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#### CONFERENCE ON FUNDAMENTAL SYMMETRIES AND FUNDAMENTAL CONSTANTS

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### FUNDAMENTAL CONSTANTS AT HIGH ENERGY (Time Variation of QCD Scale)

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### at High Energy

(Time Variation of QCD scale)

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# Role of Fundamental Constants

## What Are Fundamental Constants?

# Cosmic Accidents? Determined by Dynamics? Changing in Time? Given by Self-Consistency? Calculable?





Quantum Field Theory

 $\alpha \Longrightarrow$  function of scale (energy)

Quantum-fluctuations of fields  $(e^+e^-\text{pairs})$ 

 $e^+e^-$ : dipoles with size  $m_e^{-1}$ 

 $\rightarrow$  partial screening of bare charge of electron at distances  $> m_e^{-1}$ .

$$\alpha = (137, \ldots)^{-1} \to \alpha_{\text{eff}} \left( m_e^2 \right)$$

 $\alpha_{\rm eff}\left(q^2\right) = \frac{\alpha}{1 - \frac{\alpha}{3\pi} \ln\left(\frac{-q^2}{Am_e^2}\right)}$ 

$$A = \exp(5/3) (-q^2 > 0) \, .$$

Renormalization–Group:

$$\frac{d}{d\ln(q/M)} \bar{e}(q; e_r) = \beta(\bar{e})$$
$$\bar{e}(M; e_r) = e_r$$
$$QED: \beta(e) = \frac{e^3}{12\pi^2} + 0(e^5)$$
High Energy:
$$\mu^+\mu^-, \tau^+\tau^-, \bar{q}q, W^+W^-$$
$$e^3$$

 $\begin{array}{l} \rightarrow \beta(e) = n_{\rm eff} \cdot \frac{e^{-1}}{12\pi^2} + \dots \\ \mathcal{LEP}: \\ \alpha(200 GeV) \cong (127)^{-1} (\sim 10\% \ \text{larger}) \,. \\ \alpha(M_{\star}) \cong (127.8)^{-1} \end{array}$ 



Β



Vorcuum

Polocrizoitism







-> Sommerfeld 19/8' **Originally it was assumed:**  $1/\alpha = 137$ (integer!) **Philosophy and numerology: Eddington:** 137 > 136 = Nr of charged objects  $\alpha^{-1} = \frac{16^2 - 16}{2} + 16 = 136$ Pauli (1958): Nr 137..... Lederman: 137 Eola Road, JL Feynman: 137

Heisenberg ( $\sim 30...$ ):  $\alpha = 2^{-4} 3^{-3} \pi$ = 1/137.6...

Wyler (1971) 1ppm  $\alpha = \frac{9}{8\pi^4} \left(\frac{\pi^5}{2^4 5!}\right)^{1/4}$  The Durk Corner of HEP~

# Fermion Masses: Arbitrary



 $m_e = 0.511 MeV = 0.0000021 \cdot 246 GeV = 2.000 \cdot 10^{-30} lb$ 

me: 0.20 ppb accuracy (squantum optics)



Relations among constants?

e.g. How mixing

 $\theta_{u} \cong \sqrt{\frac{m_{u}}{m_{e}}}$ B. = Vm,

S≈ 90°

 $\theta = \frac{m_s}{m_b} \cdot \frac{m_i}{m_f}$ 

18 Pm. 14 Pm.

Cosmic time dependence of fundomental constants H. Fritzsch Munich -> CERN- Courier March 03 Eur. Journals, Phys. Joday Keck telescope ( Australia, England, U.S) (Webb, Wolfe. "many multiplet method" fine-structure of Fe, Ni, Mg, Sn, A ~150 quagars (-> 11 bu years in time)  $(-0.54 \pm 0.12)$   $\frac{\Delta \pm}{8} = (-0.72 \pm 0.18) \cdot 10^{-5}$ a= 1/137,03699933 (to day) early a= 1/137,037 - - A: 036) d ≈ 1.2.10 -15 por year Linear app .:

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Sample	Method	N <sub>abs</sub>	Redshift	$\Delta \alpha / \alpha$
Fell/MgII	MM	28	0.5 < z < 1.8	$-0.70 \pm 0.23$
NiII/CrII/ZnII	MM	21	1.8 < z < 3.5	$-0.76 \pm 0.28$
SiIV	AD	21	2.0 < z < 3.0	$-0.5 \pm 1.3$
21cm/mm	radio	2	0.25,0.68	$-0.10 \pm 0.17$

TABLE I: Summary of results for 4 independent samples. Values of  $\Delta \alpha / \alpha$  are weighted means in units of  $10^{-5}$ . MM and AD indicate "many-multiplet" and "alkali-doublet". N<sub>abs</sub> is the number of absorption systems in each sample.



Fractional look-back time

FIG. 1:  $\Delta \alpha / \alpha$  vs. fractional look-back time to the Big Bang. The conversion between redshift and look-back time assumes  $H_0 = 68 \text{ km/s/Mpc}$ ,  $(\Omega_M, \Omega_\Lambda) = (0.3, 0.7)$ , so that the age of the universe is 13.9 Gyr. 72 quasar absorption systems contribute to this binned-data plot. The hollow squares correspond to two HI 21cm and molecular absorption systems [16]. Those points assume no change in  $g_p$ , so should be interpreted with caution. The 7 solid circles are binned results for 49 quasar absorption systems. The lower redshift points (below  $z \approx 1.6$ ) are based on (MgII/FeII) and the higher redshift points on (ZnII, CrII, NiII, AlIII, AlII, SiII) [13]. 28 of these 49 systems correspond to the sample used in [4]. The hollow triangle represents the average over 21 quasar SiIV absorption doublets using the alkali doublet method [14]. 13.9 Gyr



### The Oklo Phenomenon

1.8 billion years back

**River Oklo** 

Gabžn, W-Africa

Natural reactor

High concentration of U

Natural enrichment of U235 (3% about 2bn yrs. ago, 0.7% today)

Low concentration of n absorbers

Critical size

Moderator (25% C, 75% Water)

Not commissioned by DOE ~ NRC

but operated for ~ 100 Mio years.

Strong n-absorbers found only in small quantities  $\rightarrow$  reactor activity

Sm<sup>149</sup> Eu<sup>151</sup>, GD<sup>155</sup>, Gd<sup>152</sup>

Shlyakhter

 $\rightarrow$  Damour, Dyson (96)

Neutron capture

 $\mathrm{Sm}^{149} + \mu \rightarrow \mathrm{Sm}^{150} + \gamma$ 

thermal x-section

D.D.: 57 kb  $\leq \hat{\sigma}_{149} \leq 93$  kb

Large x-section: resonance just above threshold



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E = 0.0973 eV

Resonance position:

affects strongly x-section

time shift?

 $\Delta = E_r(Oklo) - E_r \text{ (now)}$ 

 $-0.12 \text{ eV} < \Delta < 0.09 \text{ eV}$ 

 $H = H^{nucl} + H^{elm}$   $\downarrow$   $e^{2} \sum_{a,b} r_{ab}^{-1}$  (Coulomb)

$$\rightarrow \alpha \cdot \frac{dE_r}{d\alpha} < -(1,09 \pm 0.09) \text{ MeV}$$

Bethe-Weizsächer:

-1.14 MeV

$$-0.9 \bullet 10^{-7} < \frac{\alpha(Oklo) - \alpha(now)}{\alpha} < 1.2 \bullet 10^{-7}$$

# <u>Oklo</u>

Natural reactor in Gabon (Afrika) Investigated since ~ 1970 by Franch physicists. (Activa: ~ 26n years ago) Samorium: decoy depends strongly on nuclear resonance Resonance position connot have charged much Dyson - Domour: & < 10 -16 (10 -17) (if no other pms. change) -> Problem with astrophysics Change of A: afforts could curred, if signs of \$ , A different. -> Oklo construint questionalh

Change of A: Mn-Mp affected Mn-Mp = (mj-mu) const. - a · A· const. (~ 2 Mar) Mn-Mp > (-> nucleosynthesis)  $U' \rightarrow SU(2) \times U(0)$  $e^{2} = \frac{g g'}{\sqrt{g^{2} + g^{12}}}$ 9' Which g' in offected ? (Both?) GF affacted (+ nucles synthesis)  $SU(3)' \times SU(2) \times U(1) \subset SO(10)$  (2) -> g affected -> gs affected  $(\Lambda_{c})$  $\rightarrow M_{p}$ Systematic analysis needed!









$$J_{n} \quad y_{nneral:} \qquad (allows t_{i} \quad F_{i}) \\ Leany neckers, leques, blows:' \\ Mir... \\ \frac{1}{\alpha} \quad \frac{\alpha}{\alpha} = \frac{8}{3} \quad \frac{1}{d_{s}} \quad \frac{\alpha}{\alpha_{s}} - \frac{1}{2\pi} \quad (b_{2}^{s} + \frac{3}{3}b_{i}^{s} - \frac{8}{3}b_{s}^{s}) \quad \frac{\Lambda_{cat}}{\Lambda_{cat}} \\ \dot{\Lambda}_{cat} = 0 \\ \frac{1}{\alpha} \quad \frac{\alpha}{\alpha} = \frac{8}{3} \quad \frac{1}{\alpha_{s}} \quad \frac{\alpha}{\alpha_{s}} \\ \frac{\Lambda}{\Lambda} \approx \pm 38, 8 \quad \frac{\alpha}{\alpha} \\ \xrightarrow{\Lambda} \quad Magnetic Moments of Nuclei: \\ \frac{\mu_{p}}{\mu_{p}} = \frac{\mu_{N}}{\mu_{N}} \approx -38, 8 \quad \frac{\alpha}{\alpha} \quad \sim -39 \cdot 10^{-14} / y_{eur} \\ \alpha_{un} \quad inversent, \quad charge of \Lambda_{un}: \\ \frac{\Lambda}{\Lambda} \approx -31 \quad \frac{\alpha}{\alpha} \\ (sign \ charge !)$$



## Vergleich

Compension







Cs-clock

## Unterschied ca. 3 Cäsiumschwingungen 3 Cs - oscilletions / Say pro Tag Experimente am MPQ Garching und IST Boulder, USA

Quantum Optics (4 (MPQ ~ Haensch, Walther) Suppose: à 1 x = -1.2.10-15  $\hat{\Lambda} / \Lambda = 2.4 \cdot 10^{-14}$ (20) Cesium clock: 1s =: 6192631770 cyclus of microwave light ~ hf- transition of Cessum -133 the me at -> Cesium clock \$ H-clock (no dap. of A) (~ 3 cesium cyclas / day) Casium as H 1. Step: Effect ~ 30, if \$ \$ \$ 10 -14 yr-1 Jodium (trapped) ~ AG (trapped) 2. Step: > sensition to A = 10 yr-1 ·(~ 3yrs) e.g.: À /A = (2.13±0.01).10-14 (1.13

Experiment of MPQ (Munich)

Measure absolute optical fraguerry velotive to cesium atomic clock (-> F. ... , Noens 486 mm dye loser in the hydrogen spectrometer Reference: cesium fountain clock (Pharau) LPTE Paris

1999: Hydrogen 15-25 prequency accurate to 1.4.10-14

Since February US: oetave spanning comb synthesisar to measure 15-25 transition relative to PHARAO (brought to Munich ayain; 15-25 - transition: 2466 061413187127(18) Ha (~10-14 arc.!)

24 (50) Ha drift in 43 months -0.9 (2.9) -> 2.8 (S.T). 10 -15 por year (porl.) Change of a: ~ 10-15/yr -> expect: 2.10 -14 for hyperfine transitions (unlikely) Further tests are going on. (-> Boulston) Partial concellation of effect? (voivy dun + Ann!) At level ... x 10 - 15 effects should be seen! Or: Astrophysics wrong? New exp.: no effect ! 10

 $\frac{8}{3} \frac{1}{\alpha_s} \frac{\alpha_s}{\alpha_s} = \frac{1}{\alpha} \frac{\alpha}{\alpha} + \frac{1}{2\pi} (\dots) \frac{\Lambda_{Gut}}{\Lambda_{Gut}}$ 

partial concellation (s.s. th)

 $\frac{\Lambda}{\Lambda} \sim 5 \cdot \dot{\vec{a}}$ 

~ 5.10-15 / yr in Hünsch exp. (?) indications for effect or 1 ~ 4. 10-15/4r

Summery ~ Conclusions 2 18 constants in S.M. Possible relations among them. Some fund constants rather complicated (e.g. nucleon moss) Grand unif: relates a, a, a, a, Time dependence of a -> · · · · · · · · · · Oklo constraint: questionable A ~ + 35 · K > change in cesium clocks Exp.: Prequency change \$ 3.10-15 Expected: ~ 2.10-14 (problem) Concellation ? No charge of a today?