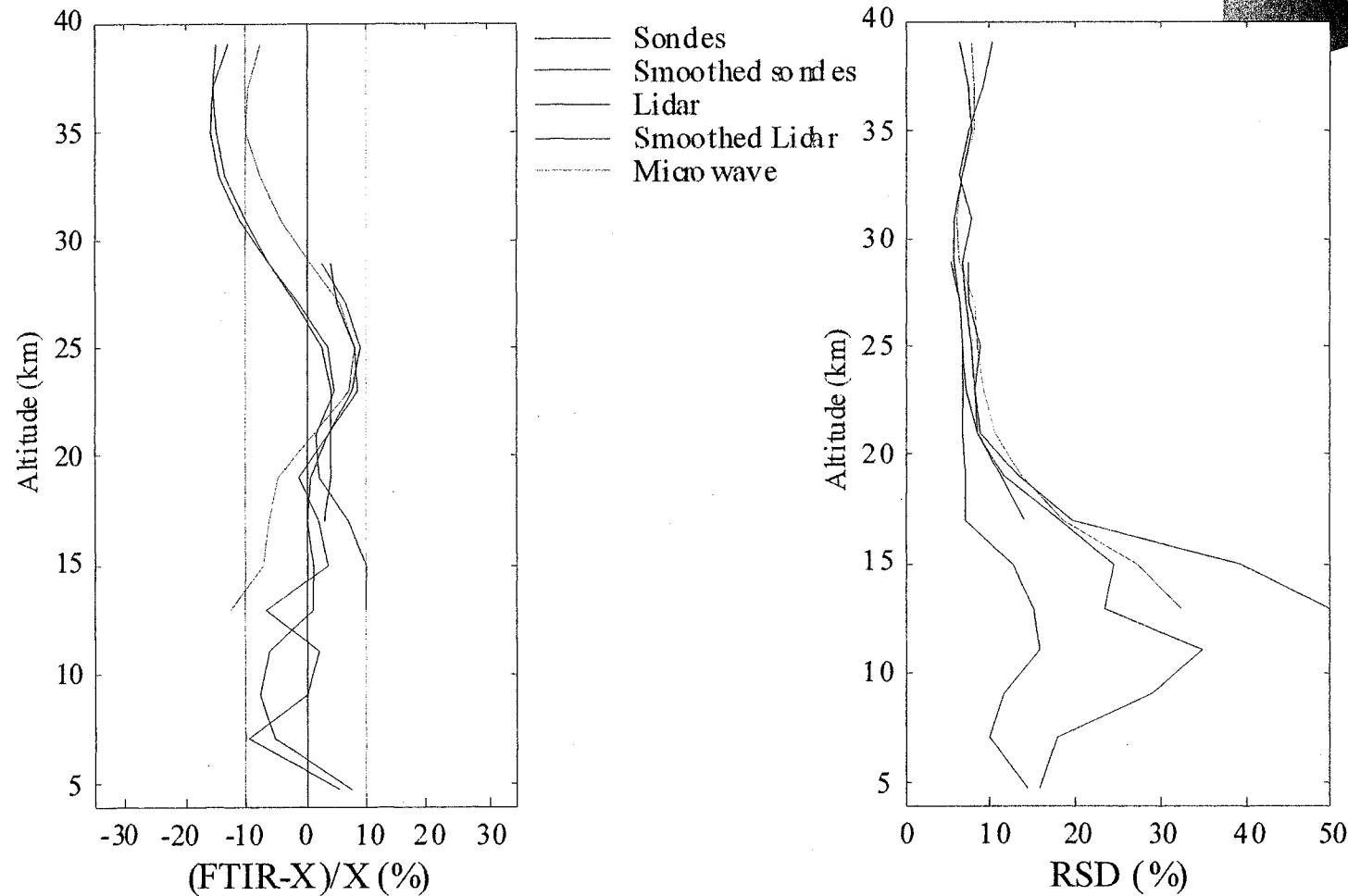
FTIR inversion: smoothing and measurement errors - 2

$$S_A = (A - I)S_x(A - I)^T$$

$$S_y = G_y S_\varepsilon G_y^T$$



FTIR inversion: smoothing error

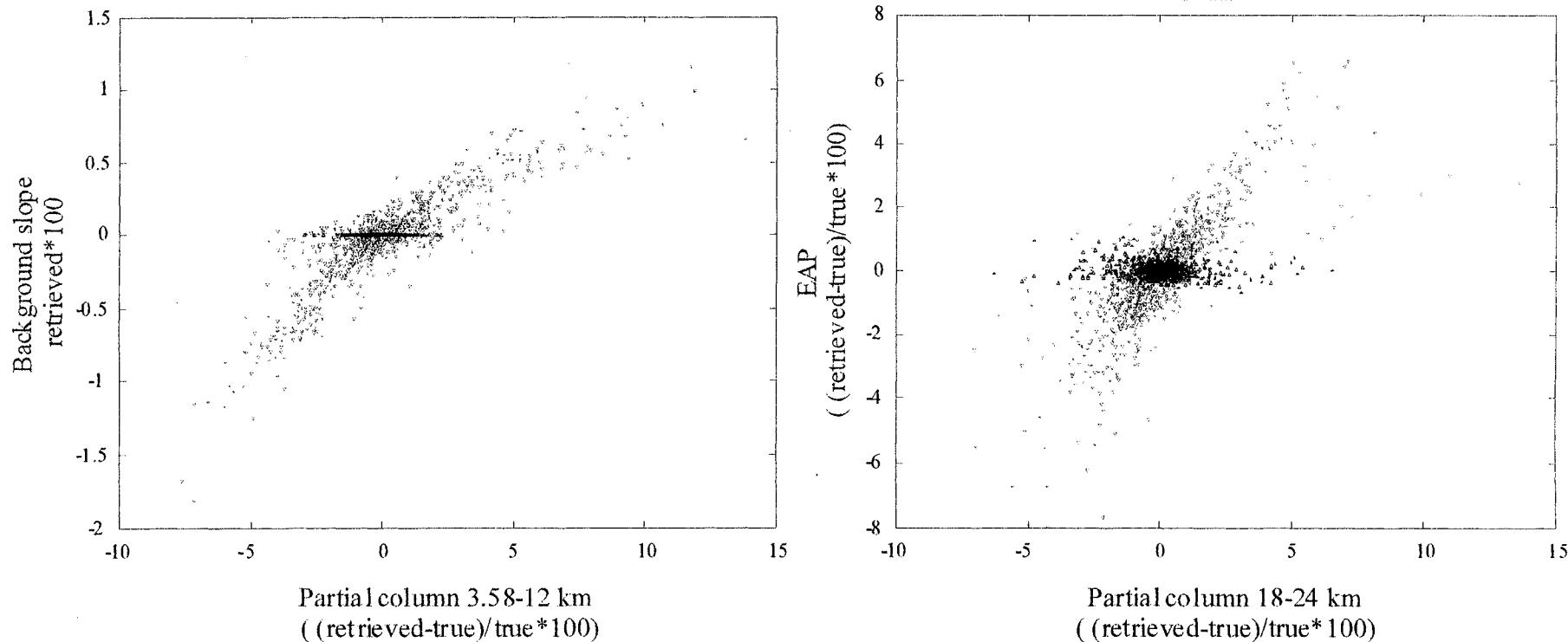


FTIR inversion: model params errors

- Spectroscopy (systematic, altitude-independent)
- P/T (random, altitude dependent)
- background slope, zero level of spectrum
- instrument parameters: EAP

estimated from Monte Carlo simulations

FTIR inversion: correlations between retrieved parameters - 1



FTIR inversion: correlations between retrieved parameters - 2

Layer	3.6-12 km		12-18 km		18-24 km		24-40 km	
Microwindow	N	B	N	B	N	B	N	B
<hr/>								
Model parameter								
Background slope	0.9	0.1	-0.6	0.1	0.2	0.0	0.0	0.0
Effective apodization parameter	0.4	0.0	-0.5	0.2	0.8	-0.1	-0.3	0.3
Zero transmission level	0.5	-0.1	-0.5	0.3	0.7	-0.3	-0.1	0.4
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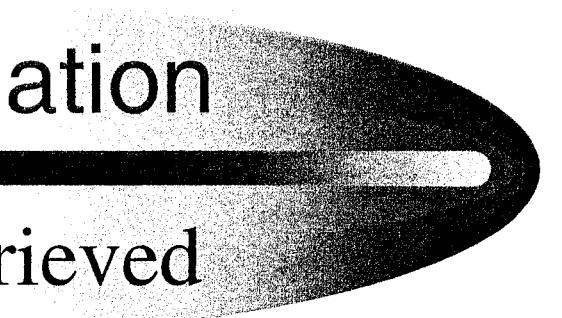
Correlation coefficients between the relative errors ((retrieved-true)/true) of the retrieved partial columns and the retrieved model parameters resulting from the Monte-Carlo study (see text) for retrievals performed in the narrow microwindow (N) and in the broad microwindow (B).

FTIR inversion: total error budget

	Layer km	3.6-12 km	12-18 km	18-24 km	24-40 km	VCA
Systematic Error Source						
Air broadening coefficient uncertainty	4.6	1.6	0.9	-3.9	0.1	
Line intensity uncertainty	2.0	2.0	2.0	2.0	2.0	
EAP uncertainty	0.8	1.9	8.4	-7.9	0.6	
Total systematic error	5.1	3.2	8.7	9.0	2.1	
Random Error Source						
Temperature uncertainty	1.9	2.8	0.3	4.9	3.3	
Measurement noise	1.3	2.2	1.7	1.1	0.2	
Null-space error	7.8	9.2	3.7	1.6	0.2	
Total random error	8.1	9.9	4.1	5.3	3.3	

Systematic and random errors budgets (%) for the retrieval of ozone partial columns in the broad microwindow $1000 - 1005 \text{ cm}^{-1}$.

FTIR inversion: validation



- Comparison between FTIR retrieved profiles at the Jungfraujoch and correlative data, from sonde, LIDAR and microwave, from June 1996 to November 2000

%

▼ FTIR

3.6-12 km

◆ Sonde

0.3 ± 15

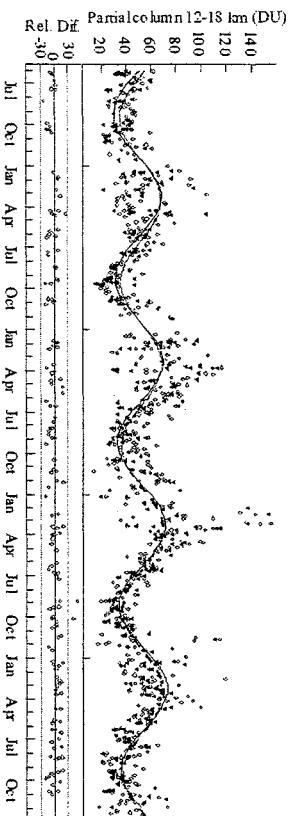
····· MW

12-18 km

● LIDAR

40
 0.1 ± 10

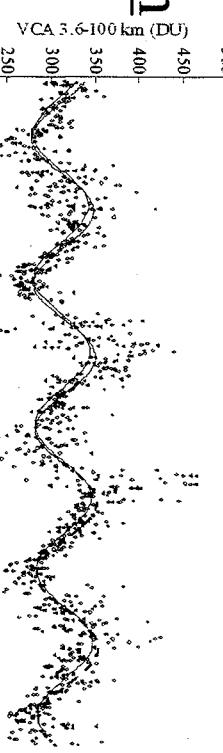
◇ Dobson



3.9 ± 6

● LIDAR

18-24 km



16
 1.5 ± 9

◇ Dobson

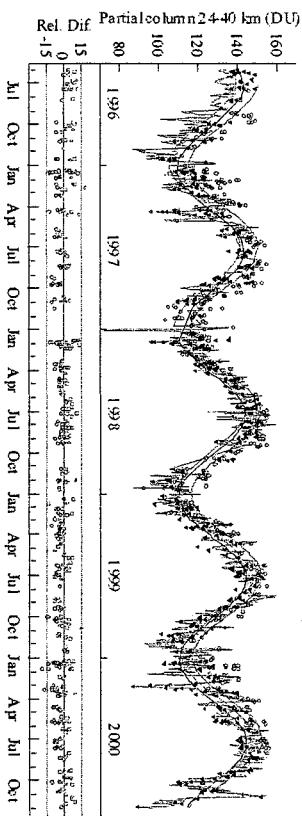
24-40 km



12

11

-0.8 ± 3

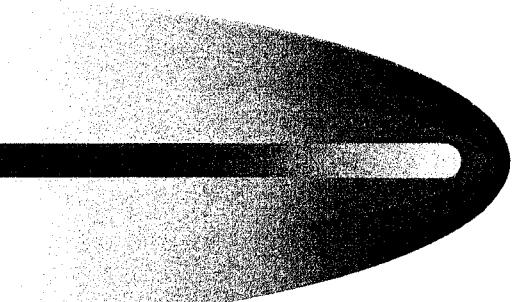


0.3 ± 5

Conclusions

- O₃ profiles can be retrieved from GB high-resolution FTIR spectra in the 1000-1005 cm⁻¹ μwindow, in 4 independent layers, incl. the troposphere, covering distinct atmospheric regimes
- The theoretical characterisation of the OEM inversion has been confirmed by a statistical comparison with independent measurements
- The inversion can cope to some extent with model and instrument uncertainties.

Conclusions



- Perspective:
application to space-based nadir viewing
FTIR experiments, like IASI/METOP-1,
TES/Aura, ...