



the
abdus salam
international centre for theoretical physics

ICTP 40th Anniversary

H4.SMR/1574-2

"VII School on Non-Accelerator Astroparticle Physics"

26 July - 6 August 2004

100 Years of Science

R. Carrigan

Fermi National Accelerator Laboratory
Batavia, U.S.A.



100 Years of Science

**Dick Carrigan
Fermilab**

Nancy Jean Carrigan



can I unweave the gauzy fabric of stars?

can I unweave the gauzy fabric of stars?

can I untangle the matted carpet of earth?

can I untangle the matted carpet of earth?

“Astrophysics and the Ant”

Acrylic on Linen

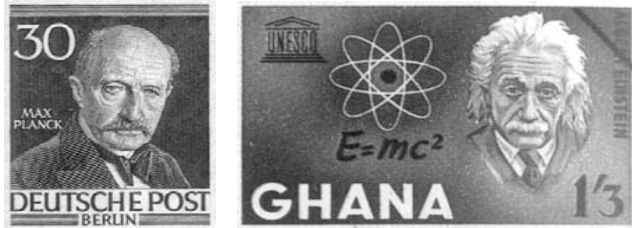
Nancy Jean Carrigan © 2000

Contact Nancy at: carrigans2@aol.com

What 100 years are we talking about?



Maxwell 1873 to 1973 – no fun



Planck – 1900,
Einstein 1905 but that's last century

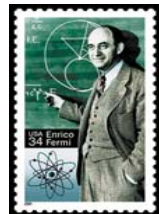


Schrödinger 1926
or maybe 1944 “What is Life?”

Take start as **1944** to get 1950s explosion: rocket science, mesons, DNA, BB

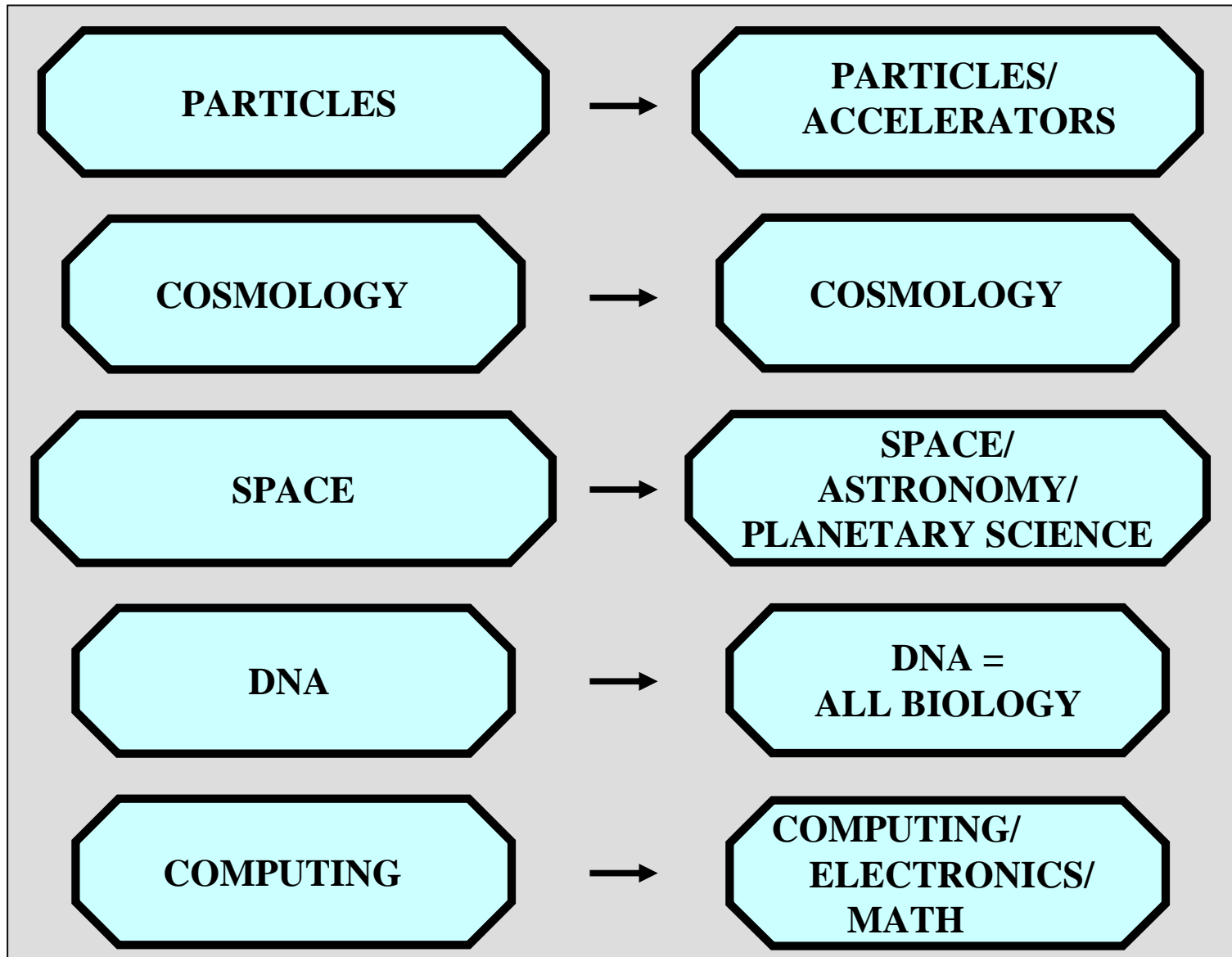


Non-Accelerator Particle Astrophysics School
D. Carrigan - Fermilab

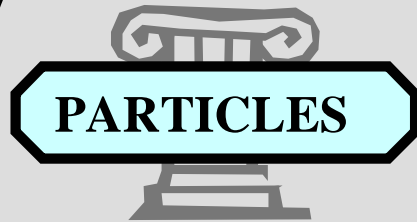


ICTP – Trieste
July 26 – Aug. 6, 2004

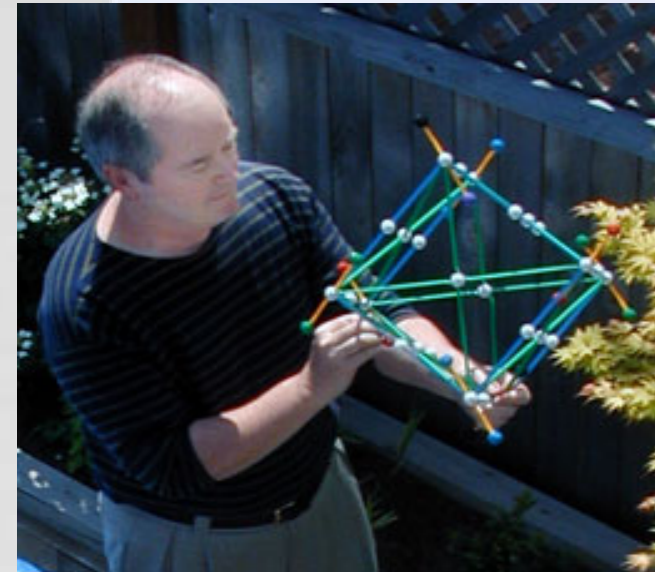
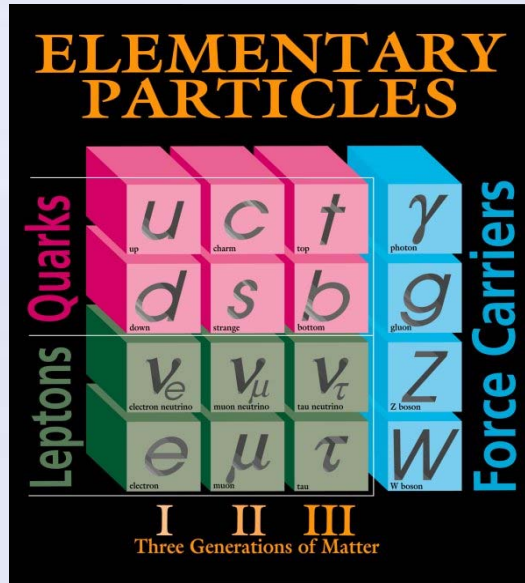
Definitions



Technical revolutions
of the last
half century



Particle physics



Chronology – mostly experimental discoveries

- 1897 Using a cathode ray tube Thomson discovers the electron at the Cavendish
- 1931 Anderson discovers the positron at Cal Tech
- 1934 [Fermi develops theory of beta decay](#)
- 1937 Neddermeyer and Anderson discover the muon in a cosmic-ray experiment
- 1951 First observation of strange particles in cosmic-ray experiments.
- 1955 Segre and Chamberlain discover the antiproton at Berkeley.
- 1956 Cowan and Reines detect the first neutrino at Savannah River.
- 1956 [Gell-Mann explains kaon lifetime with the strangeness](#)
- 1964 At Brookhaven Cronin and Fitch find kaons violate CP symmetry
- 1974 Physicists at SLAC and BNL independently discover charm quark
- 1977 The upsilon, a particle containing a bottom quark, is discovered at Fermilab
- 1979 Gluon observed at DESY.
- 1983 W and Z bosons observed at CERN
- 1995 CDF and D0 discover top quark at Fermilab

The future for particle physics

Some of Quigg's April 28, 04 "Questions for the Future"

1. Are quarks and leptons elementary?
2. What is the relationship of quarks to leptons?
3. Are there right-handed weak interactions?
4. Are there new quarks and leptons?
5. Are there new gauge interactions linking quarks and leptons?
6. What is the relationship of left-handed & right-handed particles?
7. What is the nature of the right-handed neutrino?
8. What is the nature of the new force that hides electroweak symmetry?
9. Are there different kinds of matter? Of energy?
10. Are there new forces of a novel kind?
11. What do generations mean? Is there a family symmetry?
12. What makes a top quark a top quark and an electron an electron?
13. What is the (grand) unifying symmetry?

From Quigg <http://www.fis.puc.cl/~rlineros/research/cdmexico/talks/BeyondSM/BSM1.pdf>

Accelerators



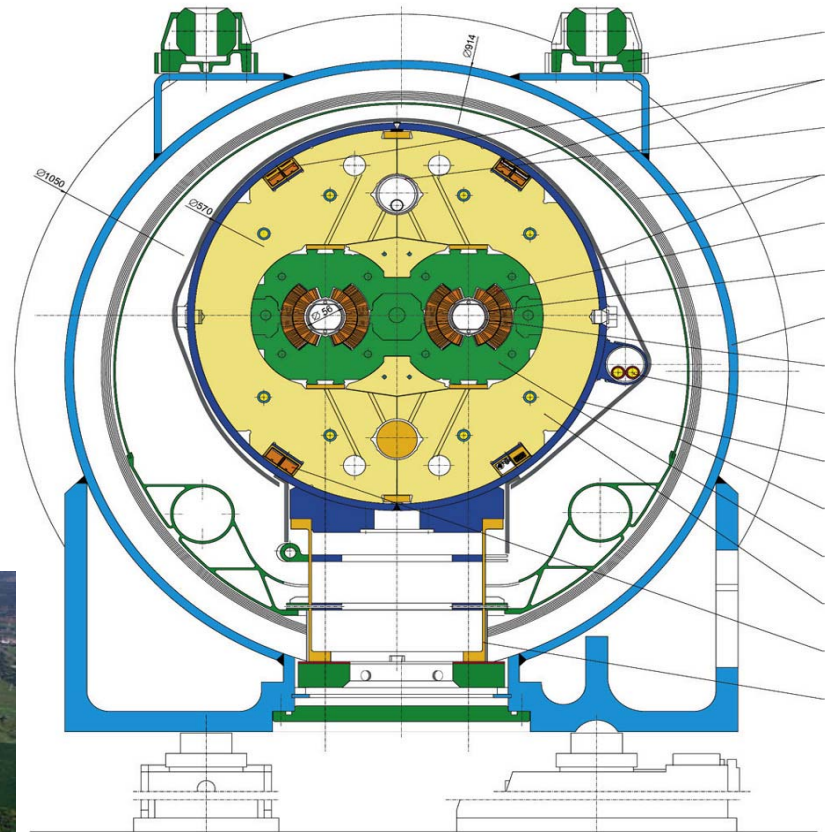
Fermi at Chicago
Synchrocyclotron - 1951



Non-Accelerator Particle Astrophysics School
D. Carrigan - Fermilab

LHC DIPOLE : STANDARD CROSS-SECTION

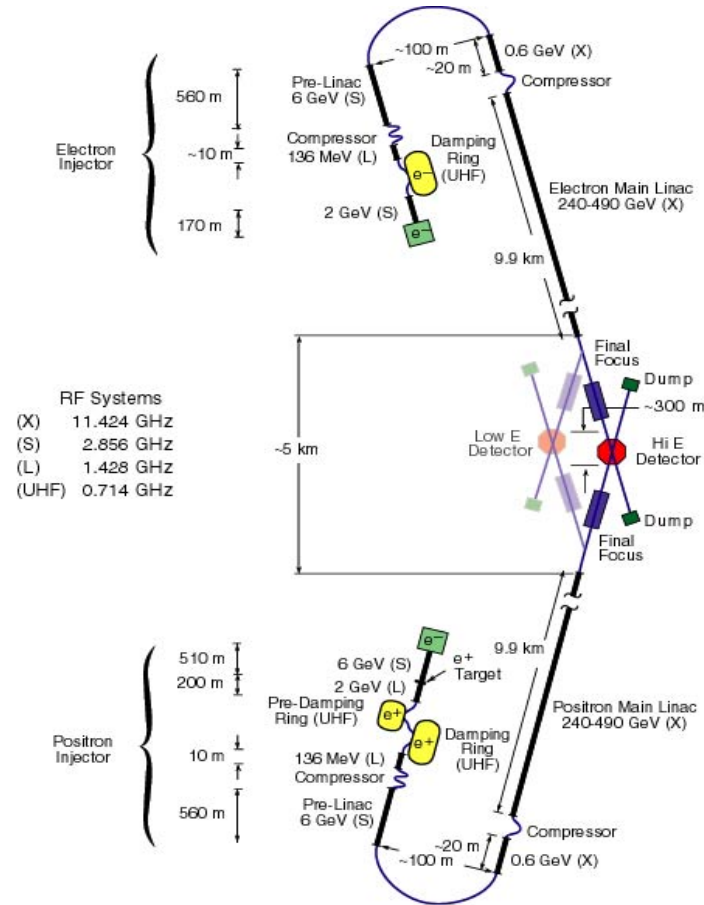
CERN AC/DI/MM - HE107 - 30 04 1999



ICTP - Trieste
July 26 - Aug. 6, 2004

Accelerators beyond LHC – everything has problems

NLC



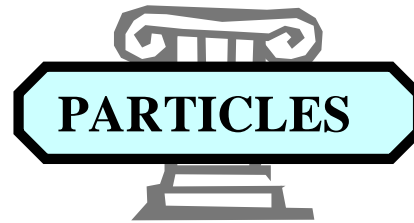
or Plasma wake field acceleration



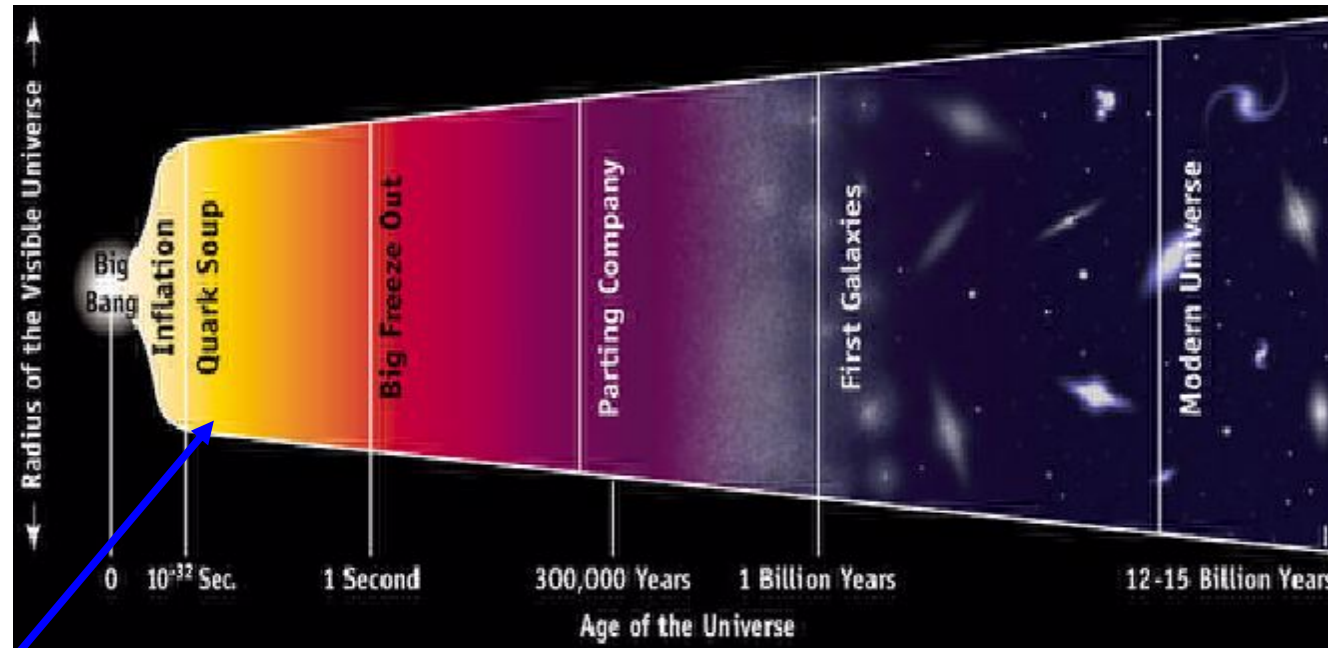
$$G = 0.96(n_0)^{1/2} \quad (\text{V/cm}) \quad n_0 \text{ is electron density}$$

RF cavity	0.0005 GV/cm
gaseous plasma	1 GV/cm
solid state plasma	100 GV/cm

Technical revolutions
of the last
half century



The Big Bang

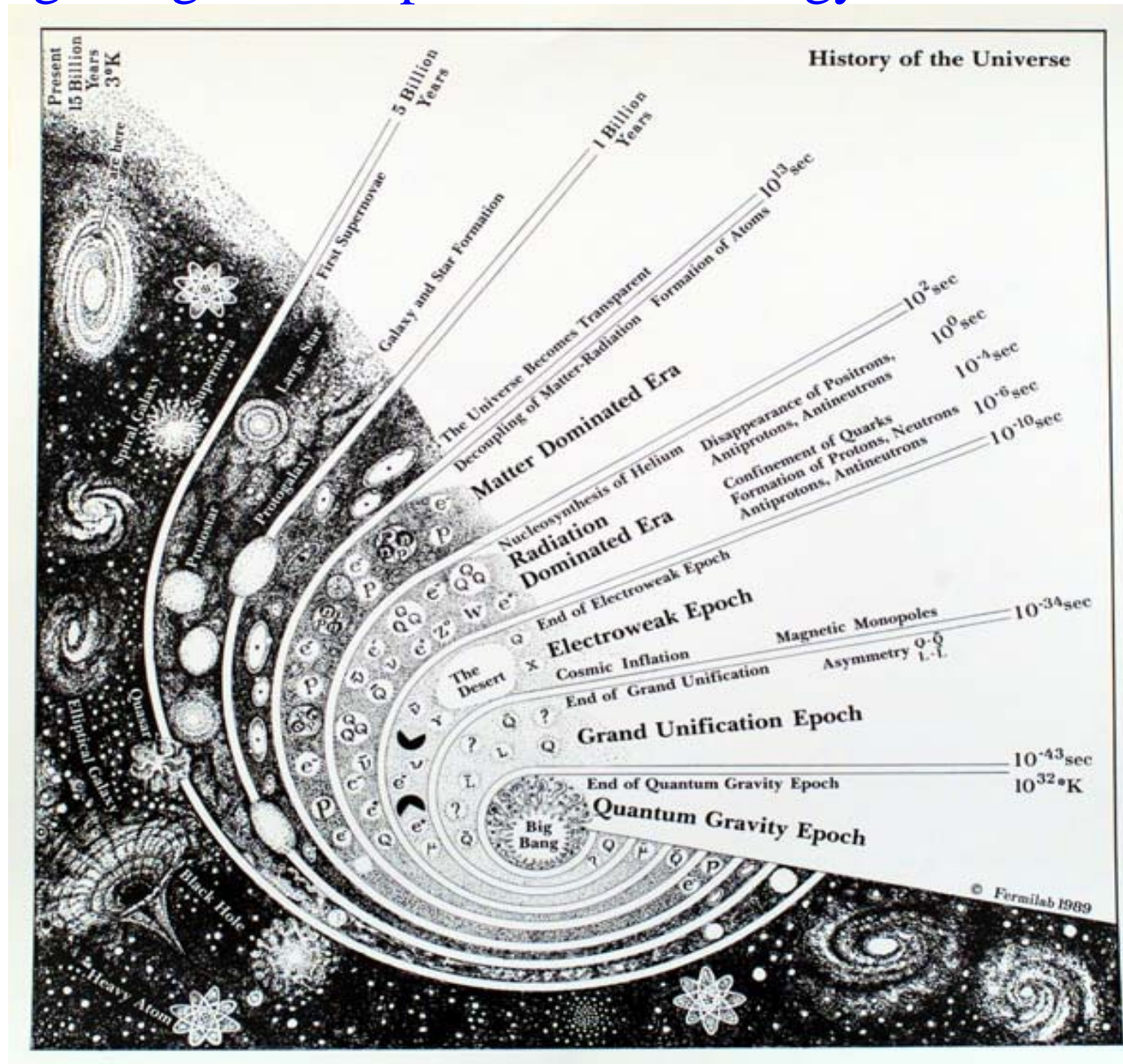


From <http://ist-socrates.berkeley.edu/~phy39/stellarE/starshine.html>

**PARTICLE
PHYSICS**

Big bang explained what came afterwards
but not the big bang itself

The Big Bang and the particle /cosmology connection



The next forty years in particle cosmology

Even more precise WMAP-type observations

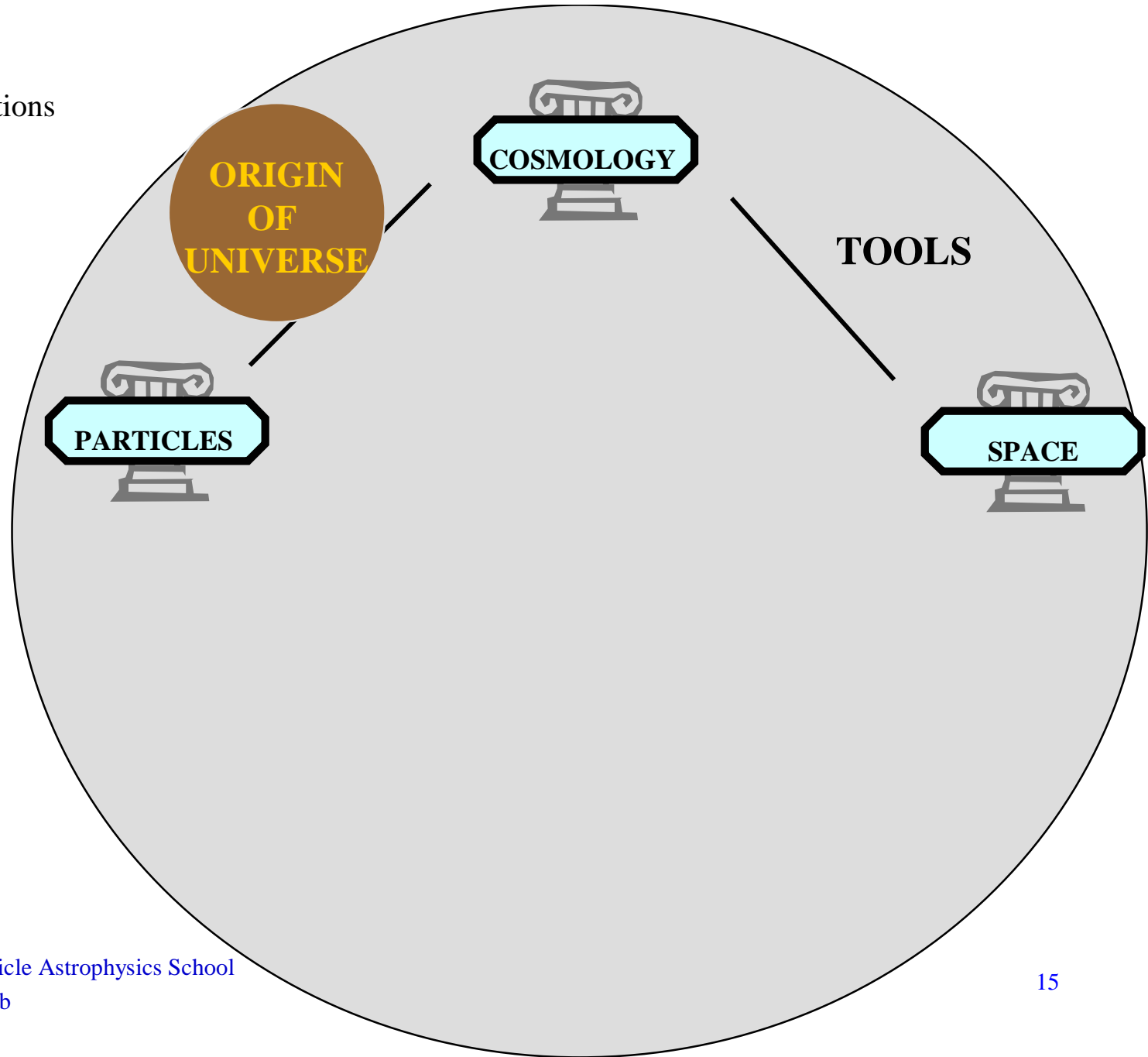
Dark matter and dark energy investigations continue

Gravity waves

The fabric of space

Particle physics finds the key?

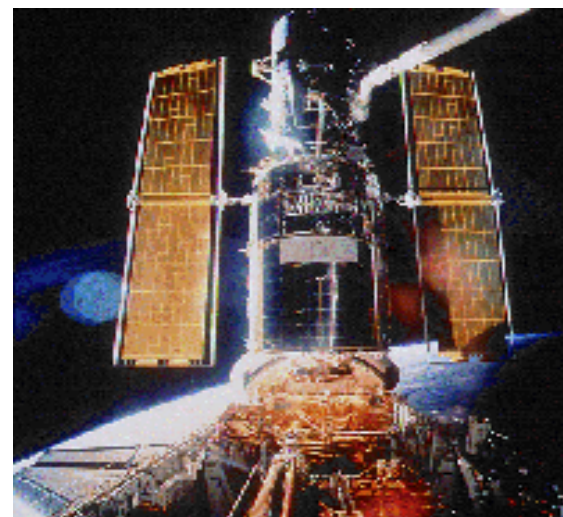
Technical revolutions
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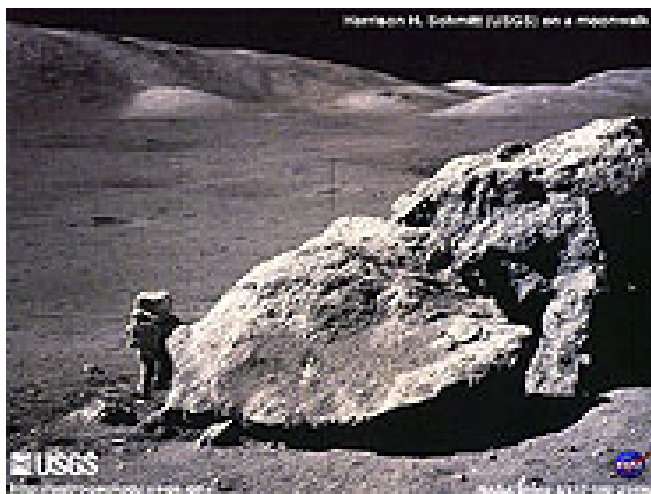
History of Space and Astronomy



V2 at White Sands
The birth of real
rockets at
Penemunde 1930s

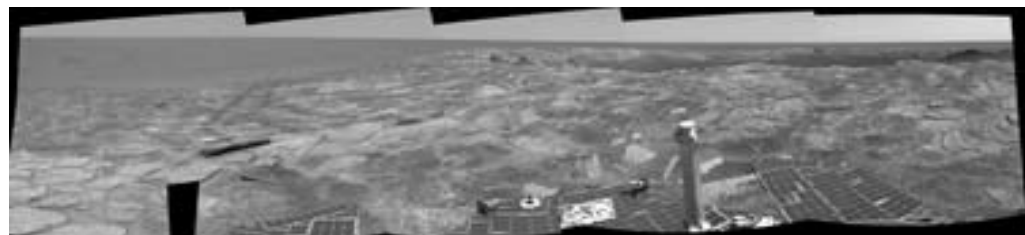


First Hubble service
mission in 1993

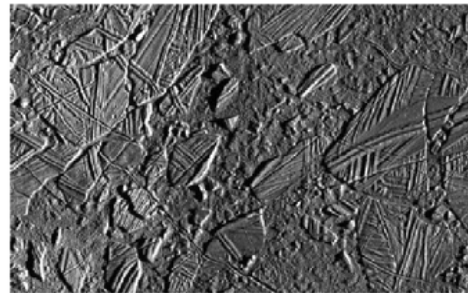
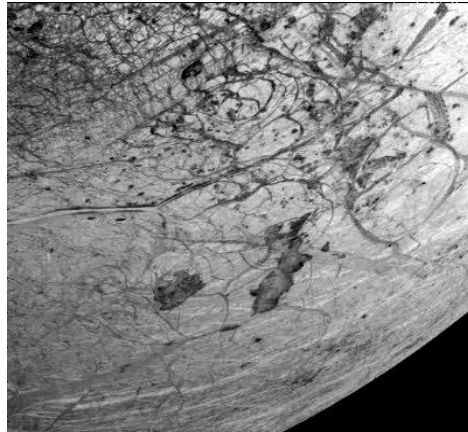


Harrison Schmitt on the Moon. NASA photo (1972)
First and only scientist on moon

NASA Rover robot on Mars 2004 looking for water

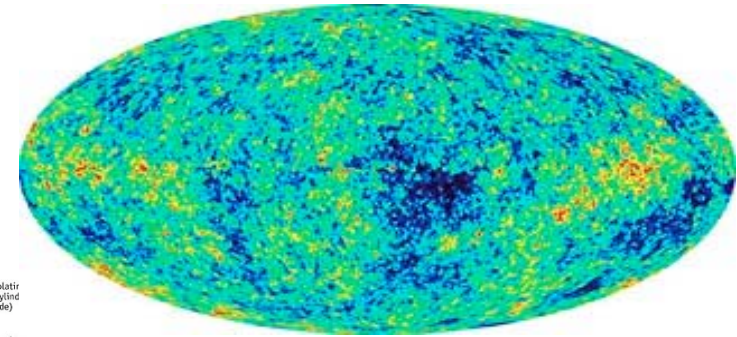
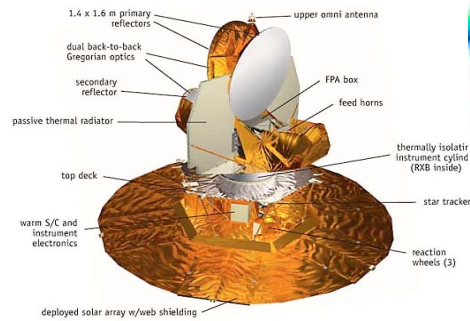


Space and Astronomy

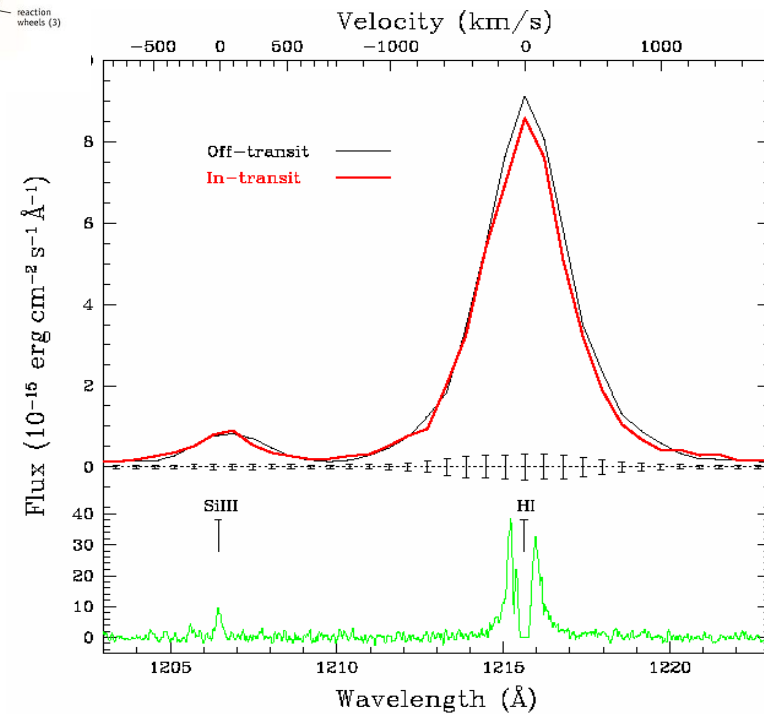


Icy surface of Europa
– an abode for life?

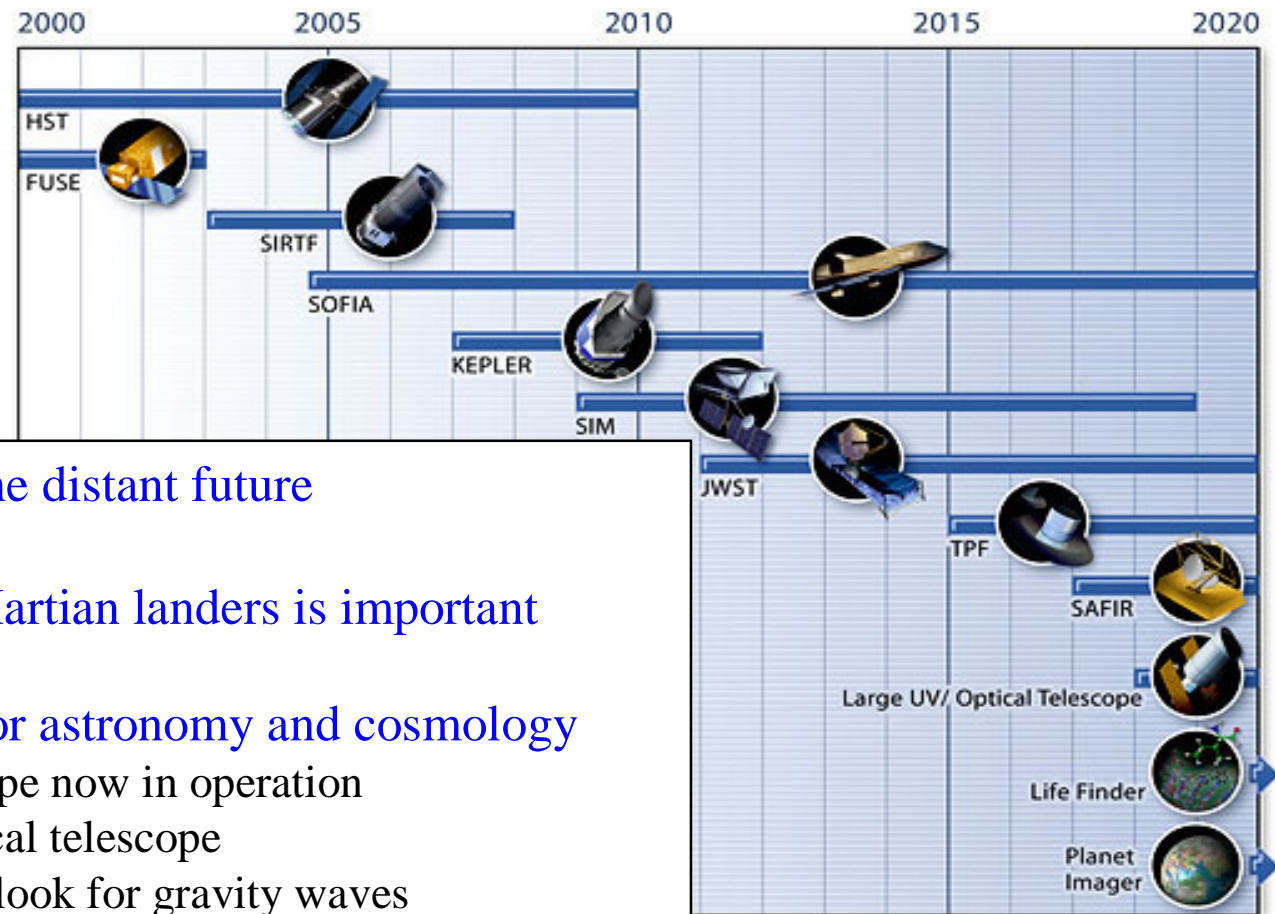
WMAP



O and C in atmosphere
of extrasolar planet
HD209458 observed
with the STIS
spectrograph on
board Hubble (2004)



The future of space



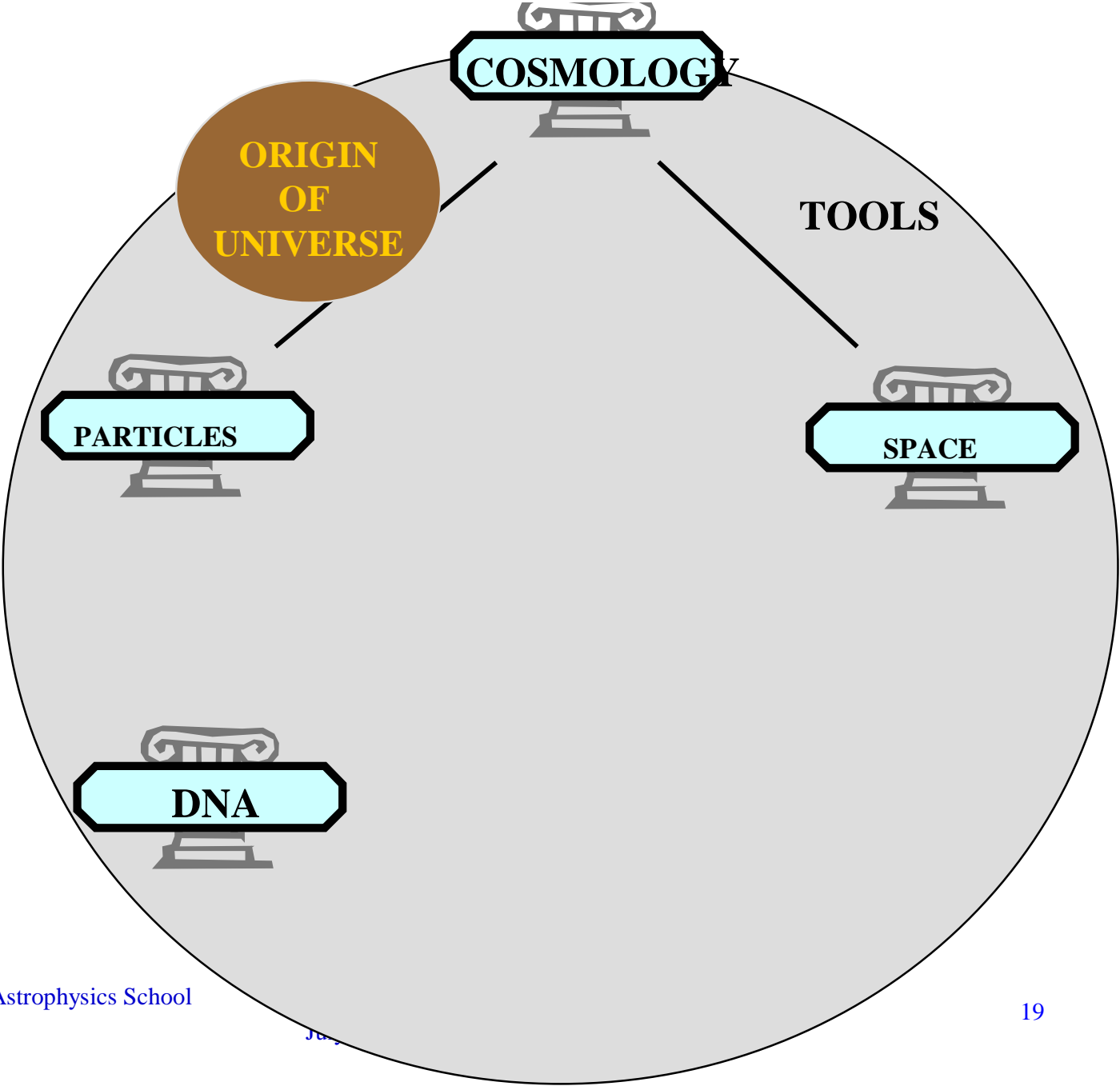
Man has a role but for the distant future

Robotics ala the 2004 Martian landers is important

Enormous importance for astronomy and cosmology

- Spitzer Infrared Telescope now in operation
- James Webb 6.5 m optical telescope
- LISA – constellation to look for gravity waves
- SIM – interferometry to look for exosolar planets
- Kepler - exosolar planets by transit

Technical revolutions of the last half century

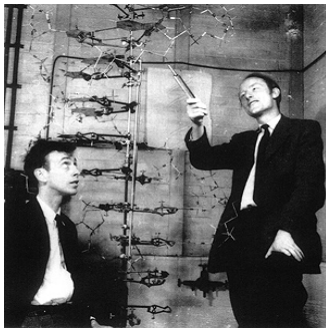


Dirt, Darwin, and DNA



The evolution of the geology of earth including impacts and the atmosphere drives the environments

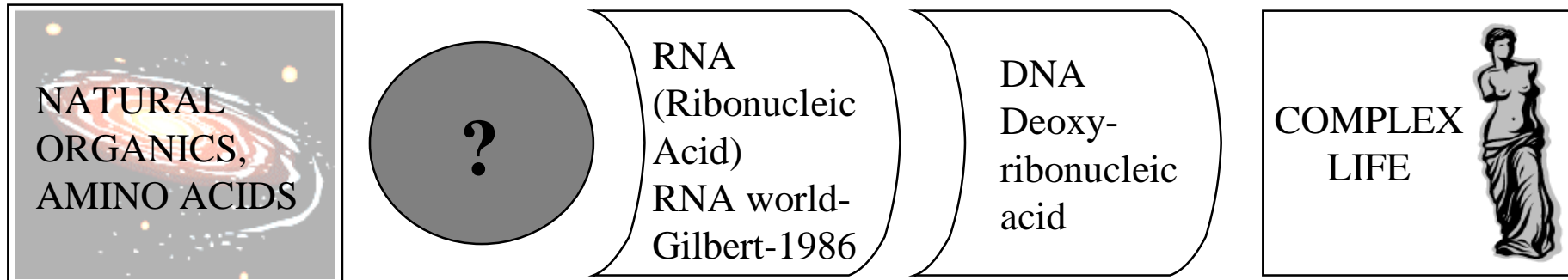
Survival of the Fittest for different environments explains evolution



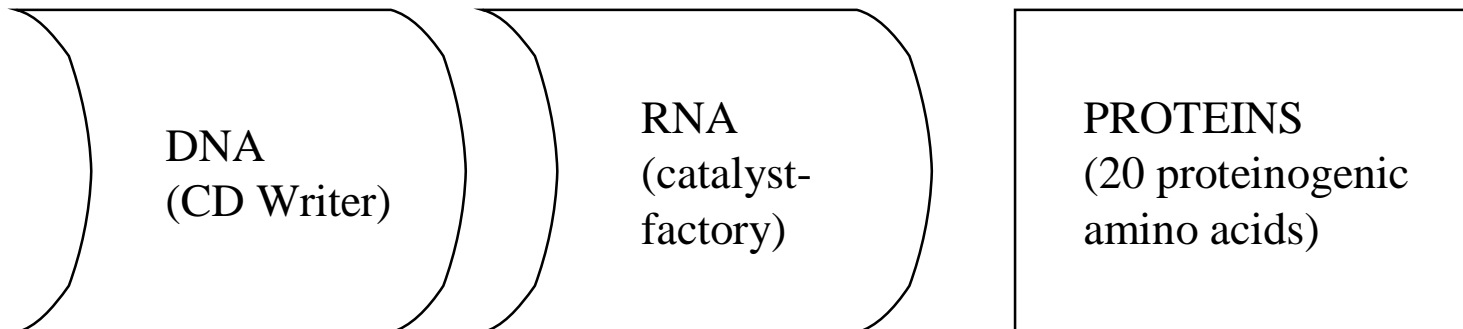
DNA (~1953) Proteins have 20 amino acids. There are 4 bases in DNA (adenine[A], guanine[G], cytosine[C], thymine[T]). Combinations are AT (2 hydrogen bonds) and GC (three hydrogen bonds). The number of bits/base pair is $4.32/3 = 1.44 \text{ bits/bp} = \ln_2(20)/3$ where 3 is the number of base pairs in a codon (a triplet of base pairs). 600 Daltons for mass of a base pair with backbone and bonds. Human genome-3 billion DNA base pairs, actual information content is on the order of 0.05 Gbytes because of junk DNA. Point- lots of information per mass.

Big Bang of Biology

ORIGIN OF LIFE

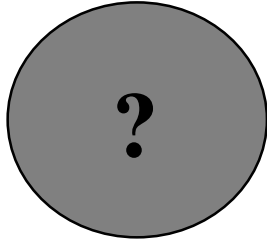


GENETIC INFORMATION FLOW

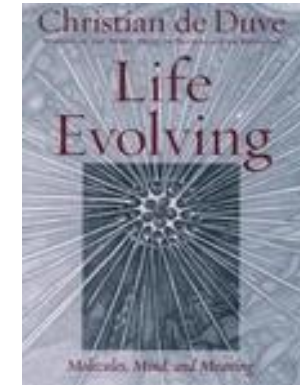


Origin of life

T



Don't know. Chirality may be a clue.
De Duve argues RNA by chance is implausible.
Needs a complex chemical environment.
Some argue catalysts are needed
but maybe not for peptides.
Maybe clays help.
Maybe a natural selection for molecules.

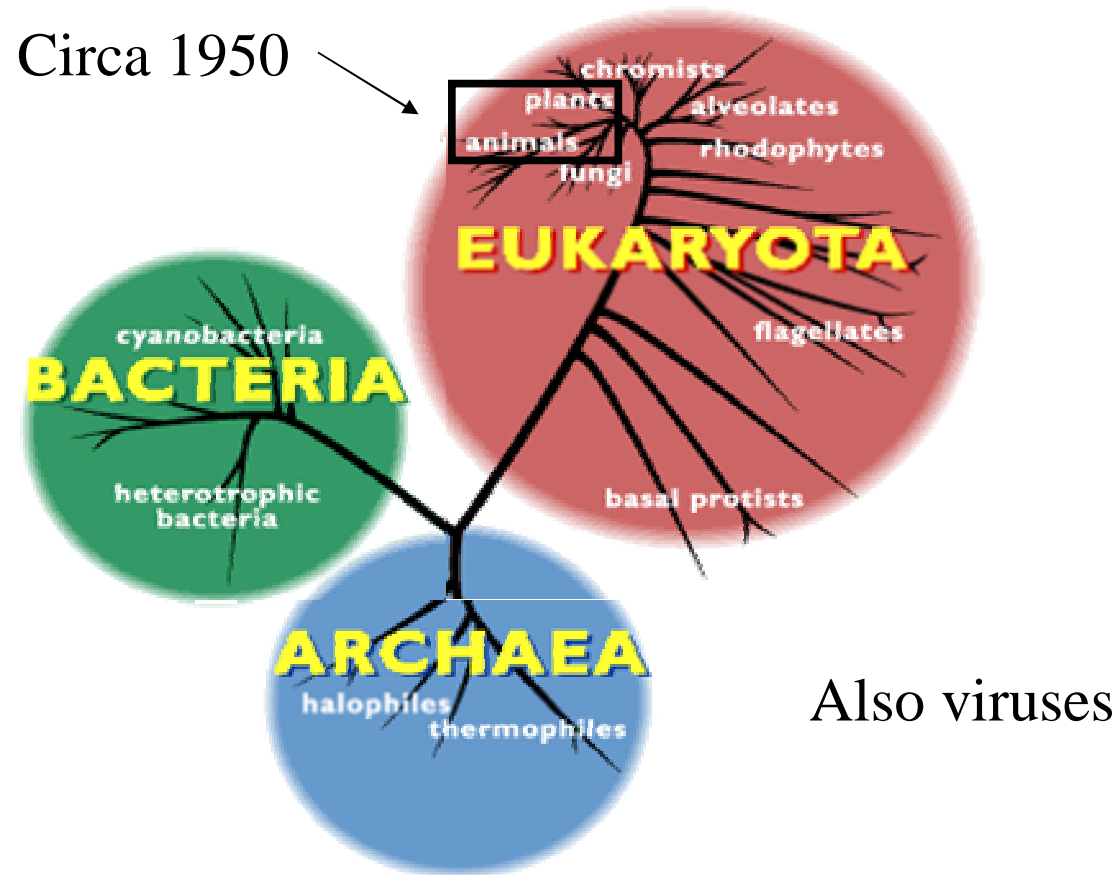


History of life on earth – the first billion years:

- Birth of the Solar System 4.6 billion years ago. Hot (molten) atmosphere of H₂O, CO₂ and CO, N₂ and H₂ until 0.1 Gyr
- Rain 0.1 to 0.3 Gyr, rocky crust 0.2 to 0.4 Gyr.
- Biologically processed carbon 1 Gyr (self-replicating, carbon-based microbial life)
- not much free oxygen - anaerobic life
- And then photosynthesis began to produce oxygen

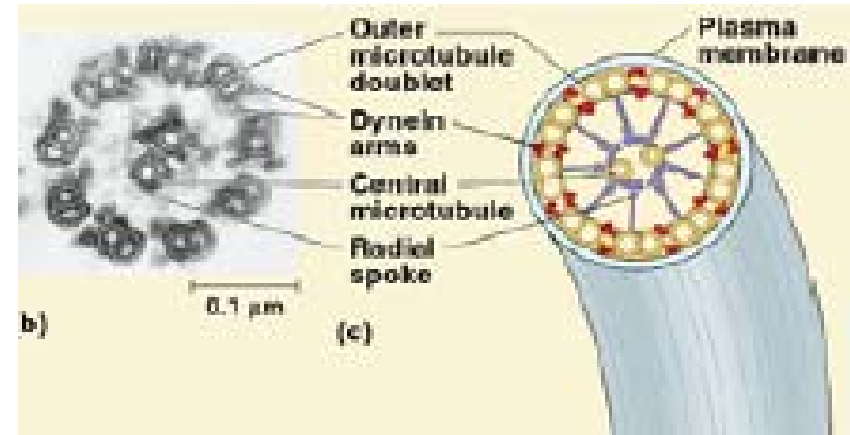
See :Astrobiology:The Quest for the Conditions of Life- Gerda Horneck Christa Baumstark-Khan (Eds.)

Tree of life – not your parents biology



From <http://www.ucmp.berkeley.edu/allife/threedomains.html>

Speeding up evolution



Extremophiles

Thermophiles-deep hydrothermal vents along mid ocean rifts - volcanic vents or "black smokers"

Salty (halophiles)

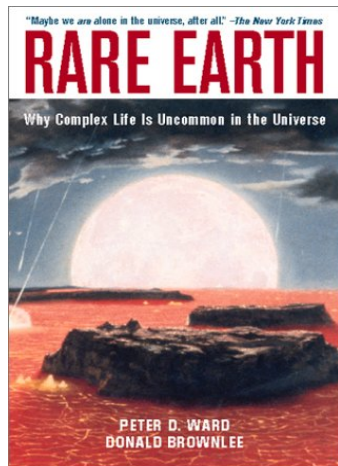
Cold, dry, high radiation, ...

Symbiosis-Margulies-speeds up

The astrobiology link

Extraterrestrial signatures of life – the next fifty years

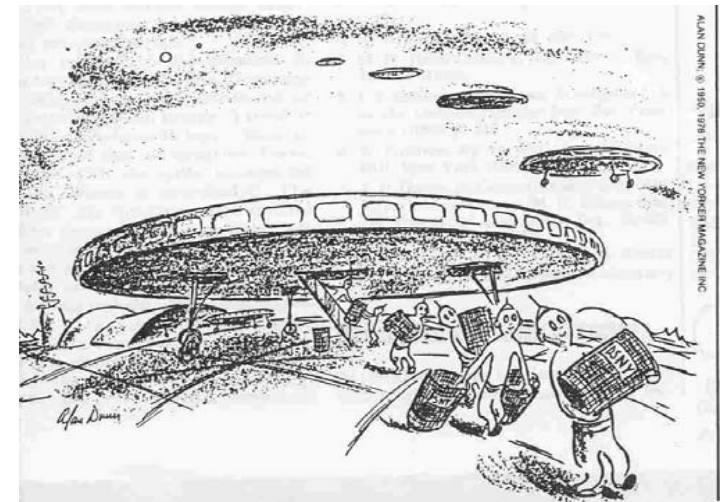
- Meteors – Antarctic collection
- Planets and satellites (Mars, Europa) - Rovers
- Atmospheres of extrasolar planets – spectroscopy, interferometry
- Planetary and galactic habitable zones



Extraterrestrial intelligence – SETI

- how frequent (if at all)?
- the Fermi question or paradox

From the New Yorker 20 May 1950,
Physics Today August 1985





**“What is Life?”
Erwin Schrödinger (1944)**

Anthropic Principle – a perfect universe for us

Brandon Carter (73)

Cosmology and the standard model

James D. Bjorken*

Stanford Linear Accelerator Center, Stanford University, Stanford, California 94309

Received 21 October 2002; published 26 February 2003

PHYSICAL REVIEW D **67**, 043508 (2003)

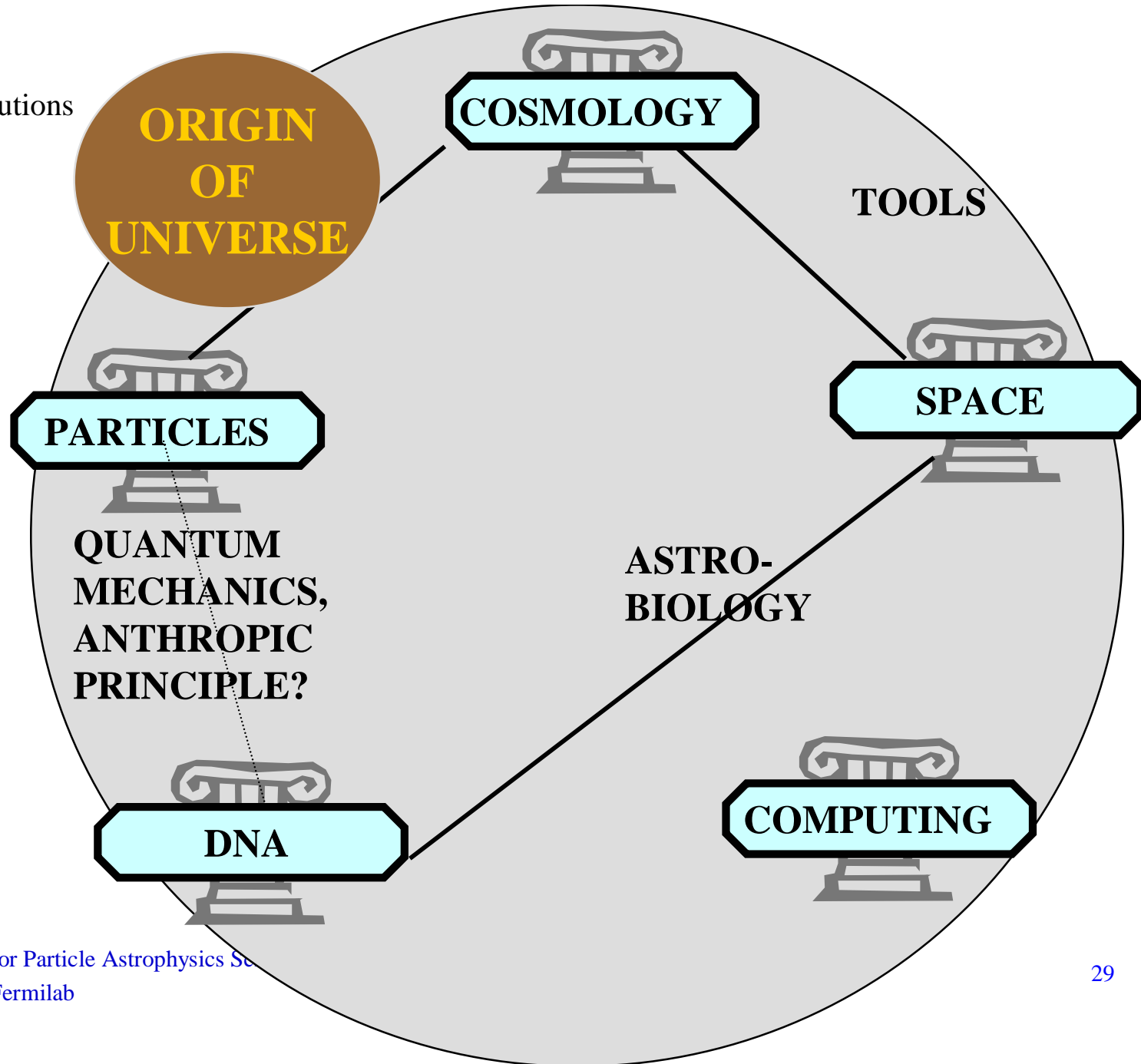
“... it is well known, ... that other properties of our universe are very finely tuned and will only exist over a quite small bandwidth. We shall pay special attention to such “anthropic” constraints, as discussed for example in the book by [Barrow and Tipler](#) and will be interested in the bandwidth in R for which they are satisfied.”

.... Crucial to the properties of nuclear and atomic matter are the values of the fine-structure constant ~here constrained to a reasonable range of values, the ratio of electron to proton mass, the ratio of pion to proton mass, and the neutron proton mass differenceFinally, we may consider the mechanism for producing carbon in stars. This depends upon the existence of the anthropically famous **triple-alpha reaction** ...with the resonance in ^{12}C predicted by Hoyle ... together with the absence of a crucial level in ^{16}O .”



**“What is Life?”
Erwin Schrödinger (1944)**

Technical revolutions
of the last
half century



Computing

This computer – 18.5 Gbytes, 650 MHz.



Tim Berners-Lee

can send questions and concerns for immediate response.

Part of the package for the new site includes templates and instructions so that anyone at Fermilab can construct a page, or redesign existing pages, to match



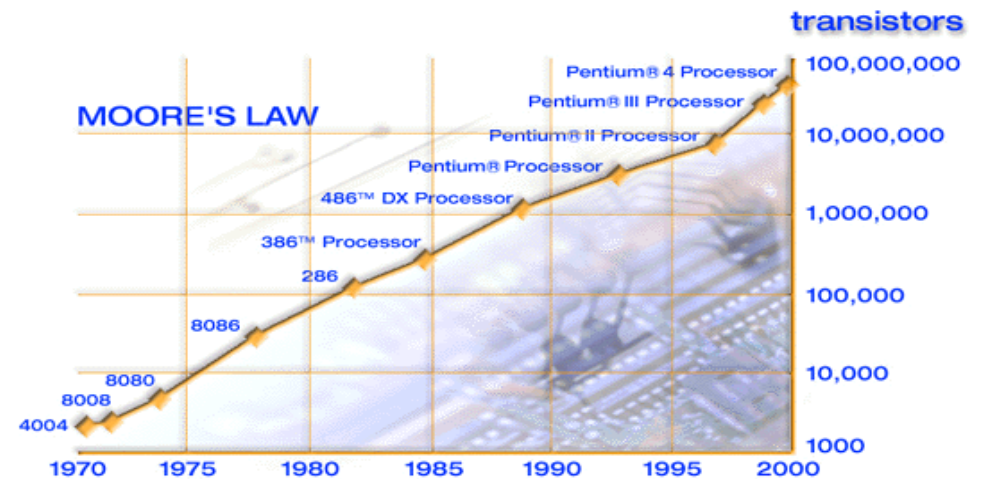
Historical (Web) Site Marker

On this site (www.fnal.gov) was established in June 1992 either the third website in the United States. The World Wide Web was born at Europe in 1991 as a tool for exchanging particle physics data. The first webserver was created at Stanford Linear Accelerator Center in Dec

In June 1992, Fermilab's Computing Division installed its first website at about the same time as a similar installation at the Massachusetts Institute of Technology. In late 1992, Computing Division staff created Fermilab's first html page. In 1993, the National Center for Supercomputing Applications at the University of Illinois launched Mosaic, a graphical interface Web browser that made the Web navigable for people without computer expertise.

In February 1994, Fermilab created the laboratory's first pages designed for the public. The public website had 12,000 hits on April 27, the day an announcement of the first evidence for the top quark.

In August 1996, the laboratory redesigned its growing volume of public web pages. A complete overhaul of the Fermilab website appears on May



<http://www.intel.com/research/silicon/mooreslaw.htm>

Size of some databases

Human genome-3 billion DNA base pairs,
actual information content is on the order
of 0.05 Gbytes (about size of Word)

Old Fermilab tape robot – 6×10^6 Gbyte –
mostly colliding beam, Monte Carlo

Sloan Telescope at Fermilab – 1000 Gbytes

10^6 Gbytes of material printed every year

Typical education through graduate school
subsumed in 1-10 Gbytes

Lifetime of images stored on DVD might be
1000 Gbytes

Knowledge base in a human brain-0.25 to 2.5
Gbyte range (10^{11} neurons). 6×10^9 people
gives 1.5 to 15×10^9 Gbytes to profile
everyone on earth



Old Fermilab tape robot

Computing and the future

Computing speed

Speed of human brain - 200×10^3 Giga computations/s (Kurzweil) or 10^5 times this computer.

Kurzweil calls the crossover a singularity
(maybe better is “phase change”)

Cross over point is 2020 to 2030 at present rate of progress

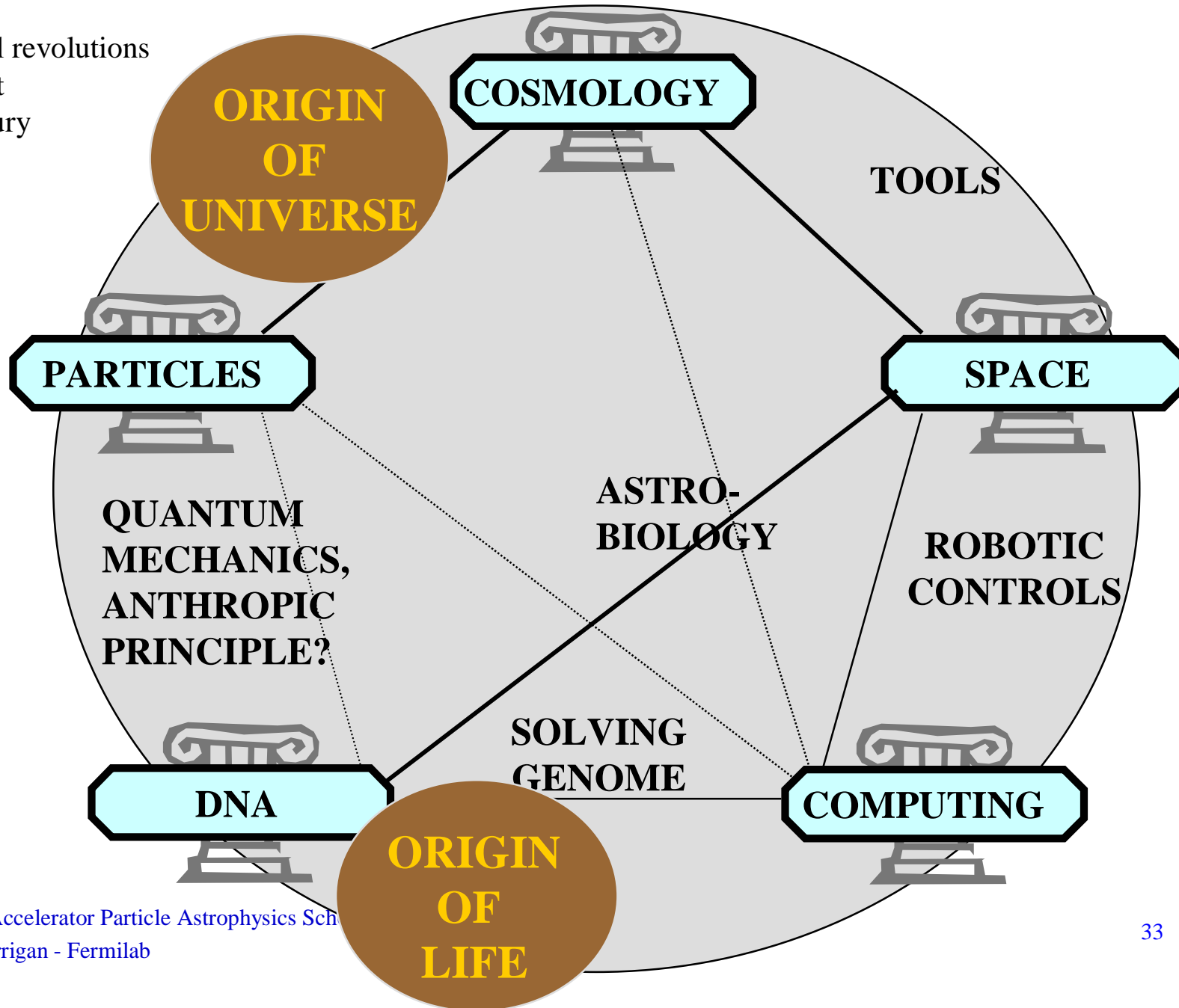
QCD - 1000 Giga Flops – getting in range

Quantum computing

Problems and challenges

- viruses
- the approach of the Kurzweil “singularity” (or phase change)
- qualifying internet and World Wide Web
- qualifying programs
- and ... **theories of computing, knowledge, mind**

Technical revolutions
of the last
half century



The bottom line



NASA Origins Program



- Particle physics and cosmology linked - explains much of the universe
- Biology has become a mathematical science
- Scientific experiments have moved into space
- Computers and computer science are reaching human capabilities
- We may find extraterrestrial life

Questions?