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SMR.762 - 23

Lectures III and IV

SUMMER SCHOOL IN HIGH ENERGY PHYSICS AND COSMOLOGY

13 June - 29 July 1994

NON-PERTURBATIVE ELECTROWEAK EFFECTS IN BARYOGENESIS

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Please note: These are preliminary notes intended for internal distribution only.

(1)

Anomalous fermionic number non-conservation.

(1) Level crossing. $l+1$

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F_{\mu\nu} + \bar{\Psi}_1 D \Psi_1 + \bar{\Psi}_2 D \Psi_2 + \text{scalars}$$

$$D \Psi_1 = i \partial_\mu (\partial_\mu - ie A_\mu (\text{c.c.})) \Psi_1$$

$$D \Psi_2 = i \partial_\mu (\partial_\mu + ie A_\mu (1-\delta_S)) \Psi_2$$

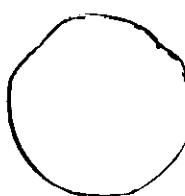
Conserved currents:

$$(i) \text{ electric current: } \bar{\Psi}_1 \gamma_\mu \Psi_1 - \bar{\Psi}_2 \gamma_\mu \Psi_2$$

$$(ii) \text{ fermionic number } \bar{\Psi}_1 \gamma_\mu \Psi_1 + \bar{\Psi}_2 \gamma_\mu \Psi_2$$

Are they conserved?

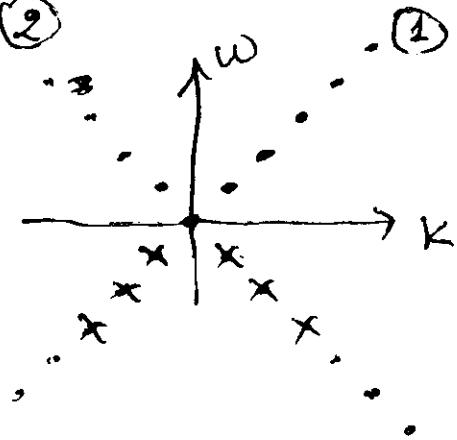
Fermionic level crossing.

 L

$$\lambda = 0 \Rightarrow \text{spectrum}$$

$$k = \frac{2\pi n}{L}; \quad \omega = k$$

(2)



$\lambda \neq 0 \Rightarrow$ (constant, time indep.)

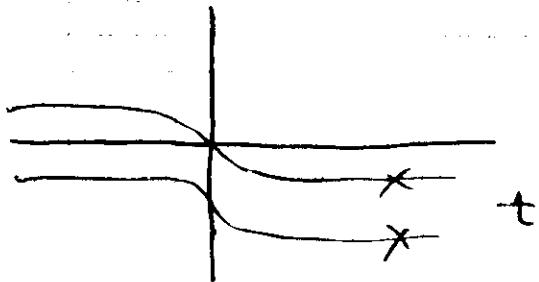
spectrum

$$\omega_1 = k - e A_1$$

$$\omega_2 = k + e A_1$$

if $A_1 = \frac{2\pi N}{eL} \leftarrow \text{integer}$ — spectrum is the same

(2)



connect

$$A_1 = 0 \quad \partial_t A_1 = \frac{2\pi}{eL}$$

Creation of fermion!

Number of created fermions:

$$F = 2 \cdot \frac{eL A_1}{2\pi} = 2 \cdot \frac{e}{2\pi} \int dx A_1 =$$

$$= 2 \cdot \frac{e}{2\pi} \int dx \int dt \partial_t A_1 = 2 \cdot \frac{e}{2\pi} \int d^2x \epsilon_{\mu\nu} F^{\mu\nu}$$

$$\partial_\mu j_\mu^F = n_f \frac{e}{2\pi} \epsilon_{\mu\nu} F^{\mu\nu} \text{ - anomaly equation}$$

$$N_{CS} = \frac{e}{2\pi} \int dx A_1 - \text{Chern-Simons number}$$

$$\Delta F = n_f \Delta N_{CS}$$

vacua with $\mathcal{A} = 0$ & $\mathcal{A} = \frac{2\pi N}{eL}$ are related by non-trivial gauge transformation

$$\varphi \rightarrow e^{i\alpha(x)} \varphi ; A_\mu \rightarrow A_\mu + \frac{1}{e} e^{i\alpha} \partial_\mu e^{-i\alpha}$$

$$\alpha = \frac{x \cdot 2\pi N}{L}$$

Mapping: $S \rightarrow U(1)$; N -winding number

Electroweak theory: everything is the same:

(3)

Fermionic spectrum is identical in any of the fields

$$\Psi = g \begin{pmatrix} 0 \\ \phi \end{pmatrix}; \quad A = g^2 g^{-1} \quad g - \text{gauge transformation.}$$

Mapping: $S_3 \rightarrow SU(2)$; winding number:

$$W(g) = \frac{1}{24\pi^2} \text{Tr} \epsilon_{ijk} S d^3x (g^2 \partial_i g^{-1}) (g \partial_j g^{-1}) (g \partial_k g^{-1})$$

integer vacuum with $W=0 \Rightarrow \sqrt{W}=1$ \rightarrow fermionic level crossing.

Selection rules: (9 quark doublets + 3 leptonie doublets) \Rightarrow creation of 9 quarks + 3 leptons

$$\partial_\mu j_\mu^B = \eta_f \frac{g^2}{32\pi^2} F_{\mu\nu} \tilde{F}_{\mu\nu}; \quad \tilde{F}_{\mu\nu} = \frac{1}{2} \epsilon_{\mu\nu\rho\sigma} F^{\rho\sigma}$$

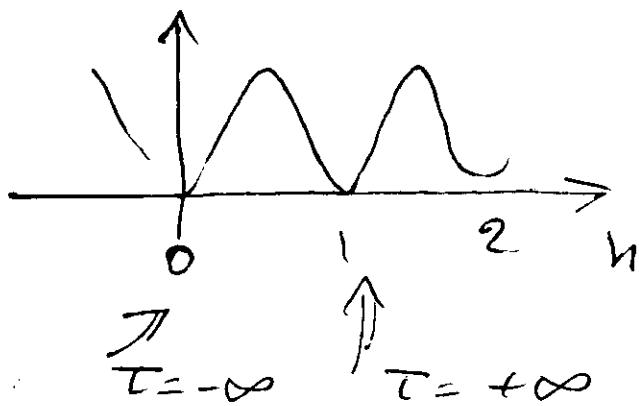
$$\partial_\mu j_\mu^L = \dots$$

$$\frac{g^2}{32\pi^2} F \tilde{F} = \partial_\mu K_\mu; \quad \text{Chern-Simons current.}$$

$$K^\mu = \frac{g^2}{32\pi^2} \epsilon^{\mu\nu\rho\sigma} \left\{ F_{\nu\lambda} A_\rho^a A_\sigma^a - \frac{2}{3} g \epsilon^{abc} A_\nu^a A_\lambda^b A_\sigma^c \right\}$$

Energy barriers

4



$$E_B = \min_{\text{paths}} \max_{\{\tau\}} \underbrace{E(\tau)}_{\text{static energy}}$$

Configuration saturating the minimax procedure: unstable solution to the classical eq. of motion.

sphaleron:



Barrier in 3+1 case:

50

sphaleron:

$$F_i^a = \frac{1}{g} \frac{\epsilon_{ijk} x_j}{r^2} \cdot f(\xi) \quad \xi = g v \sqrt{\epsilon} / |\vec{x}|$$

$$\varphi = i \frac{x^a x^a}{r} h(\xi) \begin{pmatrix} 0 \\ v \end{pmatrix} \quad f(0) = h(0) = 0 \\ f(\infty) = h(\infty) = 1$$

$$E_{\text{sph}} = \frac{2M_W}{\partial w} \delta \left(\frac{\lambda}{\partial w} \right) = \begin{cases} 8 \text{ TeV}, & M_H \ll M_W \\ 14 \text{ TeV}, & M_H \gg M_W \end{cases}$$

What is the origin of the scale $\frac{M_W}{\partial w}$?

$$E = \int d^3x \left[\frac{1}{4} (F_{ij})^2 + (\mathcal{D}_i \varphi)^2 + V(\varphi) \right]$$

dimensionless variables

$$x = \frac{1}{M_W} \xi, \quad A = \frac{M_W}{g} \tilde{A}, \quad \varphi = \frac{v}{g} \tilde{\varphi}$$

$$E = \frac{M_W}{g^2} \int d^3\xi \left[\frac{1}{4} (\tilde{F}_{ij})^2 + (\mathcal{D}_i \tilde{\varphi})^2 + \frac{\lambda}{g^2} (\tilde{\varphi}^2 - 1)^2 \right]$$

$$\sim O\left(\frac{M_W}{g^2}\right)$$

30

Important properties of the sphaleron

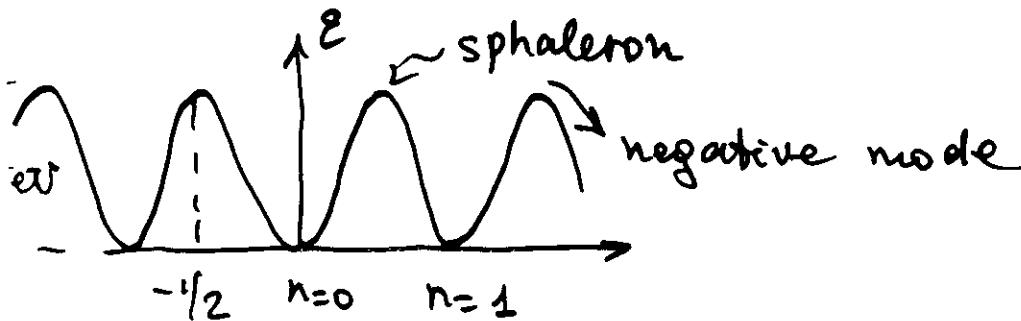
(i) topological charge = $\frac{1}{2}$

$$\frac{g^2}{32\pi^2} \int_0^{t_0} F_{\mu\nu} \tilde{F}_{\mu\nu} = \frac{1}{2}$$

$$t=0: \quad \varphi = \begin{pmatrix} 0 \\ \nu \end{pmatrix}, \quad \delta = 0$$

$$t=t_0: \quad \varphi = \varphi_{\text{sph}}, \quad \delta = \delta_{\text{sph}}$$

(ii) it has exactly one negative mode



(iii) fermionic zero modes is sphaleron background

β non-conservation under usual conditions:

tunneling through the barrier

Amplitude : $\exp\left(-\frac{2\pi}{\alpha}\right) \sim \exp(-180)$
too small

HIGH TEMPERATURES!

Rate of sphaleron transitions at
high T : (7)

sphaleron : analogy with critical bubble

Rate : $T < T_c$:

$$\Gamma \sim \exp\left(-\frac{M_{\text{sph}}}{T}\right) \approx$$

$$\approx T^4 \left(\frac{dw}{4\pi}\right)^4 N_{\text{tr}} N_{\text{rot}} \left(\frac{2M_w}{dwT}\right)^7 \exp\left(-\frac{\cancel{M_{\text{sph}}}}{\cancel{T}}\right) \propto$$

temperature dependent
sphaleron mass.

$T > T_c$:

no barrier

$$\Gamma = (dwT)^4 \quad [\text{from } M_w \approx dwT]$$

Cosmology :

compare rate with the rate of
universe expansion

$$t = \frac{M_{\text{pl}}}{T^2} \Rightarrow$$

equilibrium for $\delta S \neq 0$

$$100 \text{ GeV} \leq T \leq 10^{12} \text{ GeV}$$

Electro weak baryogenesis

①

Sakharov conditions

- (i) B - non - conservation
- (ii) CP - violation
- (iii) thermal non - equilibrium

How it can happen?



$\text{VEV} \neq 0$

sphaleron rate is
high -
no B - generation

\mathcal{T}

domain wall baryogenesis.

Out of equilibrium condition ;
upper limit on the Higgs mass

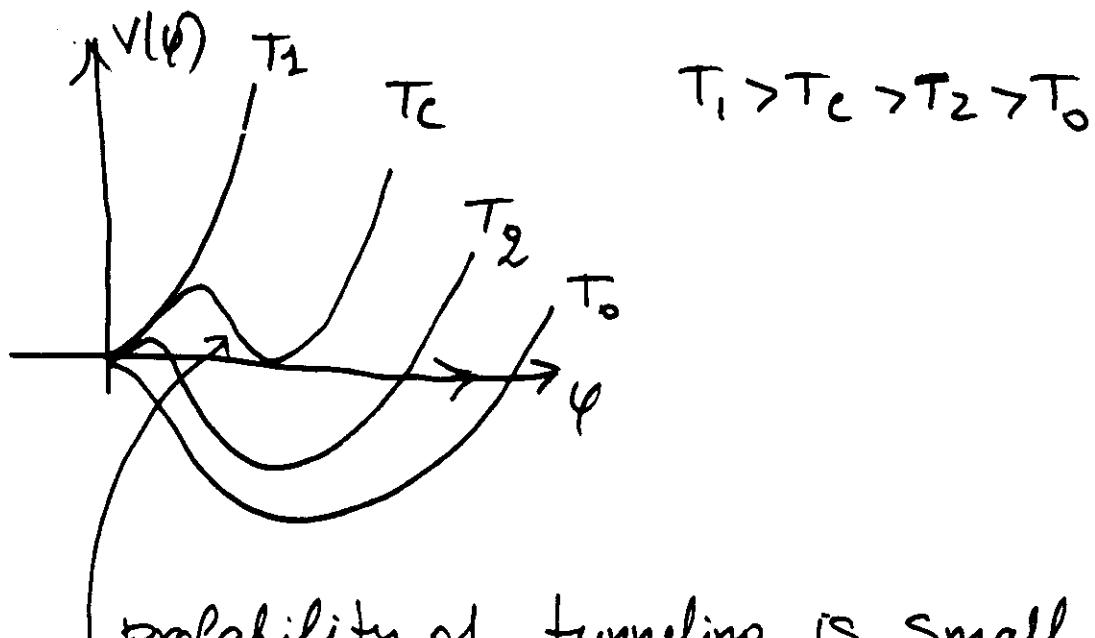
Scalar potential:

(2) ~~5~~

$$V(\varphi) = \mu^2(T)\varphi^2 - \delta T\varphi^3 + \lambda\varphi^4$$

$$\mu^2(T) = \frac{5}{16}g^2(T^2 - T_0^2), \quad \delta = \frac{3}{32\pi}g_w^3$$

Phase transition occurs near $T=T_0$



probability of tunneling is small,
at $T \approx T_0$ it is of order of 1.

or just after the phase transition:

$$\left. \frac{dV}{d\varphi} \right|_{T=T_0} = 0 \Rightarrow \varphi = \frac{3\delta T}{\lambda}$$

Sphaleron mass after phase transition

③ ~~3~~

$$M_{\text{sph}} \approx \frac{3M_W}{d_W} = \frac{3}{d_W} \cdot \frac{1}{2} g_W \cdot \frac{3\delta T}{\lambda}$$

Out of equilibrium condition:

$$\frac{M_{\text{sph}}}{T} \geq 45 \Rightarrow \lambda \leq \underbrace{\frac{3}{80} g_W^2}$$

upper bound on the Higgs mass

$$m_H \lesssim 45 \text{ GeV} \leftarrow \text{one-loop result.}$$

Theory with two Higgs doublets:

$$m_H \lesssim 120 \text{ GeV.}$$

One-loop result should not be trusted: infrared divergences

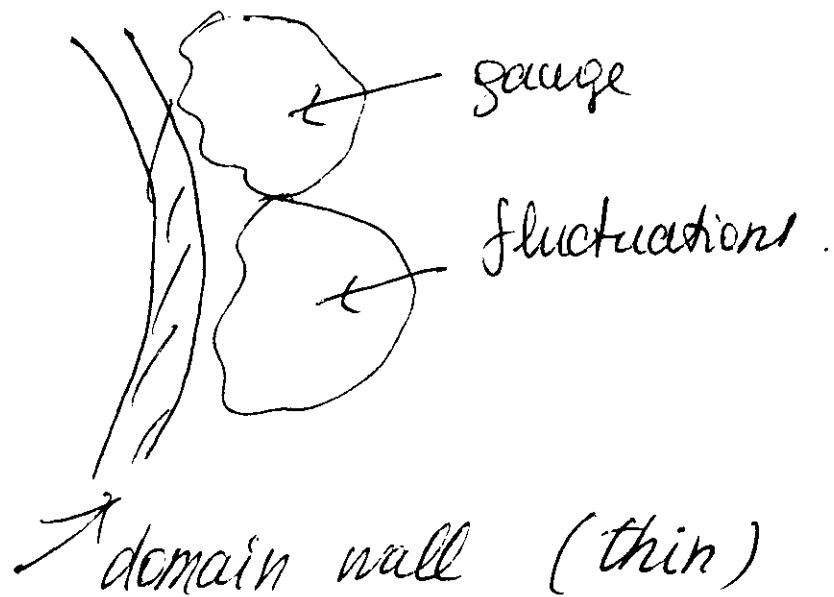
Non-perturbative effects, lattice simulations $\Rightarrow m_H^{\text{crit}}$ may be as large as $m_H^{\text{crit}} \simeq 80 - 100 \text{ GeV}$



Two general types of mechanisms

- (i) Bosonic fields are important
- (ii) fermions are important.

(i) "Bosonic" mechanism (a)



CP-violation :

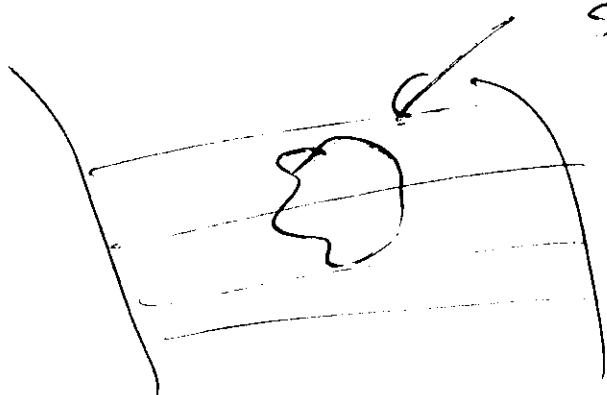
fluctuations decay and produce
non-zero $\tilde{F}\tilde{F}$

$$\begin{cases} \tilde{F}\tilde{F} > 0 \\ \tilde{F}\tilde{F} < 0 \end{cases} \text{ due to CP} \Rightarrow$$

fermions are created.

(5)

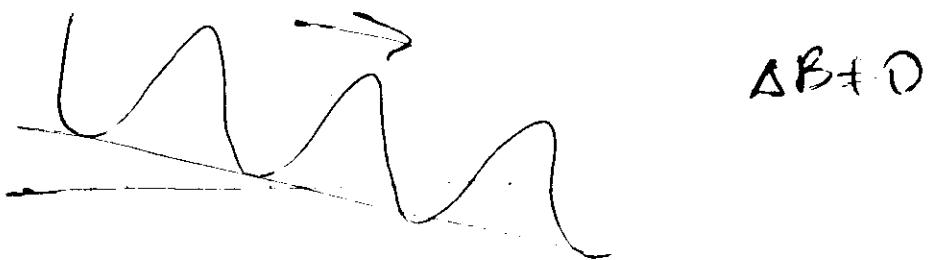
Bosonic mechanism (B)



sphaleron transitions
inside domain
wall

wall (thick)

time varying background :



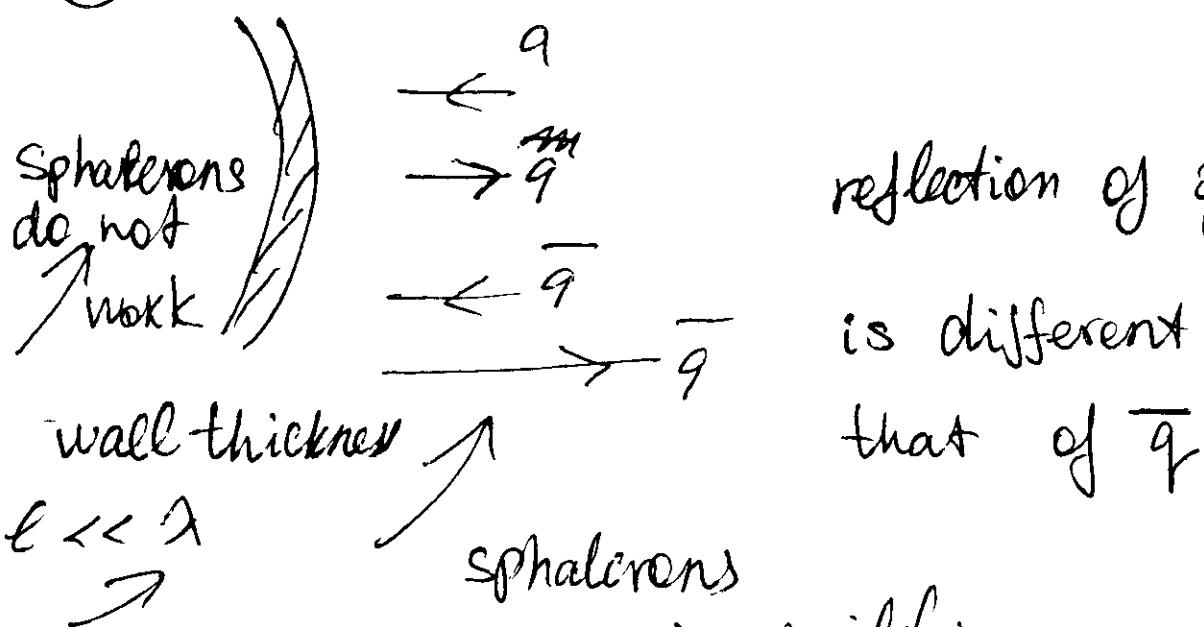
$$\Delta B \neq 0$$

② fermionic mechanism

⑥

charge transport:

① thin wall



reflection of quarks
is different from
that of \bar{q}

wall thickness

$\ell \ll \lambda$

sphalerons

mean free path are in equilibrium

if $T_q > T_{\bar{q}}$: quarks inside
survive

antiquarks outside
disappear \Rightarrow

$SB \neq 0$

② thick walls \rightarrow

interaction of quarks with gluons
inside the domain walls should
be taken into account

Results:

① MSM :

$$\frac{n_\Delta}{n_{\bar{S}S}} \approx 10^{-10} - 10^{-18}$$

Farrar, MS

may be OK.

② Extended EW theories

- (i) two Higgs doublets - OK
- (ii) SUSY SM - may be OK

Phase transitions:

Effective potential:

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D. A. K., Linde, PL 42B (1972)

D. A. K., A. D. L., Ann. Phys. 101 (1976) 195

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A. D. Linde, book.

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Linde, Phys. Lett. 188D

Gross, Pisarski, Yaffe, Rev. Mod. Phys (1981)

Appelquist, Pisarski, R.R. D23 (1981) 2385

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M.S., P.L. B316 (1993) 112

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Energy barriers (sphalerons)

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