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### *Laser Molecular Spectroscopy*

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## HIGH RESOLUTION FAR INFRARED SPECTROSCOPY

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## INTRODUCTION

The far infrared (FIR) spectral region (which we will define as 0.3 - 10 THz (1mm to 30  $\mu$ m)) plays an important role in molecular and atomic spectroscopy. The pure rotational transitions of light molecules, for example diatomic hydrides, occur in this region; FIR measurements provide the only accurate means of determining their rotational structure and associated characteristics such as line strengths, pressure broadening parameters, and permanent electric dipole moments. Heavier molecules can also be observed in the FIR. Their rotational transitions in this region involve high quantum numbers and therefore small but important centrifugal distortion effects can be measured. In atoms, transitions between fine structure levels as well as metastable levels occur at FIR frequencies.

Laboratory spectroscopic measurements in the far infrared are increasingly important for other branches of science. Radio astronomers are now making measurements in the FIR and rely heavily on the laboratory measured transition frequencies. The FIR molecular and atomic transitions are important in the characterization of interstellar molecular clouds; FIR measurements are useful for identifying species, measuring concentrations, and determining temperature profiles.<sup>1</sup> FIR spectroscopy is also important in the monitoring and modeling of the earth's atmosphere. For instance the concentration of the hydroxyl radical (an important molecule in the ozone production cycle) is monitored in balloon flights by observing its FIR spectrum.<sup>2,3</sup> Accurate determination of the concentration requires laboratory measurements of the pressure broadening coefficient at the appropriate frequency.

## FIR SPECTROMETERS

Classical spectrometers in the FIR use interferometric techniques coupled with blackbody sources and high sensitivity bolometric detectors. At present, the best available Fourier transform spectrometers provide complete coverage of the FIR region with 20 MHz spectral resolution and accuracy approaching 1 MHz.<sup>4,5</sup> They also have the advantage of performing wide spectral scans in fairly short periods and have a frequency range extending into the visible.

FIR laser spectroscopy is limited by the fact that no broadly tunable, narrow-band lasers exist in this region. However, over 1000 optically pumped, fixed frequency FIR laser lines do exist, with about 1 kHz linewidth;<sup>6</sup> the only problem is lack of tunability between lines. To overcome this problem laser magnetic resonance (LMR) was developed.<sup>7</sup> LMR uses a magnetic field to Zeeman tune a molecule or atom into resonance with the fixed frequency laser. Intracavity FIR LMR is one of the most sensitive spectroscopic techniques (it exhibits a minimum detectability of  $5 \times 10^{-10} \text{ cm}^{-1}$ ) and has wide applications in spectroscopy and chemical kinetics. For detailed LMR reviews, the reader is referred to references 8 and 9, reference 10 provides a good overview of the use of LMR in atomic spectroscopy.

As powerful and useful as FTS and LMR are in the far infrared, there are limitations to both techniques. LMR is only applicable to species with open electronic shells, has an accuracy limited to about 1 MHz, and requires the unravelling of the Zeeman splittings. With FTS, the resolution is often insufficient to observe hyperfine structure and the maximum accuracy of several MHz is only available on a few of the best spectrometers. For these reasons, other coherent sources have been developed. These all take advantage of nonlinear mixing to generate tunable, narrow-band far infrared radiation.

Harmonic multiplication of microwave oscillators was developed in the 1960's and at present provides microwave resolution and accuracy at frequencies up to 1 THz.<sup>11</sup> Frequencies above one THz have been synthesized in several different ways: 1) CO<sub>2</sub> laser difference frequency generation in GaAs<sup>12</sup> and other crystals<sup>13</sup>; 2) Generation of microwave sidebands on FIR laser radiation<sup>14-17</sup>; and 3) CO<sub>2</sub> laser difference frequency generation in metal-insulator-metal (MIM) diodes.<sup>18</sup> The first technique has never been employed for spectroscopy and will not be discussed further.

The generation of tunable FIR laser sidebands in a Schottky diode is used for spectroscopy in several laboratories and has produced fine results; however there are several serious drawbacks to this method. Although over 1000 FIR laser lines exist, sideband generation requires the most powerful FIR laser lines. This limits the spectral range to frequencies below 3 THz and also limits the coverage below 3 THz. The molecular transitions which can be measured are limited to those in the vicinity of a strong FIR laser line. The typical absolute

accuracy obtainable is  $\pm 2 \times 10^{-7}$ ,<sup>19</sup> ( $\pm 200$  kHz at 1 THz) determined by the re-settability of the FIR laser. This value is for a system designed for high frequency reproducibility; the frequency uncertainty can increase for a system designed to yield high FIR power. This limitation can in principle be overcome by measuring the FIR laser frequency simultaneously with the FIR spectra. The separation of the sideband radiation from the carrier is a less serious problem. Although these systems generate tens of  $\mu$ W sideband power, the sensitivity is limited by the presence of the stronger carrier. Even after interferometric separation of the sidebands, the power at the carrier frequency still limits the sensitivity to about  $2 \times 10^{-6}$  cm<sup>-1</sup>.

Generation of tunable FIR (TuFIR) radiation by the difference frequency between two CO<sub>2</sub> lasers in a MIM diode was first accomplished in 1984.<sup>18</sup> Later in the same laboratory microwave sidebands were generated on the difference frequency.<sup>20</sup> This method is superior in both frequency coverage and accuracy to the FIR laser sideband system and has comparable sensitivity. The frequency range extends from 0.3 to 6 THz, with greater than 95% coverage. The accuracy is limited by the accuracy of the CO<sub>2</sub> lasers, which are frequency stabilized to 25 kHz, yielding an FIR accuracy of 35 kHz. And although the FIR power generated is less than that using the FIR laser sideband method, the sensitivity is roughly the same because there is no strong carrier laser causing excess noise on the detector. The third order technique of CO<sub>2</sub> laser difference frequency plus microwave sidebands has been employed in a new spectrometer at LENS and will be described below.

### THIRD ORDER SPECTROMETER AT LENS

Figure 1 is a schematic of the third order TuFIR spectrometer. The radiation from lasers I and II are combined on a beamsplitter and coupled onto the diode by a 25 mm focal length lens. The microwave radiation is coupled onto the diode by a bias tee connected to the diode junction. The generated FIR is radiated from the diode's whisker in a long wire antenna pattern.<sup>21</sup> This FIR is then collected and collimated by a 30 mm focal length off-axis section of a parabolic mirror. After passing through an absorption cell the FIR is detected on a liquid He cooled Si bolometer. The FIR radiation is frequency modulated (via frequency modulation of the CO<sub>2</sub> lasers) and phase sensitive detected in a lock-in amplifier. The microwave frequency is swept by a personal computer, which also collects the data from the lock-in. Thus we are able to average scans for increased sensitivity. This is essentially the same experimental arrangement as described in ref. 20 only without the acousto-optic modulators isolating the lasers from the diode.

Both CO<sub>2</sub> lasers are frequency stabilized to the 4.3  $\mu$ m saturated fluorescence signal from low pressure CO<sub>2</sub> cells<sup>22</sup> (not shown in figure 1). These frequencies have been measured to an absolute frequency of better than 5 kHz,<sup>23,24</sup> but without the special locking techniques used in refs. 23 and 24 our measured stability is 25 kHz for each laser. The overall uncertainty of

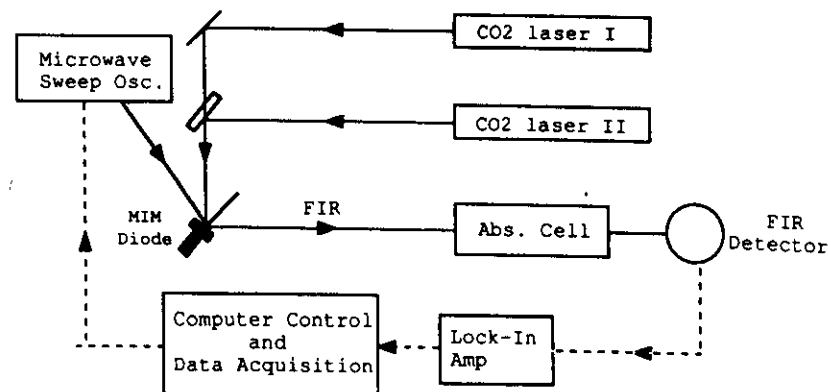


Fig. 1. Block diagram of the third order tunable FIR spectrometer.  $\nu_{\text{FIR}} = \nu_{\text{I}} - \nu_{\text{II}} \pm \nu_{\mu\text{w}}$ .

our FIR is thus  $2 \times 25$  kHz or 35 kHz. For further information on frequency stabilization of CO<sub>2</sub> lasers, the reader is referred to the article by L. Zink in this publication.

The frequency range of the spectrometer is from 0.3 to 6 THz. The lower limit is set by our bolometer and the upper limit by the largest difference frequency between <sup>12</sup>C<sup>16</sup>O<sub>2</sub> lasers. Over 100 lines oscillate on each CO<sub>2</sub> laser plus  $\pm 20$  GHz tunability of the microwave sweeper yield a 95% coverage of this region, the only gaps occur above 4.5 THz. The range could possibly be extended to 9.28 THz by substituting an ammonia laser for one of the CO<sub>2</sub> lasers.

The spectrometer sensitivity is limited by the FIR power and the sensitivity of the bolometer. 150 mW from each laser and 6-10 dBm of microwave power are applied to the MIM diode, generating about  $10^{-7}$  W of FIR radiation. Although there are two orders of magnitude less power than with the FIR laser sideband technique, the noise sources are also less. Our typical minimum detectable absorption is  $10^{-4}$  in a 1 s time constant, which for a meter long cell corresponds to  $10^{-6}$  cm<sup>-1</sup>. This is of comparable sensitivity to the laser sideband technique.

### MIM Diodes

The heart of the spectrometer is the metal-insulator-metal diode. Used as a rectifier for many years, its possible use for rectification of very high frequencies was noted as early as 1948.<sup>25</sup> The MIM diode has been used for laser frequency measurements since 1969<sup>26,27</sup> and used to measure frequencies up to 200 THz<sup>28</sup> and difference frequencies of visible light.<sup>29</sup> The measurement at 200 THz enabled the most accurate determination of the speed of light<sup>30</sup> and led to the re-definition of the meter.<sup>31</sup>

The diode consists of an electrochemically sharpened tungsten whisker (25  $\mu\text{m}$  in diameter and typically 3 to 7 mm long) contacting a metal base. The metal base has a naturally occurring thin oxide insulating layer. In the generation of the difference frequency between two  $\text{CO}_2$  lasers (second order) nickel is the best material for the base. In third order operation (difference frequency plus sidebands) cobalt appears to be superior. Good second order diodes generate very little third order FIR and vice versa. We also find that sharper tungsten whiskers produce more third order radiation. In general, third order diodes produce about 1/3 as much tunable FIR radiation as second order diodes and are more difficult to make, yet have the advantage of increased tunability. The physical properties of an ideal base material are not yet understood, hopefully in the future different materials or processing techniques will yield even greater TuFIR power.

#### Applications in Spectroscopy

The spectrometer at LENS has just been completed and the first results will be described below. However, a second order TuFIR spectrometer at the National Institute of Standards and Technology (NIST), in Boulder, Colorado has been in operation for several years and demonstrates the usefulness of this technique for high resolution spectroscopy. Spectra have been obtained for abundant, stable molecules such as  $\text{CO}$ ,<sup>20</sup> providing accurate standards for Fourier transform spectrometers. Free radicals and transient species have been observed,<sup>32</sup> including the molecular ions  $\text{H}_2\text{D}^+$  and  $\text{H}_3\text{O}^+$ . In addition to absorption spectroscopy the tunable FIR has been employed in an IR-FIR double resonance experiment;<sup>33</sup> for FIR Stark spectroscopy;<sup>34</sup> and in  $\text{O}_2$  pressure broadening measurements.<sup>35</sup>

The first molecule observed with the third order system at LENS (during the first week of this school) was  $^{13}\text{CO}$ . Traces of the  $J=6-5$  and the  $J=26-25$  lines are shown in figure 2. In all, seven transitions were measured, extending up to the  $J=30-29$  transition at 3.3 THz. New molecular parameters were determined with a 50 kHz (1 $\sigma$ ) standard deviation of the fit.

Because of the high frequency accuracy of this technique, small pressure induced shifts of the transition frequency can be measured. Pressure broadening and shift measurements of the  $\text{CH}_3\text{CN}$  molecule have been performed on transitions ranging from  $J=31-30$  at 570 GHz to the  $J=81-80$  transition at 1.5 THz. Although this data is still being analyzed, it does serve to compare our instrument with one of the FIR laser sideband spectrometers. Figure 3 illustrates different K components of the  $J=57-56$  transition measured by the two techniques; eight components were observed with video detection with a FIR laser sideband technique<sup>36</sup> and 12 components were measured with phase sensitive detection using the LENS third order spectrometer (four of which are shown in figure 3). The FIR laser sideband experiment using phase sensitive detection yielded a signal to noise of 50,<sup>17</sup> slightly less than our S/N of 90. Not only were more k components measured with our system, but the accuracy of these measurements is about one order of magnitude higher. The

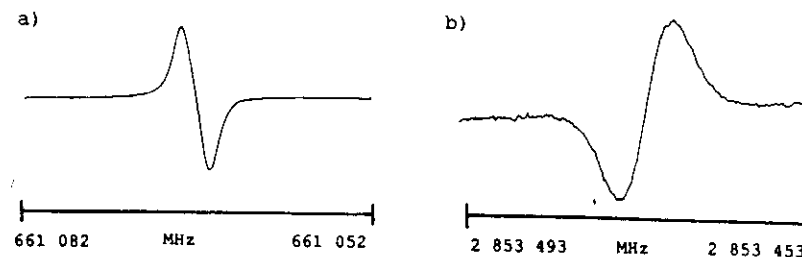


Figure 2. FIR rotational transitions of  $^{13}\text{CO}$  observed with the TuFIR spectrometer at LENS. a)  $J=6-5$  b)  $J=26-25$ .

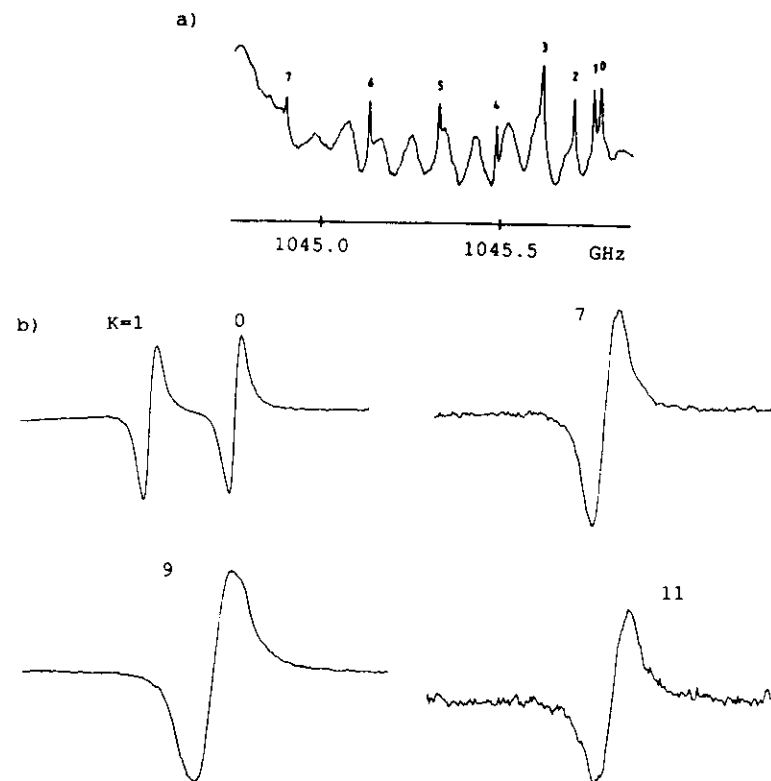


Figure 3. Observed transitions of the  $\text{CH}_3\text{CN}$  molecule;  $J=57-56$ ,  $\Delta K=0$ . a) 8 K components seen in video detection with a FIR laser sideband spectrometer (from ref. 36). b) 4 of the 12 components observed in phase sensitive detection with the third order spectrometer at LENS.

accuracy in ref. 36 is 300 - 700 kHz (determined by the re-settability of the FIR laser) compared to our accuracy of 55 - 100 kHz. Our uncertainty was limited by the S/N and the transition linewidth.

#### SUMMARY

The new third order TuFIR spectrometer at the European Laboratory for Nonlinear Spectroscopy has been described and the first results presented. This spectrometer provides nearly complete coverage from 0.3 to 6 THz, with 35 kHz accuracy and a minimum detectable absorption of 0.01% in a 1 second time constant. Plans for this system include the study of astrophysically important molecules, the study of molecular ions, and, taking advantage of the accuracy, measurements on trapped atomic ions (see the article by G. Werth in this publication).

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# High-Resolution and High-Sensitivity Spectroscopy Using Semiconductor Diode Lasers.

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## 1. - Introduction.

As clearly discussed by Prof. SCHAWLOW in his introductory lecture[1], we spectroscopists still have the problem of finding tunable lasers and, once the laser is tunable, we want it to be more and more monochromatic. Indeed, due to the central role played by the radiation source in coherent spectroscopy, the scientific achievements strictly followed the technical development of lasers.

Semiconductor diode lasers (SDL) are certainly lasers with a wide wavelength tuning range. Although they were invented more than thirty years ago, they were not used much in atomic spectroscopy until recently[2]. In the early times, diode lasers had to be operated at low temperatures, they emitted over several cavity modes and were unreliable. Recently, semiconductor diode lasers have been rapidly improving in power, spectral purity and wavelength coverage: they can now operate c.w. at room temperature and produce tens of milliwatt on a single mode. Nowadays these lasers are showing some specific properties which make them suitable sources of radiation for very-high-resolution and/or extremely-high-sensitivity experiments. In general, an extreme simplification of sophisticated experimental schemes can be obtained. As an example let us refer to the resolution of Lamb shift of  $H_\alpha$  transition in hydrogen. This measurement constitutes one of the cornerstones of atomic spectroscopy and the historical recording, obtained in 1972 at Stanford University by a dye-laser spectrometer, is reported both in Schawlow and Hänsch lectures[1,3]. The same recording is reported in fig. 1, as obtained in Munich[4] nearly twenty years later by using an InGaAlP free-running laser diode.

In this lecture, we show some of the possibilities offered by semiconductor

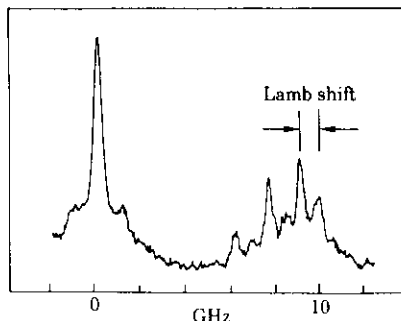


Fig. 1. - Doppler-free saturation spectrum of the red Balmer- $\alpha$  line of atomic hydrogen with resolved fine-structure components as recorded in Munich (ref.[4]) using a free-running InGaAlP laser diode.

diode lasers in atomic and molecular spectroscopy by describing experiments recently performed in the visible ((650 ÷ 690) nm) and near infrared ((750 ÷ 850) nm).

## 2. - Overview on diode laser characteristics and operation.

The diode laser basically works as a  $p-n$  junction biased in the forward direction. The recombination of electrons and holes takes place at the junction and leads to the emission of radiation at a frequency depending on the energy gap between the conduction and the valence bands. The reflectivity of the cleaved facets of the diode is sufficient to have dominance of stimulated emission.

In order to have continuous-wave operation of the laser, it is necessary to confine the carriers and the photons in a well-defined region in order to have enough gain. This is commonly obtained using double heterostructures, where the active region is sandwiched between two layers of semiconductors with a larger band gap. This prevents the carriers from escaping the active region, but also light is confined in the same active region because a higher band gap corresponds to a smaller index of refraction. In the first lasers, confinement was achieved only in the direction in which the current was injected (gain-guided lasers), while now also «index-guided» lasers are available, in which the active region is limited in all directions by a higher-band-gap material. It is evident that real devices have complex structures and require important technological advances[5]. More recently, also quantum-well and multiple-quantum-well lasers became available, with lower threshold currents and higher output powers compared with normal double heterostructures.

Important characteristics which are relevant for the use of SDL sources in atomic and molecular spectroscopy are spectral coverage, tunability, spectral linewidth and amplitude stability.

The emission wavelength depends on the band gap and in principle it should be possible to construct lasers at all wavelengths by choosing the proper stoichiometric abundance for the dopant element, provided that direct interband transitions can occur. For instance, one could obtain emission from 1100 to 1650 nm using InGaAsP, or cover the (700 ÷ 890) nm and (630 ÷ 690) nm intervals by means of GaAlAs and AlGaInP, respectively. Of course, the material must have good optical and electrical characteristics in order to minimize losses. For this reason, for instance, room temperature operation in the blu-green region was only recently demonstrated. In practice, laser diodes are commercially available only at sparsely distributed emission wavelengths and the experimentalist soon faces the problem of tuning the laser to the precise atomic or molecular absorption. At a first attempt, one can change the emitted wavelength by changing the diode temperature. Indeed, the semiconductor gap depends on the temperature and overall spans of the order of some 20 nm can easily be obtained, considering that operation at low temperatures is even more efficient and that operation down to liquid-helium temperature has been demonstrated. Tuning by temperature, however, is discontinuous, and this is a significant limitation. Changes in temperature lead to a change of the cavity length and hence of the cavity modes. Typical tuning with this effect is of the order of 0.06 nm/°C ( $\sim 30$  GHz/°C). At the same time, the gain curve shifts by about 0.25 nm/°C. Since the two temperature dependences are different, we have a continuous tuning ( $\sim 30$  GHz/°C) until the longitudinal-mode hopping caused by the gain curve shift: the general aspect is that of a staircase with sloping steps, essentially causing wavelength changes in a rather unpredictable way. From the above considerations we also learn that an extremely good thermal stability (better than 1 mK) is required in view of narrow-linewidth operation. Thermal tuning is also obtained via Joule effect by changing the current of operation. This tuning, at a rate of about 3 GHz/mA, is rather slow while a faster control, of about 300 MHz/mA, can be obtained thanks to the change in the medium refractive index caused by the change in the current density. Also in this case we deduce that an injection current stability of at least 1  $\mu$ A is required to reduce effects on the linewidth.

Let us now in fact discuss the problem of spectral linewidth. An exhaustive treatment of the different factors determining the linewidth can be found in ref.[6]. However, these lasers are so small, 100 to 500  $\mu$ m long, 0.1 to 2  $\mu$ m thick and 2 to 20  $\mu$ m wide, that the predominant broadening is caused by the quantum fluctuations of the phase as early described by SCHAWLOW and TOWNES[7], who showed the limit linewidth to be inversely proportional to the output power and to the square of the optical-mode volume. Here we report, from ref.[8], the modified formula which, by the introduction of a factor  $\alpha$ , takes into account the



dependence of the refractive index on the carrier density:

$$(1) \quad \Delta\nu_{FWHM} = \frac{h\nu}{8\pi P_0} \left( \frac{c}{nL} \right)^2 \left[ aL + \ln \left( \frac{1}{R} \right) \right] \ln \left( \frac{1}{R} \right) n_{sp} (1 + \alpha^2),$$

where  $P_0$  is the output power,  $L$  is the cavity length,  $R$  is the facet reflectivity,  $a$  gives the distributed losses in the cavity,  $c/n$  the group velocity, and  $n_{sp}$  the spontaneous-emission factor which is of the order of unity. In the case of a typical diode laser ( $P_0 = 10$  mW,  $L = 300$   $\mu$ m,  $R = 0.3$ ,  $\alpha \sim 4 \div 5$ ), a linewidth of several MHz can be estimated. An increase of the cavity length or, more generally, of the cavity quality factor  $Q$  is hence needed to reduce the linewidth. As we shall see, this is at the base of the development of external optical cavities for line-narrowing purposes.

As for amplitude stability, the intrinsic stability of the structure of the semiconductor laser reduces the sources of noise. The properties of the active medium give rise to a peculiar AM and FM noise spectrum, which has been accurately investigated by HOLLBERG at NIST[9]. However, in the low-frequency range (frequencies lower than 10 MHz), typical for simple absorption spectroscopy experiments, an amplitude noise only (10  $\div$  20) dB above the shot noise level has been measured[10], while the corresponding one for dye lasers can be even two orders of magnitude larger. At 800 nm and 5 mW power, we can compute a shot noise level of the order of  $10^{-8}$ , hence a factor 10 dB worse yields for the SDL a relative amplitude stability of the order of  $10^{-6} \div 10^{-7}$ , at 1 Hz bandwidth. This impressive stability is at the base of the development of high-sensitivity absorption spectroscopy, for instance applied to molecular forbidden transitions, as we shall discuss in the last section of this contribution.

### 3. - Line narrowing and tuning with grating feedback.

We have seen that the linewidth of commercial SDLs is too large for applications in high-resolution spectroscopy, however it can be easily reduced by means of electronic or optical feedbacks[10].

In electronic-feedback schemes, the injection current is changed and very fast electronics is needed since the FM noise extends to high frequencies. The change in the injection current produces changes not only in the emitted frequency but also in the emitted power and this introduces an additional amplitude noise. Of course, the electronic stabilization method does not provide any control of the emission wavelength.

More popular are certainly optical-feedback methods. They are based on the simple idea of increasing the  $Q$  factor by coupling the diode to another cavity. Weak-feedback and strong-feedback regimes are both successful. The first is, for instance, ensured by the coupling with a high- $Q$  Fabry-Perot interferome-

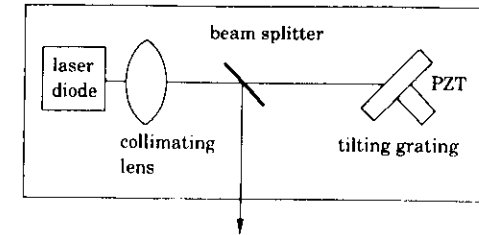


Fig. 2. - Scheme of the «pseudo-external-cavity» configuration adopted for the frequency narrowing and tuning of commercial semiconductor diode lasers.

ter, while in the second the laser is operated in an external cavity. If the feedback is selective in wavelength, it is possible consequently to control the emission. In this lecture we concentrate on the scheme using the strong feedback, which has been the most used in our laboratory after learning the art from HOLLBERG during his visit to LENS[11]. Our design, as the many others nowadays used all over the world, is a sort of compromise between easy construction, reliability, good spectral characteristics for high-resolution and high-sensitivity spectroscopy. Typically we can obtain a reduction by about two orders of magnitude of the emission linewidth and achieve a continuous wavelength tuning within a range of several nanometers at fixed temperature. In general one should use antireflection coating on the diode facet in order to increase the amount of feedback light coupled in. However, many of the best diode lasers available on the market are already provided with a high-reflectance coating on the back facet and a reduced-reflectance coating on the output facet. This allows one to operate the diode lasers in a pseudo-external-cavity configuration, as schematically shown in fig. 2. The first-order diffracted beam from a grating, mounted in the Littrow configuration, is fed back into the laser diode. An intra-cavity beam splitter ((20  $\div$  30)% reflectance) provides an additional output to the zeroth-order one and somewhat controls the feedback amount. Indeed a higher feedback reduces the out-coupled power, but also increases the wavelength tuning range. Further details concerning collimating lens and in general the actual construction can be found elsewhere[12], as well as the description of how the diode temperature and injection current stability stringent requirements can be fulfilled.

The simple rotation of the grating allows the coarse tuning of the extended-cavity lasers over a range of  $\sim 10$  nm. In particular, because of the finite reflectivity of the chip facets, the longitudinal modes of the solitary laser diode do not change, while the feedback from the grating allows one to select one of these modes and to drastically suppress other side modes. Fine tuning to the frequency of interest can then be achieved by slightly changing the temperature and the injection current of the diode. Continuous frequency scans of the order

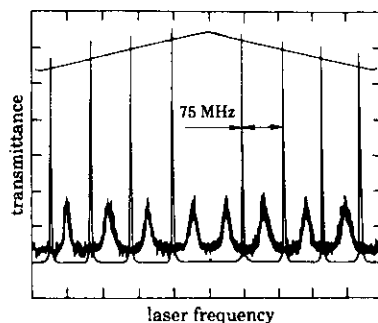


Fig. 3. – Superimposed transmission spectra from a 75 MHz free-spectral-range Fabry-Perot interferometer as the frequency of the diode laser is scanned up and down without optical feedback (broad peaks) and in the presence of optical feedback from a grating (narrow peaks). In the first case, the laser frequency was scanned up and down by sending a ramp (shown in the top of the figure) to the injection current supply. In the second case, the ramp was also sent to the PZT on the grating in order to simultaneously change the length of the external cavity.

of  $\sim 10$  GHz are accomplished by synchronously sweeping the length of the cavity, by means of a piezoelectric transducer on the grating, and the injection current. The effect of the increased  $Q$  factor in the cavity and consequent line narrowing can be evidenced by recording the transmission peaks of a Fabry-Perot interferometer as the laser frequency is scanned. A typical situation is illustrated in fig. 3. Broad peaks are recorded by blocking the return beam from the grating (*i.e.* with the laser as is commercially available), while much narrower peaks (the figure is essentially a recording of the Fabry-Perot finesse) are obtained in the presence of the optical feedback. Figure 3 has been taken using an InGaP laser at 690 nm [13], but the results are rather general as we have experi-

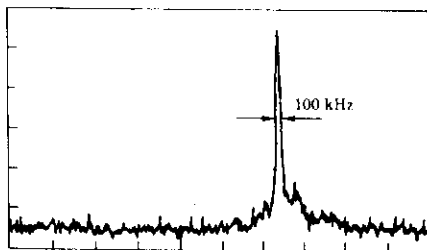


Fig. 4. – Beat note from the heterodyne between two free-running AlGaAs multiple-quantum-well lasers mounted in extended-cavity configuration. Wavelength is grating-tuned at  $\lambda = 850$  nm and the two emissions are about 2.2 GHz apart. From the recording, a value of about 50 kHz can be estimated for the laser linewidth in the millisecond time scale.

enced with AlGaAs lasers operating at 780, 820 and 850 nm, both using double heterostructures and multiple quantum wells. A more accurate analysis of the spectral properties of extended-cavity semiconductor diode lasers can be performed by observing the beat note from two of them superimposed in a nonlinear detector. This is illustrated in fig. 4 for the mixing in a fast photodiode of two extended-cavity multiple-quantum-well lasers at 850 nm. In the millisecond time scale the laser linewidth is of the order of 50 kHz, corresponding to  $\Delta\nu/\nu \approx 10^{-11}$ , while acoustic noise is responsible for the increase to about 400 kHz over longer time scale. Heterodyne measurements require particular care in the optical isolation between the lasers, to avoid injection locking which can be caused by even a slight feedback.

#### 4. – High-resolution spectroscopy of atoms.

The impressive characteristics of the conductor diode laser configurations so far described have already allowed a large number of interesting applications in atomic spectroscopy. Here we shall concentrate on very few of them particularly interesting to show the potentialities of these compact laser sources.

Let us start with the investigation of forbidden transitions. This is in general an important task for atomic spectroscopy mostly related to the possibility of developing laser-based frequency markers in the visible. We chose to detect the  $5s\ ^1S^0-5p\ ^3P_1$  intercombination transition of strontium at 689.488 nm. This wavelength is accessible to the newly available InGaP lasers and the sub-Doppler recording can provide a significant test, as the atomic line has a radiative width of only 8 kHz [13]. Strontium atoms were produced in a hollow-cathode discharge designed to allow access to the counterpropagating beams necessary for the Doppler-free recording [14]. Indeed, in spite of the weakness of the transition, a 1% absorption signal could be directly detected with a small saturation dip ( $\sim 5\%$ ) on top, thanks to the low-amplitude noise of the diode laser. The Doppler-free contribution to the signal could be consequently enhanced using a polarization spectroscopy scheme [15]. A typical recording is reported in fig. 5. Except for the radiation source, which in that case was a more conventional dye laser, a description of the experimental details can be found in [16]. The recorded linewidth was mostly determined by collisional effects due to the discharge environment and the laser jitter contribution could be estimated to be less than 1 MHz. More recently, the same intercombination transition has been observed and investigated in a Sr atomic beam [17] for a more careful analysis of the laser linewidth.

Let us now further illustrate the potentialities of grating-tuned semiconductor diode lasers showing and discussing some applications to the spectrum of atomic oxygen. Elusive to continuous-wave excitation from the ground state, this atom can conveniently be studied by means of laser spectroscopy of transi-

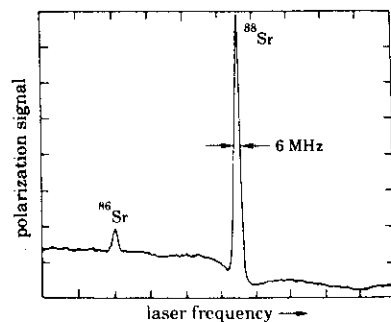


Fig. 5. - Doppler-free polarization spectroscopy of the  $5s\ ^1S_0-5p\ ^3P_1$  intercombination transition of strontium at 689.488 nm as recorded with an InGaP semiconductor laser in extended-cavity configuration.

tions starting from excited levels produced in the same Ar or He discharge used for the dissociation of  $O_2$  added in traces[18]. As evidenced in the scheme of fig. 6, transitions starting from the lowest triplet and quintet states are well suited for the investigation by means of AlGaAs lasers. The transition at 777 nm has immediately been recognized as interesting[11] because its lower level is metastable and possibly involved in a novel scheme for radiative cooling of

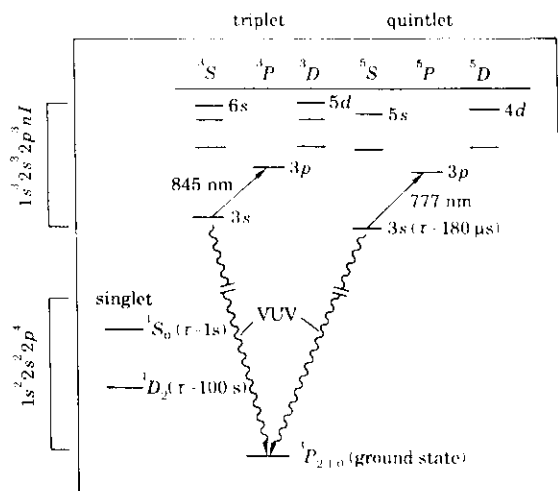


Fig. 6. - Simplified energy level scheme of atomic oxygen. Transitions studied with high resolution by means of AlGaAs diode lasers are evidenced.

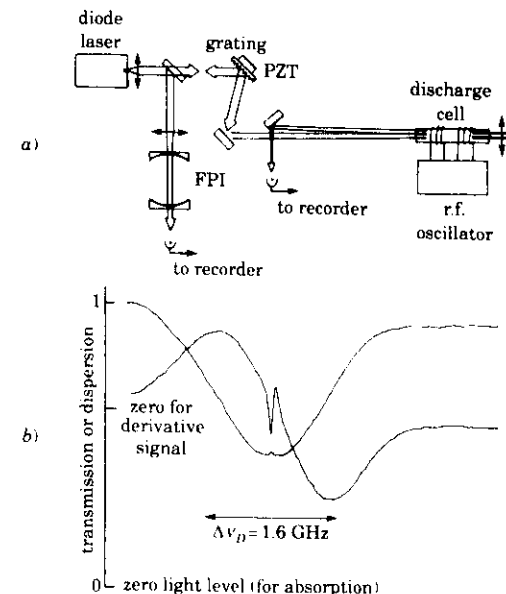


Fig. 7. - Oxygen absorption on the  $^3S_2-P_3$  transition at 777 nm as detected with a grating-tuned AlGaAs semiconductor diode laser. Two counterpropagating beams are present in the cell and a saturation dip can be recorded at the line centre. This is evident in the derivative signal (b) where phase-sensitive detection is applied.

oxygen. Also, the metastable level can be efficiently populated in a weak radiofrequency discharge and fig. 7a) reports the absorption signal as recorded on the transmitted light after passing twice through the cell: the laser beam was retroreflected by a mirror to allow saturation spectroscopy. It is worth noting the low noise level which reflects the extreme reduction in the laser amplitude fluctuations. By adding a modulation in the scanning piezo voltage, frequency modulation of the laser output can be easily achieved. This, in combination with phase-sensitive detection, allows the recording of a derivative signal and the enhancement of the saturation dip at the line centre, as displayed in fig. 7b). Let us now use this same atomic transition to illustrate more in general the versatility of the semiconductor laser as combined with different nonlinear spectroscopy techniques. The scheme of a general apparatus[19] is shown in fig. 8. A key role is played by the electro-optic modulator (EOM) which, in combination with the proper choice of the  $A$  and  $B$  optical devices, can act as a simple amplitude modulator, an alternate  $\sigma^+-\sigma^-$  polarizer, a  $\pi$  polarizer and so on. For instance, when  $A$  is not used and the EOM is placed between two parallel linear

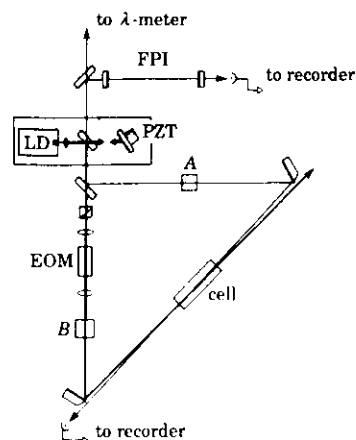


Fig. 8. - A general experimental apparatus used to combine a semiconductor diode laser with various Doppler-free spectroscopic schemes. A and B indicate different optical devices depending on the detection technique. A high-finesse Fabry-Perot interferometer (FPI) is used for the frequency scan calibration, while laser wavelength is determined by a  $\lambda$ -meter.

polarizers, the system is equivalent to a mechanical chopper and the scheme is essentially a Hänsch-type one. Consequently, a typical recording is shown in fig. 9a). As expected, only the sub-Doppler signal is recorded, while a broader pedestal is caused by the velocity-changing collisional effects[20]. The lineshape can be fitted to the sum of a Lorentzian and a Gaussian and the two contributions are separately shown in the lower part of fig. 9b). As is well known, the velocity-changing collision contribution to the signal can be reduced by changing the physical observable under investigation in the technique. In this respect an important example is provided by polarization spectroscopy, where the vapour birefringence, instead of the absorption, is recorded. For instance, the recording previously reported in fig. 5 for strontium showed no evidence for velocity-changing collisional pedestal, which on the contrary would have been rather important in a simple saturation spectroscopy scheme, the observed transition being originated from an atomic ground state. Coming back to the scheme of fig. 8, the pump beam can be manipulated in order to create an atomic orientation  $\sum_j n_{m_j} m_j \neq 0$  or an atomic alignment  $\sum_j n_{m_j} m_j^2 - (1/3)J(J+1) \neq 0$  with  $n_j$  and  $m_j$ , respectively, the relative population and the eigenvalue of the Zeeman sublevel  $J_z$ . Differently from the level population, observable involved in simple saturation spectroscopy, both the orientation and the alignment are significantly affected by collisions and the atoms emerging from a collision do no

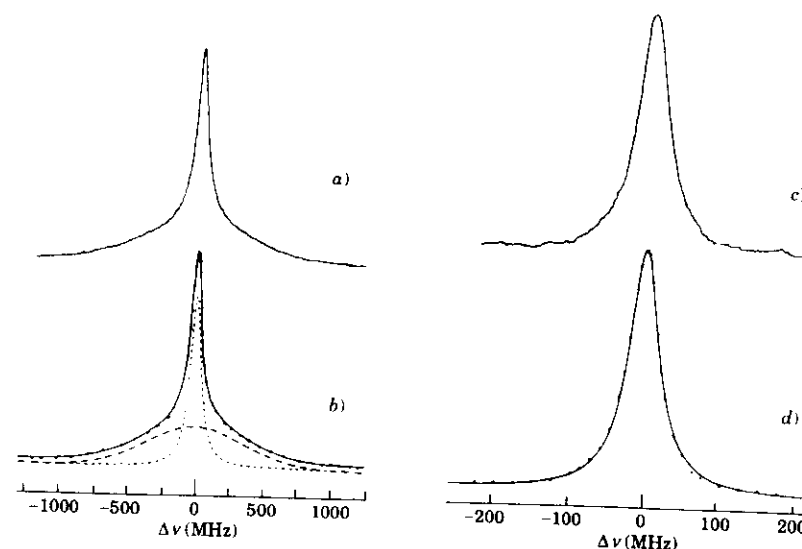


Fig. 9. - Examples of different sub-Doppler lineshapes obtained using the apparatus illustrated in fig. 8. In a), the saturation spectroscopy signal (Hänsch-type configuration) is the superposition of a Lorentzian and a Gaussian pedestal caused by the velocity-changing collisions, as evidenced by the fit in b). In c) the «alignment» spectroscopy configuration allows the elimination of the Gaussian contribution to the lineshape which in d) is fitted to a pure Lorentzian.

longer contribute to the recorded signal in the probe beam. This is clearly evident in fig. 9c), where the same transition of fig. 9a) and 7 is recorded using a configuration of «alignment» spectroscopy. The lineshape can be fitted using only a Lorentzian as shown in fig. 9d). The width of the Lorentzian is affected by some collisional broadening, since the experiment is performed in a cell. However, the sensitivity of the technique allowed the operation with low pressures ( $\sim 100$  mTorr of Ar buffer gas) and the extrapolation to zero pressure led to a determination of the radiative lifetime of the upper level of the transition ( $^5P_3$ ,  $\tau = 7$  ns). This lifetime is an important parameter for the real feasibility of the oxygen radiative cooling scheme proposed in [11] and analysed with a Monte Carlo simulation in [21].

The simplicity of the semiconductor lasers makes easy, and accessible to most laboratories, the operation of many of them in the same experiment. This is, for instance, useful when the experiments consist in the precise measurement of isotope shifts or atomic structures. Direct frequency measurements can indeed be obtained by heterodyning two independent lasers frequency-locked

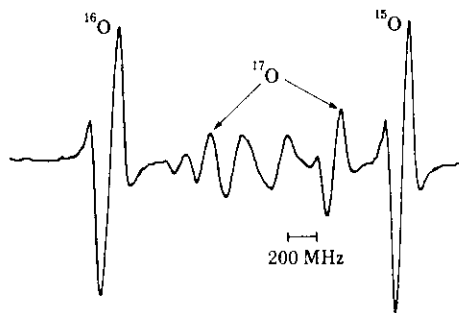


Fig. 10. - Sub-Doppler spectrum of the three stable isotopes of atomic oxygen as observed in an enriched sample ( $^5S_2$ - $^5P_3$  transition at 777 nm). The third-derivative signal is shown.

on two different atomic signals. For atomic oxygen this was the case for the  $^{16}\text{O}$ ,  $^{17}\text{O}$ ,  $^{18}\text{O}$  isotope shift and for the hyperfine structure of  $^{17}\text{O}$ . In fig. 10 the 777 nm transition using an enriched atomic sample[22] is recorded. A  $3f$  detection scheme was used to eliminate the background slope, which instead is present in the first-derivative signal. The sample absorption was  $\sim 5\%$  and the Lamb dip signal was  $\sim 5\%$  of the absorption. The recording reported in the figure was obtained with a lock-in time constant of 1 ms. A servo loop (bandwidth  $\sim 10$  Hz) was then used for frequency-locking each laser on the selected isotope signal. The scheme typically consisted of a lock-in, an integrator and a high-voltage amplifier; the output of the integrator was fed back to the high-voltage amplifier, which controlled the grating position by means of the piezoelectric transducer. The intrinsic frequency stability of such a laser system is determined by the increased  $Q$  of the optical cavity, which reduces the fast frequency fluctuations; only a relatively slow electronics is then required to correct long-term drifts. The beat note between two lasers can be observed using a spectrum analyser, as already shown in fig. 4. However, in the actual measurements with the lasers locked on the atomic third-derivative signals, a broadening of the beat note arises from the frequency modulation of the lasers and can be minimized by in-phase modulation. The accuracy of the heterodyne measurement is determined by the signal-to-noise ratio and width of the locking signal. From signals like those shown in fig. 10, isotope shifts and hyperfine structures could be measured with an accuracy of the order of 100 kHz. It is evident that heterodyne of semiconductor diode lasers can be a powerful and straightforward method for precise laser spectroscopy, in particular in combination with mixing schemes based on fast photodiodes or, even better, on GaAs Schottky diodes[9, 23, 24].

## 5. - High-sensitivity spectroscopy of forbidden molecular transitions.

As anticipated, the extremely good amplitude stability of semiconductor diode lasers opens the possibility of performing high-sensitivity absorption measurements directly on the light transmitted by the sample.

A first example is given by the detection and investigation of molecular oxygen at 760 nm. Transitions of  $\text{O}_2$  in the visible-near infrared belong to the  $\chi^3\Sigma_g^- - b^1\Sigma_g^+$  band first observed by BABCOCK and HERZBERG[27] in the atmospheric absorption from Mount Wilson. These are magnetic-dipole transitions with an absorption coefficient as low as  $\alpha \sim 10^{-6} \text{ cm}^{-1}$ . In the laboratory, they could be detected only by means of intracavity techniques involving dye lasers[28, 29]. The powerfulness of stable diode lasers for the detection of these forbidden transitions was early suggested to us by the recording of an absorption from the room atmosphere after a path length of only 20 cm[30]. This opened the possibility of performing more systematic investigations using an extended-cavity semiconductor diode laser spectrometer. The scheme of the spectrometer, which, as we shall discuss, is also used more in general for overtone molecular spectroscopy, is shown in fig. 11. The laser frequency is controlled by changing PZT voltage and injection current as done for atomic spectroscopy. A dither (fast modulation) is added when phase-sensitive detection is used. The radiation transmitted by the sample is detected by means of a photodiode, while a second photodiode is used to obtain a reference for the laser amplitude and to eliminate the baseline (due to the amplitude modulation) by means of differential detection. The absorption cell is 150 cm long. Typical measurements for oxygen are illustrated in fig. 12. An absorption recorded

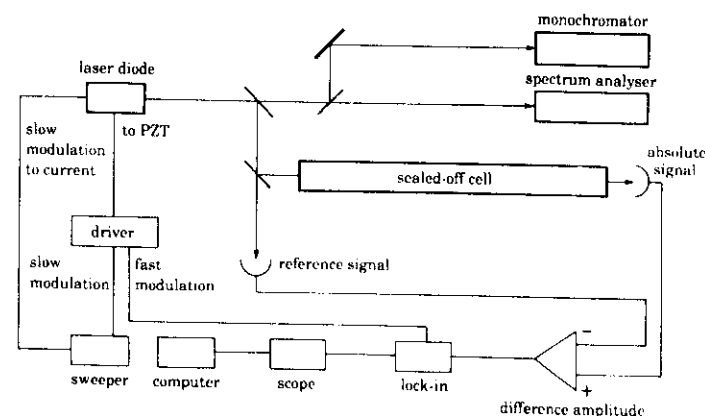


Fig. 11. - Scheme of the diode laser spectrometer used for the investigation of weak molecular absorptions.

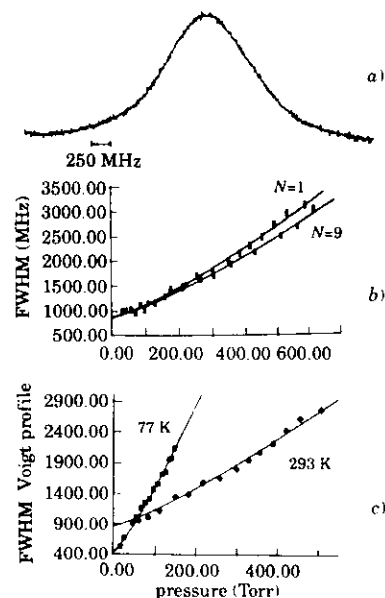


Fig. 12. - Absorption of molecular oxygen on the  ${}^1P(K''=9)$ ,  $v=0-0$  component of the forbidden  $\chi^3\Sigma_g^- - b^1\Sigma_g^+$  band (a)). In b) the dependence (on two different components) of the homogeneous contribution to the linewidth on the pressure is reported. In c) the behaviour of the collisional broadening on the same component at two different temperatures is shown.

with a gas pressure of 120 Torr is shown in fig. 12a). The lineshape is fitted to the convolution of a Gaussian and a Lorentzian. The latter contribution is due to collisional broadening and can be investigated at different pressures[31]. The results of different measurements in the (0 ÷ 1000) Torr range (on two different components) are plotted in fig. 12b), while in fig. 12c) the behaviour of the collisional broadening at two different temperatures is shown. This demonstrates how important parameters for atmospheric physics can be extracted from a relatively simple spectrometer. Wider applications can certainly be found in overtone molecular spectroscopy. Fundamental vibrations of molecules produce spectra in the medium infrared. The anharmonic contribution to the potential makes possible the excitation of overtone frequencies which can occur in the visible-near infrared, where more convenient tunable laser sources and room temperature sensitive detectors are available. Specially for polyatomic molecules, where also combination vibrations are present, overtone spectra are rich and interesting to be investigated also for a better modelling of the molecular structure. However, it is also well known that the intensity of successive

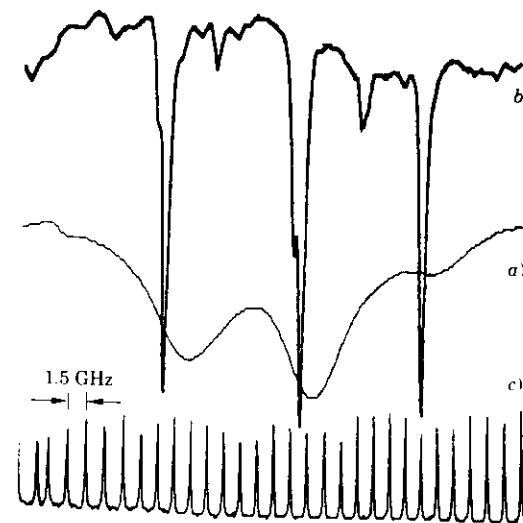


Fig. 13. - Recording of the absorption of  $\text{CH}_4$ , methane, around 886 nm by means of AlGaAs laser. Temperature and pressure are, respectively, 293 K and 100 Torr in a) and 77 K and 1 Torr in b). The total frequency scan is 42 GHz, as determined by the transmission of a reference Fabry-Perot (c)).

overtone and combination transitions decreases very rapidly with the increase of the number of vibrational modes and/or quanta involved. As a consequence, overtone molecular absorption in the visible is very weak ( $\alpha \sim 10^{-6} \text{ cm}^{-1}$  can be again a typical absorption coefficient) and constitutes an interesting challenge for the spectrometer based on our semiconductor diode lasers. A comprehensive list of the most interesting molecules and of the wavelengths at which they are accessible can be found elsewhere[32]. Let us illustrate the powerfulness of the technique using the detection of  $\text{CH}_4$ , methane, at 886 nm, where a number of combination bands are superimposed ( $3\nu_1 + \nu_3$ ,  $2\nu_1 + 2\nu_3$ ,  $\nu_1 + 3\nu_3$ ,  $4\nu_3$ ). This produces an almost continuous absorption[33] as, for instance, shown in fig. 13a). The spectrum is part of a systematic investigation performed at different temperatures to study the behaviour of the various components and possibly facilitate the line assignment. Indeed the room temperature recording, taken with  $\text{CH}_4$ , pressure of 100 Torr, is simplified when the temperature is lowered to 77 K, as shown in fig. 13b). In this case also the pressure could be reduced to 1 Torr and the linewidths are narrower because of the reduced Doppler and collisional broadenings. It is obvious that the signal-to-noise ratios observed for the detection of these overtone transitions, as well as of the oxygen forbidden

transitions, can constitute a good base for the development of high-sensitivity absorption measurements. Indeed, the transitions in methane provide a system to test and compare all the existing techniques mainly based on the modulation of the laser radiation[34]. Semiconductor lasers offer the advantage of achieving an easy and fast frequency modulation of the laser output itself by acting on the injection current. For this reason the diode lasers constitute optimal candidates for detection techniques as wavelength and frequency modulation.

In the wavelength modulation technique an a.c. component is added to the injection current of the diode laser; the frequency of modulation must be smaller than the linewidth of interest, so that the absorption is probed simultaneously by a number of sidebands. If the modulation frequency is increased to exceed the linewidth, only one sideband will be absorbed at a time, giving rise to a characteristic heterodyne beat signal at the modulation frequency. If the modulation index is chosen so that only one higher sideband and one lower sideband have appreciable amplitude, we are in the case of frequency modulation spectroscopy (one-tone FM).

A variation on the theme of the FM spectroscopy with respect to the one-tone technique, previously described, is represented by the two-tone spectroscopy: in this case the laser emission is simultaneously modulated at two distinct but closely spaced angular frequencies. Also here a heterodyne signal is obtained, but it occurs at the difference frequency between the two sidebands eliminating the need of high-speed detectors and making large-linewidth detections possible.

The trick of the modulation techniques is to observe the absorbed signal with an associated noise typical of the modulation-demodulation frequency and the electronic bandwidth chosen. Because of the  $1/f$  behaviour of the amplitude noise in the power spectrum of semiconductor diode laser emission, the possibility of obtaining more sensitive detections is bound to the choice of higher demodulation frequencies.

In fig. 14a) the direct absorbed signal of one  $\text{CH}_4$  component without any demodulation at 100 Torr is shown (corresponding to a relative absorption of 3.6%); the associated S/N is more than 20 and the bandwidth equal to 100 Hz. In fig. 14b) the demodulated signal (third derivative) is shown at 1 kHz, 100 Hz of bandwidth and a pressure of 100 mTorr, corresponding to a S/N of about 10.

Finally, in fig. 14c) the demodulated signal with the two-tone technique at frequencies around 500 MHz and 1 Hz of bandwidth is shown (pressure of 500 mTorr) corresponding to a S/N ratio of a few thousands.

A minimum detectable absorption has been demonstrated[34] (corresponding to a S/N ratio equal to 1) for the low-wavelength modulation (1 kHz), high-wavelength modulation (100 MHz) and two-tone frequency modulation  $((390 \pm 5) \text{ MHz})$  techniques, respectively, of  $4.5(1) \cdot 10^{-7}$ ,  $9.7(3) \cdot 10^{-8}$  and  $6.4(2) \cdot 10^{-8}$ .

Considering the absorption limit derived by the calculated shot noise, the

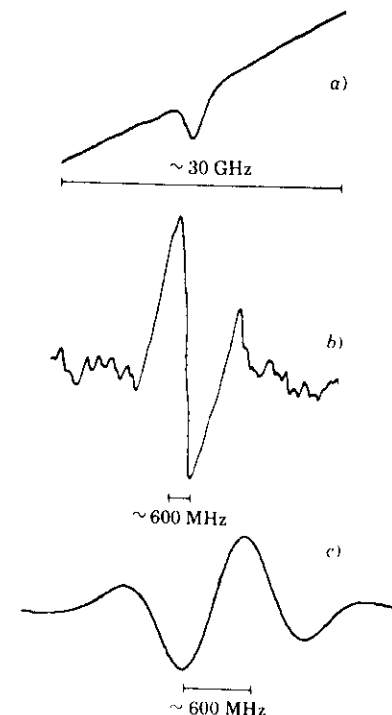


Fig. 14. - a) represents the direct absorbed signal of one  $\text{CH}_4$  component without any demodulation at 100 Torr (corresponding to a relative absorption of 3.6%); the associated S/N is more than 20 and the bandwidth equal to 100 Hz. b) represents the demodulated signal (third derivative) at 1 kHz, 100 Hz of bandwidth and a pressure of 100 mTorr, corresponding to a S/N of about 10. c) represents the demodulated signal with the two-tone technique at frequencies around 500 MHz and 1 Hz of bandwidth (pressure of 500 mTorr) corresponding to a S/N ratio of a few thousands.

minimum detectable absorptions measured in the high-frequency detection techniques are about 6 dB over this value.

## 6. - Conclusions.

This overview should have convinced that semiconductor diode lasers are important coherent sources for atomic and molecular spectroscopy. They can potentially cover the visible and the near infrared, with gaps which can be reduced with optical feedback techniques and with the development of new ma-

terials. Coverage can be extended to blue and near ultraviolet by frequency doubling. Optical feedback techniques can drastically reduce the linewidths to the level which allows also very narrow transitions to be investigated. Available powers start to be comparable with that of dye lasers and, however, the extremely reduced amplitude fluctuations are making high-sensitivity measurements possible even with low power. Further important possibilities are opened by the possibility of fast frequency control of the laser emission. Complex experimental schemes can be simplified by the use of semiconductor diode lasers and experiments which could not even be conceived before can now be performed.

Progresses in the fascinating world of laser spectroscopy have always depended on the development of new sources of radiation and we are sure that the availability of these new laser devices will allow new important chapters to be written.

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## OVERTONE MOLECULAR SPECTROSCOPY WITH DIODE LASERS

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### INTRODUCTION

All fundamental vibrational transitions of molecules are in the infrared region. At the same time only few molecules have electronic absorption bands in the visible. Such situation is not favourable for performing molecular spectroscopy measurements in the visible range which is very convenient for several reasons (tunable laser sources, sensitive detectors). On the other hand, due to transparency of molecular gases for visible radiation, the Sun light can easily penetrate through the atmosphere what is of the primary importance for all kind of life on the Earth.

The vibrational motion of a molecule besides of the fundamental frequencies, corresponding to different normal vibrational modes  $\nu_i$ , contains also overtone frequencies  $2\nu_i, 3\nu_i, \dots$ . Actually overtone frequencies are not exactly two or three times the frequency of the fundamental. They are slightly less. For convenience, however, it is conventionally written  $2\nu_i, 3\nu_i$ . In addition, in polyatomic molecules combination vibrations as  $\nu_i + \nu_k, \nu_i + 2\nu_k$  are also present.

Absorption lines corresponding to overtone and combination transitions are weak enough from the point of view of the atmosphere transparency, but at the same time they may be sufficiently strong for spectroscopic applications. However, such measurements require suitable laser sources and sensitive detection.

It has to be emphasized that the intensity of successive overtone and combination transitions decreases very rapidly with increasing number of vibrational modes and/or quanta involved. For instance, in the case of HCl molecule the intensity of the fourth overtone is more than four orders of magnitude weaker than that of the fundamental transition.

Recent progresses in semiconductor diode lasers<sup>1</sup> and in particular their output power, reliability, low cost, room temperature operation, and large spectral coverage have been of great importance for a continuously increasing use both in pure and applied spectroscopy. In particular, the extension of operation from the infrared to the visible, have really made accessible the wide and important field of absorption measure-

ments concerning the overtone molecular transitions. Various applications are possible with this sources in the field of the high resolution spectroscopy and high sensitivity detection. One of them is real time, non contact pollution measurement.

As it was already mentioned, moving from the infrared fundamental vibrational transitions to the overtones in the visible, absorption coefficients decrease several orders of magnitude. On the other hand, in comparison with laser sources operating in the infrared including lead salt diodes, semiconductor diode lasers in the visible (and in the near IR) offer the advantage of much simpler operation and much better amplitude stability. Moreover, remote sensing of atmospheric species could be more useful in the visible than in the infrared because of the reduced opacity of the atmosphere.

### TECHNIQUES DEVELOPED TO DETECT VERY WEAK ABSORPTIONS

In order to reach a satisfactory sensitivity in absorption spectroscopy several detection methods have been proposed and successfully applied. The simplest possible technique is the direct absorption measurement which consists on sweeping the frequency and detecting the signal against the constant background. The sensitivity of such a method is obviously very low. We can improve it by modulating the amplitude of the light source, but we are still limited by background contribution.

In order to understand the motivation for developing specialized techniques for direct absorption spectroscopy (i.e., the direct measurement of the optical attenuation of a light beam through an absorbing sample) let us consider some general requirements concerning the signal-to-noise ratio (SNR).

The signal is usually a detector photocurrent and may be written as  $S = kI$ , where  $k$  represents the net attenuation of the laser intensity  $I$  incident upon the absorption sample. Noise contributions may be separated into three terms:  $N_e$  - originating from the detection electronics, and then independent of  $I$ ,  $N_o$  - detector shot noise proportional to  $\sqrt{I}$  and  $N_s I$  - contribution of amplitude - fluctuation background proportional to  $I$ . The SNR may then be written<sup>2</sup>

$$SNR = \frac{kI}{[N_e^2 + (\beta\sqrt{I})^2 + (N_s I)^2]^{1/2}}$$

As the intensity  $I$  increases, the SNR increases proportionally to  $I$  until the  $I$ -dependent terms in the denominator exceed  $N_e$ . The SNR will then saturate at some value of  $I$  that depends on the relative values of  $\beta$  and  $N_s$ . As we can see the laser light amplitude modulation can not solve our problem, as this imparts a time dependence to  $I$ , which is common to both numerator and the denominator in the equation for SNR. However, by using frequency modulation, one may effectively modulate the signal  $kI$  without modulating the noise terms. In view of high sensitivity measurements this method has been developed in three different ways.

#### I. Wavelength Modulation (WM) Spectroscopy<sup>3,4,5</sup>

The single frequency laser is modulated at a relatively low frequency ( $10^3$  Hz), small compared to the width of the spectroscopic feature of interest, so that the absorption is probed simultaneously by a number of sidebands. WM is a sensitive form of derivative spectroscopy. Unlike direct absorption methods where signal is detected as a change against a constant background, in WM spectroscopy the signal arises from the differ-

ence in the absorptions of different sidebands. Therefore, the sensitivity and spectral resolution are greatly enhanced, provided that the laser shows no or very little variation of intensity with wavelength and responsivity of the detector is independent of  $\lambda$ . Consider the current  $C$  from the detector,  $C = IGT$ , where  $I$  is the laser light intensity,  $T$  is the transmissivity of the sample, and  $G$  is responsivity of the detector. We can then write for the detector signal

$$\frac{1}{C} \frac{dC}{d\lambda} \Delta\lambda = \frac{1}{I} \frac{dI}{d\lambda} \Delta\lambda + \frac{1}{G} \frac{dG}{d\lambda} \Delta\lambda + \frac{1}{T} \frac{dT}{d\lambda} \Delta\lambda$$

where  $\Delta\lambda$  is the wavelength modulation depth. Only the last term in the above expression should give a contribution to the signal. Unfortunately  $dI/d\lambda$  is not negligible in the case of diode lasers and gives an unwanted contribution to the noise level.

## II. Frequency Modulation (FM) Spectroscopy<sup>6,7</sup>

FM spectroscopy is a kind of extension of WM spectroscopy to much higher frequencies ( $\approx 10^6$  Hz). Frequency modulation produces sidebands which are widely spaced in frequency so that the spectral feature of interest can be probed by only one sideband at a time. Viewed in frequency space, the spectral distribution of the modulated laser field consists of the strong carrier at  $\omega_c$  and two sidebands of the same amplitude but  $180^\circ$  out of phase, displaced by the angular modulation frequency  $\omega_m$  from the carrier. When there is no absorption present, the beat signal at  $\omega_m$  between the carrier and the upper sideband exactly cancels with the beat signal between the carrier and the lower sideband. If, however, the laser frequency is tuned over an absorption, so that one of the sidebands is absorbed, the imperfect cancellation of two beats produces a FM signal.

If the modulation frequency does not exceed the absorption linewidth, the technique is usually called High Wavelength Modulation (HWM).

## III. Two-Tone Frequency Modulation (TTFM)<sup>8,9</sup>

If we wish to investigate broad spectral features, such as absorption lines broadened by atmospheric pressure to 2–3 GHz, we must have correspondingly high modulation frequencies. The TTFM technique was applied in order to reduce the detection bandwidth requirement since detectors with bandwidth in GHz range are not easily available.

In the TTFM, the laser is modulated simultaneously at two distinct but closely spaced angular frequencies  $\omega_1 = \omega_m + \Omega/2$  and  $\omega_2 = \omega_m - \Omega/2$  ( $\Omega/\omega_m < 10^{-3}$ ). The TTFM absorption signal arises from the difference between the absorption of the carrier and the sum of the absorptions of the sidebands. It is thus approximately analogous to a second derivative signal whereas single-tone FM is analogous to a first derivative signal. In both cases, however, there is no background signal if there is no absorption. The most attractive feature of TTFM resides in the fact that the signal is detected at  $\Omega$  instead of  $\omega_m$ .

Unfortunately pure frequency modulation is rarely achieved. There is always some residual amplitude modulation (RAM) present, especially in diode lasers where any change in the injection current changes not only the frequency of the laser but also its output power. It is evident that RAM is undesirable feature in all kinds of frequency

modulation spectroscopy. It gives rise to a background signal with accompanying noise, even when there is no absorption present, and thus limits the detection sensitivity.

It is worthwhile to mention here that a wide variety of laser techniques allowing high sensitivity and high resolution have been proposed and successively applied. Let us give few examples as laser-induced fluorescence (LIF)<sup>10</sup>, optoacoustic spectroscopy (OA)<sup>11</sup>, optogalvanic spectroscopy (OG)<sup>11</sup>, resonance ionisation spectroscopy (RIS)<sup>12</sup>, and laser-intracavity absorption<sup>13</sup>. With the exception of direct absorption methods described above, all these techniques monitor some indirect effects of the optical absorption.

Any qualitative comparison of each of these methods would be extremely difficult since their applicability depends essentially on specific aims and experimental conditions of the desired measurements. However, two important advantages of the extracavity direct absorption should be emphasized: simple calibration procedure and remote-sensing possibilities. All of them are very important from the point of view of environmental studies.

## OVERTONE SPECTROSCOPY

Several overtone transitions of various molecules have been already studied using various light sources and different detection techniques. Table 1 contains selected information concerning molecules with overtone or combination absorption lines in two spectral ranges (630–690 and 750–900 nm) covered by diode lasers. As one can see only in a few recent works diode lasers have been applied. Other data were obtained with dye lasers or even with methods of pre-laser spectroscopy. Some overtone transitions in the selected ranges refer to reported numerical values and can be considered as potential candidates for future experimental research.

Table 1. Overtone and combination transitions in spectral ranges covered by diode lasers.

| Molecule | absorp. line (nm) | transition               | laser and det. tech. | sensitivity other inf. | reference |
|----------|-------------------|--------------------------|----------------------|------------------------|-----------|
| $H_2O$   | 818               | $2\nu_1 + \nu_2 + \nu_3$ | Diode                | min. det. abs.         | 14        |
|          |                   |                          | WM                   | $3.1 \times 10^{-4}$   |           |
|          |                   |                          | FM                   | $2.8 \times 10^{-4}$   |           |
|          |                   |                          | TTFM                 | $2.4 \times 10^{-6}$   |           |
|          | 789               | $\nu_2 + 3\nu_3$         | Diode                |                        | 15        |
|          |                   |                          | FM                   | $3 \times 10^{-4}$     |           |
|          | 818-831           | $2\nu_1 + \nu_2 + \nu_3$ | Diode                |                        | 15        |
|          |                   |                          | FM                   |                        |           |
| HI       | 820               | $\nu = 6 \leftarrow 0$   | Diode                |                        | 16        |

Table 1. (cont'd)

|          |             |                             |              |                      |    |
|----------|-------------|-----------------------------|--------------|----------------------|----|
| $C_2H_2$ | 789         | $\nu_1 + 3\nu_3$            | Diode        | min.det.conc.        | 17 |
|          | 640         | $5\nu_3$                    | WM           | 0.2ppm per km        | 18 |
|          | 670         | $\nu_1 + \nu_2 + \nu_3$     | Dye          | intra-cav.           | 19 |
|          | 848         | $2\nu_1 + \nu_2 + \nu_3$    | OA           |                      |    |
| $CH_4$   | 886         | $3\nu_1 + \nu_3$            | Diode        | min.det.abs.         | 20 |
|          |             |                             | LWM          | $4.5 \times 10^{-7}$ |    |
|          |             |                             | HWM          | $9.7 \times 10^{-8}$ |    |
|          |             |                             | TTFM         | $6.4 \times 10^{-8}$ |    |
|          | 860         | $2\nu_1 + 2\nu_3$           | Diode        | $3 \times 10^{-5}$   | 21 |
|          | 682         | $2\nu_1 + 3\nu_3$           |              | White cell           | 22 |
|          | 782         | $2\nu_1 + \nu_2 + 2\nu_3$   |              |                      | 23 |
| $C_2H_6$ | 741         | $5 \leftarrow 0$<br>CH str. | OA           |                      | 24 |
| $NH_3$   | 790         | $4\nu_1$                    | Diode        | min.det.abs.         | 25 |
|          |             | $2\nu_1 + 2\nu_3$           | FM           | $3 \times 10^{-4}$   |    |
|          | 792         | $4\nu_1$                    | Diode        |                      | 26 |
|          | 647         | $5\nu_1$                    | Dye          | intra-cav            | 27 |
|          | 647         | $4\nu_1 + \nu_3$            |              |                      |    |
|          | 647         | $4\nu_1 + \nu_3$            | photoplate   |                      | 28 |
| $C_6H_6$ | 865-<br>877 | $4 \leftarrow 0$<br>CH str. | opto-thermal |                      | 29 |
| $H_2O_2$ | 793         | $5\nu_{OH} + \nu_{OO}$      | Dye          |                      | 30 |
| $HCl$    | 640         | $\nu = 6 \leftarrow 0$      | Dye          | intra-cav            | 31 |
|          | 750         | $\nu = 5 \leftarrow 0$      | OA           |                      |    |
| $CO_2$   | 869         | $5\nu_3$                    |              |                      | 32 |
|          | 789         | $2\nu_2 + 5\nu_3$           |              |                      |    |
|          | 783         | $\nu_1 + 5\nu_3$            |              |                      |    |
| $C_3H_4$ | 787         | $4\nu_1$                    | OA           |                      | 33 |
|          | 737         | $5 \leftarrow 0$            |              |                      |    |
| $HCN$    | 790         | $4\nu_3$                    |              |                      | 34 |
| $CO$     | 804         | $\nu = 6 \leftarrow 0$      |              |                      | 35 |
| $NO$     | 800         | $\nu = 7 \leftarrow 0$      |              |                      | 36 |
| $HF$     | 877         | $\nu = 3 \leftarrow 0$      |              |                      | 37 |
|          | 675         | $\nu = 4 \leftarrow 0$      |              |                      | 38 |

Results obtained for  $H_2O$  and  $CH_4$  were used for comparing the sensitivity of different frequency modulation techniques. In the first experimental work<sup>14</sup> dedicated to  $H_2O$  the authors found that the difference in sensitivity between FM and TTFM was about two orders of magnitude. In other work<sup>20</sup> the minimum detectable absorption was measured using methane - the gas of obvious interest because of environmental applications. A component of a combination band at 886nm has been chosen as a test transition, mainly because of its strength. Two recordings are shown in Figure 1:a) pure absorption signal at pressure of 100 Torr corresponding

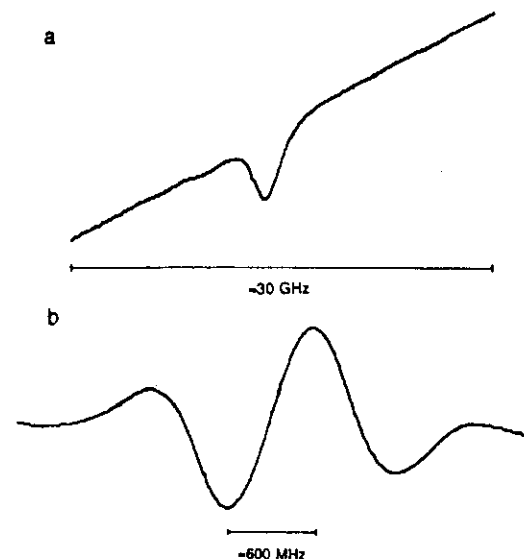


Figure 1.a) Pure absorption signal on the third overtone of methane at 100 Torr in 1.5m pathway. b) Derivative lineshape of a two-tone recording at 500 mTorr.

to an absorption of  $3.6 \times 10^{-2}$  (1.5m pathway, 100Hz detection bandwidth,  $S/N$  ratio of about 20); b) derivative lineshape of a two-tone signal at pressure of 500 mTorr corresponding to an absorption of  $1.8 \times 10^{-4}$  (1.5m pathway, 1Hz detection bandwidth) with a measured  $S/N$  ratio of a few thousands.

Minimum detectable absorption for the HWM and TTFM techniques turned out to be of the same order of magnitude, while for the value referred to LW modulation it is about one order of magnitude less, as indicated in Table 1. In the last case the detection limit value is determined by the laser amplitude  $1/f$  excess noise, while in the first two cases this noise contribution is negligible with respect to other ones as detector induce shot noise, thermal and RAM noises. The detection limits well agree with the calculated "quantum limited" values based on measured laser power, modulation index, noise of the electronic components, and other parameters of the apparatus.

In most of the measurements with diode lasers the frequency modulation was produced by modulating the injection current. In the case of  $C_2H_2$  the diode laser was

used in an extended cavity configuration. Low frequency modulation could be then produced by both changing the injection laser current and/or varying the external cavity length (by means of a PZT). With the extended cavity configuration we have measured a reduction of the linewidth (to less than 1 MHz) and better short term stability (100 ms). Another advantage of such a configuration is that wavelength modulation by varying the cavity length leads to the lower level RAM noise than direct injection current modulation.

The experimental apparatus used for  $C_2H_2$  overtone absorption measurements<sup>17</sup> is schematically shown in Figure 2. The extended cavity configuration is outlined in Figure 3. The first order beam diffracted from a grating was fed-back into the laser diode. The tilting of the grating allows to select one of the longitudinal modes within tuning range of about 20nm.

Molecular transitions could be observed either in pure absorption or in a phase sensitive detection scheme. In the latter case, the lock-in derivative signals were

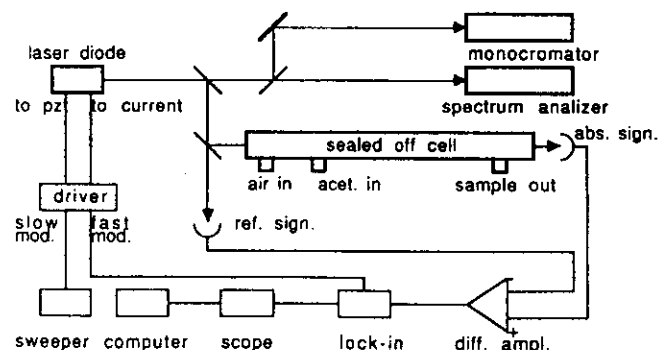


Figure 2. Schematic of the experimental apparatus for measuring the overtone absorption in acetylene.

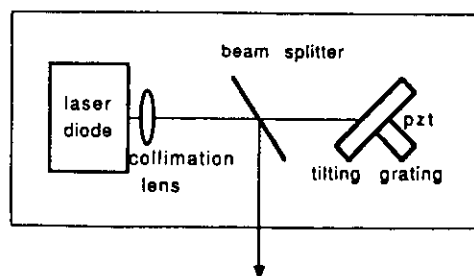


Figure 3. Schematic of the extended cavity diode laser configuration using a grating for wavelength control.

recorded by introducing a kHz modulation on the cavity length. The derivative signal for the  $P(11)$  component of the  $12646.966\text{cm}^{-1}$  transition is shown in Figure 4 at two different gas pressures. We measured a  $\text{aS/N}$  value of a few thousands and an absorption of about 6% at 30 Torr (trace a), and  $S/N = 6$  at 36 mTorr (b). The latter measurement gives a minimum relative absorption of  $10^{-4}$  which is comparable with that obtained by means of TTFM<sup>14</sup>. This result is also consistent with the recent report<sup>34</sup> where WM and TTFM techniques have been compared for lead salt diodes.

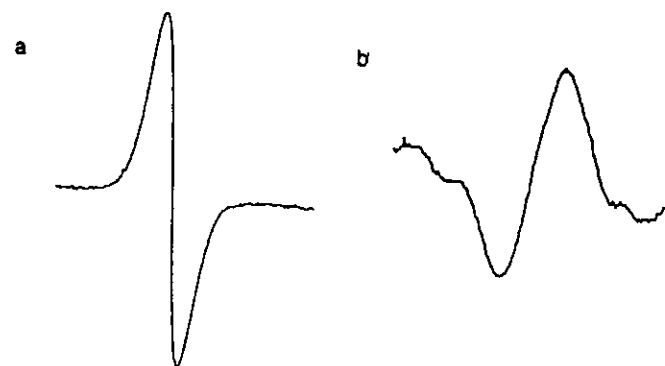


Figure 4. Derivative signal for the  $P(11)$  component at 10 Torr (a) and 36 mTorr (b).

Regarding possible environmental trace gases detection we have also extracted the minimum acetylene pressure in presence of air at atmospheric pressure still keeping a signal with a  $S/N$  ratio of a few units. We are able to detect 0.1 ppm per km of acetylene in air.

The narrow linewidth emission obtained with the extended cavity configuration has also permitted to perform the lineshape analysis measuring self and air pressure broadening parameter for two different components of the combination band. The Pade method<sup>35</sup> has been used to extract the collisional Lorentzian contribution from the Voigt profile and thus the pressure broadening parameter. Air broadening measurements have been performed on the  $P(11)$  and  $R(5)$  components. Our results are reported in Table 2 together with earlier measurements of other components of  $C_2H_2$  as well as recent measurements of pressure broadening for  $H_2O$ ,  $CH_4$  and  $NH_3$  using diode lasers. As one can see the air broadening coefficients for overtone transitions are very close to that for the fundamental transition.

For many practical applications it is requested to monitor gas traces in open-path configuration in the atmosphere. For this reason the knowledge of air broadening parameters can be useful. However, some problems arise for such experiments at atmospheric pressures. Widths of absorption lines due to pressure broadening are of the order of GHz what leads to the substantial decrease of the detection sensitivity. Moreover, the overlapping of different lines may also occur and it has to be taken into consideration.

Table 2. Pressure broadening parameters.

| Molecule | Transition                    | $\gamma$ (MHz/Torr) |      | Refer. |
|----------|-------------------------------|---------------------|------|--------|
|          |                               | air                 | self |        |
| $C_2H_2$ | $789(\nu_1 + 3\nu_3)$         |                     |      | 20     |
|          | P(11)                         | 8.0                 | 11.0 |        |
|          | R(5)                          | 9.0                 |      | 36     |
|          | $862(\nu_2 + 3\nu_3)$         |                     | 12.3 |        |
|          | R(7)                          |                     |      | 36     |
|          | $849(2\nu_1 + \nu_2 + \nu_3)$ |                     | 9.9  |        |
| $H_2O$   | R(7)                          |                     |      |        |
|          | $\nu_5$                       | 7.0                 |      | 37     |
|          | R(5)                          | 6.2                 |      |        |
| $H_2O$   | $822(2\nu_1 + \nu_2 + \nu_3)$ | 5.9                 |      | 15     |
| $CH_4$   | $860(2\nu_1 + 2\nu_3)$        | 3.2                 | 3.7  | 21     |
| $NH_3$   | $790(2\nu_1 + 2\nu_3)$        |                     | 21.7 | 21     |
|          | $792(4\nu_1)$                 |                     | 27.2 | 21     |

As for the sensitivity of diode laser systems destined to monitoring the atmosphere one should think about its improvement. The stabilization of the diode lasers operating at wavelength below  $1\mu m$  could also be important for use in conjunction with efficient pulsed solid state laser (such as Titan Sapphire) operating in the same region. This can be interesting in the frame of LIDAR<sup>36,39</sup> measurements, where amplification or injection seeding control could make possible to perform detection at larger distances with an improved sensitivity and an easier spectral control.

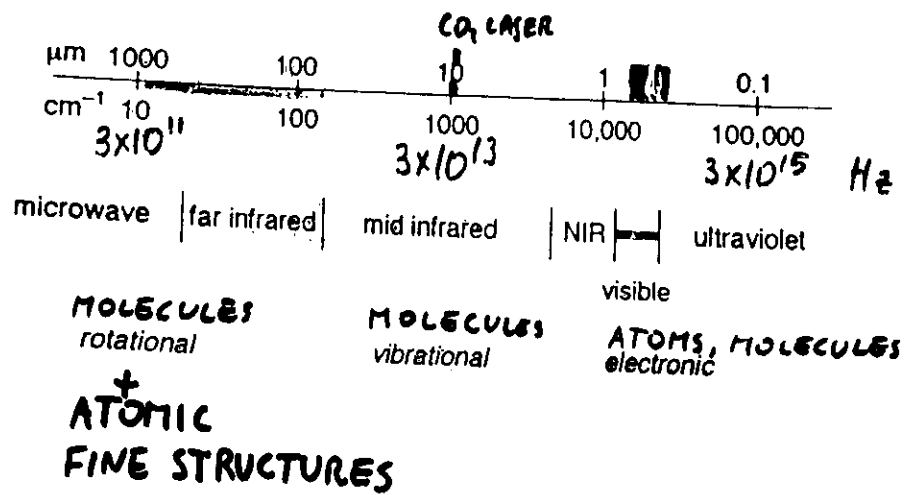
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## The electromagnetic spectrum



frequency measurements

wavelength measurements

$$c = \lambda \nu$$

$$p \sim \frac{h}{a}$$

$$E_e \sim \frac{p^2}{2m} \sim \frac{h^2}{2m a^2} \sim \frac{1}{a^2} \text{ a few eV visible } 10^3 \text{ eV}$$

$$\omega_e = \sqrt{\frac{k}{m}} \quad \hbar \omega_e \approx \sqrt{\frac{m}{M}}$$

$$\omega_N = \sqrt{\frac{k}{M}} \quad \hbar \omega_N \approx \sqrt{\frac{m}{M}}$$

$$E_N \approx \sqrt{\frac{m}{M}} E_e \quad 10^{-2} \quad 0.5 \mu \text{m} \quad 50 \mu \text{m}$$

$$I = \frac{M a^2}{2} \frac{\hbar^2}{2I} l(l+1)$$

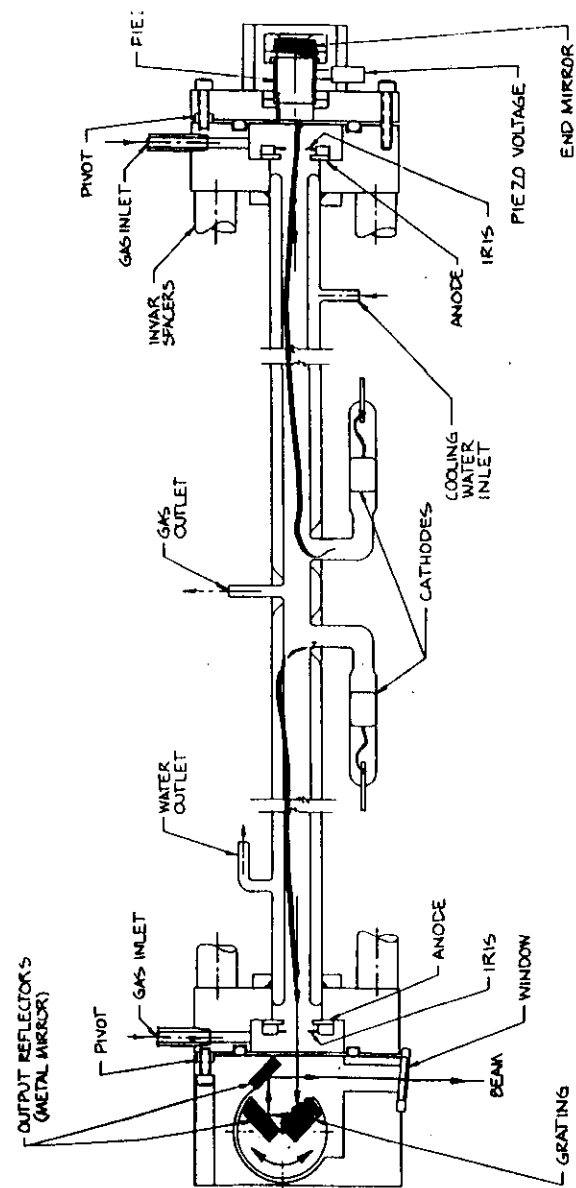
$$E_R \approx \frac{\hbar^2}{M a^2} \approx \frac{m}{M} E_e \quad 10^{-3} - 10^{-5}$$

$B J(T+1)$

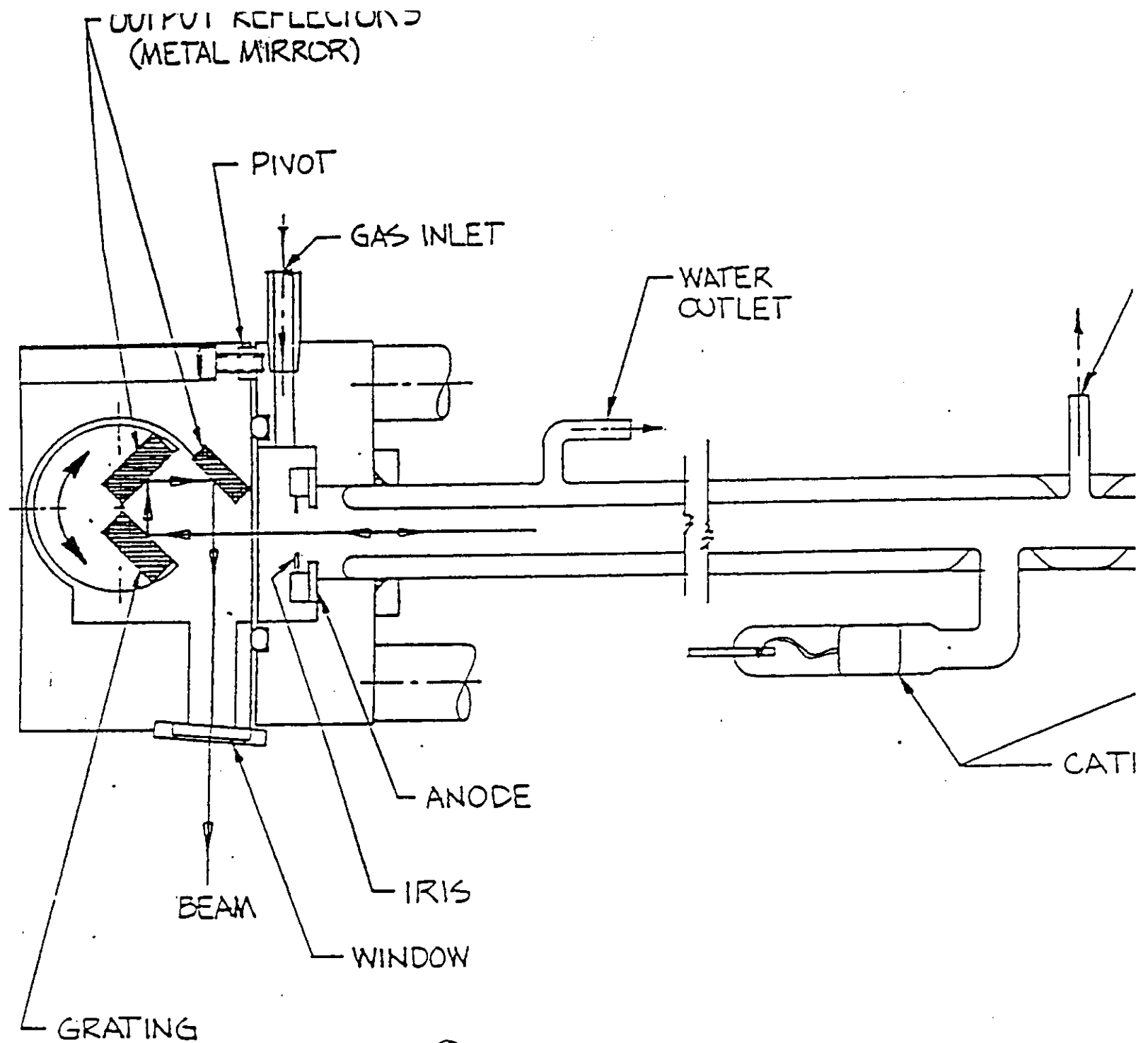
↑

$J \rightarrow J+1$

$2 B J$

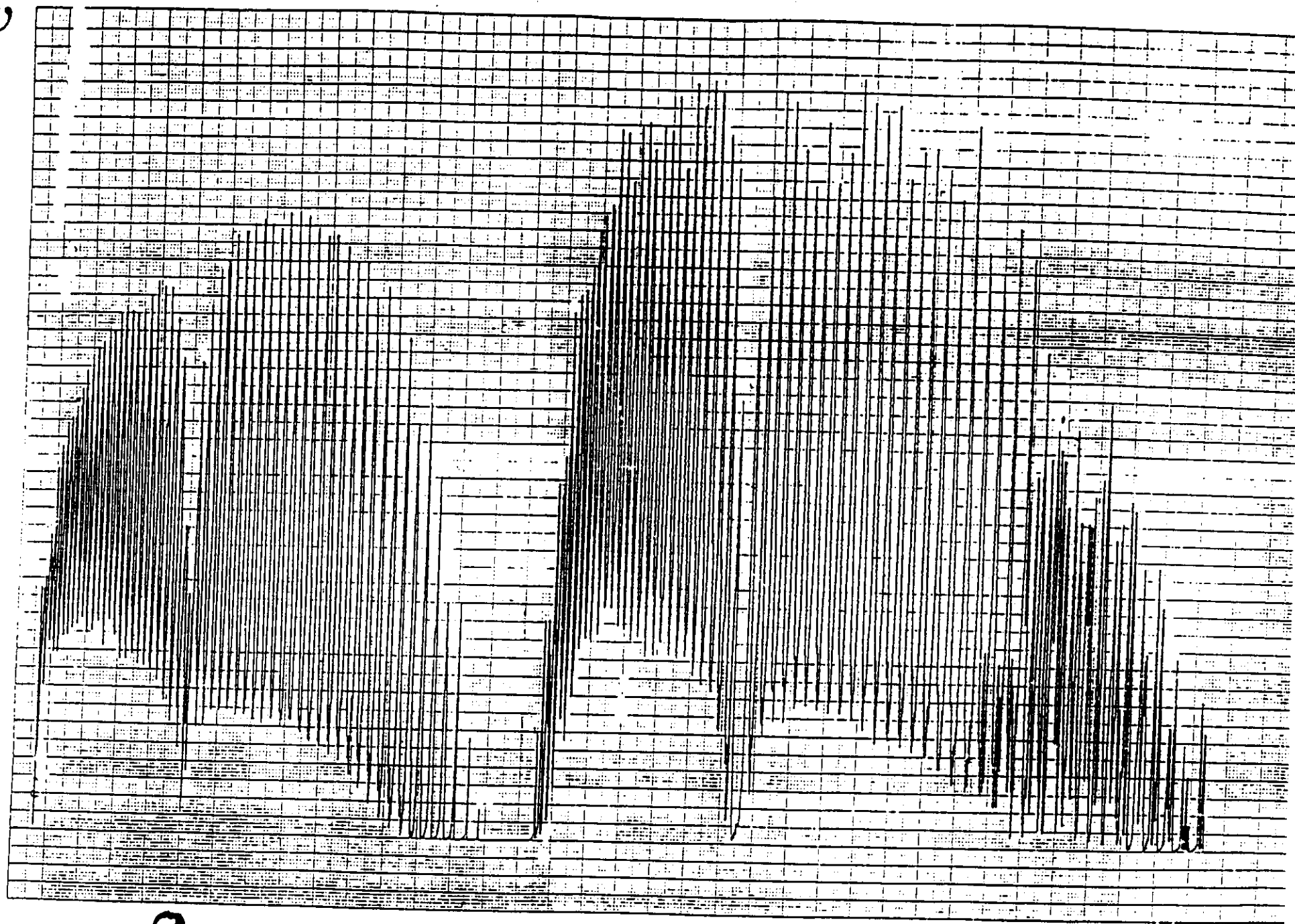






$w$   
↑

(0-9w)



9  $\mu\text{m}$

(3-11  $\mu\text{m}$ )

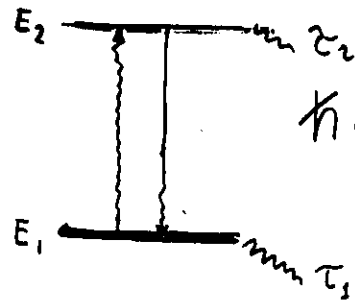
(25)

10  $\mu\text{m}$

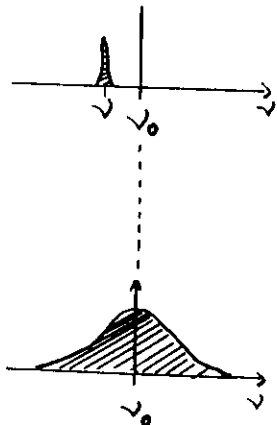
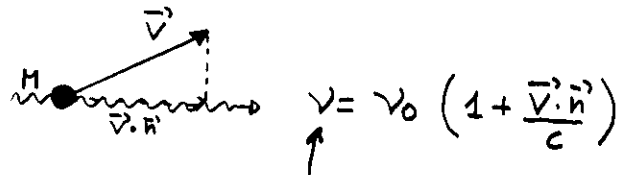
→  $\lambda$

# ALLARGAMENTO DOPPLER DI RIGHE SPETTRALI (1<sup>o</sup> ordine)

DOPPLER BROAD.  
(1<sup>st</sup> ORDER)



$$\hbar \omega_0 = E_2 - E_1$$



NOTE!  $\Delta \nu_D$  independent  
from  $\nu$

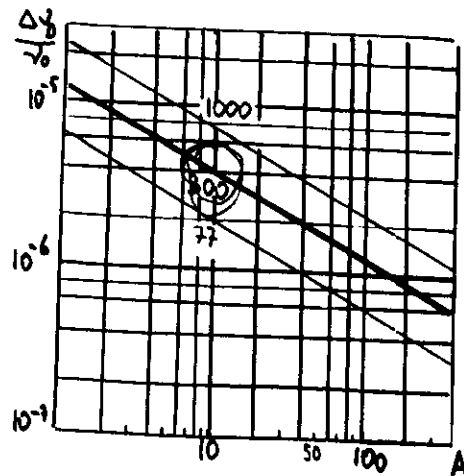
At room temp  
 $\frac{\Delta \nu}{\nu} \sim 2 \cdot 10^{-6}$

$$I = I_0 \exp \left[ -\frac{hc^2}{2KT} \left( \frac{\nu - \nu_0}{\nu_0} \right)^2 \right]$$

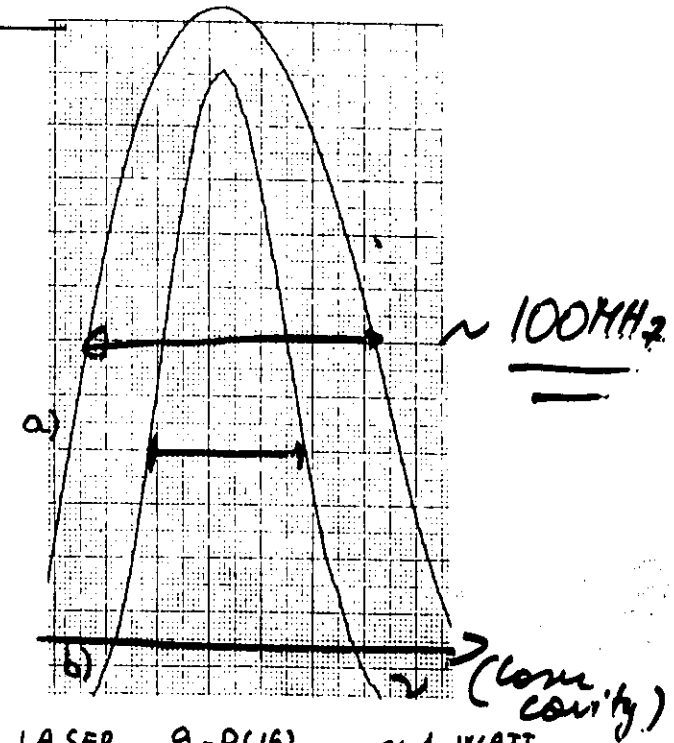
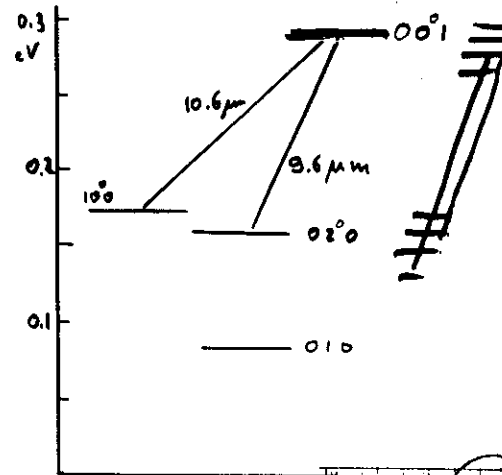
$$\Delta \nu_D = 7.163 \cdot 10^{-7} \omega_0 \sqrt{\frac{T}{A}}$$

$$10^8 - 10^{10} \text{ Hz}$$

$$\Delta \nu_D \approx 10^2 \Delta \nu_{\text{NAT}} \text{ in the VIS}$$



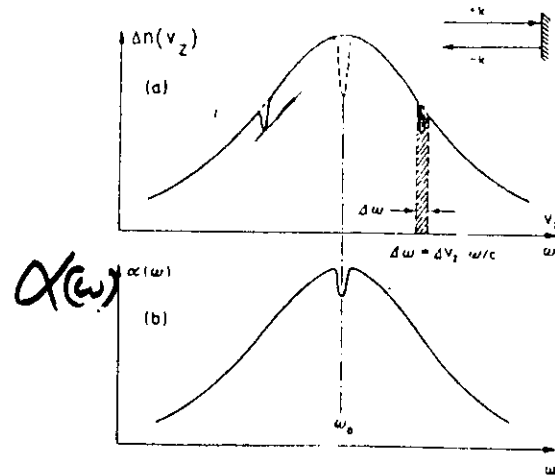
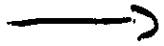
## TRANSIZIONI TRA STATI ECCITATI



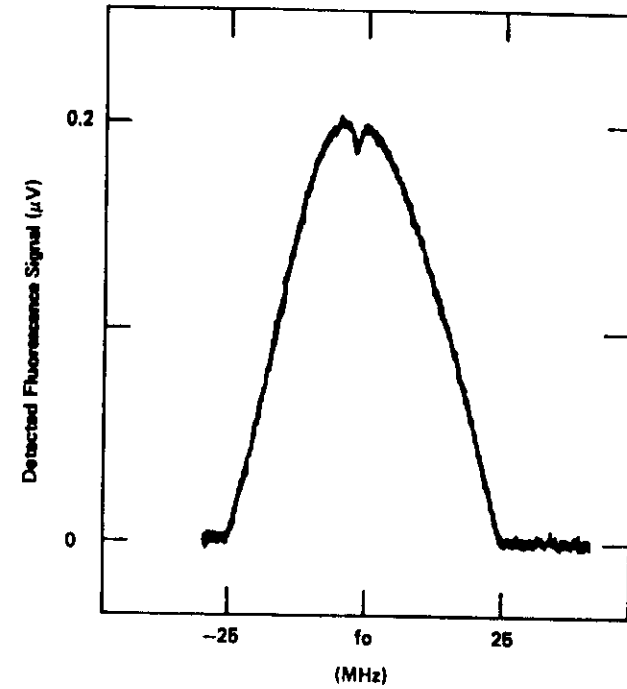
a) CO<sub>2</sub> LASER 9-P(16) ~ 1 WATT

b) SEGNALE DA DB CO<sub>2</sub> (~ 1/2 Torr)

# SUBDOPPLER ~~MA~~ SPECTROSCOPY

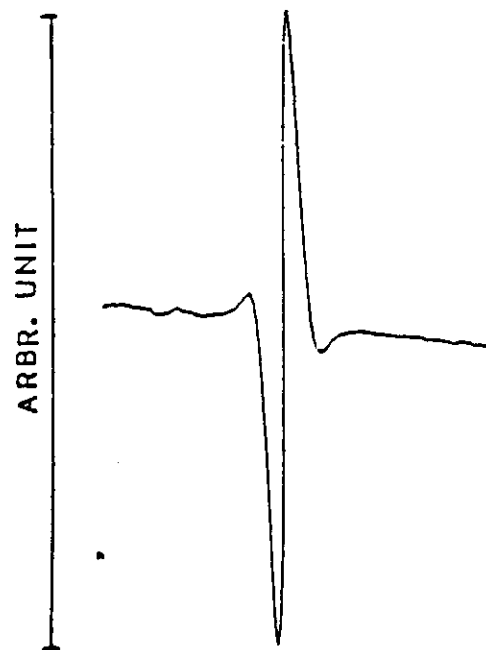


2.1



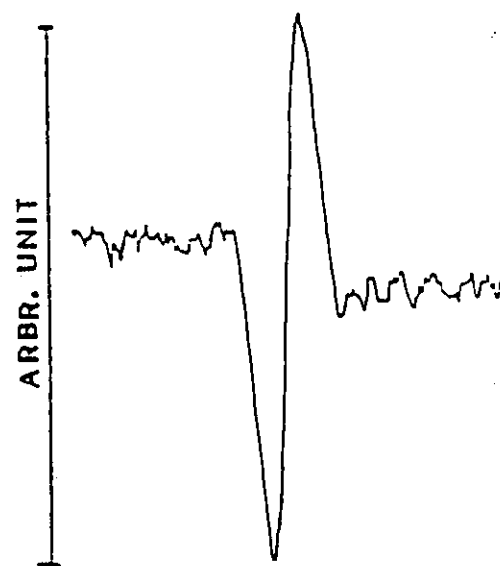
FLUOR.

OPTOGAL.



a)

11  
1 MHz



b)

11  
3 MHz

~ 3<sup>rd</sup> Der.

# NORMAL CO<sub>2</sub> LASERS

10  $\mu$ m

9  $\mu$ m

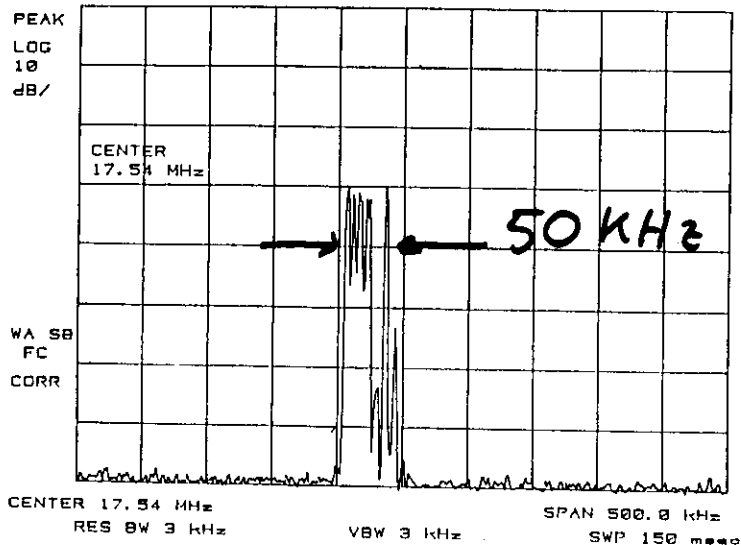
|       |      |           |          |      |             |       |      |           |           |      |
|-------|------|-----------|----------|------|-------------|-------|------|-----------|-----------|------|
| P(60) | 2707 | 7607.5077 | 903.2117 | 6484 | $\pm 3$ kHz | P(60) | 3014 | 3456.0742 | 1005.4774 | 6515 |
| P(58) | 2714 | 6404.4578 | 905.5065 | 8408 |             | P(58) | 3021 | 2223.6949 | 1007.7713 | 0607 |
| P(56) | 2721 | 4396.1809 | 907.7745 | 4384 |             | P(56) | 3028 | 0322.1201 | 1010.0428 | 2503 |
| P(54) | 2728 | 1588.8741 | 910.0158 | 5084 |             | P(54) | 3034 | 7743.7465 | 1012.2917 | 6841 |
| P(52) | 2734 | 7988.4259 | 912.2307 | 0148 |             | P(52) | 3041 | 4481.1364 | 1014.5178 | 8812 |
| P(50) | 2741 | 3600.4235 | 914.4192 | 8214 |             | P(50) | 3048 | 0527.0251 | 1016.7209 | 4183 |
| P(48) | 2747 | 8430.1601 | 916.5817 | 6938 |             | P(48) | 3054 | 5874.3277 | 1018.9006 | 9322 |
| P(46) | 2754 | 2482.6413 | 918.7183 | 3017 |             | P(46) | 3061 | 0516.1462 | 1021.0569 | 1219 |
| P(44) | 2760 | 5762.5914 | 920.8291 | 2210 |             | P(44) | 3067 | 4445.7759 | 1023.1893 | 7509 |
| P(42) | 2766 | 8274.4599 | 922.9142 | 9359 |             | P(42) | 3073 | 7656.7119 | 1025.2978 | 6496 |
| P(40) | 2773 | 0022.4271 | 924.9739 | 8407 |             | P(40) | 3080 | 0142.6555 | 1027.3821 | 7169 |
| P(38) | 2779 | 1010.4094 | 927.0083 | 2419 |             | P(38) | 3086 | 1897.5199 | 1029.4420 | 9223 |
| P(36) | 2785 | 1242.0651 | 929.0174 | 3596 |             | P(36) | 3092 | 2915.4360 | 1031.4774 | 3083 |
| P(34) | 2791 | 0720.7986 | 931.0014 | 3295 |             | P(34) | 3098 | 3190.7583 | 1033.4879 | 9917 |
| P(32) | 2796 | 9449.7656 | 932.9604 | 2043 |             | P(32) | 3104 | 2718.0700 | 1035.4736 | 1655 |
| P(30) | 2802 | 7431.8776 | 934.8944 | 9550 |             | P(30) | 3110 | 1492.1877 | 1037.4341 | 1009 |
| P(28) | 2808 | 4669.8055 | 936.8037 | 4726 |             | P(28) | 3115 | 9508.1671 | 1039.3693 | 1486 |
| P(26) | 2814 | 1165.9839 | 938.6882 | 5692 |             | P(26) | 3121 | 6761.3064 | 1041.2790 | 7402 |
| P(24) | 2819 | 6922.6147 | 940.5480 | 9793 |             | P(24) | 3127 | 3247.1518 | 1043.1632 | 3901 |
| P(22) | 2825 | 1941.6703 | 942.3833 | 3608 |             | P(22) | 3132 | 8961.5006 | 1045.0216 | 6964 |
| P(20) | 2830 | 6224.8967 | 944.1940 | 2961 |             | P(20) | 3138 | 3900.4054 | 1046.8542 | 3425 |
| P(18) | 2835 | 9773.8165 | 945.9802 | 2931 |             | P(18) | 3143 | 8060.1774 | 1048.6608 | 0978 |
| P(16) | 2841 | 2589.7314 | 947.7412 | 7060 |             | P(16) | 3149 | 1437.3897 | 1050.4413 | 8194 |
| P(14) | 2846 | 4673.7246 | 949.4793 | 1361 |             | P(14) | 3154 | 4028.8804 | 1052.1955 | 4524 |
| P(12) | 2851 | 6026.6628 | 951.1922 | 6324 |             | P(12) | 3159 | 5831.7547 | 1053.9235 | 0313 |
| P(10) | 2856 | 6649.1983 | 952.8808 | 4927 |             | P(10) | 3164 | 6843.3878 | 1055.6250 | 6805 |
| P(8)  | 2861 | 6541.7701 | 954.5450 | 8632 |             | P(8)  | 3169 | 7061.4264 | 1057.3001 | 6151 |
| P(6)  | 2866 | 5704.6061 | 956.1849 | 8202 |             | P(6)  | 3174 | 6483.7910 | 1058.9487 | 1415 |
| P(4)  | 2871 | 4137.7235 | 957.8005 | 3691 |             | P(4)  | 3179 | 5108.6771 | 1060.5706 | 6576 |
| P(2)  | 2876 | 1840.9300 | 959.3917 | 4460 |             | P(2)  | 3184 | 2934.5560 | 1062.1659 | 6536 |
| V(0)  | 2880 | 8813.8246 | 960.9585 | 9171 |             | V(0)  | 3188 | 9960.1764 | 1063.7345 | 7121 |
| R(0)  | 2883 | 2026.2225 | 961.7328 | 7396 |             | R(0)  | 3191 | 3172.5743 | 1064.5088 | 5347 |
| R(2)  | 2887 | 7902.4412 | 963.2631 | 3990 |             | R(2)  | 3195 | 8996.0672 | 1066.0373 | 6066 |
| R(4)  | 2892 | 3046.4336 | 964.7689 | 8140 |             | R(4)  | 3200 | 4017.3872 | 1067.5391 | 1025 |
| R(6)  | 2896 | 7457.0695 | 966.2503 | 6076 |             | R(6)  | 3204 | 4236.2544 | 1069.0140 | 9289 |
| R(8)  | 2901 | 1133.0097 | 967.7072 | 3331 |             | R(8)  | 3209 | 1652.6660 | 1070.4623 | 0849 |
| R(10) | 2905 | 4072.7058 | 969.1395 | 4739 |             | R(10) | 3213 | 4266.8953 | 1071.8837 | 6618 |
| R(12) | 2909 | 6274.3988 | 970.5472 | 4435 |             | R(12) | 3217 | 6079.4907 | 1073.2784 | 8423 |
| R(14) | 2913 | 7736.1185 | 971.9302 | 5845 |             | R(14) | 3221 | 7091.2743 | 1074.6464 | 9008 |
| R(16) | 2917 | 8455.6817 | 973.2885 | 1688 |             | R(16) | 3225 | 7303.3400 | 1075.9878 | 2021 |
| R(18) | 2921 | 8430.6909 | 974.6219 | 3965 |             | R(18) | 3229 | 6717.0518 | 1077.3025 | 2013 |
| R(20) | 2925 | 7658.5324 | 975.9304 | 3960 |             | R(20) | 3233 | 5334.0411 | 1078.5906 | 4423 |
| R(22) | 2929 | 6136.3740 | 977.2139 | 2224 |             | R(22) | 3237 | 3156.2043 | 1079.8522 | 5580 |
| R(24) | 2933 | 3861.1629 | 978.4722 | 8575 |             | R(24) | 3241 | 0185.7000 | 1081.0874 | 2682 |
| R(26) | 2937 | 0829.6231 | 979.7054 | 2084 |             | R(26) | 3244 | 6424.9456 | 1082.2562 | 3794 |
| R(28) | 2940 | 7038.2525 | 980.9132 | 1071 |             | R(28) | 3248 | 1876.6140 | 1083.4787 | 7831 |
| R(30) | 2944 | 2483.3197 | 982.0955 | 3089 |             | R(30) | 3251 | 6543.6298 | 1084.6351 | 4549 |
| R(32) | 2947 | 7160.8609 | 983.2522 | 4916 |             | R(32) | 3255 | 0429.1653 | 1085.7654 | 4528 |
| R(34) | 2951 | 1066.6762 | 984.3832 | 2542 |             | R(34) | 3258 | 3536.6360 | 1086.8697 | 9163 |
| R(36) | 2954 | 4196.3256 | 985.4883 | 1157 |             | R(36) | 3261 | 5869.6965 | 1087.9483 | 0644 |
| R(38) | 2957 | 6545.1250 | 986.5673 | 5137 |             | R(38) | 3264 | 7432.2354 | 1089.0011 | 1941 |
| R(40) | 2960 | 8106.1417 | 987.6201 | 8028 |             | R(40) | 3267 | 8228.3702 | 1090.0283 | 6790 |
| R(42) | 2963 | 8880.1900 | 988.6466 | 2533 |             | R(42) | 3270 | 8262.4421 | 1091.0301 | 9670 |
| R(44) | 2966 | 8855.8259 | 989.6465 | 0491 |             | R(44) | 3273 | 7539.0104 | 1092.0067 | 5790 |
| R(46) | 2969 | 8029.3420 | 990.6196 | 2866 |             | R(46) | 3276 | 6062.8469 | 1092.9582 | 1067 |
| R(48) | 2972 | 6394.7621 | 991.5657 | 9723 |             | R(48) | 3279 | 3838.9297 | 1093.8847 | 2107 |
| R(50) | 2975 | 3945.8353 | 992.4848 | 0211 |             | R(50) | 3282 | 0872.4368 | 1094.7864 | 6180 |
| R(52) | 2978 | 0676.0297 | 993.3764 | 2542 |             | R(52) | 3284 | 7168.7402 | 1095.6636 | 1206 |
| R(54) | 2980 | 6578.5263 | 994.2404 | 3971 |             | R(54) | 3287 | 2733.3987 | 1096.5163 | 5728 |
| R(56) | 2983 | 1646.2123 | 995.0766 | 0771 |             | R(56) | 3289 | 7572.1515 | 1097.3448 | 8889 |
| R(58) | 2985 | 5011.6777 | 995.8846 | 8112 |             | R(58) | 3292 | 1690.9198 | 1098.1443 | 0411 |

# FREE RUNNING

200000 Hz  
Line 10030  
3000 Hz  
(0.000000)  
200000 Hz  
1000000 Hz

1/2

REF -4.0 dBm ATTEN 10 dB

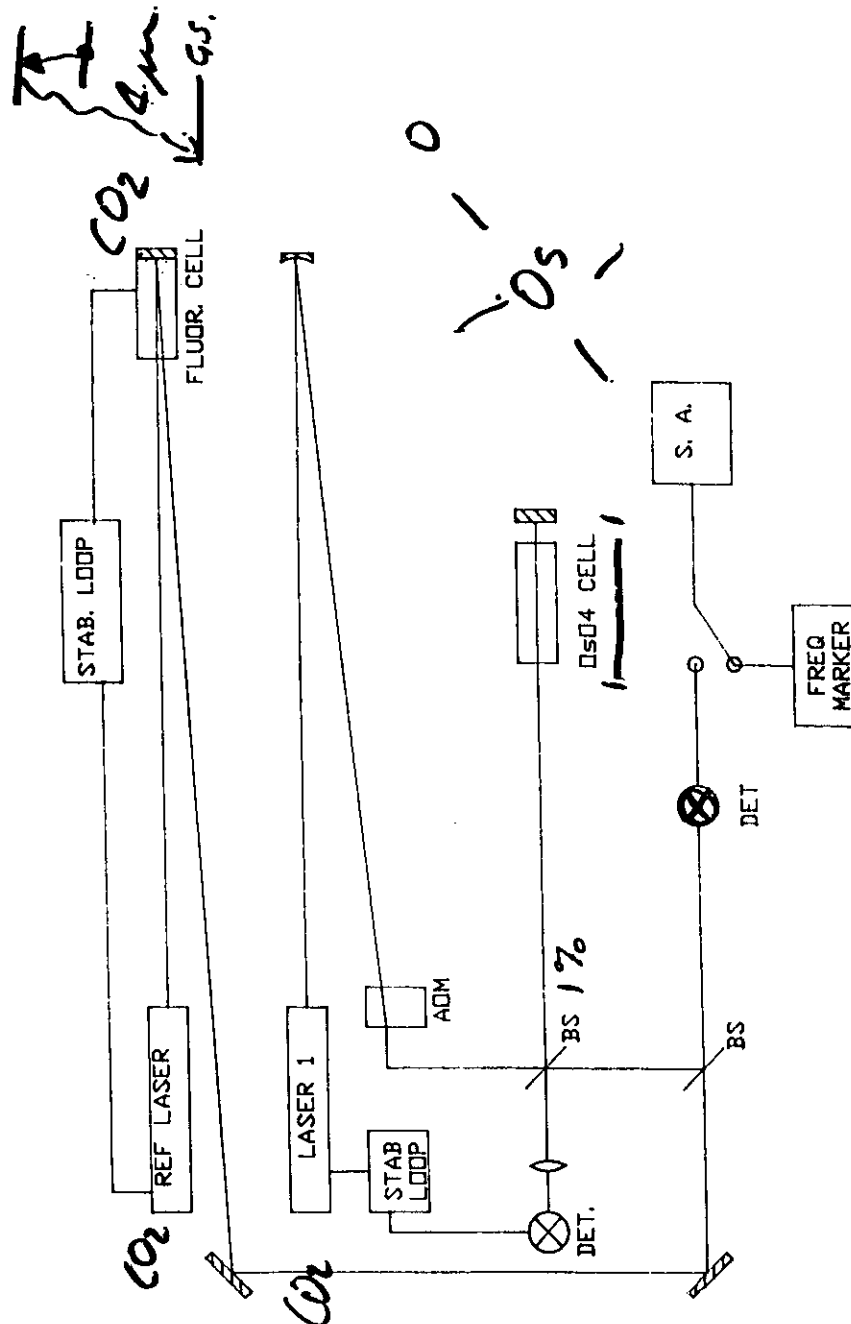
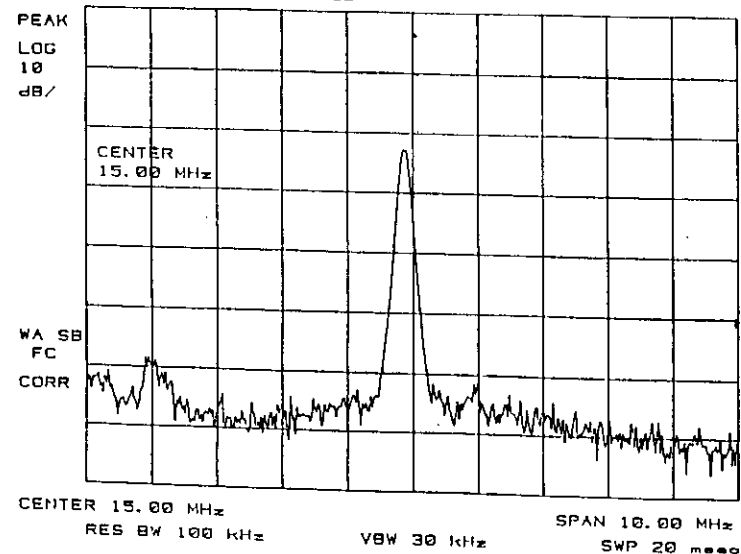


1/2

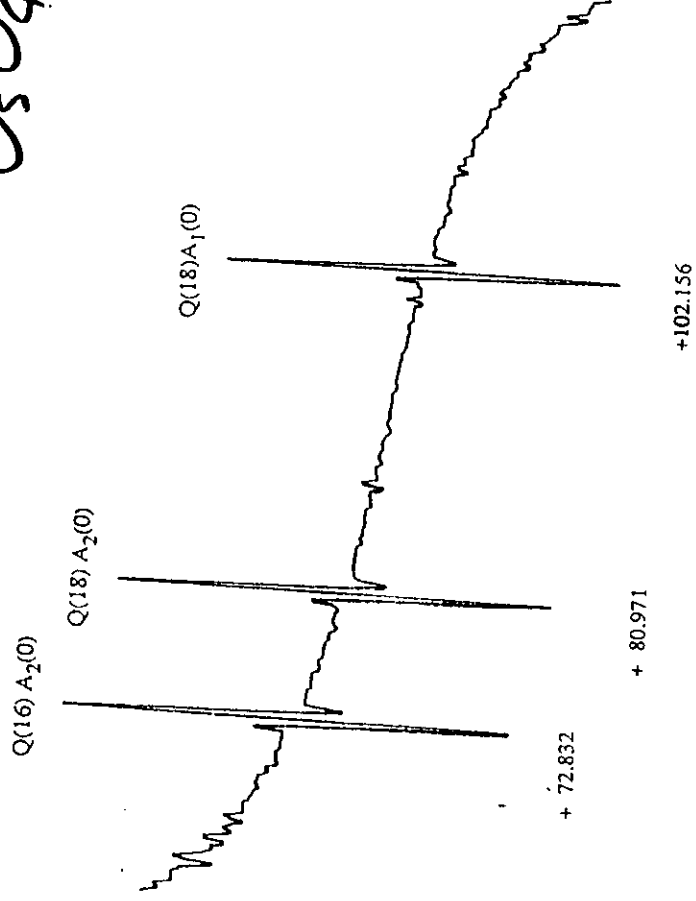
200000 Hz  
Line 10030  
(0.000000)  
200000 Hz  
1000000 Hz

1/2

REF -4.0 dBm ATTEN 10 dB



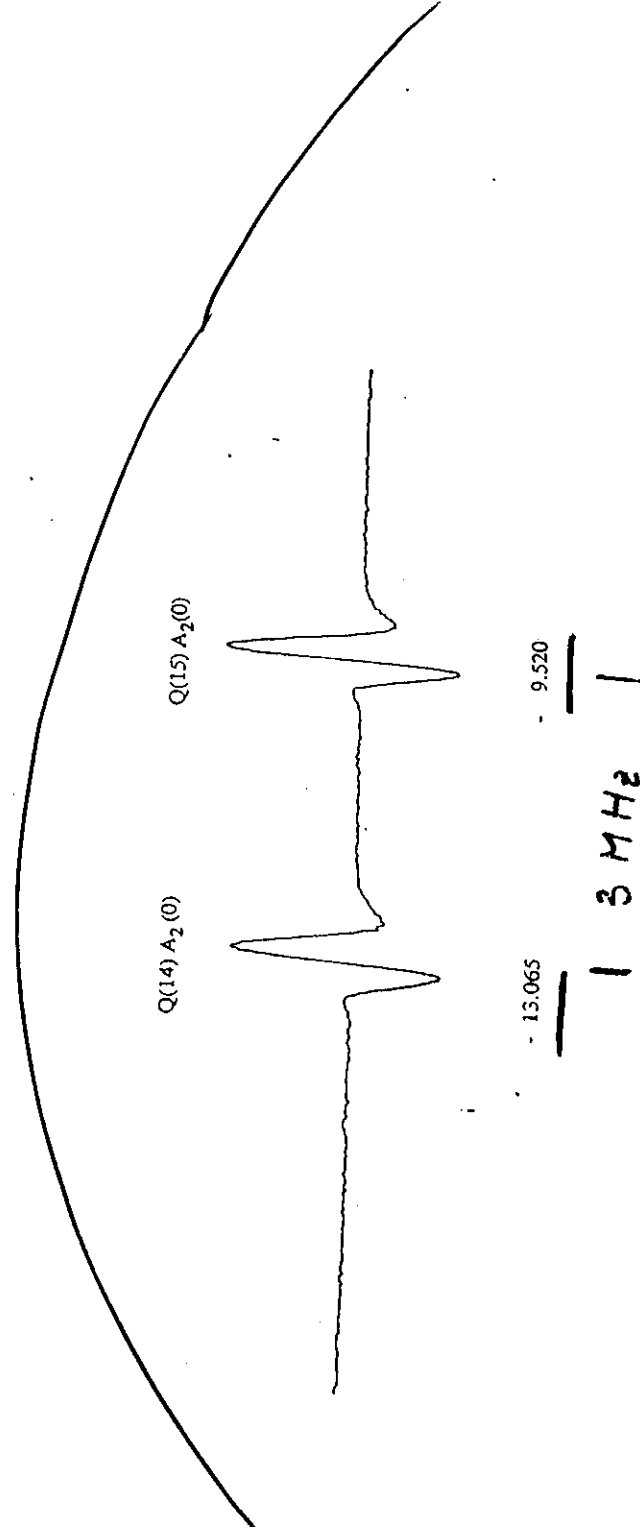
$\text{OsO}_4$



(31)

$\text{OsO}_4$

Ricci et al J.Chem. Phys. 1990





MOLECULES  $\nu = 2 B J + \dots$   
 ATOMS FINE STRUCTURES

RELATIVE SENSITIVITY (Absorption Spectroscopy)

ABSORPTION COEFFICIENT (A.C.)

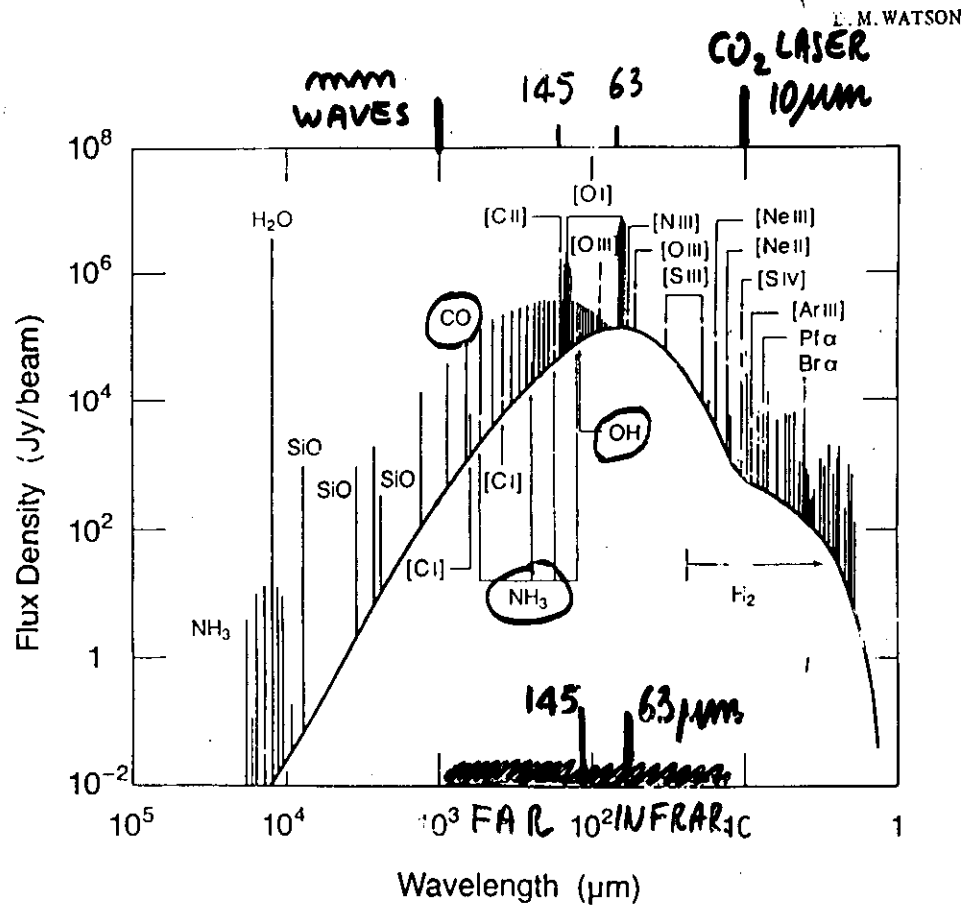


Figure 1: The spectrum of Orion-KL. Omitted for clarity are numerous molecular rotational lines in the millimeter-wavelength region which are present with antenna temperatures as high as about 20 K, the free-free continuum emission of M42, and all of the H<sub>2</sub> hydrogen recombination lines but the "alphas."

ASTROPHYSICS

HETERODYNE MEAS. ACCURACY  $< 1 \text{ MHz}$

$$\frac{(N_{\text{LOWER}} - N_{\text{UPPER}}) (\mu)^2 \nu^2}{(\text{LINE WIDTH})}$$

$$= \frac{N_{\text{LOWER}} \left(1 - e^{-\frac{h\nu}{kT}}\right) (\mu\nu)^2}{\left[\frac{\nu}{\text{DOPPL}} + \frac{1}{\tau_{\text{coll}} + \tau_{\text{rad}}}\right]}$$

| REGION   | TRANS | $\frac{h\nu}{kT}$ | $1 - e^{-\frac{h\nu}{kT}}$ | $\frac{1}{\nu_D + 1/\tau_{\text{coll}} + \tau_{\text{rad}}}$ | $(\mu\nu)^2$   | A.C.      |
|----------|-------|-------------------|----------------------------|--|----------------|-----------|
| mm waves | ROT   | .03               | .01                        | 1  | $10^{-3}$      | $10^{-5}$ |
| FIR      | ROT   | 1.0               | .2                         | 1  | 1              | .2        |
| FIR      | ROT   | 3.0               | .5                         | 1/2  | 9              | 2.5       |
| IR       | VIB   | 30.0              | 1                          | 1/60   | 9              | .15       |
| VIS      | EL.   | 500               | 1                          | 1/1000   | $3 \cdot 10^5$ | 300       |

BUT, FOR ATOMS...  
 ... magnetic dipole

# DENSITY OF FREQUENCY

$3 \times 10^{13}$  MATCHES

10  $\mu$

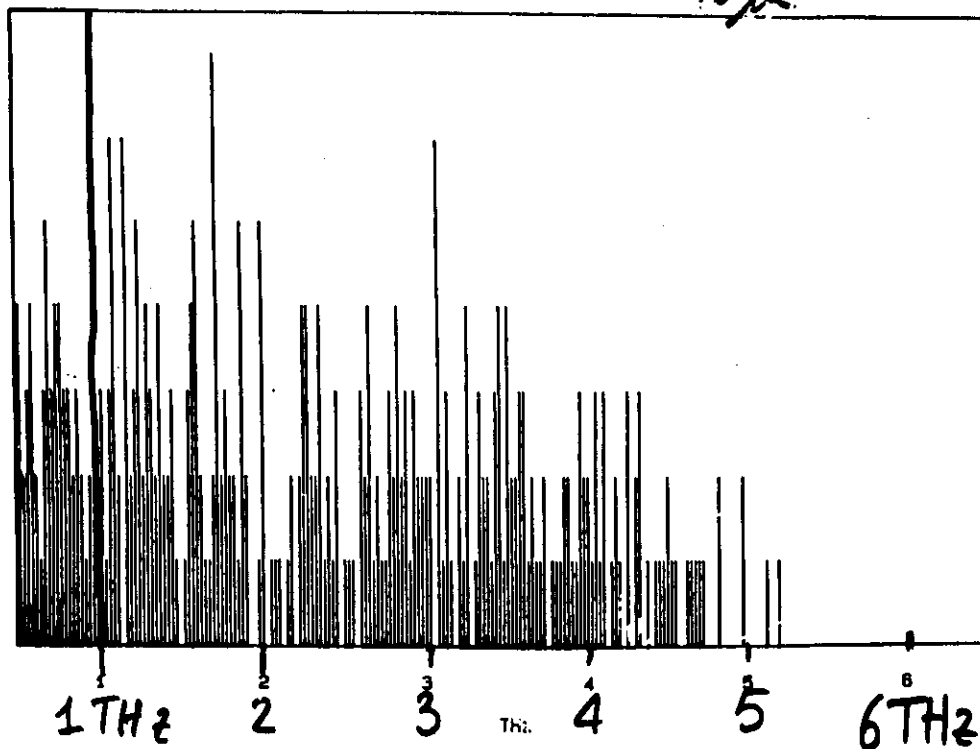
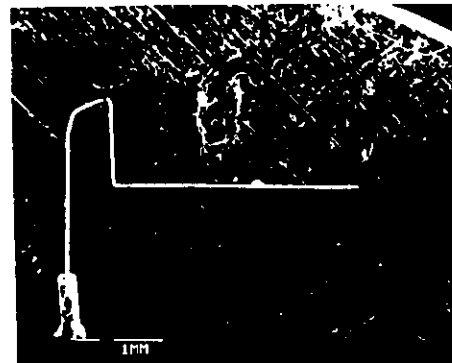


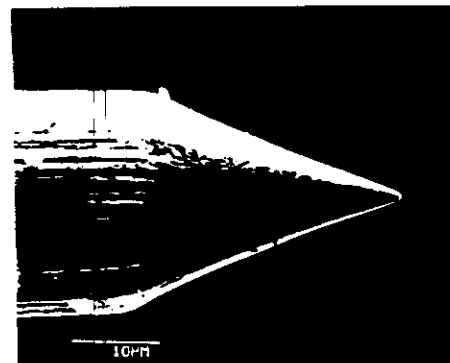
Figure 5. Density of frequency matches.

1 mm

50  $\mu$ m



LENGTH  
2.7 mm



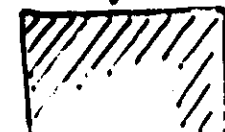
ANGLE  
 $\pm 5^\circ$



TIP RADIUS  
3000 Å



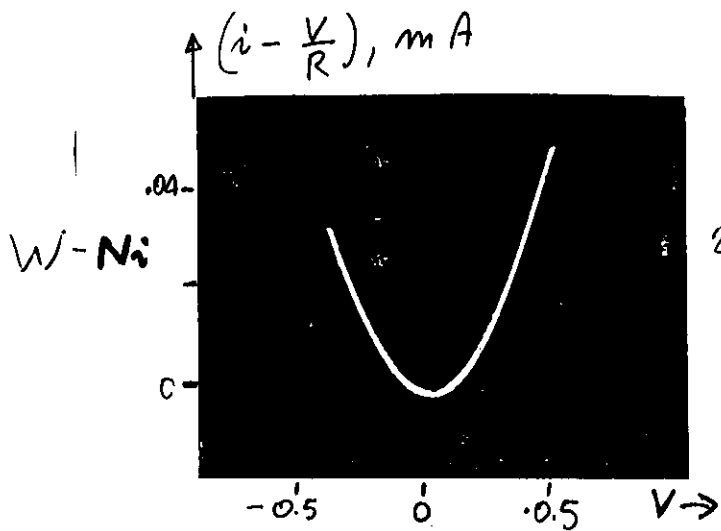
W



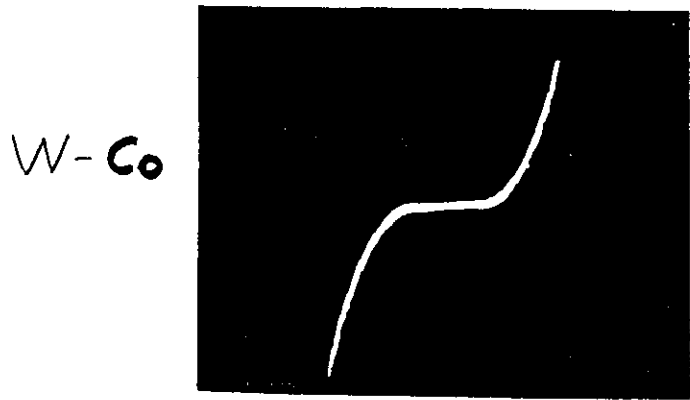
Ni

$$\dot{V}_x = \lambda \cdot \dot{V}_0 \pm \dots \pm \dot{V}_1 \pm \dot{V}_6$$

$$i = f(V) = \frac{V}{R} + \frac{i''}{2} V^2 + \frac{i'''}{3!} V^3 + \dots$$

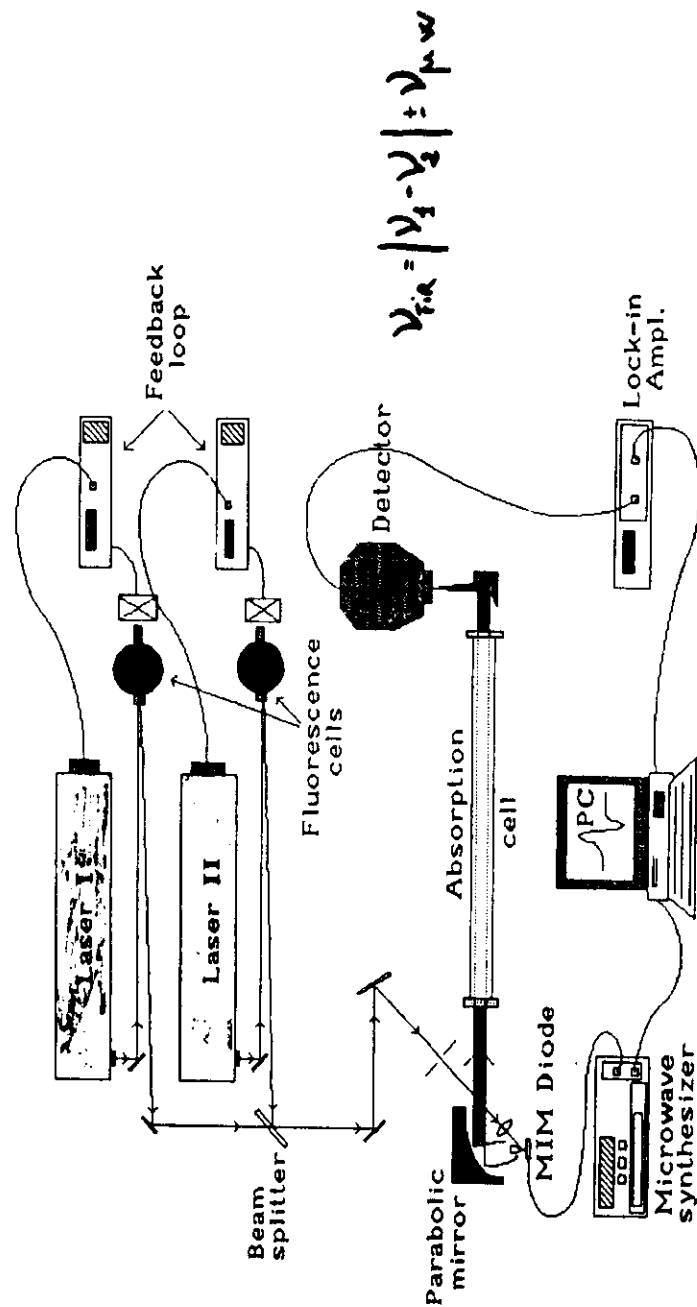


2<sup>nd</sup> ORDER



3<sup>rd</sup> ORDER  
(2<sup>nd</sup> negligible)

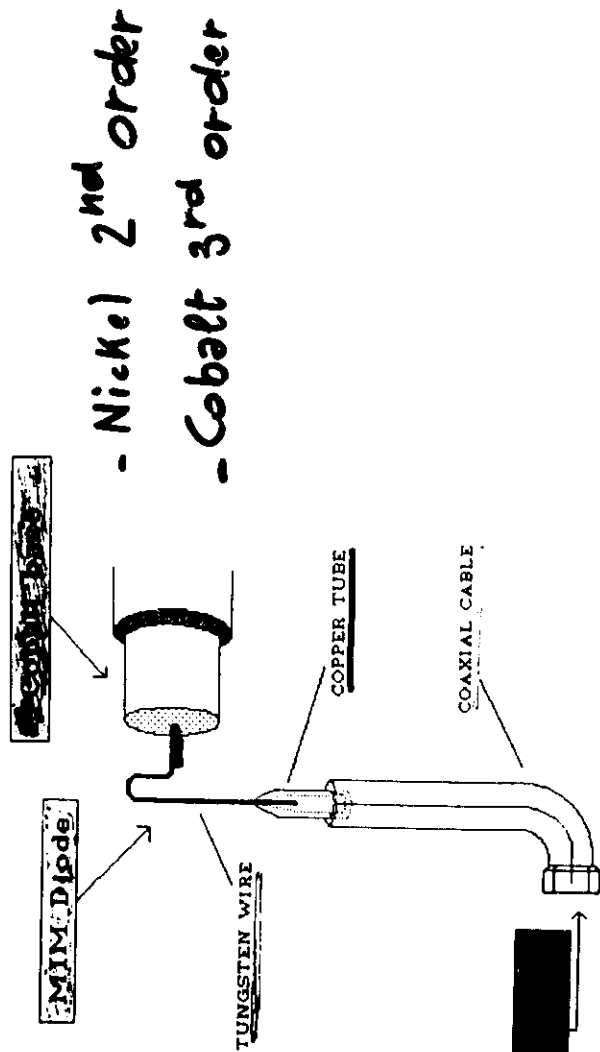
(35) "THREE WAVE" MIXING  
(ONLY TWO LASERS)



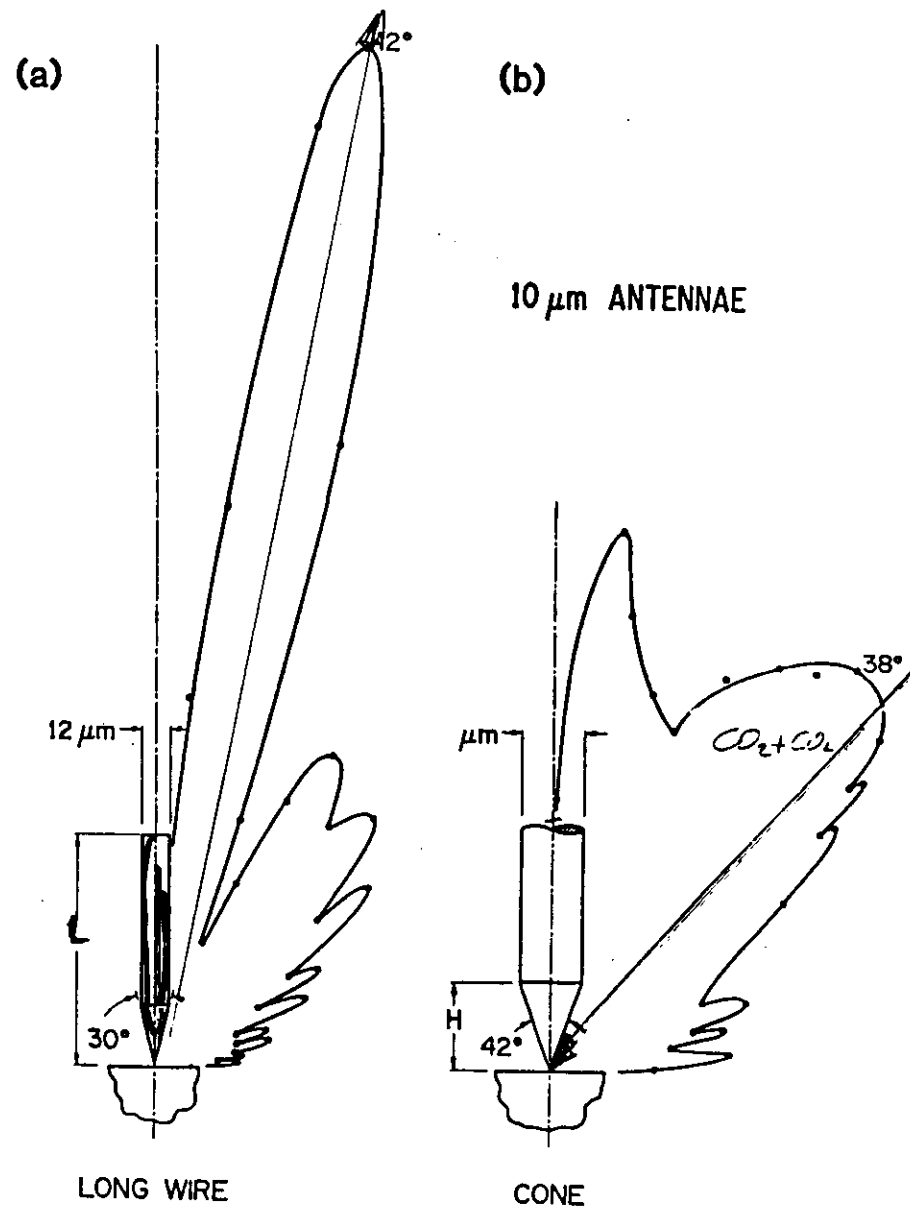
Last months also:  
TOYAMA UNIV, JAPAN  
MATSUBISHI

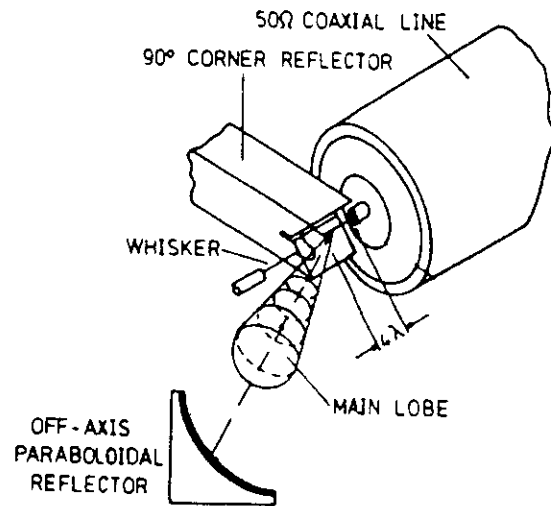
FIRE: APPARATUS

# "THE MIXER"



EVENSON, INGUSCIO, JENNINGS J. Appl. Phys. 57, 956 (85)





BETTER COLLIMATION  
LINEAR POLARIZATION

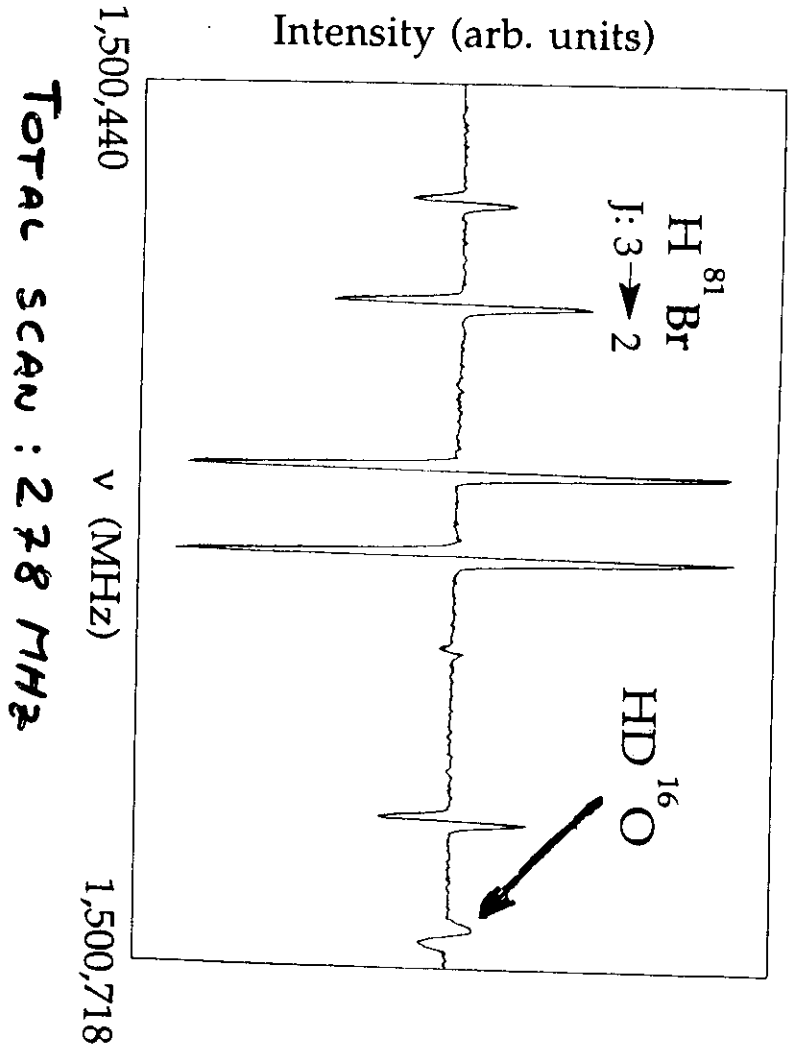
- Extension of the precision typical of the neighbouring microwave region (Heterodyne techniques)
- Use of frequency stabilized infrared sources (CO<sub>2</sub> lasers)



### Tunable Far Infrared spectrometer

- Continuous tunability (0.3 - 6 THz; 1000 - 50  $\mu\text{m}$ )
- Doppler limited resolution (FIR linewidth: a few tens of kHz)
- $1/10^4$  minimum detectable absorption (1 sec time constant)
- $1/10^8$  average accuracy

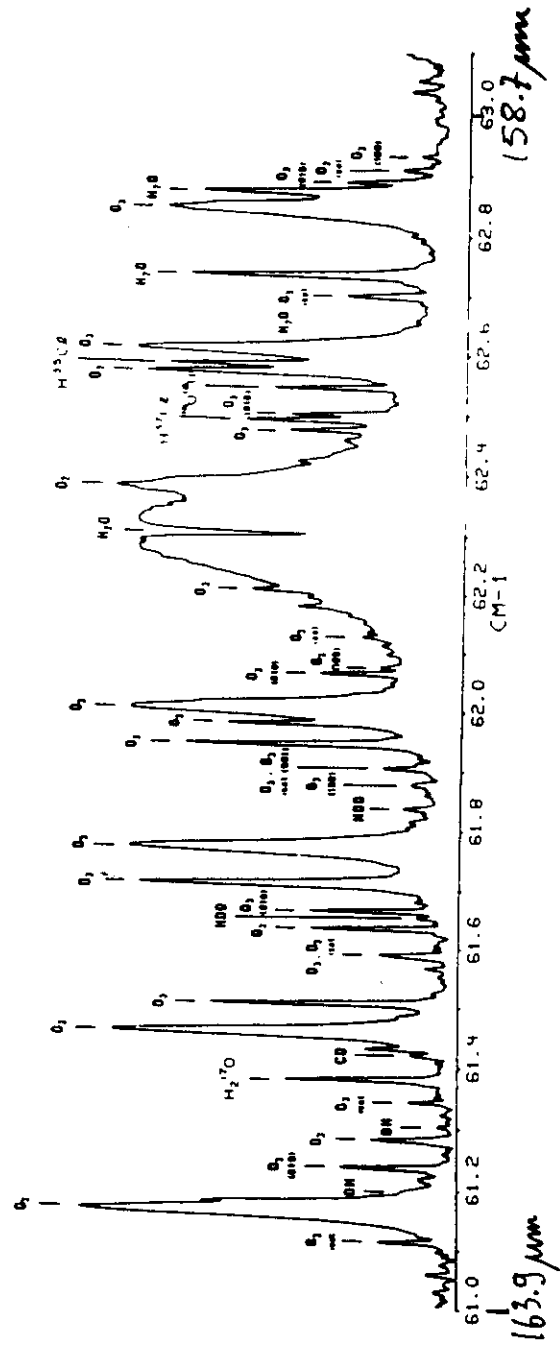
AN EASY EXPERIMENT (STABLE  
POLAR MOLECULE) TO TEST THE APPARATUS.  
J. Mol. Spectros. 1991



A spectroscopic example: hyperfine  
structure of HBr ... and more

(32)

A SMALL PORTION OF THE FAR INFRARED SPECTRUM OF  
THE ATMOSPHERE



From  
B. Carli  
J. QUANT. SPECTROS. RAD. TRANSF.  
32, 397 (1984)

SPECTRAL RADIANCE

BLACK BODY RADIANCE

$$S(\sigma, h_F, \theta) = \int_{\tau(\sigma, S_b)}^{\tau(\sigma, 0)} B[\sigma, T(s)] d\tau(\sigma, s)$$

WE NEED A HIGH SENSITIVITY  
LABORATORY TECHNIQUE TO GET

100 km  $\longrightarrow$  1 m  
unstable species

ATH. TRANSMISSION  $\tau$

$$\tau(\sigma, s) = \exp\left[-\int_0^s \sum_i [k_i(\sigma, p, T) q_i(s) p(s) ds]\right]$$

- LINE FREQUENCIES

- COLLISIONAL LINESHAPES

- LINE STRENGTHS

ABSORPTION  
COEFFICIENT

STRENGTH

LINESHAPE

DENSITY

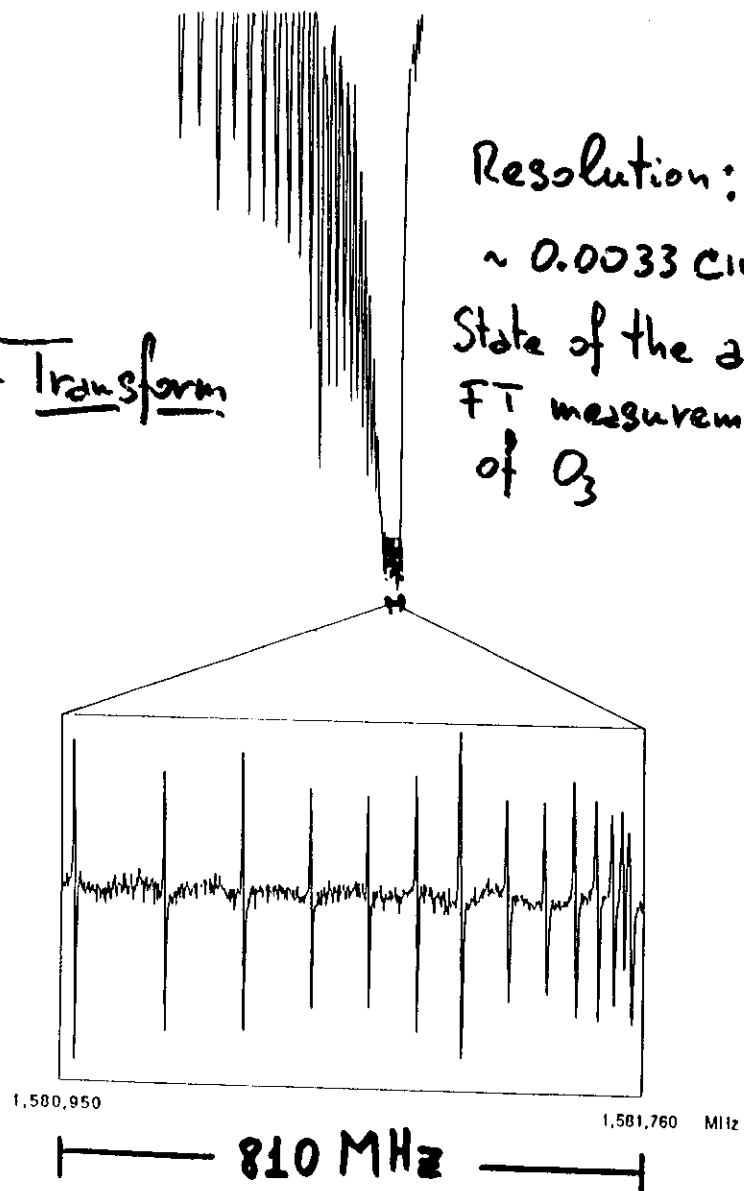
$$k_i^{lu}(\sigma, p, T) = S \cdot g(\sigma - \sigma_0) \cdot N$$

$$S \propto |\mu|^2$$

Fourier Transform

VS:

T<sub>4</sub>-FIR



Resolution:

$\sim 0.0033 \text{ cm}^{-1}$

State of the art  
FT measurement  
of  $\text{O}_3$

## FREQUENCY ACCURACY

$\text{CO}_2$  absolute frequency  $\pm 3 \text{ kHz}$

" Lock uncertainty  $\pm 25 \text{ kHz}$   
( $\pm 7 \text{ kHz}$  best)

$\Delta \nu_{\text{FIR}} = 35 \text{ kHz}$   
( $10 \text{ kHz}$  best)

$\Delta \nu_{\text{TRANSITION}}$

$\sim \Delta \nu_{\text{FIR}}$

for good S/N

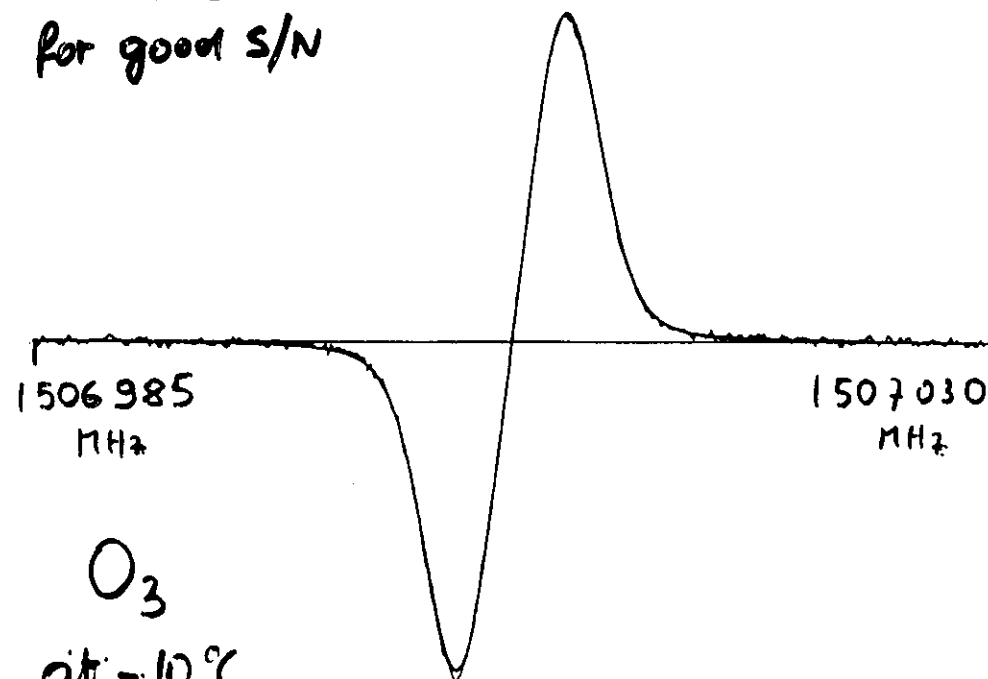


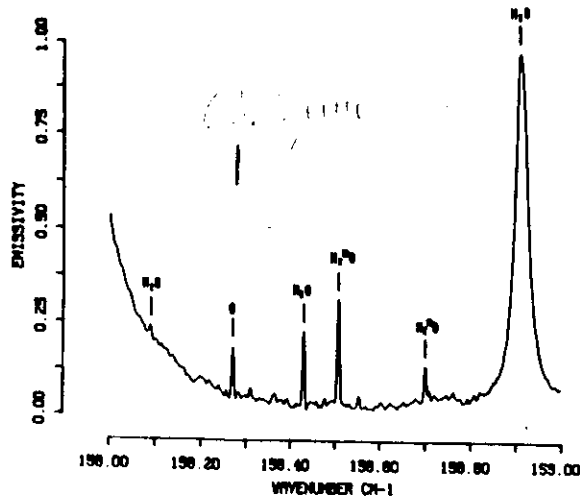
Fig. R18 T--10 p-58 I-63



# UPPER ATMOSPHERE ATOMIC OXYGEN IN THE G.S.

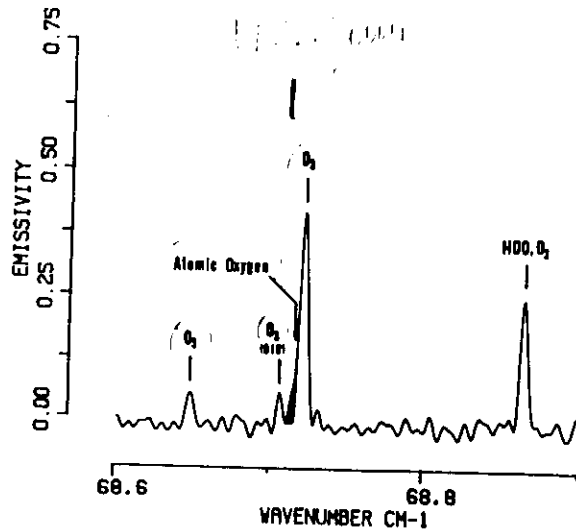
LHR  
K. EVENSON

$\tau_{\text{rad}} = 1.1 \times 10^{-9} \text{ sec}$   
 $\tau_{\text{RAD}} = 1.1 \times 10^4 \text{ sec}$



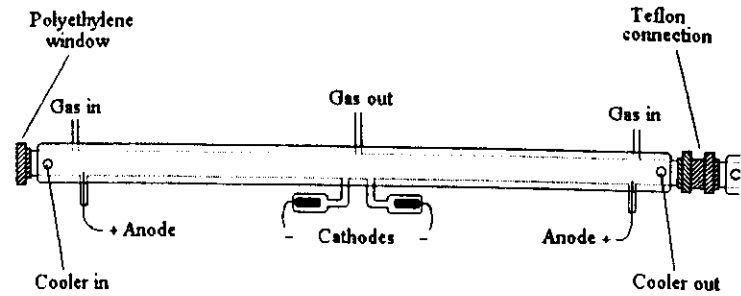
LHR  
P. DAVIES

$\tau_{\text{rad}} = 6 \cdot 10^{-9} \text{ sec}$   
 $\tau_{\text{RAD}} = 6 \cdot 10^4 \text{ sec}$

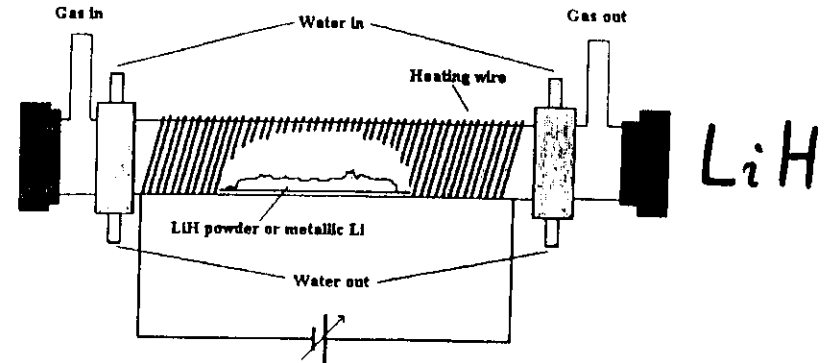


Polymers etc  
Catalytic reactions

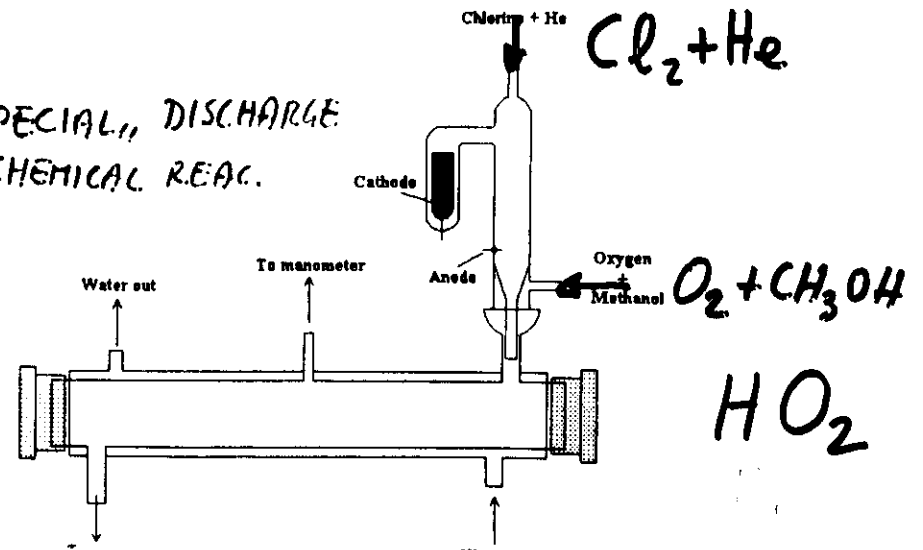
## 21 GLOW DISCHARGE



## HEAT PIPE

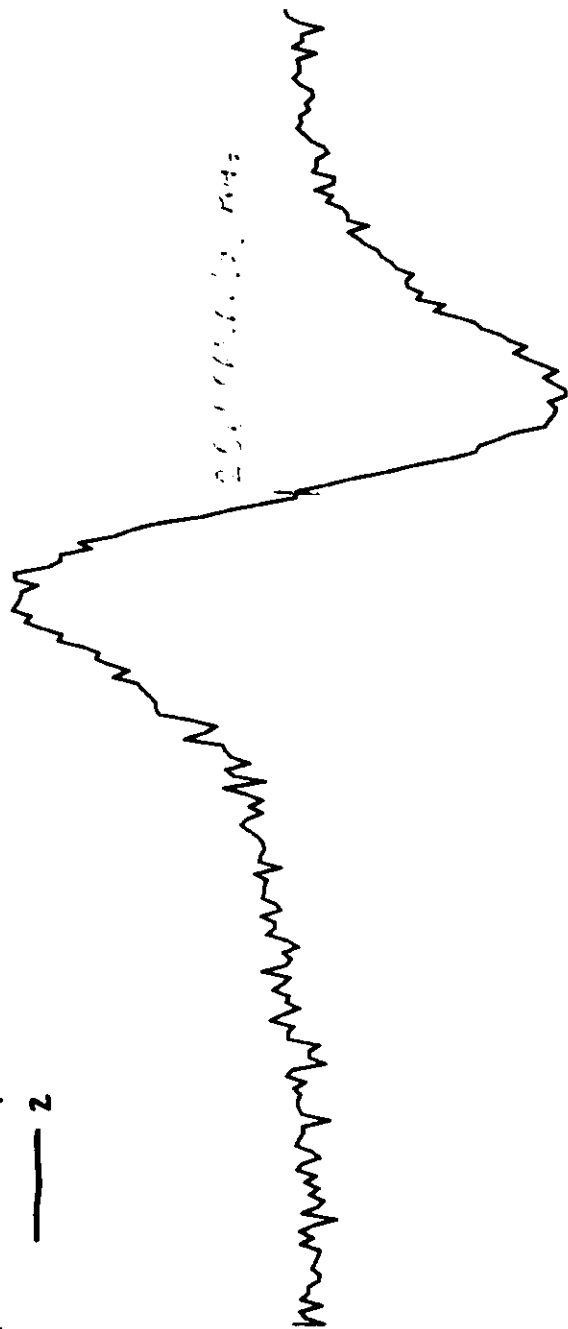


"SPECIAL" DISCHARGE  
+ CHEMICAL REAC.



ATOMIC OXYGEN

145  $\mu\text{m}$   
0  
1  
2



A: \atox2d.4

## $\text{HO}_2$ - MOTIVATION

Quantitative measurement in the atmosphere is a high priority

( $\text{HO}_2$  important constituent of  $\text{HO}_x$  chemical cycle which regulates the earth's stratospheric ozone).

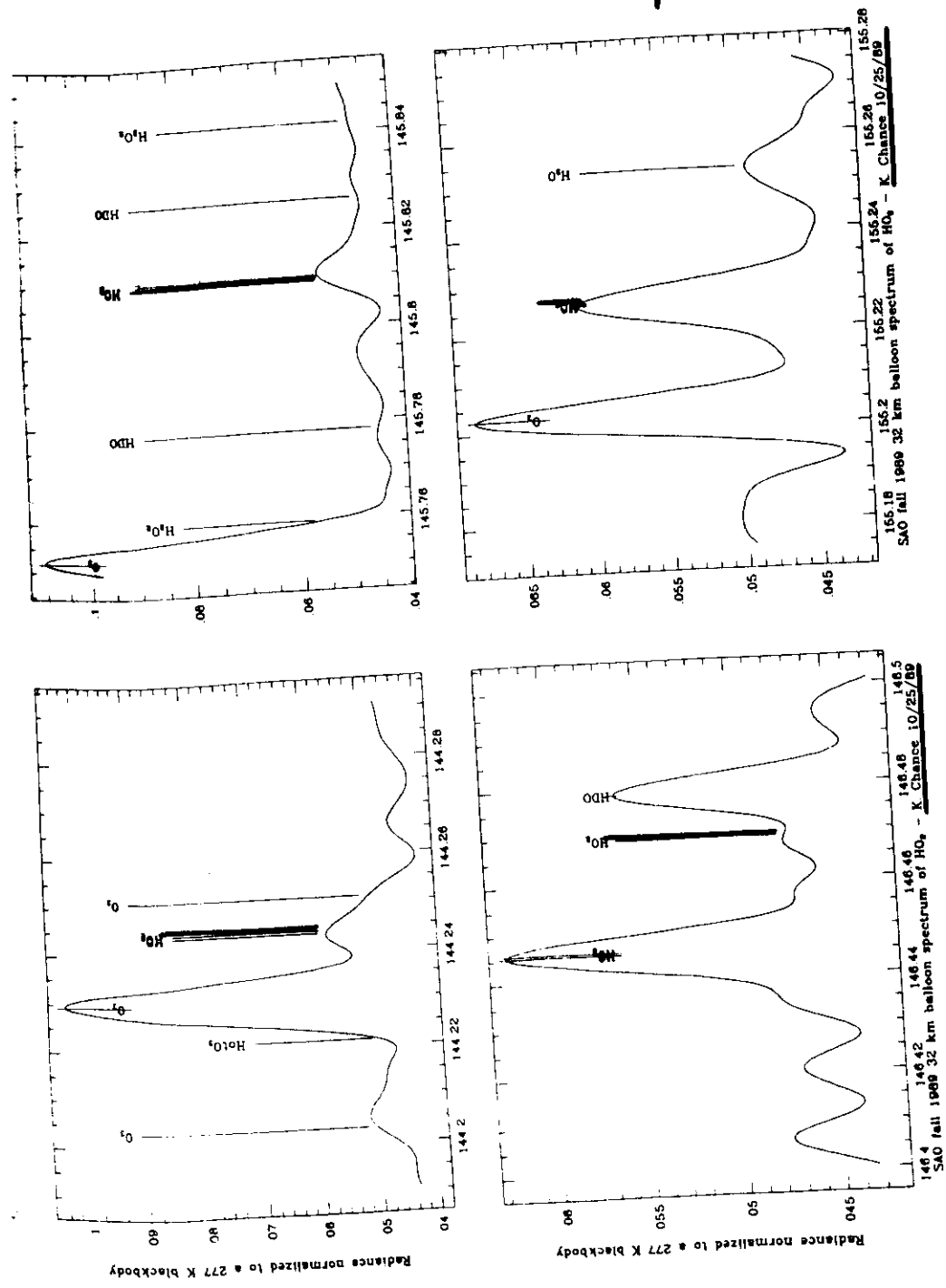
Balloon based measurements of relatively strong far infrared rotational lines in emission.

(Future developments include heterodyne detection).

Need for: line positions (OK), line strengths (OK), pressure broadening which determines the line widths in the atmosphere

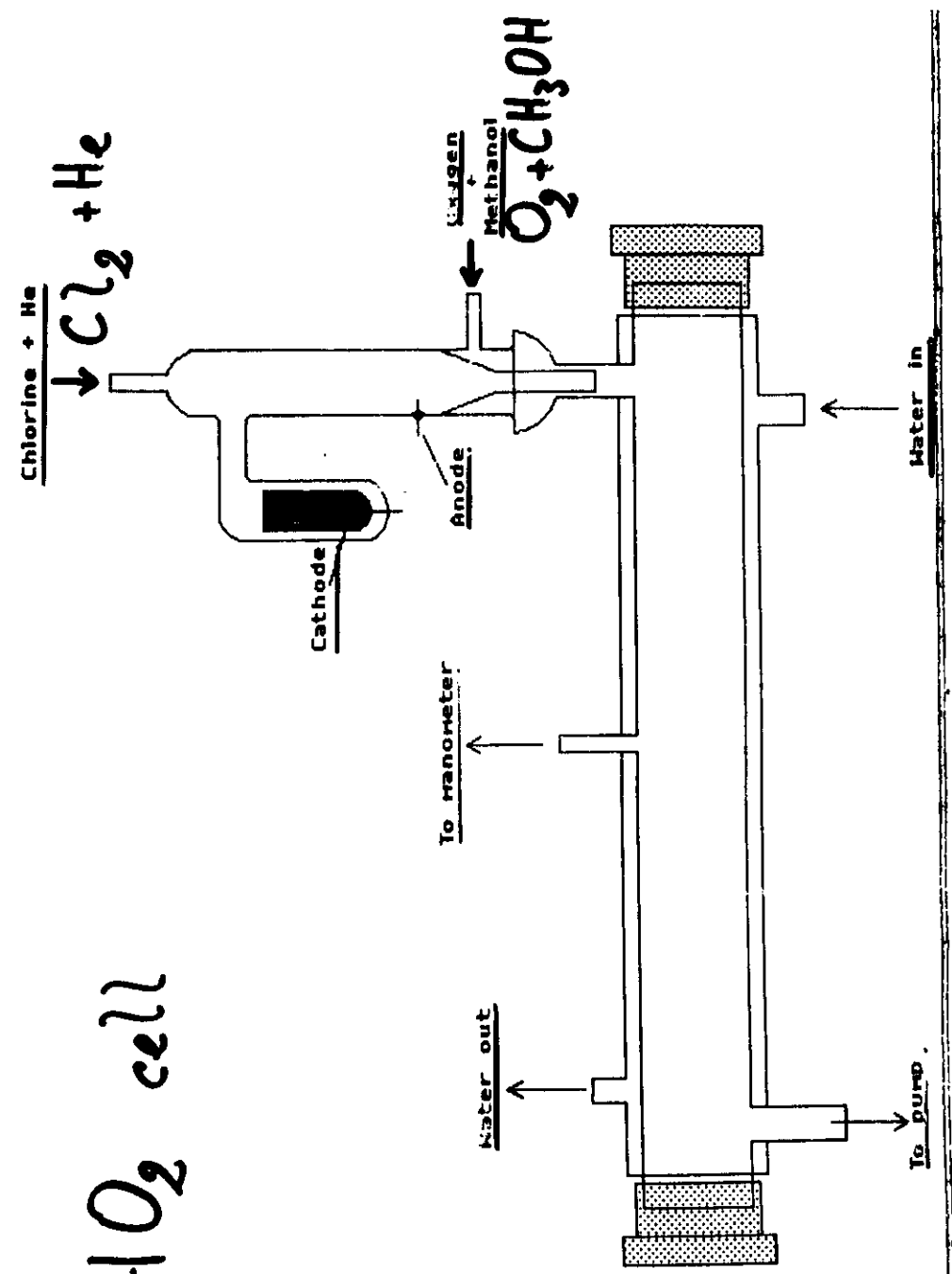
P.B. is essential in the analysis of heterodyne spectra to determine the contributions from  $\text{HO}_2$  at various pressure altitudes.

# Lines recorded in the atmosphere.....



(52)

$\text{H}_2\text{O}_2$  cell



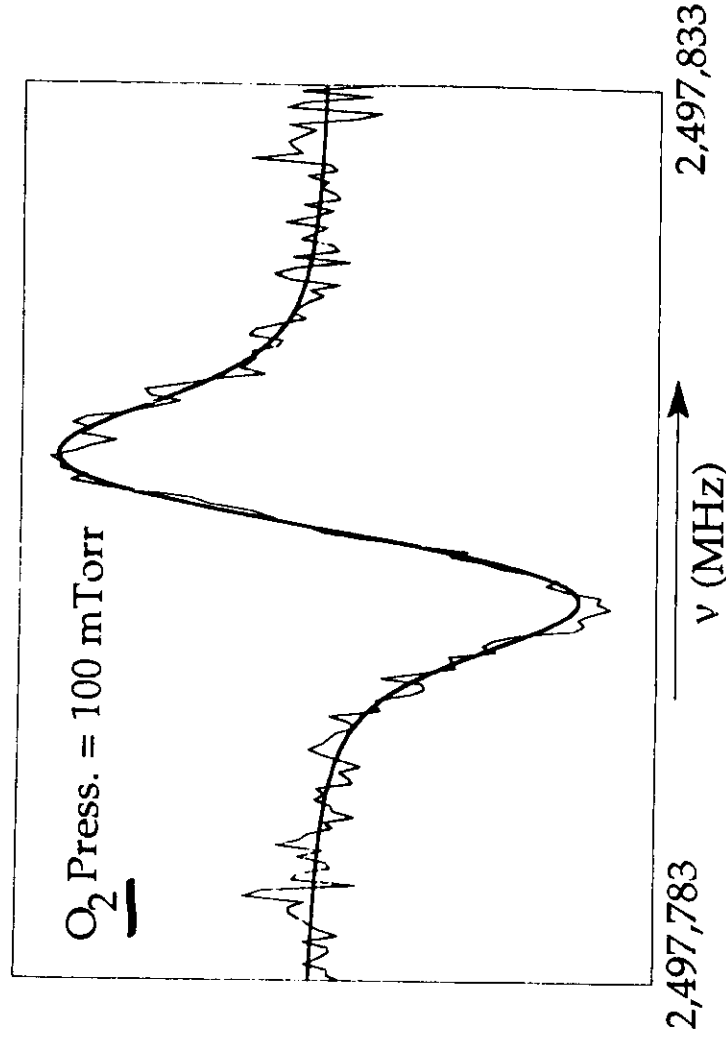
# CHEMISTRY FOR $\text{HO}_2$ PRODUCTION IN THE LABORATORY



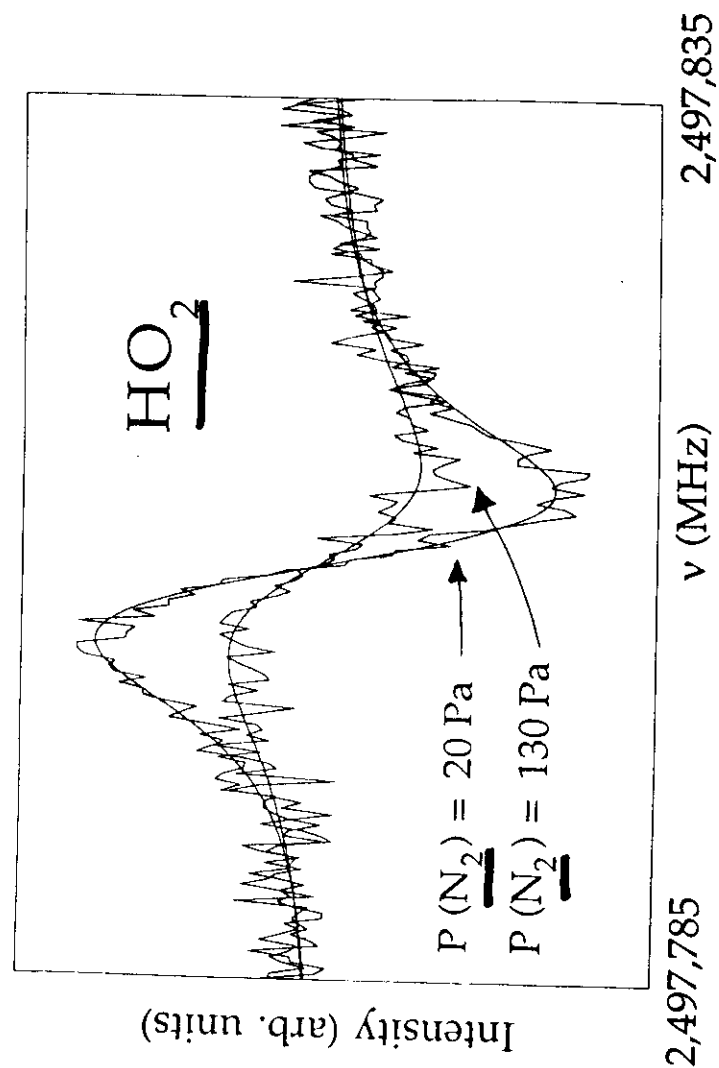
33

$\text{HO}_2$

.... and in the lab



(43)



## Results

$$\gamma(\text{N}_2) = 47.4 \pm 4.9 \text{ kHz/Pa} \\ (6.3 \pm 0.6 \text{ MHz/Torr})$$

$$\gamma(\text{O}_2) = 19.9 \pm 1.6 \text{ kHz/Pa} \\ (2.7 \pm 0.2 \text{ MHz/Torr})$$

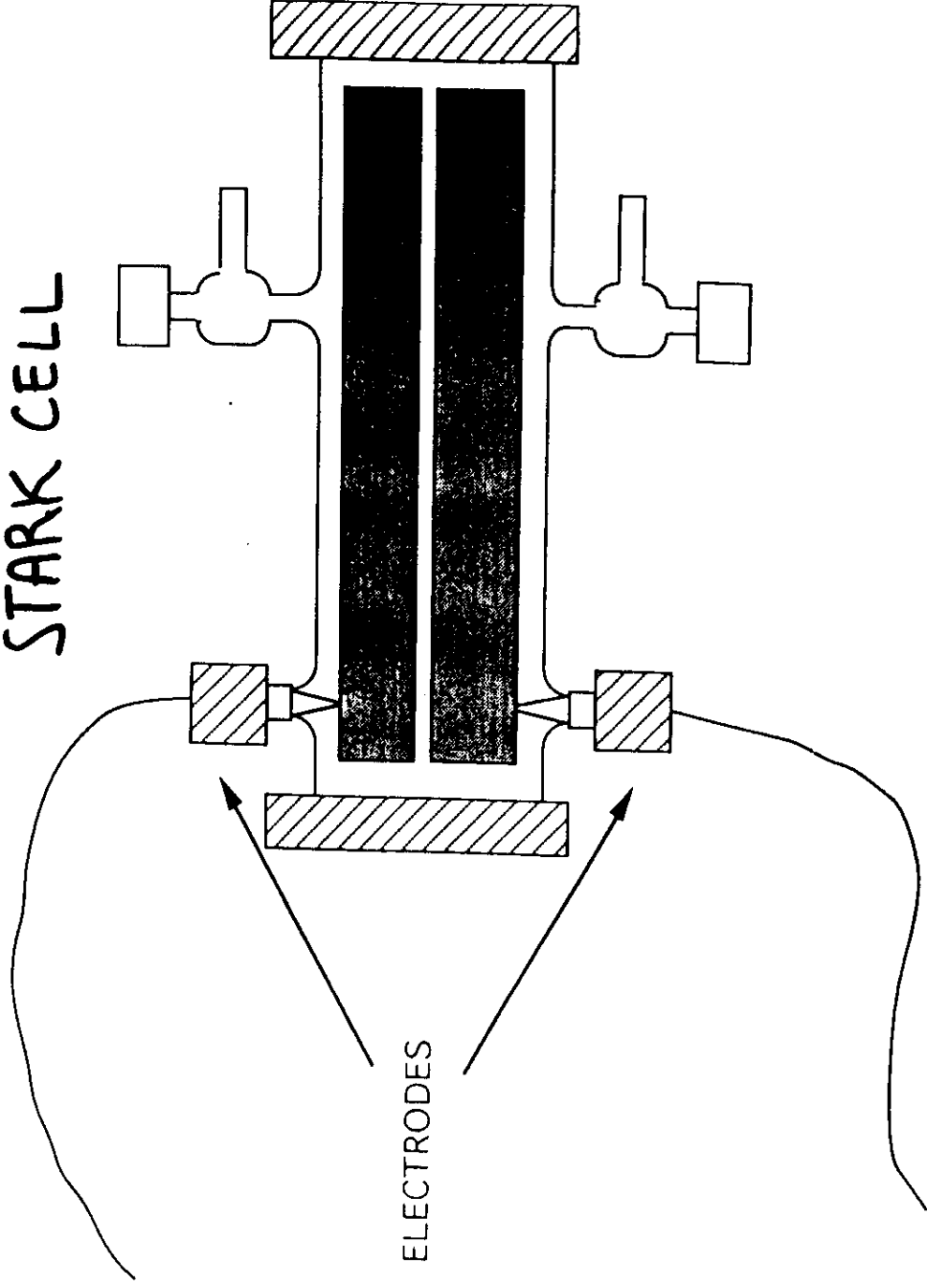
$$\gamma(\text{air}) = 41.6 \pm 3.8 \text{ kHz/Pa} \\ (5.6 \pm 0.5 \text{ MHz/Torr})$$

$$\text{at } T = 300 \pm 1 \text{ K}$$

(1σ uncertainties)

γ uncertainties still better than 10%

# STARK CELL



(45)

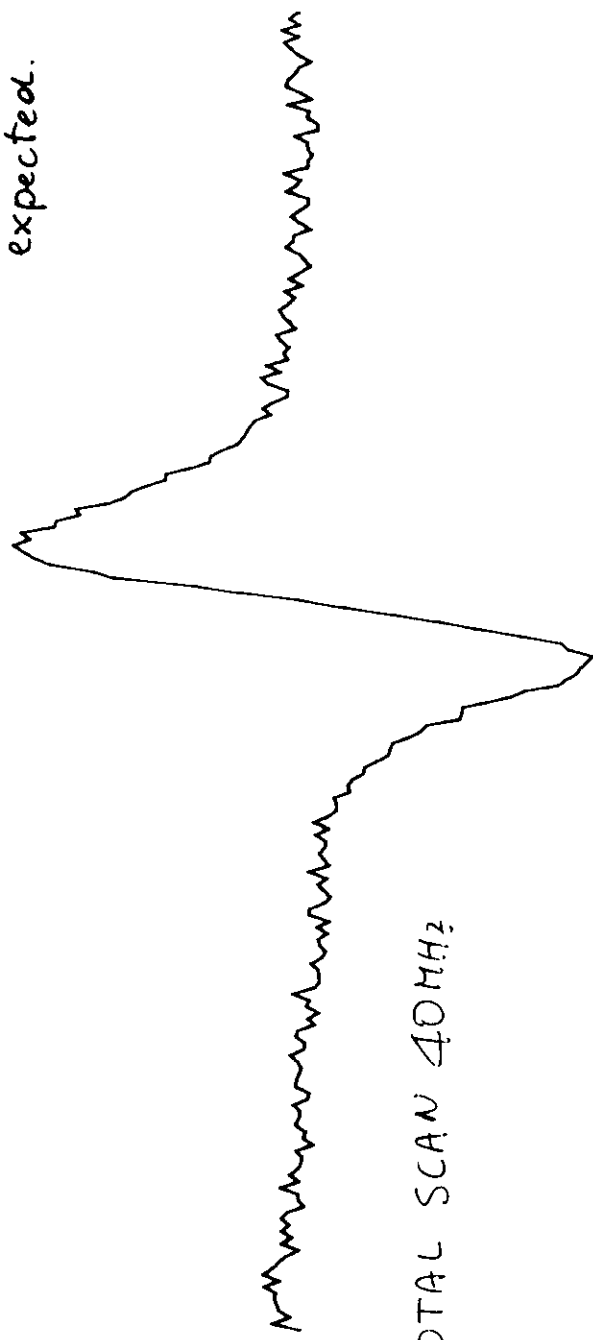
CH<sub>3</sub>OH  
at 328  $\mu$ m

3<sub>0,3</sub> - 2<sub>0,2</sub>

$\Delta J = 0$

J: 2  $\rightarrow$  3

5 components  
expected.



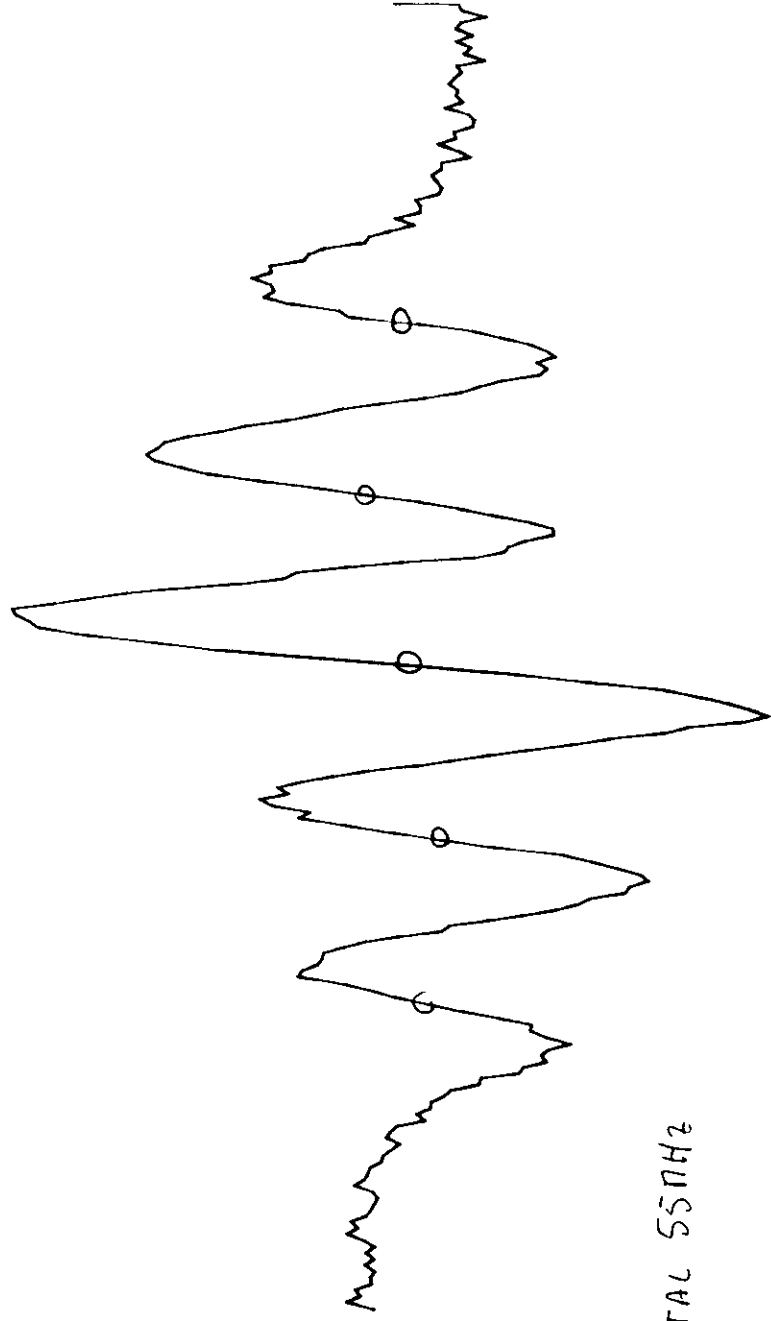
TOTAL SCAN 40 MHz

1  
912126.434 MHz

912086.434

$\bar{E} = 0$  Volt/cm

M<sub>J</sub>      -2      -1      0      +1      +2



Total 55 MHz

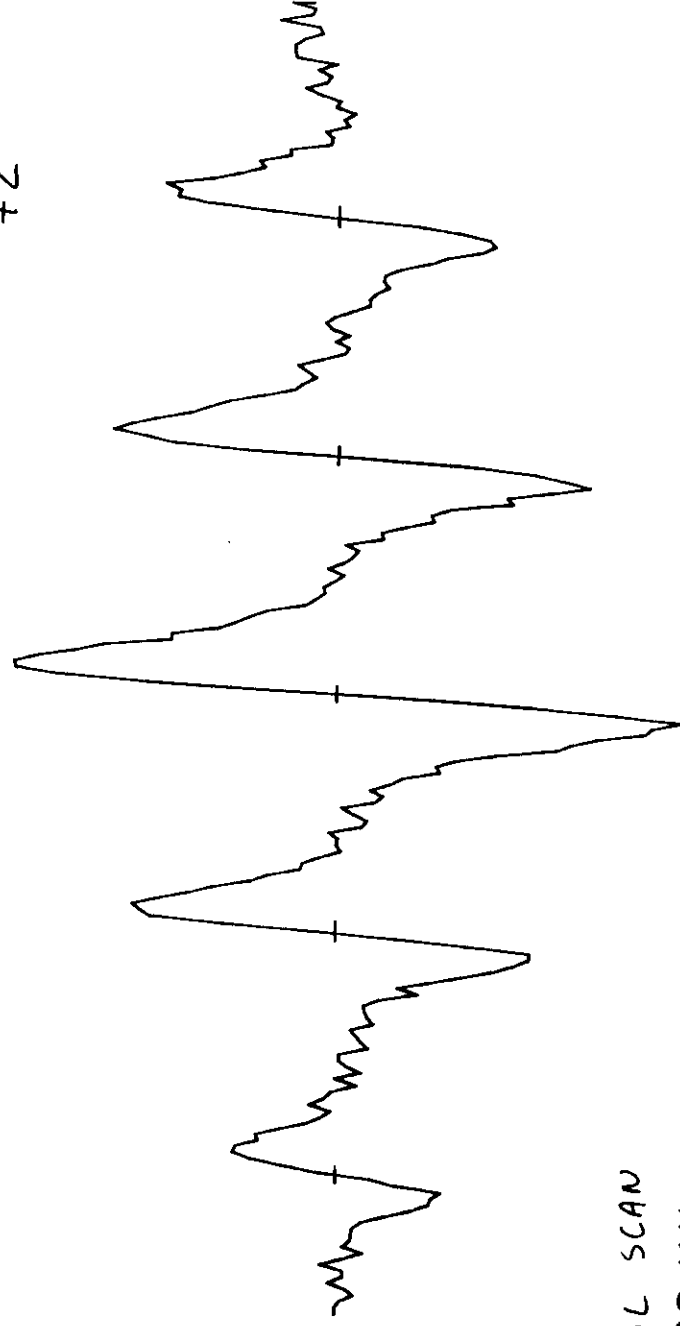
1  
912136.434

E = 200 Volt/cm

912081.434

(56)

-2      -1      0      +1      +2



TOTAL SCAN  
100 MHz

E = 500 Volt/cm

## LITHIUM HYDRIDE

- ✓ Simplest neutral heteronuclear diatomic molecule
- ✓ Test for theoretical models and breakdown of Born-Oppenheimer approx.
- ✓ Search in astrophysical environments (ISO satellite)

### Cosmological implications:

- ♦ Search in high redshift regions
- ♦ Primordial nucleosynthesis



Need for accurate values of frequencies to detect these lines or to extrapolate to uninvestigated regions.



### MOTIVATIONS

#### EXTREMELY LIGHT

(ab initio theories, Born-Oppenheimer approximation breakdown...)

#### VERY LARGE DIPOLE (S.G.D)

(Collisional: test of dipole-quadrupole)

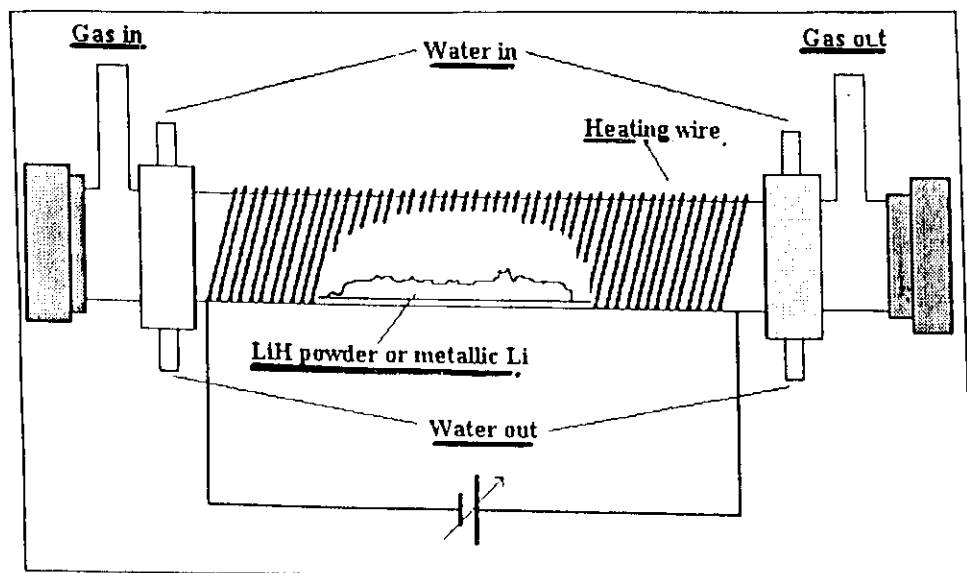
#### COSMOLOGICAL

### PROBLEMS

PRODUCTION REQUIRES HIGH  
TEMPERATURE, HIGH DISSOCIATION  
RATE



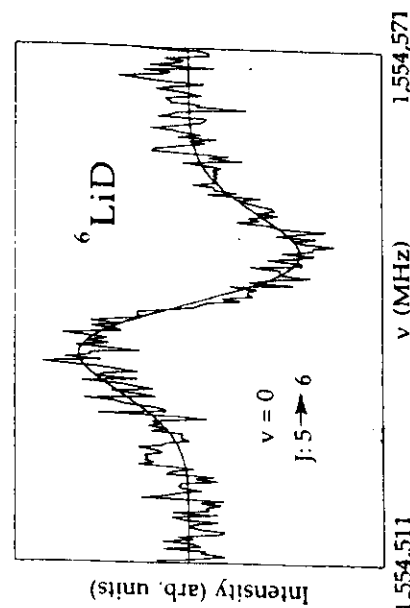
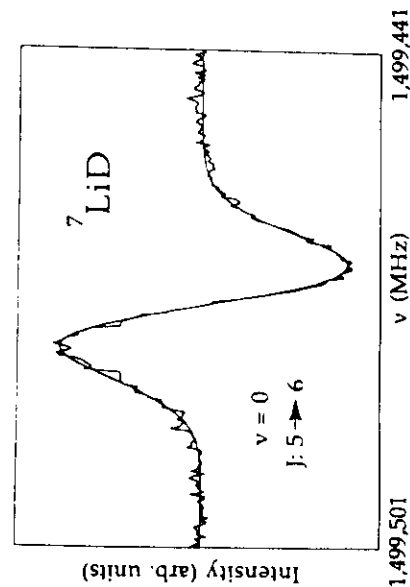
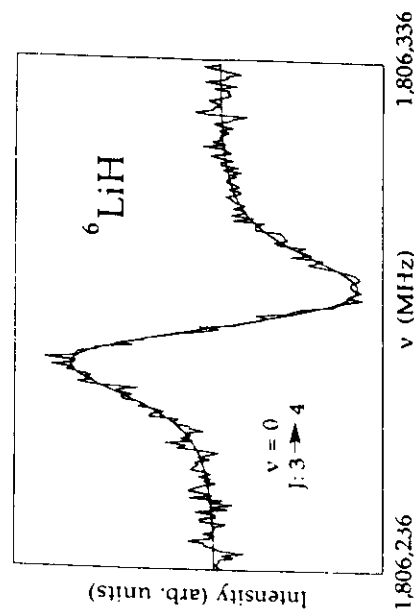
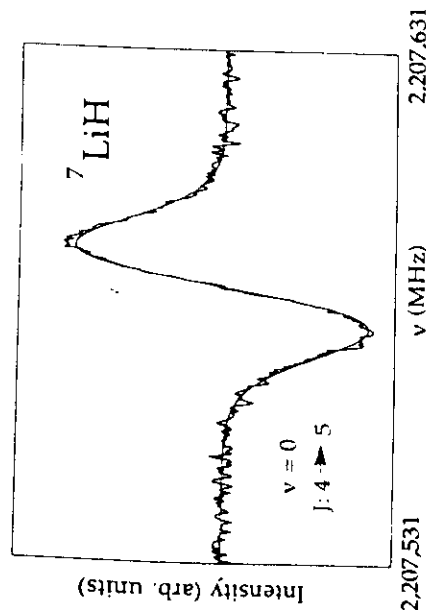
# EXPERIMENTAL SETUP



- ◆ Cell temperature:  $\sim 500 \div 600^\circ\text{C}$
- ◆ All possible combinations of  $^7\text{Li}$ ,  $^6\text{Li}$ , H, D
- ◆ Various vibrational states, from ground to  $v = 2$  excited level

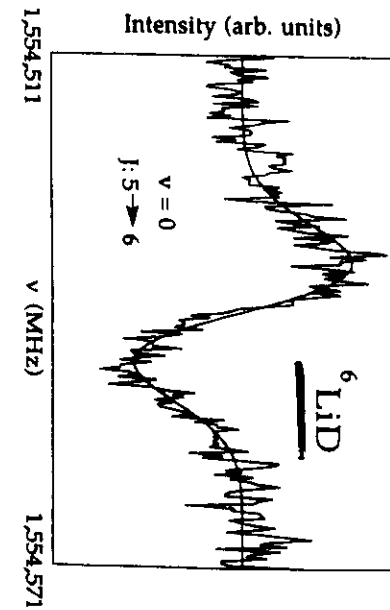
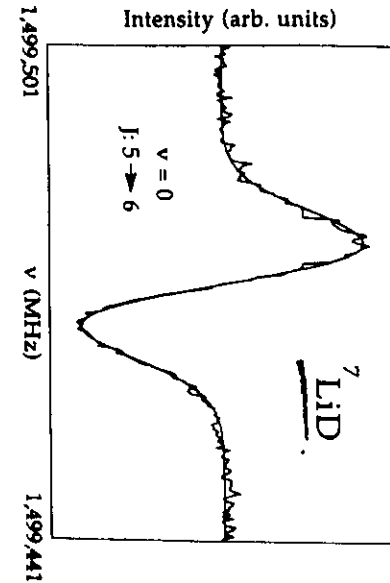
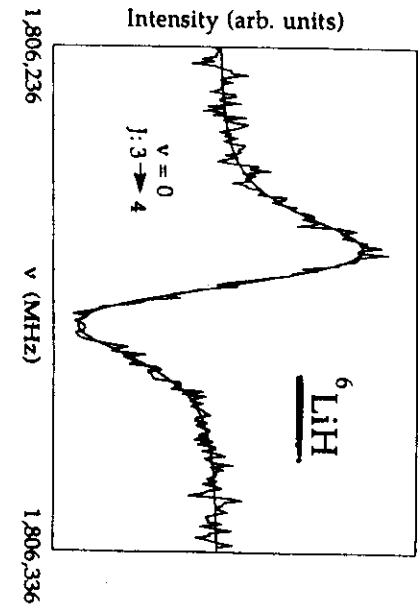
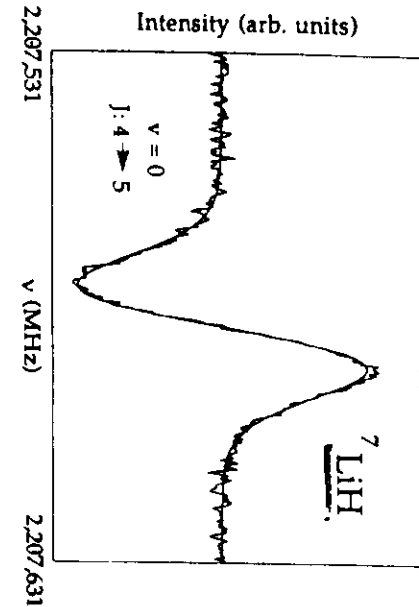
$^6\text{Li}/^7\text{Li}; \approx 0.08$  natural abundance

Deuterium enriched sample

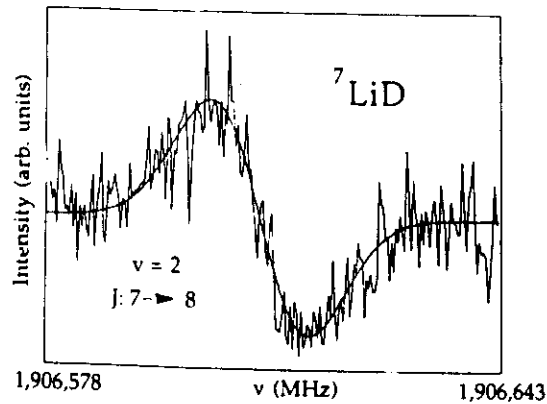
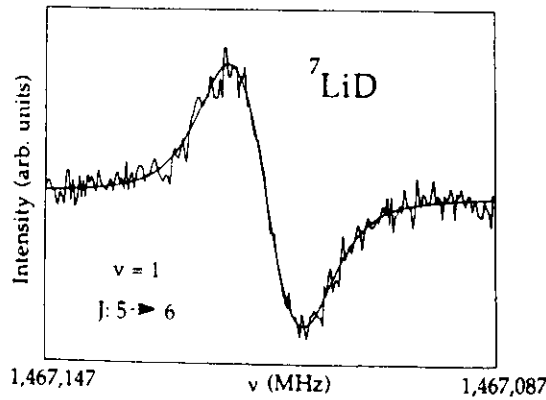
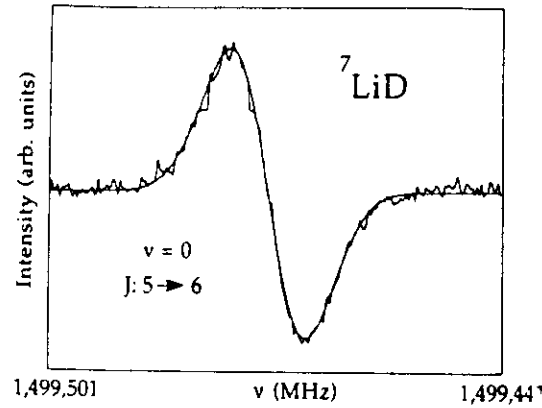


# ALL THE POSSIBLE ISOTOPE COMBINATIONS

GROUND VIBRATIONAL STATE, NATURAL ABUNDANCE  
 $^6\text{Li} \approx 7\%$



FRIDAY POSTER 039



$P = 16\%$  ( $v=0$ )

$P = 2.7\%$  ( $v=0$ )

(69)

HD

# FREQUENCY MEASUREMENT OF THE $J = 1 \leftarrow 0$ ROTATIONAL TRANSITION OF HD<sup>1</sup>

K. M. EVENSON, D. A. JENNINGS, J. M. BROWN, L. R. ZINK, K. R. LEOPOLD, M. D. VANEK, AND I. G. NOLT

Time and Frequency Division, National Bureau of Standards, Boulder

Received 1988 April 15; accepted 1988 April 22

## ABSTRACT

The frequency of the astronomically important  $J = 1 \leftarrow 0$  rotational transition of hydrogen deuteride (HD) at 2.7 THz ( $90 \text{ cm}^{-1}$ ) has been measured with tunable far-infrared radiation with an accuracy of 150 kHz. This frequency is now known to sufficient accuracy for use in future astrophysical heterodyne observations of HD in planetary atmospheres reported by Bezaud *et al.* in 1986 and the interstellar medium reported by Bussolotti *et al.* in 1975.

**Subject headings:** interstellar: molecules — laboratory spectra — molecular processes

A spectrometer using the coherent radiation resulting from the nonlinear mixing of two CO<sub>2</sub> lasers in a metal-insulator-metal diode (Evenson, Jennings, and Petersen 1984) was used to generate the far-infrared radiation. The spectrometer, previously described (Evenson *et al.* 1985), has been used to measure rotational spectra in MgH (Leopold *et al.* 1985), NaH (Leopold *et al.* 1987), OH (Brown *et al.* 1986), ArH<sup>+</sup> (Brown *et al.*

*et al.* 1988), HF (Jennings *et al.* 1987), and CO and HCl (Nolt *et al.* 1987).

The HD was observed in a 19 mm inside diameter and 3.94 m long copper tube, in single transit of about  $10^{-7}$  W of far-infrared radiation used to detect the signal shown in the figure. HD was created in a hollow cathode discharge in a mixture of equal parts of  $\text{I}_2$  and  $\text{D}_2$  by running the discharge for about 3 minutes. A 20% peak absorption was observed at a pressure of 930 Pa (7 torr). See Figure 1. This absorption is somewhat larger than that expected using a value of  $8.83 \times 10^{-4}$  debye for the dipole moment (Drakopoulos and

<sup>1</sup> Contribution of the United States Government, not subject to copyright. Work partially supported by NASA contract W15-047.

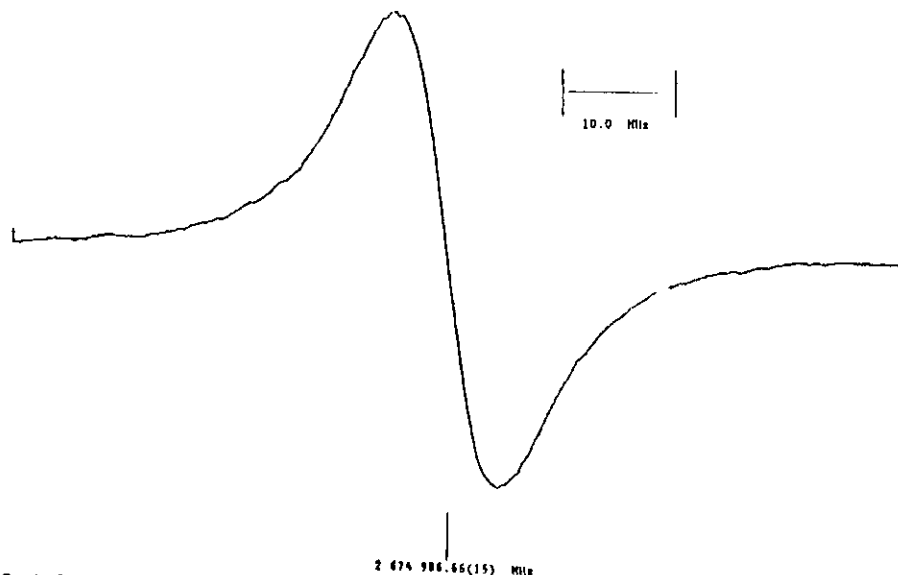


FIG. 1.—Derivative of the HD absorption line taken at 930 Pa (7 torr) with a 0.5 s time constant. A 3.5 m path length and a 5 MHz frequency modulation at a 200 kHz rate was used to observe the signal. A gallium-germanium photoconductor detector was used to observe the absorption using approximately  $10^{-7}$  W of tunable far-infrared radiation.

L135

LIGHTEST MOLECULE

HD

— 708 —

$$= i h \frac{e}{c} \frac{h}{Mc} \epsilon_{01} \left( x_1 \frac{\partial^2 A_0}{\partial x_2 \partial x_1} - x_2 \frac{\partial^2 A_0}{\partial x_1^2} \right) = - i h \frac{e}{c} \frac{h}{Mc} \epsilon_{01} \frac{\partial A_0}{\partial x_2}$$

e da altre analoghe.

Nelle (6') il 2° termine al 2° membro si ottiene con un procedimento analogo

$$= - \frac{e}{c} \frac{h}{Mc} \epsilon_{01} \left( - \frac{\partial A_0}{\partial x_1} \right) \frac{1}{2} h (-i \alpha_1 \alpha_2) - \frac{1}{2} h (-i \alpha_1 \alpha_2) \frac{e}{c} \frac{h}{Mc} \epsilon_{01} \frac{\partial A_0}{\partial x_1} =$$

$$= - i h \frac{e}{c} \frac{h}{Mc} \epsilon_{01} \frac{\partial A_0}{\partial x_1} \text{ ecc.}$$

**Fisica.** — *Sullo spettro di oscillazione e rotazione della molecola HD.* Nota di G. C. WICK, presentata (\*) dal Socio E. FERMI

**INTRODUZIONE.** — Si suole chiamare spettro di oscillazione e rotazione di una molecola quella parte del suo spettro che è dovuta a oscillazioni dei nuclei intorno alle loro posizioni di riposo e rotazioni della molecola in blocco, non accompagnate da oscillazioni elettroniche; in altre parole quella parte dello spettro che corrisponde a transizioni in cui non variano i numeri quantici elettronici.

Tale spettro non viene mai osservato in molecole biatomiche omonucleari come H<sub>2</sub>, N<sub>2</sub>, O<sub>2</sub> ecc., fatto che, com'è noto, si spiega mediante semplici considerazioni di simmetria<sup>(1)</sup>. Esso è stato invece osservato in molecole fortemente asimmetriche (molecole polari) come HCl, HBr ecc.

La molecola HD non è rigorosamente omonucleare e deve quindi possedere uno spettro di oscillazione e rotazione; tuttavia è presumibile che data l'alta simmetria della struttura elettronica della molecola, lo spettro sia estremamente debole. Vedremo infatti che lo spettro ha un'origine diversa da quella ordinaria, essendo dovuto alla asimmetria dell'oscillazione dei nuclei (la differenza essenziale tra H<sub>2</sub> e HD è questa: nella molecola di HD il punto di mezzo della congiungente dei nuclei oscilla; si ha quindi, per inerzia, uno scuotimento del sistema elettronico, che produce il momento oscillante); la sua intensità teorica risulta non nulla soltanto in un'approssimazione più elevata di quella di cui ci si contenta ordinariamente nella teoria delle molecole. Essa non è però così piccola da escludere la possibilità di osservare sperimentalmente lo spettro<sup>(2)</sup>.

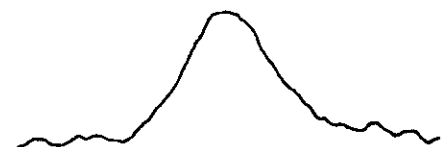
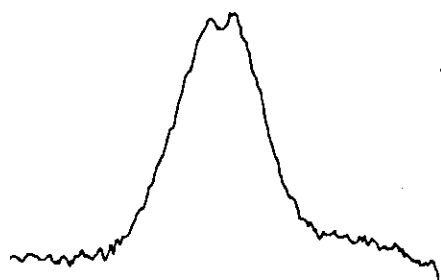
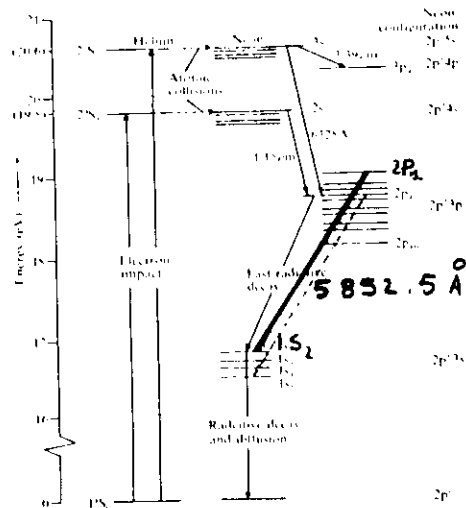
(1) Nella seduta del 19 maggio 1935.

(2) Vede per un orientamento generale, il libro di E. FERMI, *Molecole e cristalli*, la dimostrazione rigorosa — ved. R. DE L. KRONIG, *Band Spectra and Molecular Structure*.

(3) Il dott. H. DEUTLER mi ha gentilmente informato che sono in corso esperienze per questo proposito. Devo anche ringraziare il dott. DEUTLER, avendo io appreso l'esistenza del problema qui trattato da una sua lettera, cortesemente mostratami dal dott. U. FANTAUZZI.

Att: R. Accad. Naz. Lincei  
Mem. Cl. Sci. Fis. Mat. Nat.  
21, 708 (1935)

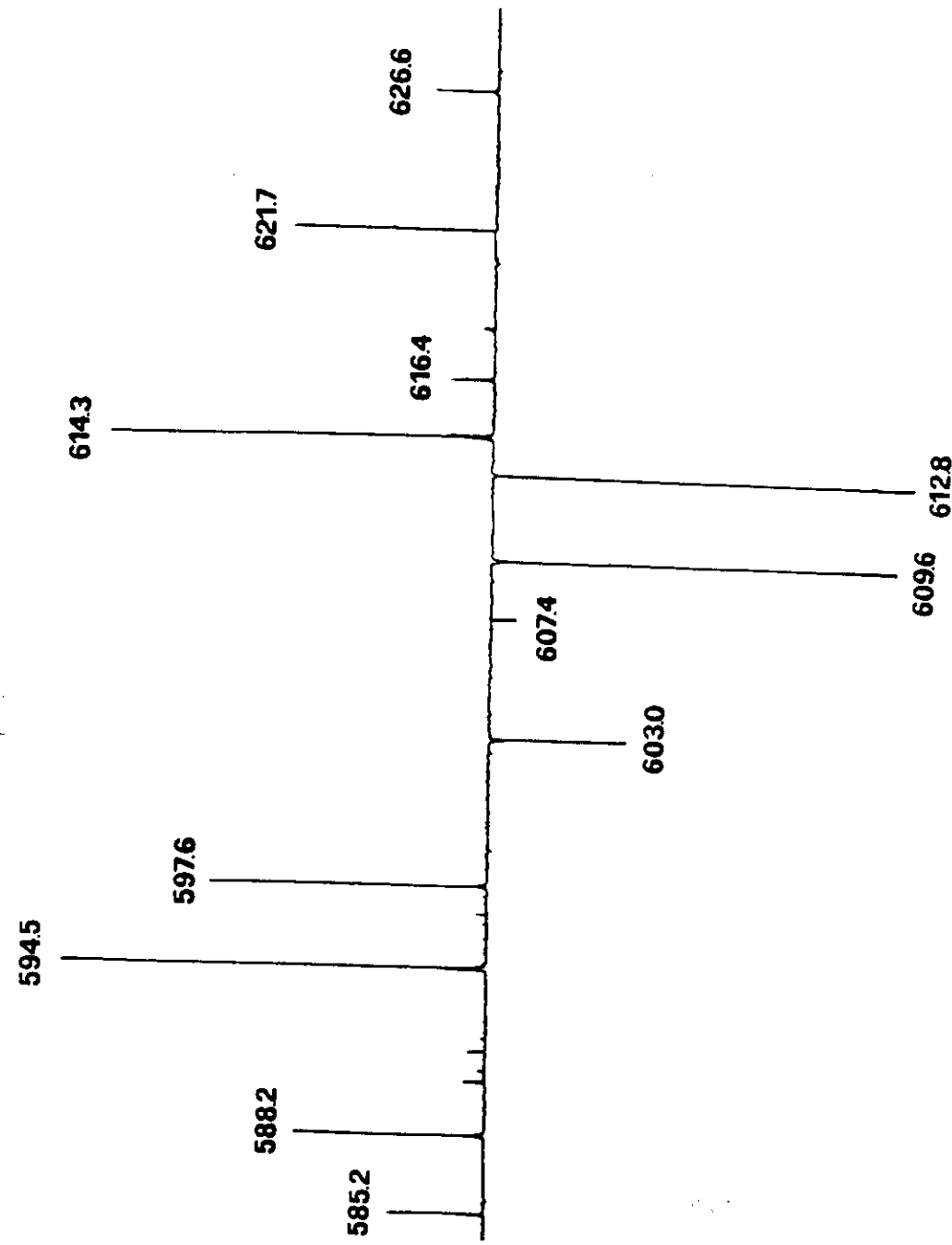
(50)

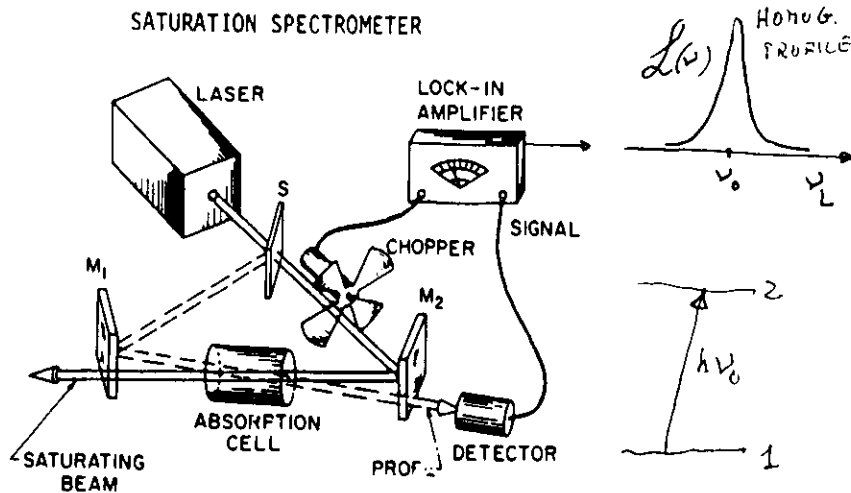
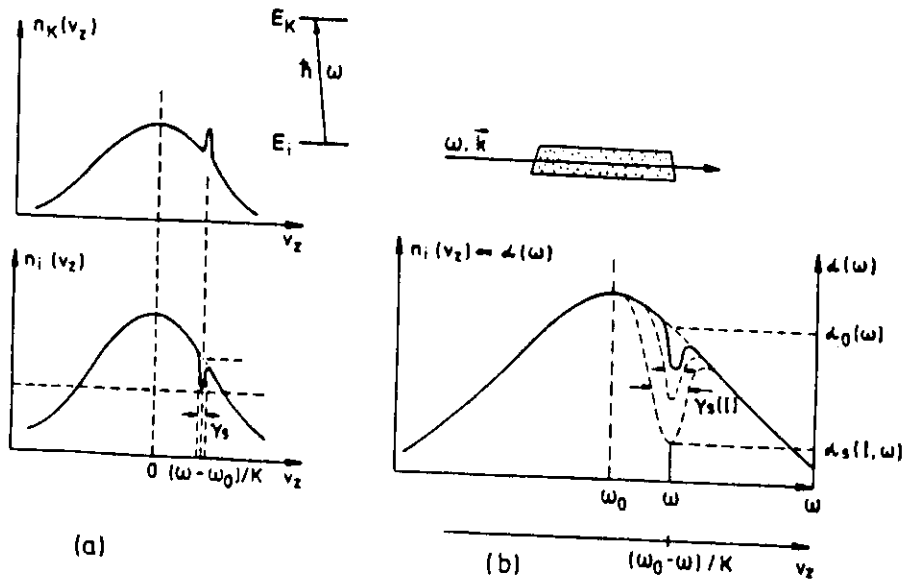


6.6 GHz

b)

a)





Neon  
 $1s_4 - 2p_3$   
 $\lambda = 6074 \text{ \AA}$

Doppler

Saturazione  
 SATURATION

Intermodulato  
 INTERMODULATED

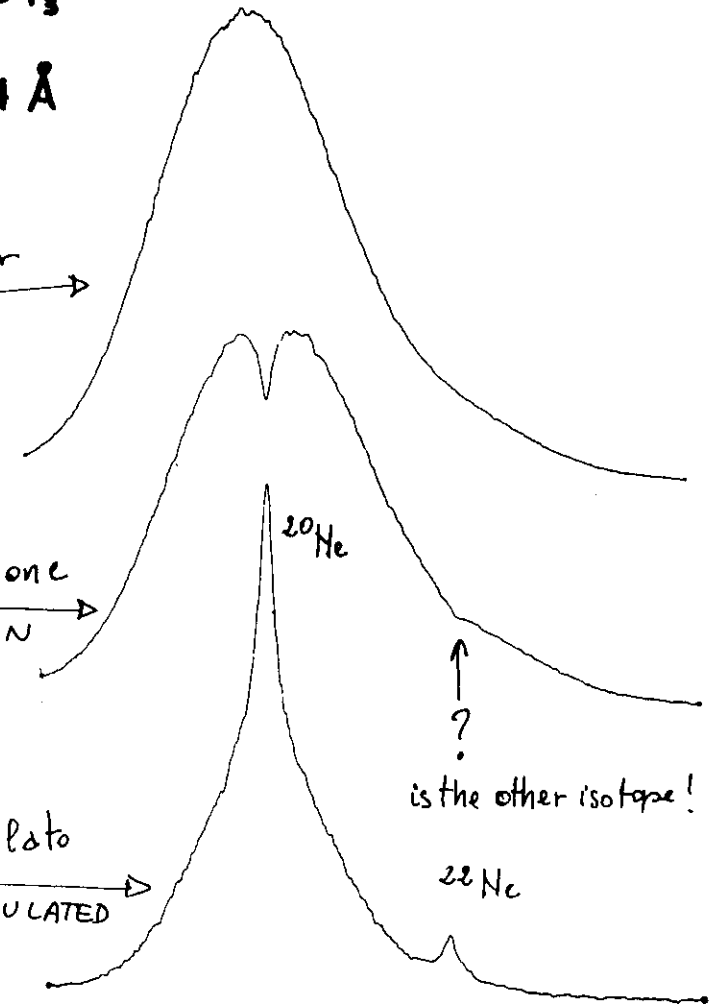


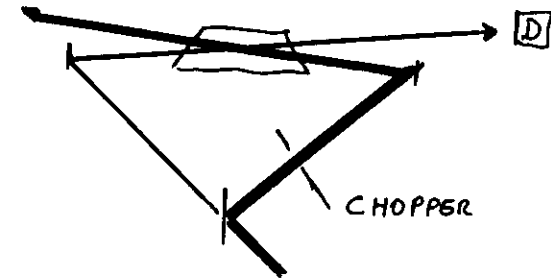
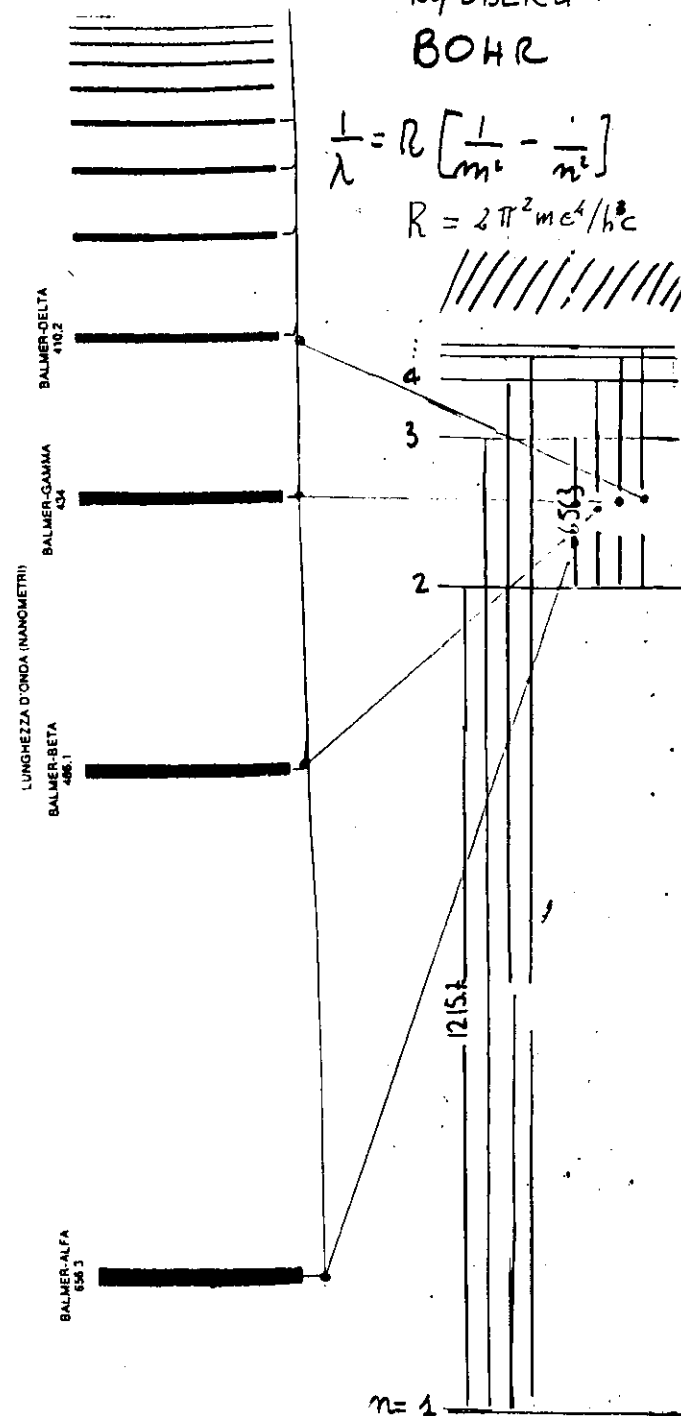
Fig. 25 Neon 6074 A. Cella a catodo cavo;  $I = 48 \text{ mA}$ ,  $P = 1.6 \text{ Torr}$ .

- a) profilo Doppler
- b) profilo con segnale di saturazione
- c) profilo in intermodulazione

RYDBERG  
BOHR

15 B. Scheumann, D. Meschede, T. W. Hänsch, J. L. Hall

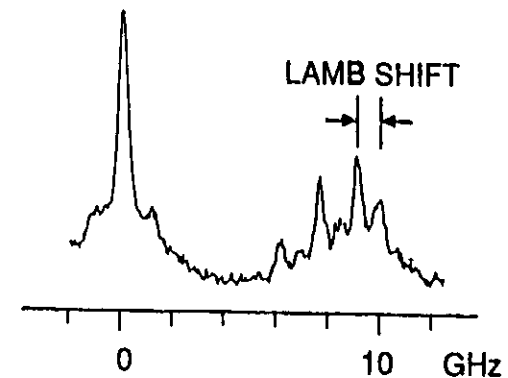
... 20 years later



## HYDROGEN BALMER- $\alpha$ LINE

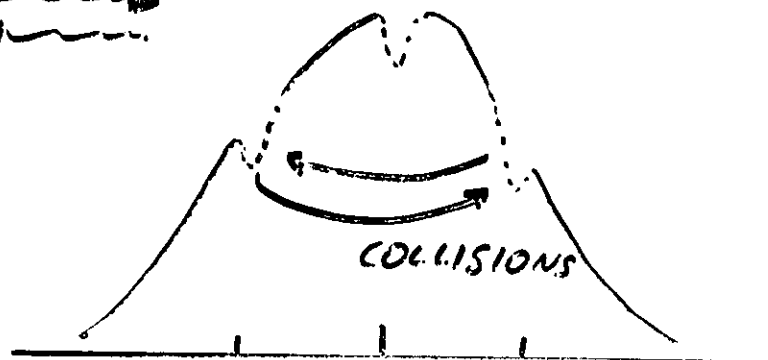
Doppler-free saturation spectrum  
recorded with free running diode laser

(InGaAlP laser diode  
Toshiba TOLD9211,  
T=-70°C, 5 mW at 656 nm)  
Index guided



# VELOCITY CHANGING COLLISIONS

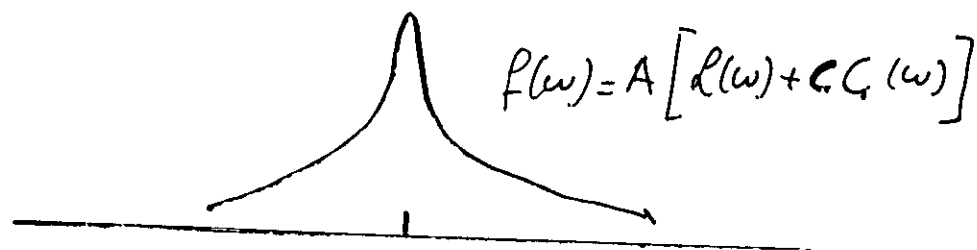
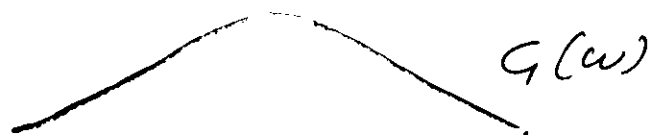
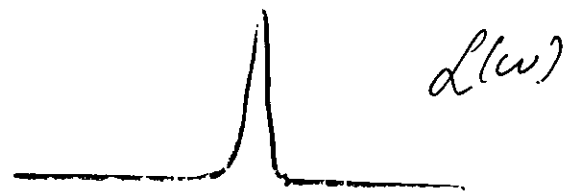
18



$$A \left[ \frac{\gamma^2/4}{-\gamma^2/4 + (\nu - \nu_0)^2} + C \exp\left(-\frac{(\nu - \nu_0)^2}{\delta \nu_D^2}\right) \right] \times \gamma$$

$$C = 2\sqrt{\pi} \log 2 \frac{\Gamma \nu}{\delta \nu_D}$$

$O_2/Ar$  (.1/5 Torr)



$$C \sim 2\sqrt{\pi} \log 2 \frac{\Gamma \nu}{\delta \nu_D}$$

$$\Delta \nu_D = 2 \text{ GHz } (520^\circ \text{K})$$

$$\gamma = \Delta \nu_{\text{hom}} = 386 \text{ MHz (FWHM)}$$

$$C = 0.39$$

K. ERNST, R. RINDIG, A. SASSO, G. TINO, M. I. I.  
Opt. Lett. 14, 554 (88)

0  
645.6 nm  
3<sup>3</sup>P<sub>1</sub> - 3<sup>3</sup>S<sub>1</sub>

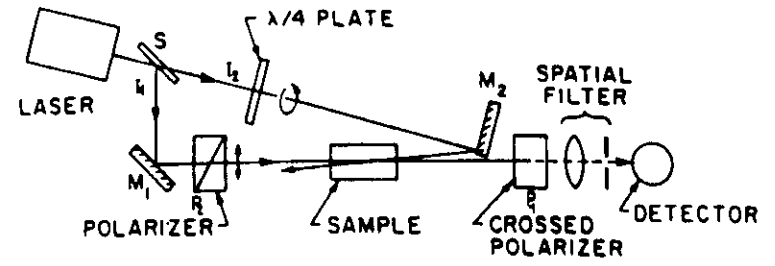
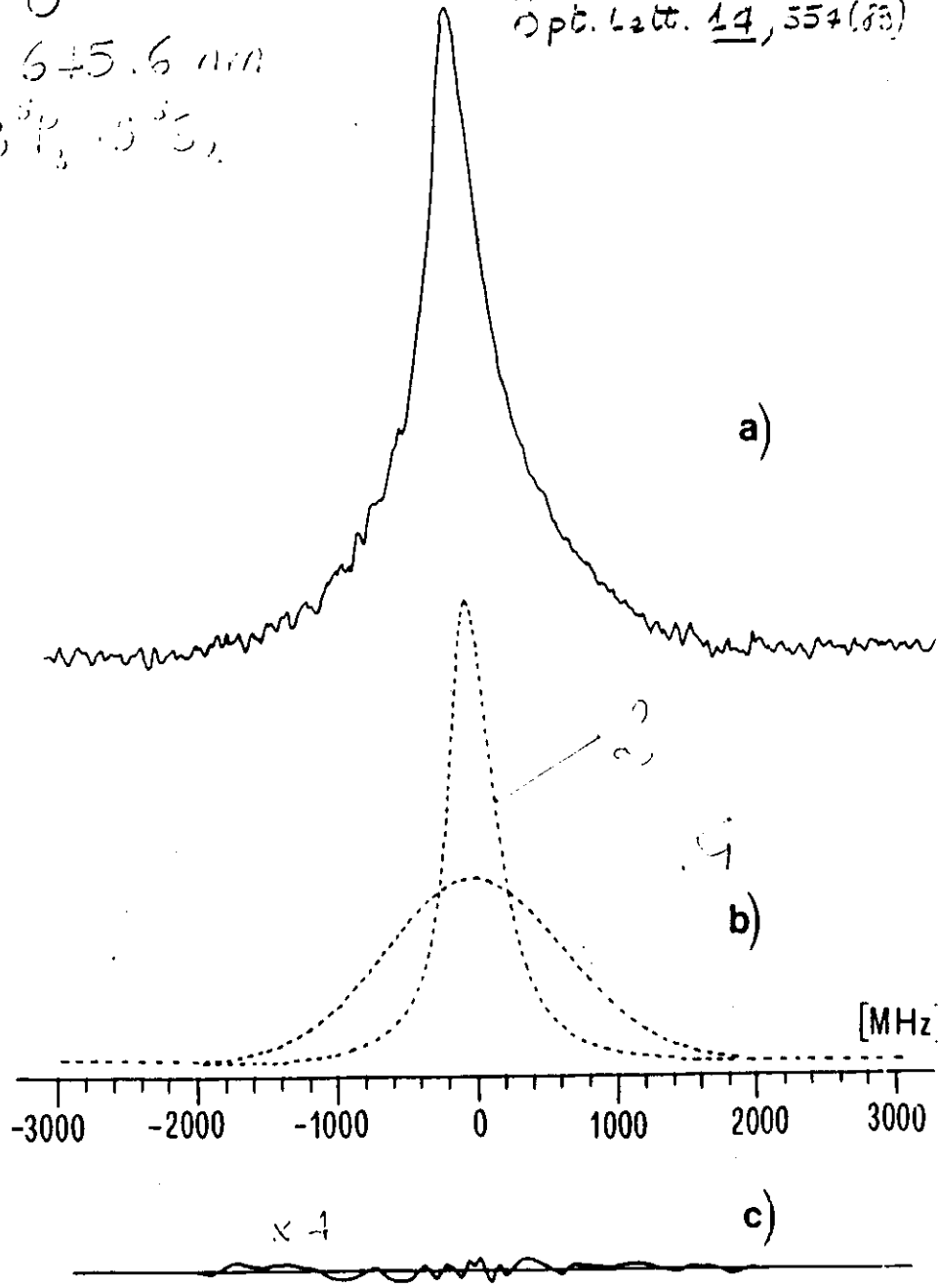
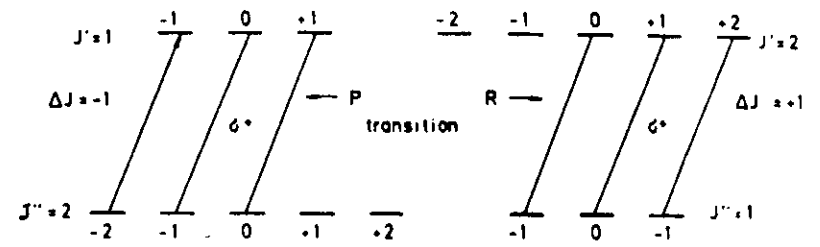
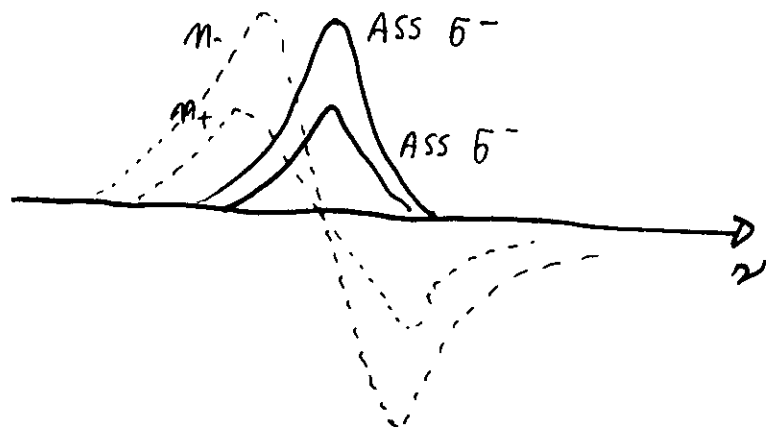
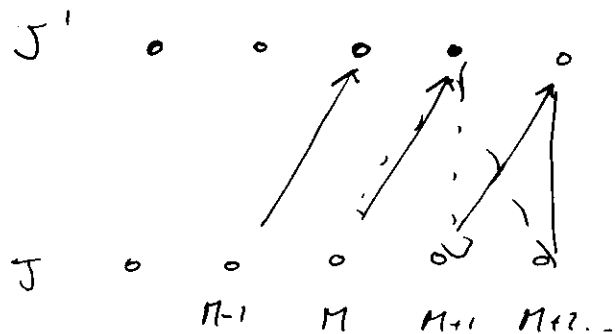


Fig.10.34. Schematic experimental arrangement for polarization spectroscopy

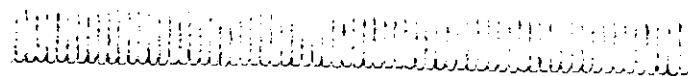
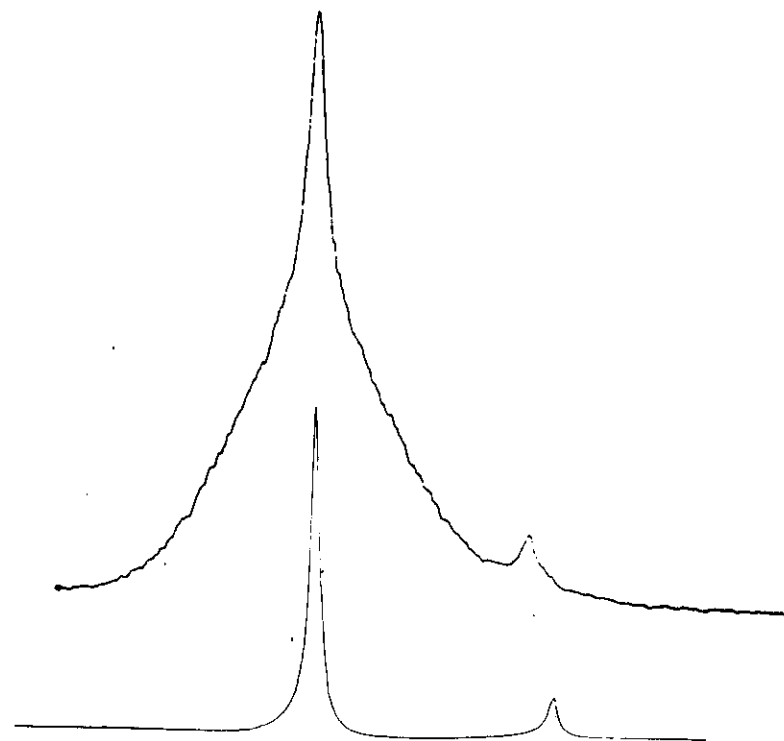
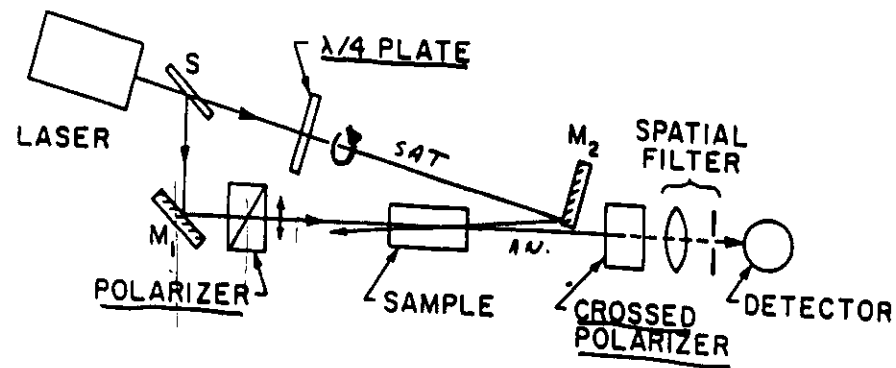




CON LUCE POL. CIRCOLARE  
 PROVOCA UN'ORIENTAZIONE  
 (SAT. SELETTIVA IN M)

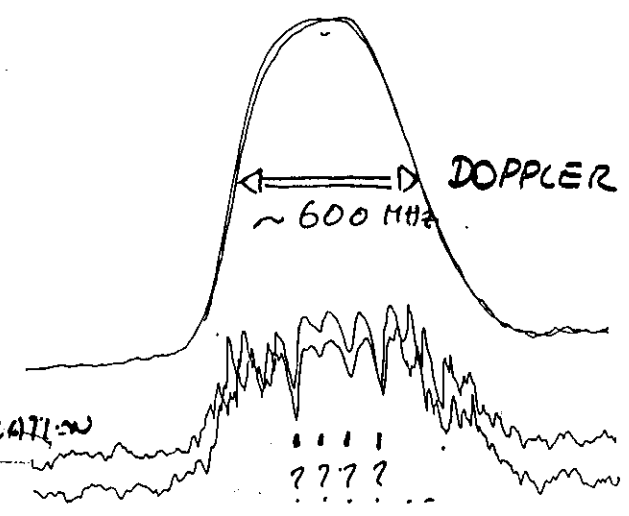
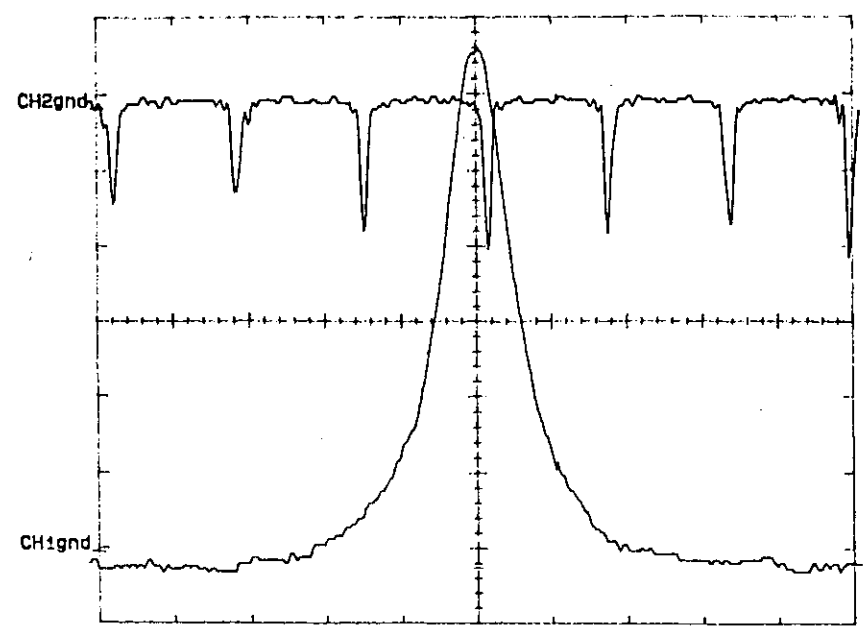


CIRCULARLY POLARIZED LIGHT  
 CREATES AN ORIENTATION  
 (SATURATION SELECTIVE IN M)



# DOPPLER + HYPERFINE

15 comp J even      UNEQUAL INTENSITY  
21 comp J odd      FOR J > 30



21 comp.

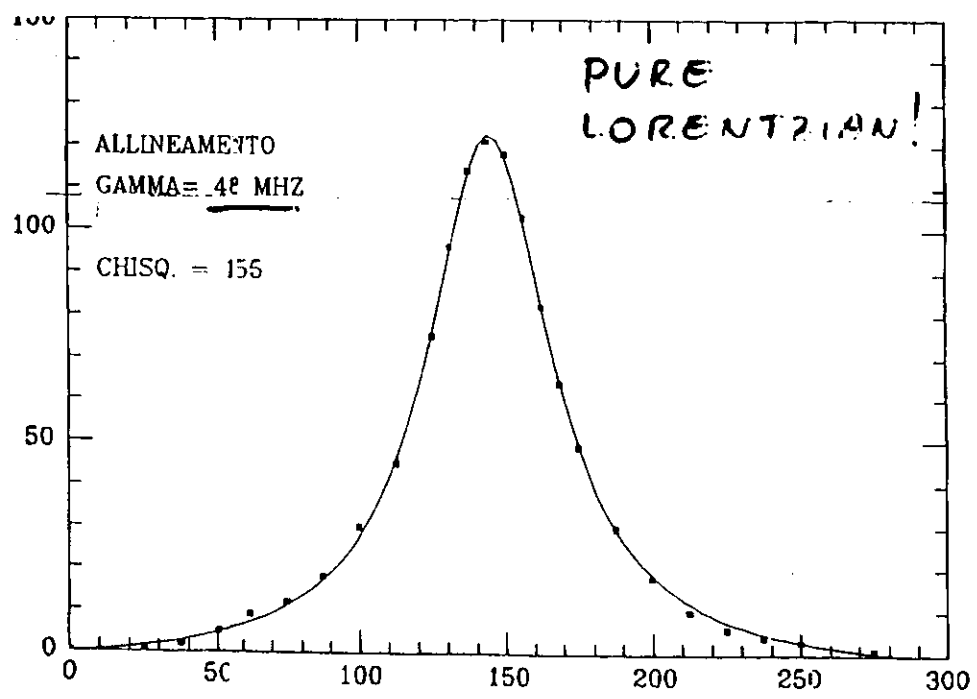
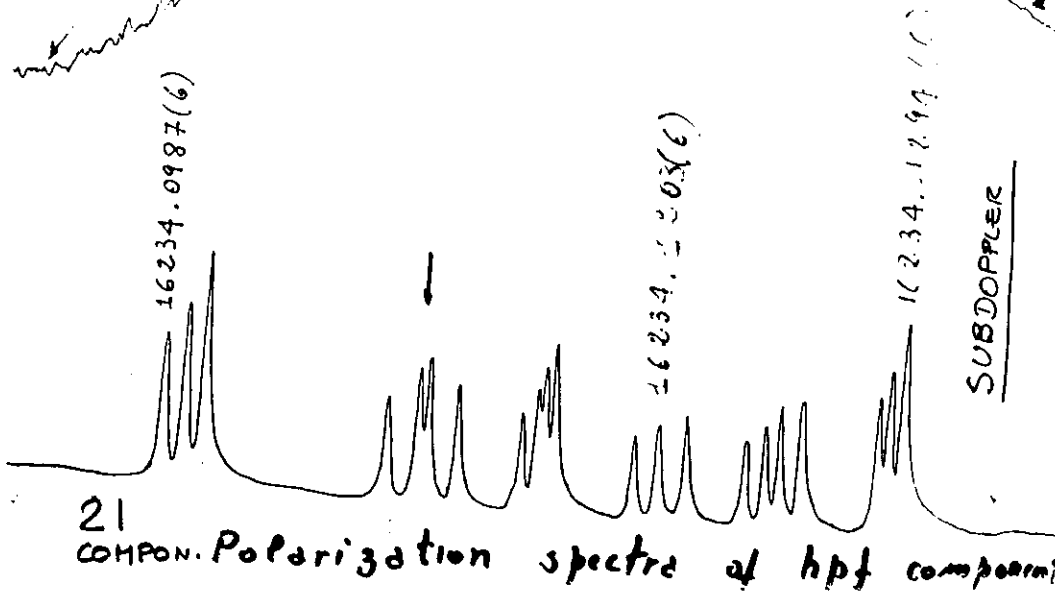
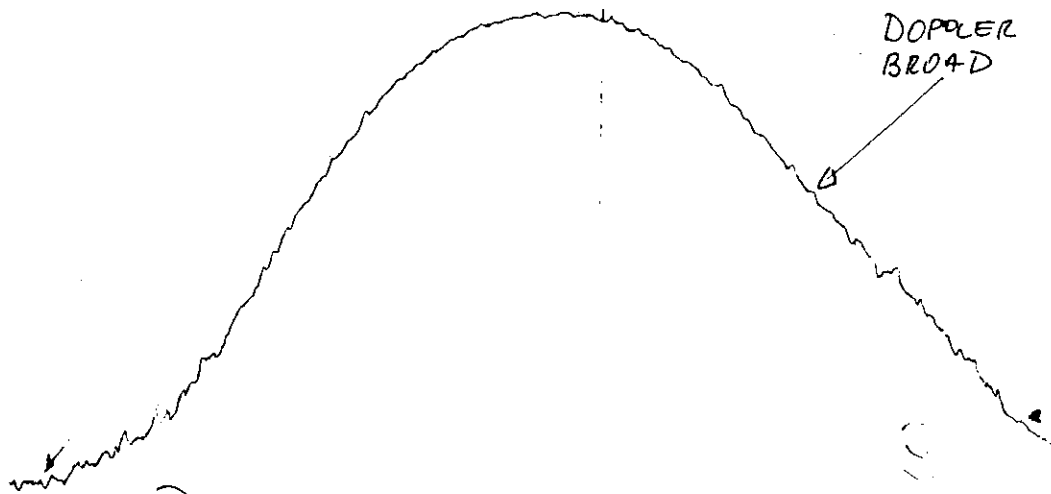
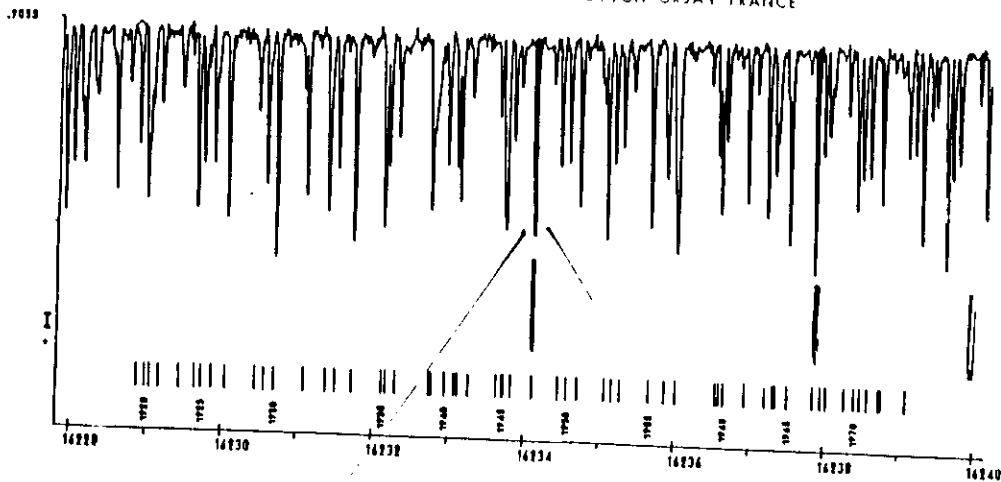


Fig. 22 Struttura iperfina della molecola di Iodio.  
 Profilo di assorbimento con fascio opposto saturante.



## SEMICONDUCTOR DIODE LASER

### DIRECT INTERBAND LASER TRANSITION (GaAl)As

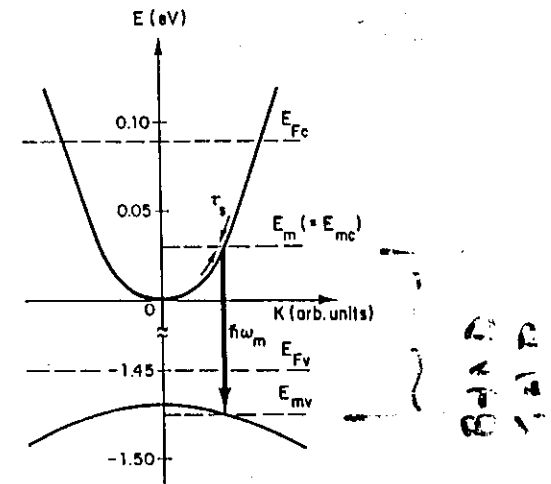
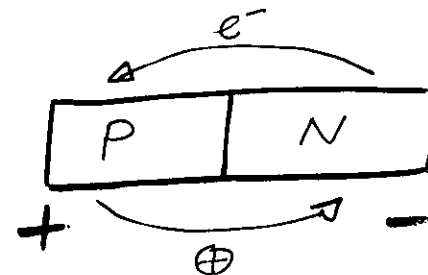
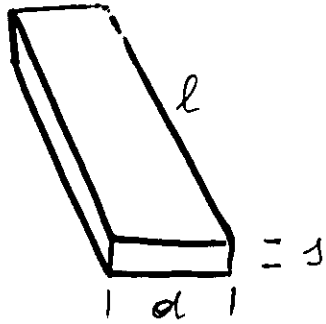


fig. 8 Schematic diagram of the simplified band structure of (GaAl)As showing the direct interband laser transition.



Also InGaP  
InGaAsP

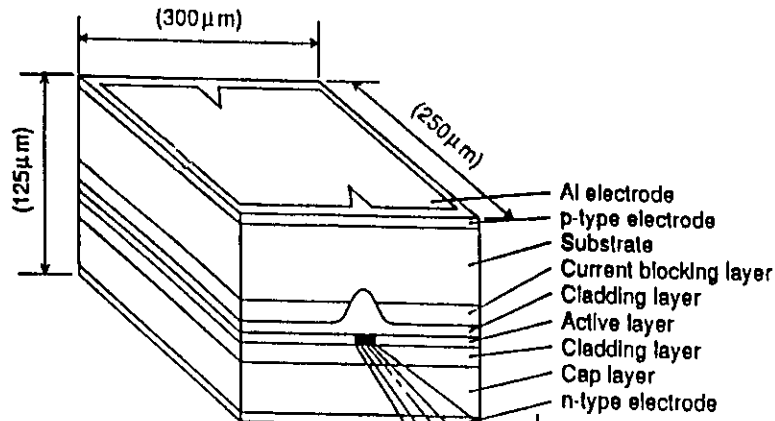
# TYPICAL STRUCTURE



$l$  100 ~ 500  $\mu\text{m}$

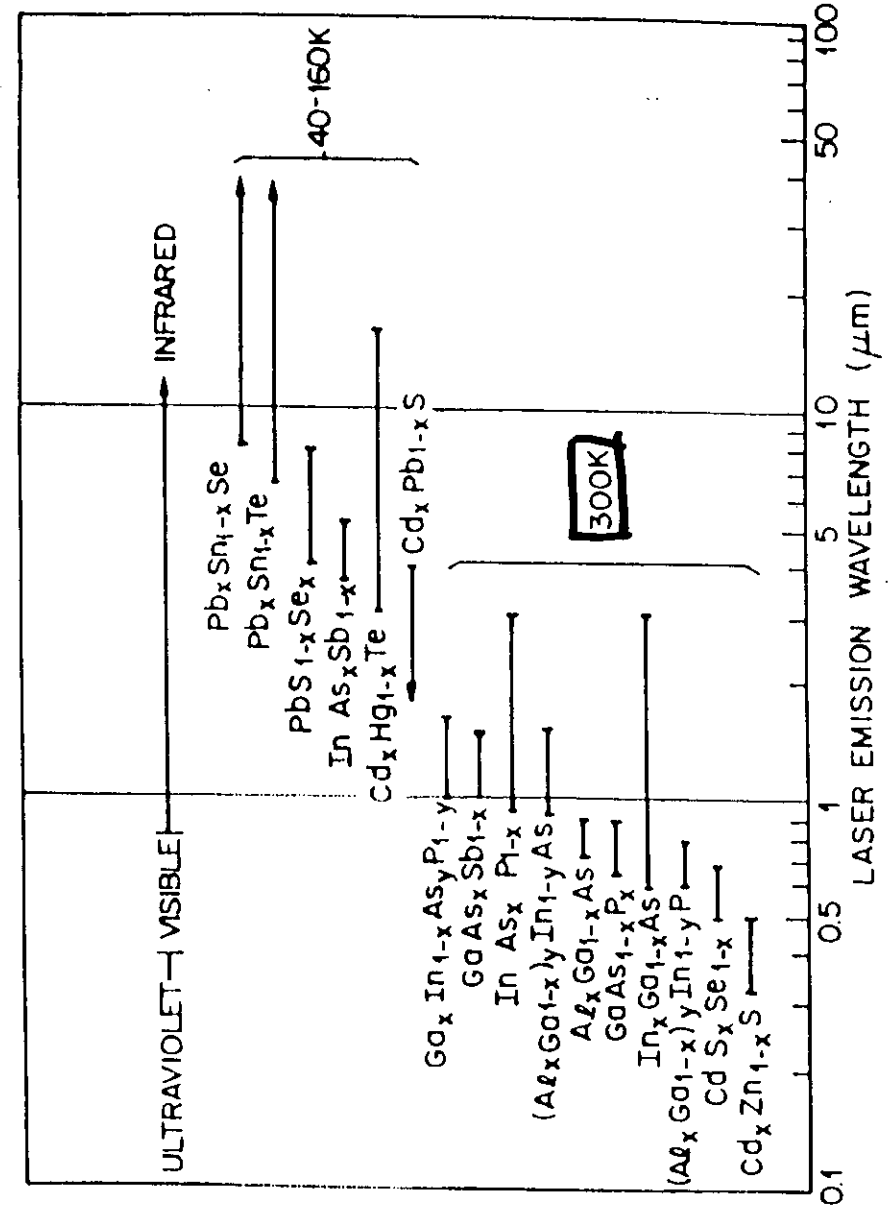
$s$  0.1 ~ 2  $\mu\text{m}$

$d$  2 ~ 20  $\mu\text{m}$

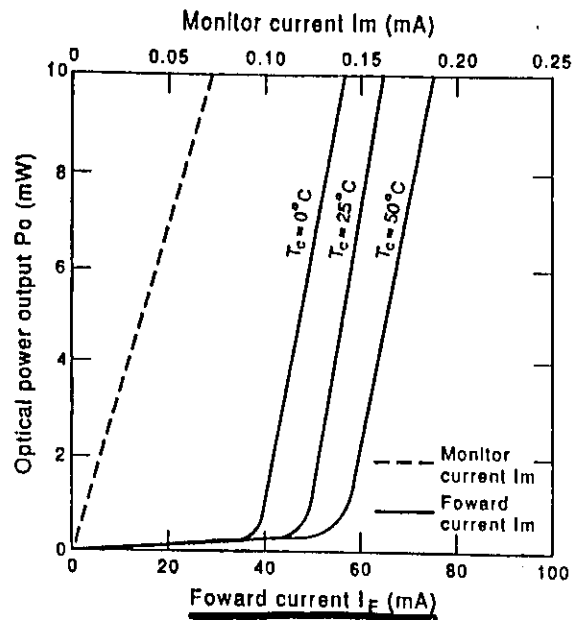


Note: Chip size may be varied

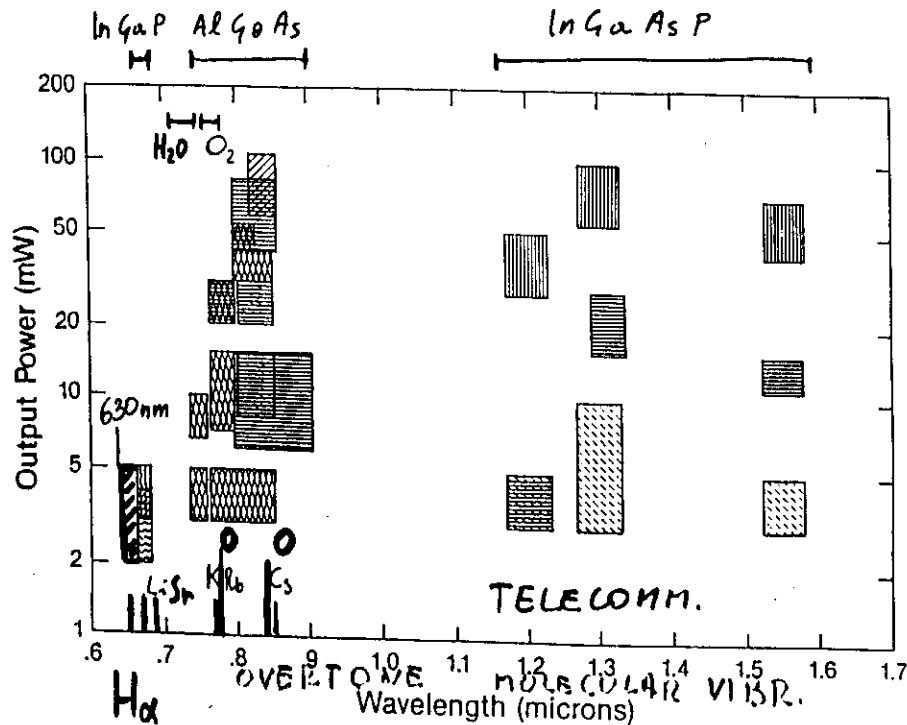
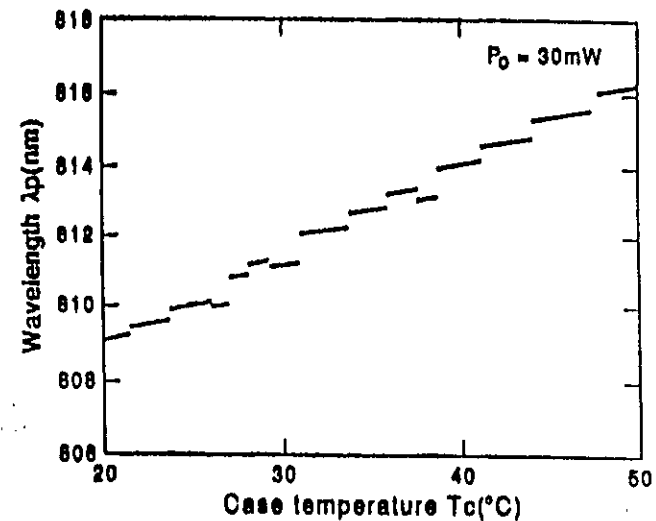
Radiation pattern



Emission wavelengths of direct semiconductors



## STAIRCASE WITH SLOPING STEPS



SLOPE OF EACH STEP:  $.06\text{ nm/K}$   
TUNING OF CAVITY MODE ( $10\text{ GHz}$ )

JUMP BETWEEN STEPS:  $.25\text{ nm/K}$   
LONGITUDINAL MODE HOPPING  
BECAUSE OF GAIN CURVE SHIFT.

BANDWIDTH OF THE GAIN  $\sim 40\text{ nm}$   
LONGITUDINAL MODE SPACING:  
 $0.3\text{ nm} \sim 10.0\text{ GHz}$

# STABILITY REQUIREMENTS

## CURRENT mA

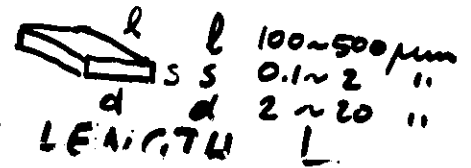
$\sim 3 \text{ GHz} / \text{mA}$  JOULE EFFECT  
SLOW

$\sim 300 \text{ MHz} / \text{mA}$  REFRACTIVE INDEX  
FAST  
(frequency chirping)

## TEMPERATURE mK

$\sim 10 \text{ GHz} / ^\circ\text{C}$

## LINEWIDTH



VERY SHORT CAVITY LENGTH

SIGNIFICANT BROADENING DUE TO  
THE QUANTUM FLUCTUATIONS OF  
PHASE (Schawlow-Townes)

$$2\Gamma = (\pi h \nu \Gamma_c^2 / P) n_{sp} \quad \text{pop. factor}$$

↑  
intrac. power

$$\Gamma_c \propto 1/L$$

PASSIVE  
RES. LINEW.

$$\Delta\nu \sim 10 \text{ MHz}$$

... for high resolution spectroscopy  
we need a narrower and  
continuously tunable source

## FEEDBACK

ELECTRONIC (MUST BE VERY FAST)

## OPTICAL

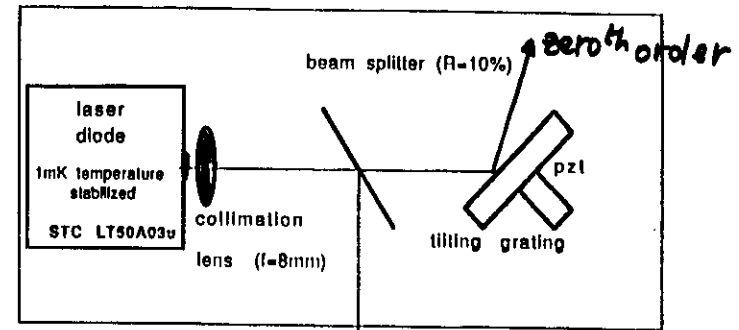
LONGER (NARROWER  $\Gamma_c$ ) CAVITY  
BECAUSE IT IS COUPLED WITH  
AN EXTERNAL "REFLECTOR,"

REFLECTORS: MIRRORS, FIBERS,  
ETALONS, FABRY PEROTS, GRATINGS

See for instance:

C. Wieman, L. Hollberg

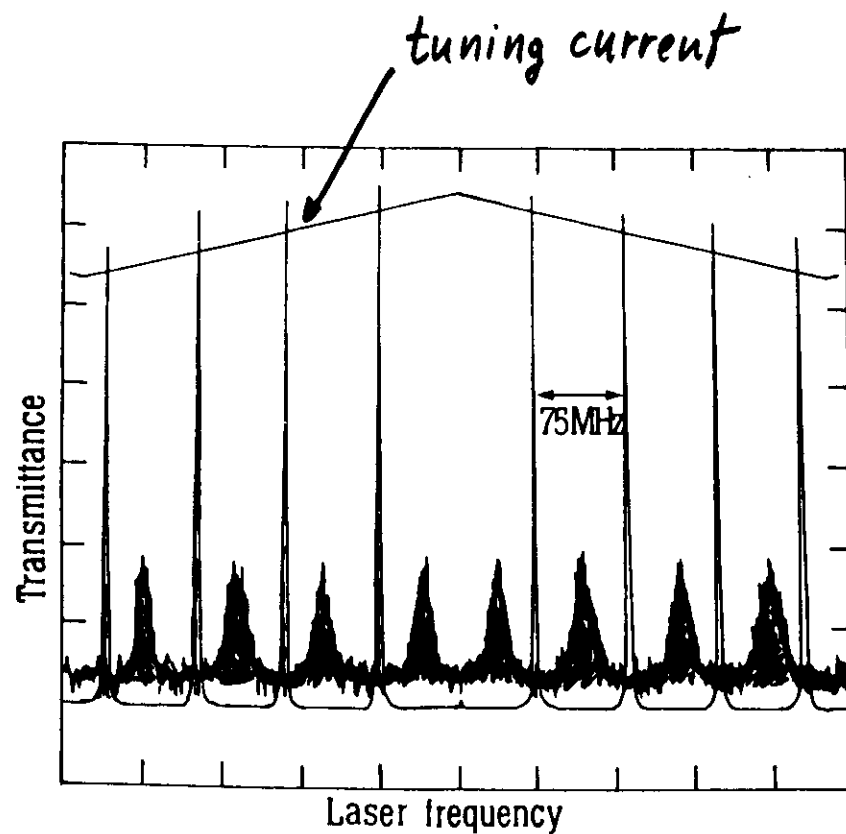
Rev. Sci. Instr. 1991



also "regular" heterostructure lasers

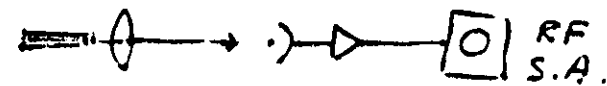
also "red" InGaP lasers

InGaP laser at 690 nm



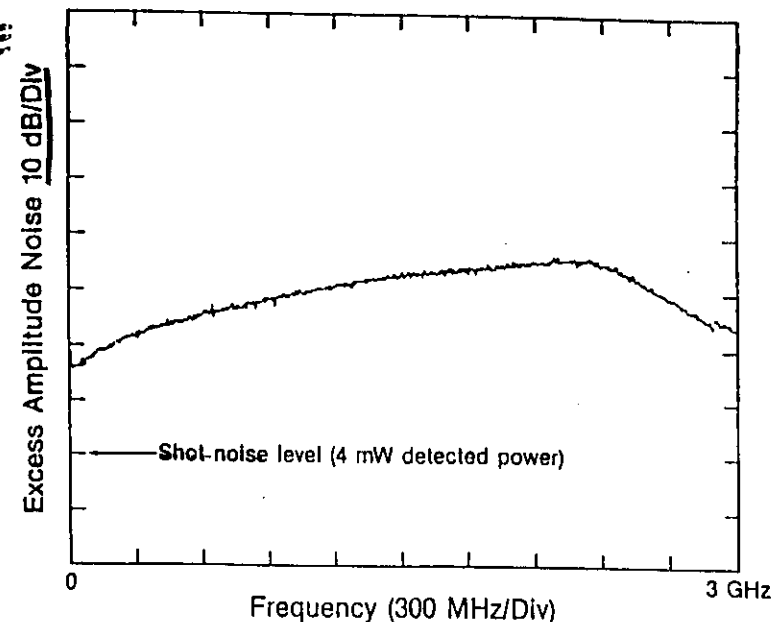
G. TINO, M. de Angelis, L. Gianfroni,  
M. Barsanti, M. I.  
Appl. Phys. B (1992) (NAPOLI + FIRENZE)

## AMPLITUDE NOISE

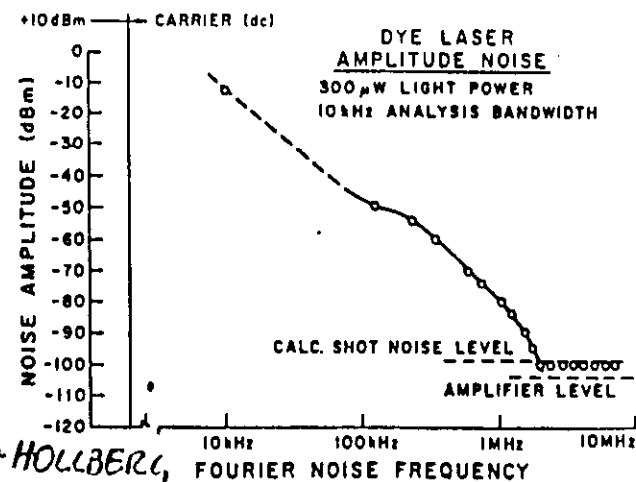


WIENAN + HOLBERG

DIODE



DYE



63

HALL + HOLBERG



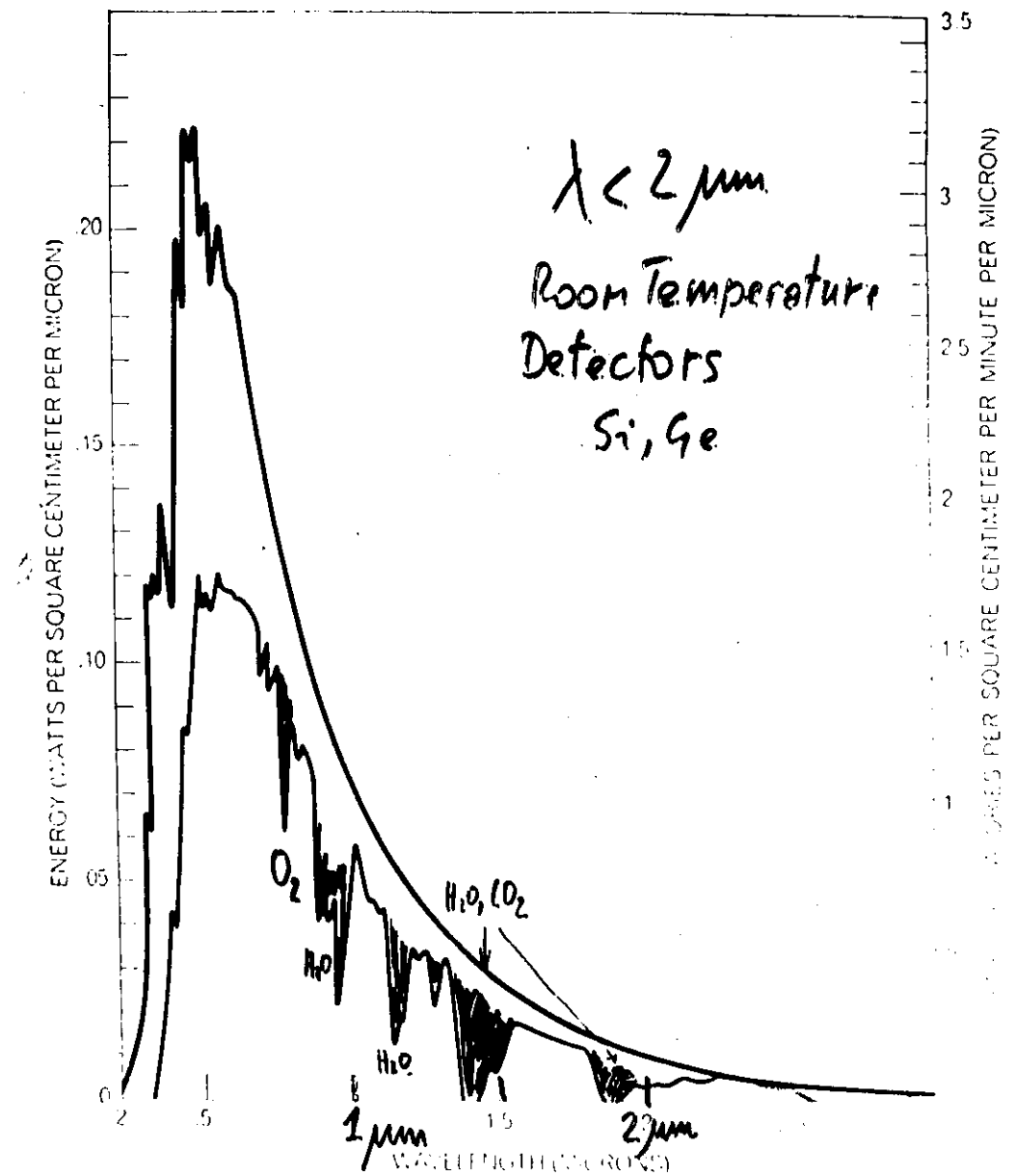
SHOT NOISE  $\propto \sqrt{N_{\text{phot}}}$

800 nm, 5 mW  $\sim 10^{-8}$

here we are a factor  
10-20 db worse, so  
we expect  $\frac{\Delta I}{I} \sim 10^{-6}$

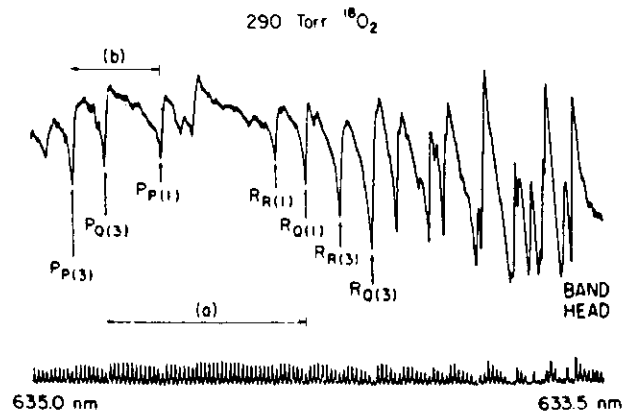
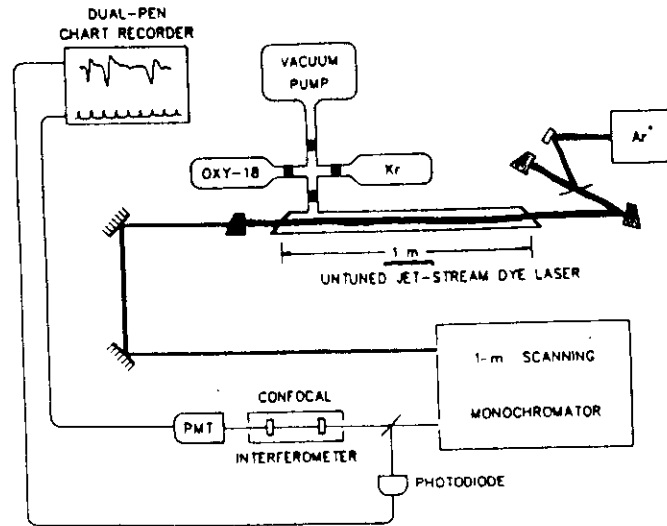


HIGH SENSITIVITY IN PURE  
ABSORPTION EVEN FOR VERY  
WEAK TRANSITIONS



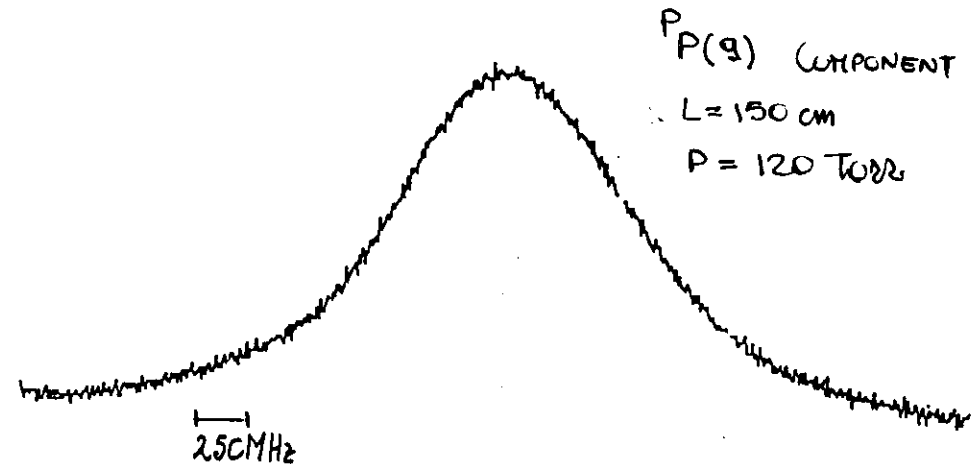
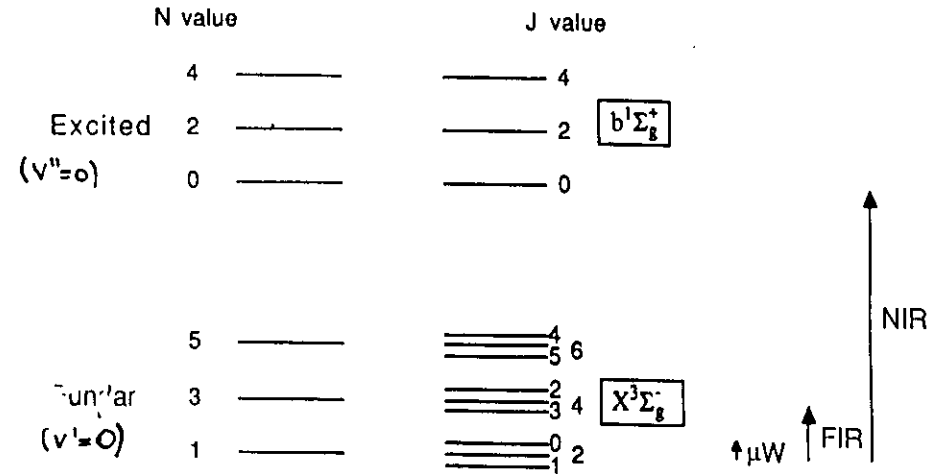
SPECTRAL DISTRIBUTION of solar radiation reaching the earth

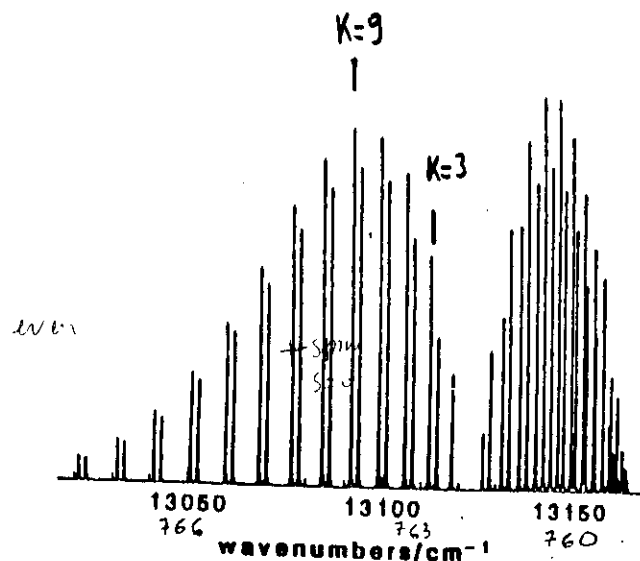
# INTRACAVITY (BROAD-DYE LASER)



Hänsch, Schawlow

# DETECTION OF MOLECULAR OXYGEN ( $\text{K}_{\text{L}_2}$ )<sup>8</sup> AT 760 nm



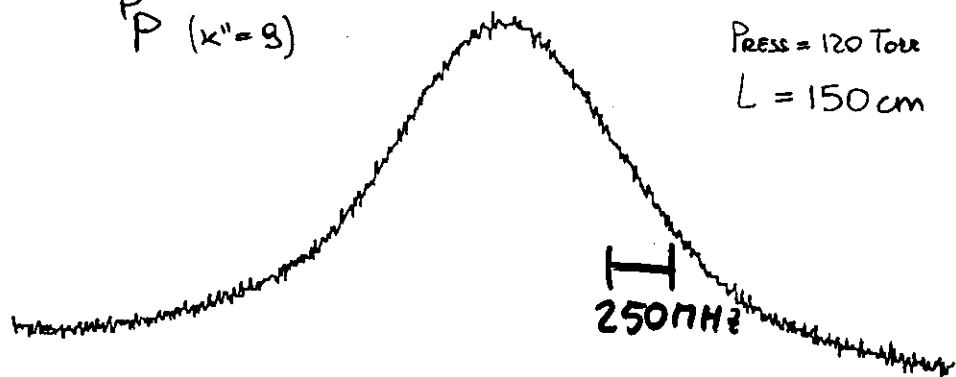


odd

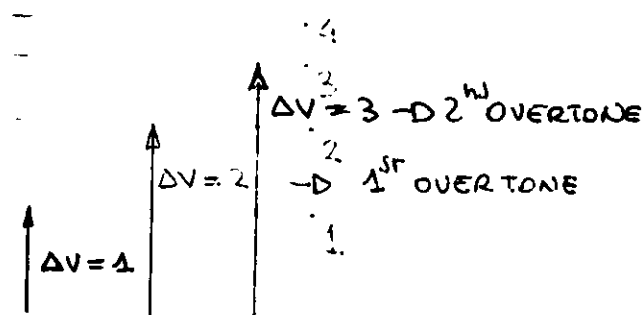
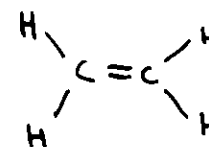
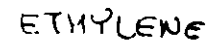
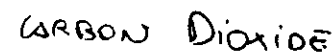
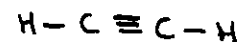
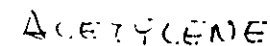
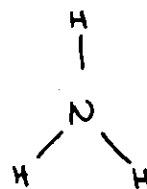
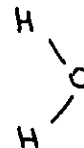
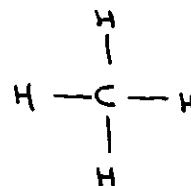
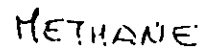
+ spin  
S=1

P (x''=9)

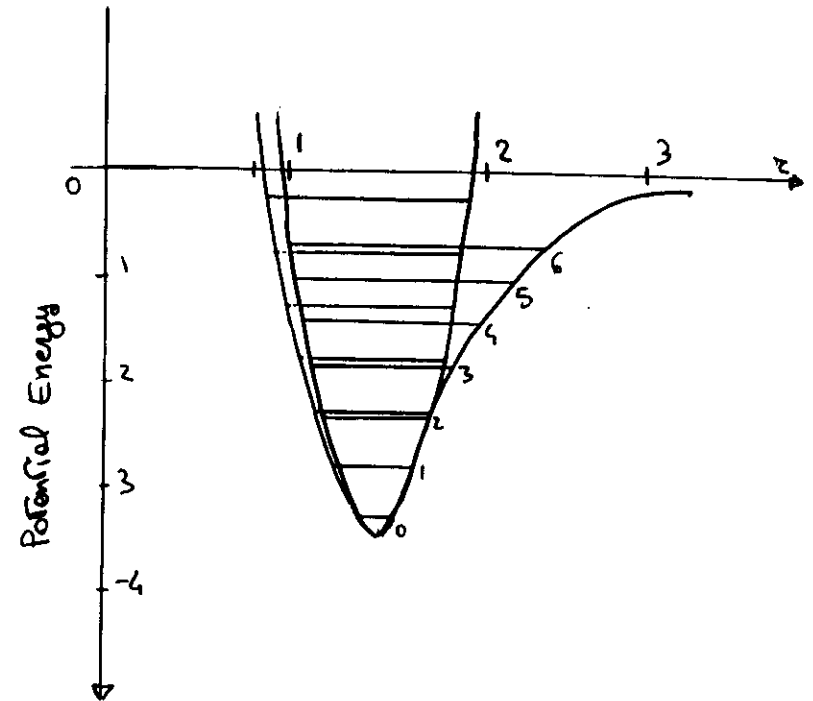
Press = 120 Torr  
L = 150 cm



Preparing molecules in the singlet,  
Pressure broad....



| MOLECULE                      | $\nu_{cm^{-1}}$ | $\lambda_{nm}$ | TRANSITION  |
|-------------------------------|-----------------|----------------|---|
| CO <sub>2</sub>               | 12672.4         | 789            | 00°0 → 01°5   |
|                               | 12774.7         | 782            | 00°0 → 10°5   |
| HCN                           | 11674.4         | 856            | 00°0 → 10°3   |
|                               | 12635.9         | 791            | 00°0 → 00°4   |
| H <sub>2</sub> O              | 12151.2         | 822            | 000 → 2.1   |
|                               | 12562.0         | 796            | 000 → 01°3  |
| C <sub>2</sub> H <sub>2</sub> | 11600.1         | 862            | 000 → 01°3  |
|                               | 11782.9         | 848            | 000 → 21°1  |
|                               | 12675.7         | 788            | 000 → 10°3  |
| C <sub>2</sub> HD             | 12263.0         | 815            | 00000 → 30100   |
|                               | 12735.0         | 785            | 00000 → 40010   |
| NH <sub>3</sub>               | 12609.2         | 793            | 4ν <sub>1</sub>   |
| CH <sub>4</sub>               | 11885.0         | 841            | 4ν <sub>3</sub>   |
|                               | 12755.0         | 782            | 2ν <sub>1</sub> + ν <sub>2</sub> + 2ν <sub>3</sub>                                      |
| HC≡C-C≡H                      | 12706.2         | 787            | ν <sub>1</sub> + 3ν <sub>4</sub>  |
| C <sub>6</sub> H <sub>6</sub> | 11500           | 869            | Δν = 4  |
| HI                            | 11950           | 836            | Δν = 6  |
| HCP                           | 13300           | 750            | Δν = 5  |
| IBr                           | 12576           | 795            | A <sup>3</sup> Π <sub>1</sub> ← X <sup>1</sup> Σ <sup>+</sup>                           |
| HOOH                          | 12600           | 793            | 5ν <sub>OH</sub> + ν <sub>OO</sub>  |
| C <sub>2</sub> D <sub>2</sub> | 12111           | 825            | 000 → 500   |
|                               | 12628           | 791            | 000 → 302   |
|                               | 11946           | 837            | 000 → 401   |
|                               | 12347           | 810            | 000 → 203   |
| CH <sub>3</sub> D             | 11602           | 861            | 2ν <sub>1</sub> + 2ν <sub>2</sub>   |
| O <sub>2</sub>                | 13100           | 763            | b <sup>1</sup> Σ <sub>g</sub> <sup>+</sup> ← x <sup>3</sup> Σ <sub>g</sub> <sup>-</sup> |



HARMONIC MOTION

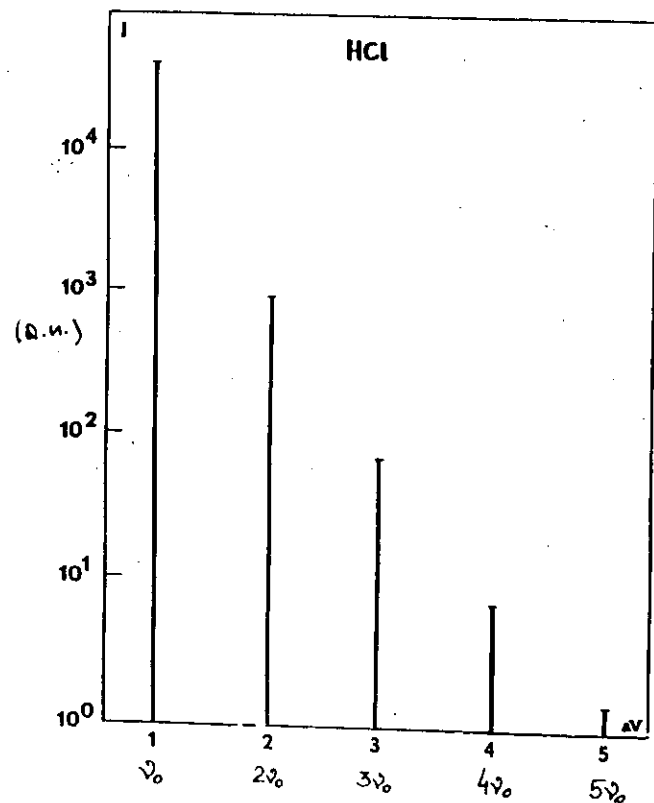
$$x = x_0 \sin 2\pi \nu t$$

$$E(\nu) = h\nu_0 \left(\nu + \frac{1}{2}\right)$$

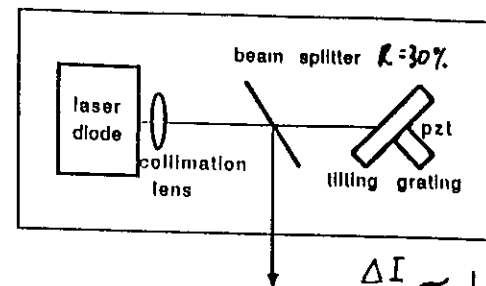
ANHARMONIC MOTION

$$E(\nu) = h\nu_0 \left(\nu + \frac{1}{2}\right) - h\nu_0 x_e \left(\nu + \frac{1}{2}\right)^2 + h\nu_0 y_0 \left(\nu + \frac{1}{2}\right)^3 + \dots$$

## TRANSITION STRENGTH



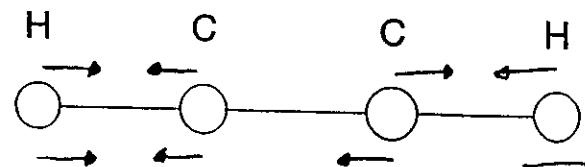
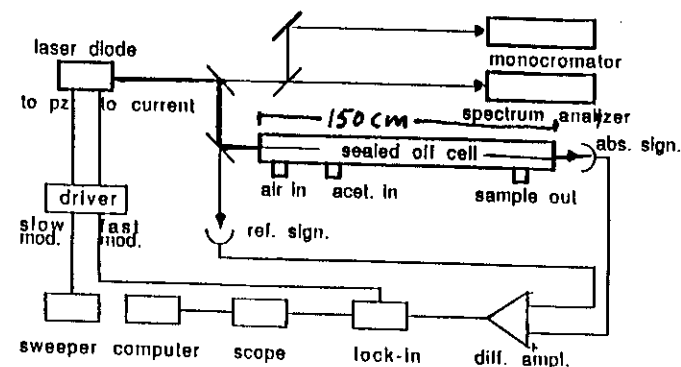
## ACETYLENE



$$\frac{\Delta I}{I} \approx 10^6 \sqrt{\text{Hz}}$$

at 1 KHz

F. PAVONE, K. ERNST, F. MARIN  
G. D. LONARDO and M. I. APPL. OPT.



$$\nu_{CH} \sim 3300 \text{ cm}^{-1}$$

$$1\nu_1 + 3\nu_3 \sim 4\nu \sim 13000 \text{ cm}^{-1} \rightarrow 780 \text{ nm}$$

(BAND HEAD  $12676 \text{ cm}^{-1}$ )

(OTHER 3 OVERTONE BANDS IN THE DIODE LASERS RANGE)

DERIVATIVE SIGNAL

COMPONENT P(11)

a)



PRESS.  $C_2H_2 = 10 \text{ Torr}$

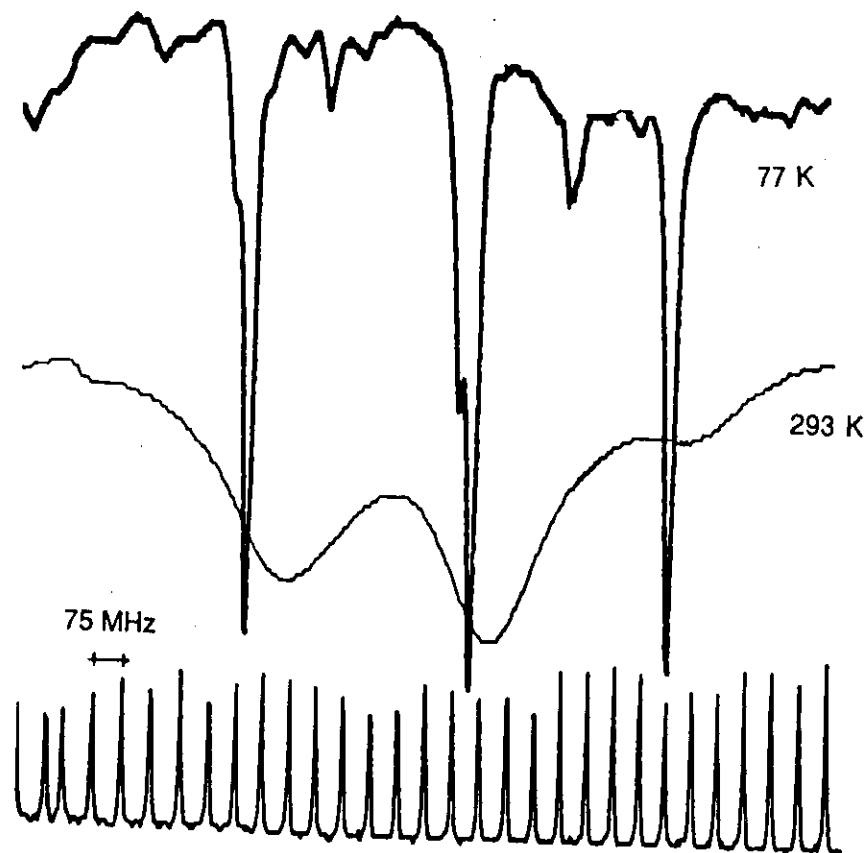
b)



PRESS.  $C_2H_2 = 30 \text{ mTorr}$

# 3rd OVERTONE METHANE ABSORPTION

$\lambda = 886 \text{ nm}$



Components of the combination bands

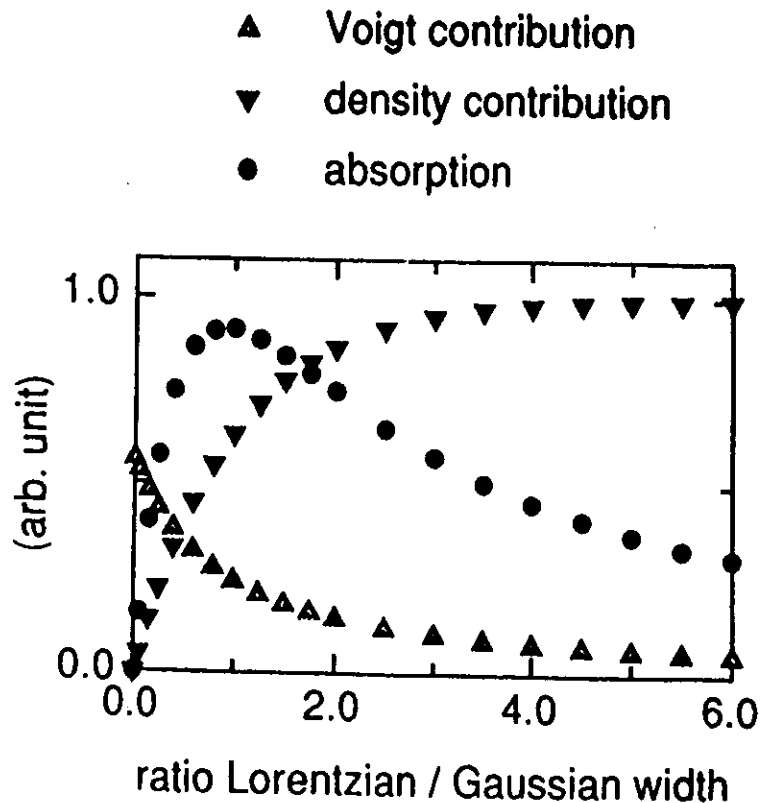
$3\nu_1 + \nu_3$

$2\nu_1 + 2\nu_3$

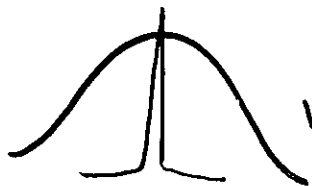
$\nu_1 + 3\nu_3$

$4\nu_3$

# STUDY OF THE BEHAVIOUR OF THE ABSORPTION WITH THE PRESSURE



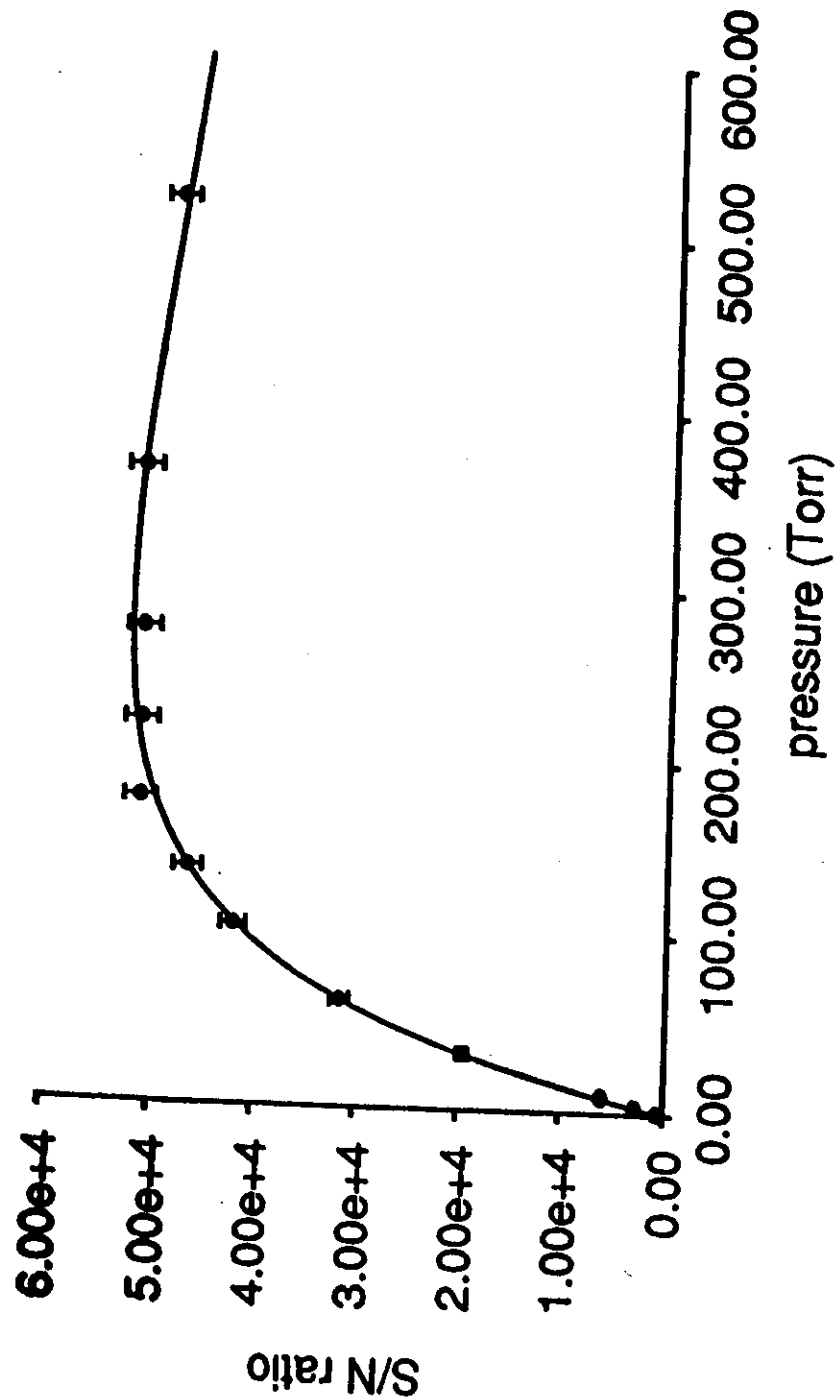
LASER (PEAK ABSORPTION)



VOIGT: GAUSSIAN + LORENTZIAN

WITH PRESSURE, THE LORENTZIAN CONTRIBUTION INCREASES WITH A BROADENING OF THE LINE AND A DECREASE OF ITS PEAK.

LWM con medie

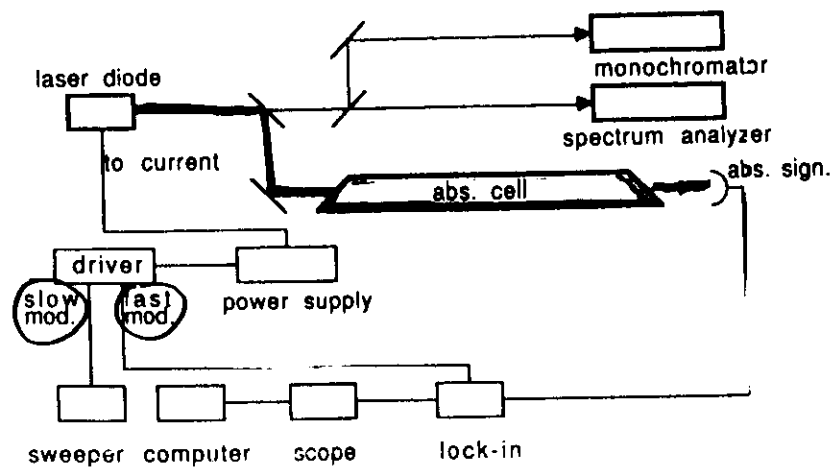


1<sup>st</sup> METHOD :

# LOW WAVELENGTH MODULATION

MODULATION FREQUENCY 1 KHz

(SMALLER THAN MOLECULAR LINEWIDTH AND THE ABSORPTION IS PROBED BY MANY SIDE BANDS)



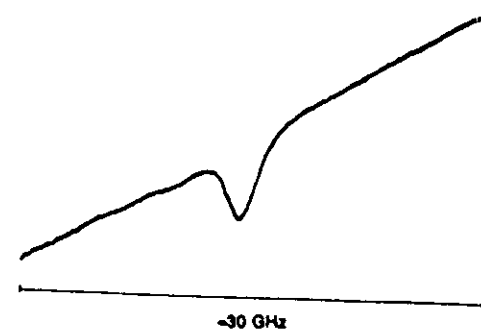
$$\alpha_{\min} = 4.5 \cdot 10^{-7}$$



CH<sub>4</sub> 150cm 0.886 μm

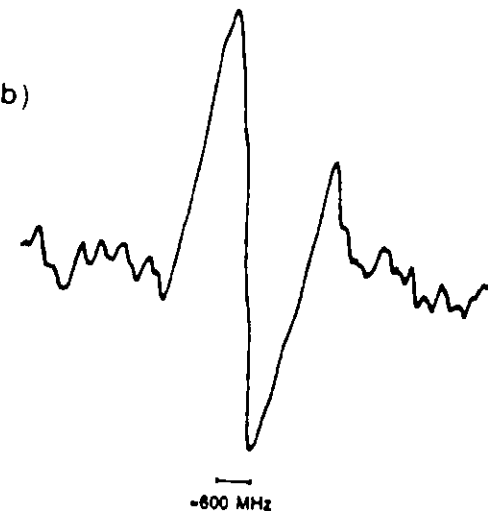
Direct

a)



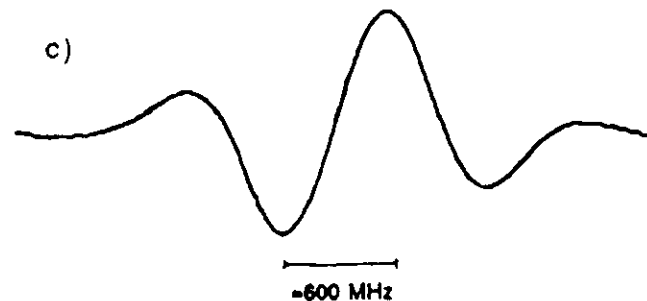
3<sup>rd</sup> Der

b)



TTFM

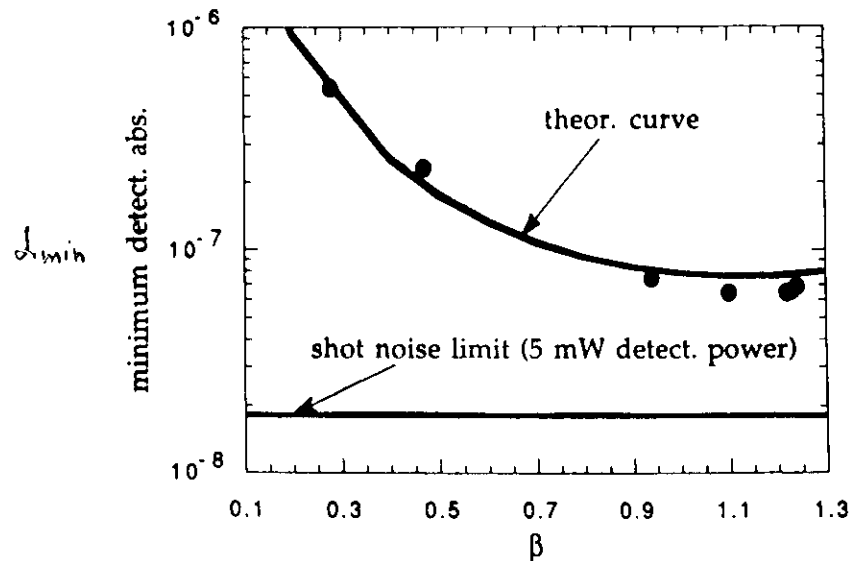
c)



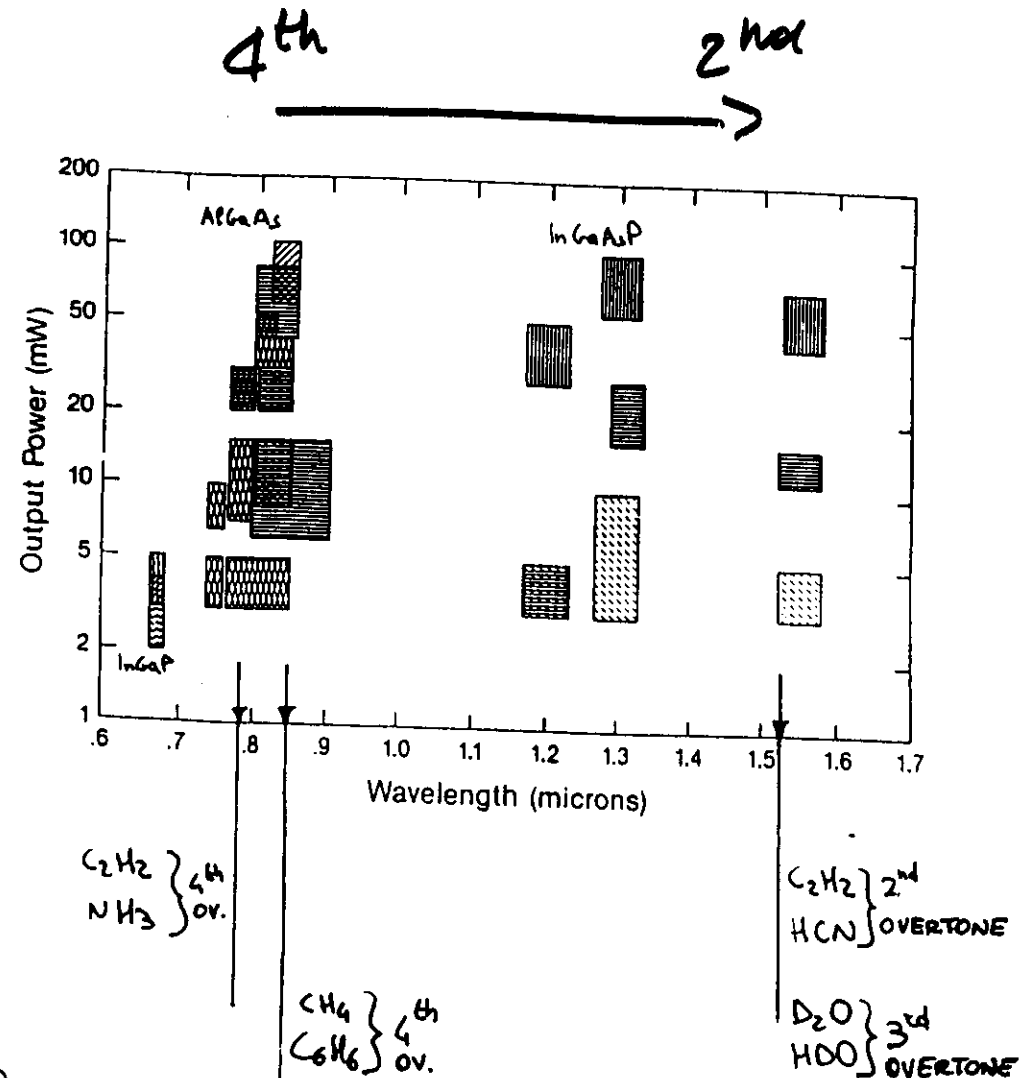


# NEAR INFRARED SPECTROSCOPY OF MOLECULES USING DIODE LASERS

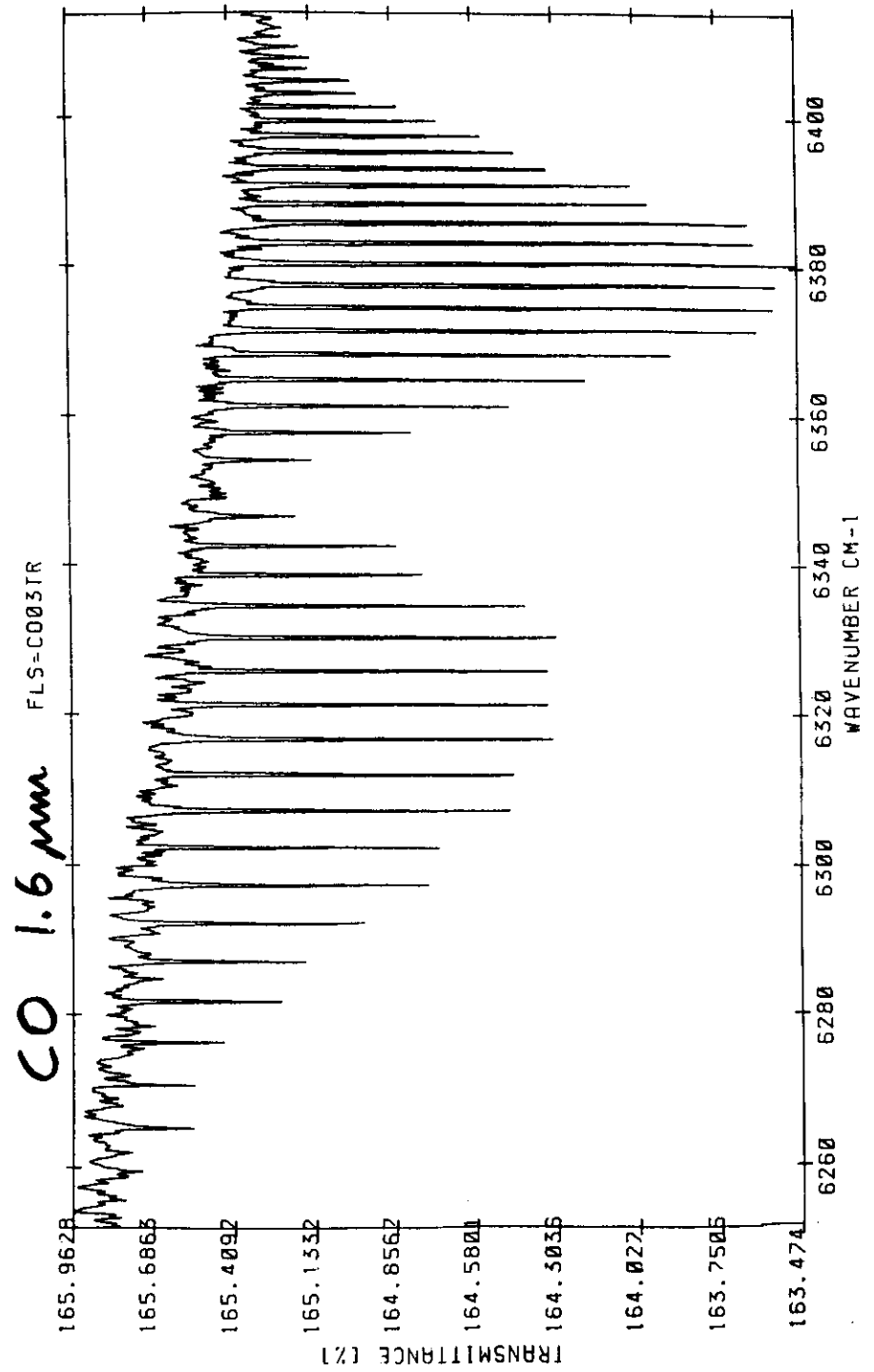
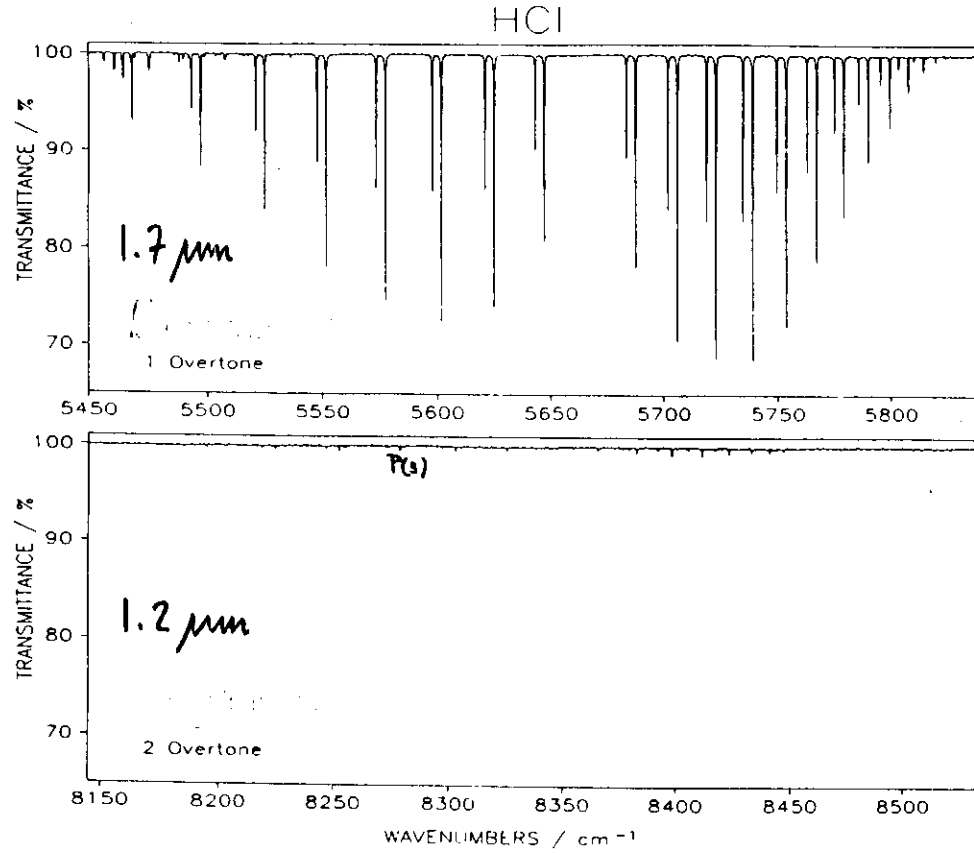
## MINIMUM DETECTABLE ABSORPTION VERSUS THE FREQUENCY MODULATION INDEX ( $\beta$ )

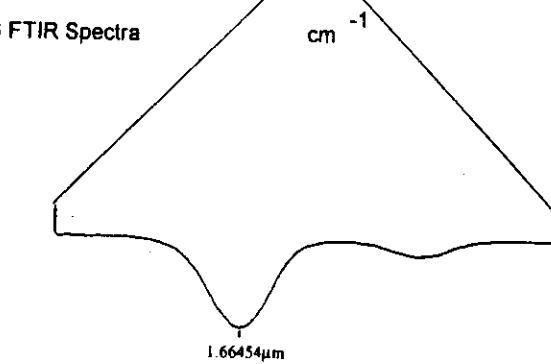
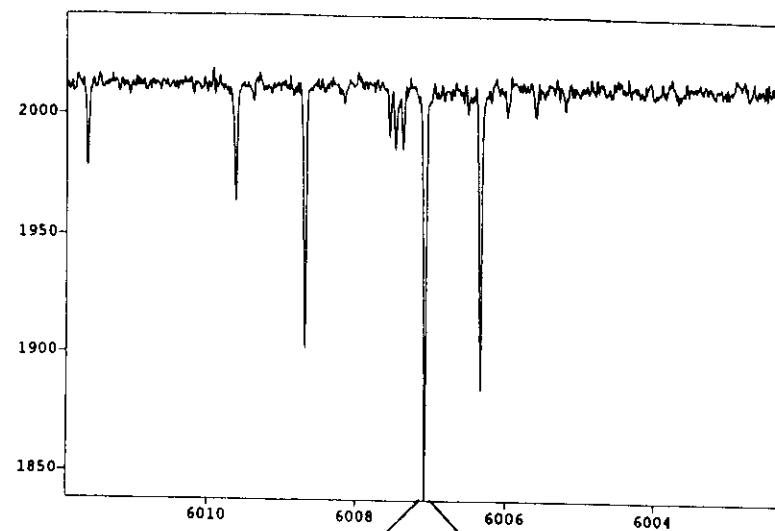
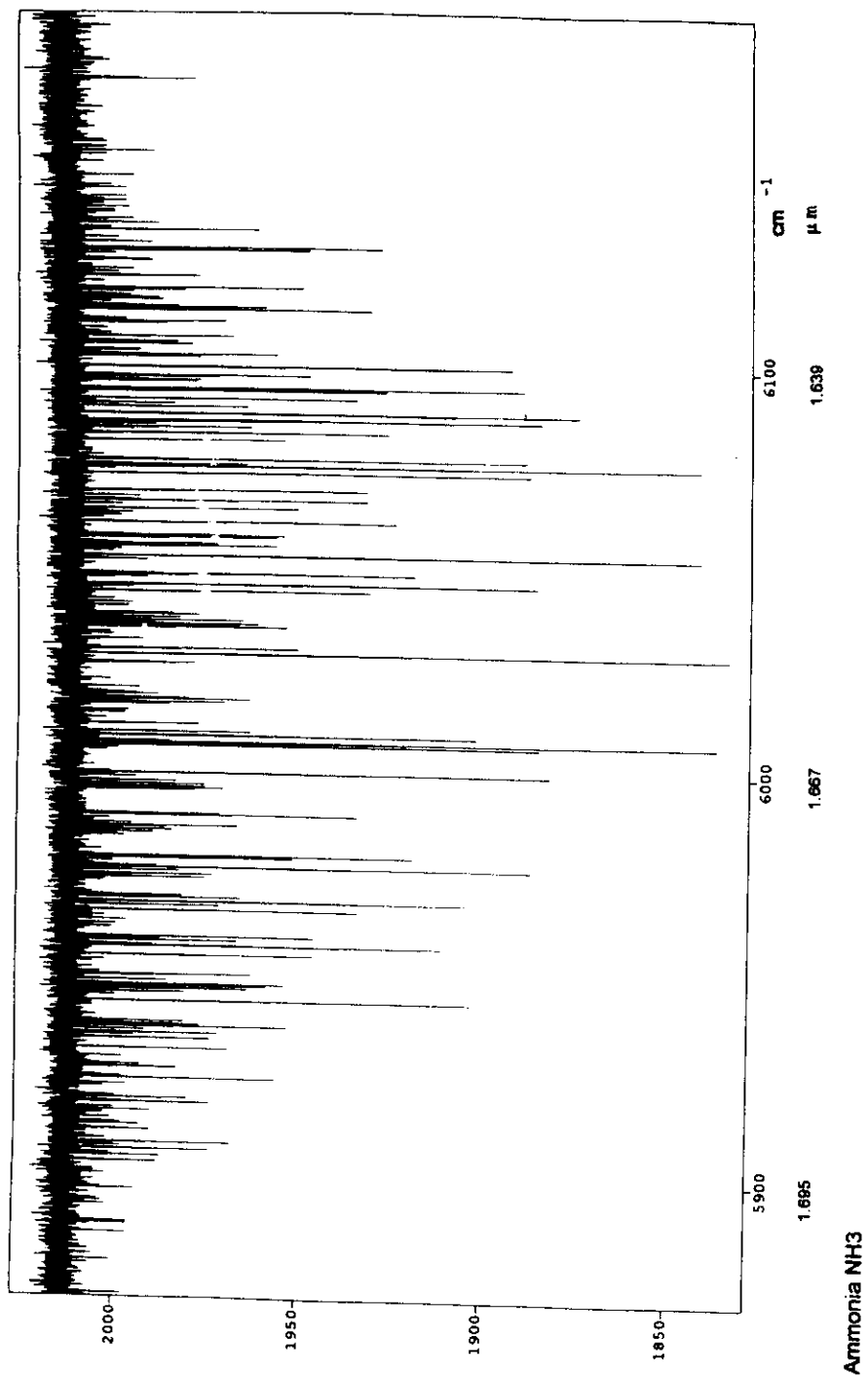


## ACTUAL EMISSION RANGE OF DIODE LASERS

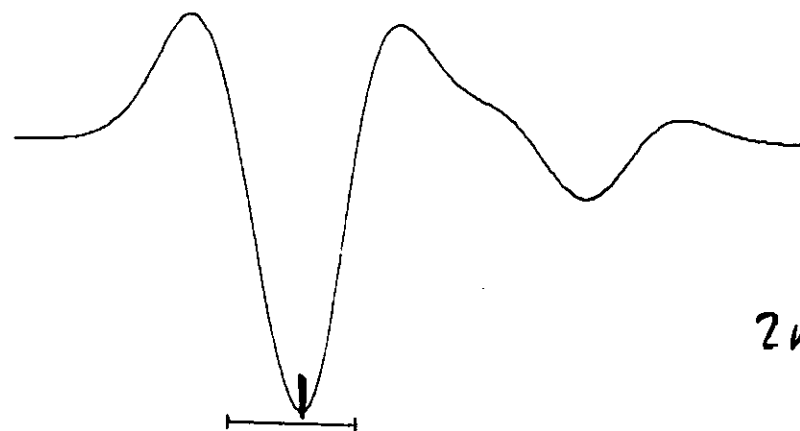
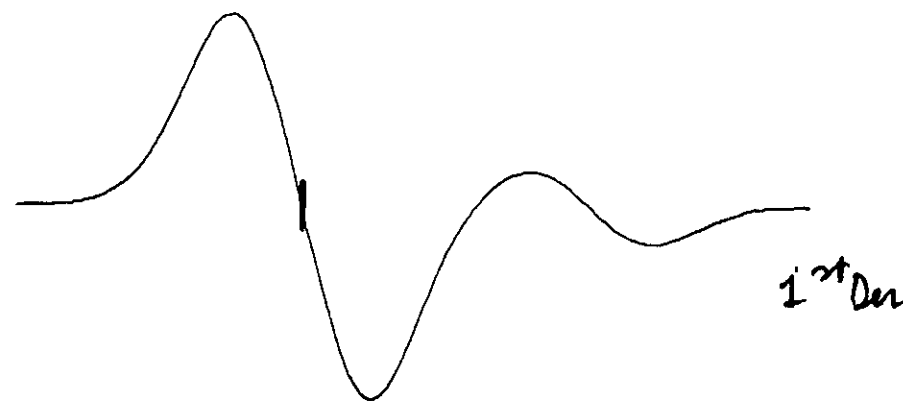
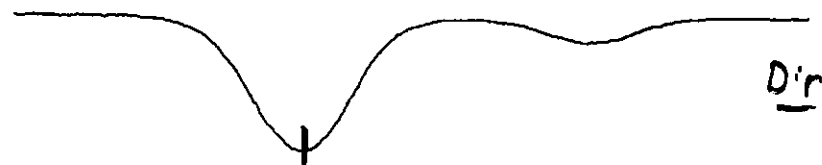


HCl





$\text{NH}_3$  1 Torr 1.64  $\mu\text{m}$  150cm FIBRES



500 MHz

3 ppm

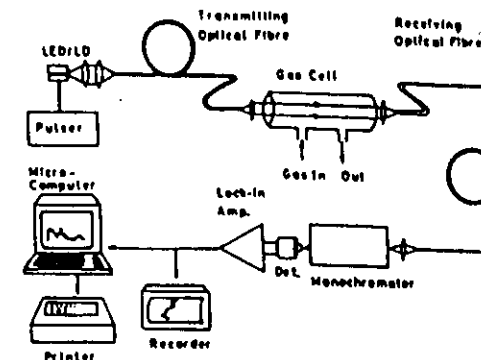


Fig. 7 Block diagram of an all-optical fiber-based remote sensing system for absorption measurement of low-level combustible, explosive, polluting and toxic gases in an absorption cell as the gas sensor head using a LED or LD operated in the near-infrared region.

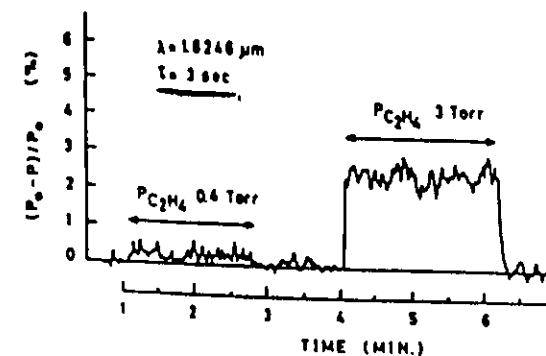
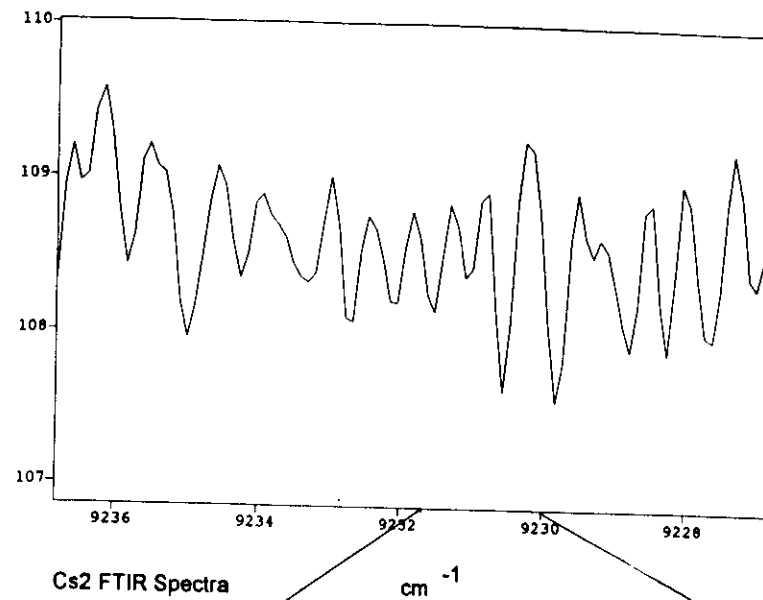
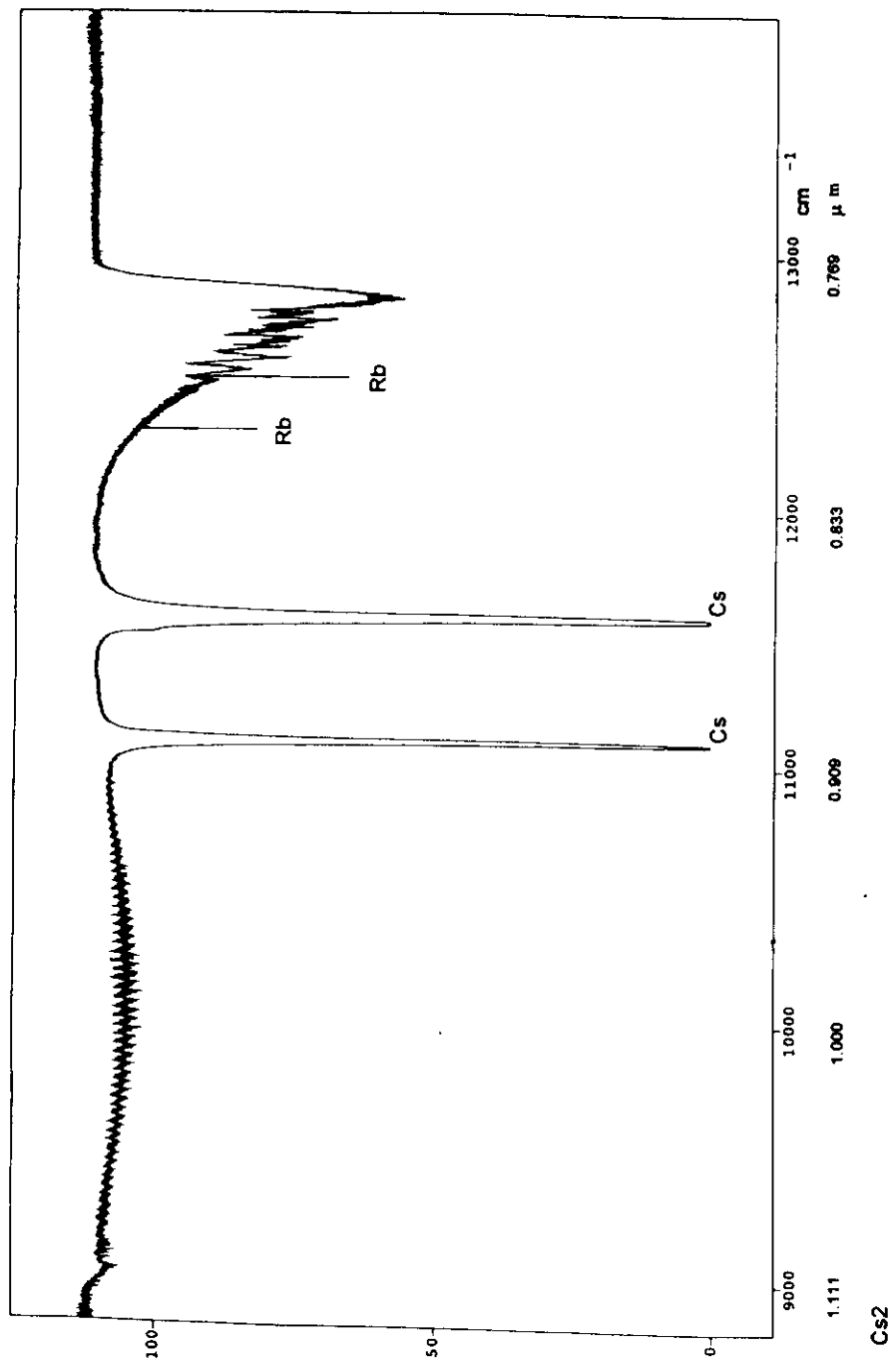
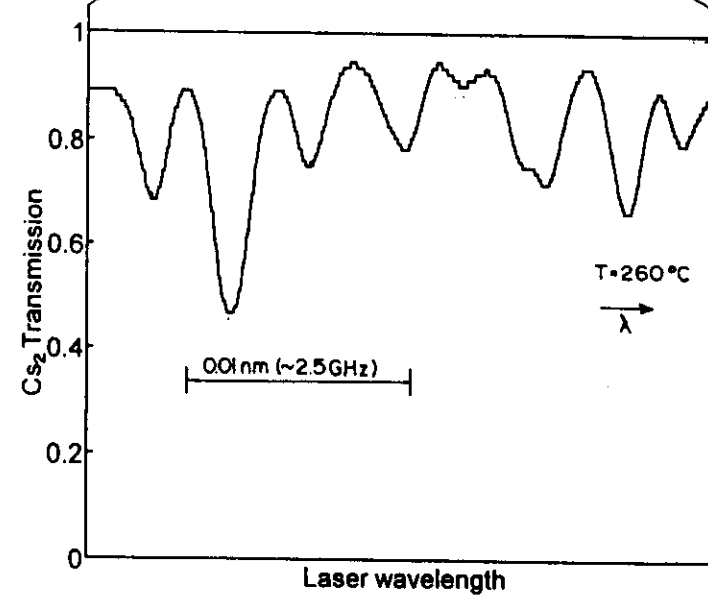


Fig. 8 Real-time detection of  $\text{C}_2\text{H}_4$  absorption signals in 1 atm.  $\text{C}_2\text{H}_4$ -air mixture contained in a remotely located 50-cm long absorption cell, utilizing a 5-km long, low-loss silica optical fiber link and a 1.64  $\mu\text{m}$  InGaAs LED.

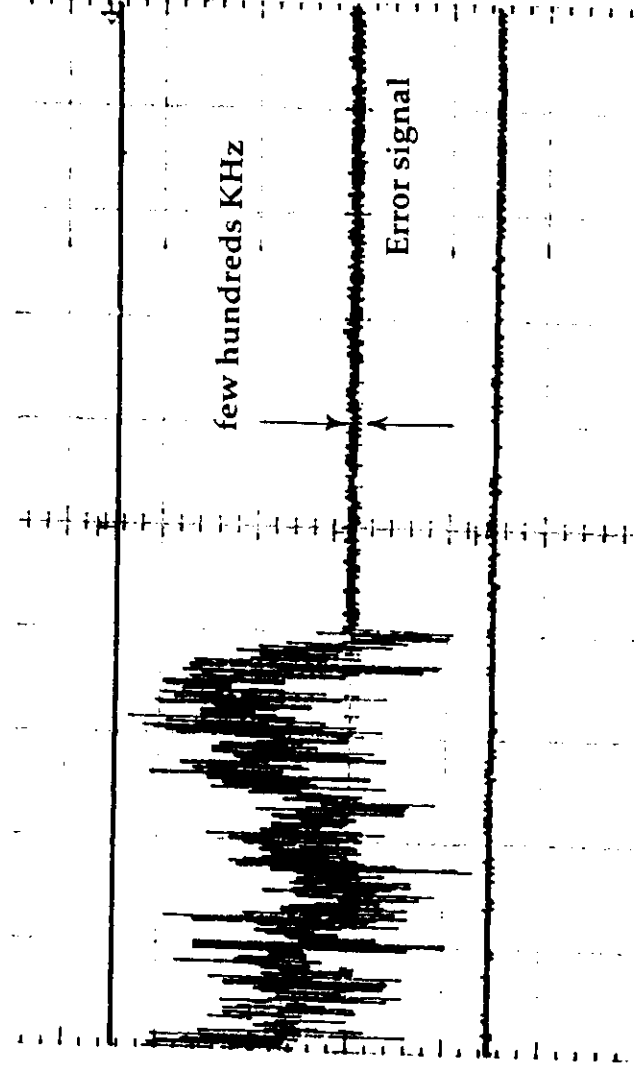


Cs<sub>2</sub> FTIR Spectra



Cs<sub>2</sub> DBR diode laser Spectra

## Frequency locking to molecular Cesium



44

## Molecular Cesium versus Helium

