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AUTUMN COLLEGE ON PLASMA PHYSICS

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Magnetic Discontinuities (Current Sheets) in Evolving Magnetic Fields

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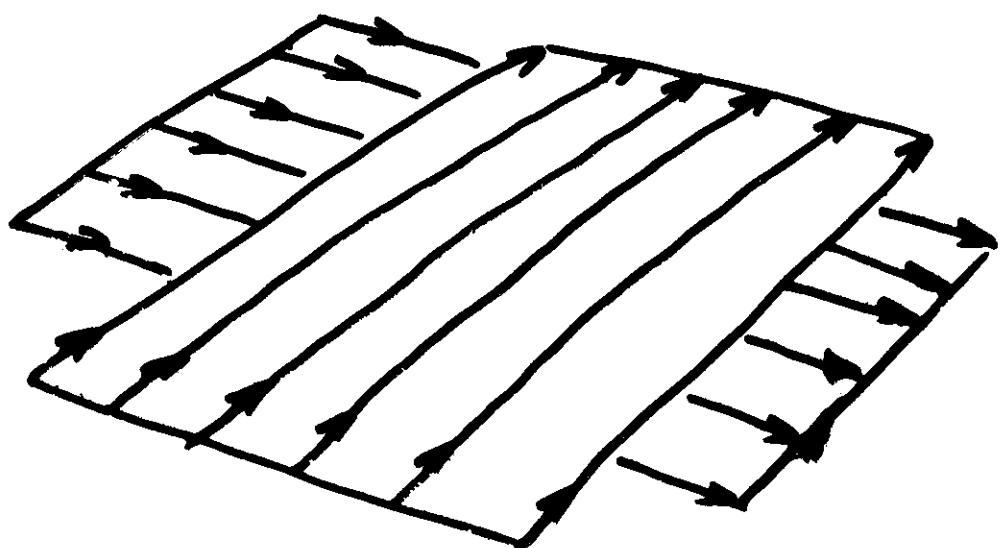
These are lecture notes, intended for distribution to participants.

BASIC MAGNETOSTATIC THEOREM

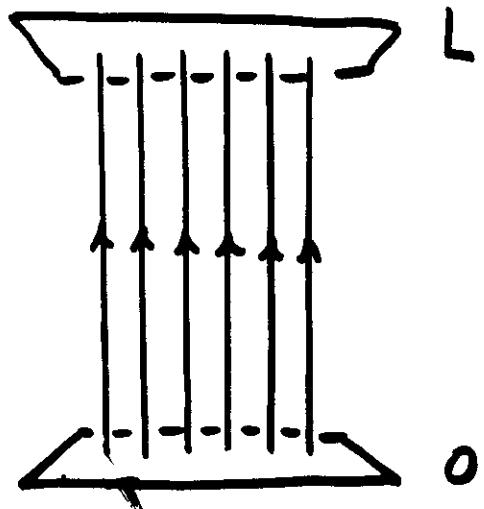
Almost all field line topologies
have tangential discontinuities
(current sheets) as an intrinsic
part of magnetostatic equilibrium.

(assuming an infinitely conducting fluid).

A tangential discontinuity is a rotational
 (directional) discontinuity, with
 continuous field magnitude.



To be clear as to the meaning of these words, consider the simple case where an initially uniform field B_0 extends through a region of infinitely conducting fluid $0 < z < L$. Then introduce the fluid motion



$$v_x = +kz \frac{\partial \psi}{\partial y}, v_y = -kz \frac{\partial \psi}{\partial x}, v_z = 0$$

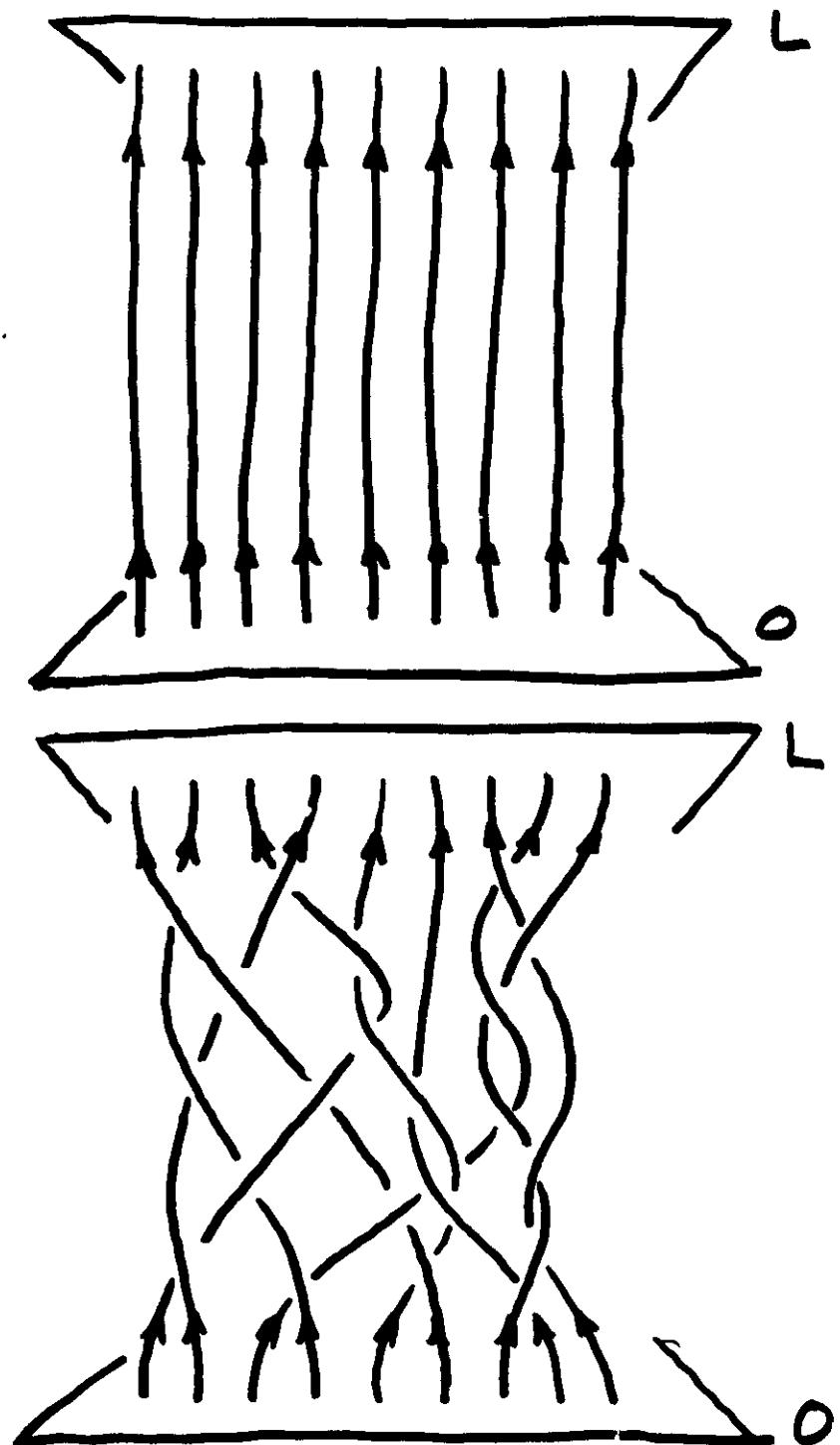
$$\psi = \psi(x, y, kz)$$

which leaves the field fixed at $z=0$ and winds and interweaves the field lines by an continuous mixing introduced at $z=L$. The field becomes

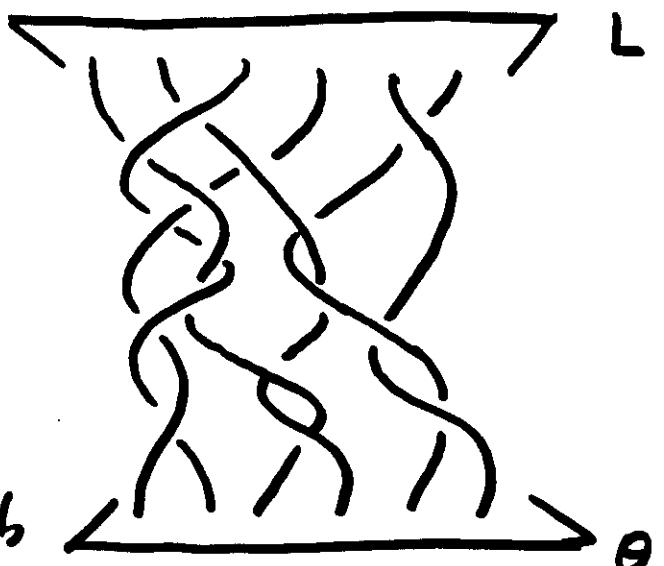
$$B_x = +kt \frac{\partial \psi}{\partial y}, B_y = -kt \frac{\partial \psi}{\partial x}, B_z = B_0$$

after a time t .

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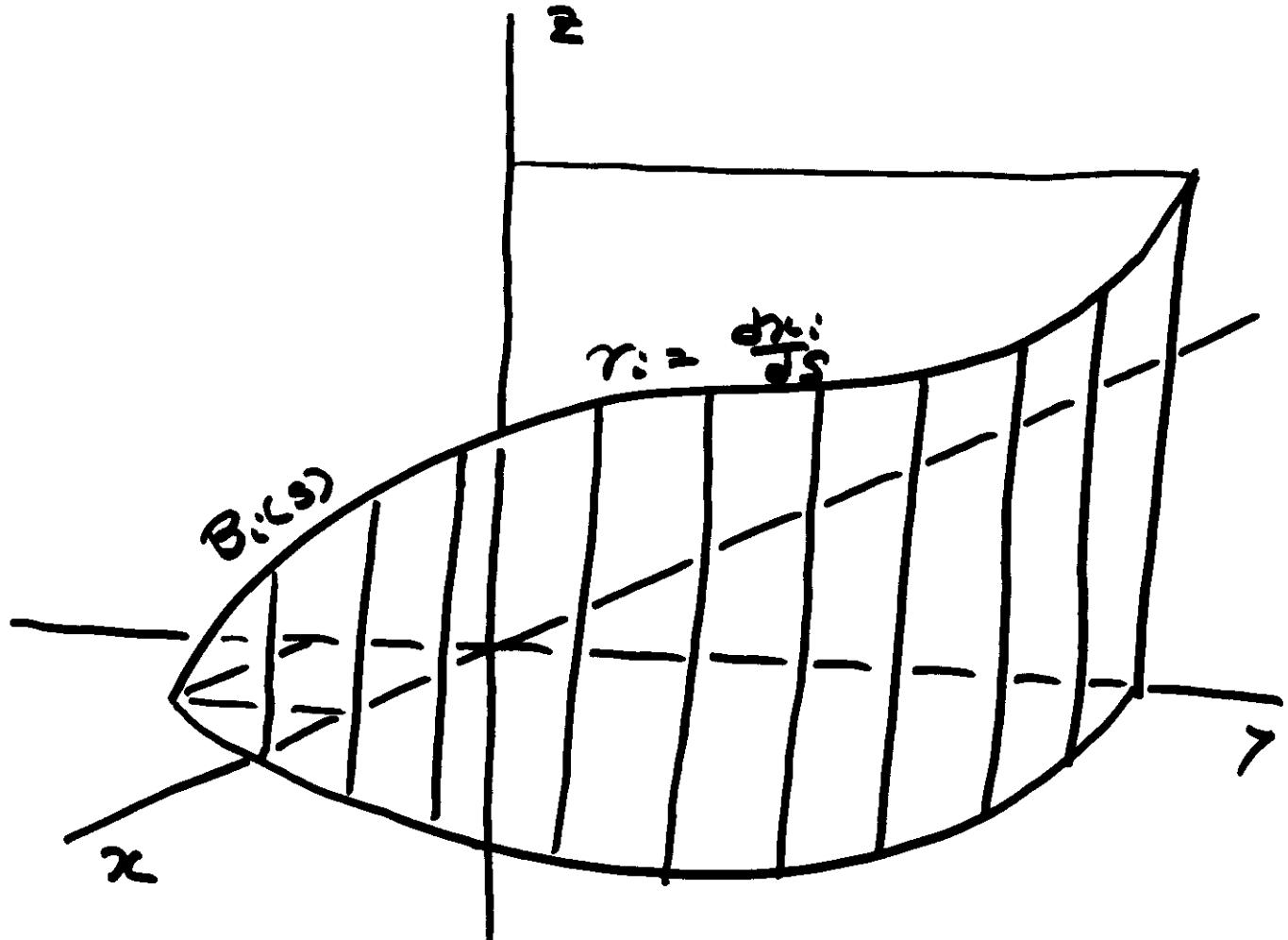
Since $\psi(x, y, kzt)$
is a continuous
bounded, well behaved
function of x, y, kzt ,
it follows that B_z
is continuous, bounded,
etc.



Fix the footpoints of the field at $t=\tau$
 $x=0, L$, and then release the fluid
so that the field relaxes to static
equilibrium. This is the lowest energy
state available for the given topology.
~~among non-potential solutions~~

The essential point is that the
magnetic field develops tangential
discontinuities as it relaxes to
equilibrium, in a finite time, or in infinite time.

In the real world, $\eta \neq 0$, Magnetic
energy is dissipated at the "discontinuities",
which have finite thickness so that they are
never in internal static equilibrium



$$B_i(s) + \delta B_i = B_i(s) + \frac{\partial B_i}{\partial x_j} \delta x_j + \dots$$

All nine $\frac{\partial B_i}{\partial x_j}$ must be available.

$$\frac{d B_i}{ds} = \gamma_j \frac{\partial B_i}{\partial x_j}, \quad B_i \left(\frac{\partial B_i}{\partial x_j} - \frac{\partial B_j}{\partial x_i} \right) = 0$$

6 eqns linear in $\frac{\partial B_i}{\partial x_j}$

Specify $B_i(s)$ and, e.g. $\frac{\partial B_i}{\partial z}$.

Then compute the remaining γ_j .

The remaining six components of $\frac{\partial \mathbf{B}}{\partial x_i}$ cannot be computed if the $\det = 0$.

$$\begin{vmatrix} 0 - B_x + B_y & 0 & B_x & 0 \\ 0 + B_x - B_y & 0 & 0 & B_z \\ 0 & 0 & 0 & -B_x - B_y \\ 1 & 0 & 0 & 1 & 0 & 0 \\ \gamma_x \gamma_y & 0 & 0 & 0 & 0 \\ 0 & 0 & \gamma_x \gamma_y & 0 & 0 \\ 0 & 0 & 0 & 0 & \gamma_x \gamma_y \end{vmatrix} = 0$$

$$(\gamma_x^2 + \gamma_y^2)(\gamma_x B_y - \gamma_y B_x) = 0$$

$$\gamma_x = \pm i \gamma_y, \quad \frac{\gamma_x}{\gamma_y} = \frac{B_x}{B_y}$$

$$\frac{dx}{B_x} = \frac{dy}{B_y}$$

for the force-free magnetic field.

In formal mathematical terms note that the curl and divergence of $\nabla \times \mathbf{B} = \alpha \mathbf{B}$ yield

$$\nabla^2 \mathbf{B} + \alpha^2 \mathbf{B} + \nabla \alpha \times \mathbf{B} = 0,$$

and

$$\mathbf{B} \cdot \nabla \alpha = 0,$$

respectively. The first of these seems elliptic (imaginary characteristics), providing continuous solutions that are uniquely determined by \mathbf{B} at the boundaries. ^{But} The second equation, for the torsion coefficient α , has shows real characteristics, which are just the field lines of \mathbf{B} . This permits TD's along the field lines, thereby avoiding the uniqueness of the continuous solutions.

OPTICAL ANALOGY

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Flux surface

defined by the lines
of force across
curve C.

$$\oint ds \cdot B = 0$$

$$\text{Force-Free Field}$$

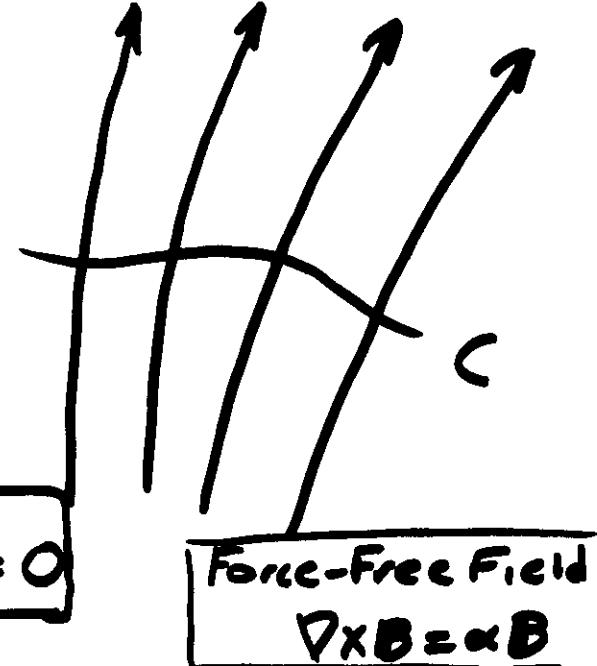
$$\nabla \times B = \alpha B$$

Note that $B = -\nabla \phi$ in any flux

surface. Lines of force described by

$$B h_1 \frac{dx_1}{ds} = -\frac{1}{h_1} \frac{\partial \phi}{\partial x_1}, \quad B h_2 \frac{dx_2}{ds} = -\frac{1}{h_2} \frac{\partial \phi}{\partial x_2}.$$

Same form as equations for optical
ray path, with eikonal ϕ , in index
of refraction B



Apply Fermat's principle

$$\delta \int_{P_1}^{P_2} ds B = 0$$

Use Euler equations to compute lines of force across a region with specified applied pressure

Regions of enhance pressure tend to exclude lines of force.

A local pressure maximum may create a gap in a free surface.

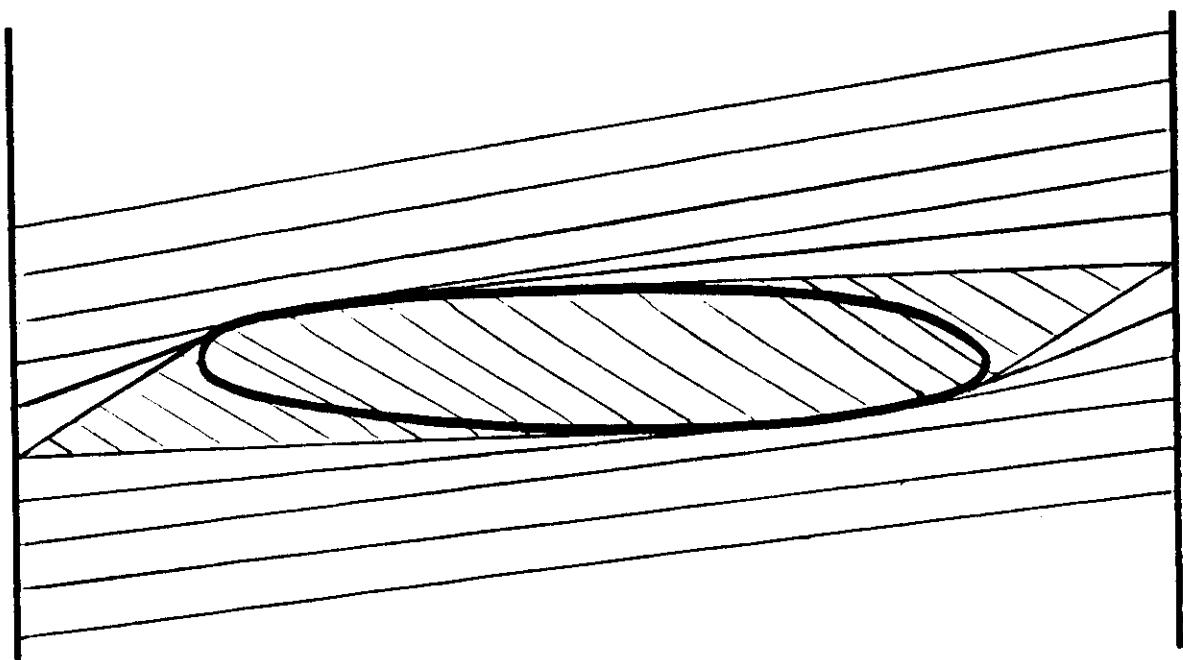
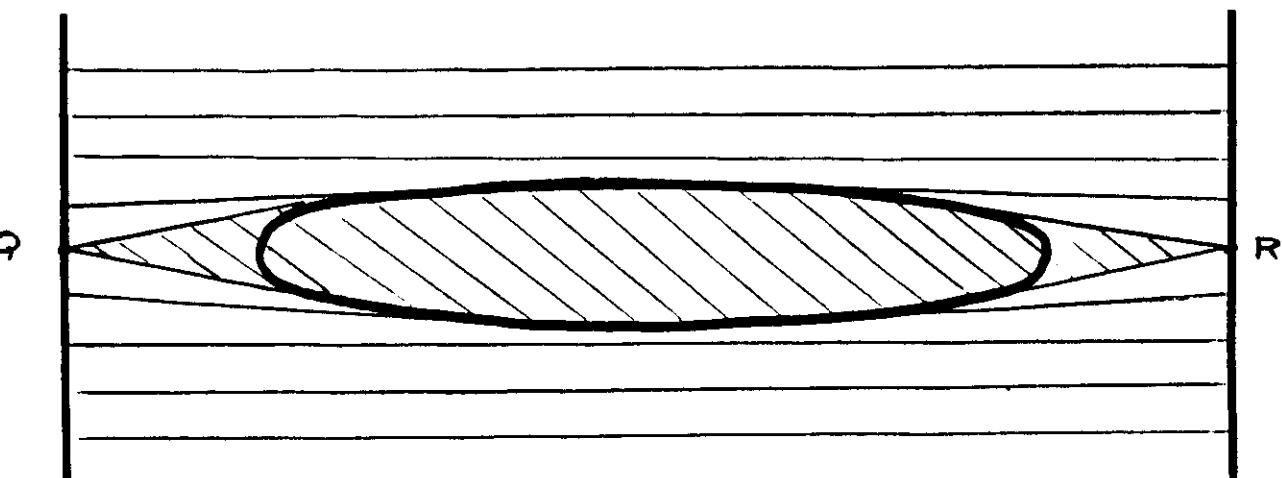


Fig. 24

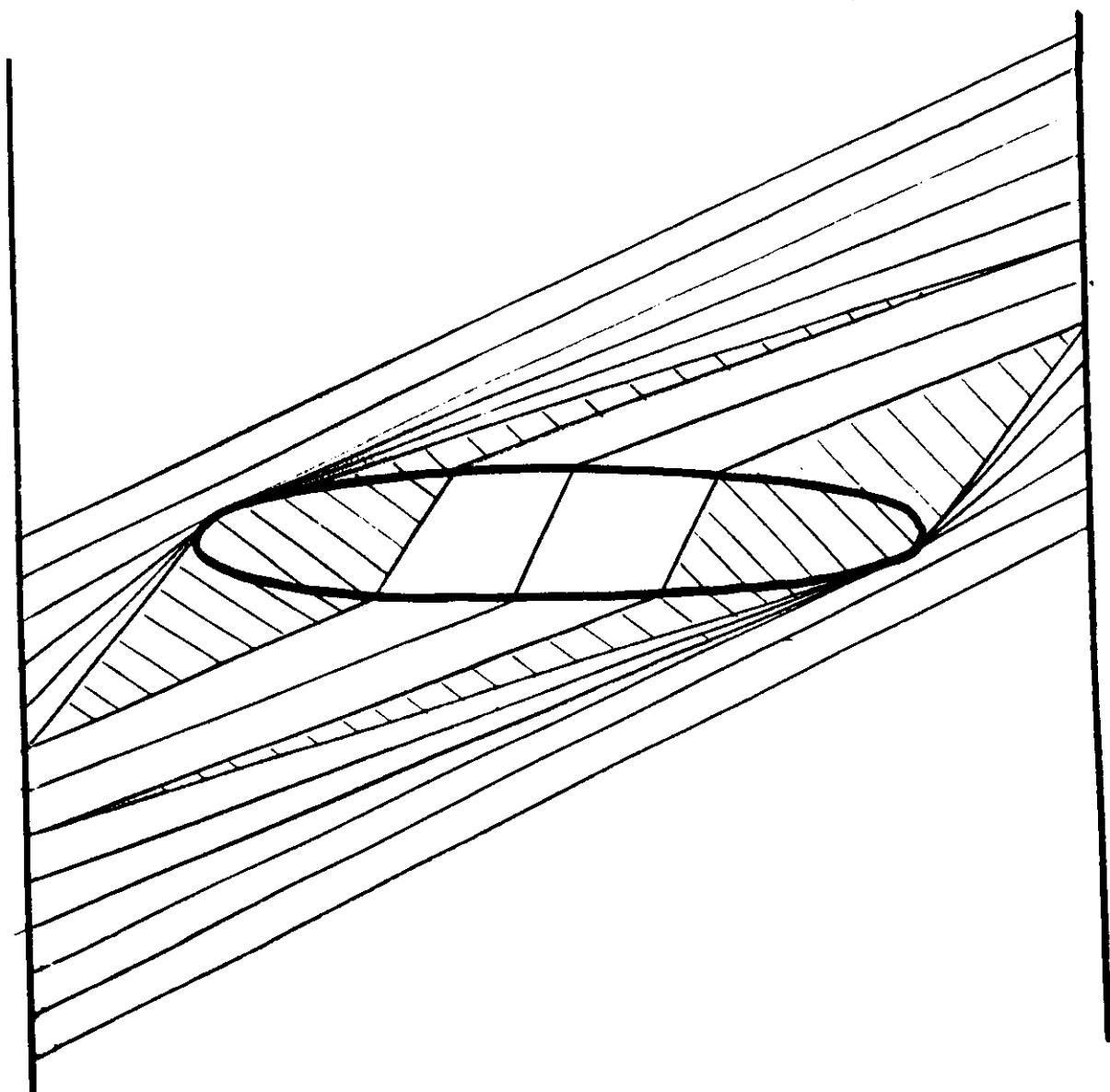
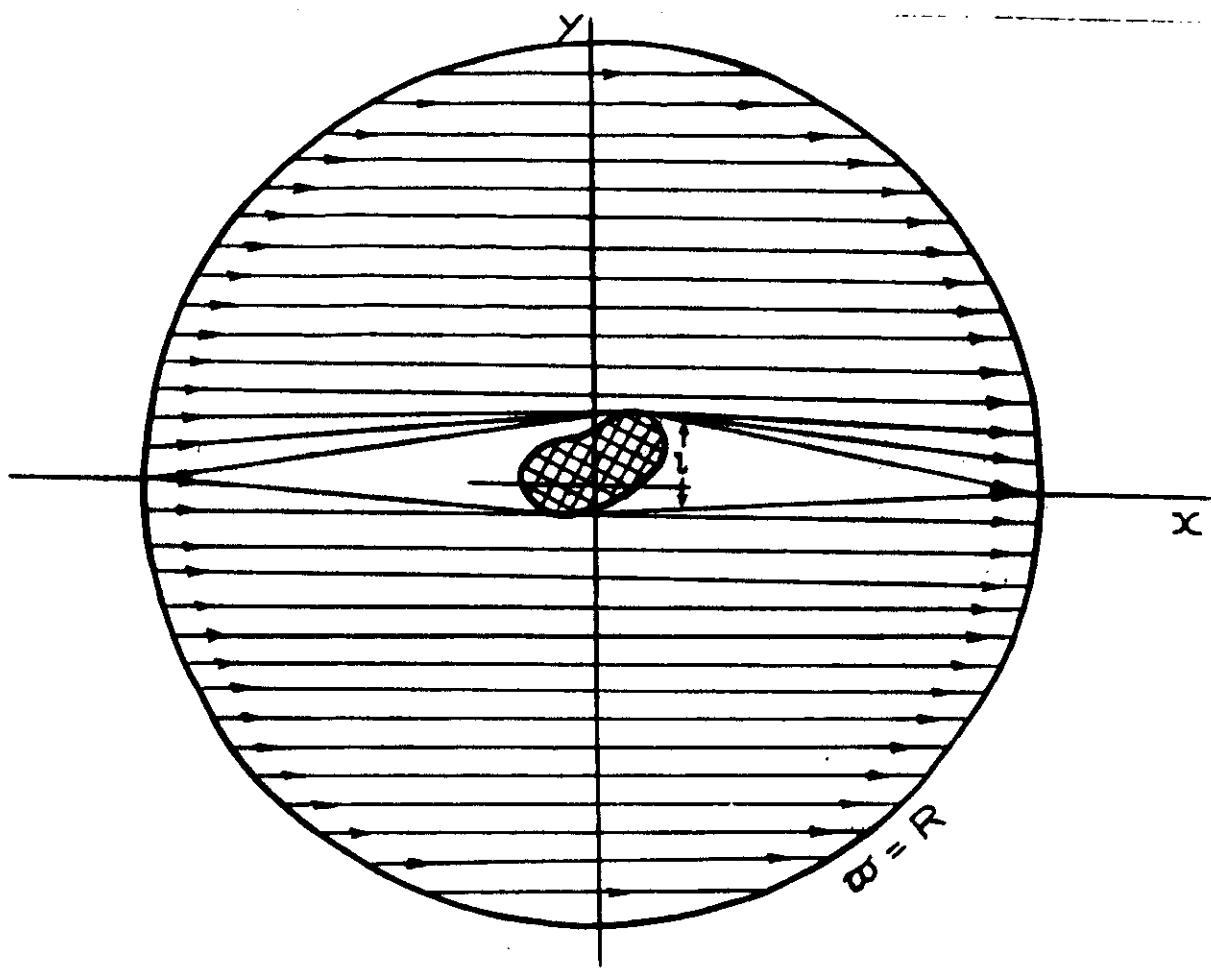
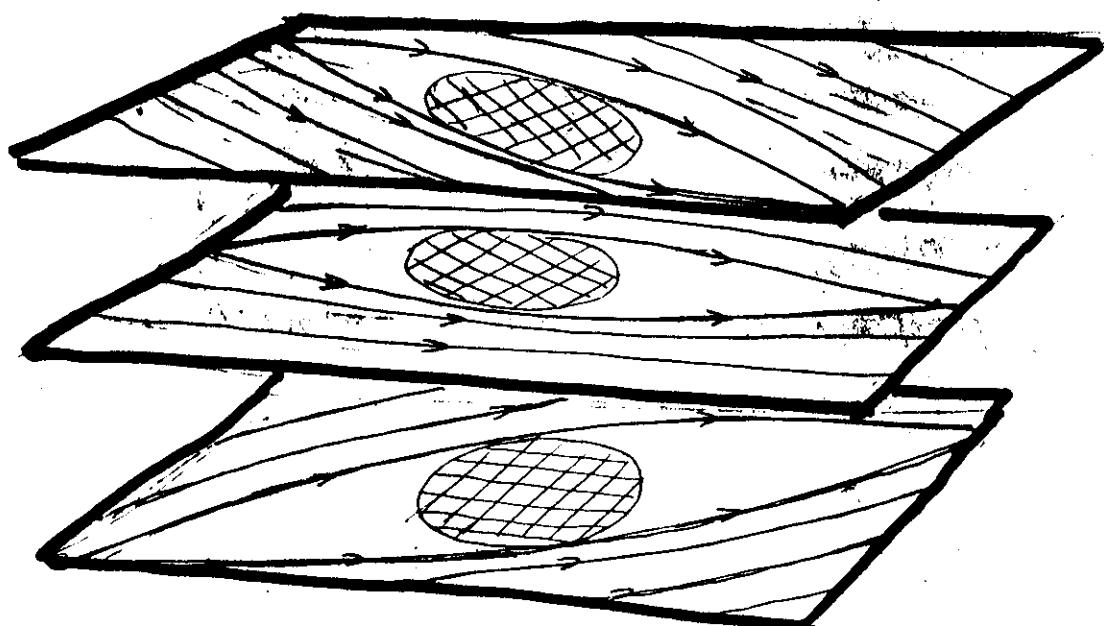


Fig. 2 c



The essential point is that the variations in the magnetic pressure in an interwoven force-free field produce gaps in the local flux-surfaces.

It follows that field lines are reflected around local regions of enhance magnetic pressure $B^2/8\pi$, creating a gap in the flux surface. The fields on either



side of S are generally not parallel where they meet each other in the gap.

See SPONTANEOUS CURRENT SHEETS IN MAGNETIC FIELDS, Oxford 1999
for illustrative examples.

A gap extending through a finite stack of flux surfaces allows otherwise separate fields to come into contact.

The separate fields are not precisely parallel where they meet, in almost all cases. The result is a tangential discontinuity, where they meet in the gap.

B^2 is continuous across the TD.

The field direction is discontinuous

$$J = (cB/2\pi) \sin \frac{1}{2}\theta.$$

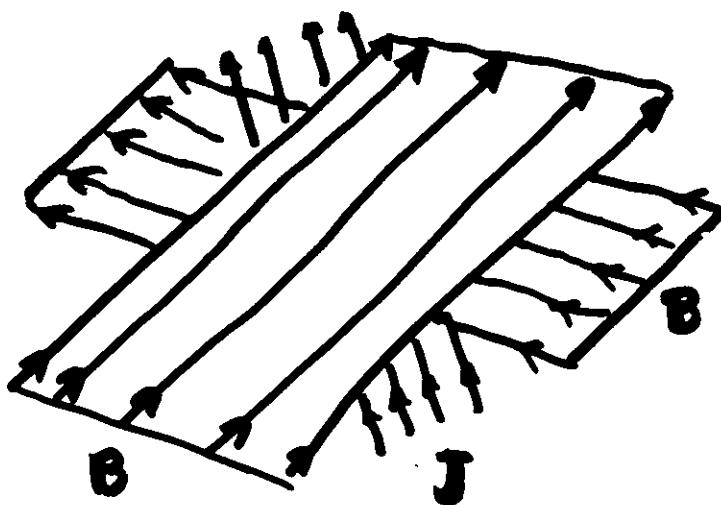
If B_+ and B_- represent the fields on opposite sides of a surface of TD, the surface current density is

$$J = \frac{c}{4\pi} (B_+ + B_-) \tan \theta , \quad B_+^2 = B_-^2 ,$$

where 2θ is the angle between B_+ and B_- . With $B = |B_{\pm}|$,

$$J = \frac{c}{2\pi} B \sin \theta .$$

Note that in static equilibrium, then, J is parallel to neither B_+ nor B_- .



Each surface of TD represents the contact between regions of continuous B where $\mathbf{B} \cdot \nabla \alpha = 0$.

If we think of the TD as a "delta fn" in \mathbf{v} , the TD does not violate $\mathbf{B} \cdot \nabla \alpha = 0$ because there is no magnetic flux in the TD.

A TD may form after only a limited interweaving of B . A couple of careless twists and a TD develops.

What happens when a small resistivity, η , is present? The TD is a current sheet, which immediately broadens.

$$j = j_m \exp\left(-\frac{\theta^2}{4\eta t}\right)$$

However, the topology of the field is such that a TD is essential for static equilibrium. So the Maxwell stress begins to drive the field and fluid back toward a TD, which it cannot achieve.

The result is a thin layer of dynamical nonequilibrium and resistive dissipation, which continues until the diffusive reconnection of the field reduces the topology to such a level that a TD is not required for equilibrium.

This dynamical state is the familiar "rapid reconnection" or "neutral point reconnection". It arises spontaneously in any magnetic field subject to convective interweaving.

Bursts of rapid reconnection limit the degree of interweaving of the field in the bipolar magnetic active regions on the Sun.

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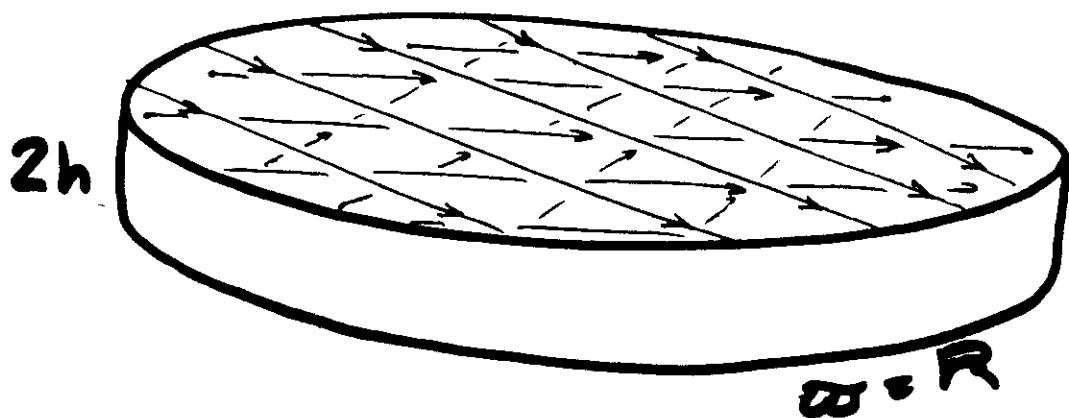
Discontinuities can be demonstrated in many ways.

IV As a formal example of the tangential discontinuity, consider the linear laminar force-free field, $\alpha = q = \text{constant}$

$$B_x = B_0 \cos q z, B_y = -B_0 \sin q z, B_z = 0,$$

$$B_x = B_0 \cos(qz + \psi), B_y = -B_0 \sin(qz + \psi).$$

confined to the thin layer $-h < z < h$ ($hq = O(1)$) and fixed in the circular infinitely conducting rigid boundary $\omega = R \gg h$.



Squeeze the layer of field gently, smoothly in a neighbourhood l of the origin.

The result is the creation of a tangential discontinuity in the field. $l = O(h)$

Work in the limit of large R .

The squeezed field becomes

$$B_\infty + b_\infty, B_\varphi + b_\varphi, b_z$$

and the torsion coefficient $\alpha_{\varphi\varphi}$ is written
 $\alpha = q(1+A).$

The force-free field equation reduces to

$$\frac{1}{\omega} \frac{\partial b_z}{\partial \varphi} - \frac{\partial b_y}{\partial z} = qAB_0, \quad (1)$$

$$\frac{\partial b_z}{\partial \omega} = q(1+A)b_y + qAB_0, \quad (2)$$

$$\frac{1}{\omega} \frac{\partial}{\partial \omega} \omega b_y = q(1+A)b_z. \quad (3)$$

If $b_y \sim \omega^{-s}$ for $\omega \gg l$, then eqn. (3) states that $b_z \sim \omega^{-s-1}$. With $\partial/\partial z = O(q)$, $\partial/\partial \varphi = O(1)$, $ql = O(1)$, eqns. (1) and (2) reduce to

$$\frac{1}{q} \frac{\partial b_y/B_0}{\partial z} + A \cos v = O\left(\frac{b_y}{B_0} \frac{1}{q^2 R}\right)$$

$$\frac{b_y}{B_0} (1+A) - A \sin v = O\left(\frac{b_y}{B_0} \frac{1}{q^2 R}\right)$$

at $\omega \asymp R$, where $B_\omega = B_0 \cos(qz + \varphi)$
 $b_\omega = 0$, $v = qz + \varphi$

Neglecting the right hand sides, integration yields

$$A = -1 \pm \sin u / (\sin^2 u - c)^{\frac{1}{2}}$$

$$b_0/B_0 = \sin u \mp (\sin^2 u - c)^{\frac{1}{2}}$$

where $u = qz + \psi$. C is a constant.

If $c < 0$, the radicals have no zeros and so cannot change sign. So if A and b_0/B_0 are small at $u = \pi/2$, they are $O(1)$ at $-\pi/2$. With $c > 0$ both A and b_0/B_0 are small $O(c^{1/2})$ as required by the physics, except that A diverges as $\sin^2 u$ decreases to c .

Since b_0/B_0 is not large than the ω^2 of a 2-D dyadic, it follows that $C = O(1/q^2 R^4)$. The ~~torsion~~
 & A is then of the form of a delta fn,

$$A = 2C^{1/2} \sum_m \delta(u - m\pi) + O(1/q^2 R^4)$$

and

$$\alpha = q(I+A) = q \left[I + 2C^{1/2} \sum_m \delta(u - m\pi) \right].$$

There is a tangential discontinuity at $u=0, \pm\pi, \text{ etc.}$ The discontinuity in field direction is $O(1/qR)$

The same result is given by the formal optical analogy, where gaps appear in the local flux surfaces in response to pressure inhomogeneities.

Note that apart from the singular surfaces, $\alpha = q$, neglecting terms $O(1/q^2 r^2)$.

Hence

$$\nabla^2 B_i + q^2 B_i = 0$$

except at singularities.

Suppose there were no singularities. Then complete general solution can be constructed. Solution is unique, given $B_\infty = B_0 \cos(\varphi z + \psi)$ at $\omega = R$.

No solutions with field connecting straight across $\omega = 0$, except original.

\therefore Continuity is unphysical.