Global Navigation Satellite Systems (GNSS) and the International GPS Service

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IAG-IASPEI Workshop on Deformation Measurements & Understanding Natural Hazards in Developing Countries

ICTP Trieste, Italy January 17 – 23, 2005

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Introduction to GNSS: GPS / GLONASS / GALILEO

- NAVSTAR-GPS: NAVigation System with Time And Ranging Global Positioning System
- GLONASS: GLObal NAvigation Satellite System (Russian pendant)

GALILEO: future European global navigation satellite system

Characteristics:

- Satellite systems for (real-time) **Positioning** and **Navigation**
- Global (everywhere on Earth, up to altitude 5000 km) and at any time
- **Unlimited** number of users
- Weather-independent (radio signals are passing through the atmosphere)
- 3-dimensional position, velocity and time information in real time

GPS/GLONASS Segments

The GPS (and the GLONASS) consists of **3 main segments**:

- **Space Segment**: the satellites and the constellation of satellites
- **Control Segment**: the ground stations, infrastructure and software for operation and monitoring of the GPS/GLONASS
- User Segment: all GPS/GLONASS receivers worldwide and the corresponding processing software

We should add an important **4th segment**:

• Ground Segment:

- **Global network** of the International GPS Service (IGS; \sim 300 stations): reference frame (coordinates and velocities), orbits and clocks, Earth rotation, atmospheric products
- Regional and local permanent networks (Europe, Japan, US, . . . , SAPOS, . . .): densification of the reference frame, positioning services

International GPS Service (IGS): Mission and Objectives

"... committed to providing the highest quality data and products as the standard for global navigation satellite systems (GNSS) in support of Earth science research, multidisciplinary applications, and education... as well as to facilitate other applications benefiting society"

- Provide the highest quality, reliable GNSS data and products, openly and readily available to all user communities.
- Promote universal acceptance of IGS products and conventions as the world standard.
- Continuously innovate by attracting leading-edge expertise and pursuing challenging projects and ideas.
- Seek and implement new growth opportunities while responding to changing user needs.
- Sustain and nurture the IGS culture of collegiality, openness, inclusiveness, and cooperation.
- Maintain a voluntary organization with effective leadership, governance, and management.

International GPS Service (IGS): Structure



Orbit Characteristics

Characteristic	GPS	GLONASS	GALILEO	
Semi-Major Axis (Radius)	26'600 km	25'500 km	29'994 km	
Orbital Period	11 h 58 min	11 h 16 min	14 h 22 min	
Orbit Inclination	55°	65°	56°	
Number of Planes	6 (60° sep.)	3 (120° sep.)	3 (120° sep.)	
Number of Satellites	29 (nom. 24)	14 (nom. 24)	0 (nom. 30)	
PRN Codes	Satellite-dep.	Satellite-indep.	Satellite-dep.	
Frequencies	Satellite-indep.	Satellite-dep.	Satellite-indep.	
Selective Availability	Yes	No	No	
Anti-Spoofing, Encription	Yes	No	Partially	

PRN Codes: Pseudo-Random Noise Codes.

GPS/GLONASS Satellite Constellation



GPS Signals



- Signals driven by an **atomic clock**
- Two carrier signals (sine waves):
 - L_1 : f=1575.43 MHz, λ =19 cm - L_2 : f=1227.60 MHz, λ =24 cm



- **Bits encoded** on carrier by phase modulation:
 - C/A-code (Clear Access / Coarse Acquisition)
 - **P-code** (Protected / Precise)
 - Satellite/Broadcast/Navigation
 Message (Information)

Pseudorange/Code Measurement

• Actual **Pseudorange** observation P_r^s :



- c : speed of light (in vacuum)
- No actual "range" (distance) because of clock errors



(Blewitt, 1996)

- Clock of receiver r reads T_r when signal is received (T_r in receiver clock time).
- Clock of satellite s reads T^s when signal is emitted (T^s in satellite clock time).
- Measurement noise: C/A-code $\sim 10~m$; P-code $\sim 1~m$

Pseudorange/Code Measurement (2)

$$P_r^s = c (T_r - T^s)$$

= $c (t_r + \delta t_r - t^s - \delta t^s)$
= $c (t_r - t^s) + c \,\delta t_r - c \,\delta t^s$
= $\rho_r^s + c \,\delta t_r - c \,\delta t^s$

 t_r , t^s GPS time of reception and emission δt_r , δt^s Receiver and satellite clock error ρ_r^s Range (distance) between receiver and satellite

Simplified model for ρ_r^s : atmospheric delay missing, exactly 4 satellites, etc.

$$P_r^{s_1} = \sqrt{(x^{s_1} - x_r)^2 + (y^{s_1} - y_r)^2 + (z^{s_1} - z_r)^2} + c \,\delta t_r - c \,\delta t^{s_1} \qquad (1)$$

$$P_r^{s_2} = \sqrt{(x^{s_2} - x_r)^2 + (y^{s_2} - y_r)^2 + (z^{s_2} - z_r)^2} + c \,\delta t_r - c \,\delta t^{s_2}$$
(2)

$$P_r^{s_3} = \sqrt{(x^{s_3} - x_r)^2 + (y^{s_3} - y_r)^2 + (z^{s_3} - z_r)^2} + c \,\delta t_r - c \,\delta t^{s_3} \tag{3}$$

$$P_r^{s_4} = \sqrt{(x^{s_4} - x_r)^2 + (y^{s_4} - y_r)^2 + (z^{s_4} - z_r)^2} + c\,\delta t_r - c\,\delta t^{s_4} \tag{4}$$

Basic Positioning and Navigation Concept

Known from the navigation message:

- satellite position $(x^{s_i}, y^{s_i}, z^{s_i})$
- satellite clock error δt^{s_i}

4 unknown parameters:

- receiver position (x_r, y_r, z_r)
- receiver clock error δt_r

More than 4 satellites: best position and clock with **least squares** or **filter** algorithms



Carrier Phase Measurement



The **satellite** generates with its clock the phase signal Φ^s . At emission time T^s (in satellite time) we have:

$$\Phi^s = f \cdot T^s$$

The same phase signal (e.g. a wave crest) propagates from the satellite to the receiver, but the receiver measures only the fractional part of the phase and doesn't know the **integer number of cycles** N_r^s (phase ambiguity):

$$\Phi_r^s = \Phi^s - N_r^s = f \cdot T^s - N_r^s$$

Carrier Phase Measurement (2)

The **receiver** generates with its clock a **reference phase**. At time of reception T_r of the satellite phase Φ_r^s (in receiver time) we have:

$$\Phi_r = f \cdot T_r$$

The actual **phase measurement** is the difference between receiver reference phase Φ_r and satellite phase Φ_r^s :

$$\psi_r^s = \Phi_r - \Phi_r^s = f \cdot T_r - (f \cdot T^s - N_r^s) = f (T_r - T^s) + N_r^s$$

Multiplication with the wavelength $\lambda = c/f$ leads to the **phase observation** equation in meters:

$$L_r^s = \lambda \psi_r^s = c (T_r - T^s) + \lambda N_r^s$$
$$= \rho_r^s + c \,\delta t_r - c \,\delta t^s + \lambda N_r^s$$

Difference to the pseudorange observation: integer ambiguity term N_r^s .

If the receiver looses the GPS signal (loss of lock), the continuous counting of the arriving wave cycles is interrupted: **jump** of an **integer** number of cycles in the phase (**cycle slip**).

Improved Observation Equation and Parameters

$$L_r^s = \rho_r^s + c\,\delta t_r + c\,\delta t_{r,sys} - c\,\delta t^s - c\,\delta t_{sys}^s + \delta\rho_{trp} + \delta\rho_{ion} + \delta\rho_{rel} + \delta\rho_{mul} + \lambda\,N_r^s + \ldots + \epsilon$$

- ρ_r^s Geometrical distance between satellite and receiver
- δt_r Station clock correction: *receiver clocks* (time and frequency transfer)
- $\delta t_{r,sys}$ Delays in receiver and antenna (cables, electronics, . . .)
- δt^s Satellite clock correction: *satellite clocks*
- $\delta t_{s,sys}$ Delays in satellite (cables, electronics, . . .)
- $\delta \rho_{trp}$ Tropospheric delay: *troposphere parameters* (meteorology, climatology)
- $\delta \rho_{ion}$ lonospheric delay: *ionosphere parameters* (atmosphere physics)
- $\delta \rho_{rel}$ Relativistic corrections
- $\delta\rho_{mul}$ ~ Multipath, scattering, bending effects
- N_r^s Phase ambiguity: *ambiguity parameters* (ambiguity resolution)
- ϵ Measurement error

Differences between Code and Phase Observation Equation

Phase:

$$L_r^s = \rho_r^s + c\,\delta t_r + c\,\delta t_{r,sys} - c\,\delta t^s - c\,\delta t_{sys}^s + \delta\rho_{trp} + \delta\rho_{ion} + \delta\rho_{rel} + \delta\rho_{mul} + \lambda\,N_r^s + \ldots + \epsilon$$

Pseudorange/Code:

$$P_r^s = \rho_r^s + c\,\delta t_r + c\,\delta t_{r,sys} - c\,\delta t^s - c\,\delta t_{sys}^s + \delta\rho_{trp} - \delta\rho_{ion} + \delta\rho_{rel} + \delta\rho_{mul} + \ldots + \epsilon$$

- The ionospheric refraction correction $\delta \rho_{ion}$ has the opposite sign for code measurements.
- There is no ambiguity term λN_r^s for code measurements.
- In the case of code measurements the so-called *differential code biases* (DCBs) are part of the system delays $\delta t_{r,sys}$ and δt_{sys}^{s} , i.e., P_1-P_2 or C_1-P_2 biases for satellites and receivers.

Geometrical Distance ρ_r^s

$$\rho_r^s = |\mathbf{r}_i^s(t^s) - \mathbf{r}_{r,i}(t_r)| = |\mathbf{r}_i^s(t^r - \tau_r^s) - \mathbf{R}(t_r) \cdot \mathbf{r}_{r,e}(t_r)|$$

- $r_i^s(t^s)$ Satellite position at emission time $t^s = t_r \tau_r^s$ in inertial system (e.g. J2000.0)
- $\tau_r^s = \rho_r^s/c$ Light travel time

 $\boldsymbol{r}_{r,i}(t_r)$ Receiver position at reception time t_r in **inertial** system

 $r_{r,e}(t_r)$ Receiver position at time t_r in **Earth-fixed** system (e.g. ITRF) with

$$\boldsymbol{r}_{r,i}(t_r) = \boldsymbol{R}(t_r) \cdot \boldsymbol{r}_{r,e}(t_r)$$

 $\begin{array}{ll} \boldsymbol{R}(t_r) & \mbox{Transformation from Earth-fixed to inertial system with } \boldsymbol{R} = \\ \boldsymbol{PNUXY}, \mbox{ where these rotation matrices contain the Earth} \\ \textit{rotation parameters} \mbox{ (IERS, geophysics, ...):} \\ \mbox{Pole coordinates } x_p \mbox{ and } y_p \mbox{ in } \boldsymbol{X} \mbox{ and } \boldsymbol{Y} \mbox{ (polar motion),} \end{array}$

UT1-UTC in the sidereal rotation matrix U, nutation in obliquity $\Delta \epsilon$ and longitude $\Delta \psi$ in N

Single, Double, and Triple Differences



Absolute and Differential/Relative Positioning

- When forming single or double differences we switch from **absolute**, **geocentric positioning** to **relative positioning** of a receiver relative to a **reference receiver** (baseline).
- In single or double differences there is almost **no information** available on the **absolute position** of the receivers (for short baseline vectors $\Delta r_{r_1,r_2}$):

$$\begin{split} \Delta \rho_{r_1,r_2}^j &= \rho_{r_1}^j - \rho_{r_2}^j \quad \approx \quad \rho_{r_1}^j - \{\rho_{r_1}^j + \Delta \boldsymbol{r}_{r_1,r_2} \cdot \boldsymbol{e}_{r_1}^j + O(\Delta \boldsymbol{r}_{r_1,r_2}^2)\} \\ &\approx \quad -\Delta \boldsymbol{r}_{r_1,r_2} \cdot \boldsymbol{e}_{r_1}^j + O(\Delta \boldsymbol{r}_{r_1,r_2}^2) \end{split}$$

- For **short baselines** various **error sources** are considerably **reduced** when using relative or differential positioning:
 - **Satellite clocks** (selective availability (SA)!)
 - Orbit errors
 - Atmosphere: tropospheric and ionospheric refraction errors
 - Antenna phase center offsets and variations (if same antenna type used)
 - Motion of the sites: solid Earth tides, ocean and atmospheric loading

Linear Combinations of GPS Measurements

Very simplified observation equations for the **4 measurement types** in GPS:

$$L_{1} = \rho' + I_{1} + \lambda_{1}N_{1}$$

$$L_{2} = \rho' + \frac{f_{1}^{2}}{f_{2}^{2}}I_{1} + \lambda_{2}N_{2}$$

$$P_{1} = \rho' - I_{1}$$

$$P_{2} = \rho' - \frac{f_{1}^{2}}{f_{2}^{2}}I_{1}$$

with the distance ρ' (incl. clocks, troposphere, relativity) and the ionospheric refraction I_1 for L_1 frequency.

General linear combination (LC) of L_1 and L_2 (or P_1 and P_2):

$$L_x = \kappa_{1,x} L_1 + \kappa_{2,x} L_2$$
$$P_x = \kappa_{1,x} P_1 + \kappa_{2,x} P_2$$

Linear Combinations of GPS Measurements (2)

Substituting the right hand sides and setting $\kappa_{1,x} + \kappa_{2,x} = 1$ (conservation of length) we get:

$$L_x = \rho' + (\kappa_{1,x} + \kappa_{2,x} \frac{f_1^2}{f_2^2}) I_1 + (\kappa_{1,x}\lambda_1N_1 + \kappa_{2,x}\lambda_2N_2)$$

The ionospheric delay changed by the factor $\alpha_I = \kappa_{1,x} + \kappa_{2,x} \frac{f_1^2}{f_2^2}$ compared to L_1 .

The ambiguity term $b_x = \kappa_{1,x}\lambda_1N_1 + \kappa_{2,x}\lambda_2N_2$, in general, is not integer anymore.

Assuming a similar noise $\sigma(L_1) \approx \sigma(L_2)$ for L_1 and L_2 the noise of the LC L_x is given by:

$$\sigma(L_x) = \sqrt{\kappa_{1,x}^2 \,\sigma(L_1)^2 + \kappa_{2,x}^2 \,\sigma(L_2)^2} \approx \sqrt{\kappa_{1,x}^2 + \kappa_{2,x}^2} \,\sigma(L_1)$$

Geometry-Free Linear Combination L_I

$$L_{I} = L_{1} - L_{2}$$

$$= (\rho' + I_{1} + \lambda_{1}N_{1}) - (\rho' + \frac{f_{1}^{2}}{f_{2}^{2}}I_{1} + \lambda_{2}N_{2})$$

$$= (1 - \frac{f_{1}^{2}}{f_{2}^{2}})I_{1} + (\lambda_{1}N_{1} - \lambda_{2}N_{2})$$

$$P_I = P_1 - P_2$$

= $-(1 - \frac{f_1^2}{f_2^2}) I_1$

All geometry information ρ' (GPS orbits, station coordinates, clocks, tropospheric refraction) is eliminated. Only ionospheric refraction remains (and a non-integer ambiguity term for phase).

Applications: information about the **ionosphere** ; estimation of regional or global **ionosphere maps**; **cycle slip fixing**.

Ionosphere-Free Linear Combination L_c

The coefficients of the ionosphere-free LC are:

$$\kappa_{1,c} = \frac{f_1^2}{f_1^2 - f_2^2}, \quad \kappa_{2,c} = -\frac{f_2^2}{f_1^2 - f_2^2}$$

and we get:

$$L_{c} = \kappa_{1,c}L_{1} + \kappa_{2,c}L_{2}$$

= $\kappa_{1,c}(\rho' + I_{1} + \lambda_{1}N_{1}) + \kappa_{2,c}(\rho' + \frac{f_{1}^{2}}{f_{2}^{2}}I_{1} + \lambda_{2}N_{2})$
= $\rho' + (\kappa_{1,c}\lambda_{1}N_{1} + \kappa_{2,c}\lambda_{2}N_{2})$
 $P_{c} = \kappa_{1,c}P_{1} + \kappa_{2,c}P_{2}$
= ρ'

Application: elimination of ionospheric refraction; most important observable for baselines longer than ca. 10 km; disadvantage: noise $\sigma(L_c) \approx 3 \sigma(L_1)$.

Widelane LC L_w and Melbourne-Wübbena LC MW

$$L_w = \frac{f_1}{f_1 - f_2} L_1 - \frac{f_2}{f_1 - f_2} L_2$$

= $\rho' - \frac{f_1}{f_2} I_1 + \frac{c}{f_1 - f_2} (N_1 - N_2)$
 $P_w = \frac{f_1}{f_1 + f_2} P_1 + \frac{f_2}{f_1 + f_2} P_2$
= $\rho' - \frac{f_1}{f_2} I_1$

where

$$\lambda_w = c/f_w = c/(f_1 - f_2) = 86 \, cm$$
; $f_w = f_1 - f_2$

Application: ambiguity resolution of widelane ambiguity $N_w = N_1 - N_2$; cycle slip fixing; advantages: integer ambiguities N_w and long wavelength λ_w .

Melbourne-Wübbena LC (phase and code):

$$MW = L_w - P_w = \lambda_w (N_1 - N_2)$$

Overview of Important LCs of Phase Measurements

LC	λ	$\kappa_{1,x}$	$\kappa_{2,x}$	Geometry		lonosphere		Noise	
	(m)	(m/m)	(m/m)	(m)	(cycles)	(m)	(cycles)	(m)	(cycles)
L_1	0.190	+1.000	0.000	1.000	1.000	+1.000	+1.000	1.000	1.000
L_2	0.244	0.000	+1.000	1.000	0.779	+1.647	+1.283	1.000	0.779
L_c	0.107	+2.546	-1.546	1.000	1.779	0.000	0.000	2.978	5.299
L_I	0.054	+1.000	-1.000	0.000	0.000	-0.647	-2.283	1.414	4.991
L_w	0.862	+4.529	-3.529	1.000	0.221	-1.283	-0.283	5.742	1.268

Ambiguity Resolution

- **Goal**: resolve phase double difference ambiguities N_{AB}^{jk} to the correct integer numbers (ambiguity fixing).
- Ambiguity resolution only possible on the **double difference level** because of the instrumental biases $\delta t_{r,sys}$ and δt_{sys}^{s} .

Gain in accuracy:

- **Short** observation times (few minutes): **Huge** gain in accuracy, typically an improvement from **1 m** (free) to **1 cm** (fixed).
- Long sessions (e.g. 1 day): Gain in accuracy much smaller, typically a factor of 2 in east-west component of site coordinates.

Resolution strategies:

- Short session (short baselines): **search algorithms** searching for best set of integer ambiguities.
- Long session: resolve widelane ambiguities $N_w = N_1 N_2$ first, then resolve N_1 (narrow lane) with the ionosphere-free linear combination.

Geometrical Distance ρ_r^s

$$\rho_r^s = |\mathbf{r}_i^s(t^s) - \mathbf{r}_{r,i}(t_r)| = |\mathbf{r}_i^s(t^r - \tau_r^s) - \mathbf{R}(t_r) \cdot \mathbf{r}_{r,e}(t_r)|$$

 $r_i^s(t^s)$ Satellite position at emission time $t^s = t_r - \tau_r^s$ in inertial system (e.g. J2000.0)

 $\tau_r^s = \rho_r^s/c$ Light travel time

 $\boldsymbol{r}_{r,i}(t_r)$ Receiver position at reception time t_r in **inertial** system

 $\boldsymbol{r}_{r,e}(t_r)$ Receiver position at time t_r in **Earth-fixed** system (e.g. ITRF) with

$$\boldsymbol{r}_{r,i}(t_r) = \boldsymbol{R}(t_r) \cdot \boldsymbol{r}_{r,e}(t_r)$$

 $R(t_r)$ Transformation from **Earth-fixed** to **inertial** system with R = PNUXY, where these rotation matrices contain the *Earth* rotation parameters (IERS, geophysics, ...): Pole coordinates x and u in X and Y (polar motion)

Pole coordinates x_p and y_p in X and Y (polar motion), UT1-UTC in the sidereal rotation matrix U, nutation in obliquity $\Delta \epsilon$ and longitude $\Delta \psi$ in N

Station Position and Modeling of Station Motion

Station position $r_{r,e}(t_r)$ (in Earth-fixed frame) is given by:

 $\boldsymbol{r}_{r,e}(t_r) = \boldsymbol{r}_{r,0} + \boldsymbol{v}_r \left(t_r - t_0 \right) + \delta \boldsymbol{r}_{r,sol} + \delta \boldsymbol{r}_{r,pol} + \delta \boldsymbol{r}_{r,ocn} + \delta \boldsymbol{r}_{r,atm} + \delta \boldsymbol{r}_{r,ant}$

$\delta m{r}_{r,sol}$, $\delta m{r}_{r,pol}$	Solid Earth and pole tides
$\delta m{r}_{r,ocn}$, $\delta m{r}_{r,atm}$	Ocean and atmosphere loading
$\delta m{r}_{r,ant}$	Antenna phase center offset and variation
$oldsymbol{r}_{r,0}$, $oldsymbol{v}_r$	Station coordinates and velocities (positioning,)

- For a high-accuracy GPS processing (e.g. for **troposphere delay estimation**) all these terms have to be modeled on the **sub-centimeter level**.
- The motion of a plate is modeled as a **rotation** of a rigid plate with the angular velocity ω according to:

$$\boldsymbol{r}_{r,e,pla}(t_r) = \boldsymbol{r}_{r,e}(t_0) + \boldsymbol{v}_{r,e} \cdot (t_r - t_0)$$
$$= \boldsymbol{r}_{r,e}(t_0) + (\boldsymbol{\omega} \times \boldsymbol{r}_{r,e}(t_0)) \cdot (t_r - t_0)$$

The direction of ω gives the direction of the pole of rotation.



Station Motion: Global Plate Velocities

GPS velocities (red) are in good agreement with long-term **geophysical velocities** (blue) derived from paleomagnetic data and hot spots.

Station Motion: Plate Tectonics



Station Motion: Postglacial Rebound



Station Motion: Daily Precision



Station Motion: Japanese Permanent Network



More than 1000 GPS permanent stations

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Station Motion: Solid Earth Tides

The non-rigid Earth is deformed by the gravitational forces of Sun and Moon.



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Station Motion: Pole Tides

Pole tides: deformation of the Earth by its rotation



- Pole tides are caused by the reaction of the elastic Earth to the change of the rotation axis
- Polar motion: variations of the Earth rotation axis with Chandler (≈ 432 days) and annual periods, "circle" with a diameter of about 15 m.
- Effect on site coordinates up to about **3 cm**.

Station Motion: Ocean Loading

- The station motion due to ocean loading is caused by the **weight of the water** on the continental plate changing with the ocean tides.
- The main effect of up to **several centimeters** for coastal sites is in **station height**.
- Even for Zimmerwald (far away from the sea) the effect in height is still **1 cm**.


Station Motion: Atmospheric Loading and Other Effects

Atmospheric loading:

- Station heights also vary (up to 2 cm) with the air pressure, i.e., the weight of the atmosphere on the continental plate.
- The size of the deformation depends on the local pressure anomaly (ca. -0.35 mm/mbar) and the mean pressure anomaly in a region of 2000 km (ca. -0.55 mm/mbar).

Regional and local effects:

- Difficult **deformations** in the neighborhood of plate boundaries; non-linear motions, **earthquakes** (coseismic and postseismic)
- **Post-glacial rebound**, i.e., the slow uplift of the Earth's crust in Scandinavia, Greenland, or Canada since the melting of the ice caps
- Instabilities of the markers due to local changes, e.g., ground water level, mining, monument stability, etc. .

Station Motion: Silent Earthquakes



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Station Motion: Episodic Tremor and Slip (Silent Earthquakes; courtesy of H. Dragert)



Station Motion: Artifacts



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Geometrical Distance ρ_r^s

$$\rho_r^s = |\mathbf{r}_i^s(t^s) - \mathbf{r}_{r,i}(t_r)| = |\mathbf{r}_i^s(t^r - \tau_r^s) - \mathbf{R}(t_r) \cdot \mathbf{r}_{r,e}(t_r)|$$

 $r_i^s(t^s)$ Satellite position at emission time $t^s = t_r - \tau_r^s$ in inertial system (e.g. J2000.0)

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$$\boldsymbol{r}_{r,i}(t_r) = \boldsymbol{R}(t_r) \cdot \boldsymbol{r}_{r,e}(t_r)$$

 $R(t_r)$ Transformation from **Earth-fixed** to **inertial** system with R = PNUXY, where these rotation matrices contain the *Earth* rotation parameters (IERS, geophysics, ...): Pole coordinates x_p and y_p in X and Y (polar motion),

UT1-UTC in the sidereal rotation matrix $oldsymbol{U}$, nutation in obliquity $\Delta\epsilon$ and longitude $\Delta\psi$ in $oldsymbol{N}$

GPS Satellite Position and Orbit Accuracies

Satellite position $r_i^s(t^s)$ (in inertial frame) is given by:

$$\boldsymbol{r}_{i}^{s}(t^{s}) = \boldsymbol{r}_{i,0}^{s}(t^{s}; a, e, i, \Omega, \omega, t_{p}; p_{1}, p_{2}, \dots, p_{d}) + \delta \boldsymbol{r}_{ant}^{s}(t^{s})$$

 $egin{aligned} m{r}_{i,0}^s, \, \delta m{r}_{ant}^s \ a,e,i,\Omega,\omega,t_p \ p_1,p_2,\dots,p_d \end{aligned}$

Center of mass position and antenna phase center offset *orbital elements* (orbit determination) *gravity field coefficients, air drag and radiation pressure parameters, ...* (dynamical parameters)

High-precision satellite positions are usually computed by numerical integration and are available from the **International GPS Service**.

IGS GPS/GLONASS Orbit Products Today

IGS/IGEX Product	Availability	Updates	Accuracy
GPS Final Orbits	10 days	weekly	<mark>3 cm</mark>
GPS Rapid Orbits	17 hours	daily	5 cm
GPS Ultrarapid Orbits	3 hours	twice a day	10 cm
GPS Predicted Orbits	real-time	twice a day	30 cm
GLONASS Final Orbits	6 weeks	weekly	10 cm
GPS Broadcast Orbits	real-time	hourly	300 cm
GLONASS Broadcast Orbits	real-time	30 min	300 cm

Quality of IGS Final Orbits



Influence of Orbit Errors on Station Coordinates

- Absolute positioning: orbit errors directly map into station position.
- Relative positioning: Bauersima (1983) computes the error Δr in a component of a baseline of length l as a function of an orbit error of size ΔR :

$$\Delta r = \frac{l}{\rho} \cdot \Delta R \approx \frac{l(km)}{25'000(km)} \cdot \Delta R$$

 ρ is the mean distance between the survey area and the satellite system.

Orbit Error	Baseline Length	Baseline Error	Baseline Error
2.5 m	10 km	.1 ppm	1 mm
2.5 m	100 km	.1 ppm	10 mm
2.5 m	1000 km	.1 ppm	100 mm
.25 m	100 km	.01 ppm	1 mm
.25 m	1000 km	.01 ppm	10 mm
.05 m	100 km	.002 ppm	- mm
.05 m	1000 km	.002 ppm	.5 mm

Influence of Orbit Errors on Station Coordinates

Baseline of about 1000 km between Onsala (Sweden) and Graz (Austria)



Improved Observation Equation and Parameters

$$L_r^s = \rho_r^s + c \, \delta t_r + c \, \delta t_{r,sys} - c \, \delta t^s - c \, \delta t_{sys}^s + \delta \rho_{trp} + \delta \rho_{ion} + \delta \rho_{rel} + \delta \rho_{mul} + \lambda \, N_r^s + \ldots + \epsilon$$

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- ϵ Measurement error

IGS GPS Satellite Clock Products Today

Satellite (and receiver) clock corrections can be estimated using **undifferenced code and phase measurements** of a global network

By convention the IGS clock estimates refer to the **ionosphere-free LC** and thus contain all system delays of this LC ($\delta t_{r,sys}$ and δt_{sys}^s).

The IGS clock products have the following characteristics:

IGS Product	Availability	Updates	Accuracy	
GPS Final Clocks GPS Rapid Clocks GPS Ultrarapid Clocks GPS Predicted Clocks	10 days 17 hours 3 hours real-time	weekly daily twice a day twice a day	< 3 cm 3 cm 6 cm 150 cm	(< 0.1 ns) (0.1 ns) (0.2 ns) (5 ns)
GPS Broadcast Clocks	real-time	hourly	200 cm	(7 ns)

Quality of IGS Final Clocks



Improved Observation Equation and Parameters

$$L_r^s = \rho_r^s + c\,\delta t_r + c\,\delta t_{r,sys} - c\,\delta t^s - c\,\delta t_{sys}^s + \delta\rho_{trp} + \delta\rho_{ion} + \delta\rho_{rel} + \delta\rho_{mul} + \lambda\,N_r^s + \ldots + \epsilon$$

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Differential Code Biases: P_1 - P_2

- Because the **system delays** in the satellite and receiver, $(\delta t_{sys}^s \text{ and } \delta t_{r,sys})$ are different for P₁ and P₂ code measurements, P₁-P₂ corrections have to be applied when processing undifferenced code measurements, so-called **differential code biases (DCBs)**. Differential: they cannot be determined absolutely, but only relatively between different satellites or stations.
- By convention the **ionosphere-free LC** is said to have **no DCBs**. All system delays are **included in the satellite and receiver clocks**, respectively.
- P₁-P₂ corrections are especially important when estimating ionosphere models or computing L₁ (or L₂) code solutions (e.g. with precise satellite orbits and clocks and an ionosphere model).

Differential Code Biases: P_1 - P_2



CODE'S 30-DAY P1-P2 DCB SOLUTION UP TO DAY 190, 2002

Differential Code Biases: P_1 - C_1

Today there are **three classes of receivers** that provide different types of code (pseudorange) measurements:

- Older C_1/P'_2 or CC receivers (Turborogue, AOA, older Trimble 4xxx) provide:
 - $\begin{array}{lcl} \mathsf{C}_1 &=& \mathsf{C}/\mathsf{A} \text{ code at L1 frequency} \\ \mathsf{L}_1(\mathsf{C}_1) &=& \mathsf{C}/\mathsf{A}\text{-based phase at L1 frequency} \\ \mathsf{P}_2' &=& \mathsf{codeless pseudorange at L2 frequency with } \mathsf{P}_2' = \mathsf{C}_1 + (\mathsf{P}_1\text{-}\mathsf{P}_2) \\ \mathsf{L}_2' &=& \mathsf{cross-correlated phase at L2 frequency with} \\ &\quad \mathsf{L}_2' = \mathsf{L}_1(\mathsf{C}_1) + (\mathsf{L}_2(\mathsf{P}_2)\text{-}\mathsf{L}_1(\mathsf{P}_1)) \end{array}$
- Modern P_1/P_2 receivers (Ashtech, older Leica) provide:

Differential Code Biases: P_1 - C_1

 Very recent C₁/P₂ receivers (Leica SR9600, Leica CRS1000, Novatel OEM4, Trimble MS750, Trimble 4700, Trimble 5700) provide:

To obtain **consistent code observations** for all these receiver types, P_1 - C_1 (and P_1 - P_2) biases have to be applied depending on the LC processed. This consistency is important for:

- Resolution of **widelane ambiguities** on baselines with receivers from different classes.
- Satellite clock corrections to be used for various linear combinations
- Estimation of global, regional or local **ionosphere models** from code measurements of different receiver classes

Differential Code Biases: P_1 - C_1



CODE'S 30-DAY P1-C1 DCB SOLUTION UP TO DAY 189, 2002

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Use of Differential Code Biases

LC	$\mathbf{P}_1/\mathbf{P}_2$	C_1/C_2	C_1/P_2	
L_1	$+1.55 \cdot B_{P_1-P_2}$	$+1.55 \cdot B_{P_1-P_2} + B_{P_1-C_1}$	$+1.55 \cdot B_{P_1-P_2} + B_{P_1-C_1}$	
L_2	$+2.55 \cdot B_{P_1-P_2}$	$+2.55 \cdot B_{P_1-P_2} + B_{P_1-C_1}$	$+2.55 \cdot B_{P_1-P_2}$	
L_c	0	$+B_{P_1-C_1}$	$+2.55 \cdot B_{P_1-C_1}$	
L_I	$-B_{P_1-P_2}$	$-B_{P_1-P_2}$	$-B_{P_1-P_2} + B_{P_1-C_1}$	
L_w	$-1.98 \cdot B_{P_1 - P_2}$	$-1.98 \cdot B_{P_1 - P_2} + B_{P_1 - C_1}$	$-1.98 \cdot B_{P_1 - P_2} + 4.53 \cdot B_{P_1 - C_1}$	
MW	0	$-B_{P_1-C_1}$	$-0.56 \cdot B_{P_1-C_1}$	

 L_c , L_I , L_w , and MW denote the ionosphere-free, geometry-free, widelane, and Melbourne-Wübbena linear combination, respectively.

Convention: DCB = 0 for L_c (ionosphere-free LC) and DCB = $-B_{P_1-P_2}$ and DCB = $-B_{P_1-C_1}$ for L_I (geometry-free LC).

The factors involved are:

$$\begin{aligned} \frac{f_1^2}{f_1^2 - f_2^2} &= 2.546 \quad , \quad \frac{f_2^2}{f_1^2 - f_2^2} = 1.546 \quad , \\ \frac{f_1 f_2}{f_1^2 - f_2^2} &= 1.984 \quad , \quad \frac{f_1}{f_1 - f_2} = 4.529 \quad , \quad \frac{f_1}{f_1 + f_2} = 0.562 \end{aligned}$$

Improved Observation Equation and Parameters

$$L_r^s = \rho_r^s + c\,\delta t_r + c\,\delta t_{r,sys} - c\,\delta t^s - c\,\delta t_{sys}^s + \delta\rho_{trp} + \delta\rho_{ion} + \delta\rho_{rel} + \delta\rho_{mul} + \lambda\,N_r^s + \ldots + \epsilon$$

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Relativistic Corrections (1)

The satellite clocks (as well as the station clocks) are affected by **special** (relative velocity of the satellite) and **general relativistic effects** (gravity field of the Earth).

Special relativity: moving clocks are **slower** than clocks at rest.

General relativity: due to the weaker gravity field at the altitude of the GPS satellites the satellite clocks are faster by about 40 μ s/d than clocks on the Earth's surface.

Resulting frequency difference $\Delta f = f - f_0$ between satellite and receiver:

$$\frac{\Delta f}{f_0} = \frac{f - f_0}{f_0} = \frac{1}{2}\frac{v^2}{c^2} + \frac{\Delta U}{c^2}$$

 $\begin{array}{ll} f_0 & & \mbox{Frequency received by the receiver} \\ f & & \mbox{Frequency emitted by the satellite} \\ v & & \mbox{Velocity of the satellite (about 4 km/s)} \\ \Delta U & & \mbox{Difference in gravity potential between satellite and receiver} \end{array}$

Relativistic Corrections (2)

Assuming a circular orbit and a spherical Earth (with mass M_E) we get approximately:

$$\frac{\Delta f}{f_0} = \frac{1}{2} \frac{v^2}{c^2} + \frac{GM_E}{c^2} \left(\frac{1}{|\boldsymbol{r}^s|} - \frac{1}{|\boldsymbol{r}_r|}\right) \approx -4.464 \cdot 10^{-10}$$

The frequency of the satellite clock is **shifted** at the ground by $\Delta f = 4.464 \cdot 10^{-10} \cdot f_0 = 4.57 \cdot 10^{-3}$ Hz to a value of **10.22999999543** MHz, so that the receiver will receive the nominal frequency of 10.23 MHz despite the relativistic effects mentioned above.

This frequency shift only corrects for a **constant** clock rate. Because the GPS orbits are not exactly circular, the satellite clock also shows **periodic variations**:

$$\delta \rho_{rel,1} = \frac{2}{c} \sqrt{GM_E a} \ e \ \sin E = \frac{2 \cdot \boldsymbol{r^s} \cdot \boldsymbol{\dot{r}^s}}{c^2}$$

with e the numerical eccentricity, a the semi-major axis and E the eccentric anomaly of the satellite orbit. This distance correction $\delta \rho_{rel,1}$ may amount to more than **10 m**.

Relativistic Corrections (3)

We also have to consider a **small delay** of the GPS signal when it **propagates through the Earth's gravity field** — a signal is slowed down when passing through the gravity field.

The corresponding correction is given by (integration of the gravity field effect from the satellite to the receiver):

$$\delta \rho_{rel,2} = \frac{2 G M_E}{c^2} \ln \left(\frac{r^s + r_r + \rho_r^s}{r^s + r_r - \rho_r^s} \right)$$

This second term has a maximum size of **18.7 mm** (at low elevations).

The total correction term $\delta \rho_{rel}$ in the observation equation is therefore

$$\delta \rho_{rel} = \delta \rho_{rel,1} + \delta \rho_{rel,2}$$

to be applied independent of whether broadcast or high-precision IGS satellite clock corrections are used.

Atmospheric Refraction



Two regions are important for the propagation of the **microwave** signals:

- Troposphere (neutral atmosphere): extending up to a height of 10 km. The signal delay of microwaves is non-dispersive, i.e. frequency-independent.
- **Ionosphere**: **dispersive**, i.e. the two frequencies (L_1 and L_2) are delayed differently. The ionosphere ranges from about **50–1000 km altitude**.

Improved Observation Equation and Parameters

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Tropospheric Refraction (signal delay)

• Delay caused by the **refractive index** n (refractivity N) along the path S:

$$\delta \rho_{trp} = \int_{S} (n-1) \, ds = 10^{-6} \int_{S} N \, ds$$

- Total delay in zenith direction: $\delta \rho_{trp}(z=0) \approx 2.3 \text{ m}$ (8 ns)
 - **Dry part:** 90% of the delay; mainly depending on pressure; easy to model.
 - Wet part: up to 40 cm; difficult to model (unknown water vapor distribution).



Modeling and Estimation of Troposphere

- Models compute the tropospheric delay $\delta \rho_{trp}$ from pressure P, temperature T, and relative humidity H.
- Meteorological data (P,T,H): measurements at the station or standard atmosphere.
- **Problems:** Standard atmosphere not representative; met data not good enough (error in temperature of **1** C^o may cause height error of **8 cm**!).
- Therefore: estimation of zenith troposphere delays from the GPS/GLONASS data.
- Long sessions needed to separate troposphere delay and height.
- **Tilting of the zenith** can be used to represent troposphere gradient parameters.



Total Delay from GPS at Tsukuba (Japan)



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Ionospheric Refraction

• The ionospheric refraction $\delta \rho_{ion}$ is caused by the interaction of the microwave signal with the **free electrons**:

$$\delta\rho_{ion} = \int_0^\infty (n_e - 1) \, ds = \pm \frac{a}{f^2} \int_0^\infty N_e \, ds = \pm \frac{a}{f^2} E$$

- *f* Frequency
- n_e Refractive index $n_e = 1 \pm \frac{a}{f^2} N_e$
- N_e Electron density (electrons/m³)
- E Number of free electrons in a column of 1 m² cross-section along the signal path (**TEC: total electron content** in TECU: 10¹⁶ free electrons/m².
- a Constant: $a = 40.3 \cdot 10^{16} \, m \, s^{-2} \, TECU^{-1}$
- **Positive sign** (delay) for pseudorange/code, **negative sign** (advance) for phase measurements (group and phase velocity).
- The number of free electrons (generated by ionization) depends on the **solar activity** (geographic latitude, time of day, 11-year solar cycle).

Impact of Ionospheric Refraction

• The **total delay** in zenith direction varies between **1–20 m** (3–70 ns):

$$\delta \rho_{ion} \ [m] \approx \frac{1}{6} E \ [TECU] \quad for \ L_1$$

Low elevation: delays of up to **100 m** possible.

- Absolute positioning: primary effect in **height** (several meters).
- Main effect on relative positioning (baselines): **scaling** of the network (neglect of ionosphere: distances are **too short**).
- Estimate of the scale factor (in ppm) for a baseline of length *l*:

$$\frac{dl}{l} \approx -0.10 \cdot E \quad [ppm] \quad (E \ in \ TECU)$$

• **Example**: Network of 4 by 4 km in Switzerland in 1989: $E \approx 50$ TECU, $dl/l \approx$ -5 ppm, and $dl \approx$ -2 cm for the baseline error. Network of 20 km: error of already -10 cm.

Modeling and Estimation of the lonosphere

• Elimination by forming the **ionospherefree** linear combination of L_1 and L_2 :

$$L_c = \frac{f_1^2}{f_1^2 - f_2^2} L_1 - \frac{f_2^2}{f_1^2 - f_2^2} L_2$$

But:

- Noise $\sigma(L_c) \approx 3 \, \sigma(L_1)$.
- Non-integer ambiguities.



- Short baselines (< 10 km): use L_1 and L_2 ; longer baselines: use L_c .
- Simple ionosphere model available in the **broadcast message**.
- Daily regional and global **ionosphere models** (maps) computed by the **IGS**.
- Assumption: all free electrons located in an **infinitesimal layer** at a height of about **450 km** height (**single layer**).
- Benefit: elimination of scale factor, improved ambiguity resolution.

Development of the lonosphere derived from GPS



Diversity of GPS Receiver Antennas



Ashtech 700936 mit Radom

Dorne Margolin T (JPL)



Leica SR399

Trimble 22020 (Compact L1/L2)

Antenna Phase Center Offsets and Variations

- The antenna phase center is not a fixed point; it varies with the direction of the incoming signal (phase center variations).
- Combining different antenna types: errors of up to 10 cm in height.



 Antenna offsets and variations available at the IGS: ftp://igscb.jpl.nasa.gov/igscb/station/general/igs_01.pcv
Elevation-Dependent Phase Center Variations (PCV)

Example: Two different Leica antenna types



Elevation- and Azimuth-Dependent PCV (Geodetic Antenna)





Elevation- and Azimuth-Dependent PCV (Rover Antenna)

GPS Satellite Antenna Offsets



(Gerry Mader, NGS)

Elevation-Dependent PCV: GPS Satellites



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GPS Block II/IIA Antenna Design





Multipath Effects

- The antenna receives (apart form the direct signal) **reflections** of the signal from **objects in the vicinity**.
- **Superposition** of direct and indirect signal.
- Systematic deviations in code up to **50 m**, in phase up to **5 cm**.
- Variations with periods of **5**-**30 min**.



- Most critical at **low elevation**, with **short** observations times.
- Package of measures: Antenna design (choke ring, ground plane, ...); selection of site (free horizon, ...); longer observation sessions (averaging of effects).