united nations educational scientific and cultural organization

international atom energy agen abdus salam
international centre for theoretical physics

School on "Exploring the Atmosphere by Remote Sensing Techniques" 18 October - 5 November 1999

1151-8

The Earth Observation Programme of the European Space Agency: "The Living Planet Programme"

C. Readings
Head - Earth Sciences Division
ESTEC
Noordwijk
The Netherlands

Please note: These are preliminary notes intended for internal distribution only.



The Earth Observation Programme of the European Space Agency

The Living Planet Programme

C.J. READINGS

Head, Earth Sciences Division Estec, Noordwijk, The Netherlands

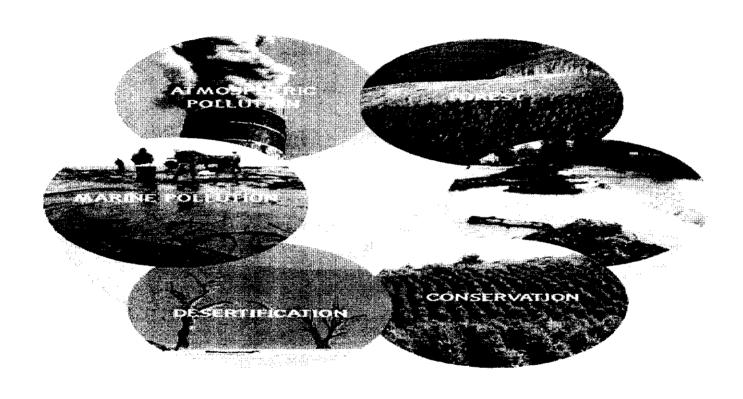


ESA's Living Planet Programme General Background

- Increasing public concern over the Earth, its environment and mankind's impact on it. Both regional and global:
 - Global concerns over Climate Warming, Ozone Depletion, Tropospheric Pollution, El Niño, etc.
 - Regional concerns over Sea Level Change, Fires in South East Asia, Floods in Europe, Droughts in the USA, etc.
- Establishment of the Intergovernmental Panel on Climate Change (IPCC) to advise on state of knowledge reports highlight general lack of knowledge in many crucial areas.
- International research initiatives include the establishment of the World Climate Research Programme (WCRP) and the International Biosphere/Geosphere Programme (IG/BP).



ESA' Living Planet Programme General Background (2)



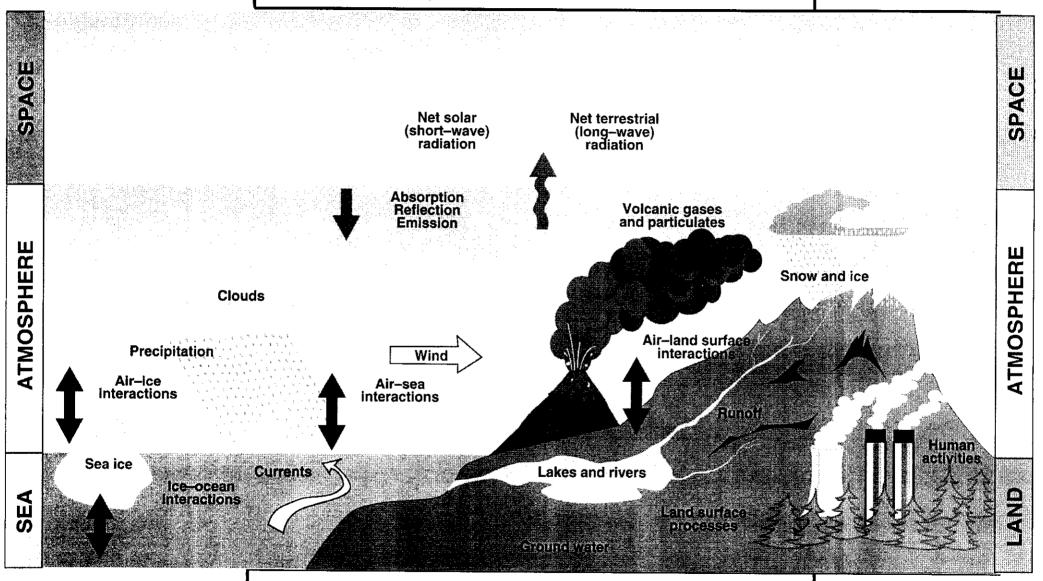


ESA' Living Planet Programme The Underlying Rationale

- The need to address public concerns about the Earth, its environment and mankind's impact on it.
- The Earth is a complex (and evolving) system which is not properly understood.
- Data required to improve knowledge of the processes involved, to develop and validate models.
- Space has a role to play in the helping to ensure the provision of the requisite data.



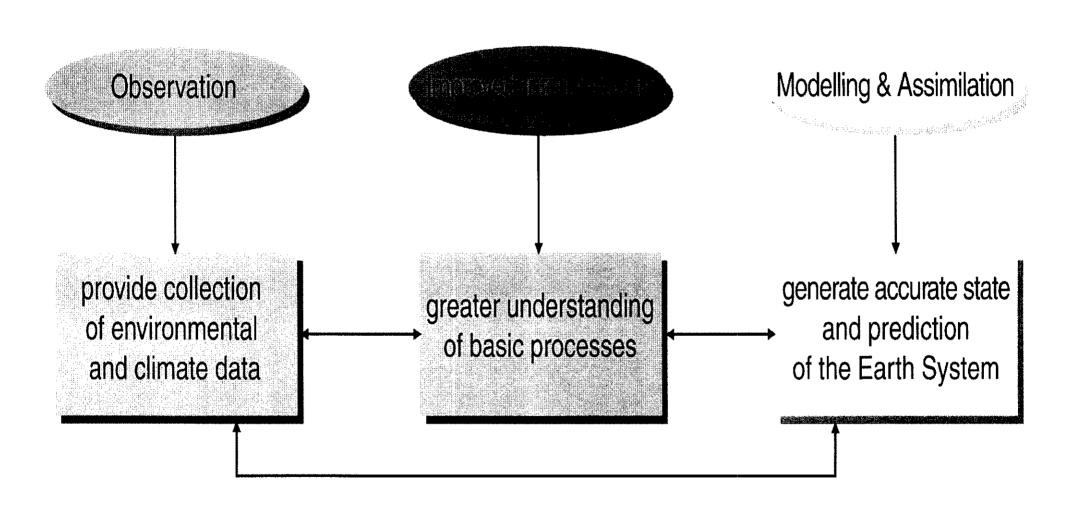
ESA' Living Planet Programme The Earth System



Exploring the Atmosphere by Remote Sensing Techniques 22 October 1999, Trieste, Italy



ESA' Living Planet Programme The Earth System Model - key modules





ESA' Living Planet Programme Earth System Models (1)

Understanding of the Earth will improve by the development and elaboration of global Earth System models which describe:

- the evolution of the state and composition of the atmosphere
- the physical state of the ocean and cryosphere
- the physical state of the top few metres of soil and dynamical interactions with the Earth's interior
- the physical state of terrestrial vegetation
- the key bio-geochemical cycles which in turn require the representation of terrestrial and ocean biota



ESA' Living Planet Programme Natural Hazards and Applications

ATMOSPHERE	Stratosphere	OZONE MONITORING UV RADIATION PREDICTION				
ATMO	Troposphere	GREENHOUSE EFFECT, POLLUTION, HURRICANES, WATER SUPPLY, WEATHER PREDICTION, AIRCRAFT NAVIGATION & ROUTING				
EARTH SURFACE		OCEAN	LAND HYDROSPHERE	LAND GEO-BIOSPHERE		
		SEA LEVEL RISE, SEA STATE, STORM SURGES, TSUNAMIS,SHIP ROUTING & NAVIGATION, FISHERIES,WATER QUALITY, POLLUTION, RESOURCES	SEA STATE, STORM SURGES, JNAMIS,SHIP ROUTING & NAVIGATION, HERIES,WATER QUALITY, WATER POLLUTION			
	INTERIOR	VOLCANIC ACTIVITIES, EARTHQUAKES, NATURAL RESOURCES, NAVIGATION, CARTOGRAPHY				



ESA' Living Planet Programme The Hierachical Set of Earth System Models

ATMOSPHERE	Stratosphere	DYNAMICS - RADIATION - CHEMISTRY				
	Troposphere	DYNAMICS - RADIATION - CLOUDS - ENERGY & WATER CYCLE - CARBON CYCLE				
		OCEAN	LAND HYDROSPHERE	LAND GEO-BIOSPHERE		
ЕАВТН	SURFACE	AIR-SEA INTERACTION, OCEAN CIRCULATION, OCEAN BIOLOGY, COASTAL ZONES, SEA ICE ENERGY TRANSPORT	HYDROLOGY, SOIL MOISTURE	LAND SURFACE PROCESSES, LAND BIOLOGY, ECOSYSTEMS, SNOW & LAND ICE		
	MERIOR	GEOID	GEODESY	GRAVITY & MAGNETIC FIELDS		



ESA' Living Planet Programme Earth System Models (2)

The formulation of these models is difficult, especially as nonlinear nature of the processes involved. Requires simultaneous progress in three different areas:

- Area 1 to identify and increase understanding of the various processes involved to the point where they can be represented in models.
- Area 2 to extend the existing hierarchy of Earth System models to include these processes.
- Area 3 to ensure the provision of the relevant data for use in these models to help address the issues highlighted above.

The provision of observations from space is of fundamental importance.



ESA' Living Planet Programme The Overall Approach: The Four Themes

Four major interdisciplinary themes identified, each of which may encompass phenomena in several of the regimes:

- Theme 1 the Earth interior
- Theme 2 -the physical climate system
- Theme 3 the geosphere-biosphere
- <u>Theme 4</u> the anthropogenic influences on the atmospheric and marine environment

Between them these four Themes span the full Earth System and recognise the need for the detailed treatment of interactions between the various regimes.



ESA' Living Planet Programme Links between the Themes and Elements of the Earth System

THEMES					
COMPONENTS		THEME 1	THEME 2	THEME 3	THEME 4
		EARTH INTERIOR	PHYSICAL CLIMATE	GEO- BIOSPHERE	ANTHROPOGENIC IMPACT
ATMOSPHERE	STRATOSPHERE		✓	✓	✓
	TROPOSPHERE		✓	✓	✓
EARTH SURFACE	LAND	✓	✓	✓	✓
	OCEAN	✓	✓	✓	✓
EARTH INTERIOR		✓	✓	✓	



The ESA Living Planet Programme The Earth Explorer Missions (1)

- Means of addressing objectives (see ESA SP-1227)
- Regular flight opportunities funded under the Earth Observation Envelope Programme
- Objectives of Earth Explorer Missions research and development focussing on specific topics/techniques
- Two complementary types of Earth Explorer missions, namely:

Earth Explorer Core Missions - larger research/demonstration missions led by **ESA**.

<u>Earth Explorer Opportunity Missions</u> - smaller research and demonstration missions not necessarily ESA led.

• Complemented by Earth Watch - thematic pre-operational missions focussing on specific emerging Earth Observation application areas



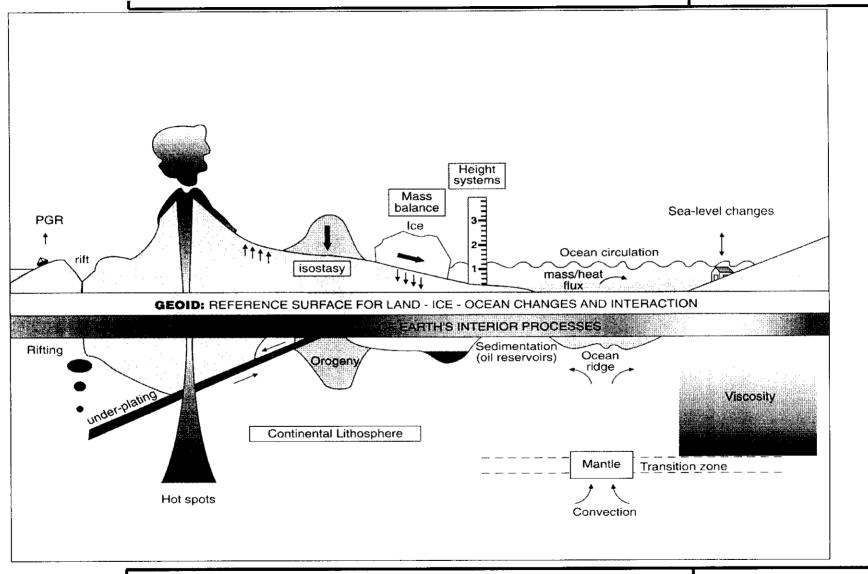
ESA' Living Planet Programme The Four Themes

Theme 1 - Earth Interior

- Marine Geoid and its Impact on Ocean Circulation
- Gravity Field and Earth Interior Processes
- Magnetic Field and Earth Interior Processes
- Geodesy

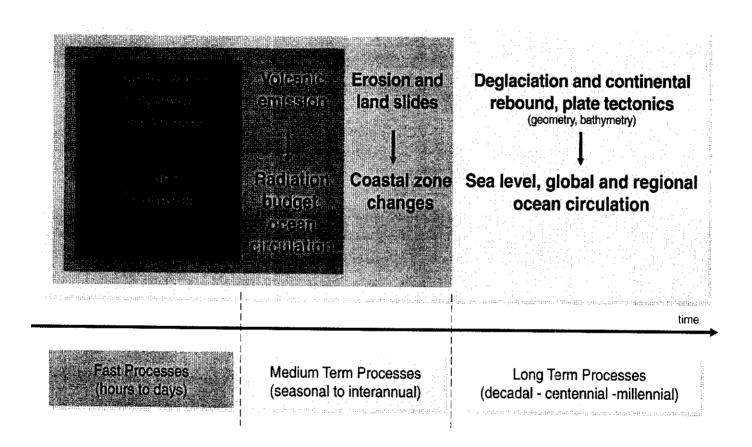


ESA' Living Planet Programme The Earth as a Dynamic System



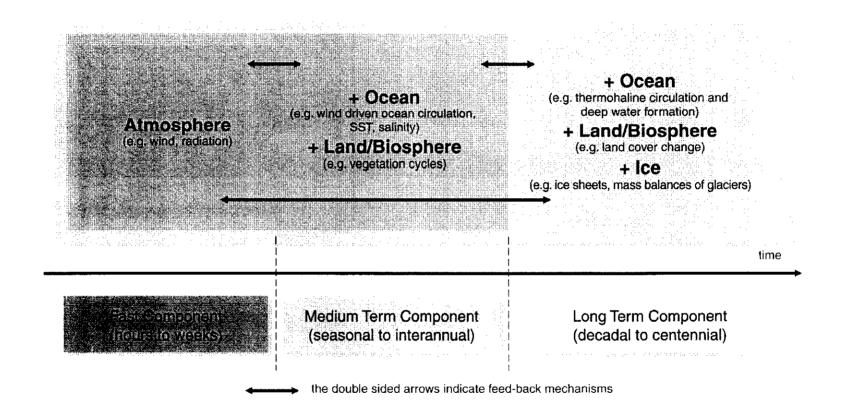


ESA' Living Planet Programme Theme 1 - Earth Interior





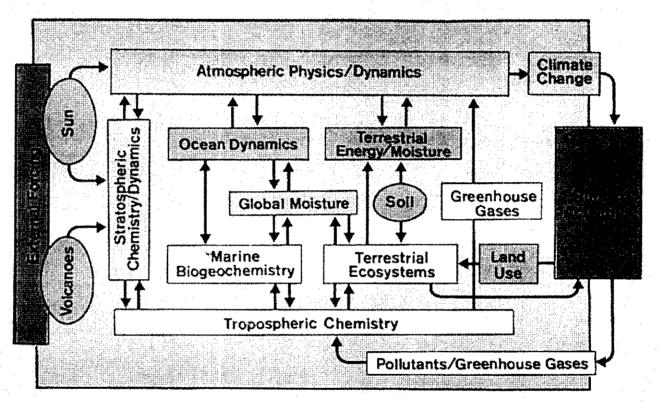
ESA' Living Planet Programme Theme 2 - Physical Climate





ESA' Living Planet Programme The Physical and Biophysical Systems

Physical Climate System



Biogeochemical Cycles



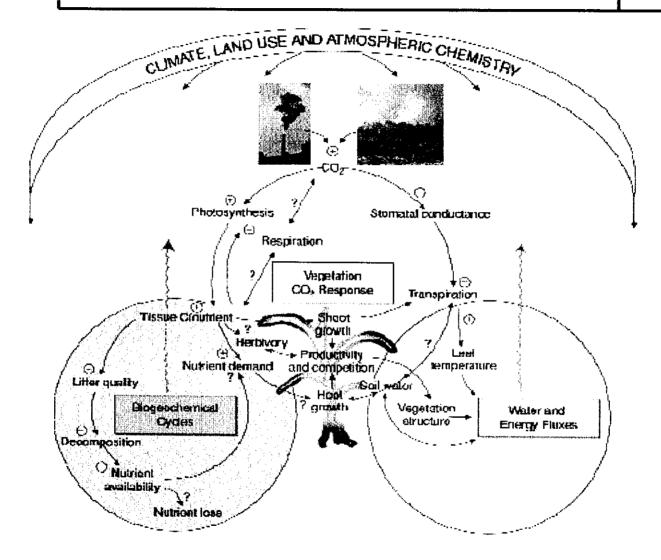
ESA' Living Planet Programme The Four Themes

Theme 3 - Geosphere/Biosphere

- The Energy and Water Cycles
- The Bio-Chemical Cycles
- The Productivity of Different Ecosystems



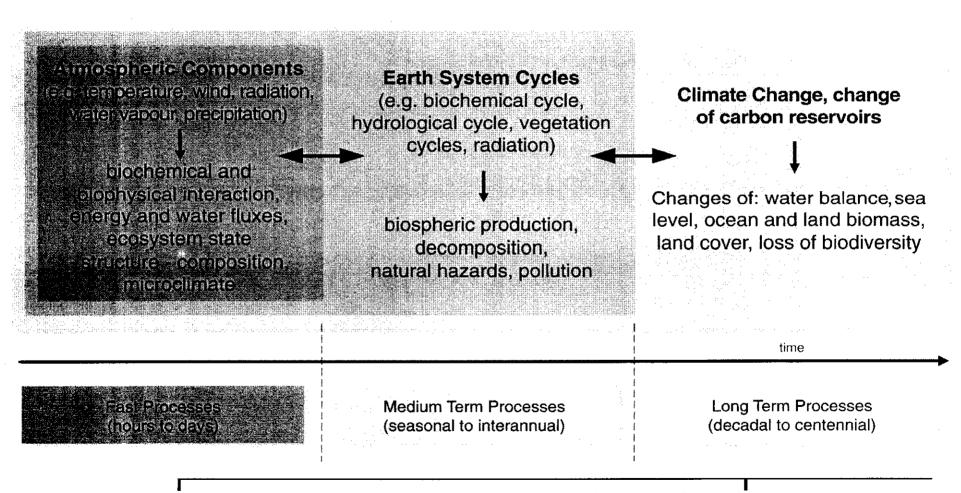
ESA' Living Planet Programme The Earth Model - Links





ESA' Living Planet Programme Theme 3 - Geosphere/Biosphere

Examples for geo-biospheric interaction: The links between the components





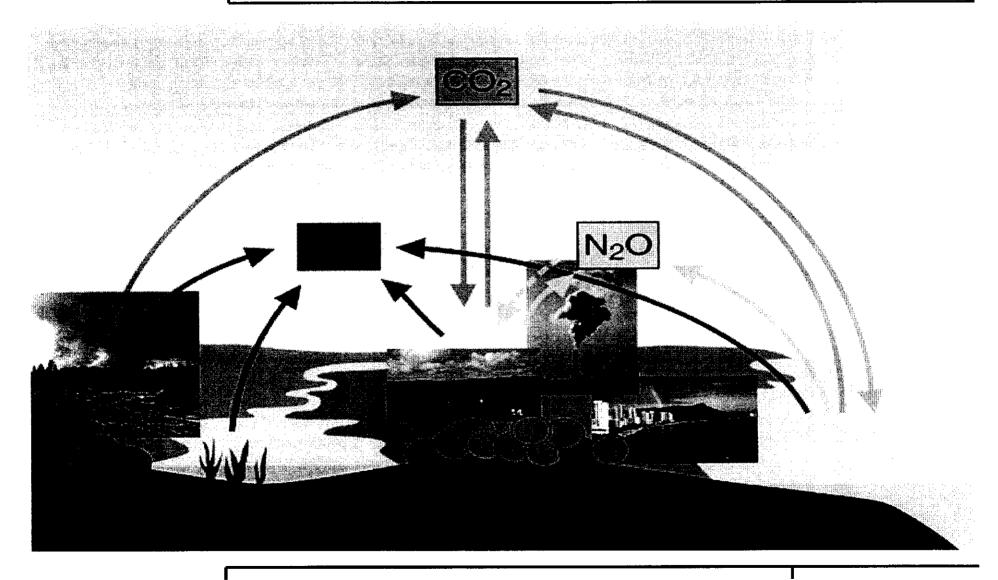
ESA' Living Planet Programme The Four Themes

Theme 4 - Atmosphere and Marine Environment: Anthropogenic Impact

- Changes in Atmospheric Composition Induced by Human Activity
- Chemical Processes in the Stratosphere and Upper Troposphere
- Marine Polution

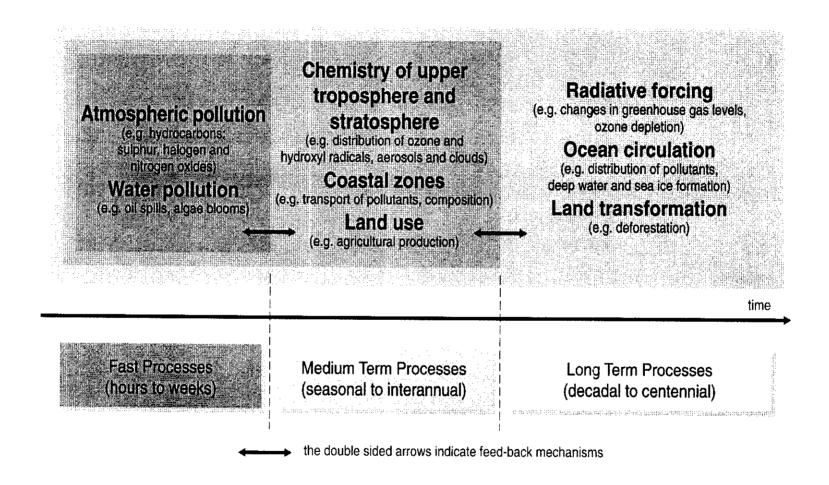


ESA' Living Planet Programme Chemical Interactions at the Surface





ESA' Living Planet Programme Theme 4 - Anthropogenic Impacts





The ESA Living Planet Programme The Earth Explorer Missions (1)

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 Complemented by Earth Watch Missions - thematic pre-operational missions focussing on specific emerging Earth Observation application areas



The ESA Living Planet Programme The Earth Explorer Missions (2)

Earth Explorer Core Missions:

<u>Granada I</u> (May 1996) - presentation of 9 missions for Phase A Study: 4 selected for Phase A study

Granada II (October 1999) - presentation of 4 missions for Phase B Study and implementation: 2 to be selected

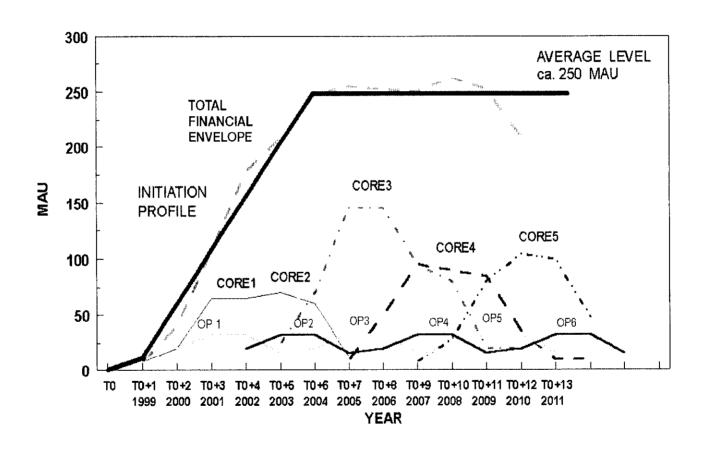
Earth Explorer Opportunity Missions:

<u>First Call for Proposals</u> (July 1998) - 27 proposals submitted: 5 missions recommended in order of priority by the Earth Sciences Advisory Committee for implementation (April 1999): recommendations endorsed by the Programme Board for Earth Observation

- Cyclic selection process; further Calls and consultation meetings planned
- Information on http://www.estec.esa.nl/explorer/



The ESA Living Planet Programme The Earth Explorer Missions (3)





The ESA Living Planet Programme The Earth Explorer Core Missions

Four candidate missions have been subjected to Phase A studies:

- Gravity Field and Steady-State Ocean Circulation Mission
- Earth Radiation Mission
- Land-Surface Processes and Interactions Mission
- Atmospheric Dynamics Mission

All four will be presented during the Granada II meeting: following Granada II two will be selected for implementation

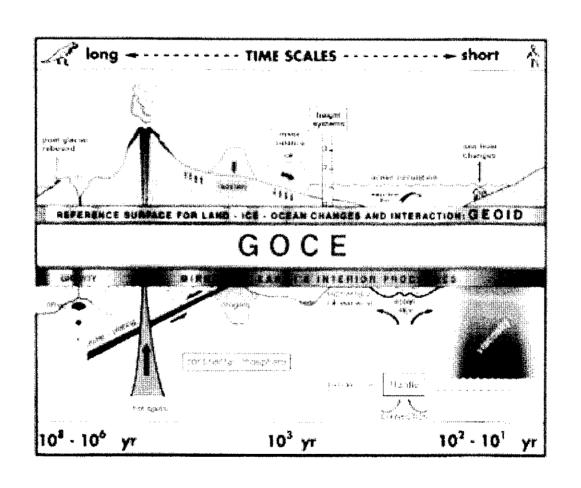


The Earth Explorer Core Missions Gravity Field and Steady-State Ocean Circulation Mission (1)

Ice Sheet Melting

Earthquakes & volcanic activities

Global unification of height system



Thermohaline ocean circulation & heat transport

Sea level change



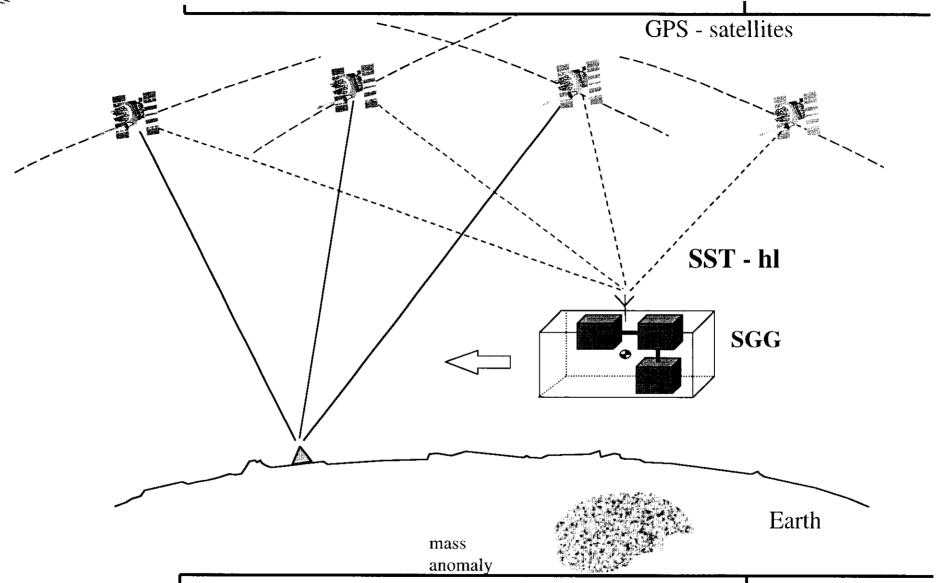
The Earth Explorer Core Missions Gravity Field and Steady-State Ocean Circulation Mission (2)

The Mission objectives are:

- determine the Earth's gravity field with an accuracy of $1 \text{ mgal } (1 \text{ mgal} = 10^{-5} \text{ m/s}^2)$
- determine the geoid (= equipotential surface for a hypothetical ocean at rest) with an accuracy of 1 cm
- achieve this at length scales down to L = 100 km (degree and order 200)

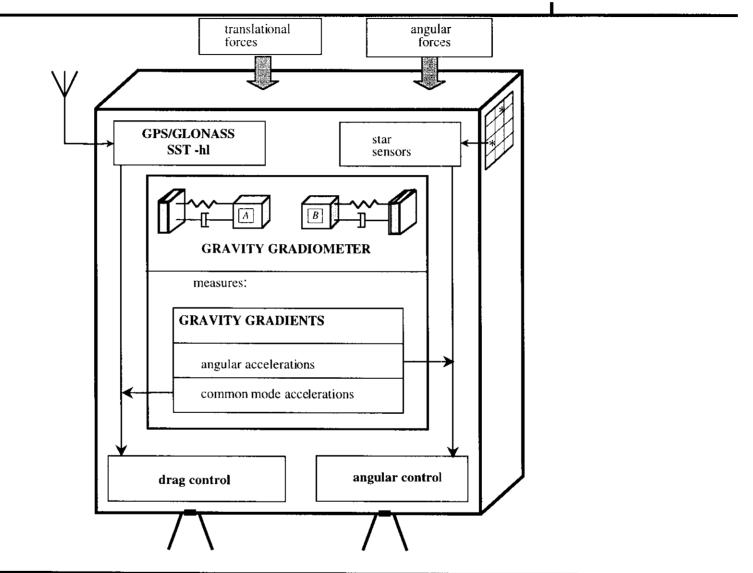


The Earth Explorer Core Missions Gravity Field and Steady-State Ocean Circulation Mission (3)





The Earth Explorer Core Missions Gravity Field and Steady-State Ocean Circulation Mission (4)





The Earth Explorer Core Missions The Earth Radiation Mission (1)

Scientific Objectives

The ERM has been specifically defined with the scientific objective of determining world wide the vertical profiles of cloud and aerosol field characteristics to provide basic (and essential) input data for numerical modelling and studies (on a global scale) of:

- the divergence of radiative energy
- aerosol-cloud-radiation interaction
- the vertical distribution of water and ice and their transport by clouds
- the vertical cloud field overlap and cloud-precipitation interactions

These objectives are the result of joint Japanese - European discussions



The Earth Explorer Core Missions The Earth Radiation Mission (2)

Instrument Performance

- ·Lidar
 - very sensitive for clouds and aerosols but attenuated by clouds
 - gate-by-gate correction of attenuation unstable
- •Radar
 - good for ice clouds potential synergy with lidar?
- •Multi-Spectral Imager
 - can provide horizontal structure
- •Broad Band Radiometer
 - can provide TOA radiance constraint

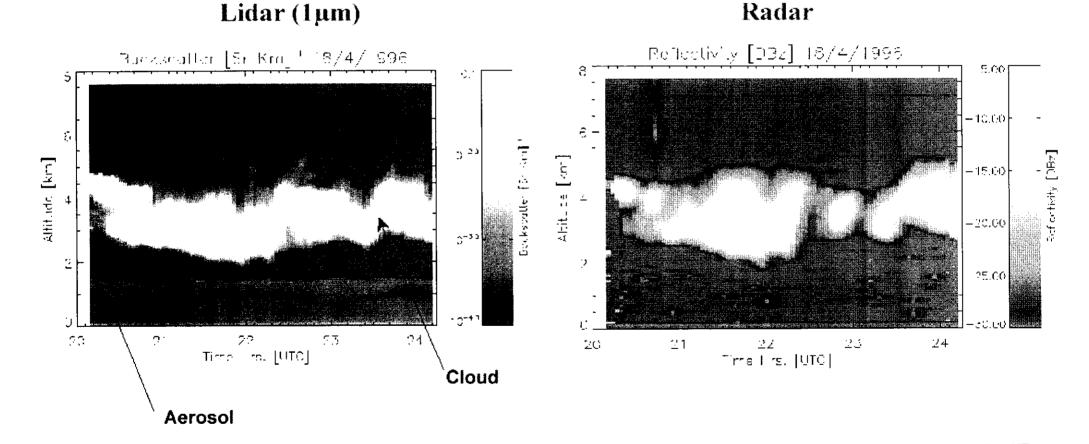


The Earth Explorer Core Missions The Earth Radiation Mission (3)

Radar/Lidar Synergy in Ice Clouds

Radar/Lidar Retrieval - Example (I)

CLARA Campaign



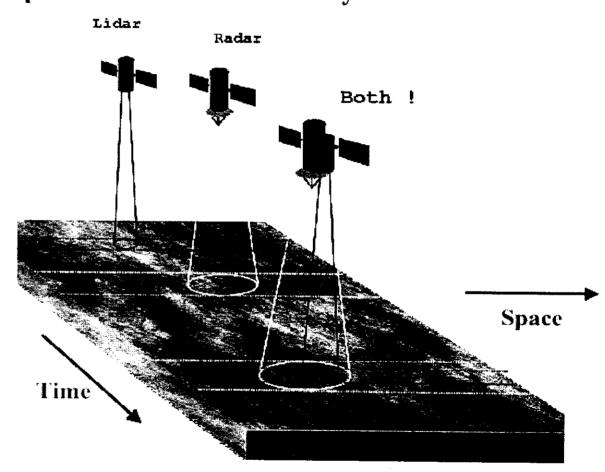


The Earth Explorer Core Missions The Earth Radiation Mission (4)

Radar/Lidar Synergy

Effect of Separation - Cloud Variability

- Ice Clouds have lots of structure:
- For a co-location error of just 2 km the ratio of the radar reflectivity to the lidar backscatter leads to unacceptable errors.
- Time separation below around 10 seconds has no significant impact on radar/lidar synergy.

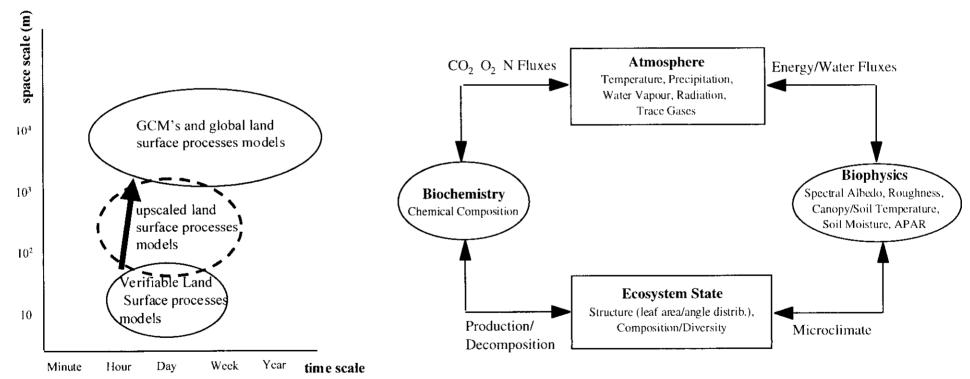




The Earth Explorer Core Missions The Land Surface Processes and Interactions Mission (1)

The prime objectives of this mission is to:

- increase the understanding of biophysical processes and land/atmosphere interactions at the local scale
- advance understanding of these interactions, on a global scale.

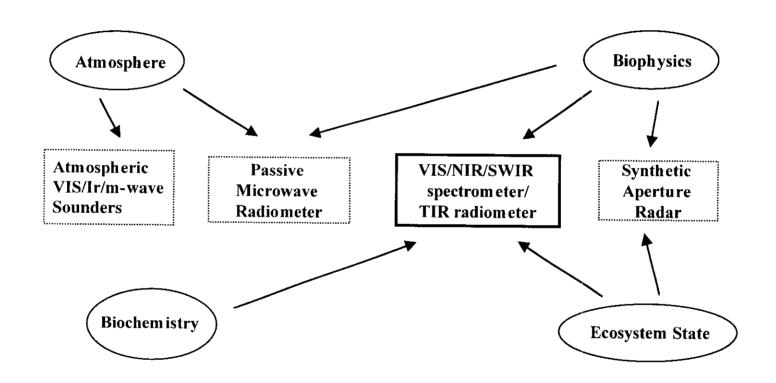


Scaling context of Land Surface Processes

Biospheric Interactions and involved Process Variabes

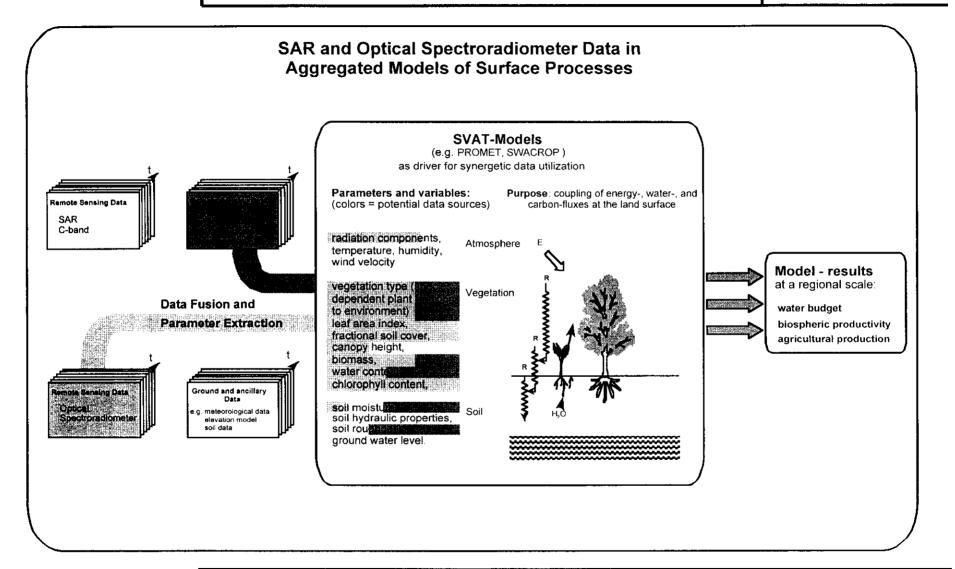
The Earth Explorer Core Missions The Land Surface Processes and Interactions Mission (2)

Links between Process Variables and Candidate Mission Elements





The Earth Explorer Core Missions The Land Surface Processes and Interactions Mission





The Earth Explorer Core Missions The Land Surface Processes and Interactions Mission (3)

The preliminary ranges for the PRISM performance parameters are:

Spectral range:

450 nm - 2350 nm at < 15 nm spectral sampling interval (spectral

absorption features of vegetation and soils)

 $8 - 8.5 \mu m$, 8.6 - 9.1 surface temperature (vegetation and soils)

Spatial sampling:

around 50 m

Swath width:

around 50 km

Pointing:

across track depointing capability of about \pm 30° for global access and

BRDF measurements

along track pointing ±70, ±60, ±45, 0 degrees for BRDF measurements

Data transmission:

Possibility of selection of bands tbd.(40 to 60 bands per scene), also

for data reduction



The Earth Explorer Core Missions The Atmospheric Dynamics Mission (1)

ADM objectives

- Improve understanding of atmospheric dynamics and global atmospheric transport
- Improve understanding of global cycling of energy, water, aerosols, chemicals
- How is this achieved?
 - Improved analysis of the atmospheric state to provide a complete threedimensional picture of the dynamical variables
- What are the benefits?
 - Improved parameterisation of atmospheric processes in models
 - Advanced climate and atmospheric flow modelling
 - Better initial conditions for weather forecasting



The Earth Explorer Core Missions The Atmospheric Dynamics Mission (2)

Observational Requirements

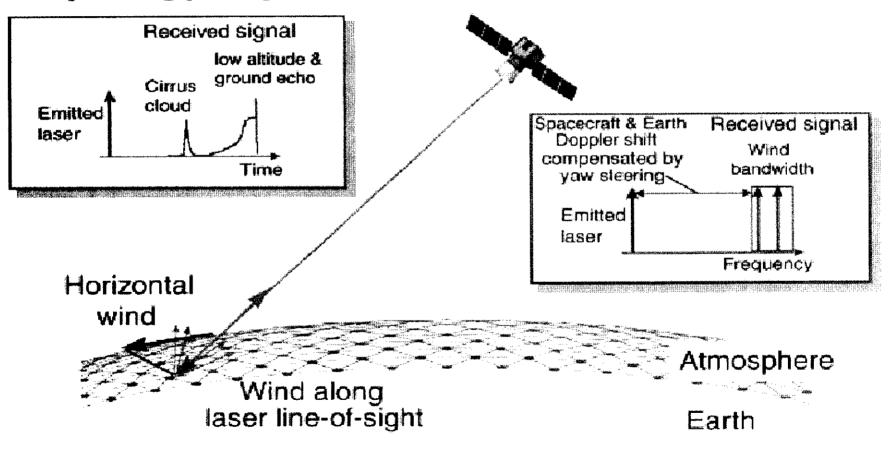
		PBL	Troposph	Siratosph	
Vertical Domain	[knj	()-2	2-16	16-20	
Vertical Resolution	[kn]	0.5	1.0	2.0	
Horizontal Domain		global			
Number of Profiles	[hour ^{:[}]	100			
Profile Separation	[km]	> 200			
Temporal Sampling	[hour]	12			
Accuracy (Component)	[ms-l]	2	2-3	3	
Horizontal Integration	[kn]	50			
Error Correlation		0.01			
Reliability	[%]	95			
Timeliness	[hour]	3			
Length of Observational Data Set	lya)	3			



The Earth Explorer Core Missions The Atmospheric Dynamics Mission (3)

Space-Borne Doppler Wind Lidar

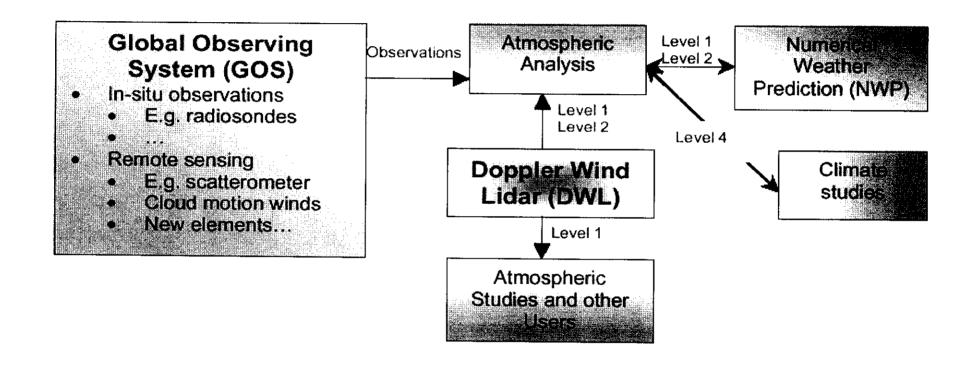
Operating principle:





The Earth Explorer Core Missions The Atmospheric Dynamics Mission (4)

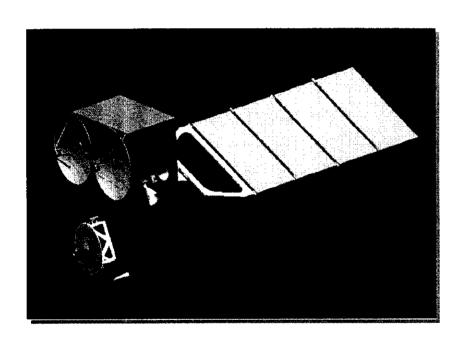
Mission Elements



The Doppler Wind Lidar is the main innovation



ESA' Living Planet Programme The Earth Explorer Opportunity Missions (1)



CRYOSAT

- For observing the ice sheets and sea ice to conduct research concerned with sea level rise and the role of fresh water in the climate system.
- Advanced microwave altimeter exploiting SAR and interferometric techniques in Ku-Band
- Selected for Phase A/B Study;



ESA' Living Planet Programme The Earth Explorer Opportunity Missions (2)

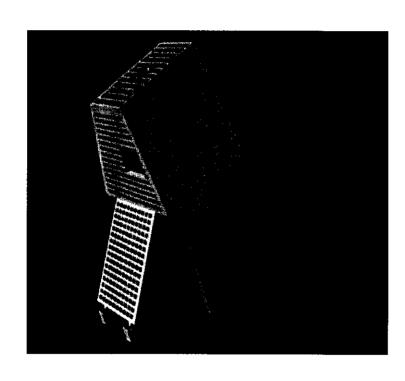


SMOS - Soil Moisture and Ocean Salinity Satellite

- To observe ocean salinity, soil moisture, sea-ice and snow cover for meteorological, climatological, hydrological, oceanographical and glaciological studies.
- A passive microwave radiometer exploiting SAR and interferometric techniques at 1.4 GHz.
- Selected for extended Phase A Study



ESA' Living Planet Programme The Earth Explorer Opportunity Missions (3)



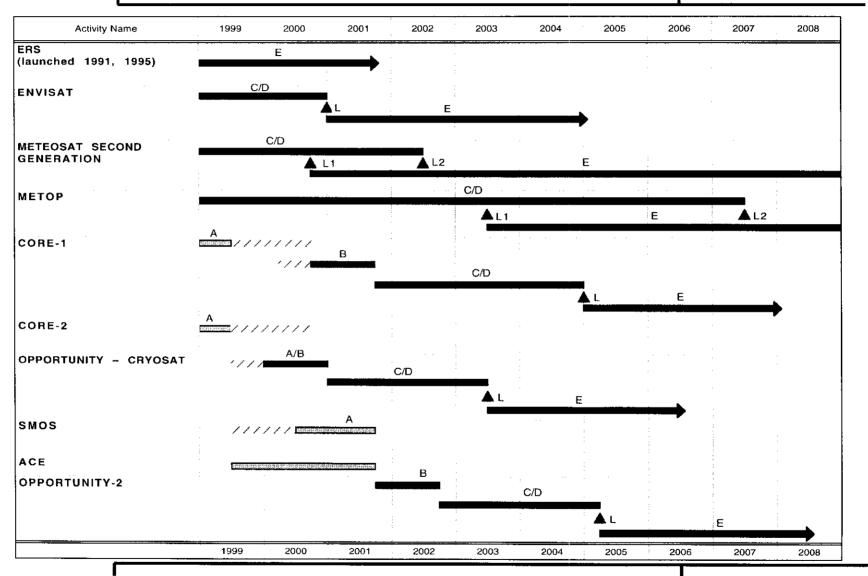
ACE - Atmospheric Climate Experiment

- For monitoring temperature and water vapour in the lower stratosphere and upper troposphere for climate change research.
- Exploits the global satellite navigation systems to measure refractive index.
- Selected as the "hot" spare;.



ESA' Living Planet Programme

Overall Schedule







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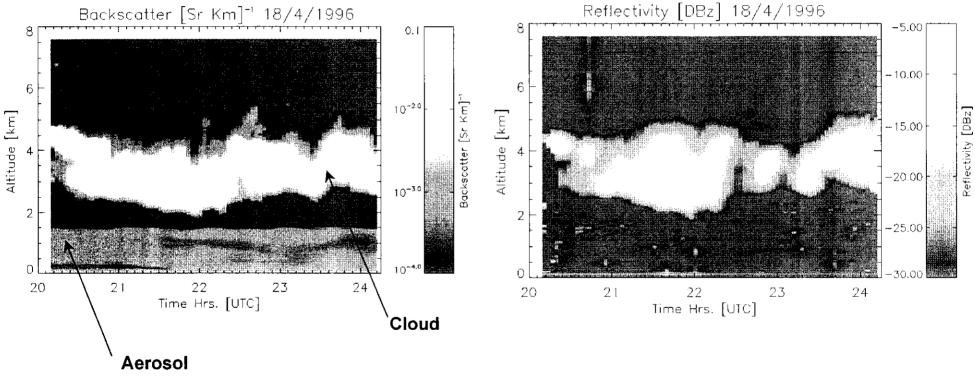
Radar/Lidar Retrieval - Example (I)

CLARA Campaign

Lidar (1µm)

- [Sr Km] - 18/4

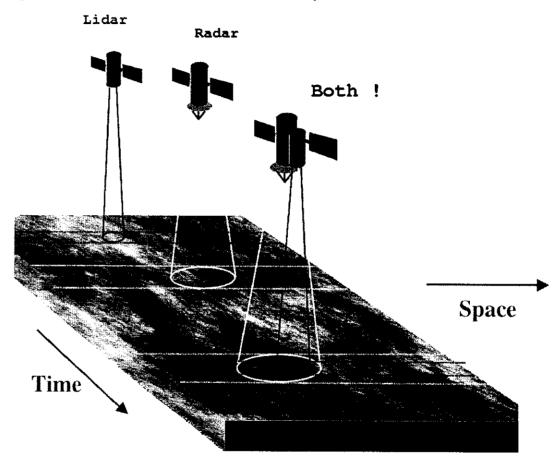
Radar





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esa Mission Requirements ADM Observational Requirements

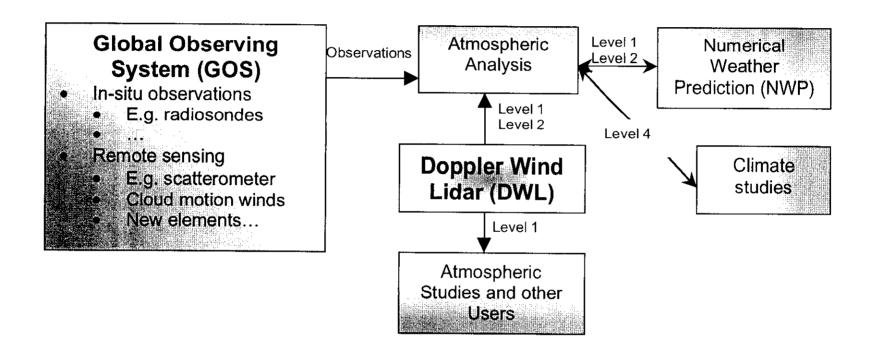


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The Four Candidate Earth Explorer Core Missions Consultative Workshop 12-14 October 1999, Granada, Spain

Mission RequirementsThe Mission Elements





The Doppler Wind Lidar is the main innovation



esa Scientific Justification: Proposed Solution A Spaceborne Doppler Wind Lidar (DWL)

ADM

Operating principle:

