



INTERNATIONAL ATOMIC ENERGY AGENCY
UNITED NATIONS EDUCATIONAL, SCIENTIFIC AND CULTURAL ORGANIZATION
INTERNATIONAL CENTRE FOR THEORETICAL PHYSICS
I.C.T.P., P.O. BOX 586, 34100 TRIESTE, ITALY, CABLE: CENTRATOM TRIESTE



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H4.SMR/645-10

140-1

**SCHOOL ON PHYSICAL METHODS FOR THE
STUDY OF THE UPPER AND LOWER
ATMOSPHERE SYSTEM**

26 October - 6 November 1992
Miramare - Trieste, Italy

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*Middle Atmosphere Measurements to Study
Variability at Different Scales*

**MIDDLE ATMOSPHERE MEASUREMENTS TO STUDY
VARIABILITY AT DIFFERENT SCALES**

140-1

Lecture 1: Ground based observations of the dynamics of the Middle Atmosphere.

Lecture 2: Long term trends in the Middle Atmospheric temperature.

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140-1

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LECTURE 1

The study of dynamics at a variety of scales with periods ranging from minutes to hours, days and years requires continuous monitoring of the atmosphere. Sporadic measurements as the ones given by balloons and rockets are biased towards long periods, while continuous long term ground-based monitoring, whether it uses passive or active techniques, is able to provide locally access to a large range of periods. I will not cover the possibility offered by satellites, as the subject is covered by the lectures of Professor M. Geller. After a brief presentation of radars and passive techniques based on the observation of natural emissions, I will present the more recently developed lidar methods, with specific emphasis on their possibilities to study dynamics: planetary waves, tides, gravity waves, stratosphere-troposphere exchange.

LECTURE 2

The middle atmosphere is the place where the largest changes of anthropogenic origin have been observed in the last decades. The discovery of the Antarctic Ozone hole is the best known of such changes. This lecture will review the trends observed in the temperature, partly under the influence of the ozone depletion, and partly as a consequence of the increase of greenhouse gases, CO₂ and H₂O. The observations will be compared with what is expected from the measured changes in atmospheric composition, using different models. In both the stratosphere and the mesosphere, the temperature has decreased faster than predicted by any model, and the fact that while the temperature is expected to increase in the troposphere, the cooling observed just above the tropopause could be a major source of perturbation in the wave propagation between the two layers, change which could have an impact on the general circulation and therefore on the climate.

To conclude this lecture, I will mention the existence of a new project within the World Climate Research Programme, on "Stratospheric Processes and their Role in the Climate" and describe its content. The SPARC project was decided in March 1992; it will aim to a better understanding of how the stratospheric changes can influence the climate, using the coordinated effort of the worldwide atmospheric community.

Anyone who wishes to be on the mailist list to receive the SPARC Newsletter and all relevant informations, should give his (her) name to Dr M.L.CHANIN during the School or later write to:
SPARC Project Office
B.P.3, 91370, Verrières-le-Buisson, FRANCE

LONG TERM TREND IN THE MIDDLE ATMOSPHERE TEMPERATURE

M-L. CHANIN

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I Model Predictions

II Temperature and O₃ trends in the low stratosphere.

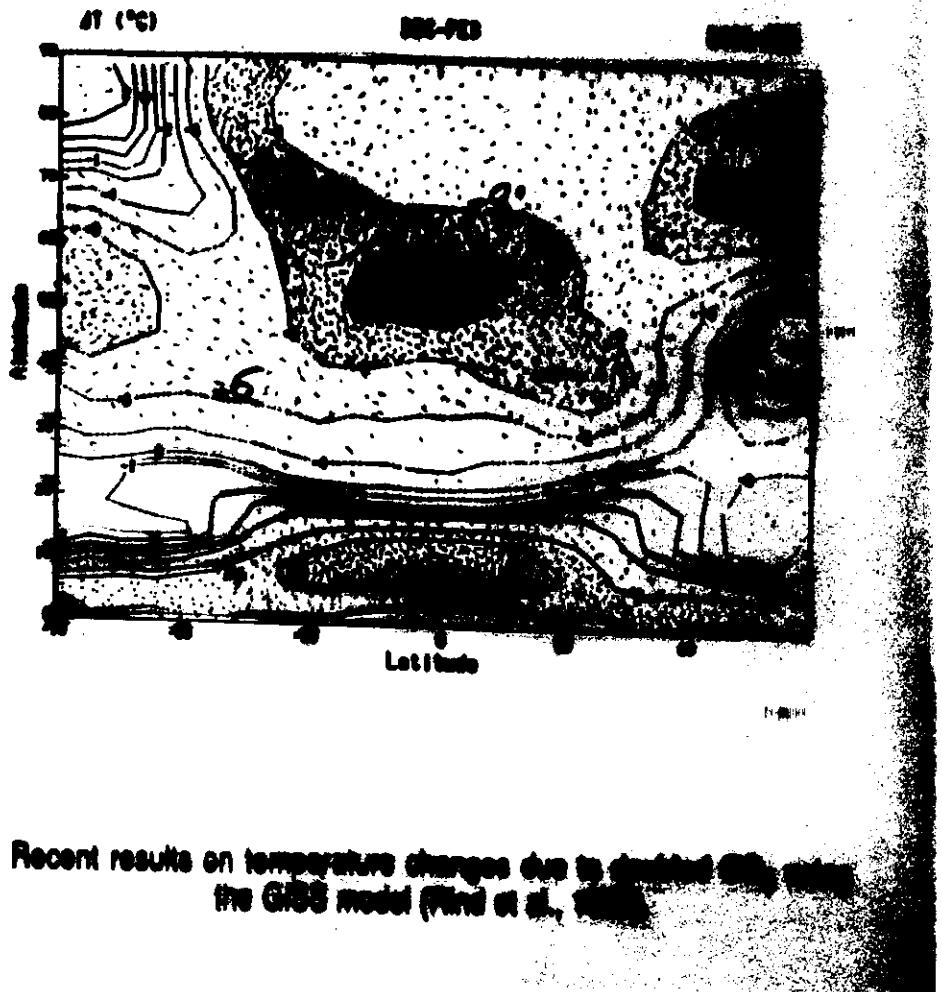
III Temperature trends in the mesosphere.

IV Comparison with Models.

V Conclusion

I

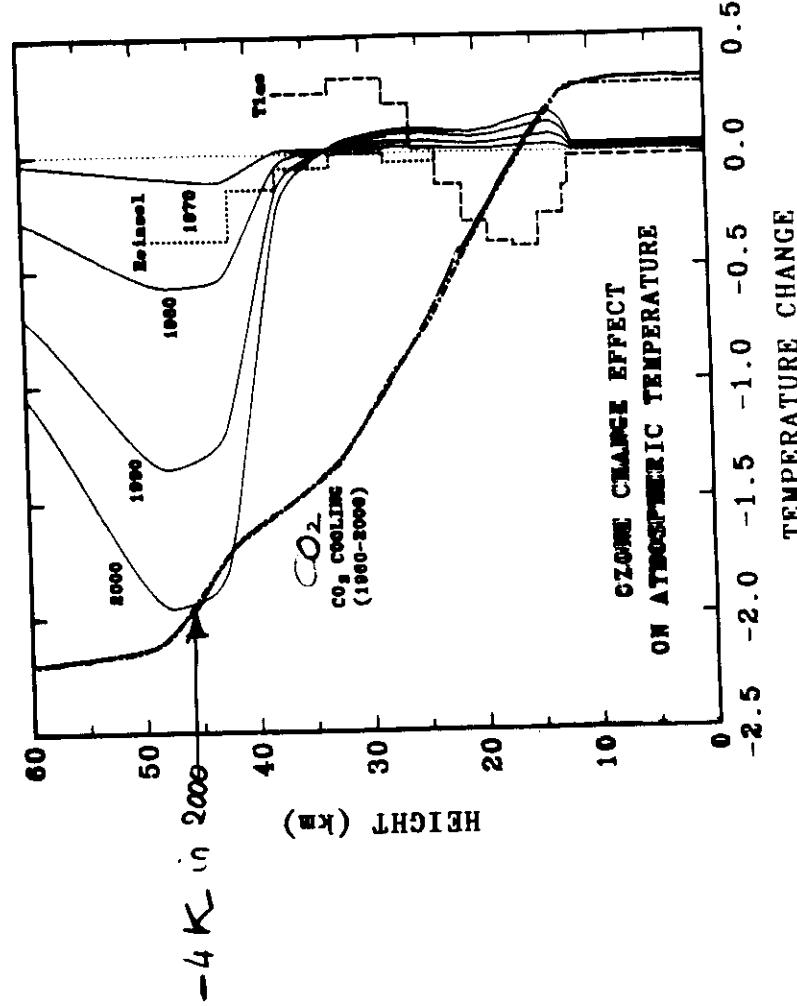
Results from the GISS model showing the warming of the troposphere (max: +8 K) and the cooling of the middle atmosphere (max: -10 K around 50 Km) for a doubling of CO₂ -



Recent results on temperature changes due to a doubling of CO₂ using the GISS model (Rind et al., 1990)

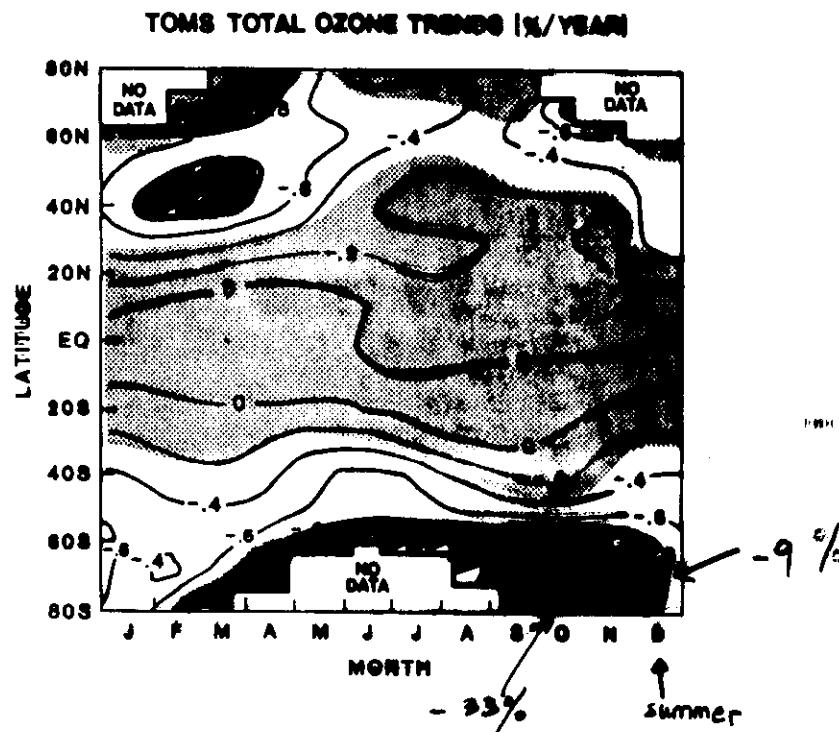
From Rind and Lacis in "Survey in Geophysics" 1992

CO₂ and O₃ forcing of the temperature.



To model more realistically the temperature trends, one should know what has been happening in the Ozone in the last decade.

This is the change observed in Total Ozone for a 11 year period.

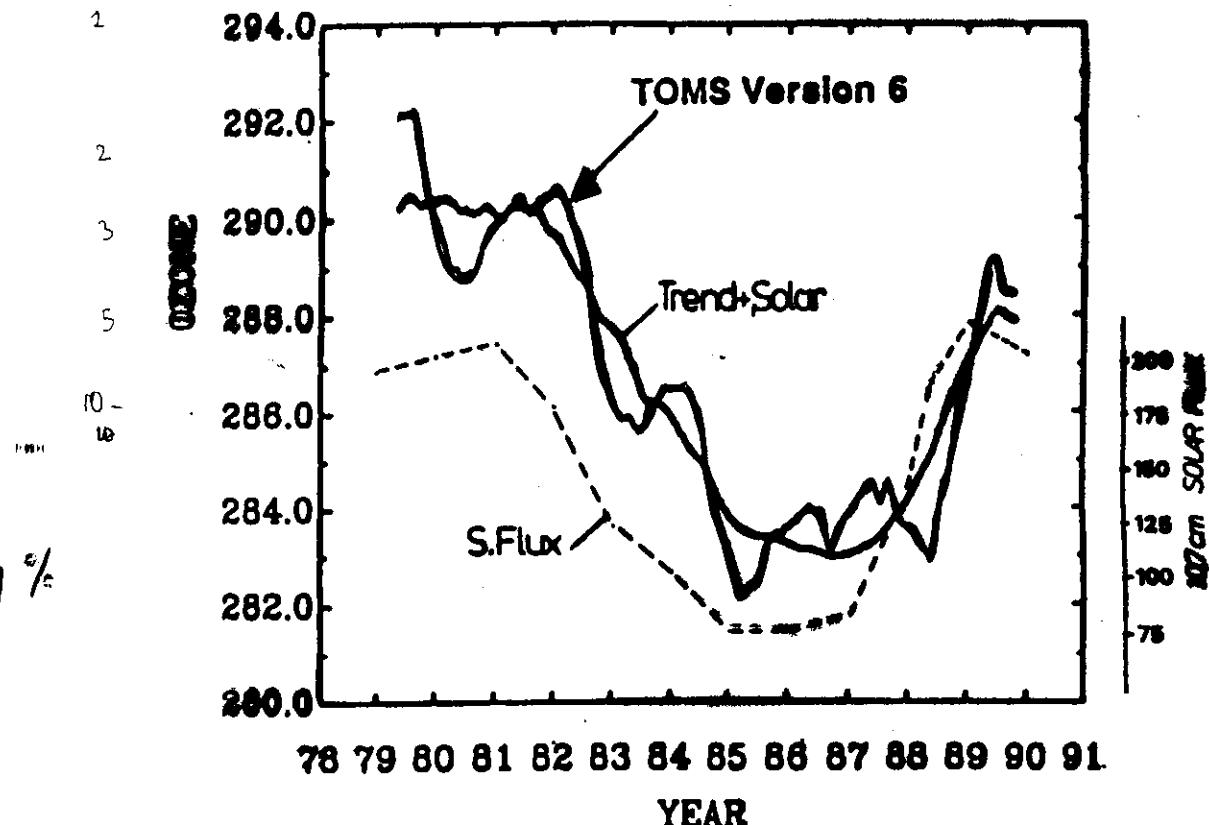


Stolarski et al GRL 1991

for 11 years of data

One should be aware that the Ozone is modulated by the 11 year solar cycle and therefore a full period of 11 years is a minimum to look at trends.

TIME SERIES OF TOMS OZONE COLUMN DENSITY 40S-40N USING 365-DAY RUNNING MEAN



Change in O₃ vertical profile from different satellites (SAGE, SBUV) and ground-based measurements (Umkehr) compared to models.

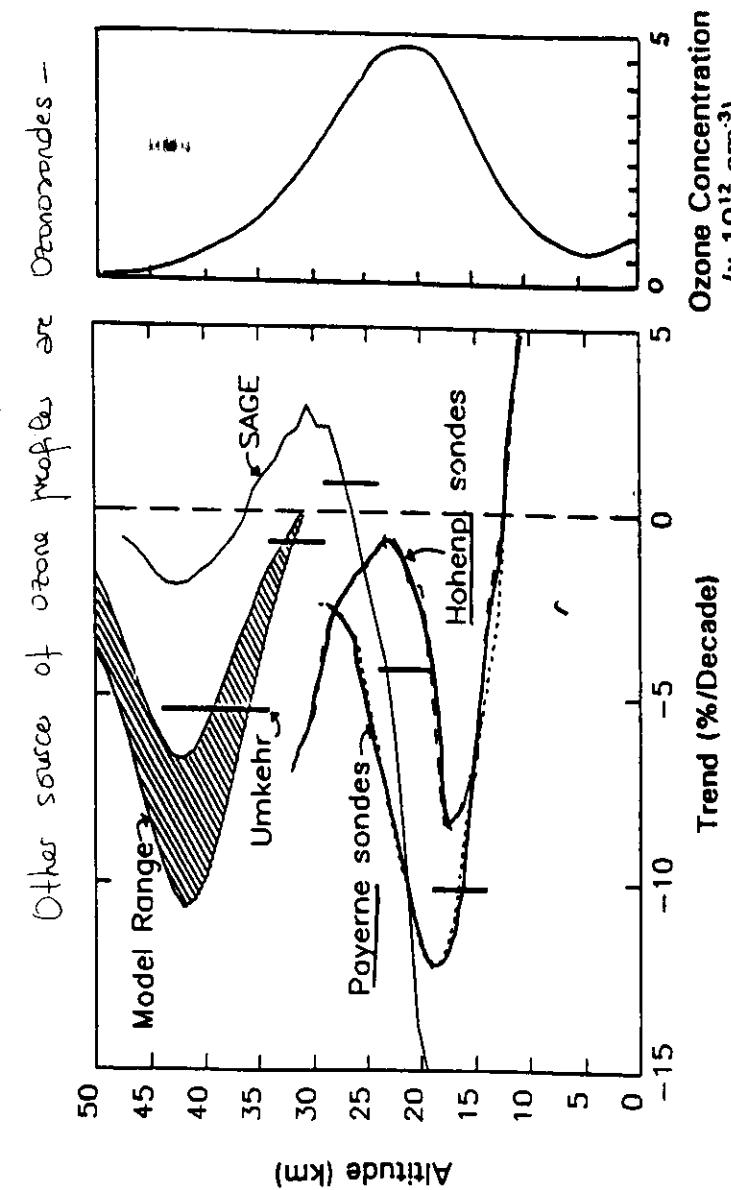
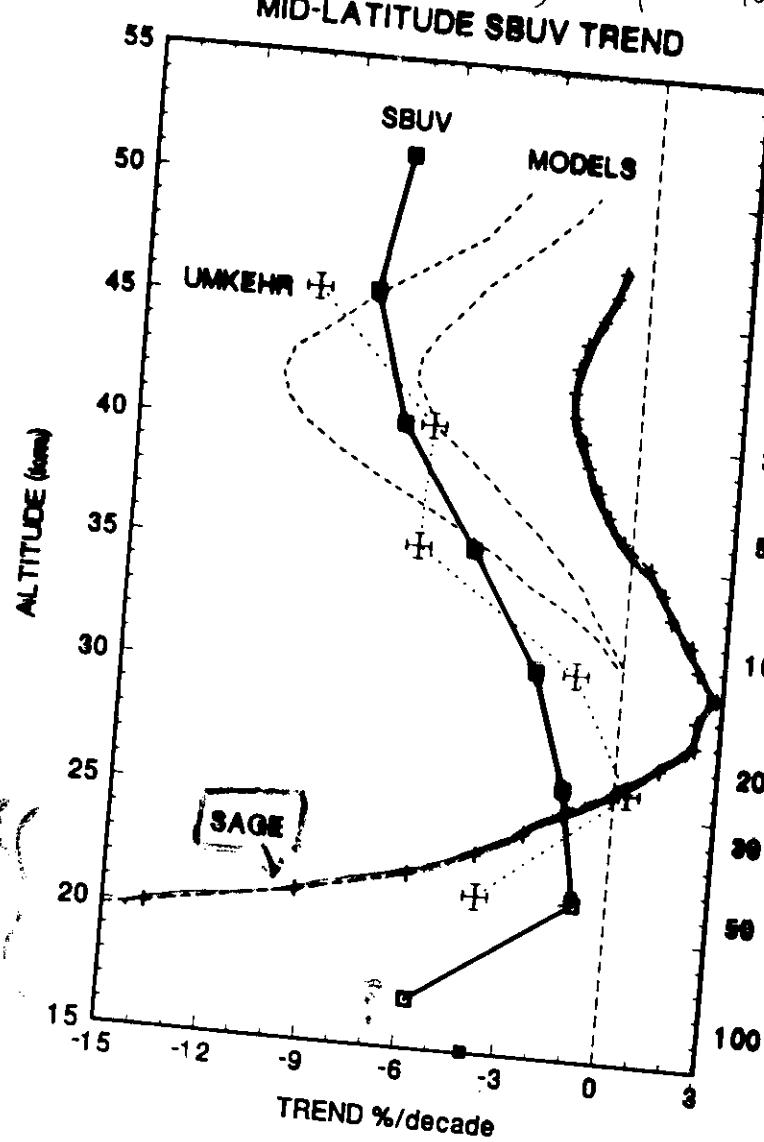


Figure 2-18: Comparison of ozone profile trend estimates from several measurement systems, SAGE, Umkehr, and two ozonsonde stations. SAGE data is an average over the latitude ranges 20-50°N and 20-50°S. The Umkehr is the average over 5 northern midlatitude stations. Shaded area shows the range of two model calculations at 50°N and 50°S. The panel on the right shows a typical ozone concentration versus altitude.

The conclusions of these measurements are that O₃ main decrease is around 15-20 Km, where it reaches 10% per decade at least. A smaller change is seen at 40 Km.

Model taking both O_3 changes into account

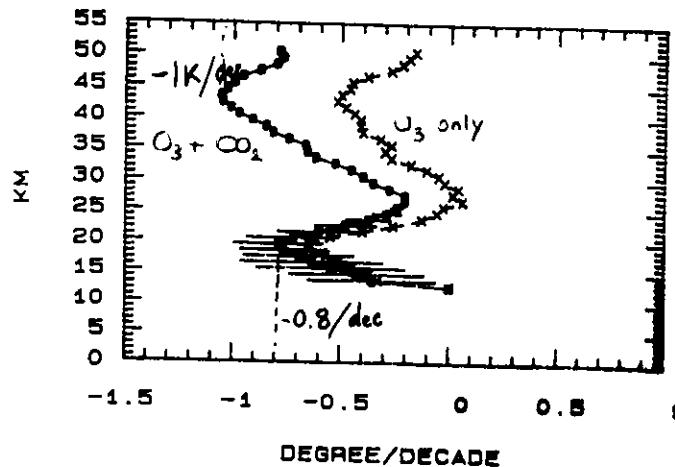


Fig. 2. Decadal temperature trend estimates derived from radiative transfer model using ozone change estimates of Figure 1. O's are estimates with CO_2 changing through the period and the horizontal lines represent the 95% confidence limits. The curve represented by X's is for the calculation with CO_2 fixed at the 1970 value.

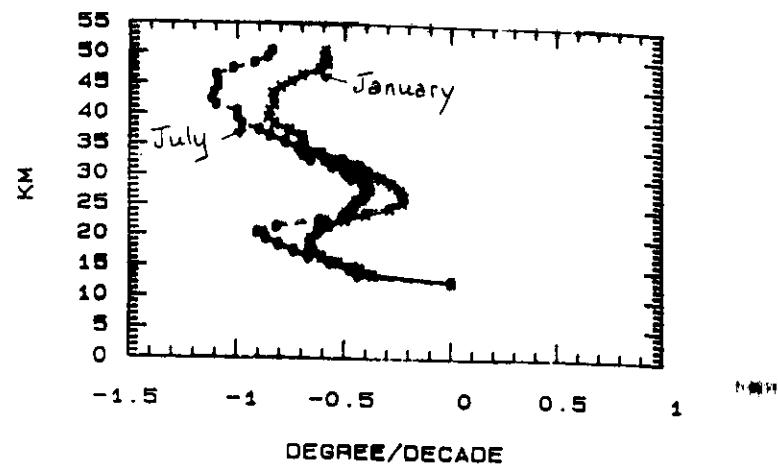
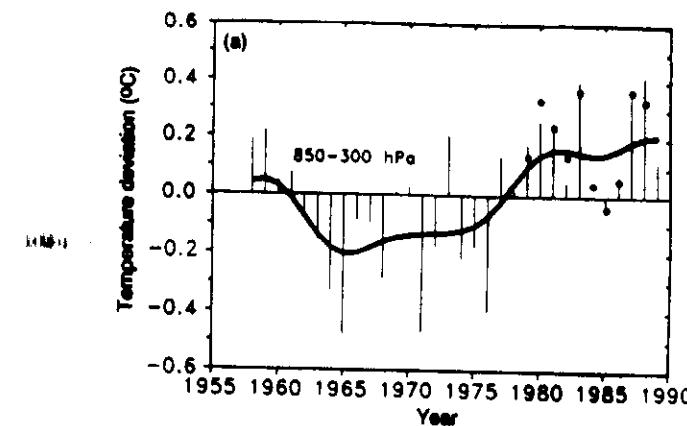
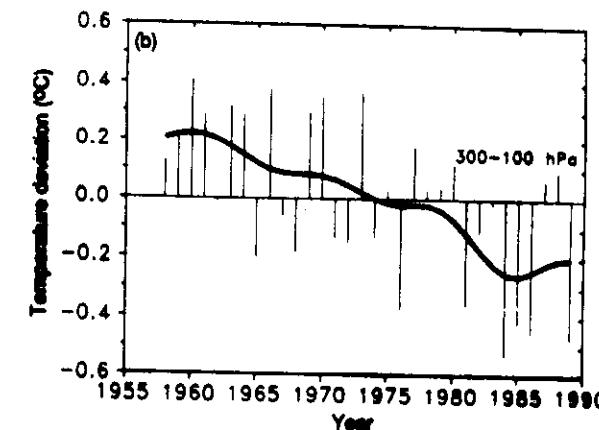


Fig. 3. Decadal temperature trend estimates as in Figure 2, but for January, X's, and July, O's. For description of model attributes, see text.

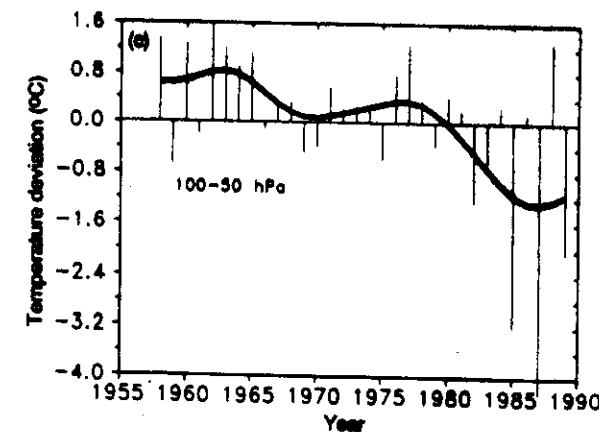
Results as for Angell (1988)



850-300 hPa

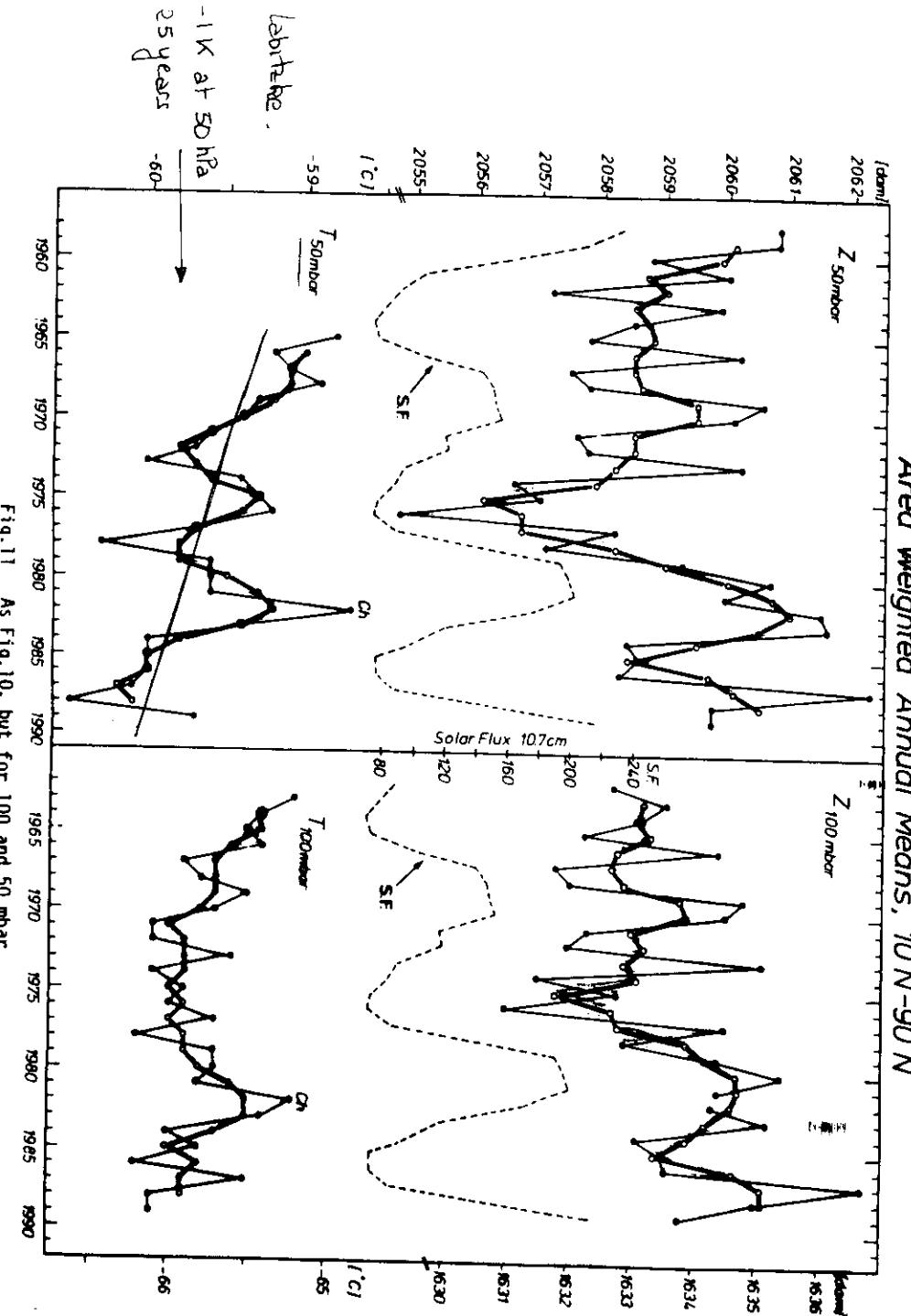


300-100 hPa
~ tropopause



0.4K / 30 years
100-50 hPa
~ 18 Km

Figure 12.—Temperature anomalies in the troposphere and lower stratosphere for 1958-1989, based on Angell (1988). (a) Annual global values for 850-300 mb. Dots are values from satellite MSU; Spencer and Christy, 1990. (b) 300-100 mb. (c) Annual Antarctic (60°S - 90°S) for 100-50 mb. Source: Houghton et al.,



Newly published analysis of the radiosonde data, by Miller et al (1991) JGR, showing that the maximum cooling took place around 15-20 km where O_3 decreased

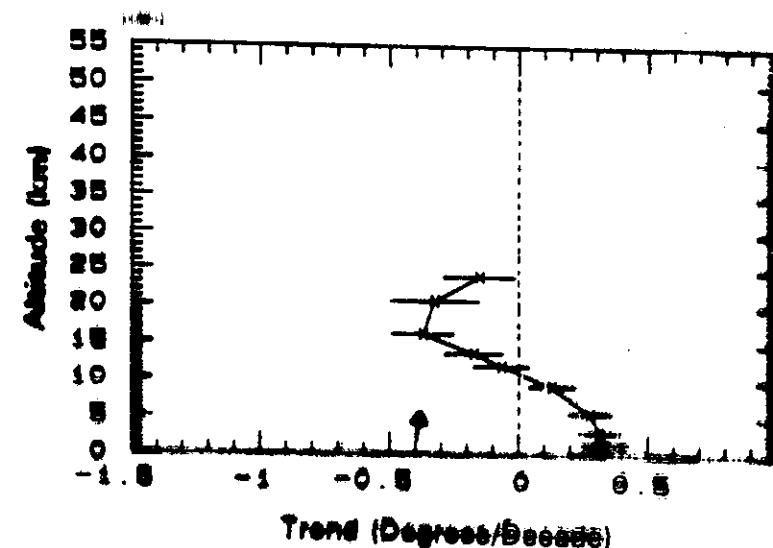
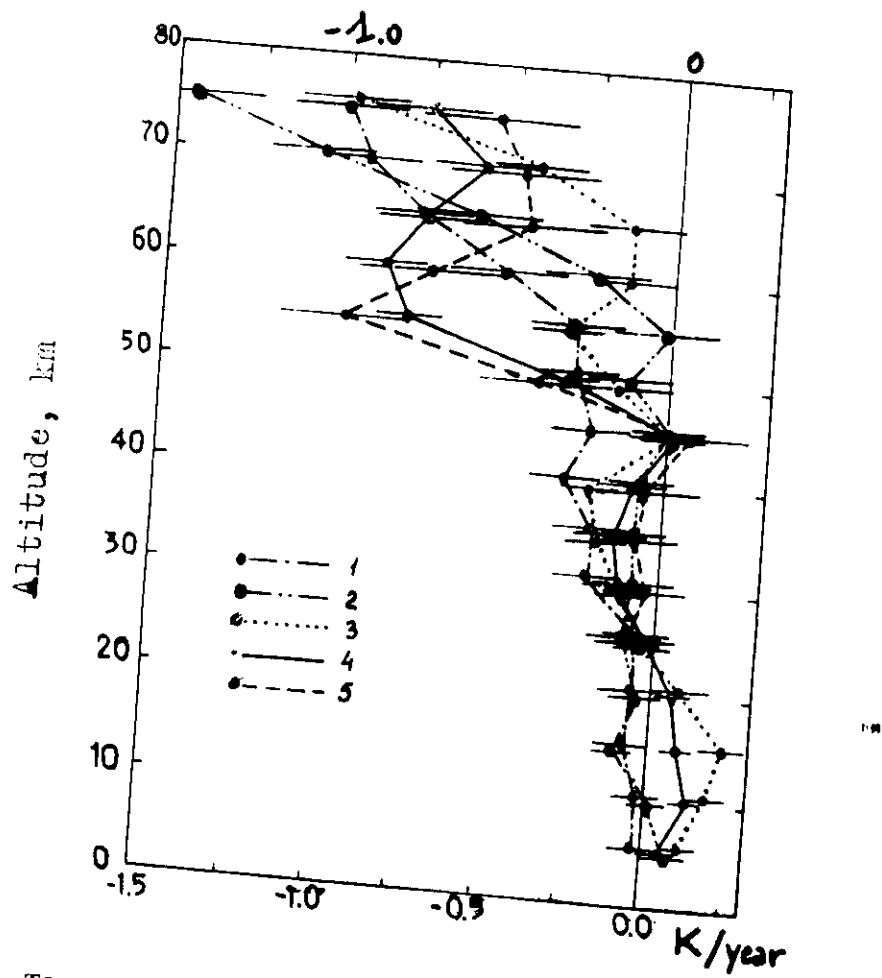


Figure 2-20: Rawinsonde temperature trend estimates in °C per decade as a function of altitude. Horizontal bars represent 95% confidence limits of estimated trends. (From [Zelinka et al., 1991](#))

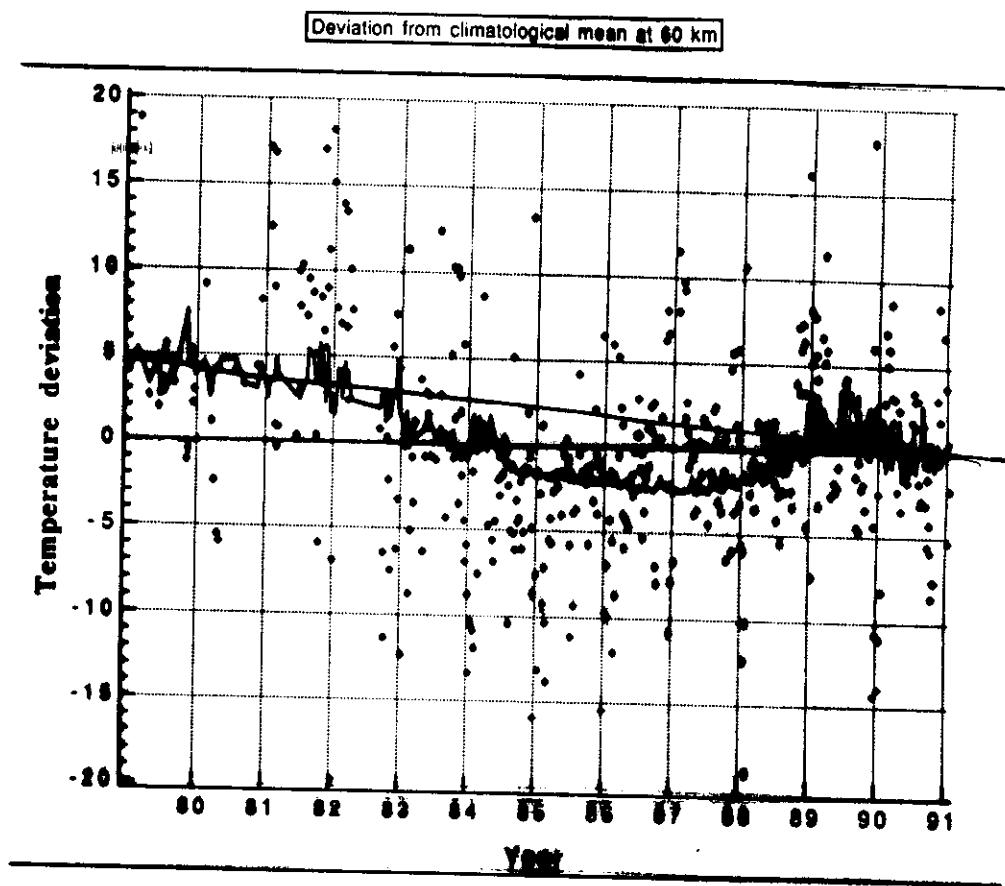
The upper middle atmosphere, as seen by rockets is shown to have cooled by 0.5 to 1 K / year around 50.70 Km.

Rayleigh Lidar data also provide a data base to study the long term trend from 30 to 80 Km. This shows the raw weekly mean value of the deviation from the climatological mean at 60 Km



Temperature trends in K/year obtained from 5 rocket sites for the period 1964-1990, by KOKIN & LYSENKO (submitted to JATP, 1992)

- 1 HEYSS ISLAND (80.6°N, 58°E)
- 2 MOLODOZHNAIA (67.7°S, 45.8°E)
- 3 THUMBA (8.5°N, 76.8°E)
- 4 VOLGOGRAD (48.7°N, 44.3°E)
- 5 BALKHASH 46.8°N, 74.6°E)



Rayleigh lidar data bank at OMP France

from Hauchecorne and Chary

DETERMINATION OF THE CONFIDENCE INTERVAL OF THE LEAST SQUARE FIT

Least square fit of the temperature:

$$T(z_i, t_j) = \bar{T}(z_i) + A_1(z_i) \cos [(2\pi t_j - j_1(z_i))] + A_2(z_i) \cos [4\pi(t_j - j_2(z_i))] \\ + B(z_i)(t_j - \bar{t}) + C(z_i) \left[\frac{F(t_j) - \bar{F}}{100} \right] + T'(z_i, t_j)$$

Variance of the temperature :

$$\sigma_T^2(z_i) = \frac{1}{n-p} \sum_{j=1}^n T^2(z_i, t_j)$$

Correlation between two successive measurements:

$$\phi(z_i) = \frac{\sum_{j=1}^{n-1} (T(z_i, t_j) T(z_i, t_{j+1}))}{\sum_{j=1}^{n-1} (T^2(z_i, t_j)/2 + T^2(z_i, t_{j+1})/2)}$$

If the autocorrelation function decreases exponentially:

$$\phi(t, t+\Delta t) = \exp(-t/\Delta t)$$

the least square estimator for the linear trend is:

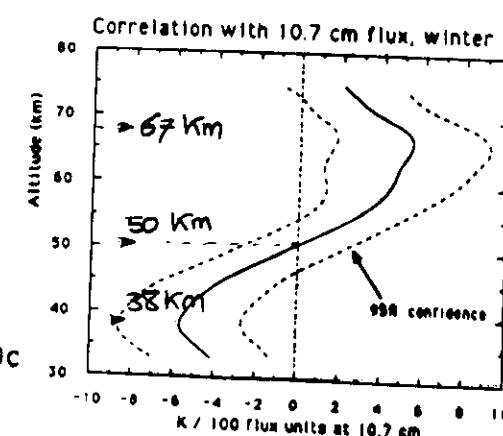
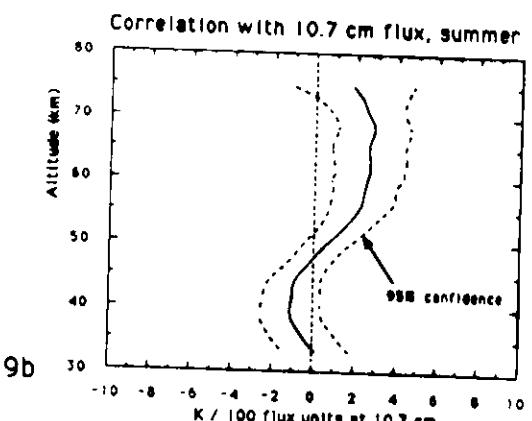
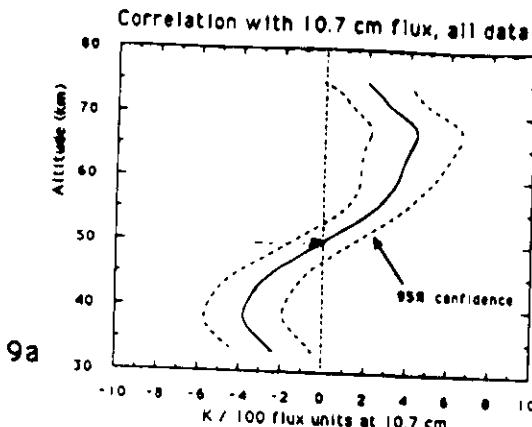
$$\sigma_B^2(z_i) = \frac{(1-\phi^2(z_i))}{(1-\phi(z_i))^2} \frac{\sigma^2(T(z_i))}{\sum_{j=1}^n (t_j - \bar{t})^2}$$

The 95% confidence interval is:

$$\delta(B(z_i)) = 1.96 \sigma(B(z_i))$$

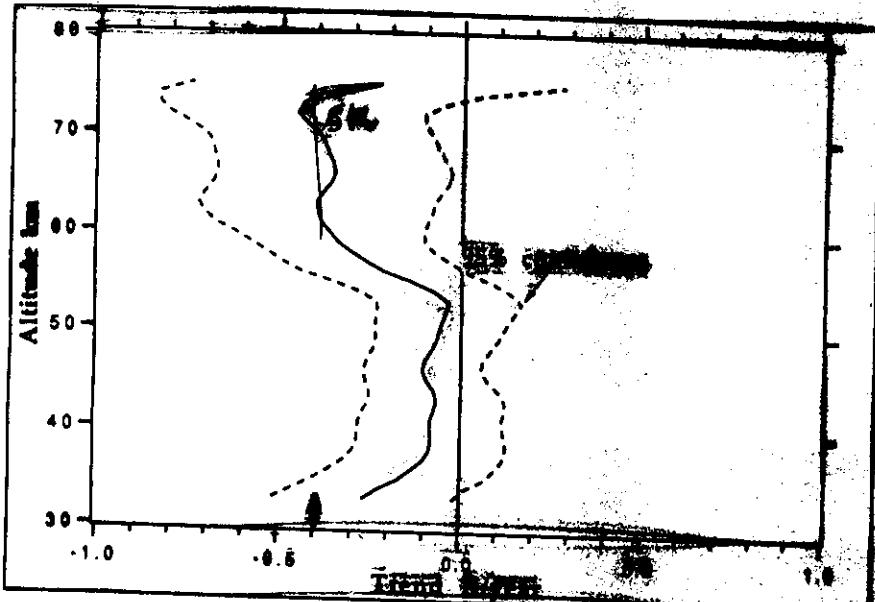
FIG. 9

Signatures of the solar cycle from the Udar data bank

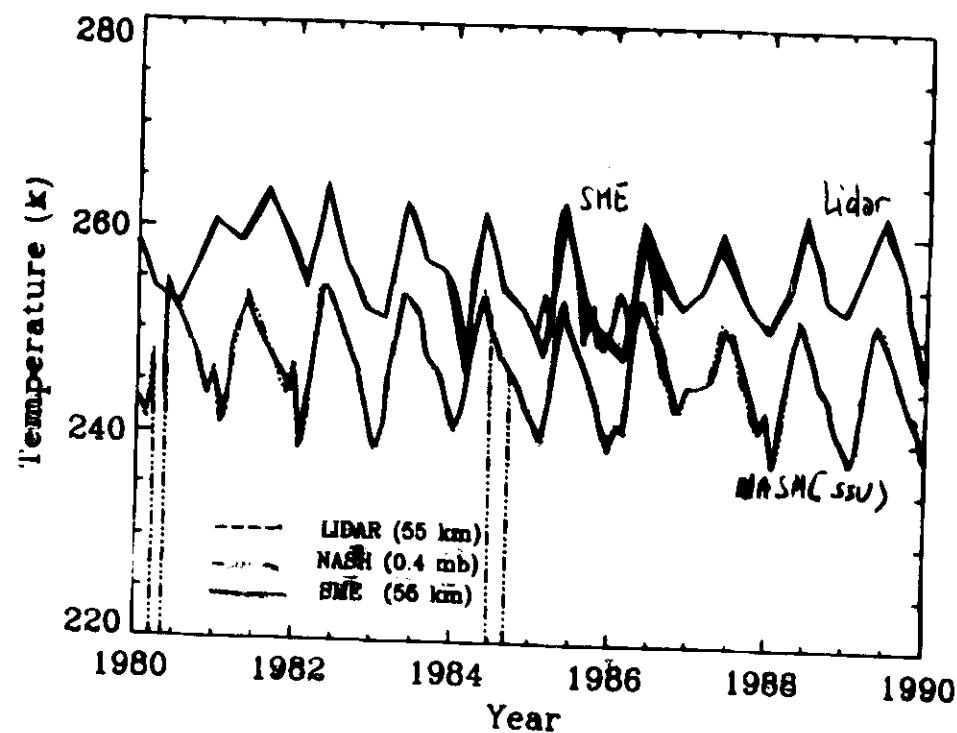


Lidar gives only local information. Comparison with satellite data show that, even though there is a systematic difference in the absolute value of the temperature measured by SSU (47x) and by lidar, the trends are quite the same. (see next page)

Lidar temperature above south of France
Trend summer 1979-91



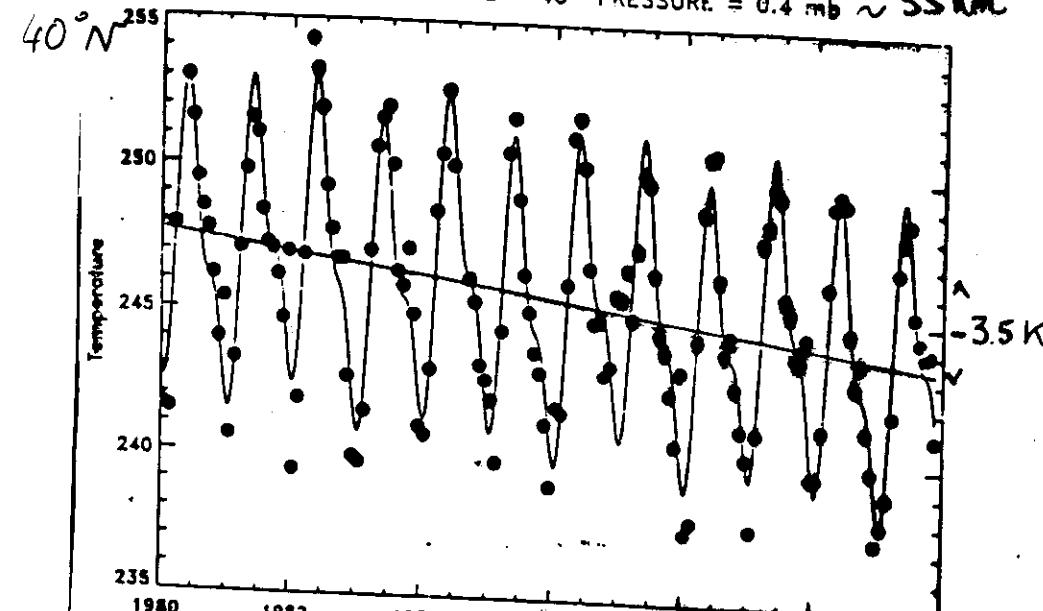
updated from Hauchecorne et al (1990)



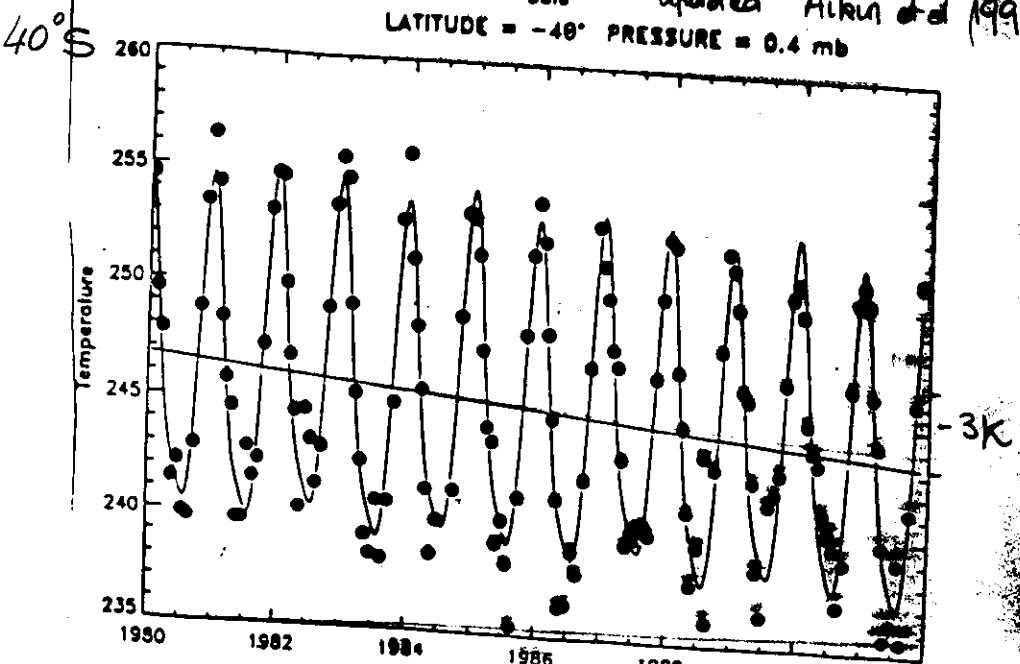
Aitken, Chanin and Nash (1990)

update from Aitken et al (1991)

LATITUDE = 40° PRESSURE = 0.4 mb \sim 55 Km



Date updated Aitken et al (1991)
LATITUDE = -40° PRESSURE = 0.4 mb



Latitudinal variation of the cooling trend.

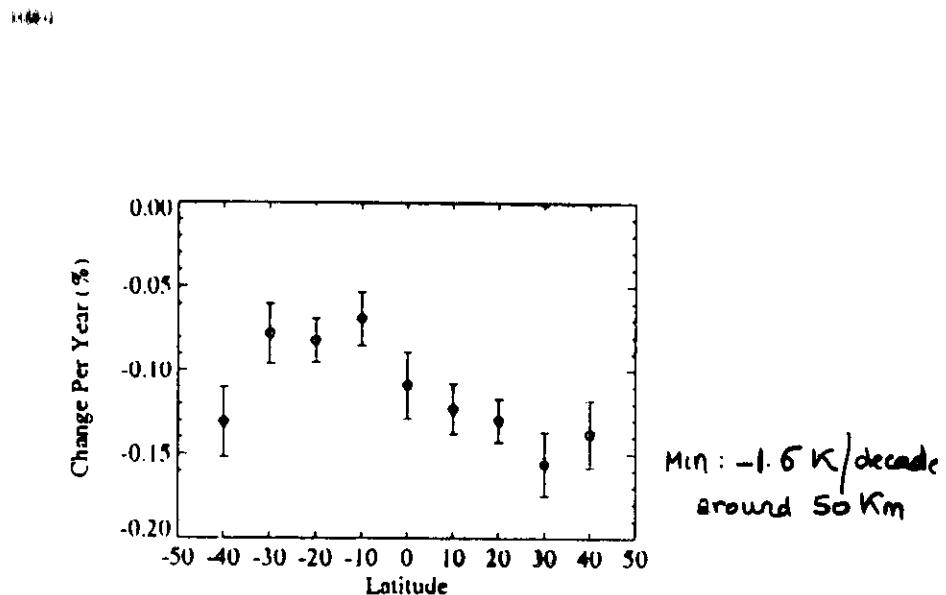
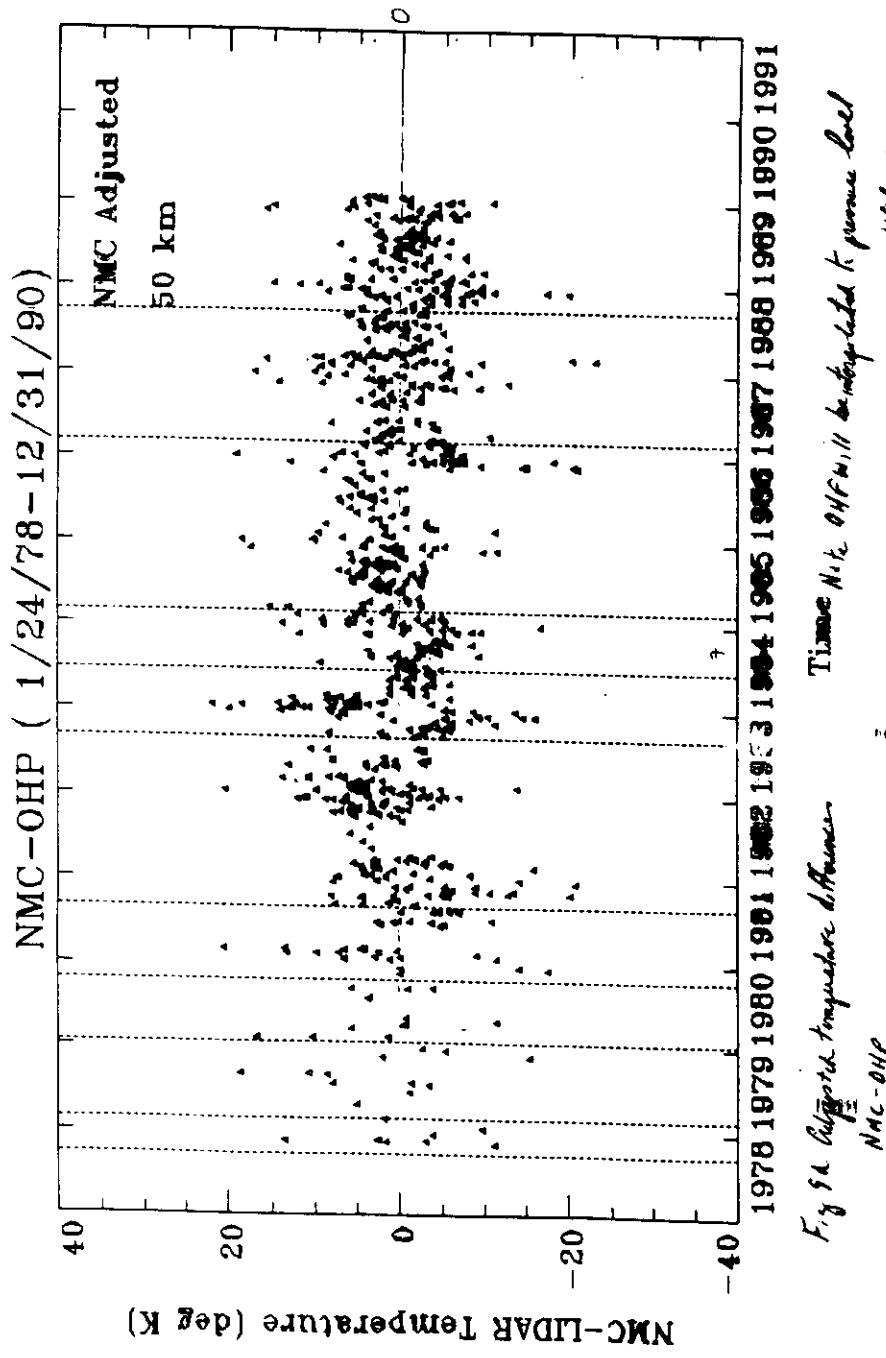


Fig. 3. Percentage change per year for the SSU 47X temperature data as a function of latitude. Error bars show $\pm 1\sigma$ uncertainties.

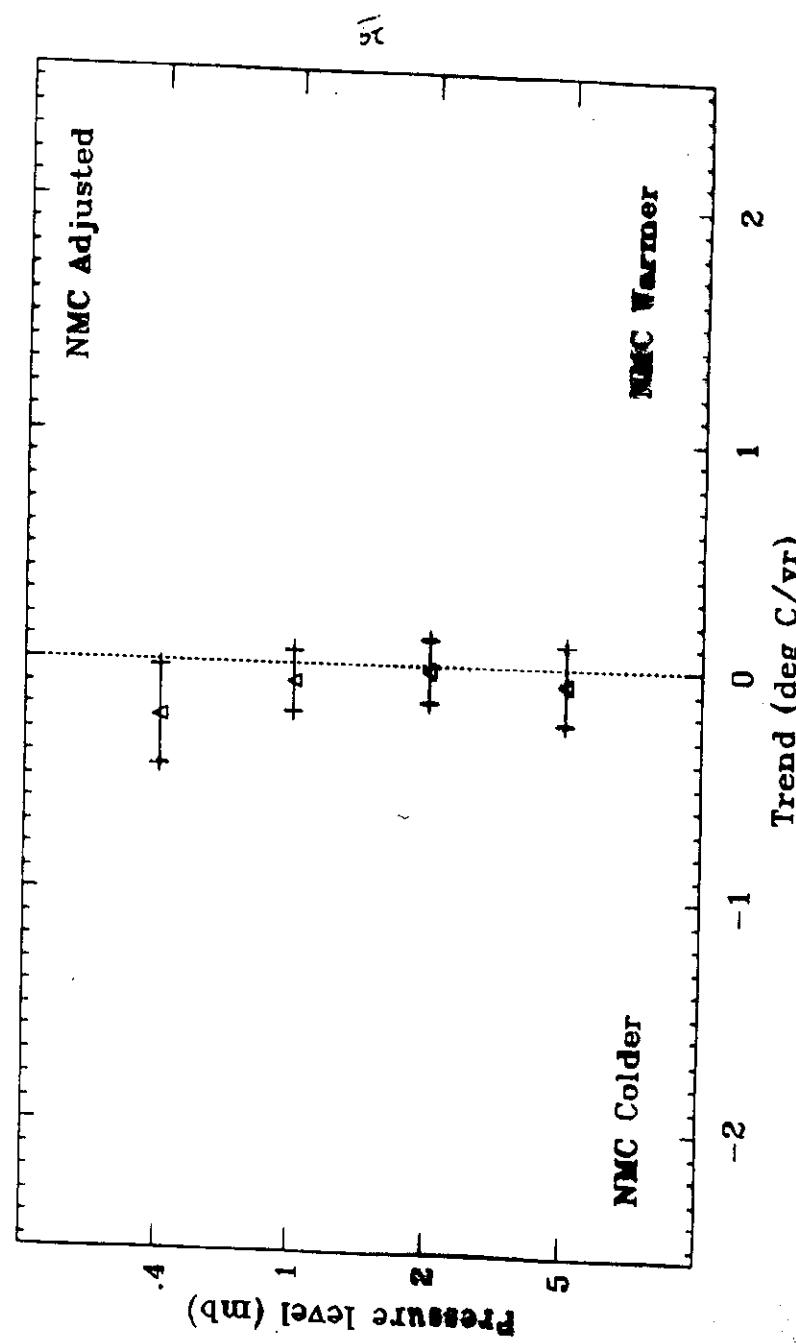
Aitken et al., GRL, 1991

Comparison between NMC and lidar data



Difference in trends observed by lidar and NMC.

Cellman et al (1992)



SUMMARY

- SMALL STRATOSPHERIC COOLING AROUND 15-20 KM.
- UNCERTAINTY ABOUT THE COOLING AROUND 40 KM WHERE CO₂, H₂O ETC SHOULD ADD THEIR EFFECTS.
- DIRECT AND INDIRECT EVIDENCE OF LARGER MESOSPHERIC COOLING.
- MORE WORK NEEDS TO BE DONE TO DETERMINE THE TEMPERATURE TRENDS.

GROUND OBSERVATIONS OF THE DYNAMICS OF THE MIDDLE ATMOSPHERE

1) Methods of observation

- Actives methods :
 - . Radars
 - . Lidars
- Passives methods:
 - . Natural emissions

2) Observational studies

- General circulation
- Planetary waves
- Tides
- Gravity waves
- Turbulence

RADARS

NATURAL EMISSIONS

1) MF / HF radars :

- Frequency : 2-6 MHz
- Process : partial reflection from ionized structures
- Altitude : 60-100 km (vertical profiles)
- Measured parameter : Doppler wind

- Process Natural emissions of OH Meinel bands (700-800 nm)
Oxygen green line (557.7 nm)
- Altitude OH : 85 km (horizontal structures)
557.7 nm : 95 km (horizontal structures)
- Measured parameters :
Airglow intensities
Rotational temperature (OH)
Doppler wind

2) UHF / VHF radars :

- Frequency : 40-50 MHz (VHF) and 440 MHz (UHF)
- Process : scattering by ionized structures (60-110 km)
and neutral structures (10-30 km)
+ meteor echoes (VHF)
- Altitude : 10-30 km and 60-110 km (vertical profiles)
- Measured parameters : Doppler wind
Turbulence

LIDARS

1) Rayleigh lidars :

- Process : Rayleigh scattering by molecules
- Laser Wavelength : 532 nm (Nd-YAG)
351-353 nm (XeF)
- Altitude : 30-90 km (vertical profiles)
- Measured parameter : air density
- Hydrostatic equilibrium : temperature

2) Sodium lidars :

- Process : resonant scattering by Na atoms
- Laser wavelength : 589-596 nm
- Altitude : 80-110 km (vertical profiles)
- Measured parameters: sodium concentration
temperature (Doppler width)

3) Raman lidars :

- Process : Rotational or vibrational Raman scattering by molecules, mostly N_2 , O_2
- Laser wavelength : 532 (Nd-Yag)
- Altitude: 2-35 Km (vertical profile)
- Measured parameter : air density
- Calculated parameter : temperature

4) Mie Lidars :

- Process : Mie backscattering by aerosols or clouds
- laser wavelength : 1,06 μm 532, 352 nm (Nd-Yag)
- Altitude: 5-25 Km
- Measured parameter : concentration of aerosols \rightarrow dynamic Doppler shift \rightarrow winds

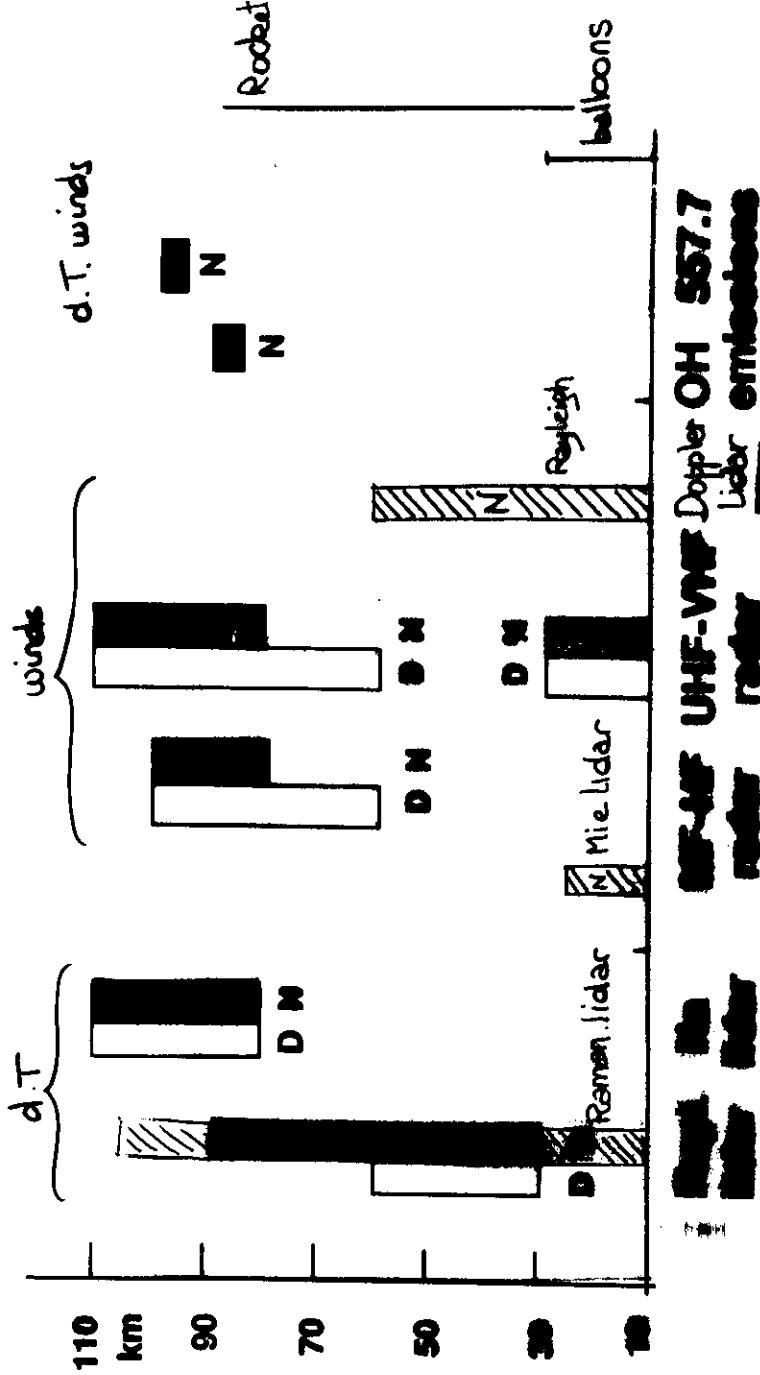
5) Doppler Rayleigh Lidar

- Process : Rayleigh backscattering by molecules
- laser wavelength : 532 nm
- altitude: 3 - 65 Km (in presence or not of aerosols)
- Measured parameter : horizontal components of the wind

6) DIAL Lidar

- Process : Differential absorption at different wavelength
 O_3 , H_2O ... \rightarrow stratosphere-troposphere exchange

DAY / NIGHT AND ALTITUDE COVERAGE



Lidar equation and principle

$$P(z, z+dz) = P_0 \times G(z) \times \tau_a(z, z+dz) \times e^{-2\tau_a(z, z+dz)}$$

Signal power linear efficiency of system
 return Power

optical thickness of the scatterer in atmosphere

2 cases :

- ① $\tau_a \rightarrow$ density of scatterers (using Rayleigh, Mie or Resonance scattering)
if $\tau_a > \tau$ of other constituents between $z, z+dz$
by adequate choice of z element \rightarrow D.I.A.L
- ② $\tau_a \rightarrow$ density of the absorbers and of a specific absorber (using Rayleigh Scattering)

Scattering processus used for Lidar studies

Elastic Scattering

Rayleigh $\lambda_r = \lambda_e$ $\sim 10^{-28} \text{ cm}^2$

Mie $\lambda_r = \lambda_e$ $\sim 10^{-10} \text{ cm}^2$

Inelastic scattering

Resonance $\lambda_r = \lambda_e$ $\sim 10^{-16}$ (atoms)

Fluorescence $\lambda_r \neq \lambda_e$ $\sim 10^{-17}$ (mol.)

Raman $\lambda_r \neq \lambda_e$ $\sim 10^{-30} \text{ cm}^2$

Resonant Raman $\lambda_r \neq \lambda_e$ $\sim 10^{-25} \text{ cm}^2$

Lidar equation for Rayleigh scattering

$$\frac{dN}{dz} = N_0 \sigma_{sr} dz$$

$$dz = c \frac{\Delta t}{2}$$

$$\Delta t = 10^{-9} \text{ s}$$

$$\Delta z = 1,5 \text{ m}$$

$$N = N_0 \cdot \frac{A}{4\pi z^2} \cdot \frac{n \sigma_{sr} dz}{dt} \cdot K T^2$$

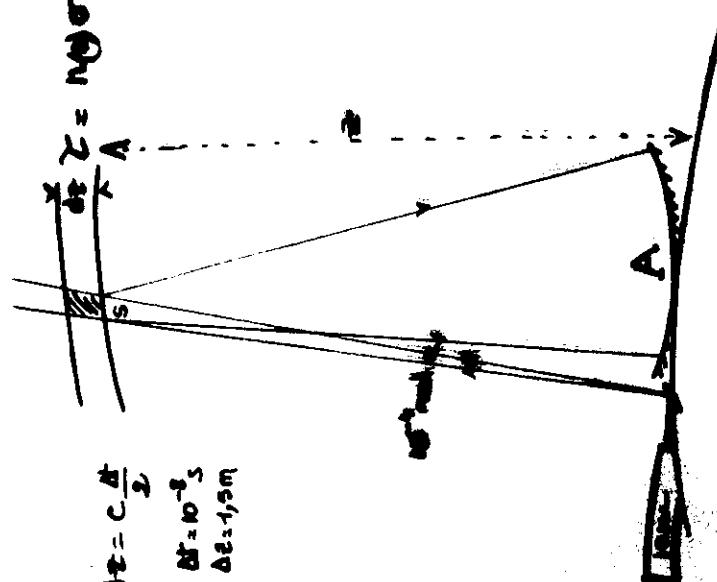
$$\int N dz = n(z) dz dt$$

Intensity

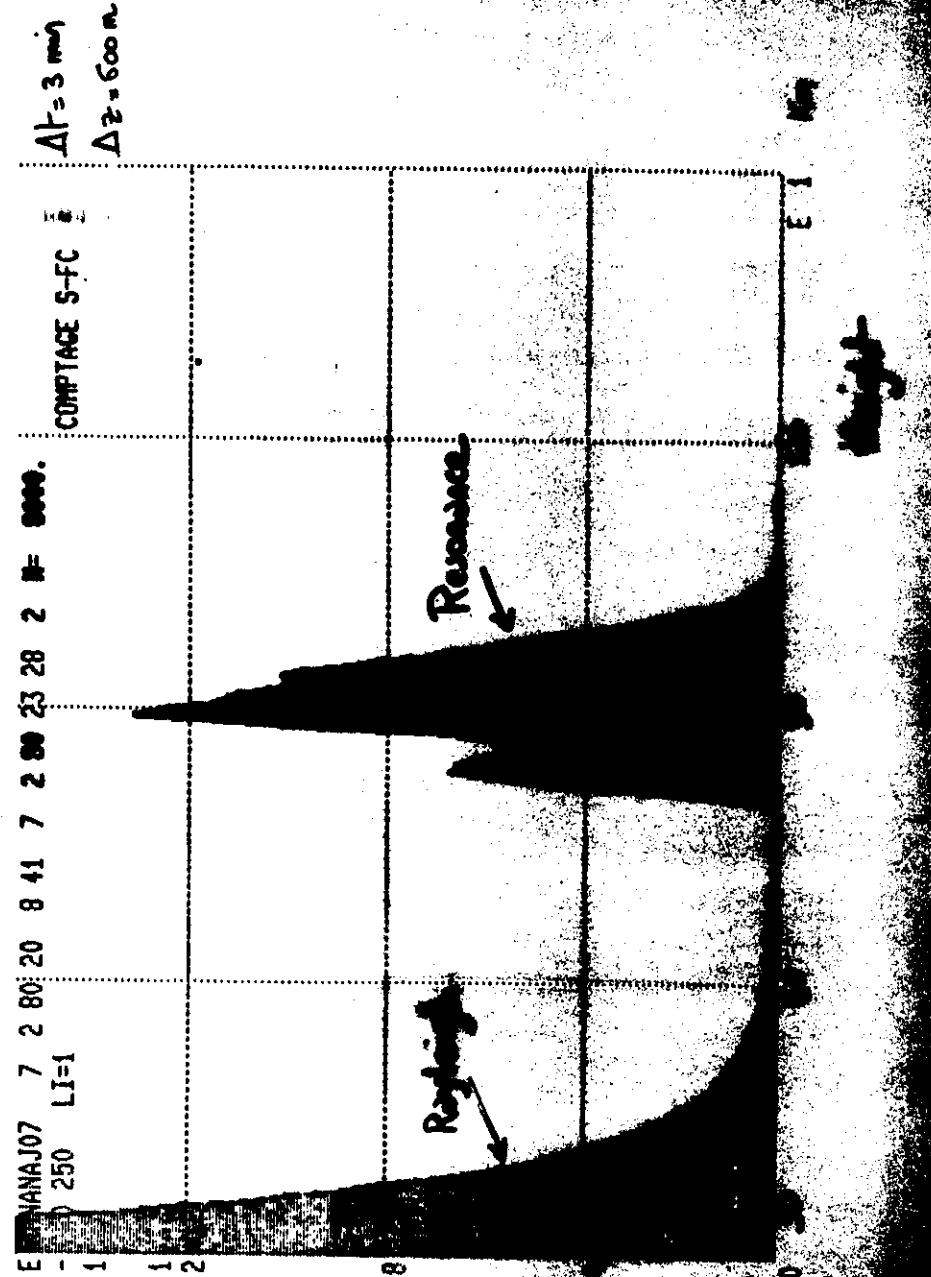
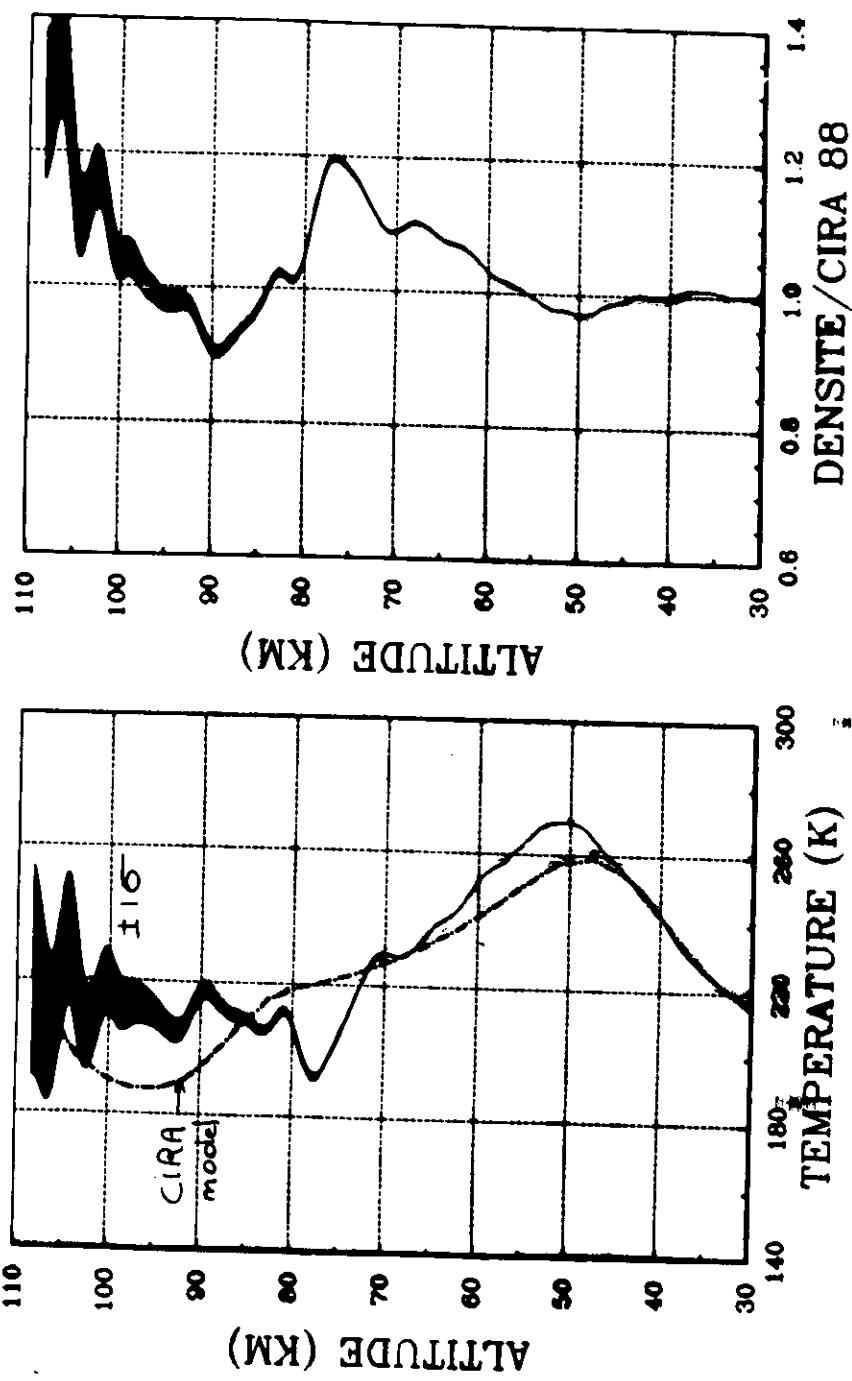
$$\frac{dN}{dz} = \frac{\sqrt{N}}{N} = \frac{1}{(dz dt)^{1/2}}$$

Resolution

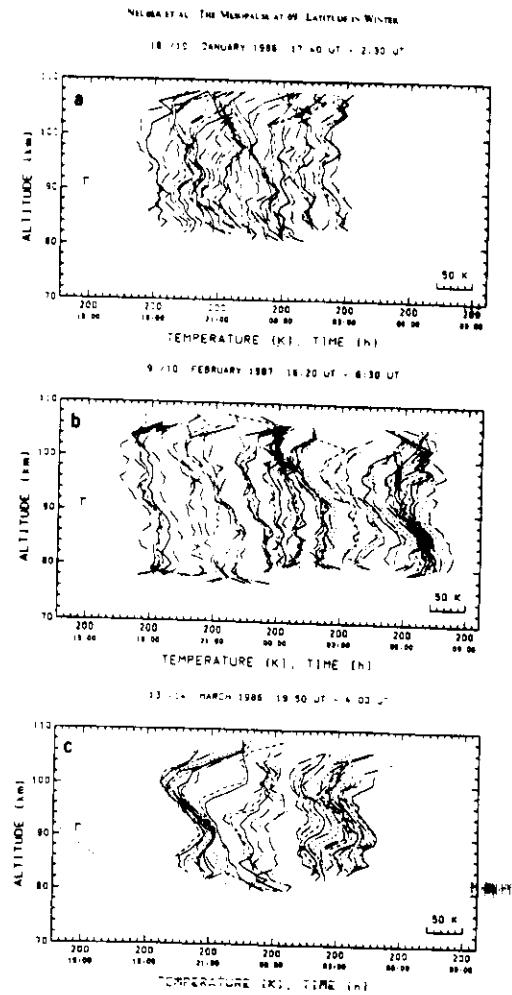
$$\frac{10^{-9}}{10^{-10}} = 10 \text{ (defined by sampling rate)}$$



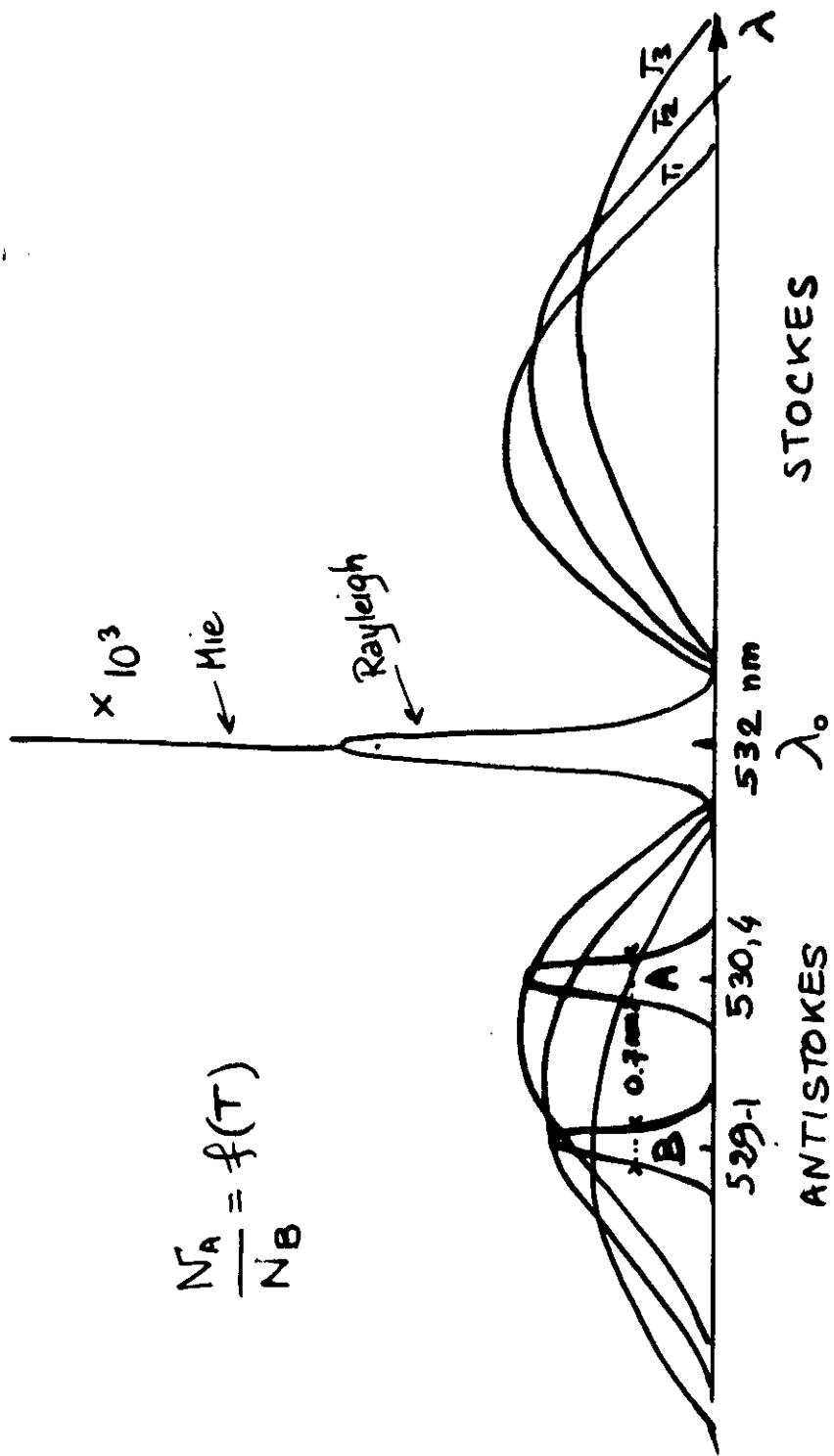
DATE 21/11/89 SITE : LIM 04h06 REFERENCE : CIMA NOW Example of Rayleigh lidar Temperature and density profiles



Neuler et al., 1988



ROTATIONAL RAMAN LIDAR



Comparaison entre Raman lidar et radiosonde

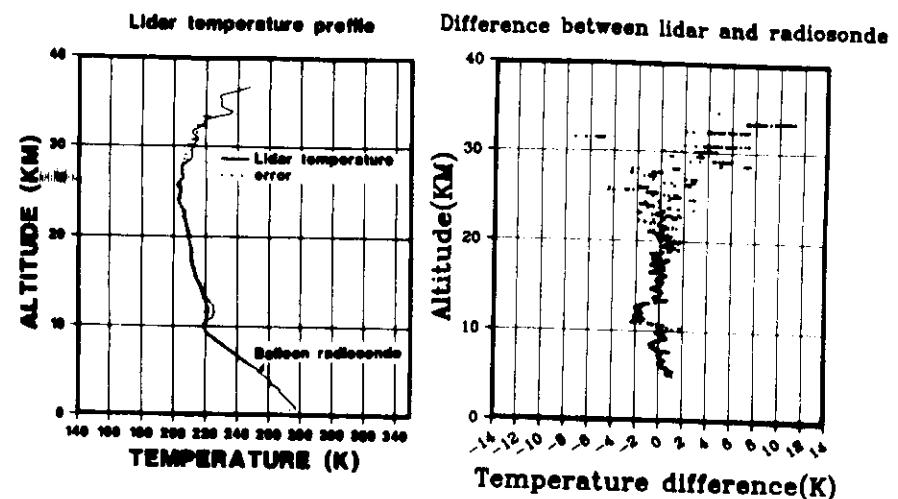
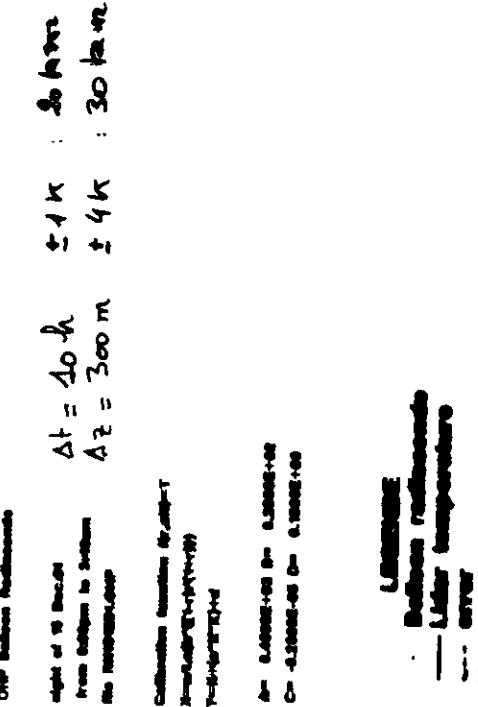
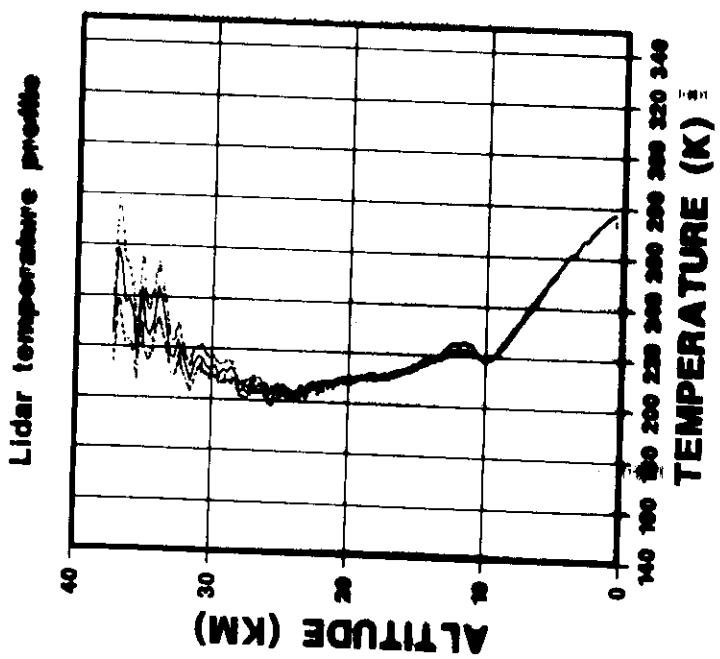


Figure 8: a) Profils de température obtenus par Lidar Raman et par radiosondage.
b) Différence entre les deux mesures

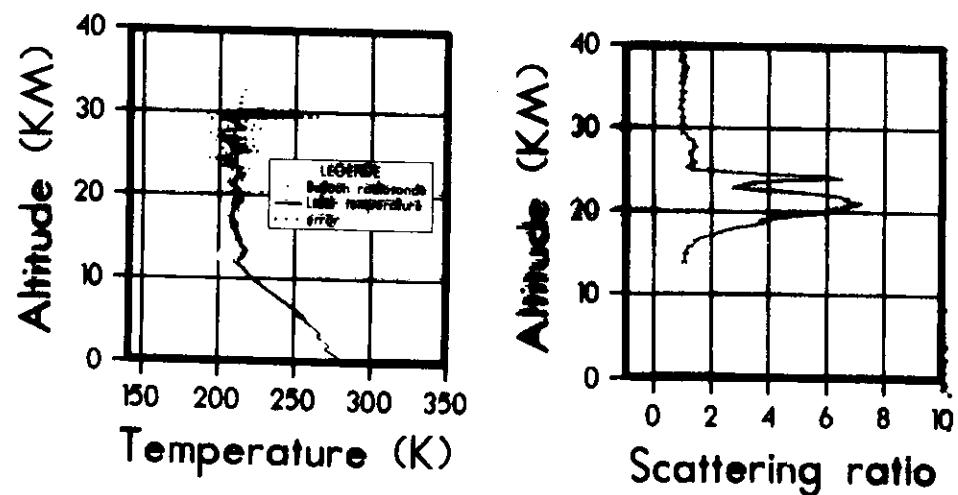
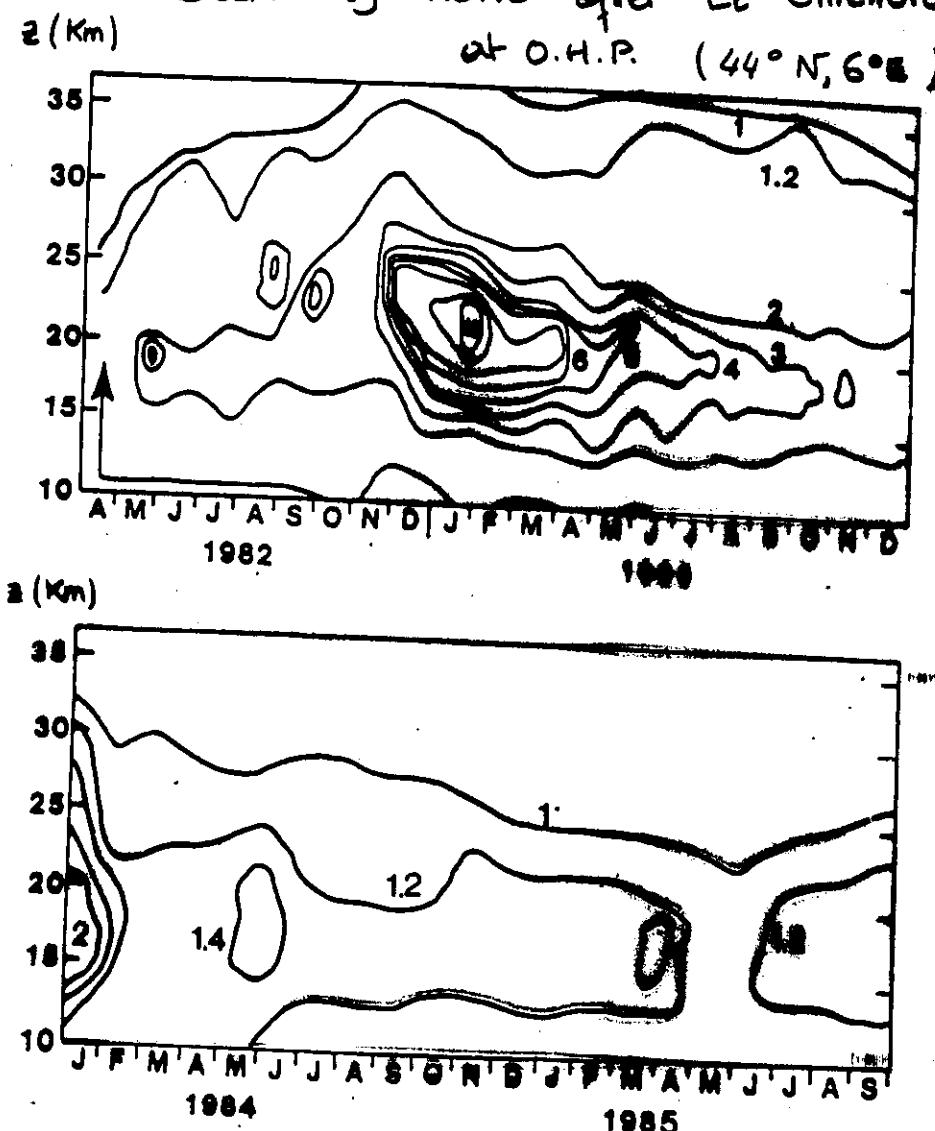


Figure 9 : a) Profil de température par Lidar Raman le 19/12/91 à l'OMP
b) profil d'aérosols obtenu simultanément le 19/12/91 à l'OMP.

Mie-Lidar

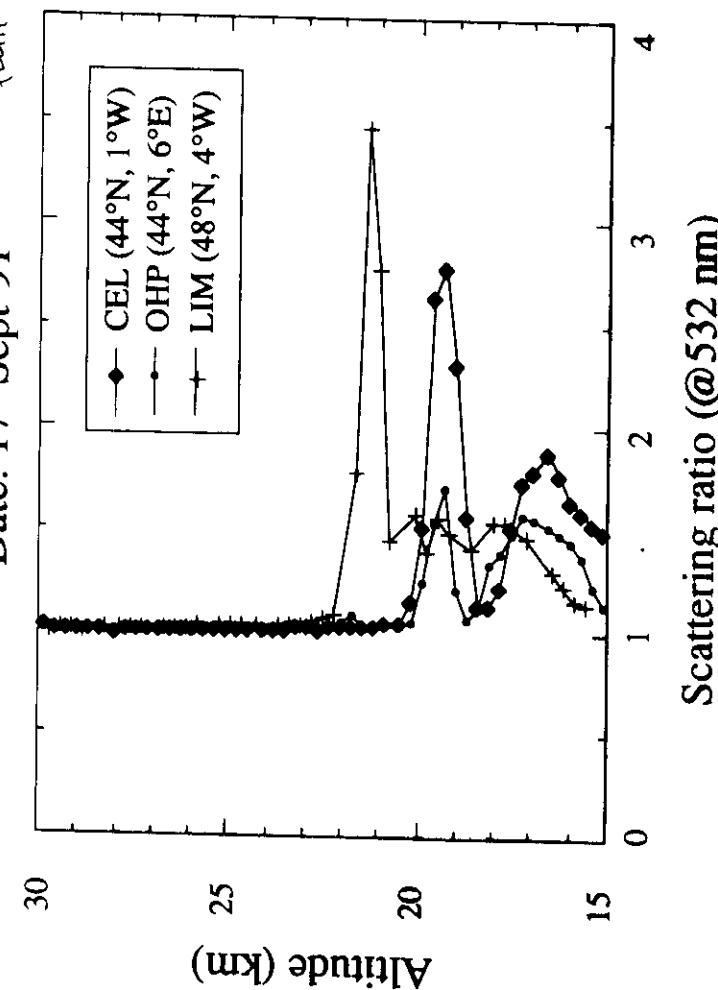
Scattering Ratio after El Chichon
at O.H.P. ($44^{\circ}N, 6^{\circ}E$)



$$R(z_i) = \frac{n_r(z_i)\beta_r + n_a(z_i)\beta_m}{n_r(z_i)\beta_r}$$

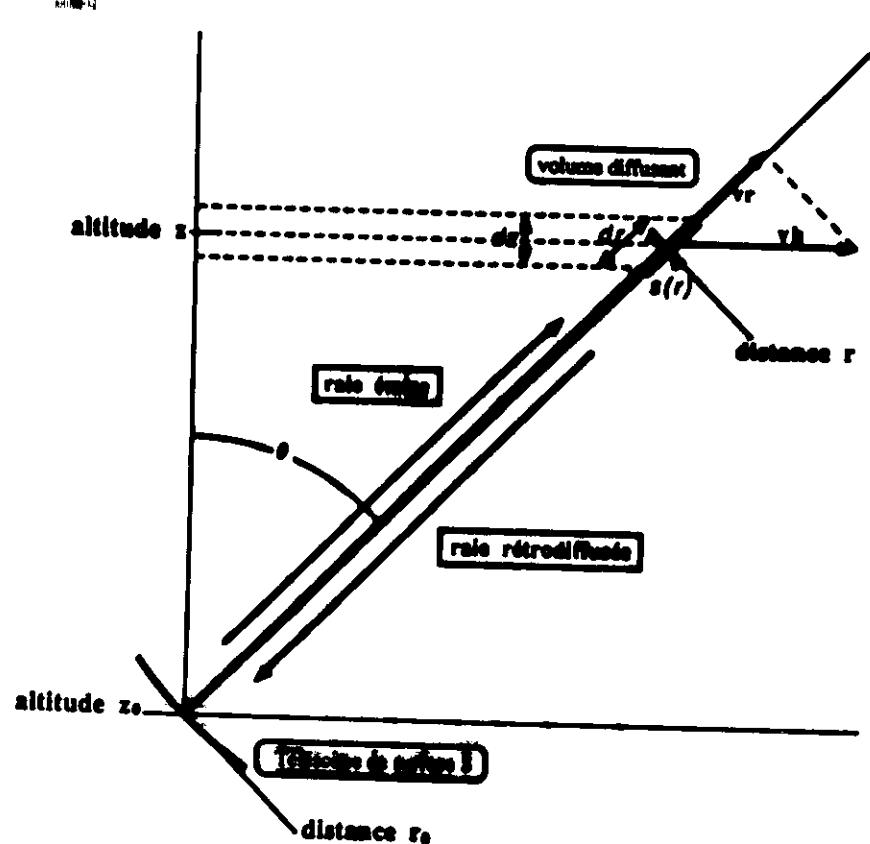
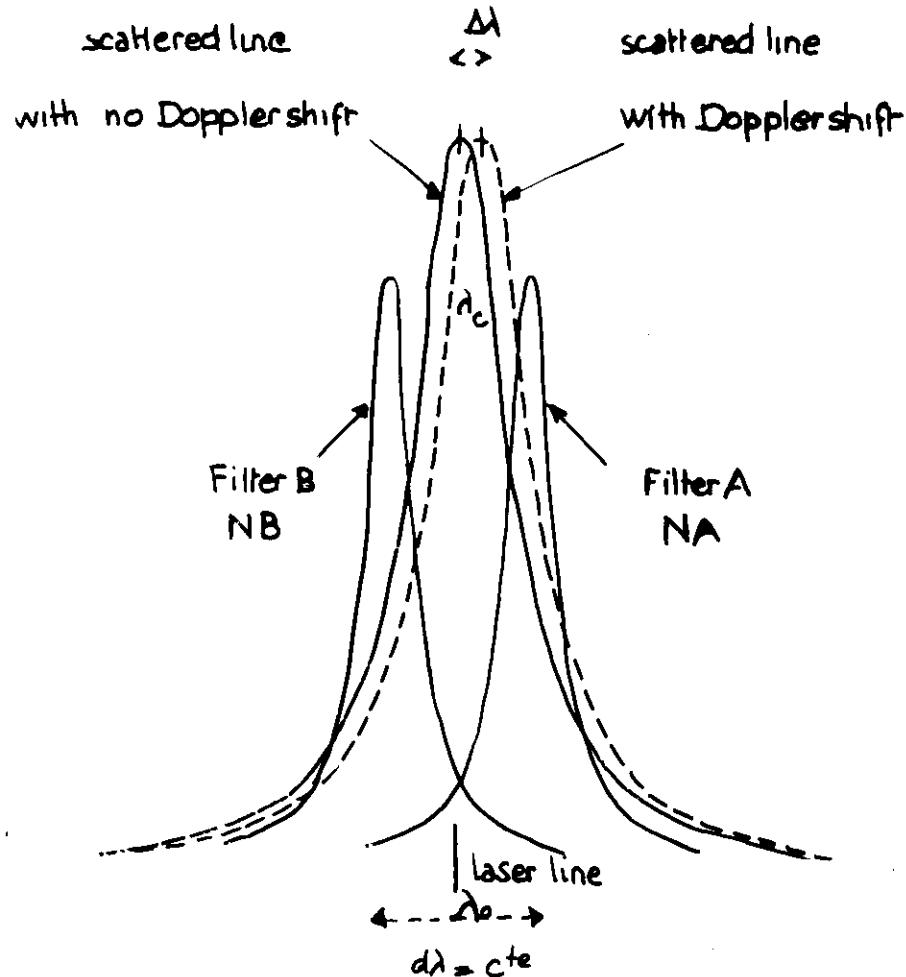
$\lambda = 532 \text{ nm}$

CNRS Rayleigh Lidars → Aerosol layers due to Pinatubo
Date: 17- Sept-91



Scattering ratio (@ 532 nm)

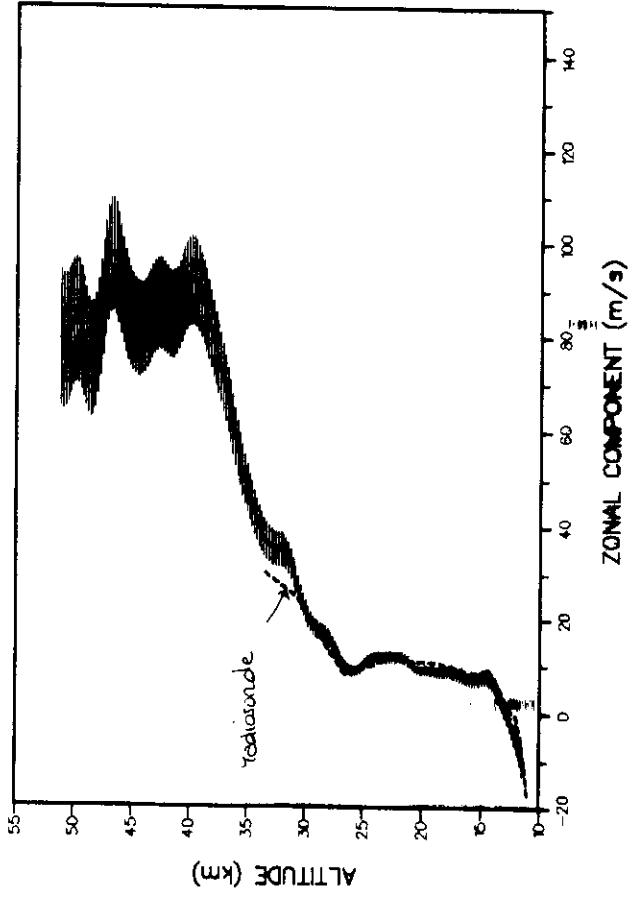
The Rayleigh Doppler Lidar principle



We measure the response R : $R = \frac{NA - NB}{NA + NB}$ successively
for a tilted (45°) line of sight and for the vertical one.

DOPPLER RAYLEIGH LIDAR

CEL 20-02-92



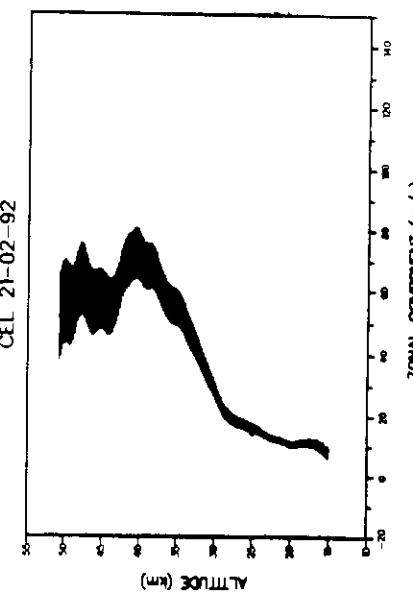
Vertical resolution : 2.00 km
Integration time : 2h 06

Dashed line :

Balloon
BOR 21-FEB-1992
00h 00

EAST
from : 0h 34
to : 2h 37
WEST
from : 0h 26
to : 1h 29

CEL aerosols lidar
21-02-92 from 00:41 to 02:51 UT

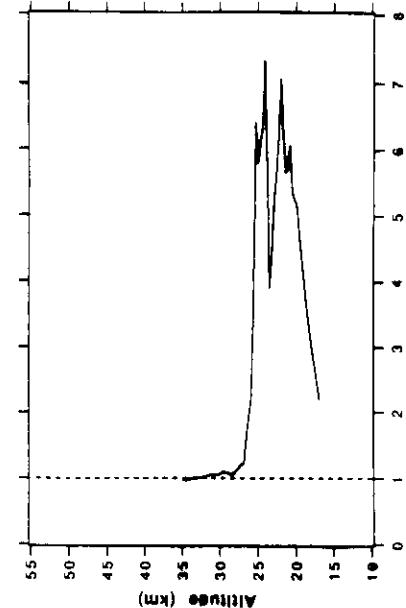


DOPPLER RAYLEIGH LIDAR
CEL 21-02-92

EAST
from : 0h 48
to : 2h 51
WEST
from : 0h 41
to : 0h 45

Vertical resolution : 2.00 km
Integration time : 2h 07

Dashed line :
Balloon
BOR 22-FEB-1992
00h 00



The wind profile is shown not to be disturbed by aerosols

STRATOSPHERIC PROCESSES AND THEIR ROLE IN CLIMATE (SPARC)

**A Project of the World Climate Research
Programme**

Chairmen: M-L. CHANIN & M. GELLER

**FOCUS 1 : THE INFLUENCE OF THE
STRATOSPHERE ON CLIMATE**

**FOCUS 2 : PROCESS STUDIES ASSOCIATED
WITH STRATOSPHERIC OZONE DECREASE**

**FOCUS 3 : GLOBAL CHANGE OF THE
STRATOSPHERE**

**FOCUS 4 : MONITORING AND MODELING
OF UV IRRADIATION CHANGES IN THE
BIOSPHERE AND THE TROPOSPHERE**

STRATOSPHERIC PROCESSES AND THEIR ROLE IN CLIMATE

SPARC

A Project of the World Climate Research Programme

**FOCUS 1 : THE INFLUENCE OF THE STRATOSPHERE ON
CLIMATE**

- Activity 1.1 Understand the Influence of the Stratosphere on Tropospheric Climate
- Activity 1.2 Assess the Potential Role of Stratosphere - Troposphere Interactions in climate change

**FOCUS 2 : PROCESS STUDIES ASSOCIATED WITH
STRATOSPHERIC OZONE DECREASE**

- Activity 2.1 Chemical and Aerosol Processes
- Activity 2.2 Dynamical Transport of Chemical Constituents
- Activity 2.3 Radiative forcing of the stratosphere
- Activity 2.4 Radiative-Dynamical-Chemical Interactions

FOCUS 3 : GLOBAL CHANGE OF THE STRATOSPHERE

- Activity 3.1 Understanding the Variability and Long-Term Trends in the Stratosphere.

**FOCUS 4 : MONITORING AND MODELING OF UV
IRRADIATION CHANGES IN THE BIOSPHERE AND THE
TROPOSPHERE**

•Activity 4.1 Modeling and Experimental Validation of UV Penetration

•Activity 4.2 Global UV Climatology and Prediction

COMPLEMENTARY PROGRAMME

**BIOSPHERIC EFFECTS AND FEEDBACKS RESULTING
FROM SOLAR UV-B CHANGE**

**TO BE RUN BY IGBP AND SCOPE IN CLOSE
COOPERATION WITH SPARC**

Activity I.1 Aquatic Ecosystem Responses

Activity I.2 Terrestrial Vegetation Responses

Activity I.3 Assessments of Altered Animal and Human Health

Activity I.4 Biogenic Trace Gas Balances

HISTORY OF SPARC

**1985-1986: W.G. on Solar-Terrestrial influence on the climate and the biosphere.

July 1987: First meeting of the SG-IGBP

August 1987: proposal of MARC by IAGA/IAMAP

1988-1989: presentation at SC-IGBP: MARC becomes STIB

Jan. 1990: STIB WG meeting at Abington UK.

June 1991: STIB Implementation Plan prepared during a meeting held at Stony-Brook, USA.

Aug. 1992: IGBP refuses to establish STIB as an IGBP Core Project.

Sept.1992:SPARC (STIB - The Biosphere) is proposed to WCRP

March 1992: The JSC/WCRP accepts SPARC as a WCRP project and appoints the SSG.

SPARC SCIENTIFIC STEERING GROUP

At this date the SPARC SSG composition is as follows.

- M.L. CHANIN (F) co-chairmen
- M. GELLER (USA)
- J. MAHLMAN (USA) members of the JSC
- T. MATSUNO (J)
- J. PYLE (U.K.) chairman of WGATC
- I. ISAKSEN (N)
- V. KHATTATOV (R) members of the ex-STIB⁺
CPPC
- H. TANAKA (J)
- R. TURCO (USA)
- D. EHHALT (G)

One member (a chemist) is still to be named by the JSC

+ ex officio members representing IGBP (GCTE,
JGOFS, LOICZ), SCOSTEP, GAW, SCOPE,
COSPAR.

SPARC MEETINGS

in the very near future

1992

September 1992: NATO ASI " The role of the stratosphere in Global Change"
Carqueiranne, France

May 1993: SPARC session at the EGS meeting,
Wiesbaden, Germany.

July 1993: SPARC Symposium at the IAMAP meeting,
Yokohama, Japan

1993: NATO ARW on Stratosphere-Troposphere exchange, Directors J.Holton, P. Haynes, Cambridge ?

1994; NATO ARW : The Effect of Pinatubo, Director:
G. Fiocco, Italy ?

June 1994: STEP Symposium in Sendai (Japan)