



UNITED NATIONS EDUCATIONAL, SCIENTIFIC AND CULTURAL ORGANIZATION
INTERNATIONAL ATOMIC ENERGY AGENCY
INTERNATIONAL CENTRE FOR THEORETICAL PHYSICS
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H4.SMR/1013-24

**SCHOOL ON THE USE OF SYNCHROTRON RADIATION
IN SCIENCE AND TECHNOLOGY:
*"John Fuggle Memorial"***

3 November - 5 December 1997

Miramare - Trieste, Italy

***The application of multilayer coatings in
synchrotron radiation research***

**Werner Jark
Area di Ricerca
Sincrotrone Trieste, Italy**

**The application of multilayer coatings
in synchrotron radiation research**

Werner Jark



**S.S. 14, km 163.5 in Area Science Park
34012 Basovizza (TS), Italy**

I will try to give answers to the following questions:

- a) Why do we need them?
- b) How do we make them?
- c) Can we simulate their performance?
- d) How do we test them?
- e) Do experiment and simulation really agree?
- f) Where do we finally use them?

We all know that using visible light

(red = $\lambda = 800 \text{ nm} = 1.5 \text{ eV}$;

violet = $\lambda = 400 \text{ nm} = 3 \text{ eV}$) *

- we can take an image with a camera by use of lenses or mirror optics!

- we can make an enlarged image in a microscope once more using either lenses or less often mirror systems!

- we can make simple polarizers or phase retarders by use of birefringent material!

What happens if we want to use these devices in the soft x-ray range (10 eV - 2000 eV)?

With increasing photon energy the light starts

- a) to be absorbed in lenses
- b) to not be dispersed anymore in lenses
- c) to find little anisotropy in lenses/filters
- d) to not be reflected anymore in normal incidence

Why this:

let's take some examples:

$n =$ index of refraction

a) focal length of lens: $(1/f) = (n-1) ((1/r_1) + (1/r_2))$ $r_1, r_2 =$ radii

b) normal incidence reflectivity: $R = ((n-1)/(n+1))^2$

The index of refraction varies as follows for glass (SiO₂):

photon energy:	3 eV	30 eV	100 eV	1000 eV
n =	1.5	0.9	0.985	0.9987
a) f =	e.g. 0.2 m	- 1 m	- 6.6 m	-80 m
b) R =	0.04	0.0028	6 10 ⁻⁵	4 10 ⁻⁷
R = (Au)	0.37	0.08	0.0014	1 10 ⁻⁶

So what can we do?

a) **Not much as far as classical lenses are concerned!**

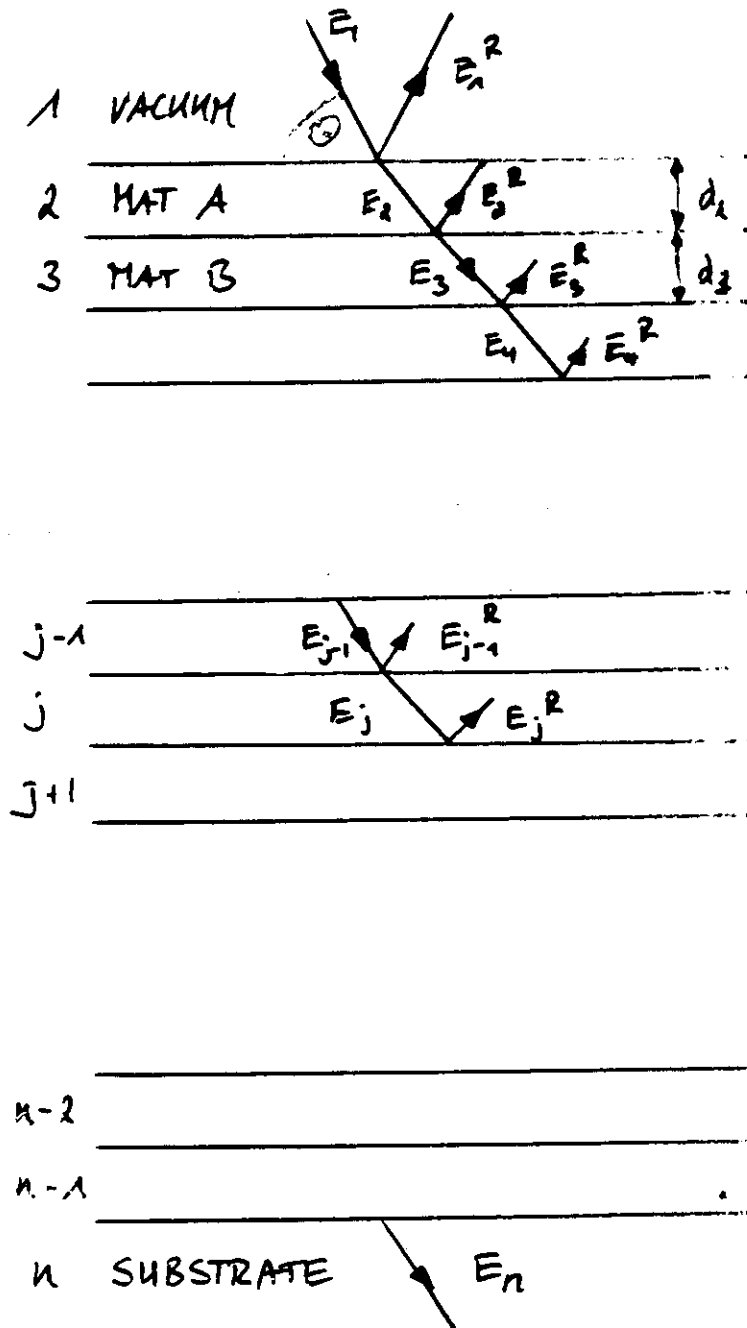
b) **However, for m interfaces in a sufficiently transparent structure we can get $R_{total} = m R_{int}$, which becomes interesting for heavier materials with better R_{int} than glass.**

But is it really so simple?

Not really, the different waves need to collaborate in phase, which is affected by the index of refraction n of a layer and which can undergo additional changes at any reflexion at and any transmission through an interface.

So let us write a program!

Multilayer simulation:



Reflectivity and transmittance can be calculated for this stack in principle for up to $(n-2)$ different materials with $(n-2)$ different thicknesses

most common are stacks of just 2 different materials with a total of 10 to >200 periods of equal spacing.

Solution of Parratt (Phys. Rev. 95, 359 (1954))

$E_{\text{tang}} = \text{const at interfaces}$

$g_j = \text{amplitude factor}$
 $= \exp(-i\pi g_j d_j / \lambda)$

$$a_j E_j + a_j^{-1} E_j^R = a_{j+1} E_{j+1} + a_{j+1}^{-1} E_{j+1}^R$$

$d = \text{thickness}$

$\lambda = \text{wavelength}$

$$g_j (a_j E_j - a_j^{-1} E_j^R) = g_{j+1} (a_{j+1} E_{j+1} - a_{j+1}^{-1} E_{j+1}^R)$$

$$g_j = \tilde{n}_j \cdot \sin \Theta_j$$

solution is a recursion equation

$$R_{j,j+1} = a_j^4 \left[\frac{R_{j+1,j+2} + F_{j,j+1}}{R_{j+1,j+2} \cdot F_{j,j+1} + 1} \right]$$

$$R_{j,j+1} = a_j^2 \cdot \frac{E_{j+1}^R}{E_j}$$

$$F_{j,j+1} = \frac{E_j^R}{E_j} = \frac{g_j + g_{j+1}}{g_j - g_{j+1}} \quad \text{for s-pol } (\delta)$$

$$= \frac{g_j / \tilde{n}_j^2 - g_{j+1} / \tilde{n}_{j+1}^2}{g_j / \tilde{n}_j^2 + g_{j+1} / \tilde{n}_{j+1}^2} \quad \text{for p-pol } (\pi)$$

finally $R = \frac{I}{I_0} = |R_{12}|^2$

more elegant to write:

($\tilde{E}_j = \tilde{n}_j^2$)

with wavevector in a medium: $\tilde{k}_j = \frac{2\pi}{\lambda} \cdot \sqrt{\tilde{E}_j - \epsilon_0 \cdot \cos^2 \Theta}$

$$R_{q1,m} = \frac{R_{q1,m} + R_{1,2,m} \cdot \exp(2i d \tilde{k}_1)}{1 + R_{0,1,m} \cdot R_{1,2,m} \cdot \exp(2i d \tilde{k}_1)}$$

$m = s, p$

$0 = \text{vacuum}$

$$R_{j,j+1,s} = \frac{\tilde{k}_j - \tilde{k}_{j+1}}{\tilde{k}_j + \tilde{k}_{j+1}}$$

$$R_{j,j+1,p} = \frac{\tilde{E}_{j+1} \cdot \tilde{k}_j - \tilde{E}_j \cdot \tilde{k}_{j+1}}{\tilde{E}_{j+1} \cdot \tilde{k}_j + \tilde{E}_j \cdot \tilde{k}_{j+1}}$$

finally $R = |R_{0,1,m}|^2$

RECURSION starts always at the last interface to the substrate \Rightarrow proceeds upwards.

A recursion equation can also easily be transferred to a matrix multiplication.

Matrix approach:

e.g. P. Lee | Opt. Commun. 43, 237 (1982)

attention: applicable only close to 45°

N = number of periods for layer pairs

d = period thickness $\frac{d}{2}$ components: $\gamma d, (1-\gamma)d$

n_i = refractive index ($n = 1 - \delta + i\beta$)

Θ = angle of grazing incidence

average unit decrement of refractive index:

$$\bar{\delta} = \gamma \cdot \delta_1 + (1-\gamma) \cdot \delta_2 \quad \bar{\beta} = \gamma \beta_1 + (1-\gamma) \beta_2$$

phase shift ψ per period

$$\psi = \frac{2\pi \cdot d}{\lambda} \cdot \sin \Theta \cdot \left(1 - \frac{\bar{\delta} + i\bar{\beta}}{\sin^2 \Theta}\right) \quad \text{instead per layer: } \psi_1, \psi_2$$

reflection coefficient for large angles:

$$r = \frac{(\Delta\delta + i\Delta\beta) \cdot P(\Theta)}{2 \sin^2 \Theta} \quad P(\Theta) = \begin{cases} 1 & \text{s-pol} \\ \cos 2\Theta & \text{p-pol} \end{cases}$$

$$\Delta\delta = \delta_1 - \delta_2 \quad \Delta\beta = \beta_1 - \beta_2$$

E_{tang} in one period with respect to the others:

$$[A] = (1-r^2)^{-1} \cdot \begin{pmatrix} e^{i\psi} (1-r^2 e^{-2i\psi_1}) & 2ir e^{i\psi_2} \cdot \sin \psi_1 \\ -2ir e^{i\psi_2} \cdot \sin \psi_1 & e^{-i\psi} (1-r^2 e^{2i\psi_1}) \end{pmatrix}$$

after extension to N periods the reflection coefficient is:

$$R(\Theta) = \frac{-2ir e^{i\psi_2} \cdot \sin \psi_1 \cdot S_{N-1}(x)}{e^{i\psi} (1-r^2 e^{-2i\psi_1}) \cdot S_{N-1}(x) - \sqrt{\Delta} \cdot S_{N-2}(x)}$$

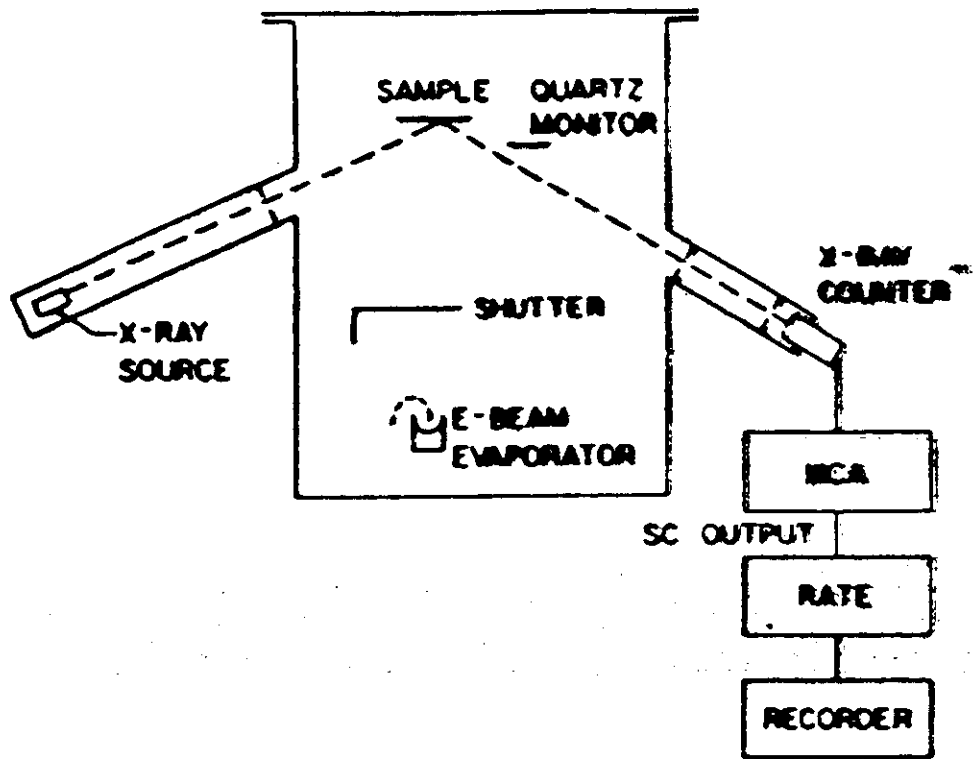
$$S_N(x) = \sin[(N+1) \cos^{-1}(x)] / \sin(\cos^{-1}(x))$$

Chebyshev polynomial of 2. kind

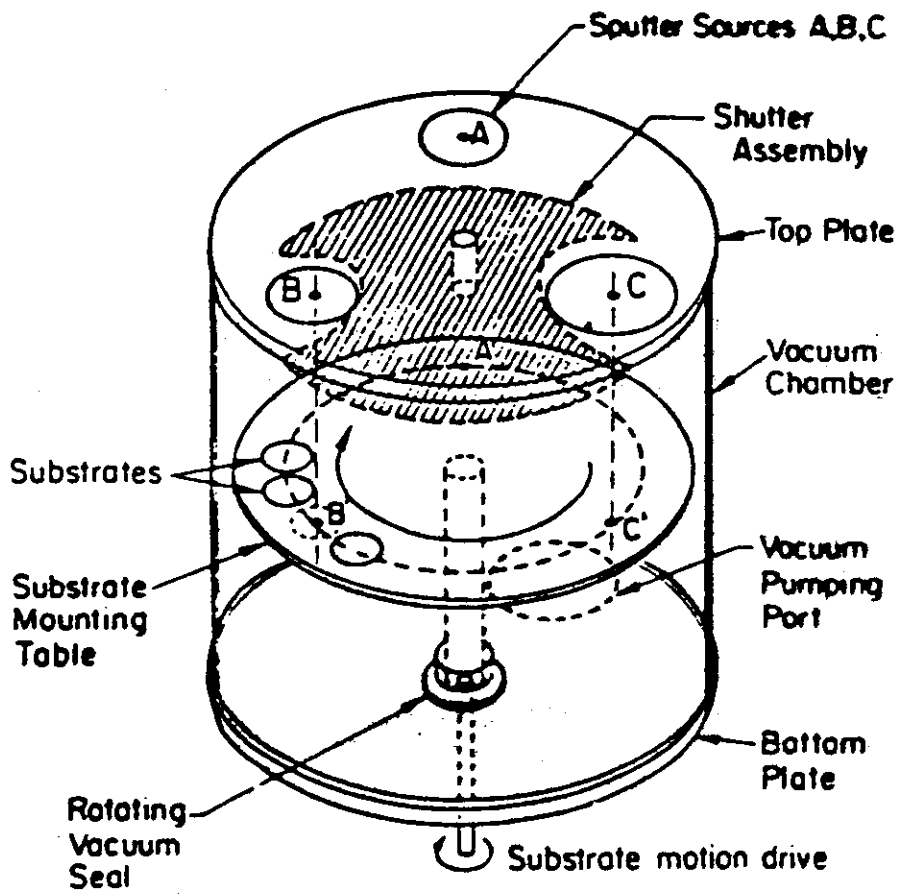
$$\Delta = (1-r^2)^2 \quad x = \frac{\cos \psi - r^2 \cos(\psi_1 - \psi_2)}{\sqrt{\Delta}}$$

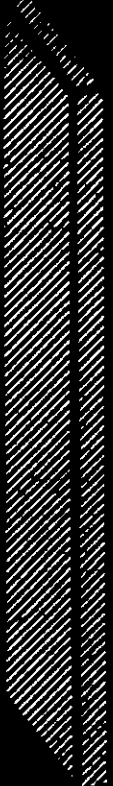
equation is not plausible anymore:

BUT, it is an analytical equation, which works and can be programmed in about $\frac{1}{2}$ hour.



Multilayer production





Evaporation:

Is it really this simple? Are there no problems?

NO!!

Evaporation:

- repeatability
- cluster evaporation
- sublimation
- monitoring of thickness
- homogeneity over large sample sizes

Sputtering:

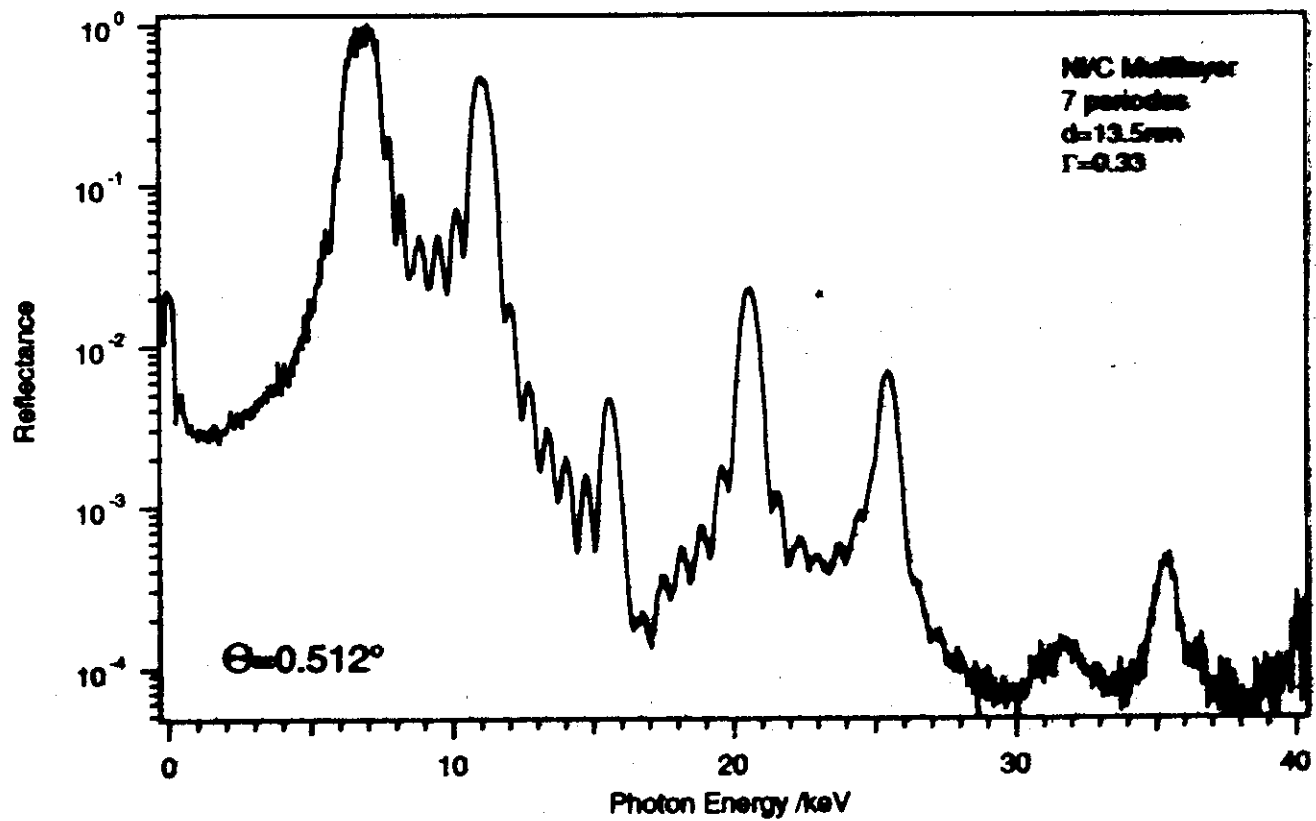
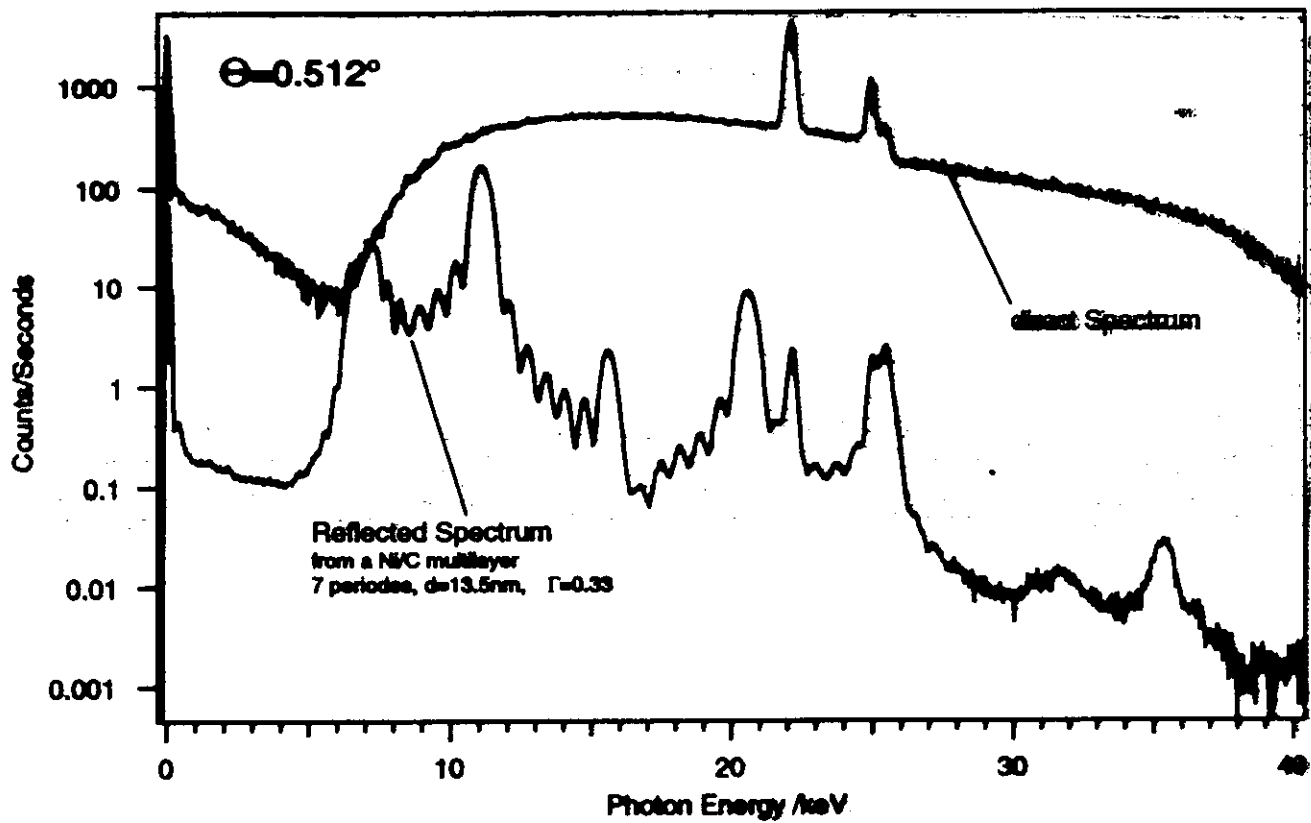
- homogeneity over large sample sizes
- simultaneous bombardment of sample with plasma electrons
- plasma gas inclusions ($p \approx 1 - 100 \text{ mbar}$)
- reactions with restgas of vacuum

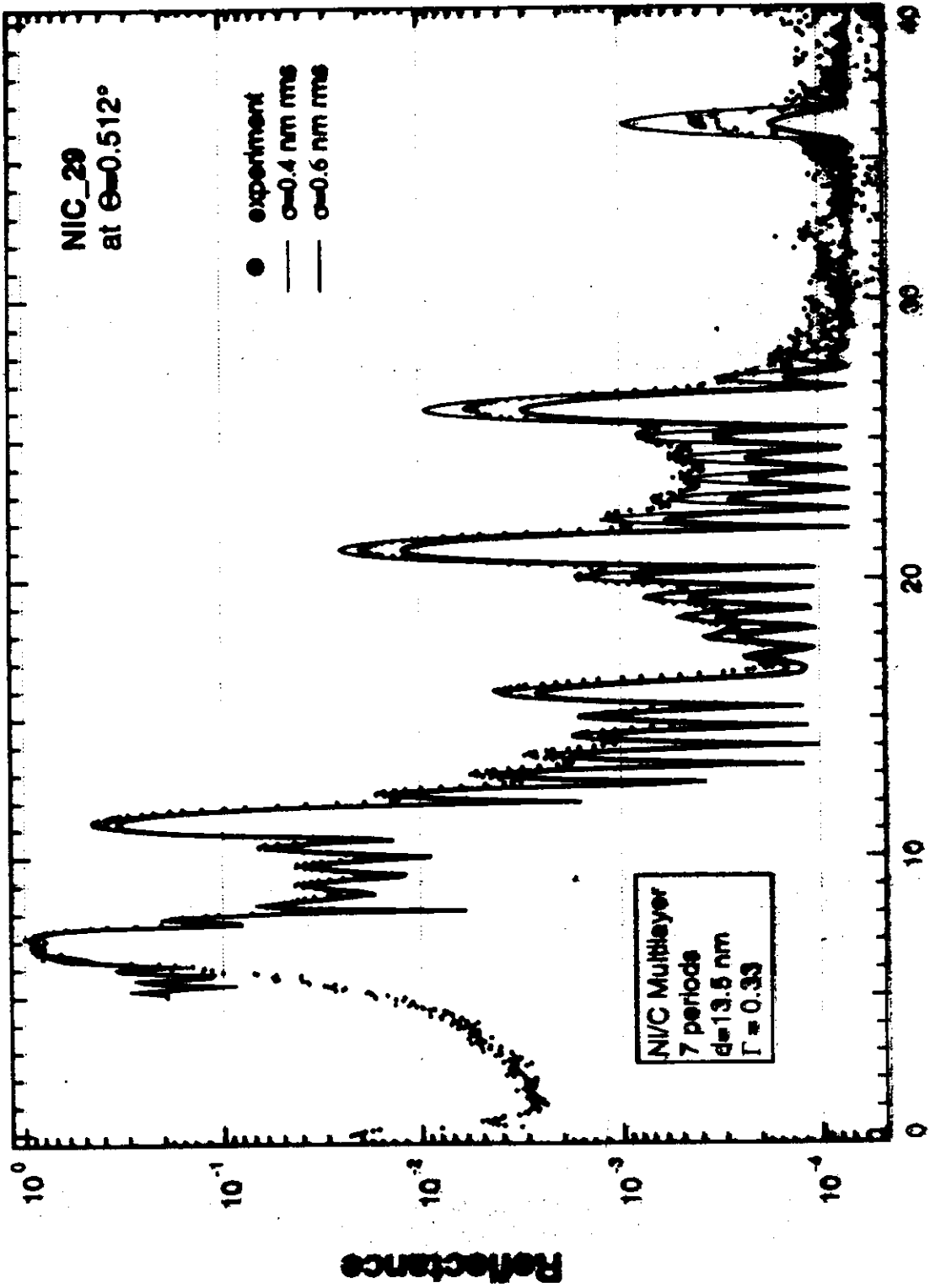
but: easy monitoring and very repeatable due to plasma stability and reproducibility

⇒ monitoring by timing
(or charge transfer)



WHITE & RAY

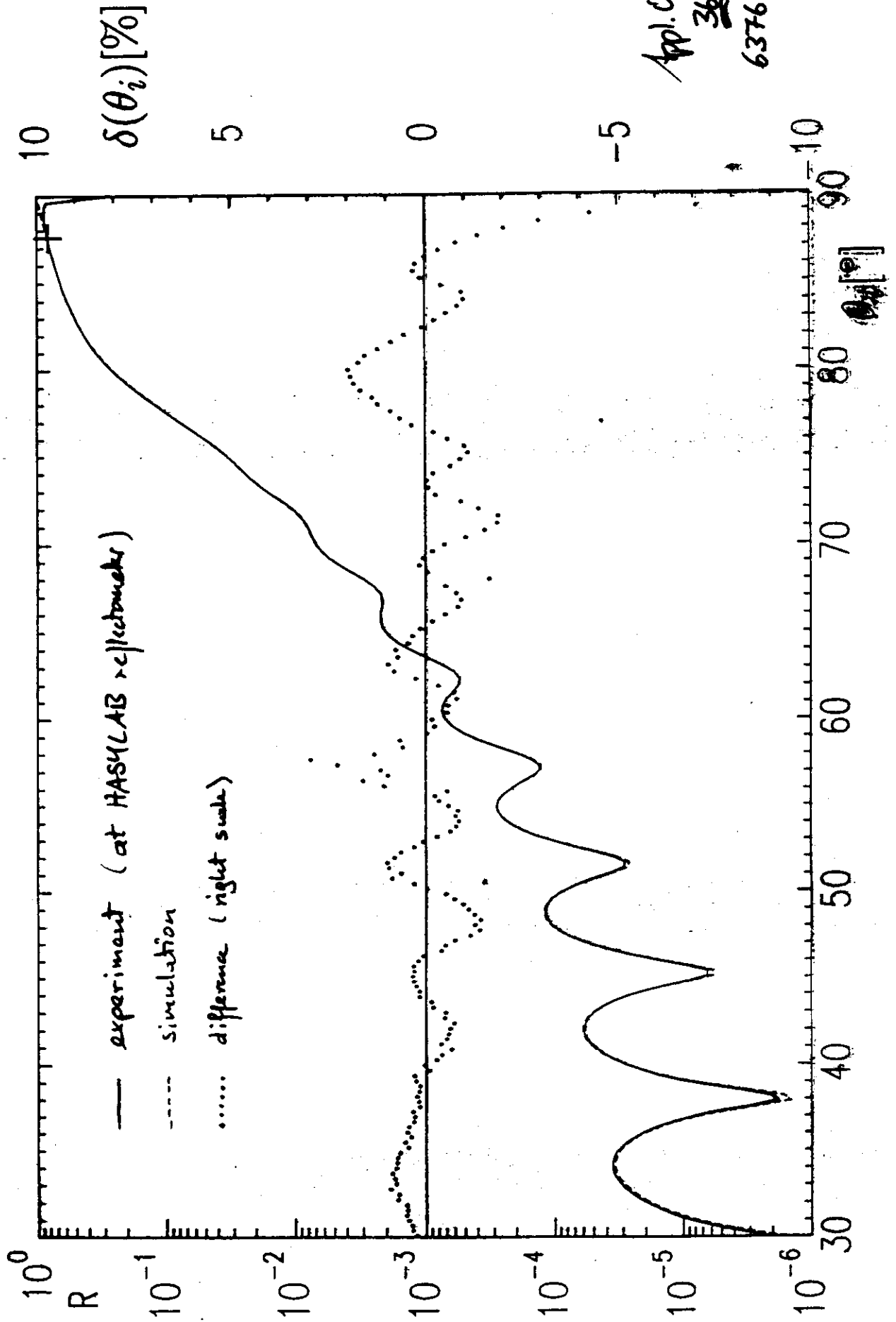




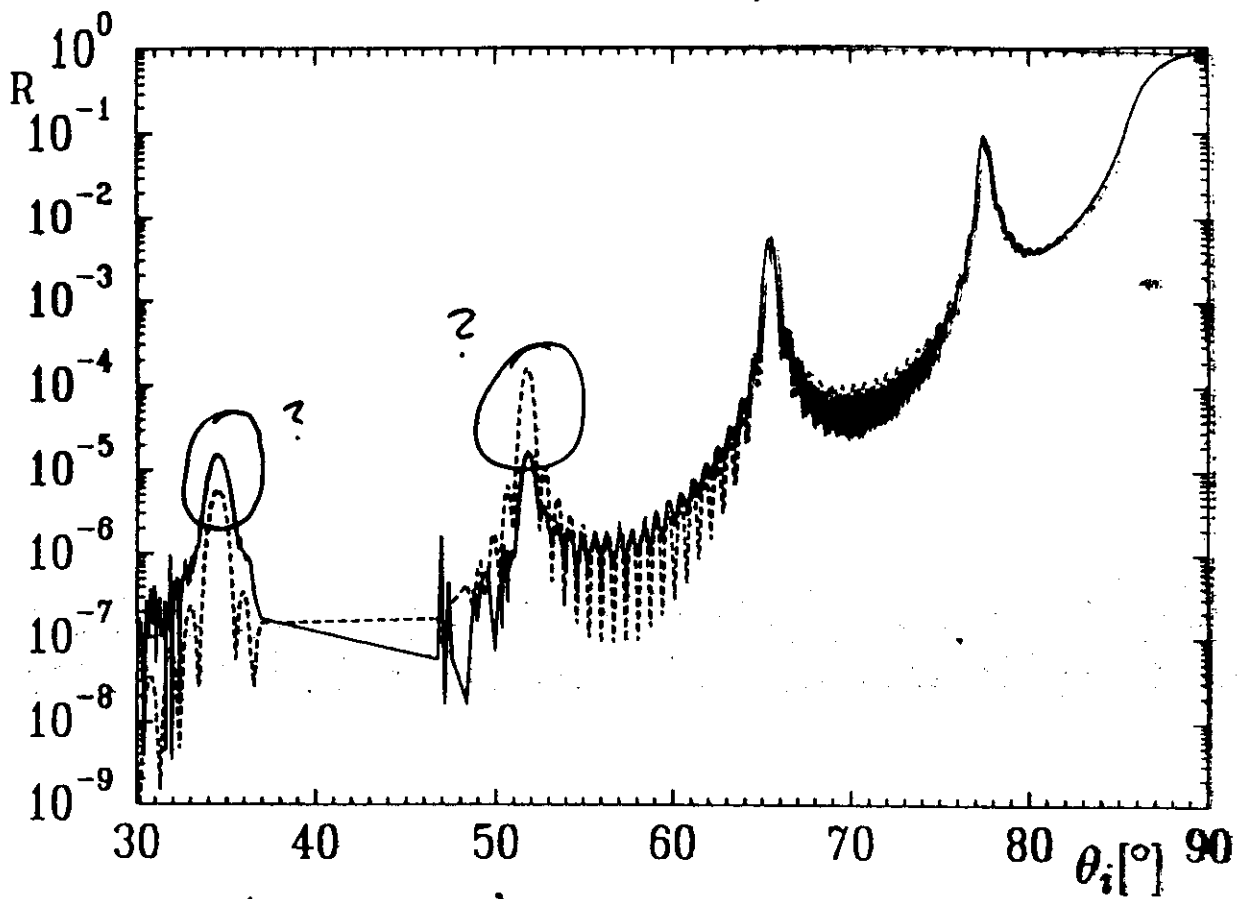
Inclusion of roughness : $\tilde{r}_m \rightarrow F_m \cdot \exp(-\frac{2\tilde{r}_m}{\lambda} \cdot \delta \sin^2 \theta)$

Real multilayers: a simple Ni-film on float glass

$d = 30.4 \text{ nm}$ $E = 260 \text{ eV}$

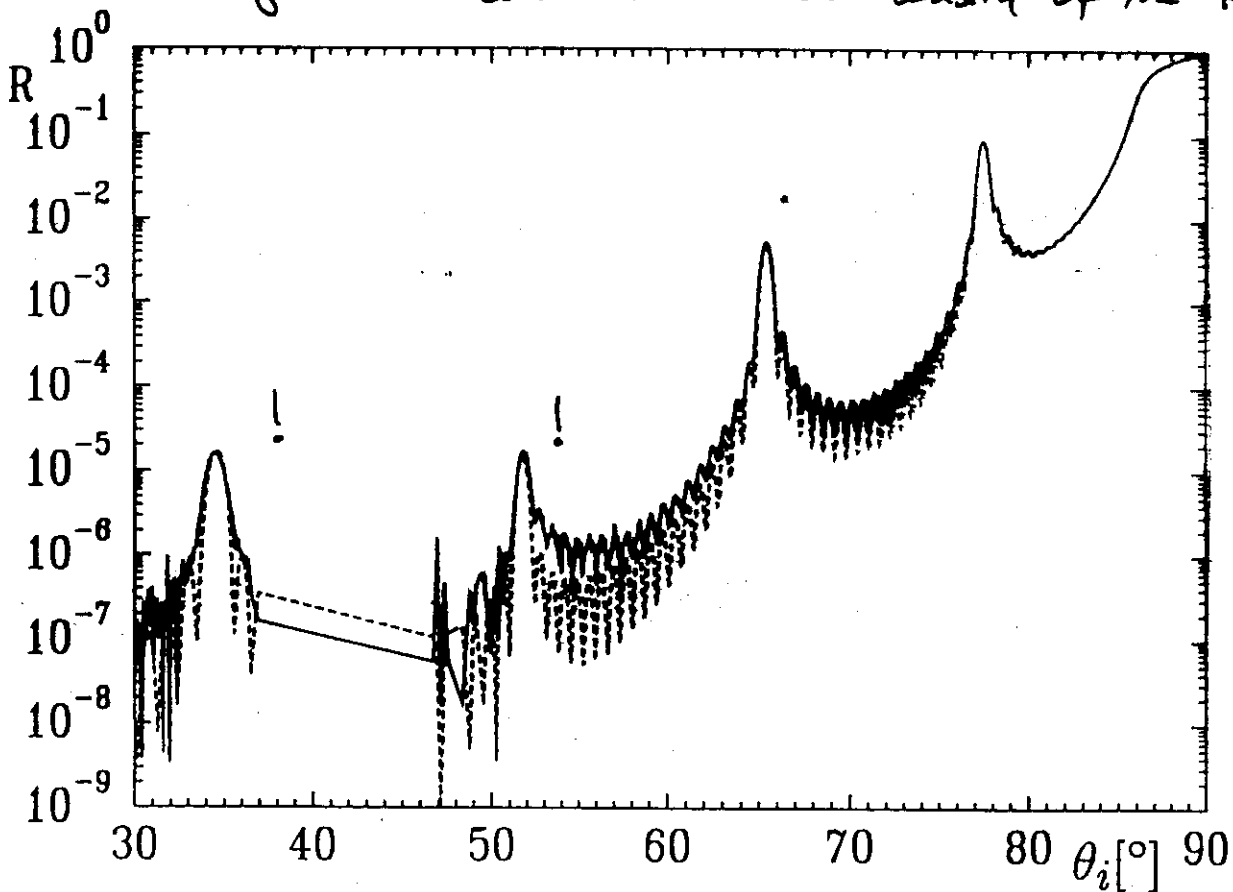


App. Opt.
38,
6376 (1988)

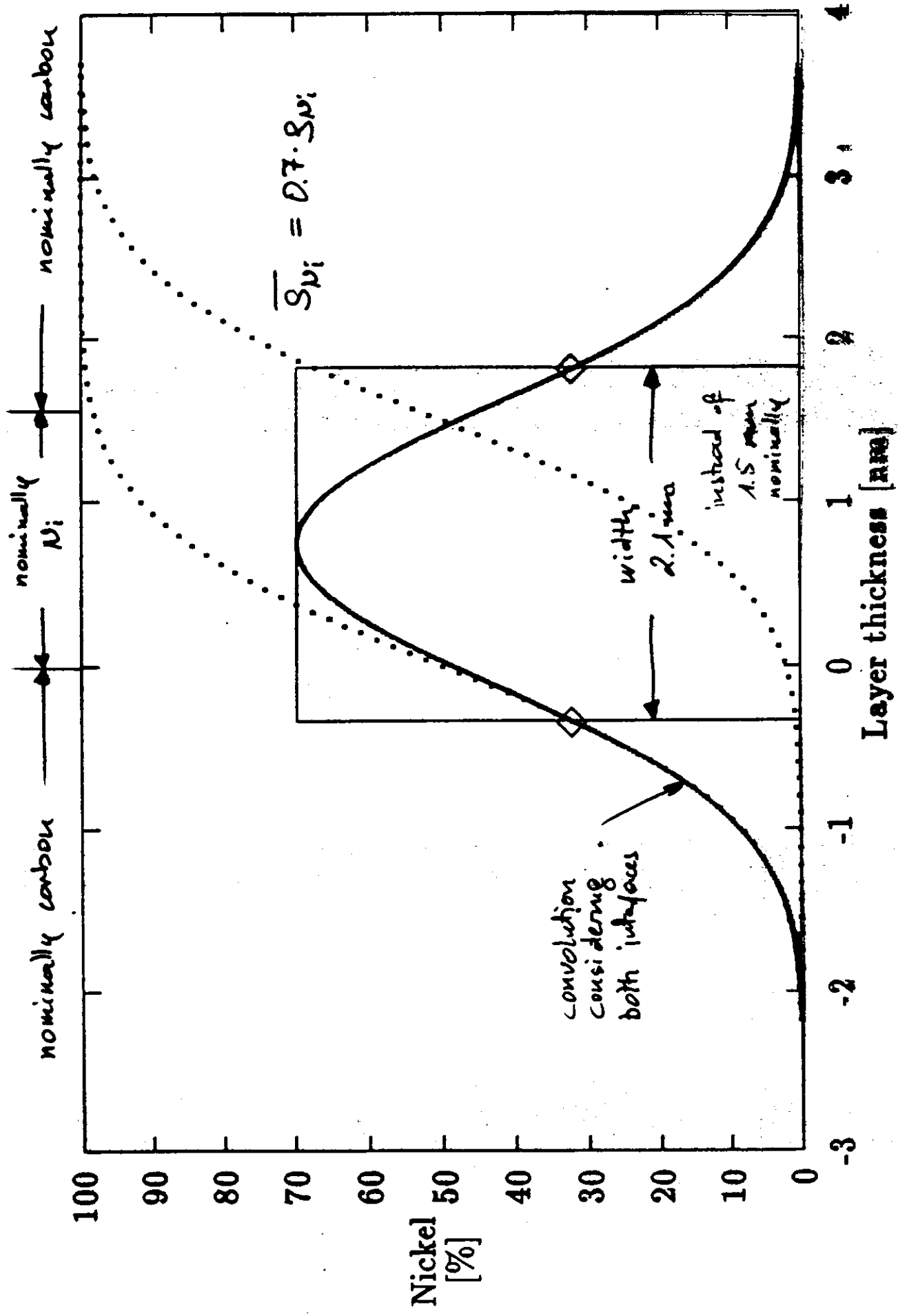


we do have $d_{Ni} = 1.5 \text{ nm}$
 with interface roughness of 0.7 nm (RMS)

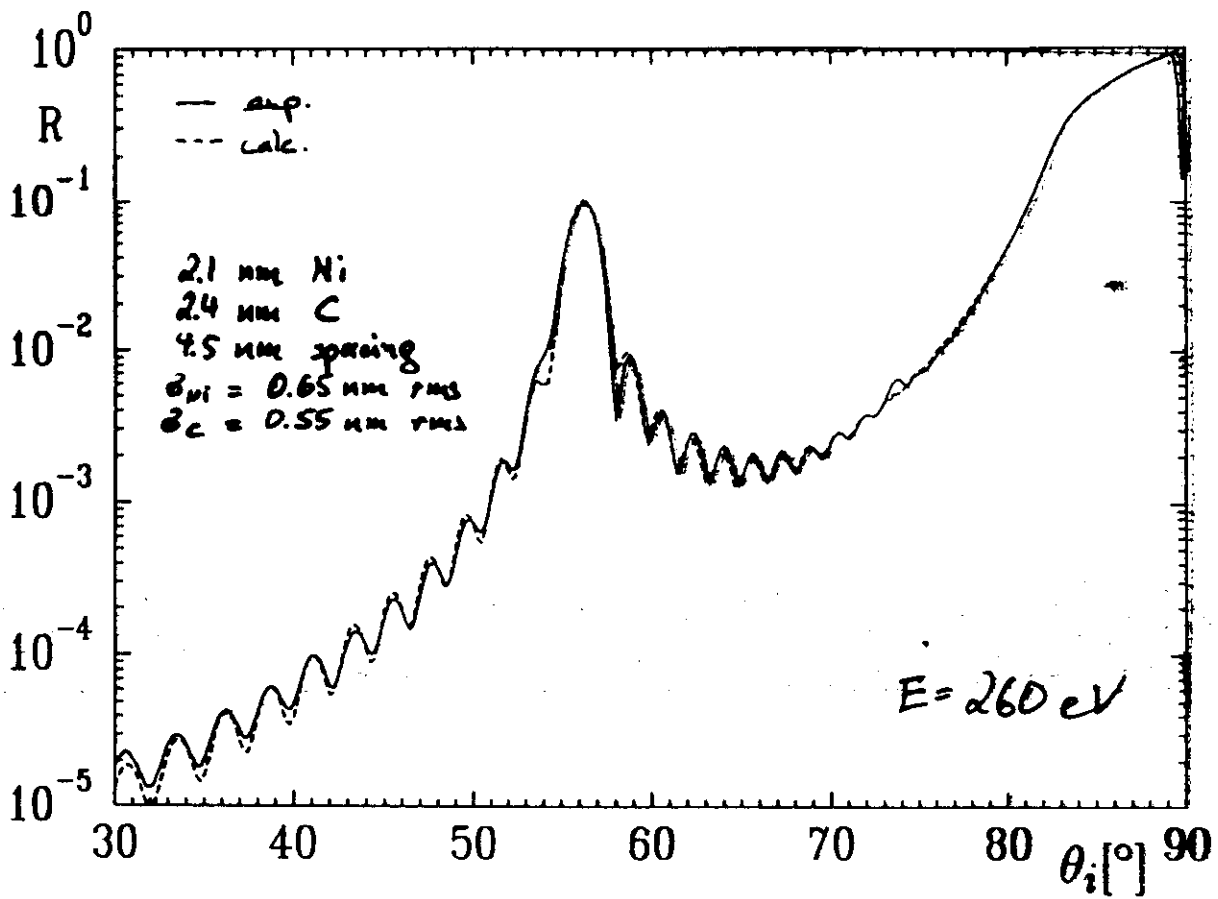
taking into account the reduced density of the Ni-layer



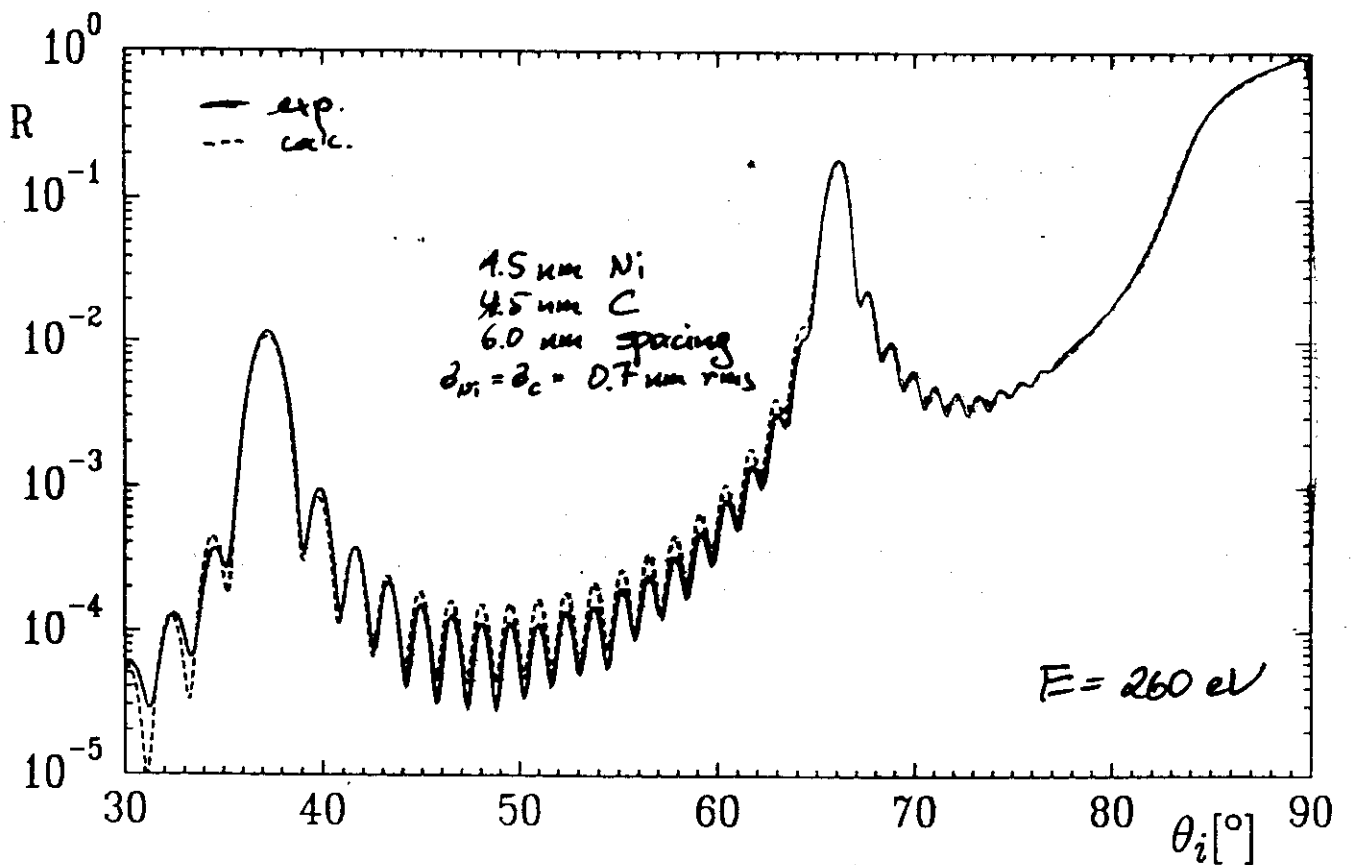
App. Opt. 36, 6328 (1987)



App. Oct. 36, 6329 (198)



Real multilayer : $N = 20$ periods Ni/C



Appl. Optics, 36, 6328 (1997)

TEMPERATURE
Optimization

Components of reflexion coefficient:

for
280 eV
photon
energy

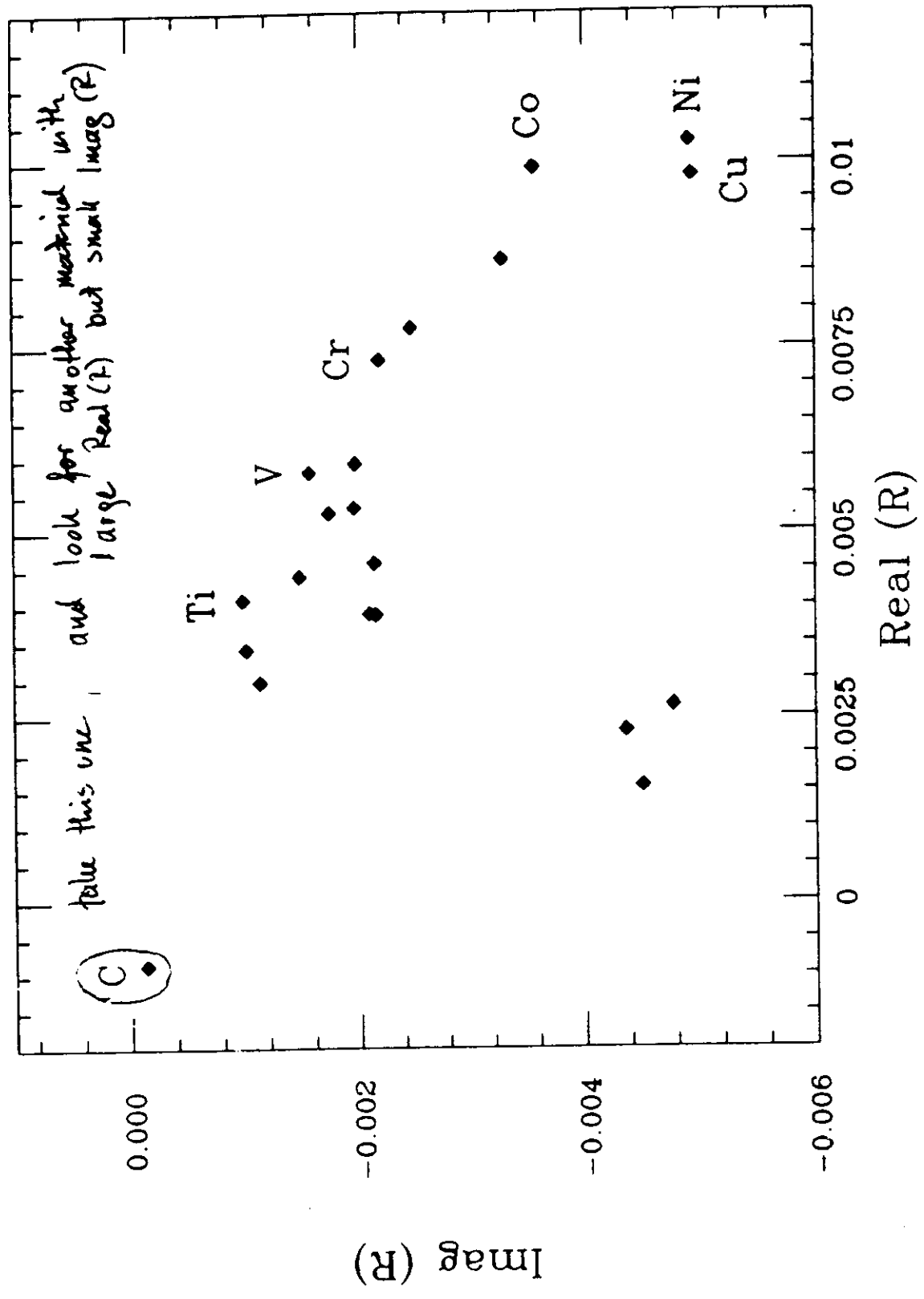


PHOTO BEAM

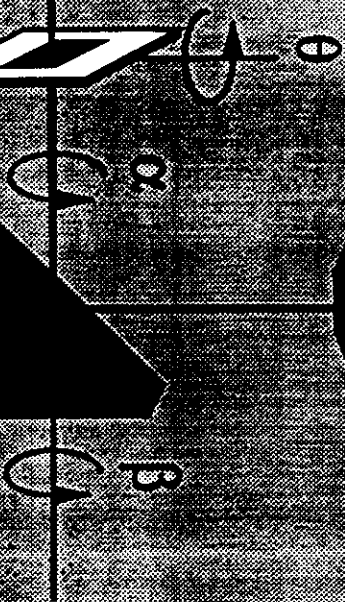
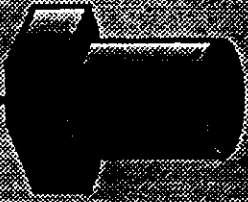
COLLIM (20)

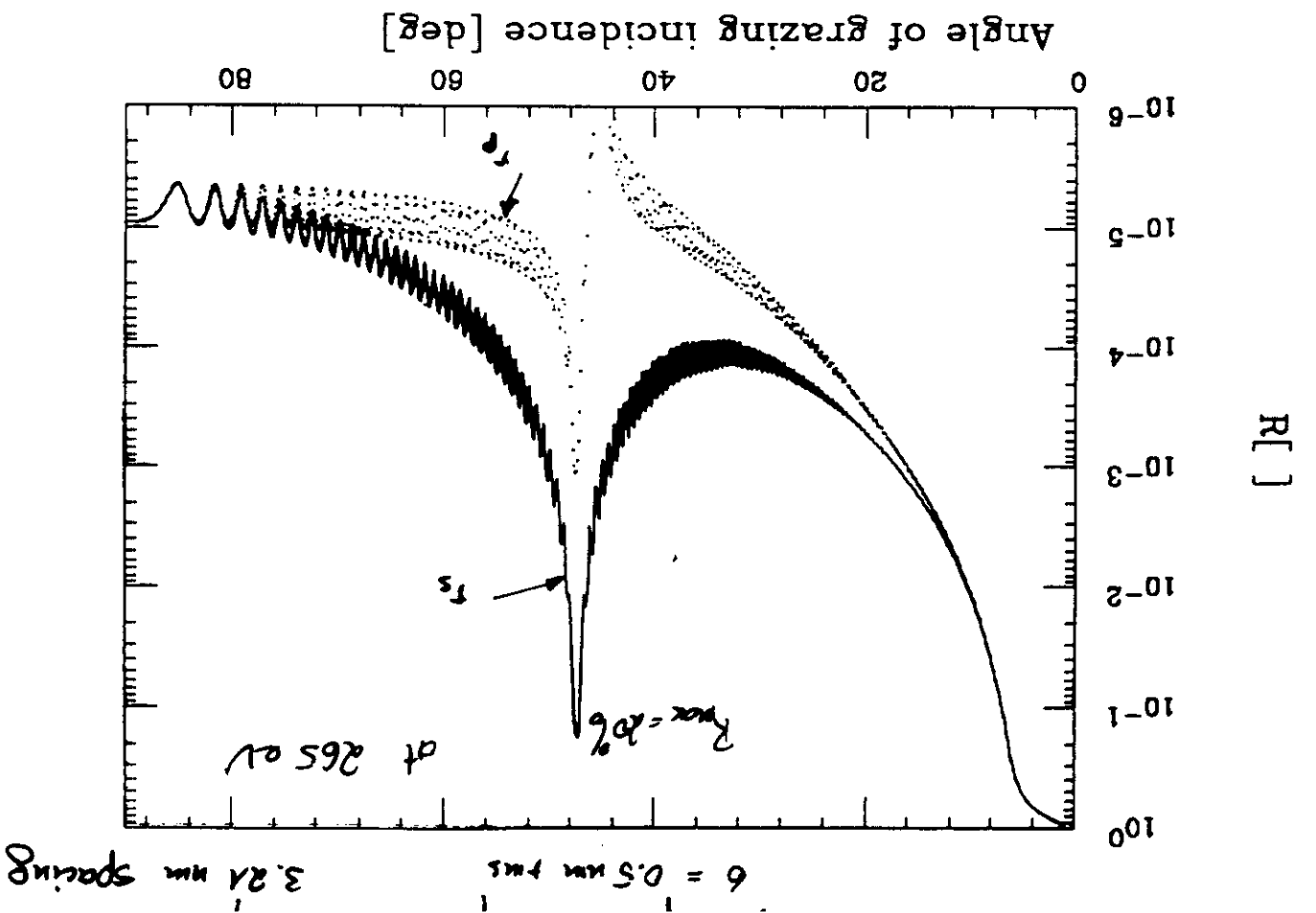
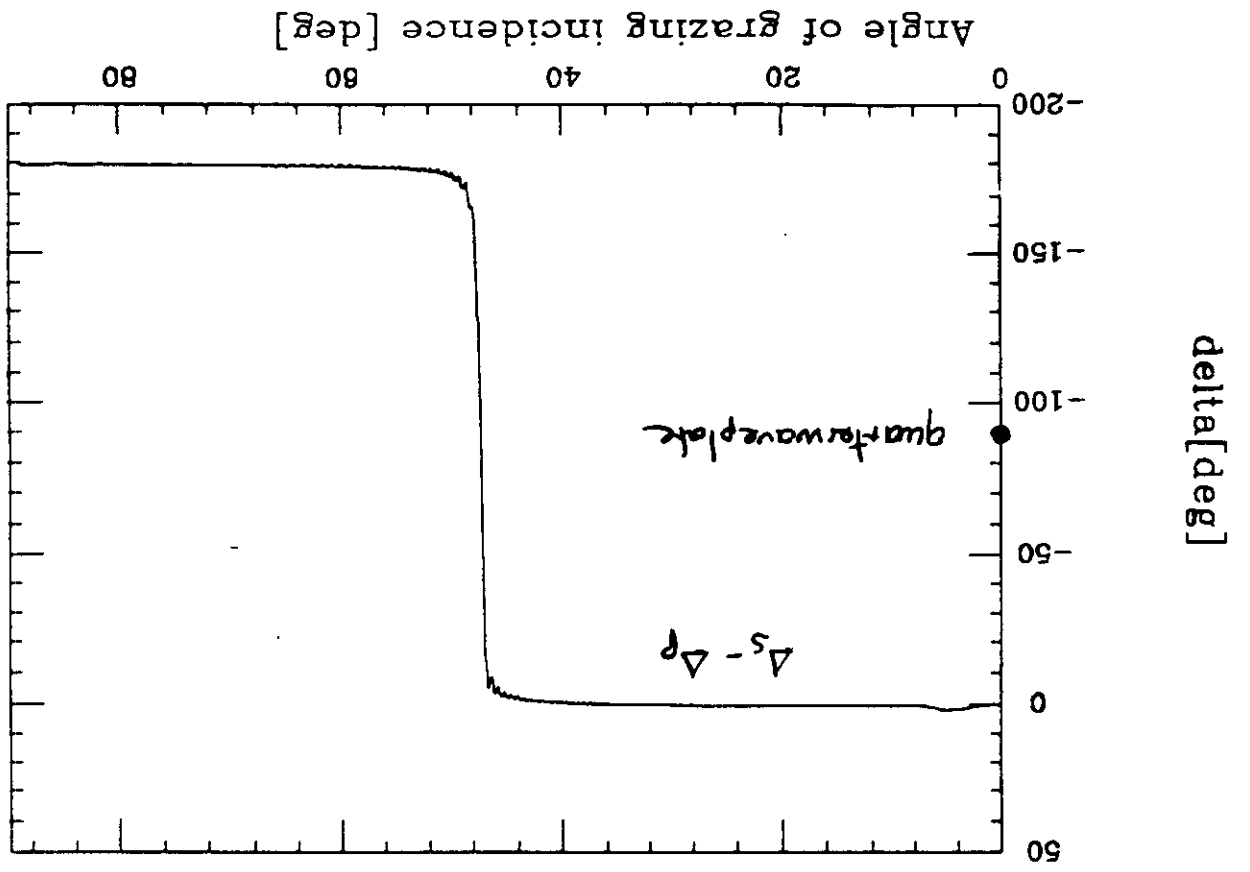
POLARIZA

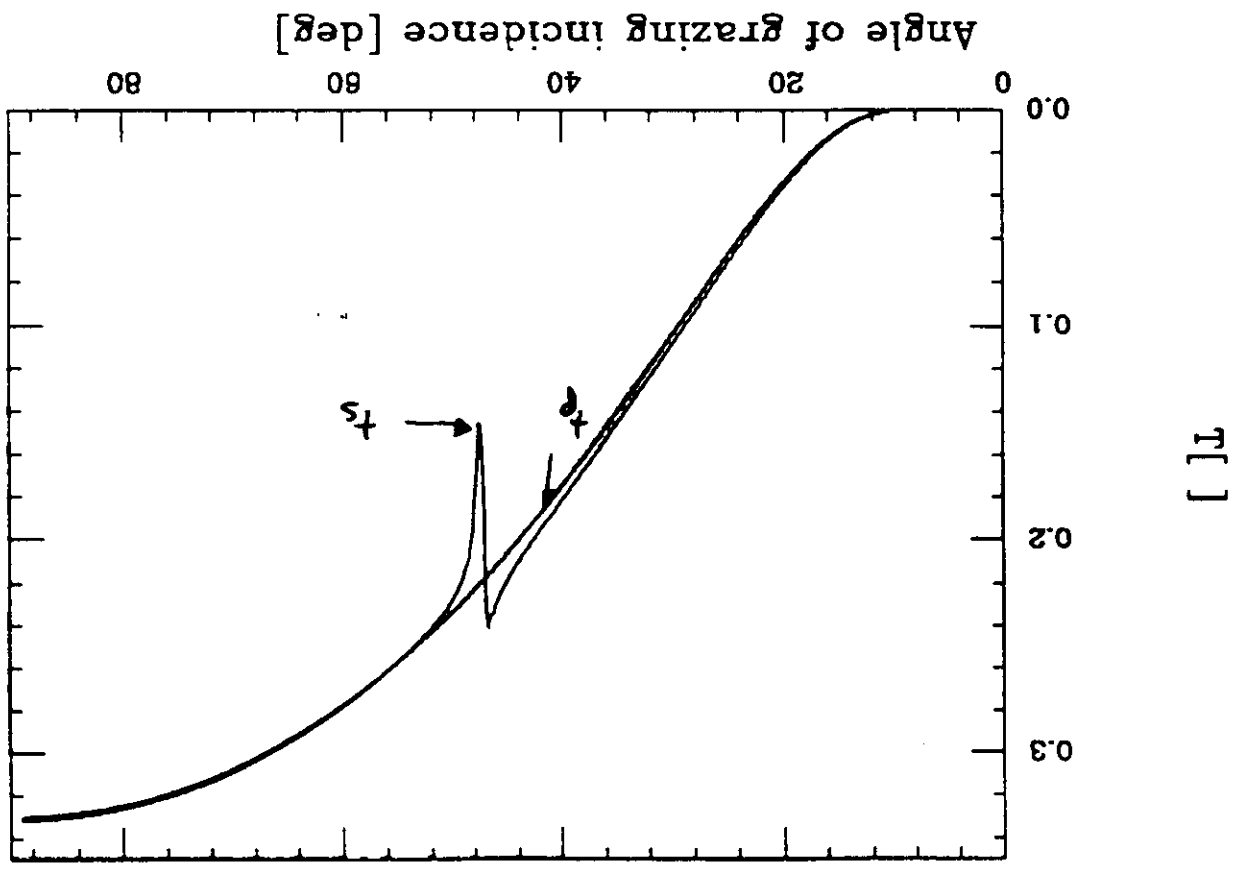
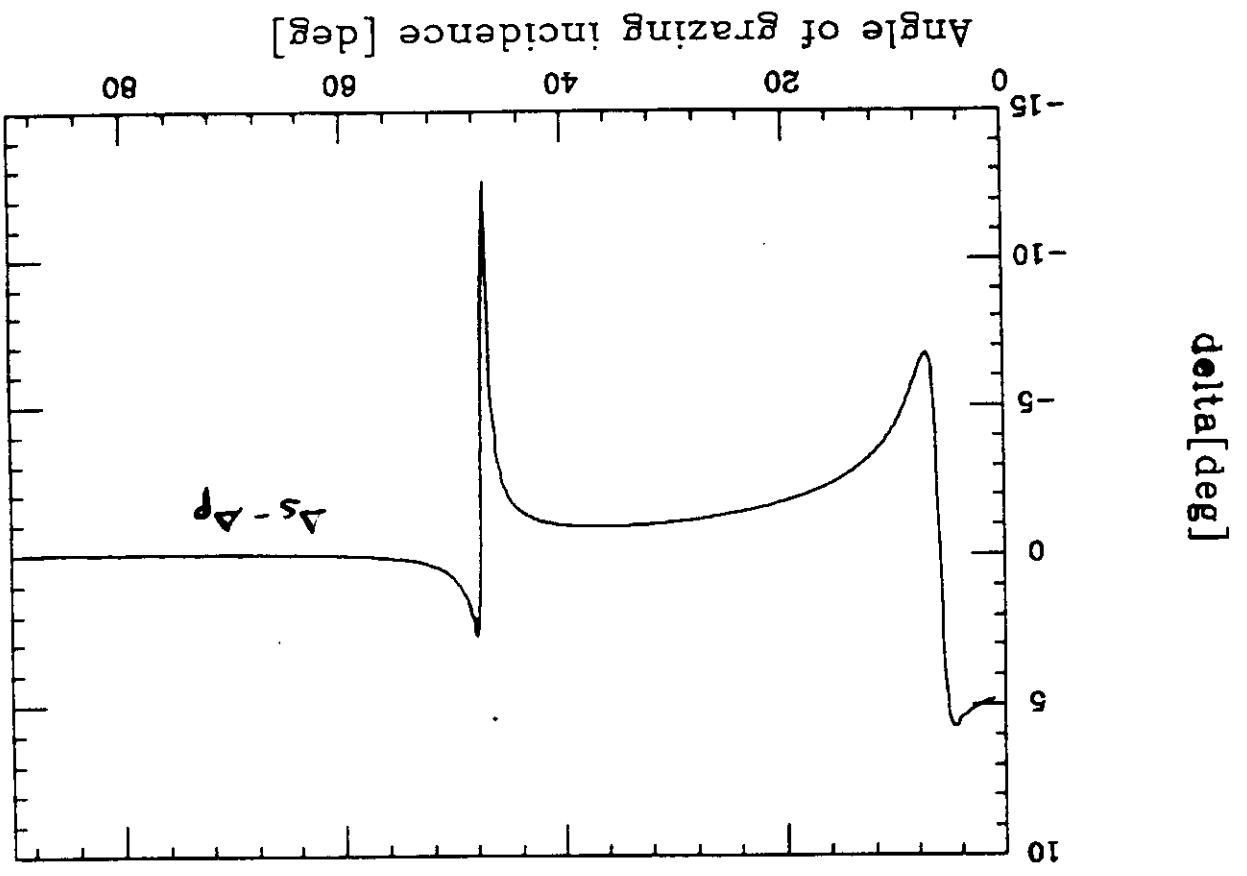
ANALYZA

DIREC(S)

REDA



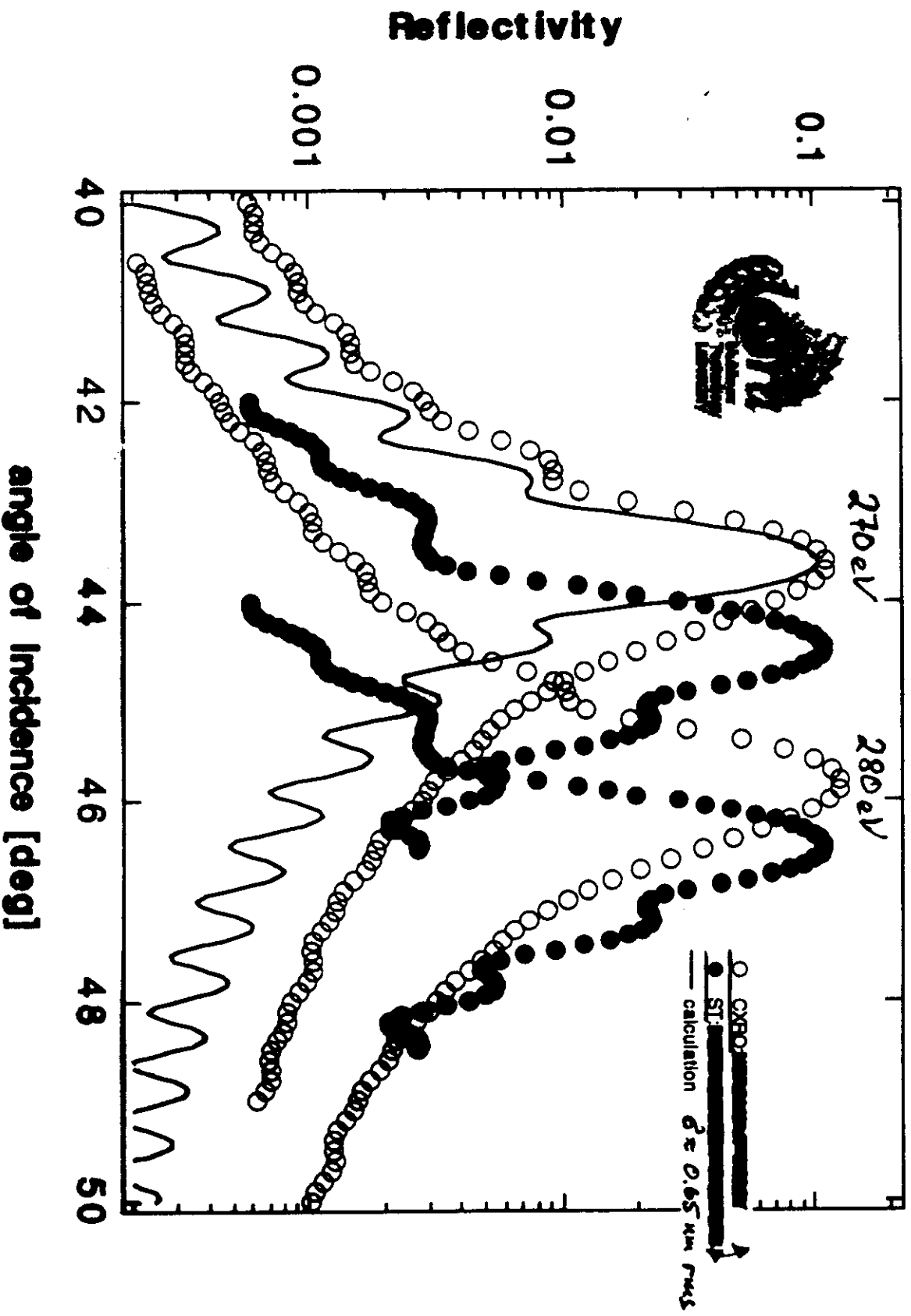




Continuation of the same figure.

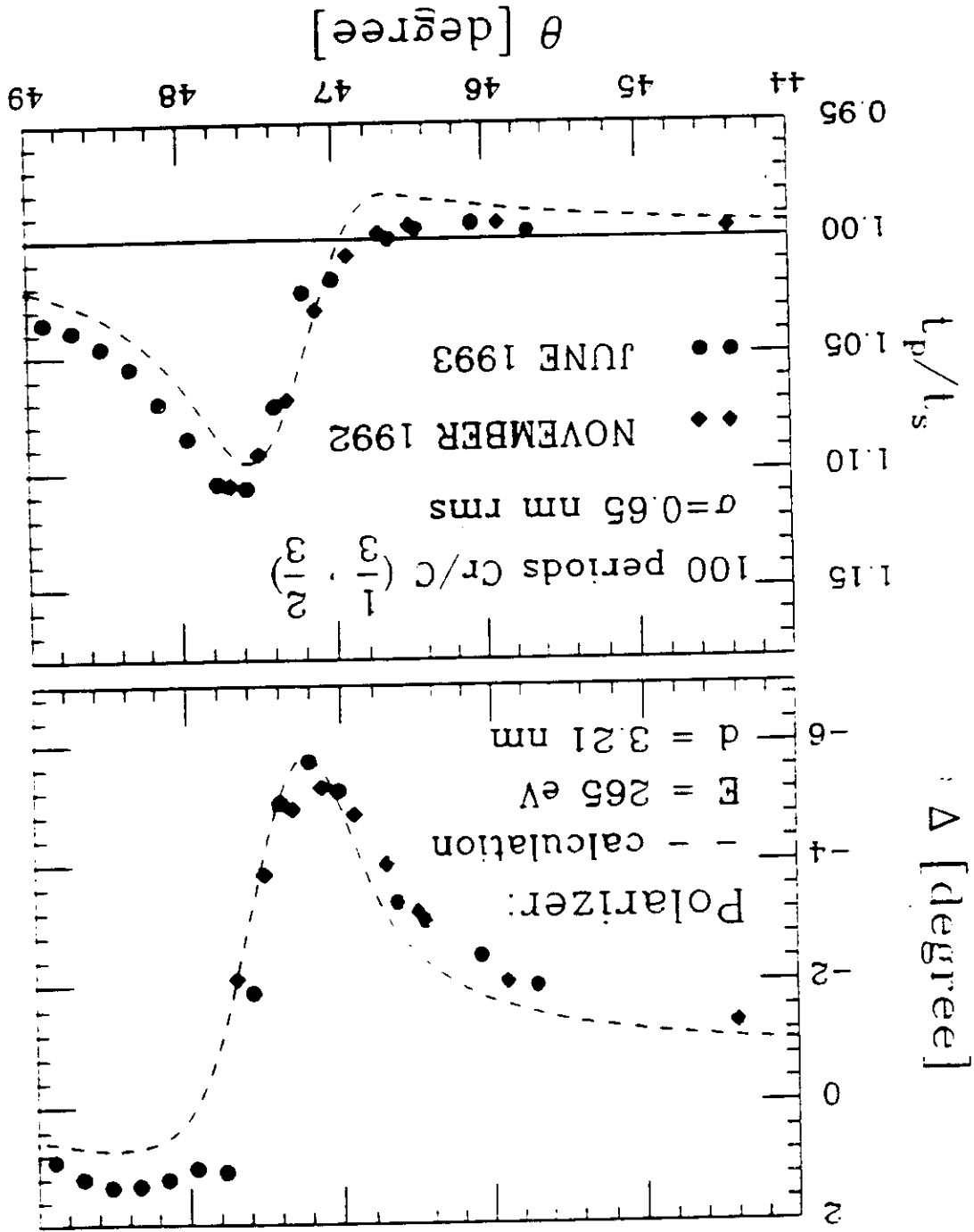
100 period multilayers with 1.07 nm Cr / 2.14 nm C sputtered onto silicon

- produced at Centre for X-ray Optics, LBL
- produced at SUCROPTROBE TRIESTE



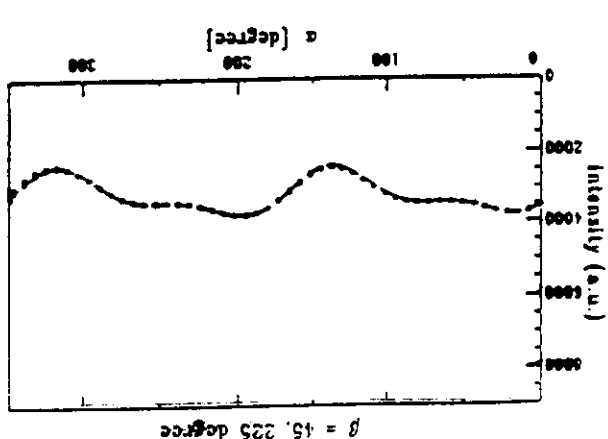
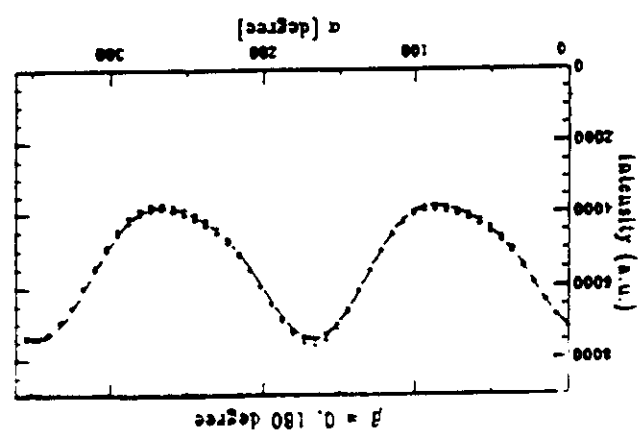
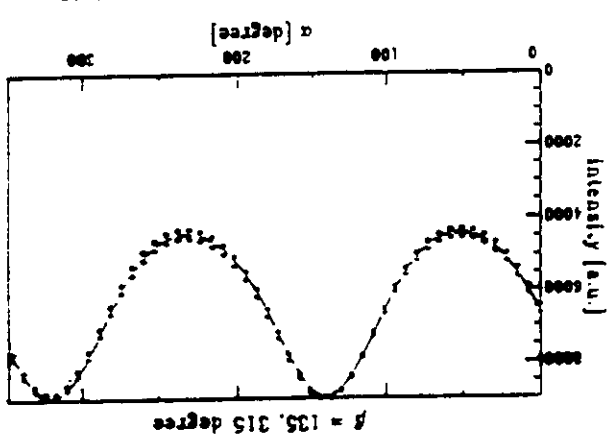
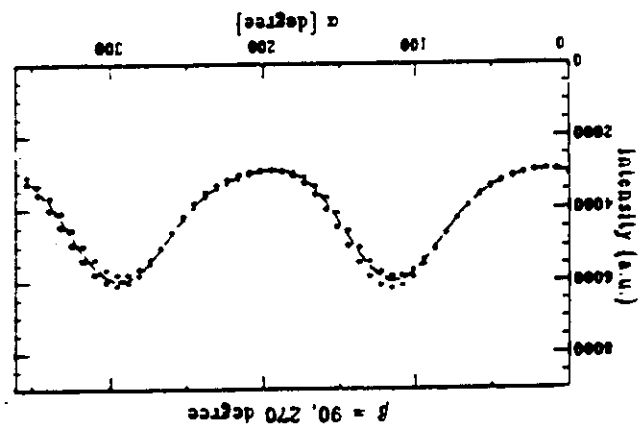
Soft x-ray phase retarders:
 SINCROTRONE TRIESTE: Silvia Di Fonzo, Bernd R. Müller, Werner Jark
 BESSY: Andreas Gaupp, Helmut Petersen, Franz Schäters
 CXRO, LBL: Jim Underwood

App. Opt. 33, 2624 (1994)

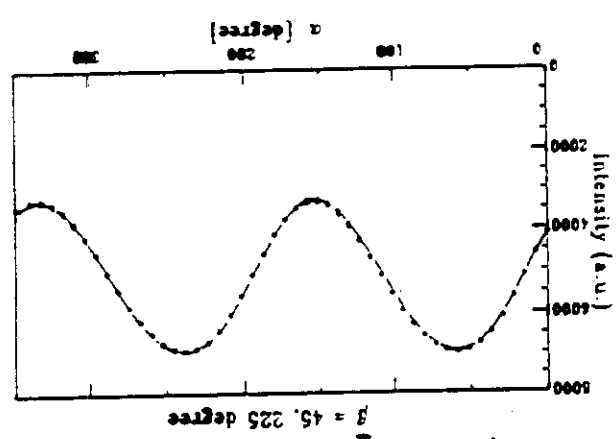
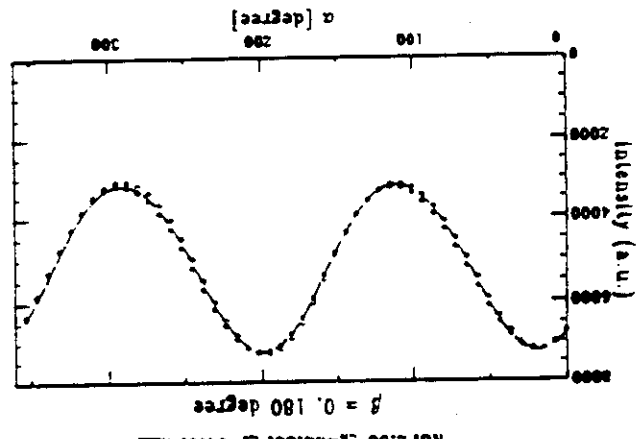
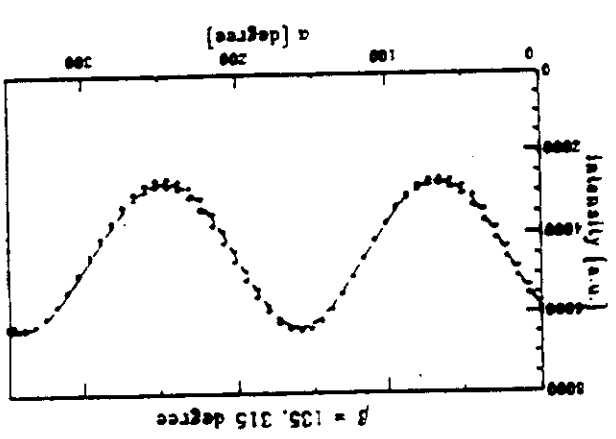
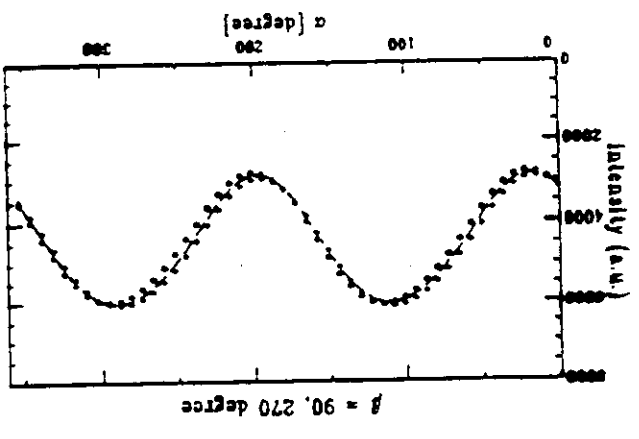


20:12:28 Camps

10-JAN-92



AGP2.02 $\epsilon_1=53.35, \epsilon_2=54.5, \Gamma=39.9$ Amp 91/09/14 $S_1 = +.13, S_2 = -.08, S_3 = .09$



AGP2.05 $\epsilon_1=53.35, \epsilon_2=54.5, \Gamma=39.9$ Amp 92/09/14 $S_1 = .10, S_2 = .09, S_3 = .45$

The BESSY polarimeter

(A. Gaupp and M. Mast: Rev. Sci. Instrum. 60, 2213 (1989))

$$t_{s1} = |t_{s1}| \exp(i\delta_{s1}) \quad r_{s2} = |r_{s2}| \exp(i\delta_{s2})$$

$$t_{p1} = |t_{p1}| \exp(i\delta_{p1}) \quad r_{p2} = |r_{p2}| \exp(i\delta_{p2})$$

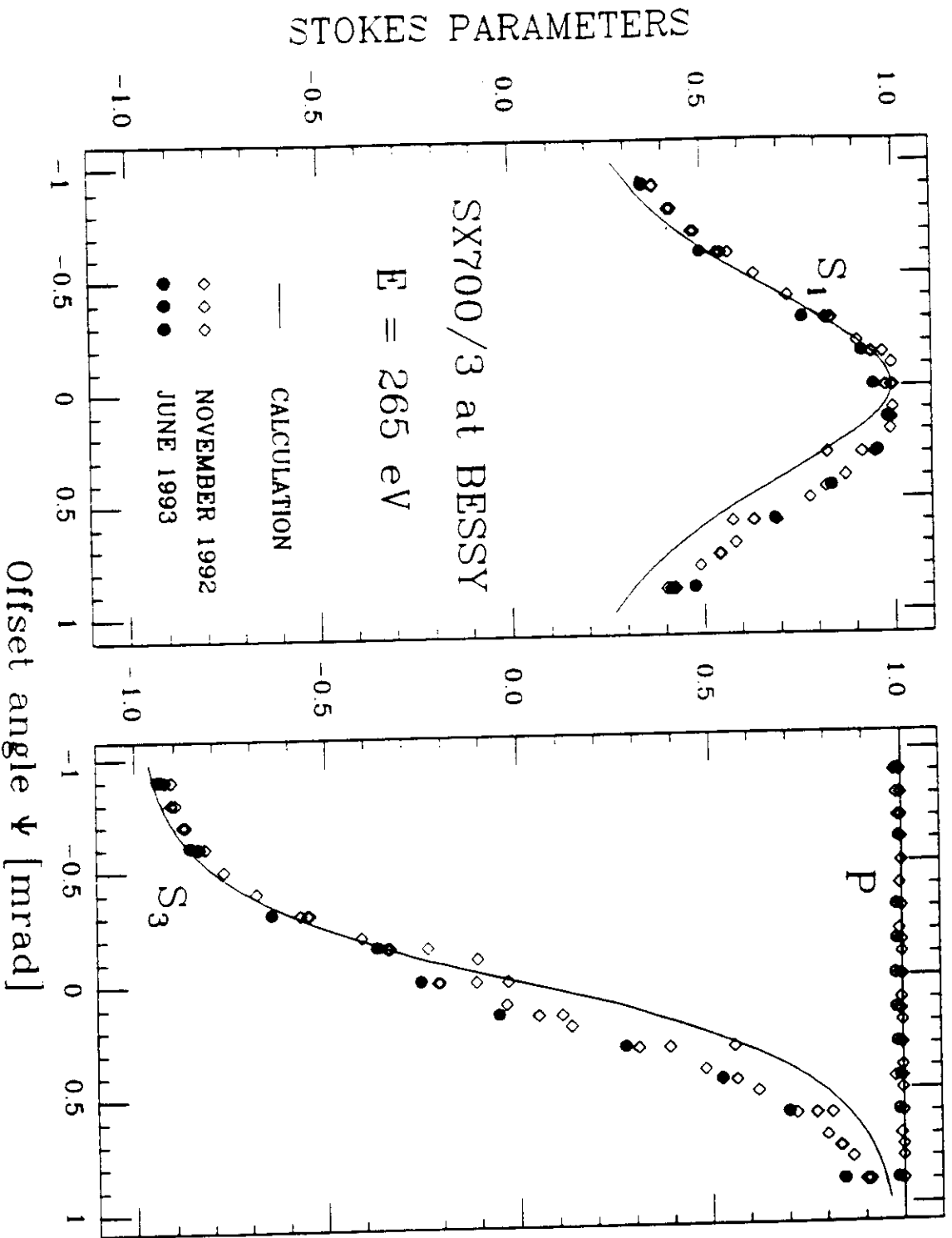
$$\Delta_1 = \delta_{p1} - \delta_{s1} \quad \Delta_2 = \delta_{p2} - \delta_{s2} = 0$$

$$\tan \psi_1 = |t_{p1}|/|t_{s1}| \quad \tan \psi_2 = |r_{p2}|/|r_{s2}|$$

Stokes-parameters: S_0, S_1, S_2 and S_3

$$I_{\text{pass}} = \left\{ \begin{array}{l} \frac{1}{2} (t_{s1}^2 + t_{p1}^2) * \frac{1}{2} (r_{s2}^2 + r_{p2}^2) * \\ S_0 \\ + \cos 2\alpha \quad [-S_1 \cos 2\psi_1] \quad + \sin 2\alpha \quad [-S_2 \cos 2\psi_1] \\ + \cos 2\alpha \quad \cos 2\beta \quad [-S_1 \cos 2\psi_2 * (1 + \sin 2\psi_1 \cos \Delta_1)/2] \quad + \sin 2\beta \quad [-S_2 \cos 2\psi_2 * (1 + \sin 2\psi_1 \cos \Delta_1)/2] \\ + \cos 2\alpha \quad \cos 2\beta \quad [+S_0 \cos 2\psi_1 * \cos 2\psi_2] \quad + \sin 2\alpha \quad \cos 2\beta \quad [+S_3 \sin 2\psi_1 * \cos 2\psi_2 \sin \Delta_1] \\ + \cos 2\alpha \quad \sin 2\beta \quad [-S_3 \sin 2\psi_1 * \cos 2\psi_2 \sin \Delta_1] \quad + \sin 2\alpha \quad \sin 2\beta \quad [+S_0 \cos 2\psi_1 * \cos 2\psi_2] \\ + \cos 4\alpha \quad \cos 2\beta \quad [-S_1 \cos 2\psi_2 * (1 - \sin 2\psi_1 \cos \Delta_1)/2] \quad + \sin 4\alpha \quad \cos 2\beta \quad [-S_2 \cos 2\psi_2 * (1 - \sin 2\psi_1 \cos \Delta_1)/2] \\ + \cos 4\alpha \quad \sin 2\beta \quad [+S_2 \cos 2\psi_2 * (1 - \sin 2\psi_1 \cos \Delta_1)/2] \quad + \sin 4\alpha \quad \sin 2\beta \quad [-S_1 \cos 2\psi_2 * (1 - \sin 2\psi_1 \cos \Delta_1)/2] \end{array} \right\}$$

SOFT X-RAY POLARIZATION
 SYNCHROTRONE TRIESTE: Silvia Di Fonzo, Bernd R. Müller, Werner Jark
 BESSY: Andreas Gaupp, Helmut Petersen, Franz Schäfers
 CXRO, LBL: Jim Underwood



calculation does
 not consider
 polarization and
 depolarization
 effects in
 waveguides

Similar object produced also by
 Kimmari (SPRINGER) and by
 Northing (Conts for x-ray optics, LBL)

Abbildung 4: Transmissionsverhältnis des TML von s- zu p-polarisierter Komponente aufgetragen über den Glanzwinkel θ

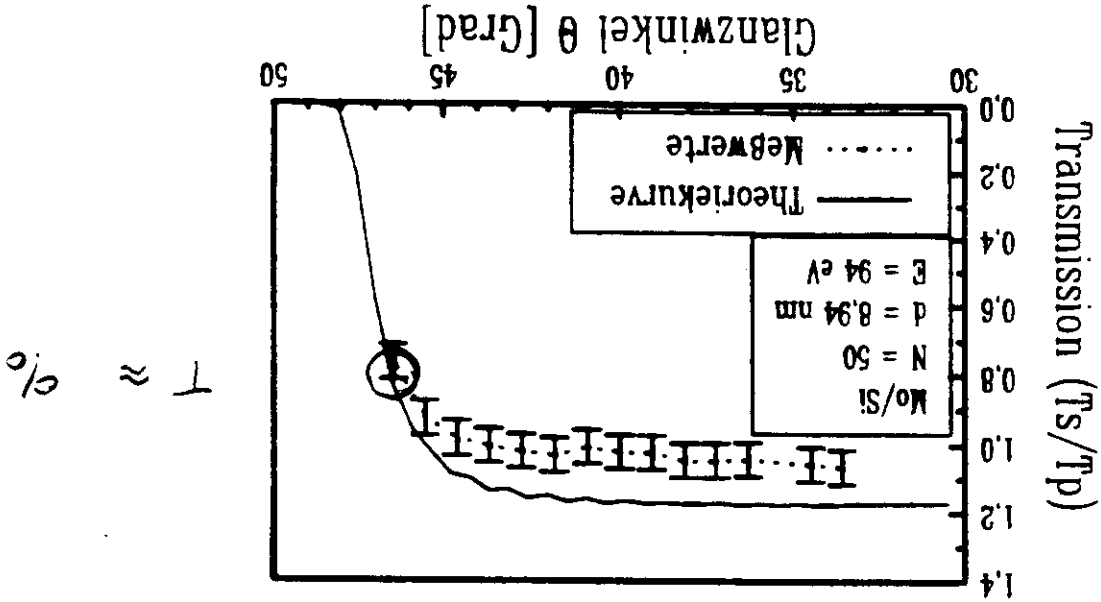
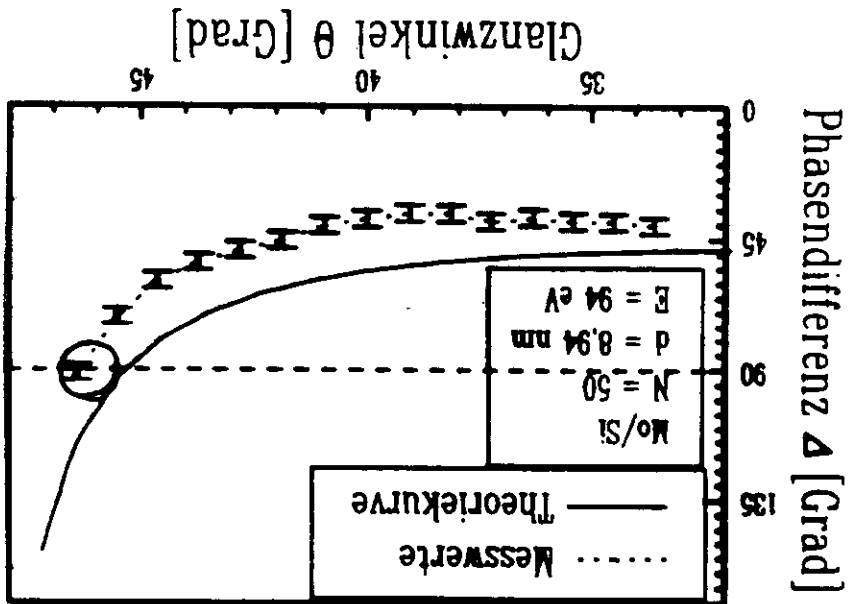
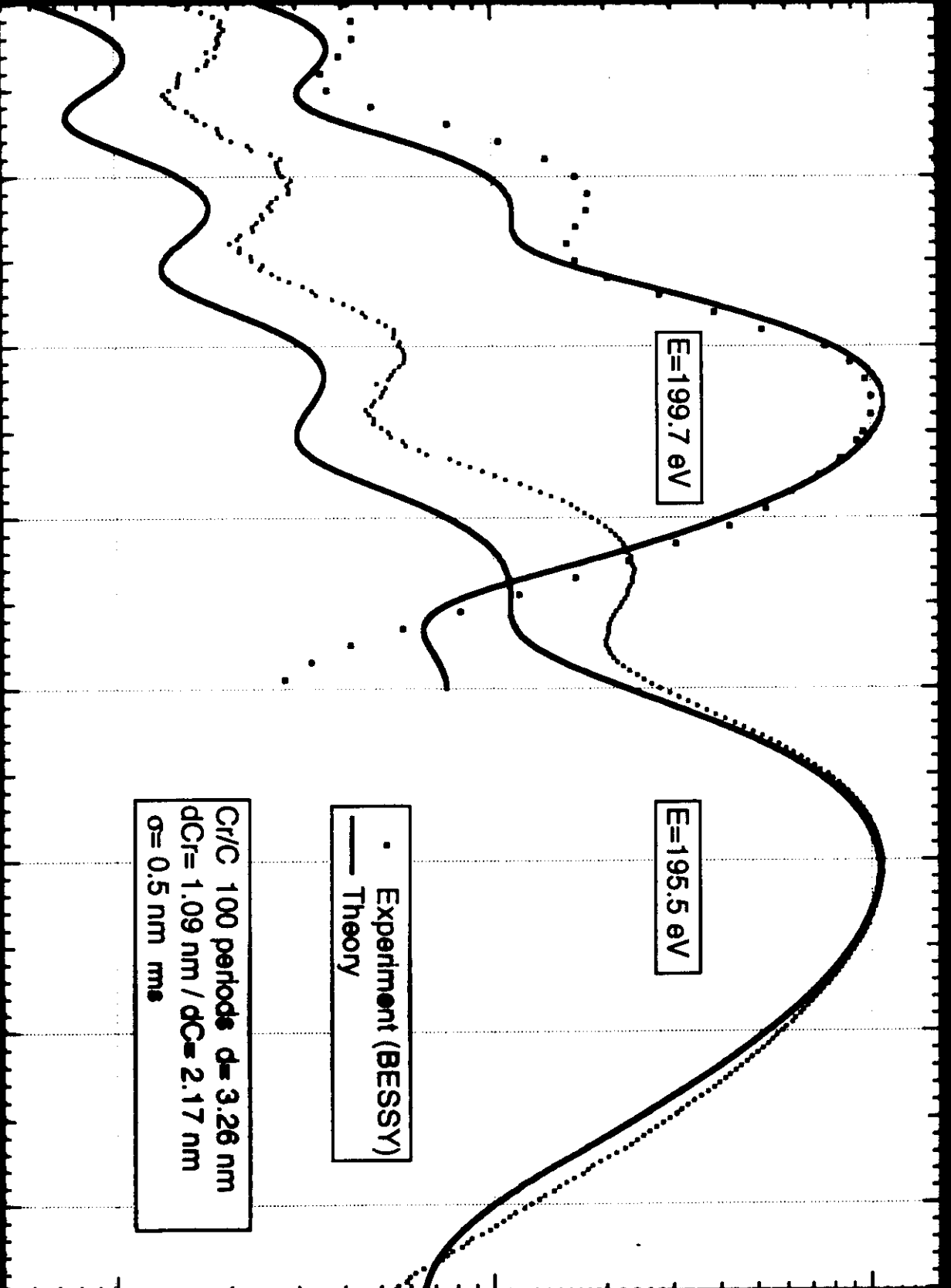


Abbildung 3: Phasenschiebung Δ des TML aufgetragen über den Glanzwinkel θ



Phase shift of group of ... (University of Bielefeld) for use of BESSY



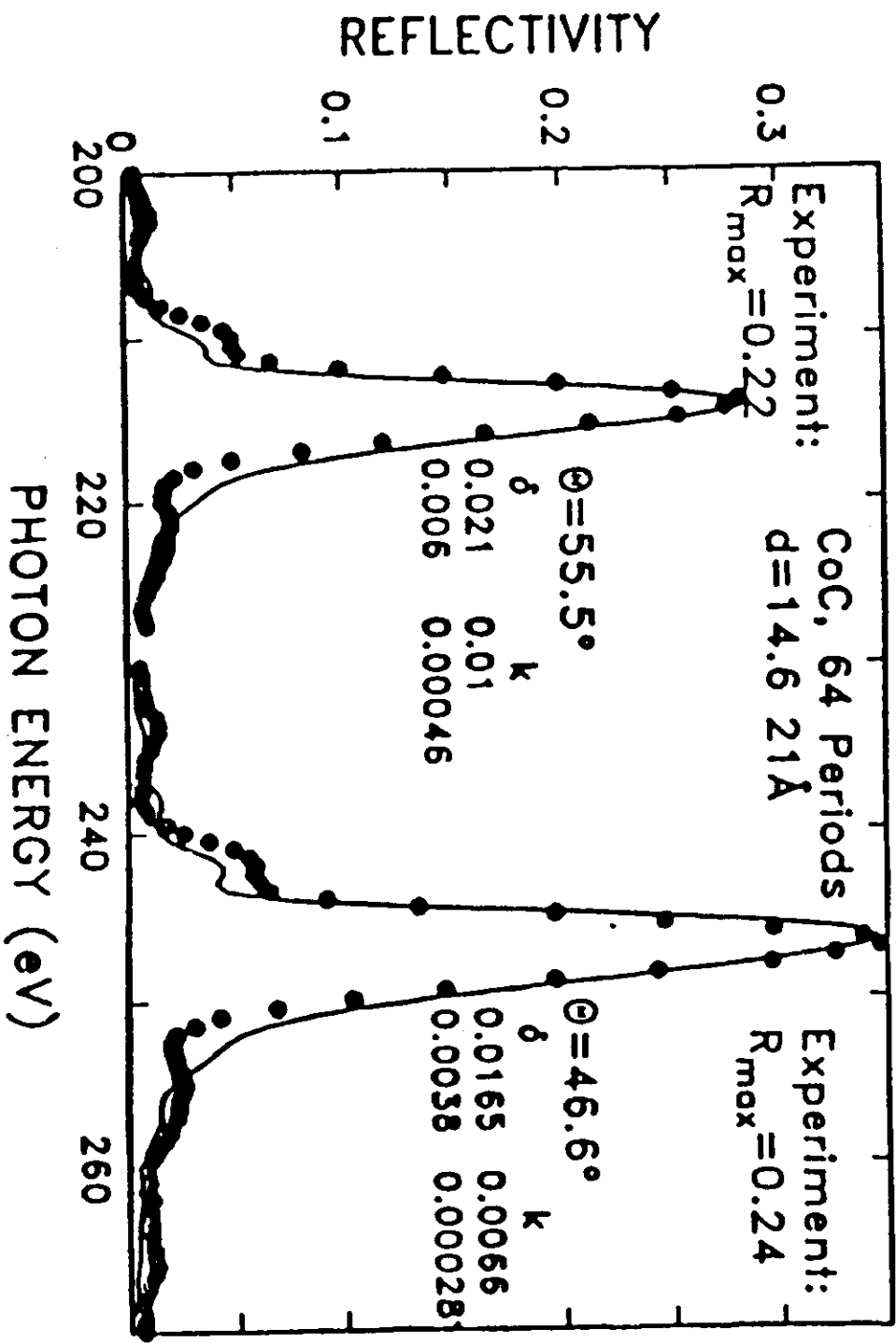
E=199.7 eV

E=195.5 eV

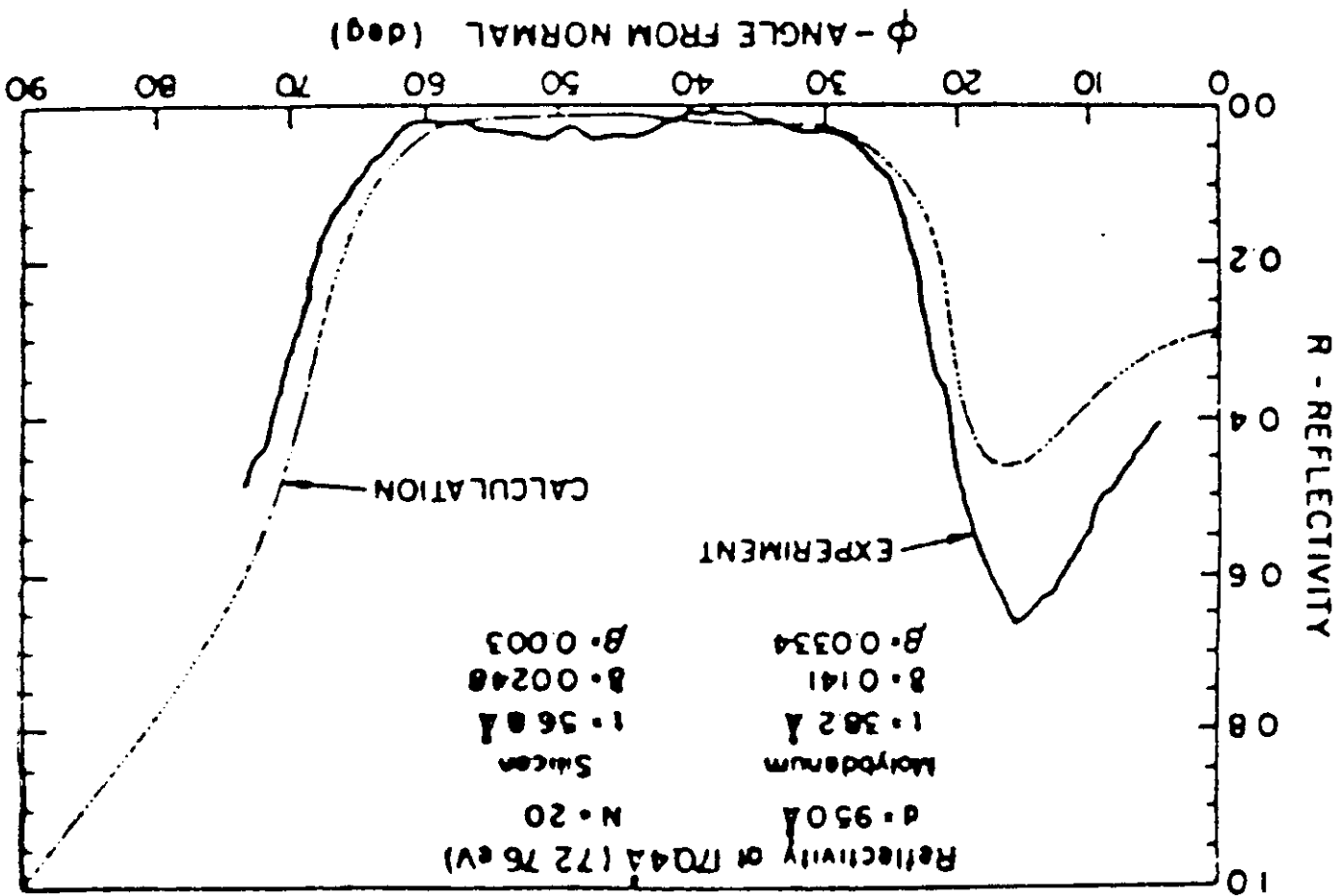
• Experiment (BESSY)
— Theory

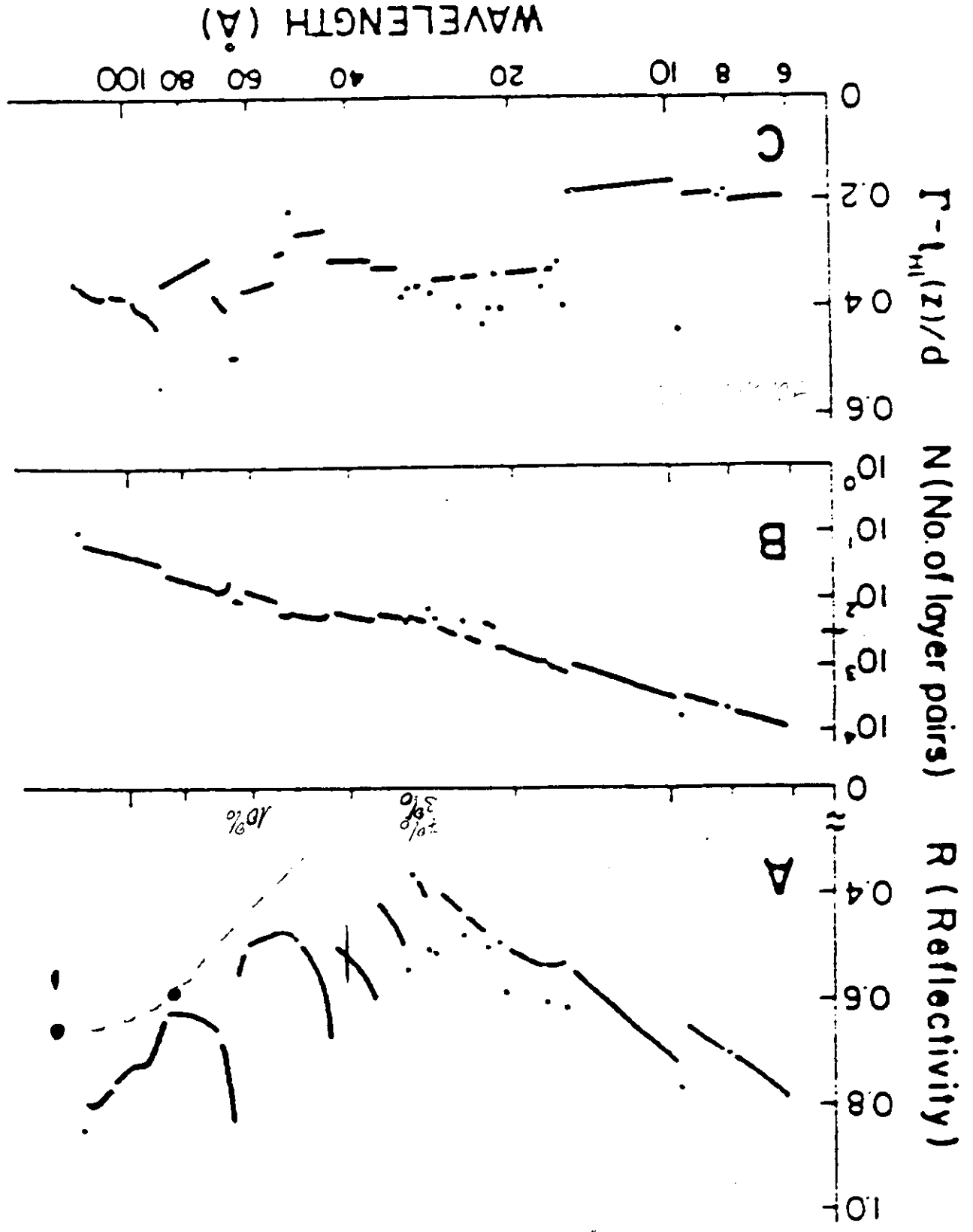
Cr/C 100 periode $d = 3.26$ nm
 $dCr = 1.09$ nm / $dC = 2.17$ nm
 $\sigma = 0.5$ nm rms

E. Spiller et al.



T. W. Barber et al





for normal incidence

KOSAMBIUM - FABRIES
 J.E. Rosenthal, Ph.D. Thesis, Rochester (1982)

transmission with normal incidence optics with multilayers

ADVANCES IN MULTILAYER X-RAY/EUV OPTICS: SYNTHESIS, PERFORMANCE, AND INSTRUMENTATION

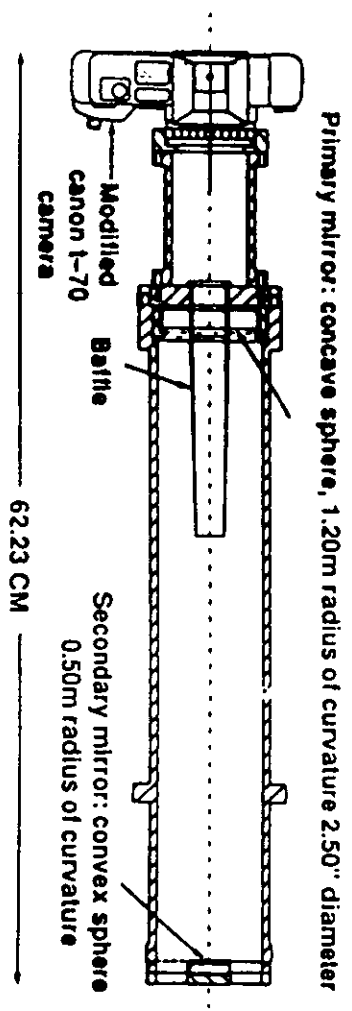


Fig. 9. Schematic of the normal incidence Cassegrain multilayer telescope.⁵⁶ Angular resolution is 1.2 arcsec, effective focal length is 2m, and bandpass is 6.5% at 175 Å.

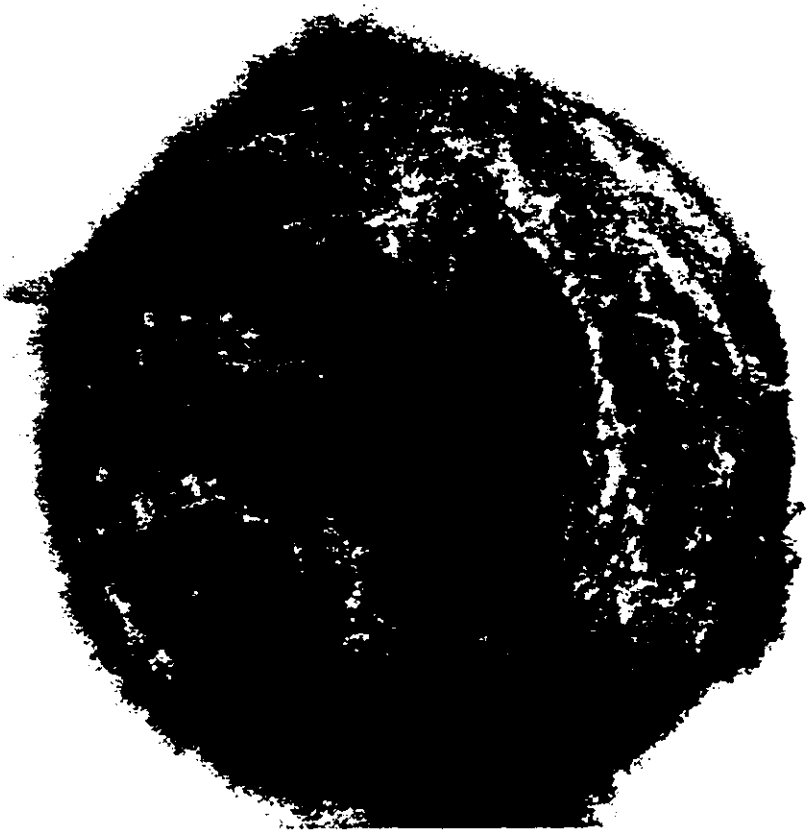


Fig. 10. Photograph of the solar corona at 1,000,000 K obtained with a multilayer Cassegrain telescope. The multilayers in molybdenum-silicon deposited using magnetron sputter for physical vapor deposition technology.

The work which has been performed at the
SINCROTRONE TRIESTE
(multilayer production and
tests with hard x-rays)
was done by:

Silvia Di Fonzo
Werner Jark
Bernd R. Müller
Gerard Soullie'

The experiments using soft x-rays were
performed at different laboratories:

at HASYLAB at DESY, Hamburg, Germany

in collaboration with
Ingo DieI
Jan Friedrich
C. Kunz

at BESSY, Berlin, Germany

in collaboration with
Andreas Gaupp
Helmuth Petersen

Franz Schäfers
James H. Underwood (LBL)
Hans Christof Mertins