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Symmetries Conservation Laws

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Symmetries - Conservation Laws. (1)

$$S = \int \mathcal{L}(\varphi, \partial_\mu \varphi) d^4x$$

Four-dimensional space-time.

$$V_\mu = \{V_0, \underline{V}\} \quad V_\mu V^\mu = -\lambda^2$$

An example of a Lagrangian scalar field

$$\mathcal{L} = \frac{1}{2} (\partial_\mu \varphi \partial^\mu \varphi) - \frac{m^2}{2} \varphi^2$$

$$= \frac{1}{2} \left((\partial_0 \varphi)^2 - (\underline{\nabla} \varphi)^2 - m^2 \varphi^2 \right)$$

⇒ Derive the equation of motion K.G equation

~~$$\frac{\partial^2 \varphi}{\partial t^2} - \nabla^2 \varphi + m^2 \varphi = 0$$~~

$$\frac{\partial^2 \varphi}{\partial t^2} - \frac{\partial^2 \varphi}{\partial x^2} + m^2 \varphi = 0$$

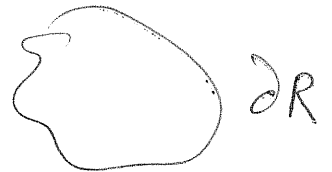
$$\Rightarrow \omega^2 = m^2 + k^2$$

Discuss → Lagrangian → why and how.

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Formalism ;

$$x^\mu \rightarrow x'^\mu = x^\mu + \delta x^\mu$$



$$\varphi(x) \rightarrow \varphi'(x) = \varphi(x) + \delta\varphi$$

all variations are zero at the surface.

Also assume

$$\mathcal{L} = \mathcal{L}(\varphi, \partial_\mu \varphi, x^\mu) \quad \text{some interaction with an external system.}$$

$\delta\varphi(x)$ is only the functional variation in φ

$$\varphi'(x') = \varphi(x) + \Delta\varphi$$

$$\begin{aligned} \Delta\varphi &= \varphi'(x') - \varphi(x') + \varphi(x') - \varphi(x) \\ &= \delta\varphi + (\partial_\mu \varphi) \delta x^\mu \end{aligned}$$

$$\begin{aligned} \delta S &= \int \mathcal{L}(\varphi', \partial_\mu \varphi', x'^\mu) d^4 x' - \int \mathcal{L}(\varphi, \partial_\mu \varphi, x^\mu) d^4 x \\ d^4 x' &= J(x'/x) d^4 x \end{aligned}$$

$$\begin{aligned} \Rightarrow \quad \frac{\partial x'^\mu}{\partial x^\lambda} &= \delta^\mu_\lambda + \partial_\lambda \delta x^\mu \\ J(x'/x) &= 1 + \partial_\mu (\delta x^\mu) \end{aligned}$$

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$$\delta S = \int (\delta \mathcal{L} + \mathcal{L} \delta x^\mu) d^4x$$

$$\delta \mathcal{L} = \frac{\partial \mathcal{L}}{\partial \varphi} \delta \varphi + \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \frac{\delta (\partial_\mu \varphi)}{\partial_\mu (\delta \varphi)} + \frac{\partial \mathcal{L}}{\partial x^\mu} \delta x^\mu$$



$$\delta S = \delta S^1 + \delta S^2$$

$$\delta S^1 = \int_R \left[\frac{\partial \mathcal{L}}{\partial \varphi} - \partial_\mu \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \right) \right] \delta \varphi d^4x$$

$$\delta S^2 = \int_{\partial R} \left[\frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \delta \varphi + \mathcal{L} \delta x^\mu \right] d\sigma_\mu$$

From assumption : At ∂R , $\delta \varphi = 0 = \delta x^\mu$
 $\therefore \delta S^2 = 0$

For stationary action,

then

$$\delta S^1 = 0$$

$$\Rightarrow \frac{\partial \mathcal{L}}{\partial \varphi} = \frac{\partial}{\partial x^\mu} \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \right)$$

But Is that All?

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An equally powerful use of the variation principle is the exploration of Conservation laws.

R arbitrary - No requirement on $\delta\varphi, \delta x^\mu$ to be zero outside R .

$$\Delta\varphi = \delta\varphi + (\partial_\nu\varphi)\delta x^\nu$$

$$\begin{aligned}\delta S^2 &= \int_{\partial R} \left[\frac{\partial \mathcal{L}}{\partial(\partial_\mu\varphi)} [\delta\varphi + \partial_\nu\varphi\delta x^\nu] \right] d\sigma_\mu \\ &= \int_{\partial R} \left(\frac{\partial \mathcal{L}}{\partial(\partial_\mu\varphi)} \partial_\nu\varphi - \eta^{\mu\nu} \mathcal{L} \right) \delta x^\nu d\sigma_\mu\end{aligned}$$

$$T^{\mu\nu} = \frac{\partial \mathcal{L}}{\partial(\partial_\mu\varphi)} \delta^\nu\varphi - \eta^{\mu\nu} \mathcal{L}$$

$$\delta S^2 = \int_{\partial R} \left[\frac{\partial \mathcal{L}}{\partial(\partial_\mu\varphi)} \Delta\varphi - T^{\mu\nu} \delta x^\nu \right] d\sigma_\mu$$

S invariant under a group of transformations

$$\Delta x^\mu = X^\mu_\nu \delta\omega^\nu$$

$$\Delta\varphi = \varphi_\mu \delta\omega^\mu$$

$\delta\omega^\nu$ are infinitesimal parameters

$$\delta S^2 = \int \left[\frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi)} \tilde{\Phi}_\nu - \eta^\mu_\alpha X^\alpha_\nu \right] \frac{d\omega^\nu}{d\tau} d\tau_\mu$$

arbitrary

$$\delta S^2 = 0 \Rightarrow$$

$$\int_{\partial R} J^\mu_\nu d\tau_\mu = 0$$

Gauss's theorem

$$\partial_\mu J^\mu_\nu = 0$$

$$J^\mu_\nu = \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi)} \tilde{\Phi}_\nu - \eta^\mu_\alpha X^\alpha_\nu$$

Consequences:

$$\partial_t J^0_\nu + \nabla \cdot \tilde{J}_\nu = 0$$

$$\frac{d}{dt} \int J^0_\nu d^3x = - \int d^3x \nabla \cdot \tilde{J}_\nu = \text{Surface term} \Rightarrow \text{far away}$$

Conservation Law

$$Q_\nu = \int J^0_\nu d^3x$$

Noether Theorem

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A Simple Example

Translation in Space - Time

$$\Delta x^{\mu} = \epsilon^{\mu} \quad \Delta \phi = 0$$

$$X^{\mu}{}_{\nu} = \delta^{\mu}{}_{\nu} \quad \phi_{,\mu} = 0$$

The conserved current now, is

$$J^{\mu}{}_{\nu} = -T^{\mu}{}_{\nu}$$

Conservation law \Rightarrow

$$\frac{d}{dt} \int T^0{}_{\nu} d^3x = 0$$

$\int T^0{}_{\nu} d^3x = P_{\nu}$ is the four moment

$$P_{\nu} = \{ E, \vec{P} \}$$

Invariance ~~due~~ on Space - Time
Translations \Rightarrow Conservation of
Energy and Momentum.

The Sine - Gordon Equation

$$\frac{\partial^2 \phi}{\partial t^2} - \frac{\partial^2 \phi}{\partial x^2} + \frac{1}{b^2} \sin b\phi = 0 \quad (1)$$

scalar - field, a nonlinear wave function.

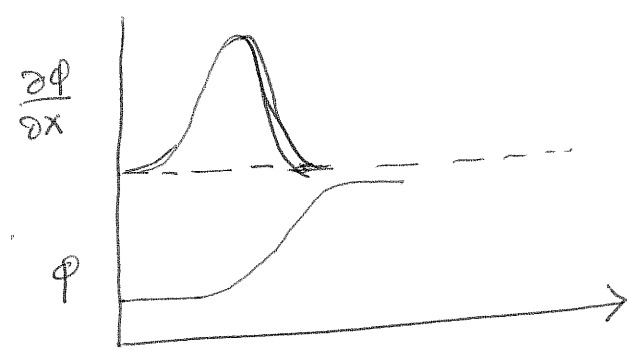
$$\phi(x, t) = f(x - vt) = f(s)$$

$$f(s) = \frac{4}{b} \tan^{-1} \left\{ e^{\pm \frac{v}{b} s} \right\}$$

$$\gamma^{-1} = \sqrt{1 - v^2}$$

Solitary wave - propagates without dissipation.

No superposition



Eq. (1) clearly has an infinite number of constant solutions

$$\phi = \frac{2\pi n}{b} \quad \text{for integral } n$$

$$\mathcal{L} = \frac{1}{2} \left(\frac{\partial \phi}{\partial t} \right)^2 - \frac{1}{2} \left(\frac{\partial \phi}{\partial x} \right)^2 - V(\phi)$$

$$V(\phi) = \frac{1}{b^2} [1 - \cos b\phi]$$

For this choice $V = 0$
 $\phi = \frac{2\pi n}{b}$

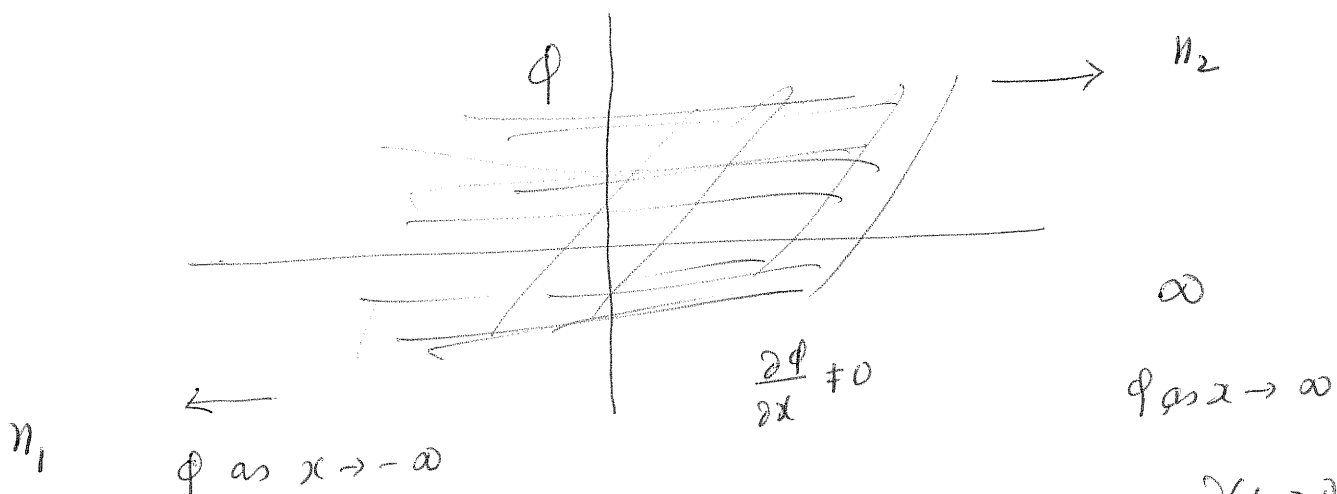
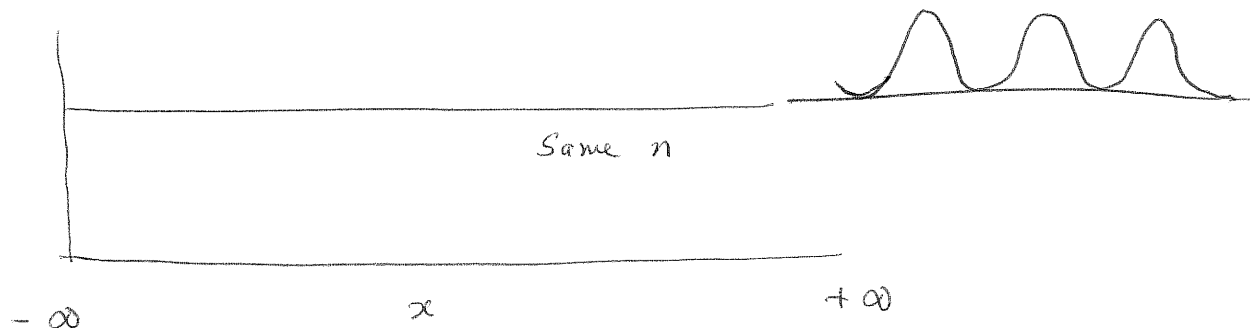
$$H = \frac{1}{2} \left(\frac{\partial \phi}{\partial t} \right)^2 + \frac{1}{2} \left(\frac{\partial \phi}{\partial x} \right)^2 + V(\phi)$$

Analysis

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$$V(\phi) = \frac{\phi^2}{2} - \frac{b^2}{4!} \phi^4$$

unit mass
 $b^2 = \lambda$



\mathcal{H} is positive definite $\rightarrow \partial/\partial t = 0$

$$\frac{\partial^2 \phi}{\partial x^2} = \frac{\partial V}{\partial \phi}$$

$$\frac{1}{2} \left(\frac{\partial \phi}{\partial x} \right)^2 = V(\phi)$$

$$E = \int \mathcal{H} dx = \int \left[\frac{1}{2} \left(\frac{\partial \phi}{\partial x} \right)^2 + V(\phi) \right] dx$$

$\left. \begin{array}{l} \eta_1 = 0 \quad \phi \rightarrow 2\pi/b \\ \eta_2 = 1 \quad \phi \rightarrow 0 \end{array} \right\}$

$$= \int_{2\pi/b}^0 2V(\phi) dx = \int_0^{2\pi/b} [2V(\phi)]^{1/2} d\phi$$

$$\equiv 8/b^2$$

Energy true definite

$$E = \frac{8m^3}{\lambda} \rightarrow \text{Energy goes as } \lambda^{-1}$$



Space is one dimensional \rightarrow
 The origin of the solution and why
 Such solutions are stable?

\Rightarrow Energy argument

\Rightarrow Mathematically argument

boundary conditions \rightarrow In one-d
 space

$(n, n' \neq n)$ solution cannot be continuously
 deformed to n, n . ($0, 0 =$ ground state)

Kink, is a ~~physical~~ topological object.

\Rightarrow The stability further implies
 - yes - a conservation law - something
 which must remain 'constant' and that
 simply prevents $(1, 0) \rightarrow (0, 0)$.

Topological Charge .

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The conserved charge in this case is an integer N .

$$J^\mu = \frac{b}{2\pi} \epsilon^{\mu\nu} (\partial_\nu \varphi)$$

$$\partial_\mu J^\mu = 0$$

$$Q = \int_{-\infty}^{+\infty} J^0 dx = \frac{b}{2\pi} \int \frac{\partial \varphi}{dx} dx$$
$$= \frac{b}{2\pi} [\varphi(\infty) - \varphi(-\infty)] = N$$

But J^μ was not derived from a symmetry — It is not a Noether current.

Quantum Numbers : Labels of a state