



INTERNATIONAL ATOMIC ENERGY AGENCY
UNITED NATIONS EDUCATIONAL, SCIENTIFIC AND CULTURAL ORGANIZATION
INTERNATIONAL CENTRE FOR THEORETICAL PHYSICS
I.C.T.P., P.O. BOX 586, 34100 TRIESTE, ITALY, CABLE CENTRATOM TRIESTE



H4.SMR/842-15

**Conference on
Ultrafast Transmission Systems in Optical Fibres**

13 - 17 February 1995

Upgrading the Installed Fibre Base

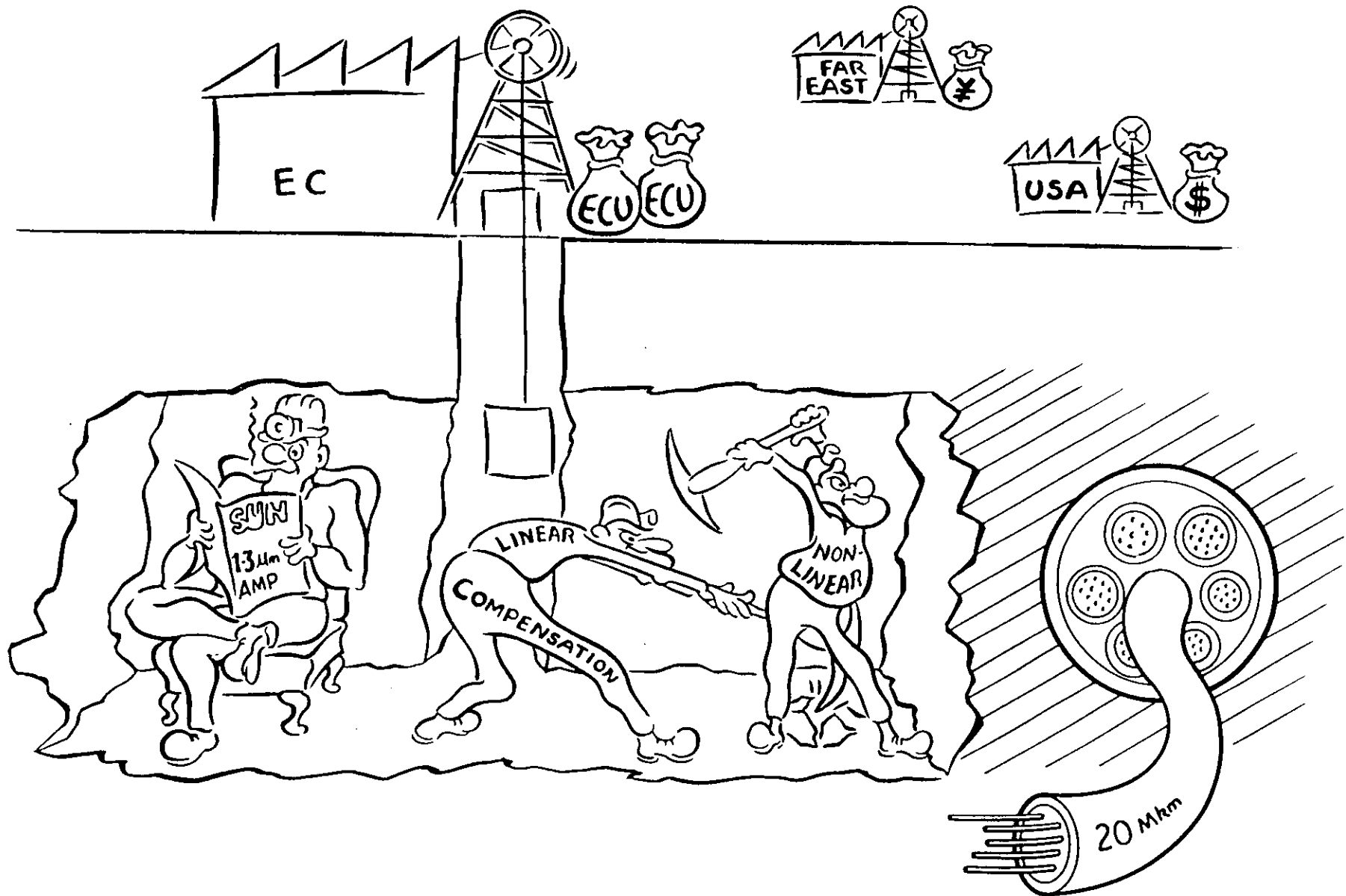
R.I. Laming

**Optoelectronics Research Centre
Southampton University
Southampton, United Kingdom**

Acknowledgements

MN Zervas
A Royset
SY Set
I Goncharenko
D Hewak
B Samson
S Barcelos
DN Payne

THE FIBRE GOLD MINE



1.3 μ m NETWORK UPGRADE

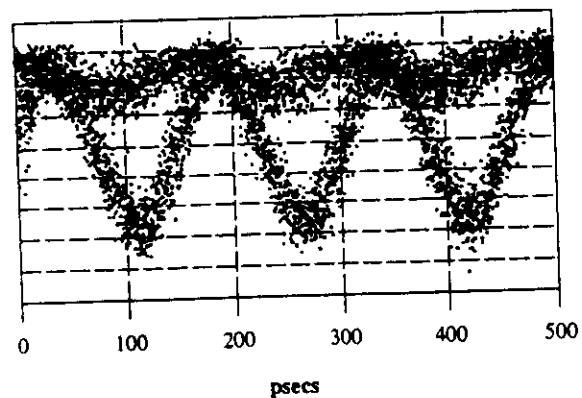
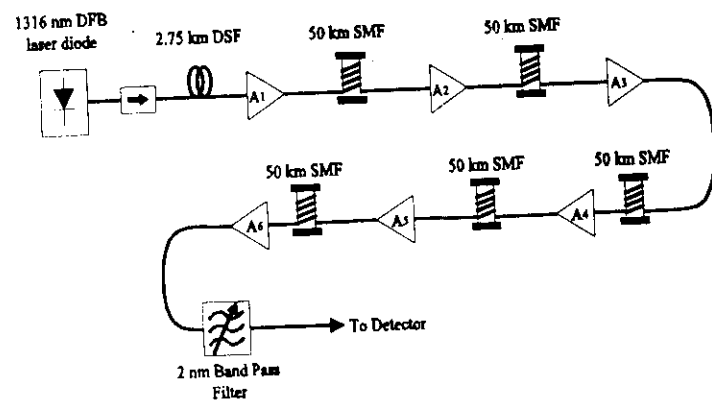
- Loss compensation at 1.3 μ m
 - Doped fibre amplifiers
 - Raman amplifiers
 - ➡ Soliton transmission at 1.3 μ m
- Dispersion compensation for 1.5 μ m
 - Solitons
 - Source pre-chirp
 - Equalizing fibre
 - Chirped fibre gratings
 - Phase conjugation

1.3 μ m SOLITON TRANSMISSION ON STANDARD FIBRE

The issues:

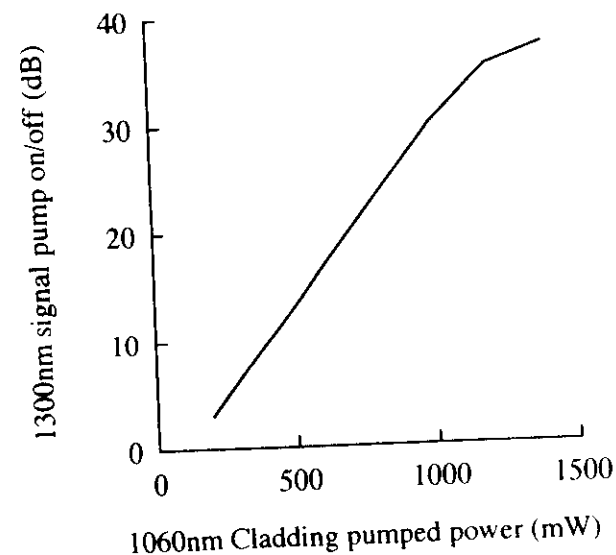
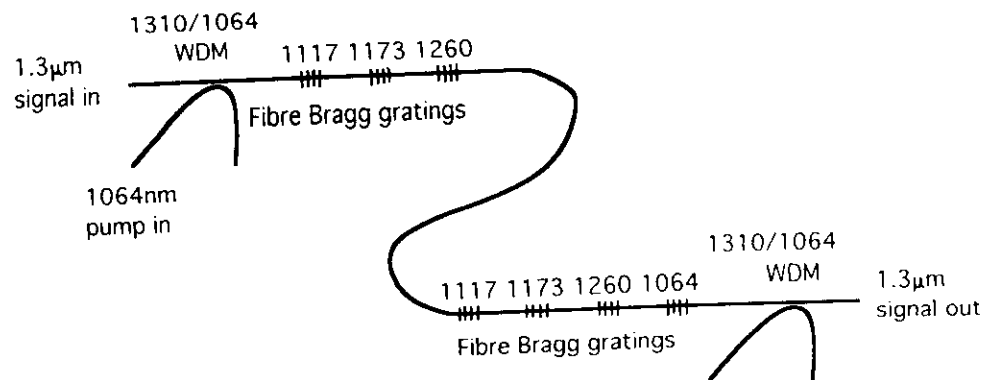
- Fibre loss
- Fibre dispersion
- Polarisation mode dispersion
- Amplifiers
 - spacing
 - power

6.5 Gbit/s, 1316nm SOLITON TRANSMISSION OVER 250km OF STANDARD FIBRE USING SEMICONDUCTOR QUANTUM WELL AMPLIFIERS



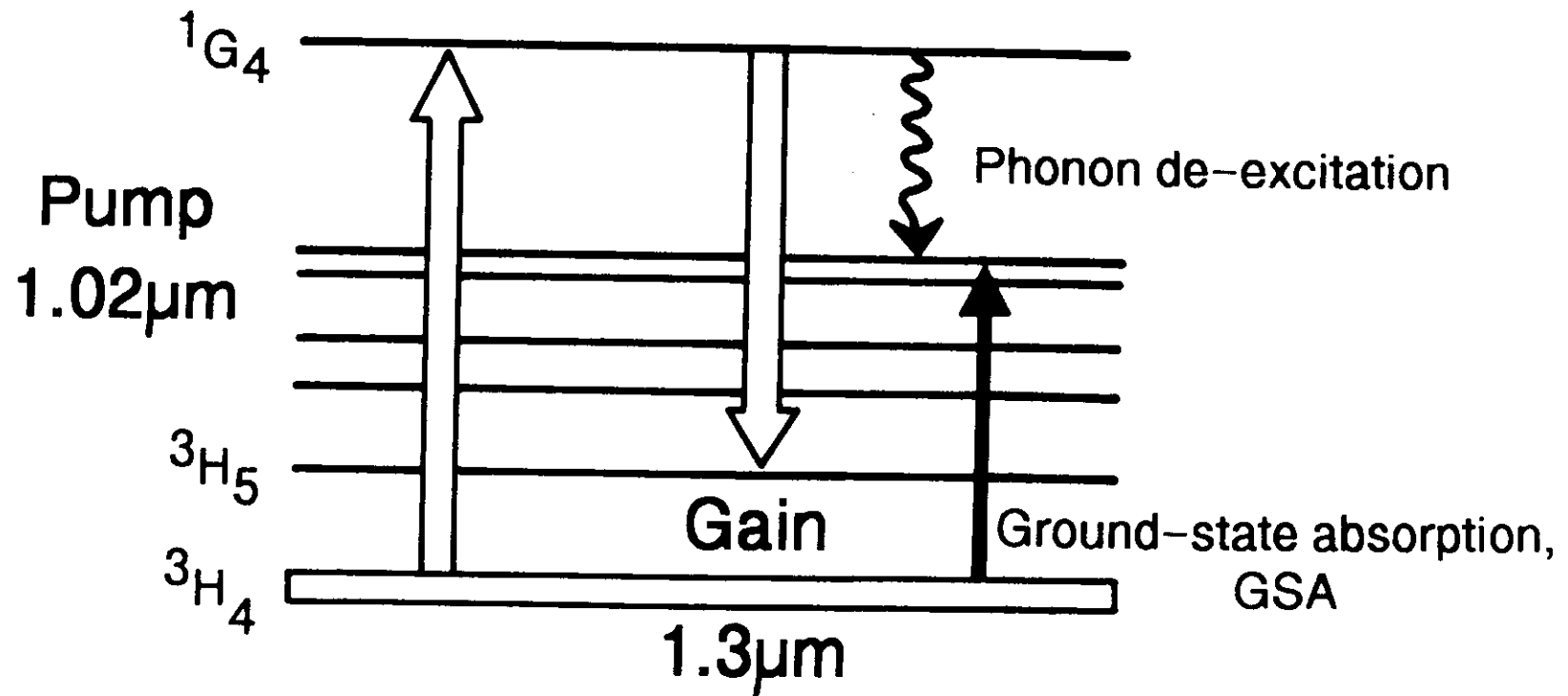
Liedenbaum et al, ECOC '94

1.3 μ m AMPLIFIERS - RAMAN CASCADE



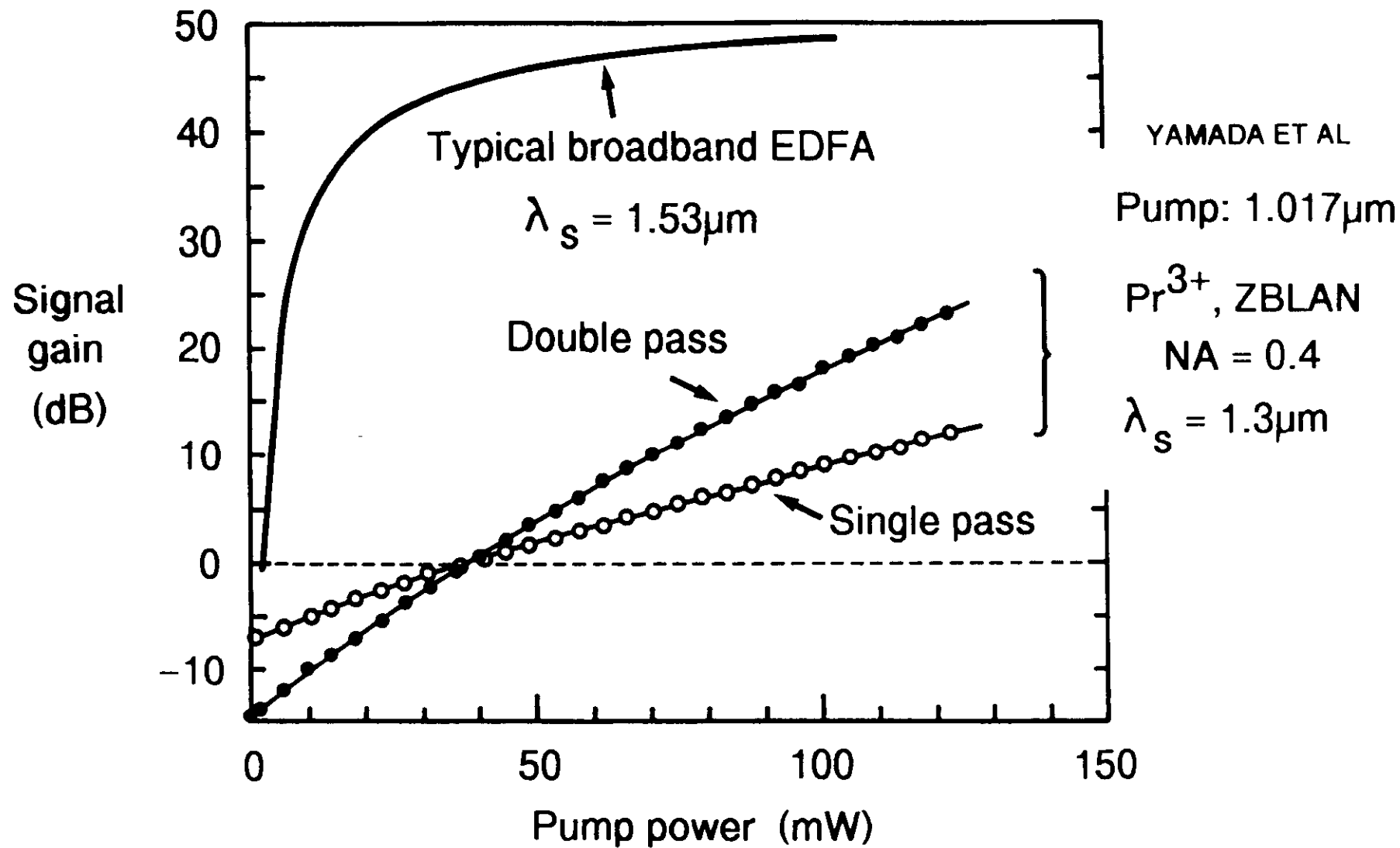
S Grubb et al, OAA '94

Pr^{3+} DOPED ZBLAN ENERGY LEVEL DIAGRAM

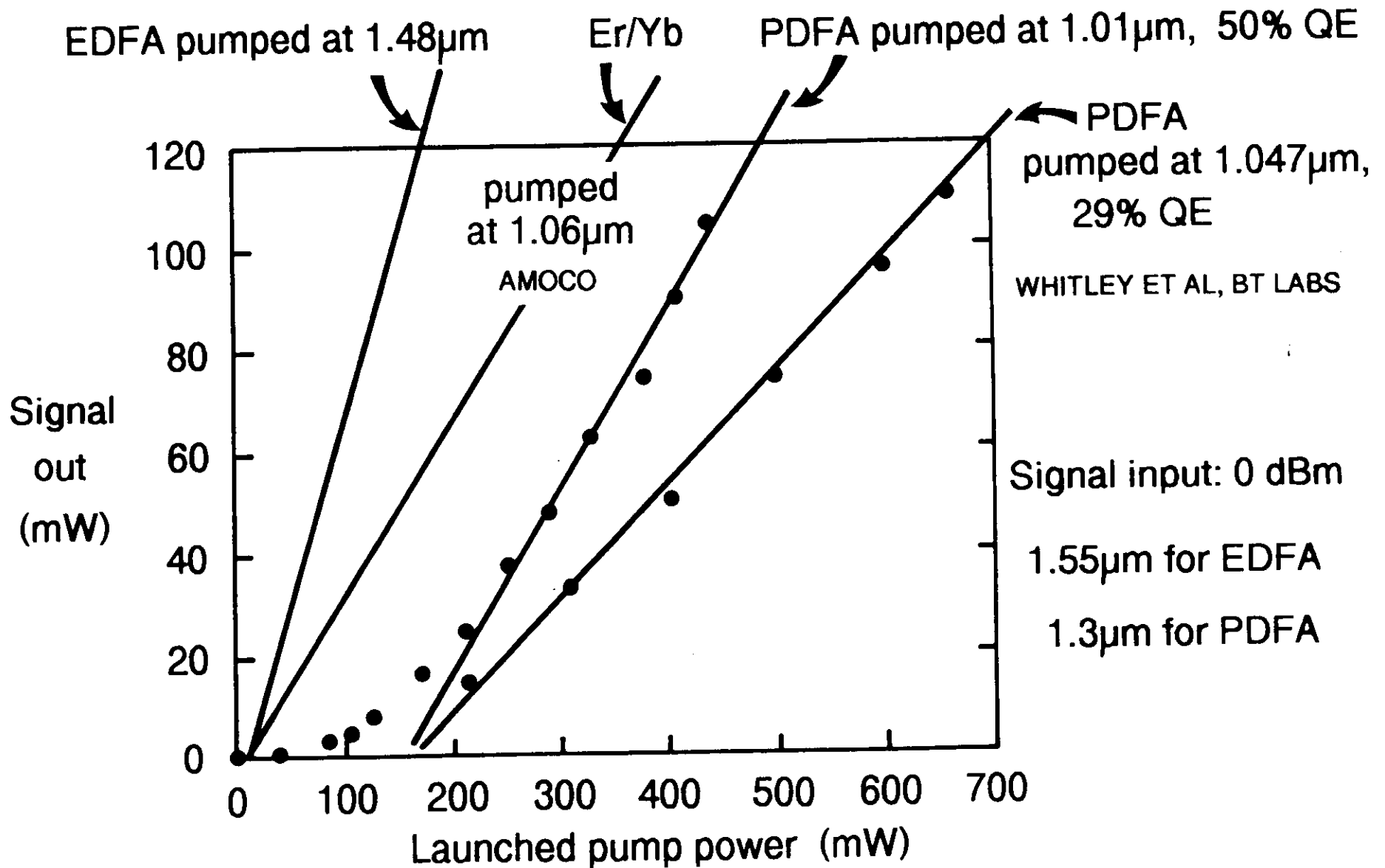


Non-radiative decay ➡ Quantum efficiency $\sim 4\%$

Pr^{3+} DOPED ZBLAN SMALL SIGNAL GAIN

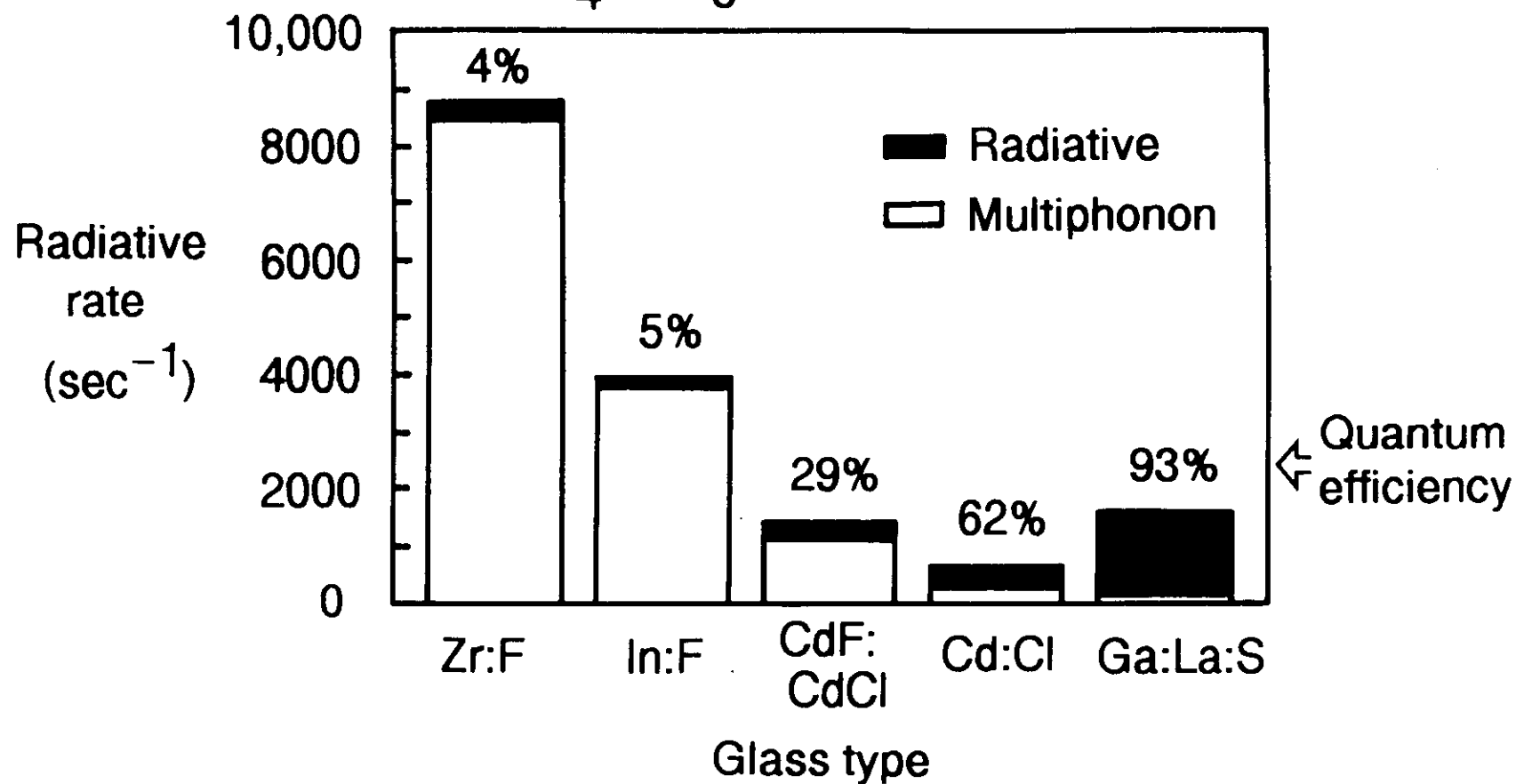


POWER AMPLIFICATION

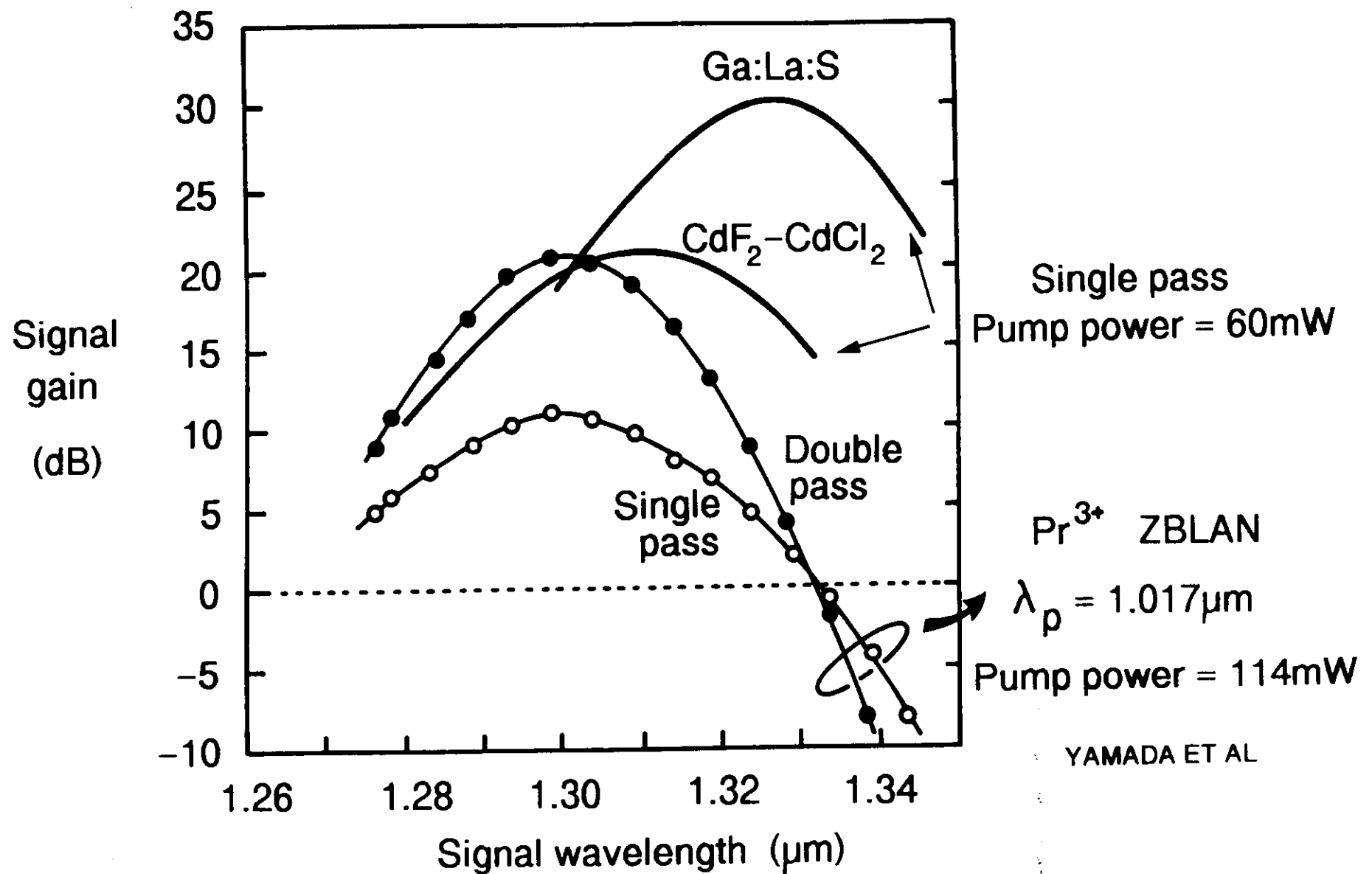


CALCULATED RADIATIVE & NON-RADIATIVE PROPERTIES

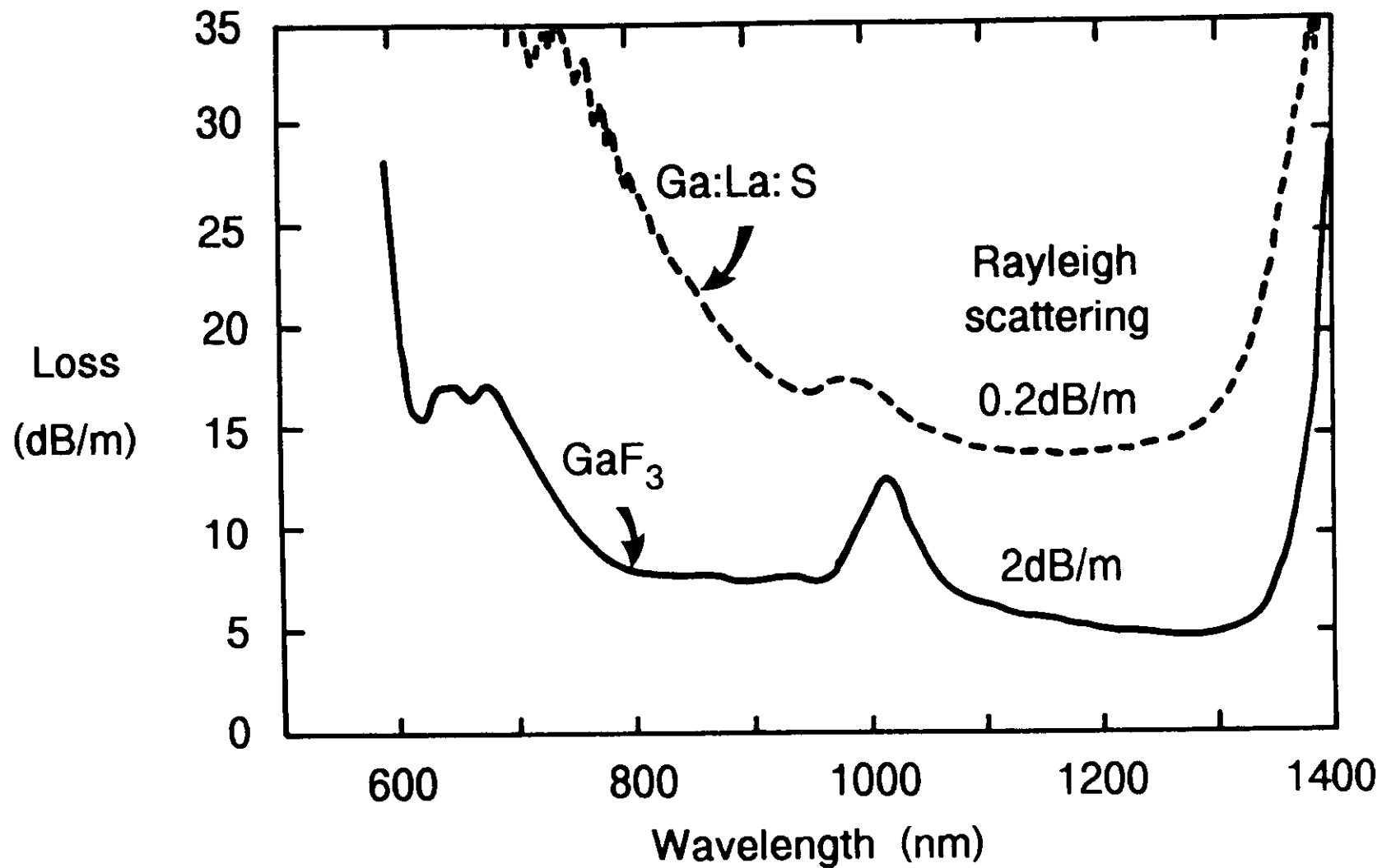
$^1G_4 - ^3H_5$ Transition in Pr^{3+}



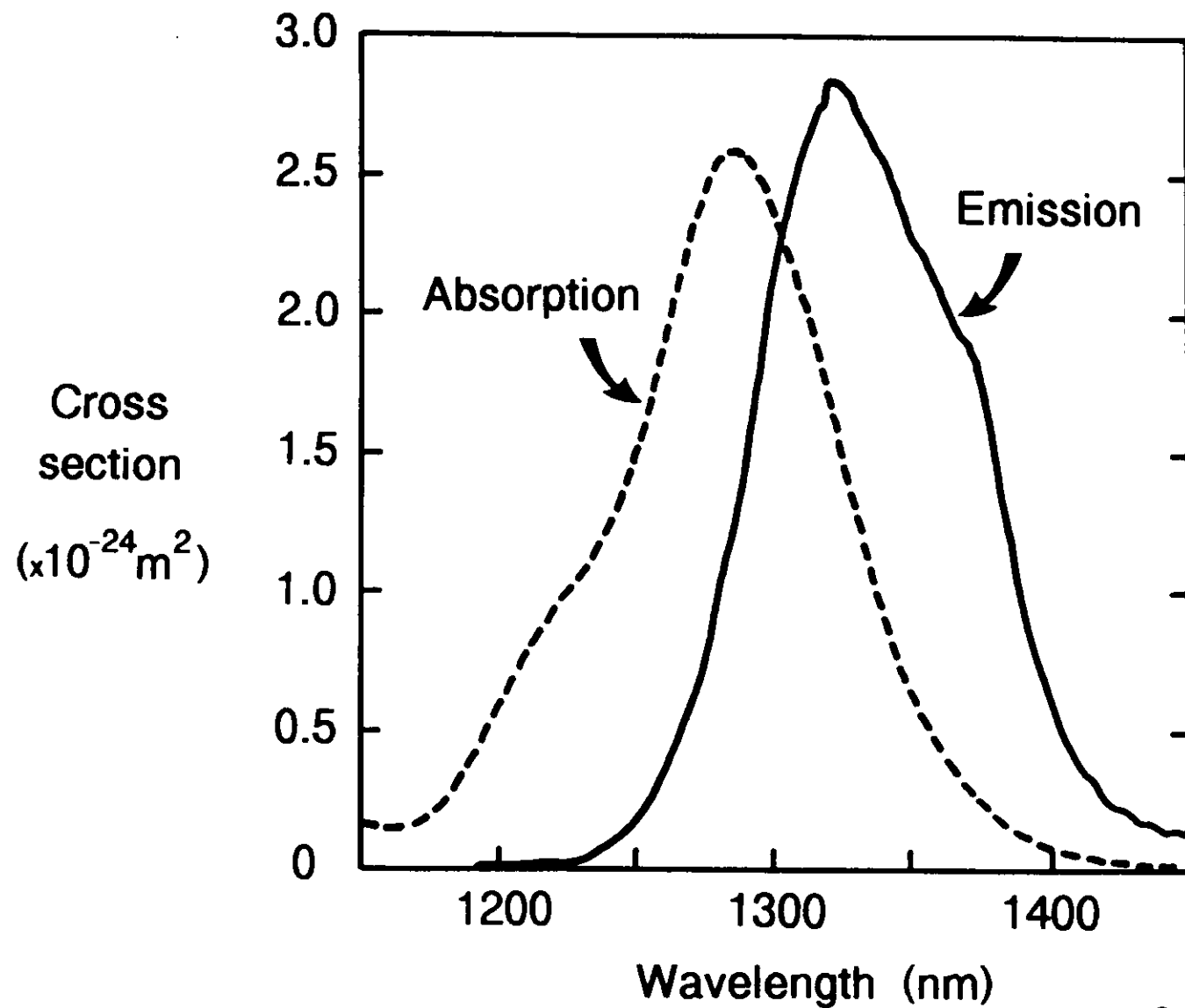
SMALL SIGNAL GAIN SPECTRA OF Pr^{3+} DOPED AMPLIFIER



LOSS SPECTRA OF EARLY SOUTHAMPTON FIBRES



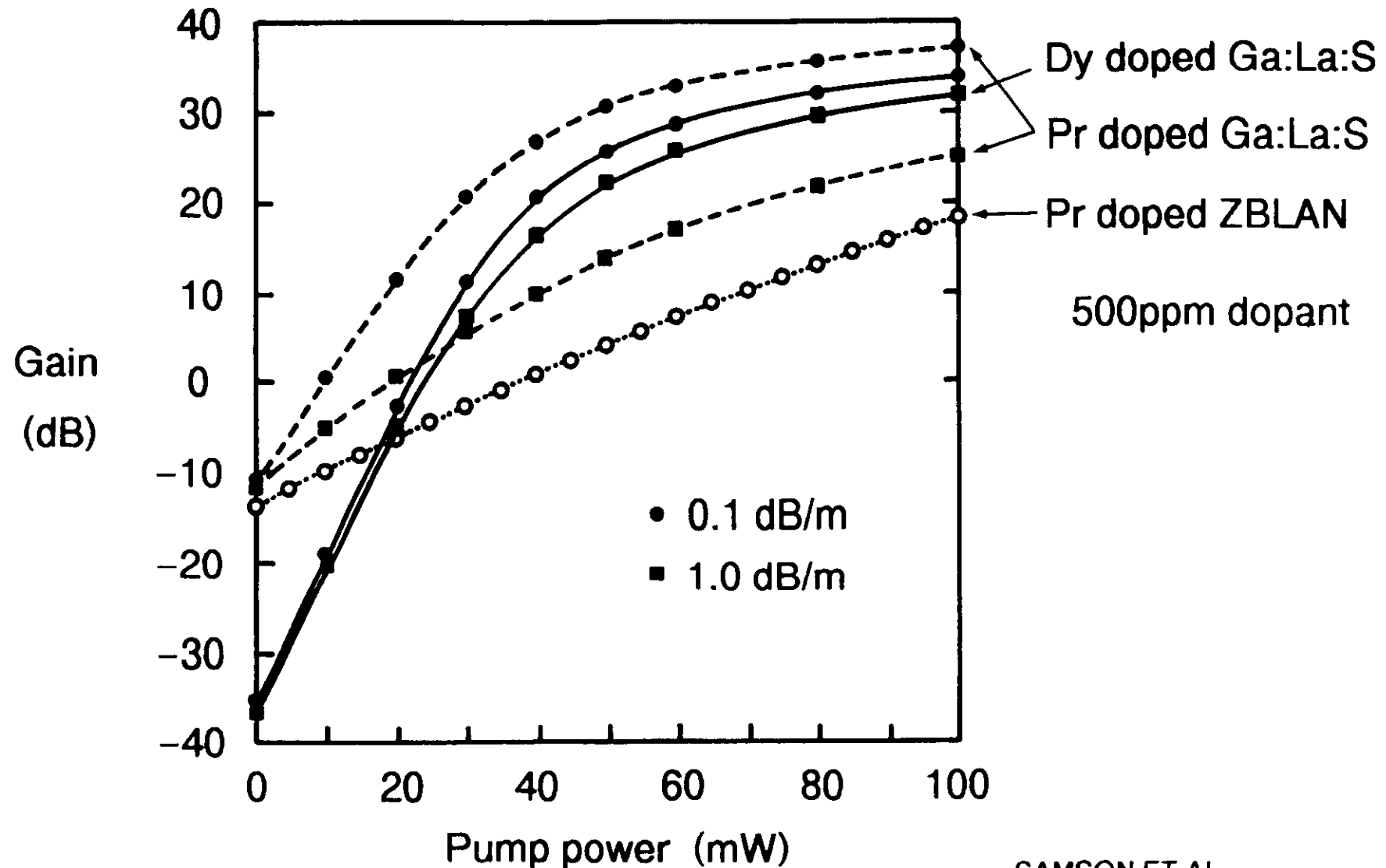
Dy³⁺ EMISSION AND ABSORPTION CROSS-SECTIONS



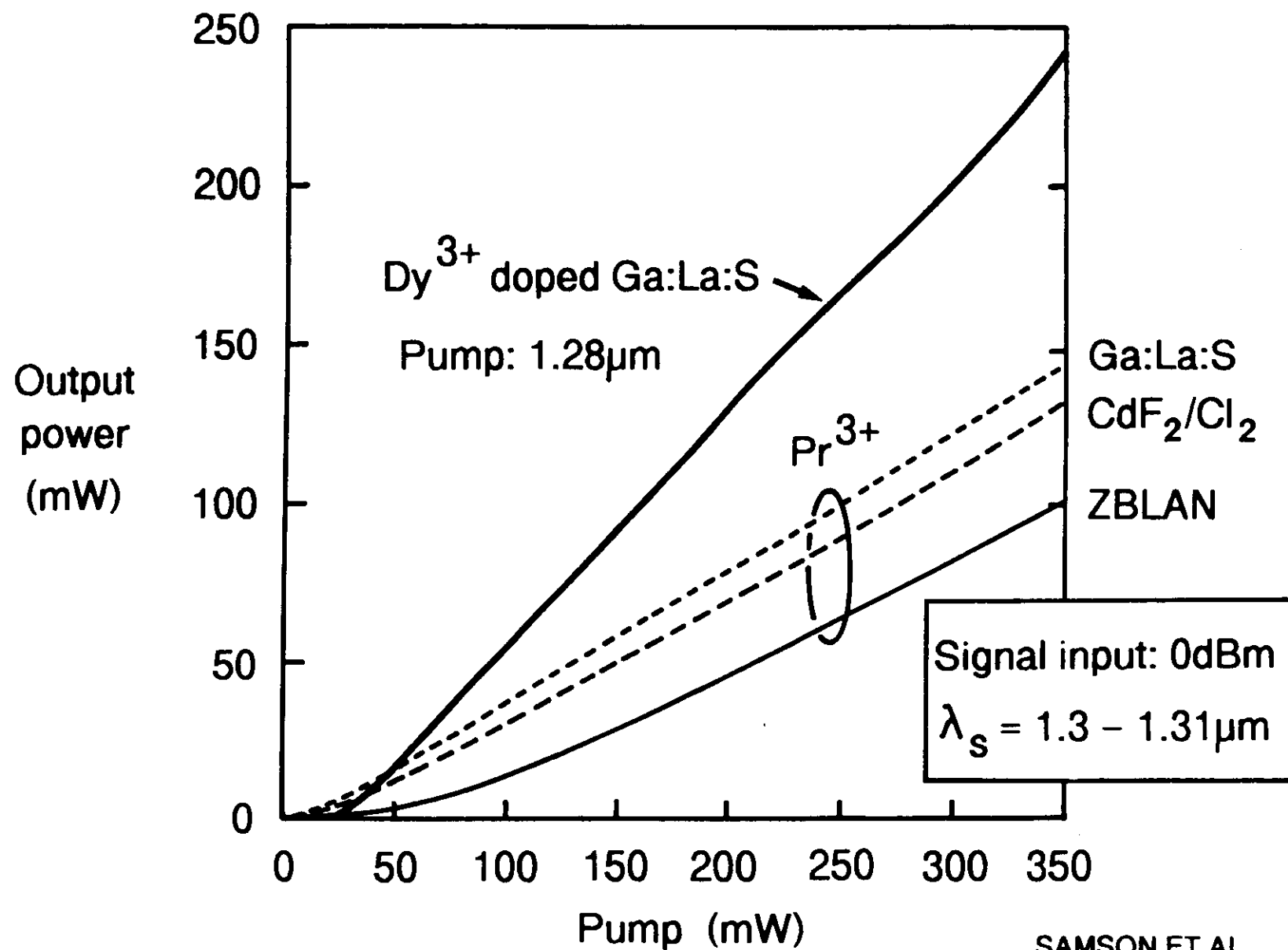
SAMSON ET AL

COMPARISON OF Dy^{3+} and Pr^{3+} DOPED GaLaS and Pr^{3+} DOPED ZBLAN

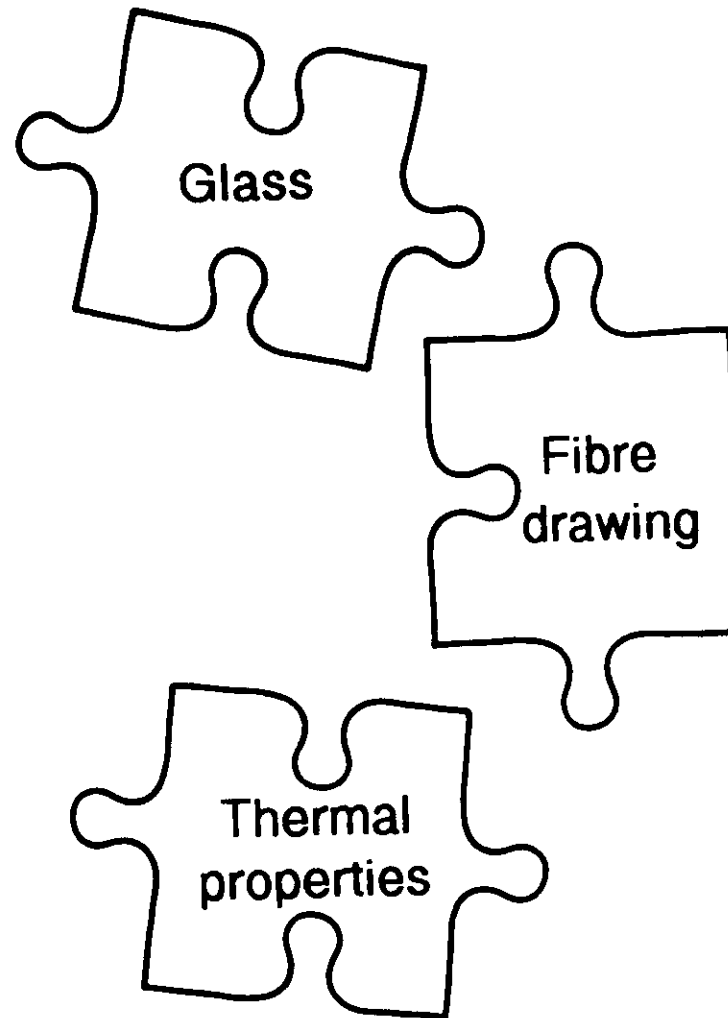
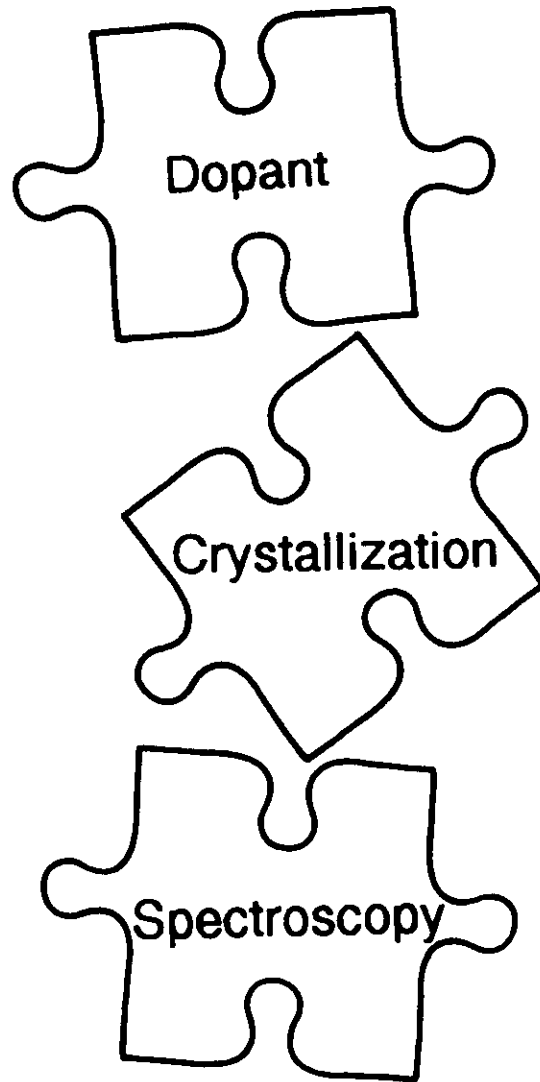
Showing the effect of background loss



Dy³⁺ DOPED Ga:La:S POWER AMPLIFIER



SAMSON ET AL



PROSPECTS FOR 1.3 μ m AMPLIFIERS

- Pr³⁺ ZBLAN - 1.02/1.047 μ m pump
Reasonable power amplifier, Poor small signal gain

Commercially available !!

- Pr³⁺ doped low phonon energy glasses;
Chalcogenides & Mixed halides - 1.02 μ m pump
Significantly improved small signal gain, Improved power amplifier

Chalcogenide preferred !!

- Dy³⁺ doped Chalcogenide - 1.28 μ m pump
Excellent power amplifier, Good small signal gain

Potential problems with bottlenecking to be overcome !!

SOLITON TRANSMISSION OVER STEP INDEX FIBRE

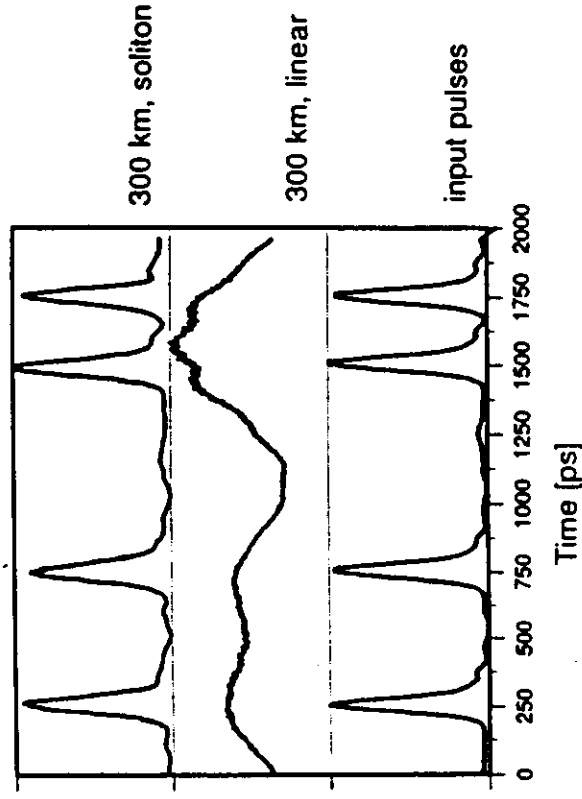
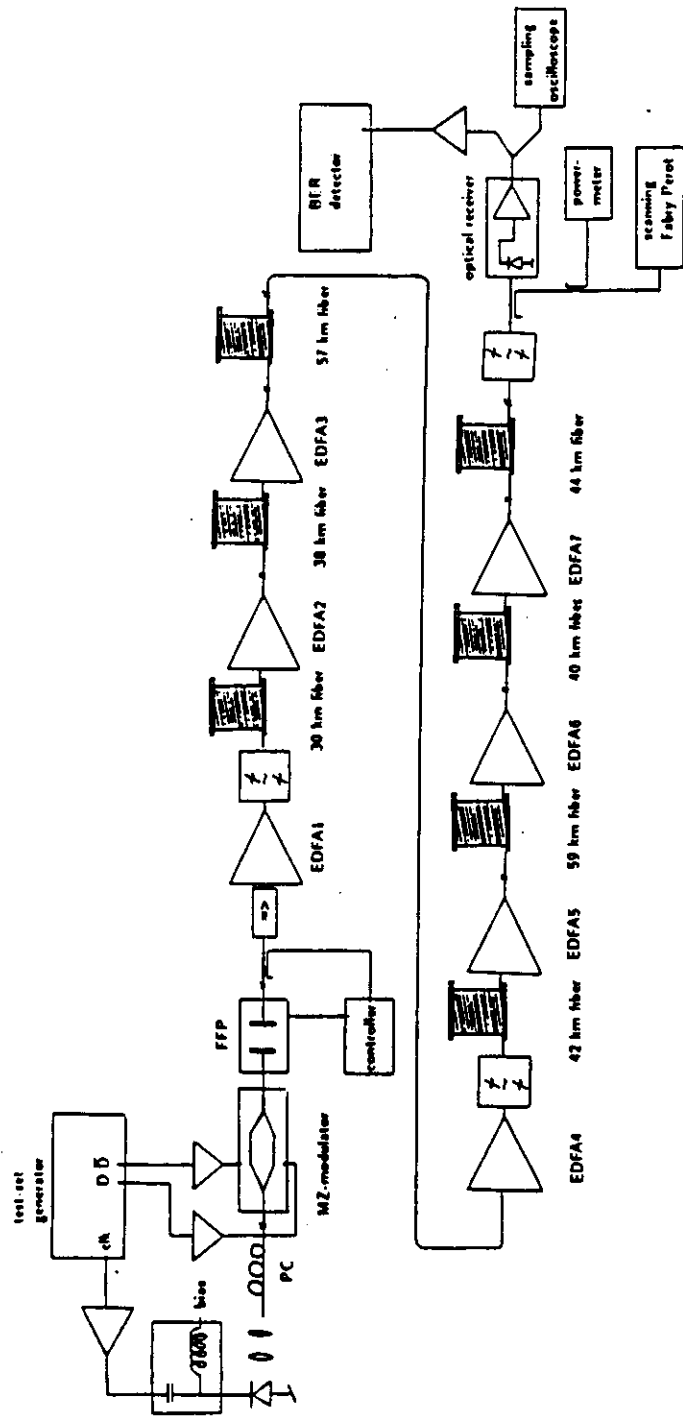
Problems

- Soliton period
- Amplifier spacing
- Gordon Haus jitter

Solutions

- Partial dispersion compensation
- Soliton filtering
- Temporal control

SOLITON TRANSMISSION AT $1.55\mu\text{m}$ IN STANDARD FIBRE



4 Gbit/s
55 ps pulses
8.6 mW peak power
70 km soliton period
40 km amp spacing

B. Christensen et al, ECOC '93

SOLITON TRANSMISSION ON STANDARD FIBRE USING SATURABLE ABSORPTION AND SPECTRAL FILTERING

Dispersion 17 ps/nm.km

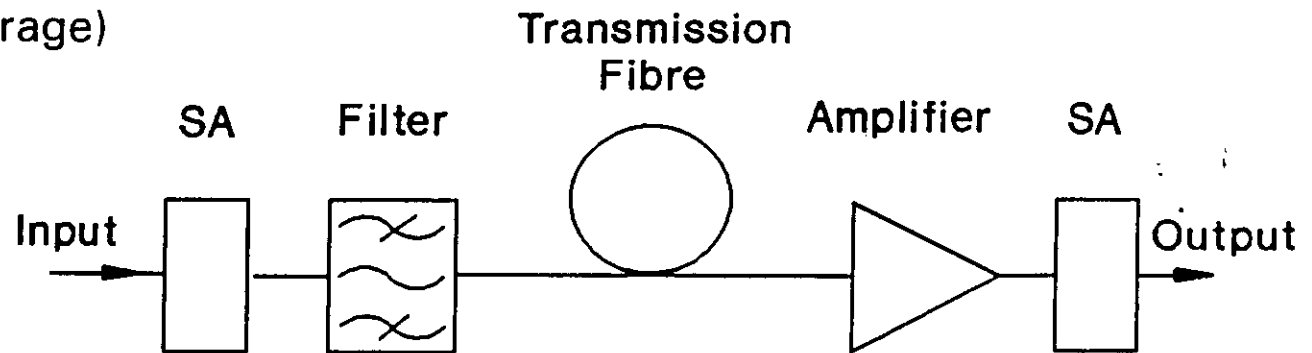
8 Gbit/s, MSR = 4.5:1
soliton pulsewidth 28ps
(40 mW peak, 10 mW average)

Amp spacing 40 km
(3 soliton periods)

Filter bandwidth 30 GHz

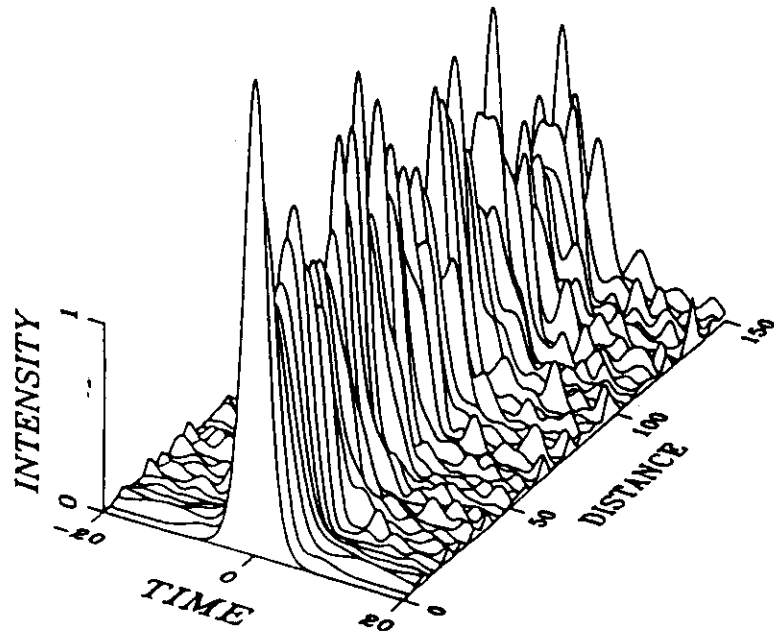
Amp gain 12 dB

For 10Gbit/s amp
spacing ~ 25 km

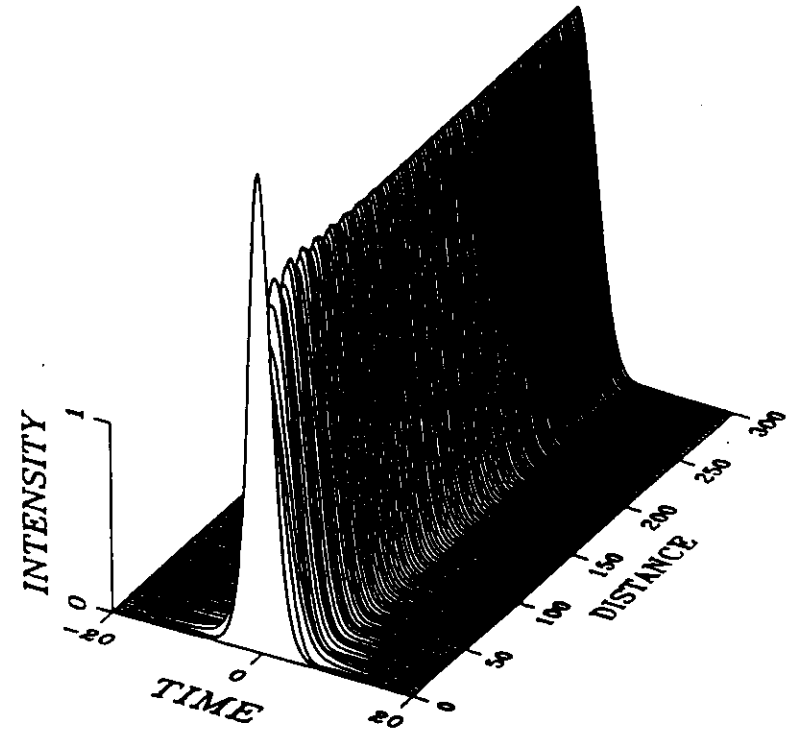


A. Grudinin et al, Southampton University

Pulse dynamics along transmission system



No control



Filter + Saturable absorber

DISPERSION EQUALISING FIBRE

Advantages

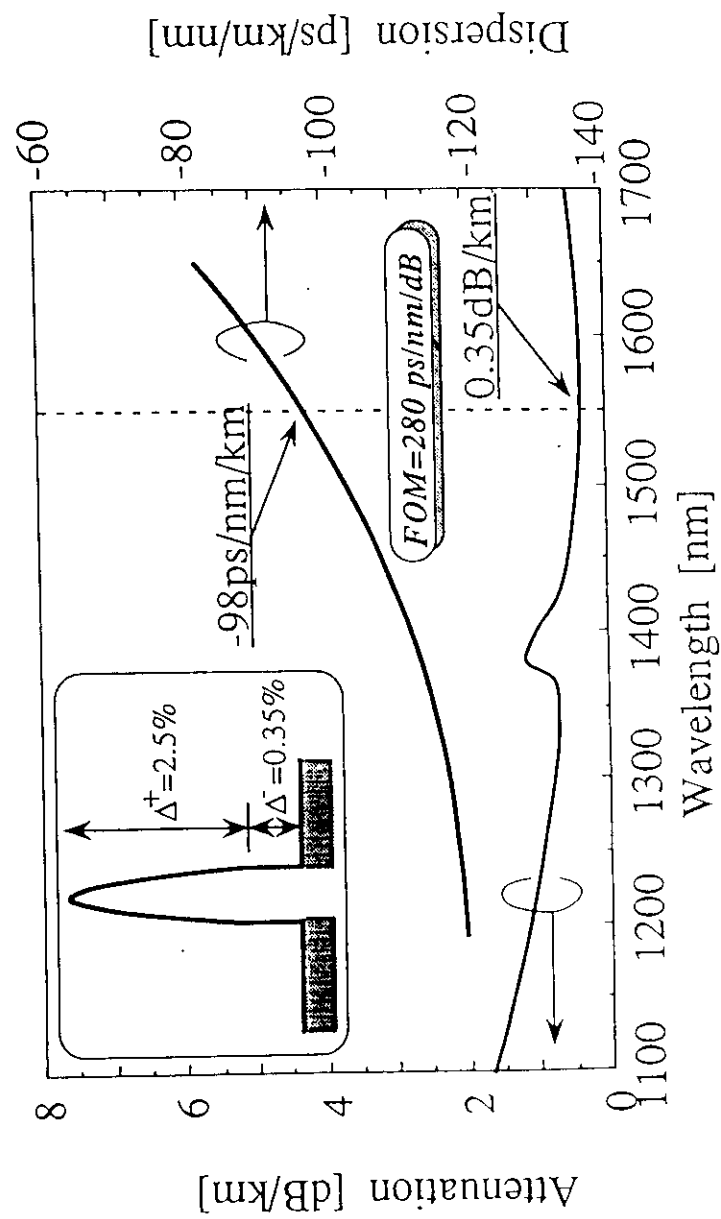
- Broad bandwidth
- Simple
- Fibre compatability

Disadvantages

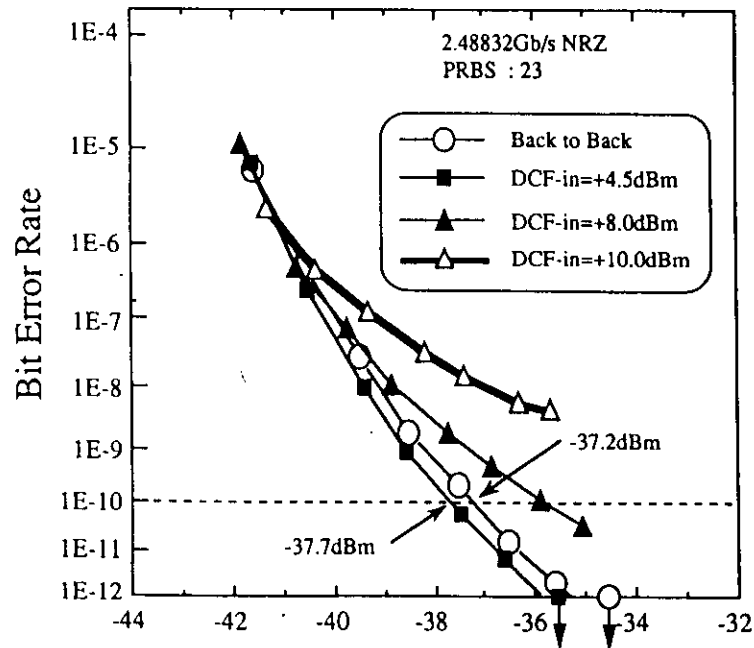
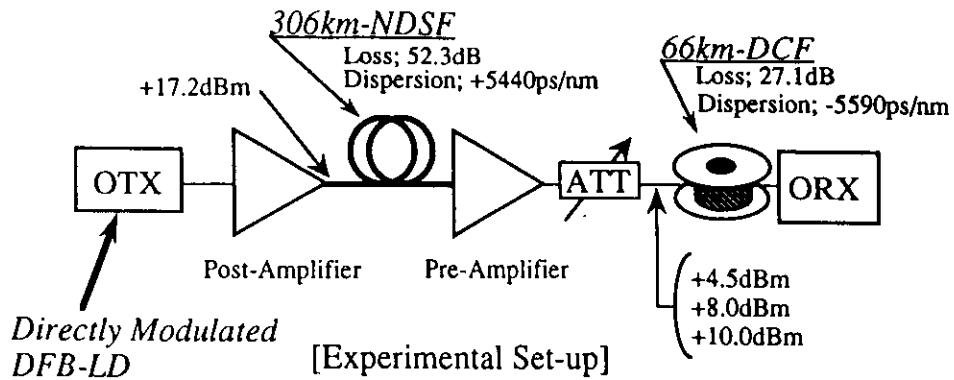
- Loss
- Size

Dispersion Compensating Fiber with Optimized Profile

SEI



2.4Gb/s, 306km Repeaterless Transmission Experiment using DCF



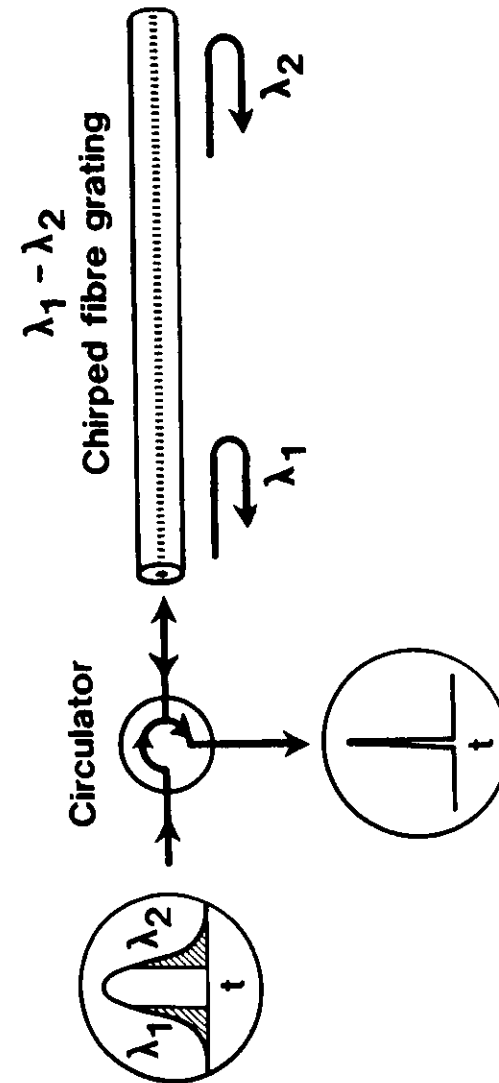
Optical Received Power [dBm]

[Bit Error Rate Characteristics]



FUTURE FIBRE DEVICES

Dispersion compensators?



- To compensate 100km of 17ps/nm.km fibre:
Need 17cm/nm
- Difficult to write gratings > 1cm at present

CHIRPED FIBRE GRATINGS

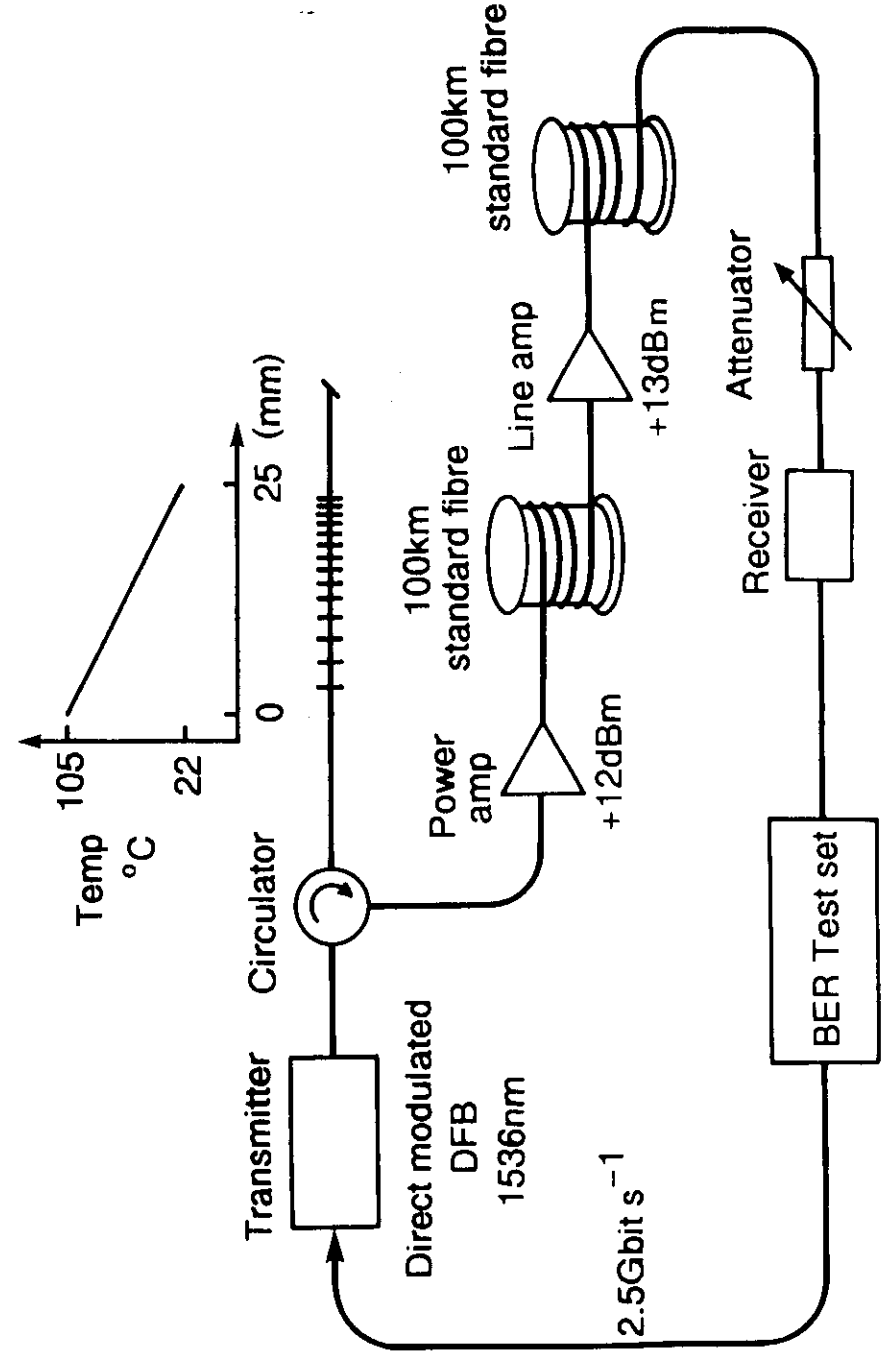
Advantages

- Low loss
- Compact
- Higher order dispersion
- Narrowband

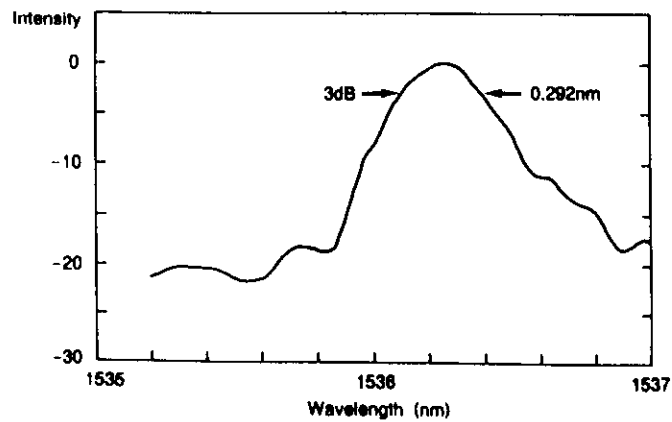
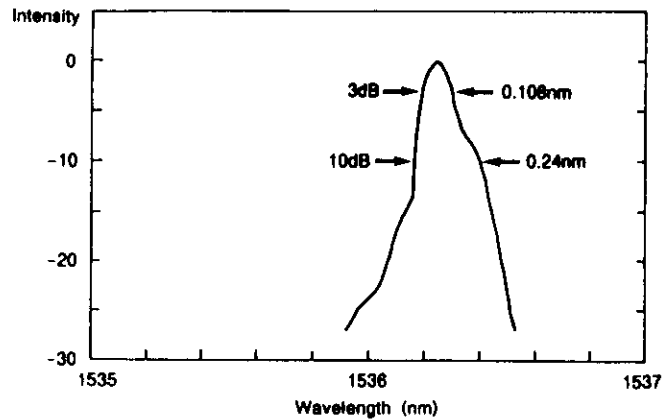
Disadvantages

- Dispersion profile
- Narrowband

Experimental setup

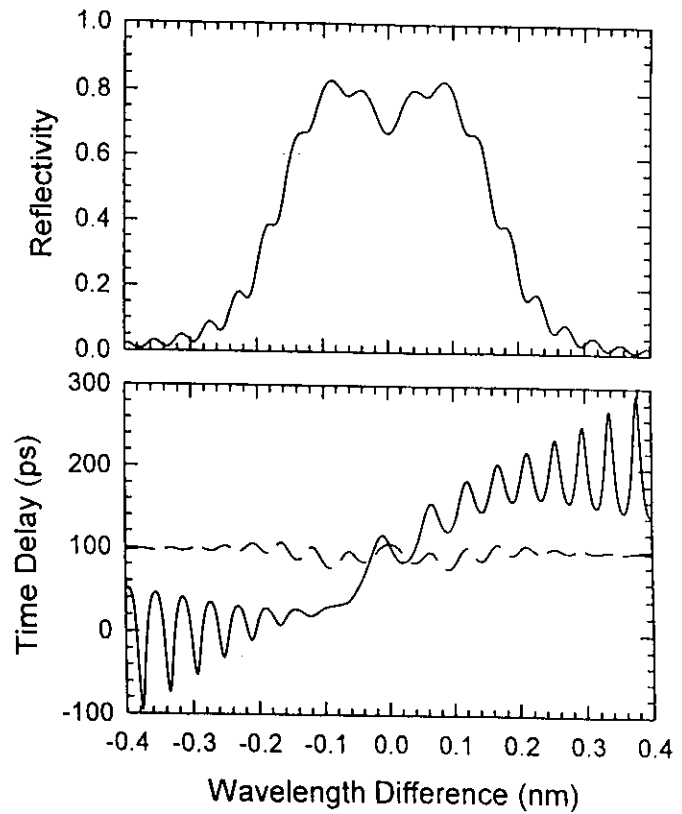


TRANSMITTER AND GRATING REFLECTION SPECTRA

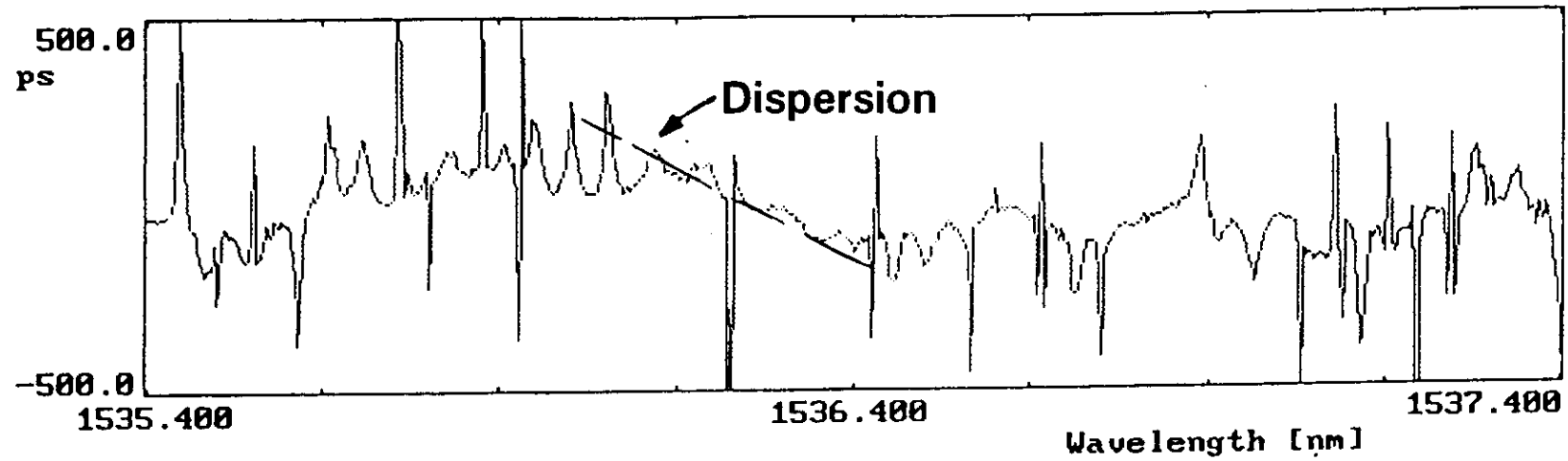


Reflection and Transmission of a Chirped Grating

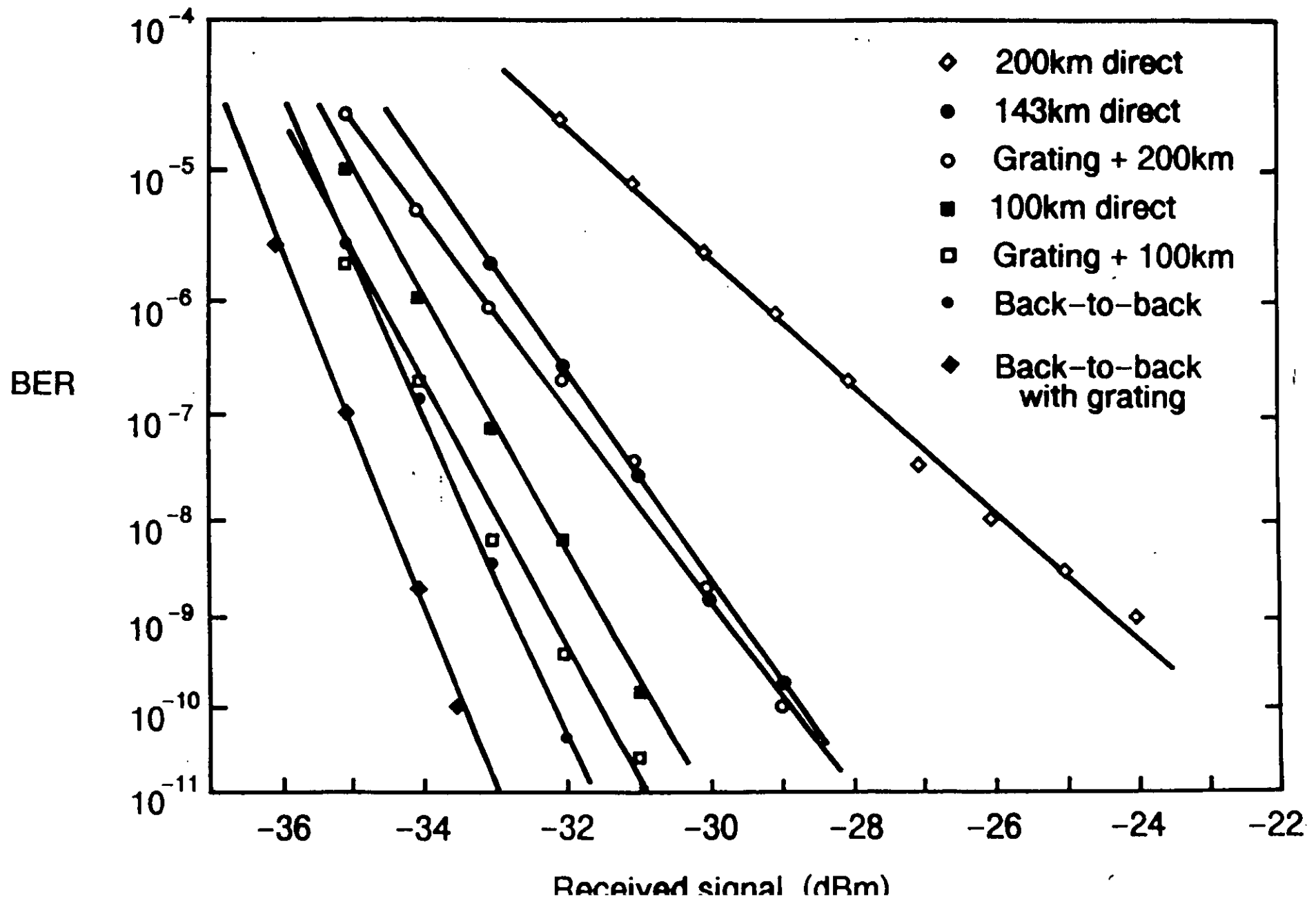
$\lambda = 1552\text{nm}$ Grating Length= 20mm $h_0 = 3 \times 10^{-5}$ $C = 0.006$



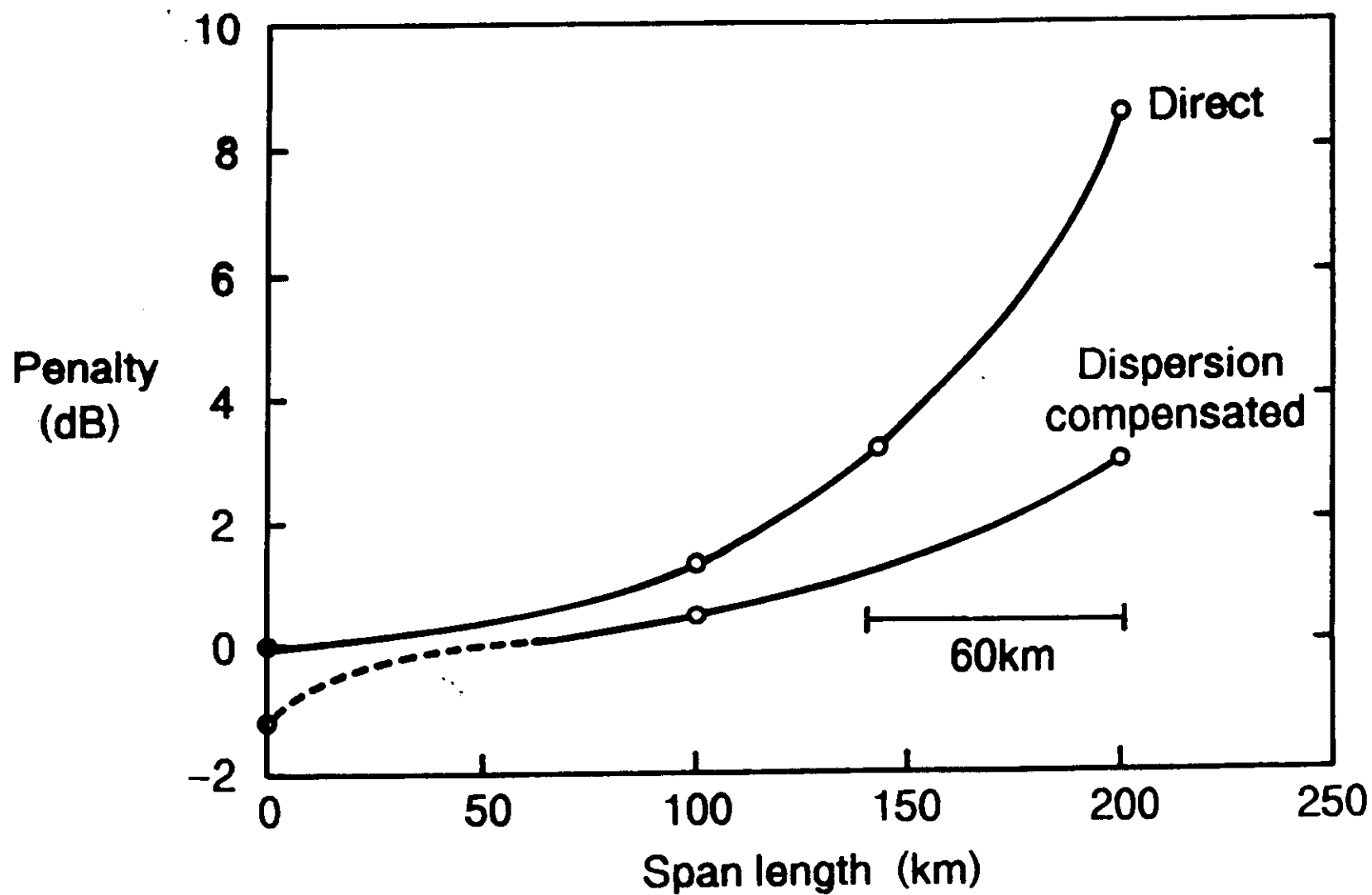
MEASURED DELAY CHARACTERISTICS OF THE GRATING



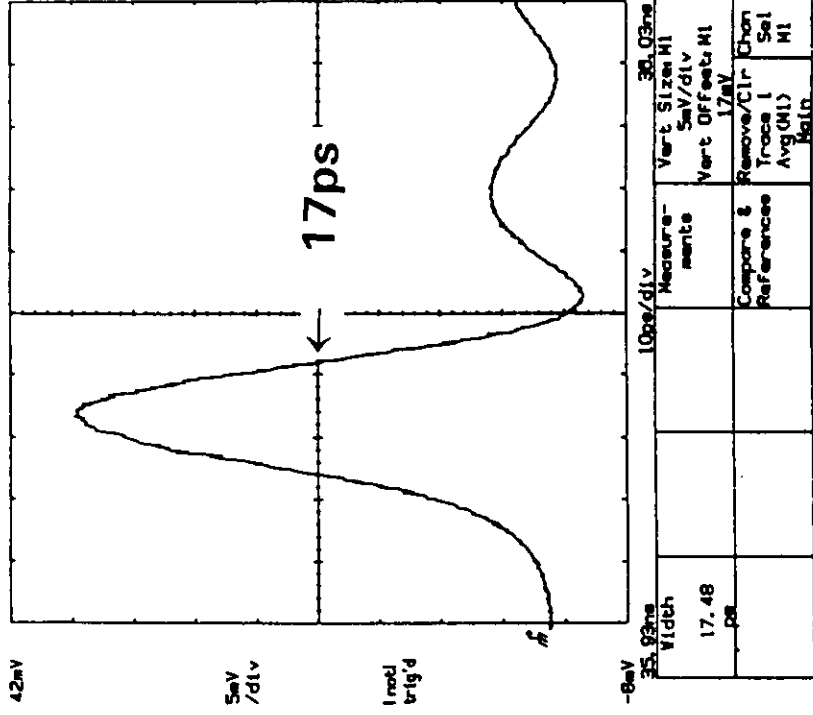
BER for 2.5Gbit/s system



Receiver penalty due to dispersion

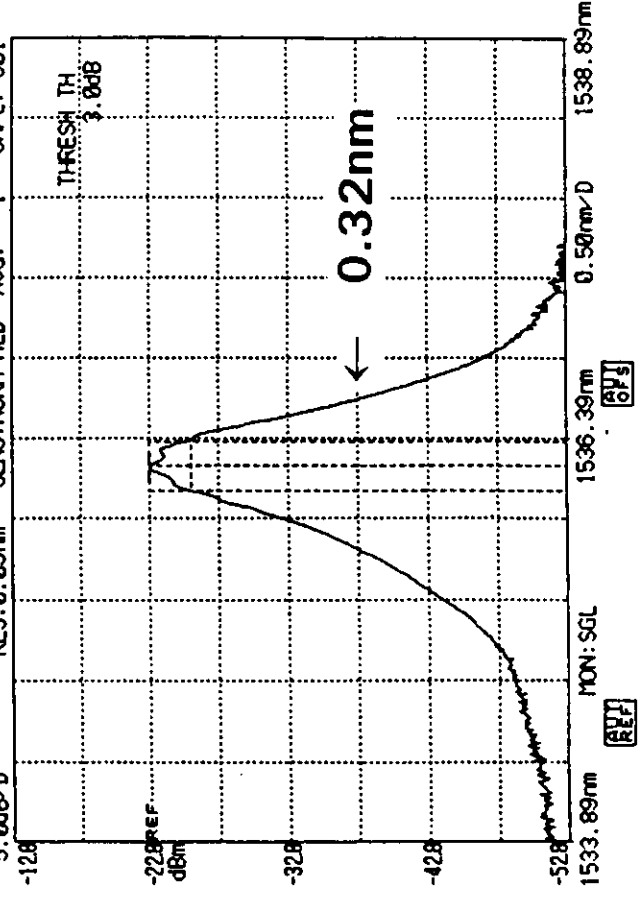


SOURCE DATA



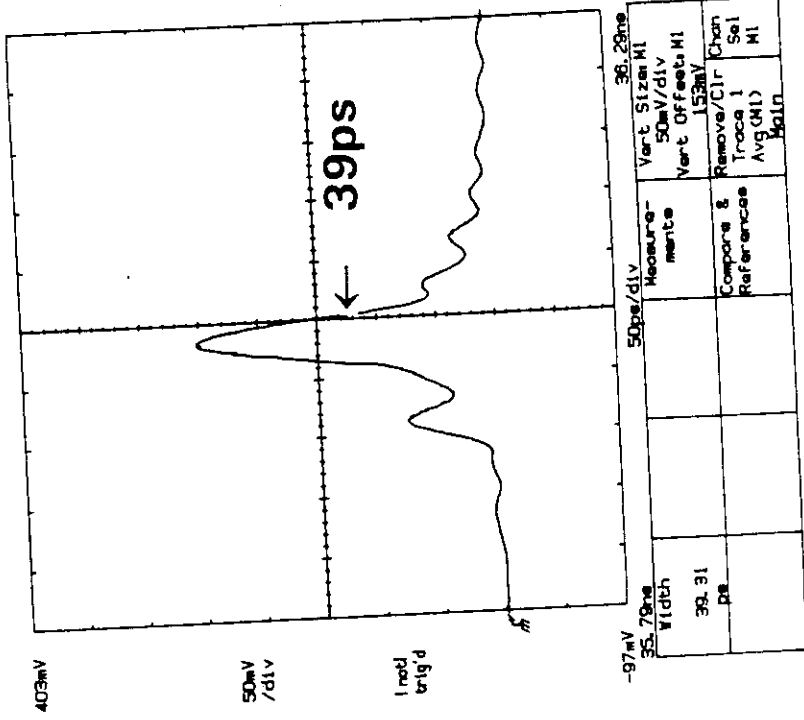
30 Sep 1994 14:40

SPECTRAL WIDTH : <THRESHOLD> MODE FIT : OFF
THRESH LVL : 3.0dB Δλ : 0.318nm
K : 1.00 MODE : 1 ΔC : 1536.220nm
5.0dB/D RES:0.05nm SENS:NORM HLD AUG: 1 SNPL: 501

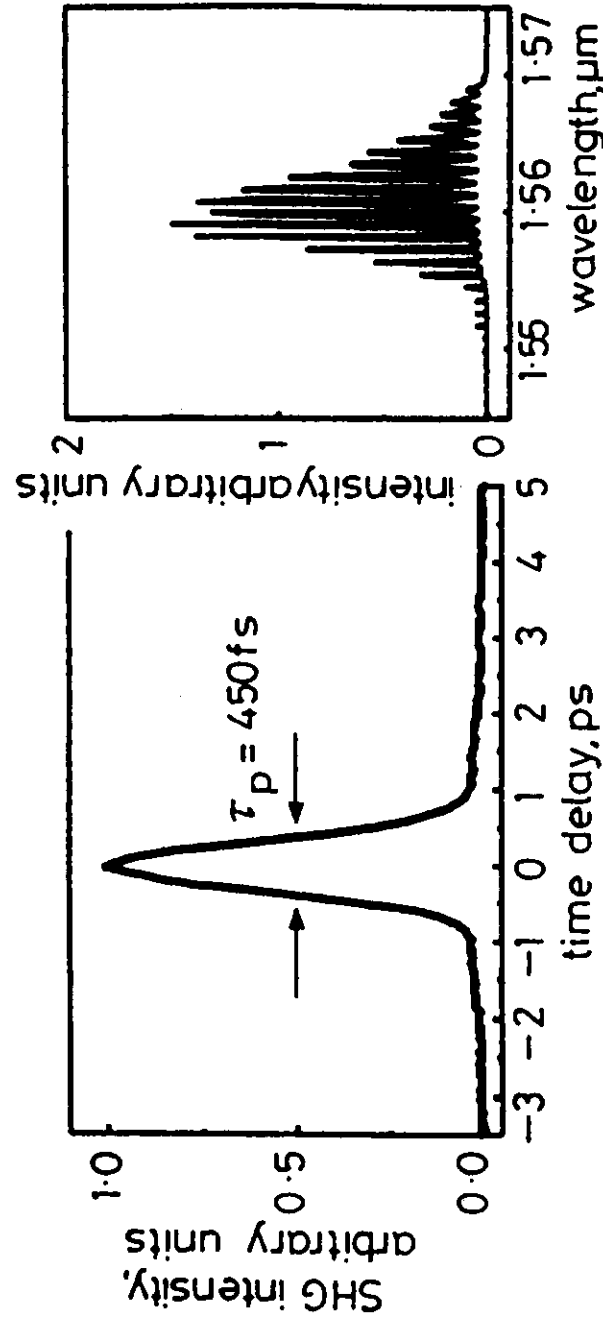
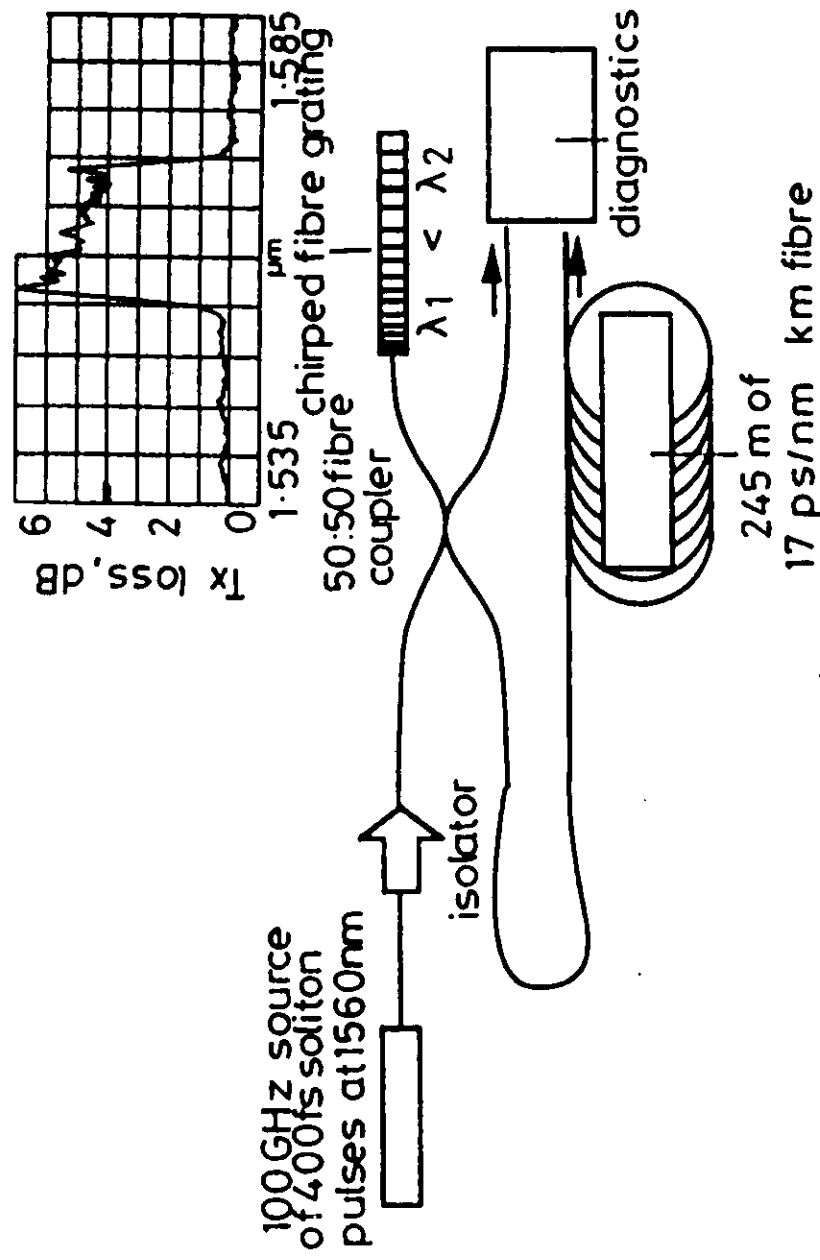


28

Compensated

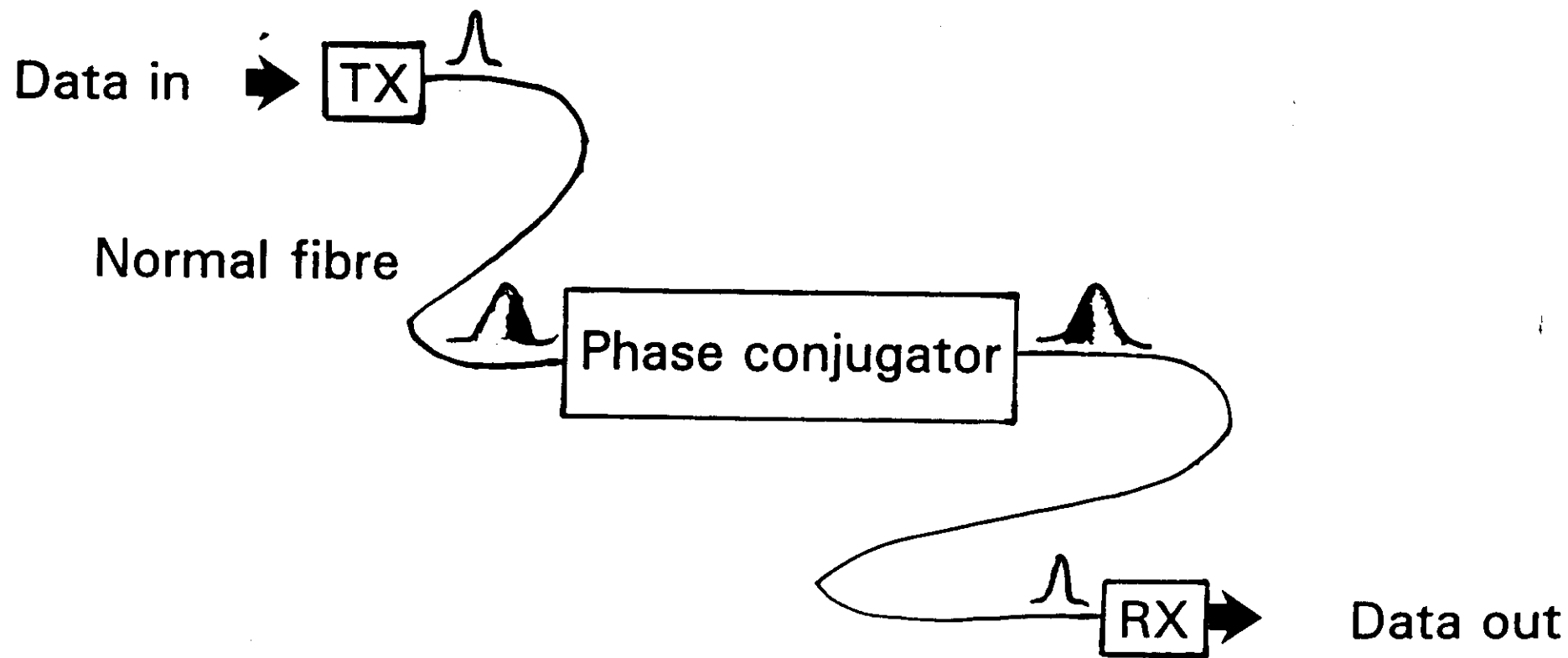


CHROMATIC DISPERSION COMPENSATION OF 400 fs PULSES



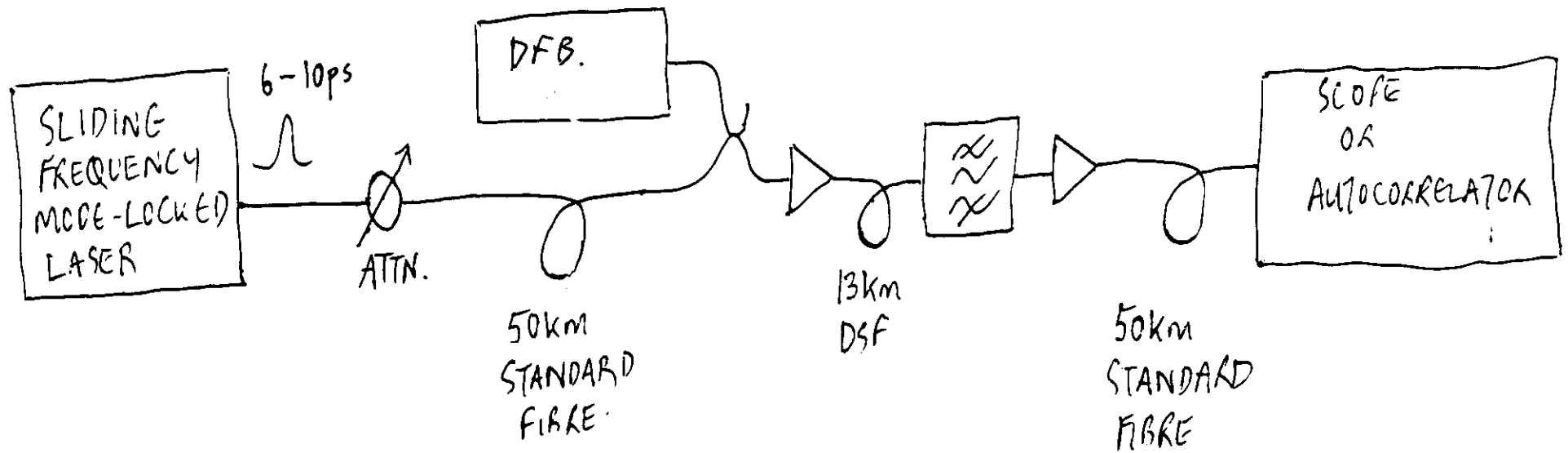
R. Kashyap et al, Electr. Lett., '94

DISPERSION COMPENSATION BY PHASE CONJUGATION



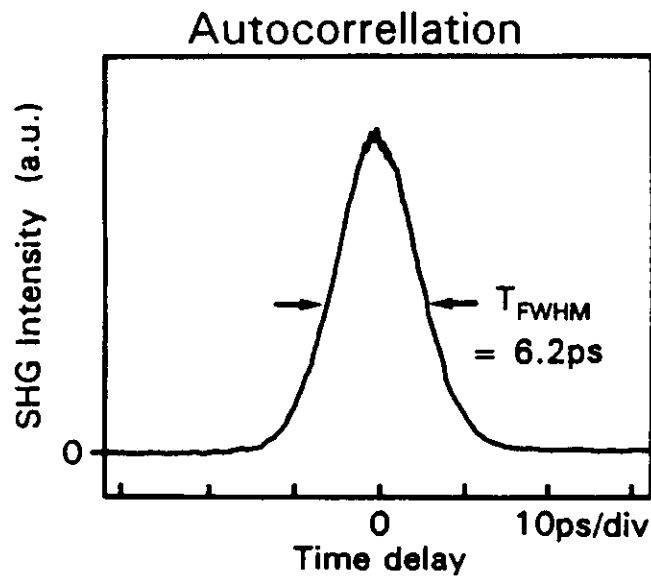
A. Yariv, Opt. Lett., 1979

MPSI EXPERIMENTS

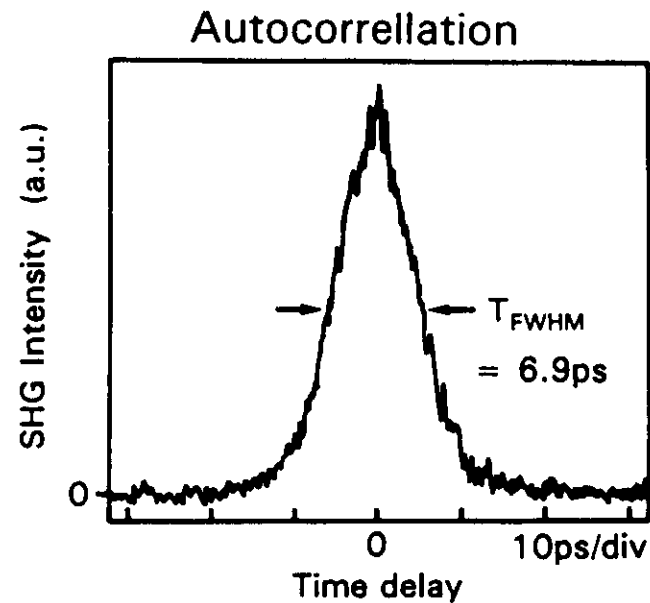


TRANSMISSION OVER 50km OF STANDARD FIBRE

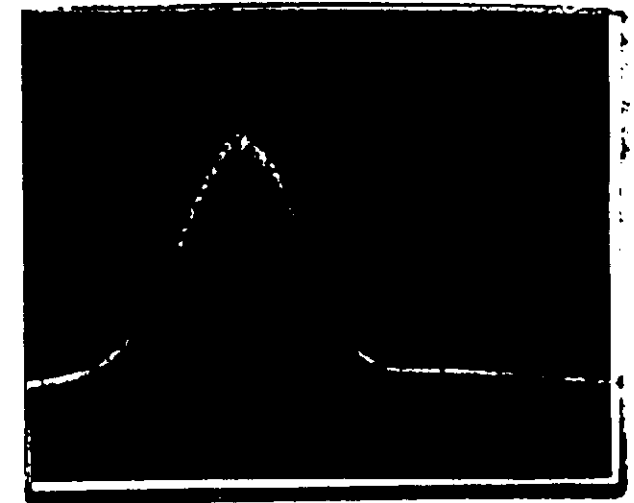
Input pulse



Output pulse
with compensation



Output pulse
without compensation

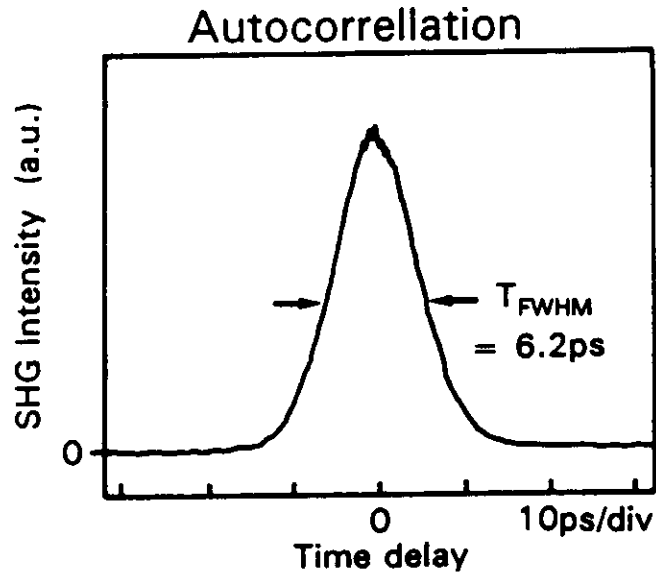


200ps/div

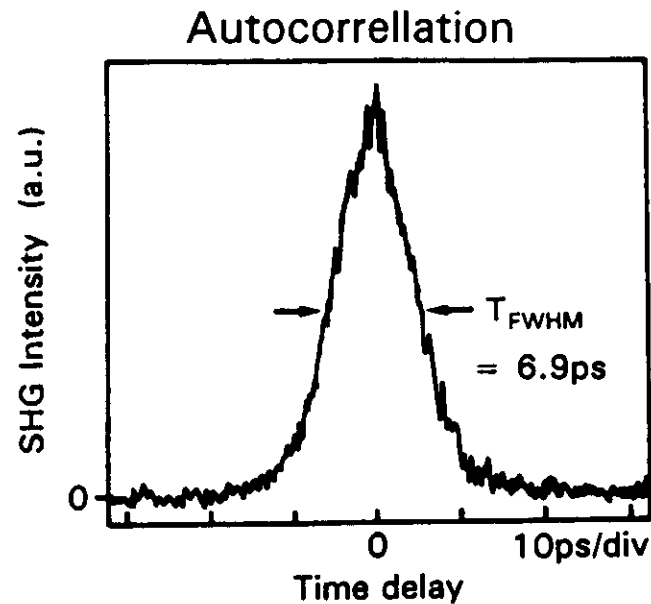
detector/scope response 320ps

TRANSMISSION OVER 50km OF STANDARD FIBRE

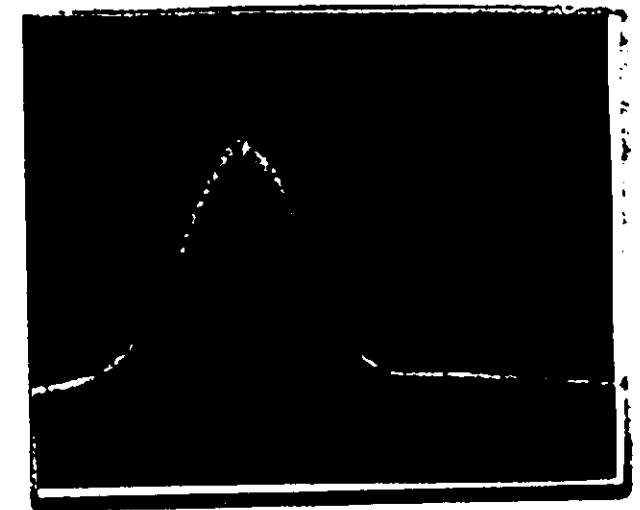
Input pulse



Output pulse
with compensation

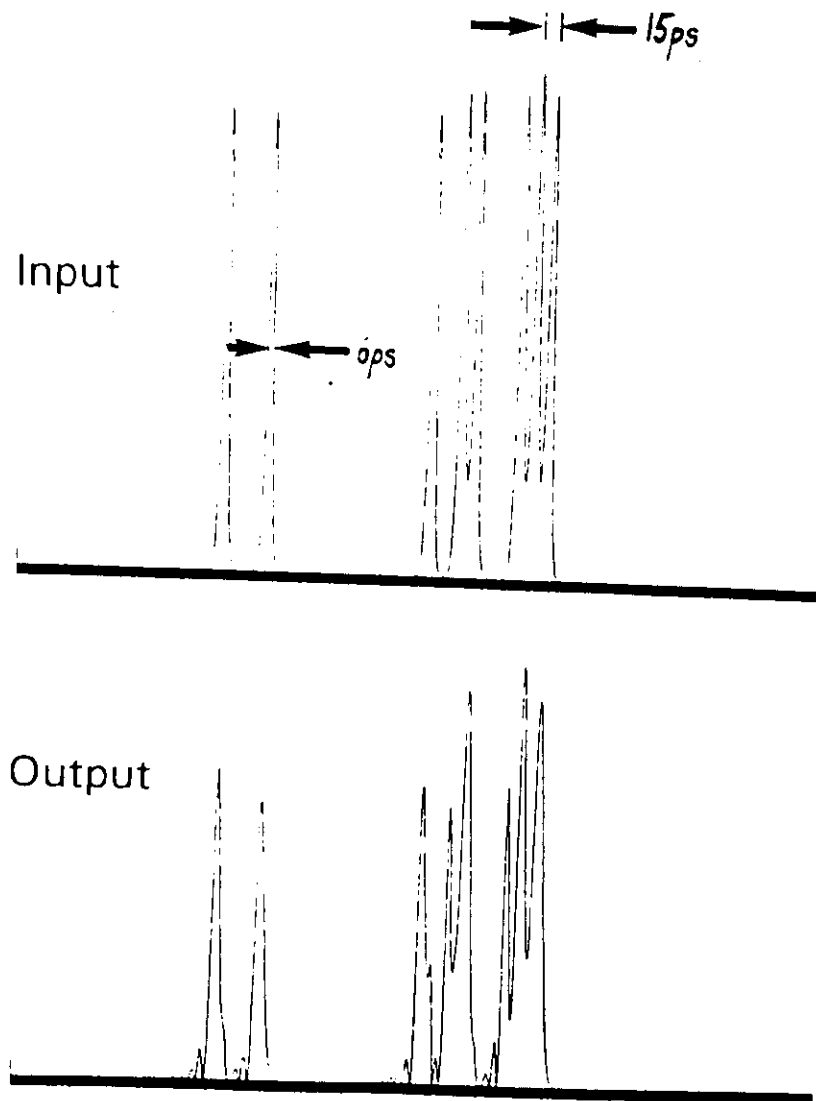


Output pulse
without compensation

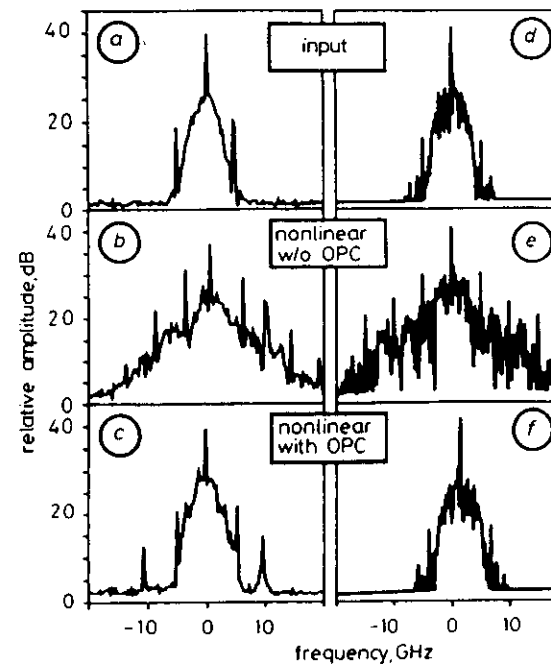
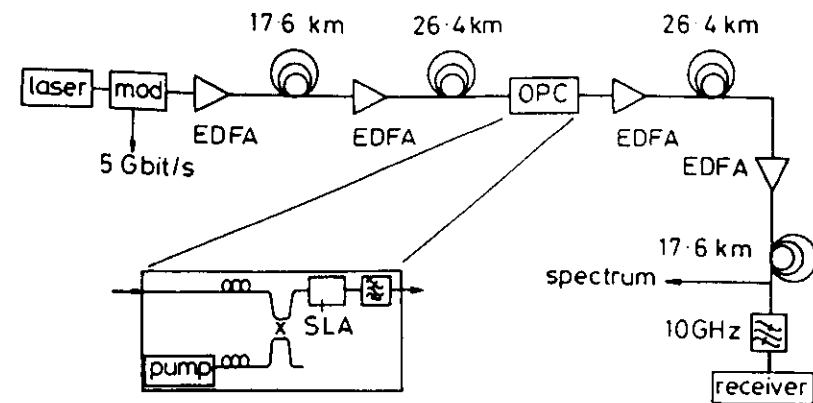


66 Gbit/s, 6ps PULSES OVER 2000km

Domino Taverner



COMPENSATION OF FIBRE NONLINEARITY WITH MPSI

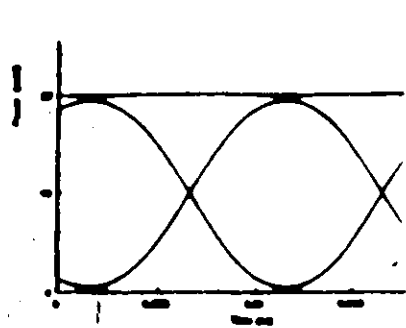


Kikuchi et al, OAA '93
Pieper et al, Electr. Lett., '94

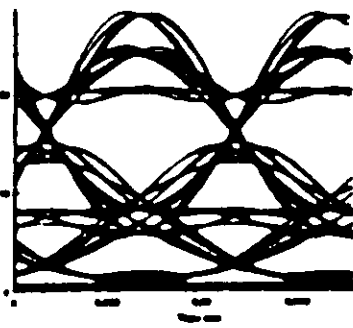
LONG-HAUL NRZ TRANSMISSION

C. Kurtzke and A. Gnauck

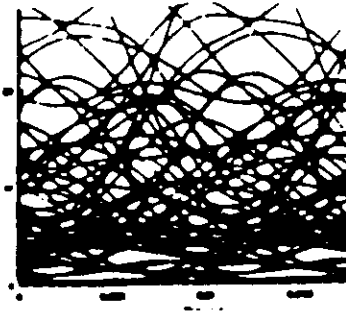
100Gbit/s



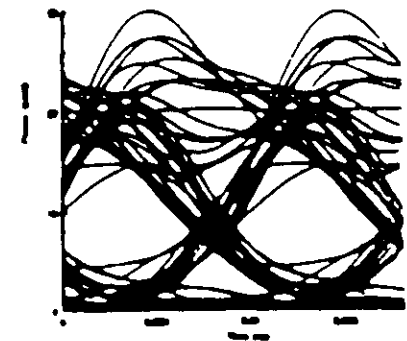
0km



10km

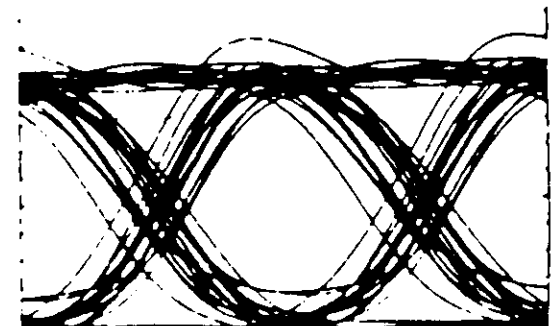


25km

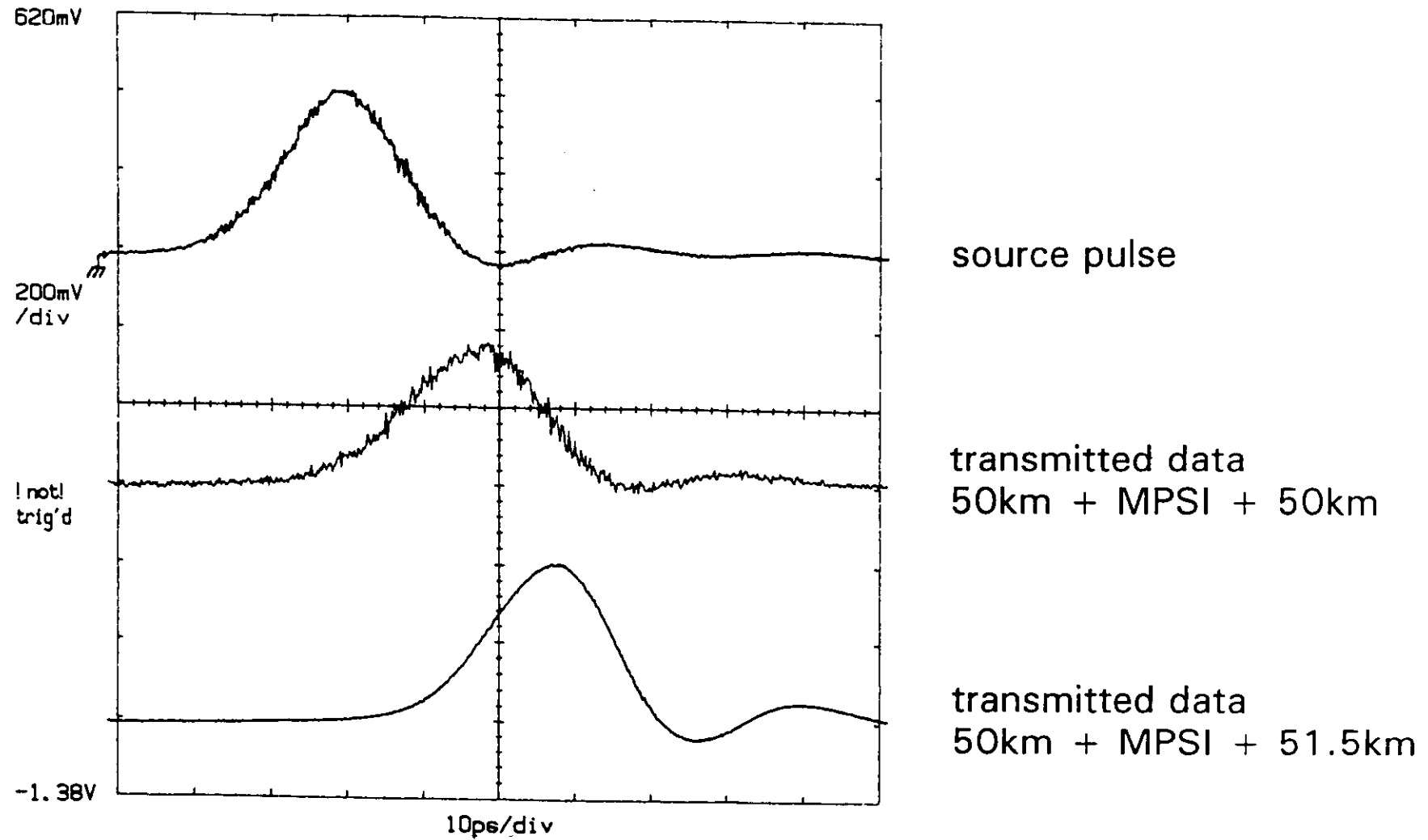


1000km
with compensation

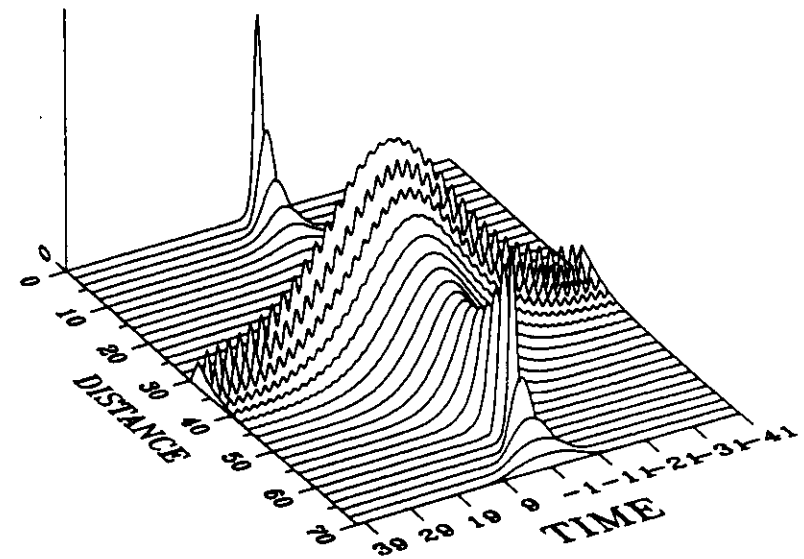
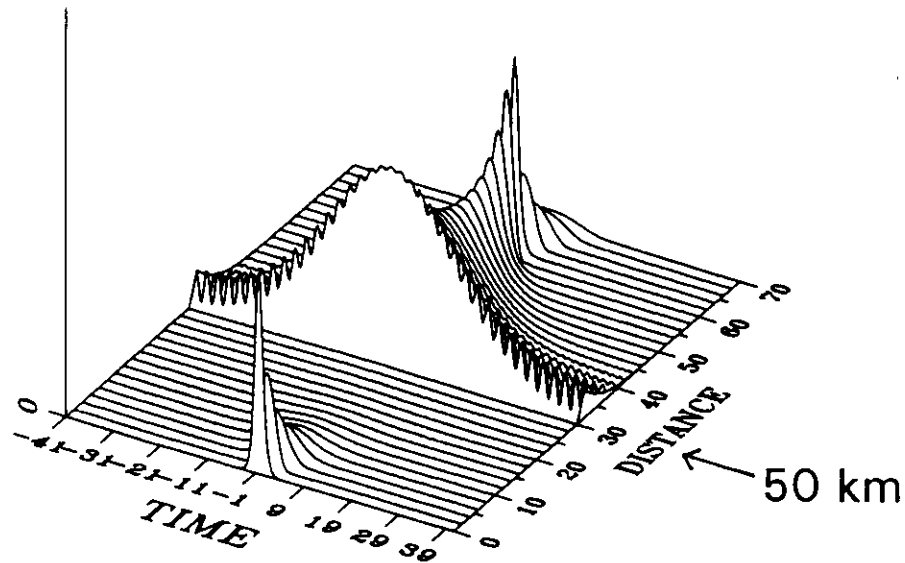
20Gbit/s over 10,000km with compensation
Including noise and dispersion fluctuations



MPSI PULSE TRANSMISSION EXPERIMENTS

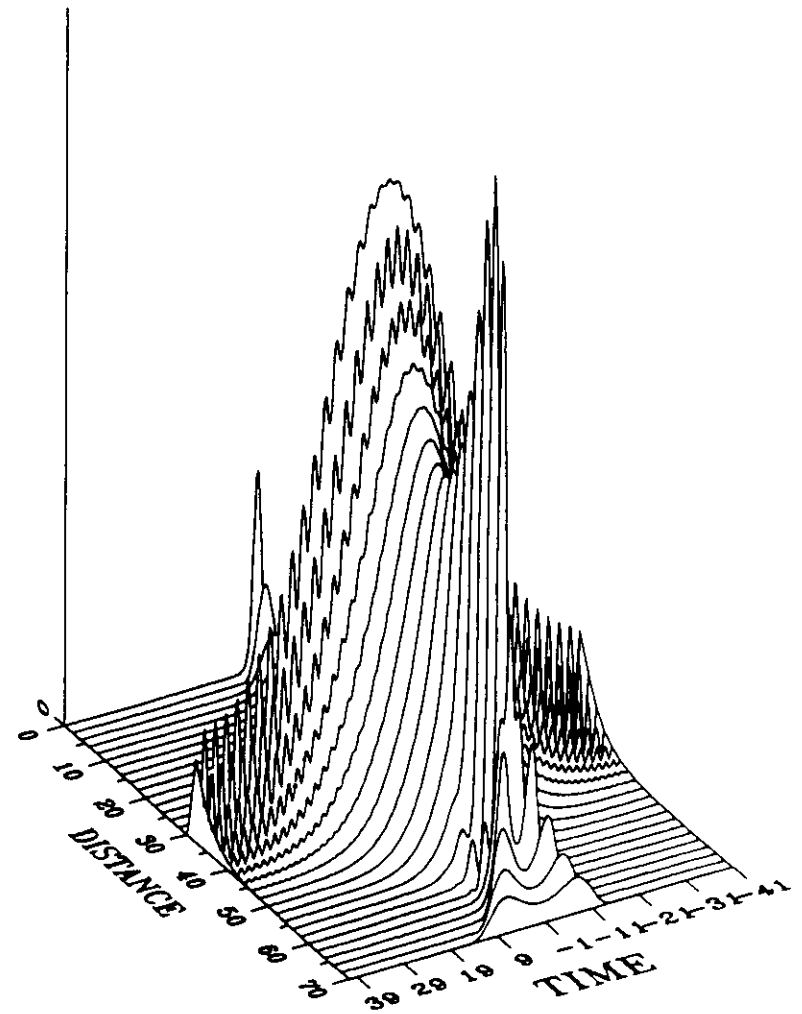
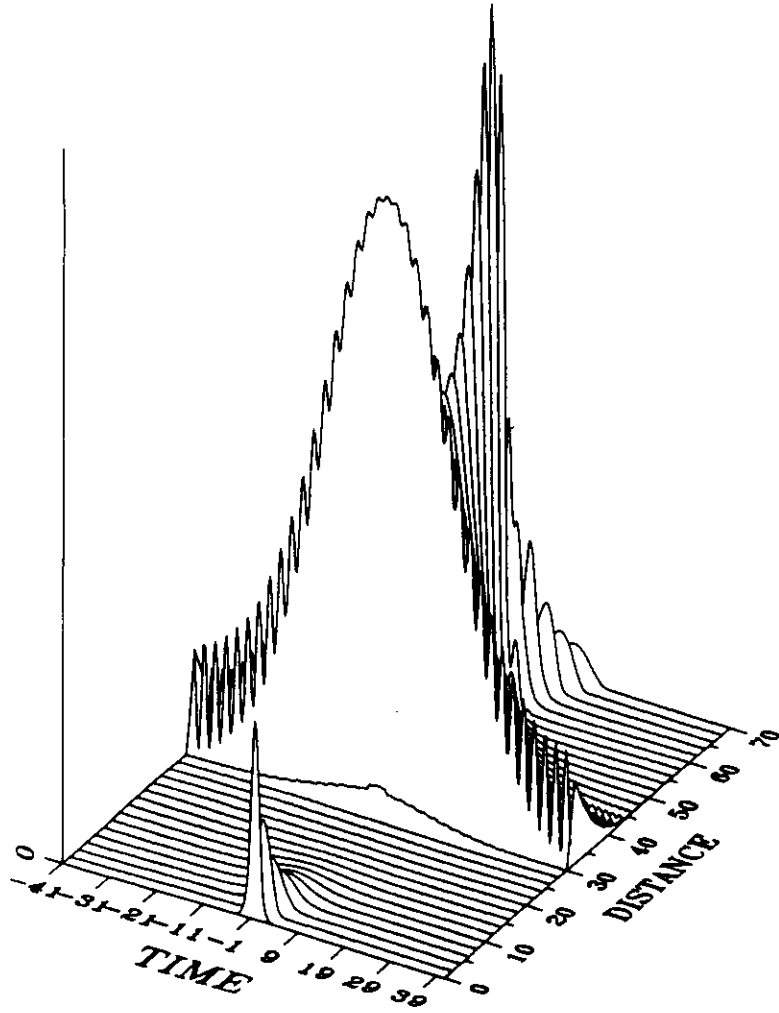


SIMULATION OF PULSE TRANSMISSION WITH MPSI; pulse powers in the linear regime

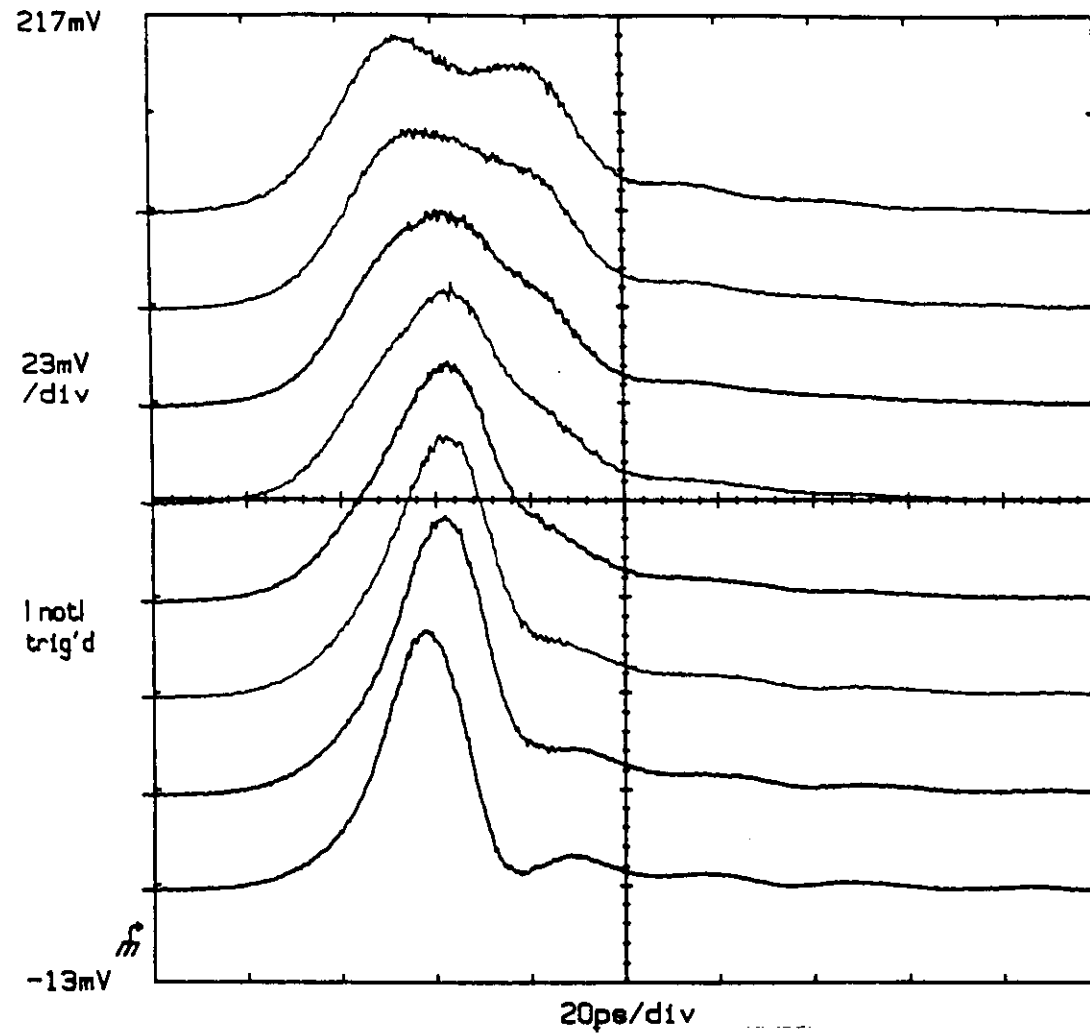


Source 10 ps FWHM @ 1535nm
Dispersion 16 ps/nm.km

SIMULATION OF PULSE TRANSMISSION WITH MPSI; pulse powers in the second span in the nonlinear regime



TRANSMITTED PULSES AFTER 101.5km WITH INCREASING POWER IN SECOND SPAN

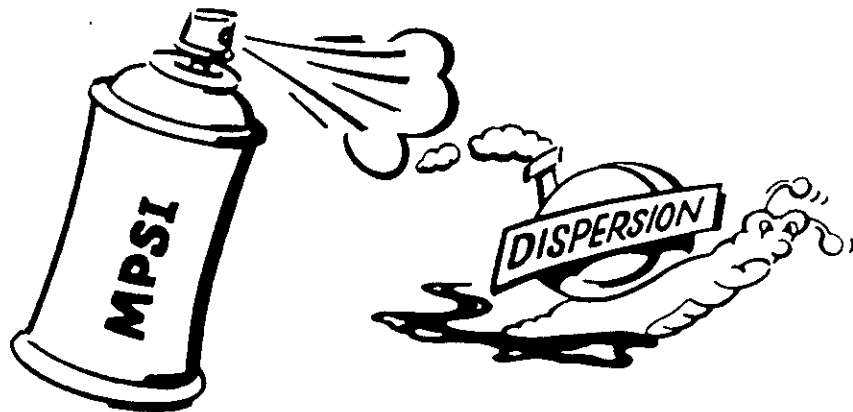


↑
Increasing
power

DISPERSION-COMPENSATION USING MPSI

CONCLUSIONS :

- Only one compensator required at link mid-point
- Compensates TX chirp and (first-order) fibre dispersion
- Tolerates some link non-linearity
- Very powerful technique (to 100 Tbit/s.km?)

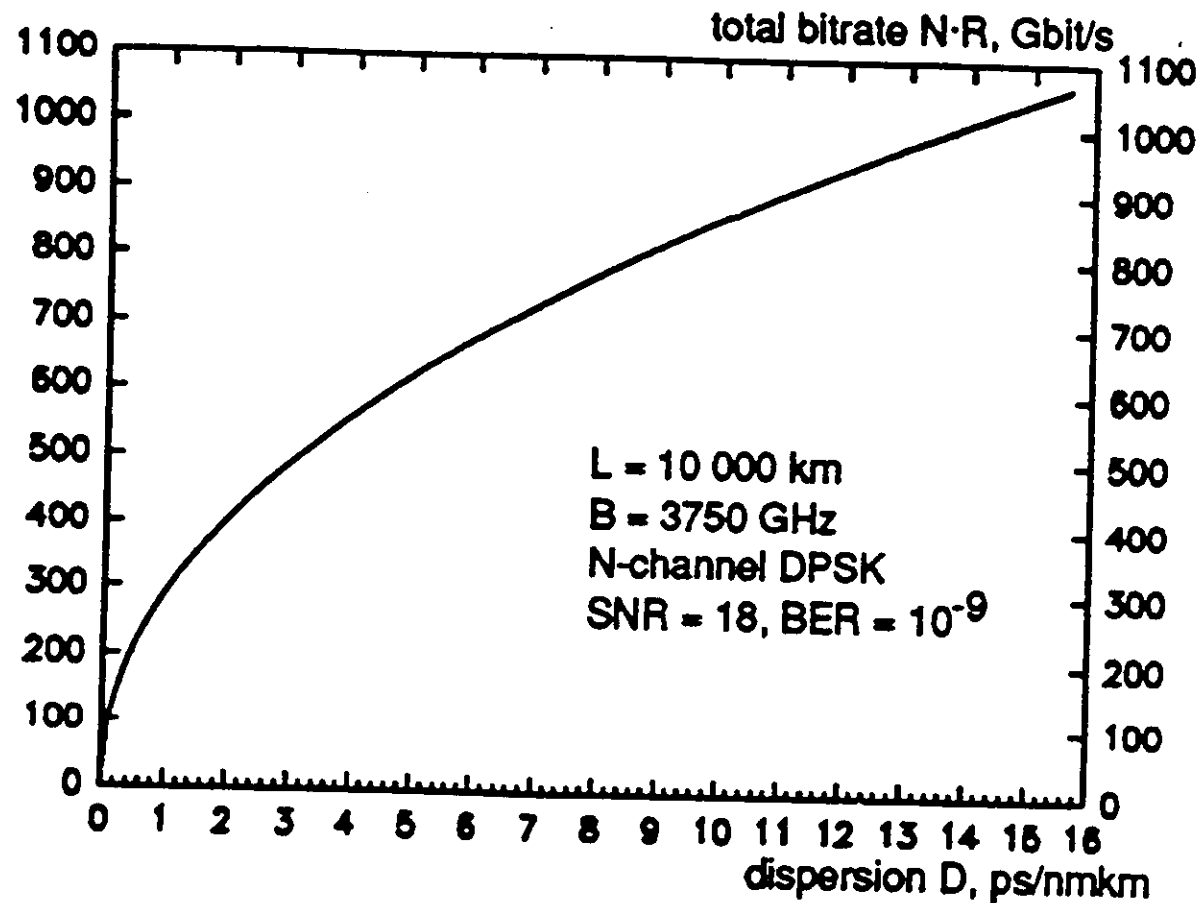


ADVANTAGES

- Standard fibre → *cheaper*
- Lower loss → *longer repeater spacings*
- High dispersion → *increased non-linear threshold*
- Phase conjugation → *minimises dispersion at all λ 's*
- Phase conjugation → *compensates fibre nonlinearity
(Kikuchi & Lorattanasane, OAA)*

Do we need solitons?

ULTIMATE TRANSMISSION CAPACITY OF OPTICALLY AMPLIFIED FIBRE SYSTEMS



A. Splett et al, ECOC '93

STANDARD FIBRE NETWORK UPGRADE; SUMMARY

■ 1.3 μ m

Semiconductor amplifiers

Fibre amplifiers

Solitons

- *crosstalk?*

- *inefficient at present*

■ 1.5 μ m

Solitons

Dispersion compensation

- *possible, but not easy*

- *fibre, gratings, MPSI*

- *allows maximum fibre bandwidth*

Do we need solitons?