



INTERNATIONAL ATOMIC ENERGY AGENCY
UNITED NATIONS EDUCATIONAL, SCIENTIFIC AND CULTURAL ORGANIZATION
INTERNATIONAL CENTRE FOR THEORETICAL PHYSICS
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H4.SMR/453-27

**TRAINING COLLEGE ON
PHYSICS AND CHARACTERIZATION
OF LASERS AND OPTICAL FIBRES**

(5 February - 2 March 1990)

**HIGH RESOLUTION OPTOGALVANIC
SPECTROSCOPY OF ATOMIC OXYGEN**

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note

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1. COLLISIONAL PHYSICS

- i) O-atom production
- ii) Anomalous Doppler broadening

2. FINE STRUCTURE MEASUREMENTS

3. ISOTOPE SHIFT ¹⁶O, ¹⁷O, ¹⁸O

4. HYPERFINE STRUCTURE ¹⁷O
DIODE LASER SPECTROSCOPY

5. PROPOSAL: LASER COOLING OF
ATOMIC OXYGEN BEAM.

LASER DIAGNOSTICS OF
RADIO-FREQUENCY OXYGEN PLASMA

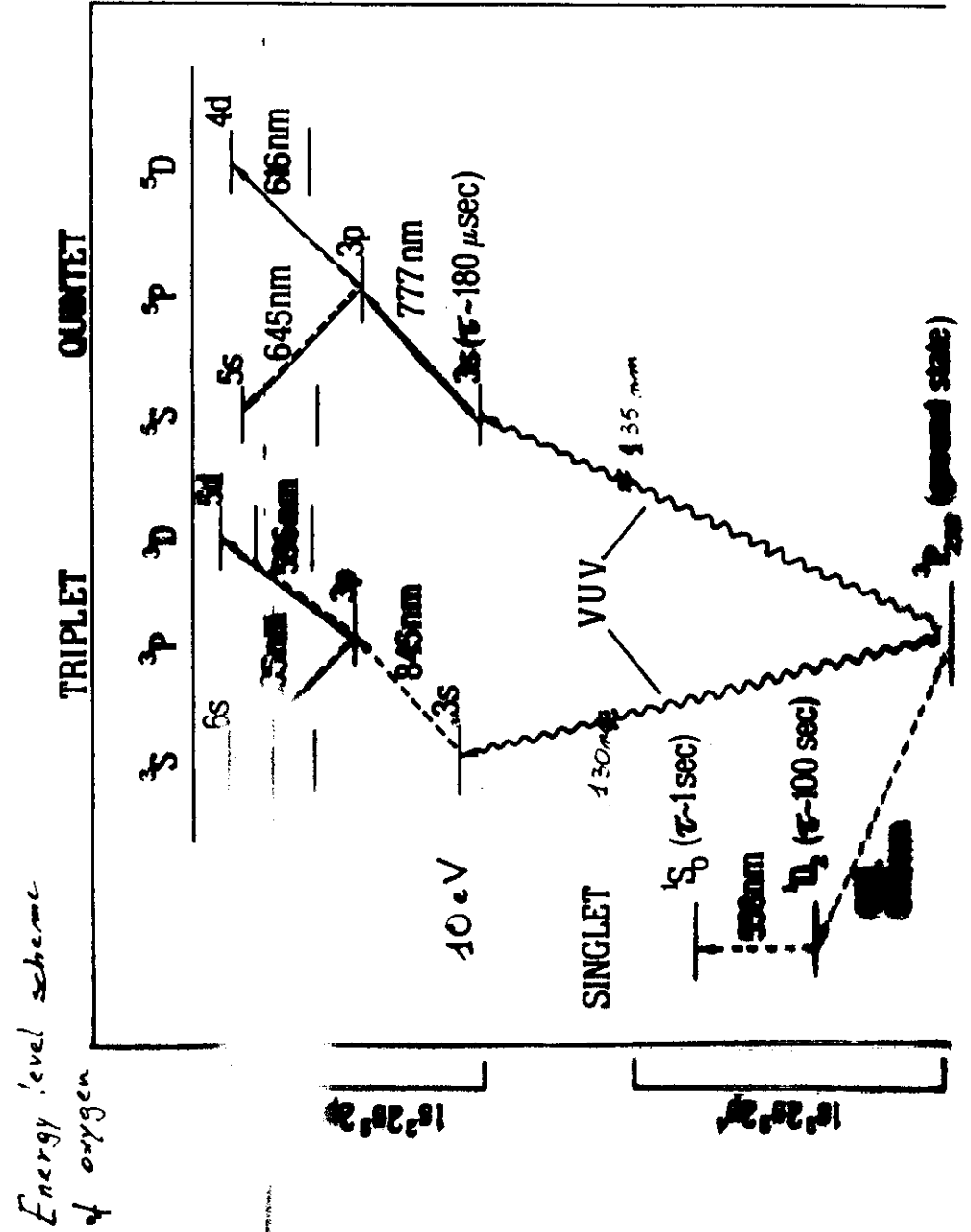
WHY OXYGEN?

- OXYGEN IS AN IMPORTANT ELEMENT IN MANY CHEMICAL AND PHYSICAL PROCESSES (upper atmosphere, combustion, ..)

but

OXYGEN WAS ELUSIVE TO HIGH
RESOLUTION SPECTROSCOPIC INVESTIGATIONS

- i) production of atomic oxygen
- ii) energy level diagram :
optical transitions between highly excited states.





$$P_{in} = P_{out} + \frac{1}{T_2}$$

Intracavity Spectroscopy

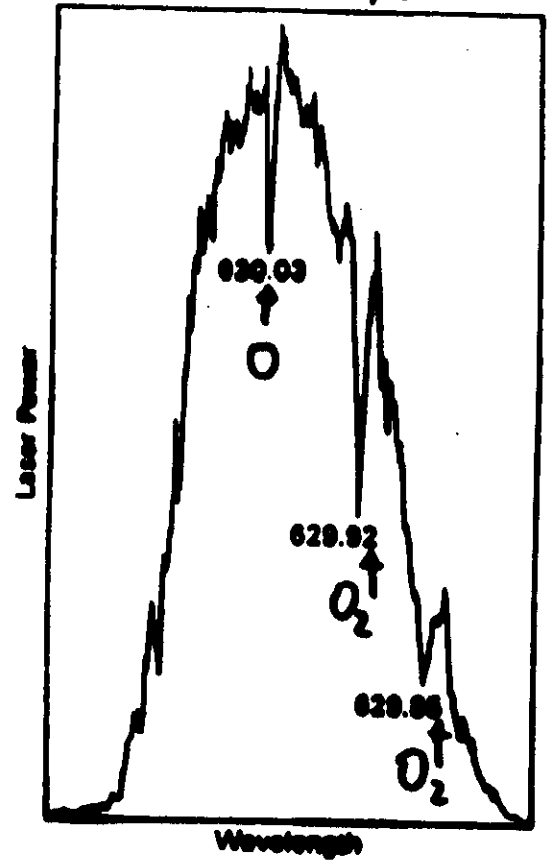


Fig. 1. Intracavity spectrum of the $O(^3P_2)$ transition at 630.03 nm, centered in the laser's output profile. A pair of O_2 lines at 629.92 and 629.85 nm is used for pumping power is about twice threshold (P^{th} element is a tuning wedge).

(1D) transition. The tuning

4 FORBIDDEN TRANSITIONS

TWO PHOTON IN FLAMES

Goldsmith J Chem Phys 78, 1660 (83)

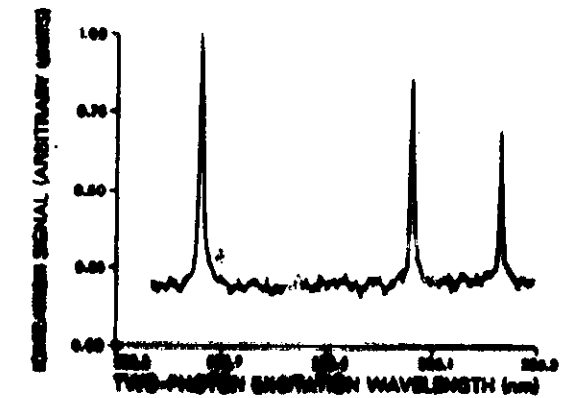
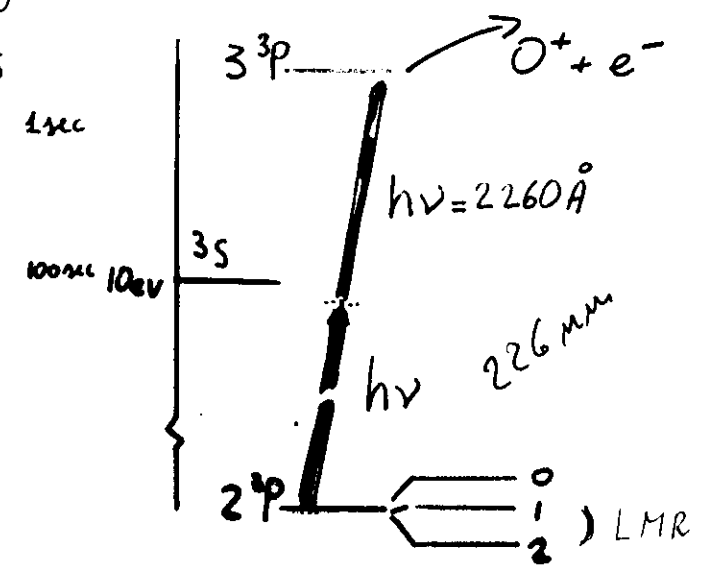
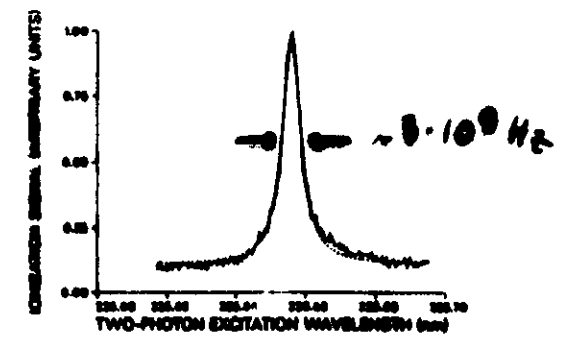


FIG. 1. Ionization signal from atomic oxygen recorded in an atmospheric pressure stoichiometric hydrogen/oxygen/argon flame. There are three peaks because of fine-structure splitting of the 2^3P state.



TWO PHOTON FOURIER TRANSFORM LIMITED PULSED LASER LINEWIDTH

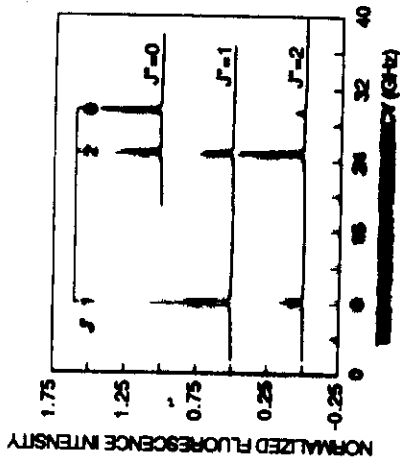


FIG. 1. Doppler-free spectra for the $3p^2P_{j=2} - 2p^2P_{j=2}$ two-photon transitions in atomic oxygen. The line of two-photon emission is modulated with frequency from the $3p^2P_{j=2} - 2p^2P_{j=2}$ transition at 800 nm, transferred to the square of the laser intensity. Doppler-broadening from a broad pulsed interferometer with peak shift spectrum has the same in repeated run. The width of the spectrum is displayed from the bottom one by 0.5 and 1.0 nmol width, respectively.

TABLE I. Experimental and theoretical (in parentheses, from Ref. 5) integrated two-photon cross sections for the $3p^2P_{j=2} - 2p^2P_{j=2}$ transition in atomic oxygen.

J''	J'	Absolute ($10^{-28} \text{ cm}^2 \text{ s}$)	Relative ^b
2	2	13.2 (6.42)	$0.706 \pm 0.077 (0.714)$
2	1	3.63 (2.44)	$0.194 \pm 0.043 (0.198)$
2	0	1.87 (1.16)	$0.100 \pm 0.040 (0.088)$
1	2	6.98 (4.34)	$0.373 \pm 0.067 (0.329)$
1	1	11.7 (8.86)	$0.627 \pm 0.067 (0.671)$
1	0	0 (0)	0 (0)
0	2	9.81 (5.79)	$0.482 \pm 0.045 (0.439)$
0	1	0 (0)	0 (0)
0	0	9.69 (7.43)	$0.518 \pm 0.045 (0.561)$

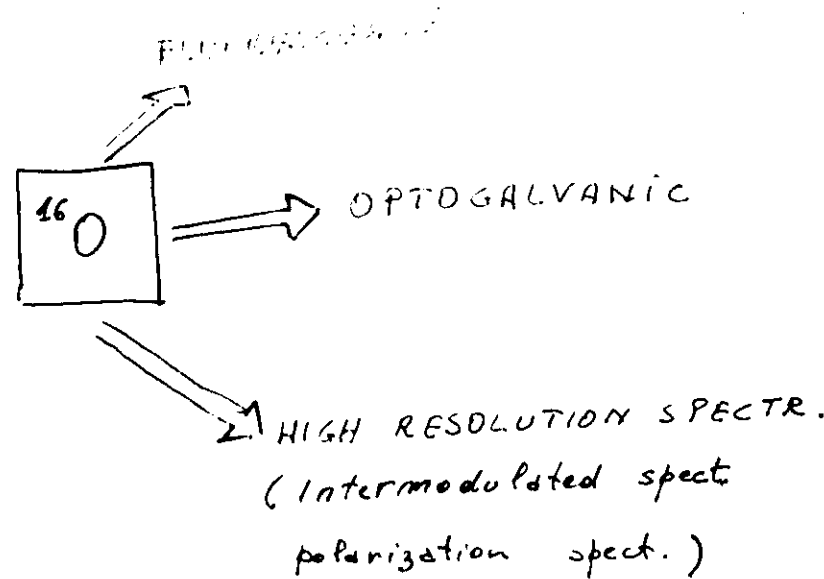
^aThese values must be multiplied by a line-shape factor $g(\omega - \omega_0)$ to get the effective cross section at any given two-photon radiated frequency, ω . If the line-shape function is a 300-K Doppler profile, $g(\omega) = 1.06 \times 10^{-11} \text{ sec}$.

^bExperimental error limits are $\pm 1\sigma$, where σ is the sample standard deviation.

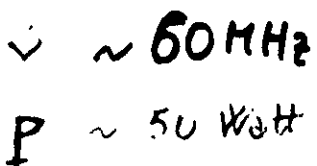
BARCEL et al. Phy. Rev. A 36, 3497 (1987)

LASER DIAGNOSTIC OF OXYGEN PLASMA:

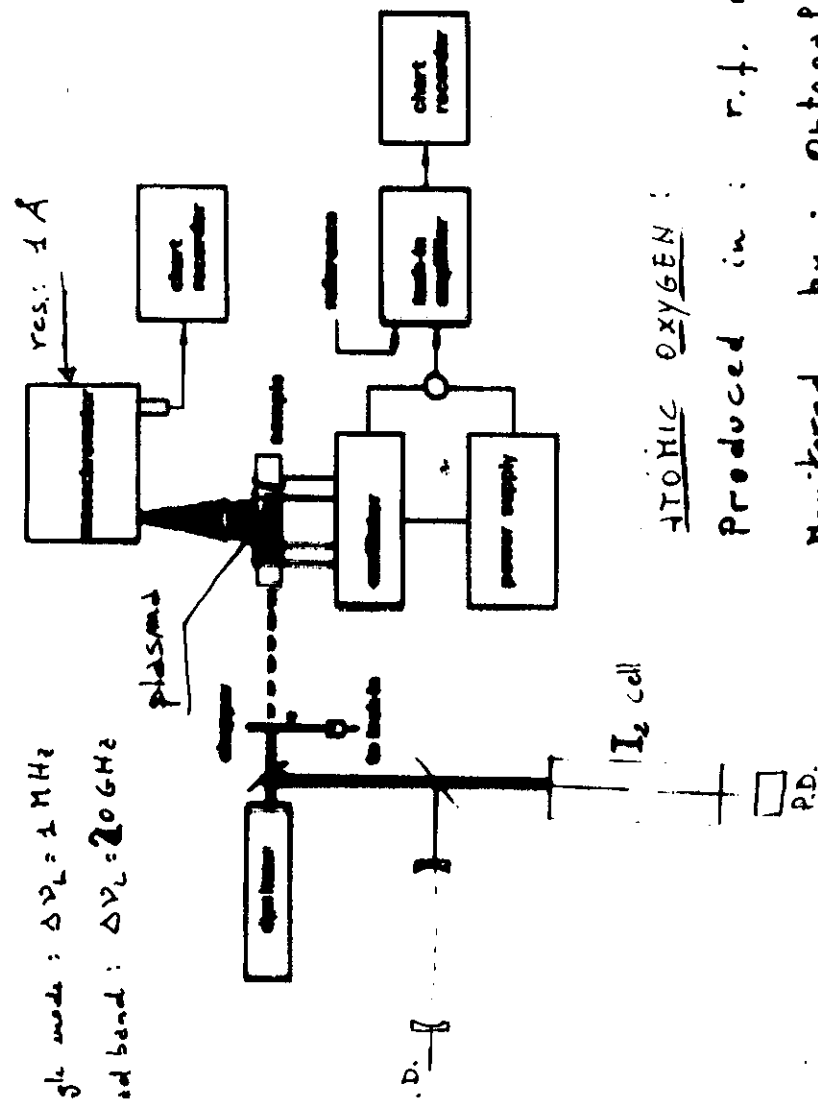
- necessary to perform spectroscopic investigation
- source of informations on the collisional processes governing the O-atoms formation.



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A. Sasso, G. Tino, M. Inguscio, N. Beverini, M. Francese
Nuovo Cimento D 10, 941 (1988)



ИТОГИ ОХУГЕН :

Produced in : r.f. discharge

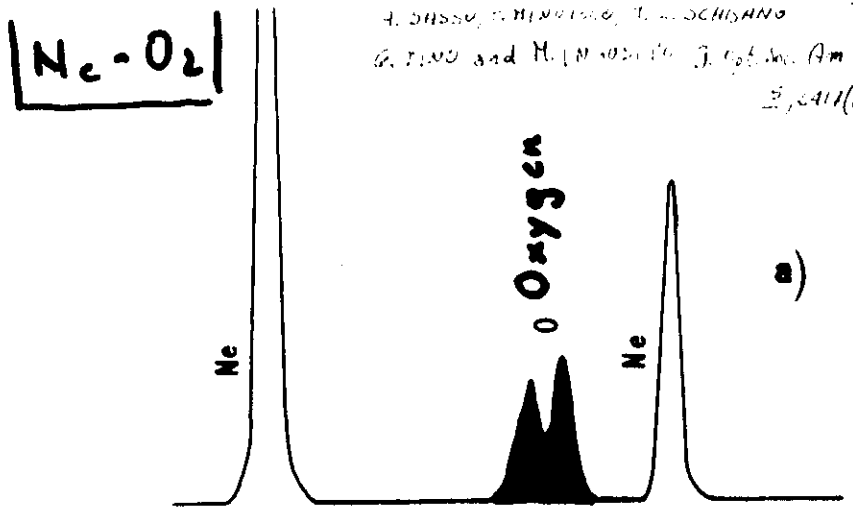
Monitored by: Optogalvanic and

fluorescence detection.

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A. DASSO, J. MINOUE, T. SCHISANO
 G. TINO and M. NORDIO, J. Opt. Soc. Am. B
 3, 2411 (86)

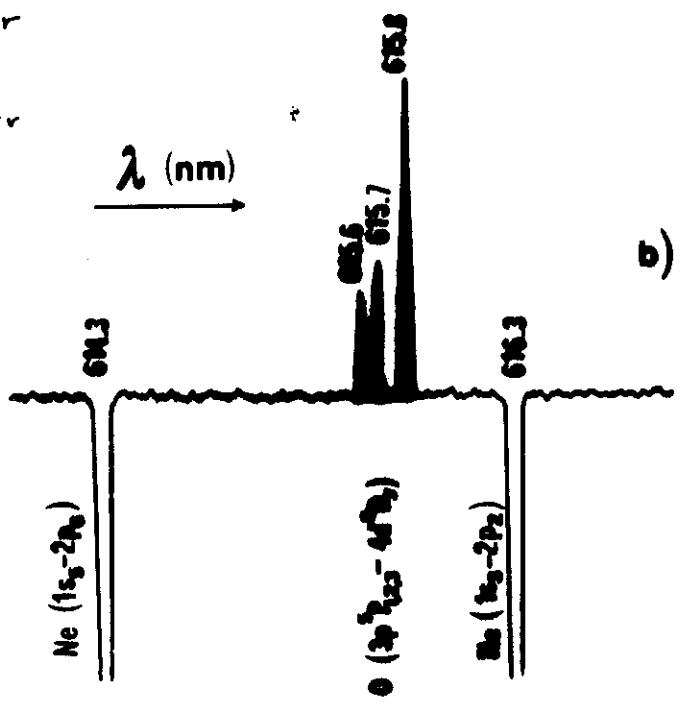
FLUORESCENCE



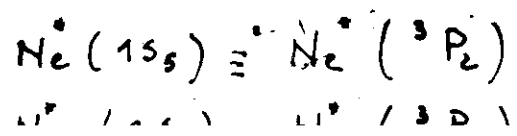
$P_{O_2} = 0.2$ Torr

$P_{Ne} = 10$ Torr

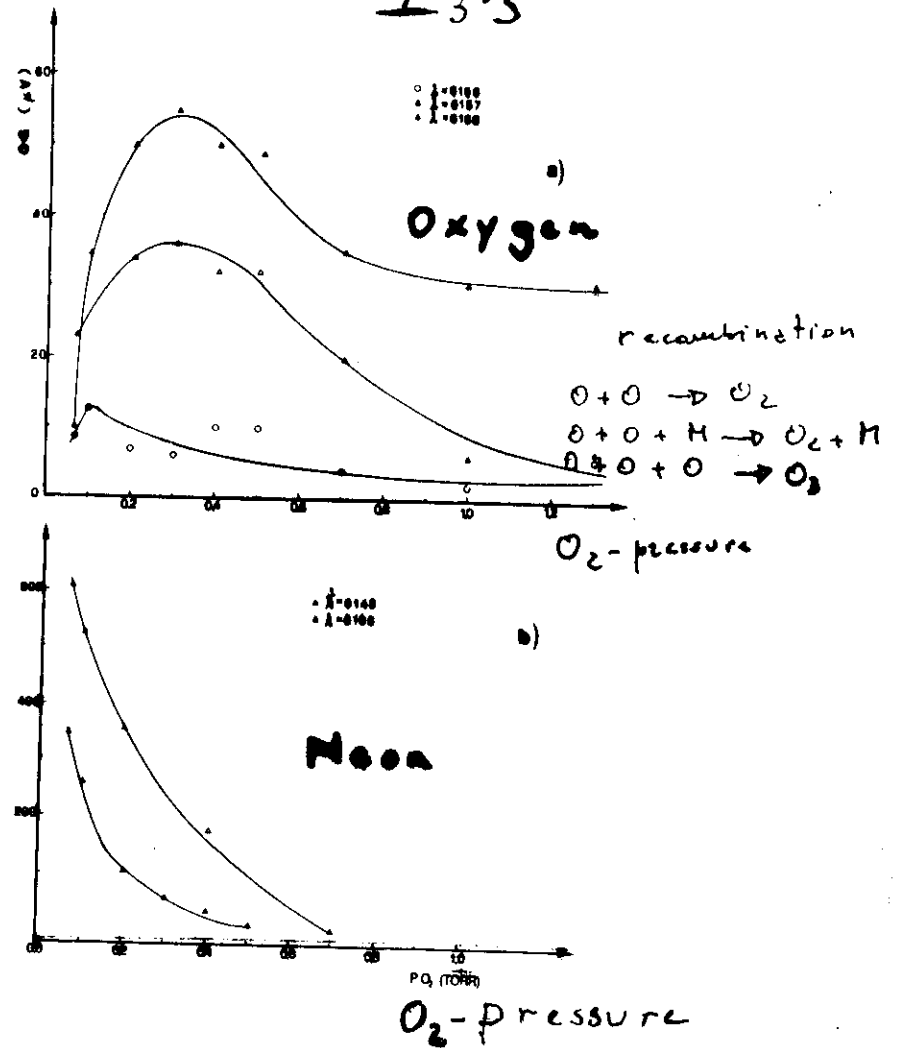
λ (nm)



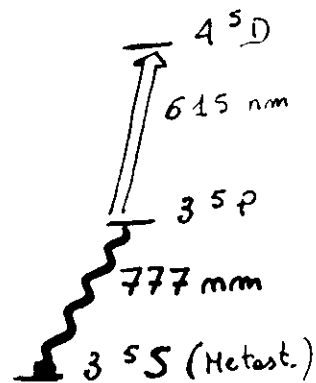
OPTOGALVANIC



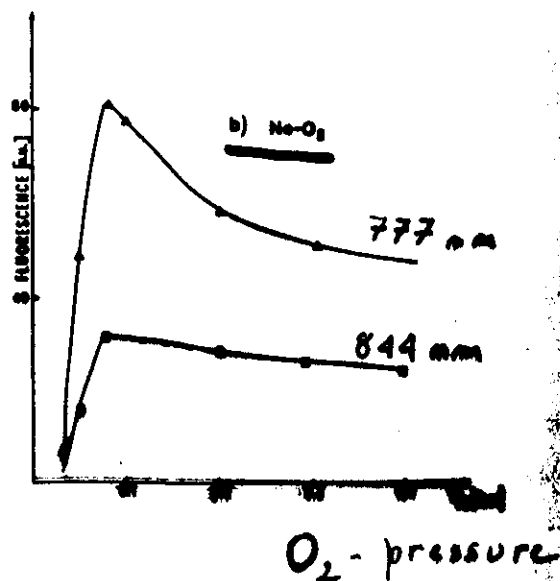
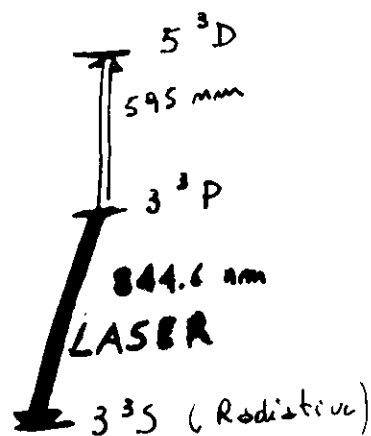
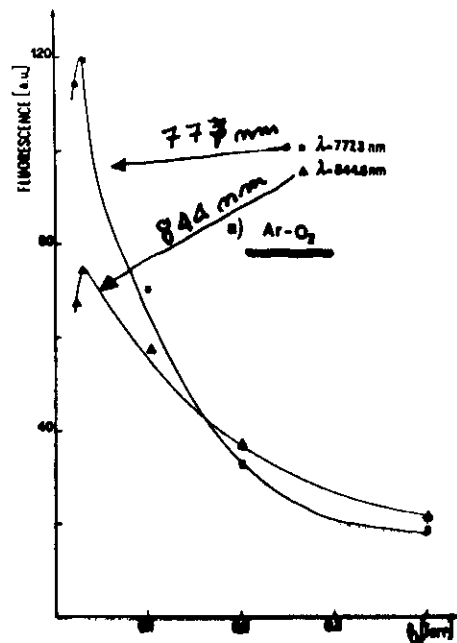
OPTOGALVANIC SIGNAL



The decrease of Ne, OG signals can be explained in terms of the reaction



FLUORESCENCE



W.R. Bennet, Jr., W.L. Faust, R.A. McFarlane, C.M.H. Rotel
P.R.L. 3, 470 (1962)

$$\frac{d[Ne^*]}{dt} = Zn_e[Ne] - \alpha[O_2][Ne^*] - w[Ne^*] - T[Ne^*]^2, \quad (4)$$

$$\frac{d[O^*]}{dt} = \alpha[O_2][Ne^*] - \beta[O_2][O^*] - \frac{[O^*]}{\tau_{O^*}} - D_{diff}[O^*], \quad (5)$$

$$\frac{d[O]}{dt} = \frac{[O^*]}{\tau_{O^*}} + 3\beta[O_2][O^*] + \alpha[O_2][Ne^*] - \gamma[Ne^*][O]^2.$$

$$[O^*] = \frac{Z\alpha[O_2][Ne]n_e}{\alpha\beta[O_2]^2 + \left(\frac{\alpha}{\tau_{O^*}} + w\beta + D_{diff}\right)[O_2] + \frac{w}{\tau_{O^*}} + D_{diff}w} \quad (7)$$

TABLE II
REACTIONS INCLUDED IN THE KINETIC MODEL.

REACTION	RATE COEFFICIENT	UNIT
$Ne + e \rightarrow Ne^+$	$k_1 = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + Ne \rightarrow Ne + Ne^+$	$k_2 = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + O_2 \rightarrow Ne + O_2^+$	$k_3 = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + O \rightarrow Ne + O^+$	$k_4 = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + Ar \rightarrow Ne + Ar^+$	$k_5 = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + N_2 \rightarrow Ne + N_2^+$	$k_6 = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + CO_2 \rightarrow Ne + CO_2^+$	$k_7 = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + H_2O \rightarrow Ne + H_2O^+$	$k_8 = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + CH_4 \rightarrow Ne + CH_4^+$	$k_9 = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + C_2H_6 \rightarrow Ne + C_2H_6^+$	$k_{10} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + C_3H_8 \rightarrow Ne + C_3H_8^+$	$k_{11} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + i-C_4H_{10} \rightarrow Ne + i-C_4H_{10}^+$	$k_{12} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + n-C_4H_{10} \rightarrow Ne + n-C_4H_{10}^+$	$k_{13} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + i-C_5H_{12} \rightarrow Ne + i-C_5H_{12}^+$	$k_{14} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + n-C_5H_{12} \rightarrow Ne + n-C_5H_{12}^+$	$k_{15} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + i-C_6H_{14} \rightarrow Ne + i-C_6H_{14}^+$	$k_{16} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + n-C_6H_{14} \rightarrow Ne + n-C_6H_{14}^+$	$k_{17} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + i-C_7H_{16} \rightarrow Ne + i-C_7H_{16}^+$	$k_{18} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + n-C_7H_{16} \rightarrow Ne + n-C_7H_{16}^+$	$k_{19} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + i-C_8H_{18} \rightarrow Ne + i-C_8H_{18}^+$	$k_{20} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + n-C_8H_{18} \rightarrow Ne + n-C_8H_{18}^+$	$k_{21} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + i-C_9H_{20} \rightarrow Ne + i-C_9H_{20}^+$	$k_{22} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + n-C_9H_{20} \rightarrow Ne + n-C_9H_{20}^+$	$k_{23} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + i-C_{10}H_{22} \rightarrow Ne + i-C_{10}H_{22}^+$	$k_{24} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + n-C_{10}H_{22} \rightarrow Ne + n-C_{10}H_{22}^+$	$k_{25} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + i-C_{11}H_{24} \rightarrow Ne + i-C_{11}H_{24}^+$	$k_{26} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + n-C_{11}H_{24} \rightarrow Ne + n-C_{11}H_{24}^+$	$k_{27} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + i-C_{12}H_{26} \rightarrow Ne + i-C_{12}H_{26}^+$	$k_{28} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + n-C_{12}H_{26} \rightarrow Ne + n-C_{12}H_{26}^+$	$k_{29} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + i-C_{13}H_{28} \rightarrow Ne + i-C_{13}H_{28}^+$	$k_{30} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + n-C_{13}H_{28} \rightarrow Ne + n-C_{13}H_{28}^+$	$k_{31} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + i-C_{14}H_{30} \rightarrow Ne + i-C_{14}H_{30}^+$	$k_{32} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + n-C_{14}H_{30} \rightarrow Ne + n-C_{14}H_{30}^+$	$k_{33} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + i-C_{15}H_{32} \rightarrow Ne + i-C_{15}H_{32}^+$	$k_{34} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + n-C_{15}H_{32} \rightarrow Ne + n-C_{15}H_{32}^+$	$k_{35} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + i-C_{16}H_{34} \rightarrow Ne + i-C_{16}H_{34}^+$	$k_{36} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + n-C_{16}H_{34} \rightarrow Ne + n-C_{16}H_{34}^+$	$k_{37} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + i-C_{17}H_{36} \rightarrow Ne + i-C_{17}H_{36}^+$	$k_{38} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + n-C_{17}H_{36} \rightarrow Ne + n-C_{17}H_{36}^+$	$k_{39} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + i-C_{18}H_{38} \rightarrow Ne + i-C_{18}H_{38}^+$	$k_{40} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + n-C_{18}H_{38} \rightarrow Ne + n-C_{18}H_{38}^+$	$k_{41} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + i-C_{19}H_{40} \rightarrow Ne + i-C_{19}H_{40}^+$	$k_{42} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + n-C_{19}H_{40} \rightarrow Ne + n-C_{19}H_{40}^+$	$k_{43} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + i-C_{20}H_{42} \rightarrow Ne + i-C_{20}H_{42}^+$	$k_{44} = 1.0 \times 10^{-16}$	cm^3/sec
$Ne^+ + n-C_{20}H_{42} \rightarrow Ne + n-C_{20}H_{42}^+$	$k_{45} = 1.0 \times 10^{-16}$	cm^3/sec

ATOMIC OXYGEN FORMATION

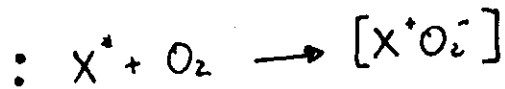
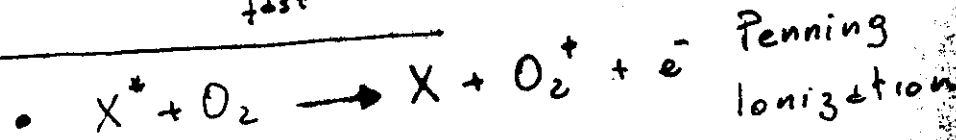
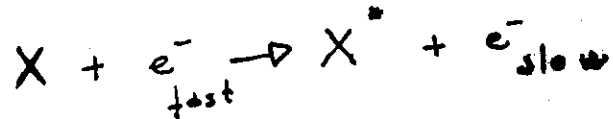
- Pure O_2 - discharge

Electron - impact dissociation



- O_2 - noble gases mixture (X)

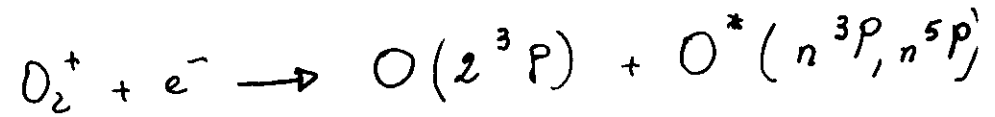
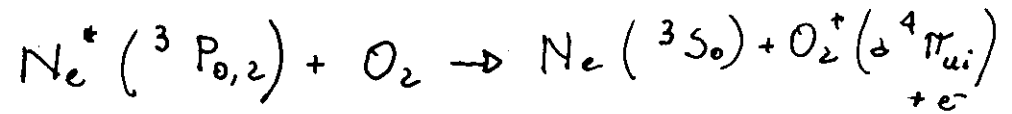
i) Metastable atoms (X) are efficiently produced and represent a reservoir of energy:



For $He^*, Ne^* \rightarrow P.I.$

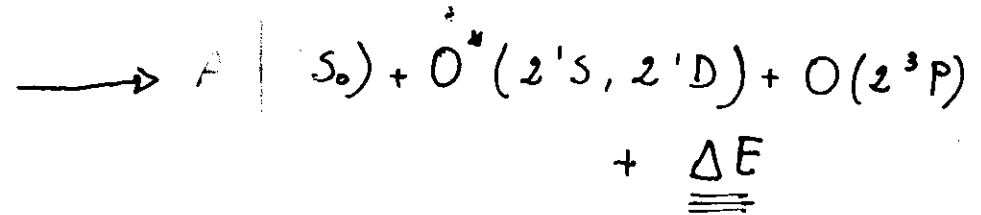
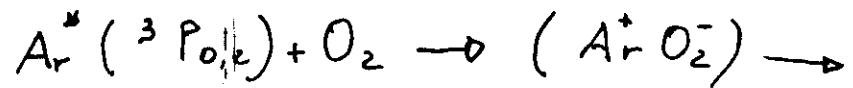
For $Ar^*, Kr^*, Xe^* \rightarrow \text{Quasi-molecules}$

- Ne - O_2 mixtures

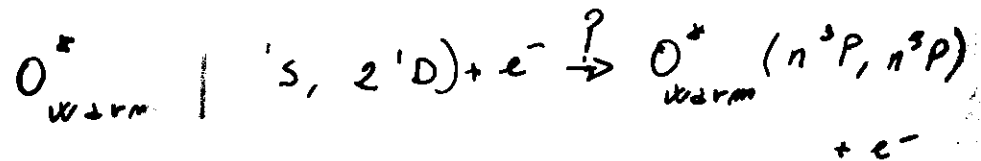


resonant dissociative recombination

- Ar - O_2 mixtures



ΔE : energy defect $\rightarrow$ kinetic energy of warm atoms



LINESHAPES OF OPTICAL TRANSITION IN R. F. DISCHARGE

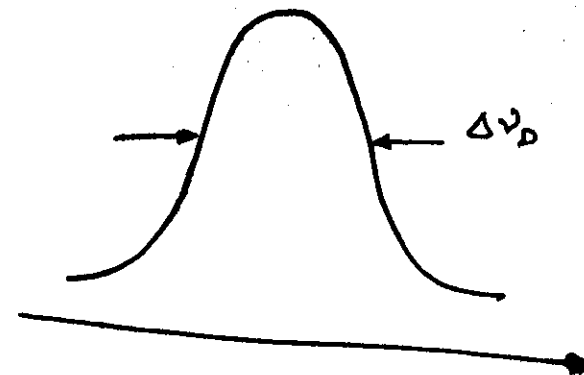
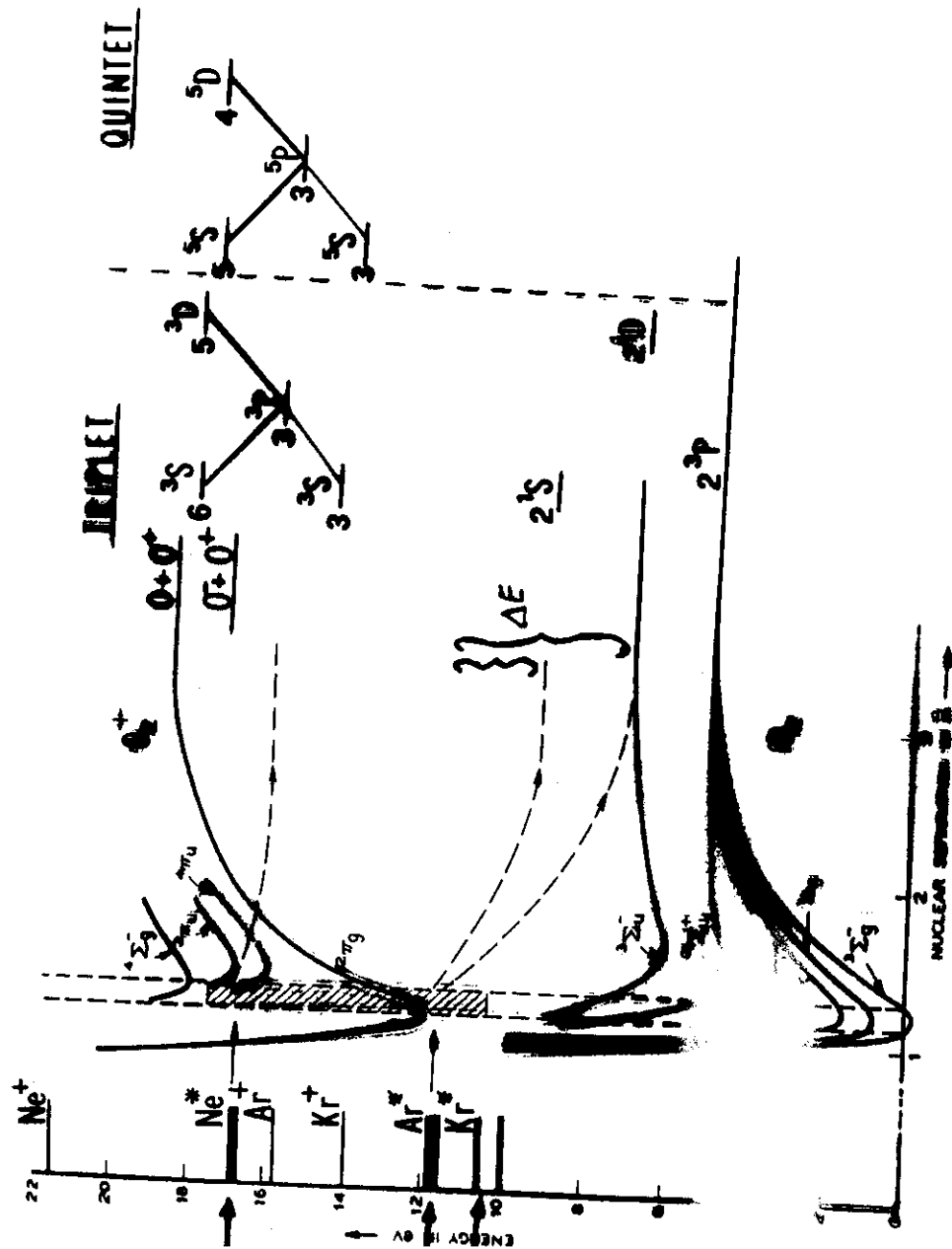
DOPLER-BROADENING IS DOMINANT
AT OUR PRESSURES ($p = 0.1 \div 10$ Torr)

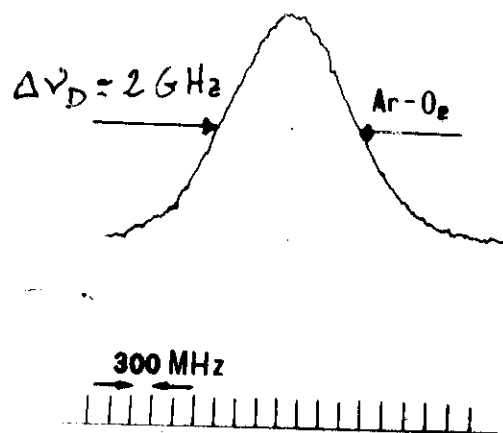
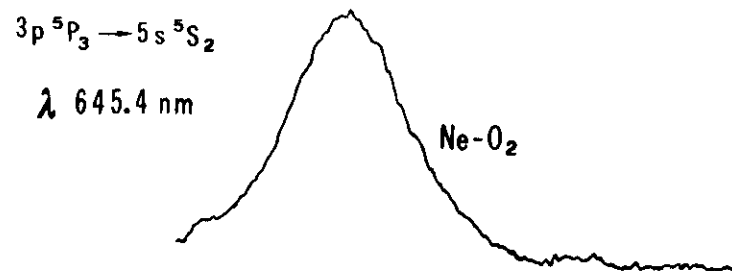
$$\Delta \nu_D = 7.16 \cdot 10^{-7} \nu_0 \sqrt{\frac{T}{M}}$$

From $\Delta \nu_D$ measurements



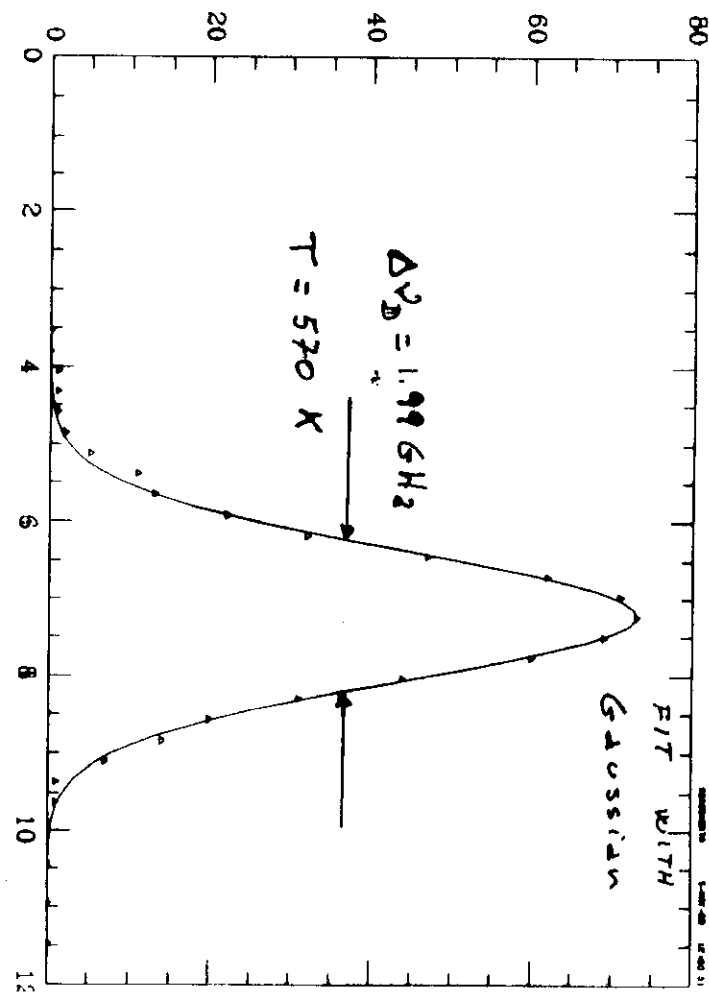
Temperature of the discharge



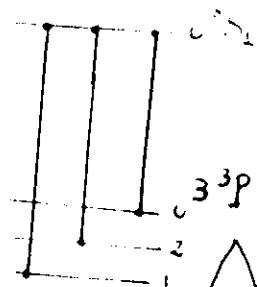


"Normal" Doppler width

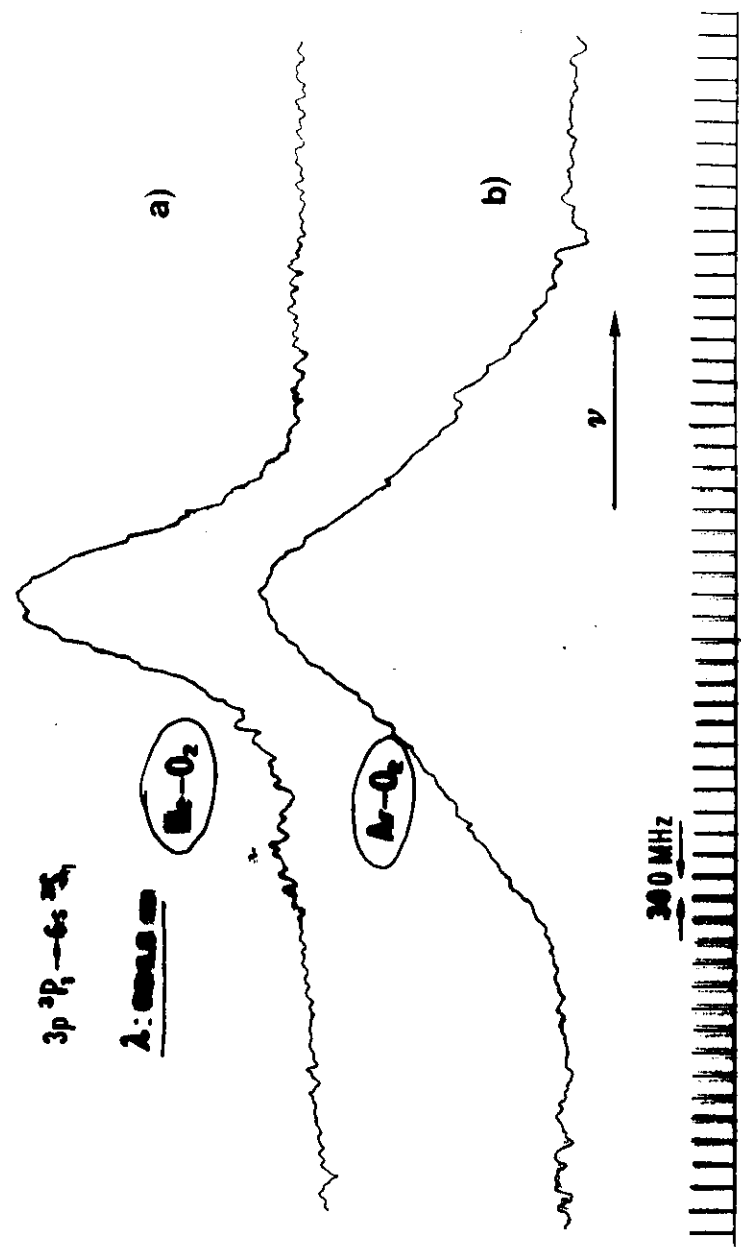
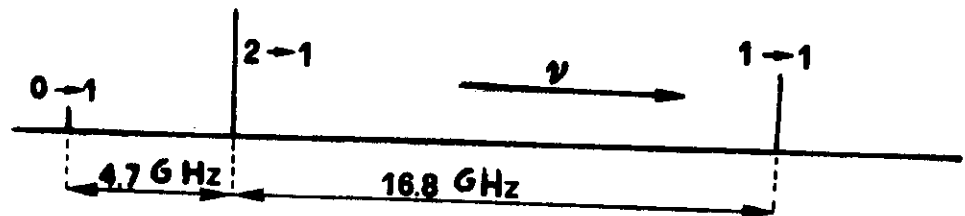
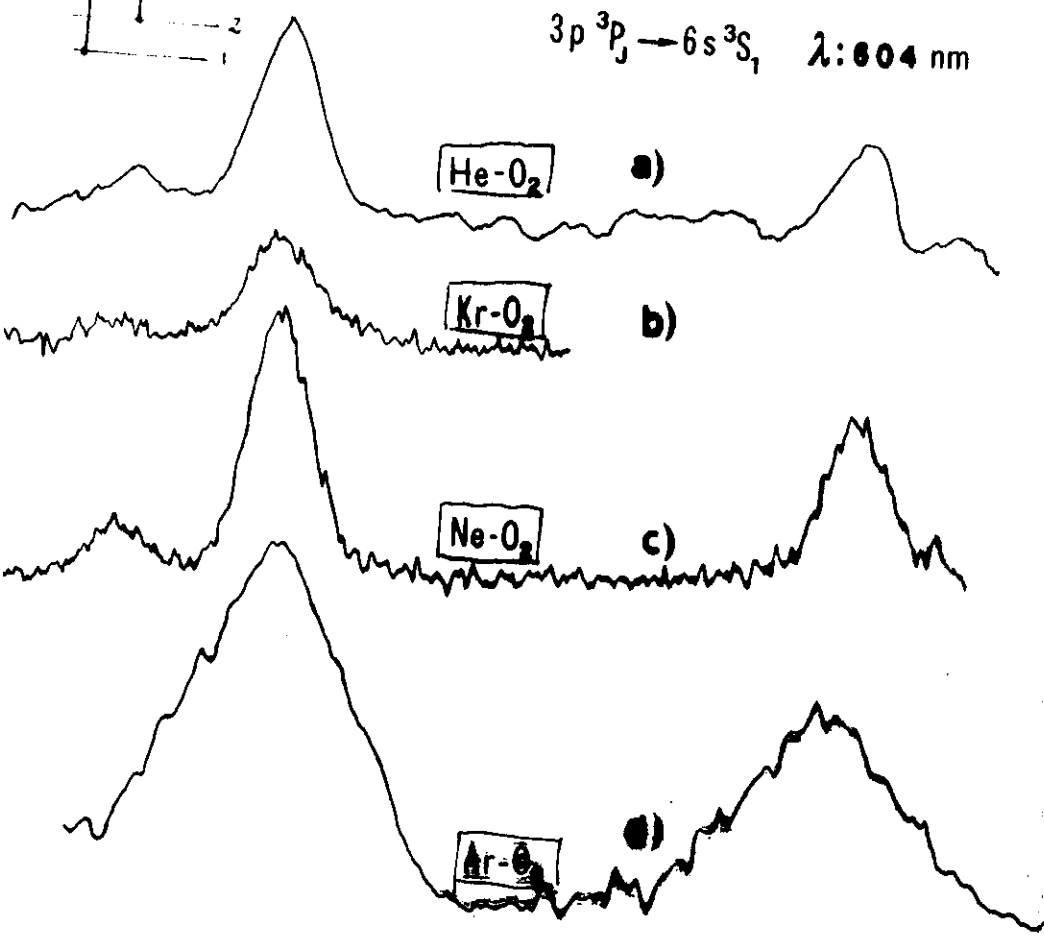
$$\Delta\nu_D = 2 \text{ GHz} \Leftrightarrow T = 570 \text{ K}$$



Ne-O₂: "pure" Doppler broadening



$3p^3P_J \rightarrow 6s^3S_1$ $\lambda: 604 \text{ nm}$



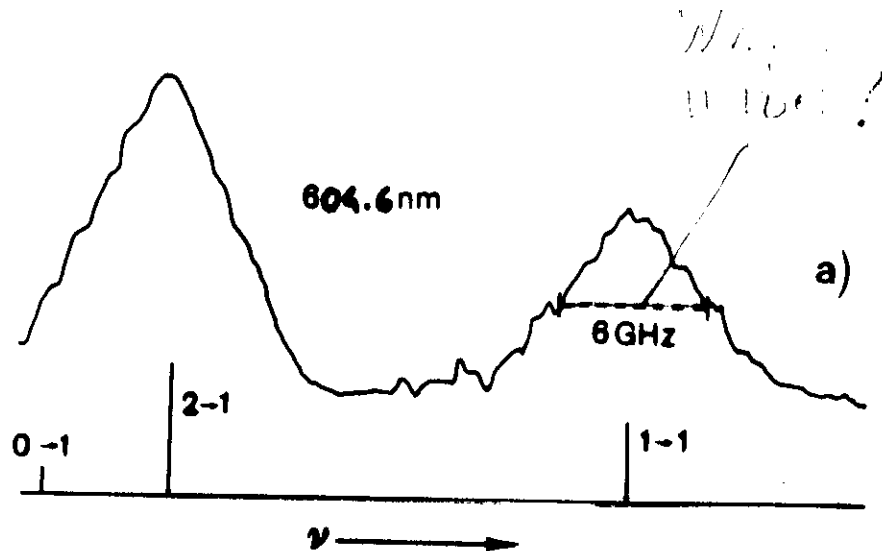
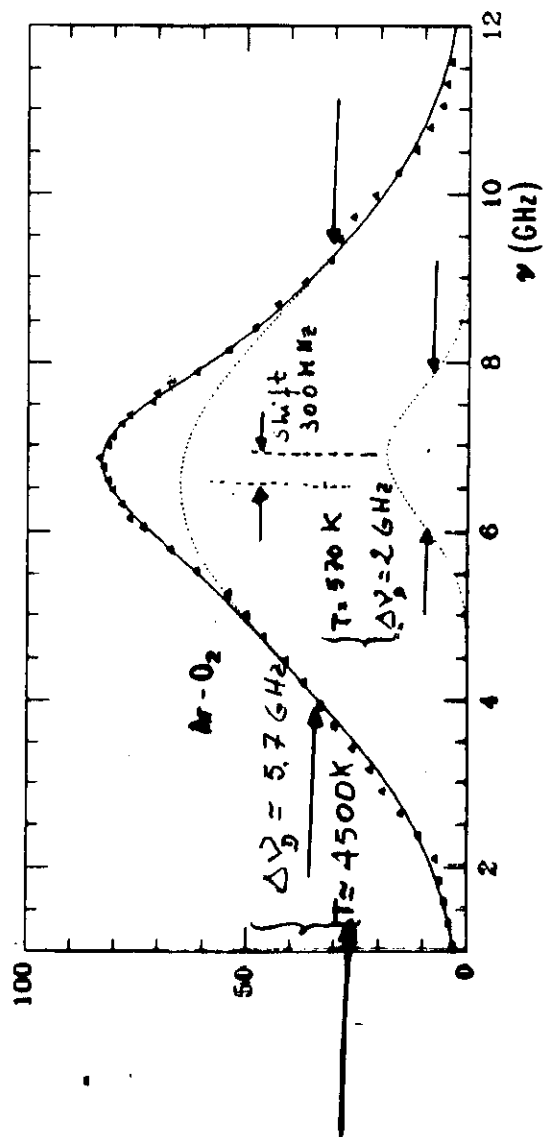
$3p^3P_J \rightarrow 6s^3S_1$

$\lambda: 604.9 \text{ nm}$

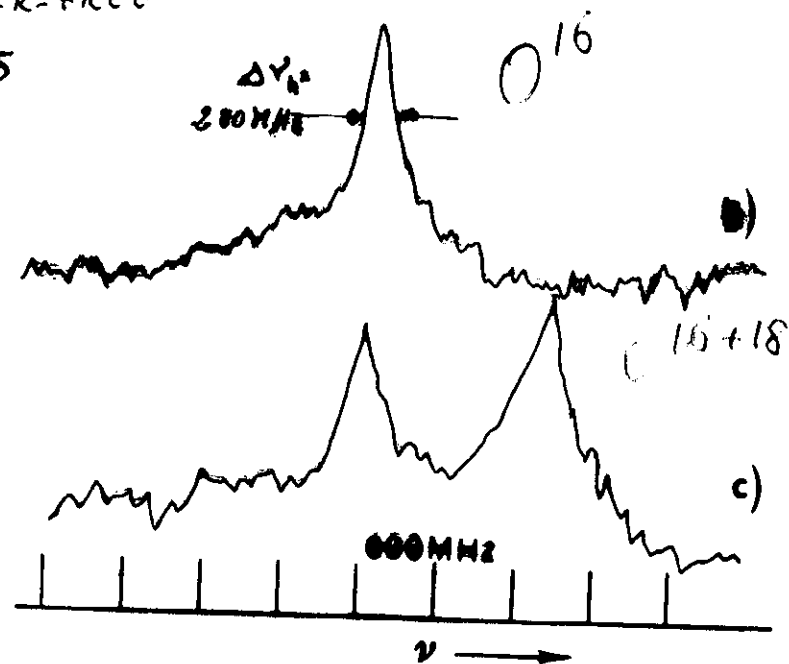
300 MHz

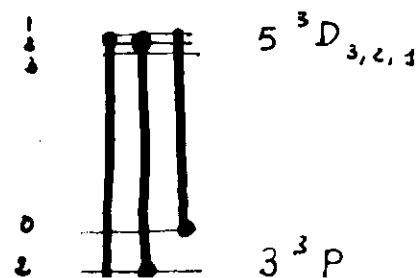
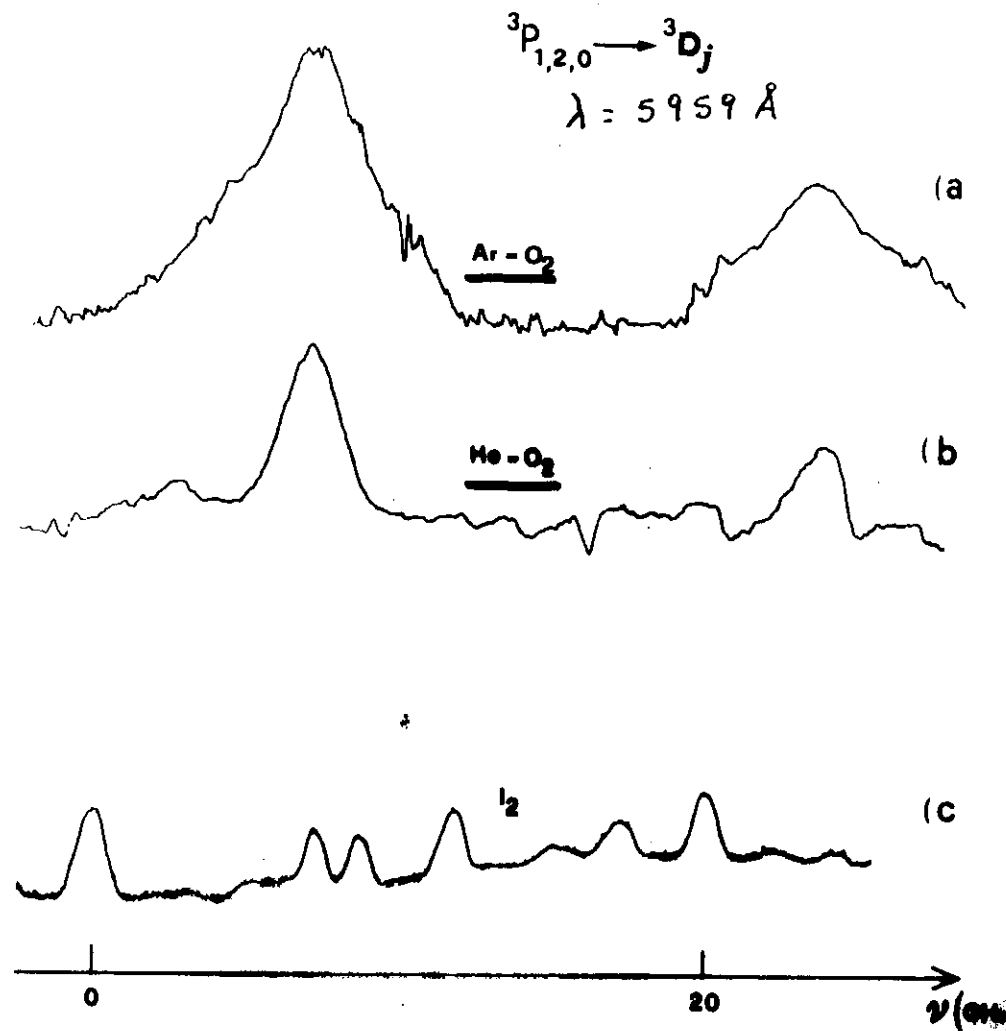
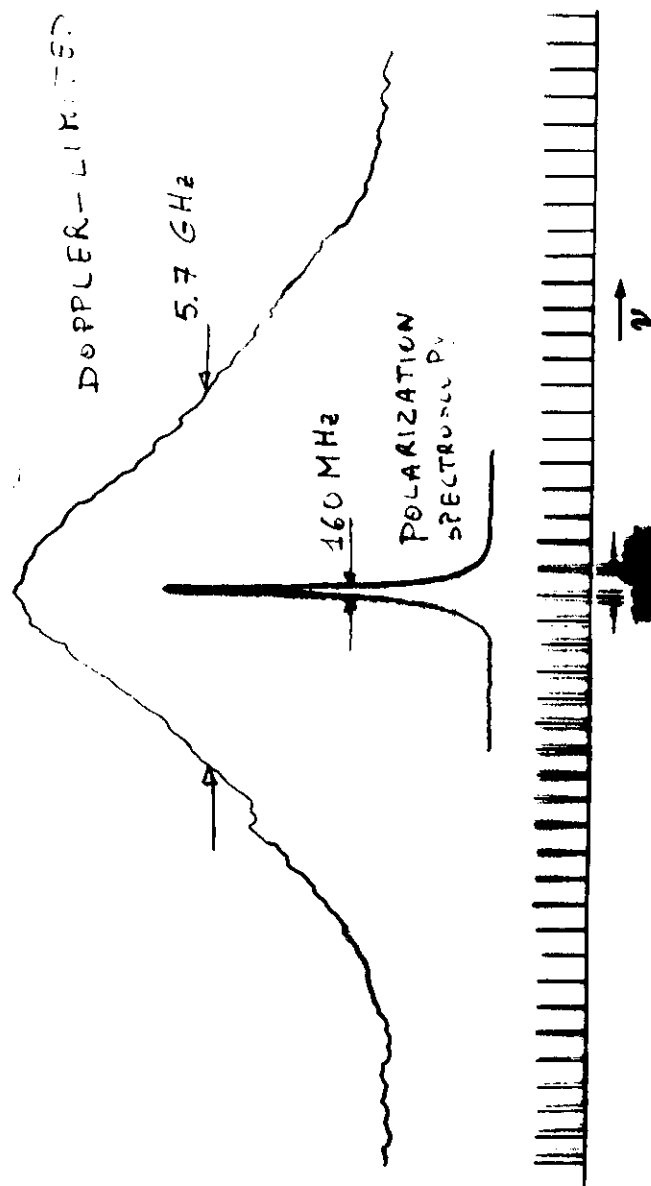
Ar-O₂: "anomalous" broadening

exp. lineshapes fitted with two Gaussians



DOPPLER-FREE
IMOGS

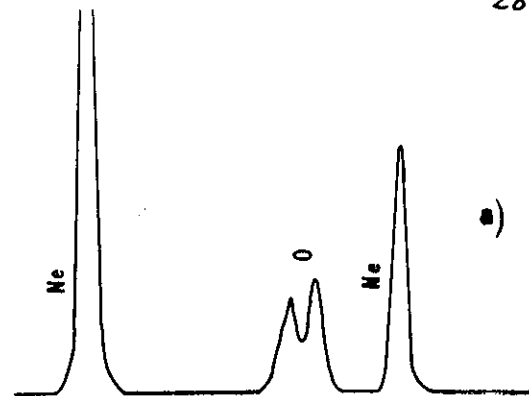




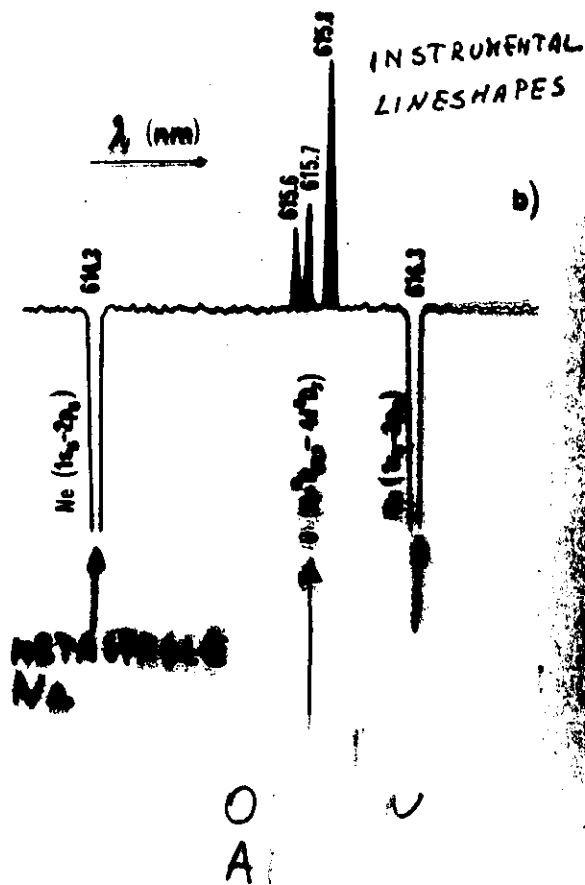
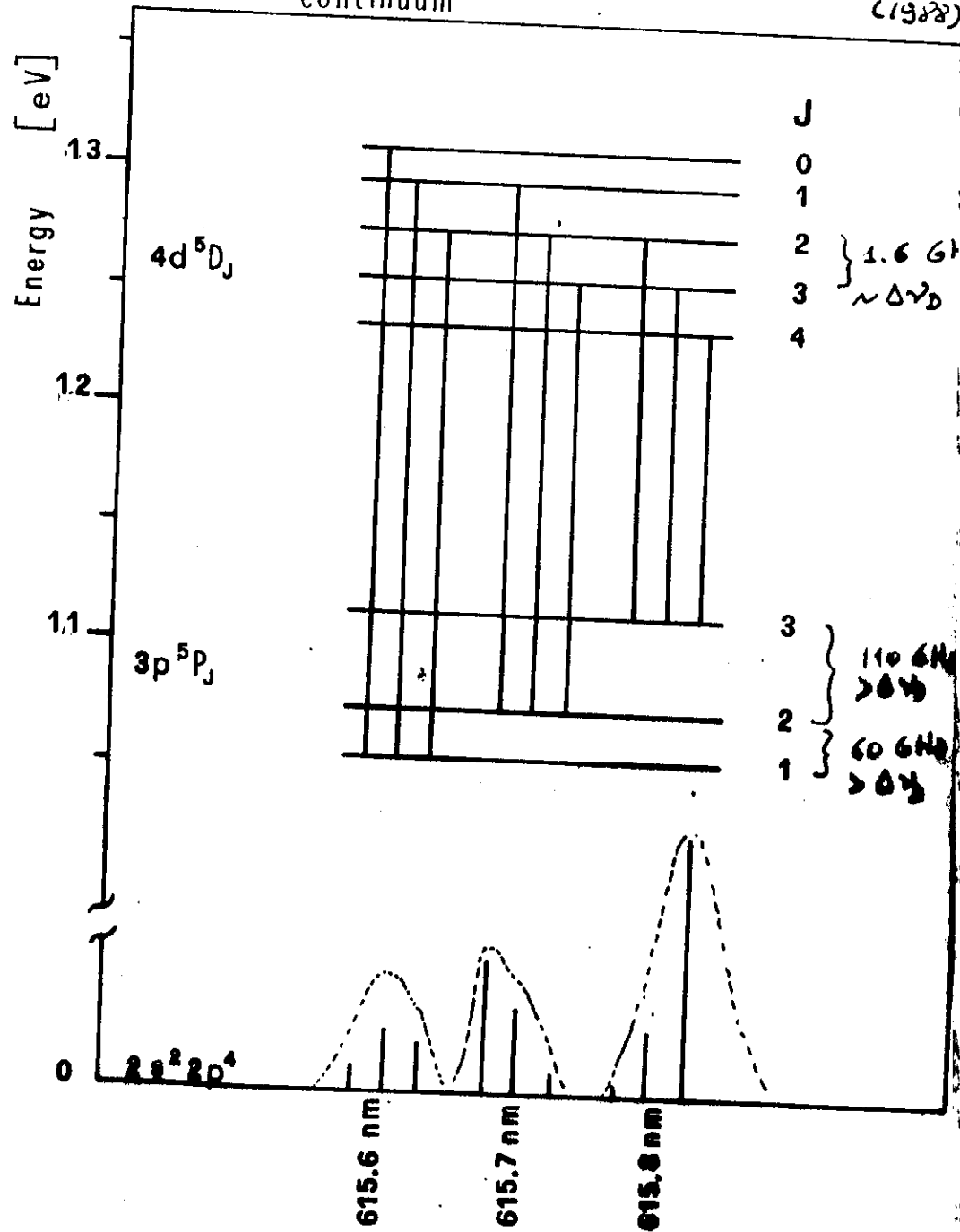
FLUORESC.

 O_2 .1 Torr

Ne 10 Torr



O.G.

MULTIMODE
LASERFINE STRUCTURE OF THE TRANSITION
M.I., P. MINUTOLO, A. SASSO, G. TINO : Phys. Rev. A 37, 4056
(1988)
continuum

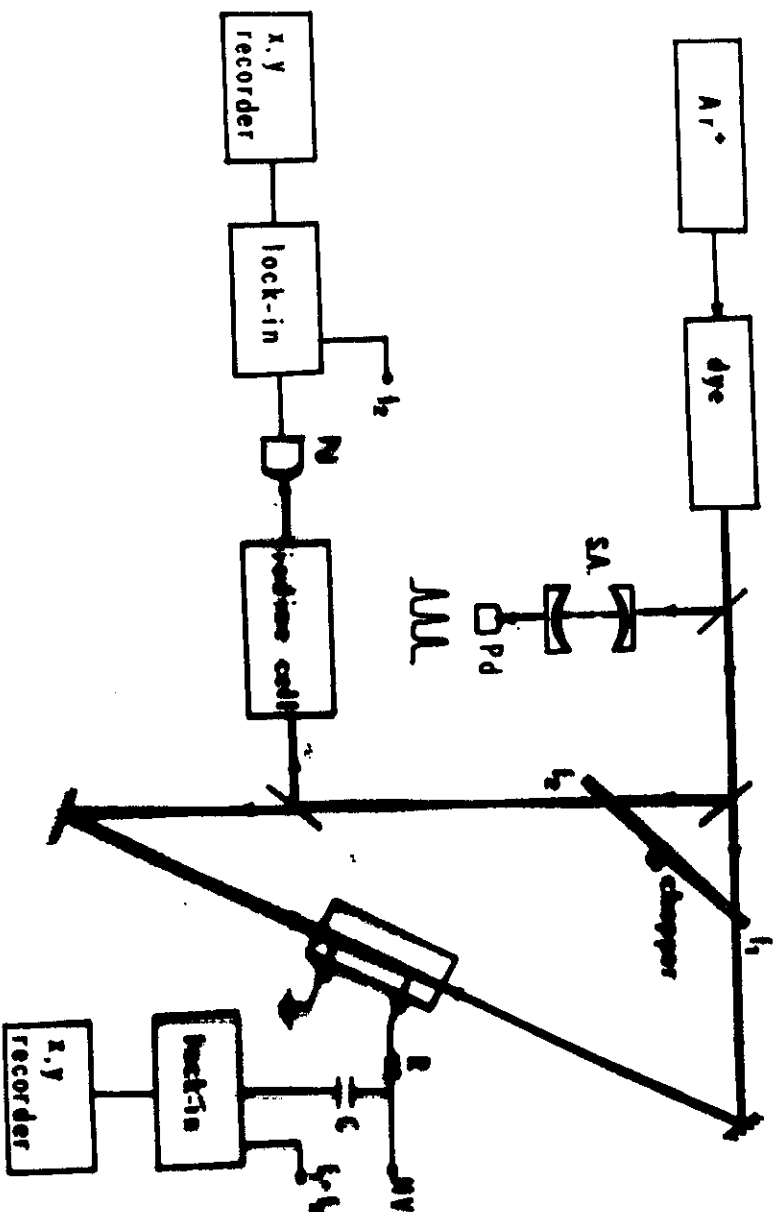
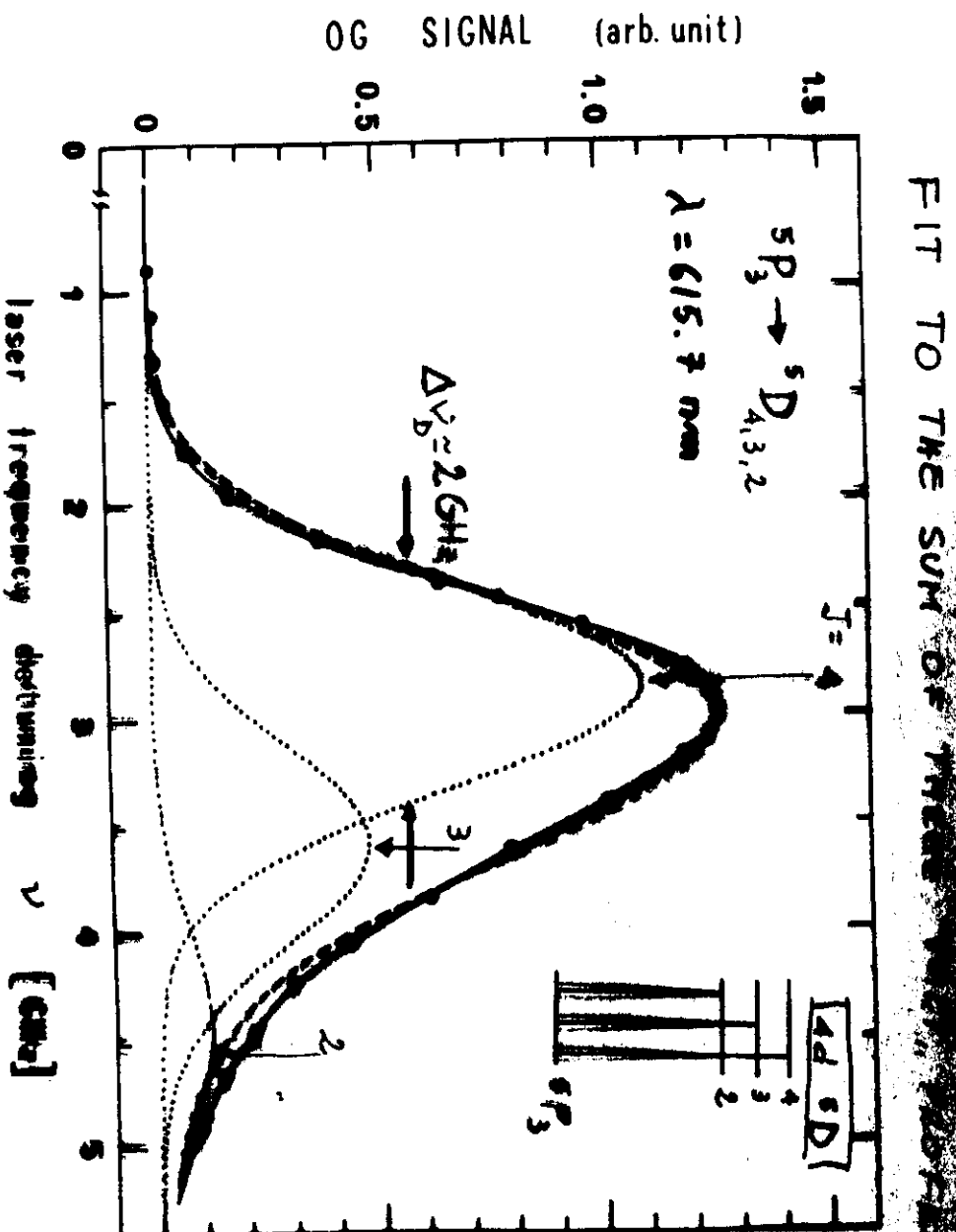
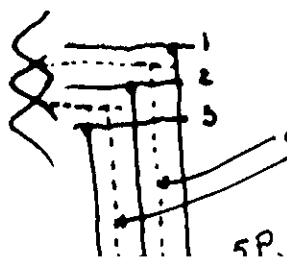
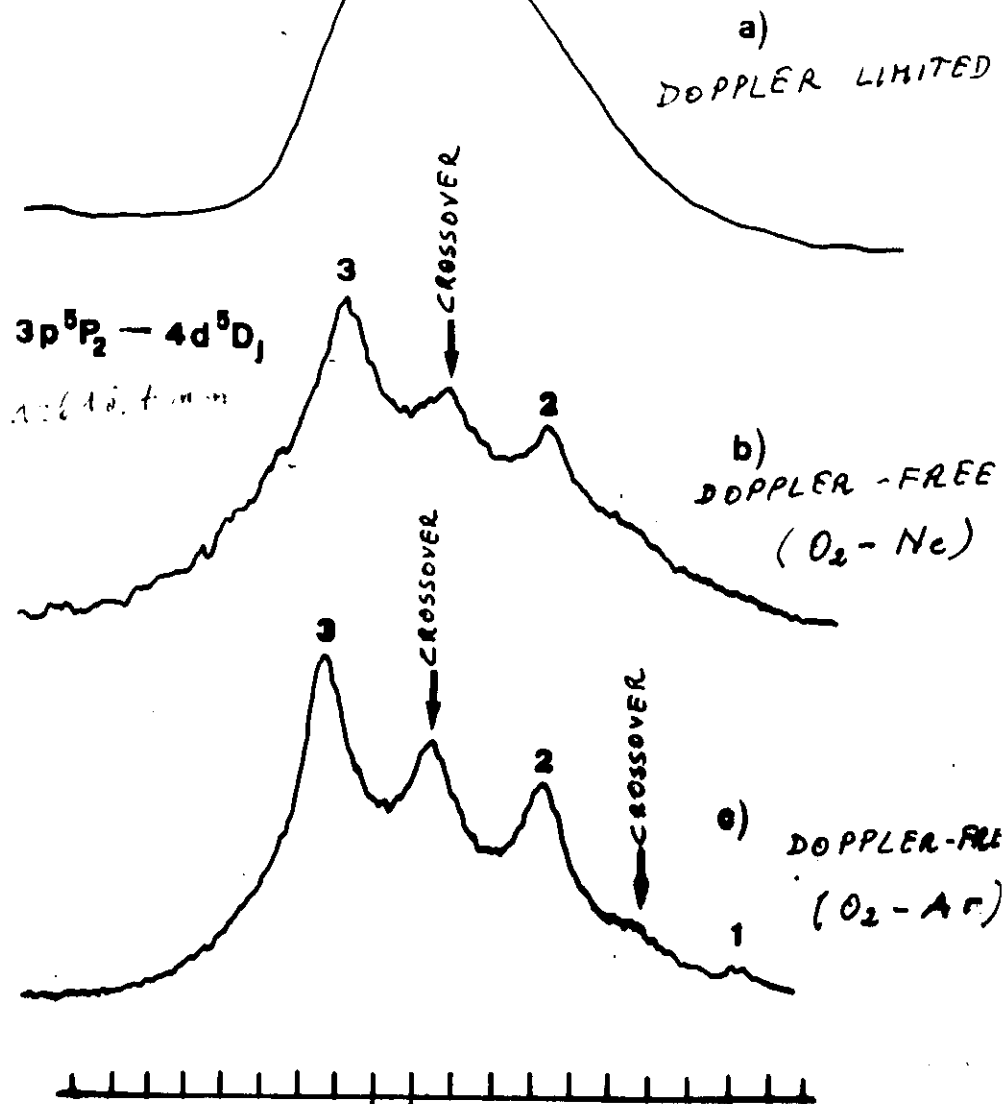


FIG. 2

APPARATUS FOR HIGH RESOLUTION.

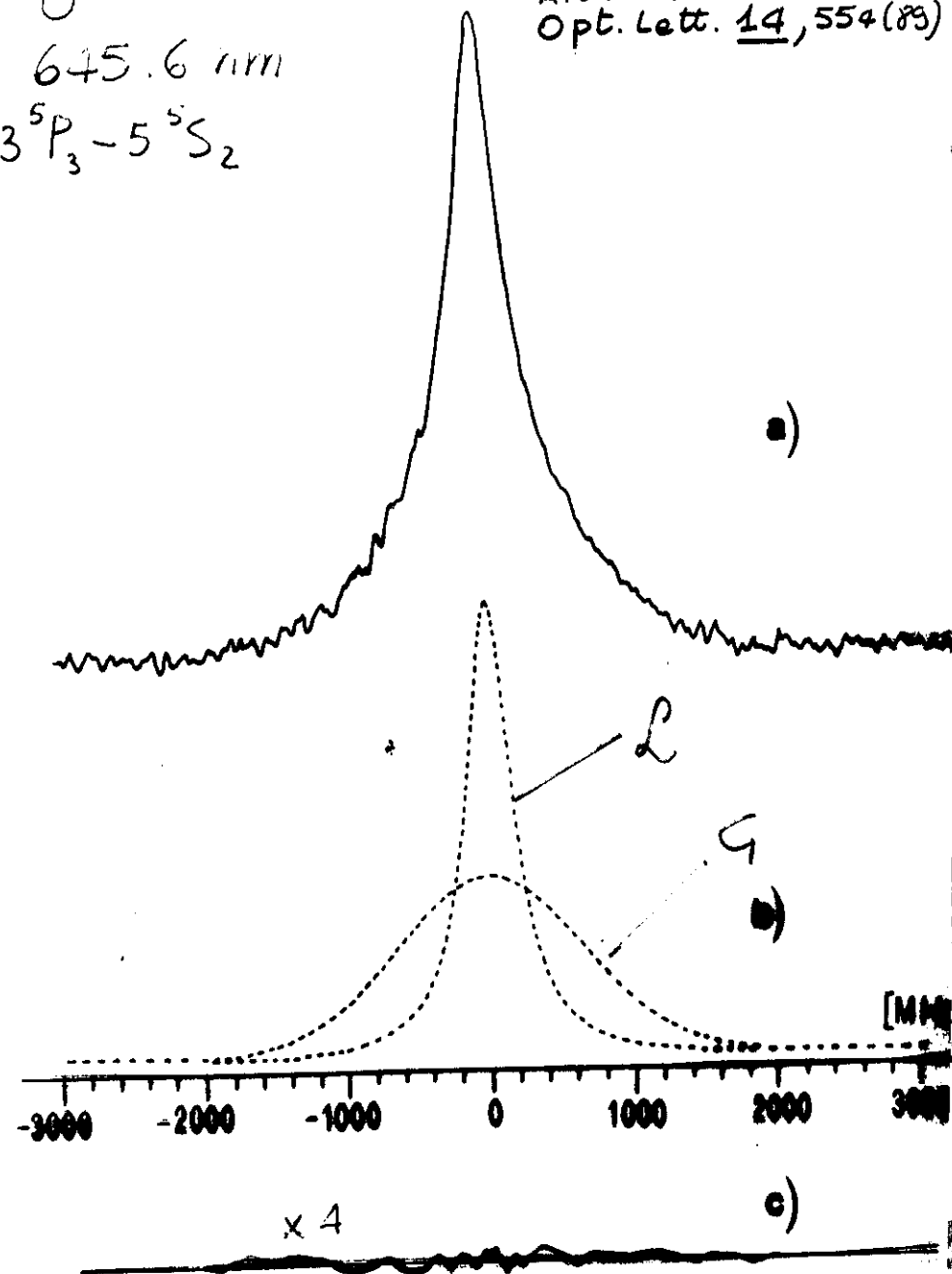




Optogalvanic recording of the O transition $^3P_1 \rightarrow ^1D_{3,1}$ using a single mode dye laser. The Doppler broadened signal is shown in a) while the intermediate scheme is used in b) and c). Different gas mixtures were used, namely O_2/Ne in b) and O_2/Ar in c). The total scan for the Doppler registration is 2.7 times broader than the Doppler-free recordings. Arrow denotes the cross-over peaks.

K. ERNST, P. MINUTOLO, 33.
 A. SASSO, G. TINO, M. I.:
 Opt. Lett. 14, 554 (89)

645.6 nm
 $3^5P_3 - 5^5S_2$



$$A \left[\frac{\gamma^2/4}{\gamma^2/4 + (\nu - \nu_0)^2} + C \exp\left(-\frac{\nu - \nu_0}{\delta \nu_D}\right)^2 \right] \times 4$$

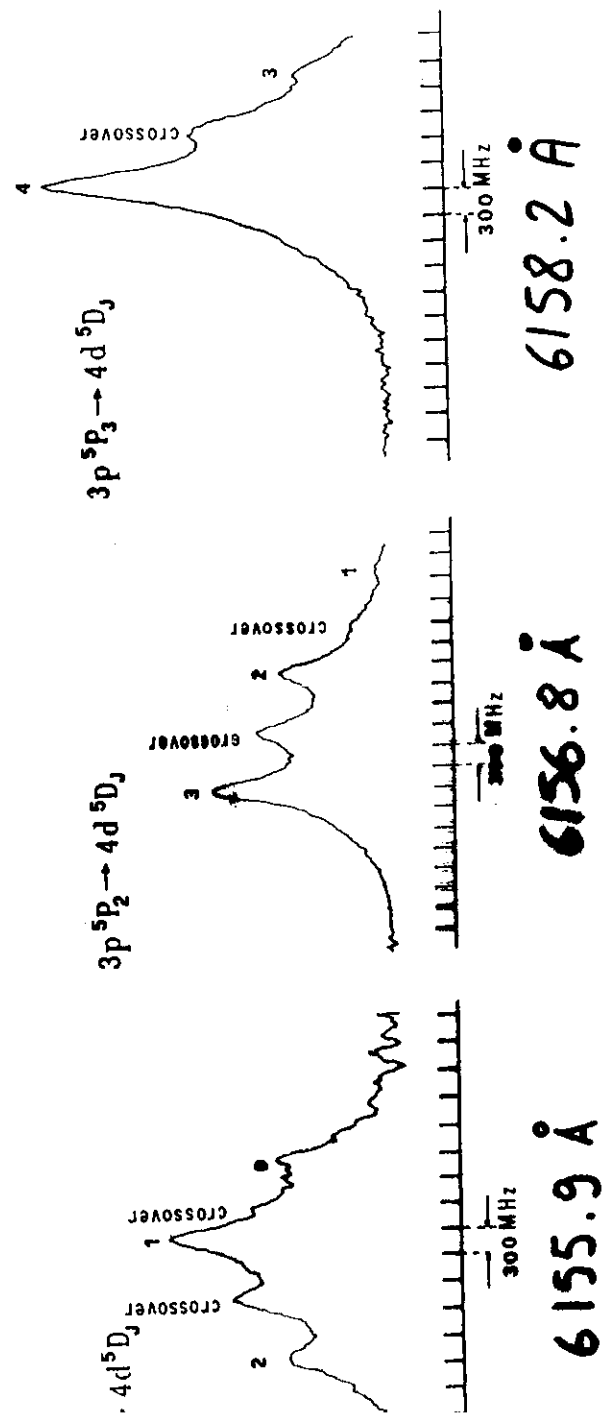
$$C = 2\sqrt{\pi} k_y 2 \frac{\Gamma \gamma}{\delta \Delta_D}$$

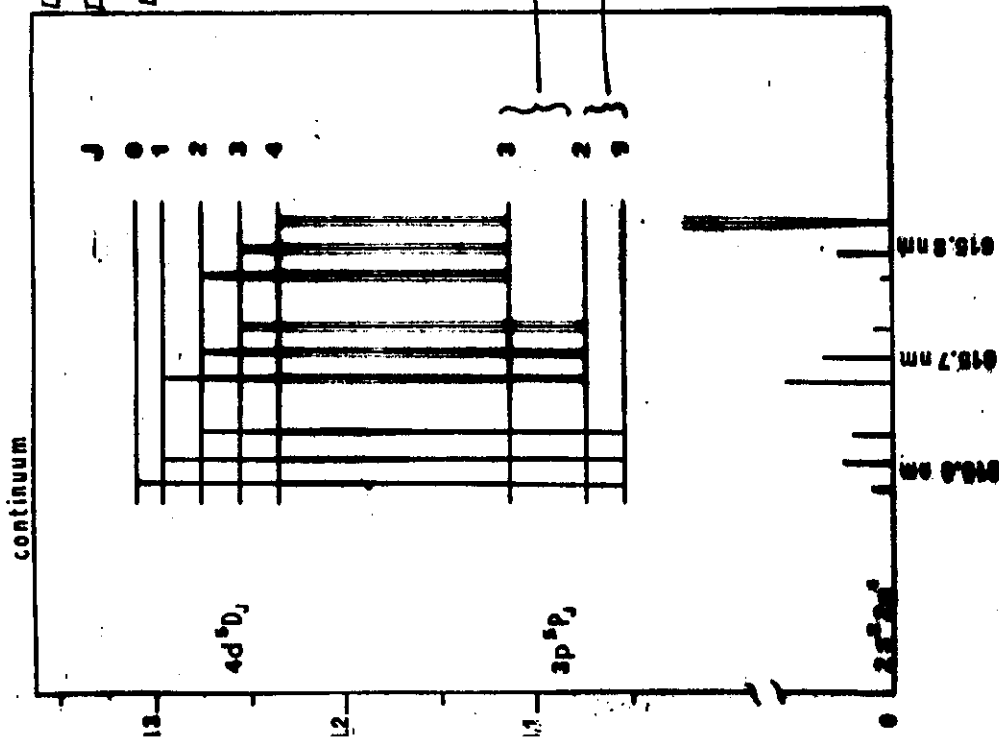
O_2/Ar (.1/5 Torr)

$$\Delta \nu_D = 2 \text{ GHz} (520^\circ K)$$

$$\gamma = \Delta \nu_{hom} = 386 \text{ MHz (FWHM)}$$

$$C = 0.39$$





[2] H. M. G. S. et al. Phys. Rev. A 37, 4056 (1986)
 [2] L. S. Chompa et al. J. Phys. B 8, 728 (1975)
 [3] K. J. Eriksen et al. Ann. Phys. 24, 549 (1963)

Table I

Δ_{ij}	Our results (MHz)	Ref. [2] (MHz)	Ref. [3] (MHz)
Δ_{0-1}	735(20)	786(51)	720(200)
Δ_{1-2}	1360(20)	1416(24)	1469(200)
Δ_{2-3}	1594(20)	1680(24)	1769(200)
Δ_{3-4}	1362(20)	1407(30)	1229(200)

$\Delta_{32} \approx 140 \text{ GHz}$
 $\Delta_{42} \approx 59 \text{ GHz}$

TABLE I. Fine structure splittings in the atomic oxygen $2p^3(4S^o)4d^5D_{4,3,2,1,0}$ levels as obtained from sub-Doppler spectroscopy. For comparison the data deduced in conventional spectroscopy are reported.

FINE STRUCTURE THEORY

$$H = \sum_i \left[\frac{p_i^2}{2m} - \frac{Ze^2}{r_i} + U_i(\vec{r}_i) \right] + \sum_{i,j} \left[\frac{e^2}{r_{ij}} - U_{ij}(\vec{r}_{ij}) \right] + \sum_i \Delta(\vec{r}_i) \vec{\ell}_i \cdot \vec{s}_i \quad (2.1)$$

• Regola dell'intervallo di Lande

$$\Delta W_J = A(\alpha, L, S) J$$

Expected (Lande)	Measured
$\Delta_{10} = 1800 \text{ MHz}$	735 MHz
$\Delta_{21} = 3120 \text{ "}$	1360 "
$\Delta_{32} = 3570 \text{ "}$	1594 "
$\Delta_{43} = 2640 \text{ "}$	1362 "

• Hamiltoniana di Breit-Dirac
 contributi relativistici
 spin-orbita, spin-altra orbita, spin-spin

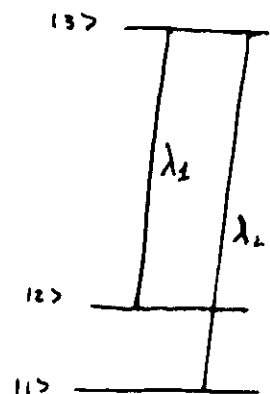
• Formula di Casimir

$$W_J = \frac{1}{2} A'(\alpha LS) + \frac{3}{8} B'(\alpha LS) \left[X(X+1) - \frac{4}{3} S(S+1) L(L+1) \right]$$

$$W = T/(T+1) - 1/(1+1) - S/(S+1)$$

ACCURATE MEASUREMENTS OF LARGE FINE STRUCTURE SPLITTINGS

38



$\Delta\nu >$ SCAN WIDTH
OF THE LASER
($\sim 30 \text{ GHz}$)

$\Delta\nu$ LARGE
SPLITTING

$$\Delta\nu = \nu_2 - \nu_1 = c \left(\frac{1}{\lambda_2} - \frac{1}{\lambda_1} \right)$$

ABSOLUTE WAVELENGTH MEASUREMENT.

1. WAVEMETER
2. IODINE (+ WAVEMETER)

39

VISIBLE $\nu \sim 5 \cdot 10^{14} \text{ Hz}$
DOPPLER WIDTH $\Delta\nu_D \sim 10^9 \text{ Hz}$
SUB-DOPPLER $\Delta\nu_H \sim 10^7$

FREQUENCY OF THE
LASER MUST BE KNOWN
WITH ACCURACY

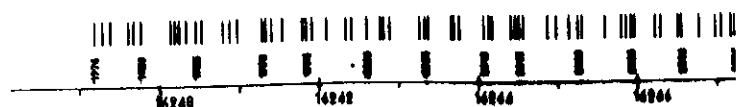
$$10^{-6} \sim 10^{-8}$$

IN THE RANGE $\sim$
11 000 - 21 000 cm^{-1}

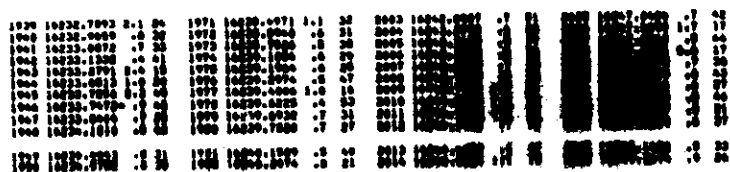
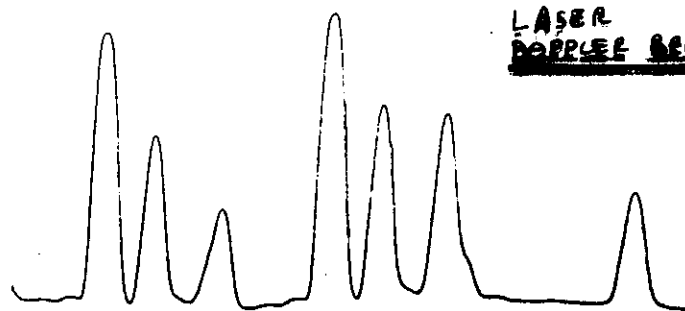
YOU CAN USE I_2
SPECTRUM



FOURIER
TRANSFORM



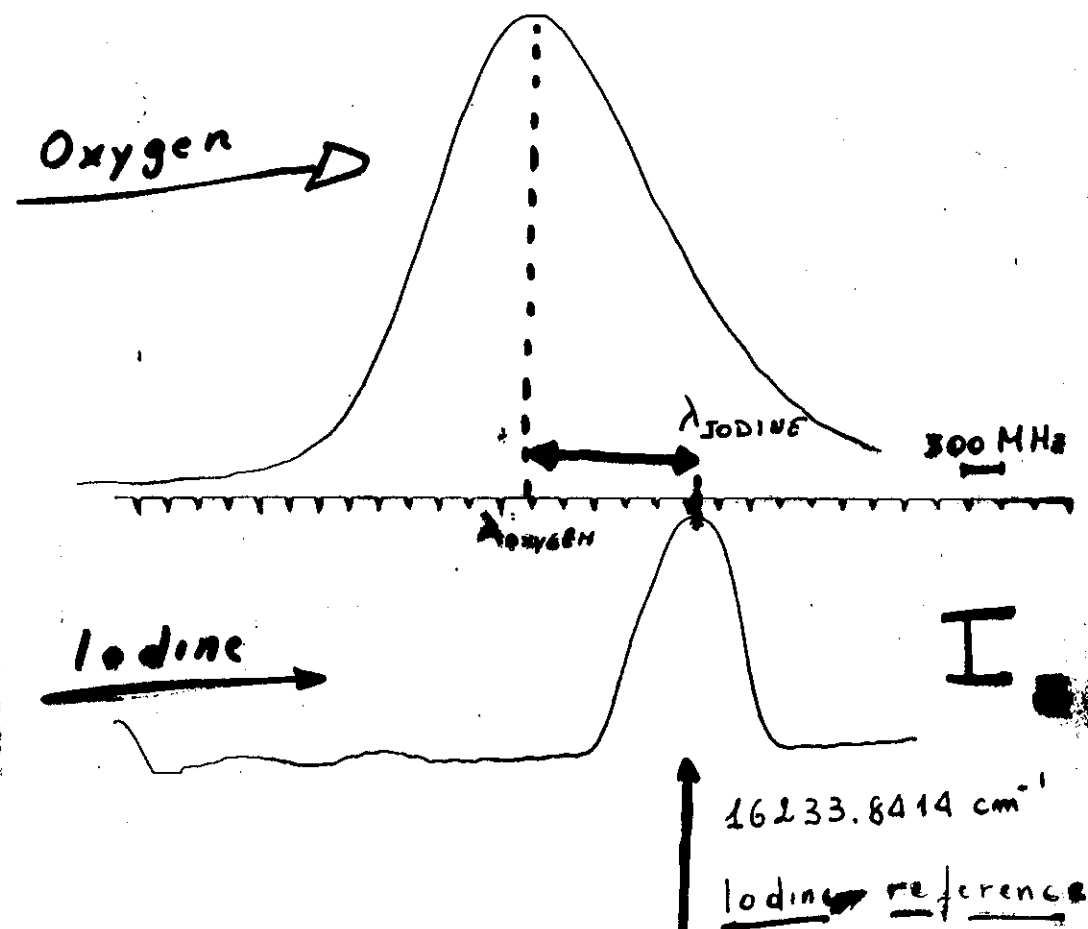
LASER
DOPPLER BEAM



$\pm .008 \text{ cm}^{-1}$
 $\pm 60 \text{ MHz}$

Fig. 15 Resolution in the spectrum of an iodine laser beam at 6158.19 μ

Doppler profile of 6158.19 μ transition



ACCURACY LIMITED BY
DOPPLER WIDTH AND HF STRUCT

DOPPLER + HYPERFINE

42

15 comp J even NEQUAL INTENSITY
21 comp J odd FOR J > 30

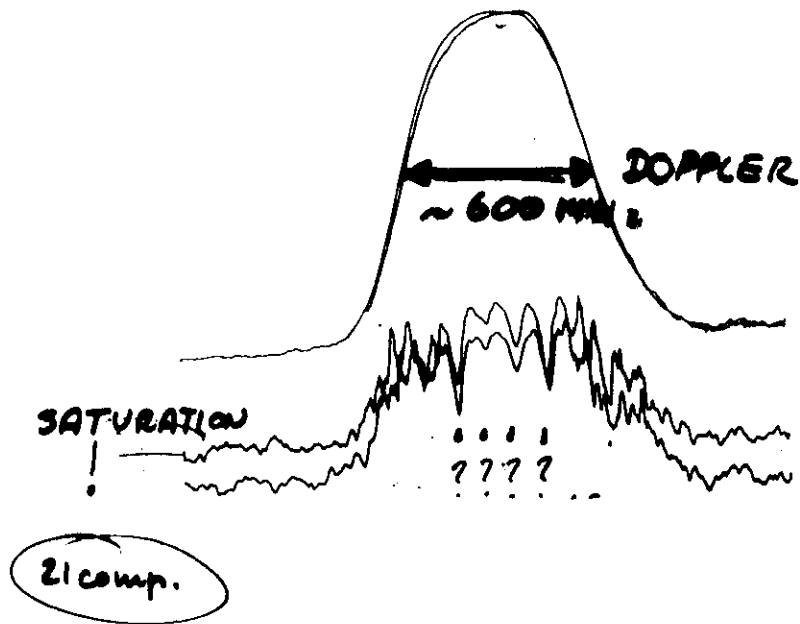


Fig. 22 Struttura iperfina della molecola di Iodio. Profilo di assorbimento con fascio appena saturato.

POLARIZATION SPECTR

43

(SATURATION COMBINED WITH OPTICAL PUMPING).

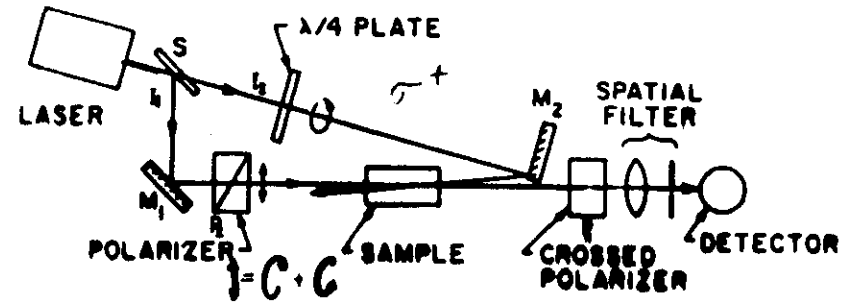
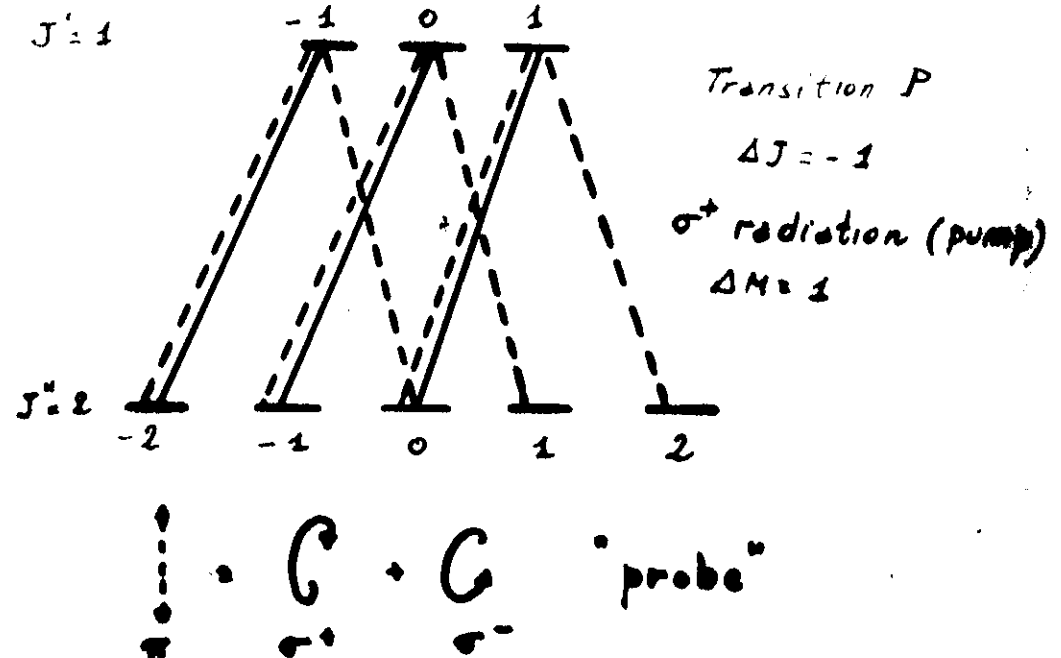
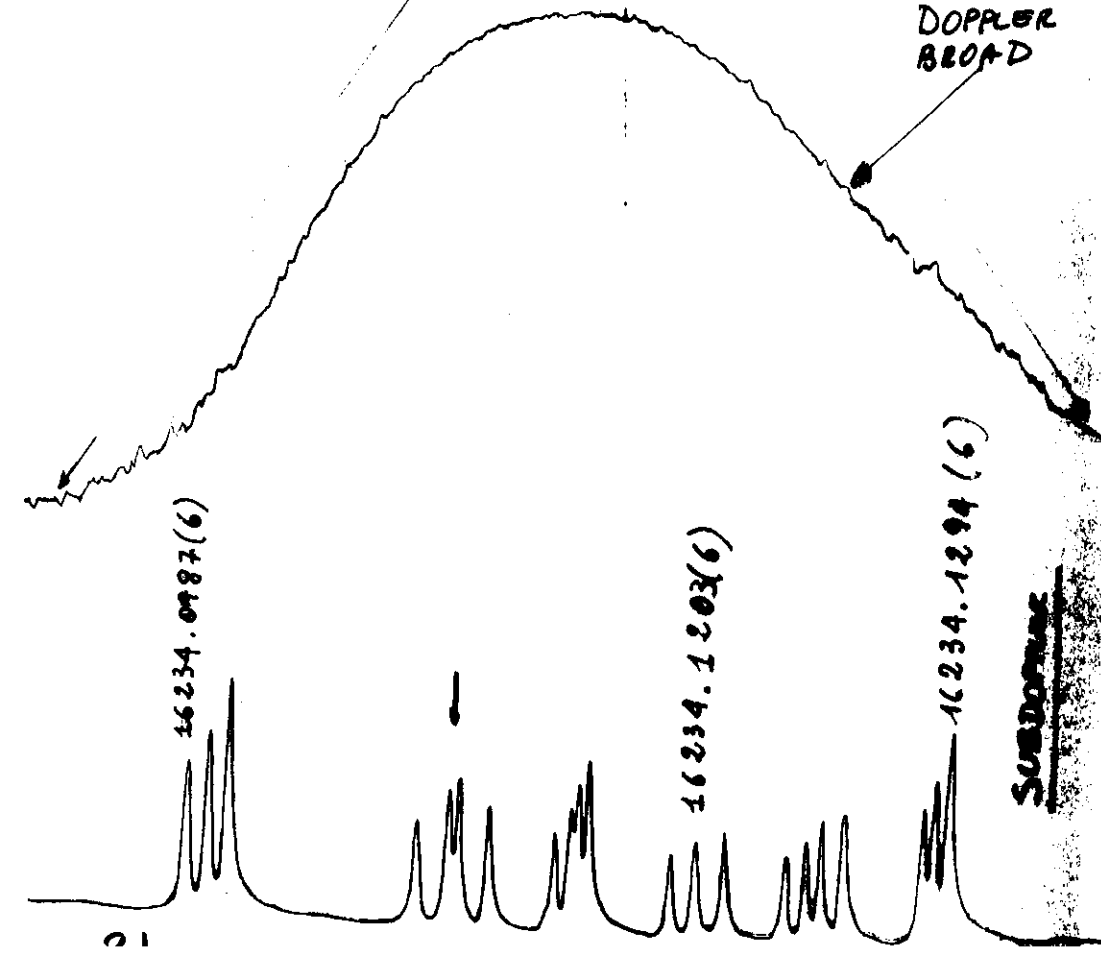


Fig. 10.34. Schematic experimental arrangement for polarization spectroscopy



$$I_T = I_0 \left[\frac{b}{2} \Delta \alpha_0 \frac{x}{1+x^2} + \frac{1}{4} (\Delta \alpha_0 L)^2 \frac{1}{1+x^2} \right]$$

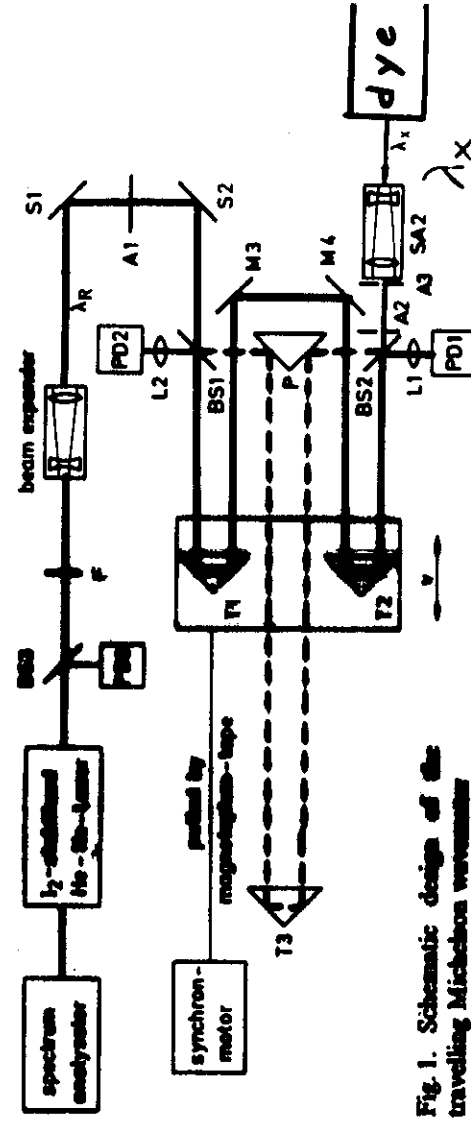


Wave meter

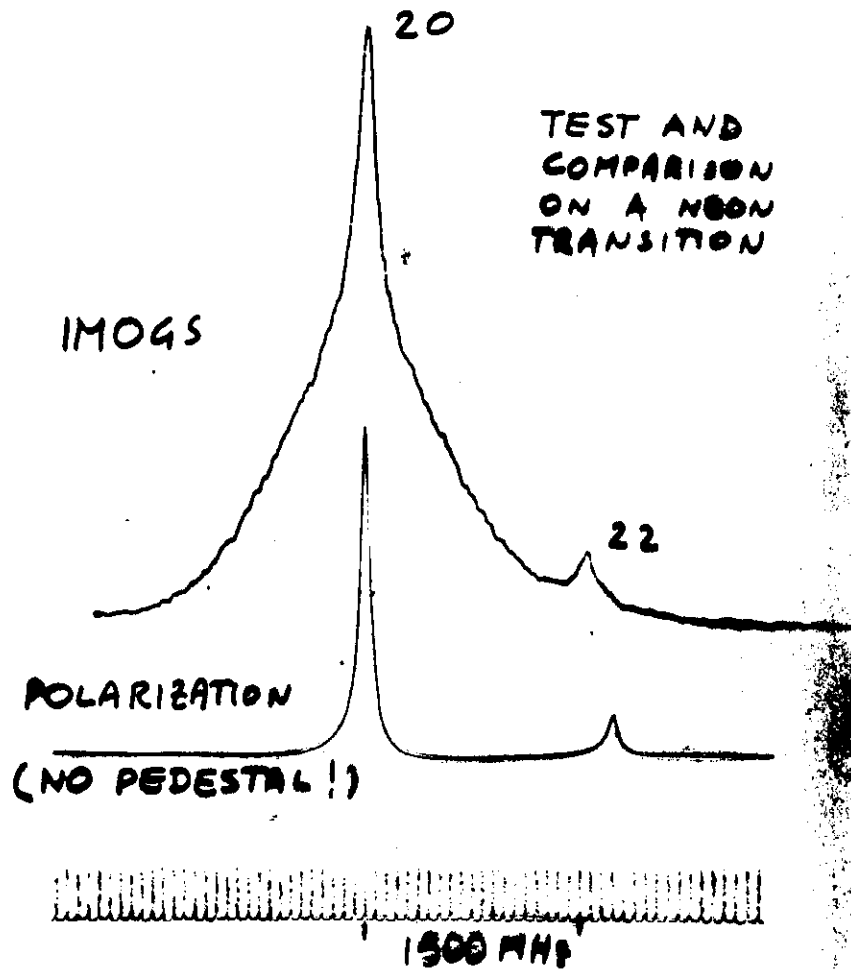
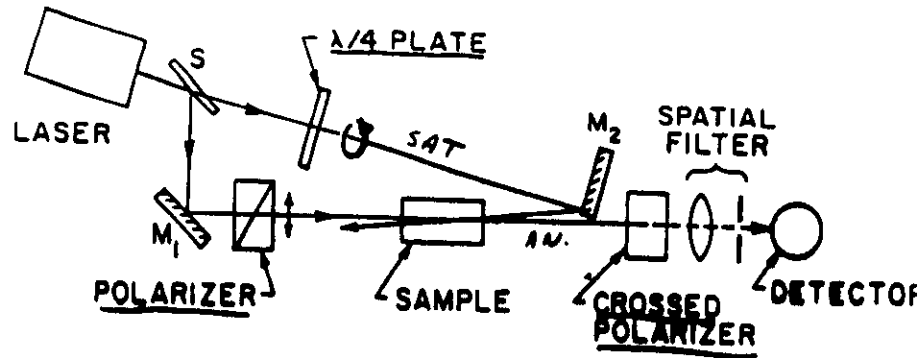
Accuracy : ~1 MHz

$$\lambda_R = 6329.906685 \text{ \AA}$$

2



$$\lambda_X = \lambda_R \cdot \frac{N_R}{N_X} \cdot \frac{n(\lambda_X, P, T)}{n(\lambda_R, P, T)}$$



- INTERMOD. FLUORES.
- INTERMOD. OPTICALLY
- POLARIZ. INTER. EXC.
- POLARIZ. SPECTROSCOPY

Experimental curves for the Ne $2p^5 3s^3 P_2 - 2p^5 3p^3 P_1$ transition (5882 Å), recorded with

- IFS, discharge current (I)=150 mA, pressure (p)=0.6 torr, time constant (τ)=3 s.
- IMOGS, I=45 mA, p=70 mtorr, $\tau=0.1$ s.
- POLINEX, I=45 mA, p=70 mtorr, $\tau=0.1$ s.
- PS, I=45 mA, p=0.1 torr, $\tau=0.3$ s.

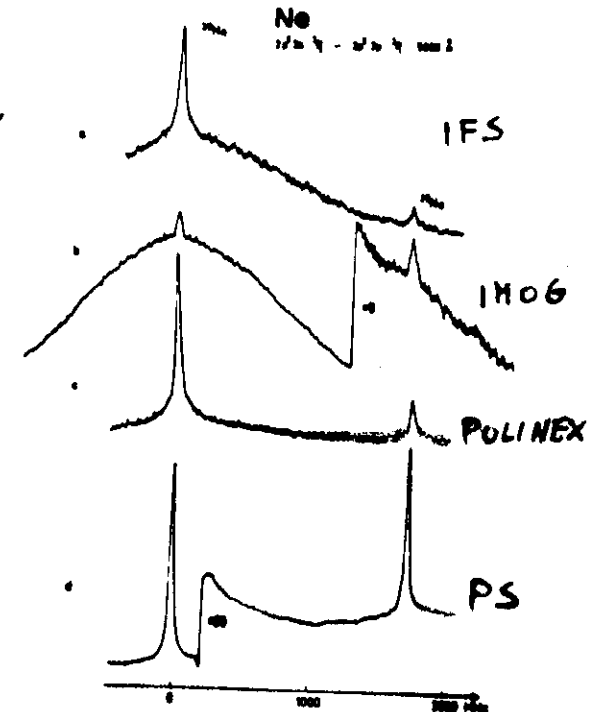


Table I

	Doppler-limited		Doppler-free			
	FS	OGS	IFS	IMOGS	POLINEX	PS
signal/noise	20	100	8	8	30	1500
signal/background	-	-	0.3	0.1	-	-

Belfrage et al. J. Phys. C 7 169 (1963)

ISOTOPE SHIFTS OF OXYGEN

Three stable isotopes:

Natural
Abundance

I

 ^{16}O

99.76 %

 ϕ "DOUBLY
MAGIC" ^{17}O

0.03 %

5/2

 ^{18}O

0.21 %

 ϕ ^{15}O unstable $\tau = 2 \text{ min}$

CLASSICAL SPECTROSCOPY INVESTIGATION

FEBRUARY, 1953

VOLUME 4, NUMBER 2

Isotope Shift in Neutral Oxygen⁴¹LEE W. FORD and JOHN E. MCKENNA
University of Southern California, Los Angeles, California
(Received October 14, 1952)

Fig. 1. Interferometric patterns of the singlet line $\lambda 8446$, showing isotope structure. 24 mm and 8 mm spectra.

P. DAVIES
et al (1978)

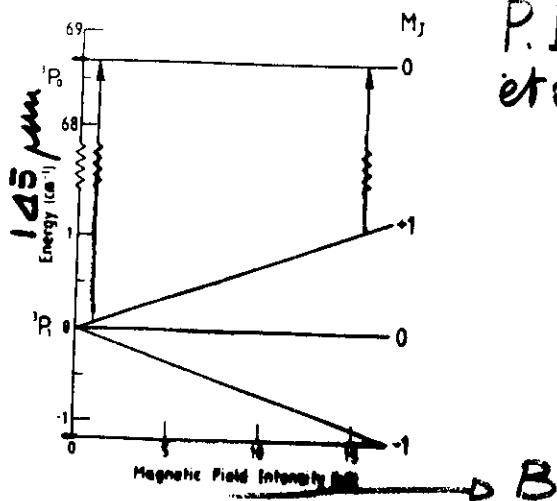
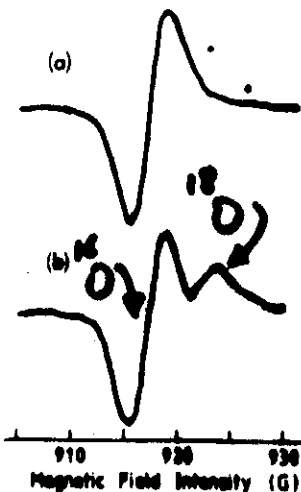


FIG. 1. LMR lines at 917 and 1040 G obtained with the CH_2OD and CH_3NM_2 lasers respectively, and their assignment to the (J, M_J) $0, 0 \rightarrow 1, +1$ transition in $\text{O}(^1P)$.

ISOTOPE SHIFT ON $3P_1 - 3P_0$



natural ^{16}O

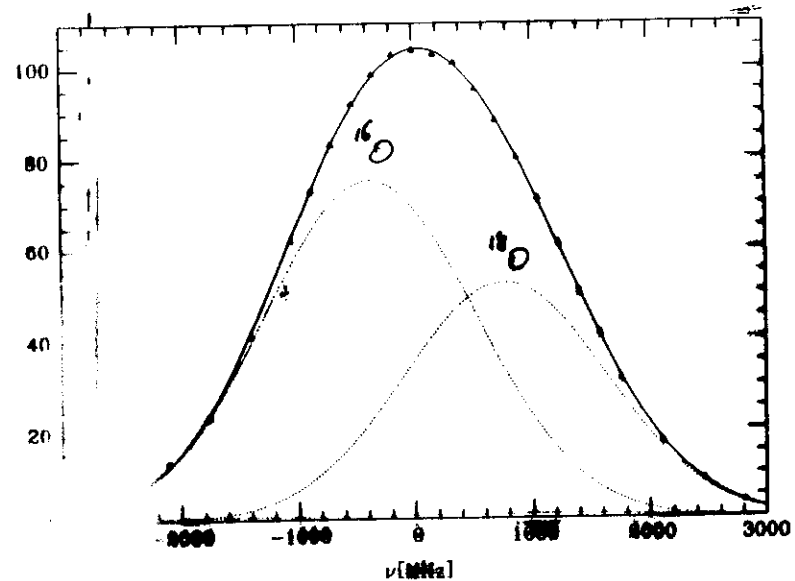
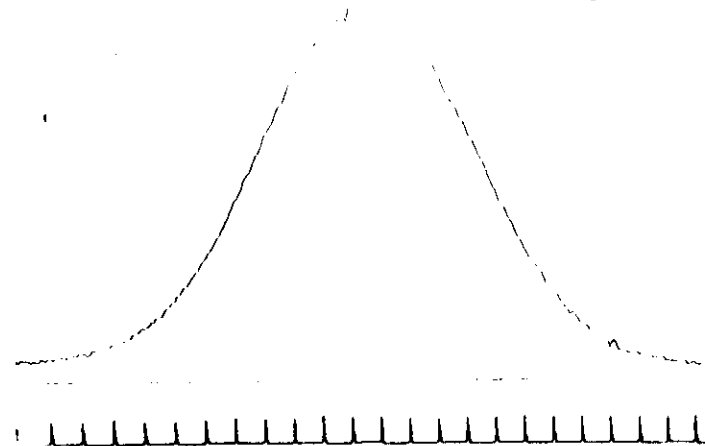
enriched (20%) ^{18}O

$$\Delta V = \frac{c}{\lambda} \frac{m(^{18}\text{O}) (M_2 - M_1)}{(M_2 + m)(M_1 + m)}$$

FIG. 2. LMR spectrum of $\text{O}(^1P)$ with the CH_2OD laser at $145.7 \mu\text{m}$ (a) microwave discharge in $^{16}\text{O}_2$; (b) discharge in O_2 enriched with 20% $^{18}\text{O}_2$.

I.S. = 10.5 MHz

Doppler



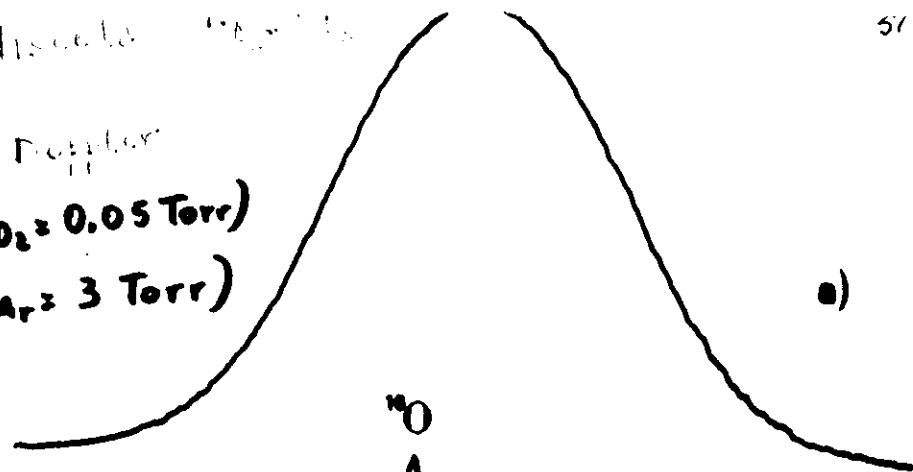
MISCELL

$3P^5P_3 -$

$1 - ^{18}\text{O}$

$\lambda = 6456 \text{ \AA}$

Doppler
($P_{O_2} \approx 0.05$ Torr)
($P_{Ar} \approx 3$ Torr)



Intermodulation

$3p^5P_3 - 4s^5S_2$

$\lambda = 6456.4$



$$\lambda, I. = (1160 \pm 1) \text{ H}_2$$

$$\lambda, I. \Big|_{\text{mass}} \approx 1740$$

S. 21

ISOTOPE SHIFT

Isotope shift is due to two effects

$$I.S. = [\text{MASS EFFECT}] + [\text{VOLUME EFFECT (FIELD)}]$$

MASS SHIFT

$$\Delta \nu_M = \frac{\bar{P}_{\text{NUCL.}}^2}{2 M_{\text{NUCL.}}} = \frac{1}{2 M_{\text{NUCL.}}} \left(\sum_{i=1}^N \bar{P}_i \right)^2$$

$$= \frac{1}{2 M_{\text{NUCL.}}} \sum P_i^2 + \frac{1}{2 M_{\text{NUCL.}}} \sum_{i,j} \bar{P}_i \cdot \bar{P}_j$$

NORMAL MASS SHIFT
NMS

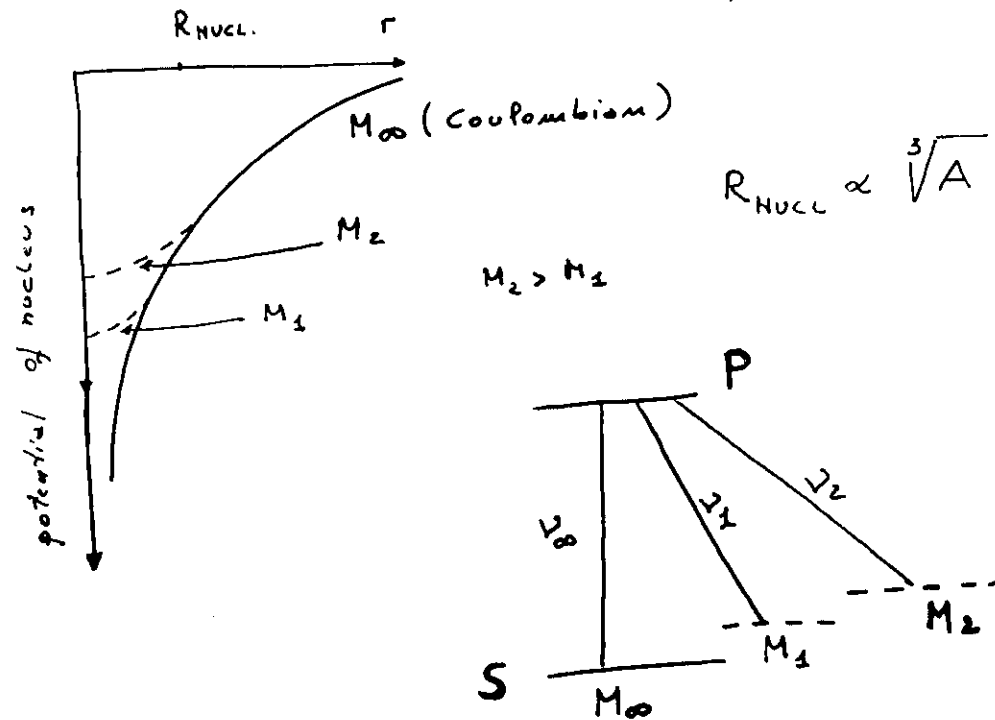
SPECIFIC MASS SHIFT
SMS

$$\Delta \nu_{\text{NMS}} = \frac{c}{\lambda} \frac{m_e}{M_2 \cdot M_1} (M_2 - M_1)$$

"difficult" to calculate
Hartree-Fock
approach + perturb.

ISOTOPE SHIFT

Z atomic number
 A mass number



- S orbital are more penetrating
- For heavier nucleus electron is less bounded.

Volume effect is negligible for light atoms;
 negligible for $A > 60$

KING-PLOT REPRESENTATION

RESIDUAL SHIFT = TOTAL SHIFT -

NORMAL MASS SHIFT

= SPECIFIC MASS SHIFT +
 VOLUME SHIFT

It can be shown that

α, β isotopes

"a" transition

$$\Delta V_R = \Delta V_{RMS} + \Delta V_{VS}$$

$$= K^a \cdot A_{\alpha\beta} + E^a C_{\alpha\beta}$$

where

$$K^a \propto \sum \bar{P}_i \cdot \bar{P}_j$$

$$A_{\alpha\beta} = \frac{A_\beta - A_\alpha}{A_\beta A_\alpha}$$

$$E^a \propto |\psi(0)|^2$$

$$C_{\alpha\beta} \propto \langle r^2 \rangle_\beta - \langle r^2 \rangle_\alpha$$

KING PLOT

MODIFIED SHIFT

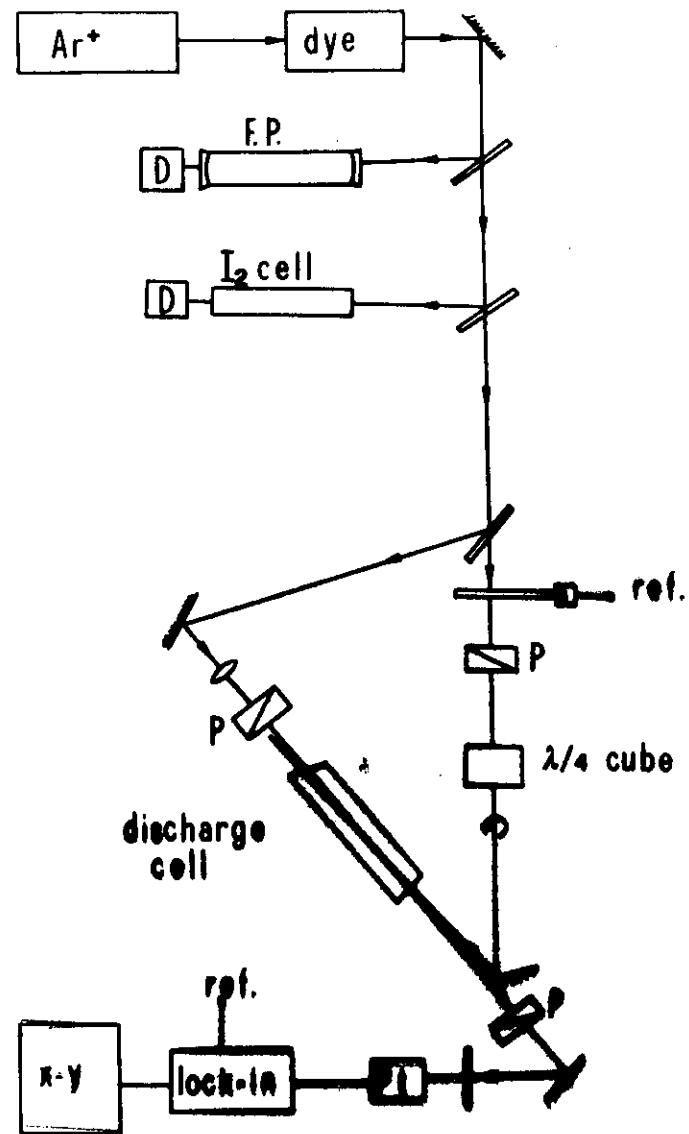
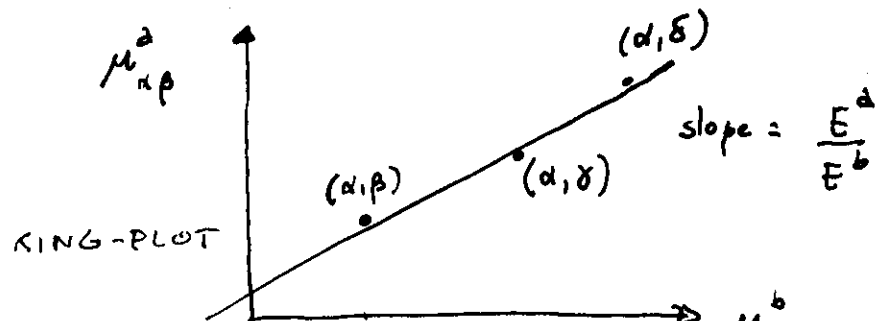
$$\mu_{\alpha\beta}^a = \frac{\Delta V_{\alpha\beta}^a}{A_{\alpha\beta}} = \Delta V_{\alpha\beta}^a \frac{A_\alpha A_\beta}{A_\beta - A_\alpha}$$

$$\mu_{\alpha\beta}^a = K^a + E^a \frac{C_{\alpha\beta}}{A_{\alpha\beta}} \quad (1)$$

$$\mu_{\alpha\beta}^b = K^b + E^b \frac{C_{\alpha\beta}}{A_{\alpha\beta}} \quad (2)$$

By combining eq. (1) and (2) we have:

$$\mu_{\alpha\beta}^a = \frac{E^a}{E^b} \mu_{\alpha\beta}^b + \left(K^a - K^b \frac{E^a}{E^b} \right)$$



Experimental set-up
for polarization spectroscopy

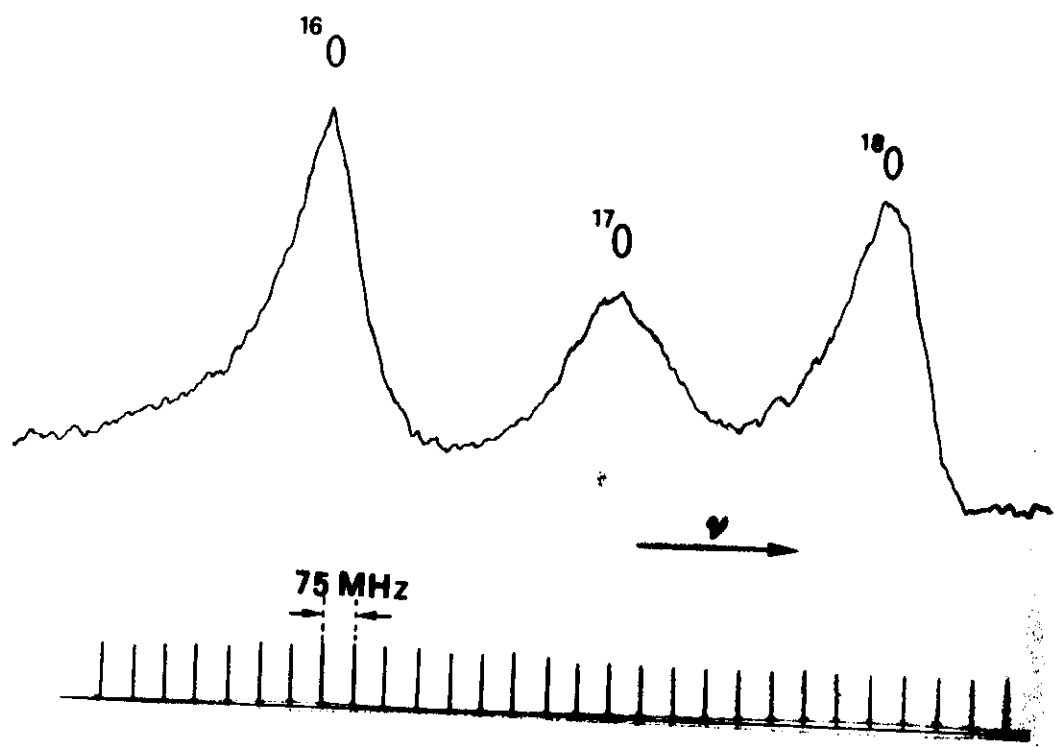
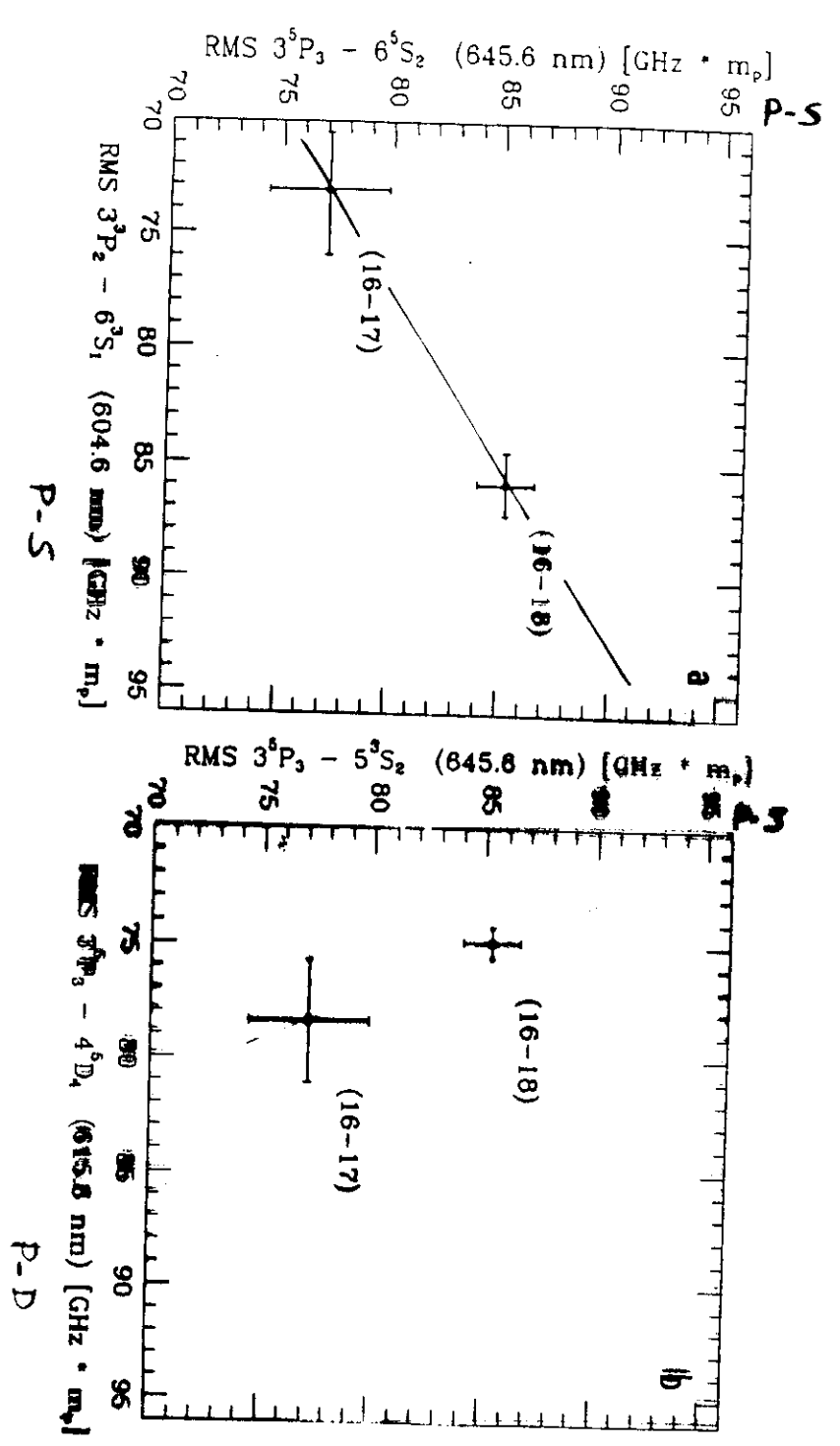


fig. 8.3. Registrazione della transizione $3p^2P_1 \rightarrow 4d^1D_2$ dei tre isotopi dell'ossigeno. Scarica a radiofrequenza 1.2 torr di Ar e 0.08 torr di O_2 .

Fig. 5



Nuclei "Magic" 59

INTEREST FOR ISOTOPE
SHIFT INVESTIGATION
OF LIGHT ELEMENTS

"MAGIC" NUCLEI

SHIFT = NORMAL MASS +
SPECIFIC MASS +
FIELD EFFECT

Ni *
⁴⁰Ca * *
¹⁶O * *
⁴He * *



Transition	Wavelength (nm)	Isotopic shift (percent work) (nm)	Previous data [ppm] (nm)
3p 5p ₃ - 4d 5D ₄	615.8	11000 (40)	< 800
3p 5p ₂ - 5s 5S ₂	645.5	11000 (20)	—
3p 5p ₃ - 5s 5S ₂	645.6	11000 (20)	< 600
3p 3p ₂ - 3s 3S ₁	604.6	11000 (40)	—

L. B. Fother and J. R. Holman
J. Opt. Soc. Am. - 43, 103 (1953)

(82)

1 h 11/2

2 d 3/2

3 s 1/2

2 d 5/2

1 g 7/2

(50)

1 g 9/2

2 p 1/2

2 p 3/2

1 f 5/2

(28)

1 f 7/2

(20)

1 d 3/2

2 s 1/2

1 d 5/2

(8)

1 p 1/2

1 p 3/2

(2)

1 s 1/2

¹⁶O : 99.76 %

¹⁷O : 0.03 %

¹⁸O : 0.20 %

DIODE LASER

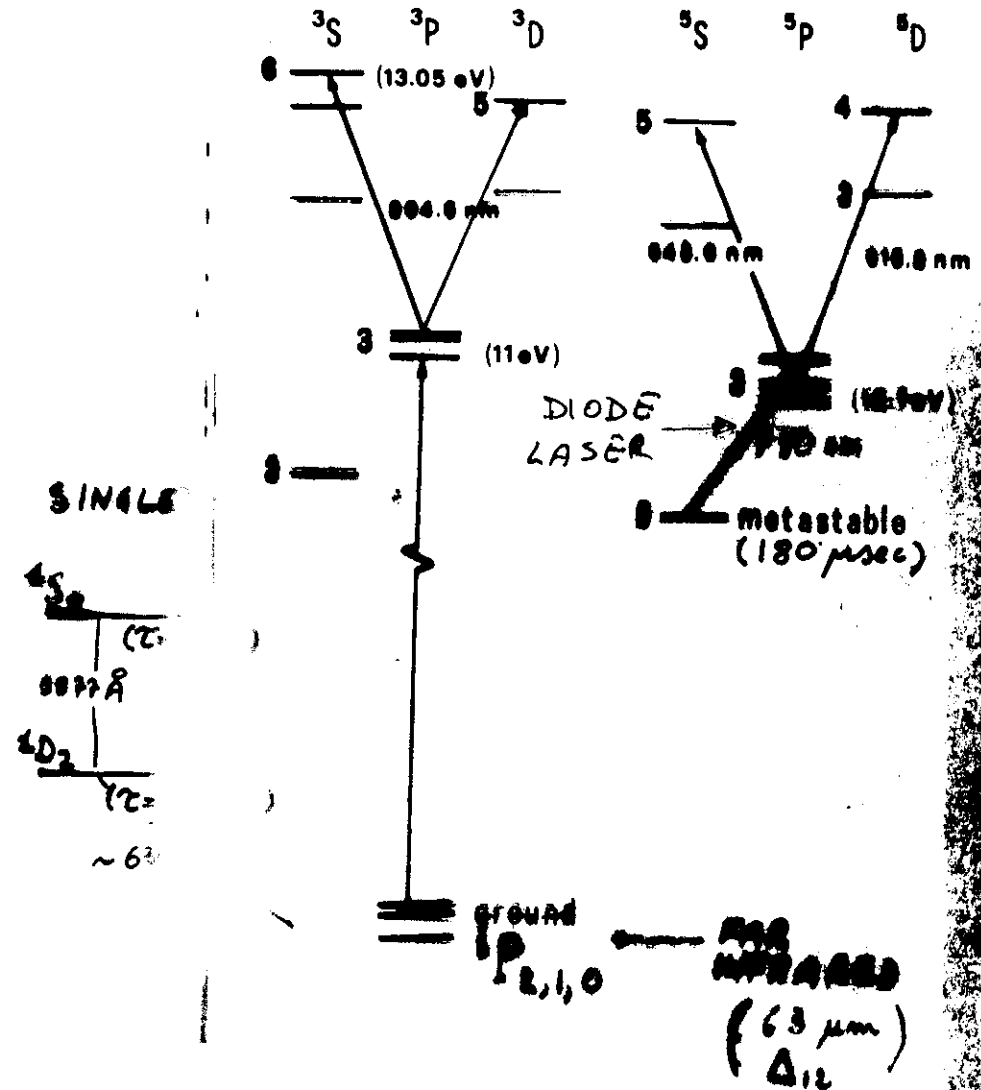
61

- WIDE SPECTRAL RANGE $0.6 \div 0.9 \mu\text{m}$
 $1.3 \div 16 \mu\text{m}$
 $2 \div 30 \mu\text{m}$
- LOW NOISE
- TUNABILITY (λ) (Temperature, pressure, current)
- BAND WIDTH $10 \text{ KHz} - 10 \text{ MHz}$
- POWER $\approx 10 \text{ mW}$
- ELECTRICAL POWER $\approx 100 \text{ mW}$
- CHEAP
- SMALL SIZE

OXYGEN

TRIPLETS

QUINTETS



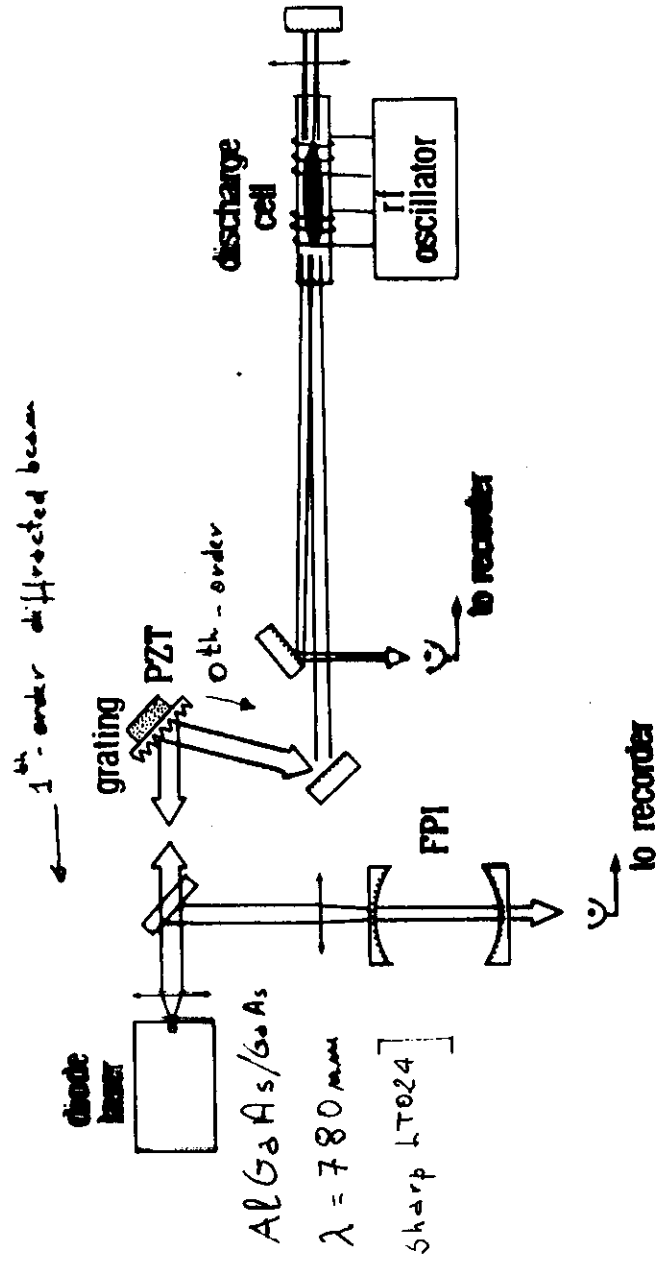
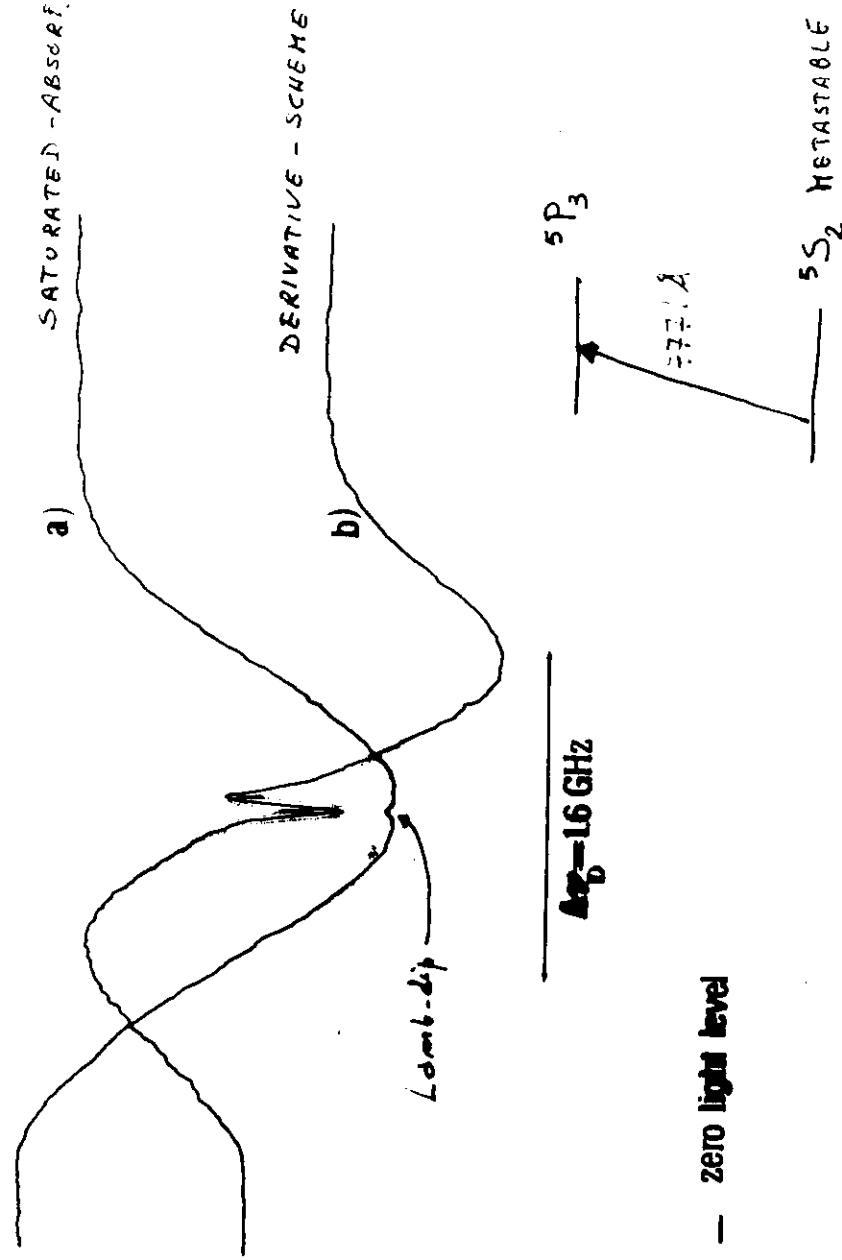
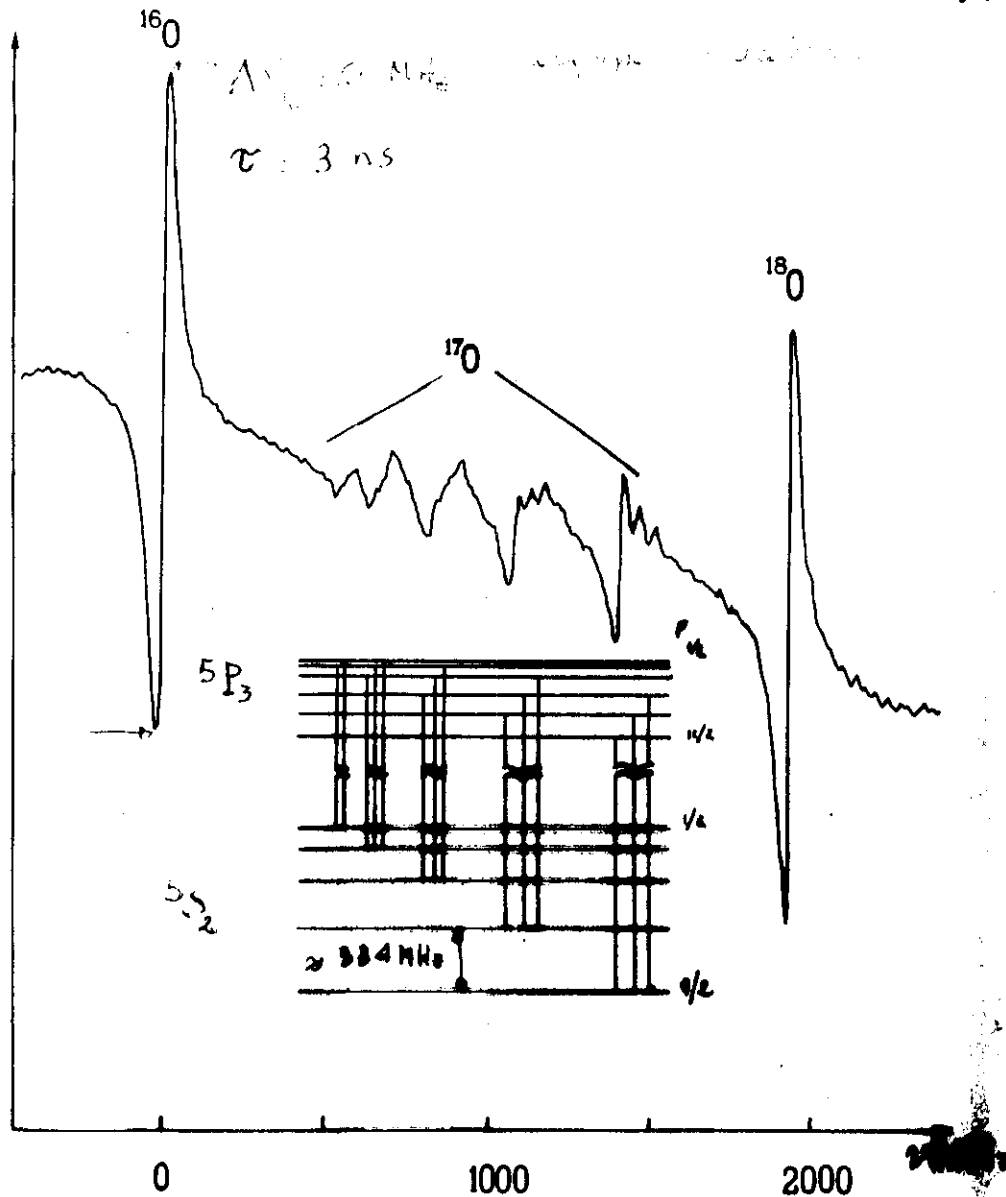


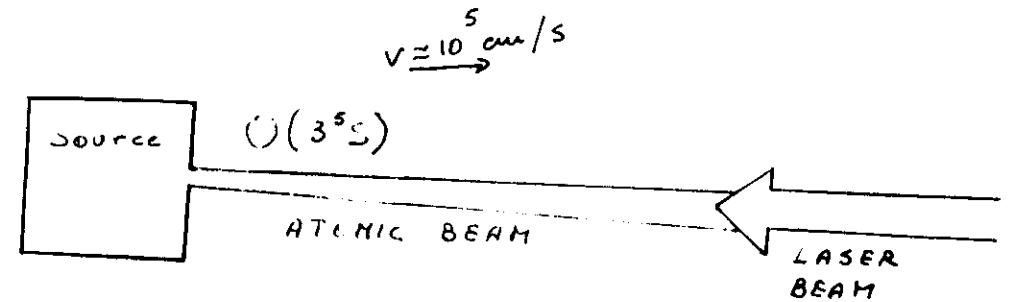
Fig. 2



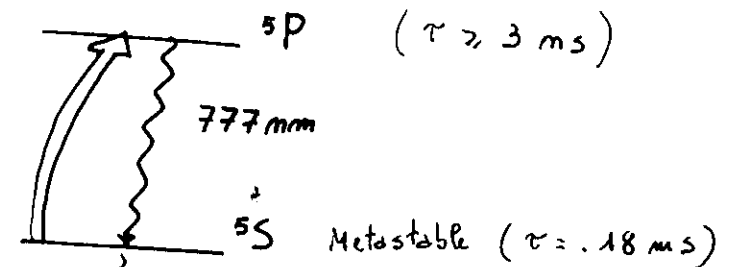


HFS constant $A = 74.3(2) \text{ MHz}$

LASER COOLING OF ^{17}O ATOMS



$$\Delta v = \frac{h\nu}{cM}$$



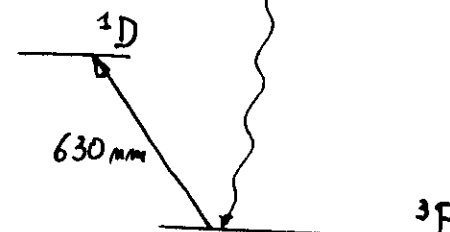
Intercombination
 $\lambda = 130 \text{ nm}$

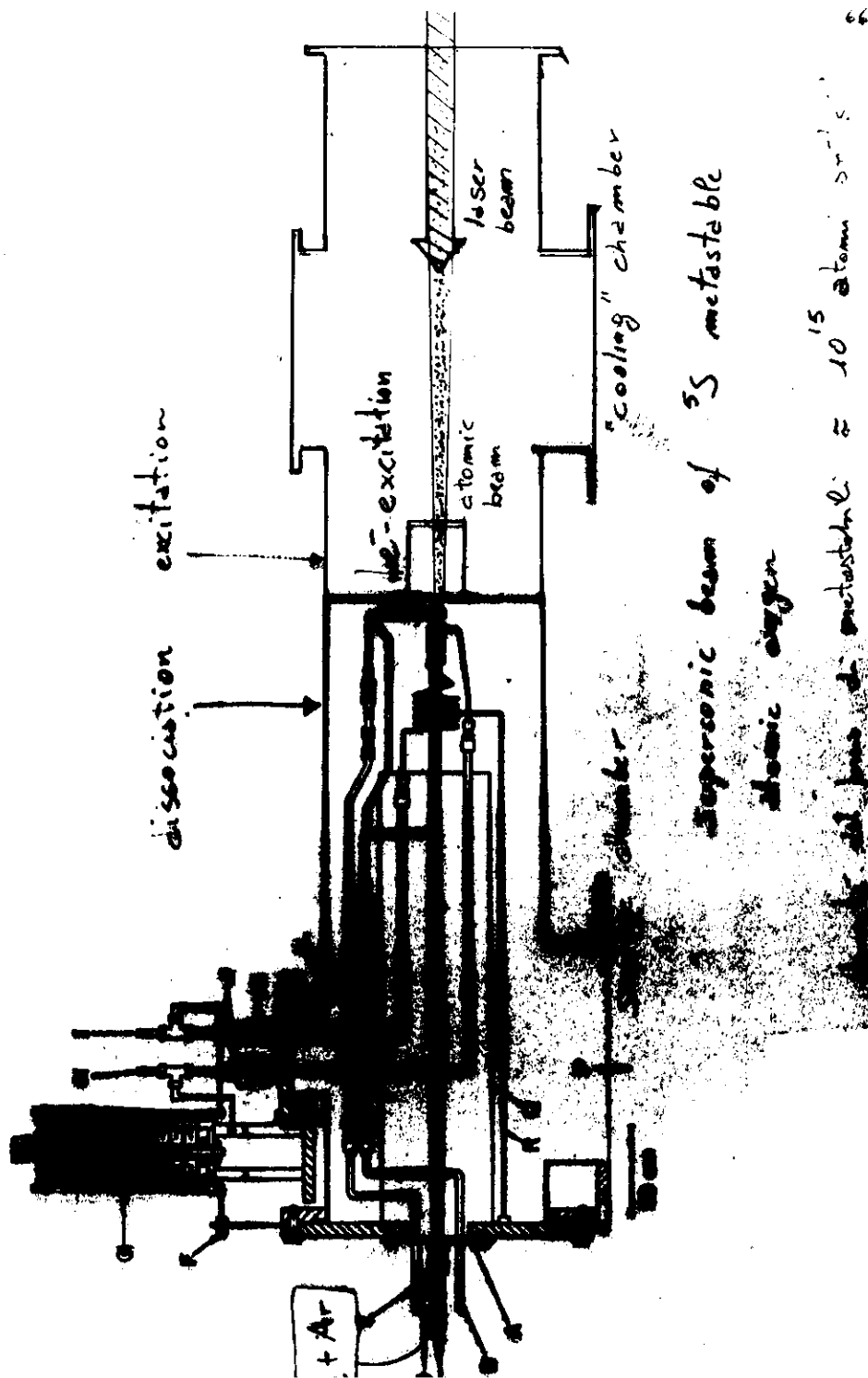
$$N_{\text{cycl}} \approx 45000$$

$$\tau_{\text{cooling}} \approx .27 \text{ ms}$$

Advantages

- CLOSED TRANSITION
- "COLD" ATOMS IN THE GROUND STATE





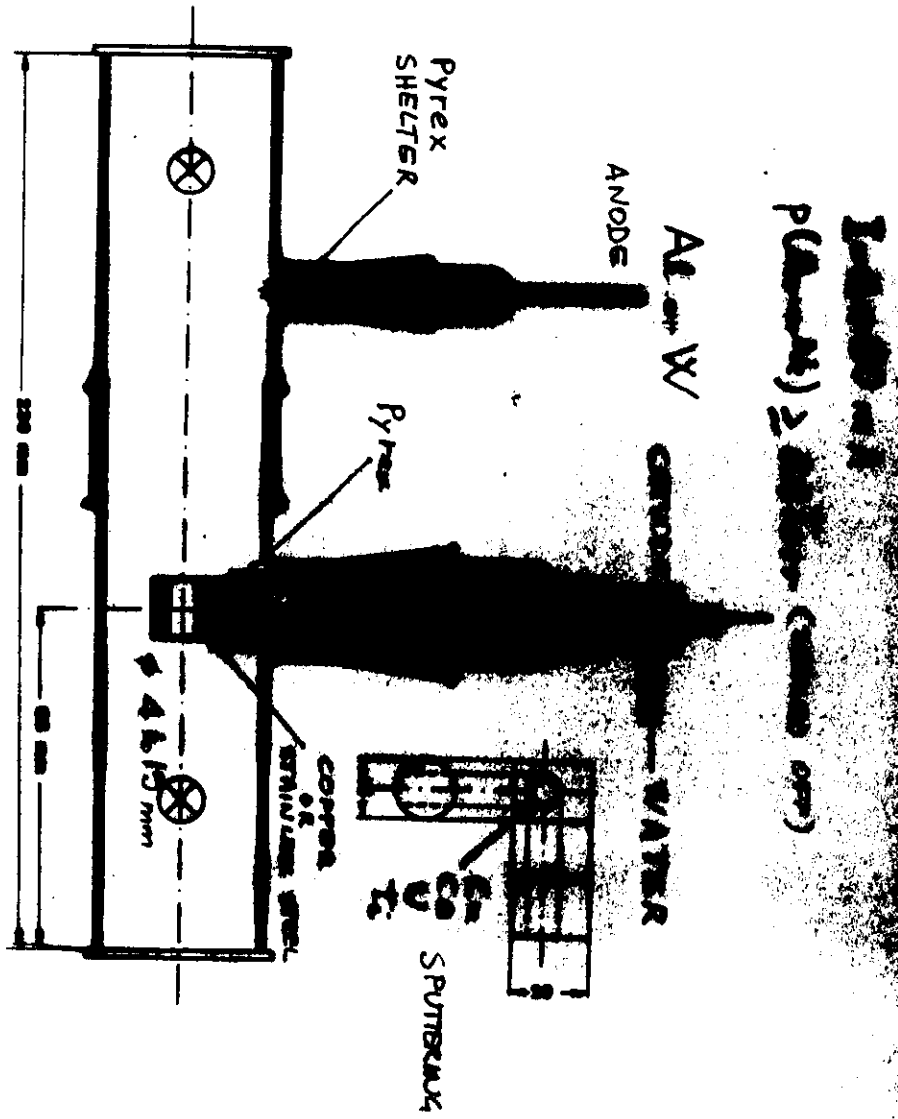
SPECTROSCOPY OF REFRACTORY ELEMENTS :

Calcium : Ca , Ca^+

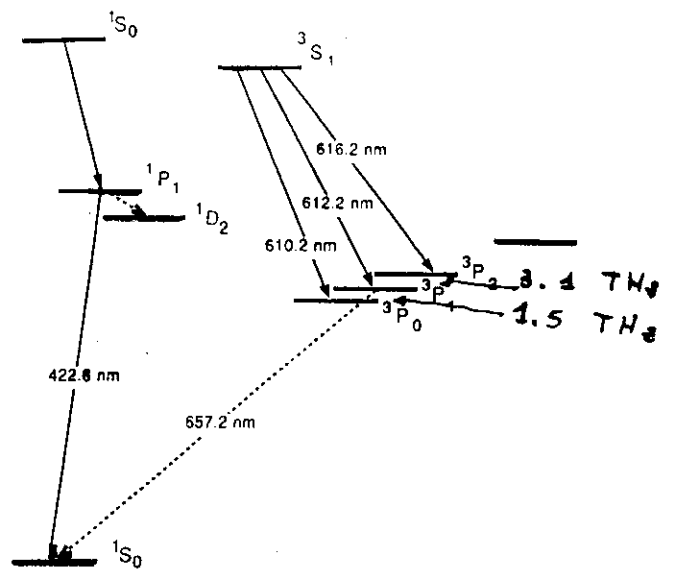
Titanium : Ti , Ti^+

Interesting for : astrophysics
metrology

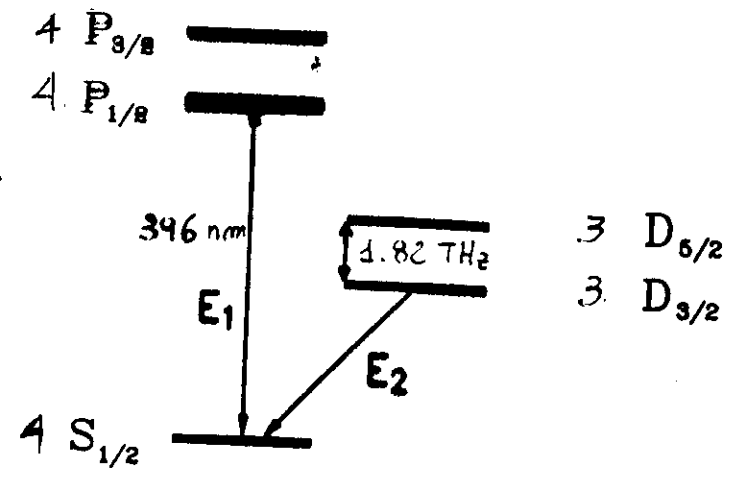
Production : hollow cathode discharge
"sputtering"



Ca



Ca⁺



FABRY-PEROT MARKER

MARIELLA
SCHISANO

 $3P_2 - 3S_1 C_d$

2:6162A

h.f.: I_2

35

Contents

८

4

6162 A

6122 A

6102 Å

32

40

+ T

48 T_i

648. 47

386.874

170.152

600

431 F

 T_4^* $4p\ ^3D'$

4p 'G'

22.14

4354

3354

384.19

300.6

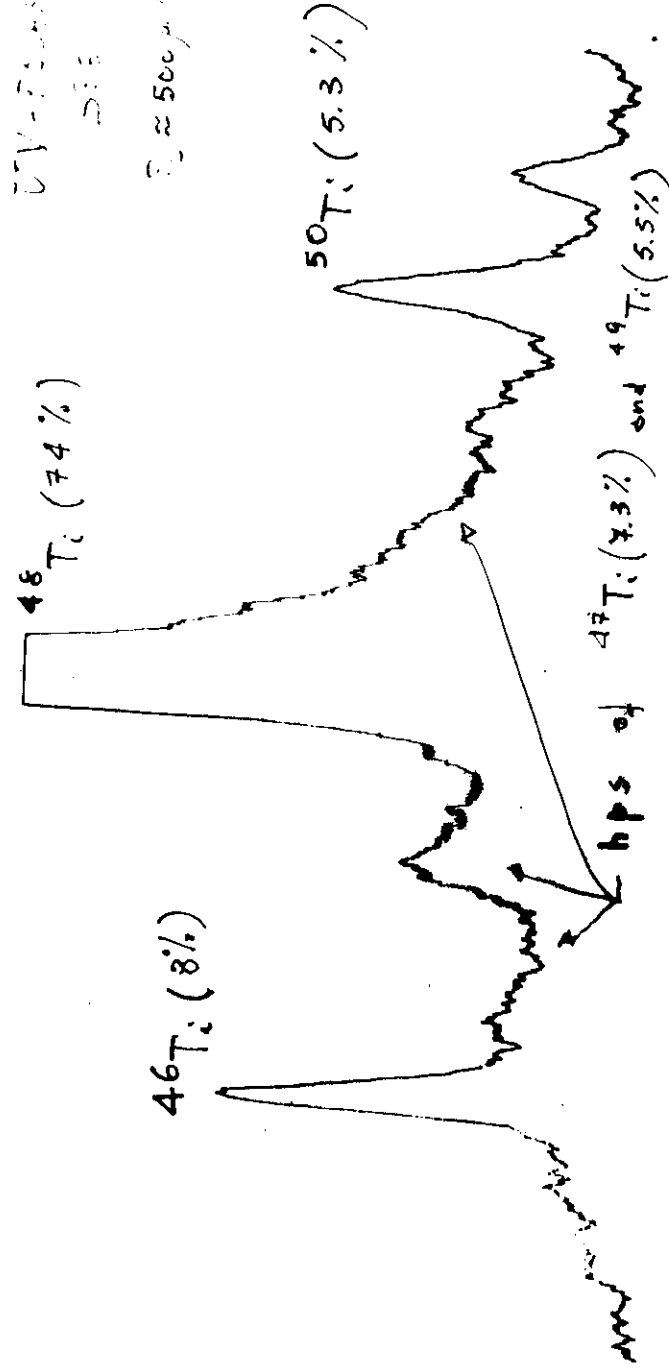
857.75

887.04

_____ 2 1 - 5TH.

Titanium: $4^3F_2 - 4p^3D_3$

$\lambda: 334.14 \text{ nm}$



150 MHz

