

Spin-Resolved Angle-Resolved Photoemission

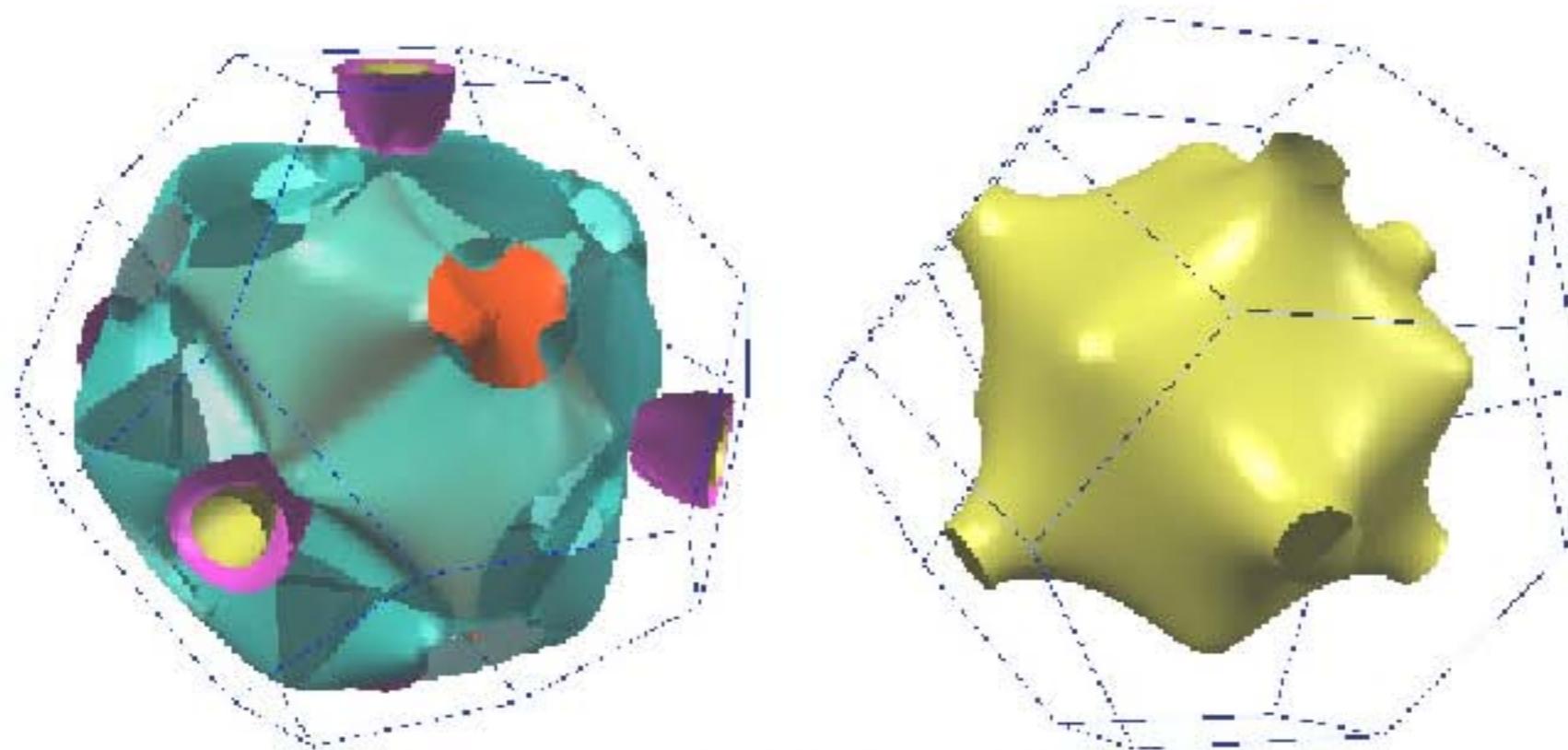
Jürg Osterwalder

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CH-8057 Zürich, Switzerland - osterwal@physik.unizh.ch
<http://www.physik.unizh.ch/groups/grouposterwalder/>*

Lecture 4

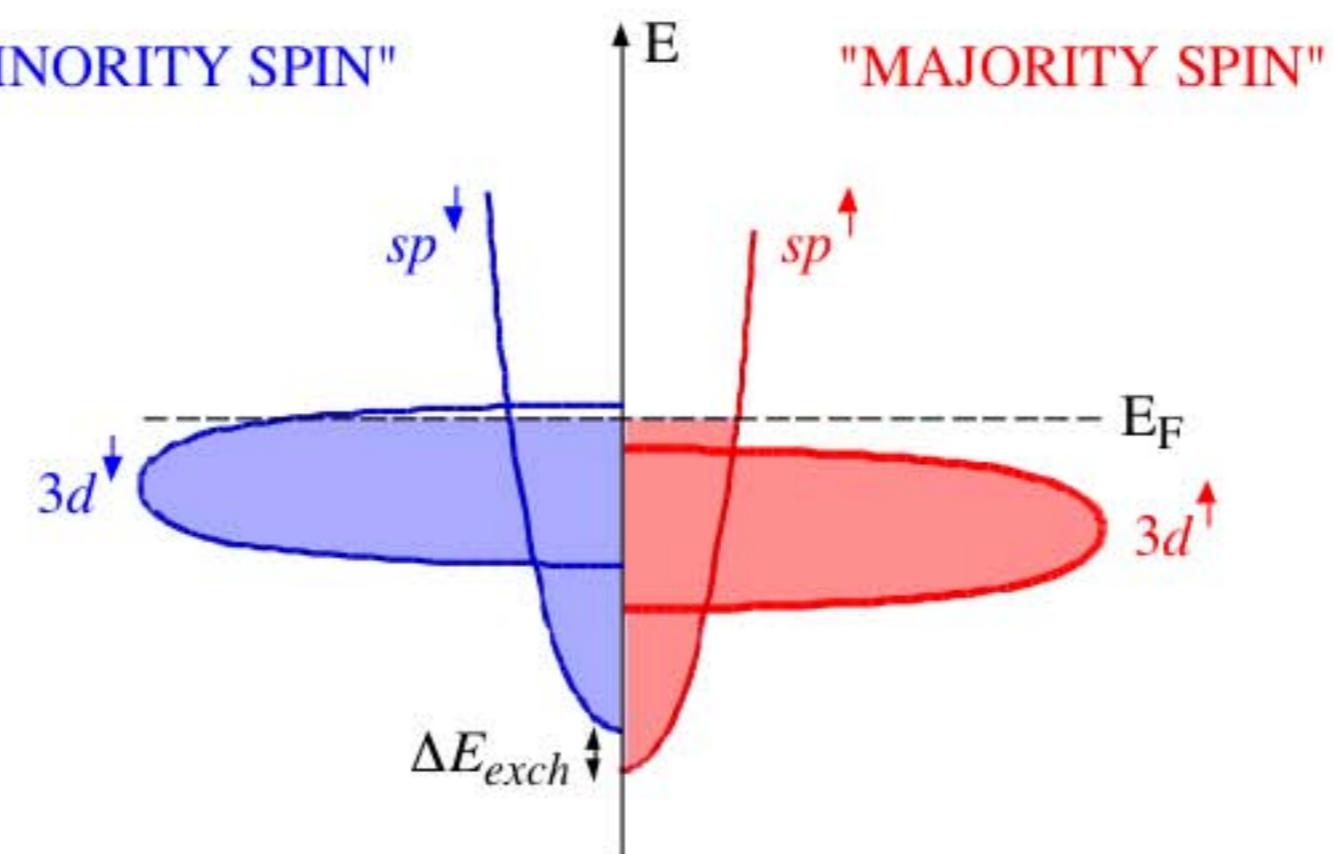
- Band Structure of an Itinerant Ferromagnet
- Measuring the Electron Spin
- Quantitative Aspects of Spin Polarization: Ni(111)
- The Magnetic Phase Transition in Ni
- Ultrathin Films of Ni on Cu(001)
- The Rashba Effect in the Surface State on Au(111)

The Fermi Surface of Nickel Metal

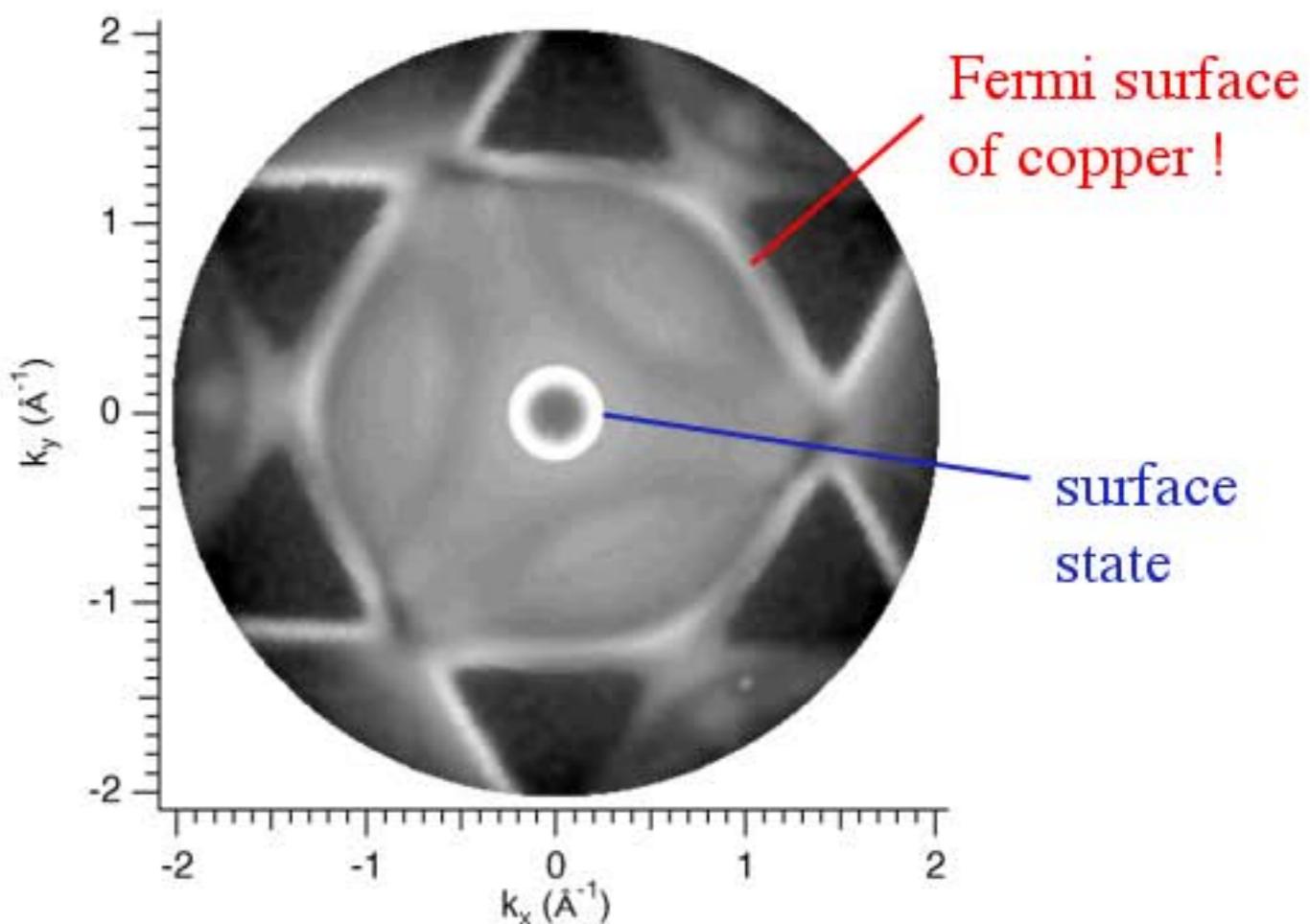
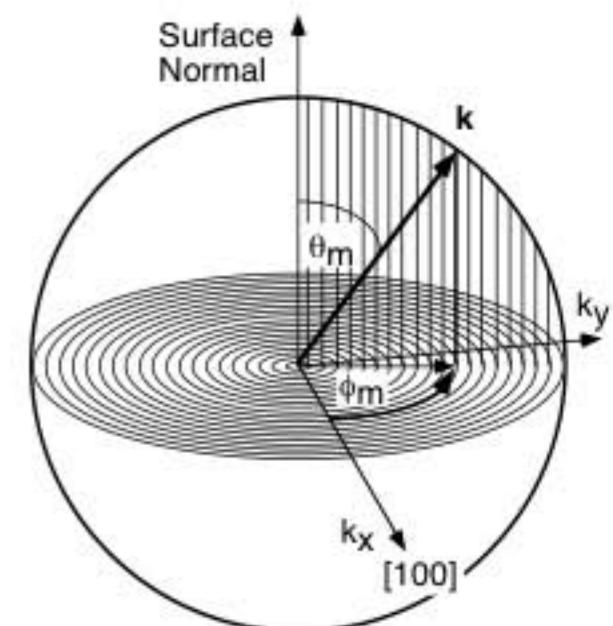
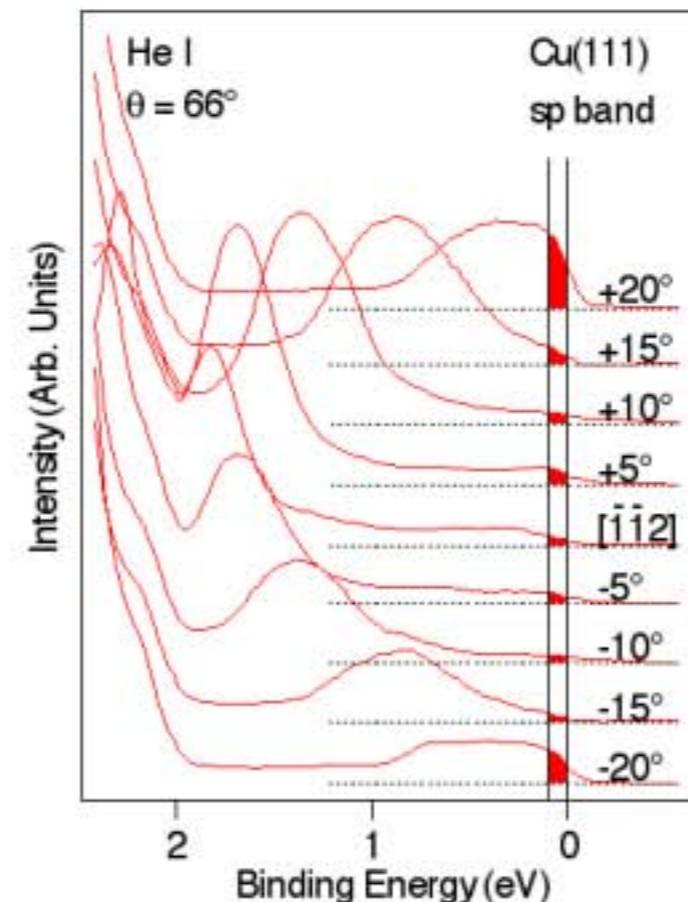
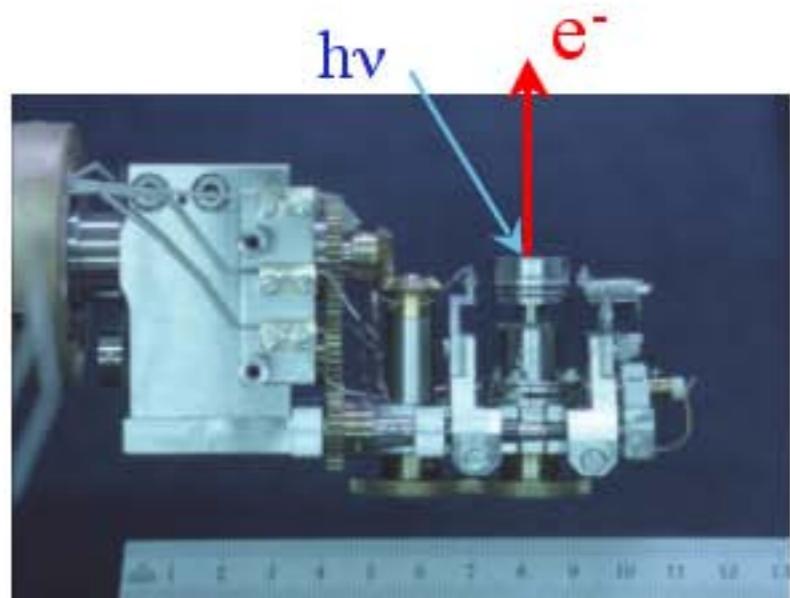
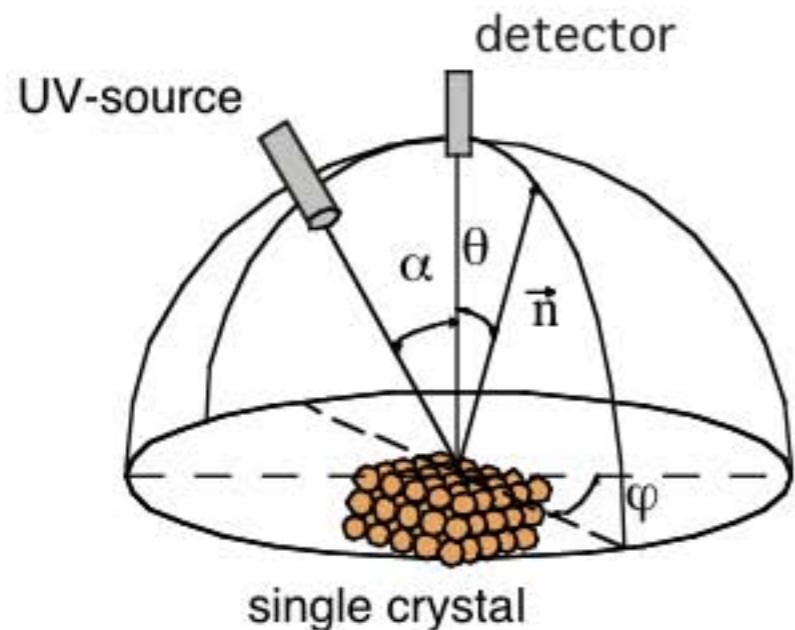


From <http://www.phy.tu-dresden.de/~fermisur/>

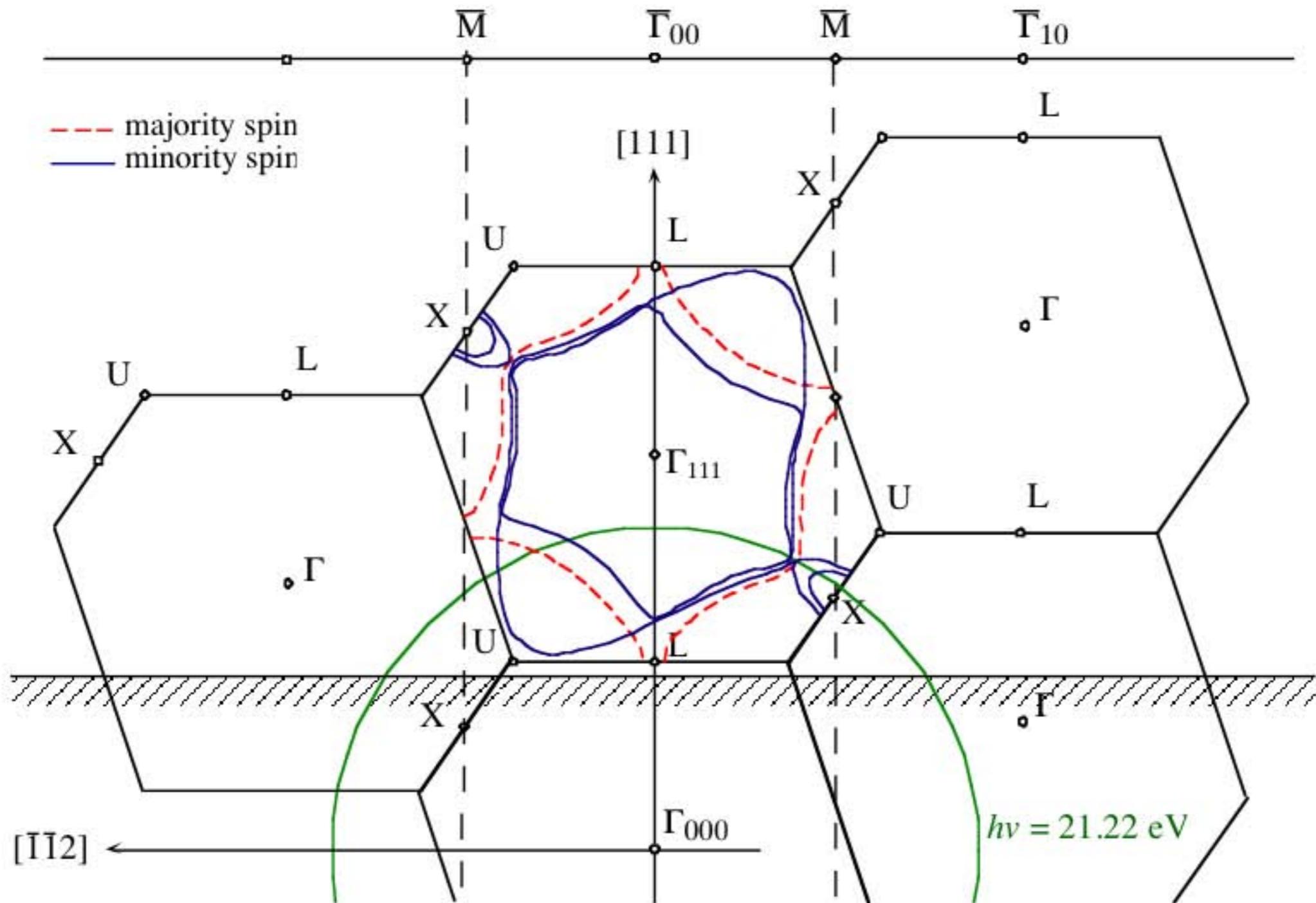
Schematic Density of States



Fermi Surface Mapping by Photoemission

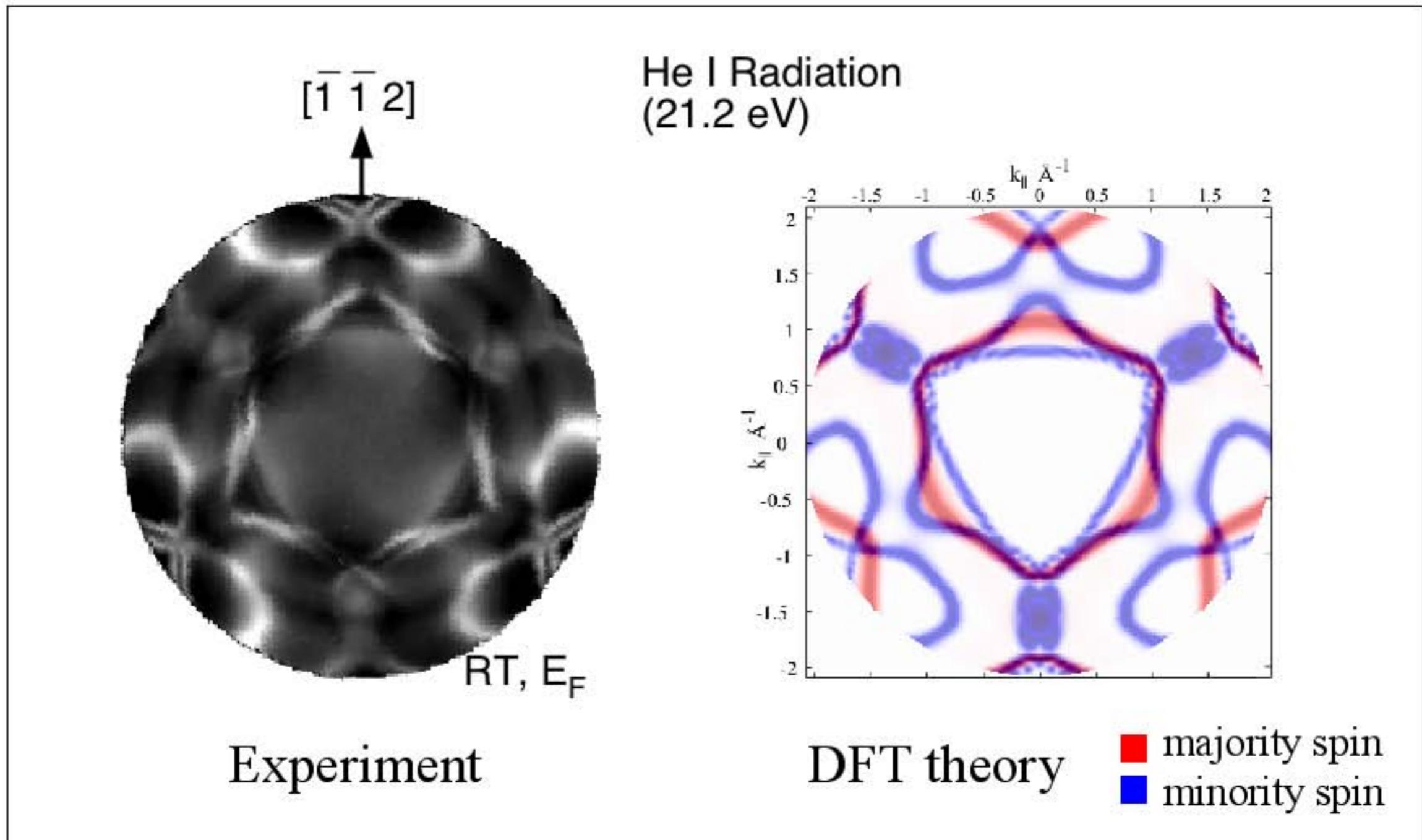


Reciprocal Space on Ni(111)



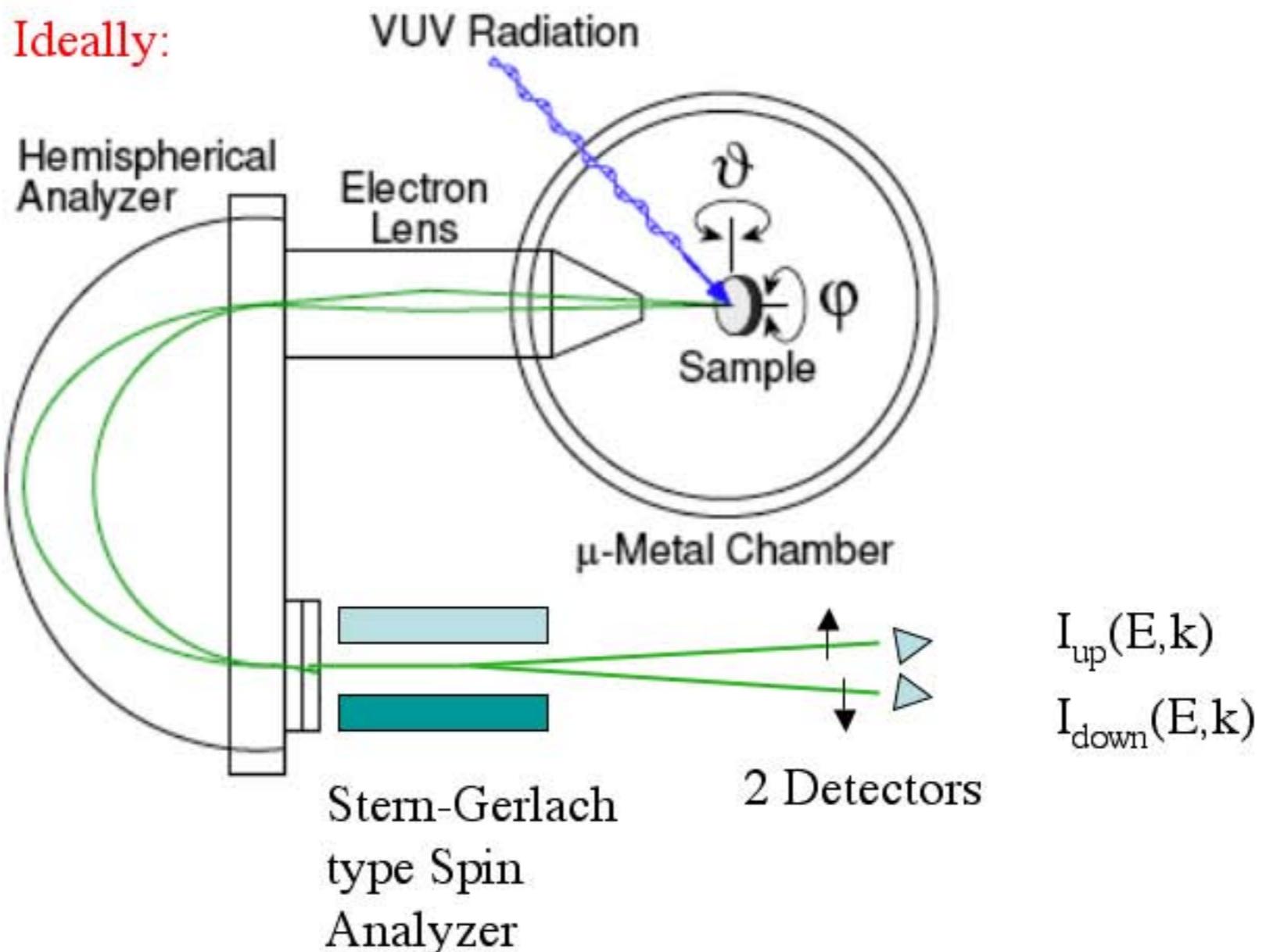
Fermi surface calculation (density functional theory, Wien 97)

Fermi Surface of Ni as Seen Through the Ni(111) Surface



M. Hoesch et al., JES 124, 263 (2002)

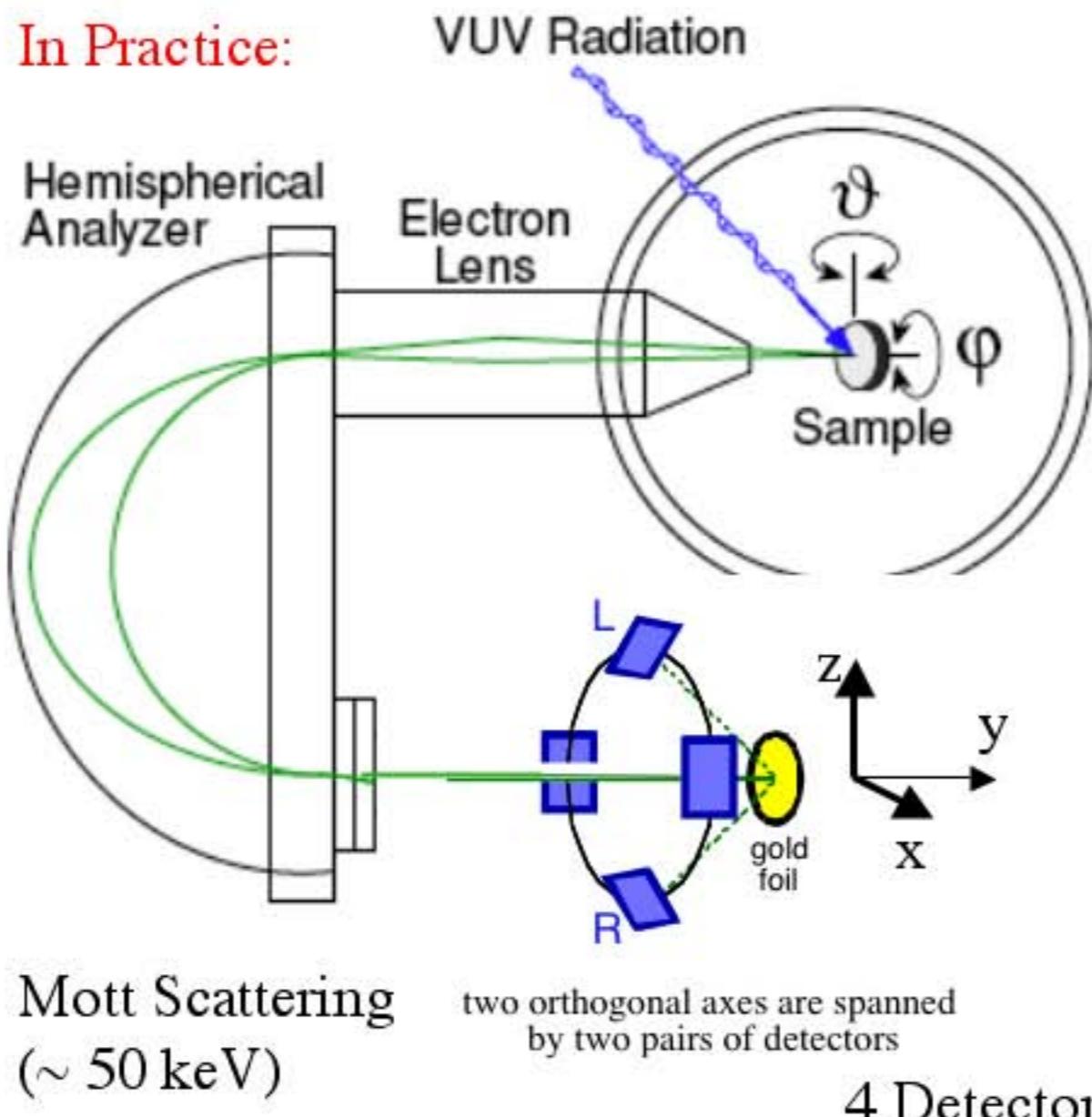
ARPES with Spin Resolution



Does not work! (Lorentz forces & uncertainty principle)

ARPES with Spin Resolution

In Practice:



Scattering
asymmetry

$$A_x = \frac{(I_L - I_R)}{(I_L + I_R)}$$

Spin Polarization

$$P_x = A_x / S.$$

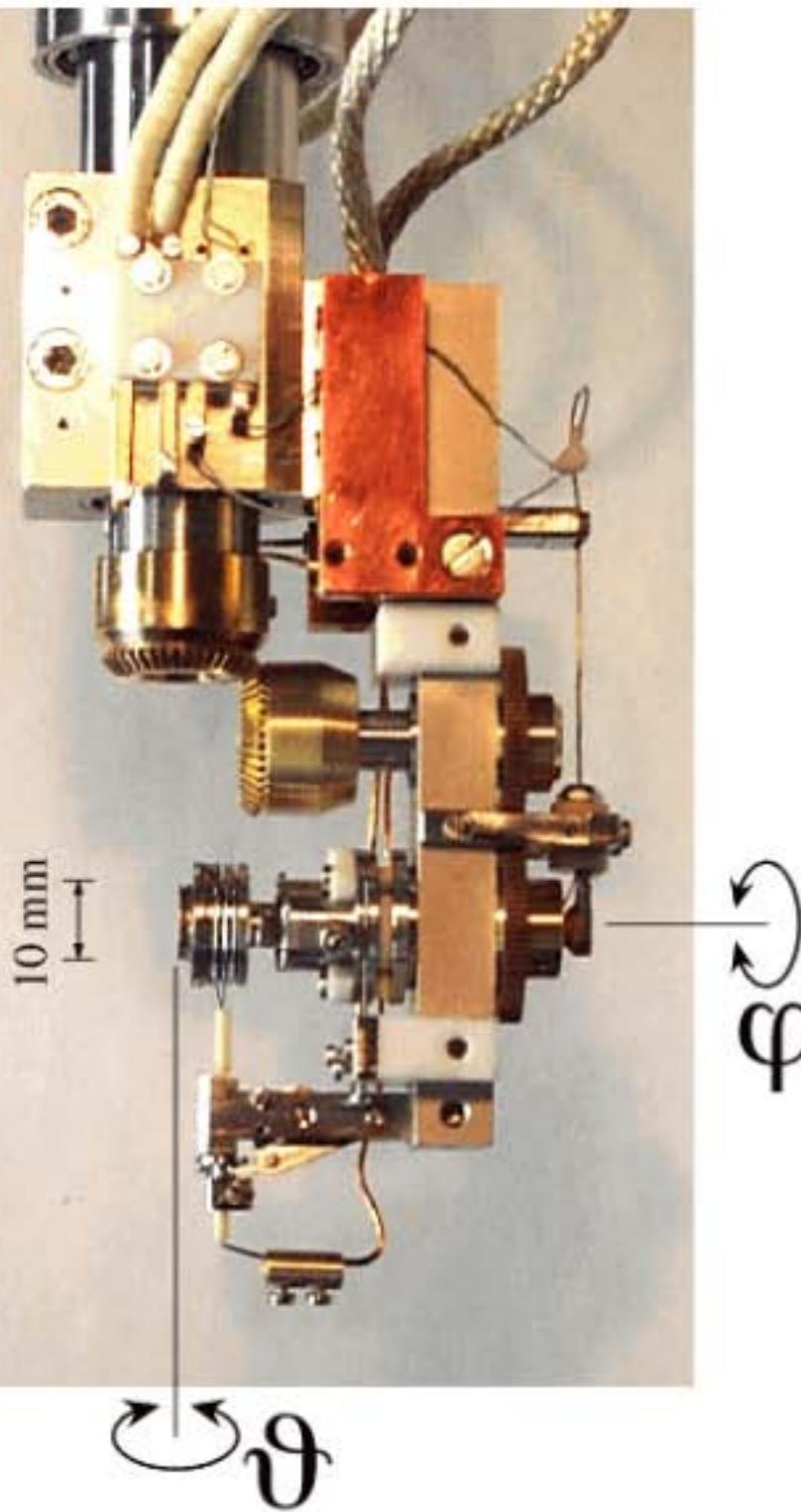
S: Sherman function

Inefficient !

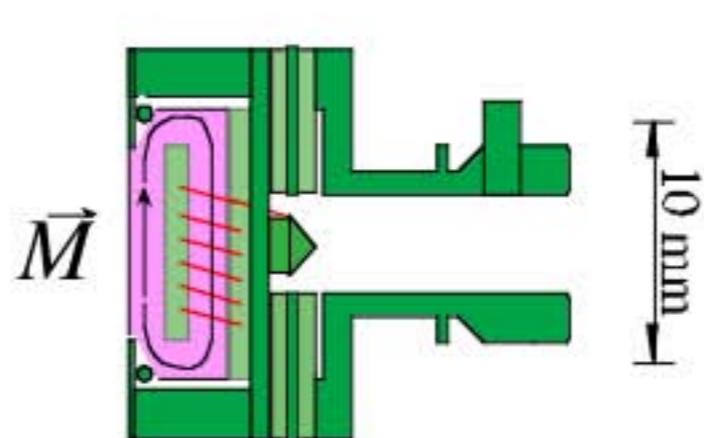
Spin Polarization: $P_I(E, k) = \frac{I_{\uparrow}(E, k) - I_{\downarrow}(E, k)}{I_{\uparrow}(E, k) + I_{\downarrow}(E, k)},$

$$I_{\uparrow,\downarrow}(E, k) = I_M(E, k)(1 \pm P_I(E, k))/2.$$

Important: Control over the sample magnetization in ARPES !



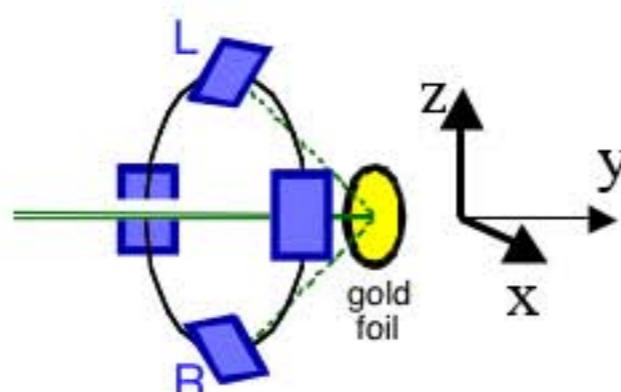
We need a magnetized sample (ideally a single domain) to measure spin polarized bands !



Switching magnetization direction ($\oplus \ominus$)

⇒ Forming cross asymmetries cancels instrumental asymmetries

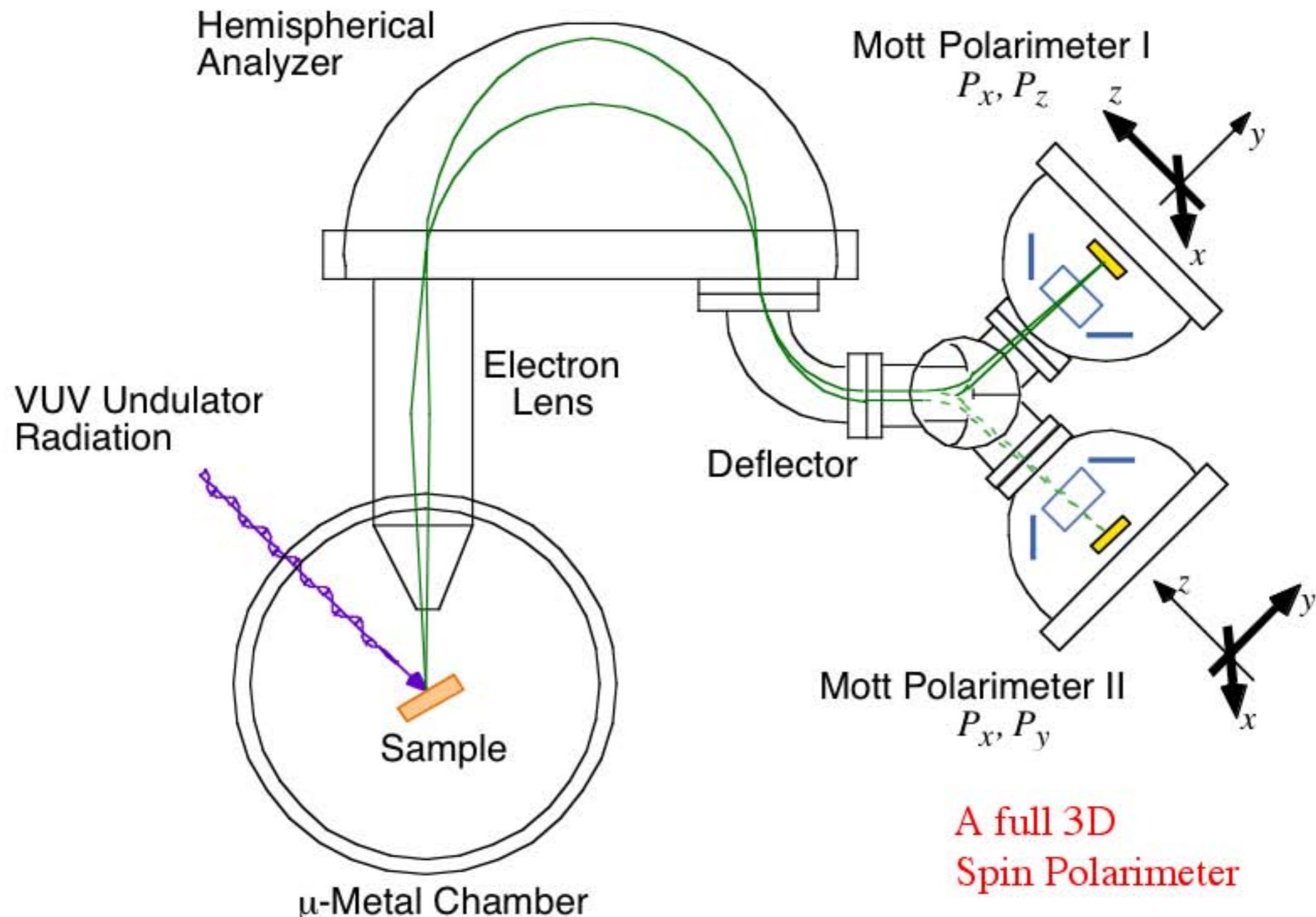
$$A^\otimes = \frac{(I_L^\oplus + I_R^\ominus) - (I_R^\oplus + I_L^\ominus)}{(I_L^\oplus + I_R^\ominus) + (I_R^\oplus + I_L^\ominus)}.$$



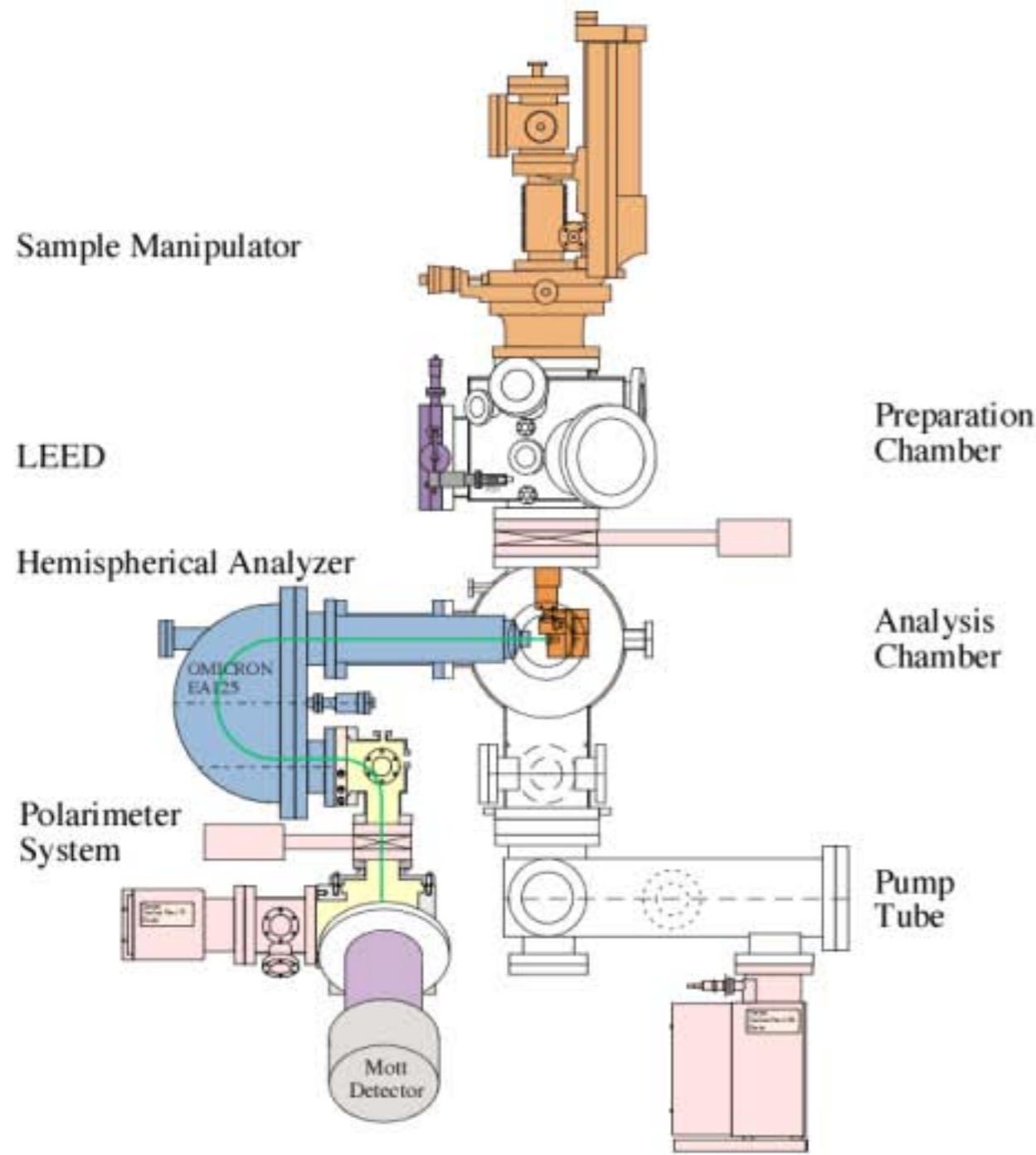
two orthogonal axes are spanned by two pairs of detectors

Polarization $P = A / S$
 $S \sim 0.15$, but not precisely known (requires calibration)
⇒ $P = 0$ is measured exactly !
⇒ P scale is known roughly

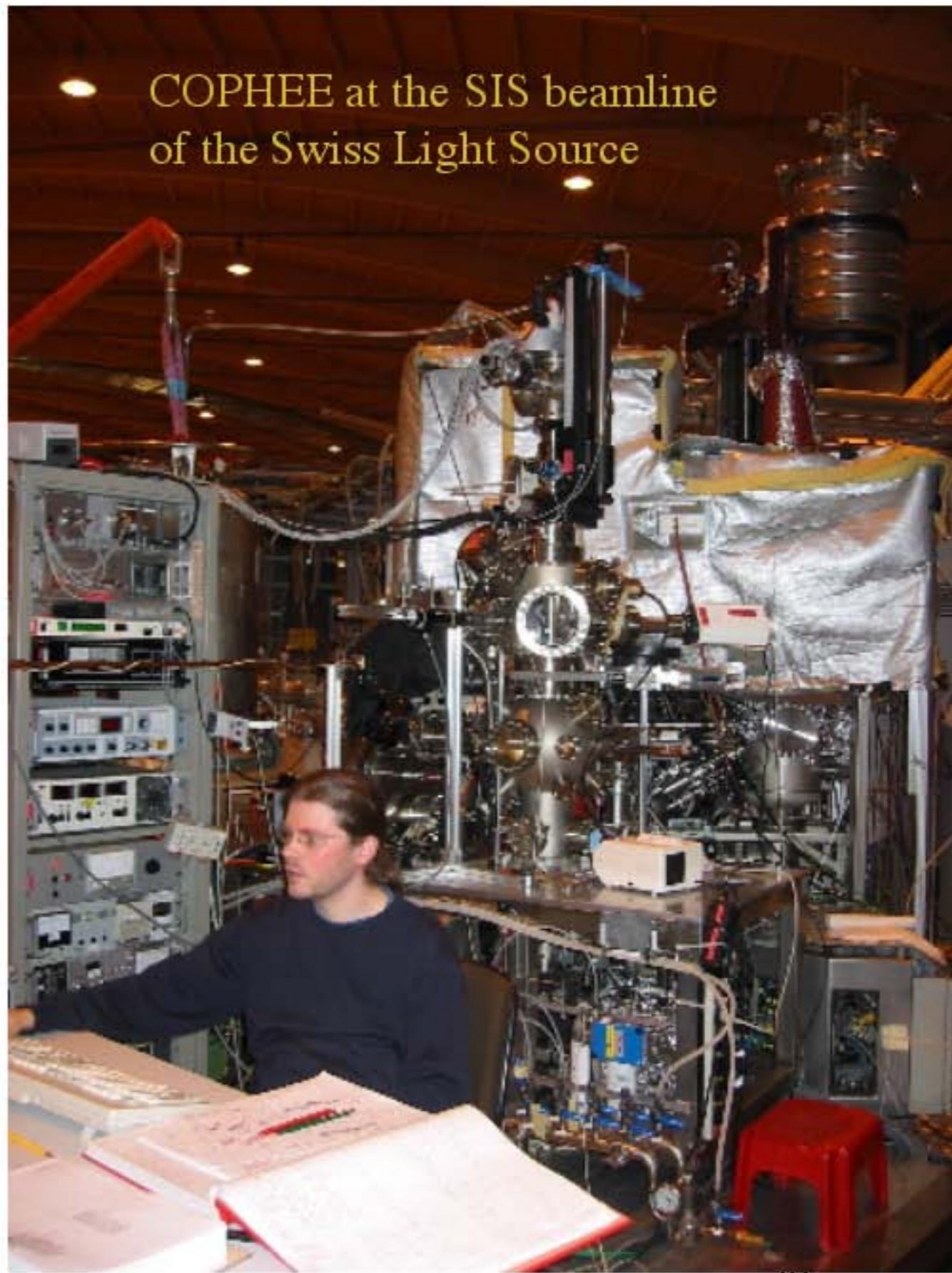
COPHEE - The Complete Photoemission Experiment



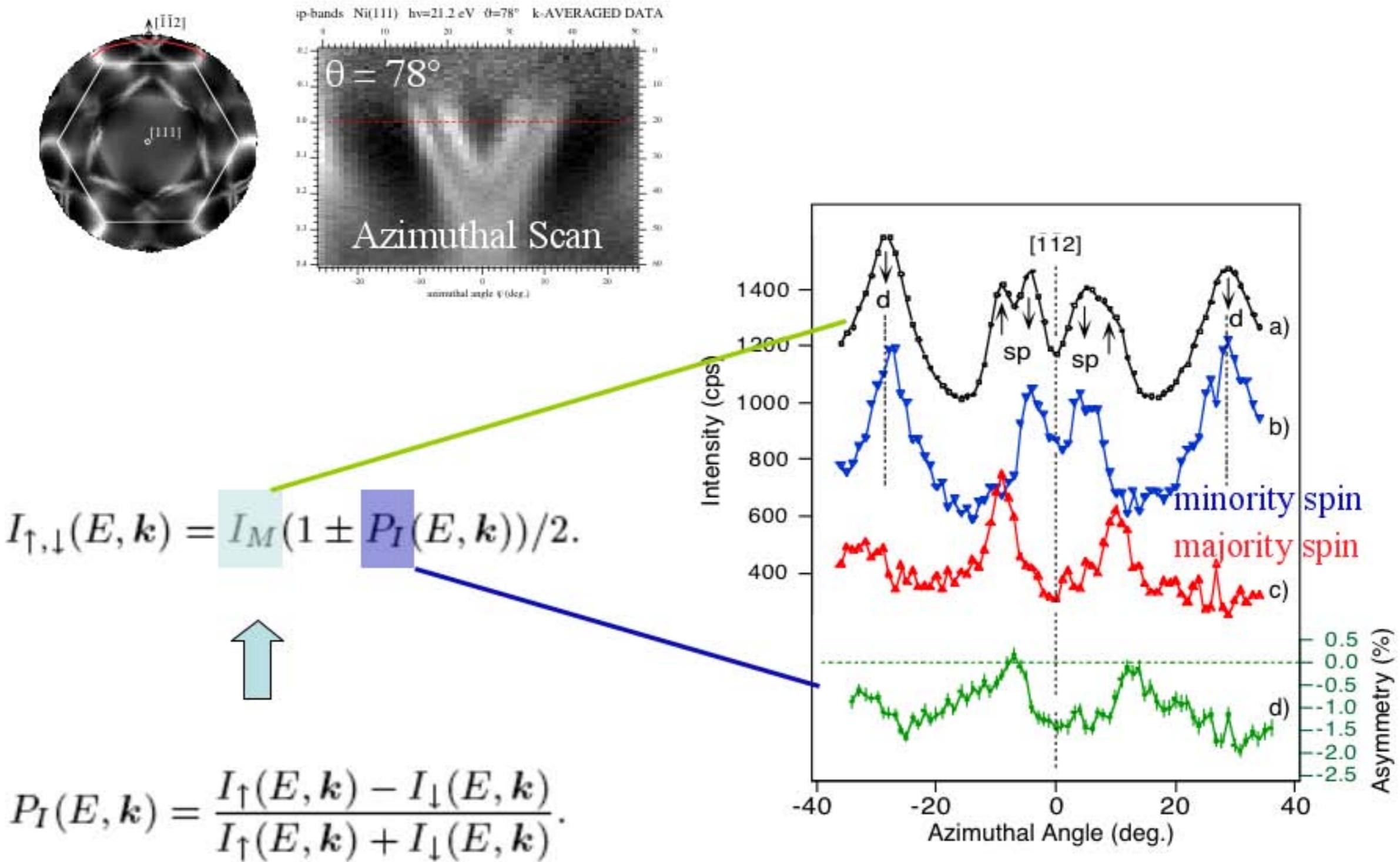
Setup of COPHEE:



COPHEE at the SIS beamline
of the Swiss Light Source

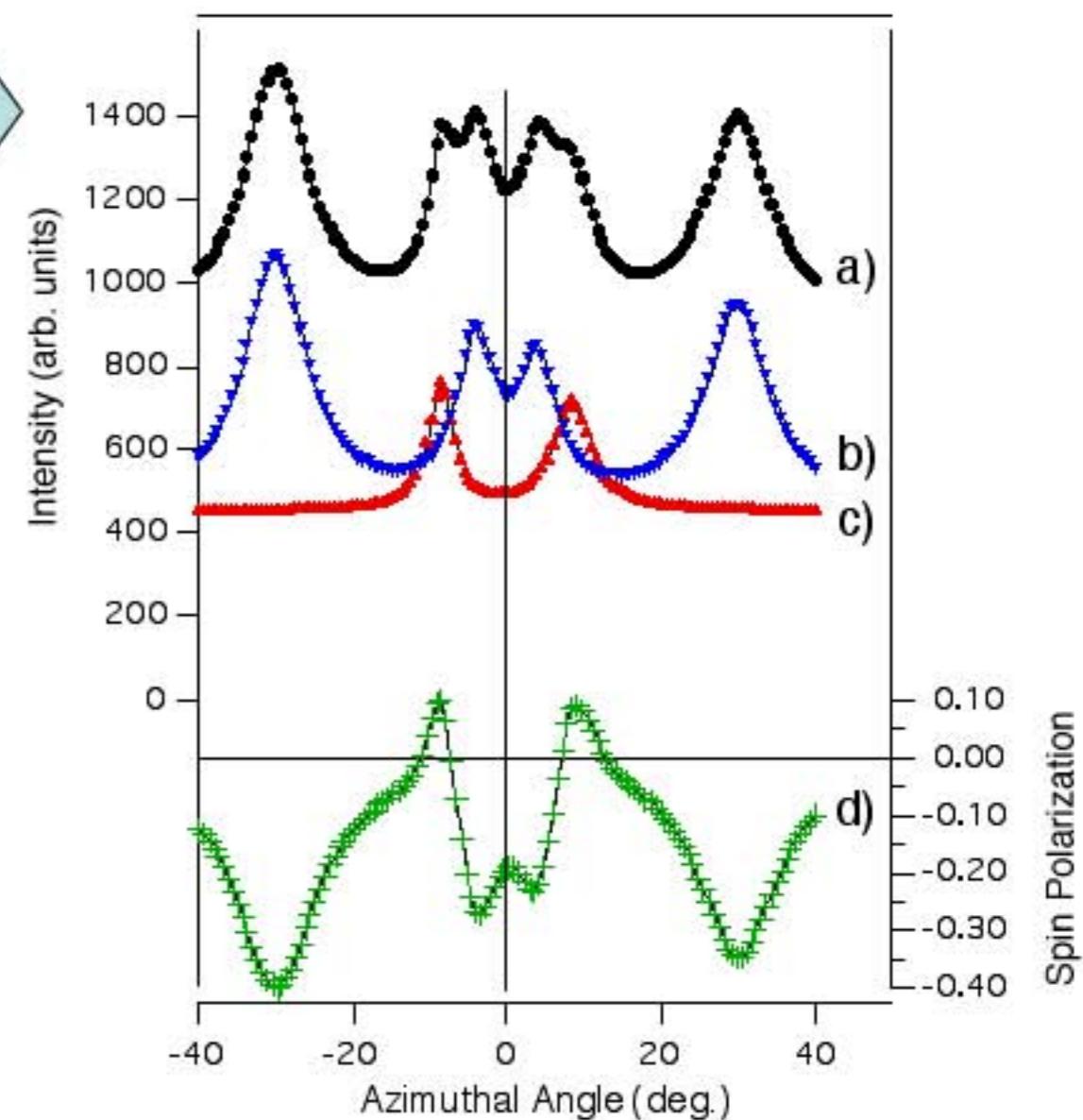
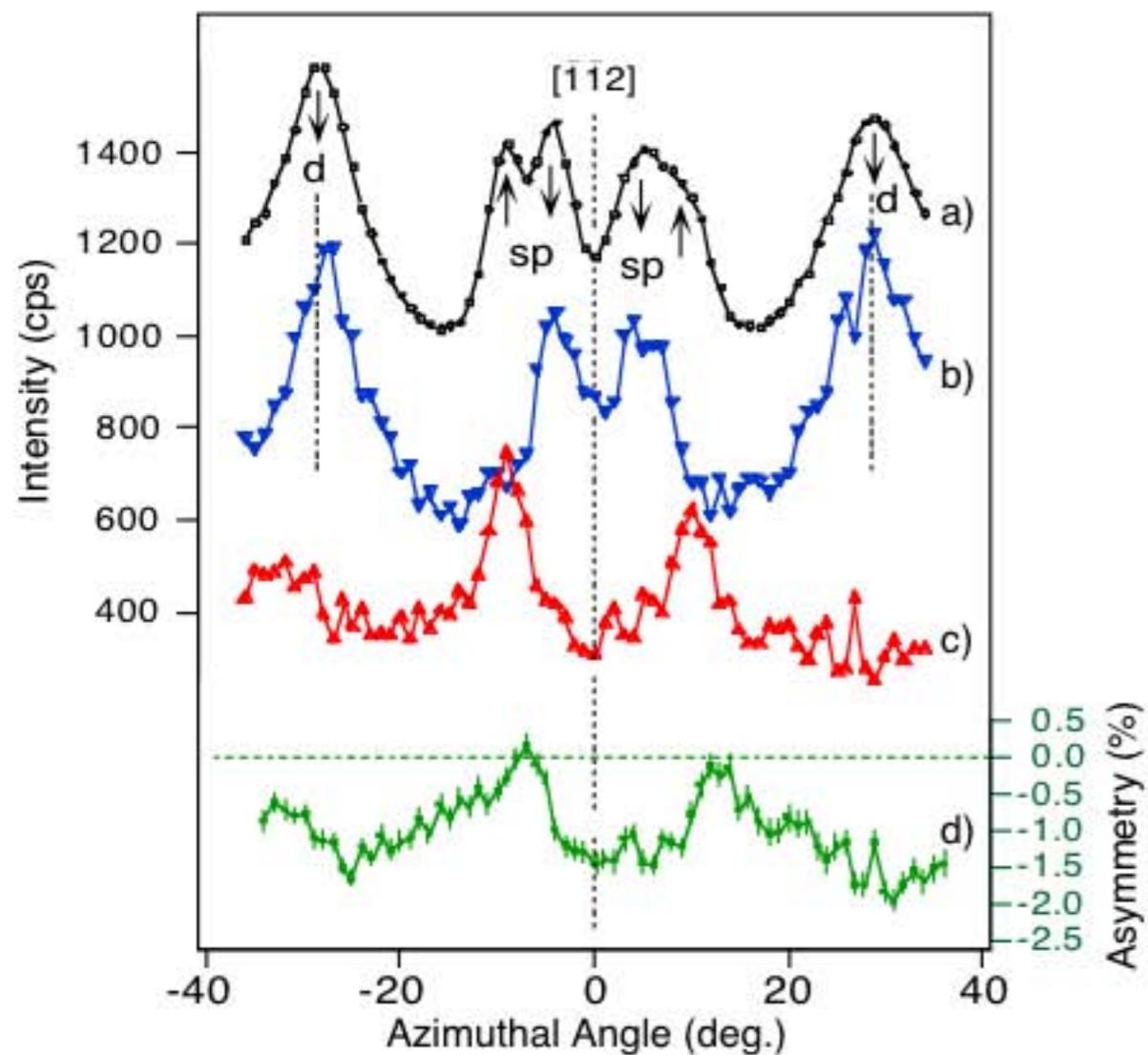
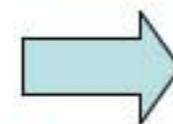


Spin-Resolved Momentum Mapping on Ni(111)



Quantitative Aspects of Spin Polarization Measurements

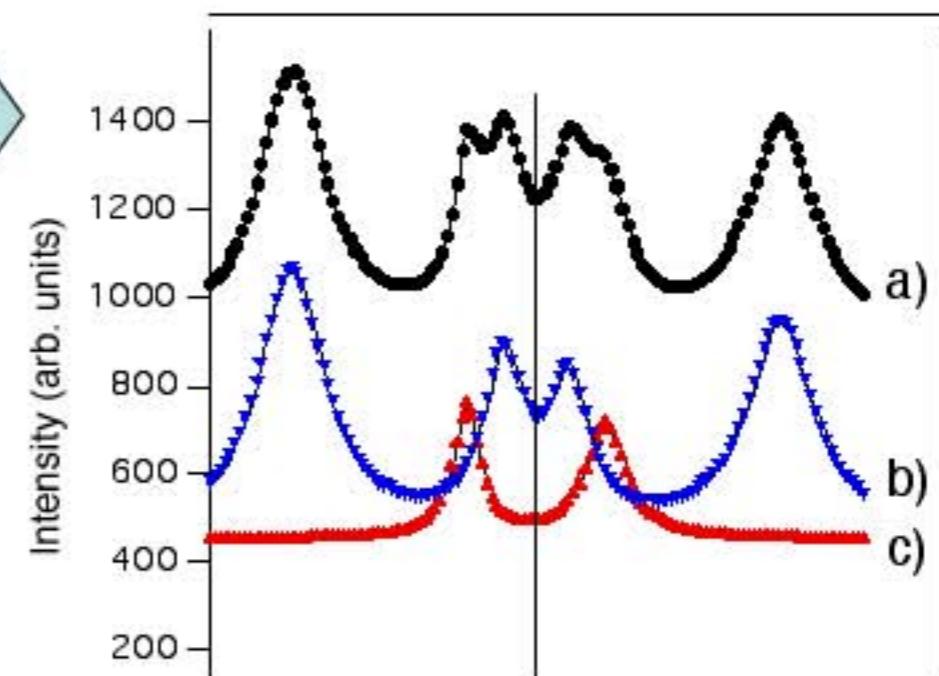
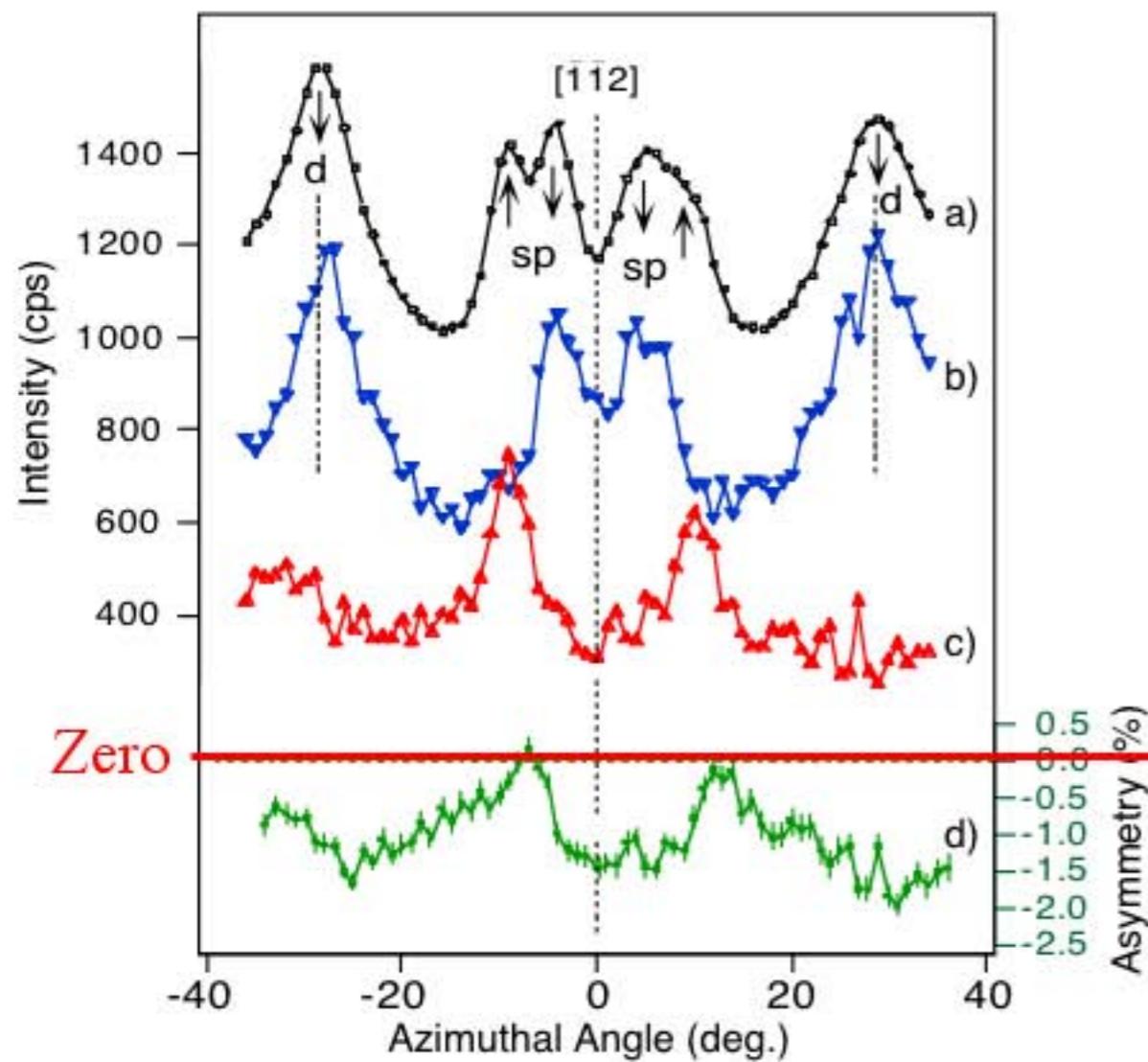
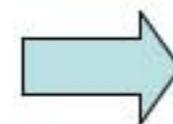
Modelling Spin Polarization:
6 Lorentzians (2 up, 4 down)
+ constant background ($S/B \sim 0.6$)



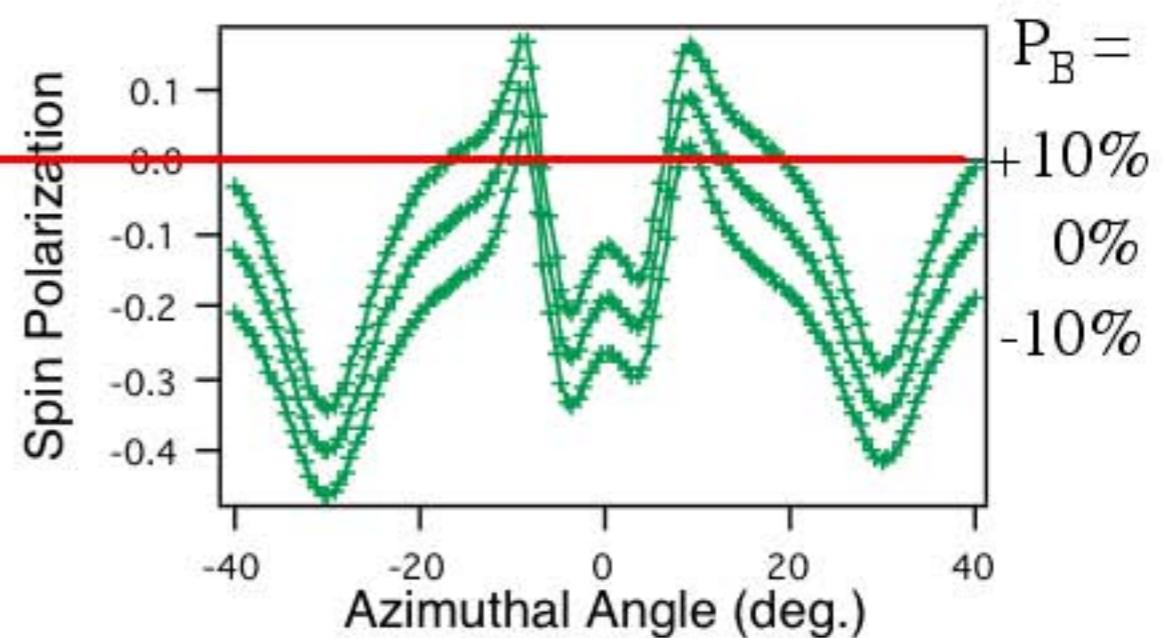
=> We can understand these asymmetry curves !

Quantitative Aspects of Spin Polarization Measurements

Modelling Spin Polarization:
6 Lorentzians (2 up, 4 down)
+ constant background ($S/B \sim 0.6$)



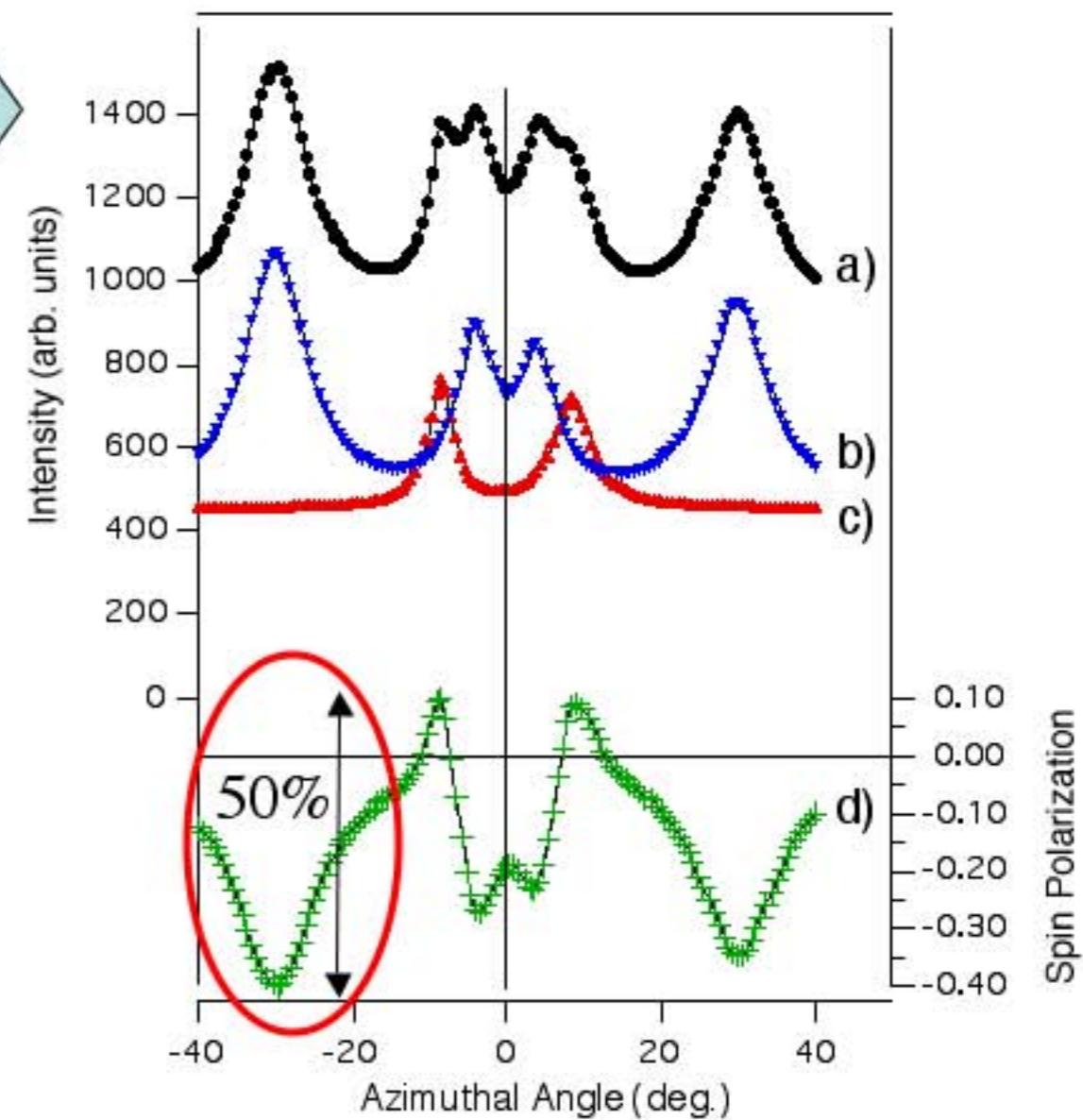
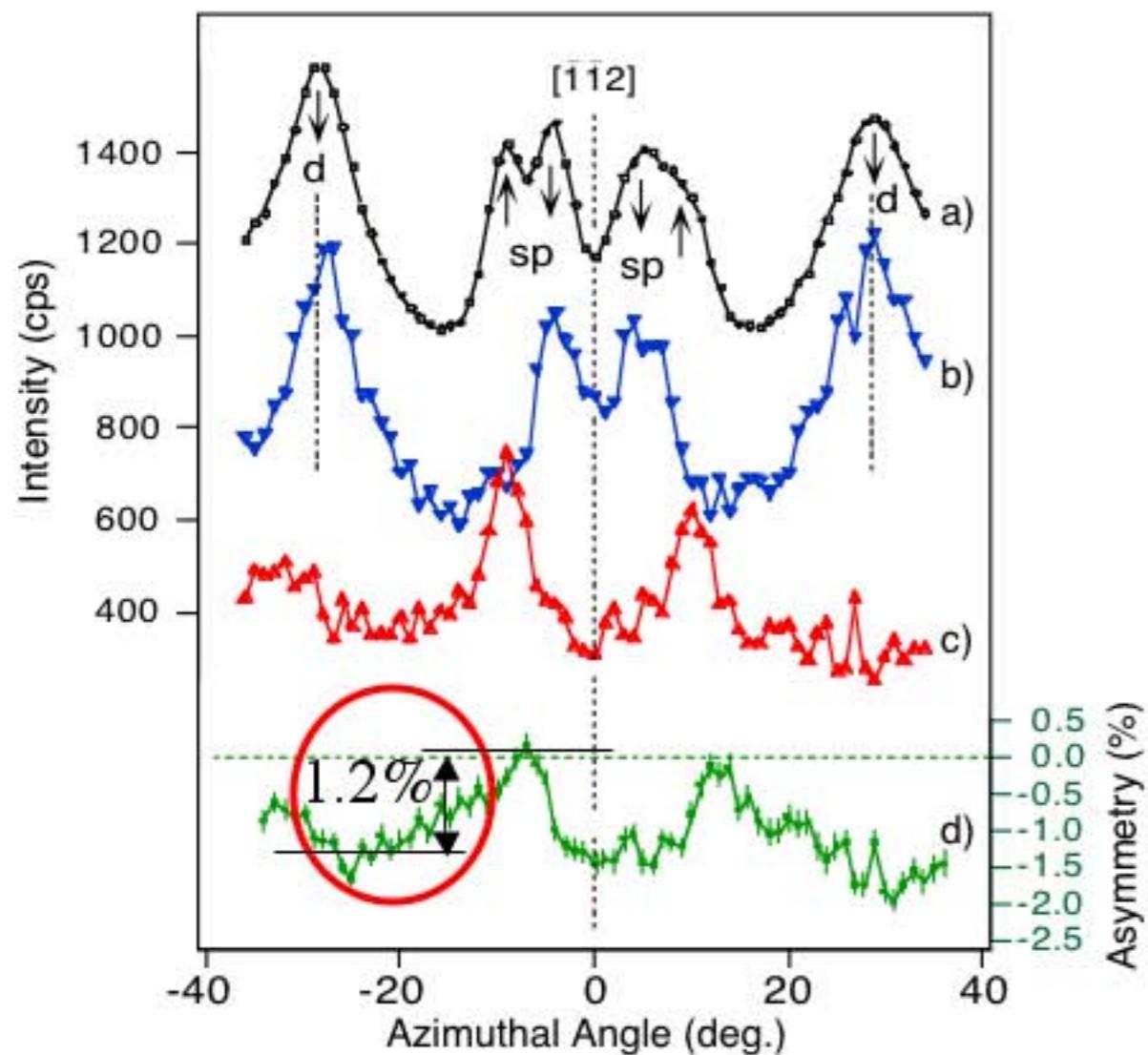
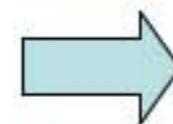
Vary polarization P_B of Background:



=> Background is ~ unpolarized !

Quantitative Aspects of Spin Polarization Measurements

Modelling Spin Polarization:
6 Lorentzians (2 up, 4 down)
+ constant background ($S/B \sim 0.6$)



Polarization $P = A / S$
 $S \sim 0.15$, but not precisely

Here: $S_{\text{eff}} = A / P \sim 0.024 \rightarrow$ Sample poorly magnetized

$M / M_S \sim 0.15$

Monitoring the Magnetic Phase Transition

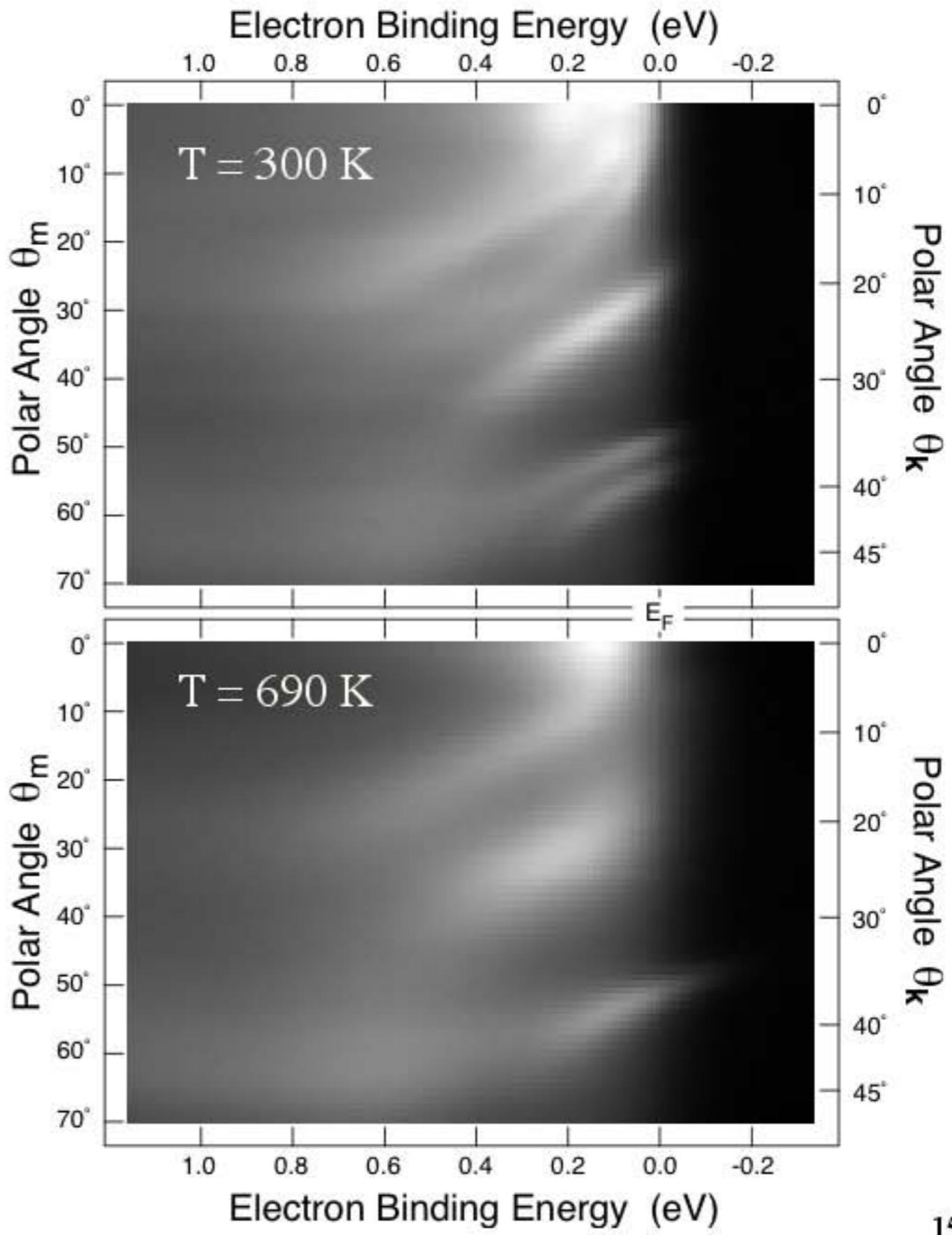
$$T_c = 631 \text{ K}$$

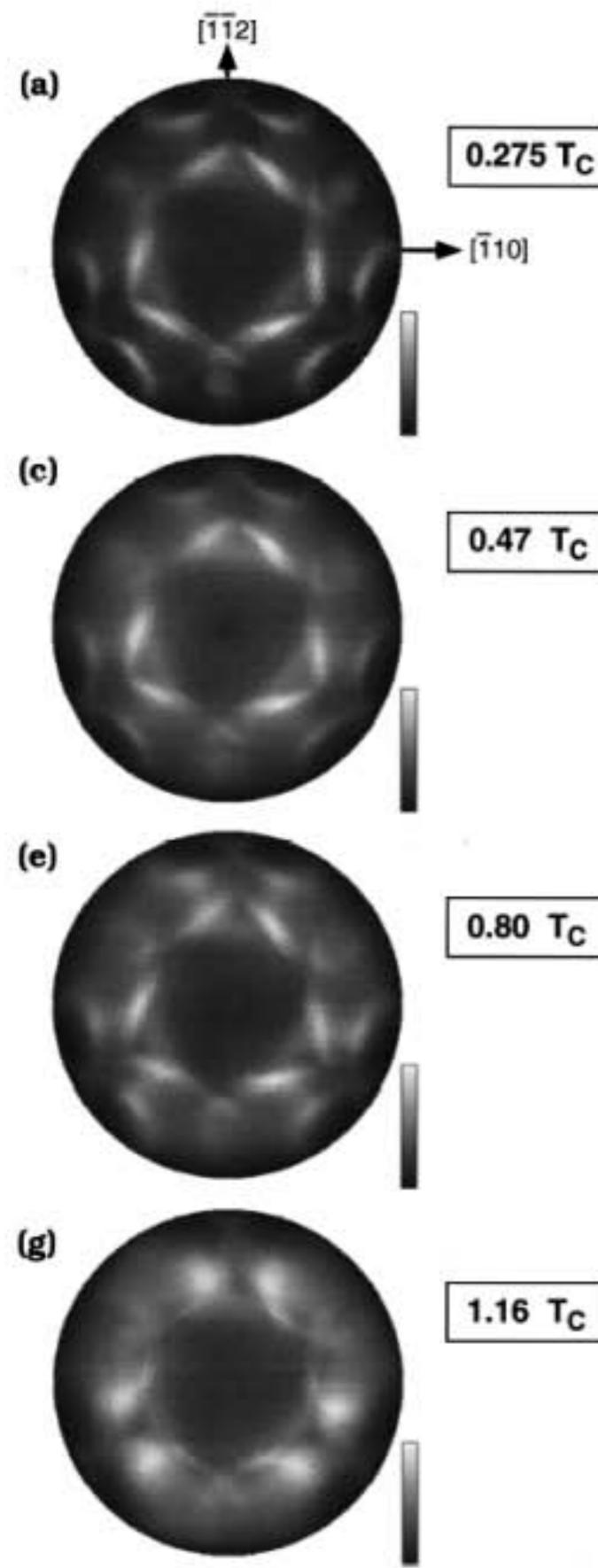
Question:

How does the band structure change when nickel goes from the ferromagnetic to the paramagnetic state?

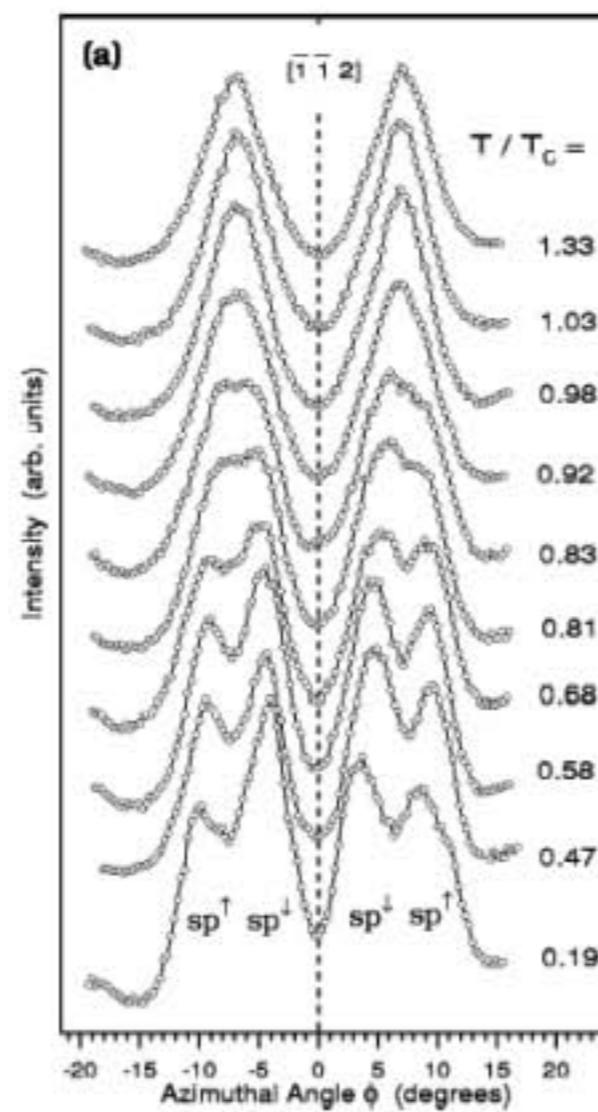
Answer:

Magnetic exchange splittings disappear

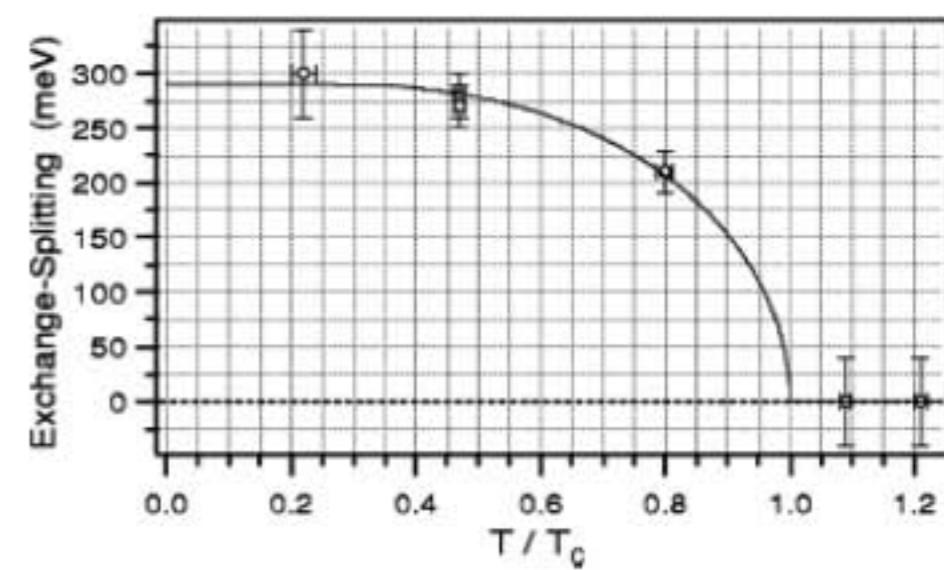




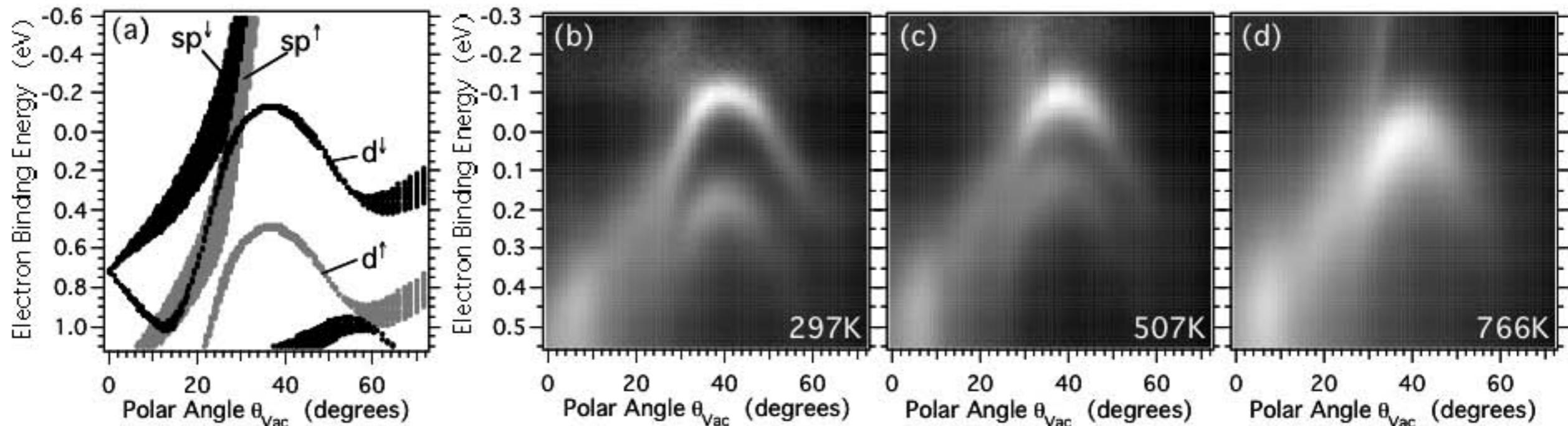
Evolution of the Fermi Surface During the Phase Transition



... and quantifying the exchange splitting:

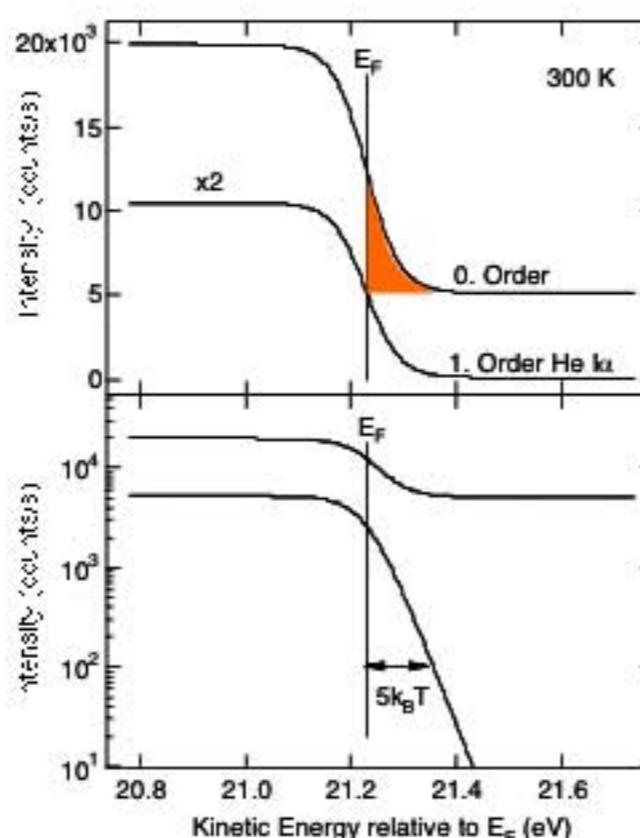


Observing the Magnetic Exchange Splitting in the d Band

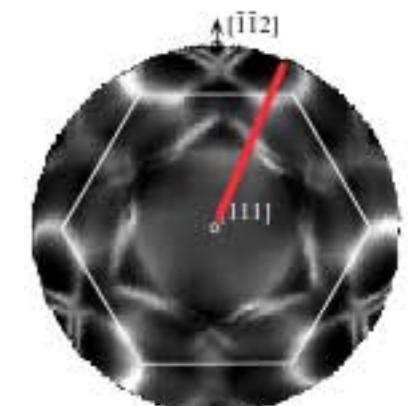


Photoemission
above the
Fermi Level ?

T. Greber et al., PRL 79, 4465 (1997)



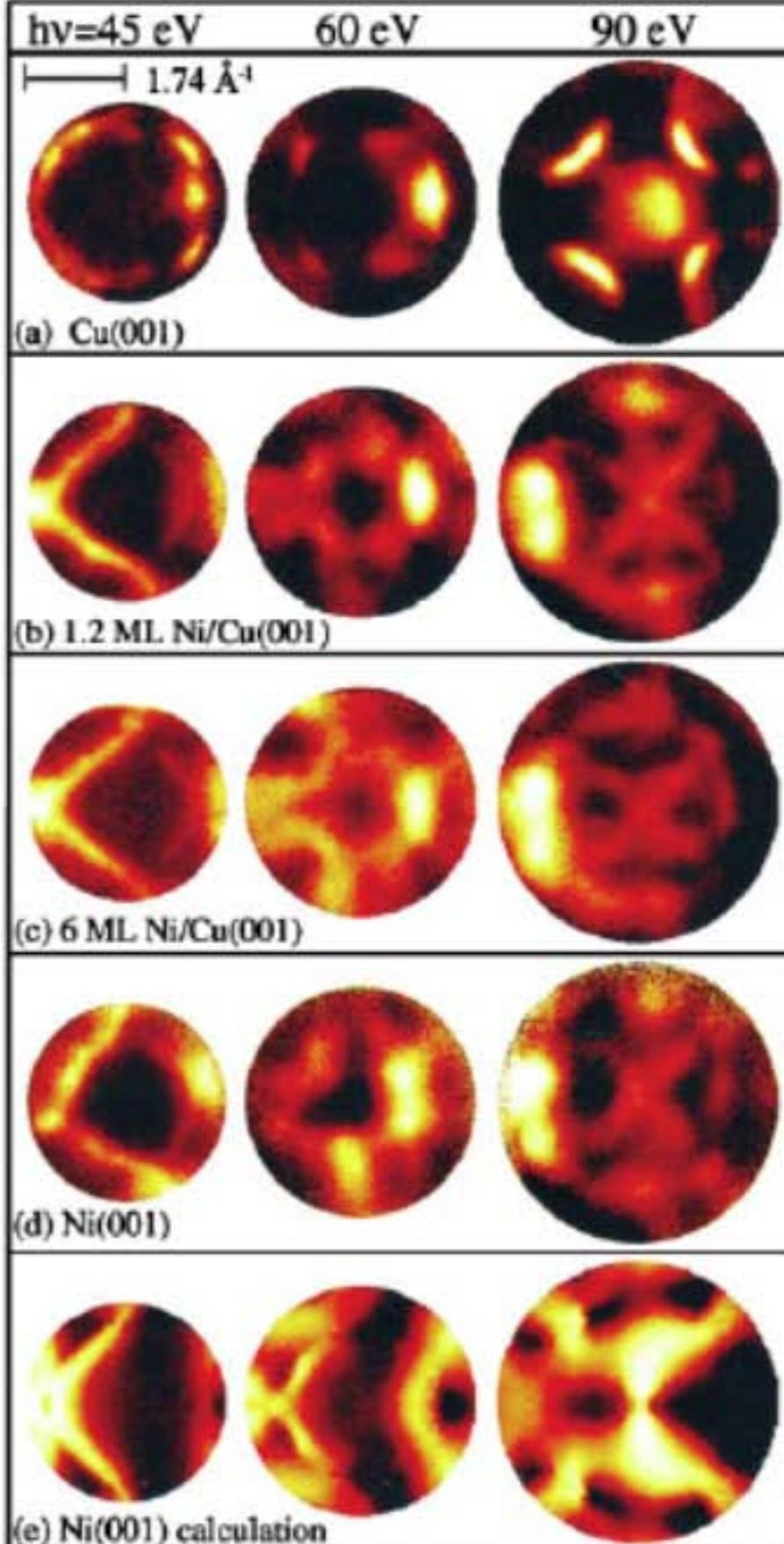
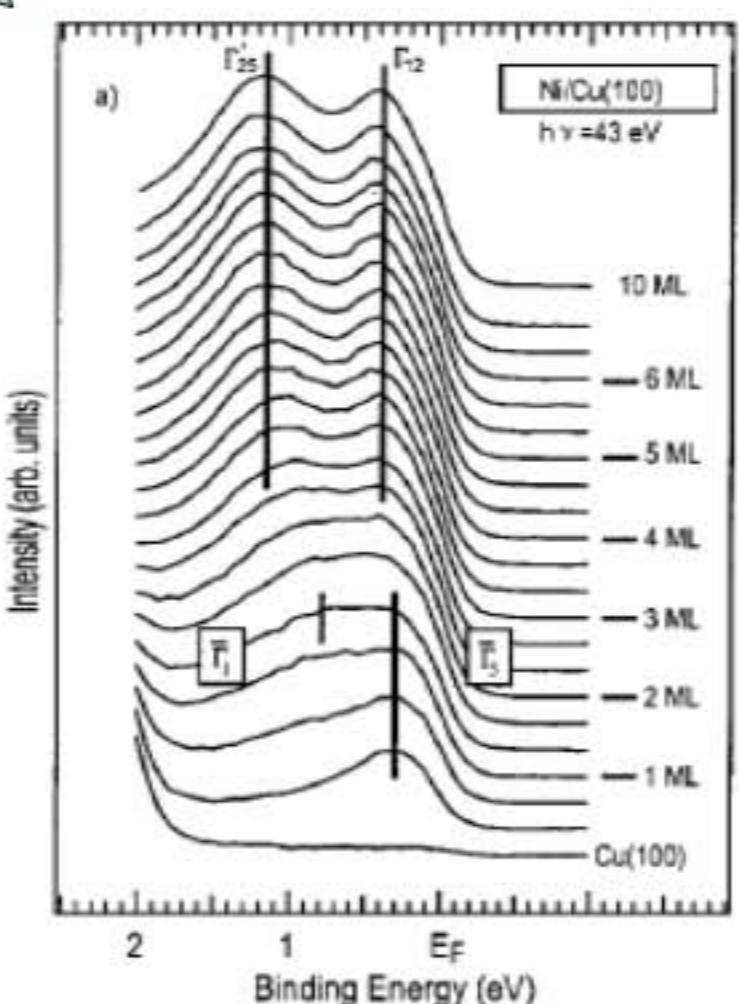
Polar Scan



Ultrathin Films of Ni on Cu(001)

- grows epitaxially
- well ordered films, some roughness
- T_c depends on film thickness
- bulk-like Fermi surface already at 1 ML
- clear changes of the band width below 2.5 ML

?



G. J. Mankey et al., PRL 78, 1146 (1997).
18

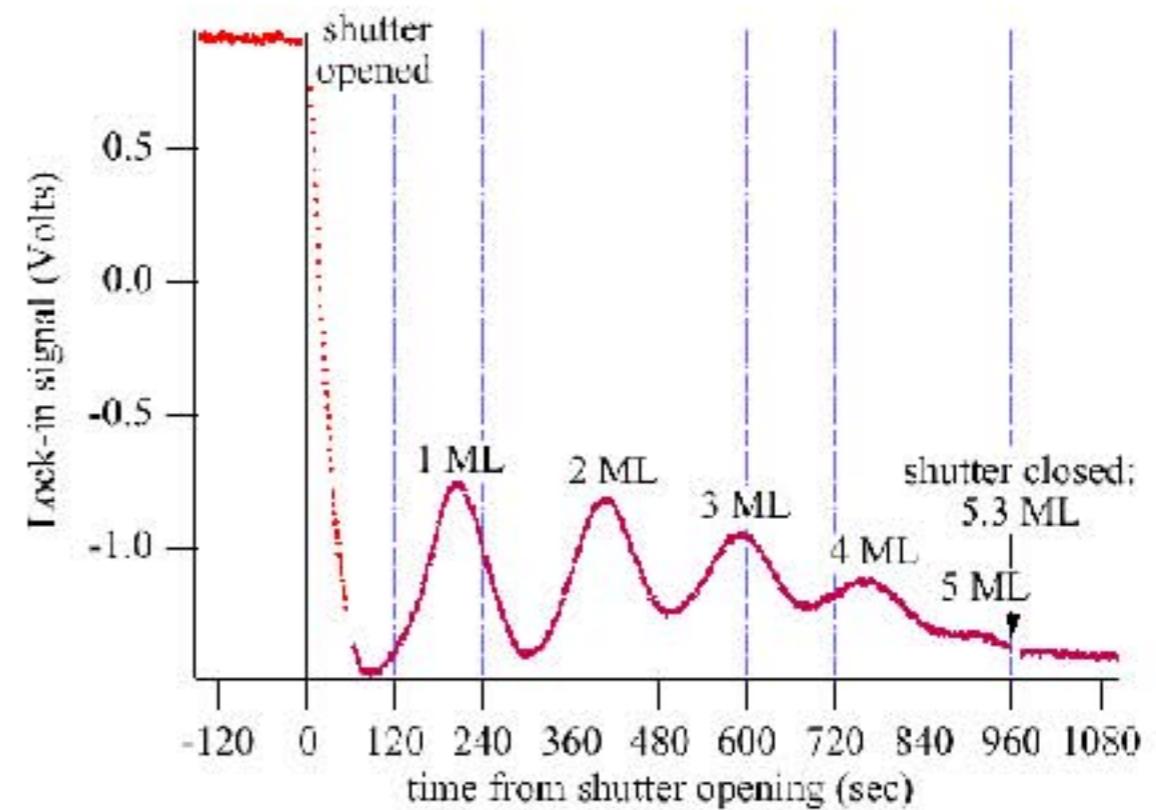
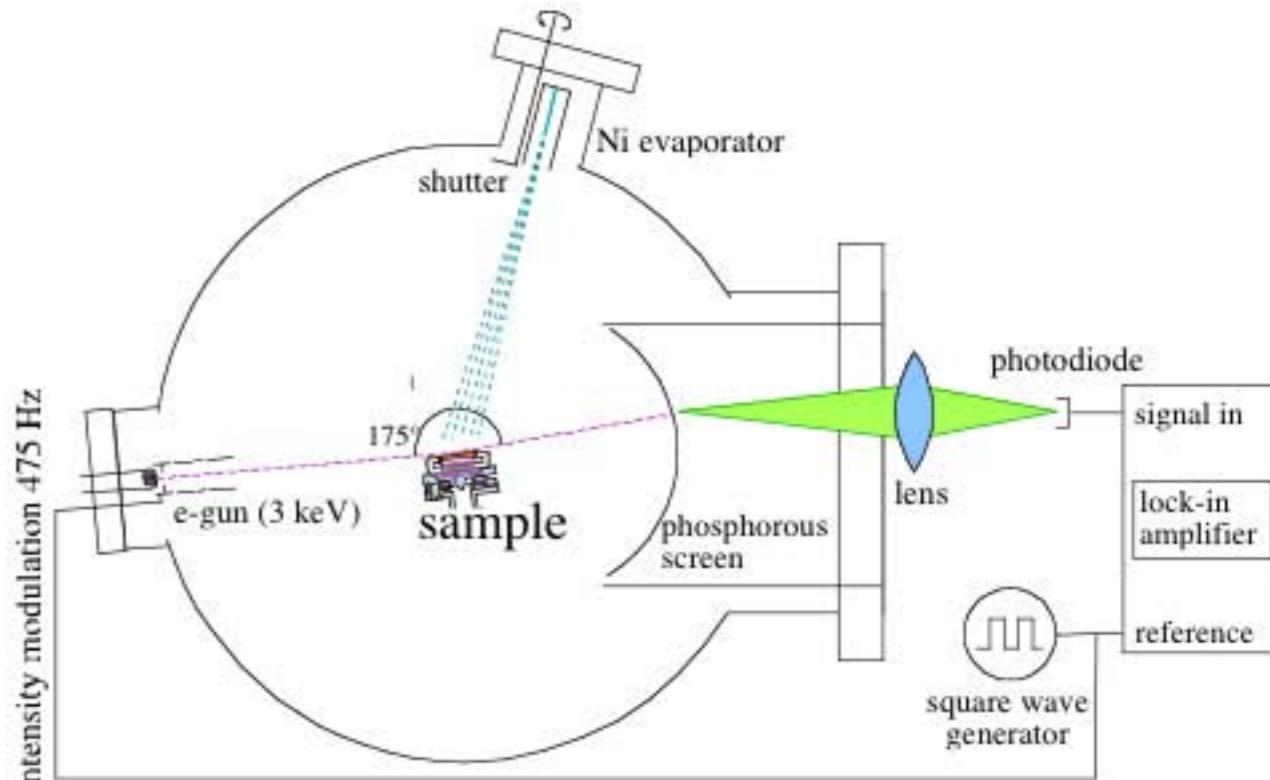
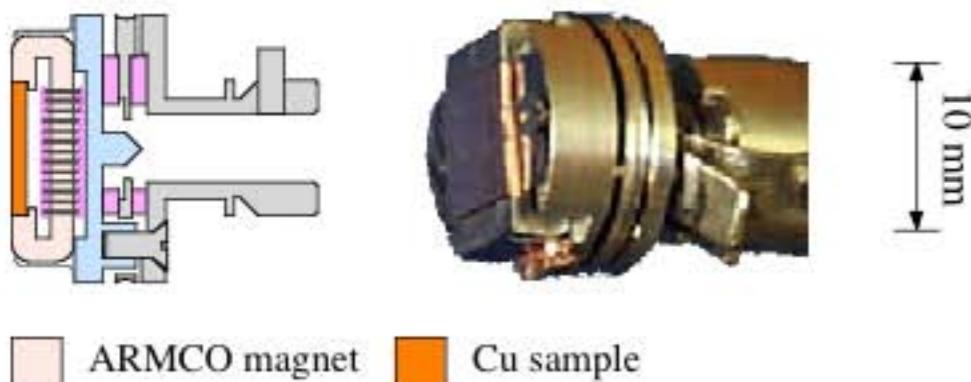
C. Pampuch et al.,
PRB 63, 153409 (2001).

Preparation of Monolayer Films of Ni on Cu(001)

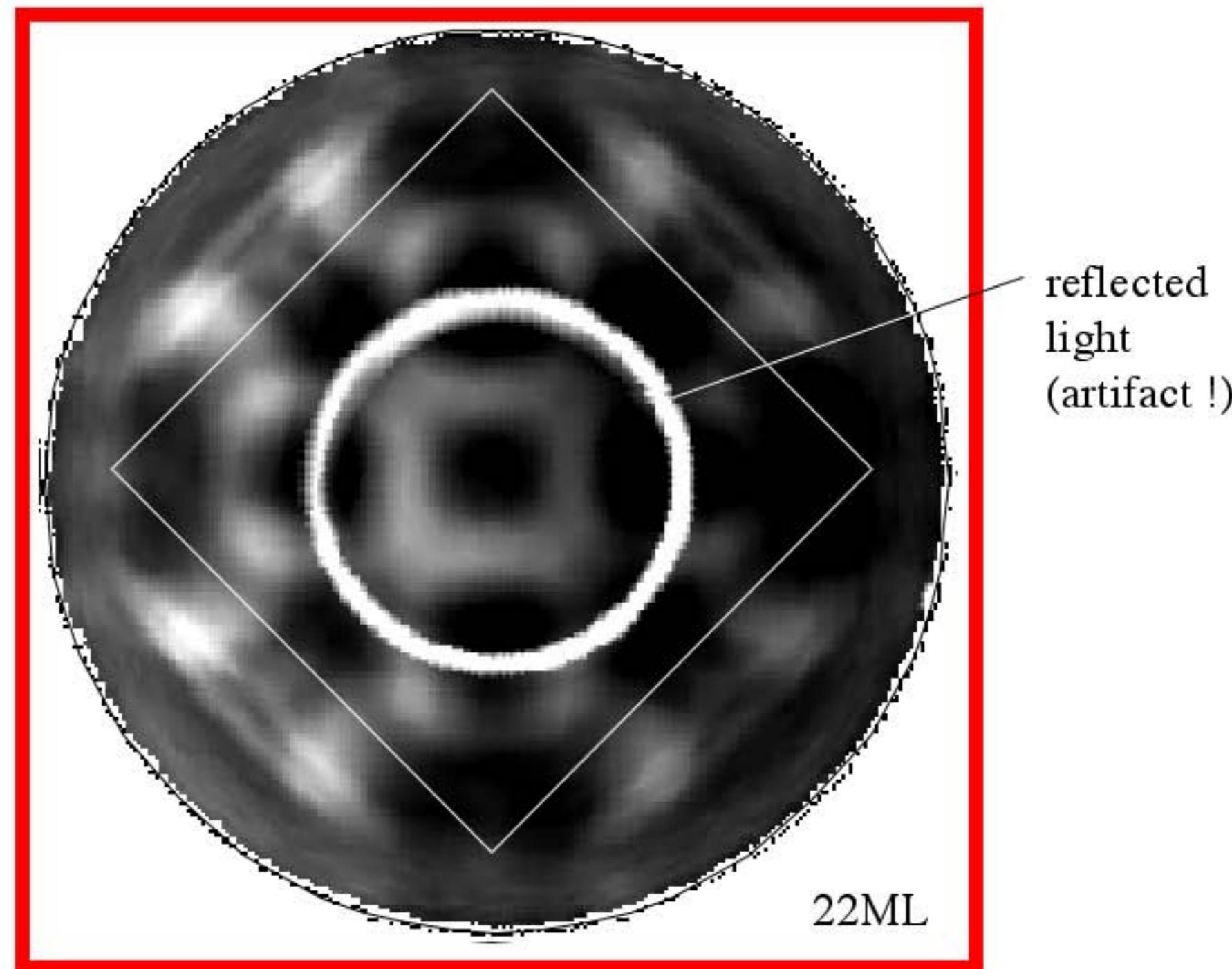
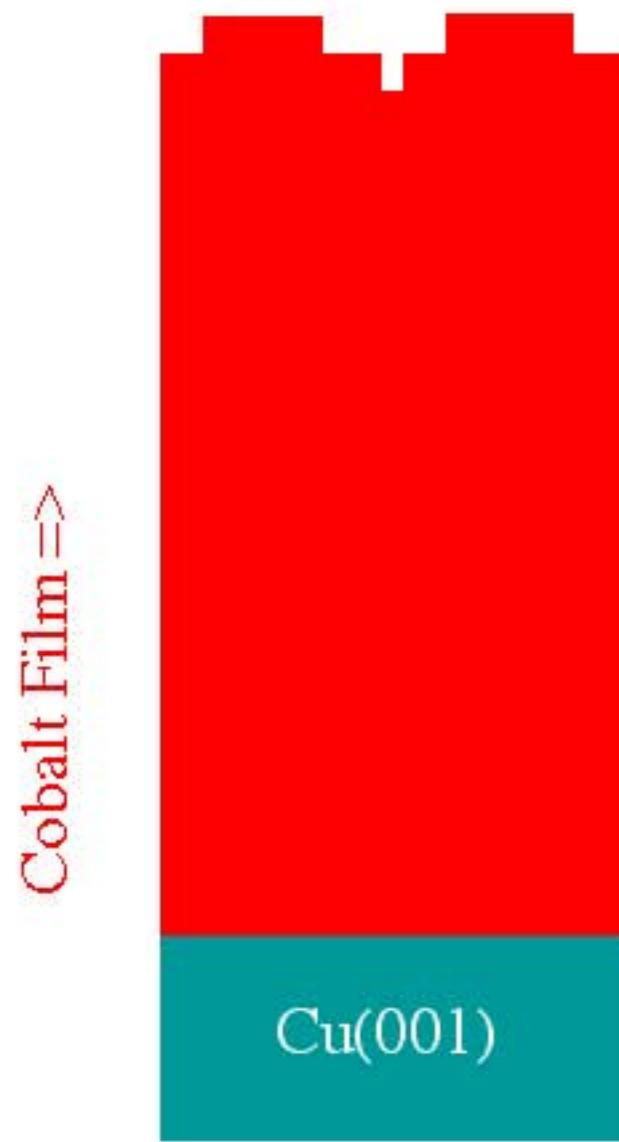
Ni deposited from e-beam heated rod at 0.33 ML/min onto clean Cu(001)

Annealing to 420 K.

Cooling to 150 K (liquid N₂).

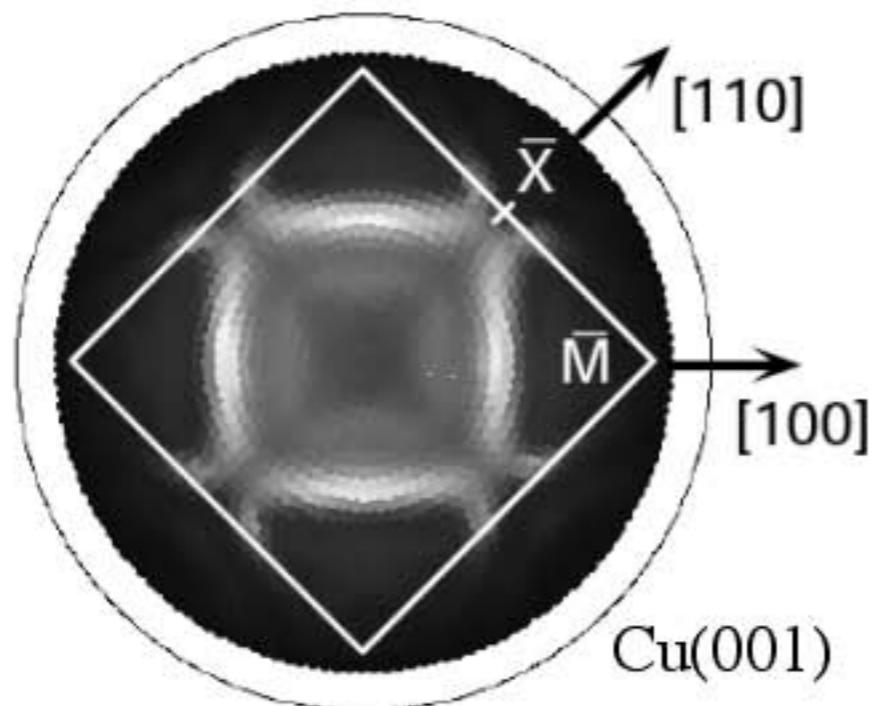


Evolution of the Fermi Surface with Film Thickness



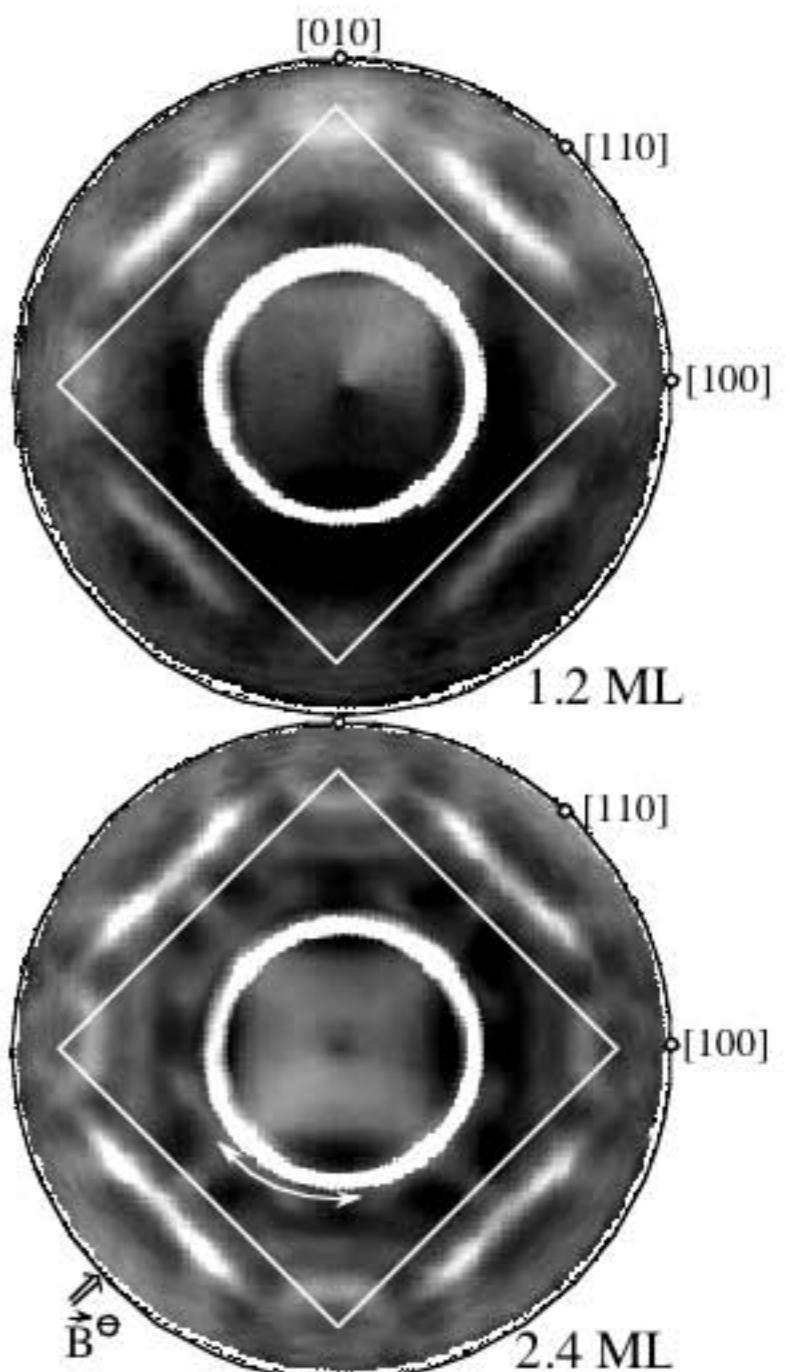
Contribution from the Cu(001) Fermi surface ...

Substrate Fermi surface:

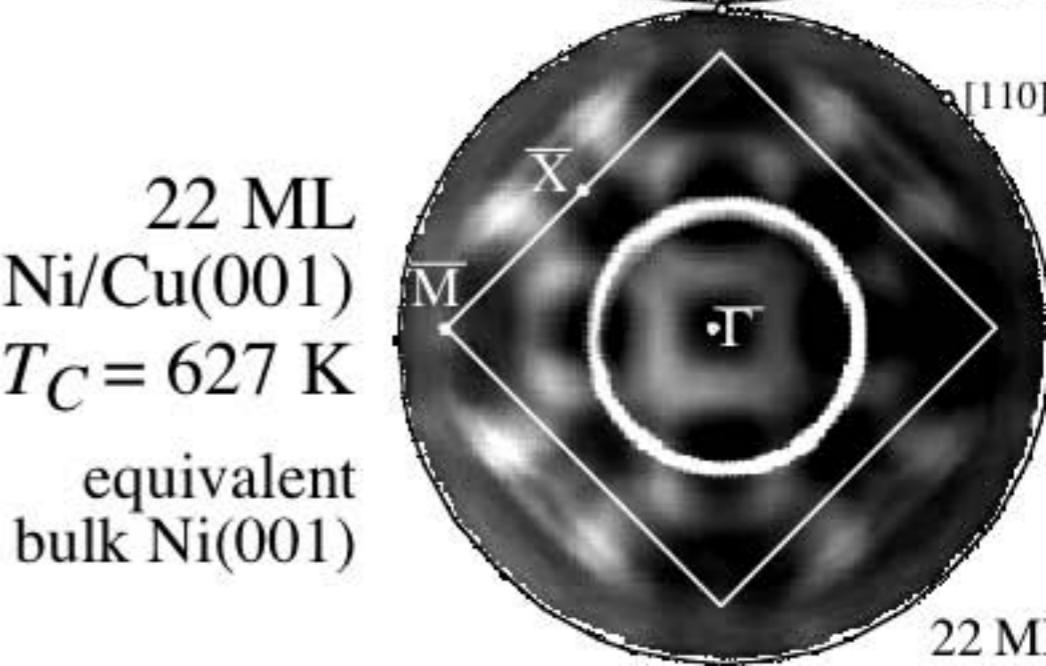
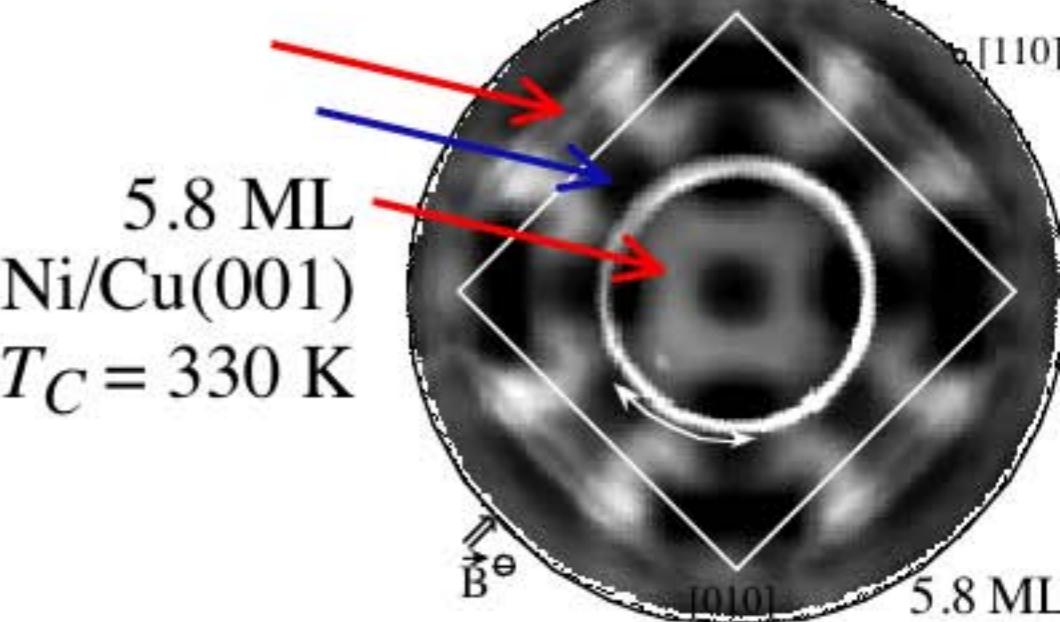
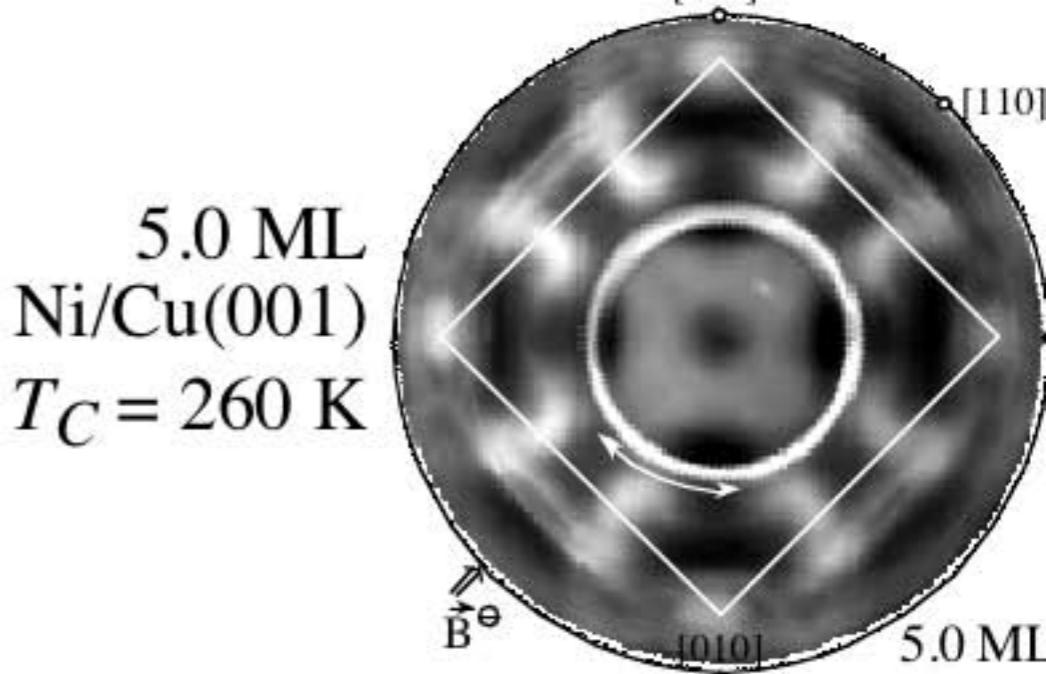
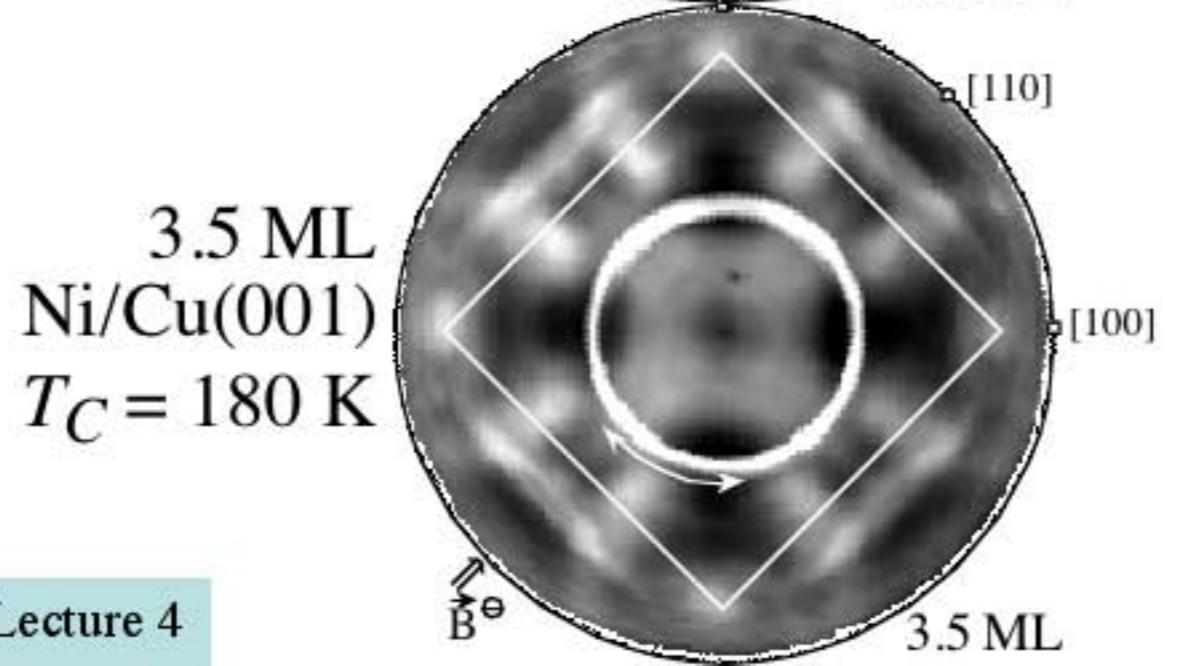
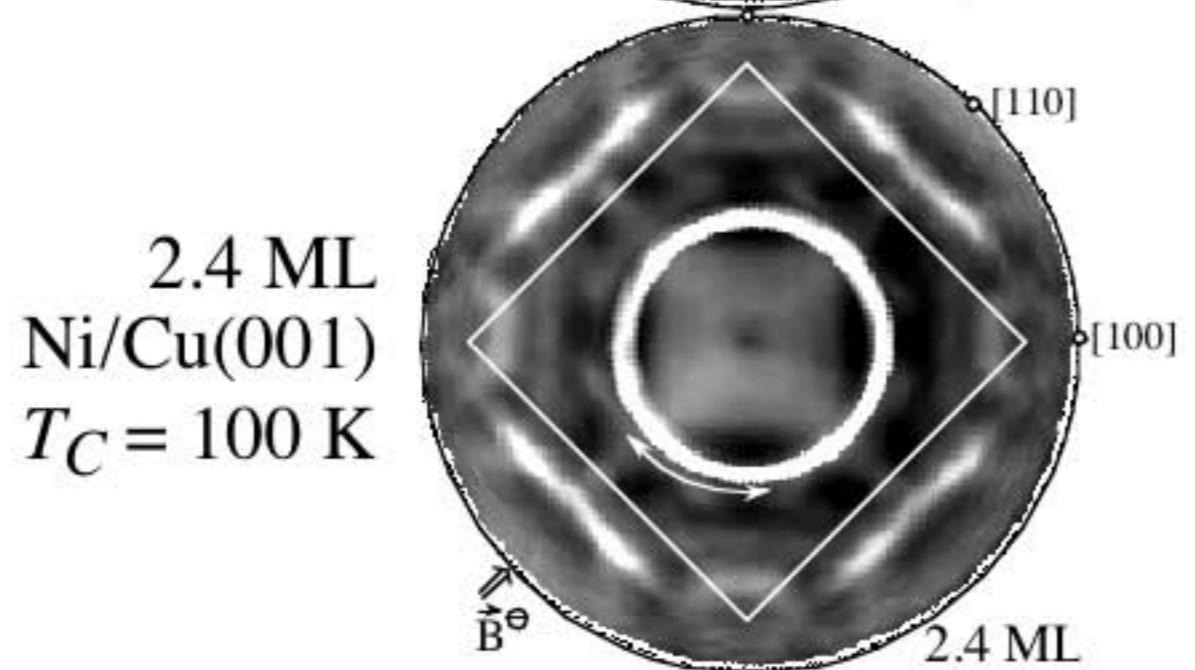
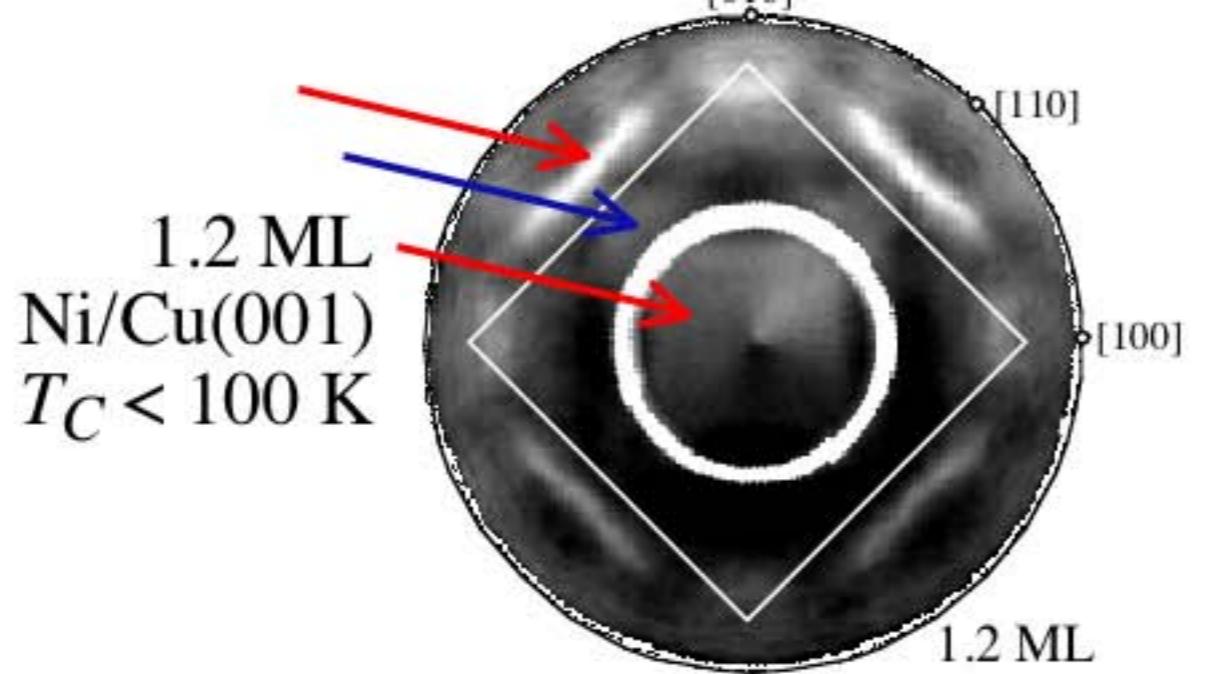


1.2 ML
Ni/Cu(001)
 $T_C < 100$ K

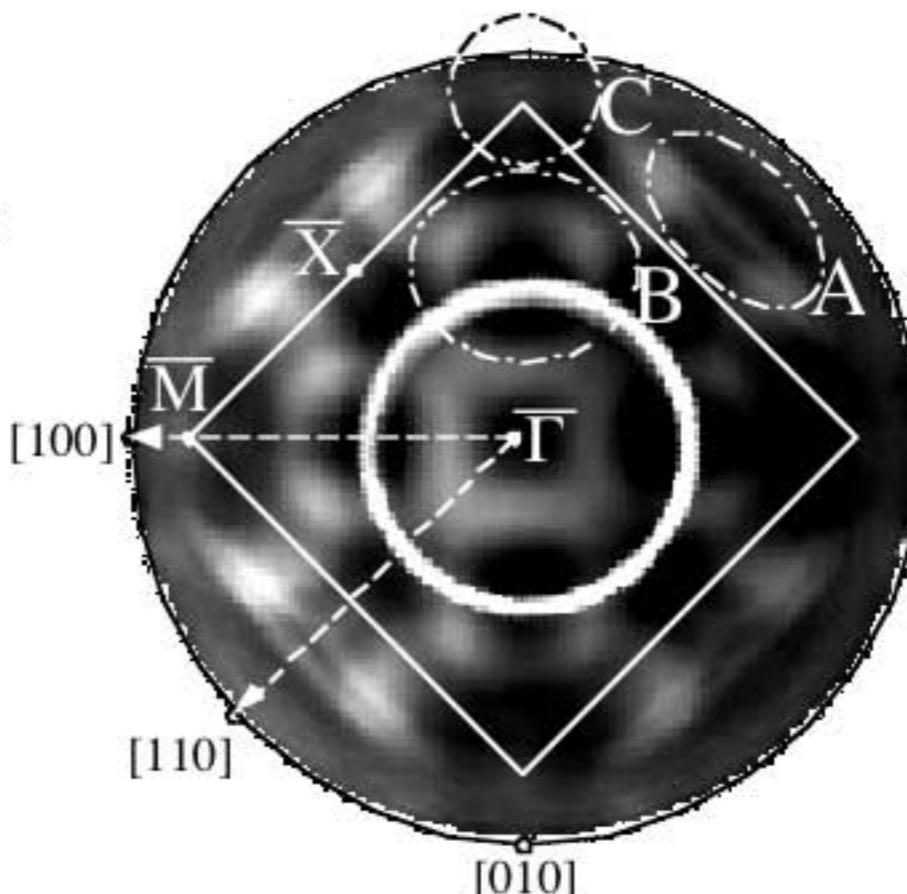
2.4 ML
Ni/Cu(001)
 $T_C = 100$ K



... is quenched from the first monolayer nickel !
(high cross section, strong inelastic scattering, ...)

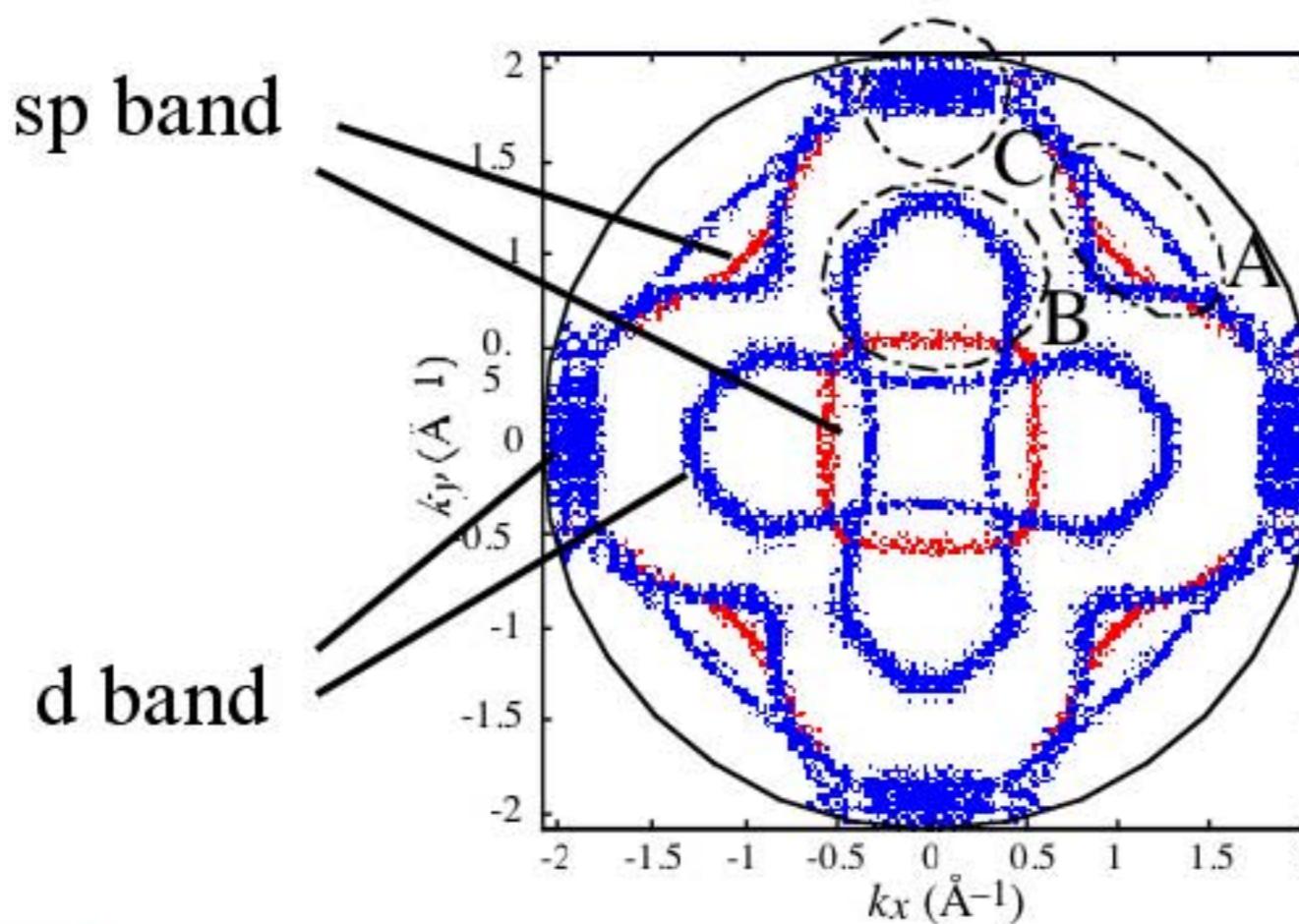


Band Character of Fermi Surface Contours (thick film \sim bulk-like)



data from
22 ML
Ni/Cu(001)

$h\nu = 21.22$ eV
equivalent
bulk Ni(001)



Fermi surface
calculation

free electron
final state
approximation

$h\nu = 21.22$ eV

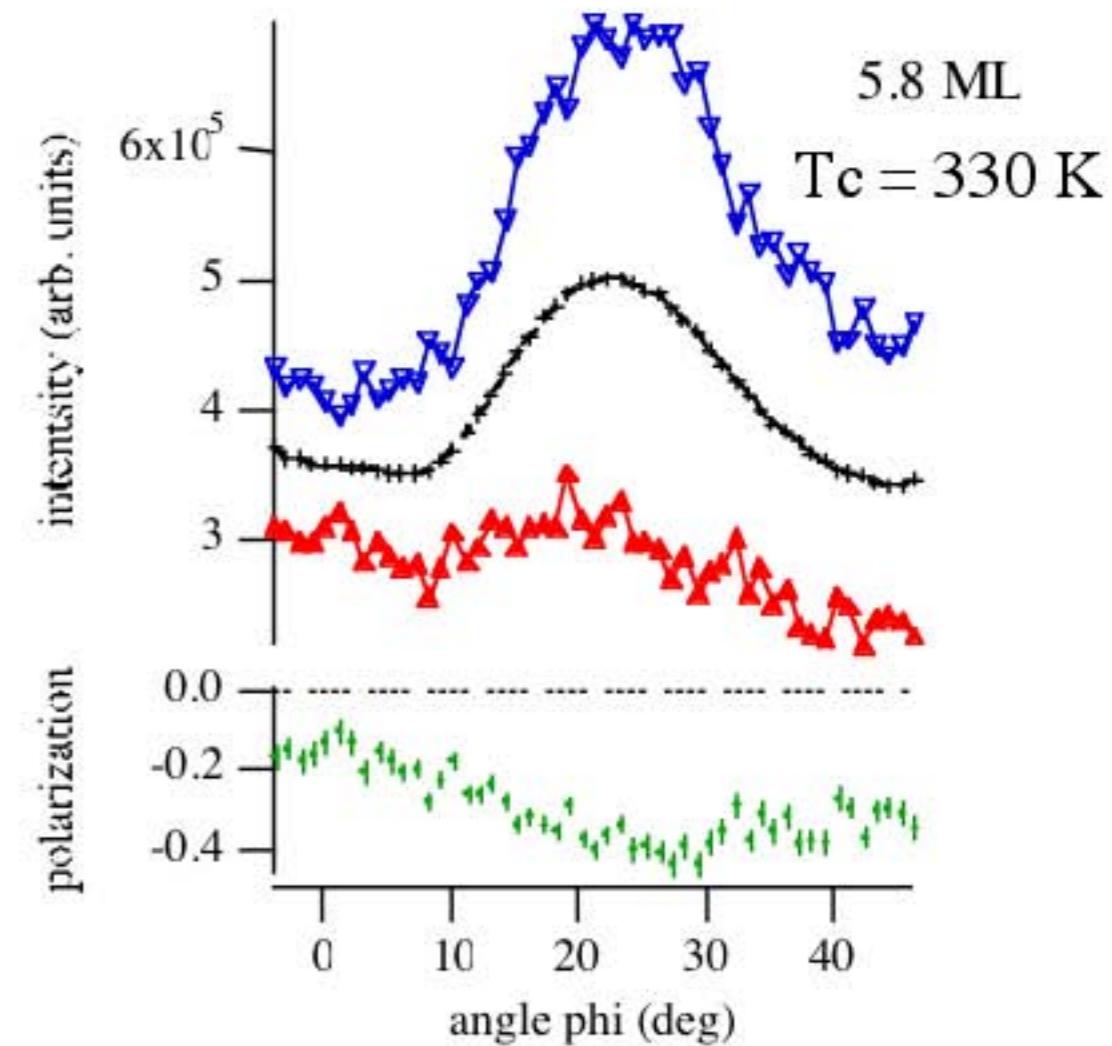
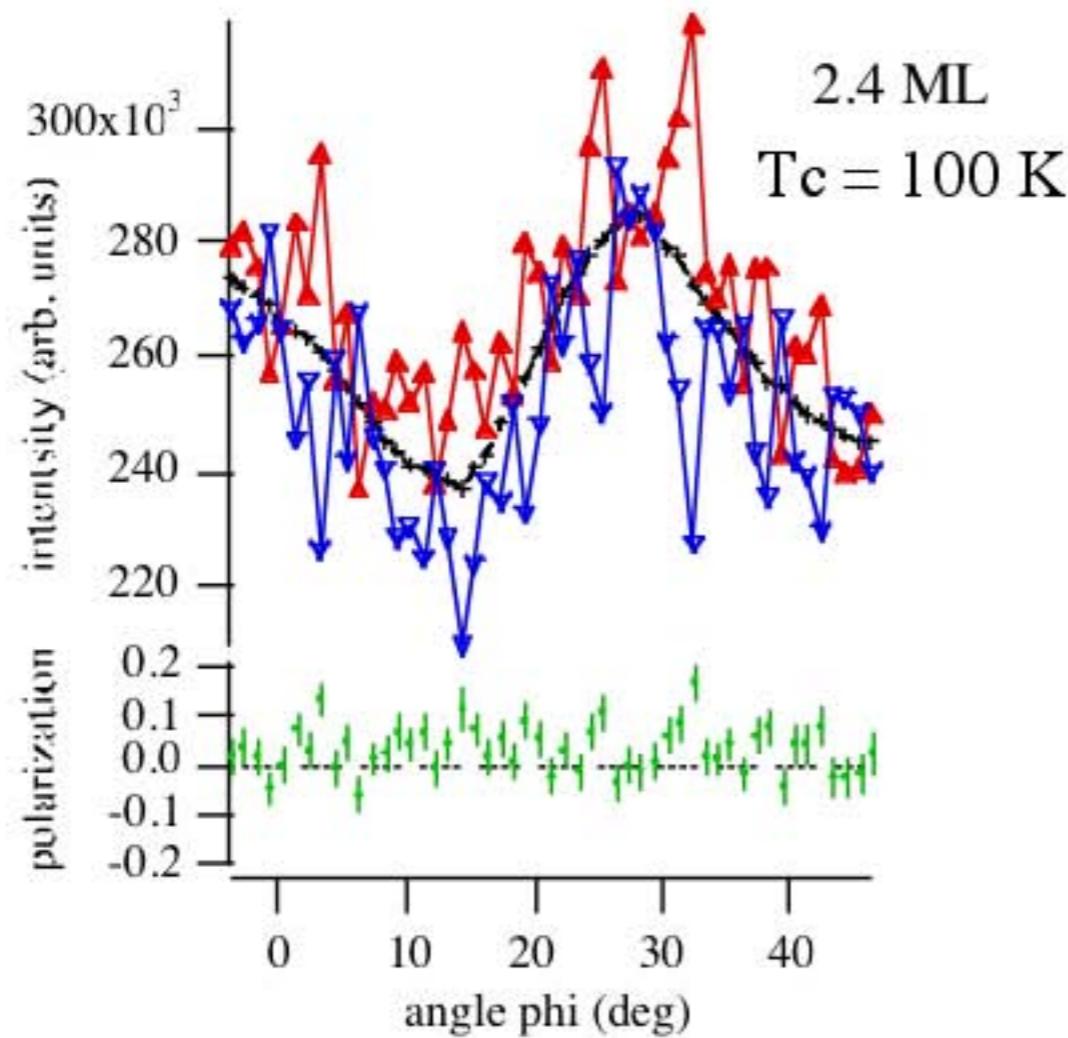
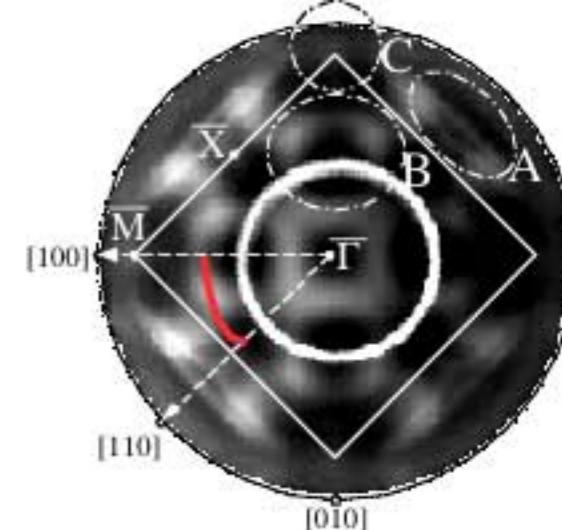
$\Phi = 5.2$ eV

$V_0 = 10.2$ eV

23

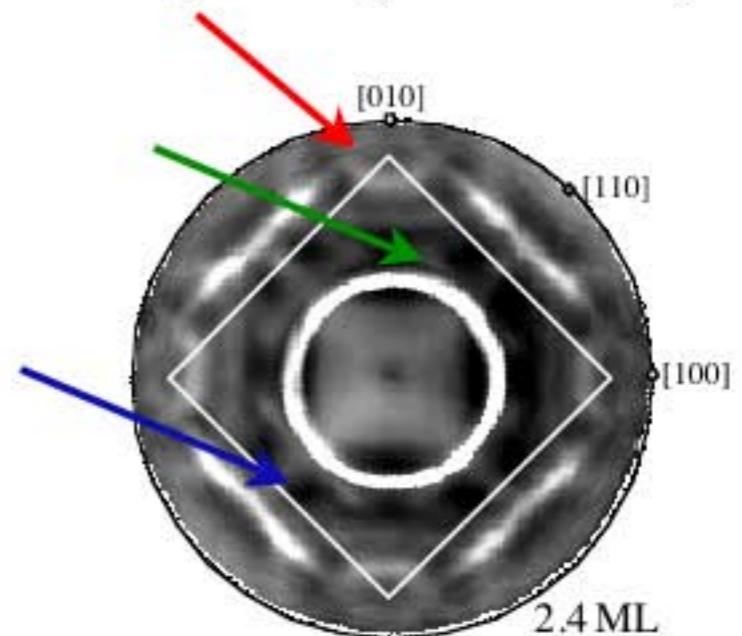
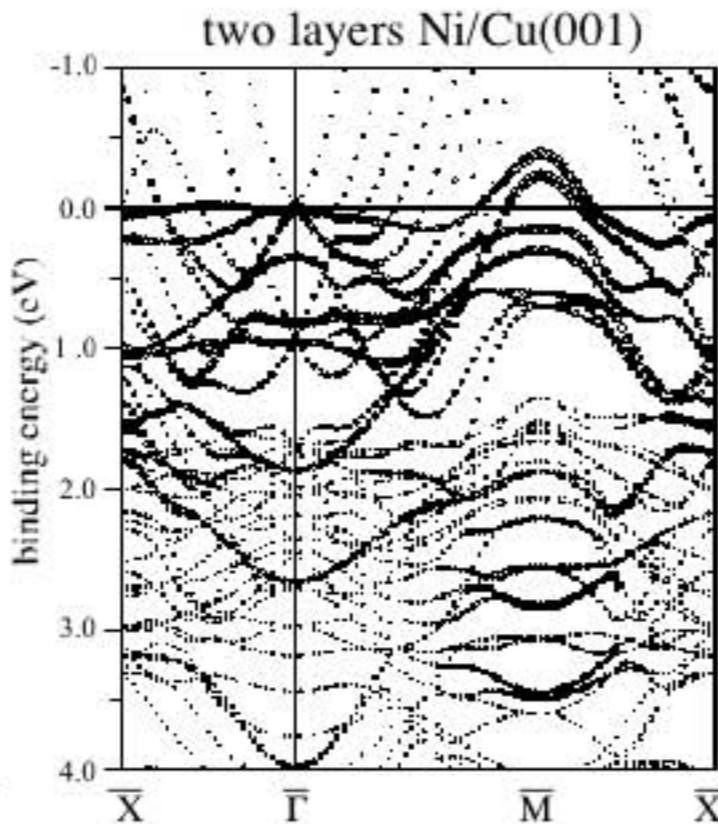
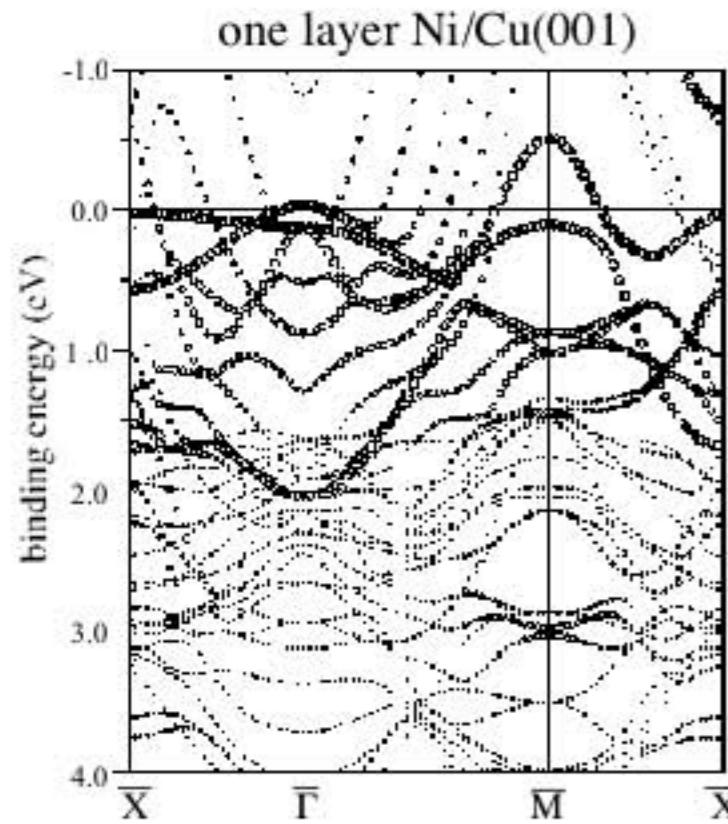
Use Spin Polarization for Band Identification

Azimuthal Momentum Scans at $\theta = 28^\circ$
($h\nu = 21.22$ eV)
 $T = 150$ K

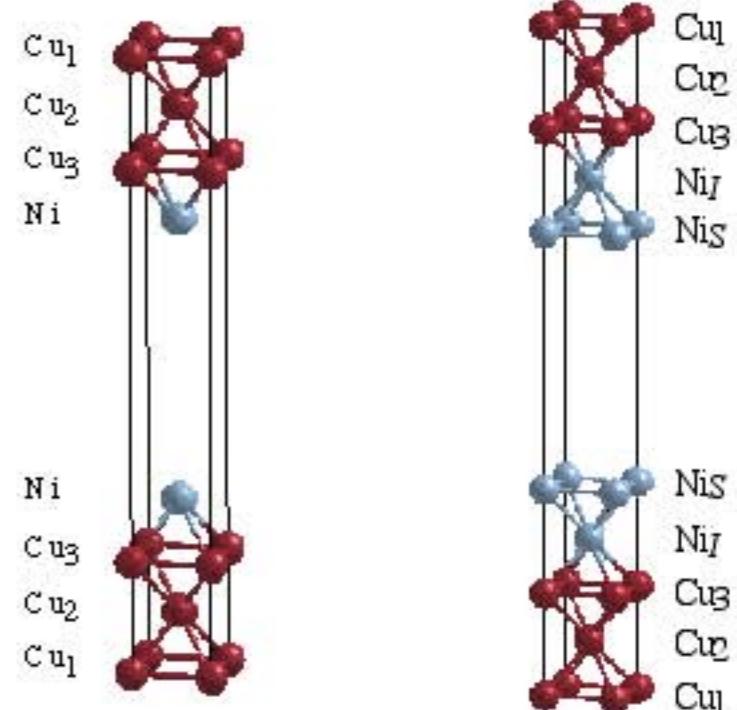
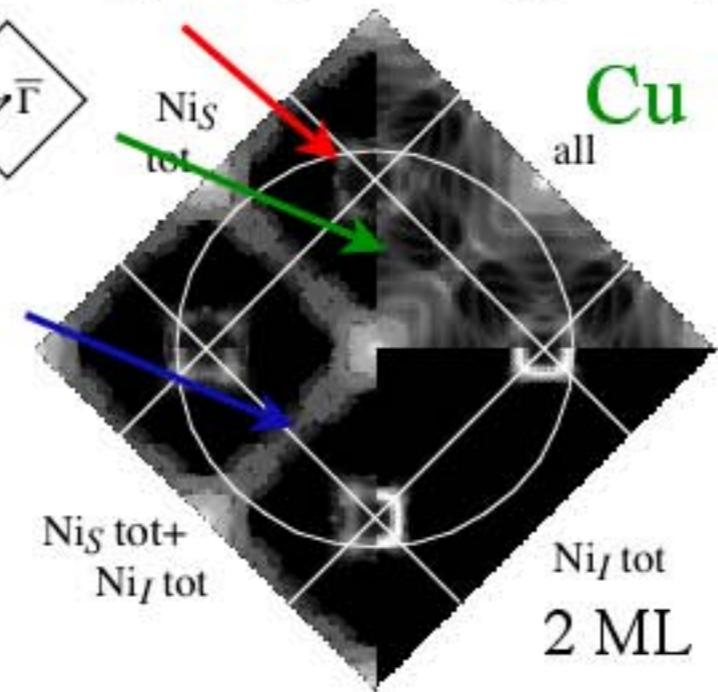
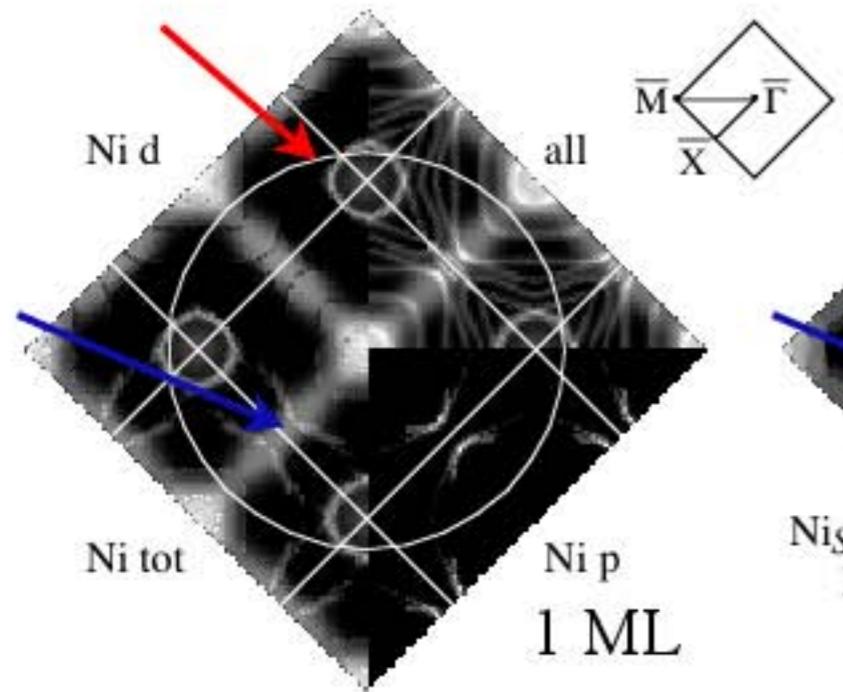


The ferromagnetic film (5.8 ML) shows strong minority spin polarization => **minority d band**

Ni/Cu(001) Slab Calculations for Monolayer Regime (Wien97)



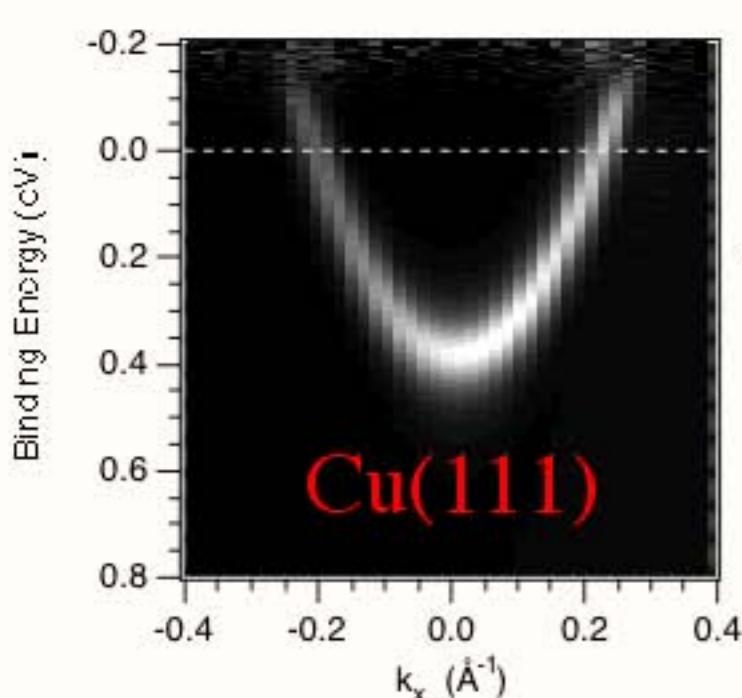
data from
2.4 ML Ni/Cu(001)
 $h\nu = 21.22$ eV



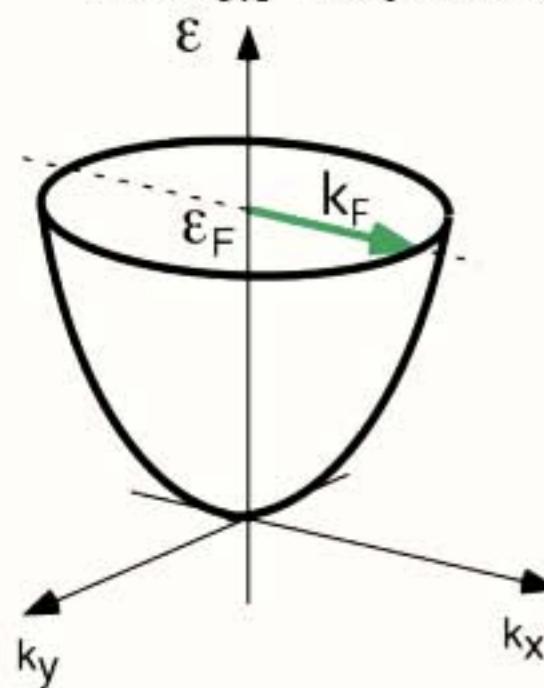
Ultrathin films of Ni on Cu(001)

- < 3 ML: paramagnetic interface band structure
- Around 3 ML two things happen:
 - $T_c > T_{\text{meas}} \Rightarrow$ Exchange splitting appears
 - Band structure becomes bulk-like (3D)
- The sp bands are bulk-like already at ~ 1 ML
- The d bands form interface states for < 3 ML

The Shockley Surface State on Noble Metal (111) Surfaces



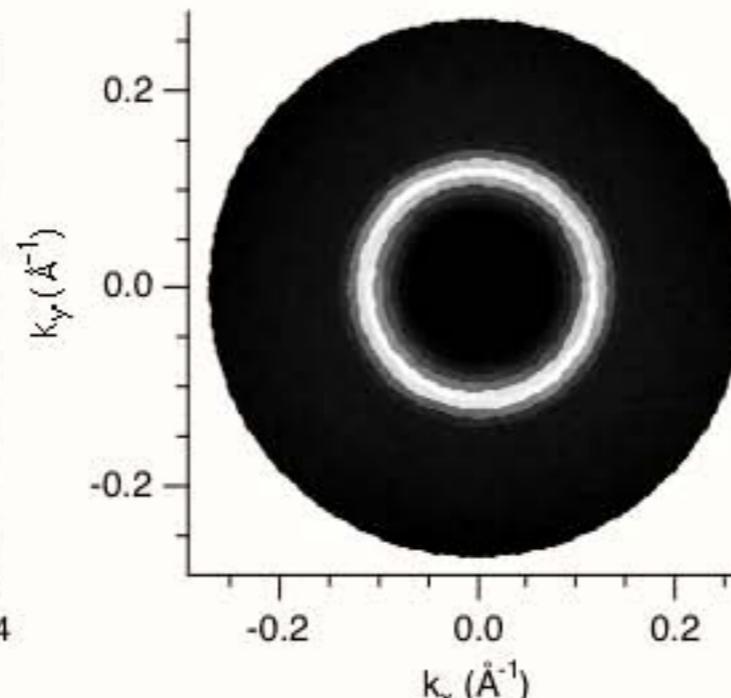
Energy dispersion



2D free electron gas

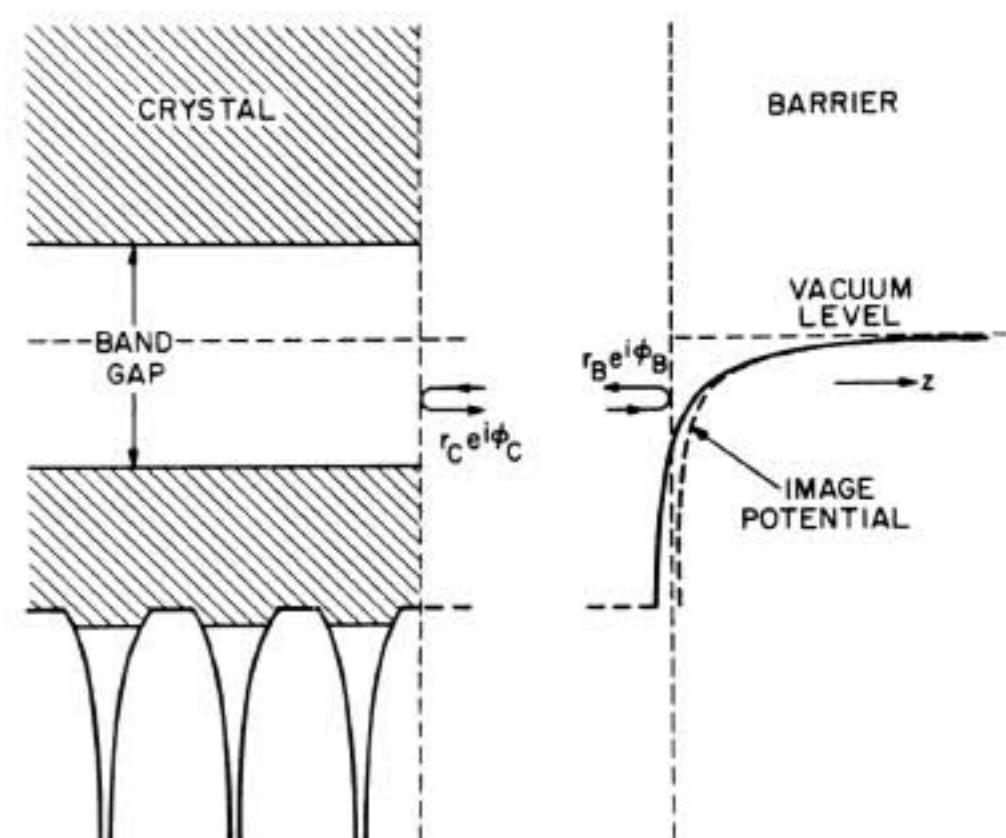
binding energy:	390 meV
effective mass:	$0.42 m_e$
Fermi wave length:	2.9 nm
coherence length:	$\sim 10 \text{ nm (300K)}$

Courtesy F. Baumberger



Fermi surface

Multiple reflection model
for surface states:



(N.V. Smith, PRB **32**, 3549 (1985))

What does this have to do
with spin and magnetism ?

The Spin-Orbit Split Surface State on Au(111)

VOLUME 77, NUMBER 16

PHYSICAL REVIEW LETTERS

14 OCTOBER 1996

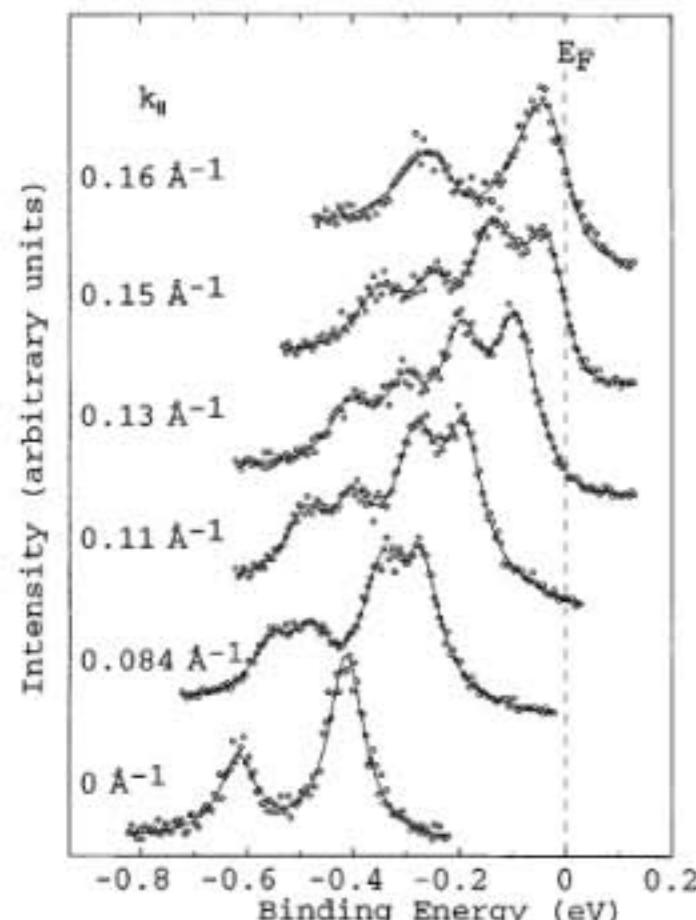
PHYSICAL REVIEW B, VOLUME 63, 115415

Spin Splitting of an Au(111) Surface State Band Observed with Angle Resolved Photoelectron Spectroscopy

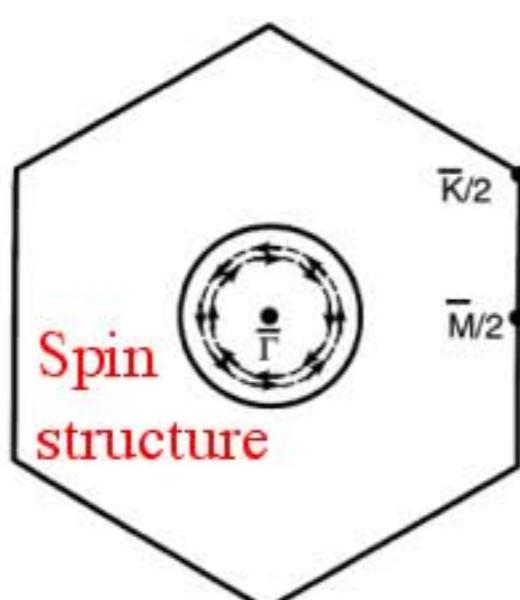
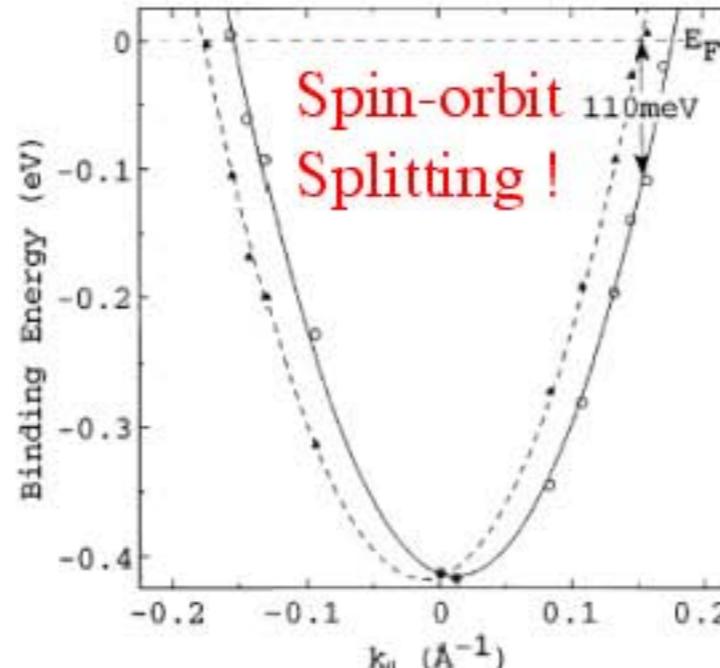
S. LaShell, B. A. McDougall, and E. Jensen

Physics Department, Brandeis University, Waltham, Massachusetts 02254

(Received 19 July 1996)



$h\nu=11.62\text{eV}$

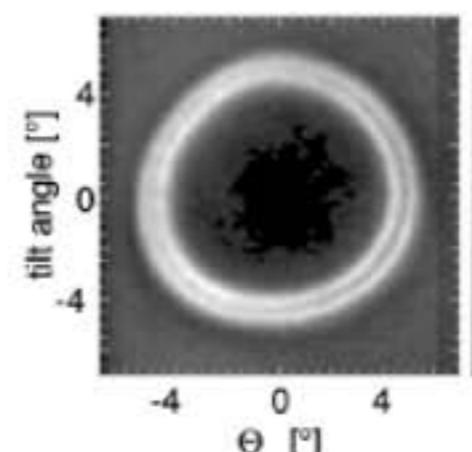
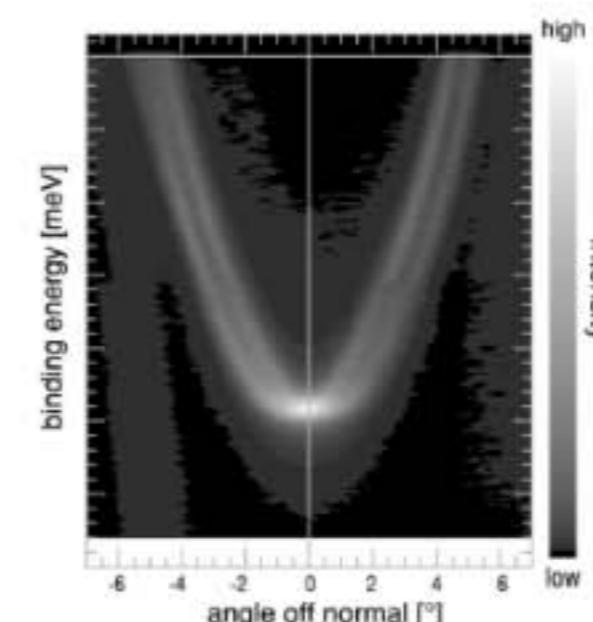


Direct measurements of the L-gap surface states on the (111) face of noble metals by photoelectron spectroscopy

F. Reinert,* G. Nicolay, S. Schmidt, D. Ehm, and S. Hüfner

Fachrichtung Experimentalphysik, Universität des Saarlandes, 66041 Saarbrücken, Germany

(Received 6 October 2000; published 1 March 2001)



Scienta 2002
spectrometer

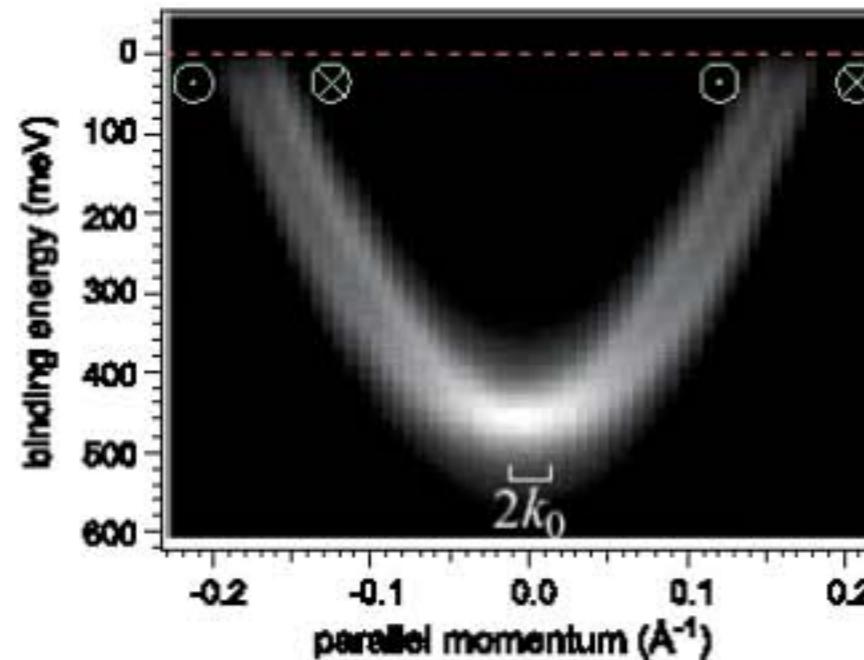
$h\nu=21.21\text{eV}$

=> spin-resolution is a
challenging task !

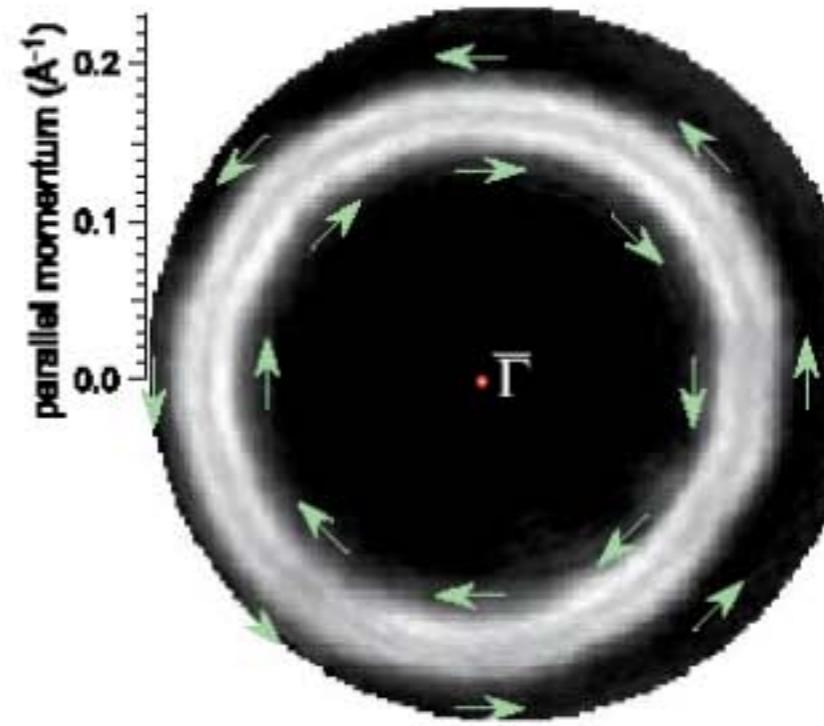
The Shockley surface state on Au(111)

spin-integrated photoemission at $h\nu = 21.1$ eV, $T = 160$ K

instrumental resolution
 $\Delta E = 25$ meV, $\Delta\vartheta = 0.5^\circ$ (FWHM)



dispersion map



Fermi surface map

$$2k_0 = 0.026 \text{ \AA}^{-1} \quad E_B = 470 \text{ meV} \quad k_F = 0.173 \text{ \AA}^{-1} \pm k_0 \quad m^* = 0.24 m_e$$

Theory: spin-orbit coupling

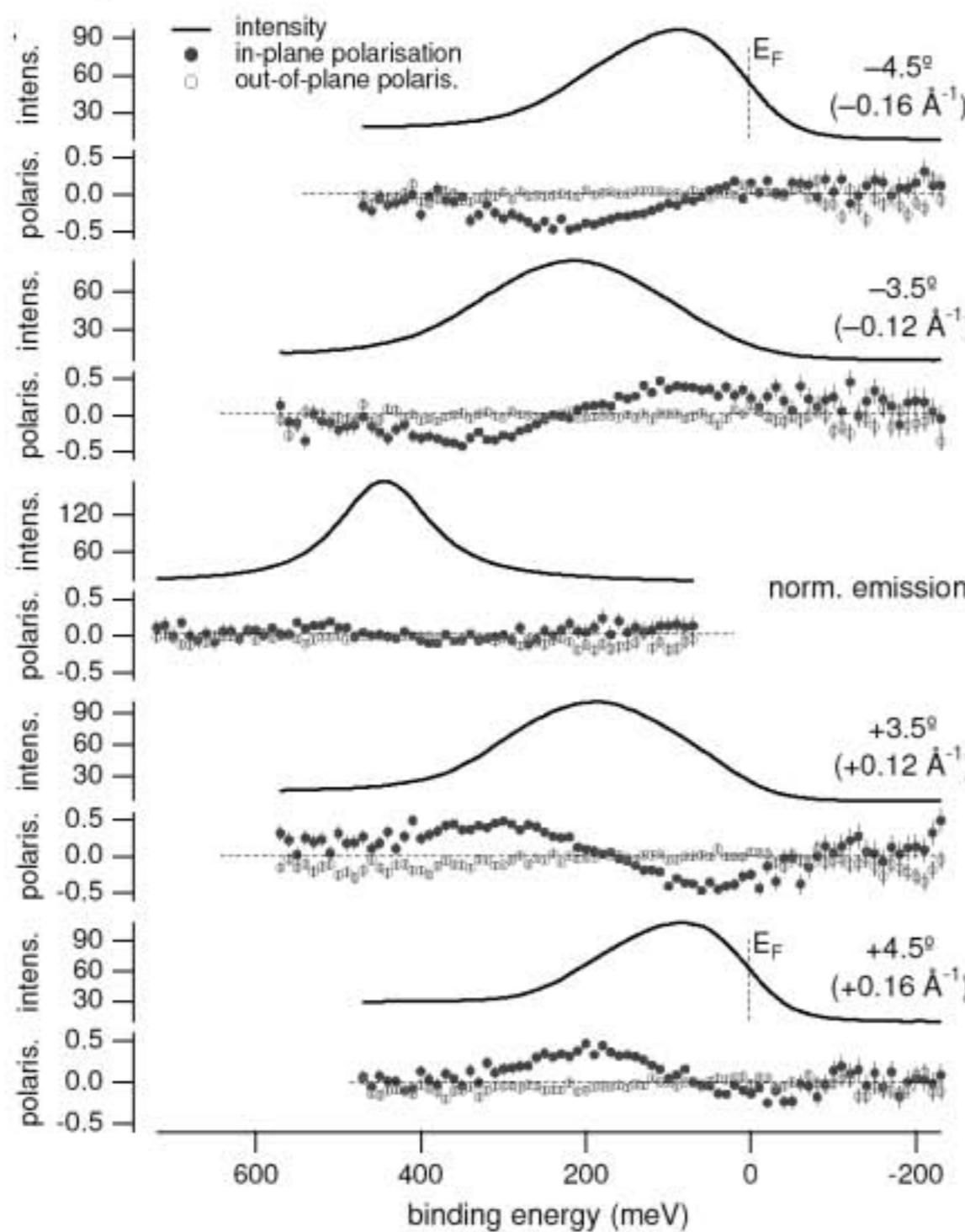
$$H_{SOC} = \frac{\mu_B}{2c^2} (\vec{v} \times \vec{E}) \cdot \vec{\sigma}$$

$$E^*(k) = E_0 + \frac{(k \pm k_0)^2}{2m^*}$$

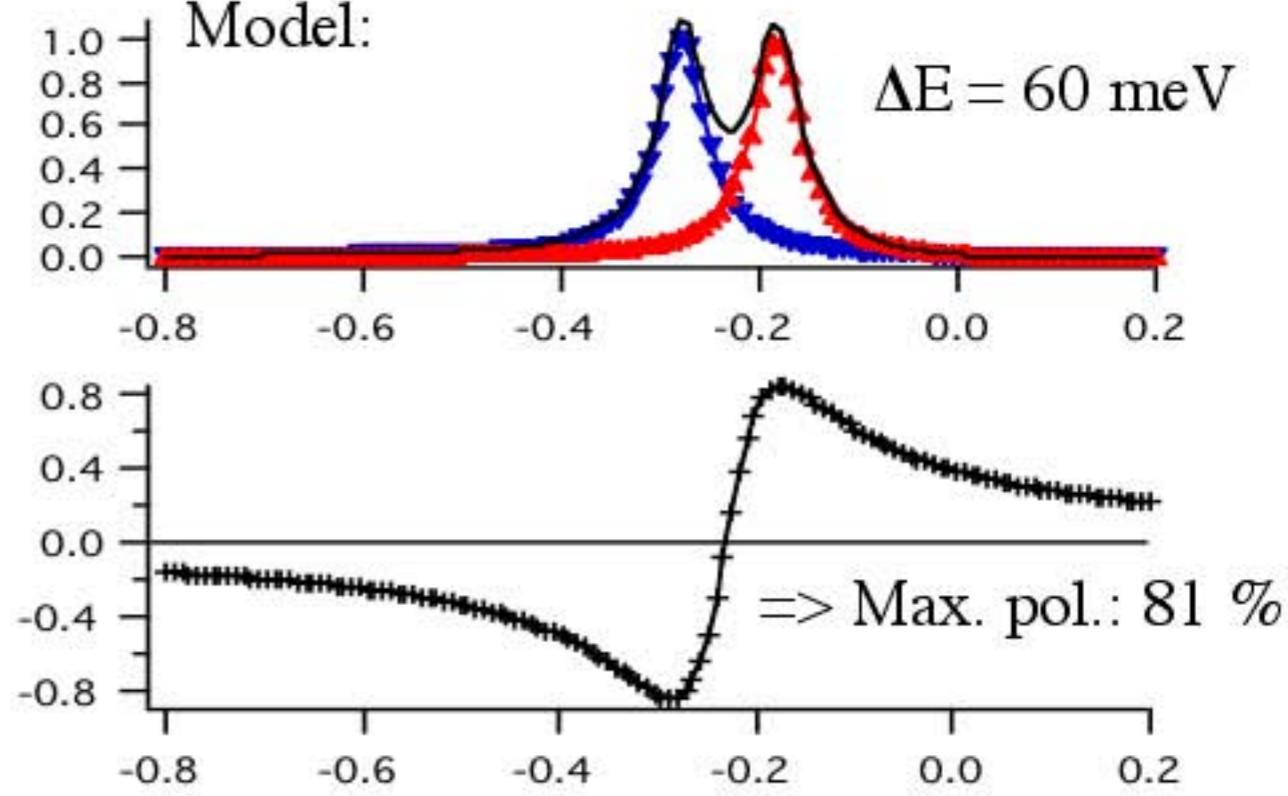
⇒ COPHEE can resolve the spin-orbit splitting
in spin-integrated mode of operation.

Quantitative Spin Polarimetry on the Au(111) Surface State

Experiment:



Model:

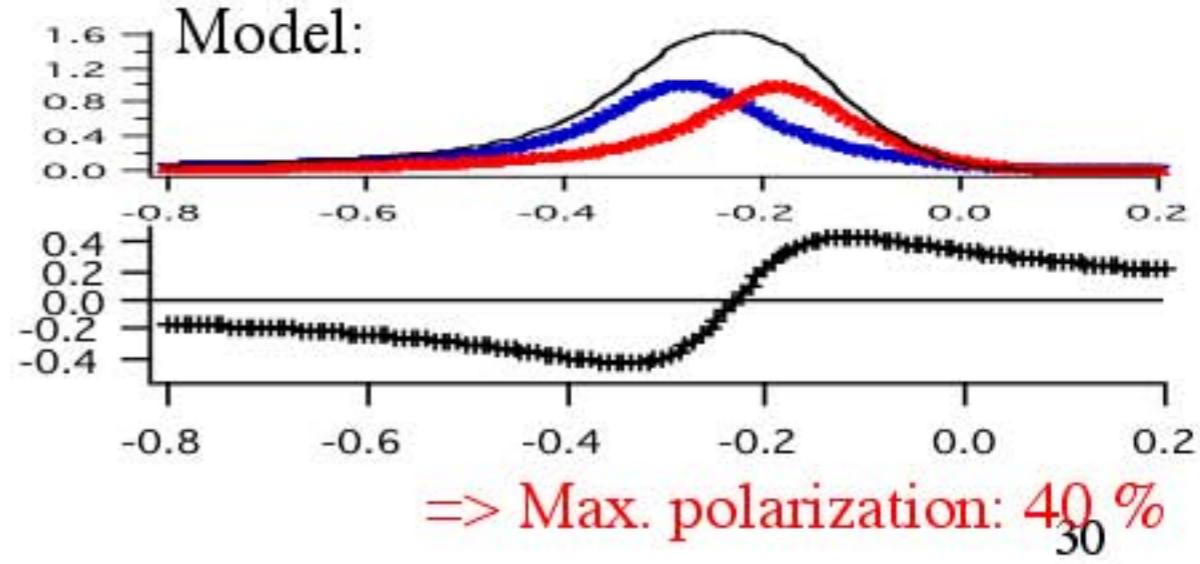


Spin-resolved experiment:

Energy resolution = 120 meV

Angular resolution = 1.2°

Model:



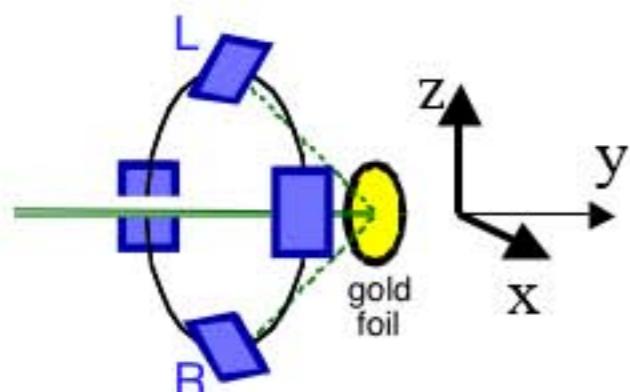
... and Recovering the Spin-Resolved Spectra

Here, sample is not magnetized !

=> Cross asymmetries cannot be obtained !

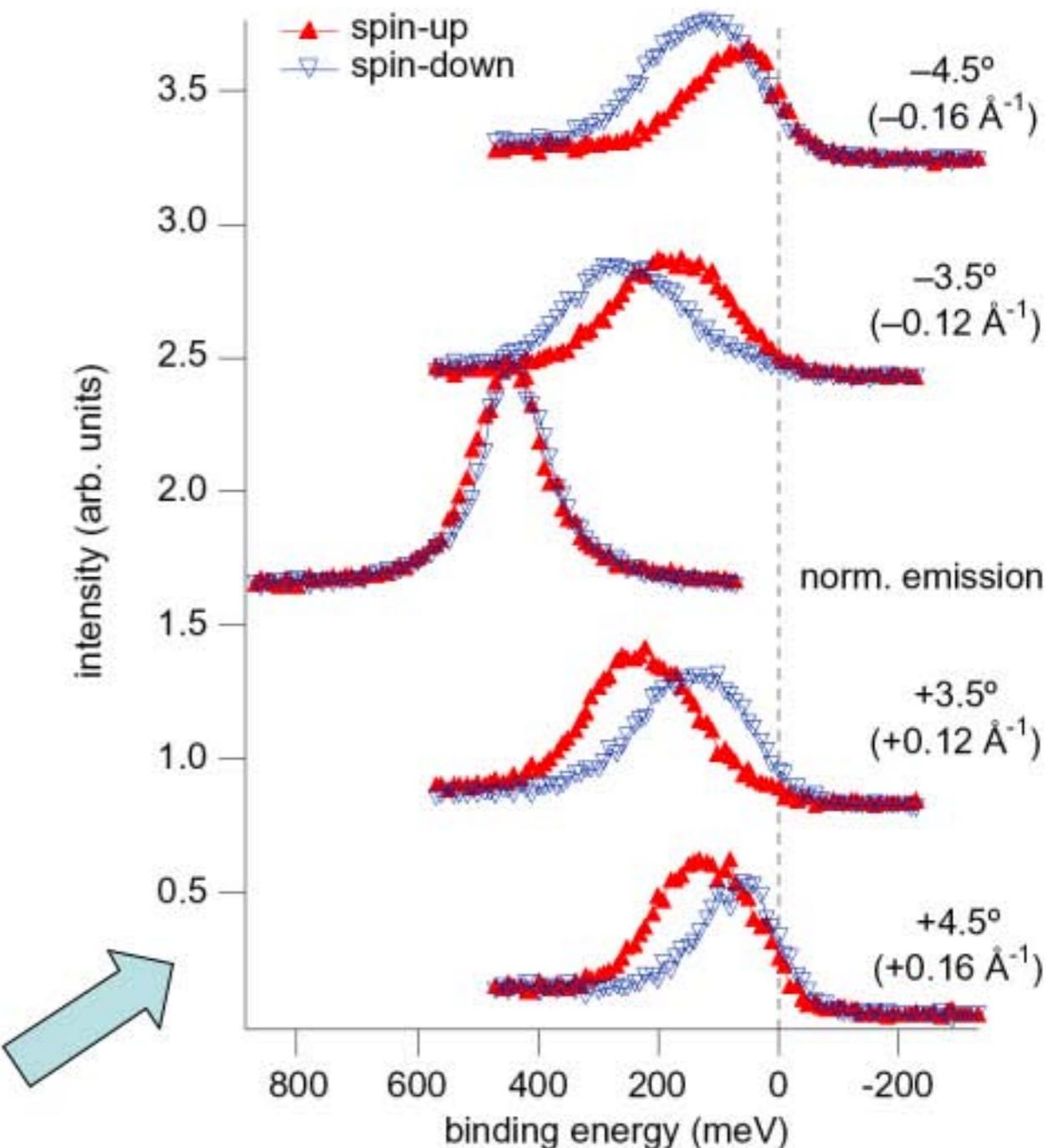
$$P = \frac{1}{S_{eff}} \frac{(I_L - \eta I_R)}{(I_L + \eta I_R)}$$

Use empirical **sensitivity factors** η
to remove instrumental asymmetries

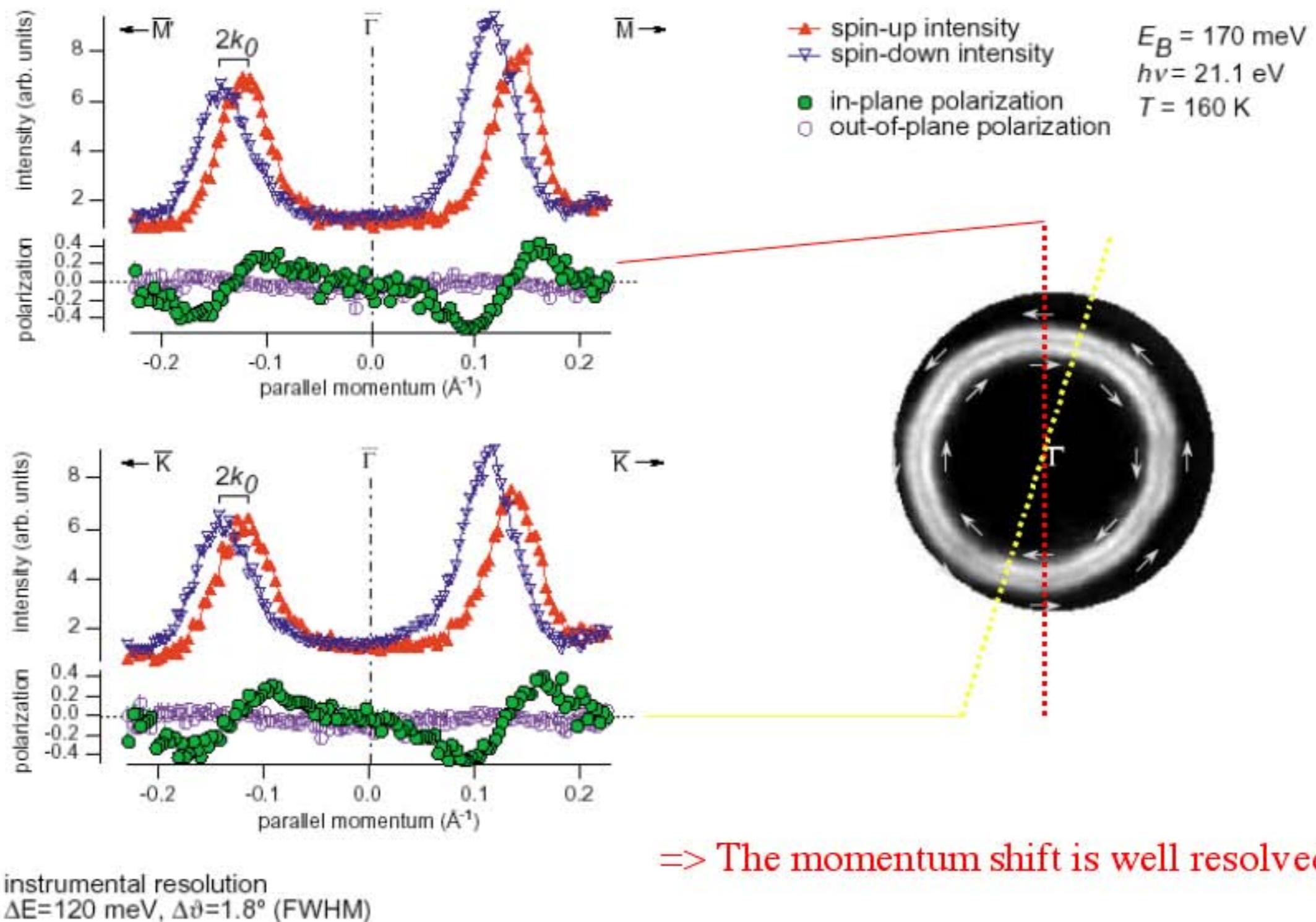


two orthogonal axes are spanned
by two pairs of detectors

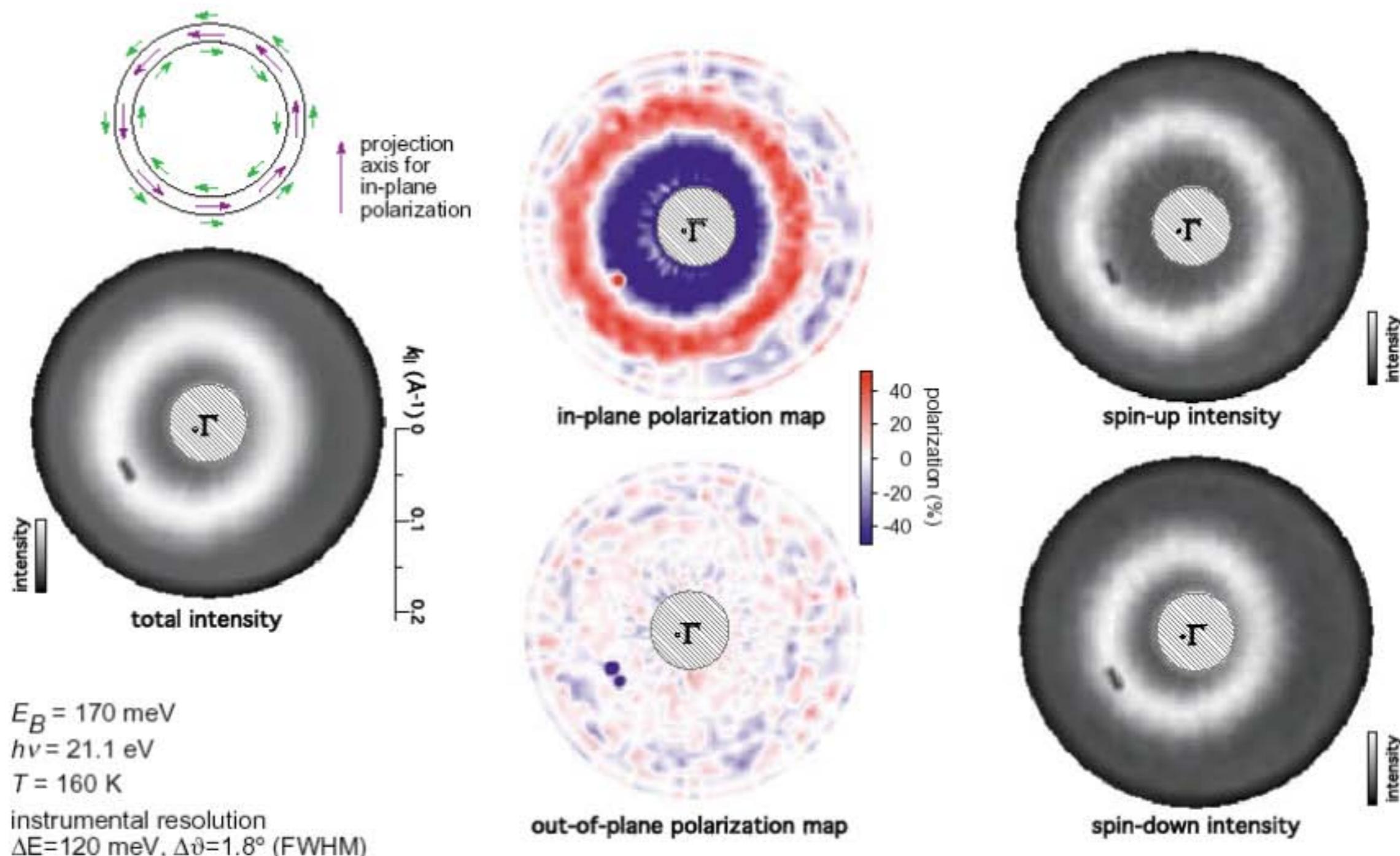
$$I_{\uparrow,\downarrow}(E, k) = I_M(1 \pm P_I(E, k))/2.$$



Spin-resolved momentum distribution curves of the surface state on Au(111)



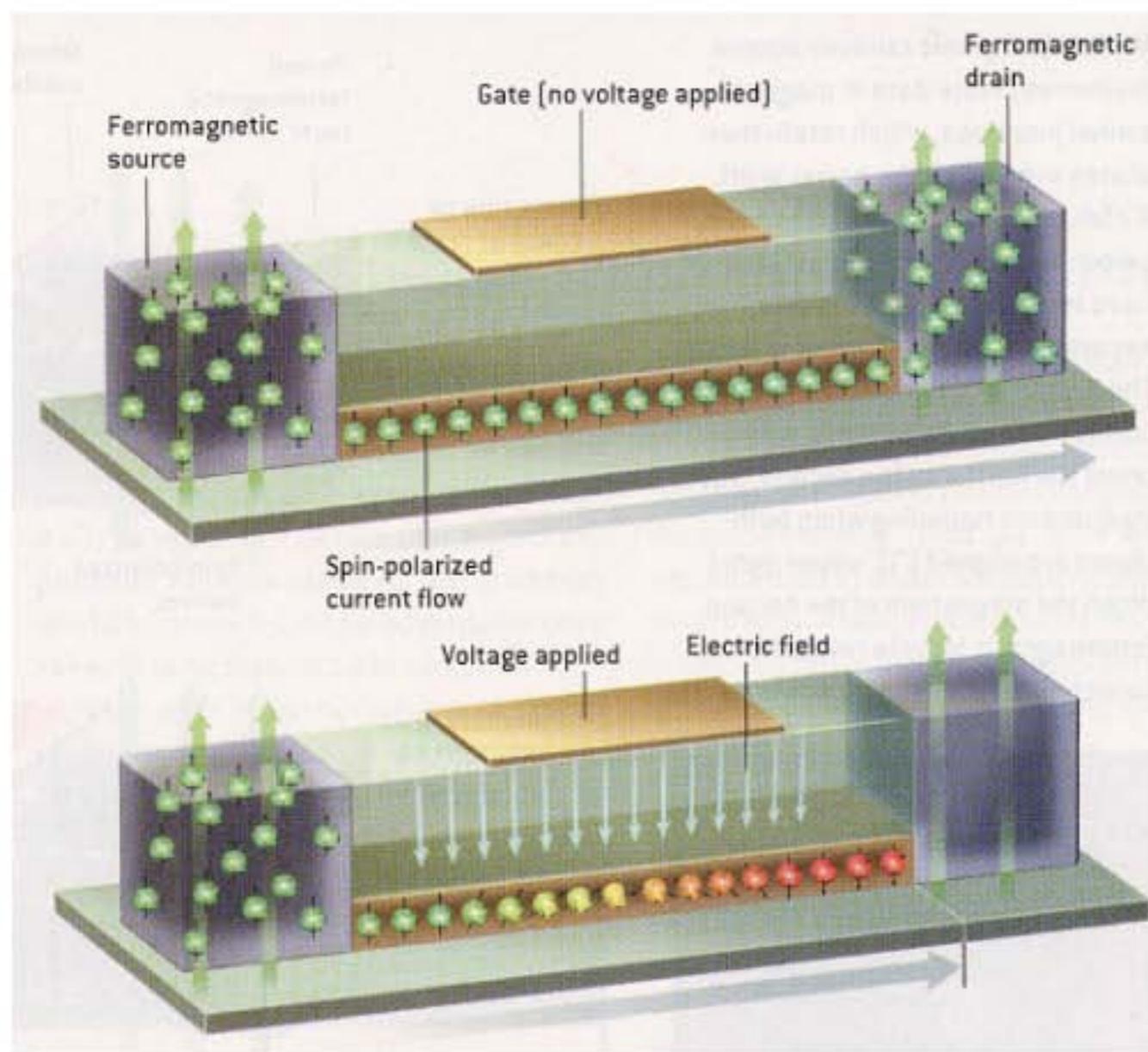
Spin-resolved momentum distribution map of the surface state on Au(111)



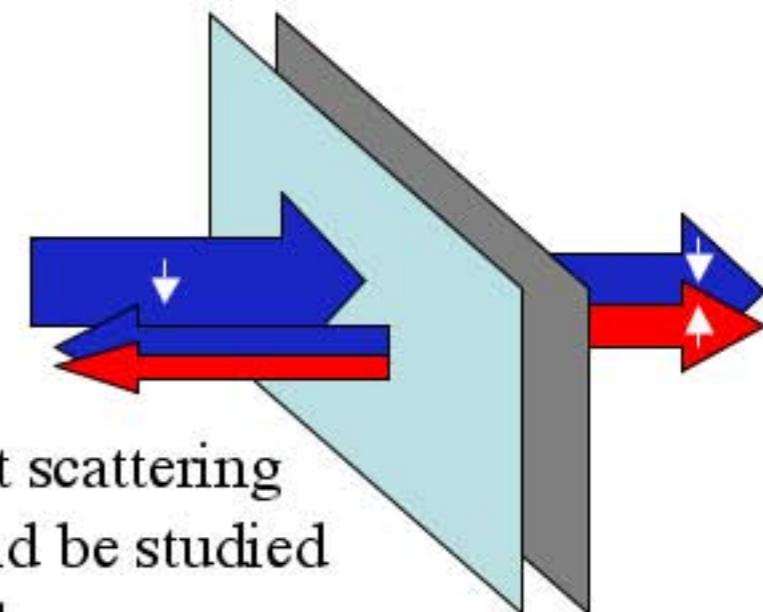
⇒ First spin-resolved “Fermi surface” map
⇒ Rotating spin structure is confirmed

Spintronics: The Importance of Interfaces and Defects

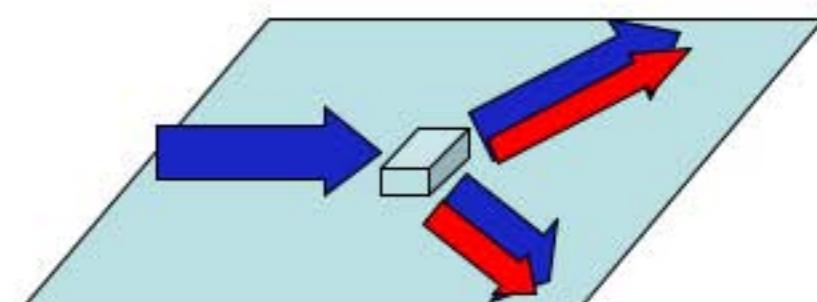
e.g. the spin field effect transistor
... largely bases on such spin-orbit effects.



But to make such a device work:



Spin-dependent scattering processes should be studied and controlled !

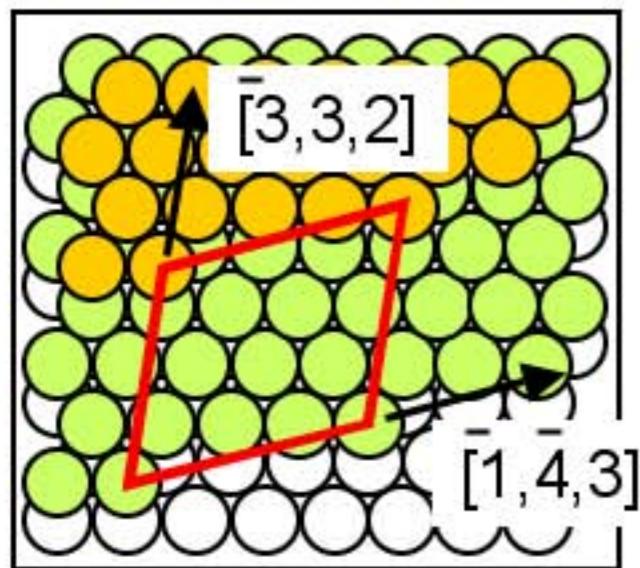


→ We currently study the effect of steps and kinks on the spin structure of Au(111).

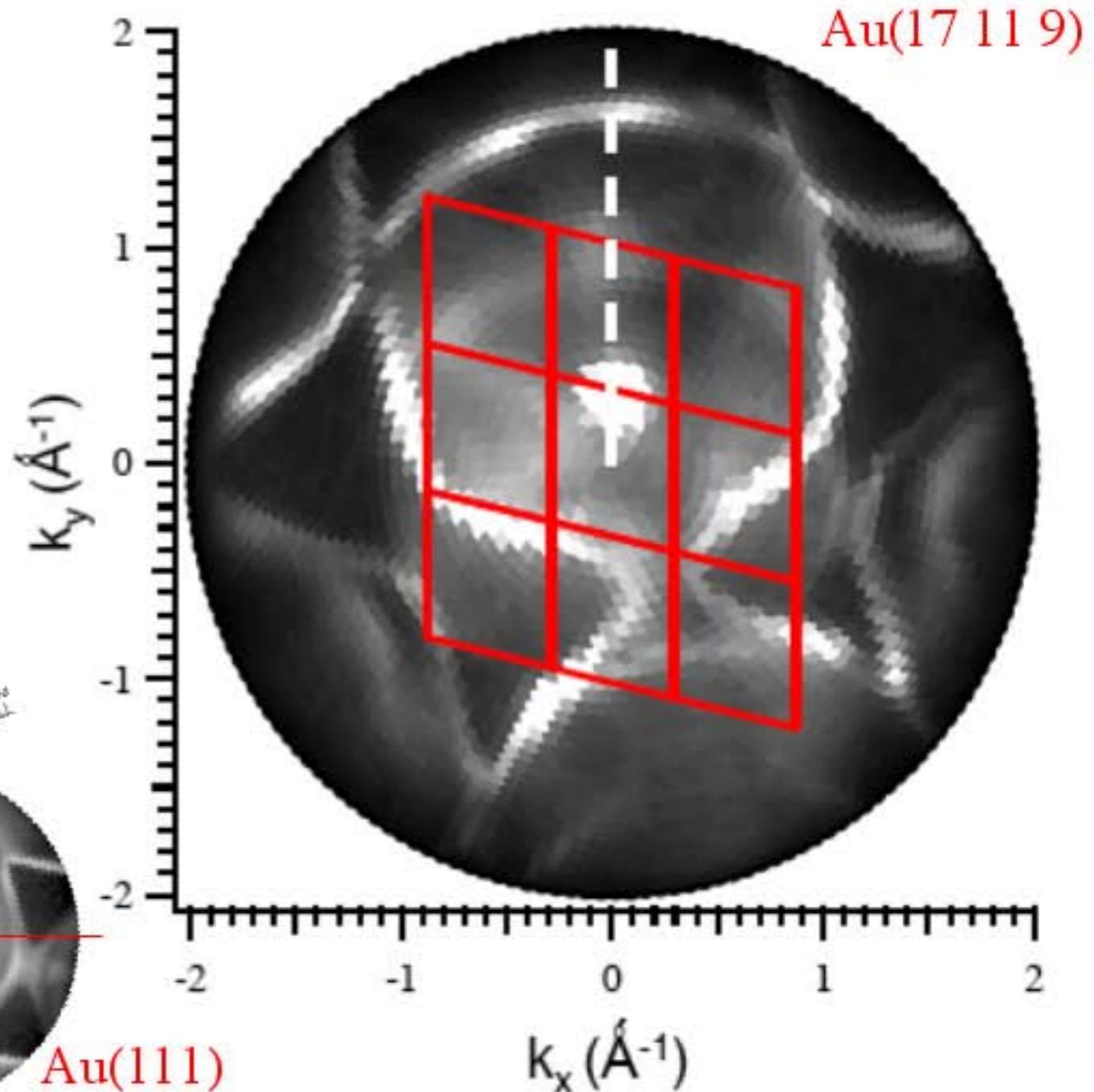
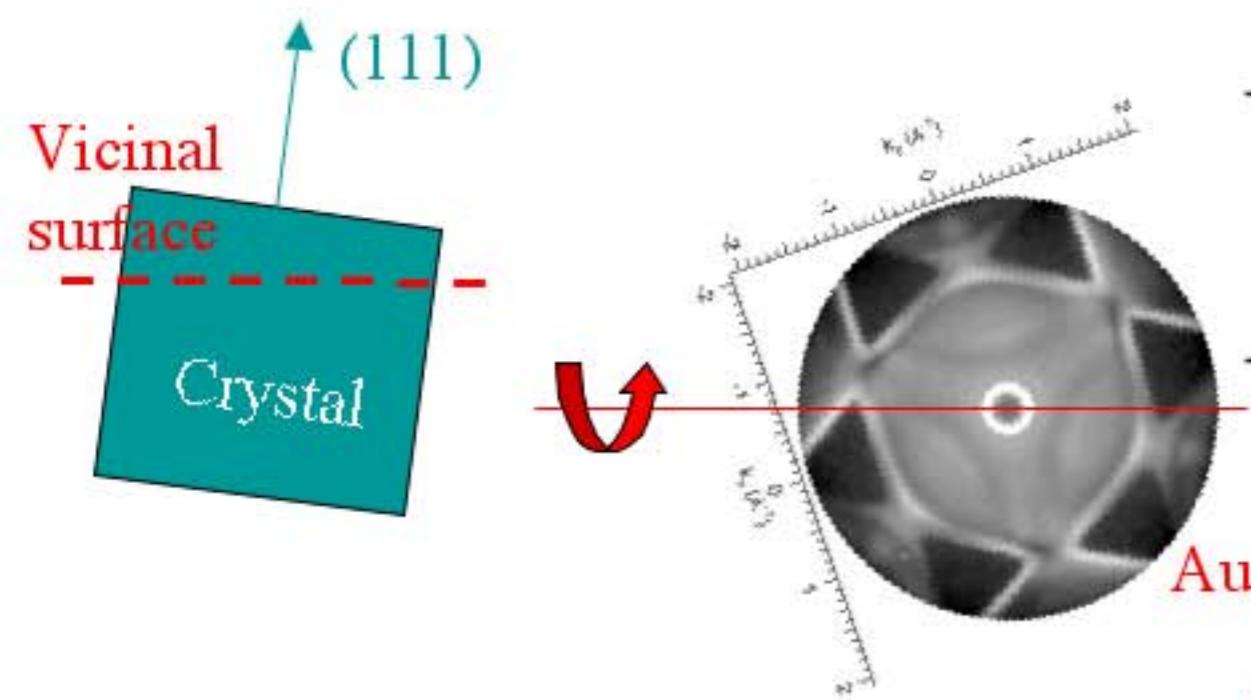
S. Datta & B. Das,
Appl. Phys. Lett. 56, 665 (1990)

Au(111) Surface State - Interaction with Steps and Kinks

Vicinal Au(17 11 9) surface:

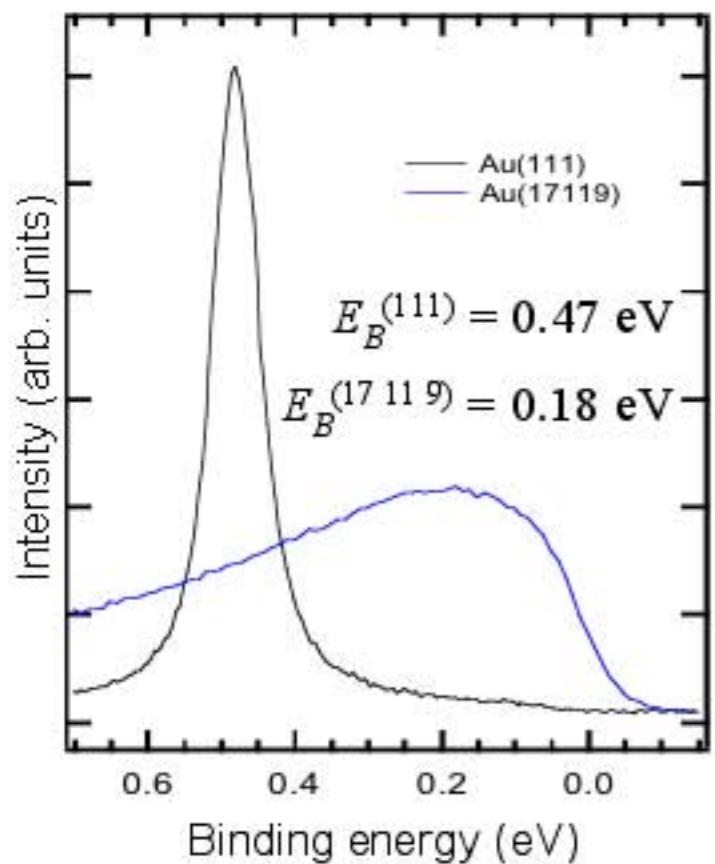


Steps and Kinks !

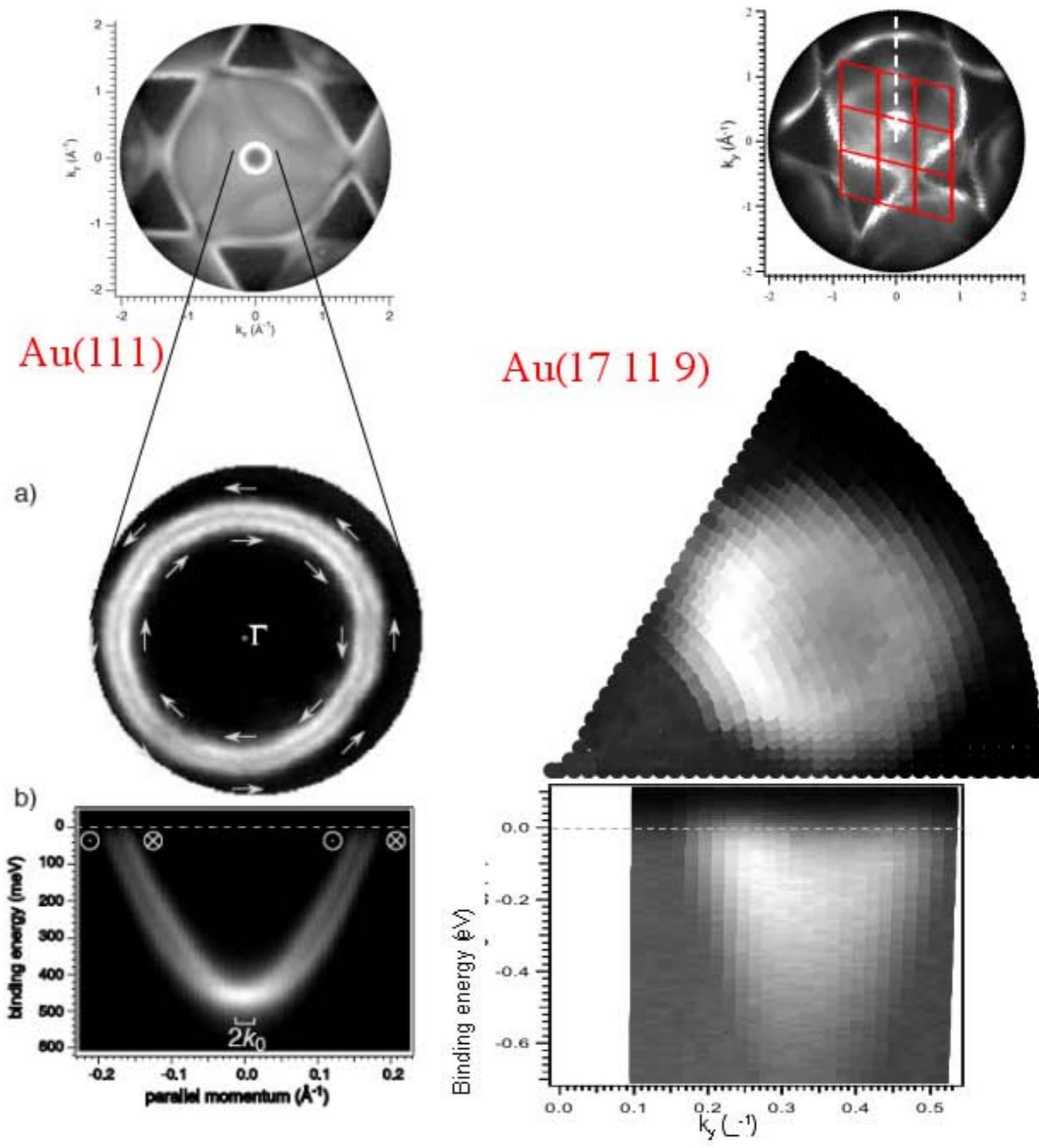


Surface state centered at zone boundary
=> propagates along the macroscopic surface

Effects on the Surface State

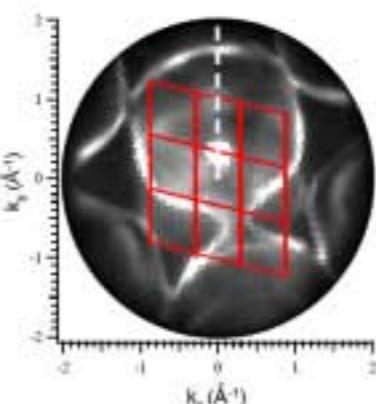


- Band bottom shifts towards E_F
- Effective mass increases (from $m^* = 0.24 m_e$ to $m^* = 0.38 m_e$)
- Peak width large due to terrace width distribution ... and **spin splitting?**

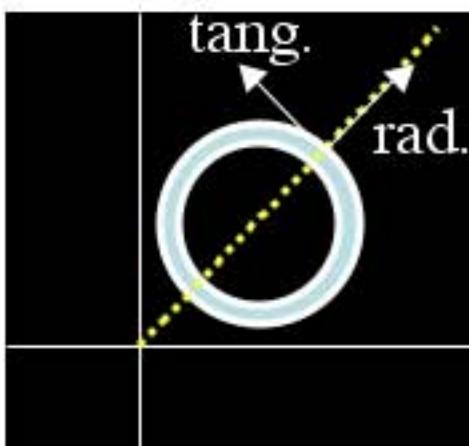


Spin-Polarized ARPES on Au(11 11 9)

MDC through surface state:
(preliminary data)



Peaks are much broader than Rashba splitting

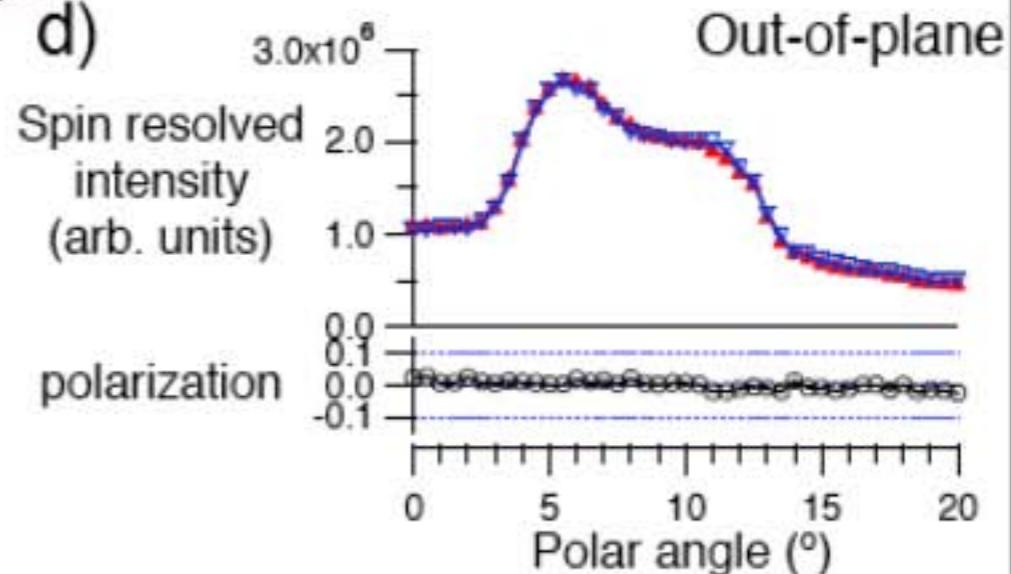
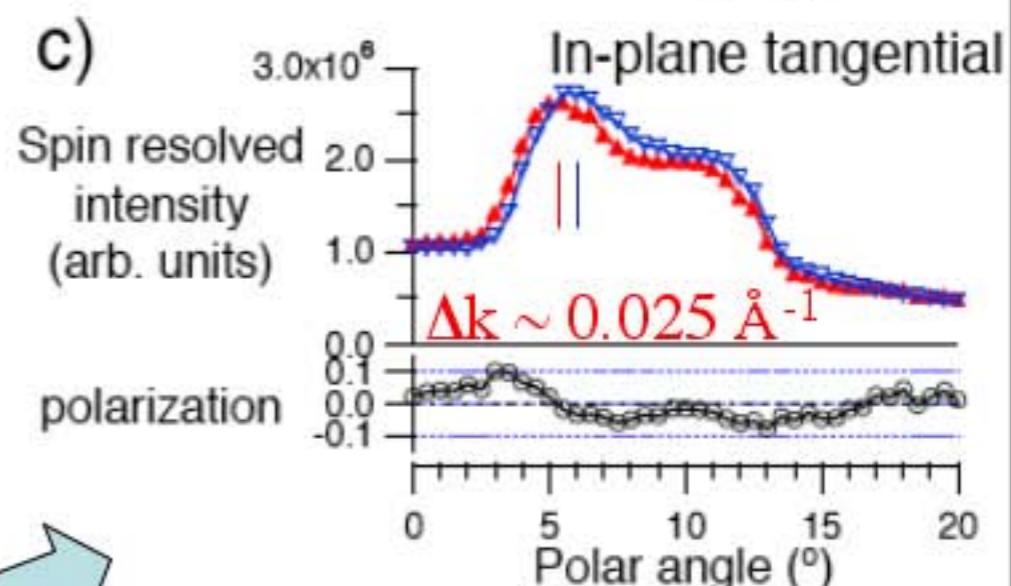
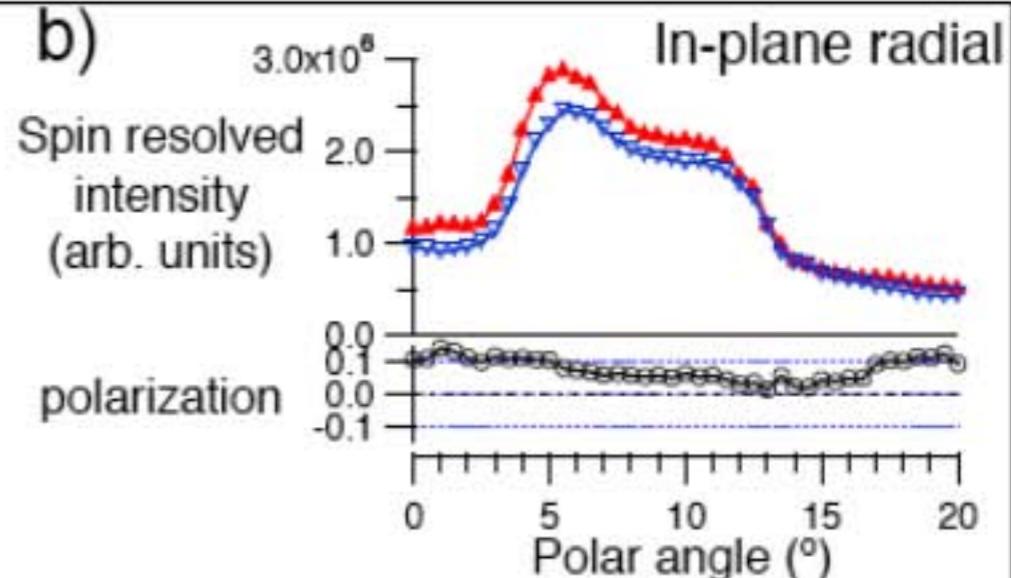


Spin polarized ARPES:

... sees polarization in the
radial in-plane component ?

... can resolve a **spin splitting**
in the tangential in-plane
component !

... sees no out-of-plane
polarization



... We are currently exploring the detailed
spin structure ...

Conclusions

- Spin-polarized photoemission is still a tedious experiment !
(ca. 10^3 times slower than ARPES)
But it is worth the effort !
- Spin-resolution can identify specific bands as minority or minority bands.
- Spin-polarized photoemission can measure spin structures in reciprocal space: e.g. Au(111) and Au(17 11 9)
- Spin resolution can markedly enhance the effective energy resolution for spin-split peaks, even if there individual peaks are broad.