



INTERNATIONAL ATOMIC ENERGY AGENCY
UNITED NATIONS EDUCATIONAL, SCIENTIFIC AND CULTURAL ORGANIZATION



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SMR/115 - 53

ICTP, Trieste, March '85
lecture #1

SPIN-POLARIZED HYDROGEN

(by D. Kleppner - M.I.T.)

- or, a study in laser-free atomic physics
(and, hopefully, molecule-free atomic physics)

"To understand hydrogen
is to understand everything"
attribute to Victor F. Weisskopff
by Gerhardt Herzberg

"I never said it -
but I wish I had!"
- V. F. Weisskopff

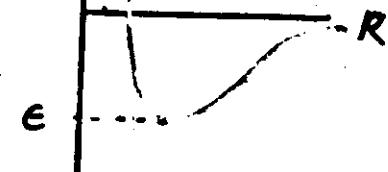
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A "Thumbnail" History of Hydrogen

- Paracelsus ~ 1510 probably saw H₂.
- Cavendish ~ 1774 isolated & identified H₂.
- Lavoisier ~ 1776 H₂ + O → H₂O: gave name "hydrogen".
- Fraunhofer ~ 1818 solar absorption - Q lines = H_α.
- Angstrom ~ 1860's lab. measurement of H spectrum.
- Huggins ~ 1870's - action. " " " "
- Balmer ~ 1883 $\frac{1}{\lambda} = R \left(\frac{1}{2^2} - \frac{1}{n^2} \right)$ n = 3, 4, 5
- Bohr ~ 1912 H theory" \Rightarrow birth of quant. mech.
- Davis ~ 1926 relativistic e theory
- Rabi ~ 1946 hyperfine structure - problems?
- Lamb ~ 1948 Lamb shift \Rightarrow Q.E.D.
- Purcell & Folsom, van Hulst ~ 1950 21 cm line in space.
- Metyger & Hoagland ~ 1965 recomb. lines n = 101 \rightarrow ∞ .
- 1970's - 80's: Lamb shift measured - proton structure limit; 1s-2s 2 photon spectroscopy, saturation spectroscopy of Balmer series, QED of 1-g(-2) "hydrogen", etc...
- 1980 - Spin-polarized H created by Silverman & Walraven, at M.I.T. [WHY?]

Solid, Liquid or Gas?

The state of matter is determined by a competition between energy order due to attractive potential E, and disorder due to kinetic energy E \sim kT.



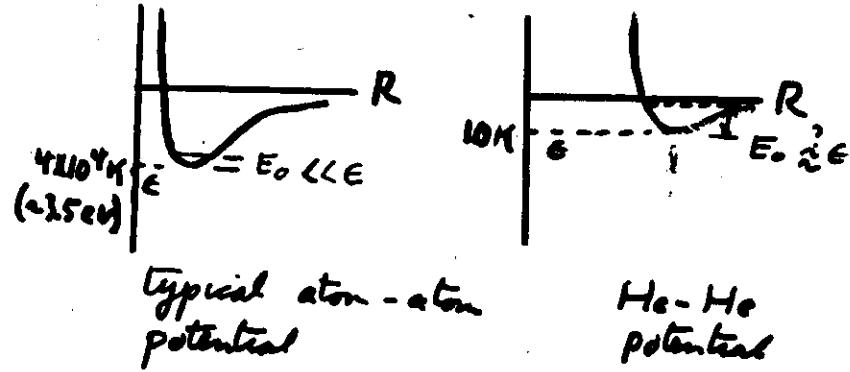
E \gg E kinetic energy wins: gas

E \ll E potential energy wins: solid

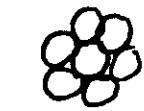
E \approx E intermediate state: liquid

E \sim kT. As T \rightarrow 0, all matter becomes solid, with one (or two) exceptions.

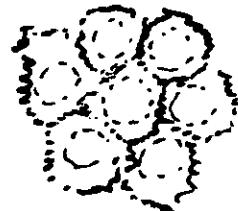
Exception - He remains liquid as $T \rightarrow 0$.



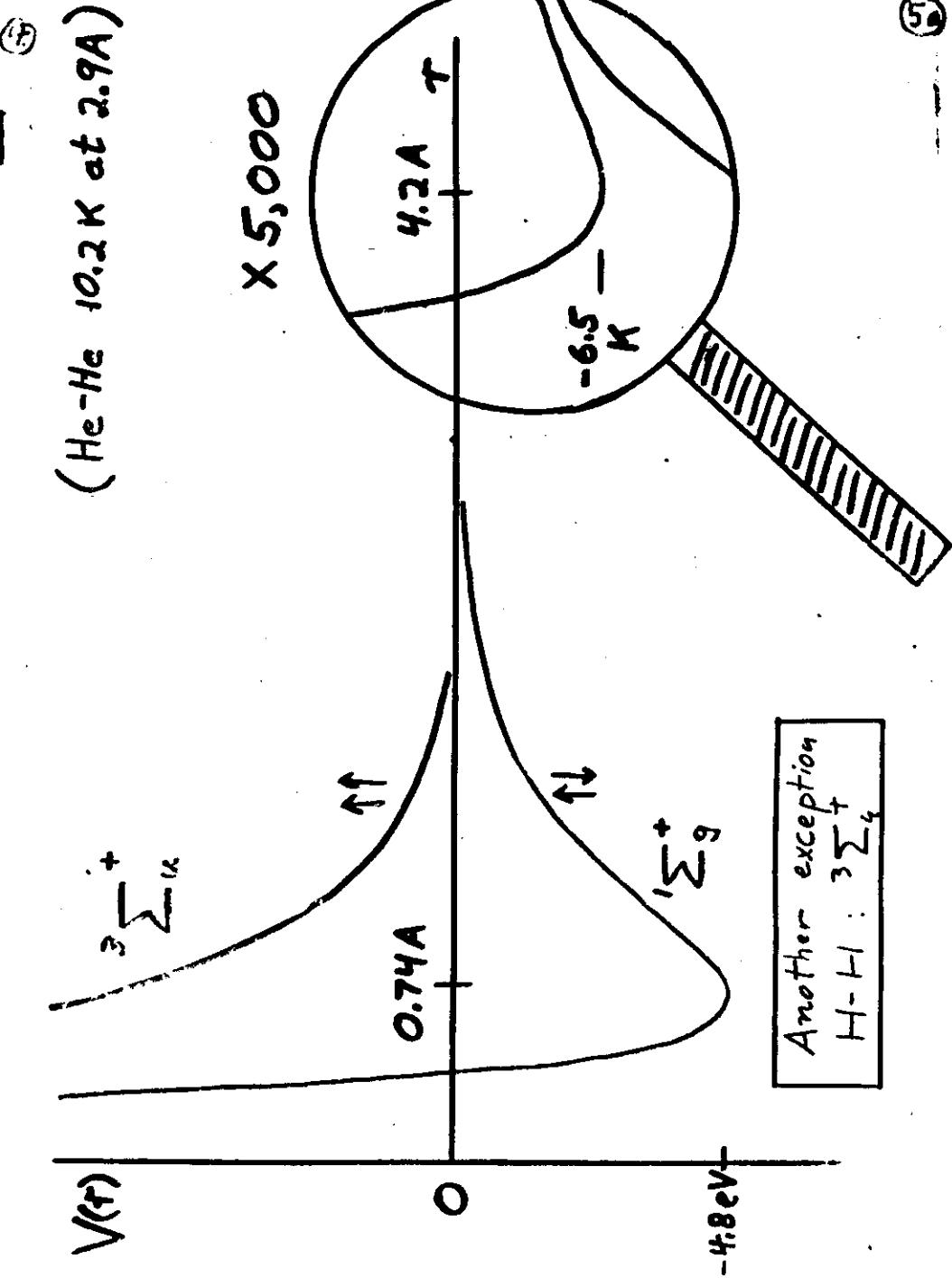
In He, the 0-point energy E_0 is so large that as $T \rightarrow 0$, $E \rightarrow E_0 \sim E$. System forms a "loose" liquid, but not a solid.



E_0 neglected
no K.E.
 $S = S_{\text{max}}$



E_0 included
much K.E.
 $S(\text{actual}) \approx \frac{1}{4} S_{\text{max}}$



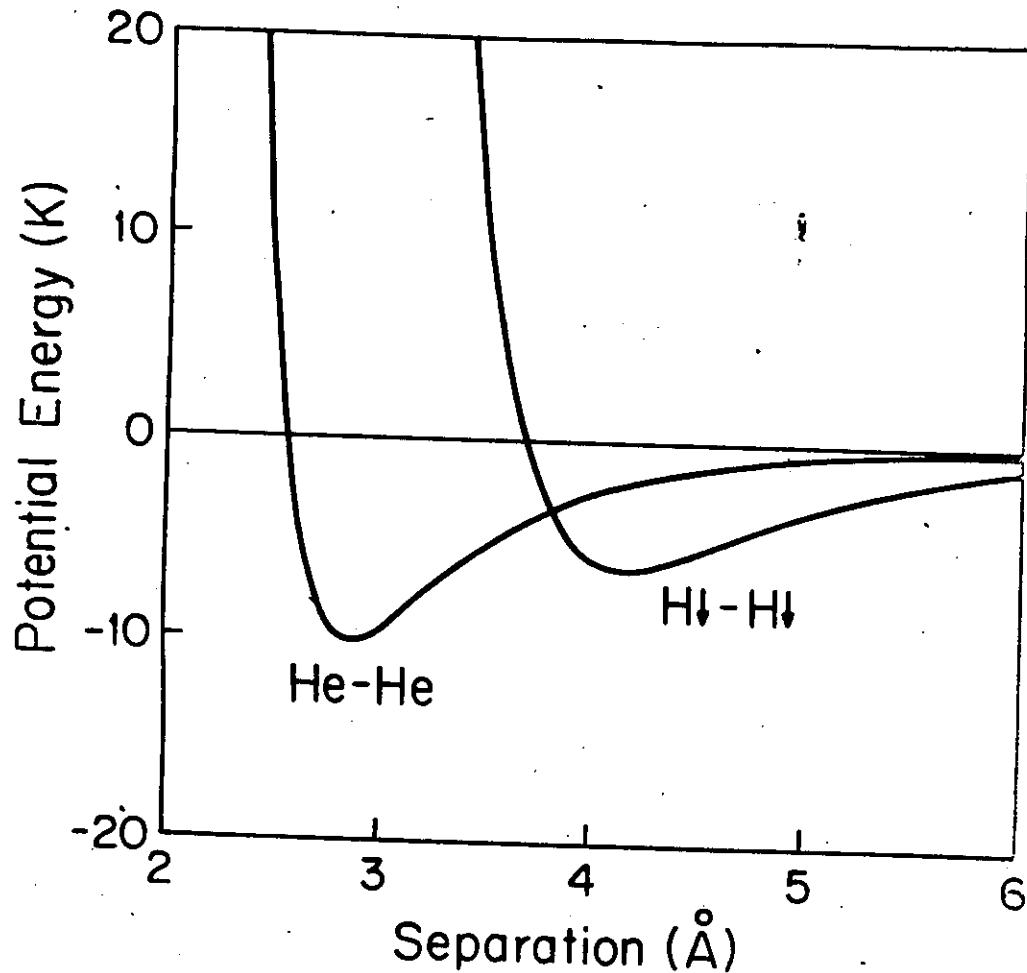
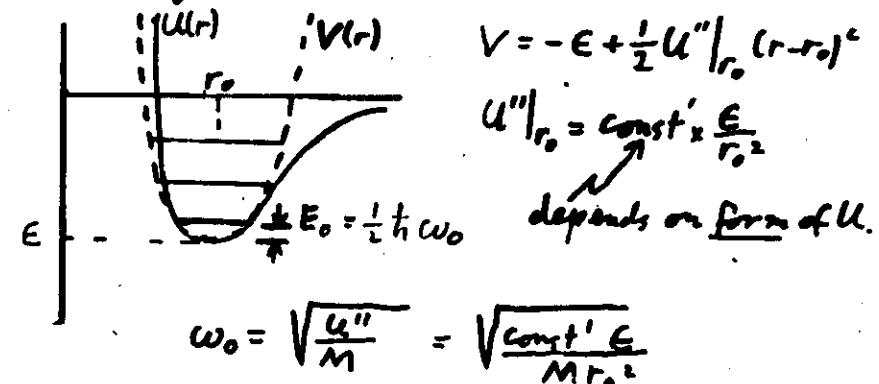


Fig. 5.1

Effect of zero-point energy
on equation of state

- can be predicted with a universal scaling law.



$$\frac{\text{Zero-point energy}}{\text{Binding energy}} = \frac{1}{2}\hbar\omega_0 = \text{const} \sqrt{\frac{\hbar^2}{Mr_0^2}}$$

$$\eta = \frac{\hbar^2}{Mr_0^2} = \text{"quantum parameter"}$$

Law of corresponding states

reduced variables: $E^* = \frac{\text{energy}}{\epsilon}$;

$$T^* = \frac{kT}{\epsilon} ; \quad r^* = \frac{\text{distance}}{r_0} ; \quad M^* = \frac{\text{mass}}{M}$$

All systems having same form for potential have same thermodynamic behavior in terms of reduced variables. Behavior depends only on η .

Why is spin-polarized hydrogen ($H\downarrow$) interesting?

- because it is predicted to remain a gas as $T \rightarrow 0$. This prediction is very firm, based on the law of corresponding states and other studies. So far, it has been confirmed to 60 mK.
- because it offers the opportunity to study matter in a new quantum regime, the regime where the translational wave packets overlap ($\lambda_{\text{deBroglie}} > \text{distance between atoms}$)
- because it has created a new technology for producing H at high density and low temperature with many potential applications to spectroscopy, atomic physics, particle & nuclear physics, etc.

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Plan of lectures

I. Introduction -

$H\downarrow$ and Bose condensation

II } strategies for making $H\downarrow$
II } what we have discovered
} how we fooled Mother Nature
} how Mother Nature fooled us
} prospects

IV applications - including spectroscopy

Reference:

Lectures on Spin-Polarized Hydrogen
T.J. Greytak & D. Kleppner

in New Trends in Atomic Physics, Vol II

G. Grymbaum & R. Stora, ed.
North Holland, 1984

Copies will be distributed in a week or two

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H₁ and the Bose-Einstein transition

A real gas, at low density, is accurately described by classical statistical mechanics. For instance, momentum described by Maxwell-Boltzmann statistics.

classical view

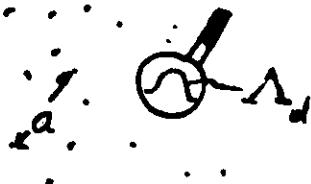


low density: $r_a \ll a$

For most atoms, $\Lambda_d = \frac{\hbar}{MV_{\text{thermal}}} \ll r_a$.

So - if gas is at low density classically, it is also at low density quantum-mechanically.

quantum view



low density: $\Lambda_d \ll a$

Quant. mech. effects important when

$$a \sim \Lambda = \frac{\hbar}{p(\text{thermal})} = \frac{\hbar}{\sqrt{2\pi k T M}}$$

density $n = \frac{1}{a^3}$ atoms cm⁻³.

$$n_0 = (\text{const}) \times \left(\frac{2\pi k T M}{\hbar^2} \right)^{3/2}$$

For Bose particles, phase transition occurs when

$$n_0 = (2.612\dots) \left(\frac{2\pi k T M}{\hbar^2} \right)^{3/2}$$

A very quick summary of Bose-Einstein transition.

Note: H obeys Bose statistics. To interchange 2 H atoms, first interchange electrons, then protons. Each interchange reverses sign of wavefunction. $(-1)(-1) = +1$. This is true so long as the atoms are not electronically excited. (In such a case, one would have 4 fermions, instead of 2 bosons.) The situation is analogous to the deuteron, which is composed of 2 fermions - a proton and a neutron - but behaves like a simple boson at low energy.

(summary of Bose transition - continue).

⑩

Mean occupation number, \bar{N}_p

- non interacting particles
- momentum p
- energy ϵ_p .

$$\bar{N}_p = \begin{cases} \frac{1}{\exp(\beta(\epsilon_p - \mu)) - 1} & \text{Bose particles } (\beta = \frac{1}{kT}) \\ \frac{1}{\exp(\beta(\epsilon_p - \mu)) + 1} & \text{Fermi particles} \\ \frac{1}{\exp(\beta(\epsilon_p - \mu))} & \text{Boltzmann law - classical} \end{cases}$$

$$\sum \bar{N}_p = N \quad (\text{total } \# \text{ particles})$$

of states: bos quantization $V=L^3$

$$p_x = \frac{2\pi\hbar}{L} j_x; \quad p_y = \frac{2\pi\hbar}{L} j_y; \quad p_z = \frac{2\pi\hbar}{L} j_z$$

j_x, j_y, j_z integers

$$\sum_{j_x, j_y, j_z} \rightarrow \left(\frac{L}{2\pi\hbar}\right)^3 \iiint dp_x dp_y dp_z$$

Change from integral over momenta to integral over energy.

$$\epsilon_p = \frac{p^2}{2M} \cdot \iiint dp_x dp_y dp_z \rightarrow 4\pi \int p^2 dp$$

summarizing:

$$\sum_{j_x, j_y, j_z} \rightarrow \frac{V}{(2\pi)^3} \left(\frac{2M}{\hbar^2}\right)^{3/2} \int \Gamma(\epsilon) d\epsilon = \int D(\epsilon) d\epsilon.$$

$$D(\epsilon) = \frac{V}{(2\pi)^3} \left(\frac{2M}{\hbar^2}\right)^{3/2} \Gamma(\epsilon) = \text{density of states in energy.}$$

Require:

$$\int \bar{N}(\epsilon) D(\epsilon) d\epsilon = N.$$

$$\bar{N}(\epsilon) = \frac{1}{\exp(\beta(\epsilon - \mu)) - 1} \cdot \mu \text{ is adjustable parameter (chemical potential).}$$

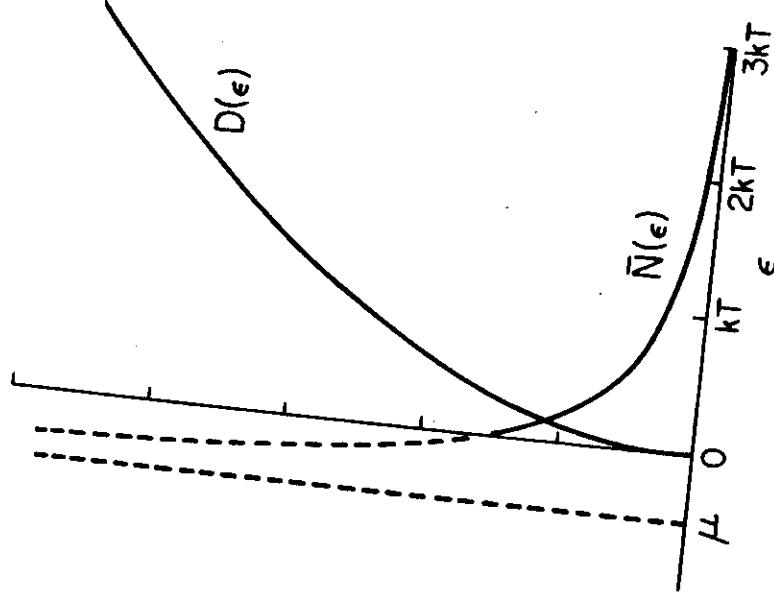
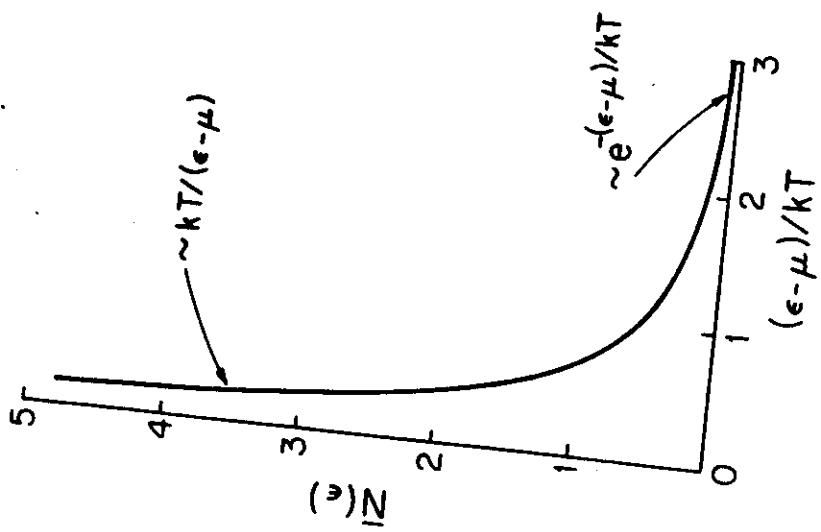


FIG. 1.1



As T decreases, μ increases toward 0
so that $N = \int N(\epsilon) D(\epsilon) d\epsilon$ is satisfied.
Maximum temperature: $\mu = 0$, $T = T_0$

$$N = V \left(\frac{M k T_0}{2 \pi \hbar^2} \right)^{3/2} \frac{2}{\sqrt{\pi}} \int_0^\infty \frac{\sqrt{x} dx}{e^x - 1}$$

$$= V \frac{1}{A(T_0)^3} \times 2.612 \dots$$

$$n = \left(\frac{M k T_0}{2 \pi \hbar^2} \right)^{3/2} \times 2.612 \dots$$

What happens if $T < T_0$?

Finite number of particles drop into the 0-energy state.

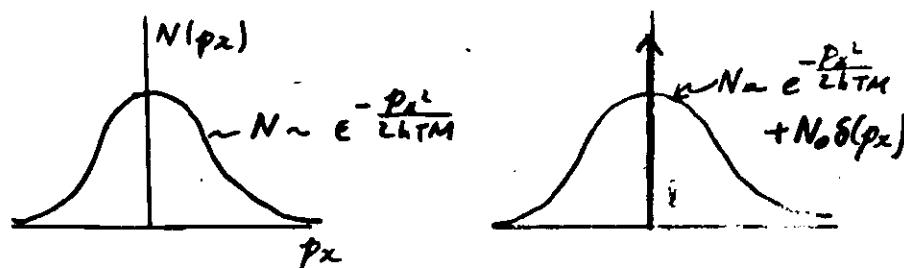
(This state was discarded in converting sum over states into integral over energy.)

Can show that the number of particles in 0-energy state is

$$\bar{N}_0 = N (1 - T/T_0)^{3/2}$$

$T \leq T_0$
 $T > T_0$

Evidence for Bose-Einstein condensation



Most systems: solid before T_0 is reached.

Liquid ${}^4\text{He}$: $n = 2.20 \times 10^{22} \text{ atoms cm}^{-3}$
(liquid density)

$$T_0 = 3.15 \text{ K}$$

Close to superfluid temp. 2.24
Is superfluid ${}^4\text{He}$ a

Bose condensate?

$$T_0 = \frac{2\pi\hbar^2}{kM} \left(\frac{n}{2.612}\right)^{2/3}$$

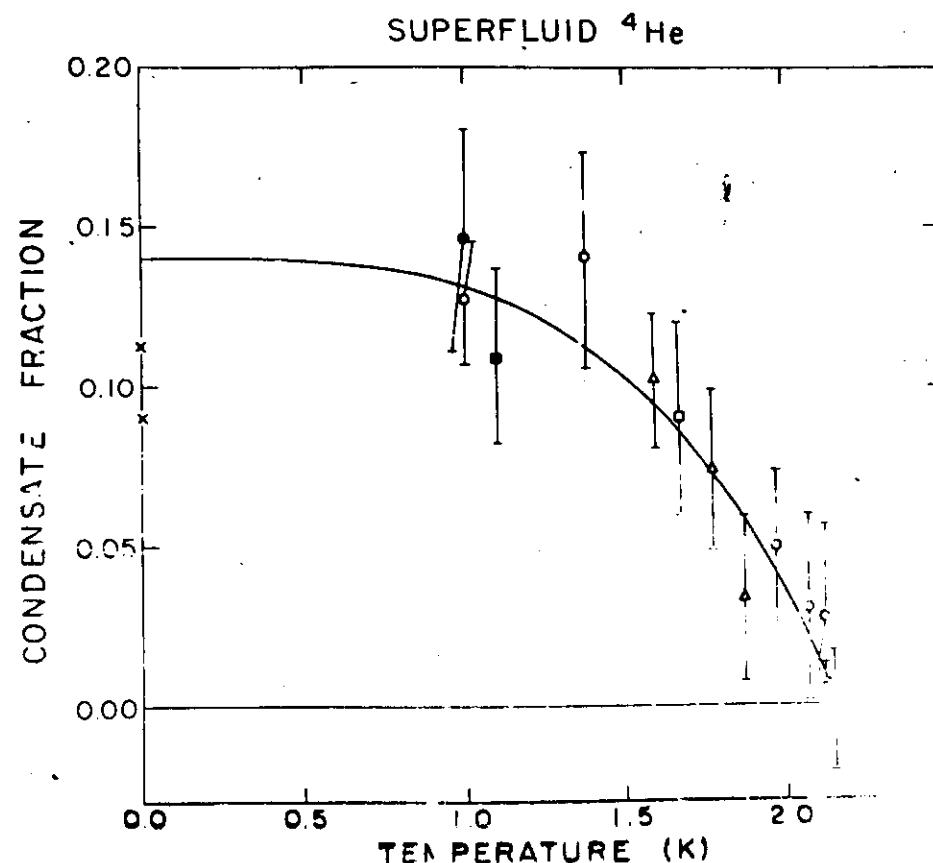
require minimum M !

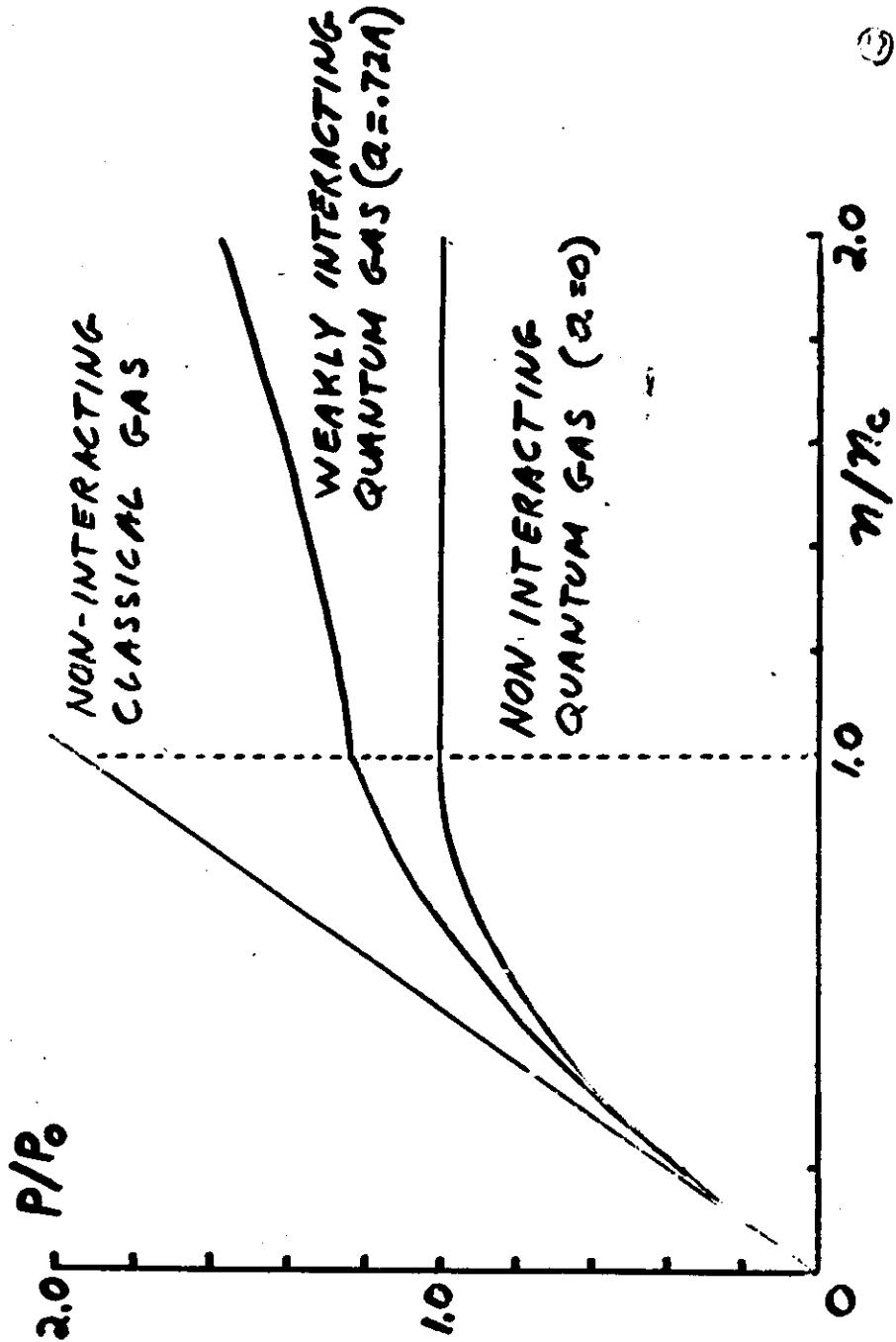
Hydrogen: for $T_0 = 0.3 \text{ K}$, $n = 6 \times 10^{19} \text{ cm}^{-3}$?

Can B-E condensation be achieved?

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Quest to observe B-E transition in H₂

- Required:
- atomic hydrogen
 - at temperature < 1K
(previous lowest temperature,
78K)
 - at density > 10^{19} cm^{-3}
(previous maximum density for
H (without H₂) $\sim 10^{12} \text{ cm}^{-3}$)
 - highly spin polarized
(a few percent of wrong
spin state would rapidly
destroy system)
 - stabilized against any sort
of spin relaxation

