

the **abdus salam** international centre for theoretical physics

ICTP 40th Anniversary

H4.SMR/1574-1

"VII School on Non-Accelerator Astroparticle Physics"

26 July - 6 August 2004

Future Accelerators, Muon Colliders, and Neutrino Factories

R. Carrigan

Fermi National Accelerator Laboratory Batavia, U.S.A.

# Future Accelerators, Neutrino Factories, and Muon Colliders

Dick Carrigan Fermilab

## Six great questions

- SM, SUSY, Higgs
- extra dimensions-may be related to SUSY
- neutrino mass and oscillations
- Missing mass and energy
- origins of CP
- none of the above

#### Plan of Talks

Current status Tevatron, LHC accelerator technologies Midrange possibilities Linear colliders gamma-gamma colliders free electron lasers Neutrino Facilities NUMI proton drivers neutrino factories muon colliders Very large hadron colliders Visionary possibilities laser accelerators plasma accelerators

Non-Accelerator Particle Astrophysics School D. Carrigan

Accelerators and storage rings in operation or underway

The Tevatron LHC-this is the 800 lb gorilla of the future (about 2007)

Non-Accelerator Particle Astrophysics School D. Carrigan

#### The Tevatron



Non-Accelerator Particle Astrophysics School D. Carrigan

#### **Tevatron schematic**

#### FERMILAB'S ACCELERATOR CHAIN



Non-Accelerator Particle Astrophysics School D. Carrigan

### **Tevatron luminosity**



Peak as of July 7, 2004 is 92.1E30 cm<sup>-2</sup>s<sup>-1</sup>

Non-Accelerator Particle Astrophysics School D. Carrigan

ICTP – Trieste July 26 – Aug. 6, 2004

7



Non-Accelerator Particle Astrophysics School D. Carrigan

#### LHC magnet

#### LHC DIPOLE : STANDARD CROSS-SECTION

CERN AC/DI/MM - HE107 - 30 04 1999



Non-Accelerator Particle Astrophysics School D. Carrigan

ICTP – Trieste July 26 – Aug. 6, 2004

9

## Physics and technology of accelerators

#### References

Edwards and Syphers, Tigner and Chao

<u>ion sources</u> (anti proton accumulation a special case)

<u>acceleration</u>-via RF (a limit for linear colliders)

<u>bending</u>-via magnets-typically superconducting at high energy (a limit)

focusing

<u>colliders</u>-luminosity important

an enclosure (limit)

#### Ion sources

#### Fermilab Cockcroft-Walton



Non-Accelerator Particle Astrophysics School D. Carrigan

ICTP – Trieste July 26 – Aug. 6, 2004

#### Fermilab Anti-proton source



11

#### Acceleration



#### Fermilab Linac RF cavity

Non-Accelerator Particle Astrophysics School D. Carrigan

#### An enclosure



#### Tevatron tunnel with old Main Ring in place

Non-Accelerator Particle Astrophysics School D. Carrigan

#### Mid-range possibilities

Linear colliders gamma-gamma colliders free electron lasers Neutrino facilities NUMI Proton drivers neutrino factories muon factories Very large hadron colliders

Non-Accelerator Particle Astrophysics School D. Carrigan

## Mid range possibilities: linear colliders

done at SLAC with 50 GeV on 50 GeV need

500 –1000 GeV total energy L=  $10^{34}$  /cm<sup>2</sup>s... luminosity <u>physics reach</u>-must compliment LHC Glashow-Lane skeptical (DOE 4/2001) polarization important Butler-sit, span (pk ener), scan-possible 300 fb<sup>-1</sup> and 250 fb X sec gives 75 K Higgs O(1 yr)

Non-Accelerator Particle Astrophysics School D. Carrigan

linear colliders - continued

possibilities

JLC (Japan), NLC (US), TESLA (DESY), CLIC (CERN) <u>political process</u> – technology choice end of 2004 build a test bed somewhere? <u>gamma-gamma collider</u>

Non-Accelerator Particle Astrophysics School D. Carrigan

#### Linear Collider luminosity

#### PERFORMANCE LIMITATIONS IN ELECTRON-POSITRON LINEAR COLLIDERS

The performance of a high energy electron-positron collider is characterized by the <u>energy</u> and the <u>luminosity</u>.

Because of the inverse square relationship between reaction cross sections and center-of-mass energy, a next generation electron-positron collider must target a luminosity in the range 5-50x10<sup>33</sup>cm<sup>-2</sup>sec<sup>-1</sup>.

⇒Nearly all technical issues in electron colliders are related to the achievement of high luminosity.

The luminosity in any collider is given by

$$L = \frac{fN_1N_2}{4\pi\sigma_x\sigma_y}F$$

where f is the frequency of collisions between bunches,  $N_1$  and  $N_2$  are the number of particles in the colliding bunches,  $\sigma_x$  and  $\sigma_y$  are the transverse beam dimensions, and F is a form factor (usually ~1) related to the specifics of the collision geometry.





#### CLIC - Compact Linear Collider – 3 TeV



From http://clic-study.web.cern.ch/CLIC-Study/Layout/OverallCLIC3.html

## Linear collider comparisons

#### From http://clic-study.web.cern.ch/CLIC-Study/Parameters/ILC\_TRC/Table1\_1W.pdf

8/13/02

	TES	SLA		JLC (C)	JLC/NI	$LC^{*}(X)$	CI	IC
Center of mass energy	$500 { m GeV}$	$800  {\rm GeV}$	$500~{\rm GeV}$	$1000 { m ~GeV}$	$500 { m GeV}$	$1000~{\rm GeV}$	$500~{\rm GeV}$	$3000~{\rm GeV}$
RF frequency of main linac (GHz)	1	.3	5.7	$5.7/11.4^{\P}$	11	4	3	0
Design luminosity $(10^{33} \rm cm^{-2} s^{-1})$	34.0	58.0	16.8	25.0	25.0(20.0)	25.0(30.0)	21.0	80.0
Linac repetition rate (Hz)	5	4		100	150(120)	100(120)	200	100
No. of particles/bunch at IP $(10^{10})$	2	1.4		0.75	0.	75	0	.4
No. of bunches/pulse	2820	4886		192	19	92	1	54
Bunch separation (nsec)	337	176		1.4	1.	.4	0.	67
Bunch train length $(\mu sec)$	950	860		0.267	0.2	267	0.1	102
Beam power/beam (MW)	11.3	17.5	5.8	11.5	8.7(6.9)	11.5(13.8)	4.9	14.8
Unloaded/loaded gradient <sup>†</sup> (MV/m)	$23.4 \ / \ 23.4$	35 / 35	41.8/31.5	$41.8/31.5 \ / \ 70/54$	70 ,	/ 54	172 ,	/ 150
Total two-linac length (km)	30	30	17.1	29.2	12.6	25.8	5.0	28.0
Total beam delivery length (km)	÷	3		3.7	3	.7	5	.2
Proposed site length (km)	3	3		33	3	2	10.2	33.2
Total site AC power <sup>‡</sup> (MW)	140	200	235	310	215~(185)	280(320)	175	410
Tunnel configuration <sup>#</sup>	Sin	gle		Separate	Sepa	arate	Two-	Beam
		Tesla			NLC			
selling fe	selling features long pulse			high grad				
_	low freq			expansion				
		reduced wal	ka fialds					
		around moti	ion					
			1011					
expansion paths double RF								
		2 deg K						
		electro polis	sh					
flat beam	flat beam development @ A0 may eliminate 1 damping ring @ Tesla							
Non-Accelerator Particle Astrophy	sics School	- ICT	· P – Trieste		-			
D. Corrigon	Sies Senoor	July 26		04			21	
D. Camgan		July 26	– Aug. 6, 20	04				

#### Gamma-gamma collider

Proposed by Ginzburg et al. (1982) for producing a photon collider

Collide a high power laser pulse with an electron beam to produce a high energy photon beam

```
laser: 1 J, 1.8 ps FWHM, 1 micron, > 1 TW, 10 KW av
```

Two body process

Correlation between outgoing photon angle and energy  $(1/\gamma)$ 

Maximum energy when the photon is co-linear with the incoming electron (about 80% of e)



#### Free electron laser

# SASE-self amplified spontaneous emission first proposed at SLAC Brilliance is 10<sup>8</sup> times current light sources, 1 Å x-rays, pulse length of 100 fs

#### TTF demonstrated SASE @ 80-180 nm



From Tesla DR 3.6 The X-ray Free Electron Laser (XFEL)  $\ensuremath{\mathsf{I}}\xspace$ -37

Non-Accelerator Particle Astrophysics School	ICTP – Trieste
D. Carrigan	July 26 – Aug. 6, 2004

Process: small xsec, high current e beam synchrotron radiates in undulator. For correct energy, undulator period get resonance in longitudinal charge density modulation

-micro bunching. Number of photons grows exponentially

Mid range possibility - hot neutrino facilities

Conventional beams Neutrino factories Reactors-not accelerators

Non-Accelerator Particle Astrophysics School D. Carrigan



NUMI – neutrinos at the Main Injector





**FERMILAB #98-1321D** 

The NUMI project

120 GeV Protons from Fermilab Main Injector

 $10\mu s$  pulse, every 1.9s

Proton Intensity:

- 4x10<sup>13</sup> protons/pulse design
- 2.5×10<sup>13</sup> p/p expected at startup

Hadrons focused with 2 horns

 Select beam energy spectrum by adjusting horn and target positions

Project will be complete/ commissioning starts Dec. 2004

0.2 MW first year 0.4 MW design

Non-Accelerator Particle Astrophysics School D. Carrigan

#### NUMI – layout



#### NUMI - construction

Target Hall



Decay Pipe Endcap at Absorber Hall



Non-Accelerator Particle Astrophysics School D. Carrigan

#### Some neutrino beam naming jargon

NUMI0.4 MWSuperbeam2 MWNeutrino factorypower about like superbeammore flexible on unlike sign suppressioncan do ve

Beta-beam neutrinos via radioactive nuclei <u>complicated</u>

Non-Accelerator Particle Astrophysics School D. Carrigan

## Going to superbeam J. Hylen

# Difficulty of handling 2 MW instead of 0.4 MW depends significantly on:

•Proton beam energy still 120 GeV? (windows more problematic if lower)

•Increase repetition rate? (more cooling/electrical power for magnets)

•Increase protons per spill? (more stress on target, windows, horn)

# General issues with higher beam power:

•Getting the average beam heating load out

•Average thermal stress limit

•Radiation safety

Groundwater Protection

Airborne Activation

**Prompt Radiation** 

Residual Activation

•Radiation damage lifetime of materials

•Thermal shock

•Mis-alignment from thermal expansion

•Ionization leakage current

## Possible Fermilab proton drivers



SYNCHROTRON (yellow)
Sited West of the existing booster
Re-uses existing linac enclosure
8 GeV LINAC (red)
Baseline Site injects at MI-30
 straight section

Others possible

http://www.fnal.gov/orgs/fermilab\_users\_org/users\_mtg/2004/foster.pdf0

Non-Accelerator Particle Astrophysics School D. Carrigan

#### Fermilab 8 GeV Superconducting Linac

#### New idea incorporating concepts from both the Spallation Neutron Source (SNS) and Tesla.

Copy SNS Linac up to 1.3 GeV

Use Tesla cryomodules from 1.3 to 8 GeV

H<sup>-</sup> injection at 8 GeV in Main Injector

2 MW beam power at BOTH 8 GeV and 120 GeV

Small emittances so that there are small losses in Main Injector

#### A draft design study exist. Cost comparable to Main Injector.

ICTP – Trieste

July 26 – Aug. 6, 2004

http://www.fnal.gov/orgs/fermilab\_users\_org/users\_mtg/2004/foster.pdf0

Non-Accelerator Particle Astrophysics School D. Carrigan

32



http://www.fnal.gov/orgs/fermilab\_users\_org/users\_mtg/2004/foster.pdf0

## Questions?

Non-Accelerator Particle Astrophysics School D. Carrigan

#### Future Accelerators, Neutrino Factories, and Muon Colliders continued

Dick Carrigan Fermilab

So far have talked about: Current status Tevatron, LHC accelerator technologies Midrange possibilities Linear colliders gamma-gamma colliders free electron lasers Neutrino Facilities NUMI proton drivers



## Mid range possibility muon collider

Muon storage ring-old concept recent BNL result for g-2 physics and Higgs  $\sigma_H \sim (m_\mu/m_e)^2$ problem-collecting the muons muons equal neutrinos-gave rise to idea of neutrino factories

Non-Accelerator Particle Astrophysics School D. Carrigan

#### Muon collider-concept

Produce pions with proton accelerator

pions decay:  $\pi \rightarrow \mu + \overline{\nu}$  30 MeV/c muons decay:  $\mu \rightarrow e + \nu + \overline{\nu}$   $\tau = 2.2 \,\mu s$ 

focus pions, muons-need a lens

cool muons-ionization cooling6 dimensional phase space

#### accelerate very fast

Non-Accelerator Particle Astrophysics School D. Carrigan



# **Muon collider schematic** (taken from Bruce King-BNL, 4th International Conference on the Physics Potential and Development of mu+mu- Colliders, San Fransisco, 1997.)





#### VLHC super ferric magnet



Superconducting coil in center energizes both fields 2 cm gap, 2T return loop is below

(from VLHC design report-2001)

Non-Accelerator Particle Astrophysics School D. Carrigan

#### Accelerator status recapitulation

this decade-Tevatron, LHC NUMI next decade-ambitious future linear collider Neutrino factory? muon factory? VLHC? 2020 New technology, need a vision

Non-Accelerator Particle Astrophysics School D. Carrigan

#### Visionary possibilities for acceleration

Would like much higher accelerating gradients

**Two thoughts:** 

Lasers

R. Palmer, Particle Accelerators V11, 81 (1980). Recent progress Kimura et al. PRL **92**, 054801 (2004). See also LEAP at Stanford (Colby)

Plasmas

Tajima and Dawson PRL **43**, 267 (1979)

E. Esarey, et al., IEEE Trans. On Plasma Sci, 24, 252 (1996).

J. Dawson, Scientific American March, 1989 (p. 54)

Non-Accelerator Particle Astrophysics School D. Carrigan

#### Lasers

basic laser challenge

good news: can get very high fields

bad news: vectors transverse to particle direction

ways to defeat

gratings, maybe boundary conditions, special modes

R. Palmer, Particle Accelerators 11, 81 (80)

Inverse free electron laser IFEL-next transparency

Non-Accelerator Particle Astrophysics School D. Carrigan

#### Cascading laser stages

[from W. Kimura et al, PRL 86, 4041 (2001)]

Inverse free electron laser (IFEL)

electrons oscillate in undulator and absorb energy from laser

Gradients not on a scale with plasma accelerators



FIG. 1. Schematic layout for the STELLA experiment. For size reference, the distance separating the two IFELs is 2.3 m and the laser beams enter the beam line  $\approx 6$  m apart.

Require fs micro bunches, very good timing

24 MW first stage, 300 MW second

This demonstrated rephasing, not acceleration

#### Plasma wake field acceleration



Photo S. Carrigan

 $G= 0.96(n_0)\frac{1}{2}$  (V/cm) RF cavity gaseous plasma

Non-Accelerator Particle Astrophysics School D. Carrigan

ICTP – Trieste July 26 – Aug. 6, 2004 n<sub>0</sub> is electron density 0.0005 GV/cm 1 GV/cm

## Plasma model

(from Lawson, Scientific American-1989)



Pendulum cluster moves to the right



Plasma snapshot: red plasma electrons cluster and make field. Electrons in red ball are trapped.

Non-Accelerator Particle Astrophysics School D. Carrigan

## Creating a plasma wave with a laser



Ponderomotive force on electron

Generated field is:

 $E = m_e c\omega_p / e \qquad (e.g. sin dist of charge)$  $\omega_p = (4\pi n_0 e^2 / m_e)^{\frac{1}{2}} \quad \text{plasma freq}$ 

plasma wave has a phase velocity

$$v_{p} = c(1-\omega_{p}^{2}/\omega^{2})^{1/2}$$

wake generated best when photon packet is half wavelength of plasma

find  $\gamma^{\text{max}} = 2 \omega^2 / \omega_p^2$ 

acceleration length: smallest of dephasing length, pump depletion length, or depth of field (optical channel helps)

2004

Non-Accelerator Particle Astrophysics School	ICTP – Trieste
D. Carrigan	July 26 – Aug. 6, 2

#### Characteristic field strengths



highly relativistic laser driven plasma. Laser pulse length is .03 cm, pulse moves to right, fast oscillations are laser freq. Density  $(n_0)$  is  $10^{16}/\text{cm}^3$ . Moderate case would be more sinusoidal.)

(from Sprangle, et al.) Non-Accelerator Particle Astrophysics School D. Carrigan

A wakefield accelerator - E157 at SLAC



Head of beam generates plasma wakefield,

tail is accelerated by 80 MeV. Also do  $e^+$  - E162.

(E-164 later version ,  $n_e O(3*10^{15})$ , 100 micron bunches - see 2003 Particle Acc. Conf, p. 1530)

M. Hogan Phys. Plasmas 7, 2241 (2000)

Non-Accelerator Particle Astrophysics SchoolICTP – TriesteD. CarriganJuly 26 – Aug. 6, 2004

#### Results from SLAC E-157



Acceleration

Time

Barov and Rosenzweig (UCLA) see similar results at Fermilab. 100 MeV/m using A0 14 MeV photoinjector. 6-8 nC,  $n_e \sim 10^{14}/cc$ .

M. Hogan Phys. Plasmas 7, 2241 (2000). See also Muggli, et al. PRL 93, 014802-1 (2004)

Non-Accelerator Particle Astrophysics School D. Carrigan

## Bob Hofstadter "The Atomic Accelerator" HEPL 560 (1968)

"To anyone who has carried out experiments with a large modern accelerator there always comes a moment when he wishes that a powerful spatial compression of his equipment could take place. If only the very large and massive pieces could fit in a small room!"



Non-Accelerator Particle Astrophysics School D. Carrigan

#### Hofstadter wanted a crystal accelerator!

A table top accelerator ("miniac")

The first solid state accelerator use channeling for focus maybe an after-burner scheme excite atoms coherently with 1 keV-xray

Get out 1 keV/Å

in 1 cm would get 100 GeV

Need an x-ray laser (1968)

#### Problem-transit time

Non-Accelerator Particle Astrophysics School D. Carrigan

Photo S. Carrigan

# Plasma wake field acceleration – solid state

 $G= 0.96(n_0)\frac{1}{2}$  (V/cm) $n_0$  is electron densityRF cavity0.0005 GV/cmgaseous plasma1 GV/cmsolid state plasma100 GV/cm

Non-Accelerator Particle Astrophysics School D. Carrigan

#### Pseudo solid state accelerators

At least four groups see high energy ions, electrons from intense lasers hitting foils

Livermore PRL 85, 2945 (2000)

Michigan APL <u>78</u>, 595 (2001)

Rutherford PRL 90, 064801 (2003) - discussion of mechanisms, target evolution



Livermore





 $3*10^{20}$  W/cm<sup>2</sup>, 1000 TW,  $10^{13}$  proton beams with E to 58 MeV, electrons

protons can be focused by curving target

process: electrostatic fields produced by ponderomotively accelerated hot

electrons act on protons from absorbed hydrocarbons rear side (downstream)

Non-Accelerator Particle Astrophysics School	ICTP – Trieste	55
D. Carrigan	July 26 – Aug. 6, 2004	55

#### **Basic Crystal Accelerator Concept**

excite plasma wake field in solid with density a thousand times gas use channeling to reduce energy loss, focus, and maybe even cool Chen-Noble Tahoe (1996), p. 441

Positives

very high power, femtosec lasers radiative damping (Huang, Ruth, Chen)

Big problems!

blow away material dechanneling

Non-Accelerator Particle Astrophysics School D. Carrigan

## The Fermilab A0 photoinjector

- built as Tesla injector prototype in the late 1990s by Helen Edwards' group
- essentially a gigantic phototube powered by a laser followed by a so-called 3.5 MeV warm RF gun and second stage of a Tesla superconducting nine-cell RF cavity
- beam energy 14.4 MeV.
- very large picosecond electron pulses of 10 nanocoulombs or 10<sup>6</sup> A/cm<sup>2</sup>

#### So what did the Fermilab A0 photoinjector do?

studied channeling nearer extreme conditions needed for

a channeling accelerator Could we make a crystal accelerator or do

#### unique channeling studies?

Ferminic CO-635

Non-Accelerator Particle Astrophysics School D. Carrigan



## Crystal survivability?

#### Process

#### excite electronic plasma

tunnel ionization

partial or total lattice ionization

#### electronic plasma decay

via interband transitions lifetime: (plasma frequency)-O(fs) excitation of phonons in lattice

#### crystal disorder, fracture, or vaporization

lattice dissociation via plasmon absorption lifetime: (ion plasma frequency)<sup>-1</sup> vaporization O(10-100 fs) hydrodynamic heating O(1-10 ps) [Livermore]

Non-Accelerator Particle Astrophysics School D. Carrigan

ICTP – Trieste July 26 – Aug. 6, 2004

$$\omega_{p} = (4 \pi n_{0} e^{2} / m_{e})^{1/2}$$

$$\omega_{pi} = (m_e / m_i)^{1/2} \omega_p$$

58

## Crystal destruction

#### ACCELERATION

G (gradient) proportional to  $(n_0)^{1/2}$ , P (power) prop to  $n_0$ for G = 1 GeV/cm P =  $10^5$  J/cm<sup>3</sup>  $10^{19}$  W/cm<sup>3</sup> for O(10 fs) @ 1 GeV/cm

#### LASER

 $10^{11}$  W/gmBelotshitkii & Kumakhov (1979)or  $10^6$  a/cm² for particle beam $10^{12}$  W/cm³ns long pulses $10^{13}$  W/cm³Chen-Noble (1987)fracture thresholdO(0.1 ns) ref 16Skin depth < 0.1 mm</td>

#### LATTICE IONIZED

 $10^{15}$ - $10^{16}$  W/cm<sup>2</sup> Chen & Noble (1996)/laser

#### **PARTICLE BEAM**

 $10^{11} \text{ A/cm}^2$ 

Chen & Noble (1987) (crystal OK for 10 fs)

Non-Accelerator Particle Astrophysics School

D. Carrigan

#### Situation for Fermilab A0 photoinjector

#### **A0 RF GUN FOR COMPARISON**

 $I/cm^2 = 10 \text{ nc}/1 \text{ ps in } 1 \text{ mm}^2$  or  $10^6 \text{ A/cm}^2$  (OK driver @ 1GeV)

#### **A0 LASER FOR COMPARISON**

$10 \text{ W/cm}^3$	slap ruptured (continuous, 10 <sup>15</sup> W/cm <sup>3</sup> for 10 fs)
$10^9 \text{ W/cm}^2$	damage on lens
$10^{18} \text{ W/cm}^2$	1 Joule on 10 µm spot in 1 ps (OK driver)

Non-Accelerator Particle Astrophysics School D. Carrigan

#### Fermilab A0 schematic layout



R. Carrigan, et al. Phys. Rev. A68, 062901 (2003)Non-Accelerator Particle Astrophysics SchoolICTP – Trieste

D. Carrigan

ICTP – Trieste July 26 – Aug. 6, 2004

61

## A0 at the goniometer



10 nC peak, ε typically 10 mm\*mrad, 10 ps

Non-Accelerator Particle Astrophysics School D. Carrigan

#### Summary of high charge measurements



#### The Future Beyond the Fermilab A0 Experiment

get into 10 fs regime
ne 10<sup>3</sup> to 10<sup>5</sup> larger (small beam size important)
higher energy might be better for channeling, beam size
But new experimental geometry, channeling approaches needed

Possibilities:

SLAC E164 geometry for channeling radiation at 30 GeV

Livermore

Toronto – studying laser melting with sub picosec electron diffraction

Non-Accelerator Particle Astrophysics School D. Carrigan

#### Using SLAC E164 to study channeling



Add crystal, goniometer, x-ray det. (integrating). Now at FFTB (final foc TB) for big q. Channeling radiation ala N. A. Filatova, Phys. Rev. Lett. **48**, 488 @ 12 GeV, (1982), K. Kirsebom, et al., NIMB **119**, 79 (96) @ 150 GeV.

Beam:

charge:  $2*10^{10}$ /bunch (< A0), size 25 µm.

```
time: 100 \text{ mm/c} = 300 \text{ fs}
```

 $I/cm^2$ : 50\*10<sup>6</sup> A/cm<sup>2</sup> (500 times better than A0)

This could take channeling measurements nearly to the plasma regime.

C. Barnes et al., Proc. 2003 Particle Acc. Conf. 1530 (03)

Non-Accelerator Particle Astrophysics School	ICTP – Trieste	۲. ۲.
D. Carrigan	July 26 – Aug. 6, 2004	05

#### The Far Future?



Non-Accelerator Particle Astrophysics School D. Carrigan

ICTP – Trieste July 26 – Aug. 6, 2004

66

## Summary

#### 2001-2009

Tevatron-completing a decade has a proton driver been added? LHC-in operation for pp (15 yr future) 2010-2019 linear collider completed somewhere, running underway upgrade path? Neutrino factories? 2020-2029 VLHC? Muon collider? Can exotic accelerators step up?

## Questions?

Non-Accelerator Particle Astrophysics School D. Carrigan