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UNITED NATIONS EDUCATIONAL SCIENTIFIC AND CULTURAL ORGANIZATION



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AUTUMN COURSE ON GEOMAGNETISM, THE IONOSPHERE
AND MAGNETOSPHERE

(21 September - 12 November 1982)

TROPICAL IONOSPHERE AND TROPOSPHERE

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These are preliminary lecture notes, intended only for distribution to participants.
Missing or extra copies are available from Room 230.

Documents available here

- CCIR Document
- Davies book on "Ionospheric Radio Propagation"
- Indian book on "Tropospheric Communication and Antenna Systems"
- Indian Atlas on Radio Refractivity
- Indian Atlas on Water Vapor
- Argentinian Document on Ionogram Analysis
- Vol 1 of Boulder Symposium
- ITS 78 Volume

Tropical Troposphere

- Frequent occurrences of ducting in some zones

S.E. Asia

- High rain rate

S.E. Asia
South America

- Desert conditions
Duct problem

Middle East

- Mountain
Knife-edge diffraction
Obstacle gain

How are Tropics Different?

1

- LOW GEOGRAPHIC LATITUDES (X)
- LOW GEOMAGNETIC LATITUDES (Appleton Anomaly Field-aligned irregularities)

But Geogr lat $\leftrightarrow$ Geomag. lat. relation different in three zones:

ASIAN ZONE

AFRICAN ZONE

AMERICAN ZONE

2

CONSEQUENCES

F-Region

- Generally within Appleton Anomaly belt
- High Spread-F Incidence

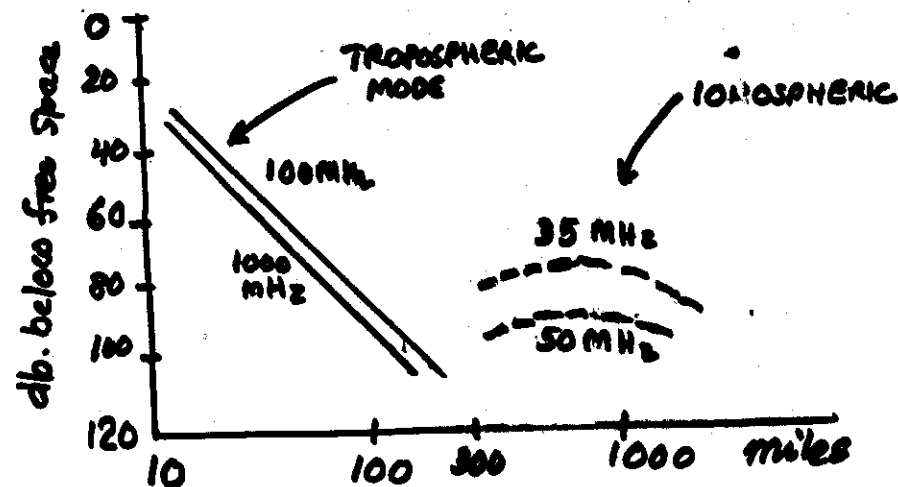
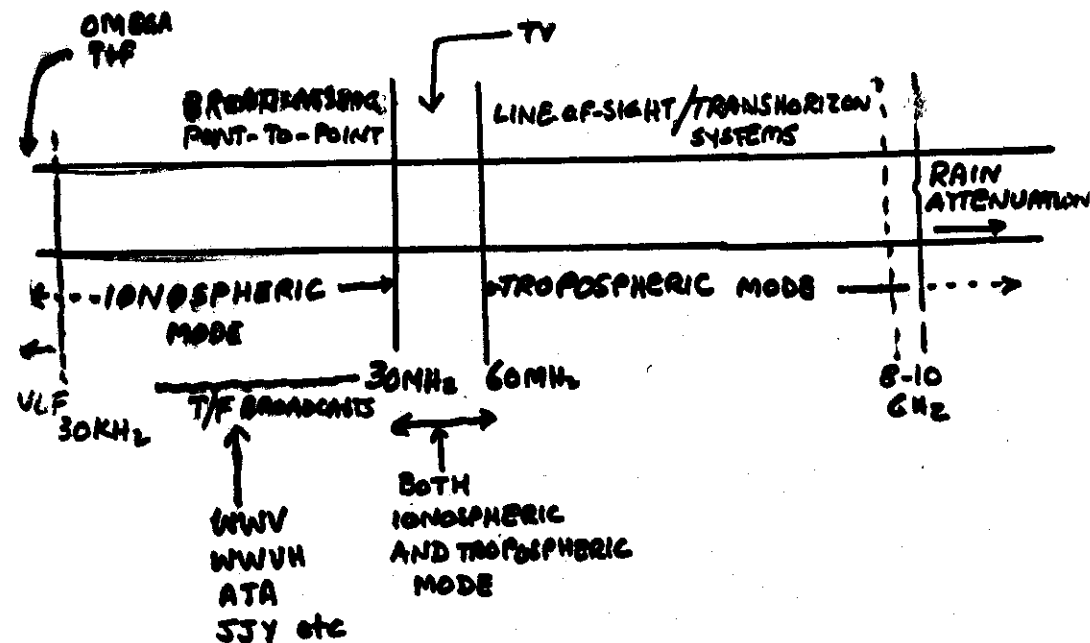
E-Region

- Equatorial Electrojet
- Equatorial Es & Es prop. beyond MUF

D-Region

Not subject to PCA's
 ↑
 Prop. relatively undisturbed

3



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Some Simple Observational Techniques

A. Ionosphere

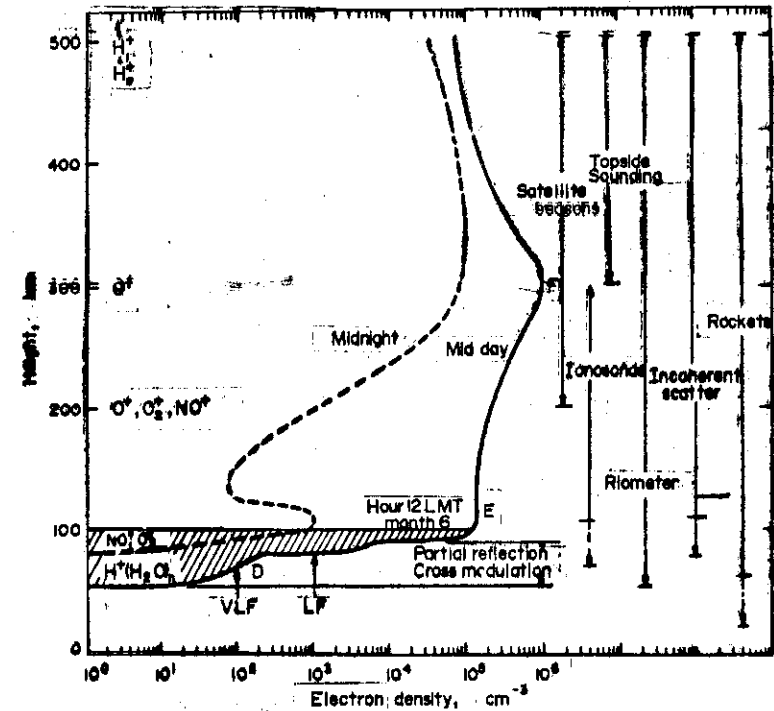
- Radio wave field strengths
- **IONOSONDES** (list distributed - note non-uniform distribution)
- **A1, A2, A3**
- TEC

PROFILE MODELLING

- IRI
- FROM BASIC CONSIDERATIONS

B. Troposphere

- **Radiosondes**
- Sodar
- Use of link data
- Radar



ENC-2.9-1

Parameters of Interest

R

11 years
22 years ← Hale Cycle
180 years ?

Solar
constant

≤ 1%
(~ 0.1%)

F10.7

UV

limited period only

PCA

ΔA , ΔNO_x

SID's

ΔA (Riometer)
Phase Changes (VLF)

PROTON FLUXES

> 10 MeV
> 30 MeV
> 60 MeV

Kp, Ap, Ci, Cp

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SOLAR RADIATION

S

$1373 \pm 20 \text{ W m}^{-2}$

• UNLIKELY TO HAVE
VARIED MORE THAN
10-15% IN PAST 3 billion
YEARS.

FAR IR

$10 \mu\text{m} - 1000 \mu\text{m}$

4950K $10 \mu\text{m}$
4450K ~ $150 \mu\text{m}$
6000K ~ $1000 \mu\text{m}$

OPTICAL
&
IR

$0.3 - 10 \mu\text{m}$

99% OF TOTAL SOLAR
RADIATION

UV

$1200 - 3000 \text{ \AA}$

3000 $\text{\AA}$ 1%
2000 $\text{\AA}$ 5%
1200 $\text{\AA}$ 25% 27d.
L₂ 15% 27d.
100% M → MX.
25% 16% BFL.

EUV

$10 - 1200 \text{ \AA}$

$\lambda > 170 \text{ \AA}$ CAN BE X5
FOR CORONAL
LINES

X RAYS

$< 10 \text{ \AA}$

x 1000-100 OVER
SOLAR CYCLES AND
DISTURBANCES

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Range and Altitudes

XUV	0.1 nm	45 km.
	2.0 nm	100 km
	20 nm	125 km
	150 nm	80 km
	220-280 nm	35 km
PROTONS	1 MeV	95 km
	100 MeV	30 km
	1000 MeV	8 km
α -PARTICLES	1 "	95 km
	1000 "	25
ELECTRONS	0.02 "	90
	100 "	10
BREMSTRALLUNG	0.005 "	75 km
	0.20 "	30
IR (5700 nm)		SURFACE

VARIABILITIES

LONG-TERM

SHORT-TERM

foE

WITH SEASON; SOL. CYCL. DAY-TO-DAY
: $\pm 6\%$
FLARES: $\pm 5-10\%$

PRED. METHOD: a. MUGGLETON
1944-1973 $\rightarrow$ (1975) $\leftarrow f(F_{10.7}, L, \text{standard})$
550 km
b. LEFTIN (1976) $\leftarrow f(R, M, UT, L, \text{IOWA})$
c. Nightingale (Wakai, NH)

F-REGION

foF1

: DUCHARME ETAL (1973)
ROISCH AND JONES (1973)

foF2

: CCIR ATLAS (CCIR 340-3) DAY-TO-DAY
COEFFICIENTS FROM JONES
et al (1968) : AS MUCH AS
15-20%
PREDICTIONS

• hmf2 from M3000F2

• LONGTERM PRED. OF
M3000F2 AVAILABLE
THROUGH COEFF IN
ITS-78 PROGRAMS

• SRI MODEL

• AIRFORCE GLOBAL
WEATHER CENTRE
CCIR $\rightarrow$ UPDATED
WITH NOAA REALTIME
IONOSPHERES AND TEC.

Propagation by E_s

1. ZONES

AUTOTOTAL ZONE : NIGHTTIME, LITTLE SEASONAL VARIATION
 MIDLATITUDE : BOTH DAY/NIGHT, SUMMER
 EQUATORIAL : DAYTIME, FREQUENT

2. Role of E_s in Communication

- Extension of role/range of usable HF radio waves - particularly imp. at S. min. when range limited
- May increase multipath problems for high HF radio system.
- Permits prop. of interfering signals to a larger extent

3. Global/Regional Distr.

LEFTIN ETAL (1968):

- SOLAR CYCLE MAX 4 MIN.
- INTERSECT A MODIFIED LATITUDE COORDINATE

E_s-MUF

SMITH (1981)

$$f_o(F_2) = 4 \text{ MHz} \sec^2$$

$$\log_{10} \left(\frac{f_o(F_2)}{f_o(F_1)} \right) = b(f_1 - f_2)$$

$$b \approx 0$$

SAXENA
 IZMIRAN (1964)

FOR INE
 GLOBAL

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D-Region Problem

1. Communication and Broadcasters

Those using MF and HF bands.
 Depend upon accurate signal-to-noise ratio

2. Users of ELF/VLF or LF navigation Systems

eg. Omega
 Depend upon stability of reflection region for time and position determination

3. T&E Systems

One of the most vital way of calibrating atomic clocks $\Rightarrow$ through VLF transmissions.

LOW LATITUDE PROBLEM

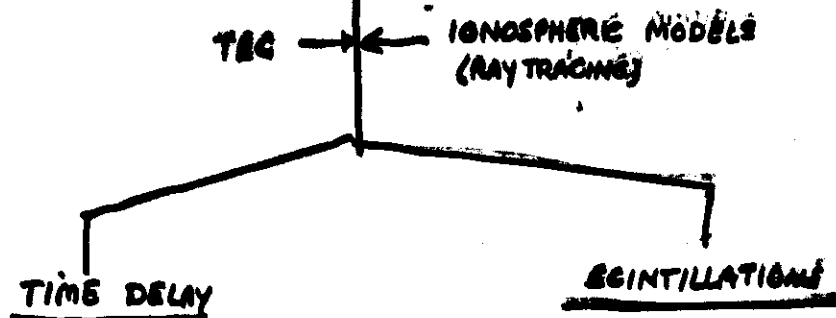
- Essentially no data to predict propagation of Omega and other VLF signals in this zone - however improving

- Very few automated MUF absorption equipment - practically dependent on log/loss data

Is reality of absorption anomaly?
 However, simpler to model

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TRANS-IONOSPHERIC PROPAGATION



PREDICTION: $\pm 20\%$ of TEC MONTHLY
MEAN (DAYTIME)
 $\pm 30-35\%$ AT NIGHT

$$S_4 = \left\{ \frac{\langle I^2 \rangle - \langle I \rangle^2}{\langle I \rangle^2} \right\}^{1/2}$$

PROVEN TO BE MOST
USEFUL FOR THEORETICAL
ANALYSIS AND MODELLING
PARAMETERS NEEDED TO
DEFINE PERFORMANCE
CHARACTERISTICS:

a) LUNAR
TIDAL EFFECTS

$$I(t) = S(t) \{1 + M \cos t\}$$

b) Improved: EUV (?)
Solar
flux

c) Geomag.: Mendillo-
Klobuchar
(STP, 1979)
storm
effects

- S_4
- τ_c (COHERENCE TIME)
- COHERENCE BANDWIDTH

MODELS: RINO (RADIO SCIENCE,
1979)
FREEMAN-RINO,
AGARD CONF 238
(1978)

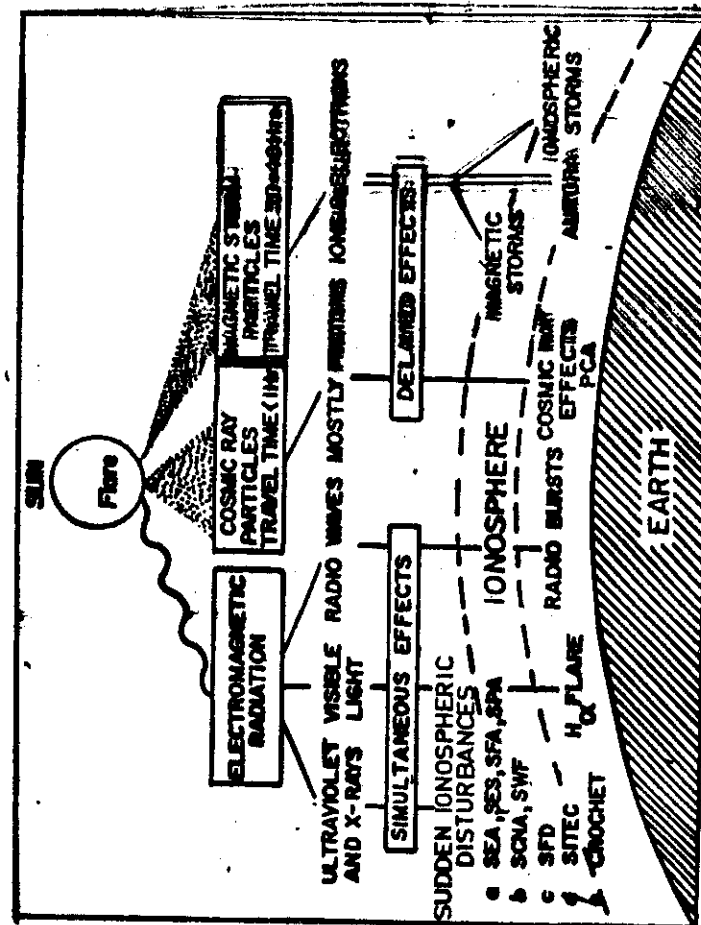
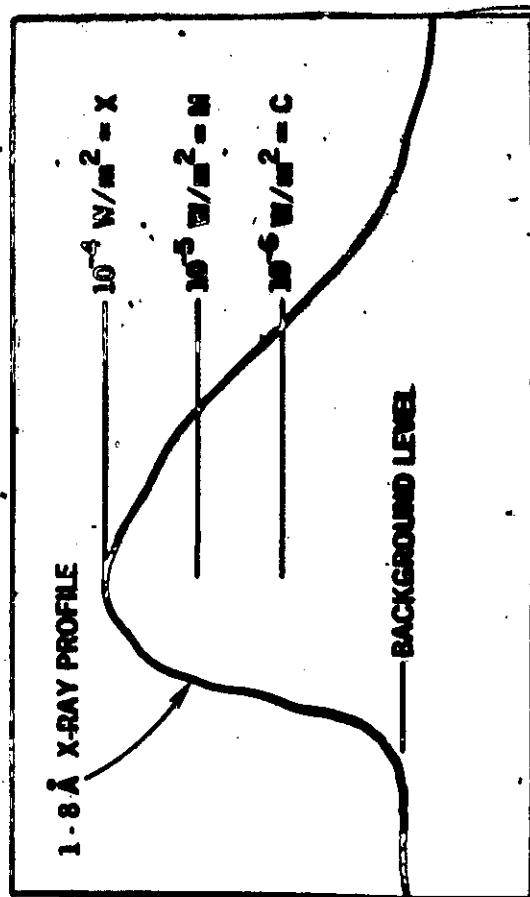


Fig. 1. An illustrative sequence of ionospheric effects associated with solar flares.

SESC FLARE CLASSES



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DISTURBANCES

	<u>E. MAX</u>	<u>SMALL</u>	<u>THIN</u>	<u>LONG</u>	<u>PHASE</u>
SID	2/WEEK	2/YEAR	3/4 hr	5-30 Cec	-1 to +5 db
PCA	1/MO	6	3d	40 db	
STORM (MAG.)	26/YR	22/YR	days	10-20 Cec	5-10 db fade
W.A.	20/YR	20/YR	days	usually phase abnorm	5-10db fade
STRATWARM	2/YR	2/YR	weeks	unknown	unknown

Cec = Centicycles; At 10KHz, 1Cec = 1ms interval
 ≈ 0.1 mile

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CHARACTERISTICS OF MAJOR T+F SYSTEMS

	ACCURACY (FREQ. STABILITY)	ACCURACY (TIME TRANSFER)	AMBIGUITY	COVERAGE
VLF : GBR, NBR, OMEGA ETC	1×10^{-8}	ENVELOPE 1-10ms	1 CYC	NEARLY GLOBAL
LF : STAND. FREQ. RANGE	1×10^{-8}	ENVELOPE 1-10 ms	DAY OF YEAR	USA-LIMITED (MNVB)
HF : STAND. FREQ. RANGE (eg. MNVB, RTA)	1×10^{-8}	1000 μ s	CODE - DAY OF YEAR	HEMISPHERAL
TV : PULSE LINE-TO (VHF/UHF)	1×10^{-8}	1 μ s	33 MS	NET
PORTABLE CLOCKS :	1×10^{-8}	1000 μ s	NA	COVERAGE
	1×10^{-8}	1000 μ s	NA	LIMITED BY TRANSMISSION

DISTRIBUTION OF SESC SERVICES

TYPE	TIME OF AVAILABILITY	PHENOMENA OF INTEREST
FORECAST	30 DAYS 6 MONTHS, 7 and 27 DAYS	SUNSPOT NUMBER GENERAL LEVEL OF SOLAR and GEOMAGNETIC ACTIVITY
FORECAST	1.2 DAYS FORECAST	FLARES, PROTON EVENTS, GEOMAGNETIC INDICES, SOLAR FLUX INDEX
FORECAST	20 MINUTES	FLARES
ALERT	IMMEDIATE	X-RAY, RADIO, OPTICAL EMISSION FROM FLARES
FORECAST	30 MINUTES - 120 HOURS	PROTON EVENT, GEOMAGNETIC STORM, SUBSTORMS
ALERT	IMMEDIATE	PROTON EVENT, GEOMAGNETIC STORM, SUBSTORMS
DATA (RELAYED)	REAL-TIME	REL. DATA
DATA (STANDARD)	DAILY and WEEKLY	SUMMARY DATA

INSTITUTIONS REQUIRING THE SERVICES IN INDIA.

P+T	IN FUTURE FEW <u>MS</u>	1×10^{18} (HAVE OWN CS STANDARD)
ITI	FEW <u>MS</u>	1×10^{12} (OWN CS STANDARD)
R.A	FEW <u>MS</u>	1×10^{-4}
ISRO	FEW <u>MS</u>	1×10^{10}
Sat. Tracking		1×10^{14}
LIGHTHOUSE	NOMINAL	1×10^8
BROADCASTING	GET TIME-PIP CALIBR. DIRECTLY FROM ATA	

Some Important Parameters in Tropospheric Propagation

- $N = \left(\frac{726 P}{T} + \frac{375,000 e}{T^2} \right) \approx 1000$
 $\uparrow$ N_{dry} $\uparrow$ N_{wet} $\uparrow$ $N = (n-1) \times 10^6$
 $200-400 \text{ units}$
- $\frac{dN}{dh} \leftarrow \begin{cases} 0 \text{ to } -40 N \text{ subrefr.} \\ -40 \text{ to } -157 N \text{ Superrefr} \\ -157 N \text{ Ducting} \end{cases}$
- C_N^2
- $a' = ka$

$$k = \frac{1}{1 + \frac{a}{n} \frac{dn}{dh} \cos \theta} \quad \theta = 0^\circ$$

Standard atm: $k = \frac{4}{5}$

N-Kinetic

Very short
Long
Longer

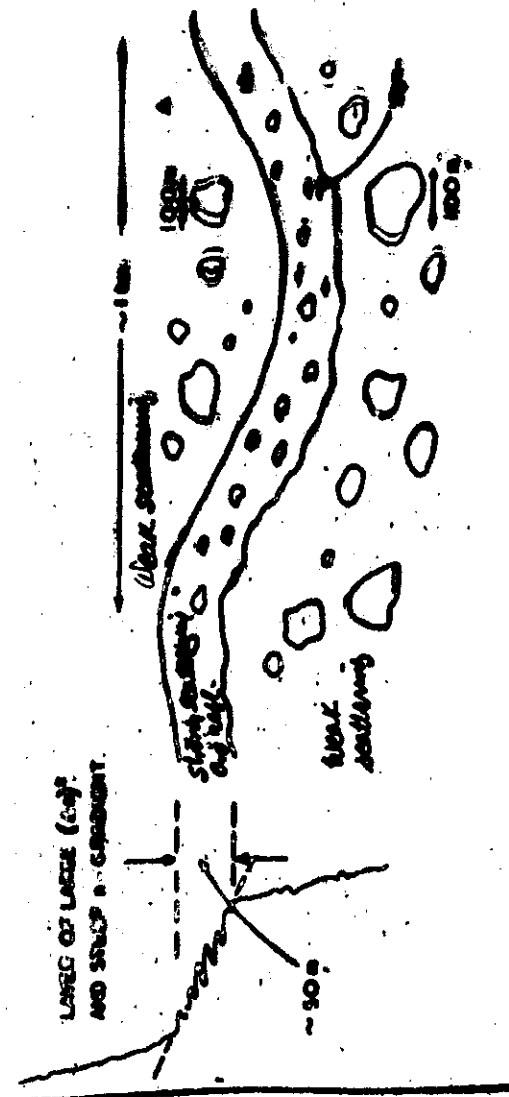
Ducting
Ref.
Random Perturbation

$$d = 118 \lambda^{2/3} \text{ max}$$

$$= 20 =$$

Characteristics of Climatic Types

	ANNUAL MEAN N.	ANNUAL RANGE	GEOGRAPHICAL AREAS
1. CONTINENTAL TEMPERATE	~320 (40-50 in S. COAST U.S.A)	20-40	± 10-60° EXTREMES OF TEMP. IN A LARGE LAND MASS
2. MARITIME TEMPERATE OVER LAND	~320	SMALL 26-38	A 20-50° NEAR THE SEA WHERE PREVAILING WIND UNOBSTRUCTED BY MOUNTAINS CARRY MOIST MARITIME AIR. U.K., COASTS OF USA, EUROPE, AFRICA
3. MARITIME TEMPERATE OVER SEA	AS IN (2) BUT DUCTING COMMON		DUCTING COMMON BETWEEN UK-EUROPE, WEST COAST OF USA-MEXICO
4. MARITIME SUBTROPICAL OVER LAND	~370	30-60	± 10-30° LOW LANDS NEAR SEA WITH DEFINITE RAINY AND DRY SEASONS
5. " OVER SEA	AS IN 4		COASTAL AREAS
6. DESERT SAHARA	~280	20-80	ROUND-THE-YEAR SEMI-ARID CONDITIONS
7. EQUATORIAL	~360	0-30	MARITIME CLIMATE 20°N-20°S MONOTONOUS HEAVY RAINS HIGH AV. SUMMER TEMP.
8. CONTINENTAL SUBTROPICAL	~320	60-160	HOT, SEASONAL EXTREMES IN RAIN. 20-40°N LAT



Scale of temperature layer in temperature diagram
Diagram 1: very much heating
Diagram 2: strong heating or reflection
Diagram 3: weak heating

