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SMR.762 - 9

SUMMER SCHOOL IN HIGH ENERGY PHYSICS AND COSMOLOGY

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NEUTRINO PHYSICS

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ICTP

Please note: These are preliminary notes intended for internal distribution only.

IN THREE FEATURES:

NEUTRINO
PHYSICS
OR
LOOKING FOR NOTHING

SUM KNOW 1 (1)

IN SEARCH FOR NOTHING

15 minutes?

WHAT IS MAIN, BASIC?

WHAT WE KNOW?

SUMMARIZE

DEF (1) Neutrino is

→ LIGHT OR MASSLESS $m \approx 0$

→ NEUTRAL: $Q = 0$; $\bar{Q} = 0$

→ WITH SPIN $1/2$

PARTICLE

INTERACTIONS

~~NEUTRAL LEPTON~~ → ~~WEAK INTERACTIONS~~ ~~GRAVITY~~

EM INTERACTIONS
IN HIGH
ENERGIES

3rd
LEPTON

(2) THREE NEUTRINOS: ν_e, ν_μ, ν_τ
AND CORRESPOND
TO CHARGE e, μ, τ .

→ THIS CORRESPONDENCE IS DESCRIBED
BY LEPTON NUMBER

EM

(3) NEUTRAL LEPTON → weak interaction.

NO EM, STRONG →

• HUGE PENETRATION ABILITY, ESPECIALLY
AT LOW E

• rise of σ with E
(cross section)

• THE INTERACTIONS ARE WELL DESCRIBED
BY STANDARD MODEL
Moreover

ν played a crucial role in
building and check of SM

→ NC. in ν interactions

(4) STABLE (1)

$\nu \rightarrow \nu \gamma$
 ϕ

BOUND

• Has uniparticle
 $\bar{\nu}_\mu$

~~WATSON~~

MAJORANA?

$\nu = \bar{\nu}$?

ABUNDANCE

UNIVERSAL
 MAIN COMPONENT IN THE
 Universe:

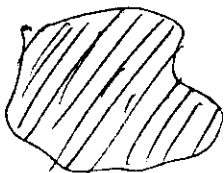
ν :
 DENSE DENSITY: $\rightarrow$ MORE THAN
 NUMBER OF PHOTONS

UNIVERSAL

NATURAL
 SOURCES

- RELIC NEUTRINOS.
- PRODUCED
 SUN.
- ~~PRODUCED~~
- COLLAPSING STARS
- APOG.

EARTH'S ATMOSPHERE



$6 \cdot 10^{20} \approx 10^4$
 $\approx 10^{16} \nu \text{ set cm}^{-2}$

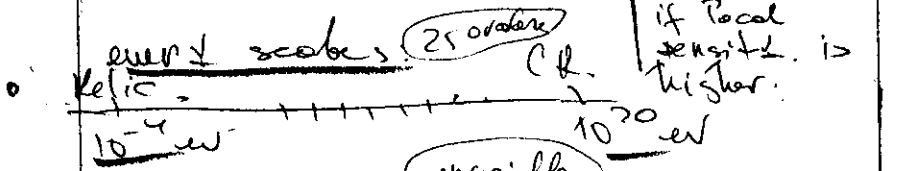
ELEMENTARY PROC.
 $\mu \rightarrow \nu \mu$ e^- capture
 $p \rightarrow \nu \pi^0, \pi^+, \mu, \dots$
 ν annihilation
 ... NEUTRINOS.

ARTIFICIAL:
 * ACCELERATORS
 * SOURCES $=$ 3H 60Co
 * NEUTRON FACTORIES 235U

Normal:
 $2.5^\circ C$.

RELIC

Relic $4 \cdot 10^2 \cdot 3 \cdot 10^{10} \cdot 10^4 \approx 10^{17} 10^{-4}$



ν give appreciable (main) contribution
 TO ENERGY DENSITY IN THE
 universe $\rightarrow$ DYNAMICS OF
 THE UNIVERSE

$\gamma = \frac{E}{h\nu}$
 $\gamma = \frac{1}{\beta}$
 $\beta = \frac{v}{c}$
 $\gamma = \frac{1}{\sqrt{1 - \beta^2}}$
 $\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}$

IN MINIMAL VERSION:

No ν_R

~~→ NO MASSLESS~~

NO coupling ~~like~~ with Higgs which
GENERATES MASSES of CHARGED LEPTON

~~$m_\nu = 0$ NO LEPTON NUMBER~~
~~but even in SM one can~~

Lepton number conservation

l_e, l_μ, l_τ

INTERACTION

idle

$\nu_e \rightarrow e \quad \nu_e$

$\nu_e, \nu_\mu \rightarrow e \quad \nu_e$

THE LARGER THE SCALES THE MORE IMPORTANT ROLE OF NEUTRINOS:

IMPORTANT FOR

ASTROPHYSICS
COSMOLOGY

→ Responsible for
→ LARGE SCALE structure of the
universe

→ PLAY CRUCIAL ROLE IN THE
EVOLUTION OF STARS. !

→ ~~STAR~~, ~~GRAVITATIONAL~~ COLLAPSE
SN

→ cooling of stars

→ Nucleosynthesis of the elements
in the universe

- Direct
- Indirect → expansion of the
universe.
2 - 4 neutrinos → t_e
- Nucleosynthesis in SN.

Problems

✓ AFTER t - quak.
• π

→ NEUTRINO MASSES

(NATURE
OF NEUTRINO)

→ Lepton Mixing -

→ b - conservation.

→ New interactions

* ~~LABORATORY~~ experiments →
Bounds

(FROM TIME TO TIME
SOME INDICATIONS:

* More steady (difficult to
check.

• solar neutrinos

• ATMOSPHERIC NEUTRINOS

• Cosmological - large scale

- ITED
- RUGBY
- 17 keV
- 7 pA

- oscillation at Los Alamos

NEUTRINO PHYSICS

PURPOSE

Base what happens in the sec

CONSERVATION

COVERS THE FOLLOWING TOPICS

MASS AND LEPTON MIXING

I. INTRINSIC NEUTRINO PROPERTIES

(WHAT WE KNOW AND WHAT IS NOT KNOWN)

II. PROPAGATION OF MIXED NEUTRINOS IN VACUUM AND MATTER

III. "NEUTRINO ANOMALIES" ASTROPHYSICS HINTS AND NEUTRINO MASS SPECTRUM

NEUTRINO (2)

1. LIGHT OR MASSES
2. NEUTRAL $Q_0 = Q_c = 0$
3. SPIN $1/2$

ν AND $\bar{\nu}$ LEPTON NUMBER VIOLATION.

THREE TYPES

NEUTRINOS ANTINEUTRINOS

EXPERIMENTAL BASIS:

- ALL THE ACCELERATORS
- LOW ENERGY REACTOR
- NUCLEON FACTORIES

THE NEUTRINOS IN THE ATMOSPHERE

RICHNESS AUG

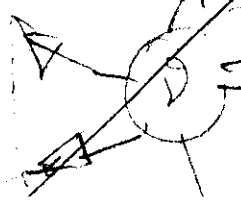
10^{-7} eV 10^{21} eV

COSMOLOGY LARGE SCALE STRUCTURE

NUCLEOSYNTHESIS

UNDERSTANDING EXPERIMENT

A REVIEW AND A SUMMARY



I. NEUTRINO PROPERTIES

I. ν IN THE STANDARD MODEL

• $\begin{pmatrix} \nu_e \\ e \end{pmatrix}_L, \begin{pmatrix} \nu_\mu \\ \mu \end{pmatrix}_L, \begin{pmatrix} \nu_\tau \\ \tau \end{pmatrix}_L$

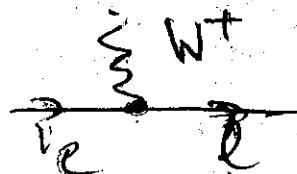
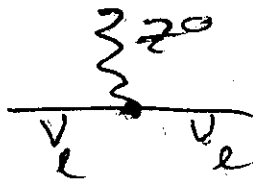
• DOUBLETS $\xrightarrow{\text{FORM}} SU_2$ Doublets

• THREE NEUTRINOS

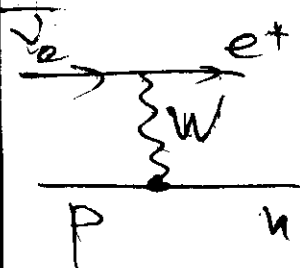
• CAN BE CHARACTERIZED BY

• $T_3 \nu = \frac{1}{2} \leftarrow \text{weak isospin}$
 $Y = -1 \leftarrow \text{hypercharge}$

$Q = 0 \leftarrow \text{No direct couplings with photon}$



fix "flavor" of NEUTRINO STATE



IN SUSY?

• lepton $\rightarrow$ singlet of SU_3

link to what
REKIND

FORM DOUBLET

THIS MEANS

CHARGES fix Interactions

THIS INTERACTION
DETERMINES THE FLAVOR

• VERTEXES CRUCIAL ROLE
IN CONSTRUCTING
ESTABLISHING THE SM

• NOW CHECK WITH
VERY HIGH PRECISION

• NEUTRALITY
INTERPRETING
DOES PLACE
OF ν IN NATURE

RH. NEUTRINOS?

PH. Neutrinos
STERILE
B-L

ν_R ?

DO THEY EXIST?

Why not? SYMMETRIES
Why charges conserved? HAVE

Suppose ν exist:

$$T_3 = 0$$

$$Y = 0$$

$$Q = 0$$

Why it is so crucial

In contrast with CHARGED FERMIONS
NO ELECTROMAGNETIC INTERACTIONS W/ (Z, γ)

~~STERILE~~

• INTERACTIONS with Higgses

→ MASSES OF NEUTRINOS

Why they are suppressed.

ACTUALLY:

THE EXISTENCE OF ν_R
AND THEIR PROPERTIES
(MASSES, CHARGES)

→ MAIN ISSUE OF ν -physics

→ NO ν_R
→ VERY HEAVY
→ SYMMETRY

WHY THIS SYMMETRY DOES NOT DETECT MASSES OF FERMIONS

The problem of ν -physics
is can be reduced to

ν_P

Pauli: had solved the problems of p -neutrinos
→ something very bright and
very difficult to observe

Now we reduce the problem of ν to
something which has even more weak interactions

interaction with Higgses
 ν -INTERACTIONS

BEYOND

STOCHASTIC CONVERGENCE

WHETHER ν_P BEYOND

As we will see.

DO THEY EXIST
• WHAT IS THE MASS
• HOW THEY INTERACT WITH "active" sector

FLAVORS

FLAVOR STATES: fixed by CC weak interactions
 ν_e, ν_μ, ν_τ correspond to Charge Lepton Numbers

$\rightarrow$ up to now $\rightarrow$ no lepton number violation

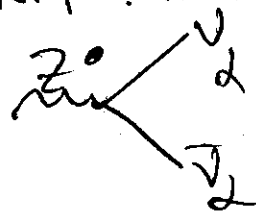
$\nu_e \leftrightarrow \nu_\mu$ affects no transition between neutrinos

difference with q -sees

NUMBER OF NEUTRINO SPECIES: N_ν

ALREADY DISCUSSED

KEP: invisible width of Z^0



Γ_{inv} or N_ν
 $N_\nu = \frac{\Gamma_{inv}}{\Gamma_\nu}$

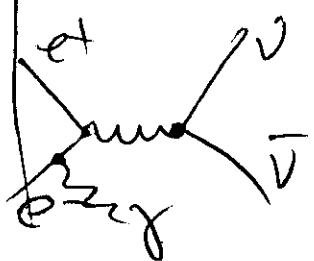
$\Gamma_{inv} = \Gamma_{tot} - \Gamma_{vis}$

$N_\nu = 2.985 \pm 0.023$ (LIVE SHAPE from all 4 exp. 90% CL)

IMPLICATIONS: THREE NEUTRINOS NO NEW

$M_\nu > 97$ eV

- THREE generations
- BOUNDS ON new physics
 - $\rightarrow$ new additional channels
 - "NON-SINGLET MAJORONS" - excluded
 - $\rightarrow$ MIXTURE OF HEAVY LEPTONS



$N_\nu = 3.04 \pm 0.16$

radiative corrections

$N_\nu = \frac{\Gamma_{inv}}{\Gamma_\nu} = \frac{\Gamma_{inv}}{\Gamma_e} \left(\frac{\Gamma_e}{\Gamma_\nu} \right)_{SM}$
 $= \left(\frac{\Gamma - \Gamma_e - \Gamma_h}{\Gamma_e} \right) \left(\frac{\Gamma_e}{\Gamma_\nu} \right)_{SM}$
 MANY implications
 basic figures

implications $\rightarrow$ to ν physics
 SHAPE

⊗ NUCLEOSYNTHESIS

Bound on number of light degrees of freedom

$$(n < 40 \text{ MeV})$$

$$N_V < 3.3$$

$$3.01$$

criticized

NUCLEOSYNTHESIS AND NUCLEOSYNTHESIS

→ direct / indirect INFLUENCE

COSMOLOGY

MISSSES

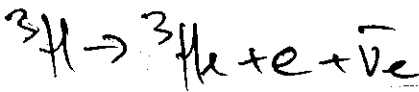
→ ROOUNDS

direct KINEMATICS

① ν_e

$$W \leq 4-5 \text{ eV}$$

(MAINZ) INE-KINE



18 keV

$$W(\nu_e) / W_0 < 10^{-5}$$

$$W^2 = -8 \pm 12 \text{ eV}^2$$

$$(-39 \pm 34 \pm 15)$$

SYSTEMATICS, $W < 4 \text{ eV}$ 3-

② ν_μ



PST:

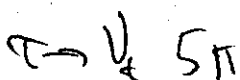
rest $\uparrow$ W_π P_μ W_μ

$$W < 220 \text{ keV}$$

$$W(\nu_\mu) / W_\pi \leq 2 \cdot 10^{-3}$$

PST

③ ν_e



$$W(\nu_e) < 31 \text{ MeV ARGUS}$$

$$< 32.6 \text{ CLEO}$$

$$\rightarrow 29 \text{ MeV}$$

Nucleosynthesis!

⊗

$$W_0 < 0.1 \text{ MeV} \quad \tau > 10^3 \text{ s} \quad \tau > 10^3 \text{ s.}$$

Cosmology: (energy density) $\rightarrow$ expansion rate $\rightarrow$ age

(11)

$$\sum_i m_i < 50 \text{ eV}$$

20-

DIRAC

MAJORANA

(diff. pend)

← THESE BOUNDS DO NOT DEPEND ON NATURE

SN 187 A

$m < 20 \text{ eV}$

spread of signal

RESUME

→ MUCH SMALLER THAN CHARGE LEPTON MASS

→ SPLITTING

CHARGE

↓

NEUTRINO

↓

→ SINGLED OUT

SWALLIES OF ν_{μ}

9

NEUTRALITY

→ OTHER BOUNDS ON ν

→ LEPTON NUMBER VIOLATION

→ NOT PURE BOUND BUT

✓

MODEL FOR ν -PMNS VIOLATION

MODEL FOR ν -PMNS VIOLATION

MASS (THEORY - UE)

(14/7)

THIS SECTION NUMBERS -

To have further some more THEORETICAL CONSIDERATION

① IN SM: ν_L - WEIL SPINOR
MASSLESS.

WEIL $m_\nu = 0$

CP: $\bar{\nu}$ (two component)
L-CONSERVATION

WEIL - DIRAC
MASS TERM

② If there are ν_R : suppose there is ν_R

Dirac particle is

$$\mathcal{L} = \dots - m_D (\bar{\nu}_L \nu_R + \bar{\nu}_R \nu_L)$$

$$\Psi = \nu_L + \nu_R \quad m_D (F.4)$$

4 components

③ Can two components neutrino
HAVE A MASS?

MASS TERM CHANGE
CHIRALITY

$$\nu_L \xrightarrow{m_D} \nu_R$$

ν_L

$$C: \quad C(\nu_L)^c = \nu_R^c$$

CHARGE CONJUGATION

$$(\nu_L)^c \rightarrow C(\bar{\nu}_L) \rightarrow \nu_R^c$$

$$C = i\gamma^2 \gamma^0$$

$$-\frac{1}{2} m (\bar{\nu}_L \nu_R^c + \bar{\nu}_R^c \nu_L) = m \bar{\nu}_M \nu_M$$

$$\nu_M \equiv \nu_L + \nu_R^c$$

WHAT IS THE SENSE?

Allow to introduce

NOTE

C - operator

change the sign
of the mass

$$(1+\gamma^5)$$

$$\gamma^5 = -\gamma^5$$

Properties:

$$1. \nu_M^c = (\nu_L)^c + \nu_L = \nu_M$$

NEUTRINO $\equiv$ ANTI-NEUTRINO

true Neutral $\Rightarrow$ MAJORANA

2. It is impossible to introduce conserving lepton numbers

3. $L(\nu_L) = 1$ then
MASS TERMS: $|N| = 2$
 $\Rightarrow$ effect of lepton number violation

$$4. \tilde{\nu}_\mu = \nu_L - \nu_R$$

MASS TERM $\rightarrow +\frac{1}{2}m \tilde{\nu}_M \tilde{\nu}_M$

properties $\tilde{\nu}_\mu^c = -\tilde{\nu}_\mu$ opposite CP PARITY
 $\equiv$ Different sign of MASS

$$\left(\begin{aligned} \nu_L &\rightarrow \nu_L = i\nu' = e^{i\pi} \nu' \\ \nu_R^c &= -i\nu_R^c \\ \tilde{\nu}_M &= i(\nu_L' + \nu_R^c) = \\ \tilde{\nu}_M &= i\nu_M' \end{aligned} \right) \text{ - relation}$$

for ν_M - positive mass. but the interaction will be changed.

$$m(\nu_L^T \nu_L + h.c.)$$

For those who know
 $\rightarrow$ INTRODUCTION OF NOTATIONS

can exist for neutral fermion

suppose all terms but MASS conserve L

$\tilde{\nu}$ - HAS

$\rightarrow$ different sign of mass $\Rightarrow$ works to different CP parity

see below

sign of mass can be removed by redefinition of the respective field sign is important.

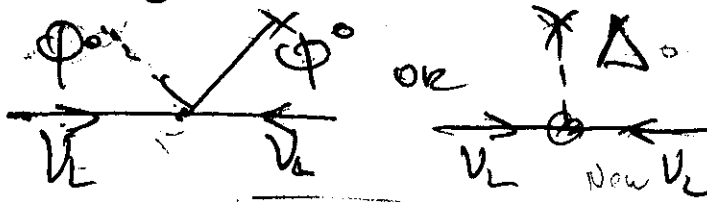
change the coupling

CAN BE written check.

NEUTRINO MASSES AND SM

• NO RH.

$\psi_l(\nu)$



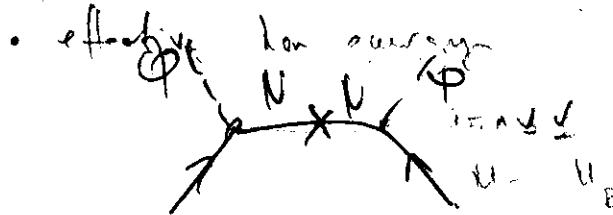
$$T_3 = 1$$

effective operator

$$\frac{1}{M} \psi^T \psi \cdot \Phi \Phi$$

GAUGE INVARIANT
BUT NON RENORMALIZABLE
← origin?

something beyond



• NON RENORMALIZABLE
⇒ GRAVITY?

$$M = M_R$$



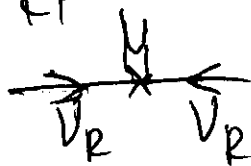
← FROM UNITARITY

• WITH RH: DIAC



$$T_3 = \frac{1}{2}$$

• RH



$$T_3 = 0$$

← origin?

MOST GENERAL
 BOTH MAJORANA AND DIRAC

$$\frac{1}{2} \begin{pmatrix} \nu_L \\ \nu_L^c \end{pmatrix} \begin{pmatrix} \nu_R^c & \nu_R \\ m_L & m_D \\ m_D & m_R \end{pmatrix}$$

$\begin{pmatrix} \nu_L \\ \nu_L^c \end{pmatrix} M \begin{pmatrix} \nu_R^c \\ \nu_R \end{pmatrix}$ MASS MATRIX

DEPENDENT ON RELATIONS BETWEEN THE ELEMENTS
 $m_L, m_R, m_D \rightarrow$ diagonal

$$\begin{pmatrix} \nu_L & \nu_L^c \\ \nu_L^c & \nu_R \end{pmatrix} \frac{1}{2} \begin{pmatrix} m_L & m_D \\ m_D & m_R \end{pmatrix} \begin{pmatrix} \nu_R^c \\ \nu_R \end{pmatrix}$$

SPECIAL CASES:
 $m_L, m_R \ll m_D$
pseudodirac neutrino

*

SEE SAW

1) suppose there is no MINIMAL STRUCTURE:

$$\Delta \rightarrow \boxed{m_L = \infty}$$

induced $\rightarrow$ small

AND NEUTRINOS HAVE

2) USUAL DIRAC MASS TERM

$$\boxed{m_D \sim m_e, m_\mu}$$

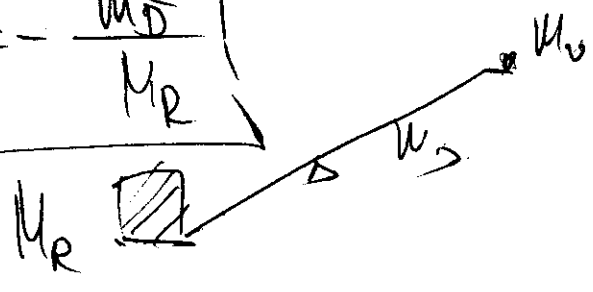
NATURAL

• ν_R HAVE MAJORANA MASS TERMS

$$\boxed{m_R \gg m_D}$$

$\Rightarrow$ DIAGONALIZATION gives

$$\boxed{m_e = - \frac{m_D^2}{m_R}}$$



see-saw relates

SHALLNESS
OF
MASSES
IN COMPARISON

NEUTRALITY

ν -Neutral

ν_R
MAY HAVE
MAJORANA
MASS

ν_R IS
SINGLET of $SU_2 \times U_1$
 ν_R is unprotected

ν_R :

M_D ?

AGAIN

unprotected

b - is violated by mass term
OF ν_R

→ explicit (BARE MASS TERMS)

→ spontaneous
 $\nu_R^T C \nu_R S$

MAJORANA

(spontaneous violation)

→ MAJORANA

→ CAN BE GAUGED AS IN SO_{10}

[in SUSY]

gravitational effects

→ BOUNDS ON
SCALE OF
SPONTANEOUS
VIOLATION

mixing with LH.

$$\frac{m_D}{M_R} \ll 1$$

if ν_R light → mixing
bound,