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Lecture I & II

## SUMMER SCHOOL IN HIGH ENERGY PHYSICS AND COSMOLOGY

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### ELECTROWEAK PHASE TRANSITION AND BARYOGENESIS

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Please note: These are preliminary notes intended for internal distribution only.

# I. Electroweak phase transition

(1)

1. Effective potential: vacuum ~~is~~ case

Action of Standard Model (as an example)

$$S = \int d^4x \left[ (D_\mu \varphi)^\dagger D_\mu \varphi - V_0(\varphi) + \text{gauge terms} + \text{fermions} \right]$$

$$V_0(\varphi) = -\mu^2 \varphi^\dagger \varphi + \frac{\lambda}{2} (\varphi^\dagger \varphi)^2 \quad - \text{scalar potential}$$

$\langle \varphi \rangle$ : minimum of  $V(\varphi)$

This is true at classical (= tree) level only.

Is  $V_0(\varphi)$  the whole story?

No: loop corrections.

One loop:  $V_{\text{eff}}(\varphi) = V_0(\varphi) + V_1(\varphi)$

$V_1(\varphi)$  = zero point energy (density) of all species of particles in background field  $\varphi$ .

$$m_{\text{gauge}} = g \varphi, \quad m_{\text{fermion}} = h \varphi, \text{ etc}$$

$$V_1(\varphi) = \left[ \sum_{\text{bosons}} \sum_{\text{states}} \frac{1}{2} E - \sum_{\text{fermions}} \sum_{\text{states}} E \right] / \text{Volume}$$

i.e.

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$$V_1(\varphi) = \sum_{\text{bosons}} \sum_{\text{spins}} \frac{1}{2} \int \frac{d^3 k}{(2\pi)^3} \sqrt{\vec{k}^2 + m_i^2(\varphi)}$$

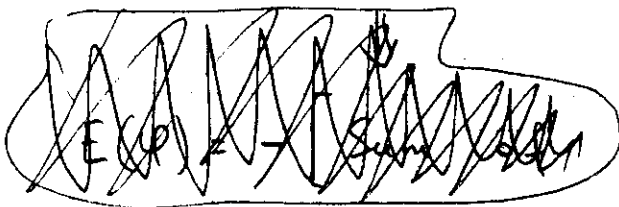
$$- \sum_{\text{fermions}} \sum_{\text{spins}} \int \frac{d^3 k}{(2\pi)^3} \sqrt{\vec{k}^2 + m_f^2(\varphi)}$$

Sometimes non-trivial contribution (Coleman, E. Weinberg)

More systematic way

$$e^{-E(\varphi)(\text{time})} = \int d(\text{fields}) e^{-S[\varphi, \text{other fields}]}$$

$$\delta\left[\int \varphi d^4x - (\text{Volume})(\text{time})\right]$$



$\varphi \rightarrow \varphi + \delta\varphi$  ; expand integral in loops

One loop

$$e^{-E(\varphi) \cdot \text{time}} = e^{-V_0(\varphi)(\text{volume})(\text{time})}$$

$$\times \prod_{\text{bosons}} \left[ \text{Det}(-\partial^2 + m_i^2(\varphi)) \right]^{-1/2}$$

$$\times \prod_{\text{fermions}} \left[ \text{Det}(\not{\partial} + m_f(\varphi)) \right]$$

$\Downarrow$

$$V_1(\varphi) = \sum_{\text{bosons}} \frac{1}{2} \int \frac{d^4 k}{(2\pi)^4} \ln(k^2 + m_i^2) - \sum_{\text{fermions}} \int \frac{d^4 k}{(2\pi)^4} \ln(k^2 + m_f^2)$$

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Actually, this is the same zero-point energy. (Integrate over  $k^0$ ), up to divergent ~~terms~~ terms to be renormalized anyway.

$\langle \varphi \rangle$  : minimum of  $V_{\text{eff}}(\varphi)$  !

2. Temperature-dependent effective potential = free energy <sup>density</sup> in background field  $\varphi$

"One loop" : free energy of particles in plasma; their interactions neglected

$$V_T(\varphi) = V_{T=0}(\varphi)$$

$$+ \left[ \sum_{\text{bosons}} \sum_{\text{states}} T \ln(1 - e^{-\beta E}) \right] \quad \beta = \frac{1}{T}$$

$$+ \sum_{\text{fermions}} \sum_{\text{states}} T \ln(1 + e^{-\beta E}) \Big] / \text{Volume}$$

(textbooks on STATISTICAL MECHANICS)

$$= V_{T=0}(\varphi) + \sum_i T \int \frac{d^3 k}{(2\pi)^3} \ln \left( 1 \mp e^{-\frac{1}{T} \sqrt{\vec{k}^2 + m_i^2(\varphi)}} \right)$$

$\mp$  : bosons/fermions

$\sum$  : sum over species and spins

Generally: RATHER complicated function of  $\varphi$  and  $T$

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- High temperature expansion: " $T \rightarrow \infty$ "

Bosonic contribution to leading order:

$$T \int \frac{d^3 k}{(2\pi)^3} \cdot \frac{1}{2} \ln [k^2 + m_i^2(\varphi)]$$

NB: same thing as one-loop contribution to eff. potential at zero temperature in three dimensions.

Anyway, MSM with  $\sin \Theta_w = 0$ :

$$V_T(\varphi) = \underbrace{-\frac{\mu^2}{2} \varphi^2 + \frac{\lambda}{4} \varphi^4}_{V_0(\varphi)} + \frac{3T^2}{32} g^2 \varphi^2 - \frac{3g^3}{32\pi} \varphi^3 \cdot T$$

+ "unimportant" terms

also top Yukawa, with positive sign

$g = SU(2)_{\text{weak}}$  gauge coupling.

First order phase transition always.

Remember: this is one loop approximation!

Structure of 1-loop  $V_T(\varphi)$  is generic, only coefficients change (MSSM, two-Higgs-doublet, ...)

Second critical point :  $\varphi^2$  term in  $V_T$

disappears at

$$\frac{3}{16} g^2 T^2 = \mu^2 \Rightarrow T = \sqrt{\frac{8}{3}} \frac{m_H}{g}$$

At this point

$$V_T(\varphi) = \frac{\lambda}{4} \varphi^4 - \frac{3}{32\pi^2} g^3 T \varphi^3$$

in fact  $\frac{m_H}{\sqrt{g^2 + h_{top}^2}}$

Minimum at  $\varphi = \frac{g}{32\pi} \frac{g^3}{\lambda} T = \frac{3}{2\pi} \frac{g^2}{\lambda} \mu$

Compare to zero temperature  $\langle \varphi \rangle_{vac} = \frac{\mu}{\sqrt{\lambda}}$

At critical temperature  $\varphi \ll \langle \varphi \rangle_{vac}$

As temperature drops down,  $\varphi$  ~~too~~ moves to

$\langle \varphi \rangle_{vac}$ .

All this is true only in 1-loop approximation. Whether this approximation is valid, remains to be understood!

4. More systematic approach:  
euclidean ~~field~~ field theory with  
finite time. (6)

Static properties

$$\text{Tr} \left[ \hat{O} e^{-\frac{H}{T}} \right] ; \hat{O} \text{ depends only on } \vec{x}.$$

Example: free energy

$$e^{-F(T)/T} = \text{Tr} e^{-\frac{H}{T}}$$

etc.

Note: 1)  $e^{-H/T}$  // imaginary time  
= euclidean evolution operator  
for time  $1/T \equiv \beta$ .

$$2) \text{Tr} e^{-H/T} = \int \mathcal{D}\phi \, e^{-\int_0^\beta \mathcal{L}(\tau) d\tau}$$

Its kernel in generalized coordinate representation  
(bosonic fields  $\phi$  only for simplicity)

$$(e^{-H/T})(\phi_f(\vec{x}); \phi_i(\vec{z})) = \int d\phi \, e^{-\int_0^\beta \mathcal{L}(\tau) d\tau}$$

$$\phi(\tau=0) = \phi_i$$

$$\phi(\tau=\beta) = \phi_f$$

Trace : set  $\phi_i = \phi_f = \phi$  and integrate over  $\phi$

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$\Downarrow$

$$e^{-F(T)/T} = \int d\phi e^{-\int_0^\beta \mathcal{L}(\tau) d\tau}$$

periodic fields:  
 $\phi(\beta) = \phi(0)$

Similarly,

$$\text{Tr} \left[ \hat{O} e^{-H/T} \right] = \int_{\text{periodic } \phi} d\phi O(\phi) e^{-\int_0^\beta \mathcal{L}(\tau) d\tau}$$

Fermions: ANTI - periodic in euclidean time.

$$\text{Now, } \phi(\vec{x}, \tau) = \sum_{n=0, \pm 1, \dots} e^{i \cdot 2\pi n T \tau} \phi_n(\vec{x}) \quad - \text{bosons}$$

$$\psi(\vec{x}, \tau) = \sum_{n=0, \pm 1, \dots} e^{i 2\pi (n + \frac{1}{2}) T \cdot \tau} \psi_n(\vec{x}) \quad - \text{fermions}$$

$\phi_n(\vec{x}), \psi_n(\vec{x}) = \text{Matsubara modes}$

$\Downarrow$

$$\int d\phi e^{-\int_0^\beta d\tau \int d\vec{x} \mathcal{L}(\vec{x}, \tau) \phi} =$$

$$= \int \prod_n d\phi_n e^{-\beta \int d\vec{x} \tilde{\mathcal{L}}(\vec{x}; \phi_n)} \Rightarrow \text{Dimensional reduction to 3dim!}$$



Most non-trivial: "static" fields

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$$S = \int_0^\beta d\tau \int d^3x \left( (\partial_0 \phi)^2 + (\partial_i \phi)^2 + V(\phi) + \dots \right)$$

$$\int d^3x \left\{ \beta \left[ \partial_i \phi_{n=0} \right]^2 + V_0(\phi_{n=0}) \right.$$

$$\left. + \sum_{n \neq 0} \beta \left\{ (\partial_i \phi_n)^2 + (2\pi T)^2 \cdot n^2 \phi_n^2 + \dots \right\} \right.$$

heavy at high T

$\phi_{n=0}$ : light at high T. (NB) Fermions always heavy!

Ignore heavy fields (integrate them out)

get 3d theory! - at zero temperature

$$S_{3d} = \frac{1}{T} \int d^3x \left[ (\mathcal{D}_i \varphi)^\dagger (\mathcal{D}_i \varphi) + \frac{1}{4} (F_{ij}^a)^2 + V(\varphi) \right]$$

(plus  $A_0$  terms!)

Scale T out

$$S_{3d} = \int d^3x \left[ |\mathcal{D}\varphi|^2 + \frac{1}{4} F_{ij}^2 + V(\varphi) + A_0\text{-terms} \right]$$

with three-dim couplings:

gauge  $g_3 = g\sqrt{T}$

Higgs:  $\lambda_3 = \lambda T$

$A_0$  massive  
due to Debye  
screening!  
 $m_{A_0} \sim gT$

## 5. Infrared problem.

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Massless 3-dim gauge theories are "sick" in Pert. Theory in infrared!

Gauge coupling is dimensionfull:

$$n\text{-th loop} \propto (g_3^2)^{N-1}$$

If 3-d fields are massive ("broken phase")  $\Rightarrow$  series in  $\frac{g_3^2}{M^2}$  [if external momenta  $k \neq 0$ ]  
 $\Downarrow \frac{g_3^2}{k^2}$

If massless  $\Rightarrow$  infrared divergencies.

Free energy in unbroken phase infrared divergent in four loops ( $F \propto T^3$  by dimensions)

$$F_{4\text{-loop}} \propto (g_3^2)^3 = g^6 T^3 \cdot \log \Lambda_{\text{infrared}}$$

Higher loops: linear, quadratic, ... divergencies in infrared.

Good behavior:  $M^2 \gg g_3^2 = g^2 T$

Recall at critical temperature: one loop result

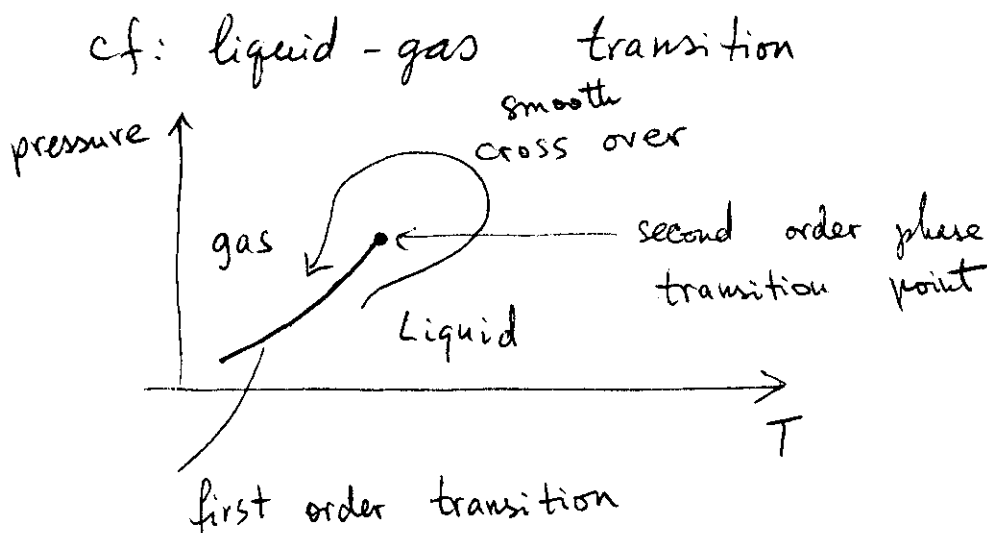
$$\varphi \propto \frac{g^3}{\lambda} T \Rightarrow M_W \propto \frac{g^4}{\lambda} T$$

Need  $\frac{g^2}{\lambda} \gg 1$ . Not valid at large  $\lambda$ ,  
i.e., at large  $m_H$ !

## Intermediate conclusions:

- \* Small  $m_H$  : first order phase transition; perturbation theory reliable, except nearby  $\varphi=0$ , at critical temperature and at lower  $T$
- \* Large  $m_H$  : perturbation theory unreliable at critical temperature. Order of phase transition ~~can~~ cannot be found perturbatively.
- \* Above critical temperature: perturbation theory unreliable at  $k \lesssim g^2 T$

Complication. No order parameter in MSM or MSSM (local, gauge invariant)



Gauge "symmetry" is not quite a symmetry  
Higgs mechanism is not symmetry breaking

## 6. What to do?

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Non-perturbative studies : lattice

full 4d simulations

Ilgenfritz et. al.  
Karsch et. al.

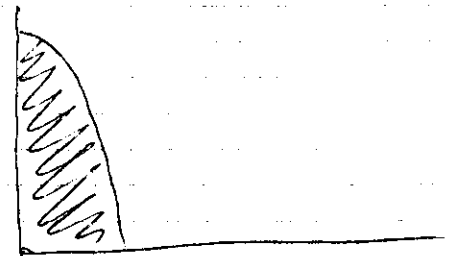
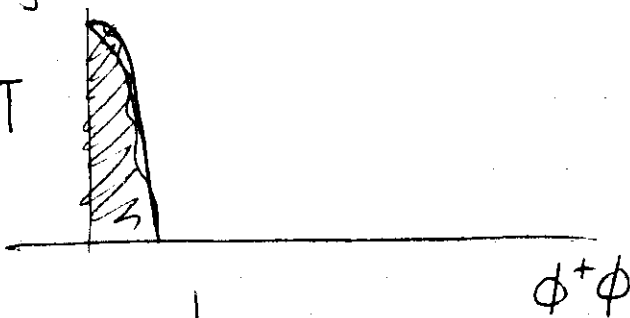
3d simulations

with analytic treatment  
of heavy modes  
Kajantie et. al.

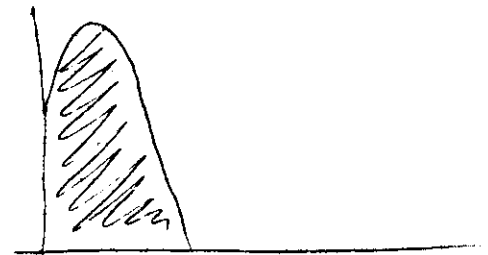
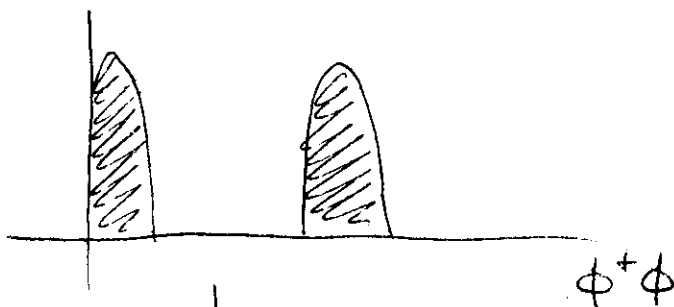
Example: calculate probability of observing  $\langle \phi^\dagger \phi \rangle$  on lattice

Probability

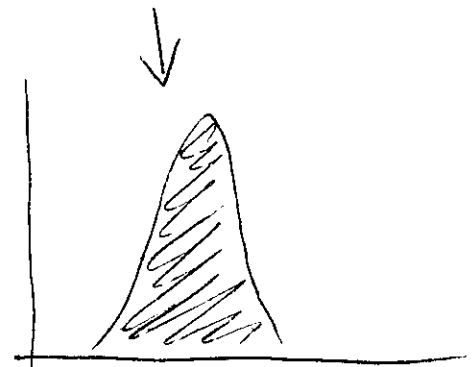
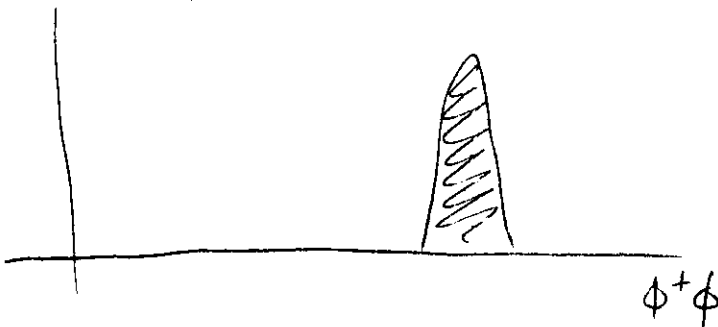
High T



$T = T_c$



$T > T_c$



First order transition:  
coexistence of phases at  $T = T_c$

Second order  
or  
cross over

Similarly: latent heat, surface tension.

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Simplification in 3d simulations:

3d theory is the same (near phase transition)  
for all  $SU(2) \times U(1)$  theories (MSM, MSSM, ...).

Because

Only one scalar field becomes massless  
at phase transition temperature

Outcome FOR MSM:

Phase transition is first order at  $M_H \leq 60 \text{ GeV}$   
For realistic Higgs masses smooth crossover,  
second order or very weakly first order.

MSSM: Depends on  $t_{\text{tp}}$ , etc., but mostly  
2nd order, cross over or very weakly first order

Perturbation theory works quite well

until at critical point  $m_W(\varphi) \approx gT$

## II. Electroweak baryon number non-conservation.

1. Simple example of non-conservation of fermion quantum numbers due to level crossing:  
massless fermions in (1+1) dimensions.

Free Dirac equation in (1+1) dimensions:  
fermions = two-component ~~spinors~~ spinors

$$\psi = \begin{pmatrix} \chi \\ \eta \end{pmatrix}$$

Dirac equation

$$i\gamma^\mu \partial_\mu \psi = 0$$

$$\gamma^0 = \tau^1, \quad \gamma^1 = i\tau^2$$



$$i(\partial_0 - \partial_1)\chi = 0$$

$$i(\partial_0 + \partial_1)\eta = 0$$

$\chi = \chi(x_0 + x_1)$  moves left with speed of light

$\eta = \eta(x_0 - x_1)$  moves right with speed of light

Concentrate on right fermions  $\eta$ .

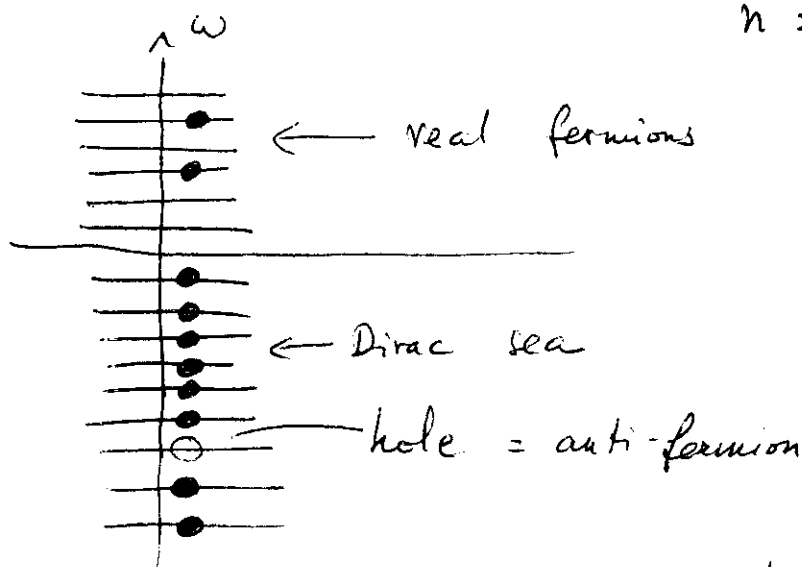
Levels of free right fermions

Space = circle of length  $L$  (with periodic boundary conditions)

Eigenfunctions of Dirac hamiltonian

$$\eta_\omega = e^{-i\omega x^0 + ikx^1} \quad ; \quad k = \omega = \frac{2\pi}{L} n$$

$n = \text{integer}$



Suppose now that fermions are <sup>electrically</sup> charged

Impose background electric field for a while (in positive = right direction)

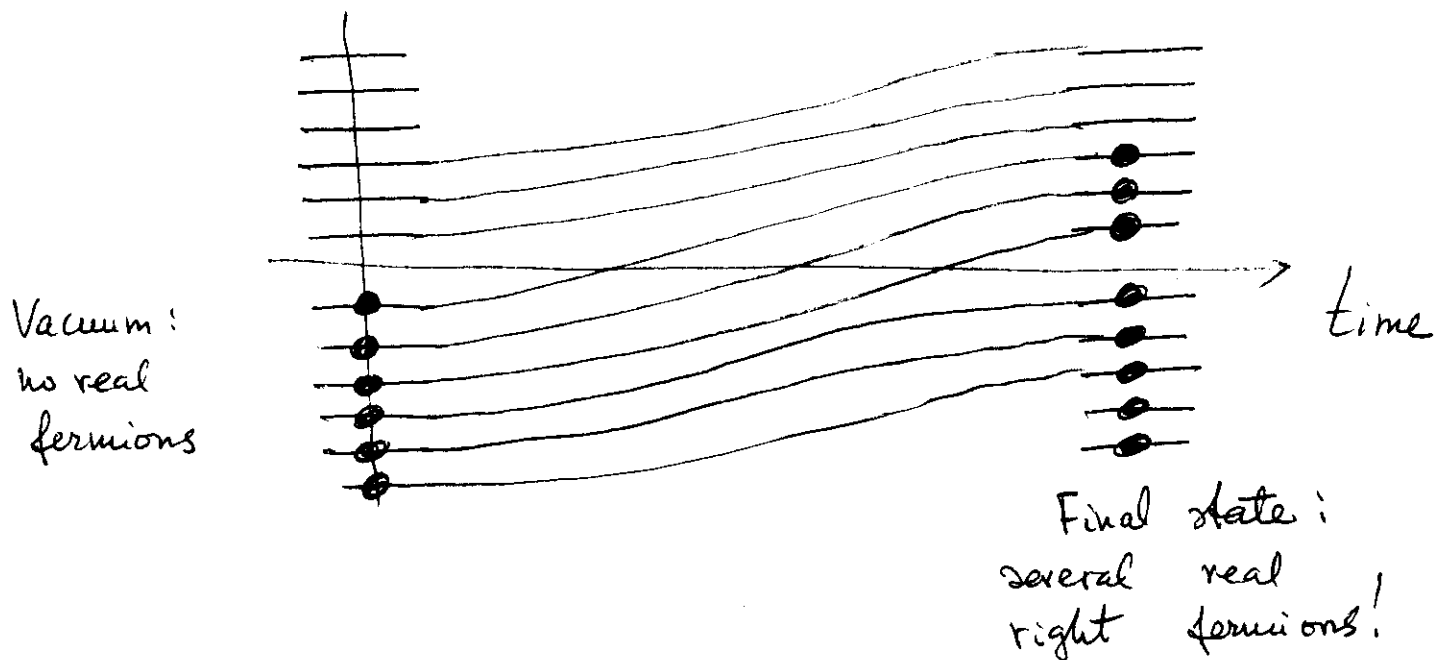
Fermions accelerate in this field

No ~~fermions~~ right fermions move left  $\Rightarrow$  no decelerated right ~~electro~~ fermions.

Energy of all <sup>right</sup> fermions increases.

Negative energy levels become positive energy!

When electric field is switched off, fermions remain on their levels  $\Rightarrow$  real fermions are born from Dirac sea! Number of right fermions is changed!



How many?

Dirac equation in background electric field;  
 $A_0 = 0$  gauge

$$i\partial_0 \eta = -i(\partial_1 - i\cancel{A_1})\eta$$

include into  $A_1$

Adiabatically changing field  $A_1(x, t)$  [for a moment]

$$\eta = e^{-i \int^t \omega(t) dt} \eta_{\omega(t)}(x^1)$$



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$$-i \left[ \partial_1 - ie A_1(x^1, t) \right] \eta_{\omega(t)} = \omega(t) \eta_{\omega(t)}$$

$$\Downarrow$$

$$\eta_{\omega(t)} = e^{i \omega(t) x^1 + i \int_{-L/2}^{x^1} A_1(x^1, t) dx^1}$$

Periodicity :  $\omega_n(t) \cdot L + \int_{-L/2}^{L/2} A_1(x^1, t) dx^1 = 2\pi n$

$$\Downarrow$$

$$\omega_n(t) = \frac{2\pi n}{L} - \frac{1}{L} \int_{-L/2}^{L/2} A_1(x^1, t) dx^1$$

Now, electric field  $E = -\partial_0 A_1$

If  $E$  is turned on for a while, then

$$A_1(t \rightarrow -\infty) = 0 \quad ; \quad A_1(t \rightarrow +\infty) \neq 0$$

$$\Downarrow$$

$$\omega_n(t \rightarrow -\infty) = \frac{2\pi n}{L} \quad \underline{\text{but}} \quad \omega_n(t \rightarrow +\infty) = \frac{2\pi}{L} (n + q)$$

with

$$q = -\frac{1}{2\pi} \left[ \int A_1 dx^1 \right]_{t \rightarrow +\infty}$$

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Level #  $n$  becomes level #  $(n+q)$

Number of levels crossing zero from below =  $q$

$\Downarrow$   
Change in number of right fermions  
 $\Delta N_R = q$

NB1 Adiabaticity is not essential:

in non-adiabatic case only pair creation of and anti-fermions / can occur.

$$N_R = \left( \begin{array}{c} \# \text{ of right} \\ \text{fermions} \end{array} \right) - \left( \begin{array}{c} \# \text{ of their} \\ \text{anti-fermions} \end{array} \right)$$

changes by  $\Delta N_R = q$

NB2 If there are also left fermions in theory, then

$$\Delta N_L = -q$$

Total  $N_F = N_L + N_R$  is conserved

Chirality  $Q^5 = N_L - N_R$  is not conserved.

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Meaning of  $q = -\frac{1}{2\pi} \int_{-L/2}^{L/2} A_1 dx^1$   
 topologically  
 distinct vacua.

Vacuum (classical) = pure gauge

$$A_1 = \frac{1}{i} \omega \partial_1 \omega^{-1}$$

$$\underline{A_0 = 0}; \omega = \omega(x^1)$$

$$\omega = e^{i\alpha(x^1)} \in U(1)$$

$\omega$  = periodic in space  $\Rightarrow \alpha(+\frac{L}{2}) - \alpha(-\frac{L}{2}) = 2\pi q$

$q$  = integer = topological # of vacuum =

= degree of mapping  $S^1 \rightarrow U(1)$

space  $\nearrow$  gauge group

$$-\frac{1}{2\pi} \int_{-L/2}^{L/2} A_1 dx^1 = -\frac{1}{2\pi i} \int_{-L/2}^{L/2} (-i \partial_1 \alpha) dx^1$$

$$= \frac{1}{2\pi} \left[ \alpha\left(\frac{L}{2}\right) - \alpha\left(-\frac{L}{2}\right) \right]$$

$$\Downarrow$$

$$q = -\frac{1}{2\pi} \int_{-L/2}^{L/2} A_1 dx^1 = \text{topological \# of final vacuum}$$

(if  $q=0$  in initial vacuum).

So, in this simple model :

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- 1)  $\Delta N_R = q \neq 0$
- 2)  $\Delta N_R$  does not depend on form of background gauge field, but only on its topology
- 3) Selection rule :  $\Delta N_R$  is the same for all ~~fermion~~ right fermion species;  
 $\Delta N_L = -q$  for all left species.

NB1. For process considered,  
define

$$Q = -\frac{1}{4\pi} \int \epsilon_{\mu\nu} F_{\mu\nu} d^2x = \frac{1}{2\pi} \int_{-\infty}^{+\infty} dt \int_{-1/2}^{1/2} E dx^1$$

Since  $E = -\partial_0 A_1$ , this is full derivative integral  $\Rightarrow Q = -q$

$$\Delta N_R = -Q ; \Delta N_L = Q \quad \text{for each species.}$$

NB2 Euclidean zero mode

Consider euclidean Dirac (Weyl) equation  
( $A_0 = 0$  gauge)

$$-\partial_0 \eta = H_D(\tau) \eta$$

where levels of  $H_D$  move like above.

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In adiabatic case: solution

$$\eta(\tau, x^1) = e^{-\int \omega(\tau) d\tau} \eta_{\omega(\tau)}(x^1)$$

$\eta_{\omega(\tau)}$  = same eigenfunctions as above.

If  $\omega(\tau \rightarrow -\infty) < 0$

$$\omega(\tau \rightarrow +\infty) > 0$$

Then  $\eta(\tau, x^1)$  decays both as  $\tau \rightarrow -\infty$   
and  $\tau \rightarrow +\infty$

$\Downarrow$

Euclidean Dirac operator has ~~eigenmode~~  
eigenfunction with zero eigenvalue.

$$\partial_0 \eta + H_D(\tau) \eta = 0 \Leftrightarrow \text{euclidean Dirac equation has normalizable solution in euclidean space-time}$$

In fact, # of normalizable zero eigenvalues of euclidean Dirac operator, ~~#~~  $\text{Ker}(\not{D})$ , also depends only on topology; namely

$$\Delta N_R = \text{Ker}(\not{D}_{\text{right}}) - \text{Ker}(\not{D}_{\text{right}}^+) = -Q = q$$

This is another way to study fermion number non-conservation: count # of fermion zero modes in euclidean space-time  
(in fact, it is original way of  $\rightarrow$  +1/2 hofst)

## 2. Vacuum structure of non-abelian gauge theories in $(3+1)$ dimensions

Classical vacuum = pure gauge

$$A_0 = 0 \quad \text{gauge}$$

$$\vec{A}_\mu = \omega \vec{\partial}_\mu \omega^{-1}$$

$\omega(\vec{x}) \in \text{Gauge group}$

$$\psi = \omega \psi_0 \quad (\text{Higgs field})$$

Take space to be a large sphere  $S^3$   
(not essential!)

$\omega$ : mapping  $S^3 \rightarrow \text{Gauge group}$

$$\pi_3(S^3) = \mathbb{Z}$$

For simple gauge groups,  $\pi_3(G) = \mathbb{Z}$

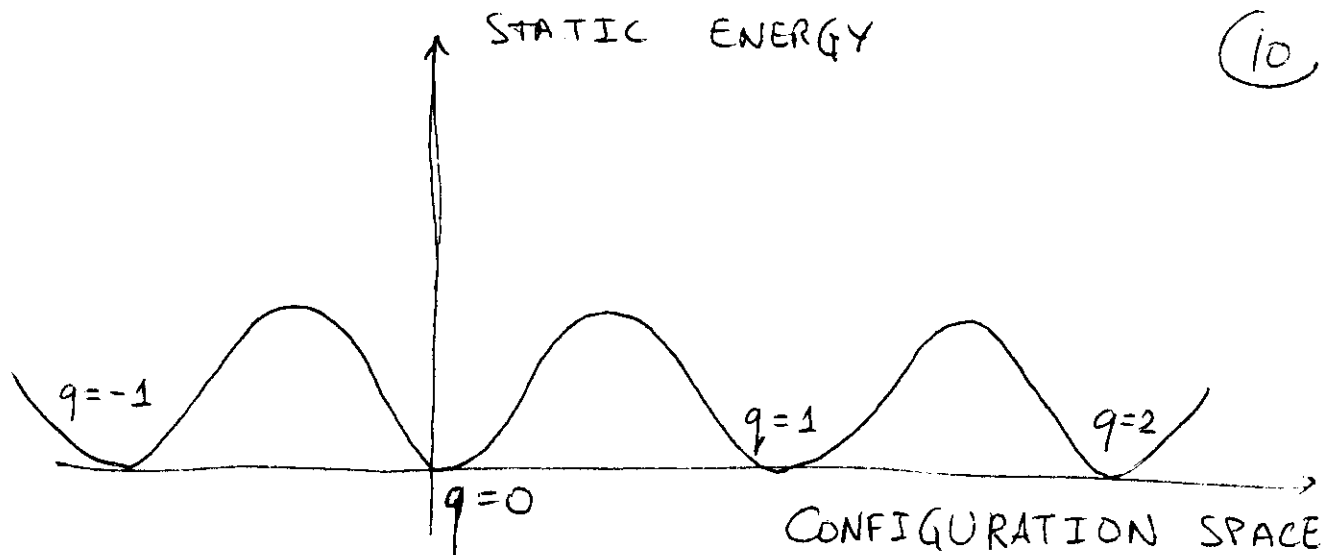
(say,  $G = SU(2)$  - 3-sphere)

$\Downarrow$

Topologically distinct vacua, numbered by  $q = 0, \pm 1, \pm 2, \dots$

If one moves from one vacuum to another, one should pass configurations with non-zero static energy.

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If a configuration  $\vec{A}(\vec{x}, t)$  interpolates between different vacua, then

$$q_{\text{final}} - q_{\text{initial}} = Q = -\frac{1}{16\pi^2} \int d^4x \operatorname{Tr}(F_{\mu\nu} \tilde{F}_{\mu\nu})$$

$$\tilde{F}_{\mu\nu} = \frac{1}{2} \epsilon_{\mu\nu\lambda\rho} F_{\lambda\rho}$$

Exactly in the same manner as in  $(1+1)d$

$$\Delta N_L = Q \quad \text{for fundamental fermions in } SU(N).$$

$$\Delta N_R = -Q$$

For each species!

Fermion quantum numbers are not conserved  
iff one travels from ~~one~~<sup>a</sup> vacuum to  
topologically distinct one.

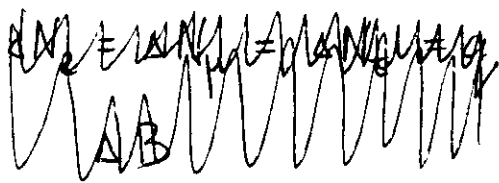
NB Nothing like this for abelian theories  
in  $(3+1)$  dimensions.

### 3. Electroweak theory: selection rules.

- Only left-handed fermions interact with  $SU(2)$  gauge fields.
- $U(1)$  does not produce violation of fermion numbers.

$$\Delta N = q \quad \text{For each species (doublet)}$$

In physical terms



$$\Delta L_e = \Delta L_\mu = \Delta L_\tau = q$$

$$\Delta B = 3 \cdot 3 \cdot \frac{1}{3} q = 3q$$

color      generation      B of each quark



1)  $B$  and  $L_i$  are not conserved separately

2)  $(B-L)$ ;  ~~$(L_e - \frac{1}{3}B)$~~   $(L_e - \frac{1}{3}B)$ ;  $(L_\mu - \frac{1}{3}B)$ ;  $(L_\tau - \frac{1}{3}B)$

are all conserved;  $(L_e - L_\mu)$ , etc also.

Out of four  $B, L_e, L_\mu, L_\tau$  three

are conserved after non-perturbative effects are taking into account.

NB QCD with massless quarks:  $U_A(1)$  is not meson conserved  $\Rightarrow$  no 9th light pseudoscalar ~~meson~~



#### 4. Rates.

(12)

a)  $T=0$  : tunneling;

$$\Gamma \propto e^{-16\pi^2/g_2^2} \sim e^{-160} \quad - \text{forget.}$$

$g_2 = \text{SU}(2) \text{ gauge coupling}$

b) Intermediate temperatures : thermal jumps

- Most naive estimate : Boltzmann

$$\Gamma \propto e^{-\frac{\text{Height of barrier}}{T}}$$

Saddle point between vacua = static unstable solution to field equations = sphaleron

$$\text{Height of barrier} = E_{\text{sph}} = \frac{2m_W}{\alpha_W} B\left(\frac{m_H}{m_W}\right) \sim 10 \text{ TeV}$$

$\alpha_W = \frac{g_2^2}{4\pi}$

Most naively  $\Gamma \propto e^{-\frac{E_{\text{sph}}}{T}}$

- Less naively :  ~~$m_W$~~   $\langle \varphi \rangle = \langle \varphi \rangle(T) \Rightarrow$

$$m_W = m_W(T) \Rightarrow E_{\text{sph}} = E_{\text{sph}}(T)$$

Reasonable estimate!

$$\Gamma = (\text{Pre-exponent}) \cdot e^{-\frac{2m_W(T)}{\alpha_W T} B\left(\frac{m_H}{m_W}\right)}$$

Pre-exponent calculated, but not greatly important.

c) High temperatures

$$\alpha_W T \gtrsim M_W(T) \Rightarrow \text{no exponential suppression}$$

$\ll$   
Semiclassical analysis not applicable.

This applies, in particular, to "unbroken phase".

"Unbroken phase": coupling constant  $= g_3 = g\sqrt{T}$   
T always comes with  $g^2 \approx \alpha_W$

Sphaleron rate/unit time · unit volume

$$\Gamma \propto T^4 \Rightarrow \Gamma = \text{const} (\alpha_W T)^4$$

Shaposhnikov

Not necessarily true! (Arnold, Son, Yaffe),  
maybe, extra  $\alpha_W$ .

Numerical simulations must decide.

Early Universe: compare rate to rate of evolution of Universe.

High T:

$$\underbrace{\text{const.}}_{1 \pm 0.1?} \alpha_W^4 T \gtrsim \frac{T^2}{M_{\text{pl}} \cdot N_*} \Rightarrow T \lesssim 10^{12} \text{ GeV}$$

( $10^{10} \text{ GeV}$  if  $\alpha^5$ )

Baryon & lepton numbers definitely  
not conserved before EW phase transition.

After phase transition:

Sphaleron rate small if

$$\frac{2 m_W(T)}{\alpha_W T} B\left(\frac{m_H}{m_W}\right) > \ln \frac{M_{\text{pl}} N_*}{T}$$

OK if  $\langle \varphi \rangle$  is large after phase transition  $\Rightarrow$  strongly first order phase transition

MSM:  $m_H < 50 \text{ GeV}$  - ruled out experimentally

MSSM: corners of parameter space.

Two-Higgs-Doublet models: part of parameter space.

Generically, sphaleron processes are in equilibrium immediately after phase transition

