



H4.SMR/1058-17

WINTER COLLEGE ON OPTICS

9 - 27 February 1998

The Fundamentals of Speckles and their Applications

Part II

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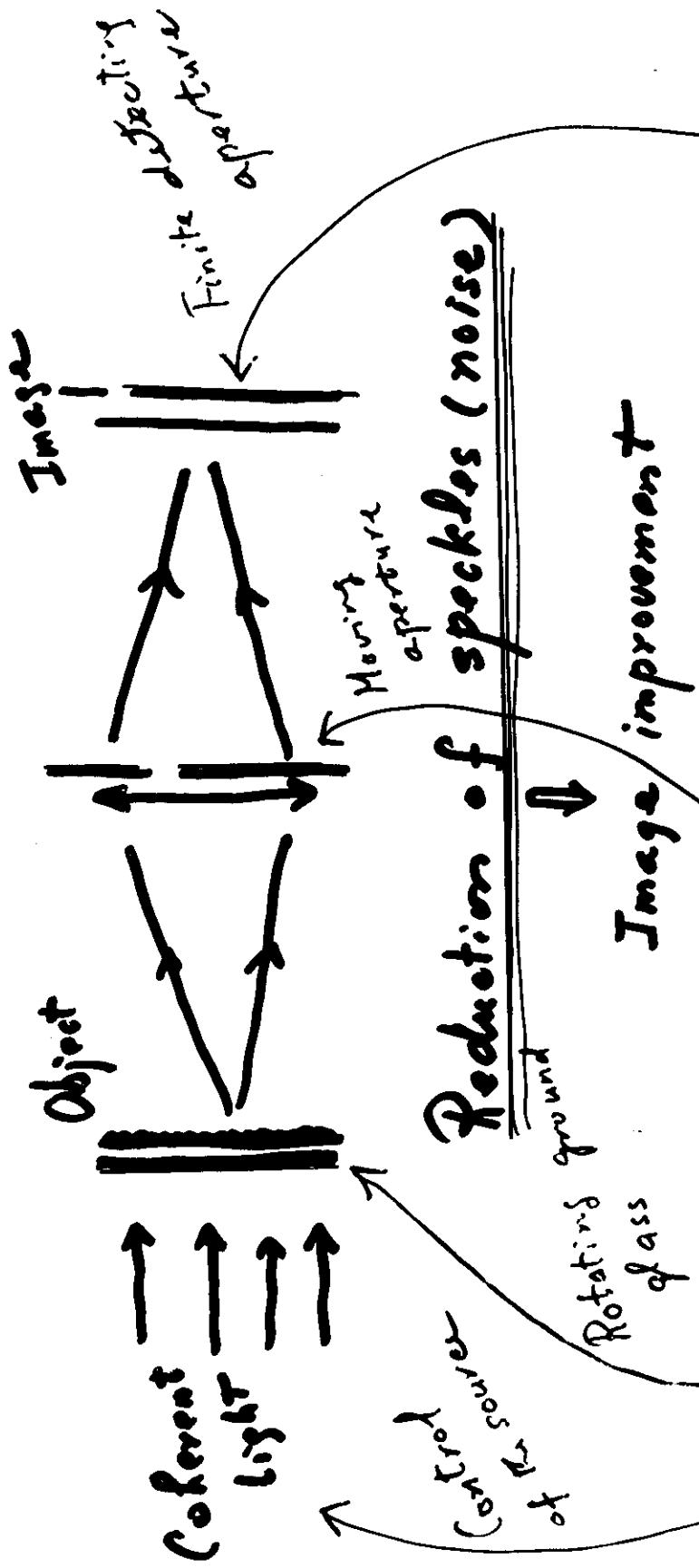
Superposition of speckles

noise

Speckle reduction



Image improvement



Reduction of speckles (noise)

Image improvement

- ① Temporally partially coherent light
- ② Spatially partially coherent light

- ③ Time-averaging by means of the moving aperture
- ④ Space-averaging by means of the finite detecting aperture

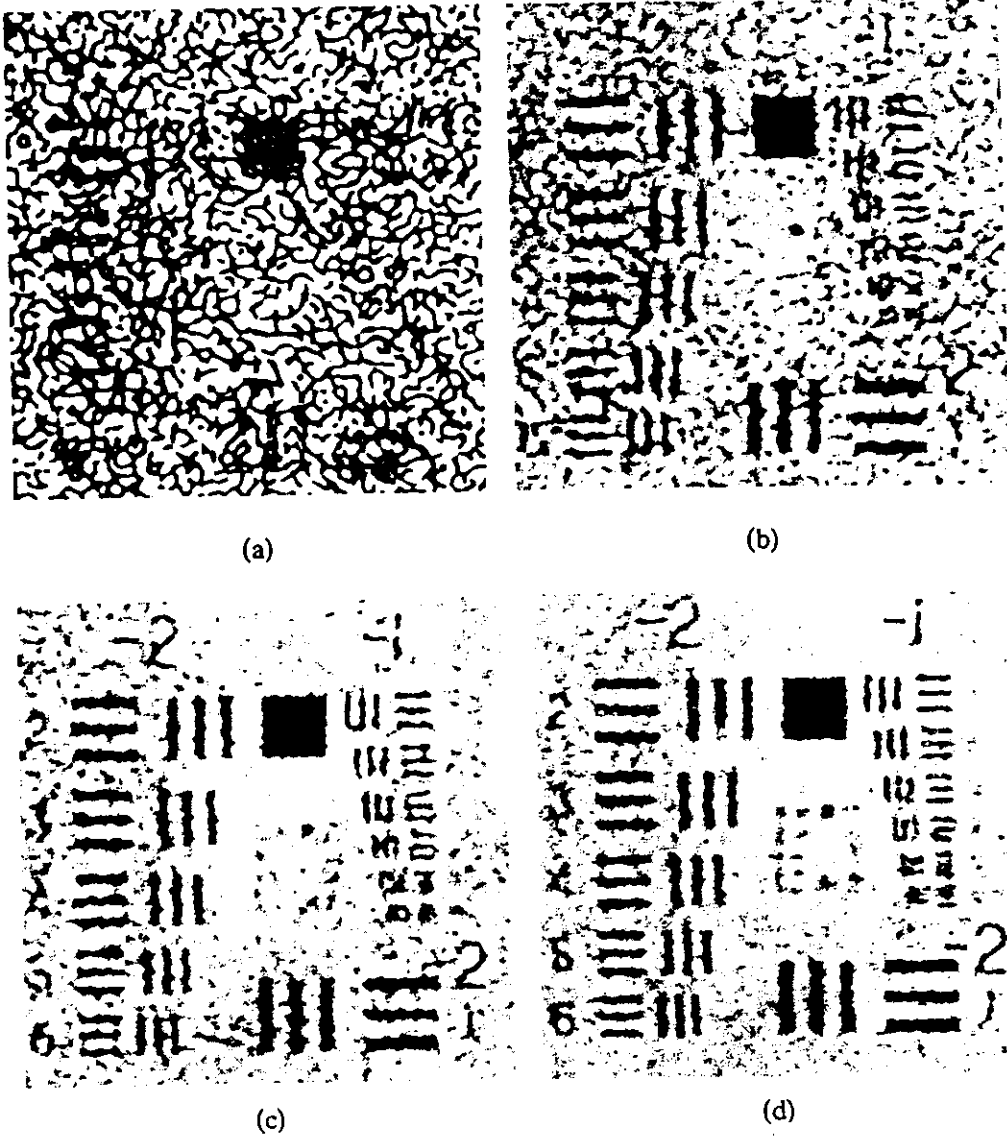


Fig. 13. Photographic images obtained for exposure time $T = 50$ s through a stationary aperture of width $W = 0.35$ mm with radius $R = 1.0$ mm (a), and through a rotating aperture of width $W = 0.35$ mm with $R = 1.0$ mm (b), 3.0 mm (c), and 8.0 mm (d).

Dynamic speckles

Two fundamental motions of speckles

Boiling } Mixed state
Translation } of boiling and translation

Characteristics dependence of dynamic speckles

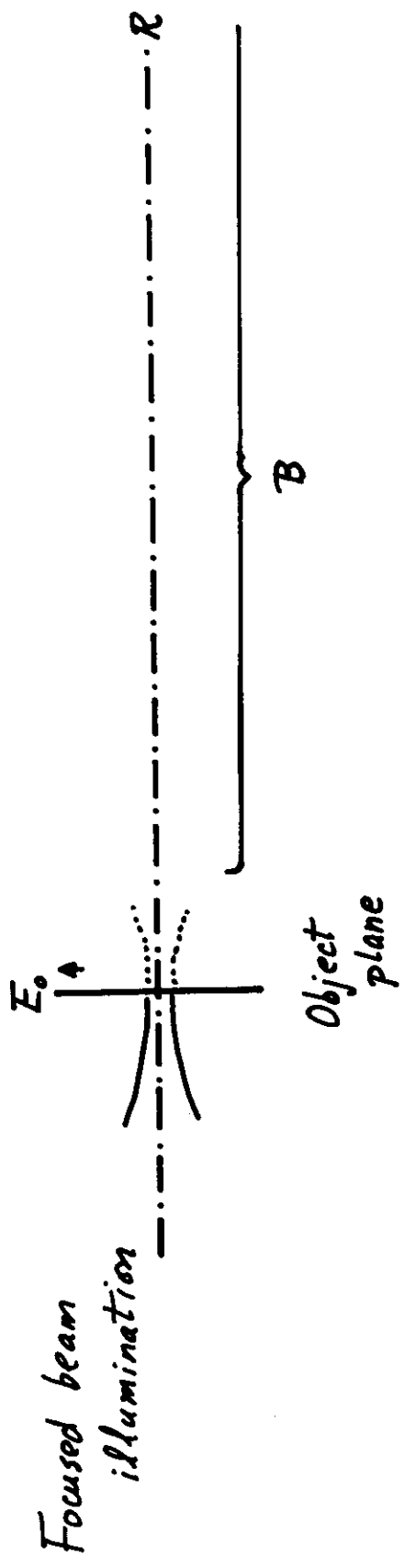
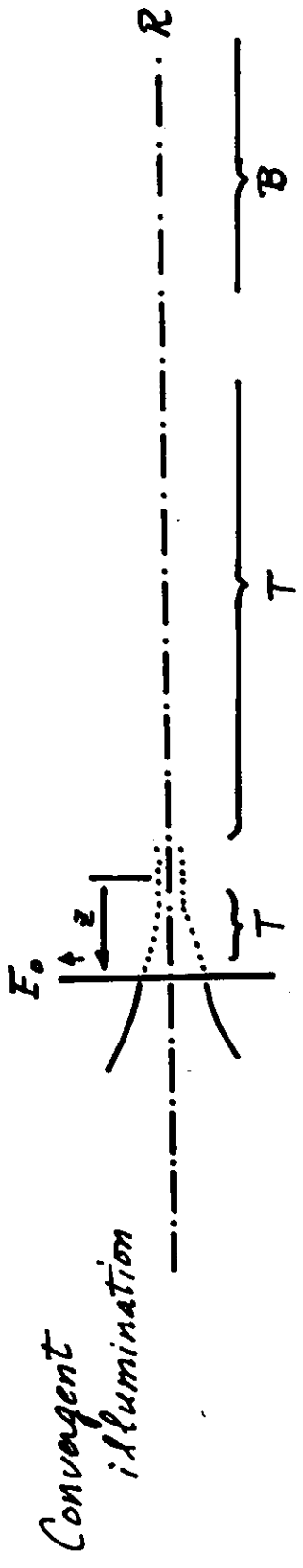
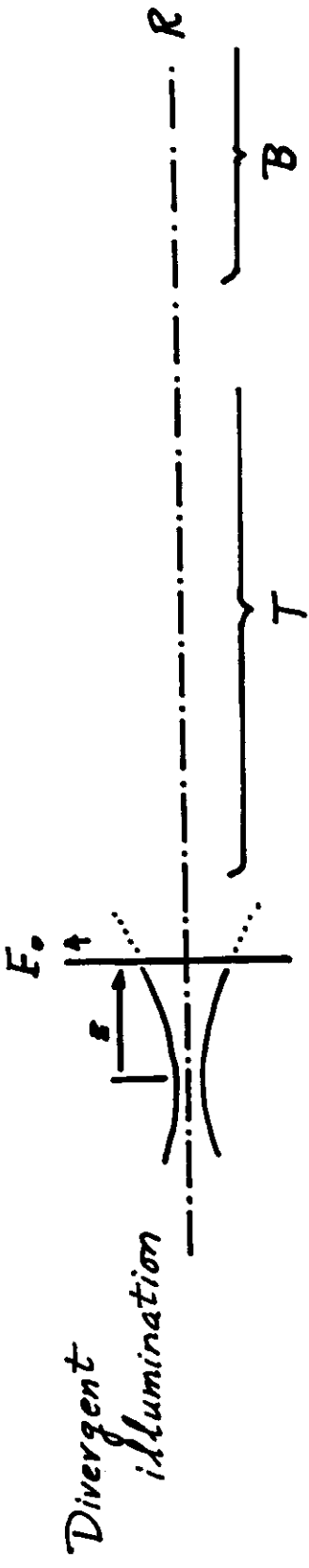
Motion of the diffuse object → Velocity measurements

Parameters of the optical systems including

the conditions of illumination and detection

Fundamental physical quantity: Space-time correlation of the speckle field

Regions of boiling [B] and translation [T]



Velocity measurement of the diffuse object using dynamic speckles

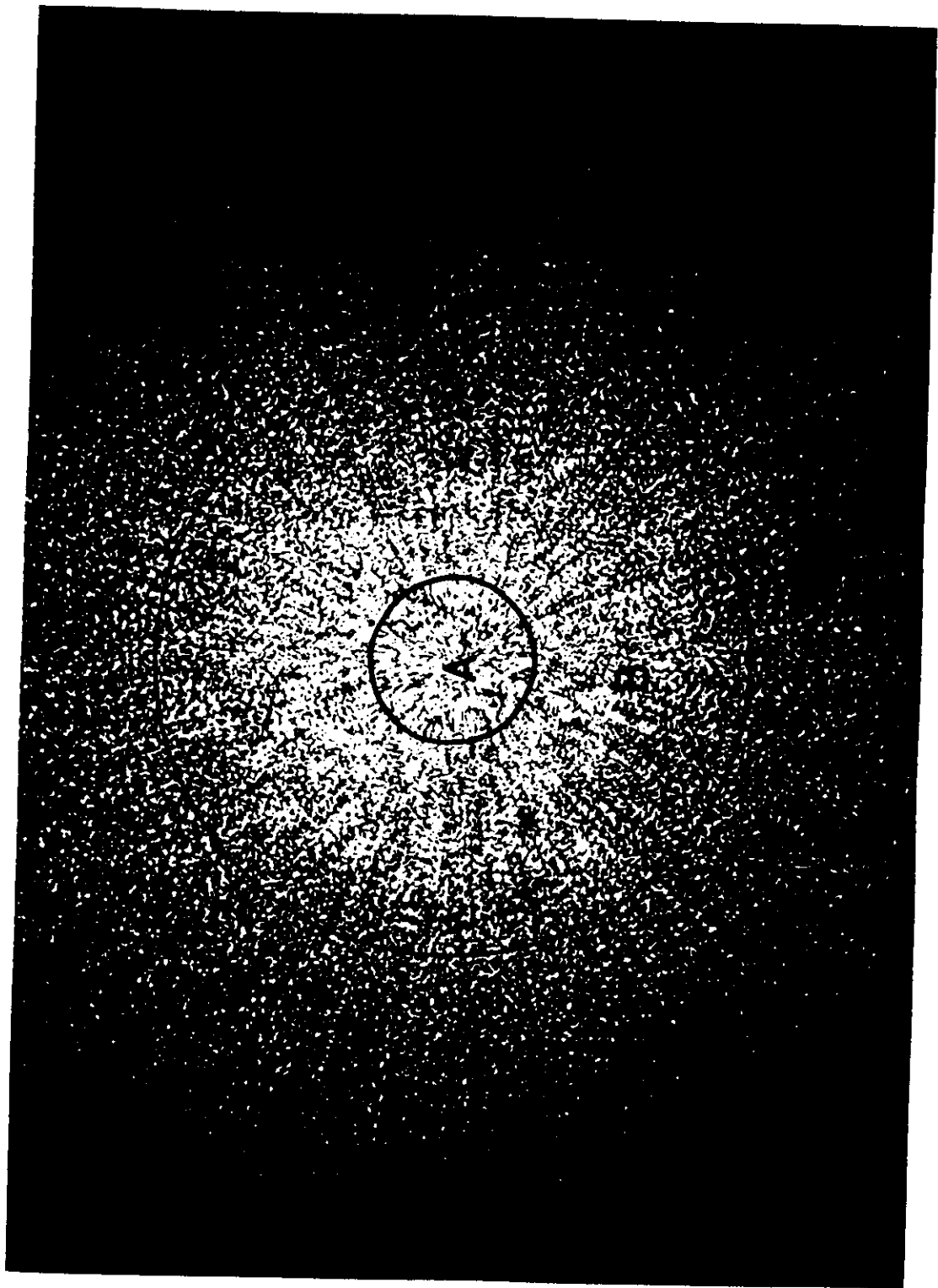
Measuring method	Number of detectors	Boiling	Translation
Autocorrelation method	1	0	0
Cross-correlation method	2	x	0
Spatial filtering method	1	x	0
Zero-crossing method	1	0	0
circular aperture		x	0
directional aperture			0

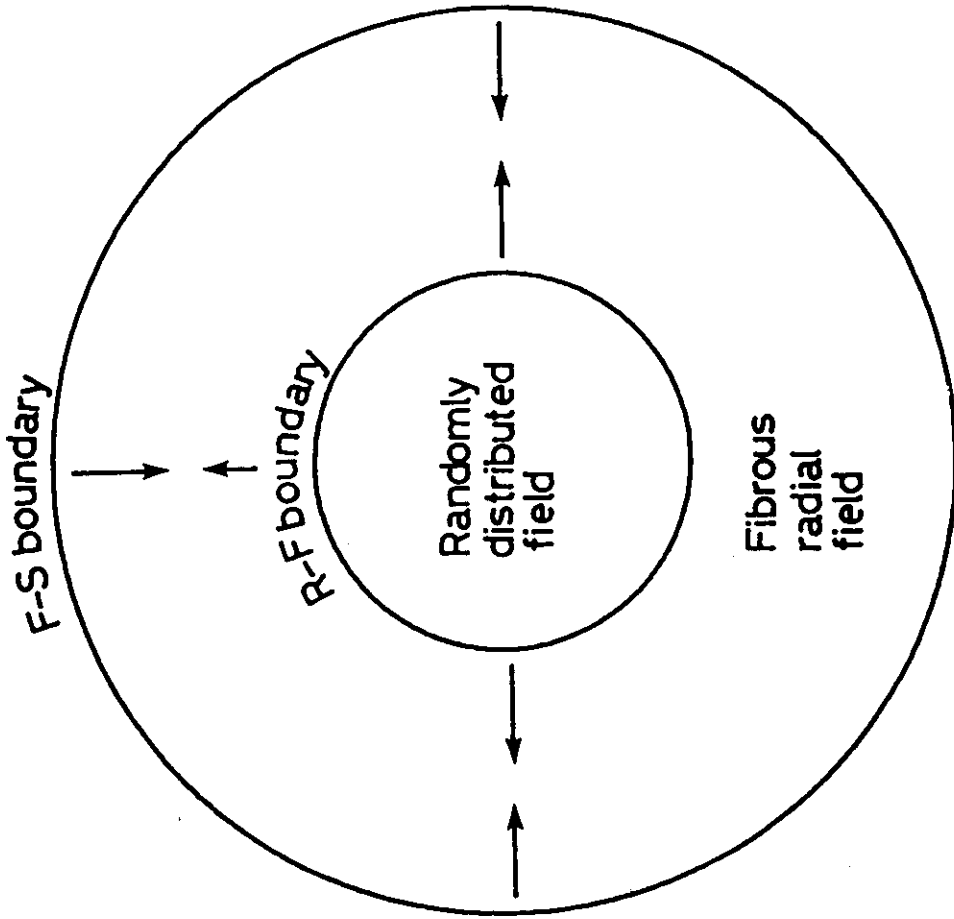
White light . polychromatic light

speckle

Image field speckle : Color separation

Diffraction field speckle : Fibrous structure





Radially
separated
field

Fiber speckle

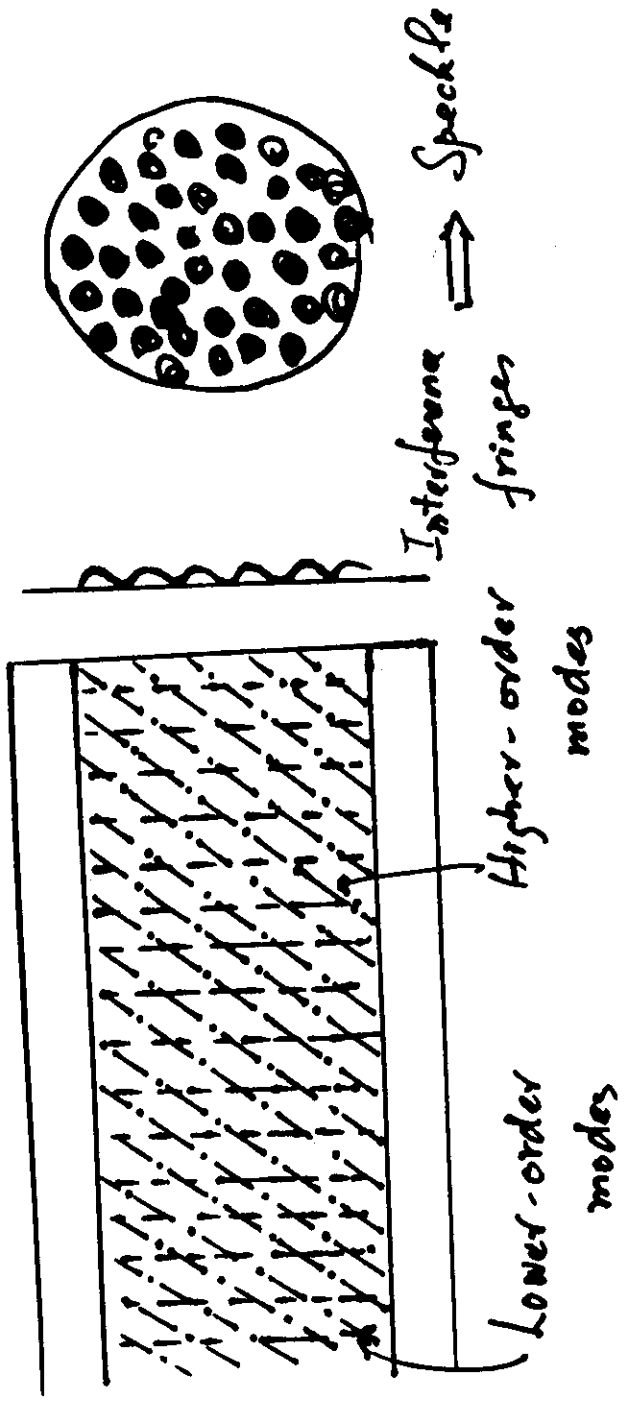
Modal noise



Fiber sensor

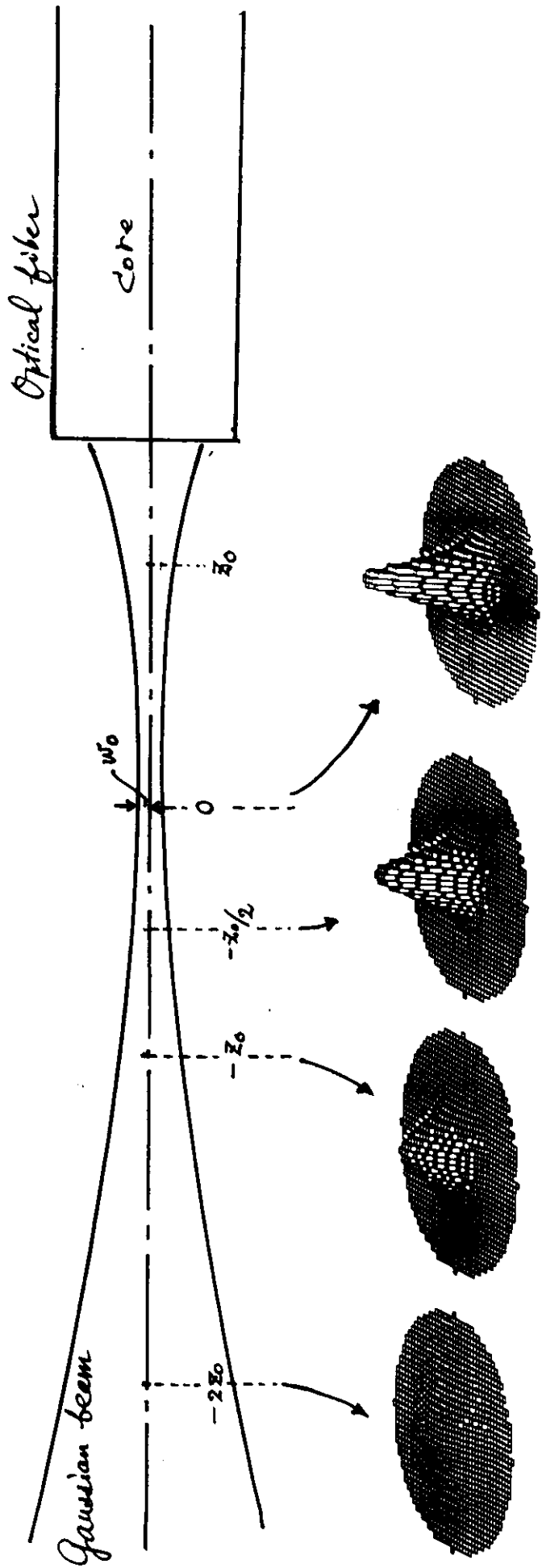
Modal noise \rightleftharpoons Fiber speckle

Interference phenomena of fiber modes



Causes of fiber speckle

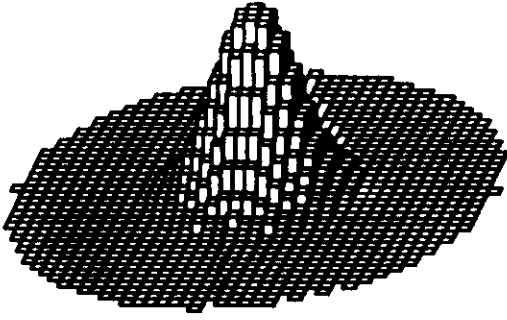
- ① High coherence properties of the source (laser)
- ② Selectivity of modes
- ③ Small numbers of modes
- ④ Low mode dispersion



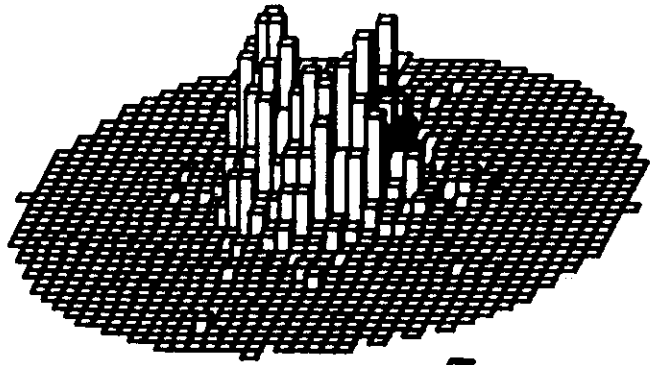
$$z_0 = \frac{\pi w_0^2}{\lambda}$$

Focusing misalignment

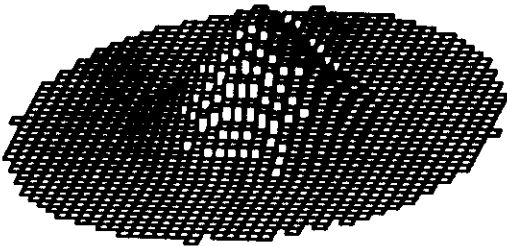
$$\sqrt{\langle \Delta \theta^2 \rangle} = \begin{cases} 0.01 & \text{(inside)} \\ 0.05 & \text{(outside)} \end{cases}$$



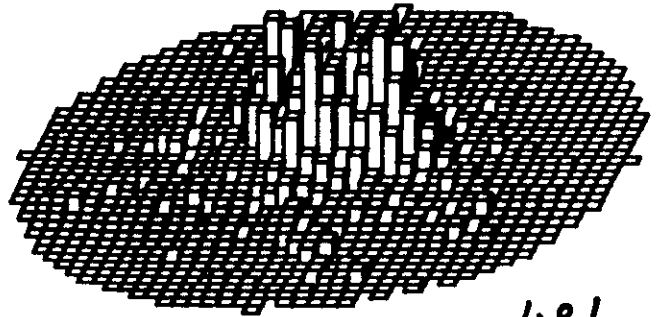
$\Delta z / \sigma_0 = 0.5$



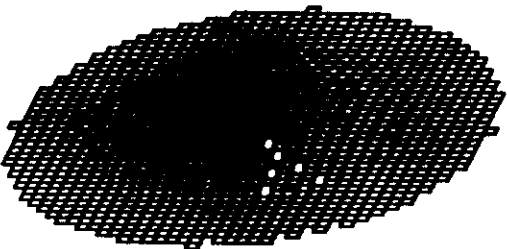
$\sigma_y / \sigma_z = 1.03$



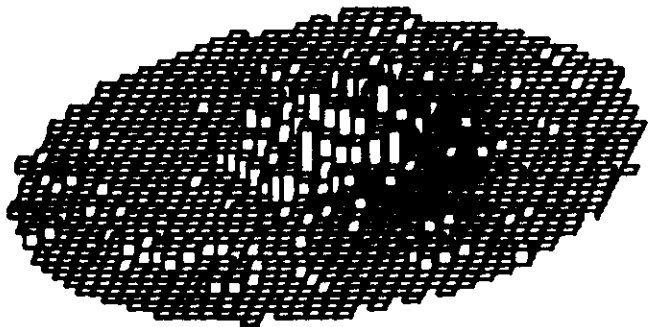
1.0



1.01



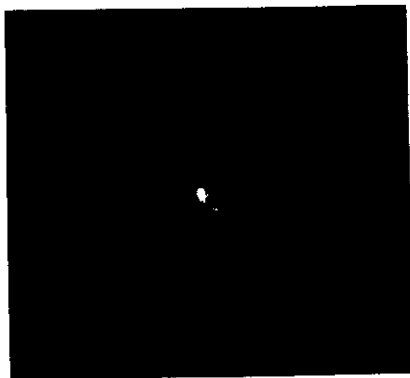
2.0



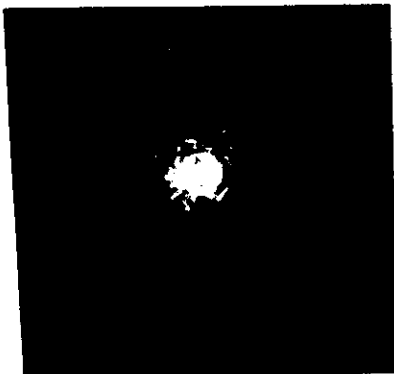
1.01

Focusing
misalignment

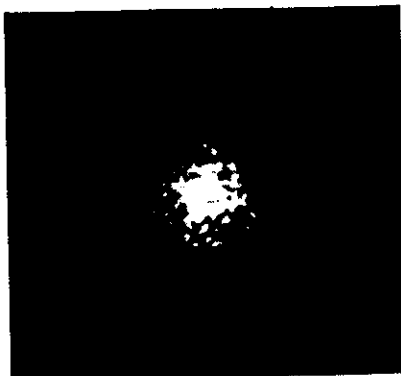
$$\frac{\Delta z}{z_0} = 0$$



-0.22



+0.22



-0.45



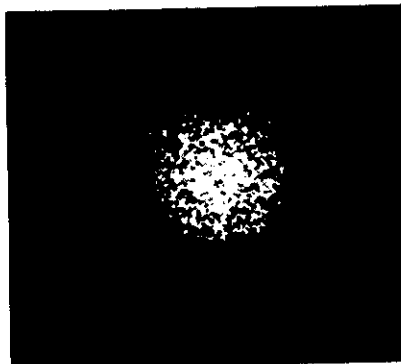
+0.45



-0.67

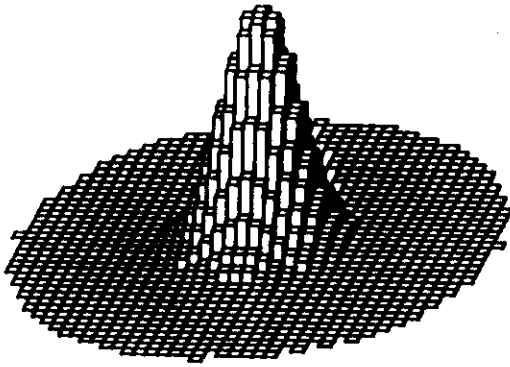


+0.67

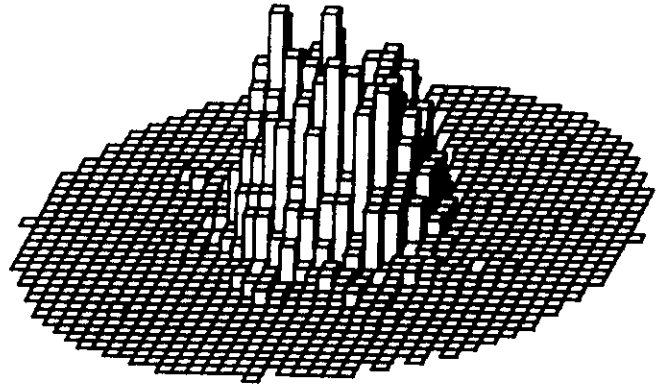


Axial misalignment

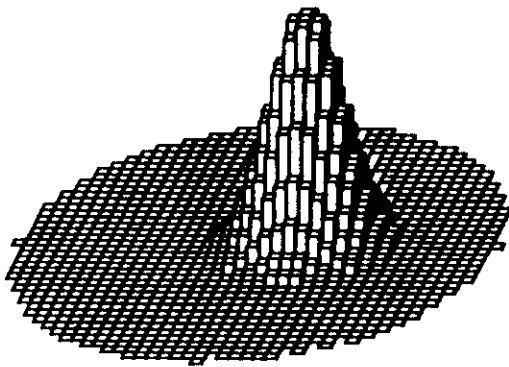
inhomogeneous : $\sqrt{\langle \Delta \sigma^2 \rangle} = \begin{cases} 0.01 & \text{(inside)} \\ 0.05 & \text{(outside)} \end{cases}$



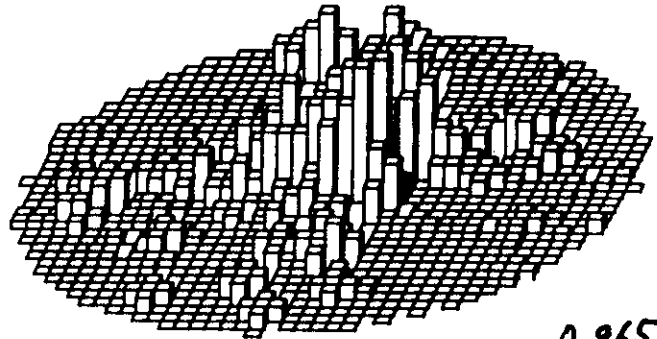
$$\frac{\Delta a}{a} = 0.0$$



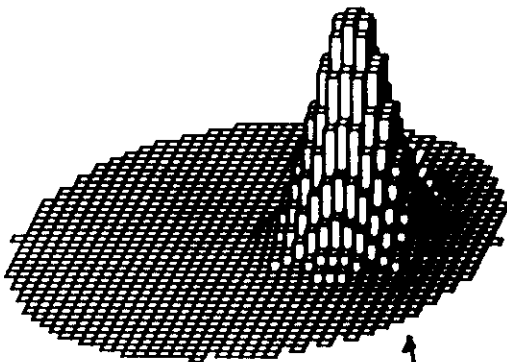
$$\sigma_y/\sigma_x = 0.988$$



0.2

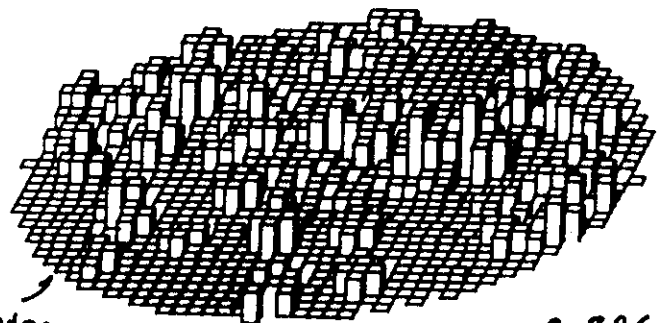


0.965



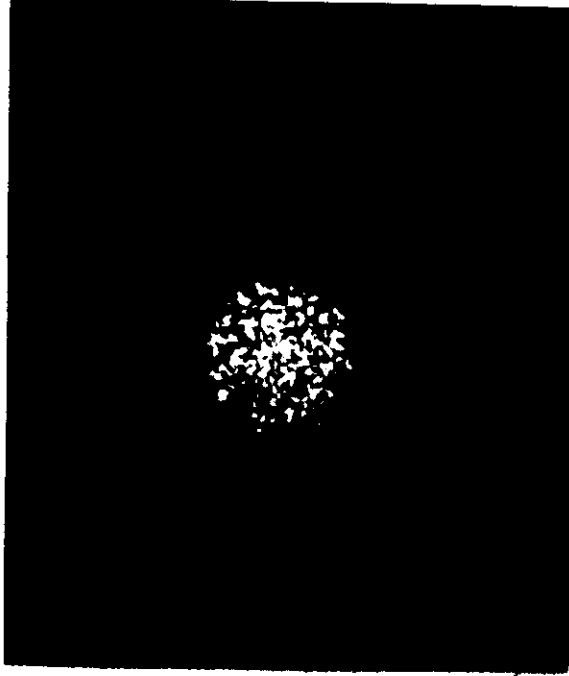
0.4

core region

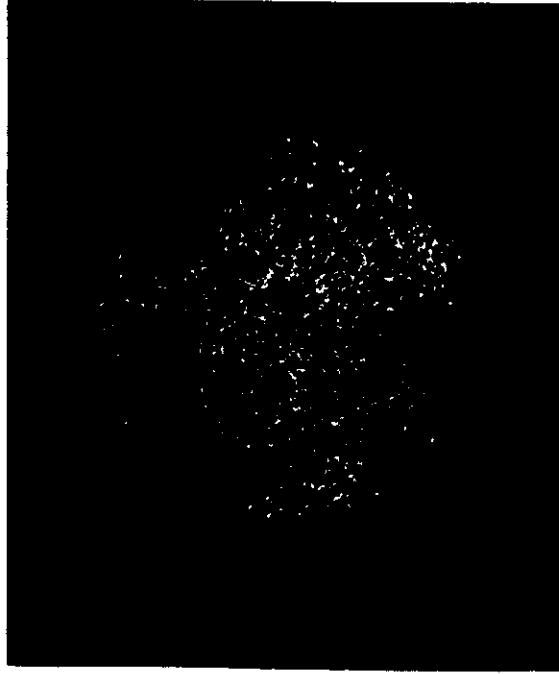


0.786

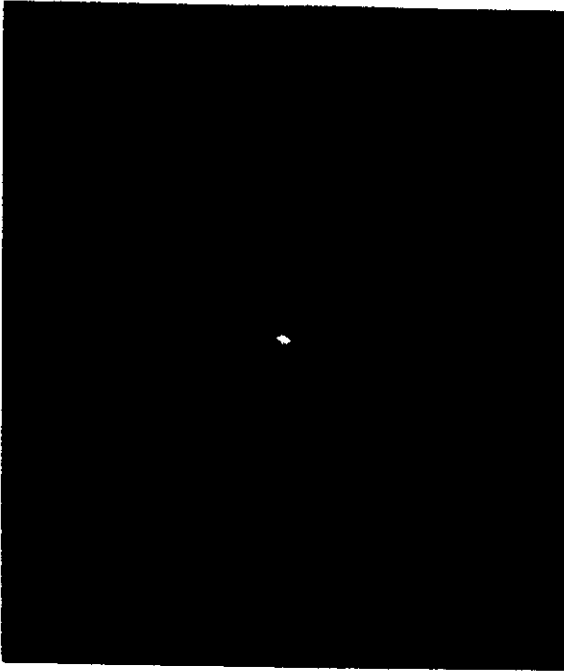
Axial
misalignment



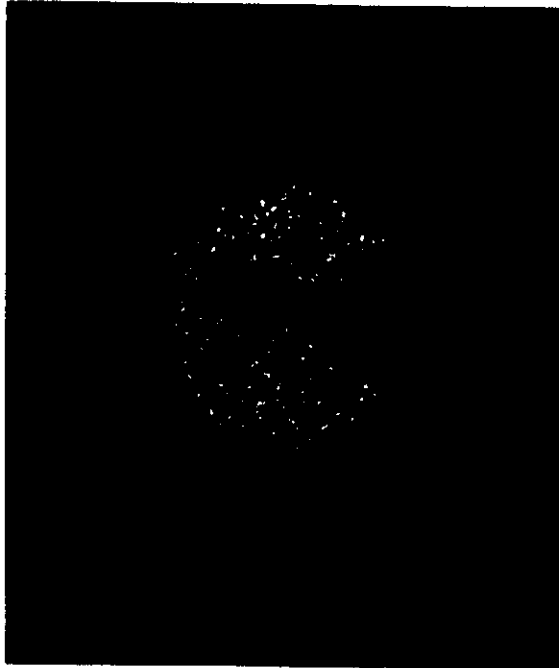
0.25



0.75



$\Delta\sigma/\sigma = 0$



0.50

Bio-speckles

Larger speckles produced from living objects



"Bio-speckles"

Time-varying properties of bio-speckles



Physiological movements of living objects

Spectral and correlation properties of bio-speckles

from

- skin surface

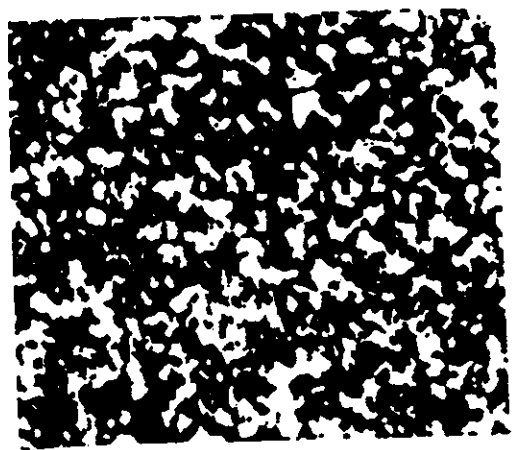
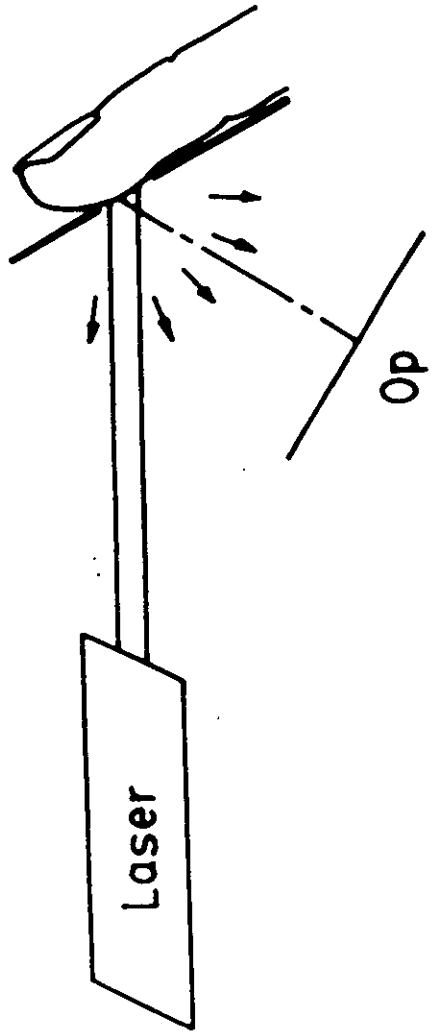
- gastric mucous membrane ←

Stomach surface

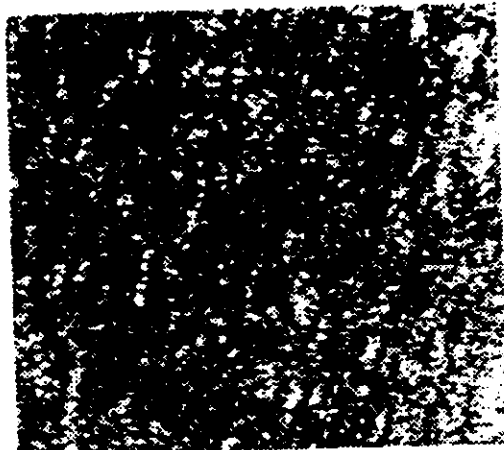
- ocular fundus



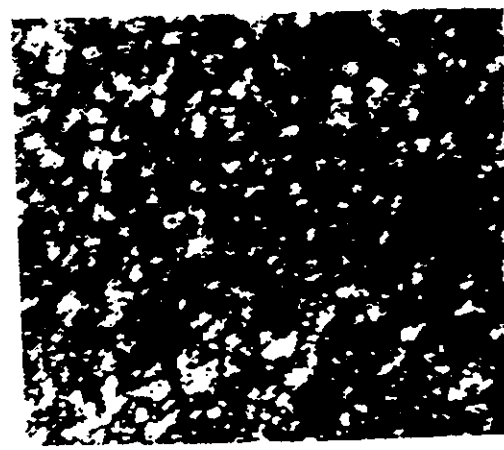
Measurements of blood flow



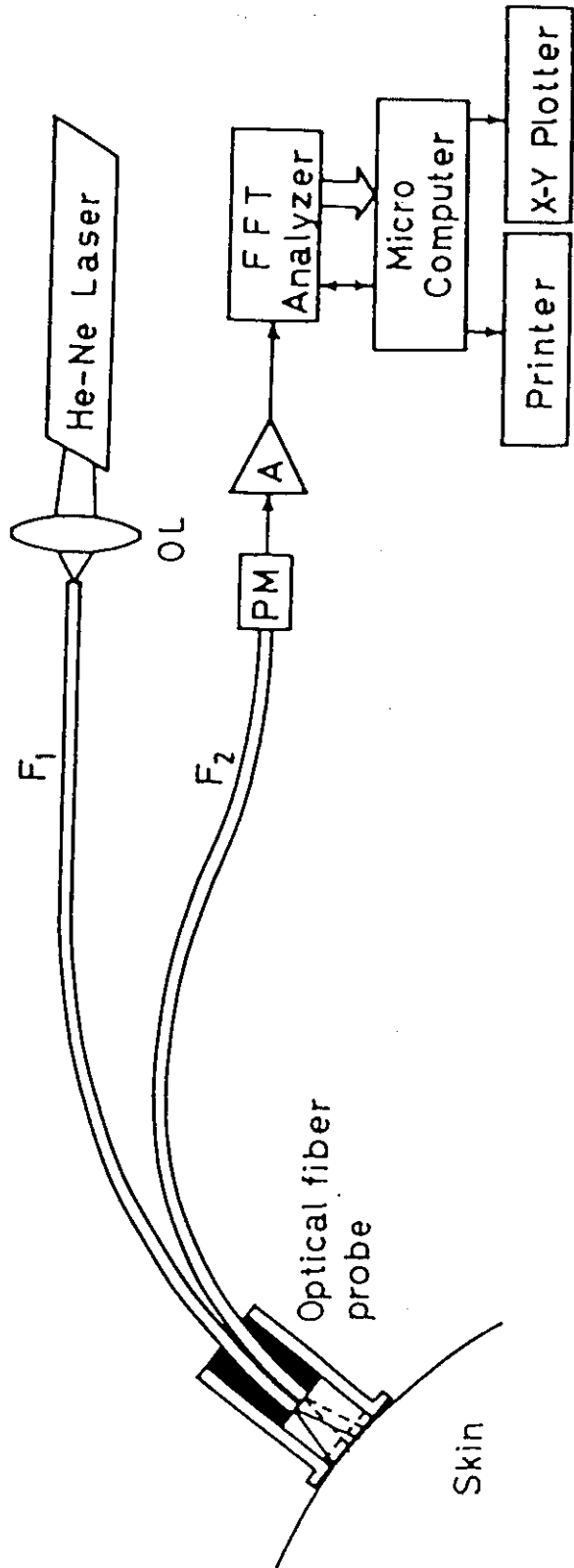
(c)

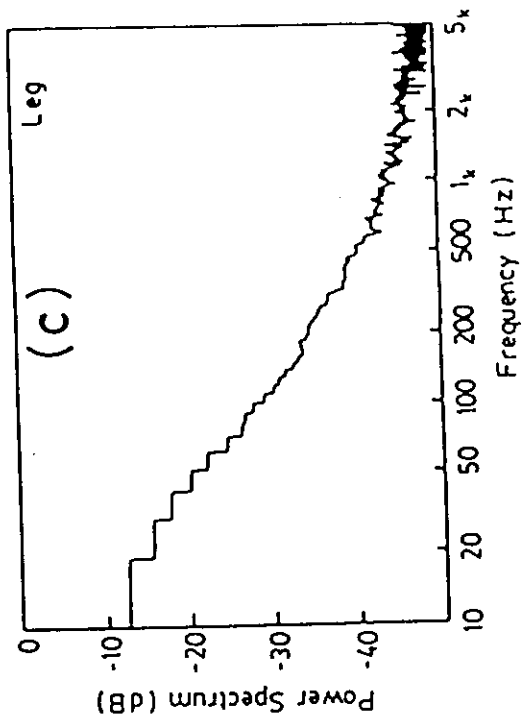
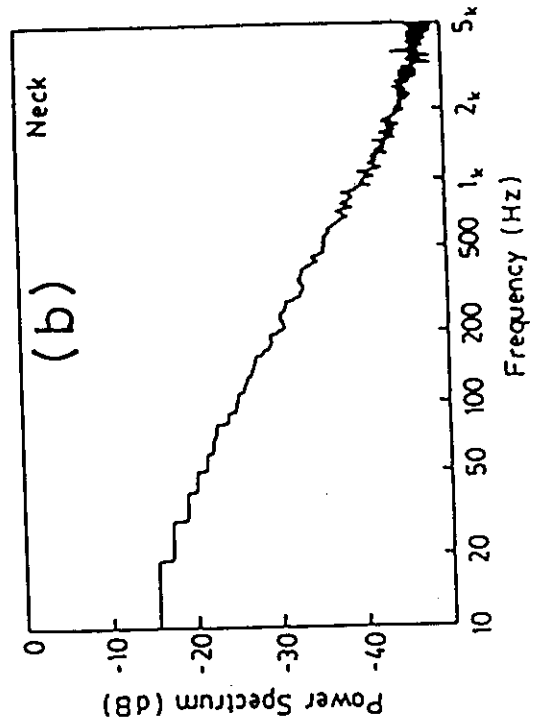
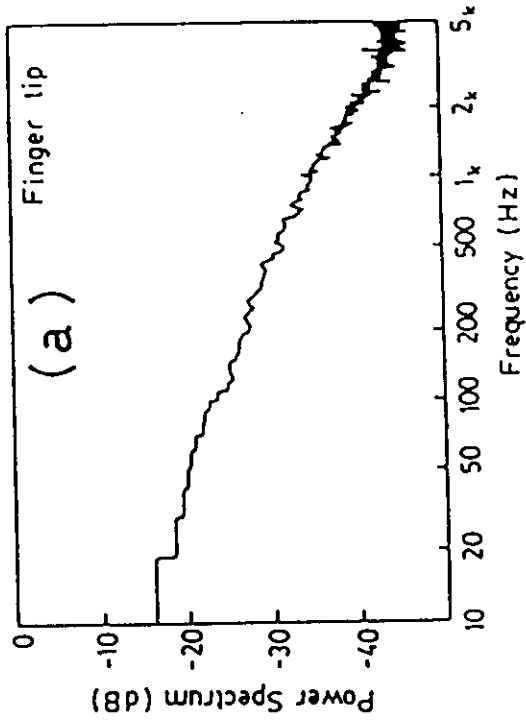


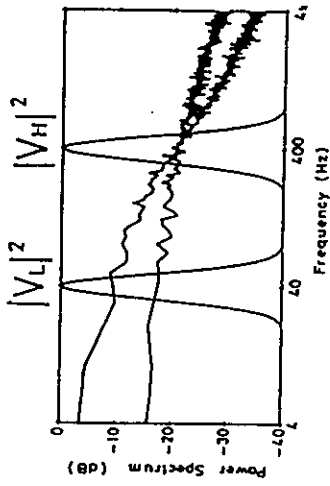
(b)



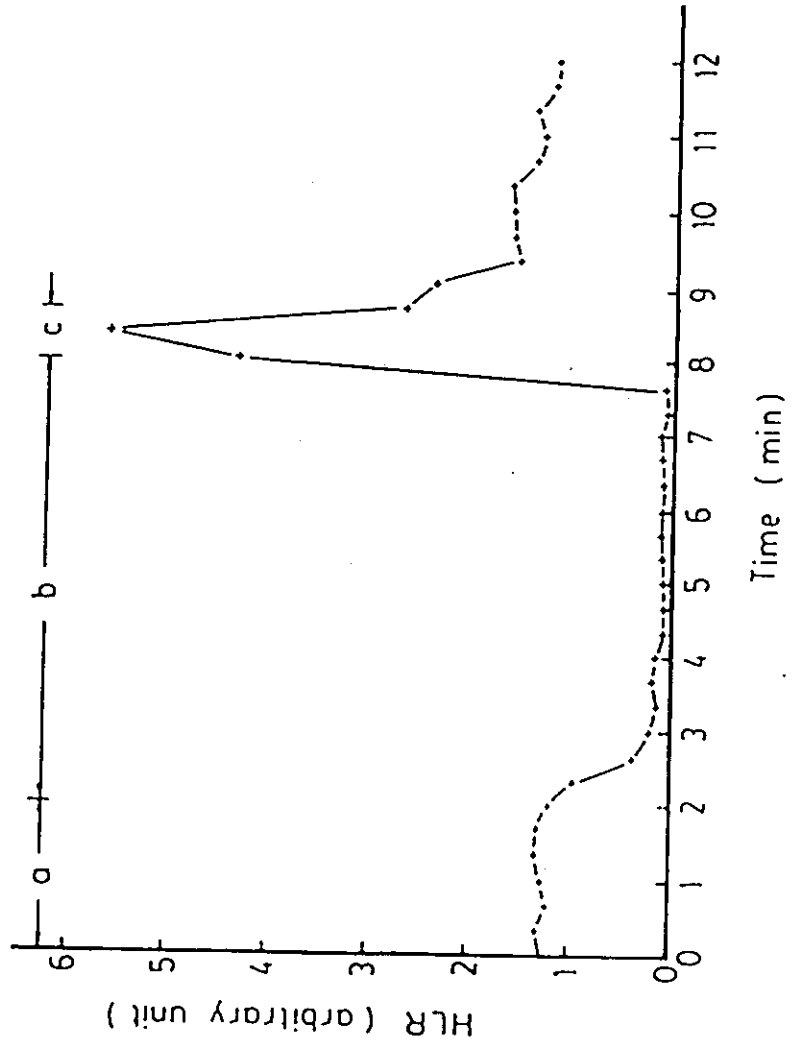
(a)



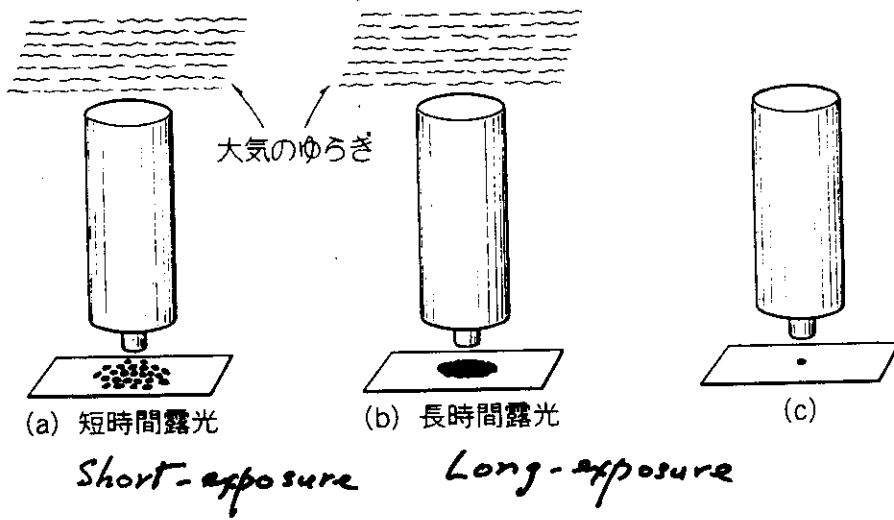




$$HLR = \frac{|V_H|^2}{|V_L|^2}$$



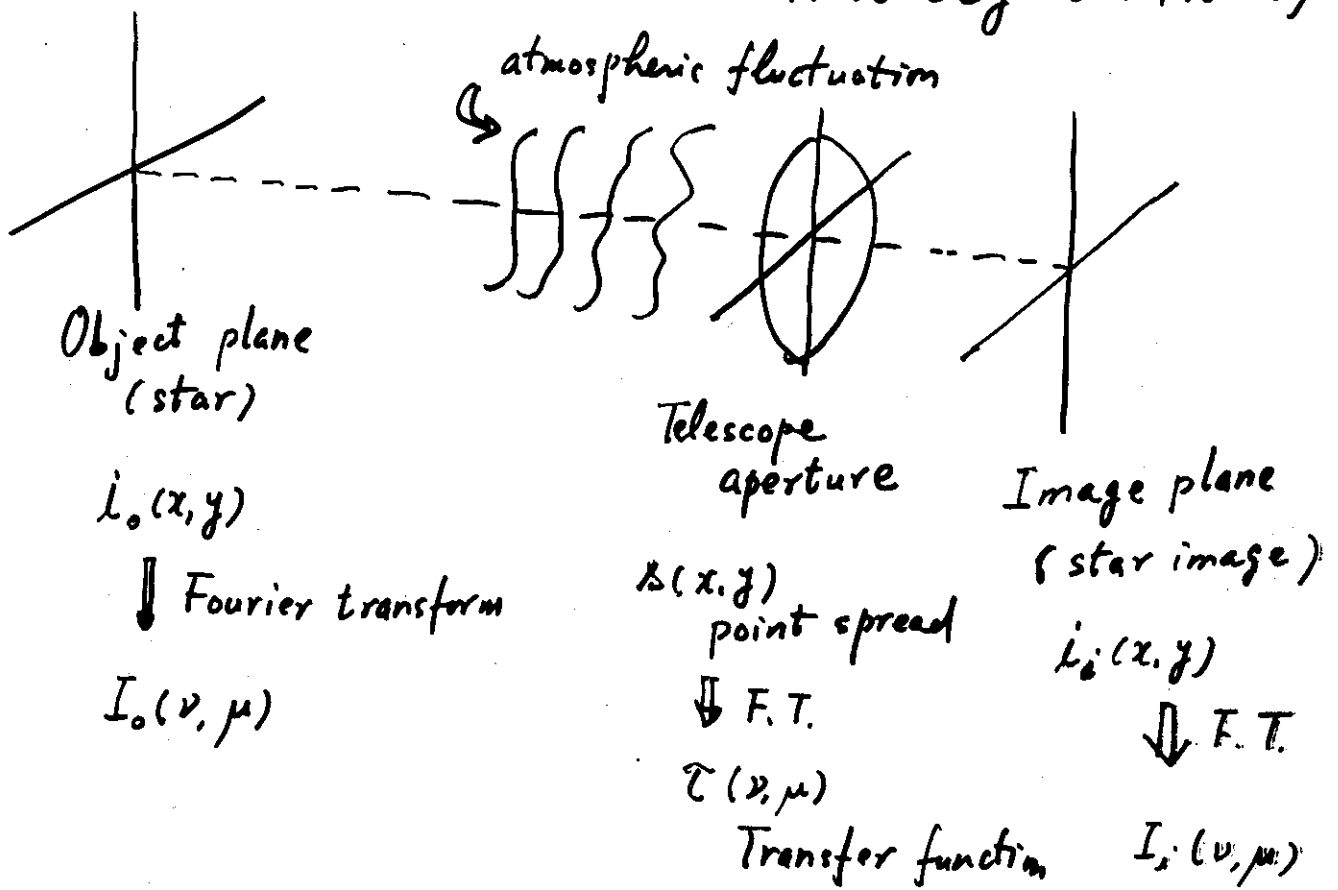
Atmospheric turbulence



Star speckles taken with the short time exposure

Stellar Speckle Interferometry

A. Labeyrie (France)



One step many photos $I_i(x, y) = I_0(x, y) \otimes \Delta(x, y)$: \otimes convolution integral
 \Downarrow Fourier transform (optically)

Second step superposition

$$I_i(\nu, \mu) = I_0(\nu, \mu) \tau(\nu, \mu)$$

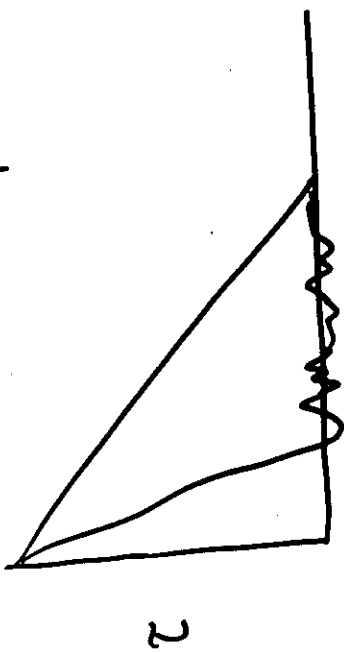
\Downarrow intensity recording

$$|I_i(\nu, \mu)|^2 = |I_0(\nu, \mu)|^2 \cdot |\tau(\nu, \mu)|^2$$

\Downarrow averaging

$$\langle |I_i(\nu, \mu)|^2 \rangle = \langle |\tau(\nu, \mu)|^2 \rangle |I_0(\nu, \mu)|^2$$

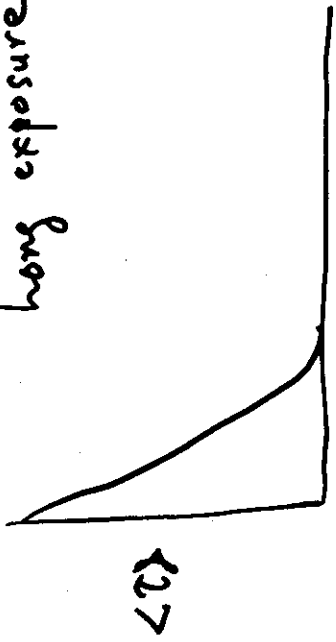
Short exposure



2

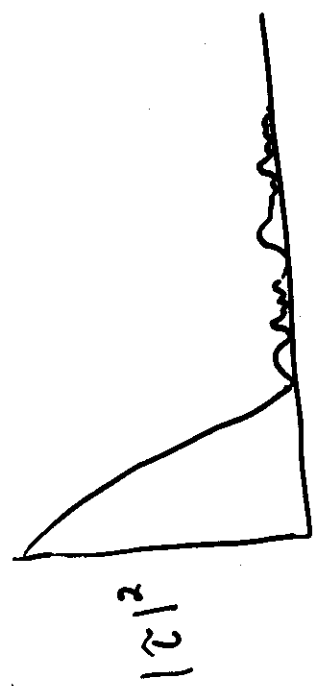
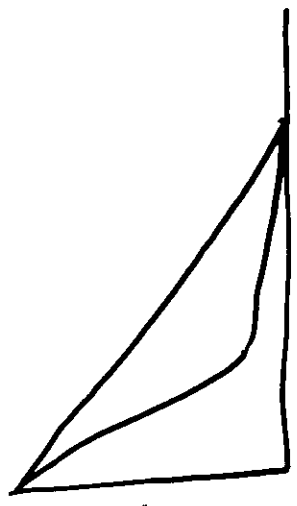
Stellar speckle interferometry

long exposure



$\langle 2 \rangle$

$\langle 2 \rangle$



$2^{1/2}$