



The Physics of Spallation Processes Theory and Experiments

Frank Goldenbaum

- What is spallation?
- Why are spallation reactions of interest? (applications, fund. physics)
- Theoretical models describing spallation reactions (INC and evap. model)
 - Limits and constraints
 - validation
- **Experimental investigations**
(here: NESSI, PISA & JESSICA at COSY)
- **Comparison between models and experiments**
- **Conclusion**

Physical observables and deduced quantities of interest

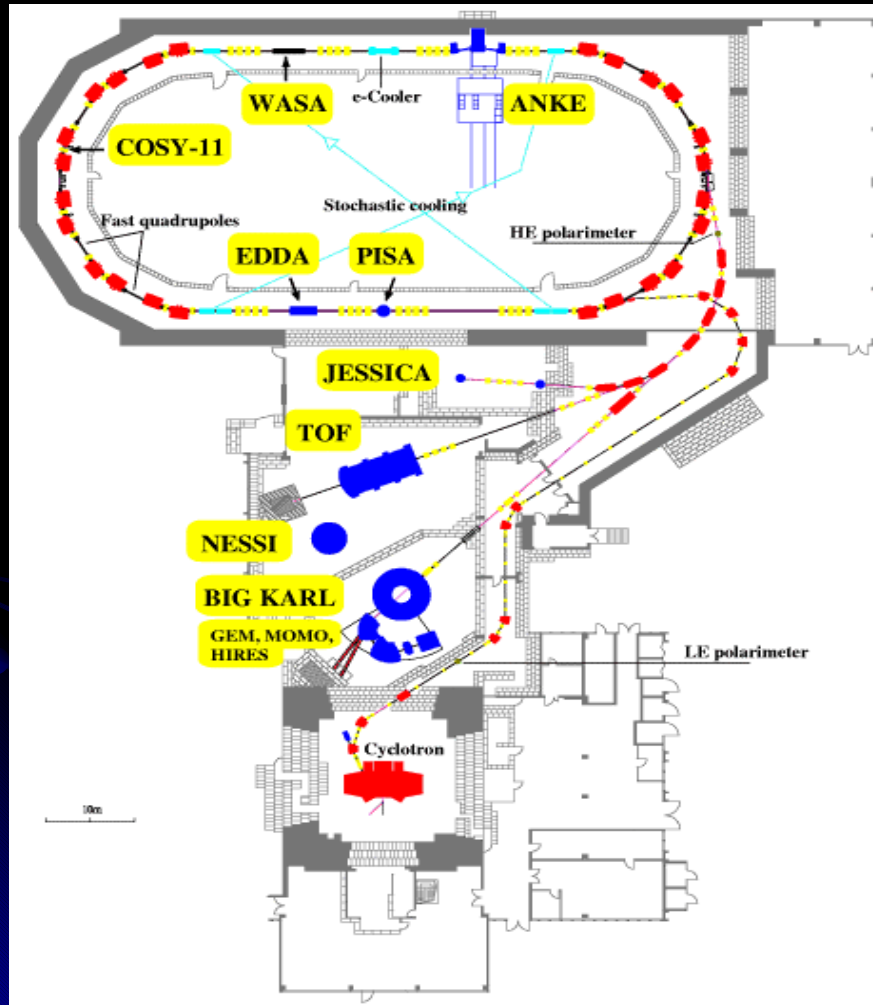


most important parameters

- survival probabilities/reaction rates
- average ejectile multiplicities per incident proton and per nuclear reaction
- ejectile ($n, p, d, t, \text{He}, \text{Li}, \dots, \pi, \gamma, \dots$) multiplicity distribution concerning transport in complex geometries note
 - difference thin/thick target scenarios
 - not only one nucl.reac., further multiplication in multiplying medium
 - particles originate from both: **inter/intra** nuclear cascade
- angular and energy distributions of neutrons and charged particles ($\pi, n, p, d, t, {}^3\text{He}, {}^4\text{He}, \text{Li}, \text{Be}, \dots, \text{IMF}, \text{remnants}$)
- excitation energy distributions
- energy deposition, dpa
- ...



COSY (COoler SYnchrotron) 150 MeV- 2.5 GeV



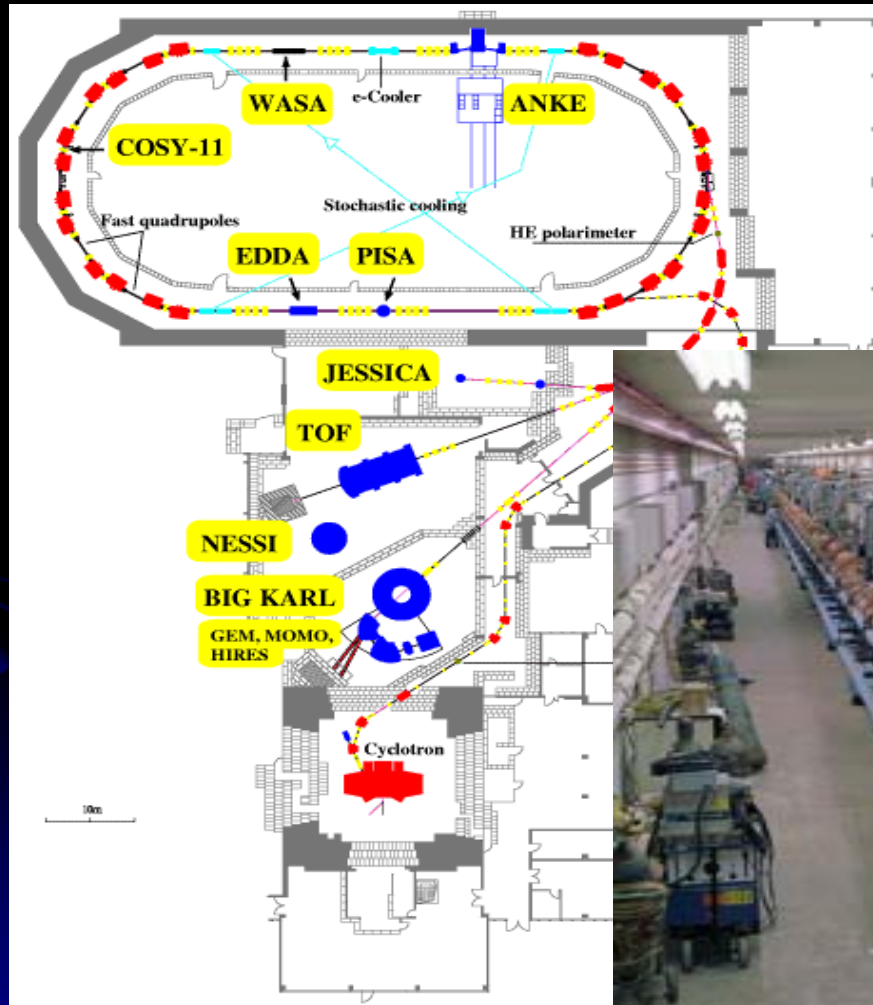
JESSICA (Jülich Experimental Spallation Target Setup in COSY Area)
external beam

NESSI (Neutron Scintillator and Silicon detector)
external beam

PISA (Proton Induced Spallation)
-internal beam
-luminosity: $6.6 \cdot 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$
-low absorption and small energy loss of ejectiles due to target thicknesses (down to $50 \mu\text{g}/\text{cm}^2$)



COSY (COoler SYnchrotron) 150 MeV- 2.5 GeV

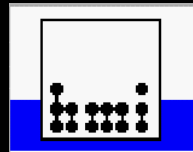


JESSICA (Jülich Experimental Spallation Target Setup in COSY Area)
external beam

NESSI (Neutron Scintillator and Silicon detector)
external beam



Induced Spallation)
m
 $6 \cdot 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$
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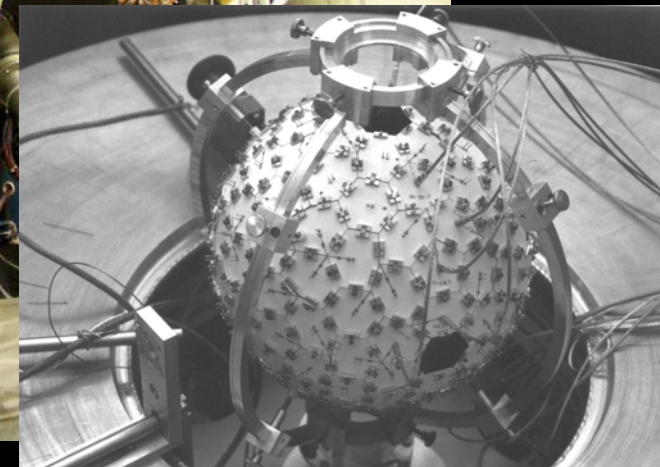
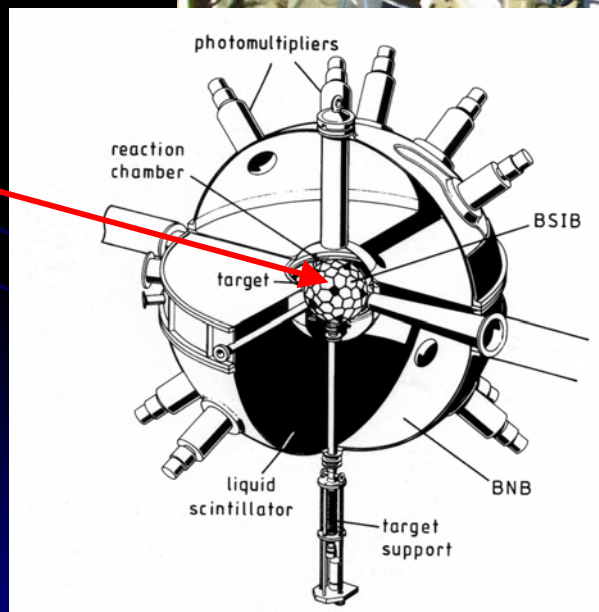
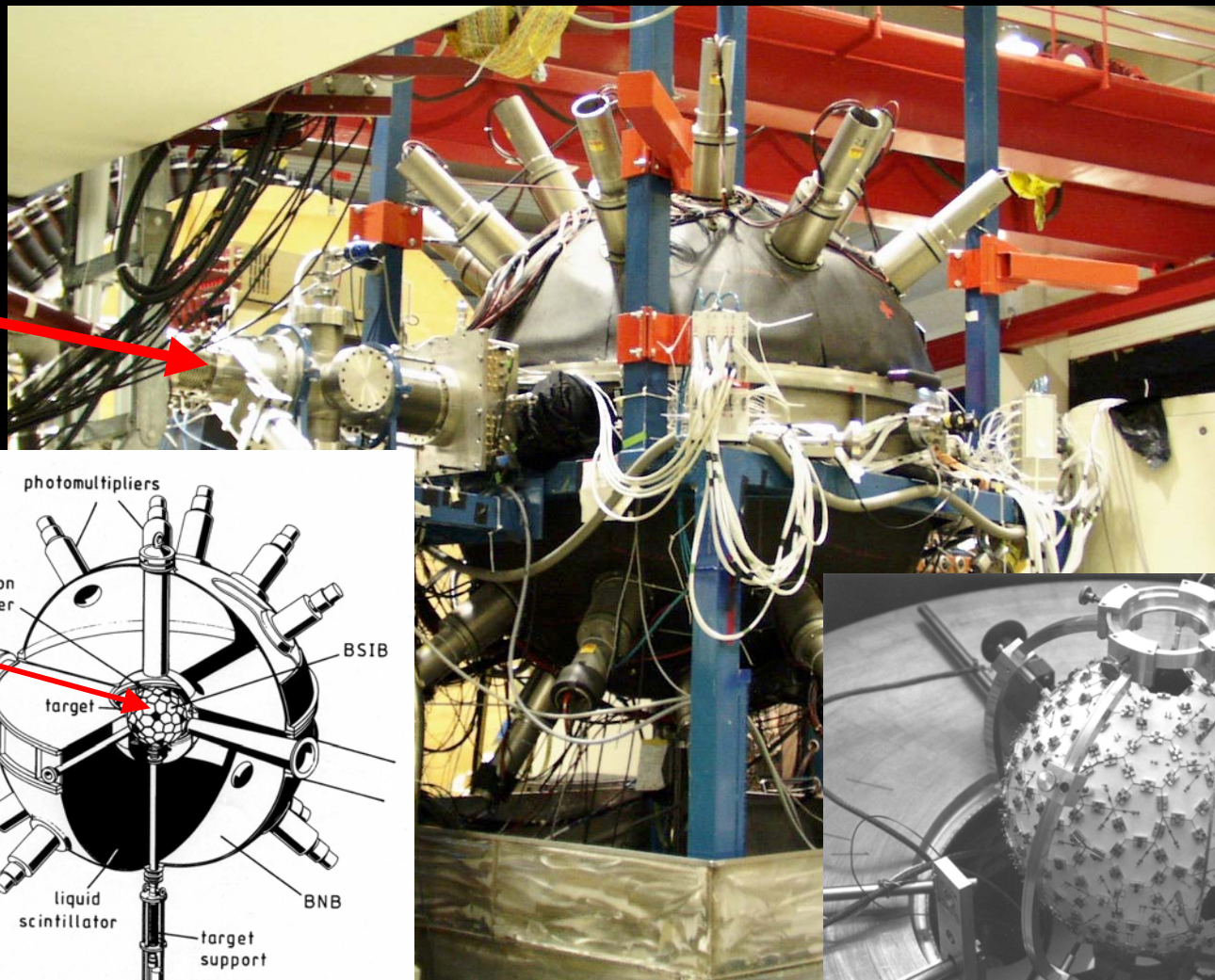
NESSI @ COSY

Forschungszentrum Jülich
in der Helmholtz-Association

GANIL
GRAND ACCELERATEUR NATIONAL D'IONS LOURDS



GeV p



05. April 2005

Frank Goldenbaum
The Physics of Spallation Processes --- Experiment and Theory



Light charged particle and cluster production in 1.2 GeV p-induced reactions on Al-Th

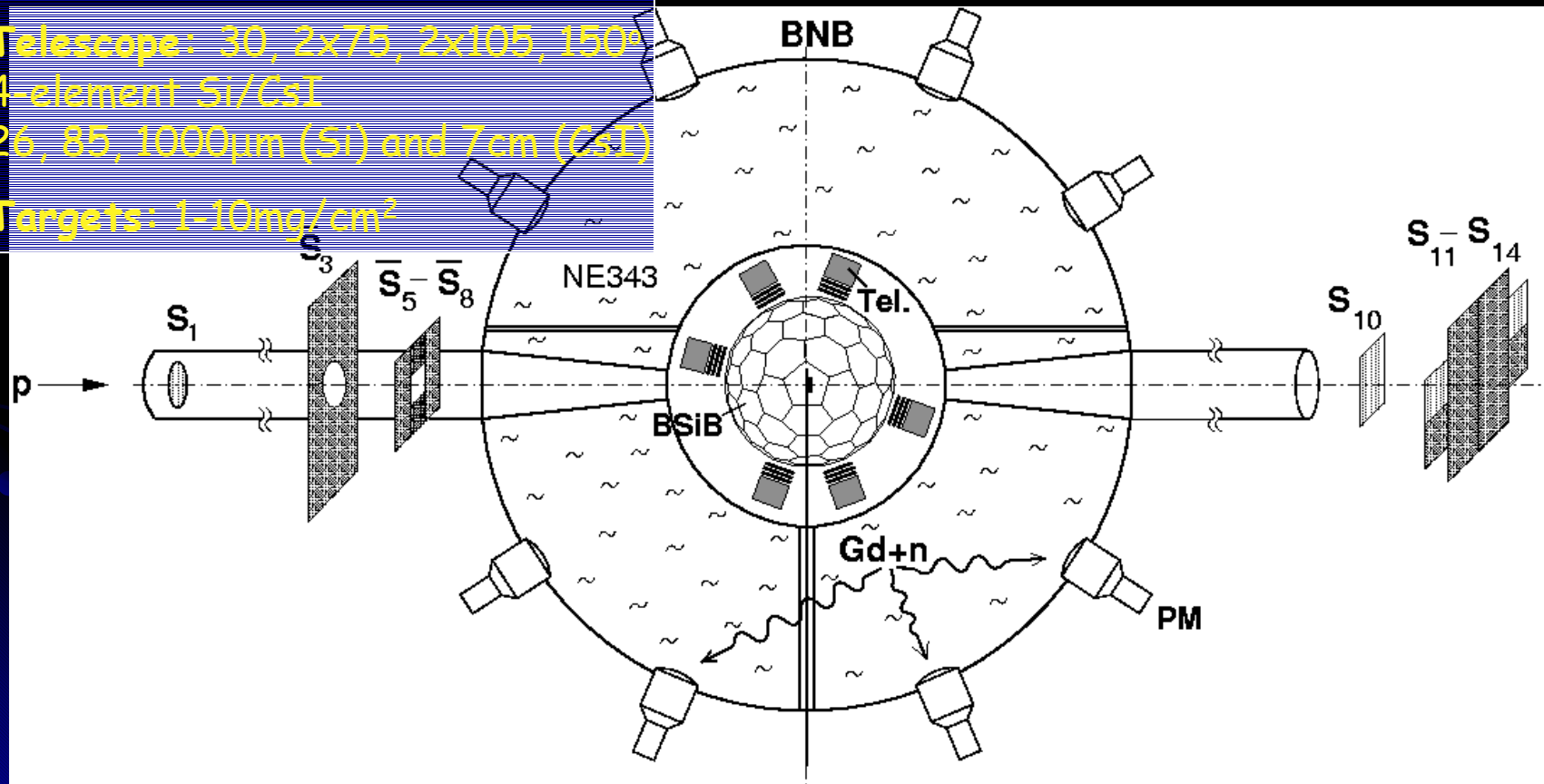
BSiB: 20cm, 151, 500 μ m Si, 67msr

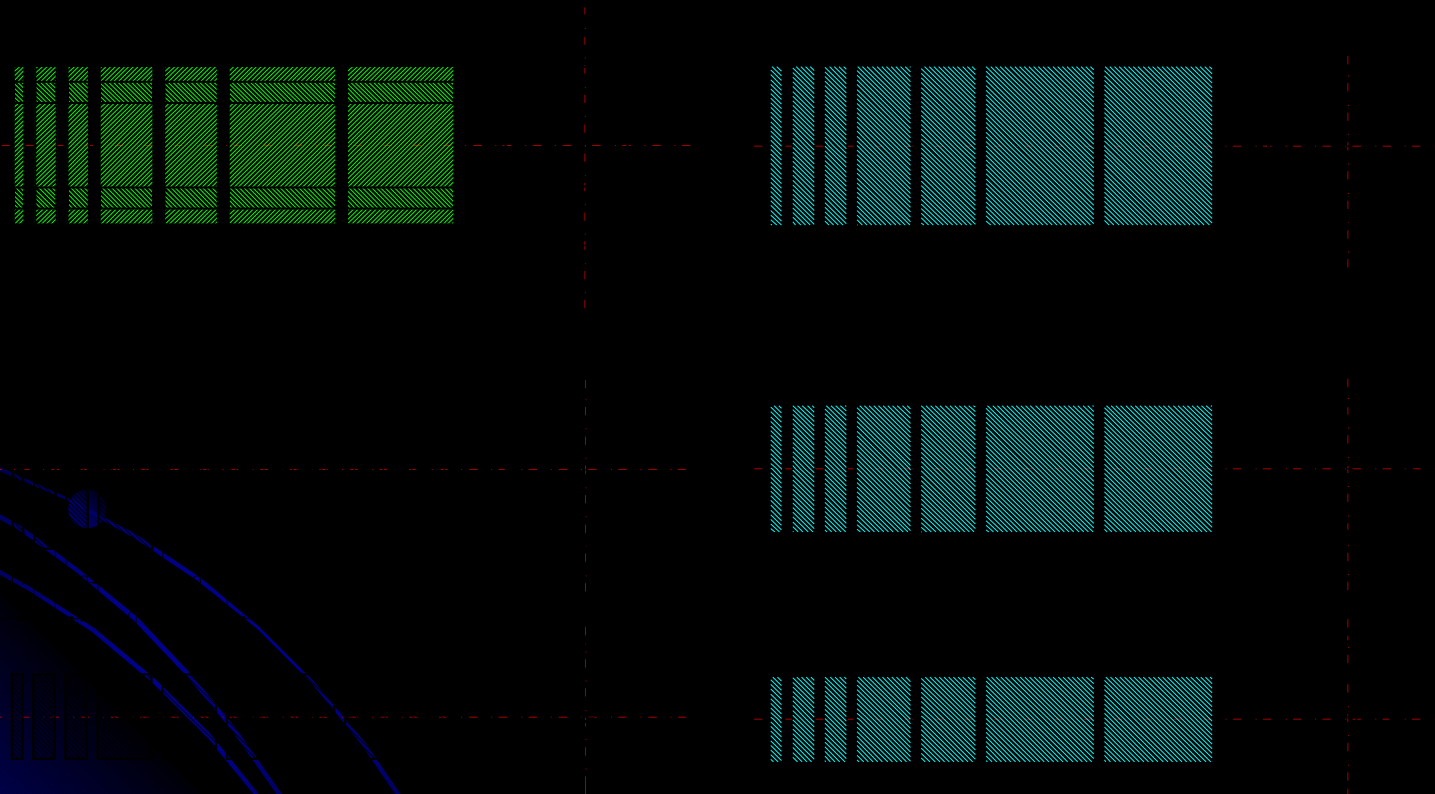
Telescope: 30, 2x75, 2x105, 150 $^\circ$

4-element Si/CsI

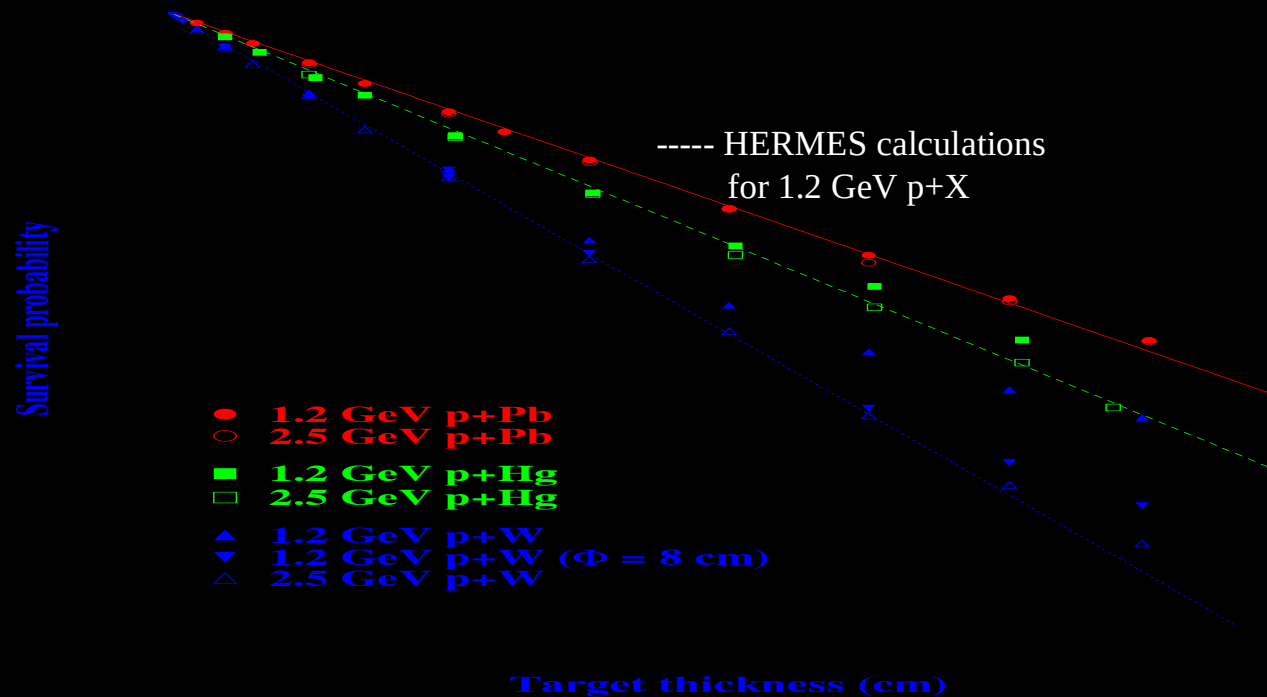
26, 85, 1000 μ m (Si) and 7cm (CsI)

Targets: 1-10mg/cm 2

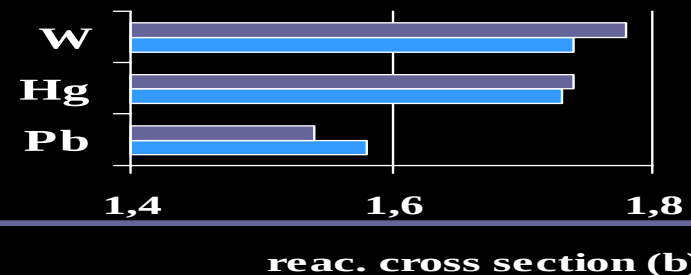
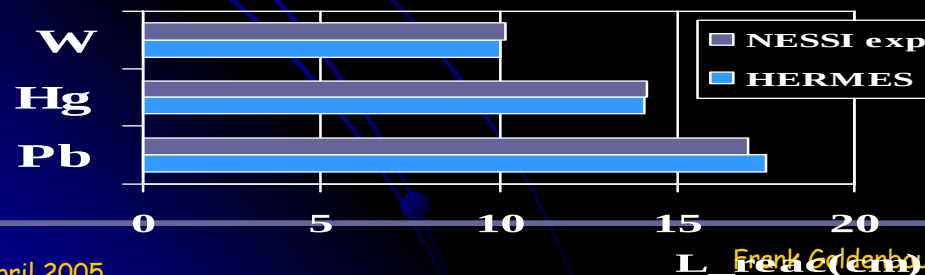




Comparison measured and calculated survival probability



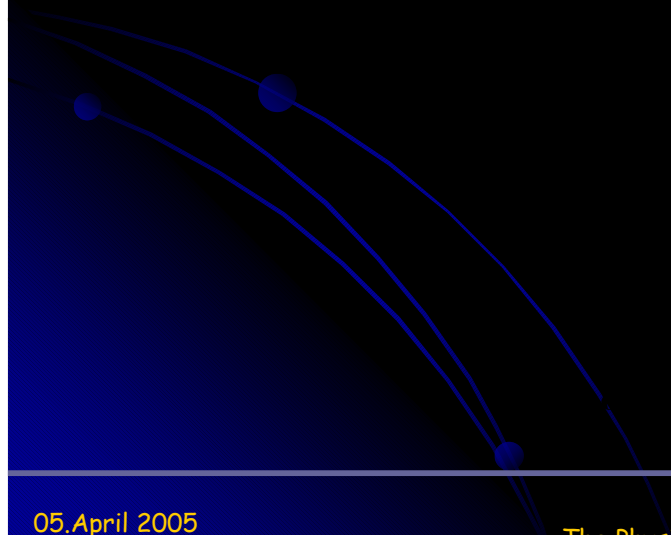
HERMES calculations follow perfectly the exponential law
 $(1-P_{\text{reac}}) = \exp(-L/L_{\text{reac}})$



Average neutron multiplicity M_n/p for 1.2 and 2.5 GeV p+Hg, Pb and W

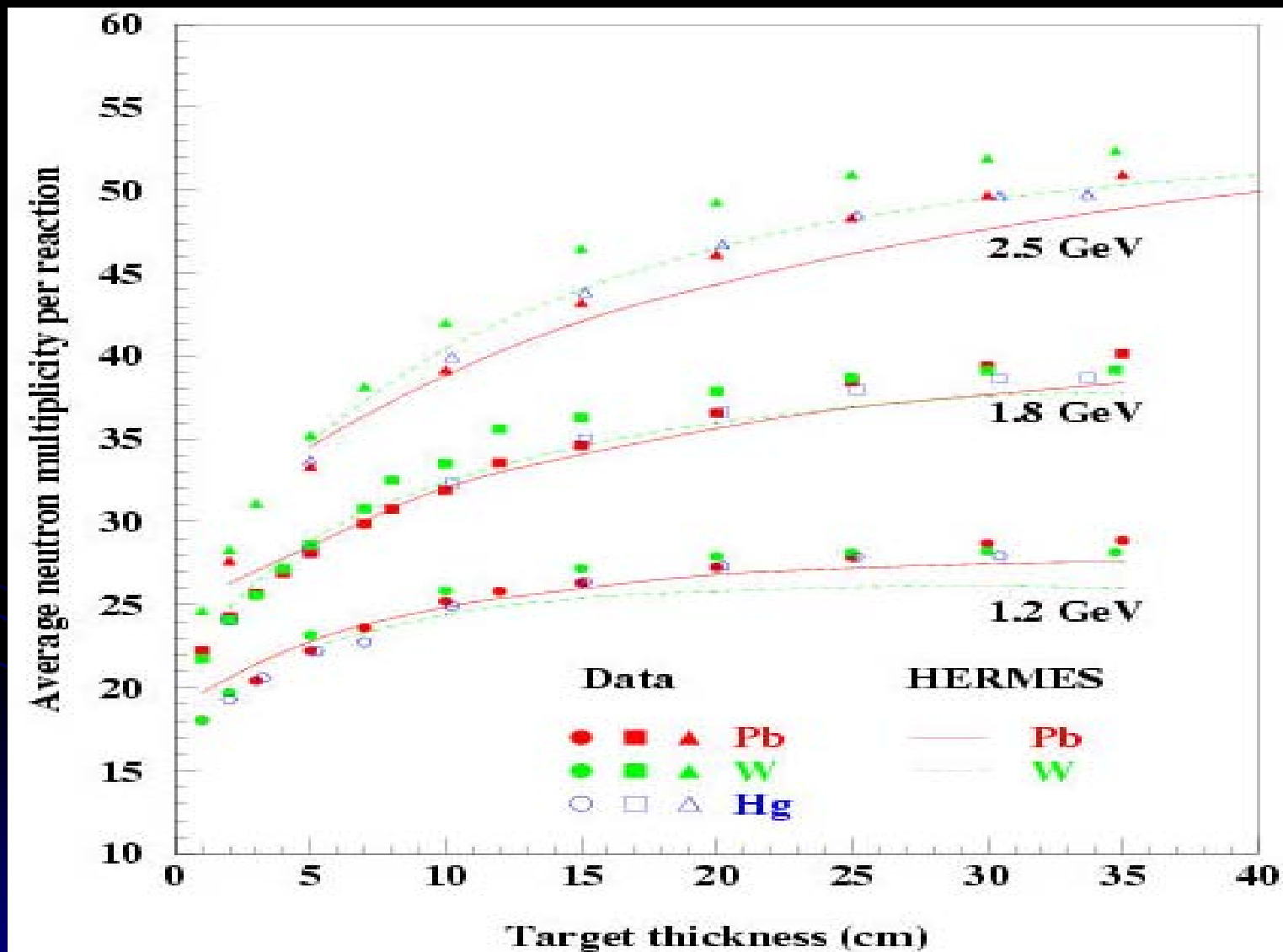


Average...



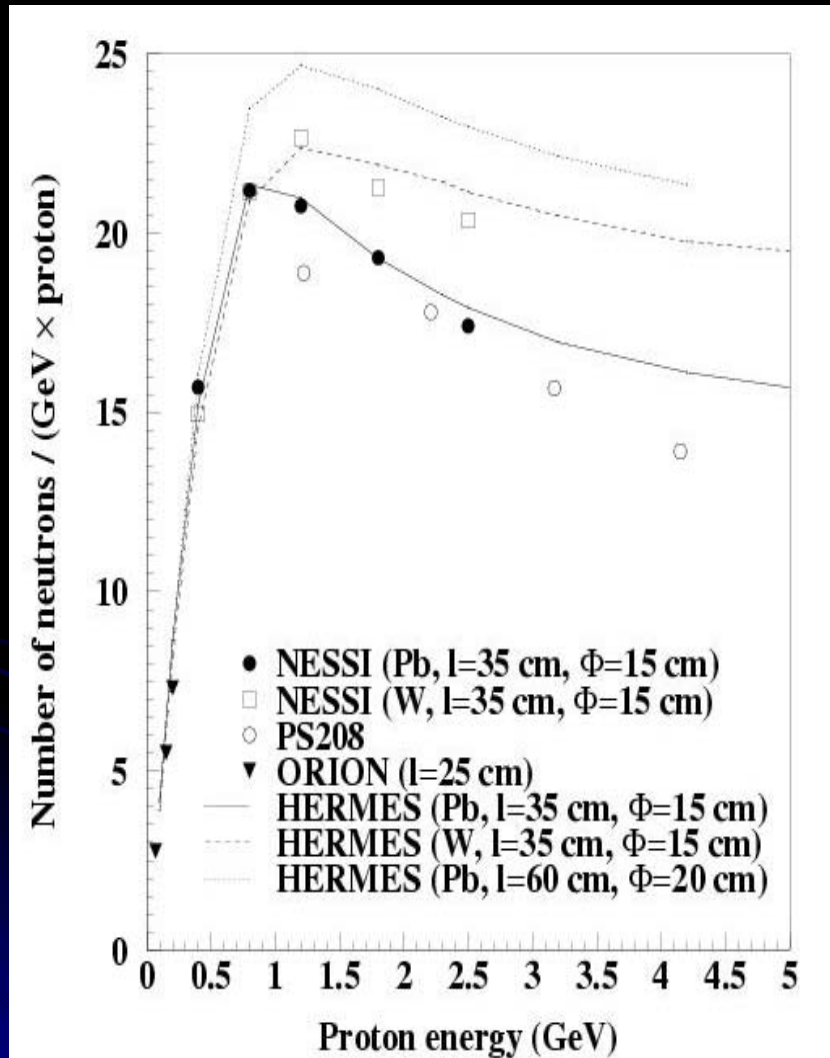
Neutron multiplicities in thick targets

Forschungszentrum Jülich
in der Helmholtz-Gemeinschaft





Neutron multiplicities in thick Pb and W targets

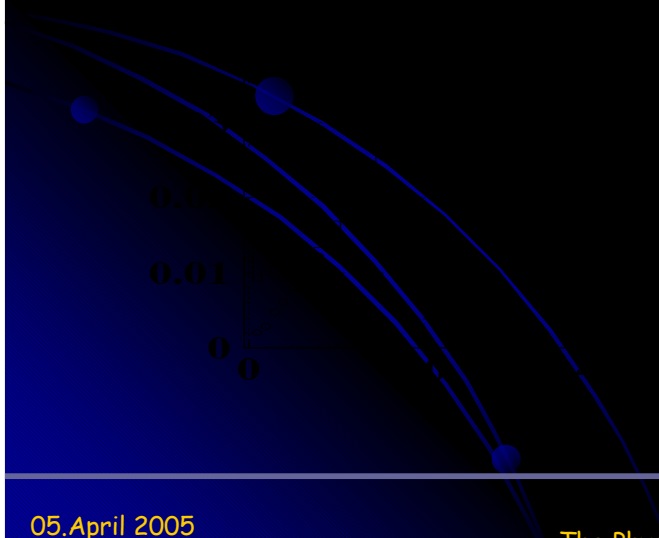


NIM B 170 (2000), 299



HERMES/MCNPX 1.2 GeV p+Hg

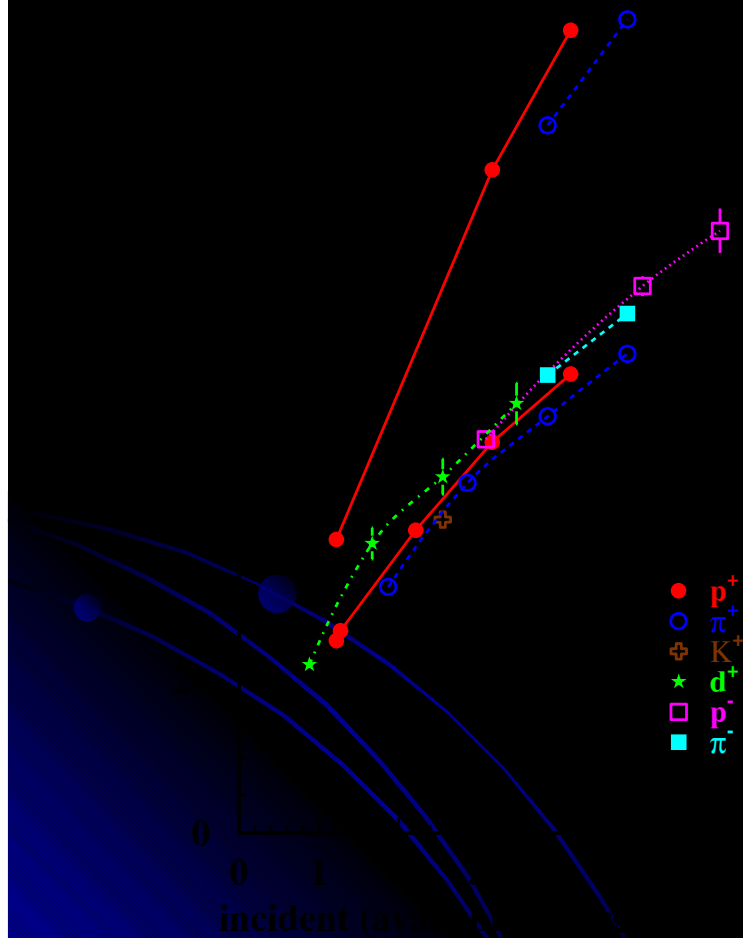
Distributions!
putting much
stronger constraints
to the models...



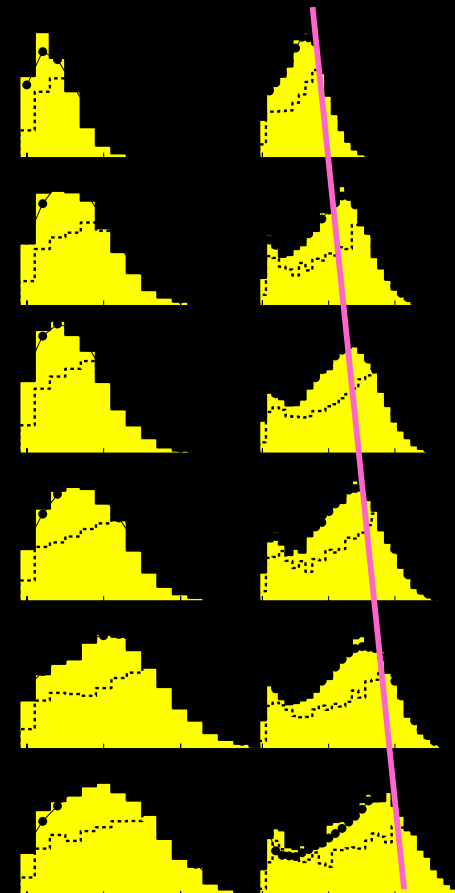


Neutron multiplicity distributions for 1.2 GeV p + Al ... U

Dickes Target

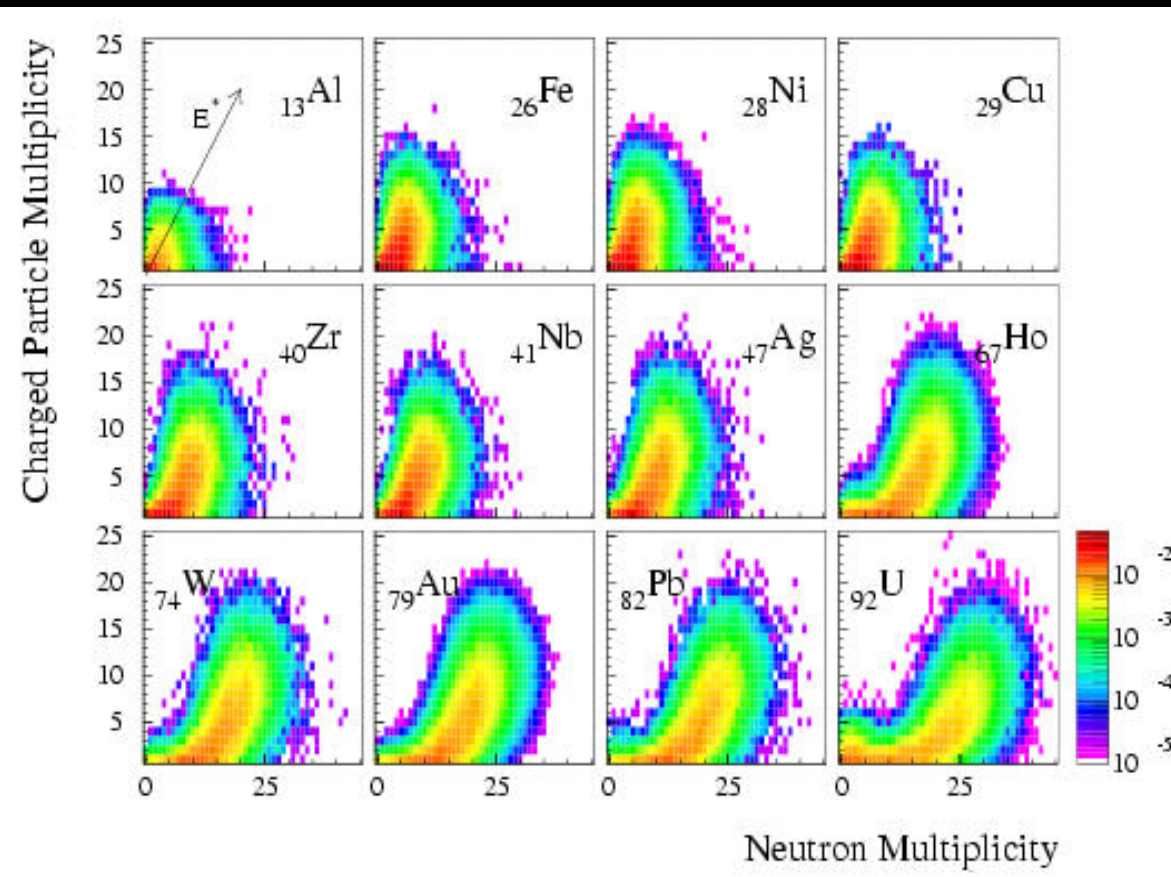


Thin target 1.2 GeV p + X





Light particle production in spallation reactions induced by protons of 0.8 to 2.5 GeV incident kinetic energy (NESSI)



Correlation of measured LCP-vs. N-multiplicity for 2.5 GeV proton-induced spallation reactions

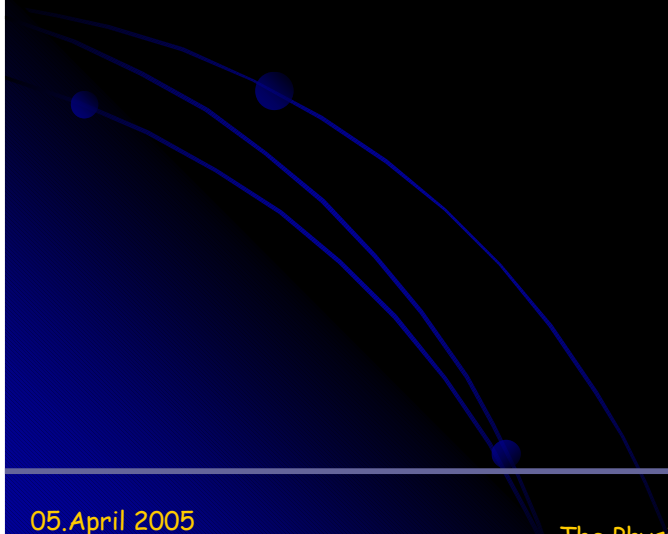
The color scale gives the relative yield for each target per multiplicity bin

Thermal excitation is following indicated arrow



Distributions of E^* for 0.4, 1.2 and 2.5 GeV p+Au induced reactions

Discrepancy found between the models among themselves...

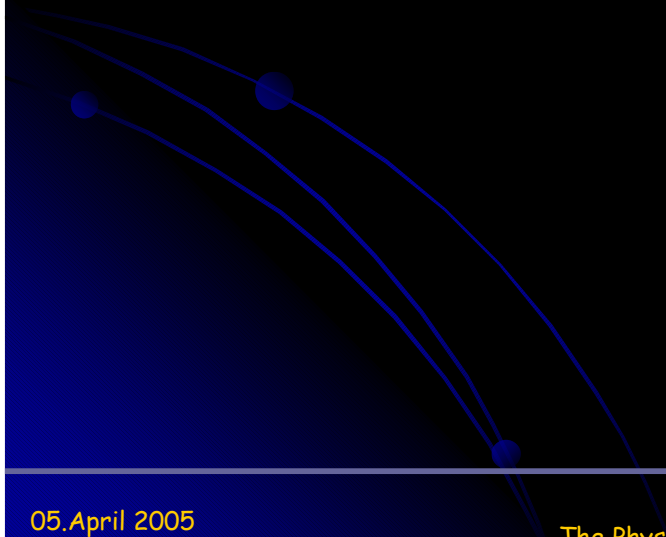




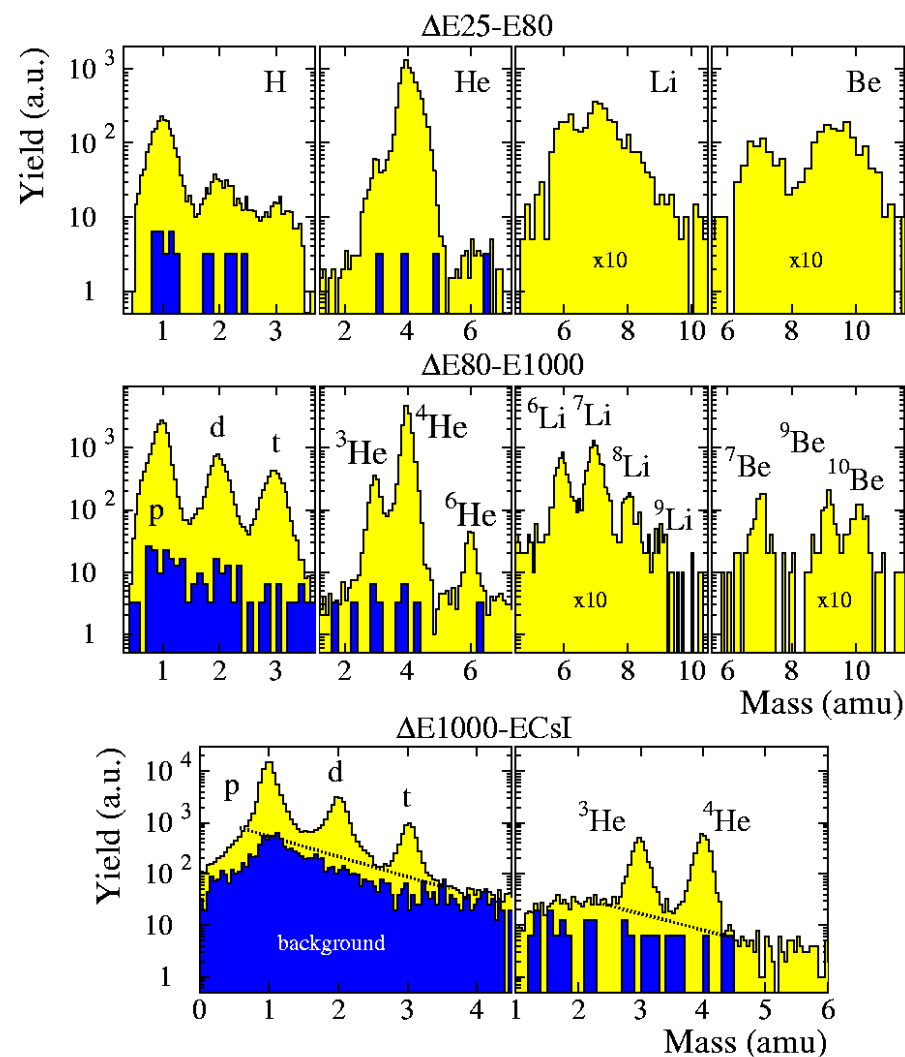
Kinetic energy spectra of $\pi^\pm$ 1.2 GeV p+Au BERTINI/INCL2.0

Explanation for discrepancy...

Bertini:
less pions, lower energetic...



Mass spectra of H, He, Li and Be



measured with different combinations of ΔE -E telesc.

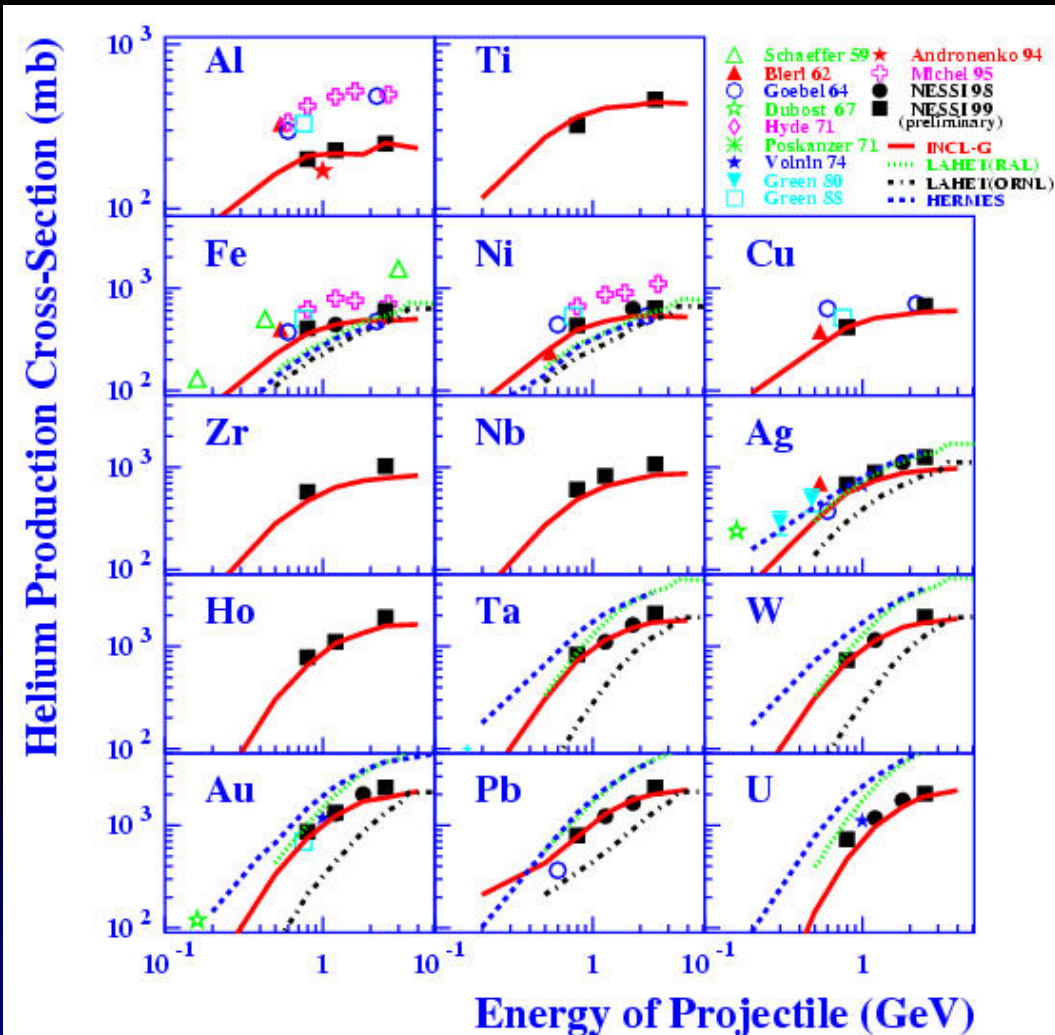
blue: background empty target frame

For CsI high background due to sensitivity and high light output for neutrons and gammy rays

Energy calibration of CsI for individual particles by reference to the energy loss in the 1000 μm Si detector in front; for high energy H ions CsI punch-through energies were used for calibration (160, 210, 250 MeV for p,d,t)



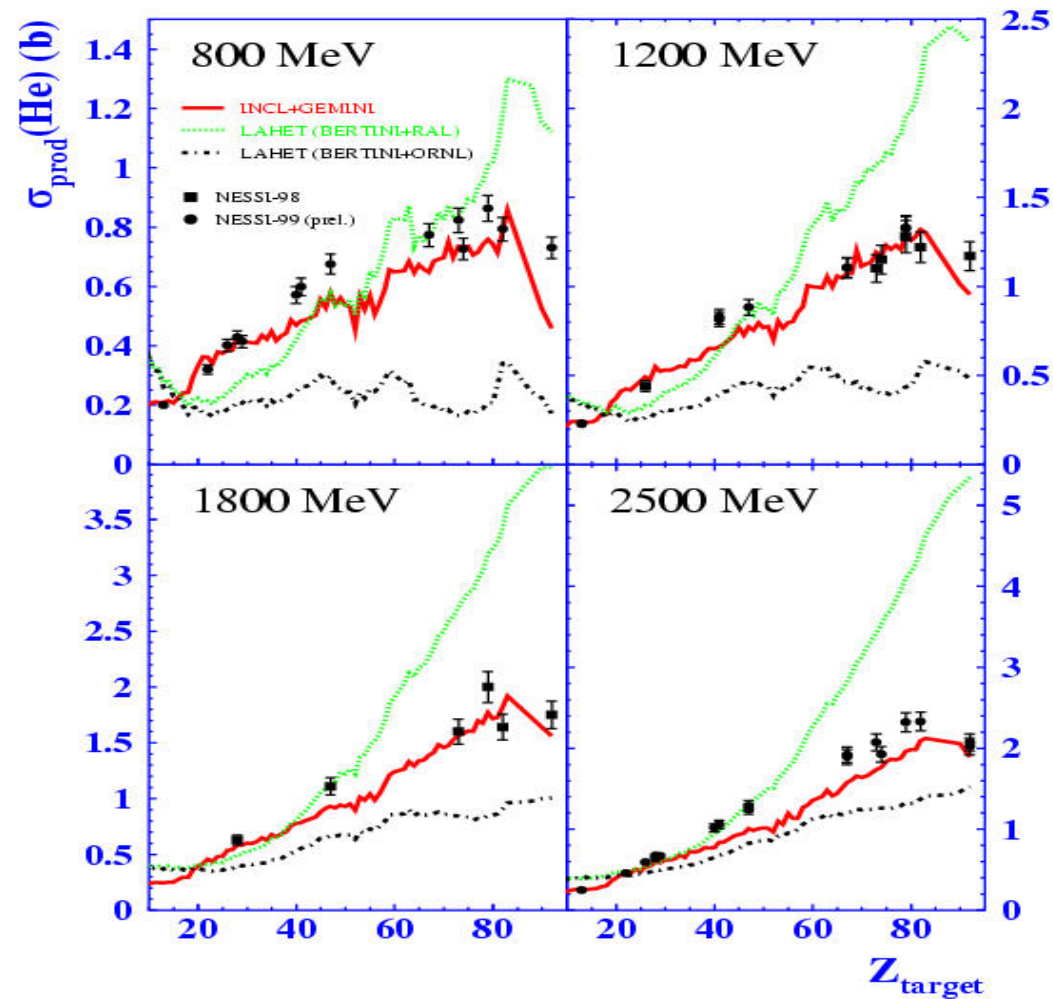
H production cross sections as a function of E_p for p+Al,...,U



Comparison of experimental He production cross sections (symbols) and calculations (INCL, LAHET, HERMES as lines) as a function of proton energy for various targets.

The first authors and year of the publications of previous experiments are also indicated.

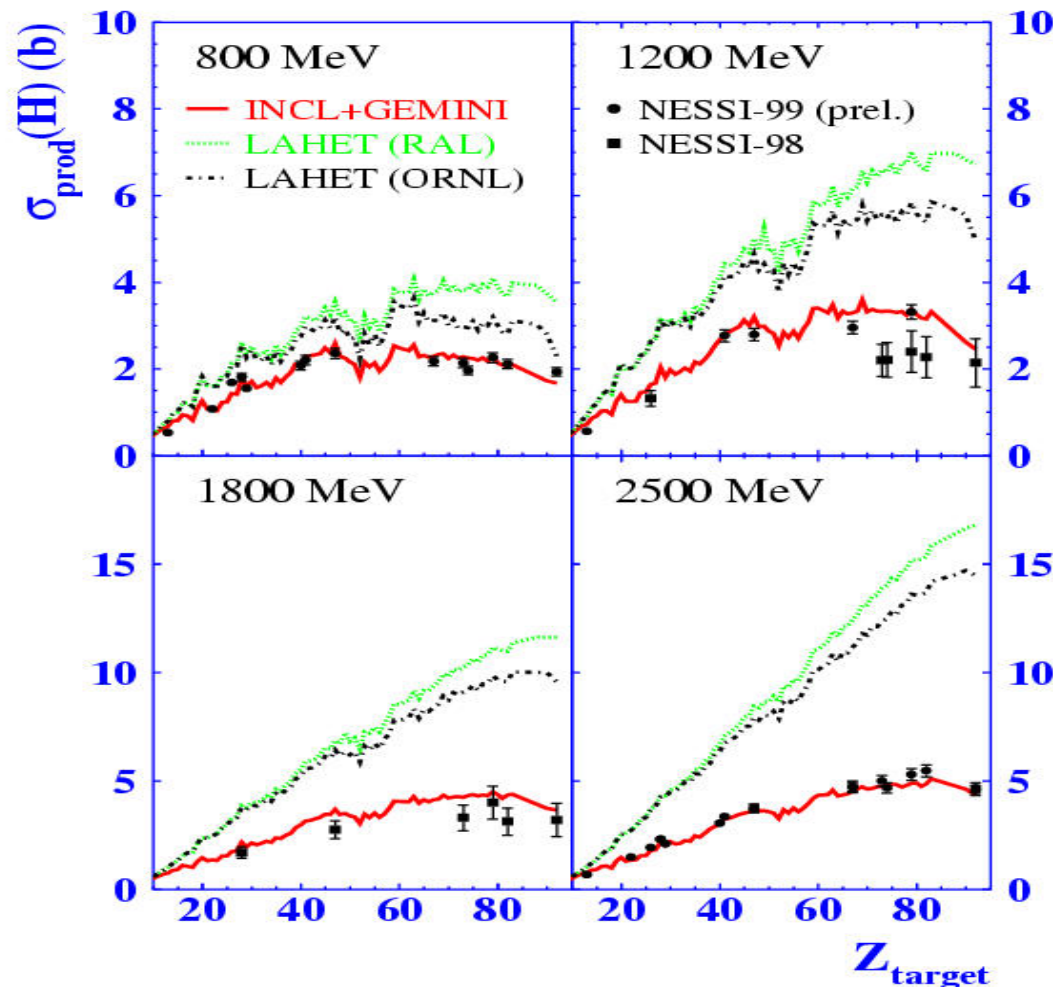
He production cross sections (Z, E_p)



Note the different scales of the left and the right panels.



H production cross sections (Z, E_p) for p,d,t with thresholds of 2.2-26,49,76 MeV



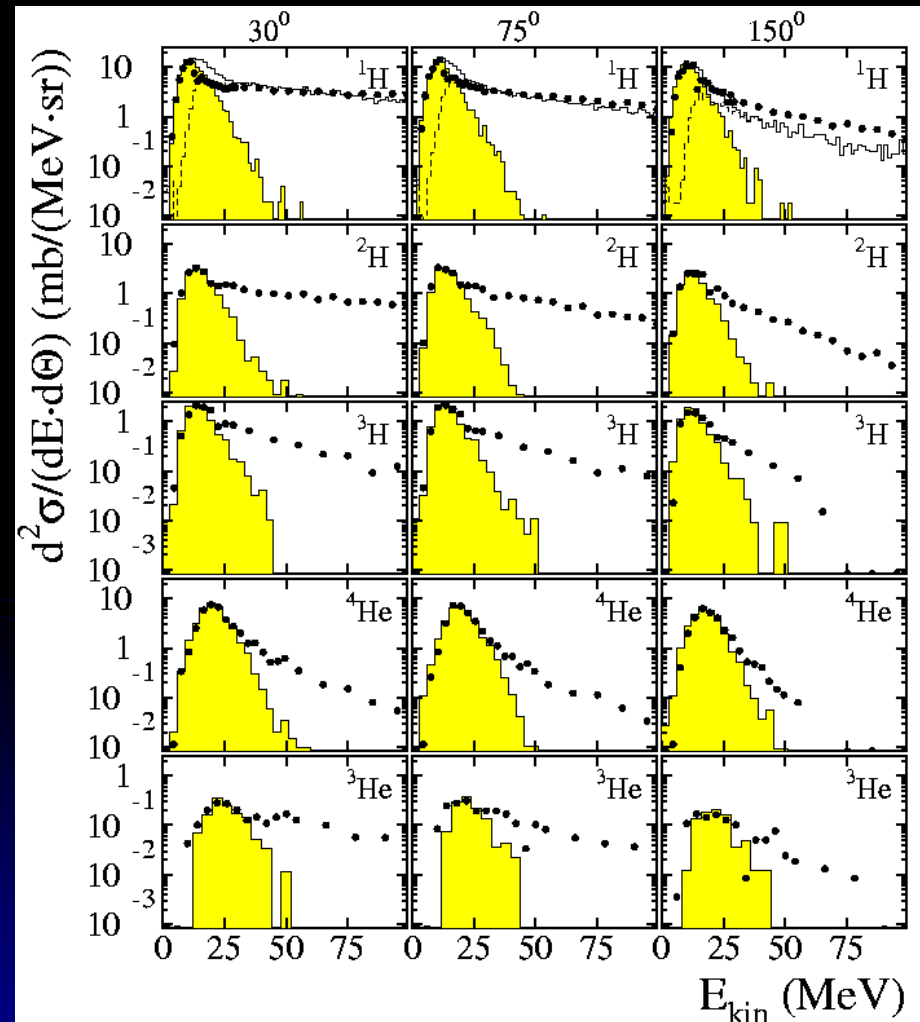
-the INCL code combined with the GEMINI evaporation code describes reasonably well the H- and He- production cross sections for light and heavy nuclei.

-the LAHET code with RAL option overestimates the H production for all Z and the He production for heavy nuclei while for light nuclei and low bombarding energies the He production is underestimated.

-the ORNL option for the evaporation code (no E^* dependence of Coulomb barriers underestimates the He production and still overestimates the H-production when compared to the experimental results of NESSI.



Energy spectra of $^1,2,3\text{H}$ and $^3,4\text{He}$



@30, 75 and 150° for the
reaction 1.2 GeV p + Ta

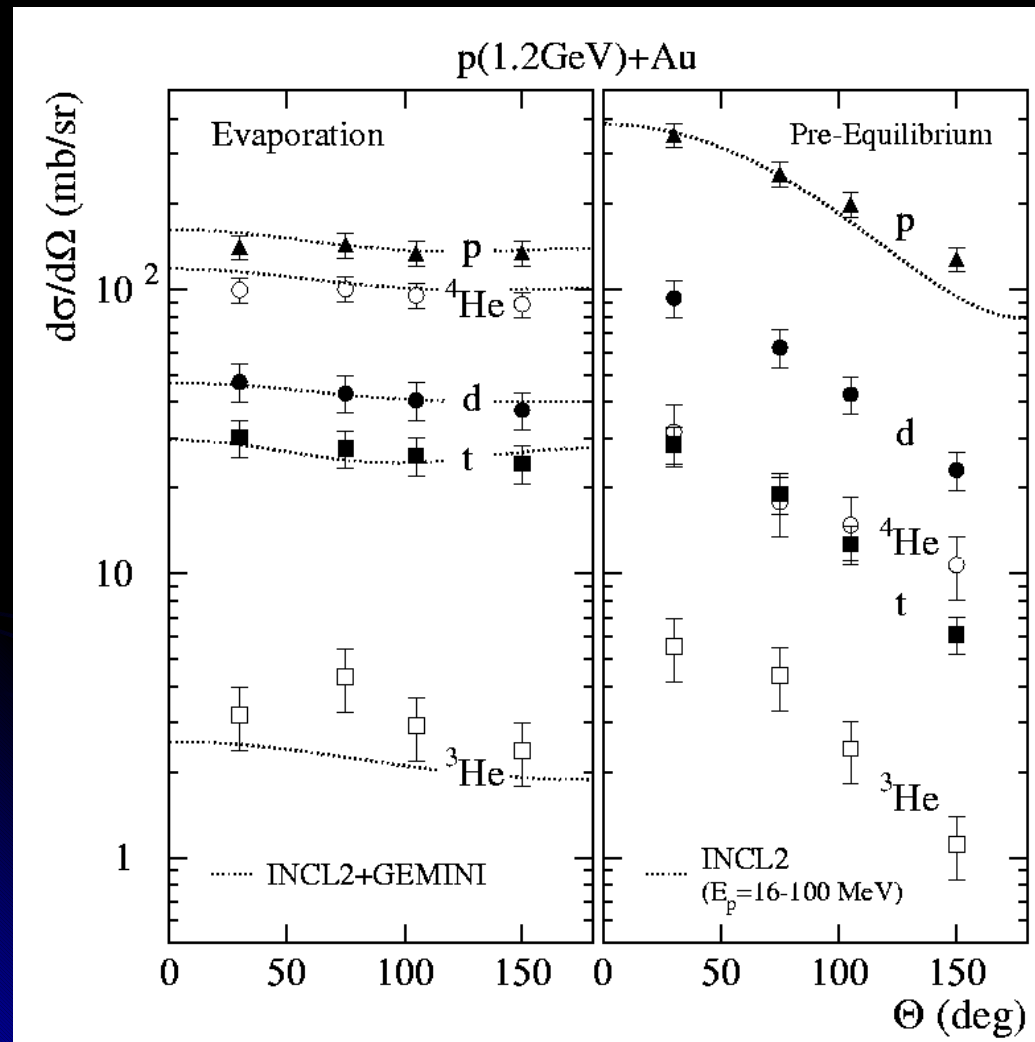
dots: experimental data
→ clearly feature two
components!

shaded histos: calculated
evaporation spectra
normalized to experimental
spectra

dashed histo: INC-protons by
INCL2.0



Angular distributions of LCPs in the 1.2 GeV p+Au



Symbols exp.

Lines: INCL2.0+GEMINI

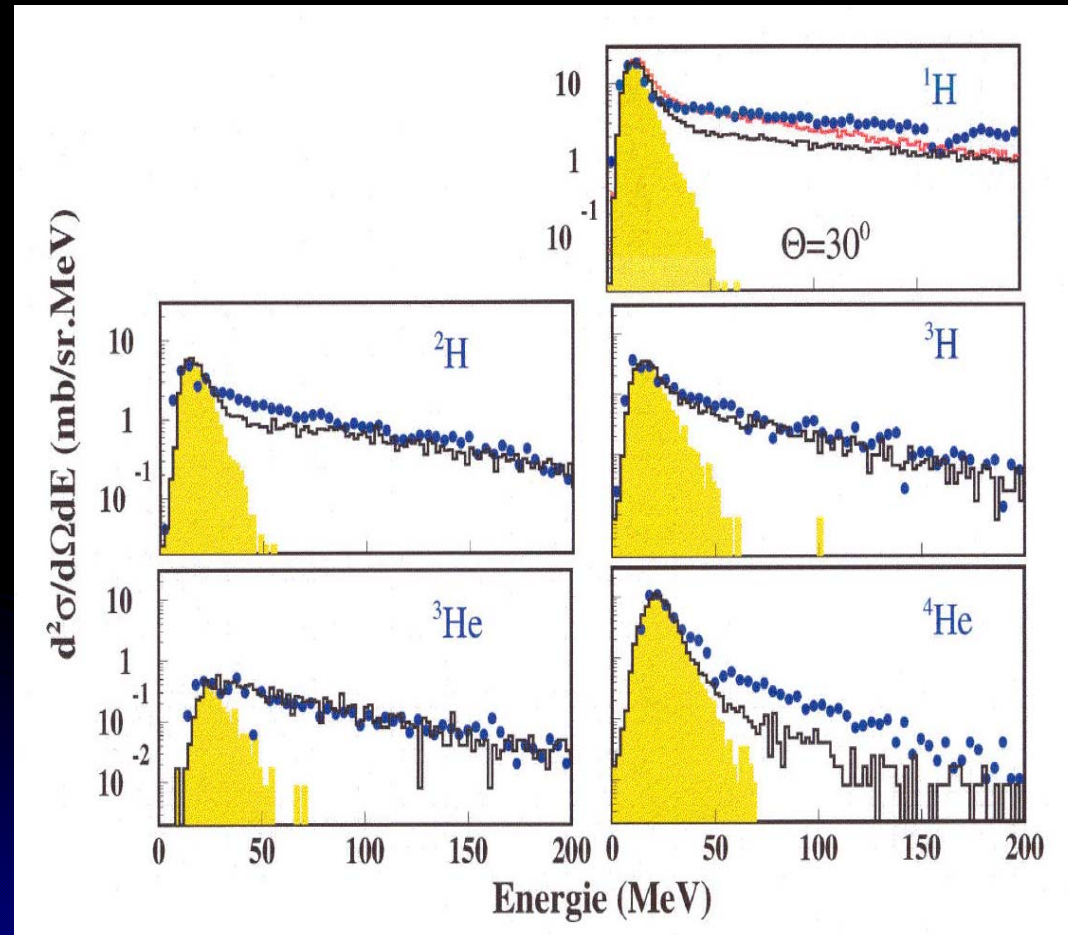
Left panel isotropic evap

Right panel: more directly emitted particles exhibit larger forward/backward asymmetry;

Protons are well described also in pre-equilibrium process...



$1,2,3$ H and $3,4$ He spectra for 2.5 GeV p+Au at 30 degree



Kinetic energy spectra for $1,2,3$ H and $3,4$ He measured by telescopes for 2.5 GeV p+Au at 30 degree (symbols)

Simulations with INCL including a coalescence model for the pre-equilibrium stage are shown as the black histograms.

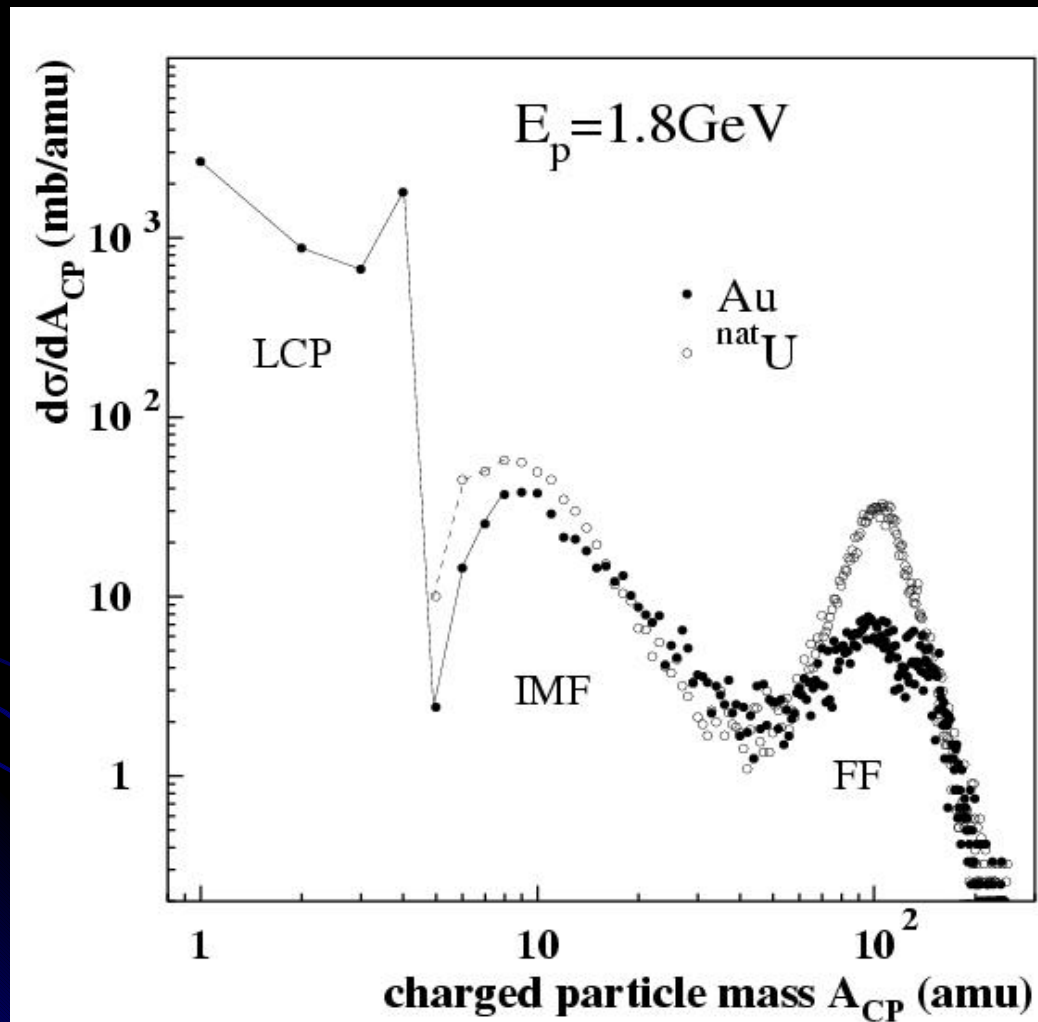
The yellow shaded histograms represent the evaporation component.

The red histogram for protons (upper right panel) displays the proton spectra of the original INCL calculations without preequilibrium.

4 He mostly evaporative!



Mass distribution for 1.8GeV p+Au,U



M.Enke et al.