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## **Workshop on Grand Unification and Proton Decay**

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**Unravelling the Missing Link of Grand Unification: Proton Decay**

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# UNRAVELLING THE MISSING LINK OF GRAND UNIFICATION: PROTON DECAY

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ICTP Workshop (July 22-26, 2007)

## REFERENCES:

JCP: Review Talk: "Grand Unification As  
a Bridge..." Aug, 2006, hep-ph/0606089

# I. Introduction

## A) Improved Studies of

① Proton Decay

②  $\nu$  oscillations

2 indispensable tools to probe nature at the shortest distance  $\sim 10^{-30}$  cm.

In a broader sense, these two processes get intimately connected with each other within Grand Unification.

In fact both are predictions of a well-motivated class of Grand Unification symmetries.

I will therefore begin by first listing evidence which has built up over the years in favor of Grand Unification, in particular for a certain class of Grand Unification Symmetries.

## SM Now Brilliantly Successful

Since 1972, both the  $SU(2) \times U(1)$  EW Theory & QCD have been confirmed by numerous experiments

→ A Triumph of Gauge Principle & SSB

→ Waiting To see Higgs at LHC/Fermilab

But  $\exists$  Clear Evidence For Physics Beyond SM

- |   |   |   |
|---|---|---|
| <p>1) Masses (<math>\sqrt{\Delta m^2(\nu)_{23}} \sim 1/20 \text{ eV}</math>)</p> <p>2) Higgs Mass Fine Tuning</p> <p>3) Need For Inflation <math>\left\{ \begin{array}{l} \rightarrow \text{Horizon} \\ \rightarrow \text{Flatness } \Omega = 1 \end{array} \right.</math></p> <p>4) Dark Matter (Cold): <math>\Omega_{DM} \approx 0.23</math></p> <p>5) Baryogenesis <math>\leftrightarrow</math> Leptogenesis</p> | } | <p>All Five<br/>Go Well<br/>With <u>SUSY</u><br/><u>Grand</u><br/><u>Unification!</u></p> |
|---|---|---|

$\Downarrow$   
 There exists PHYSICS BEYOND SM

- |  |   |                                       |
|--|---|---------------------------------------|
| <p>6) Dark Energy: Cosm. Const.<br/><math>\Omega_{\Lambda} \approx 0.72</math></p> | } | <p>A MYSTERY FOR<br/>ALL THEORIES</p> |
|--|---|---------------------------------------|



## II. Grand Unification: Alternative Routes & Evidence

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### A) Main Idea (1972-74)

① Unify (Q & L)

&

② Unify (Weak, EM & Strong) Forces  
 $SU(2)_L \times U(1) \times SU(3)_C$

Purely on aesthetic grounds

## B) Alt. Routes To Grand Unification

$$G(213) = SU(2)_L \times U(1)_Y \times SU(3)_C$$

$$\left( \begin{array}{ccc} u_r & u_y & u_b \\ d_r & d_y & d_b \end{array} \right)_L^{1/3}, \left( u_{r,y,b} \right)_R^{4/3}, \left( d_{r,y,b} \right)_R^{-2/3}, \left( \nu_e \right)_L^{-1}, \left( e^- \right)_L^{-2}, \left( e^- \right)_R^{-2}$$

5 disconn. multiplets // Arb.  $Y_W, SU(2), SU(3)_C$  Q. NOS.

$$G(224) = SU(2)_L \times SU(2)_R \times SU(4)_{L+R}^C \times (L+R)$$

$$F_{L,R}^e = \begin{bmatrix} u_r & u_y & u_b & \nu_e \\ d_r & d_y & d_b & e^- \end{bmatrix}_{L,R} \quad \begin{array}{l} F_L^e = (2, 1, 4) \\ F_R^e = (1, 2, 4) \end{array}$$

$$Q_{em} = I_{3L} + I_{3R} + \frac{B-L}{2}$$

### Advantages of G(224)

- All 16 in one L-R Conj. multiplet // L-R Symm
- Explain  $Y_W$  & All Q. NOS. // Quantize  $Q_{em}$  //
- $Q_{e^-} = -Q_p$  //  $(\nu_R // B-L)$  ← Need For seesaw & Leptogenesis

All Advantages of G(224) retained by SO(10), not SU(5)

$$\downarrow$$

$$SO(10): 16 \text{ (one coupling)}$$

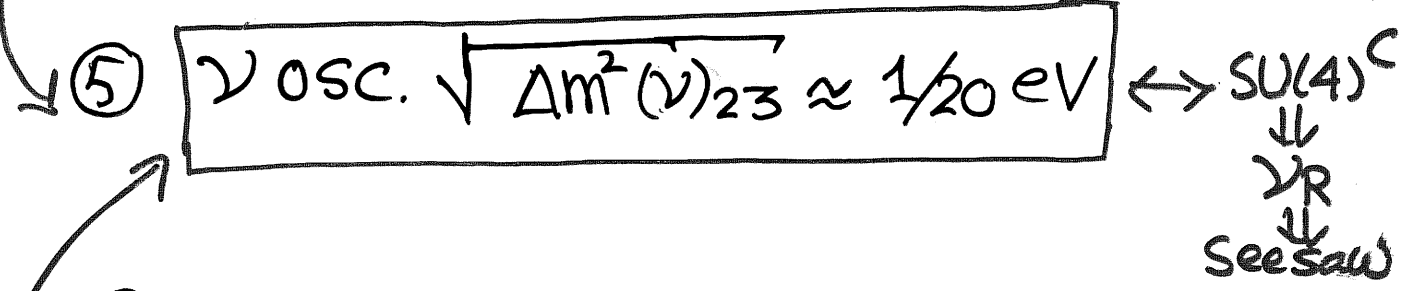
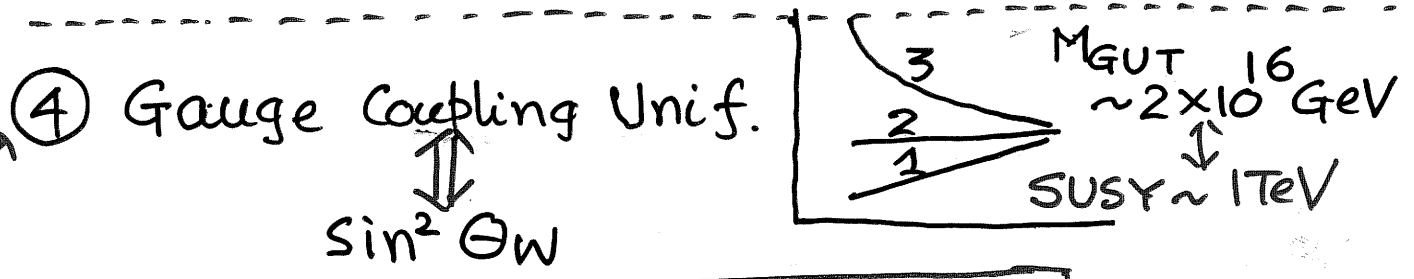
$$SU(5): \bar{5} + 10$$

NO  $\nu_R$ , NO B-L

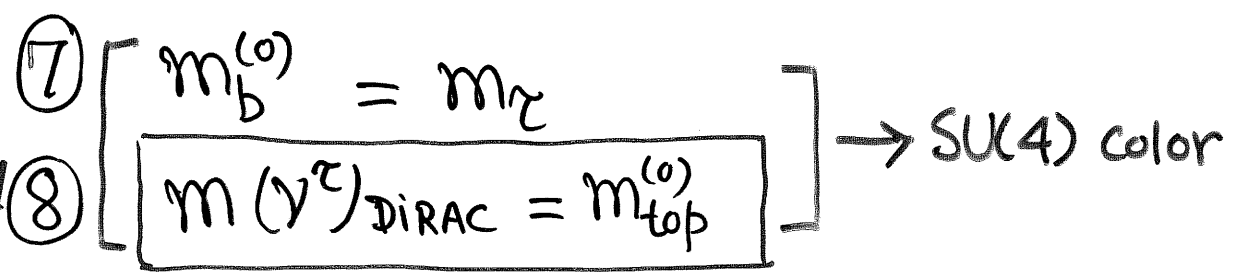
} Problem with  $\nu$  Masses & Leptogenesis.

# II C. EVIDENCE FOR GRAND UNIFICATION

- ① Quantum nos. of members in a Family
- ② Quantization of  $Q_{em}$
- ③  $Q_{e^-} = -Q_p$



⑥ Baryogenesis  $\leftrightarrow$  Leptogenesis  $\rightarrow \nu_R, B-L$



Agreement of all eight non-trivial  
 $\Downarrow$   
 strong support for SUSY Grand Unif.  
 & an Eff. Symmetry like  $SO(10)$  or minimally  
 a string-derived  $G(224)$ .

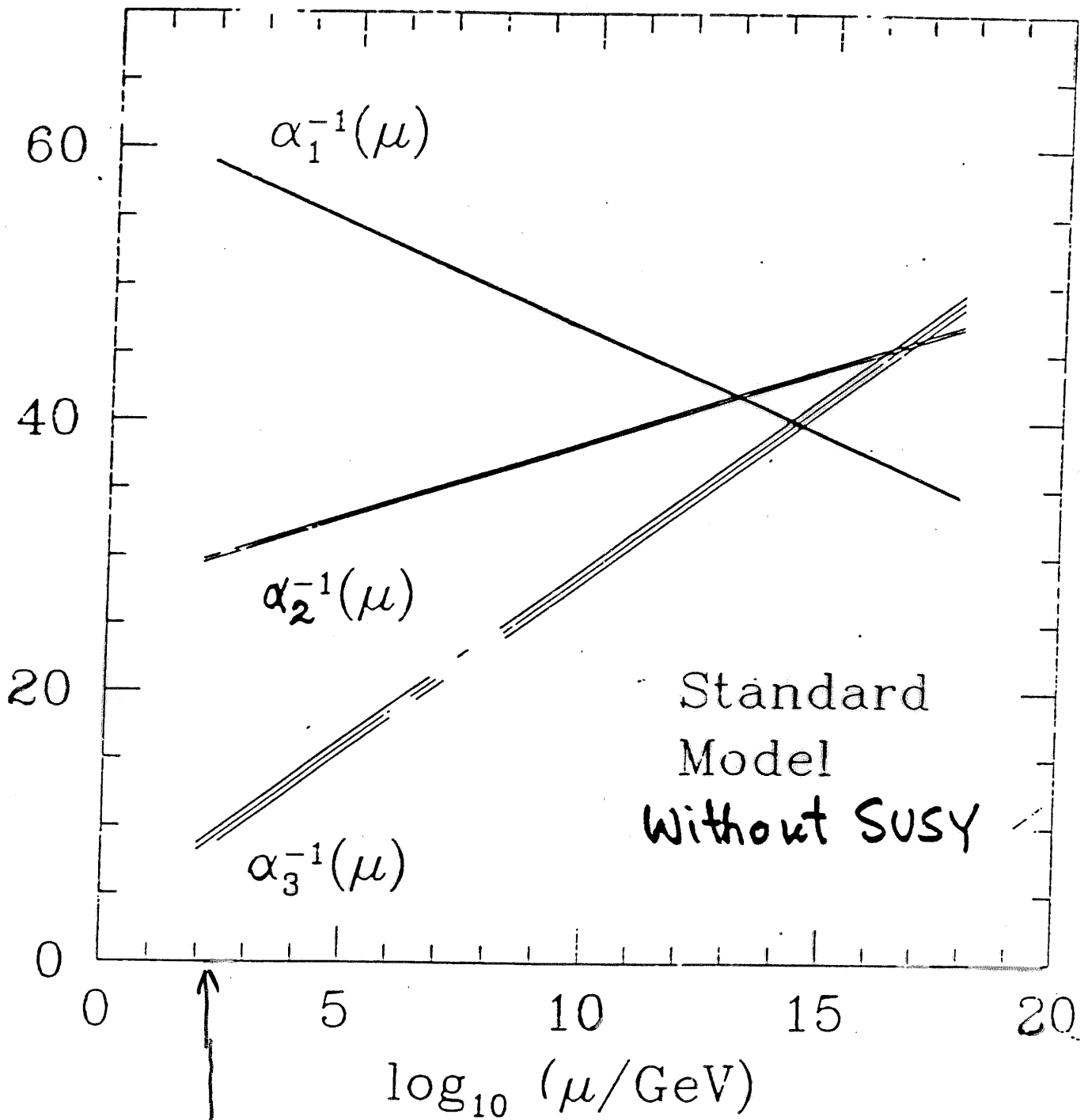
## II. Evidence Favoring Grand Unification

- 1) Family-Multiplet Structure // All 5 sub-multiplets  $\rightarrow$  ONE MULTIPLET }  $\Rightarrow$   $G(224)/SO(10)$   
 $SU(5)$
- 2) Quantization of  $Q_{em}$   $\rightarrow$   $SU(5)/G(224)/SO(10)$
- 3) Coupling Unification  $M_X \sim 2 \times 10^{16}$  GeV)  $\rightarrow$   $SU(5)/SO(10)$  / String  $G(224)$   
 $\leftarrow$  SUSY  $\xrightarrow{G(224)}$
- 4)  $m(\nu_\tau) \sim 1/20$  eV (superk)  $\rightarrow$  String  $G(224)/SO(10)$   
 $SU(5)$
- 5)  $\theta(\nu_\mu - \nu_\tau)$  Large  $\leftrightarrow$   $V_{bc}$  small  $\rightarrow$   $G(224)/SO(10)$
- 6)  $m_b^0 \approx m_\tau^0$   $\rightarrow$   $G(224)/SO(10)/SU(5)$
- 7) Lepto/Baryogenesis  $\rightarrow$  Spont  $\rightarrow$   $G(224)/SO(10)$   
 $B-L$   $SU(5)$

THE SUCCESS OF THESE SEVEN FEATURES SEEM NON-TRIVIAL. TOGETHER THEY MAKE A STRONG CASE FOR SUSY GRAND UNIFICATION & SIMULTANEOUSLY FOR A STRING-DERIVED -  $G(224)/SO(10)$  ROUTE TO SUCH UNIFICATION.

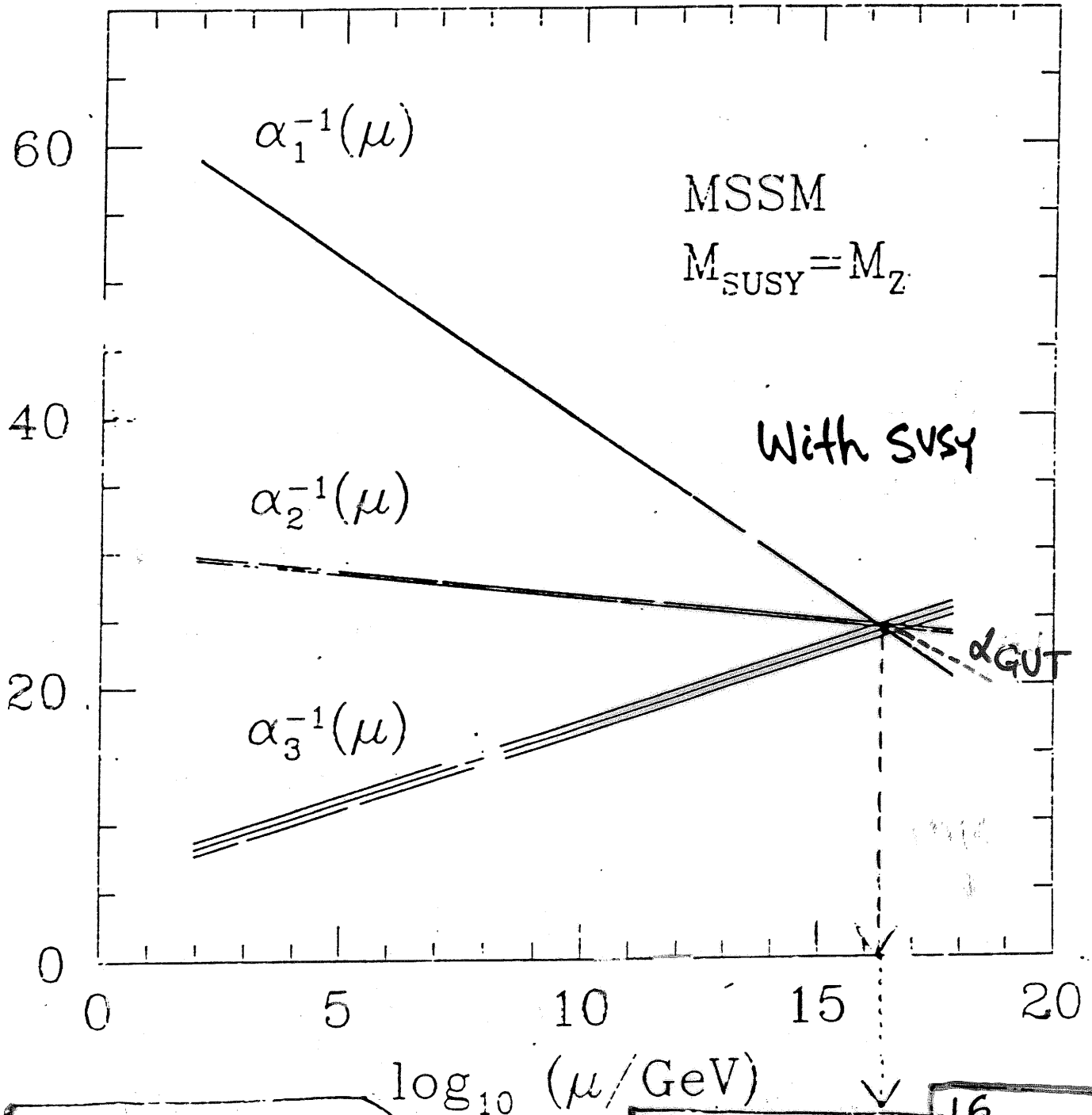
CONSIDER PROTON DECAY WITHIN THIS FRAMEWORK.

# Coupling Unification Without SUSY



$\alpha_i$ 's Measured at LEP  
at  $\sim 100$  GeV

# Gauge Coupling Unification With SUSY



Supports SUSY Unification

$M_{\text{GUT}} \approx 2 \times 10^{16} \text{ GeV}$

$$\sin^2 \Theta_W(m_Z)_{\text{th}} = 0.2315 \pm 0.003$$

$$\sin^2 \Theta_W(m_Z)_{\text{EXPT}} = 0.23124 \pm 0.00017$$

$$\alpha_3(m_Z)_{\text{theory}}^0 \approx 0.125 - 0.13$$

$$\alpha_3(m_Z)_{\text{obs}} = 0.118 \pm 0.003$$

GUT THRESHOLD CORRECTION

## TWO GENERAL PREDICTIONS OF GUT SO(10)/ SUSY G(224)

- ①  $\nu$  OSC
- ② PROTON DECAY



Correlated With Fermion Masses  
& Mixings

$\therefore$  Discuss First Fermion Masses  
& Mixings within  $G(224)/SO(10)$

### III SO(10) Breaking: Alternative Higgs

#### ① Low Dim. Higgs

{ $45_H, 16_H, \bar{16}_H, 10_H$ }

Babu, Barr // Albright, Babu, Barr //  
Albright, Barr (Lopsided) // Dermisek, Mafi,  
Raby // Babu, Pati, Wilczek (Hierarchical)

#### ② Large Dim. Higgs

$126, \bar{126}, (210 // 120 // 54), 10$

Aulukh, Mohapatra // Babu, Mohapatra //  
Aulukh, Bajc, Melfo, Rasin, Senjanovic //  
Clark, Kuo, Nakagawa // Chang, Mohapatra,  
Parida // Chen, Mahanthappa // Aulukh //  
Bando et al // Fukuyama et al //

A Good Comparative analysis:

Melfo, Senjanovic (2005).



## Low Dim. Higgs

 $45_H, 16_H, \bar{16}_H, 10_H$ 

$$16_i 16_j 10_H ; 16_i 16_j 10_H \frac{45_H}{M}$$

$$16_i 16_j 16_H \frac{16_H}{M} ; 16_i 16_j \bar{16}_H \frac{\bar{16}_H}{M}$$



- Non-Ren. Terms // should be kept if allowed by Symm.

$$|\Delta(B-L)_{16_H}| = 1$$



R-Parity Broken, but discrete Symm naturally Present

- GUT Threshold Corr & RG effects incl. D-T Splitting Well under Control // Coupling Unif. demonstrated Sign & Magn.

$$m_b^{(0)} = m_\tau, \text{ Yet } m_\mu \neq m_s$$

&  $V_{cb} \ll \Theta_{23}^2$

Fermion Mass Correl  
Proton Decay Concretely Studied  
(BPW)

## Large Dim. Higgs

 $126, \bar{126}, (210 // 120 // 54), 10_H$ 

$$16_i 16_j (10_H, 126, 120)$$



- All Ren. Terms

$$|\Delta(B-L)_{126_H}| = 2$$



R-Parity Preserved

- Threshold Corr. Large // RG studies & Coupling Unif  $\leftrightarrow$  MGUT ?

$$m_b^{(0)} = m_\tau \quad \text{Lost}$$

?

c) Insight From SuperK Result:  $\sqrt{\Delta m_{23}^2} \approx 1/20 \text{ eV}$  <sup>9</sup>

SeeSaw  
ignore mixing  
for a moment

$$m(\nu_L^c) \approx \frac{m(\nu_{\text{Dirac}}^c)^2}{M(\nu_R^c)}$$

(a)  $m(\nu_{\text{Dirac}}^c) \approx m_t (M_X) \approx 120 \text{ GeV}$  ← SU(4) - Color  
SU(5), [SU(3)]<sup>3</sup>

$m_b \approx m_c$

(b) Get  $M(\nu_R^c)$  from SUSY uni f. Scale:  $M_X \approx 2 \times 10^{16} \text{ GeV}$

$$f_{33} \frac{16_3 16_3 \langle \bar{16}_H \rangle \langle \bar{16}_H \rangle}{M \rightarrow \sim 10^{18} \text{ GeV}} \Rightarrow M(\nu_R^c) \sim \frac{(2 \times 10^{16} \text{ GeV})^2}{10^{18} \text{ GeV}} \approx \{4 \times 10^{14} \text{ GeV} (\frac{1}{2}-2)\}$$

( $\approx 1$ )

$$m(\nu_L^c) \sim \frac{(120 \text{ GeV})^2}{4 \times 10^{14} \text{ GeV}} \approx \left(\frac{1}{30} \text{ eV}\right) \left(\frac{1}{2} \text{ to } 2\right)$$

Also get  $m(\nu_L^{uc}) \sim \frac{m(\nu_L^c)}{10} \Rightarrow \sqrt{\Delta m_{23}^2} \approx \left(\frac{1}{30} \text{ eV}\right) \left(\frac{1}{2}-2\right)$

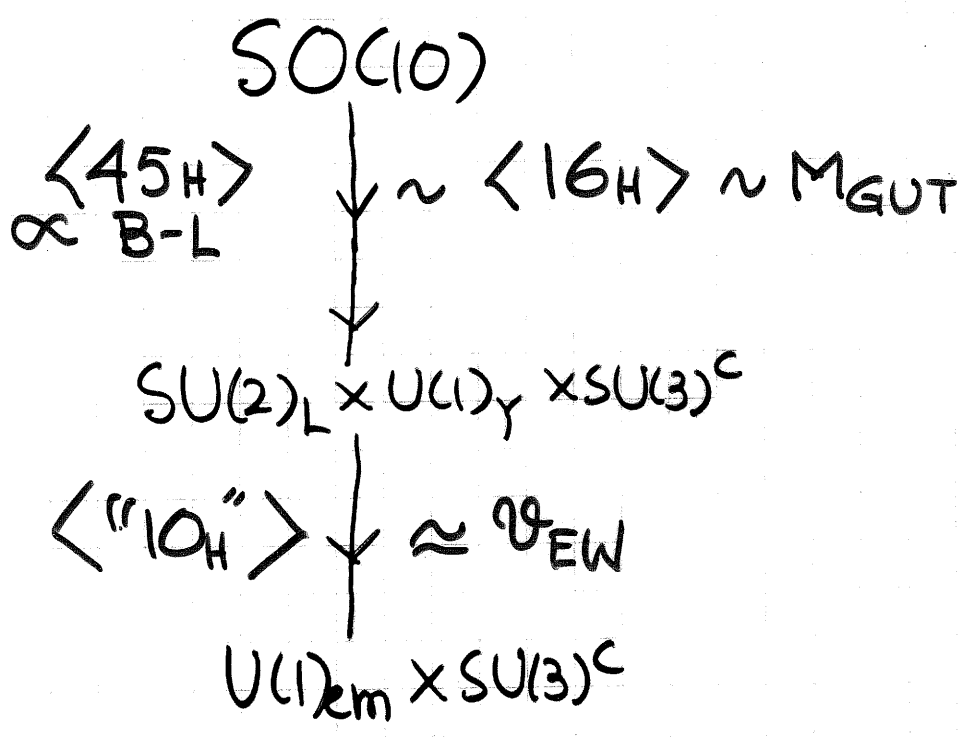
Thus SuperK result brings to light the existence of  $\nu_R$  // reinforces the ideas of  
a) SeeSaw // (b) SU(4) color // & (c) SUSY Unif

IV Assume  $(9, l, \nu)$  Masses & Mixings in a predictive  $G(224)$  or  $SO(10)$ -Framework

Babu, Pati, Wilczek (98, 2000)

Minimal Higgs For  $SO(10)$  Breaking

$\{45_H, 16_H, \bar{16}_H, 10_H\} \quad 126_H, 126_H, \dots$



Only allowed cubic coupling

$16_i 16_j 10_H \rightarrow B-L \text{ Indep // symmetric // } \begin{matrix} i \leftrightarrow j \\ \text{No CKM} \end{matrix}$

These arise from  $\leftarrow$  Higher Dim operators  
 $16_i 16_j 16_H 16_H / M, \quad 16_i 16_j 10_H 45_H / M$

# FLAVOR SYMMETRY: $SO(10) \times U(1)_F$

Assume Hierarchical Yukawa Couplings

$$"33" \gg "23" \sim "32" \gg "22" \gg "12" \text{ etc.}$$

Due to Flavor Symmetry  $U(1)_F$

$16_3$	$16_2$	$16_1$	$10_H$	$16_H$	$\bar{16}_H$	$45_H$	$S$
$U(1)_F \rightarrow a$	$a+1$	$a+2$	$-2a$	$-a-1/2$	$-a$	$0$	$-1$

Thus,  $16_3 \begin{smallmatrix} a \\ a \end{smallmatrix} 16_3 \begin{smallmatrix} a \\ a \end{smallmatrix} 10_H \begin{smallmatrix} -2a \\ -2a \end{smallmatrix} \rightarrow \text{Allowed} \sim 1$

$$16_3 \begin{smallmatrix} a \\ a \end{smallmatrix} 16_2 \begin{smallmatrix} a \\ a \end{smallmatrix} 10_H \langle S/M \rangle \rightarrow \left( \frac{M_{GUT}}{M_{st}} \sim 1/10 \right)$$

$$\langle S \rangle \sim M_{GUT}, \quad M \sim M_{string}$$

# A Concrete Example: Minimal Higgs: $\{45_H, 16_H, \bar{16}_H, 10_H\} + "5"$

Consider only  $\mu$  &  $\tau$  Families (Flavor sym:  $\mu \neq \tau$ )

- ①  $L_{mass} = h_{33} 16_3 16_3 \langle 10_H \rangle \rightarrow$  3rd Family:  $m_b^0 = m_\tau^0$
- ②  $+ \tilde{h}_{23} 16_2 16_3 \langle 10_H \rangle \langle S/M \rangle \rightarrow$  2nd Family
- ③  $+ a_{23} 16_2 16_3 \langle 10_H \rangle \langle 45_H \rangle \rightarrow m_\mu^0 \neq m_s^0$   
ANTISYMMETRIC,  $\propto B-L$
- ④  $+ g_{23} 16_2 16_3 \langle 16_H \rangle \langle 16_H \rangle_{EW/M}^D \rightarrow$  CKM  $\neq 1$

$10_H \rightarrow (2 \leftrightarrow 2) \ll (2 \leftrightarrow 3) \ll (3, 3)$  Flavor Sym

$$U = \begin{pmatrix} c & t \\ 0 & \epsilon + \sigma \\ -\epsilon + \sigma & 1 \end{pmatrix} m_U^0 \quad D = \begin{pmatrix} s & b \\ 0 & \epsilon + \eta \\ -\epsilon + \eta & 1 \end{pmatrix} m_D^0$$

$$N_{Dirac} = \begin{pmatrix} \nu_\mu & \nu_\tau \\ 0 & -3\epsilon + \sigma \\ +3\epsilon + \sigma & 1 \end{pmatrix} m_U^0 \quad L = \begin{pmatrix} \mu & \tau \\ 0 & -3\epsilon + \eta \\ 3\epsilon + \eta & 1 \end{pmatrix} m_D^0$$

Note  $q-l$  correlation (SU(4) - color)  
 up-down                      )                      (SU(2)<sub>L</sub> x SU(2)<sub>R</sub>)

$\eta \equiv \hat{\eta} + \sigma$

# Dirac Masses (3 Families)

$$U = \begin{pmatrix} 0 & \epsilon' & 0 \\ -\epsilon' & 0 & \epsilon + \sigma \\ 0 & -\epsilon + \sigma & 1 \end{pmatrix} m_{\nu}^0 ; \quad D = \begin{pmatrix} 0 & \epsilon' \eta' & 0 \\ -\epsilon' \eta' & 0 & \epsilon + \eta \\ 0 & \epsilon + \eta & 1 \end{pmatrix} m_{\nu}^0$$

$$N = \begin{pmatrix} 0 & -3\epsilon' & 0 \\ 3\epsilon' & 0 & -3\epsilon + \sigma \\ 0 & 3\epsilon + \sigma & 1 \end{pmatrix} m_{\nu}^0 ; \quad L = \begin{pmatrix} 0 & \epsilon' \eta' & 0 \\ -\epsilon' \eta' & 0 & -3\epsilon + \eta \\ 0 & +3\epsilon + \eta & 1 \end{pmatrix} \times m_{\nu}^0$$

2 New param ( $\epsilon', \eta'$ ), but 5 new observables just in  $(\nu, \bar{\nu})$  System  $\Rightarrow$  3 New predictions for  $(\nu, \bar{\nu})$  // With  $\epsilon' = 0 \rightarrow m_{\nu} \rightarrow 0$

## $\nu$ Majorana Masses

$$M_R^{\nu} = \begin{pmatrix} x & 0 & z \\ 0 & 0 & y \\ z & y & 1 \end{pmatrix} M_R$$

Saw before,  
 $M_R \approx 10^{15} \text{ GeV}$   
Expect  $y \sim 1/10$

$$f_{ij} \nu_{iR} \nu_{jR} \langle \overline{16}_H \rangle \langle \overline{16}_H \rangle / M_{st}$$

$$f_{ij} \nu_{iR}^T \bar{C} \nu_{jR} \langle \overline{16}_H \rangle^2 / M_{st}$$

$$(M_R^{\nu})_{ij} = f_{ij} \langle \overline{16}_H \rangle^2 / M_{st}$$

Note same hier. pattern as in Dirac Sectn

Including  $m_U^0 \rightarrow$  7 param ( $\eta, \epsilon, \delta, \eta', \epsilon', m_U^0, m_D^0$ ) <sup>15</sup>  
 describing  $9 \times 4 = 36$  entries  $\rightarrow$  Will it work?

Input! Assume all param real for a moment

$$m_t^{\text{phys}} = 174 \text{ GeV}; m_c(m_c) = 1.37 \text{ GeV},$$

$$m_s(1 \text{ GeV}) = 116 \text{ MeV}, m_u, m_\tau, m_u(M_X) = 1.5 \text{ MeV}, m_e$$



$$\delta \approx 0.110, \eta \approx 0.151, \epsilon \approx -0.095,$$

$$\epsilon' = \sqrt{m_u/m_c} (m_c/m_t) \approx 2 \times 10^{-4}; \eta' = \sqrt{m_e/m_u} (m_c/m_t) \approx 4 \times 10^{-3}$$

$$m_U^0 = m_t(M_X) \approx 110 \text{ GeV}; m_D^0 \approx 1.5 \text{ GeV}$$

Majana Mass of  $\nu_R^i$ 's:  $f_{ij} 16_i 16_j \bar{16}_H \bar{16}_H / M$

$$\Rightarrow f_{ij} (\nu_R^{iT} \bar{\nu}^i \nu_R^j) < (16_H)^2 / M$$

$$M_R^{\nu} = \begin{pmatrix} x & 0 & z \\ 0 & 0 & y \\ \cancel{z} & y & 1 \end{pmatrix}$$

6 New observables.

$\rightarrow$  determined by  $m_{\nu_2}/m_{\nu_3} \approx 1/6$

$\downarrow$  Calculated  $\approx 5 \times 10^{14} \text{ GeV}$ .

# Summary on Fermion Masses & Mixings (Babu, Pati Wilczek)

## Predictions

$$m_b(m_b) \approx (4.7 - 4.9) \text{ GeV}$$

$$m(\nu_\tau) \sim \left(\frac{1}{24} \text{ eV}\right) \left(\frac{1}{2} - 2\right)$$

$$V_{cb} \approx 0.043$$

$$\text{Sin}^2 2\theta_{\nu_\mu \nu_\tau}^{\text{osc}} \approx \boxed{0.92} \leftrightarrow \boxed{0.99}$$

SMA                      LMA

$$V_{us} \approx 0.21$$

$$|V_{ub}| \approx 0.0032$$

$$m_d(1 \text{ GeV}) \approx 8 \text{ MeV}$$

$$m(\nu_\mu) \approx (2 - 10) \times 10^{-3} \text{ eV} \leftrightarrow \left\{ \begin{array}{l} \text{SMA} \sim 3 \times 10^{-3} \text{ eV} \\ \text{LMA} \approx 7 \times 10^{-3} \text{ eV} \end{array} \right\}$$

$$m(\nu_e) \sim (1 \text{ to few}) \times 10^{-3} \text{ eV}$$

consistent with the framework

$$M(\nu_R^{\tau}, \nu_R^{\mu}, \nu_R^e) \approx (10^{15}, 2 \times 10^{12}, (\frac{1}{3} - 3) \times 10^{10}) \text{ GeV}$$

Just right for leptogenesis

## Observations

$$\approx 4.2 \text{ GeV}$$

$$\approx (1/15 - 1/25) \text{ eV} \otimes$$

$$\approx 0.04$$

$$\approx 0.92 \leftrightarrow 1$$

$$\approx 0.22$$

$$\approx 0.003 - 0.004$$

$$\approx 8 - 10 \text{ MeV}$$

NOTE: Masses Necessarily Hierarchical.



16.1  
(cont'd)

$$M_{3R} \approx M_R \approx 10^{15} \text{ GeV} (\frac{1}{2} - 1)$$

$$M_{2R} \approx |y^2| M_{3R} \approx 2.5 \times 10^{12} \text{ GeV} (\frac{1}{2} - 1)$$

$$M_{1R} \approx |x - z^2| M_{3R} \approx (\frac{1}{2} - 2) \times 10^{-5} M_{3R}$$
$$\approx 10^{10} \text{ GeV} (\frac{1}{4} - 2)$$

Predictions

Writing only for  $2 \times 2$  (for simplicity)

$$U = \begin{pmatrix} c & t \\ 0 & \epsilon + \delta \\ -\epsilon + \delta & 1 \end{pmatrix} m_U^0 \quad D = \begin{pmatrix} s & b \\ 0 & \epsilon + \eta \\ -\epsilon + \eta & 1 \end{pmatrix} m_D^0$$

$$N = \begin{pmatrix} 0 & -3\epsilon + \delta \\ 3\epsilon + \delta & 1 \end{pmatrix} m_U^0 \quad L = \begin{pmatrix} \mu & \tau \\ 0 & -3\epsilon + \eta \\ 3\epsilon + \eta & 1 \end{pmatrix} m_D^0$$



$$m_b^0 \approx m_\tau^0 (1 - 8\epsilon^2) \Rightarrow m_b(m_b) \approx 4.7 \text{ TeV}$$

$$V_{cb} = \left| \sqrt{\frac{m_s}{m_b}} \left( \frac{\eta + \epsilon}{\eta - \epsilon} \right)^{1/2} - \sqrt{\frac{m_c}{m_t}} \left( \frac{\delta + \epsilon}{\delta - \epsilon} \right)^{1/2} \right| = |\delta - \eta| = 0.042$$

(0.156) (1/2.2)

Suppressed

ENHANCED  $\approx 1.8$

$$\theta_{\nu_\mu \nu_\tau}^{osc} = |\theta_{\mu\tau}^l - \theta_{\mu\tau}^{\nu}| \approx \left| \sqrt{\frac{m_\mu}{m_\tau}} \left( \frac{\eta - 3\epsilon}{\eta + 3\epsilon} \right)^{1/2} + \sqrt{\frac{m_{\nu_2}}{m_{\nu_3}}} \right|$$

$$= 0.437 + \sqrt{m_{\nu_2}/m_{\nu_3}} \approx 0.3$$

Expt

$$0.92 \leftrightarrow 1$$

$$\Rightarrow \sin^2 2\theta_{\nu_\mu \nu_\tau}^{osc} = 0.92 \leftrightarrow 0.99$$

$$m_{\nu_2}/m_{\nu_3} = 1/15 \leftrightarrow 1/7$$

SMA OR LMA ?

fig 16i 16j  $\bar{16}_H \bar{16}_H / M$

Just with standard See-saw  $\nu_L$ -masses,  
SMA rather generic

$$m(\nu_L^e) \sim 2 \times 10^{-5} - 2 \times 10^{-6} \text{ eV} // m(\nu_L^\mu) \sim 3 \times 10^{-3} \text{ eV}$$

$$\Theta_{\nu_e \nu_\mu}^{\text{osc}} = \Theta_{e\mu}^L - \Theta_{e\mu}^\nu \approx 0.05$$

-----  
situation alters once allow for direct Maj. masses of  $\nu_L$ 's - Most likely to arise through Higher Dim. op. involving GUT & EW VEV's - Through tiny  $\sim 10^{-3} \text{ eV}$  entries  $\rightarrow$  IMPORTANT For  $(\nu_e - \nu_\mu)$

$$W \supset g_{12} \underbrace{16_1 16_2}_{\nu_L^e \nu_L^\mu} \underbrace{16_H 16_H}_{\langle \nu_{RH} \rangle / M_{\text{GUT}}} \underbrace{10_H 10_H}_{\nu_u^2} / M_{\text{GUT}}^3$$

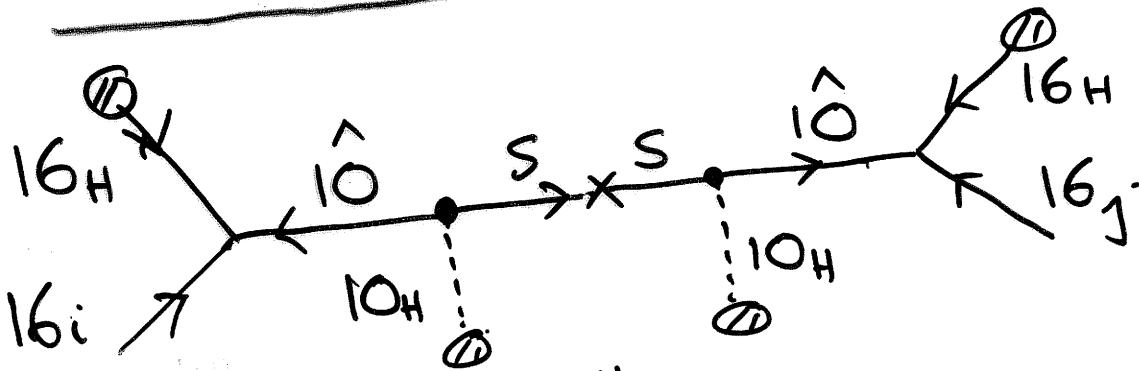
$$\sim g_{12} (\nu_L^e \nu_L^\mu) (1.5 - 6) \times 10^{-3} \text{ eV} \quad (\langle 16_H \rangle \approx (1-2) M_{\text{GUT}})$$

$$\begin{bmatrix} \nu_L^e & \nu_L^\mu \\ \approx 0 & (3-4) \end{bmatrix} \times 10^{-3} \text{ eV} \Rightarrow \begin{bmatrix} \Theta_{\nu_e \nu_\mu}^\nu \approx 1/2 \\ \sin^2 2\Theta_{\nu_e \nu_\mu}^{\text{osc}} \approx 0.7 \end{bmatrix} \text{ quite plausible}$$

Thus LMA not strictly a prediction, but perfectly plausible within the framework.

16.4.24.3

# QUANTUM-SCALE EFFECTIVE OPERATORS



$$g_{ij} |16i\rangle |16j\rangle |16H\rangle |16H\rangle |10H\rangle |10H\rangle / (M_{10}^{\wedge 2} M_S)$$



$\theta_{13}$

$$m(\nu_L^e \nu_L^e)_{\text{Non-Seesaw}} \sim (2-6) \times 10^{-3} \text{ eV}$$

$$\theta_{13} \sim \frac{(2-6) \times 10^{-3} \text{ eV}}{5 \times 10^{-2} \text{ eV}}$$

$$\sim 0.03 - 0.1$$

$\nu$ -less  $2\beta$  decay:  $\Delta L = \pm 2$

$$m_{ee} = \left| \sum_i m_i U_{ei}^2 \right|$$

$$m_1 \sim \text{few} \times 10^{-3} \text{ eV}, \quad m_2 \approx (6-8) \times 10^{-3} \text{ eV}$$

$$m_3 \approx 5 \times 10^{-2} \text{ eV}$$

$$\theta_{12} \approx 1/2, \quad \theta_{13} \sim 0.03 - 0.1$$

$\Downarrow$

$$m_{ee} \sim (1 \text{ to } 6) \times 10^{-3} \text{ eV}$$

## Summary on Fermion Masses & Mixings in the $E(6)/SO(10)$ Framework

---

Given the bizarre pattern of masses & mixings of quarks, charged leptons and neutrinos, it seems remarkable that the simple pattern of fermion mass matrices\*, motivated in large part by the group th of  $E(6)/SO(10)$  and the assumption of minimality of the system of  $\wedge$  Higgs, makes  $\gamma$  predictions in agreement with observation.

→ Study Proton Decay // Leptogenesis // CP within this framework.

\* Need to understand the origin of flavor symmetries.  
→ Hierarchical entries.

# V CP & Flavor Violations

Prepared by:

19

Date:

Question! Can observed CP & Flavor Viols. emerge consistently within the SUSY  $G(224)/SO(10)$ -Framework, while preserving its successes wrt fermion masses &  $\nu$ -oscills?

→ A Non-trivial Challenge (Many  $SO(10)$ -models do not satisfy both sets of constraints).

## ⑤ RESULTS (Babu, Pati, Rastogi, <sup>Ph/0410200</sup> $\wedge$ )

Find For natural phases ( $\sim 10-40\%$ ) in the Dirac mass-parameters ( $\sigma, \eta, \epsilon', \dots$ ), of the same  $G(224)/SO(10)$ -framework, can get observed CP & Flavor Viols while preserving the successes in fermion masses &  $\nu$ -oscillations!

(a) For phases in a natural range, get

$$\hat{\eta}_W \approx 0.30 - 0.37; \hat{\rho}_W \approx 0.15 - 0.18 \quad \text{BPR-Values}$$

close to SM CKM }  $\bar{\eta}_W \approx 0.34 \quad \bar{\rho}_W \approx 0.18$  SM CKM Values

New With Phases

Babu, Pati, Rastogi

$$\eta = 0.12 - 0.05i$$

$$\sigma = 0.1 - 0.012i$$

$$\epsilon = -0.0954$$

$$\epsilon' = 1.8 \times 10^{-4} e^{i[0.4\pi]}$$

$$\zeta_{22}^d = 1.1 \times 10^{-2} e^{-i[0.8\pi]}$$

$$\eta' = 2.5 \times 10^{-3}$$

No Phases

Babu, Pati, Wilczek

$$\eta = 0.15$$

$$\sigma = 0.11$$

$$\epsilon = -0.095$$

$$\epsilon' = 2 \times 10^{-4}$$

$$\zeta_{22}^d \approx (1/3) \times 10^{-2}$$

$$\eta' = 4.4 \times 10^{-3}$$

Preserves all the Predictions of BPW  
for fermion masses/mixings &  $\nu$ -OSC

AND USING SUSY YIELDS RIGHT CP  
& Flavor Viols:

$\Delta m_K, \epsilon_K, \Delta m_{B_d}, \Delta m_{B_s},$   
 $S(B_d \rightarrow J/\psi K_S), S(B_d \rightarrow \phi K_S), b \rightarrow s\gamma,$

→ MANY PREDICTIONS:  $\mu \rightarrow e\gamma, \tau \rightarrow \mu\gamma, (EDM)_{e,n}$



## Ample Further Tests

- ①  $B(\mu \rightarrow e\gamma) \approx 10^{-9} - 10^{-13}$  ;  $< \overset{\text{Expt}}{10^{-11}}$
- ②  $B(\tau \rightarrow \mu\gamma) \approx 10^{-7} - 10^{-9}$  ;  $< 7 \times 10^{-8}$   
(BABAR)
- ③  $(edm)_n \approx (1.6 - 1) \times 10^{-26} \text{ ecm}$   $< 6.3 \times 10^{-26} \text{ ecm}$   
 $\tan\beta = 5, 10$
- ④  $(edm)_e \approx \frac{1.1 \times 10^{-28} \text{ ecm}}{\tan\beta}$

## The Two Notable Missing Pieces

⑤ SUSY Particles  $\lesssim 1 \text{ TeV} \rightarrow \text{LHC}$

⑥ PROTON DECAY

$$\left. \begin{aligned} \bar{p}^+ (p \rightarrow e^+ \pi^0) &\approx 10^{35 \pm 1} \text{ yrs} \\ \bar{p}^+ (p \rightarrow \bar{\nu} k^+) &\lesssim 10^{34} \text{ yrs.} \end{aligned} \right\}$$

NEED MEGATON SIZE DETECTOR

## VI $\nu$ Masses $\leftrightarrow$ Matter-Antimatter Asymmetry

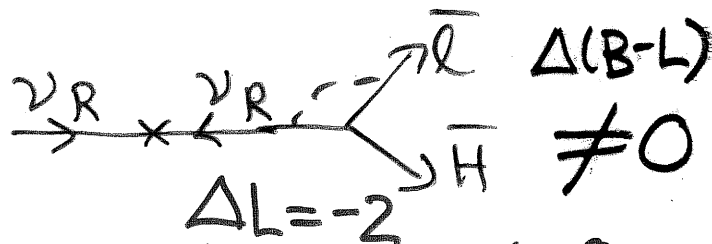
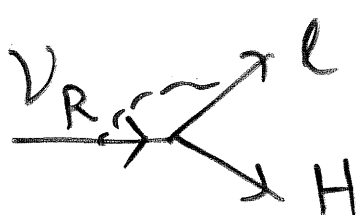
Starting With a baryon-Symmetric Universe at  $t=0$  (Big Bang) or just at the beginning of reheating following inflation, what <sup>eventually</sup> caused

$$Y_B \equiv \frac{n_B - n_{\bar{B}}}{s} = 9 \times 10^{-11} \quad (\text{WMAP})$$

Sakharov (Need (1967) )  $\left\{ \begin{array}{l} \textcircled{1} C \ \& \ CP \ \text{Non-Conservation} \\ \textcircled{2} \Delta B \neq 0 \ \text{Processes} \\ \textcircled{3} \text{Thermodynamic Inequilibrium} \end{array} \right.$

### Most Promising Explanation

$$M(\nu_R) \nu_R^T \nu_R + \text{h.c.} \quad (\Delta L = \pm 2)$$



Mukugita,  
Yanagida

$$\Rightarrow \boxed{n_l \neq n_{\bar{l}}} \quad \text{With } T_{RH} \sim 10^6 - 10^8 \text{ GeV}$$

$$Y_L \neq 0 \xrightarrow{SU(2) \times U(1)} Y_B \neq 0 \quad (\text{Kuzmin, Rubakov, Sapozhnikov})$$

G(224)/SO(10) Theory

JCP, Phys. Rev 2003, hep-ph/0209160

Note All Vertices & Dirac Phases determined

$$\nu_R \rightarrow l + H ; \bar{l} \bar{H} \} \Delta(B-L) \neq 0$$

$$\Delta B = \left[ \frac{\Delta(B+L)}{2} \right] + \frac{\Delta(B-L)}{2}$$

ERASED BY EW SPHALERONS

$$\frac{(Y_B)_{\text{Thermal}}}{\sin^2 \phi_{21}} \approx (10-30) \times 10^{-11} \quad \text{THEORY}$$

$$\text{Non-Thermal} \approx (10-100) \times 10^{-11} \quad \text{THEORY}$$

$$\text{WMAP} \approx 9 \times 10^{-11} \quad \text{OBSERVED}$$

Agrees with obs. For natural values of  $\phi_{21} \approx 1 - 1/10$  !

A Simple & Unified Picture of fermion masses,  $\nu$  oscillations, CP Violation & Baryo-Leptogenesis IN FULL ACCORD WITH OBSERVATIONS !

EXISTENCE OF  $\nu_R$ , B-L & SU(4)-Color CRUCIAL ! AS ALSO EW SU(2)<sub>L</sub> x U(1)<sub>Y</sub> !

## Summary

2 Alt. Scenarios  $\begin{cases} \rightarrow \text{Thermal} \\ \rightarrow \text{Non-Thermal} \end{cases}$

within  $G(224)/SO(10)$  Framework

↓  
 Unified description of not only fermion masses &  $\nu$  oscill (consistent with maximal atmospheric & LMA) but also of baryogenesis via leptogenesis.

- The Exist of RH neutrinos
- B-L Local Symmetry
- q-l unif  $SU(4)$ -Color
- See Saw Mechanism

& • SUSY unif Scale

Crucial roles in this unified description.

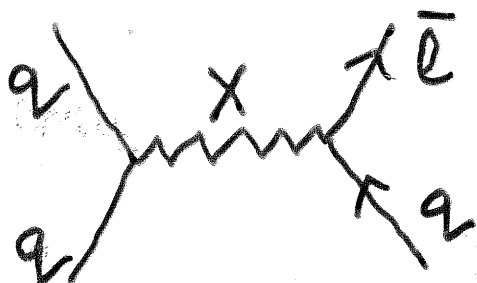
→ Relevance of  $G(224)/SO(10)$  in 4D below String Scale.

→ CP Phases → Sign of  $Y_B$  a Challenge

# VII PROTON DECAY : THREE MECHANISMS

## ① d = 6 Gauge Med. SUSY SU(5)/SO(10)

$M_X \approx M_Y$   
 $\approx 10^{16} \text{ GeV}$   
 $(1 \pm 25\%)$



$\Rightarrow \text{Amp } (p \rightarrow e^+ \pi^0)$   
 $\propto g^2 / M_X^2$

Chiral Lag Param (D+F)  $\approx 1.25$  // AR  $\approx 3.4$

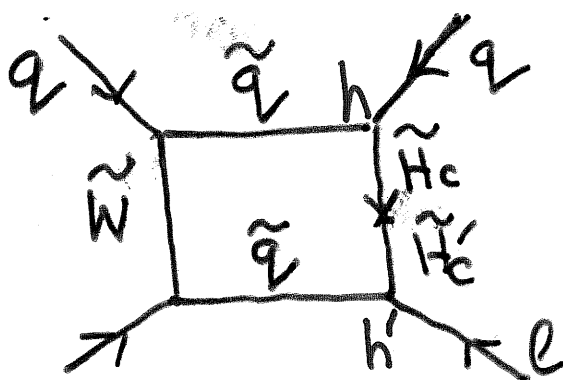
$H \approx 0.01 \text{ GeV}^3$   
 $(\frac{1}{\sqrt{2}} - \sqrt{2})$

$\tau^{-1}(p \rightarrow e^+ \pi^0) \approx 10^{35 \pm 1} \text{ yrs (Theory)}$

Can be as short as  $\sim 10^{34} \text{ yrs.}$

## ② SUSY SU(5) / SO(10) Standard d=5

Color Triplet Higgsino Med Sakai, Yanagida, Weinberg



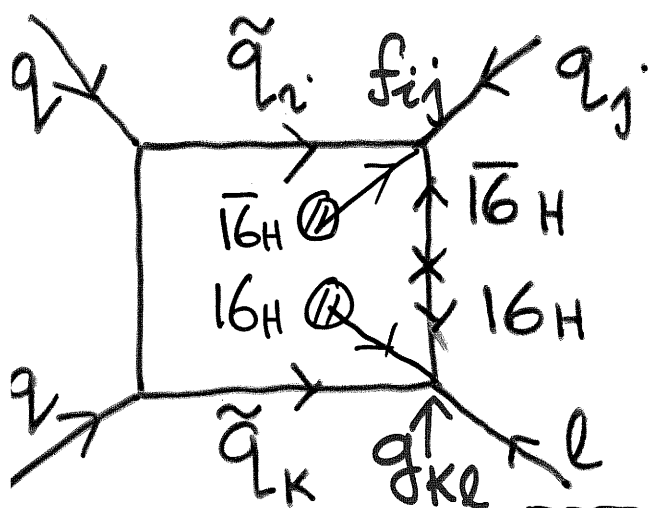
$\tilde{H}_c \subset (10_H)_{SO(10)}$   
 $= (2, 2, 1) + (1, 1, 6)$

$\text{Amp} \propto (h h' / M_{H_c}) (m_{\tilde{W}} / m_{\tilde{q}}^2) \propto 2$

- $p \rightarrow \bar{\nu} k^+$  (dominant)
  - $\rightarrow \mu^+ k^0$  (Suppressed, BR  $\sim 10^{-5}$ )
- Std. d=5

### ③ New $d=5$ $\nu$ -Mass Related operators

Babu, Pati, Nilczek



$$f_{ij} 16_i 16_j 16_H 16_H / M$$

Major Masses of  $\nu_R$ 's

Generically  $16_i 16_H$  in  $45$  &  $1$  of  $SU(10)$

$$\Rightarrow g_{kl} 16_k 16_l 16_H 16_H / M$$

Indep of  $\tan\beta$

$$\Rightarrow V_{CKM} \neq 1$$

$$\Gamma^{\bar{l}}(p \rightarrow \bar{\nu} K^+)_{\text{New } d=5} \approx 10^{-33} - 10^{-34} \text{ yrs} \quad \text{SUSY } SU(10) \text{ or } G(224)$$

$$BR(p \rightarrow \mu^+ K^0)_{\text{New } d=5} \approx (10 - 30)\%$$

Note these contributions from new  $d=5$  would generically be present, even if "standard"  $d=5$  (with color triplets  $C 1Q_H$ ) absent, as in SUSY  $G(224)$ !

The  $\mu^+ K^0$  mode a signature of this mechanism.

# d = 5 proton decay rate calculation\*

Babu, Pati, Wilczek (99) // Dermisek, Raby (2000) //  
 Lucas & Raby // Pati (2001 - Ph/0106082) //  
 June 2003 - Ph/0305221 // KEK talk Ph/0407220 //

Incorporate Dep. of d = 5 p-decay amp. on

- Masses & Mixings of All Fermions

- [Unif. Scale Threshold Effects Incl. D-T  
 Splitting  $\leftrightarrow$  Natural Coupling Unif.]

- Take

$$\textcircled{1} |\beta_H| \approx |\alpha_H| = (0.01 \text{ GeV}^3) (\sqrt{2} - \sqrt{2})$$

$$\textcircled{2} \tan \beta \gtrsim 3$$

$$\textcircled{3} A_L \approx 0.32 \quad (2 \text{ Loop})$$

$$\textcircled{4} A_S \approx 0.93$$

$$\textcircled{5} m_{\tilde{q}} \approx 1.2 \text{ TeV} (\sqrt{2} - 2) \quad (\text{Focus Point})$$

$$\textcircled{6} (m_{\tilde{u}}/m_{\tilde{q}}) \approx 1/6 (\sqrt{2} - 2)$$

Lattice // Domain  
 Wall Fermions +  
 Non-Pert Ren  
 Quenching error small  
 $|\alpha_H| \approx |\beta_H| \approx 0.01 \text{ GeV}^3$   
 Aoki et al (2006)

\* Dimopoulos, Raby, Wilczek (82) // Ellis, Nanopoulos, Rudaz //  
 Nath, Chamseddine, Arnowitt // Nath, Arnowitt (97) //  
 Hisano, Murayama, Yanagida (93) // Hisano (2004) //  
 Babu, Barr (95) //

# Doublet-Triplet Splitting in SO(10) & $M_{eff}$

Dimopoulos, Wilczek // Babu, Barr  
Babu, Pati, Wilczek

$$W_H = \lambda 10_H 45_H 10'_H + M_{10} 10'^2_H + M_{16} 16_H \bar{16}_H + \lambda' \bar{16}_H \bar{16}_H 10_H$$

$$\langle 45_H \rangle = (a, a, a, 0, 0) \times \tau_2 \quad ; \quad a \sim M_{GUT}$$

$$\langle 45_H \rangle \propto B-L \rightarrow \langle (1, 1, 15) \rangle = (a, a, a, -3a)$$

$$(\bar{5}_{10H} \quad \bar{5}_{10'H} \quad \bar{5}_{16H}) \begin{bmatrix} 0 & \lambda \langle 45_H \rangle & \lambda' \langle \bar{16}_H \rangle \\ -\lambda \langle 45_H \rangle & M_{10} & 0 \\ 0 & 0 & M_{16} \end{bmatrix} \begin{pmatrix} 5_{10H} \\ 5_{10'H} \\ 5_{16H} \end{pmatrix}$$

$\Rightarrow$  One pair of Higgs doublets light, triplets heavy

$$H_u = 10_u \quad ; \quad H_d = \cos\gamma 10_d + \sin\gamma 16_d$$

$$1) \quad \tan\gamma = \lambda' \langle \bar{16}_H \rangle / M_{16} \Rightarrow \boxed{\tan\beta = (\cos\gamma) (m_t/m_b)}$$

$$b) \quad \boxed{A(d=5)_{std} \approx (h^2_{33}/M_{eff}) (2 \times 10^5)}$$

$$\boxed{M_{eff} \equiv (\lambda a)^2 / M_{10'} \approx M_{GUT}^2 / M_{10'}}$$

$$\Rightarrow M_{eff} \neq M_{HC} \quad ; \quad \text{If } M_{10'} < M_{GUT} \Rightarrow \boxed{M_{eff} > M_{GUT}}$$



$M_{eff}$  Gets Bounded From above by 25.2  
Threshold Corr. & Demand of Natural Coupling Unif

$$\Delta\alpha_3(m_Z)_{DT} = \frac{\alpha_3(m_Z)^2 (9/7)}{2\pi} \ln \frac{M_{eff} \cos\theta}{M_X} > 0$$

$$\Delta\alpha_3(m_Z)_{45H} \approx -4\%$$

$$\Delta\alpha_3(m_Z)_{\text{gauge multiplet}} \approx -1.5 \text{ to } -2\%$$

$$\Delta\alpha_3(m_Z)_{16_H + \bar{16}_H} \approx \pm 2\%$$

For Coupling Unif need

$$\begin{aligned} \Delta\alpha_3(m_Z)_{\text{Net}} &\approx DT + 45H + \text{Gauge} + (16, \bar{16})_H \\ &\approx -5 \text{ to } -8\% \end{aligned}$$

$\therefore$  For coupling ~~Unif~~ need

$$\Delta\alpha_3(m_Z)_{DT} < 2\%$$

$$\boxed{M_{eff} < 2 \times 10^{18} \text{ GeV}}$$

UPPER LIMIT ON PROTON LIFETIME

# The Case of ESSM (Extended Supersymm.

Standard Model)

Babu, Pati (96,97)

Babu, Pati, Stremnitzer (92,93)

Introduce 1 Pair of Vector-Like Families  
( $16_V + \bar{16}_V$ ) at TeV-Scale & Superpartners.

→ [3 chiral  $16_i$  + ( $H_u + H_d$ ) + ( $16_V + \bar{16}_V$ )  
[at  $\sim$  TeV] + SUSY Partners at 1 TeV

consistent with (i)  $\nu$  Counting // (ii) Precision  
EW Tests

A priori Motivations:

(i) Removal of Unif. Mismatch

$$M_{\text{GUT}} \sim (\frac{1}{2} - 2) \times 10^{17} \text{ GeV}$$

$$M_{\text{string}} \approx \downarrow 3.6 \times 10^{17} \text{ GeV (Pent.)}$$

(ii) Dilaton stabilization

(iii)  $(g-2)_\mu$  ?

MSSM  
0.125 - 0.13

(iv)  $\alpha_3(m_Z^0)_{\text{ESSM}} \approx 0.112 - 0.118$   
EXPT  $\approx 0.118 \pm 0.003$

(v) Enhances Proton Decay Rate  
 $d=5$

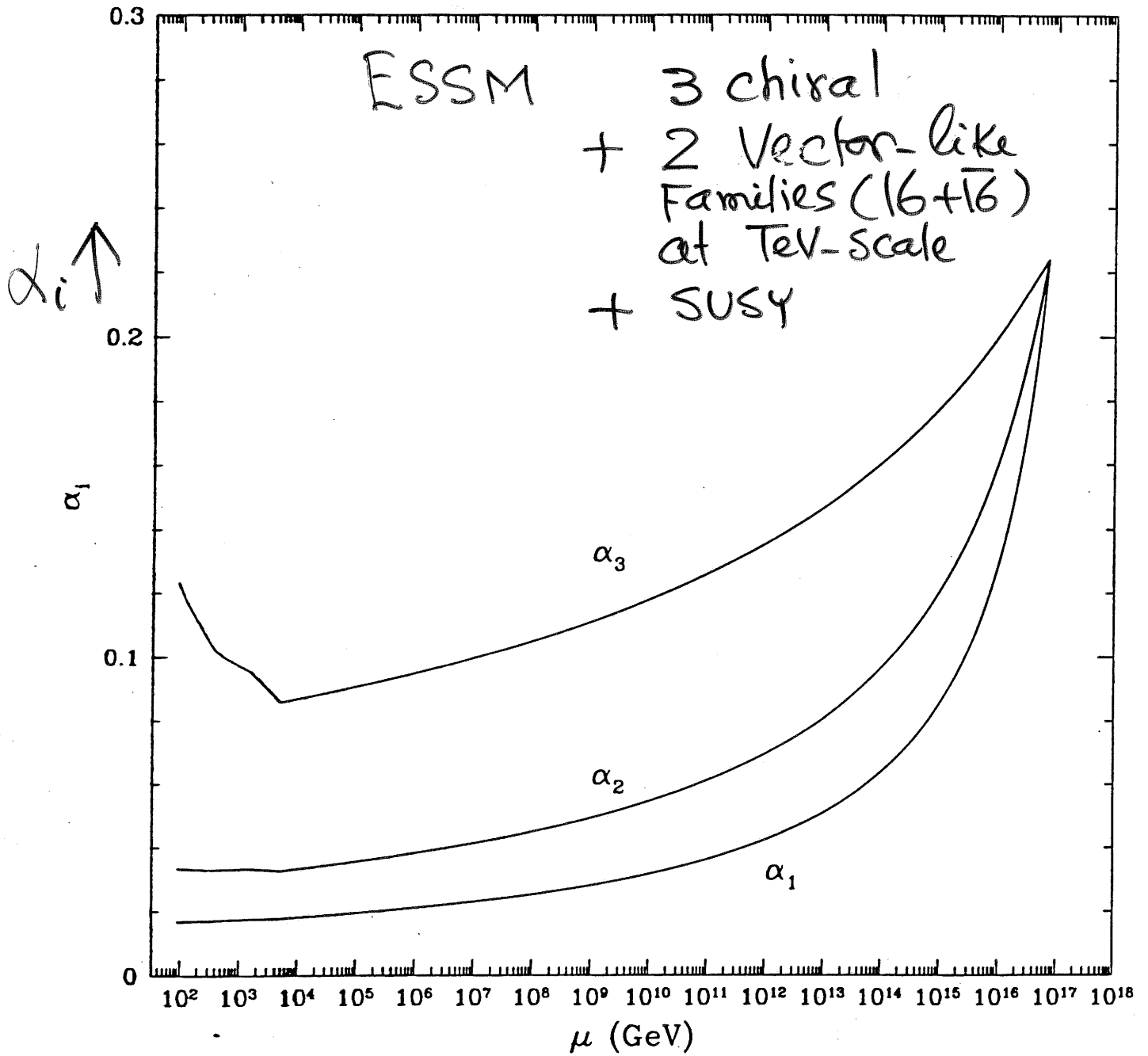


Fig. 3 (Case 3)

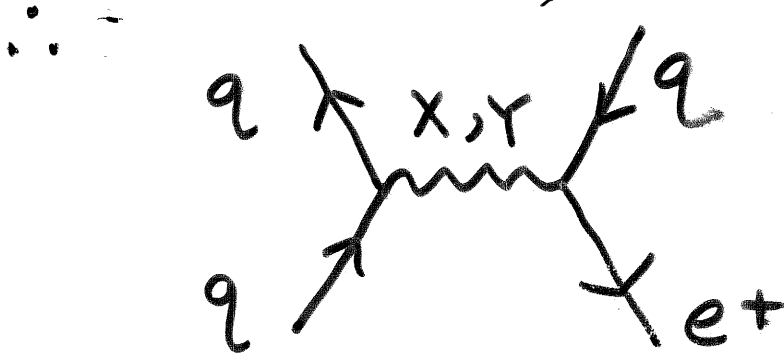
Babu, Pati. Phys. Lett (96)

SUSY SO(10)/ESSM (Also SU(5))

$\alpha_{GUT}$  increases by 6-7 (Rel to MSSM)

$M_{GUT}$  " by 2.5 - 5

$$\tilde{\Gamma}^{\prime}(p \rightarrow e^+ \pi^0) \underset{\substack{\text{ESSM/} \\ \text{SO}(10) \\ \propto \text{SU}(5)}}}{\approx} (1-17) \tilde{\Gamma}^{\prime}(p \rightarrow e^+ \pi^0) \underset{\substack{\text{MSSM/} \\ \text{SO}(10) \\ \propto \text{SU}(5)}}{}$$



SuperK Limits (50 KT Water Cherenkov) <sup>27</sup>  
(2005)

$$\bar{\Gamma}^1 (p \rightarrow \bar{\nu} K^+) < 2.3 \times 10^{33} \text{ yrs}$$

$$\bar{\Gamma}^1 (p \rightarrow e^+ \pi^0) < 6.5 \times 10^{33} \text{ yrs}.$$

---

# Summary of Results

$$1) \Gamma^{-1}(\bar{\nu} K^+) \lesssim 0.33 \times 10^{33} \text{ yrs} \left\{ \begin{array}{l} \text{MSSM} \\ \downarrow \\ \text{SU}(5) \end{array} \right\} \rightarrow \text{Excluded.}$$

---

$$2) \Gamma^{-1}(\bar{\nu} K^+) \lesssim \frac{2.5}{1.5} \times 10^{33} \text{ yrs} \left\{ \text{MSSM} \rightarrow \text{SO}(10) \right\} \begin{array}{l} \text{Tightly} \\ \text{Constrained} \\ \downarrow \\ \text{Disfavored.} \end{array}$$

---

$$3) \Gamma^{-1}(\bar{\nu} K^+) \approx (10^{33} - 10^{34}) \text{ yrs.} \left\{ \begin{array}{l} \text{Most Plausible} \\ \text{range For} \\ \text{ESSM} \rightarrow \text{SO}(10) \end{array} \right\}$$

Can realize even for nearly central values of  
Param // & also with  $m_{\tilde{q}} < 1 \text{ TeV}$  //  
 $\tau_{\beta} \approx 3-20$  //

---

④ MSSM or ESSM  $\rightarrow$  String G(224) ONLY  
New op.  
Contribute.

$$\Gamma^{-1}(\bar{\nu} K^+) \begin{array}{l} \text{Nearly Central} \\ \text{New } d=5 \end{array} \approx (1-6) \times 10^{33} \text{ yrs.}$$

$$\Gamma^{-1}(\bar{\nu} K^+) \lesssim 10^{34} \text{ yrs.}$$

---

Thus Conclude  $\Downarrow$

$$\Gamma^{-1}(\bar{\nu} K^+) \lesssim (\frac{1}{2} - 2) \times 10^{34} \text{ yrs.}$$

SUSY SO(10)/G(224)



$$\tau^{-1}(p \rightarrow e^+ \pi^0) \approx 10^{35 \pm 1} \text{ yrs}$$

$$\tau^{-1}(p \rightarrow \bar{\nu} K^+) \approx 10^{30-34} \text{ yrs.}$$

SuperK

$$> 6 \times 10^{33} \text{ yrs}$$

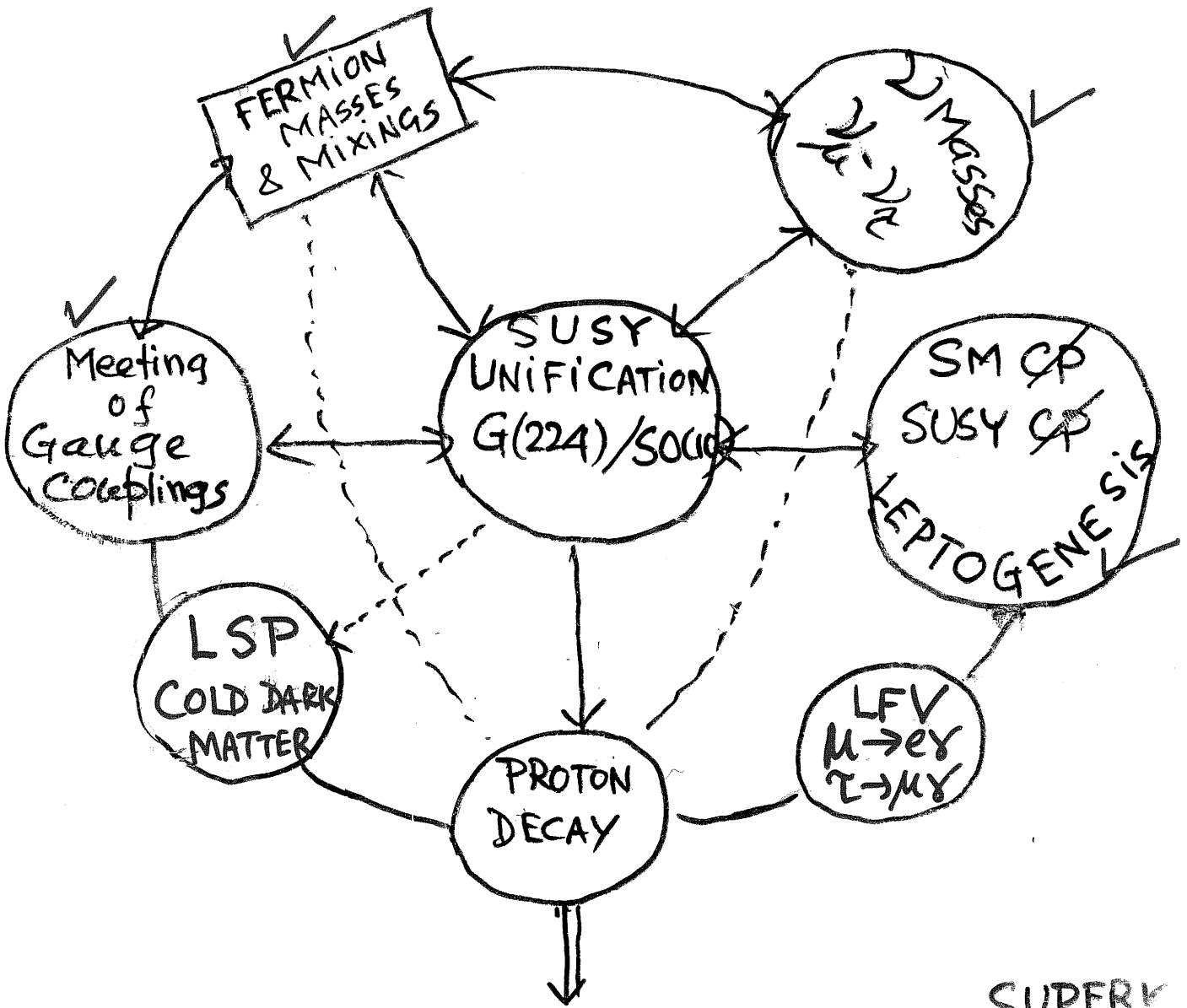
$$> 2.3 \times 10^{33} \text{ yrs}$$

Need Improvement by Factor 10



A Megaton Size Detector

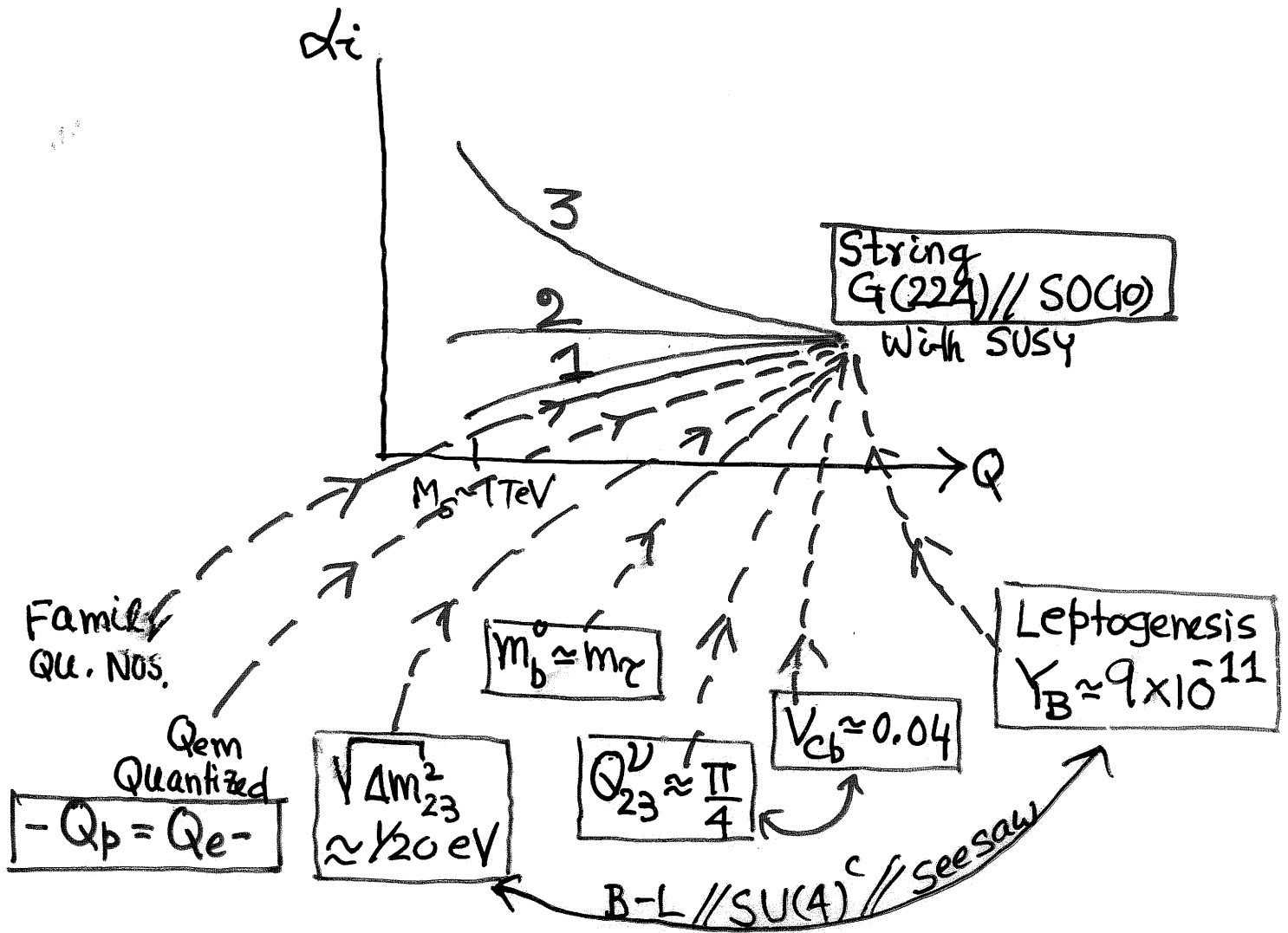
# UNIFICATION LINKS



$$\begin{aligned}
 \tau(p \rightarrow e^+ \pi^0) &\approx 10^{35 \pm 1} \text{ yrs} > 6 \times 10^{33} \text{ yrs} \\
 \tau(p \rightarrow \bar{\nu} K^+) &\approx 10^{30-34} \text{ yrs} > 2 \times 10^{33} \text{ yrs}
 \end{aligned}$$

PROTON DECAY AND SUPERSYMMETRY  
 ARE THE TWO MISSING LINKS  
 → NEED A MEGATON SIZE DETECTOR





All these features hang together neatly within a single unified framework → Hard to believe this can be a mere coincidence.

Rendezvous Incomplete → Two Missing Links

- |                  |                            |
|------------------|----------------------------|
| 1) Supersymmetry | → LHC                      |
| 2) Proton Decay  | → Need Next Gen. Detector. |

# 3 Important New Ideas Beyond Grand Unif

① Supersymmetry

② Cosmology // Inflation //

GUT  $\downarrow$  Symm Breaking

Monopole Soln

Leptogenesis  $\rightarrow$  Baryogenesis

③ String/M Th  $\rightarrow$  Extra Dim  $\rightarrow$  Compactification

Grand Unif A VERY USEFUL BRIDGE

BETWEEN STRING THEORY ( $D=10/11$ )



( $D=4$ )

PLANCK/string scale ( $10^{18}$  GeV)



3 Families,  
Hierarchical  
Yukawas,  
CP PHASES

----- $\rightarrow$  **GUT OR EFF. SUSY  $G(224)$**

$\downarrow$  SSB

$G(213)$  (SM)

PHENOMENOLOGY

----- $\rightarrow$  ( $\sim 2 \times 10^{16}$  GeV)

⊗ MANY INTERESTING PIECES OF WORK  
DERIVING EFF. SUSY  $G(224)$  with 3 Fam.

# Puzzles & Challenges

To realize a TRULY UNIFIED THEORY  
WITH GOOD QUANTUM GRAVITY - That  
is predictive, explaining at least  
Some of the major puzzles

Be it String/M Th or something  
not yet known

- 1) Why 3 Families?
- 2) Hierarchical Masses/Mixings/CP?
- 3) Why live in  $D = 4$ ?
- 4) Dark Energy  $\rightarrow$  Cosmological  
Constant  
 $(\Lambda_{\text{Cosm}}/M_{\text{Pl}})^4 \sim 10^{-120}$  ?