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international centre for theoretical physics

SMR/1328/3 bis

**School on the Physics of Equatorial Atmosphere**  
**(24 September - 5 October 2001)**

Appendix to  
*Gravity Waves*

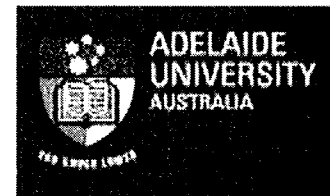
R. A. Vincent  
(University of Adelaide)



**Gravity Waves**  
**International School on the**  
**Physics of the Equatorial Atmosphere**  
**and Ionosphere**  
**ICTP Trieste**  
**September 2001**



**R. A. Vincent**



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**Adelaide University**

# Basic Wave Parameters

- **The following information is required to assess the importance of gravity waves in determining the state of the tropical atmosphere**
  - **Wave amplitudes or energy as a function of time and space**
  - **Propagation directions in both the horizontal and vertical**
  - **Relative importance of different wave sources**
  - **Energy and momentum fluxes as a function of frequency and wavenumber**
- **How do we obtain this information from basic observations of wind, temperature and other observed quantities?**

# Wave Energy

- **Filter wind or temperature field to obtain wave perturbations or fluctuations**
- **Compute mean square amplitudes and hence energy**

$$U = (u, v, w) \text{ where } u = \bar{u} + u' \text{ etc}$$

$$E = E_k + E_p$$

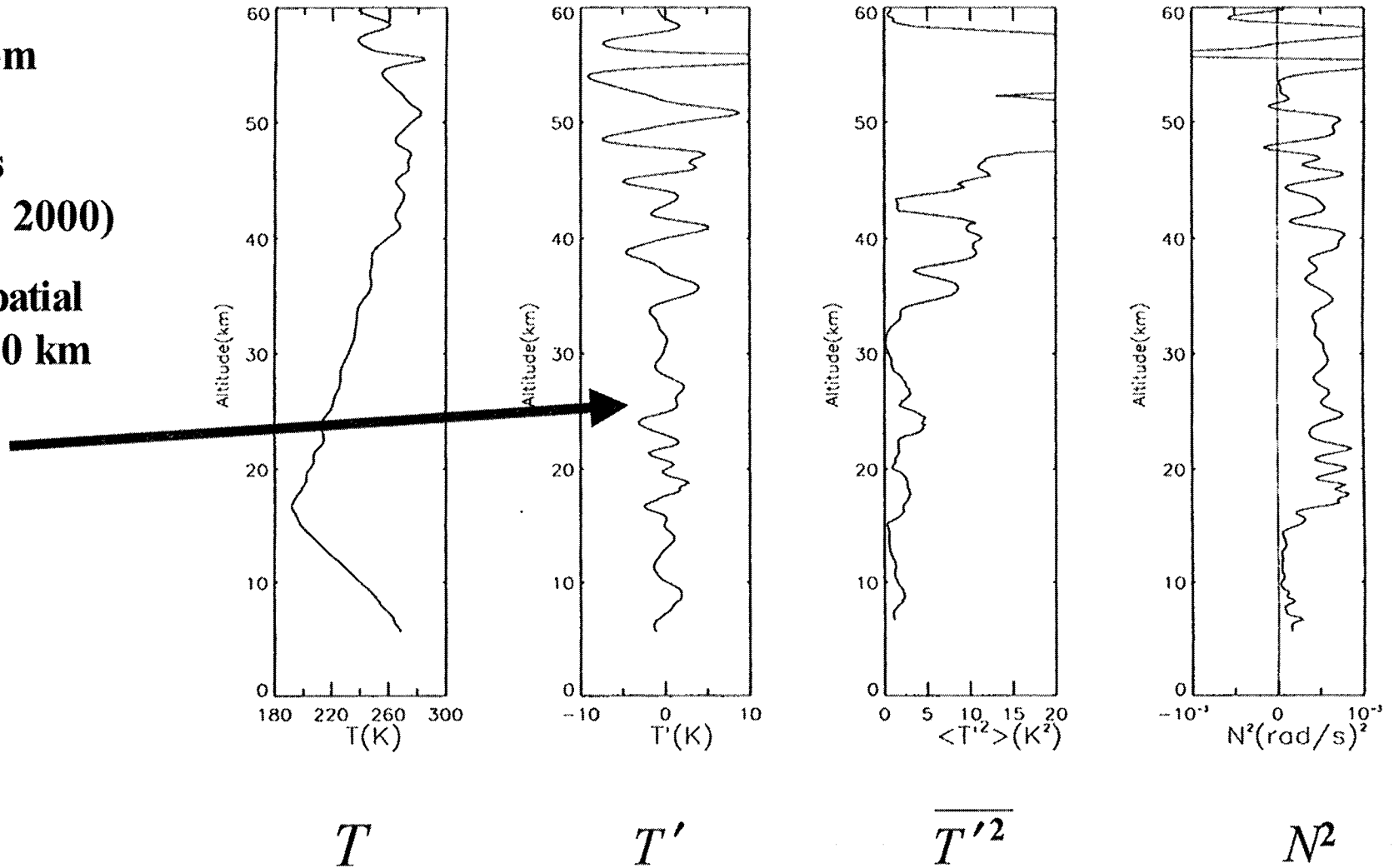
$$E = \frac{1}{2} \left( \overline{u'^2} + \overline{v'^2} + \overline{w'^2} + \frac{g^2}{N^2} \overline{\hat{T}^2} \right)$$

$$\text{where } \hat{T} = \frac{T'}{\bar{T}}$$

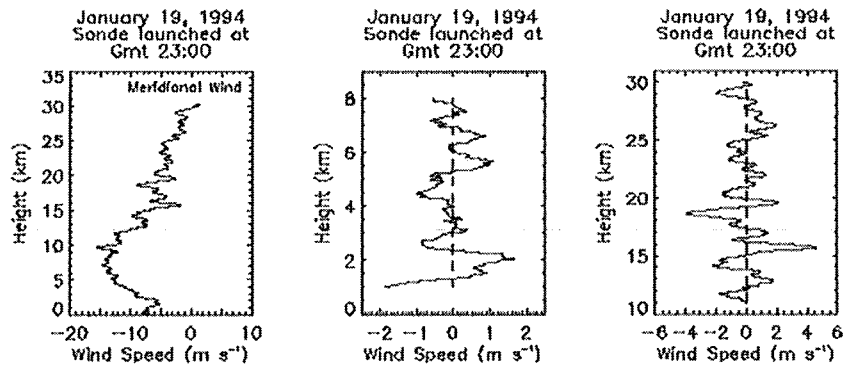
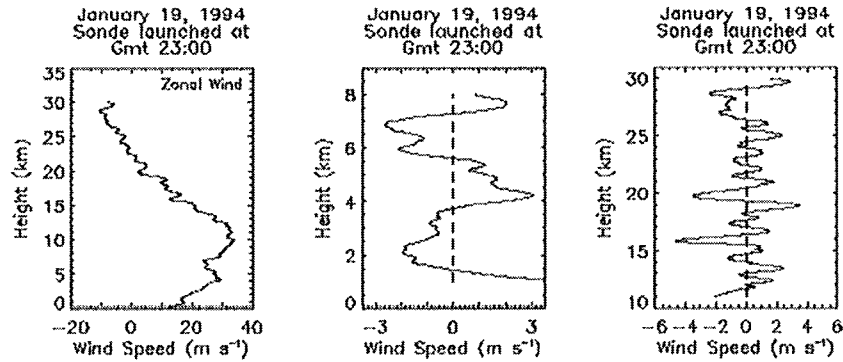
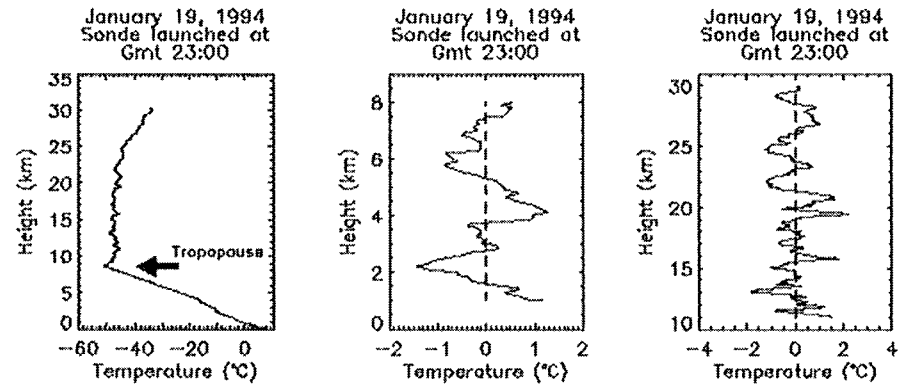
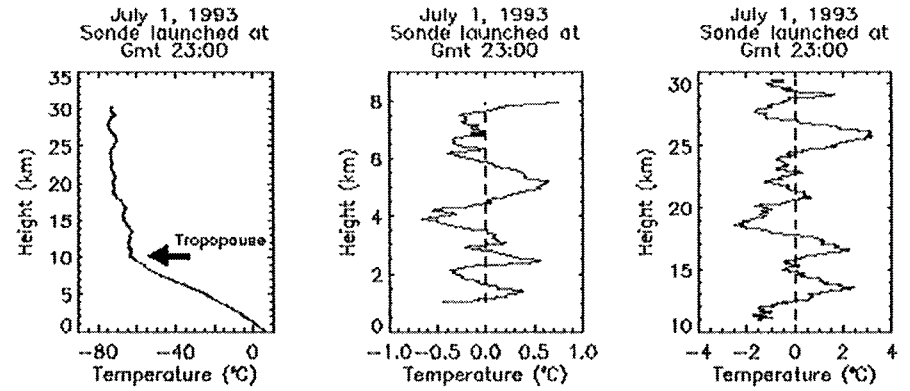
04:18:44 UT on Oct 20, 1995 at 11.08S 98.48E

**Example from  
GPS/MET  
observations  
(Tsuda et al, 2000)**

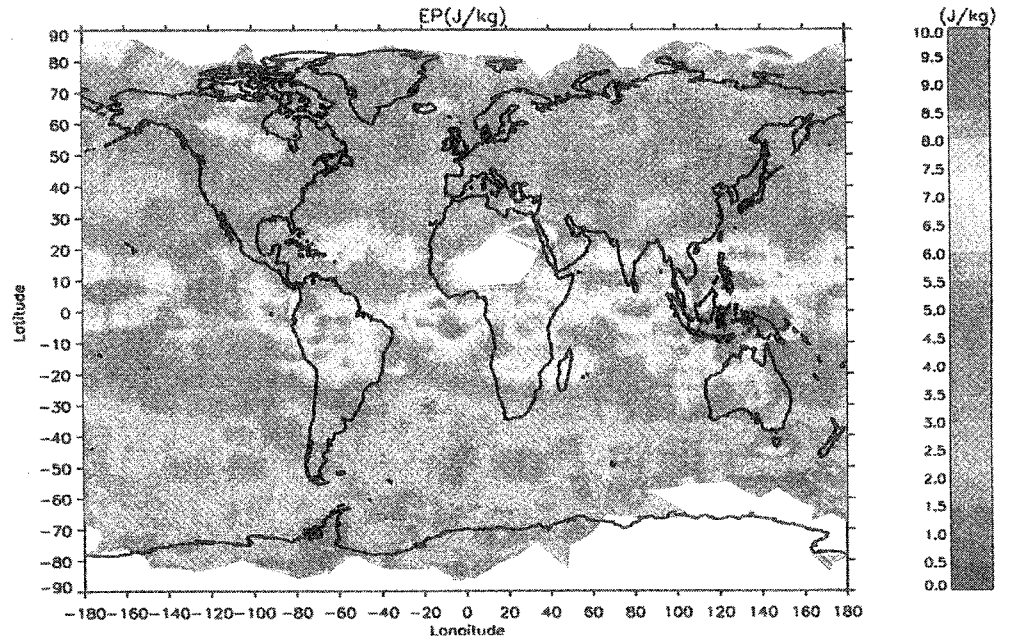
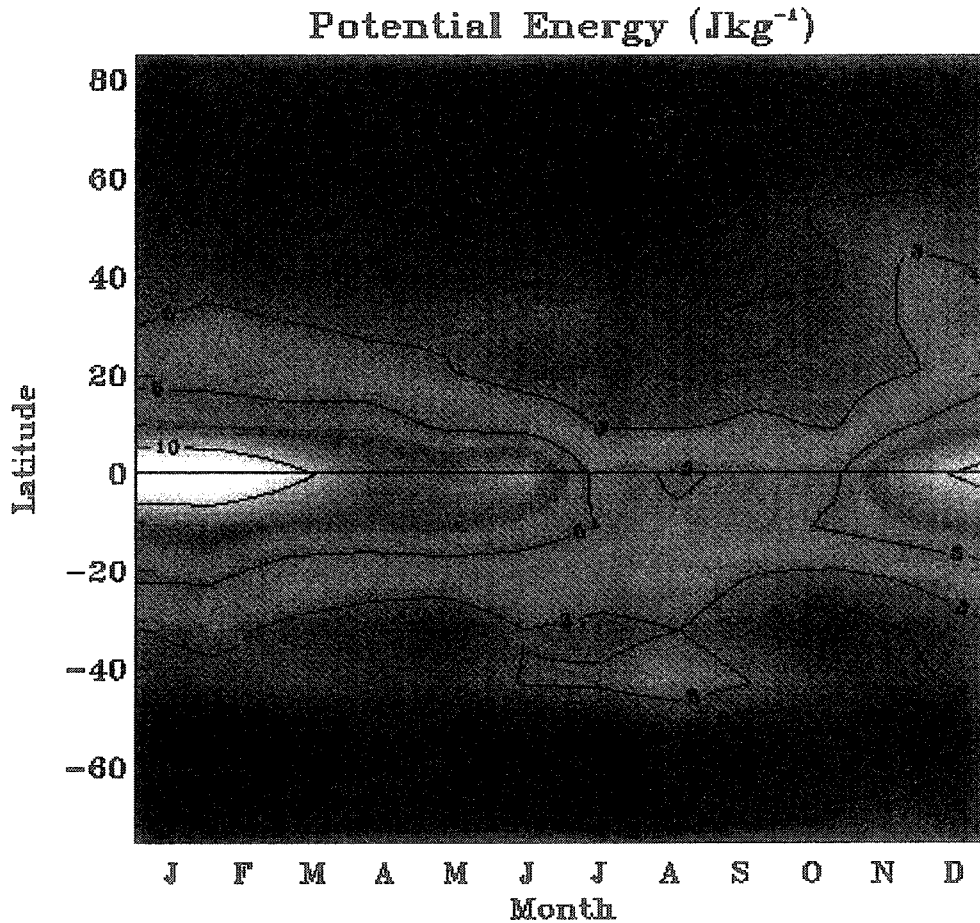
**High pass spatial  
filter with 10 km  
cut-off**



- Synoptic radiosonde observations are a valuable resource
- Fit polynomial to remove background



# $E_p$ Climatologies



GPS/MET Nov-Feb 20-30 km

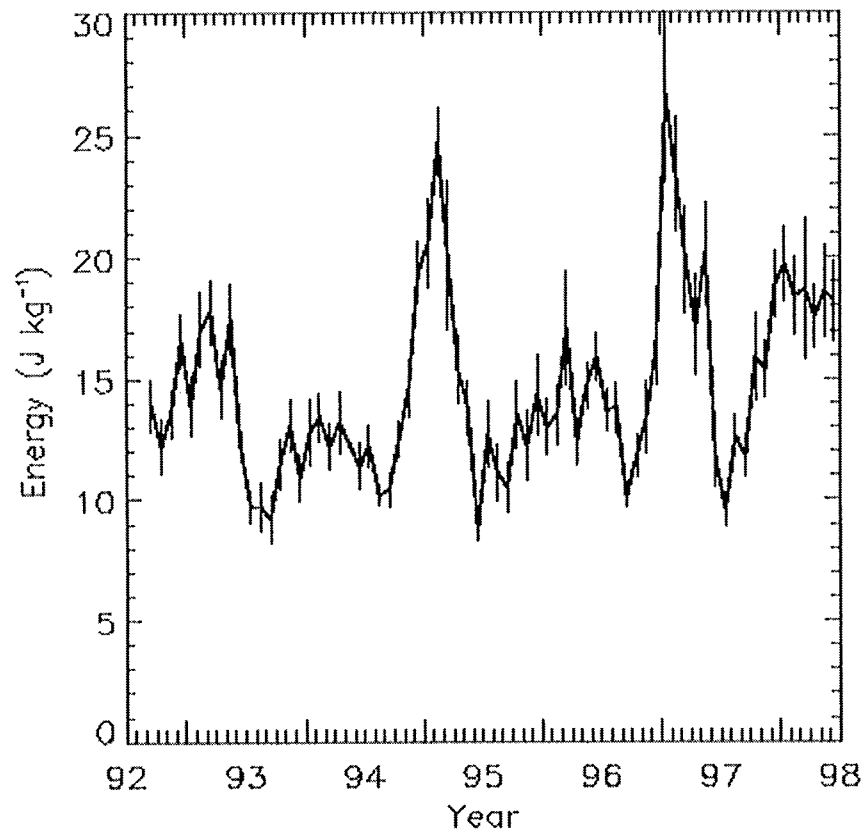
SPARC radiosonde climatology

17-24 km

Note large equatorial amplitudes.  
Source effect or observational  
bias?



# Tropical GW

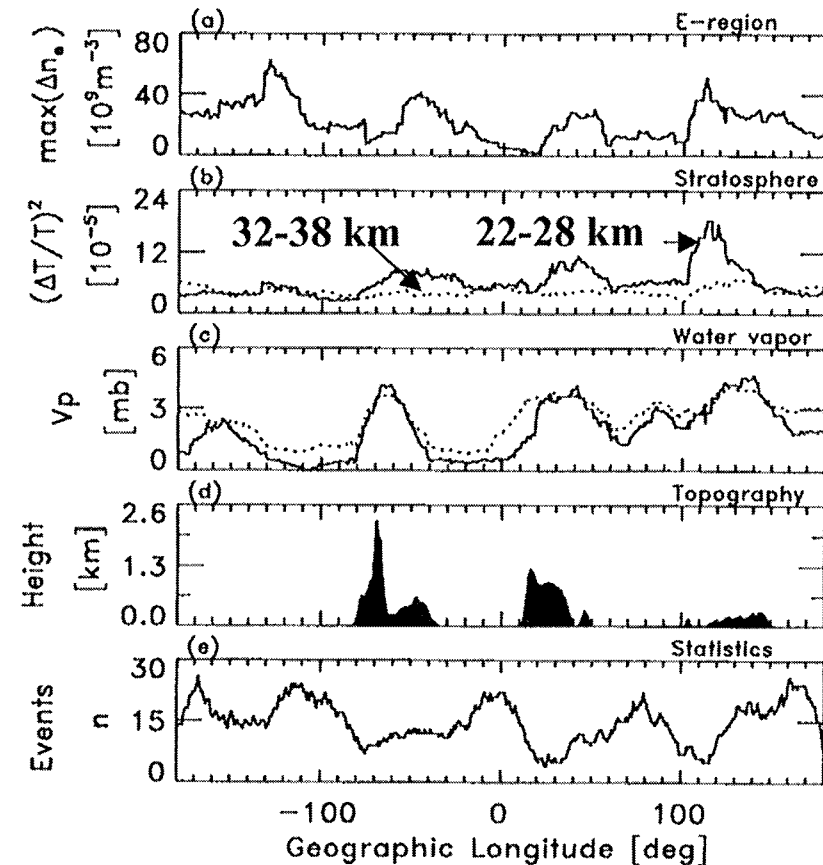


**Six year time series of  $E$  in 18-25 km range at Cocos Is ( $12^\circ\text{S}$ ). Note annual and QBO-like cycles. Source or wind filtering effects? (Vincent and Alexander, 2000)**

9/25/01

International School on Equatorial Atmosphere: Gravity Waves

## GPS/MET observations (Hocke and Tsuda, 2001)



**Note apparent geographic relation between location of enhanced fluctuations in  $N$ ,  $T$ , and humidity (convection?).**

# Definitions

To determine direction of propagation in both the vertical and horizontal use the polarization relations that relate the perturbation quantities,  $u'$ ,  $v'$ ,  $w'$ , and  $T'$ .

Assume that we are dealing with harmonic waves in a rotating fluid

$$u', v', w', T' \propto \rho^{-1/2} e^{i(kx+ly+mz-\hat{\omega}t)}$$

$$v' = \frac{\hat{\omega}l - ifk}{\hat{\omega}k + ifl} u'$$

$\hat{\omega}$  = intrinsic frequency

$\omega$  = ground - based (observed) frequency

$$\hat{\omega} = \omega - \vec{k}_h \cdot \vec{U}$$

$f$  = Coriolis parameter

$$m^2 = \frac{N^2 - \hat{\omega}^2}{\hat{\omega}^2 - f^2} k_h^2 - \frac{1}{4H^2}$$

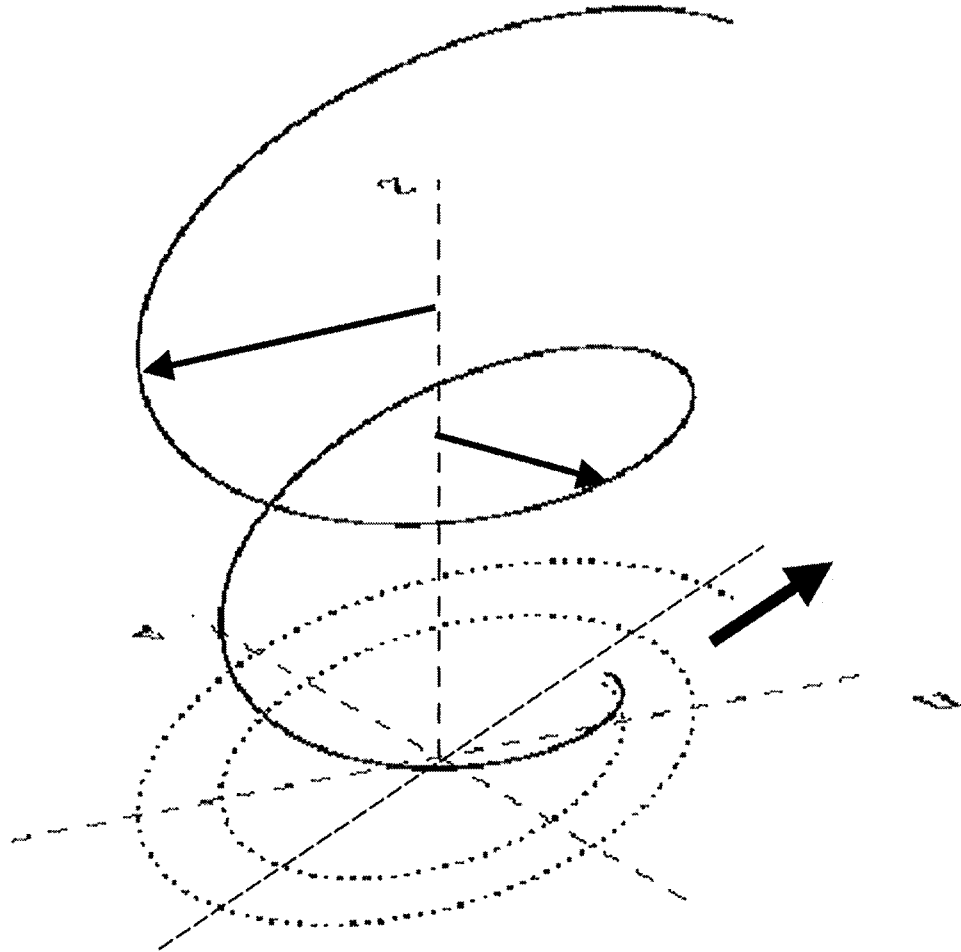
$\hat{c} = (\hat{\omega}/k, \hat{\omega}/l) =$  intrinsic phase speed

For many waves it is possible to assume

$$m^2 \gg \frac{1}{4H^2}$$

$$m = \pm \sqrt{\frac{N^2 - \hat{\omega}^2}{\hat{\omega}^2 - f^2}} k_h$$

# Propagation Directions: Anisotropy



- Hodograph analysis
- For upward energy propagation
  - Clockwise rotation in NH ( $f > 0$ )
  - Anticlockwise in SH
- Circular polarization as  $\hat{\omega} \rightarrow f$
- Linear polarization as  $\hat{\omega} \rightarrow N$

# Vertical Propagation

- **Rotary decomposition provides a statistical approach to hodograph analysis**

$$V(z) = u'(z) + iv'(z) = ae^{-imz}$$

$$V(z) \leftrightarrow \Psi(m)$$

$$\text{Clockwise component } C(m) = \frac{1}{2} \Psi(-m) \Psi^*(-m)$$

$$V(z) \leftrightarrow \Psi(m)$$

$$\text{Anticlockwise component } A(m) = \frac{1}{2} \Psi(m) \Psi^*(m)$$

$$\text{Total Energy, } E(m) = C(m) + A(m)$$

$$\text{In NH (SH) fraction of upward energy} \quad r_N = \frac{\overline{C(m)}}{\overline{E(m)}} \quad \left( r_S = \frac{\overline{A(m)}}{\overline{E(m)}} \right)$$

# Horizontal Direction

- If axes are aligned parallel and perpendicular to major and minor axes of hodograph ellipse then

$$v'_{\perp} = -i \frac{f}{\hat{\omega}} u'_{\parallel}$$

$$\text{Axial ratio, } AR = \frac{u'_{\parallel}}{v'_{\perp}} = \frac{\hat{\omega}}{f}$$

- Hodograph gives
  - direction of propagation (with  $180^{\circ}$  ambiguity)
  - intrinsic frequency.

Statistical approach is better.

# Stokes Parameters

- Note similarity between gravity wave and electromagnetic wave
- Both are transverse oscillations that are in general elliptically polarised
- Stokes parameters are relations that determine sense and degree of polarisation

$$I = \overline{u'^2} + \overline{v'^2}$$

$I$  = total intensity

$$D = \overline{u'^2} - \overline{v'^2}$$

$D$  = intensity difference

$$P = \overline{2u'v' \cos \delta}$$

$P$  = Linear polarization parameter

$$Q = \overline{2u'v' \sin \delta}$$

$Q$  = circular polarization parameter

# Stokes Parameters II

- **$Q$  gives sense of rotation**
  - $Q > 0 \rightarrow$  Anticlockwise polarisation
  - $Q < 0 \rightarrow$  Clockwise polarisation
- **Degree of polarisation**

$$d = \frac{(D^2 + P^2 + Q^2)^{1/2}}{I}$$

$d = 1 \Rightarrow$  perfect polarisation

$d = 0 \Rightarrow$  unpolarised

# Polarisation Ellipse

- **Polarisation ellipse parameters**

$$\delta = \tan^{-1}\left(\frac{Q}{P}\right)$$

$$2\phi = \tan^{-1}\left(\frac{P}{D}\right)$$

$$AR = \cot \varepsilon \text{ where } 2\varepsilon = \sin^{-1}\left(\frac{Q}{dI}\right)$$

**Apply in frequency domain using radar winds data acquired at a single height**

**Apply in vertical wavenumber domain using rocket or radiosonde wind profiles**



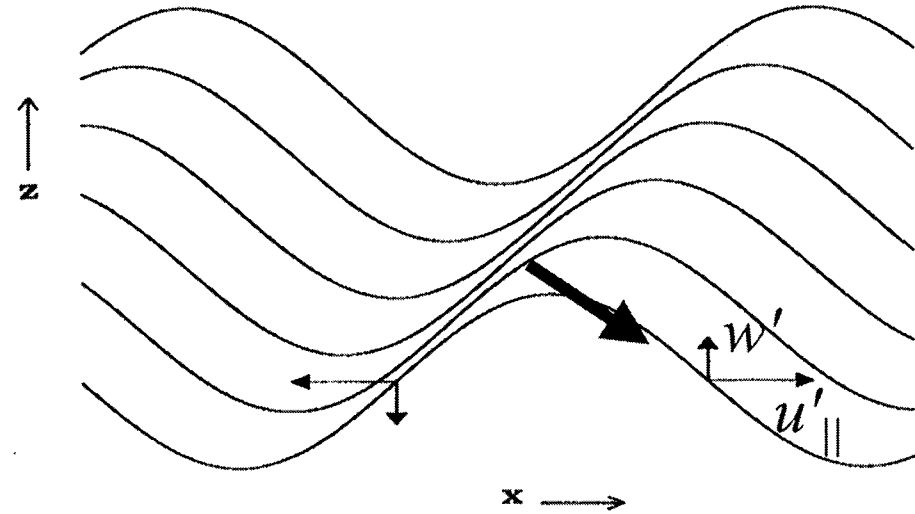
# Horizontal Directions

- Require extra information to resolve 180° directional ambiguity from hodograph
- Use either vertical velocity or temperature

$$w' = -\frac{k_h}{m} u'_{\parallel}$$

$w'$  in - phase with  $u'$  for  $k_h > 0$  and  $m < 0 \Rightarrow \overline{u'_{\parallel} w'} > 0$

$w'$  out - of - phase with  $u'$  for  $k_h < 0$  and  $m < 0 \Rightarrow \overline{u'_{\parallel} w'} < 0$



# Horizontal Directions II

- Temperature polarisation relation

$$\hat{T}' = \mp i \frac{N^2}{g\hat{\omega}} \frac{k_h}{m} u'_{\parallel}$$

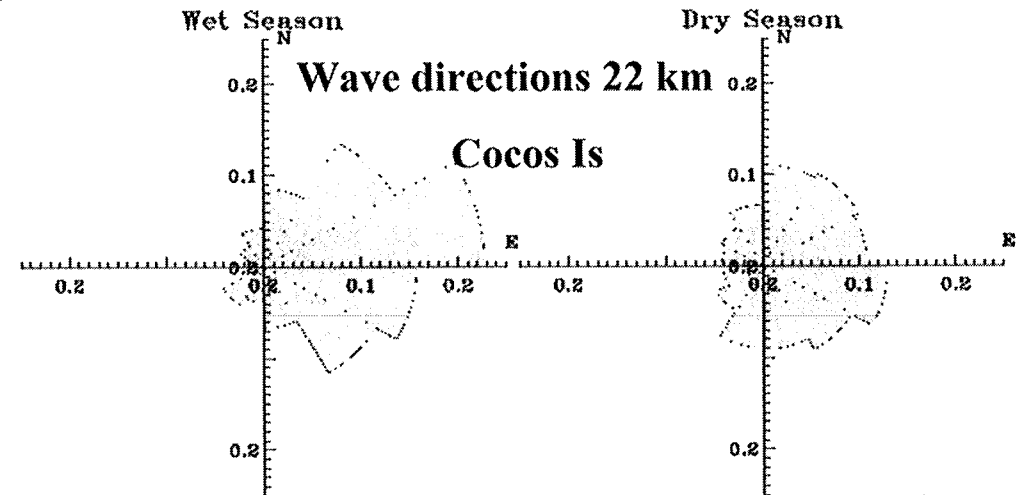
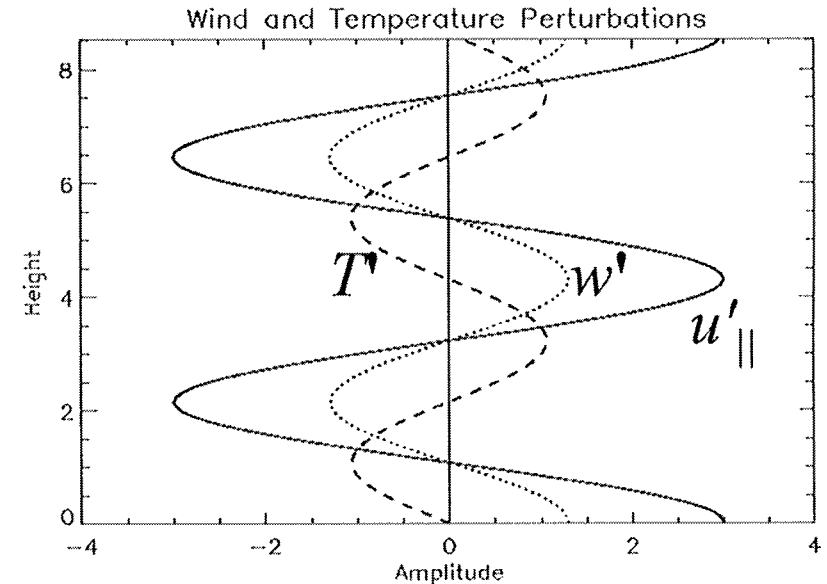
– sign chosen when  $km > 0$

+ sign chosen when  $km < 0$

Note 90° phase shift between  $T$  and  $u$

$$\overline{u'_{\parallel} w'} = \frac{\hat{\omega} g}{N^2} \overline{u'_{\parallel} \hat{T}'_{+90}} \quad \text{where } \hat{T}'_{+90} = i\hat{T}'$$

Sign of  $\overline{u'_{\parallel} \hat{T}'_{+90}}$  determines  
direction of propagation



# Momentum Fluxes

$$\overline{u'w'}, \overline{v'w'}$$

- Important quantities, but difficult to measure directly  
( $w' \ll u', v'$ )
- Indirect estimates using temperature require knowledge of  $\hat{\omega}$

If  $\overline{U} = 0$  then  $\hat{\omega} = \omega$ , the ground - based frequency

# Radar Measurements

$$\overline{u'w'}, \overline{v'w'}$$

- **Dual-beam Doppler technique**
- **Measure radial velocity at range,  $R$ , along beams 1 and 2**

$$v_1(\theta, R) = u' \sin \theta + w' \cos \theta$$

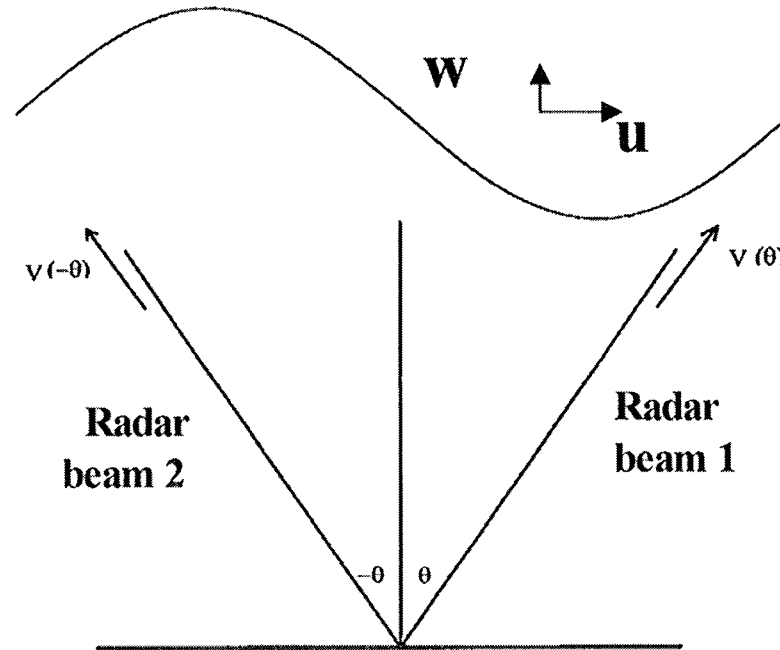
$$v_2(-\theta, R) = u' \sin \theta - w' \cos \theta$$

$$v_1^2(\theta, R) = u'^2 \sin^2 \theta + 2u'w' \cos \theta \sin \theta + w'^2 \cos^2 \theta$$

$$v_2^2(-\theta, R) = u'^2 \sin^2 \theta - 2u'w' \cos \theta \sin \theta + w'^2 \cos^2 \theta$$

**Subtract and rearrange**

$$\overline{u'w'}(z) = \frac{\overline{v_1^2(\theta, R)} - \overline{v_2^2(-\theta, R)}}{2 \sin 2\theta}$$



**Strengths:**

- **Robust**
- **Direct**

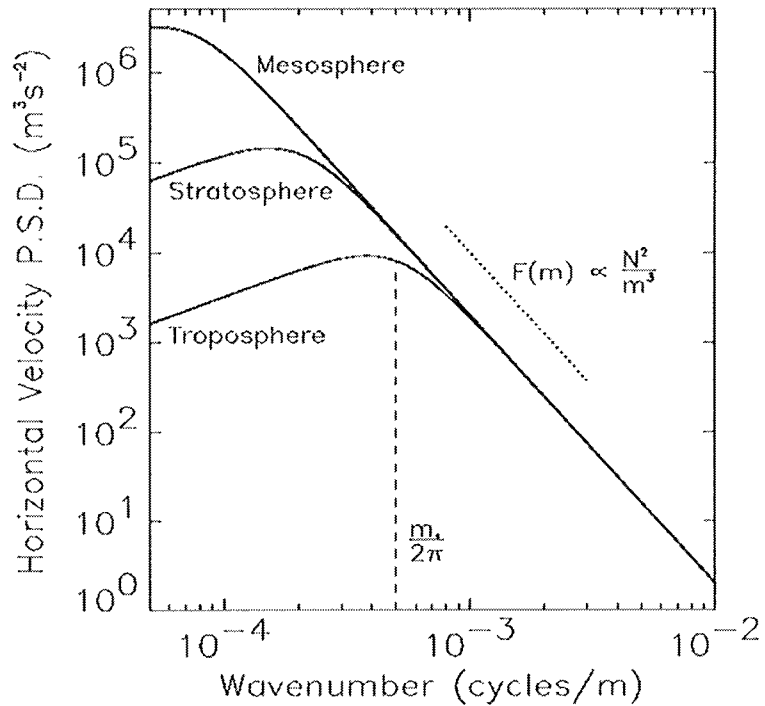
**Limitations:**

- **Require accurate beam angles**
- **Small differences**

# Spectra

- Frequency and wavenumber spectra are a convenient way of summarizing wave information
- Suggestion that there is a “universal” wave spectrum
- Spectral shapes and amplitudes set by instability processes
- Many different theories for producing spectral features
- Observations show that
  - Frequency spectra derived from ground-based radar and lidar observations of horizontal winds show  $S_u(\omega) \propto \omega^{-p}$  with  $p \sim 5/3$
  - Recent *constant pressure* balloon observations provide first *intrinsic frequency spectra*  
$$S_u(\hat{\omega}) \propto \hat{\omega}^{-2}$$
  - Vertical wavenumber spectra derived from radiosonde, rocket, and radar observations of winds and temperatures show  $S_{u,T}(m) \propto m^{-t}$  with  $t \sim 3$  for large  $m$

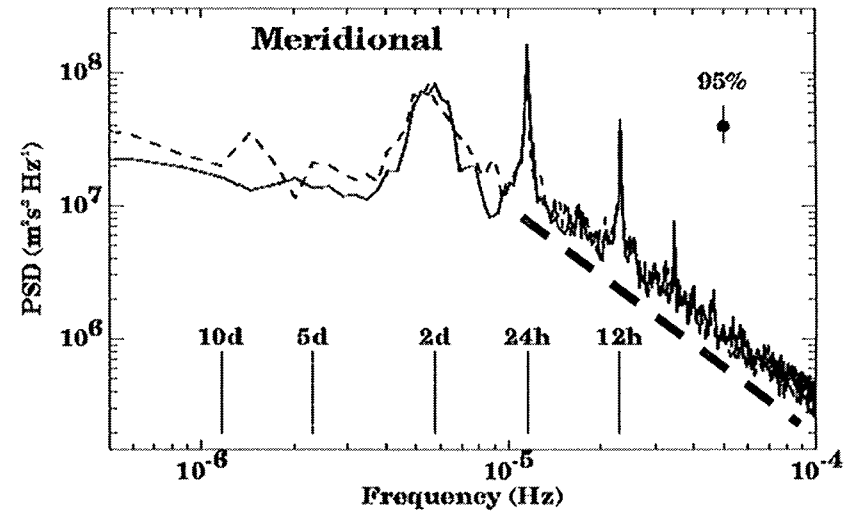
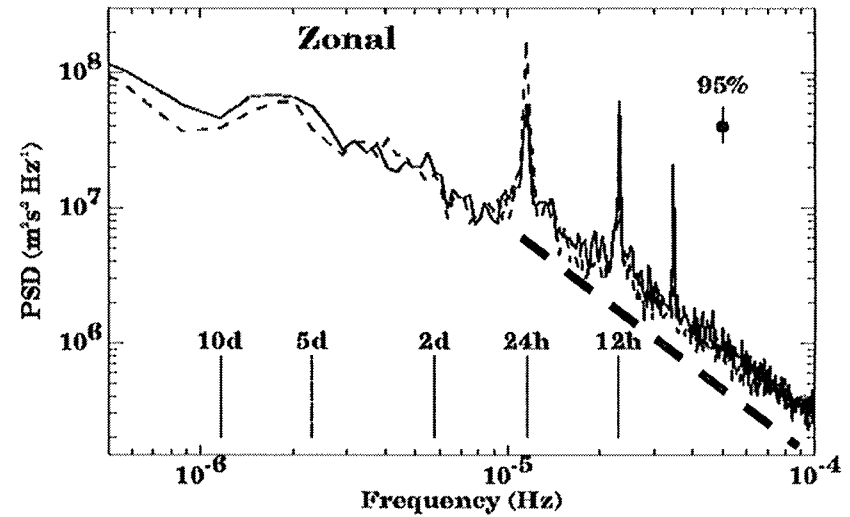
# Spectra: Examples



**Modified Desaubies spectrum**

$$F(\mu) = \frac{\mu^s}{1 + \mu^{s+t}}$$

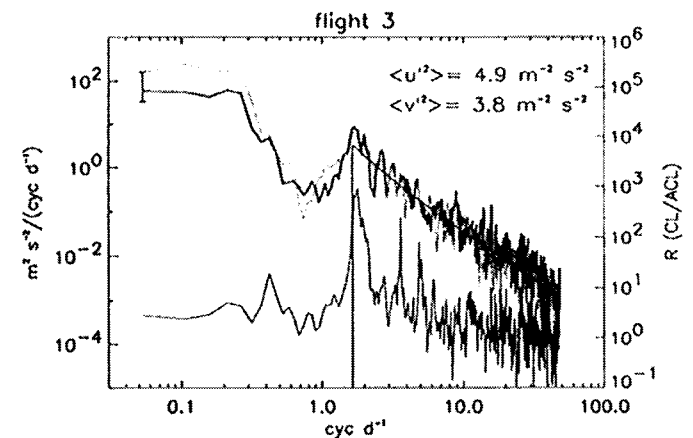
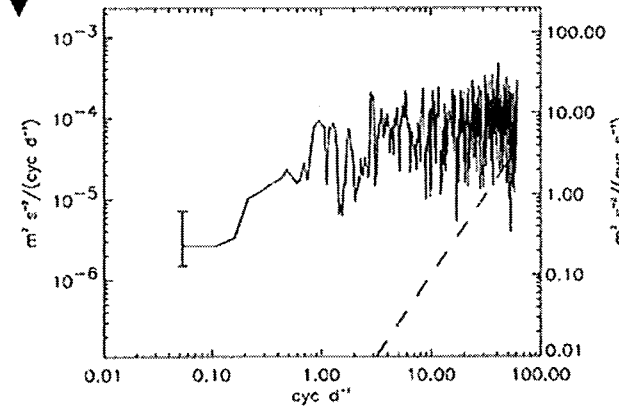
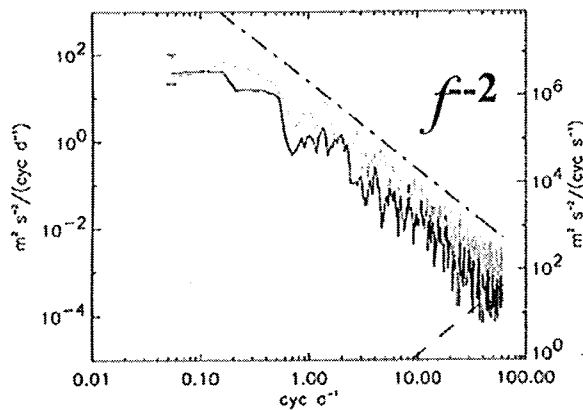
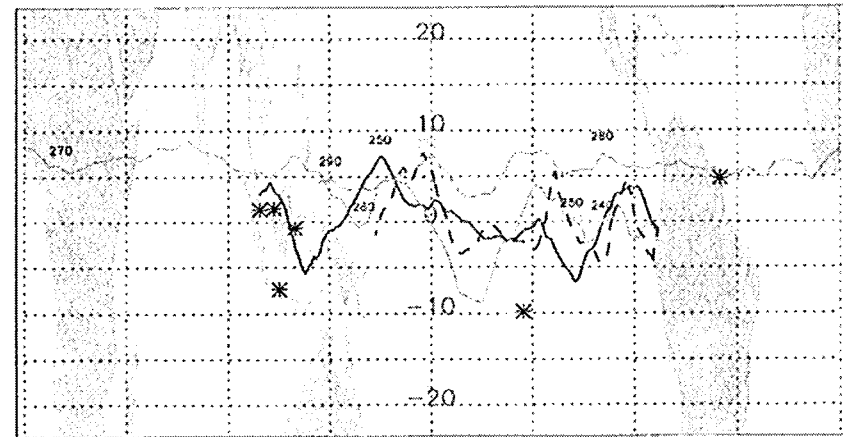
$$\mu \equiv \frac{m}{m_*}, t \sim 3, s \sim 1(?)$$



**Equatorial mesospheric frequency spectra**

# Intrinsic Frequency Spectra

- Trans-Pacific constant-pressure balloon flights
- 19-20 km float altitude
- Zonal, meridional, vertical velocity spectra

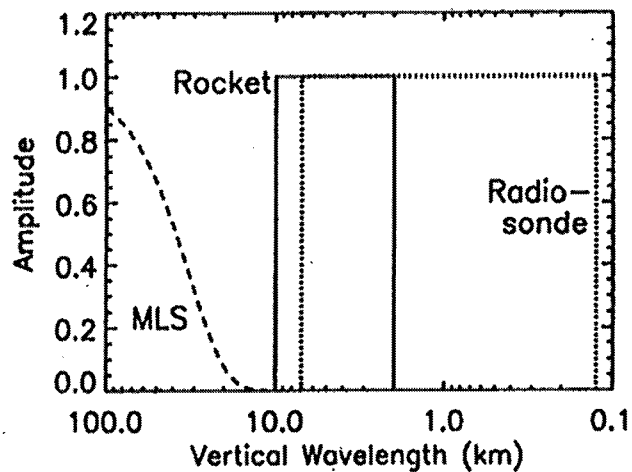


## Circumpolar spectra

# Observational Selection

- **All observational techniques have some bias in which part of the wave spectrum they can observe**
- **No one technique can cover the complete wave spectrum**
- **Important to recognize and quantify the possible biases**





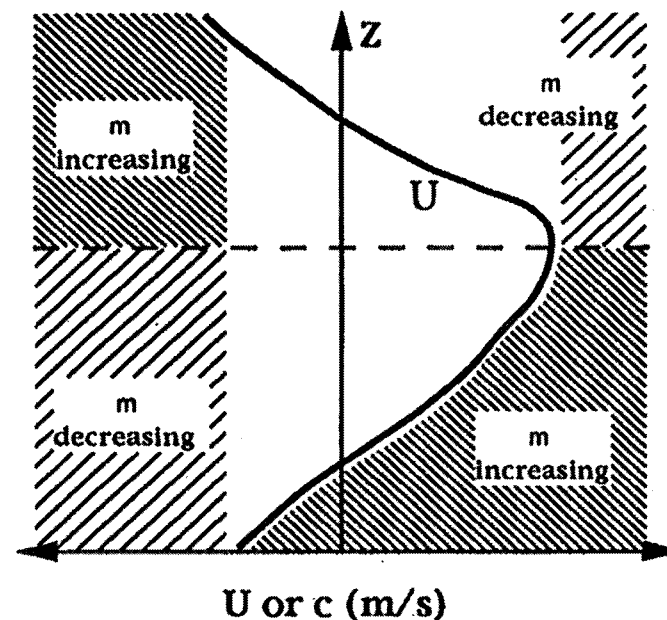
**Observational filters for  
MLS, rockets, (GPS/MET),  
radiosondes**

**Modelling of observations  
helps with interpretation**

**e.g. Alexander and Vincent  
(2000); McLandress et al  
(2000)**

**Simple dispersion relation**

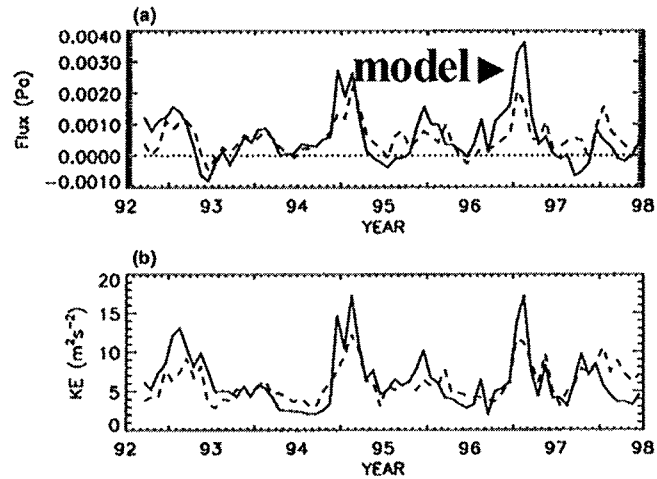
$$|m| \sim \frac{N}{\hat{\omega}} k_h \sim \frac{N}{|u - c|}$$



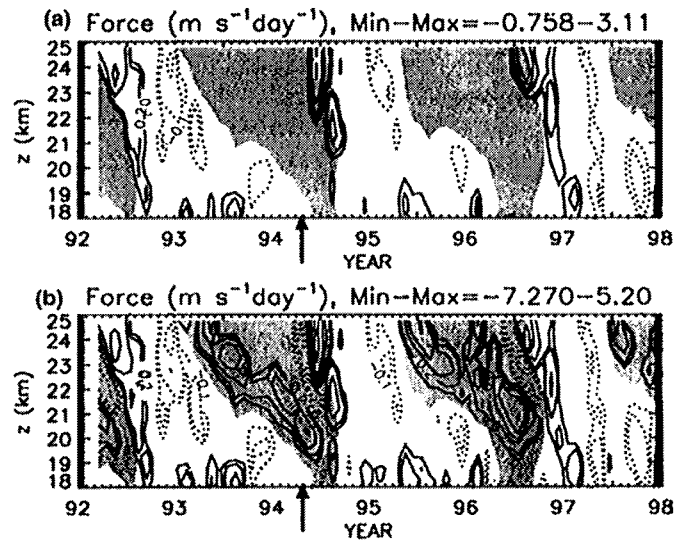
**Waves move in or out of each  
“window” as  $m$  changes due to height-  
varying background winds**

# Modelling of observations helps with interpretation

e.g. Alexander and Vincent (2000); McLandress et al (2000)



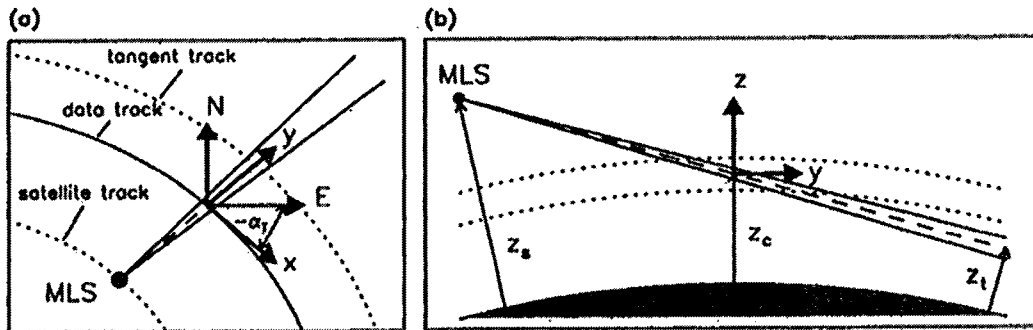
Radiosonde observations constrain model



Gravity-wave driven force

Force estimate from "best-fit" model

Force estimated without allowing for observational selection



## Understanding MLS

